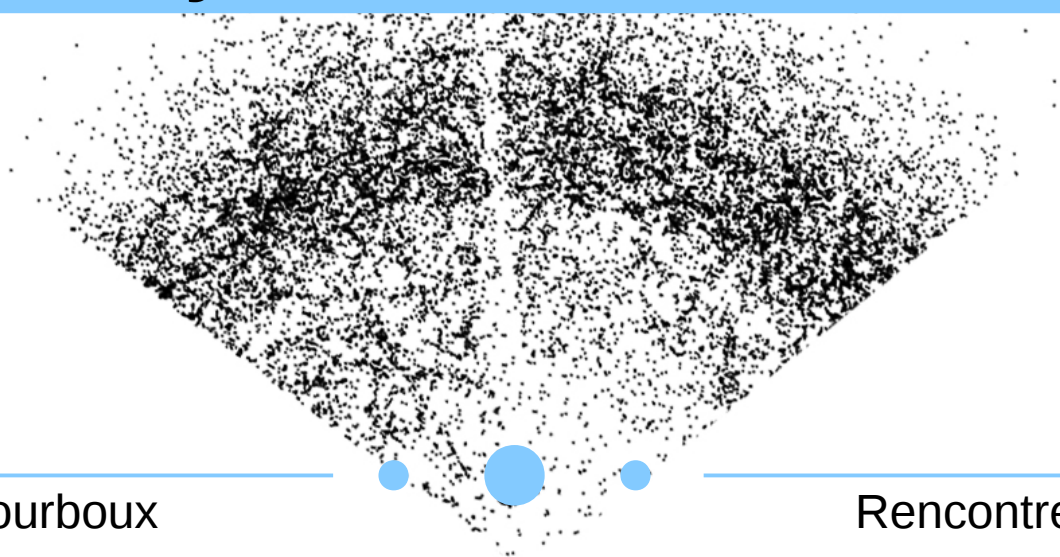


# Baryonic Acoustic Oscillation Correlations at $z=2.3$ with SDSS-III Lyman- $\alpha$ Forests



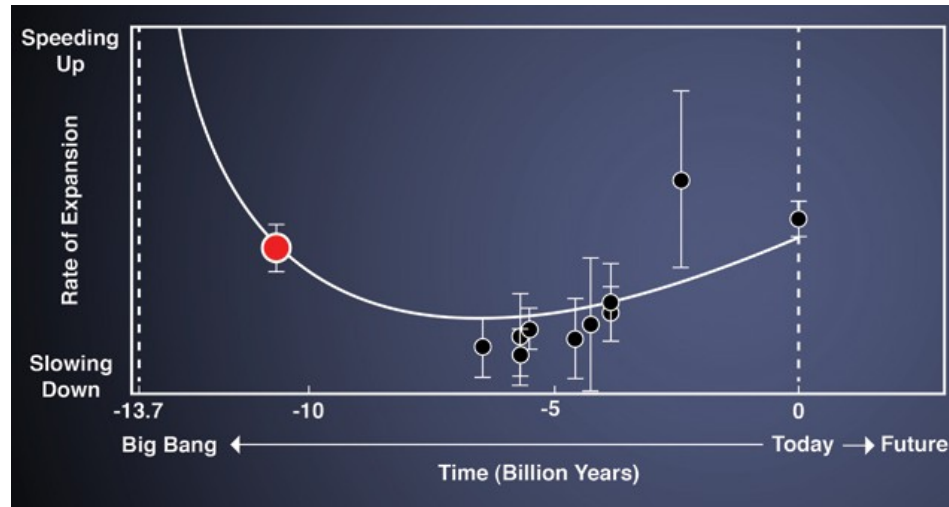
- introduction to BAO and BOSS
- Lyman- $\alpha$  BAO

# BAO

(Baryonic Acoustic Oscillations)

# BAO

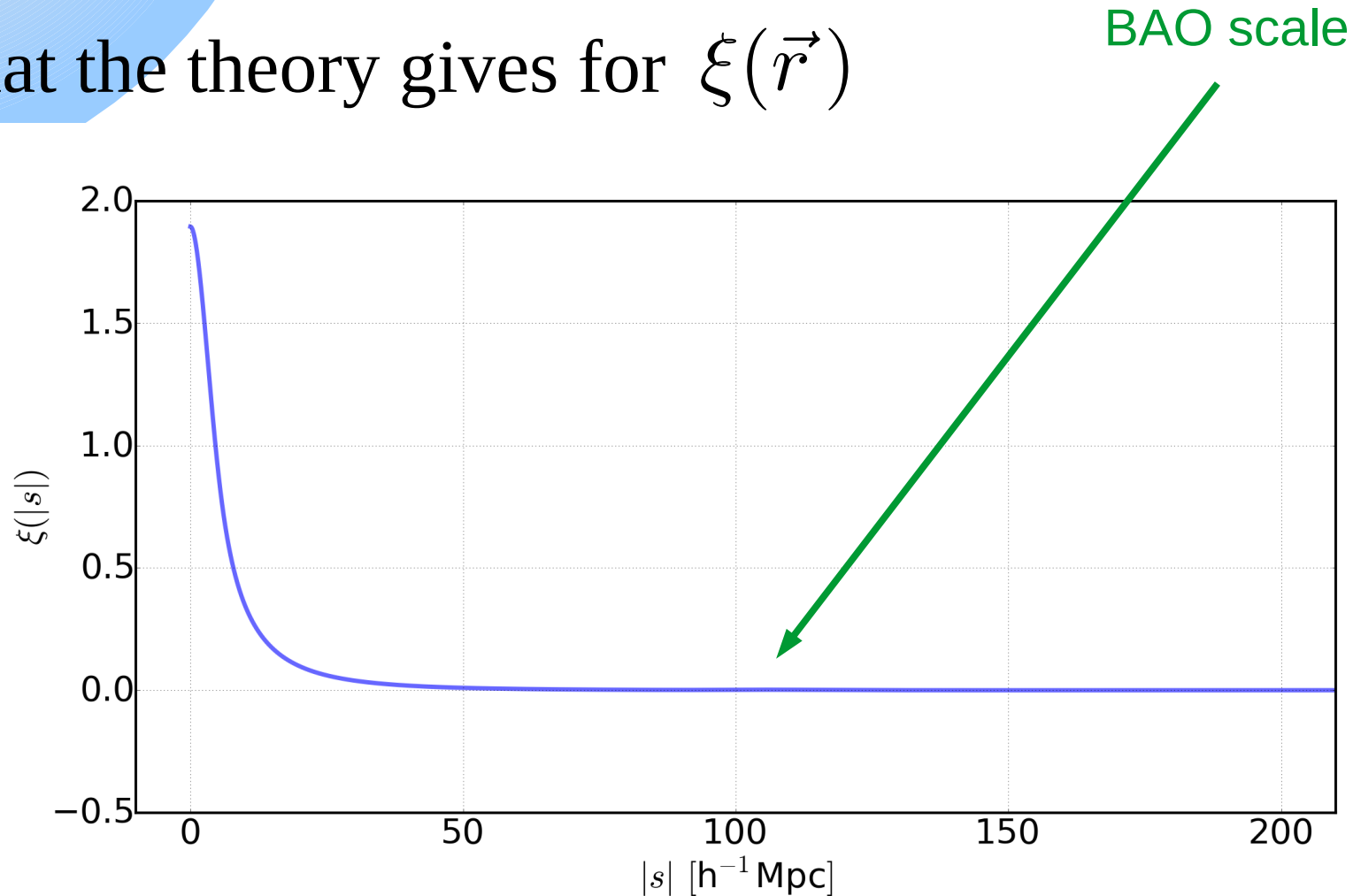
To understand what is Dark Energy we need to measure distances at different redshifts



BAOs measure the expansion rate and angular diameter distance at a given redshift

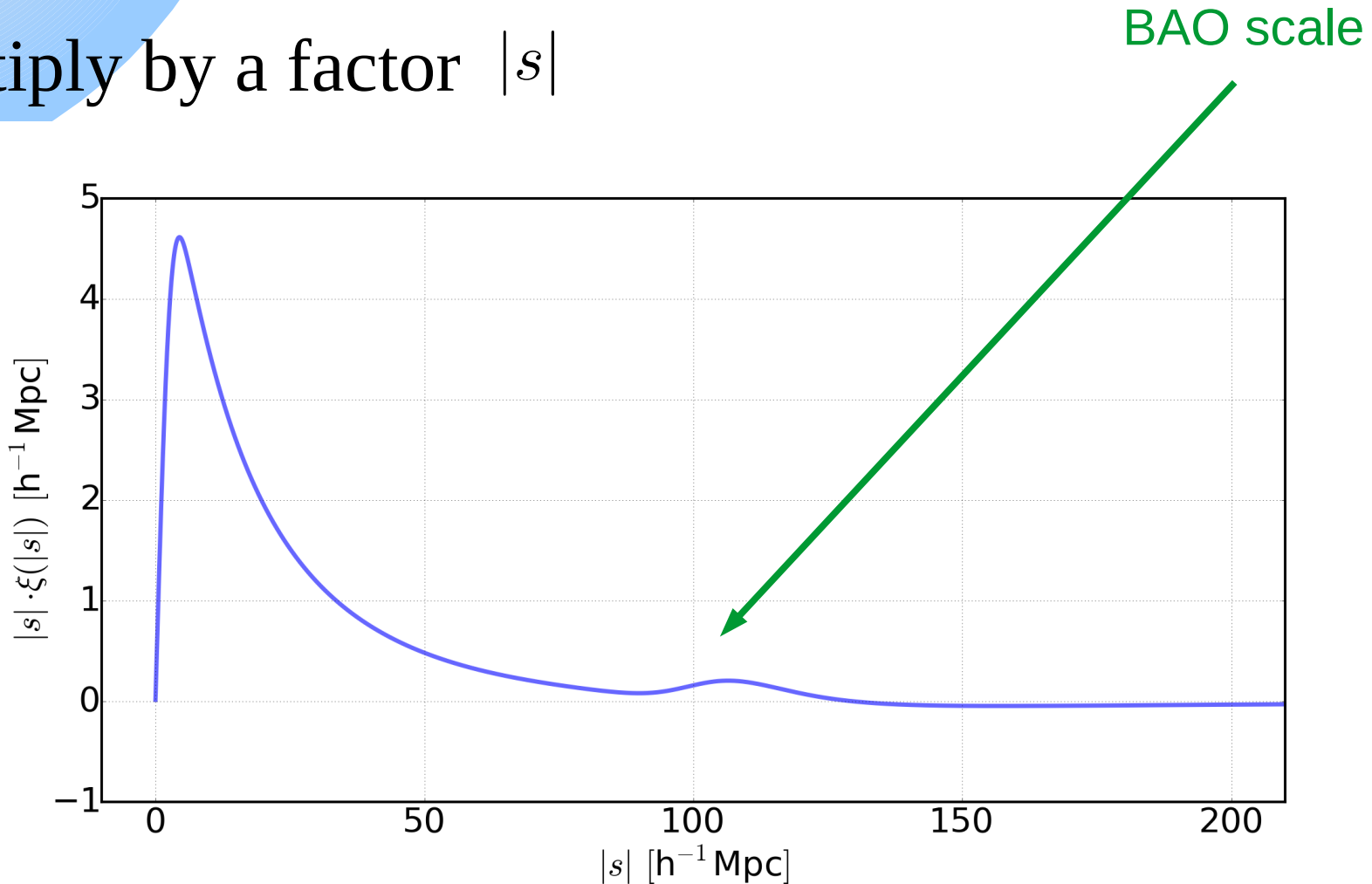
# Measuring the BAO scale

What the theory gives for  $\xi(\vec{r})$



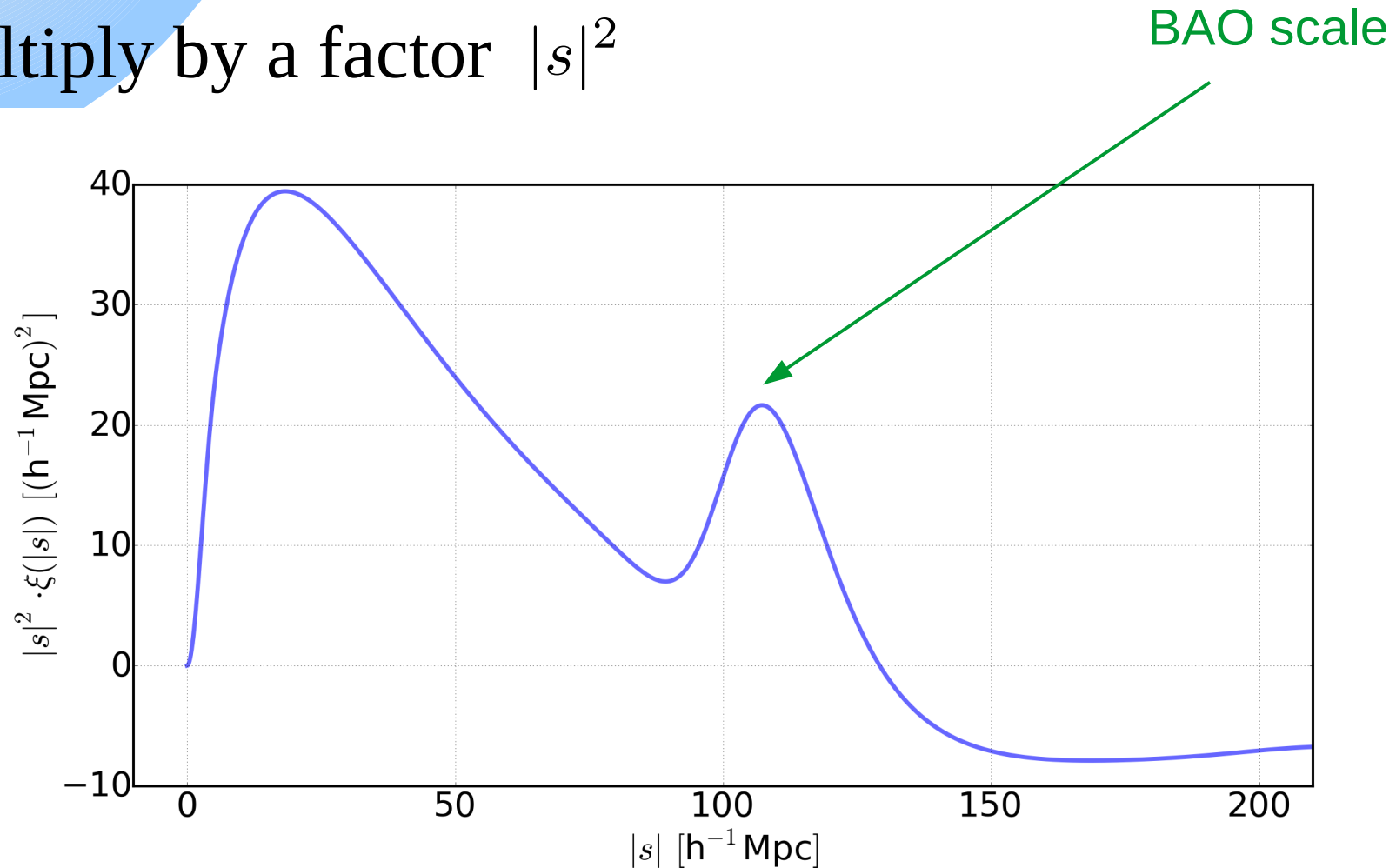
# Measuring the BAO scale

Multiply by a factor  $|s|$



# Measuring the BAO scale

Multiply by a factor  $|s|^2$

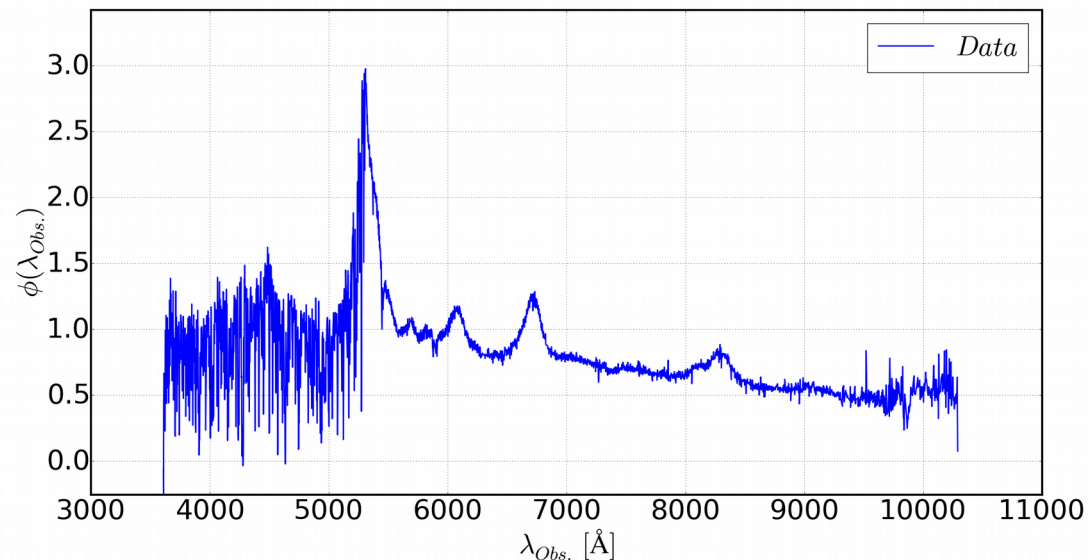


# Measuring the BAO scale



# Quasar

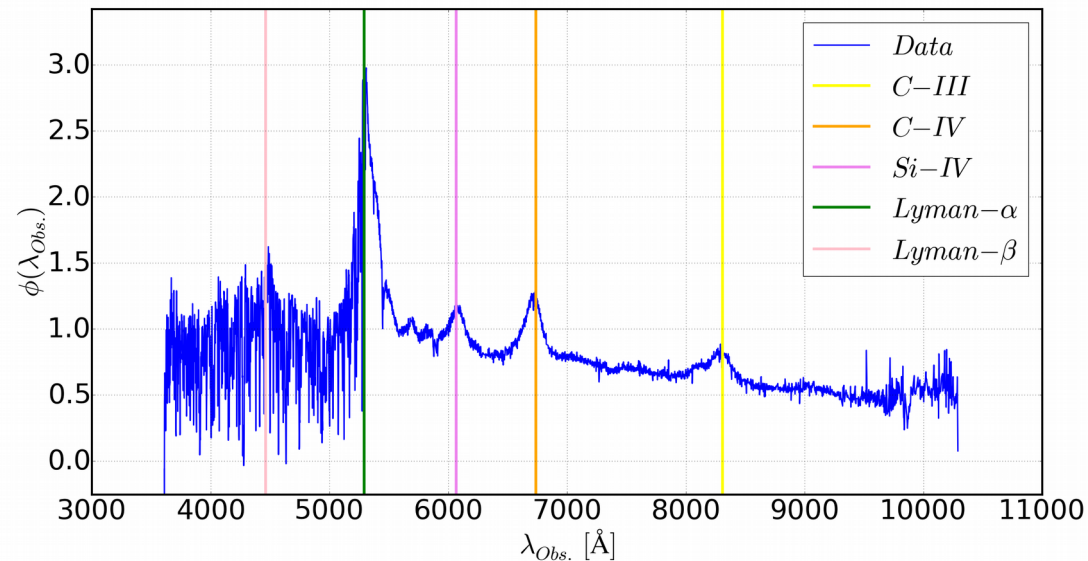
Quasar flux originates from the surrounding of a super-massive black hole



Spectrum of a BOSS Quasar at redshift  $z = 3.35$ , the Universe was only 2 billion years old

# Quasar

Get redshift from emission lines

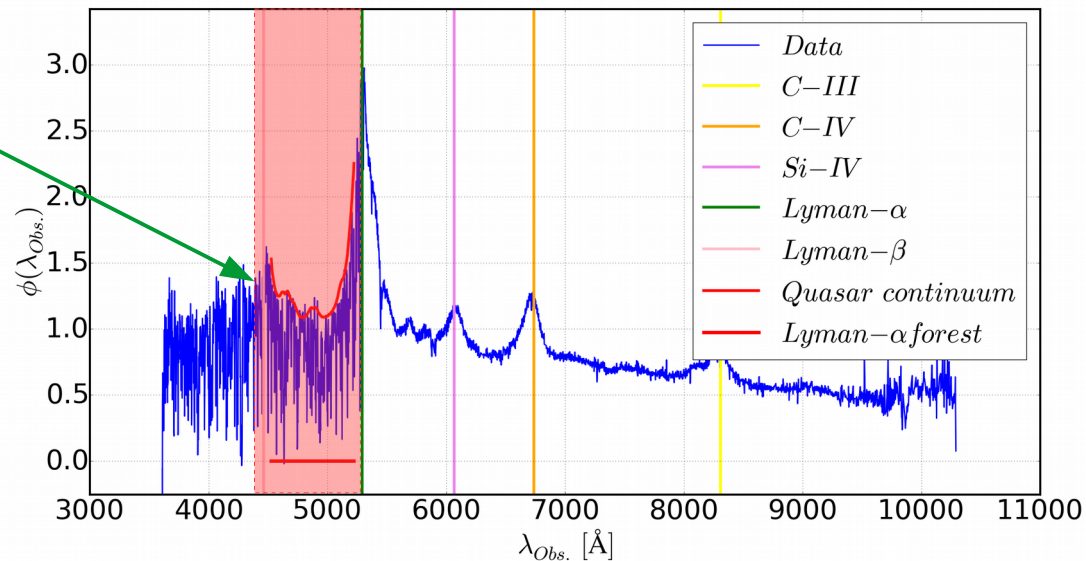


A Quasar is a boolean matter density tracer

# Lyman- $\alpha$ forest

Absorption lines from Hydrogen continuum in the Intergalactic Medium (IGM)

Lyman- $\alpha$  forest



A Lyman- $\alpha$  pixel gives a continuous matter density tracer

# Lyman- $\alpha$ forest

Matter density fluctuation

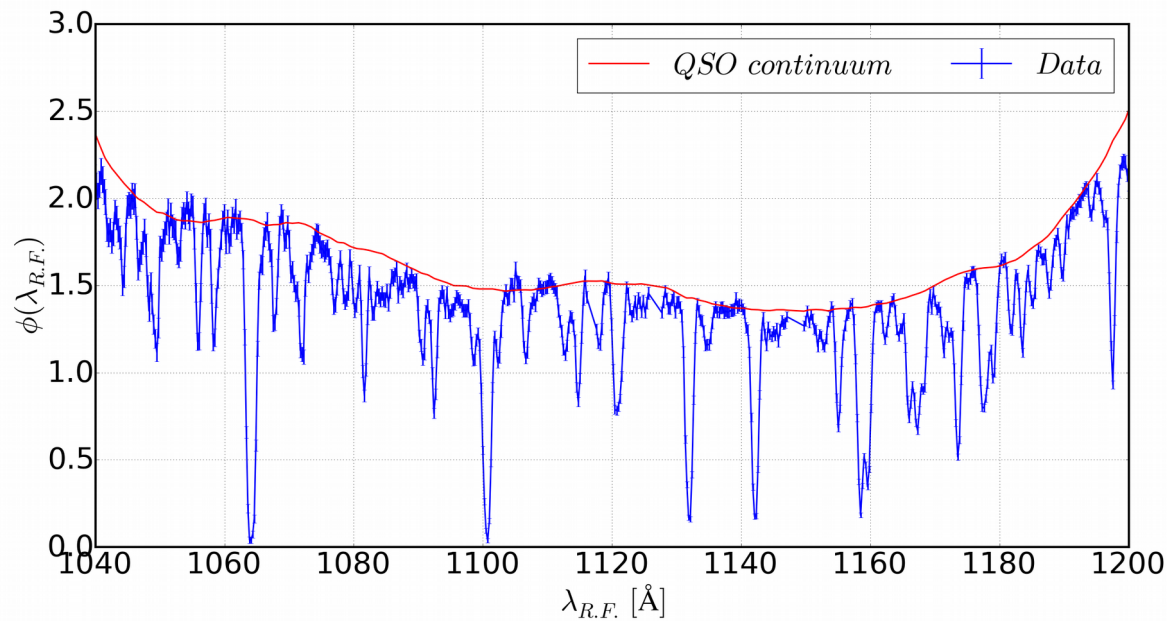
$$\delta_{\alpha,i} = \frac{f_{\alpha,i}}{C_{\alpha}(\lambda_{R.F.}) \cdot \bar{F}(\lambda_{Obs.})} - 1$$

Normalized Data  
+ correction

QSO physics

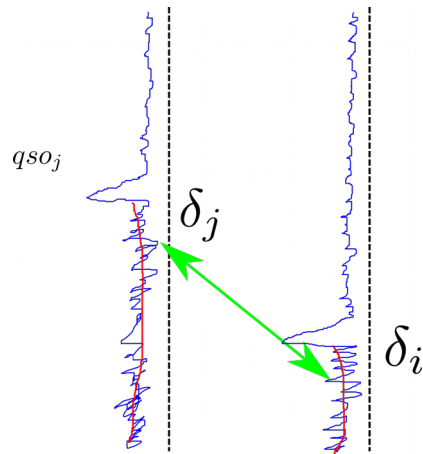
$$z_{\alpha,i} = \frac{\lambda_{Obs.,i}}{1215.67} - 1$$

Sky + cosmology  
physics

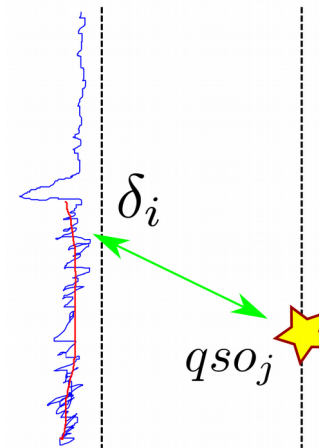


# Two matter density tracers

- We have two matter density tracers:
  - Quasars
  - Lyman- $\alpha$  forest
- We can estimate two different correlation functions



Lyman- $\alpha$  auto-correlation

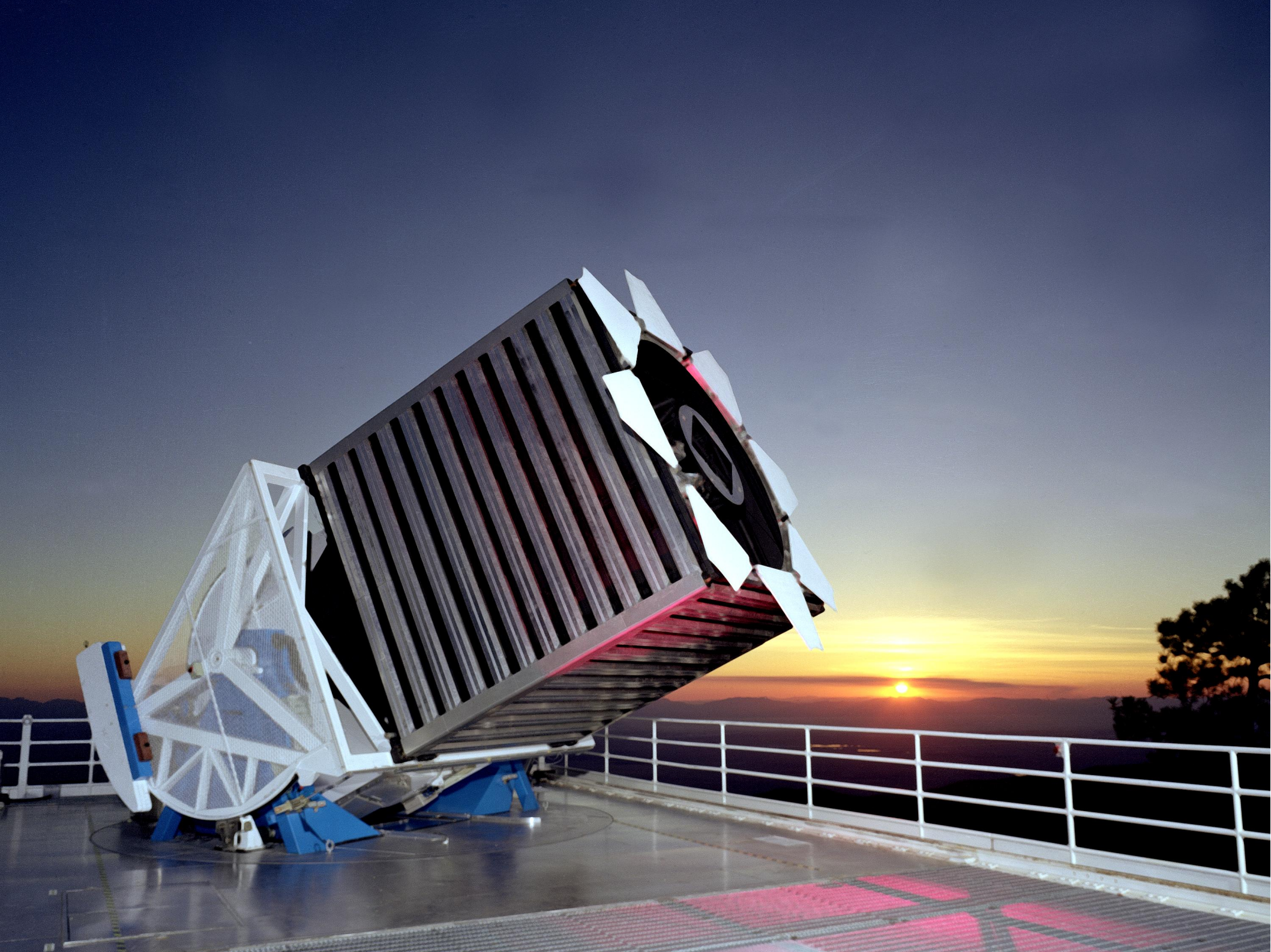


Lyman- $\alpha$  – QSO cross-correlation



# BOSS

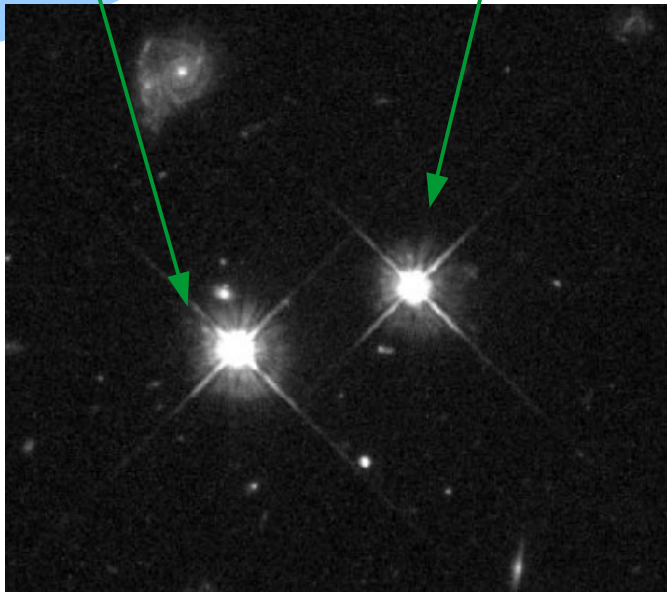
(Baryon Oscillation Spectroscopic Survey)



# From Photometry to Spectroscopy

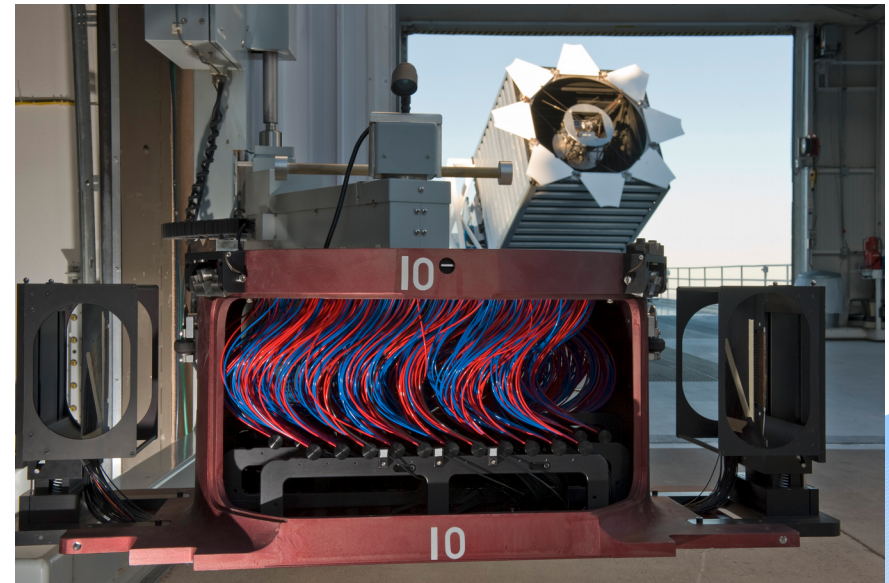
Quasar

Star



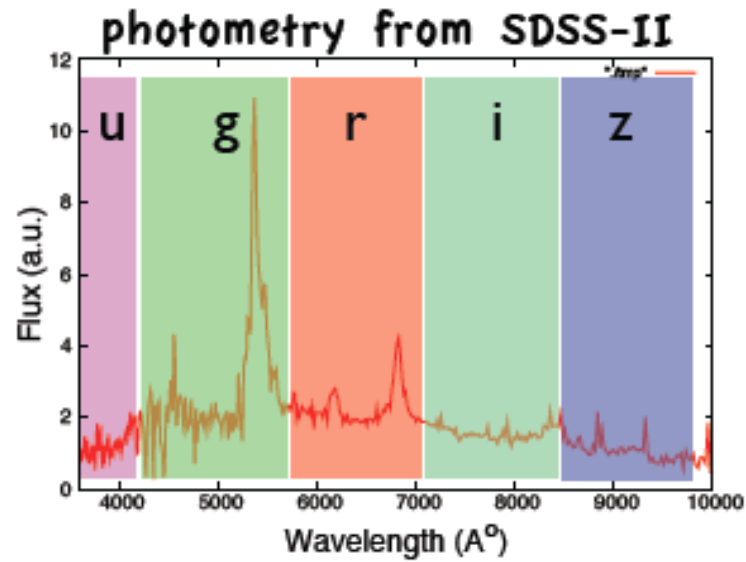
- Sloan Digital Sky Survey (SDSS)
- 2.5-m wide-angle optical telescope
- Spectroscopy with the Baryonic Oscillation Spectroscopic Survey (BOSS)
- 1000 fibers
- Run: 2009-2014

List of targets from photometry sent to the BOSS spectrograph.

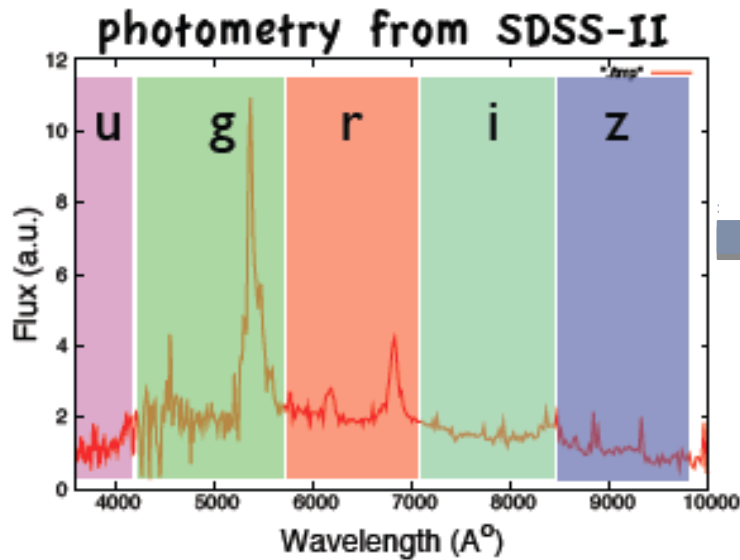




# From Photometry to Spectroscopy



# From Photometry to Spectroscopy

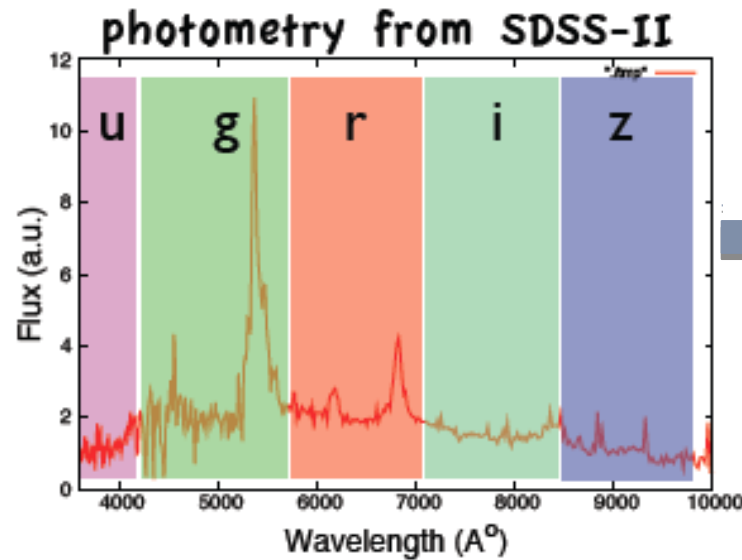


QSO selection

## List of targets

SDSS J112253.51+005329.8  
SDSSp J120441.73-002149.6  
SDSSp J130348.94+002010.4  
SDSSp J141205.78-010152.6  
SDSSp J141315.36+000032.1  
....

# From Photometry to Spectroscopy

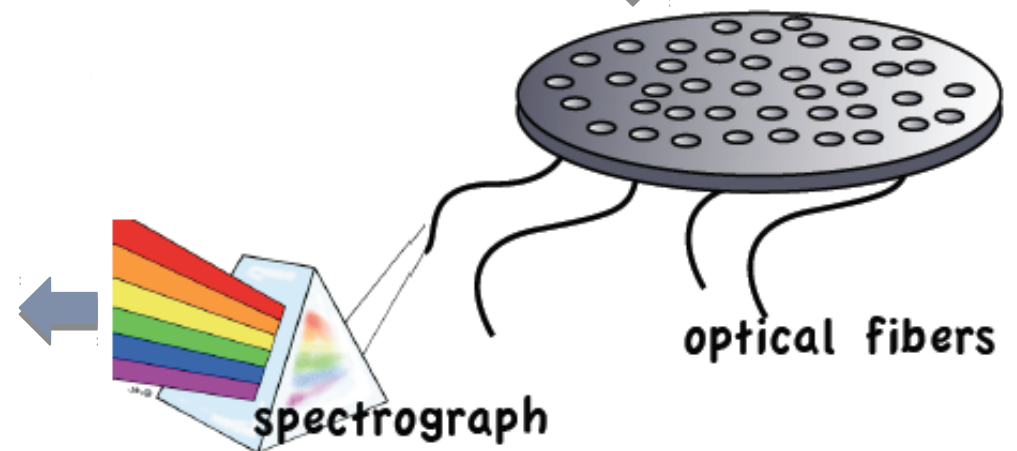


QSO selection

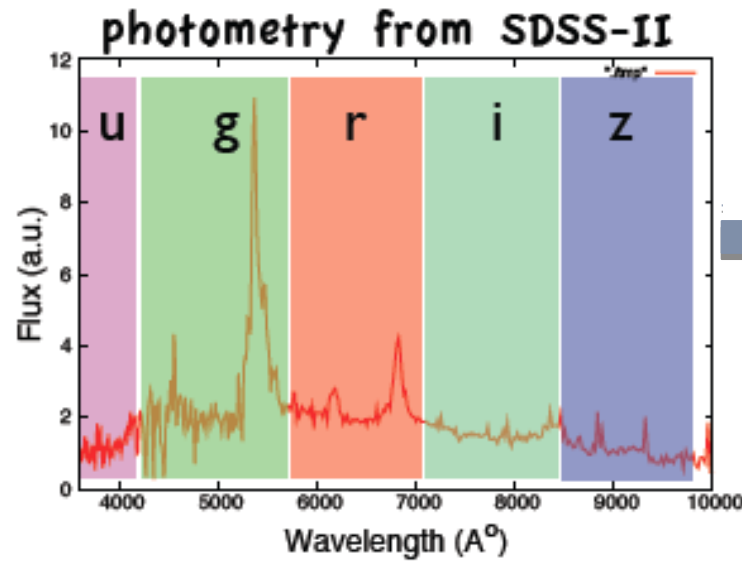
List of targets

```
SDSS J112253.51+005329.8
SDSSp J120441.73-002149.6
SDSSp J130348.94+002010.4
SDSSp J141205.78-010152.6
SDSSp J141315.36+000032.1
....
```

1000 fibers  
per plate



# From Photometry to Spectroscopy

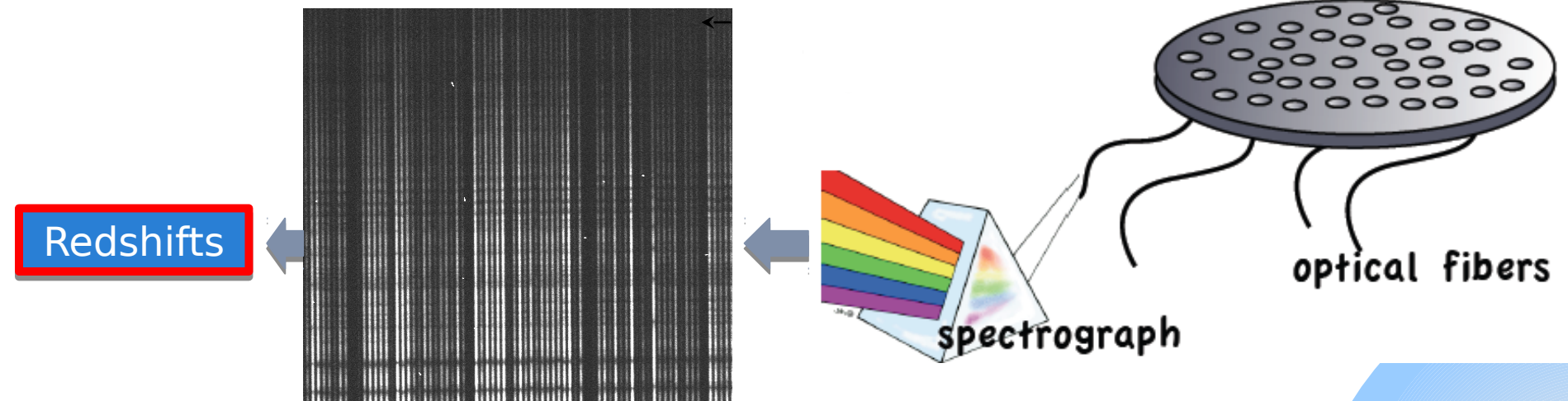


QSO selection

List of targets

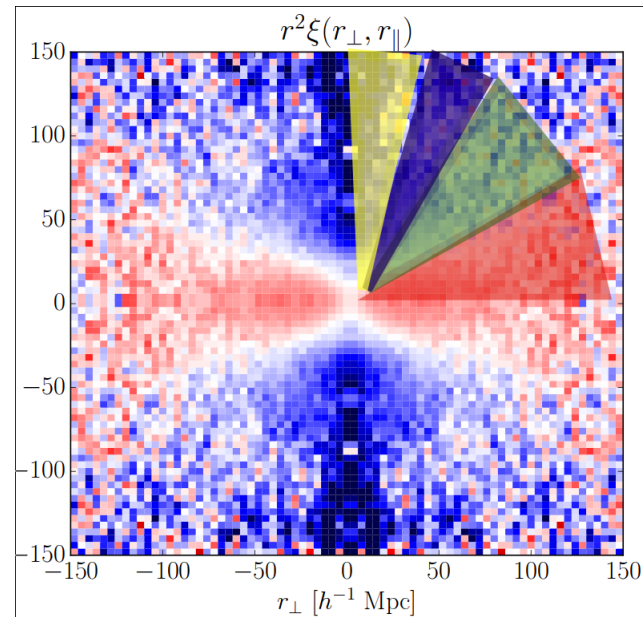
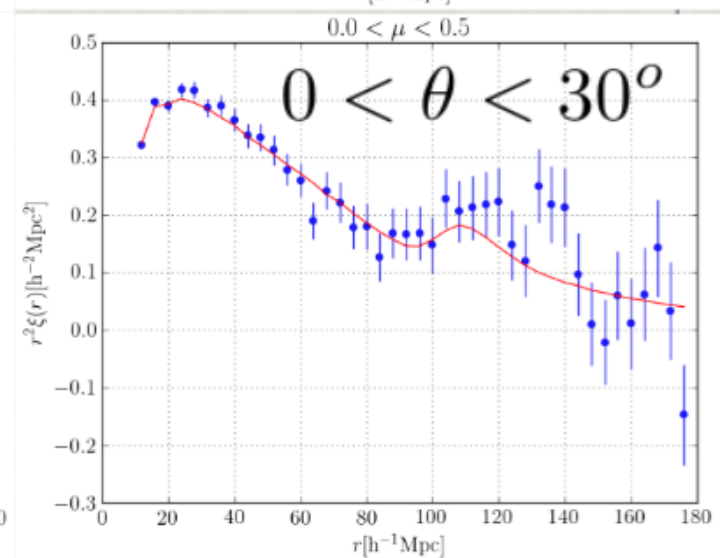
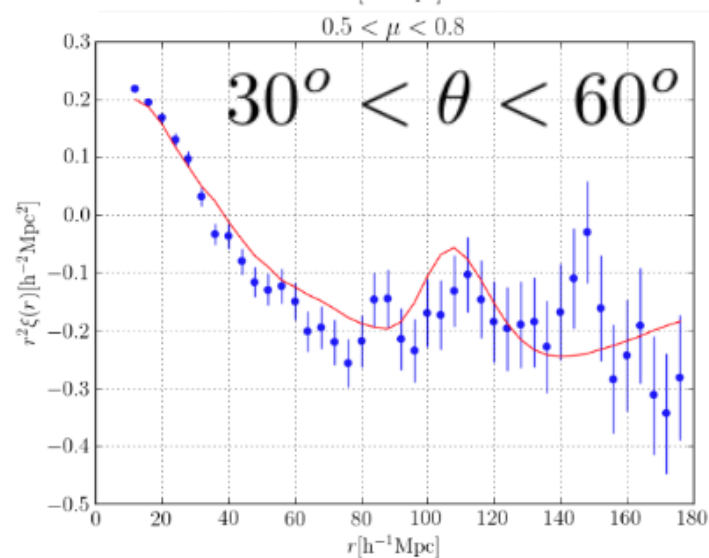
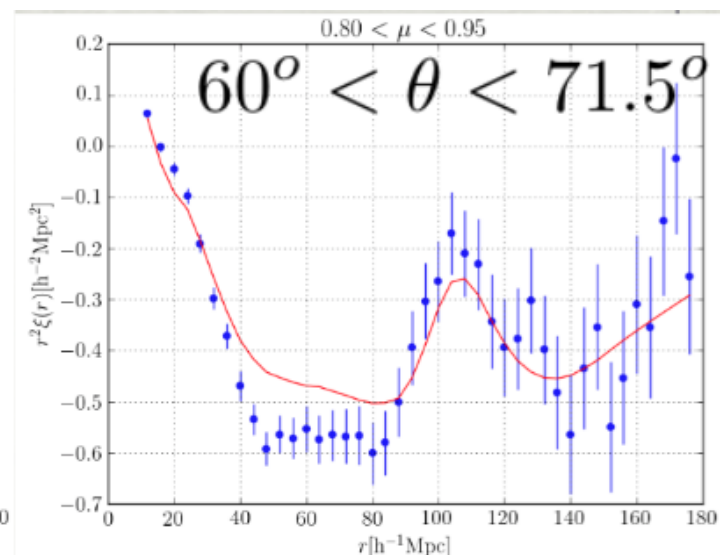
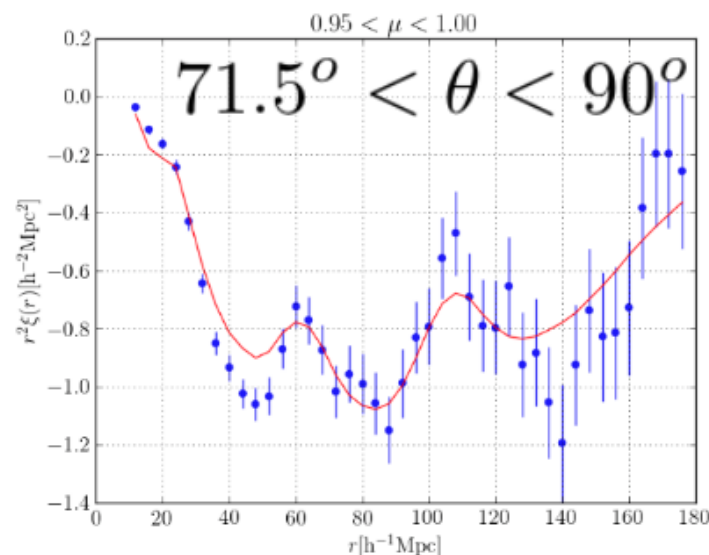
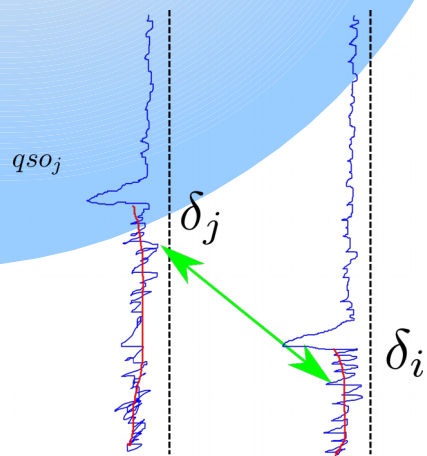
```
SDSS J112253.51+005329.8
SDSSp J120441.73-002149.6
SDSSp J130348.94+002010.4
SDSSp J141205.78-010152.6
SDSSp J141315.36+000032.1
....
```

1000 fibers  
per plate



# $\text{Ly}\alpha$ BAO Results

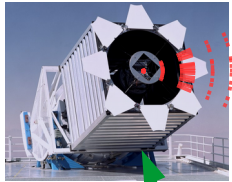
# Auto-correlation



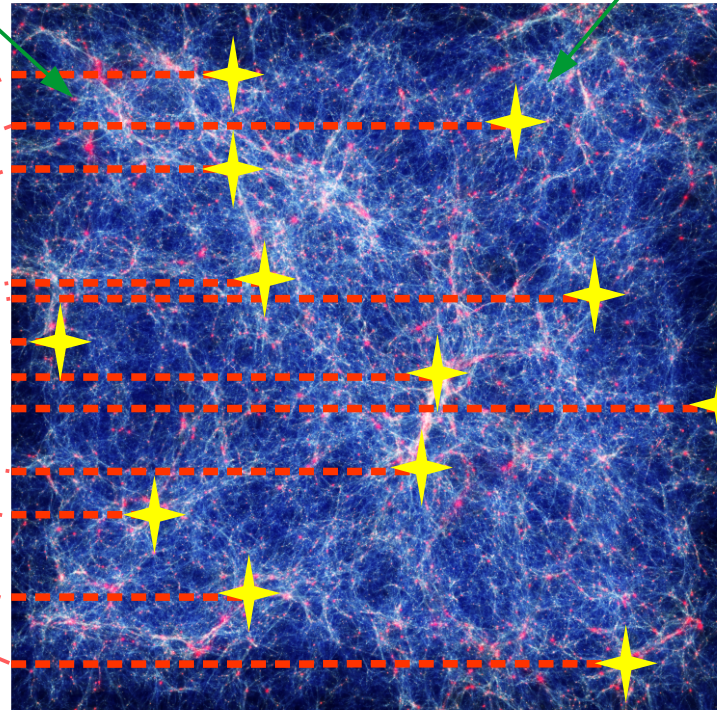
# Gaussian Random Field Simulations

Ly $\alpha$  forest along the line-of-sight

QSO set on big Over-density

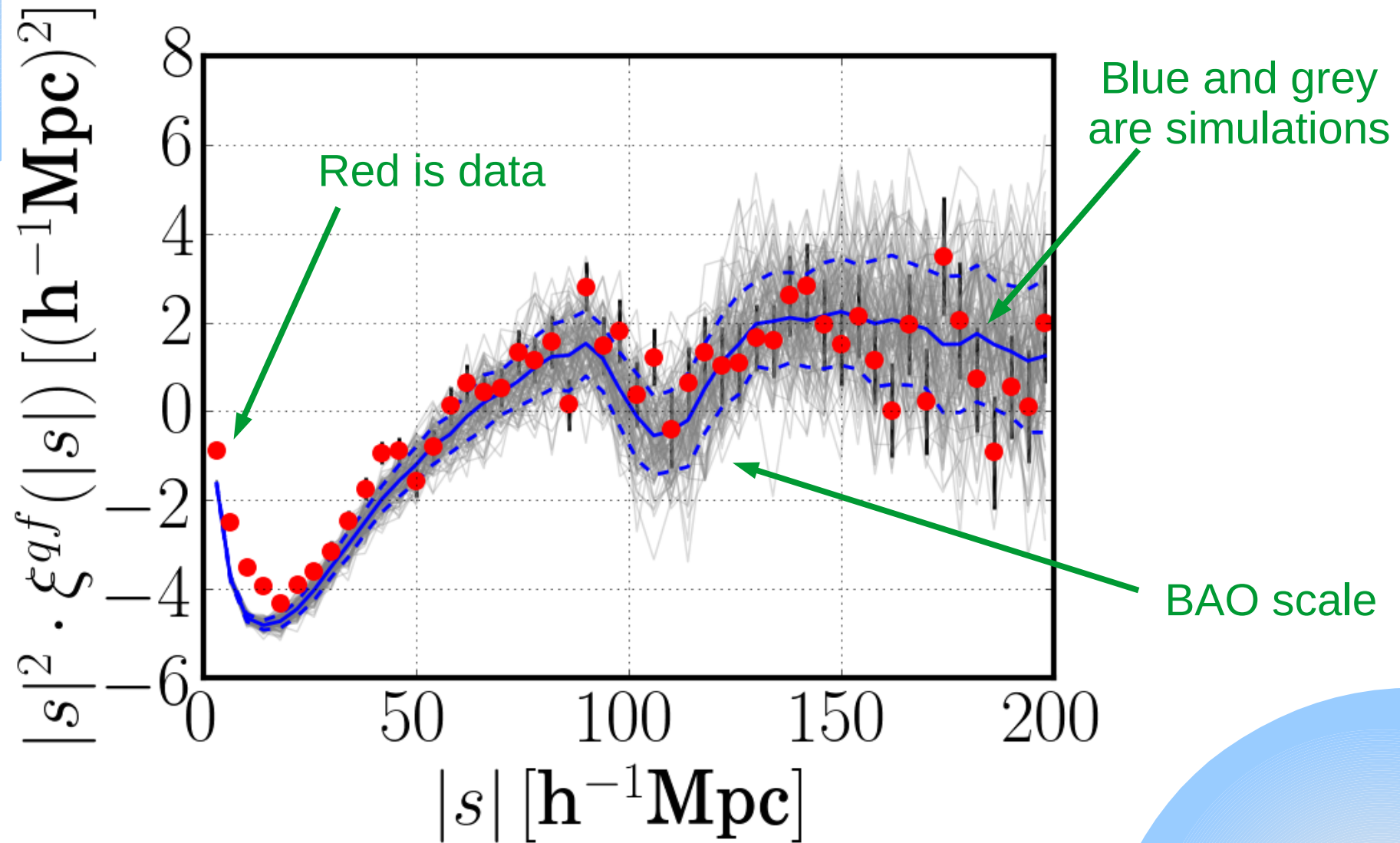


Apply telescope properties



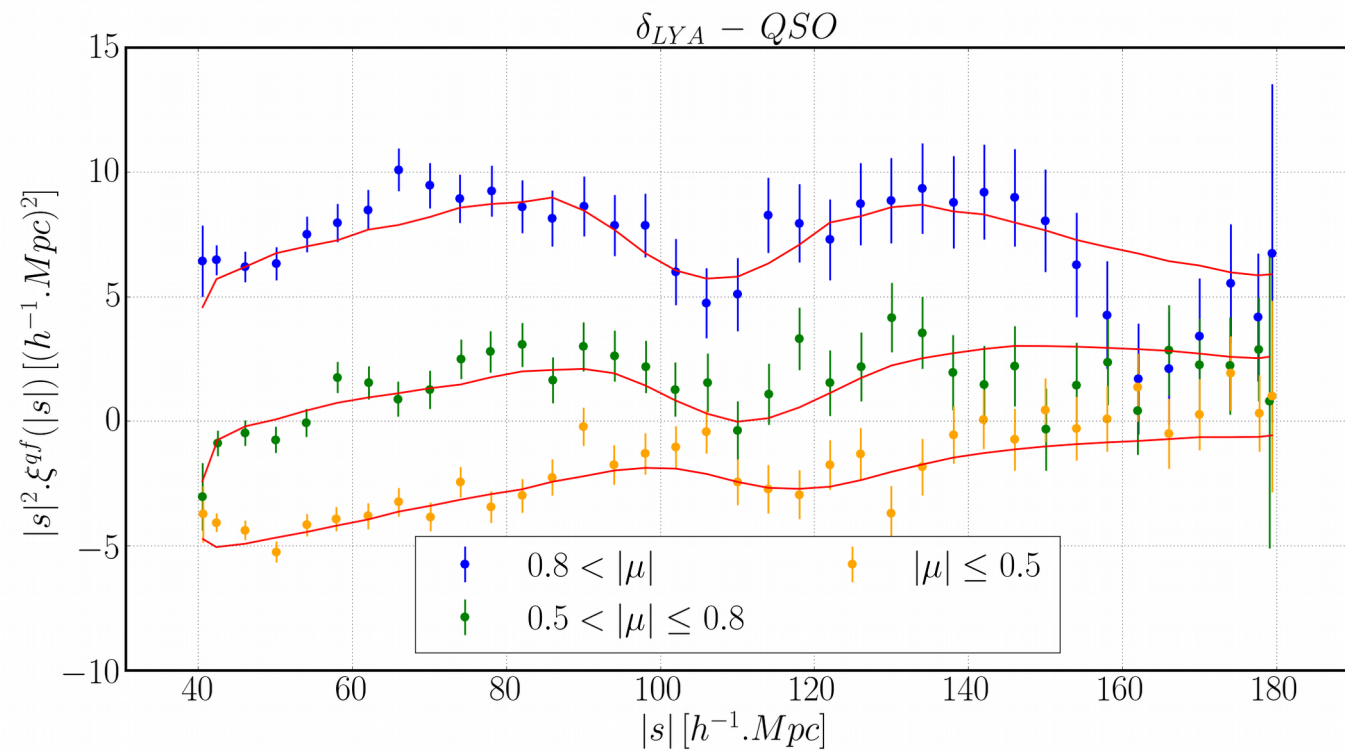
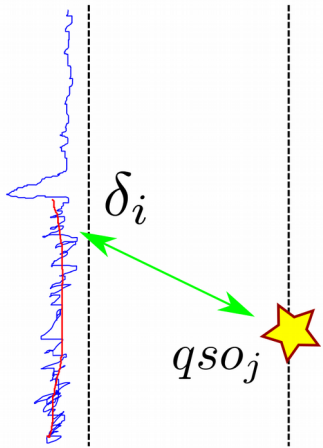
IGM image provided by Julien Baur

# Gaussian Random Field Simulations





# Cross-correlation



# BAO Results

## Radial BAO

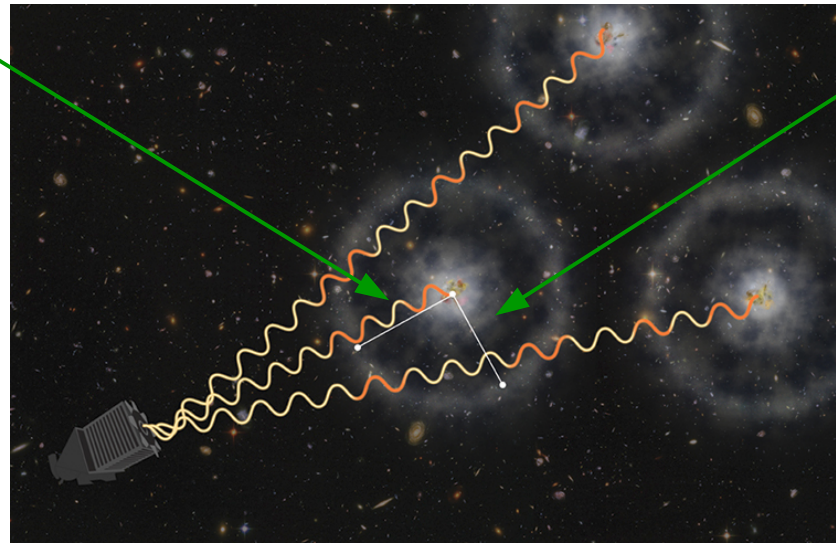
$$\alpha_{\parallel} = \frac{D_H(\bar{z})/r_d}{[D_H(\bar{z})/r_d]_{fid}}$$

Hubble scale factor

## Transverse BAO

$$\alpha_{\perp} = \frac{D_A(\bar{z})/r_d}{[D_A(\bar{z})/r_d]_{fid}}$$

Angular size



# BAO Results

- Auto-correlation DR12: (JB++ in prep.)

$$\chi^2/dof = 1630.43/(1589 - 10)$$

$$\alpha_{\parallel} = 1.028 \pm 0.028$$

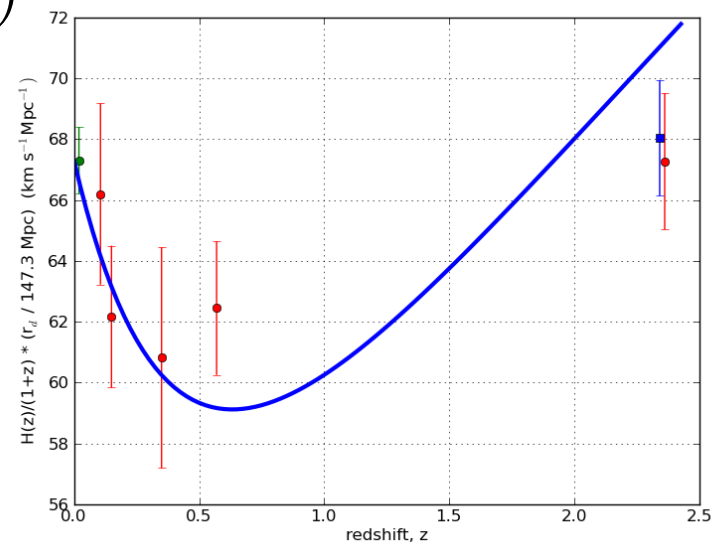
$$\alpha_{\perp} = 0.983 \pm 0.048$$

- Cross-correlation DR12: (HdMdB++ in prep.)

$$\chi^2/dof = 3115.24/(3030 - 14)$$

$$\alpha_{\parallel} = 1.045 \pm 0.032$$

$$\alpha_{\perp} = 0.913 \pm 0.038$$



# BAO Results

- Auto-correlation DR12: (JB++ in prep.)

$$\chi^2/dof = 1630.43/(1589 - 10)$$

$$\alpha_{\parallel} = 1.028 \pm 0.028$$

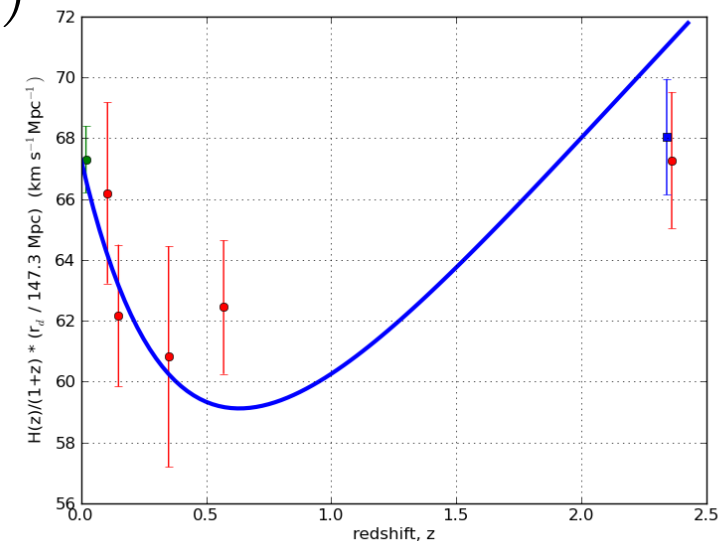
$$\alpha_{\perp} = 0.983 \pm 0.048$$

- Cross-correlation DR12: (HdMdB++ in prep.)

$$\chi^2/dof = 3115.24/(3030 - 14)$$

$$\alpha_{\parallel} = 1.045 \pm 0.032$$

$$\alpha_{\perp} = 0.913 \pm 0.038$$



# Improvements

- Better model for contamination by carbon, silicon ...
- Better model for the distortion caused by continuum fitting
- Better data-reduction and calibration
- Better understanding of spurious correlations induced by instrument/data-reduction
- First simulations of the cross-correlation.

# Conclusion

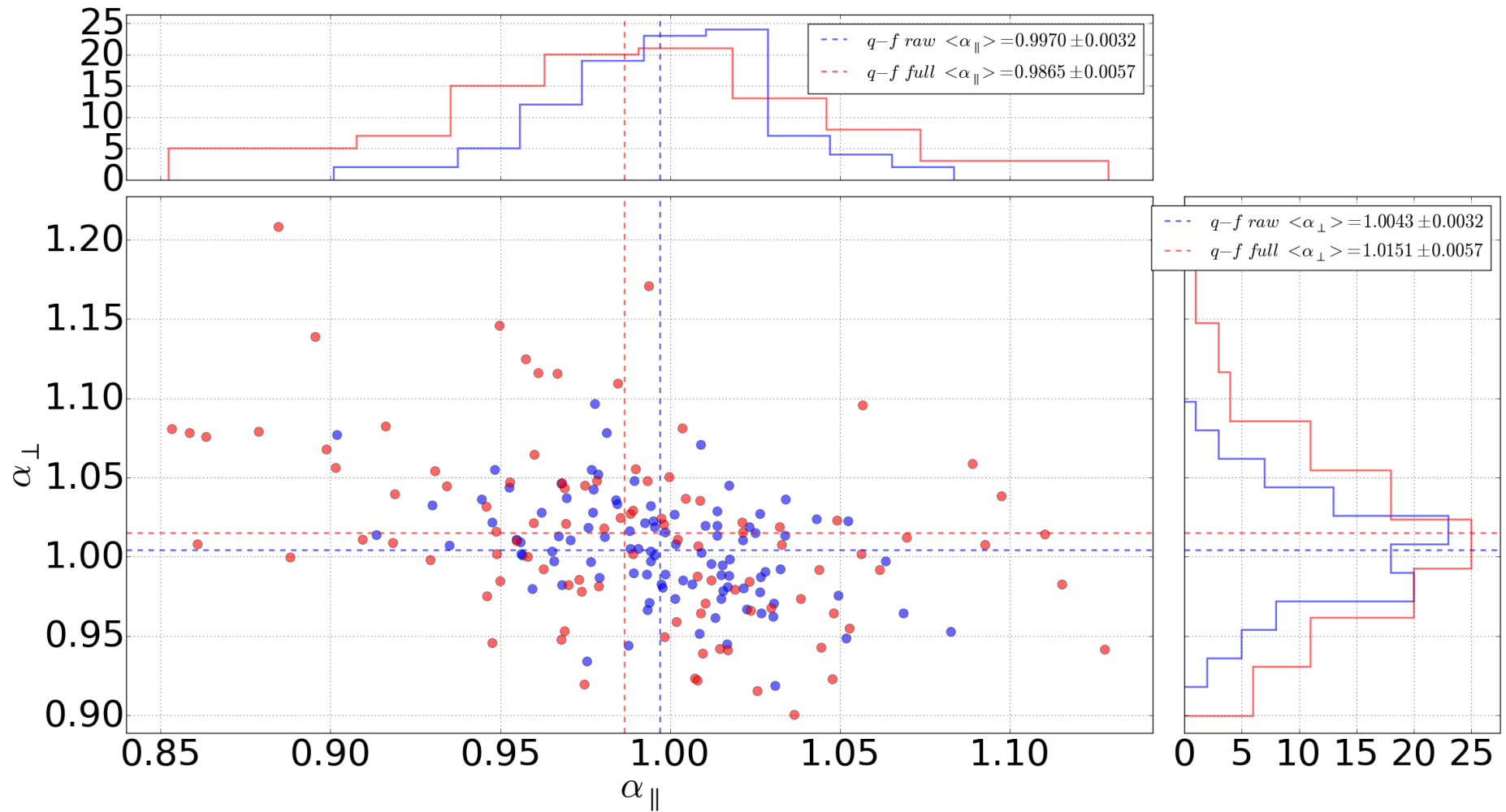
- A  $\sim 2.8$  % measurement of the expansion rate at  $z = 2.3$ .
- Robust measurement against systematics
- First simulations of the cross-correlation.



# BACKUP slides



# Gaussian Random Field Simulations



# Metal templates

