



*InVisibles 16 workshop  
Padova, Sep. 12, 2016*

# *Status and prospect of neutrino oscillation experiments*

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# Outline

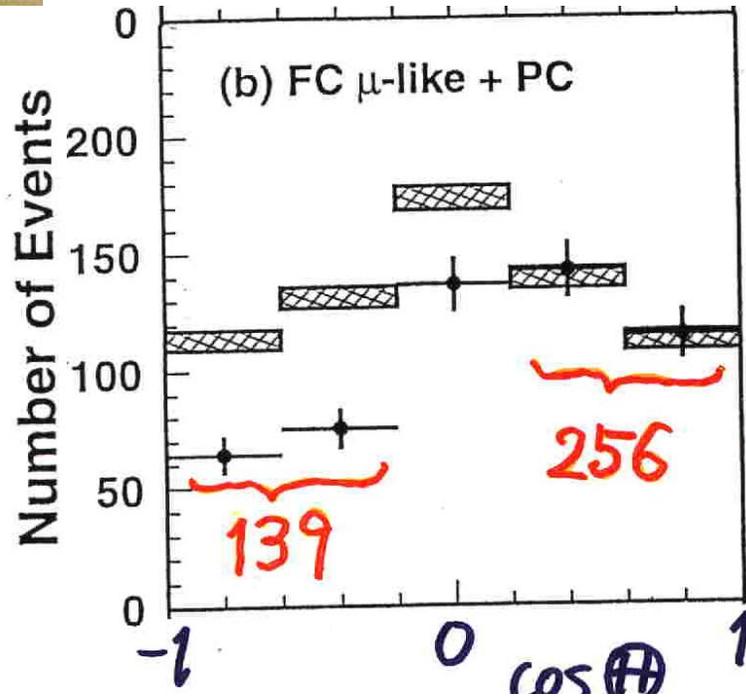
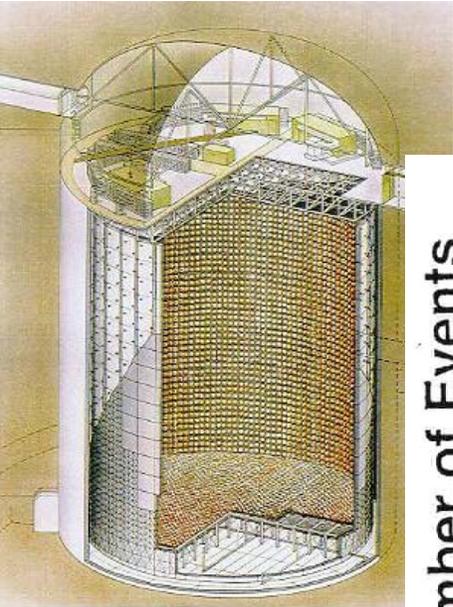
- *Introduction and present status of neutrino oscillations*
- *Future oscillation studies*
- *Other physics with future experiments*
- *Summary*

# *Introduction and present status of neutrino oscillations*

# Introduction: Discovery of neutrino oscillations

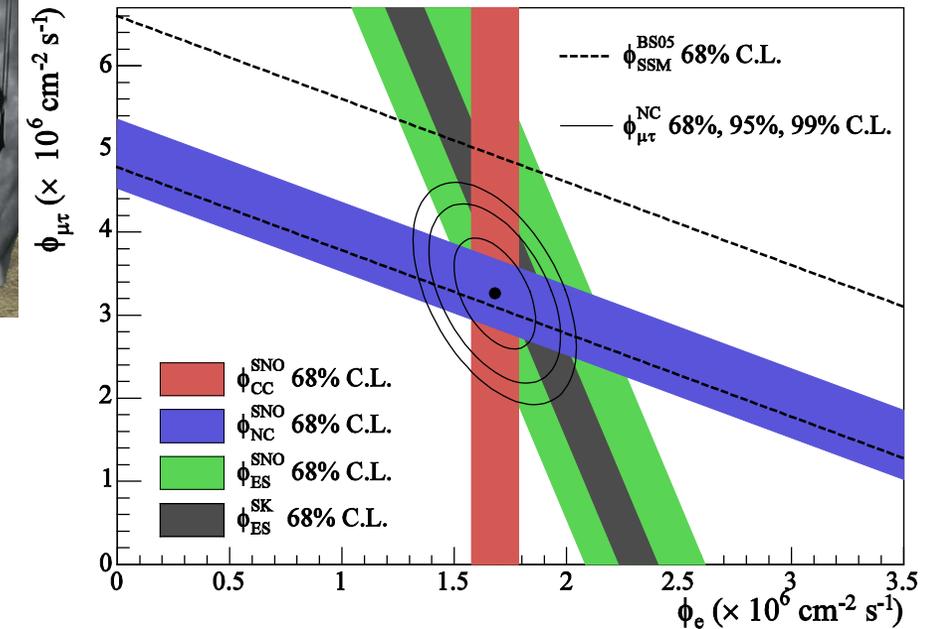
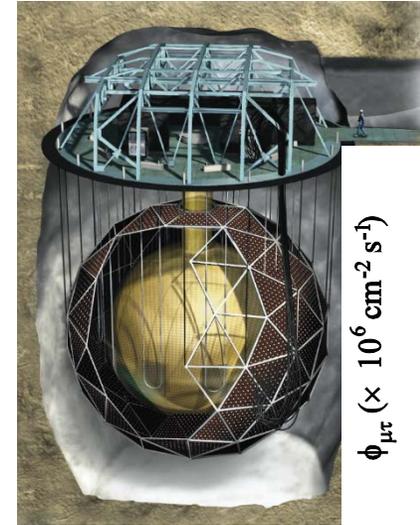
Super-K, PRL 81 (1998) 1562

## Atmospheric neutrino oscillations



## Solar neutrino oscillations

SNO PRL 89 (2002) 011301  
SNO PRC 72, 055502 (2005)



*The very small neutrino masses opened a window to study the physics beyond the Standard Model of particle physics!*

# Status of neutrino oscillation studies

$\nu_\mu \rightarrow \nu_\tau$  oscillations ( $\Delta m_{23}, \theta_{23}$ )

Atmospheric: Super-K, Soudan-2,  
MACRO IceCube/Deepcore, ...

LBL: K2K, MINOS, OPERA, T2K, NOvA, ...

$\nu_e \rightarrow (\nu_\mu + \nu_\tau)$  oscillations ( $\Delta m_{12}, \theta_{12}$ )

Solar: SNO, Super-K, Borexino, ...

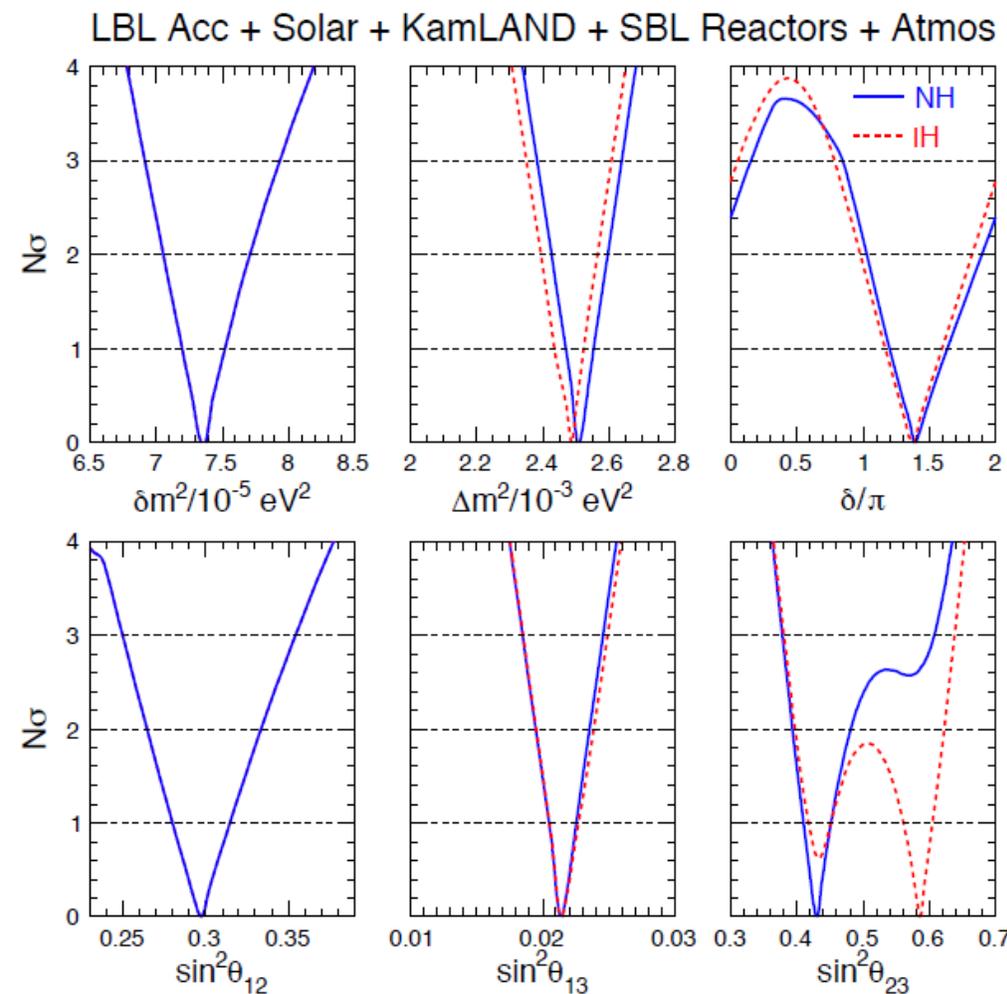
Reactor: KamLAND

$\theta_{13}$  experiments

LBL: MINOS, T2K, NOvA, ...

Reactor: Daya Bay, Reno, Double Chooz

**Status** (A. Marrone, @Neutrino 2016)



**Basic structure for 3 flavor oscillations has been understood!**

# Some tensions ( $\nu_\mu$ disappearance)

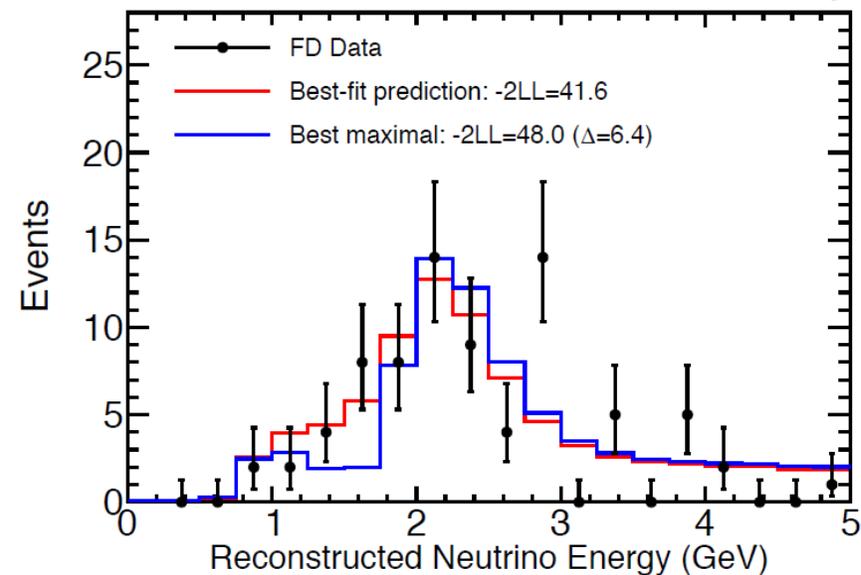
## NOvA @Neutrino2016

## T2K

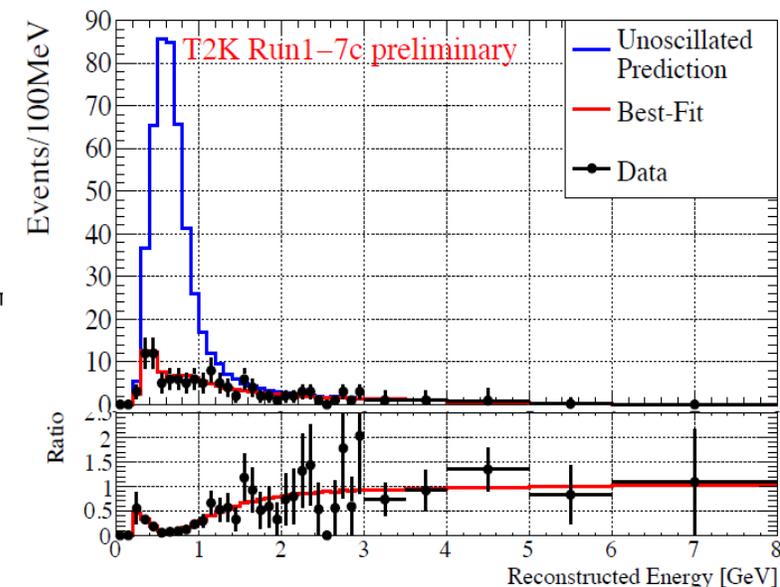
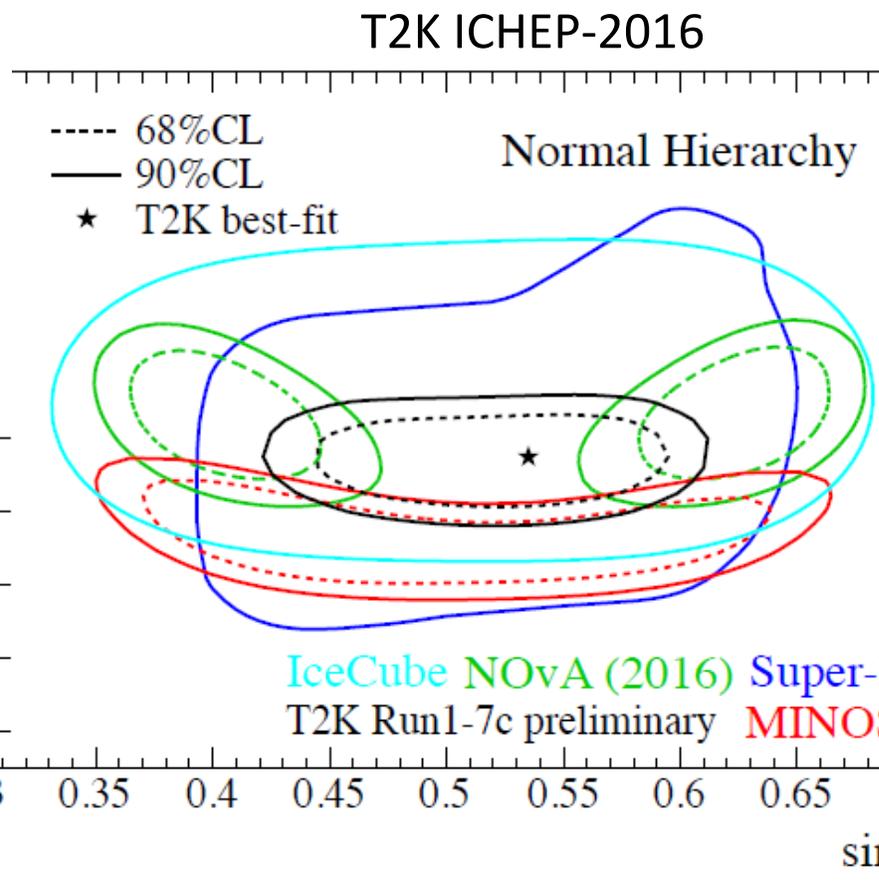
T2K, PRL 112 (2014) 181801

T2K@ICHEP 2016

NOvA Preliminary



Maximal mixing excluded at  $2.5\sigma$ .

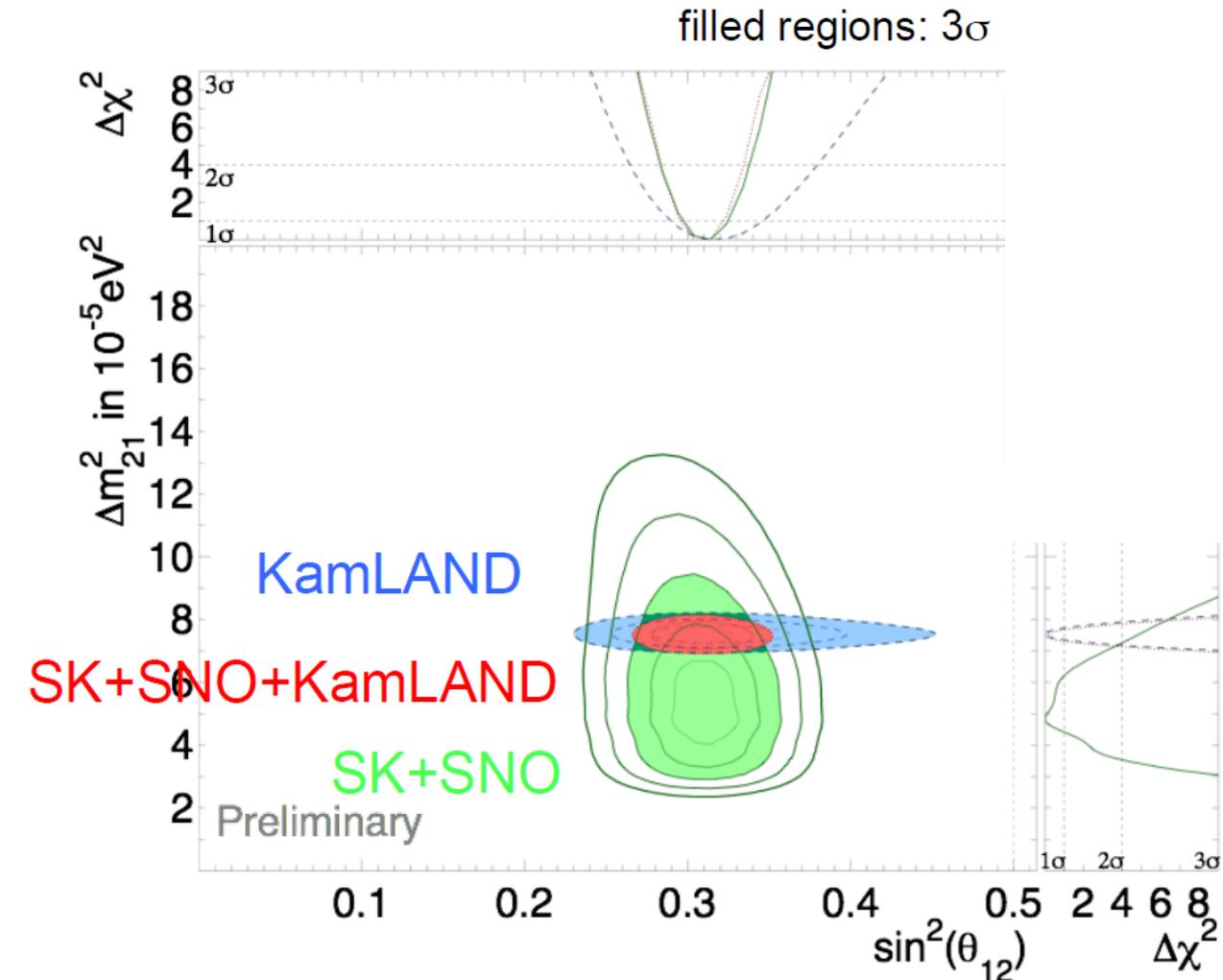


Consistent with maximal mixing...

# Some tensions (solar neutrino oscillations and day/night effect)

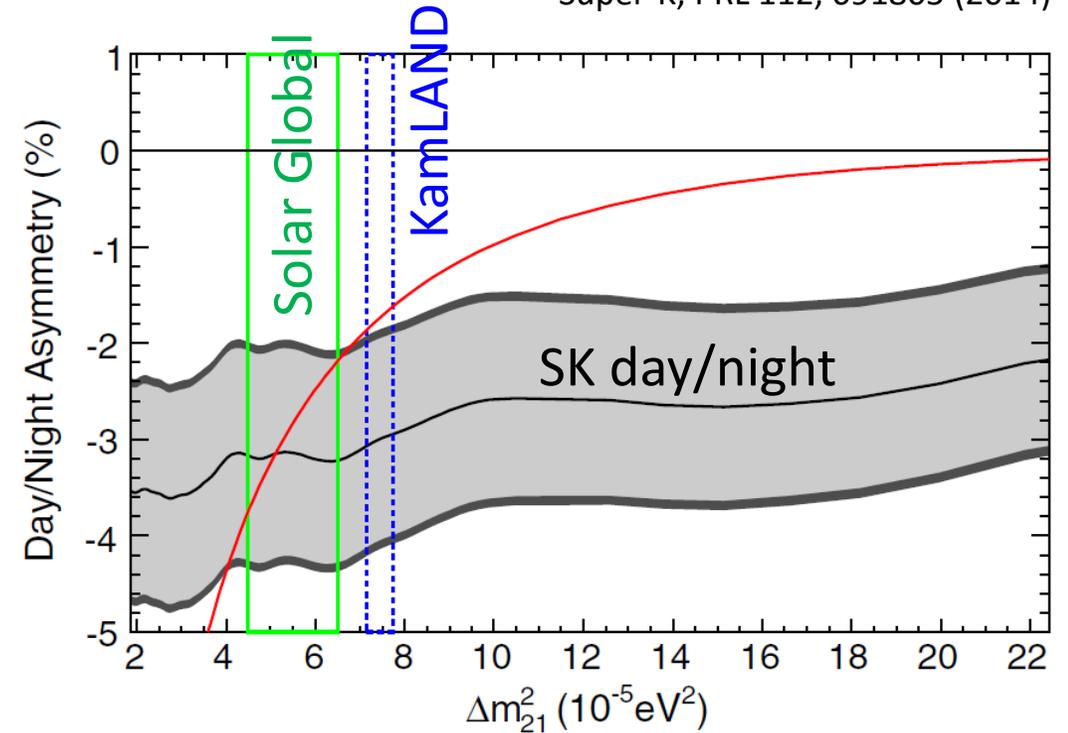
## Status of the 12-parameter measurements

S. Moriyama (Super-K), Neutrino 2016



## SK day/night and $\Delta m_{12}^2$

Super-K, PRL 112, 091805 (2014)

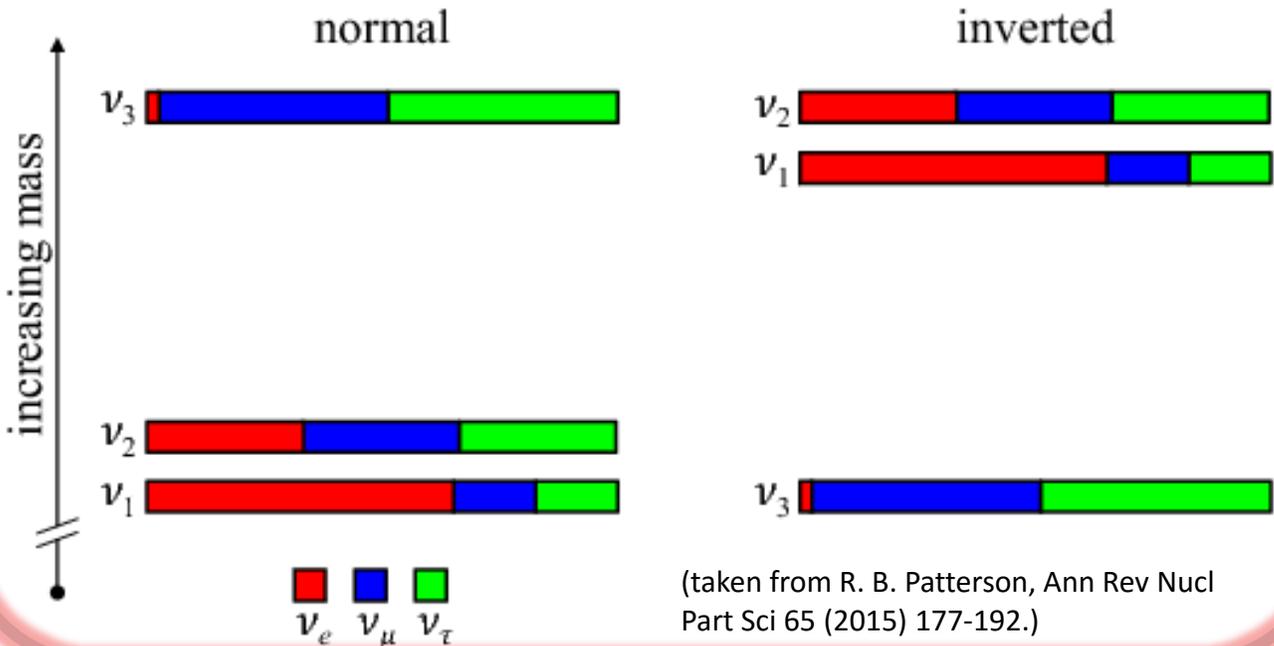


- ✓ The data might indicate that there is something interesting going on in solar neutrinos.... (will come back later.)

*Future oscillation studies*

# Agenda for the future neutrino measurements

## Neutrino mass hierarchy?



## Absolute neutrino mass?

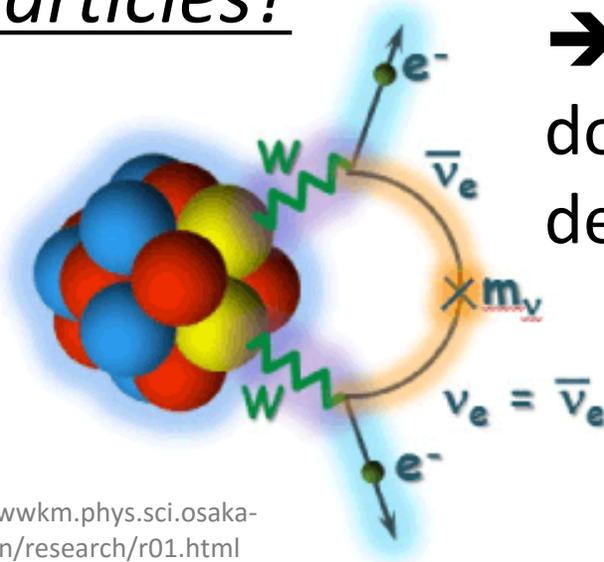
## Beyond the 3 flavor framework? (Sterile neutrinos?)

## CP violation?

$$P(\nu_\alpha \rightarrow \nu_\beta) \neq P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta) ?$$

Baryon asymmetry of the Universe?

## Are neutrinos Majorana particles?

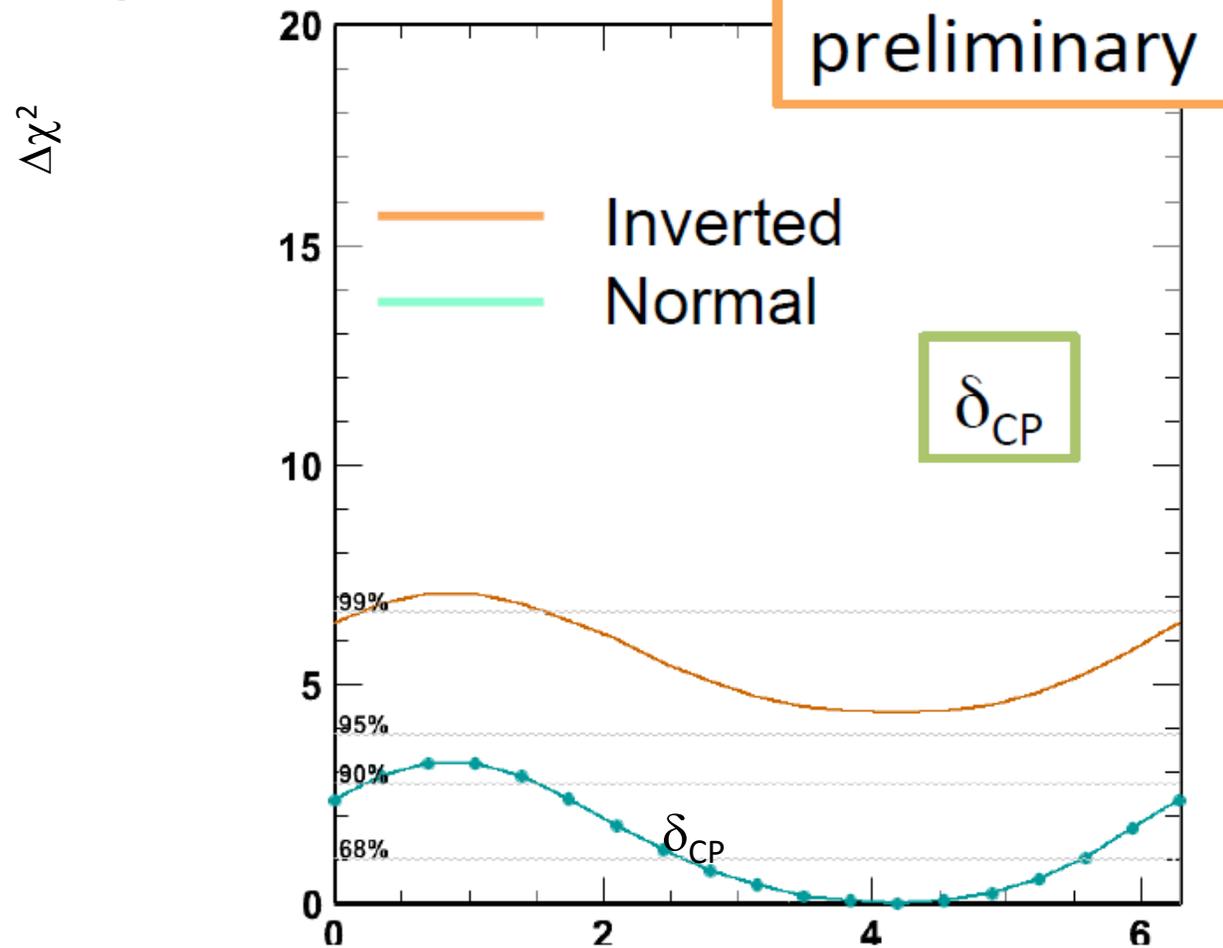


→ Neutrinoless double beta decay

<http://wwwkm.phys.sci.osaka-u.ac.jp/en/research/r01.html>

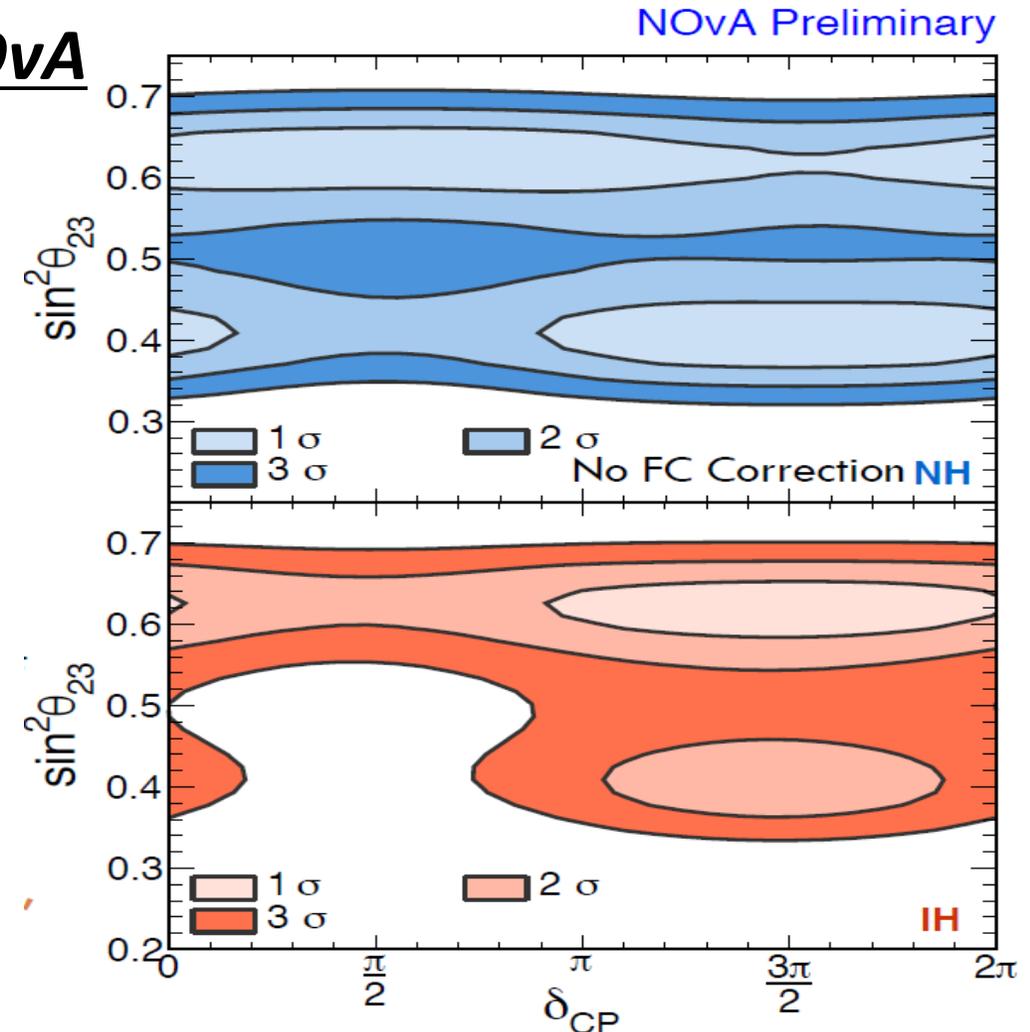
# Mass hierarchy and CP violation measurements @Neutrino 2016

## Super-K atmospheric



Best fit at  $\delta_{CP} = \sim 1.33\pi$  (NH)  
 $\Delta\chi^2$  (IH-NH) = 4.3 (expt'd = 3.1)

## NOvA



Best fit at  $\delta_{CP} = 1.49\pi$  (NH)  
 $\Delta\chi^2$  (IH-NH) = 0.47

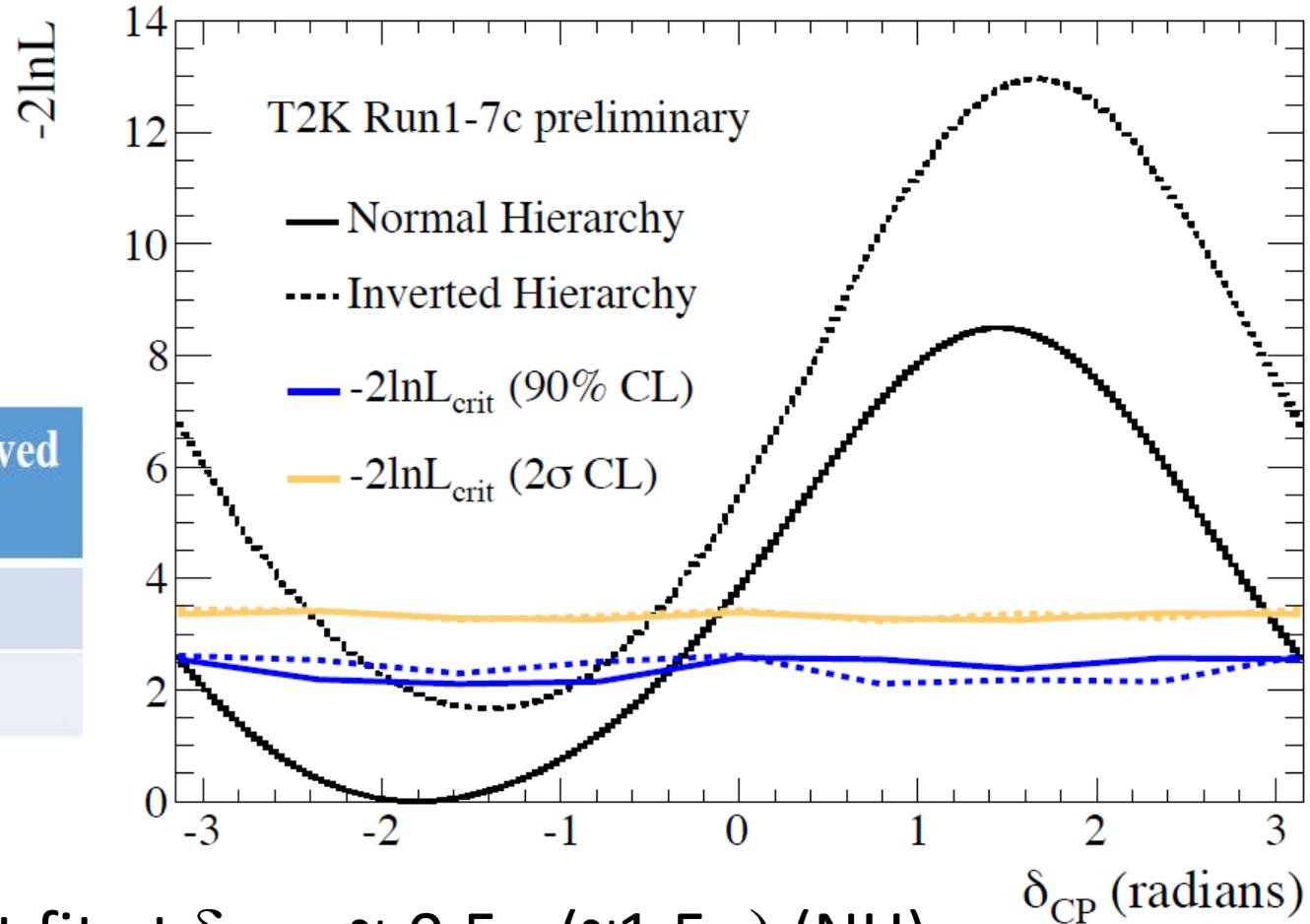
# Mass hierarchy and CP violation measurements @ICHEP-2016

## T2K

32  $\nu_e$  events observed

4 anti- $\nu_e$  events observed

	$\delta_{cp} = -\pi/2$ (NH)	$\delta_{cp} = 0$ (NH)	$\delta_{cp} = +\pi/2$ (NH)	$\delta_{cp} = \pi$ (NH)	Observed
$\nu_e$	28.7	24.2	19.6	24.1	32
$\bar{\nu}_e$	6.0	6.9	7.7	6.8	4



$\delta_{CP} = 0$  and  $\pi$  disfavored at 90%CL. Best fit at  $\delta_{CP} = \sim -0.5\pi$  ( $\sim 1.5\pi$ ) (NH)

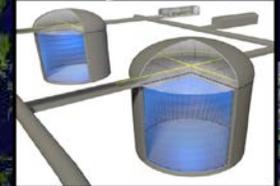
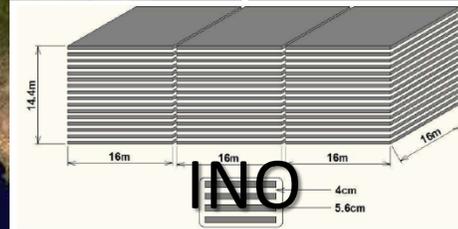
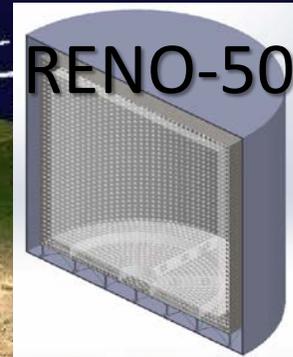
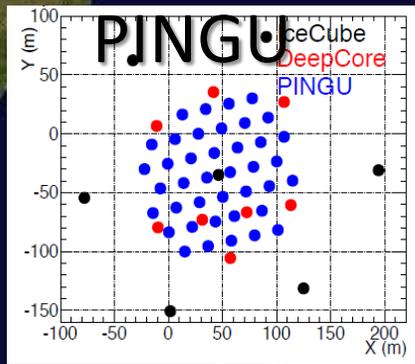
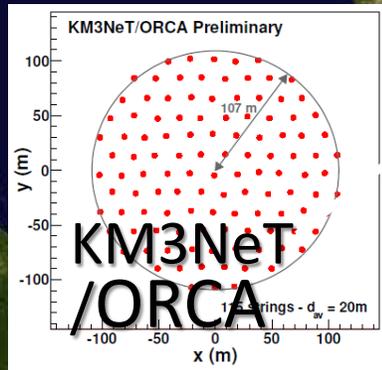
$\Delta\chi^2$  (IH-NH) =  $\sim 1.5$

***Already some interesting hints from 3 experiments!***

# Neutrino mass hierarchy measurements



RENO-50

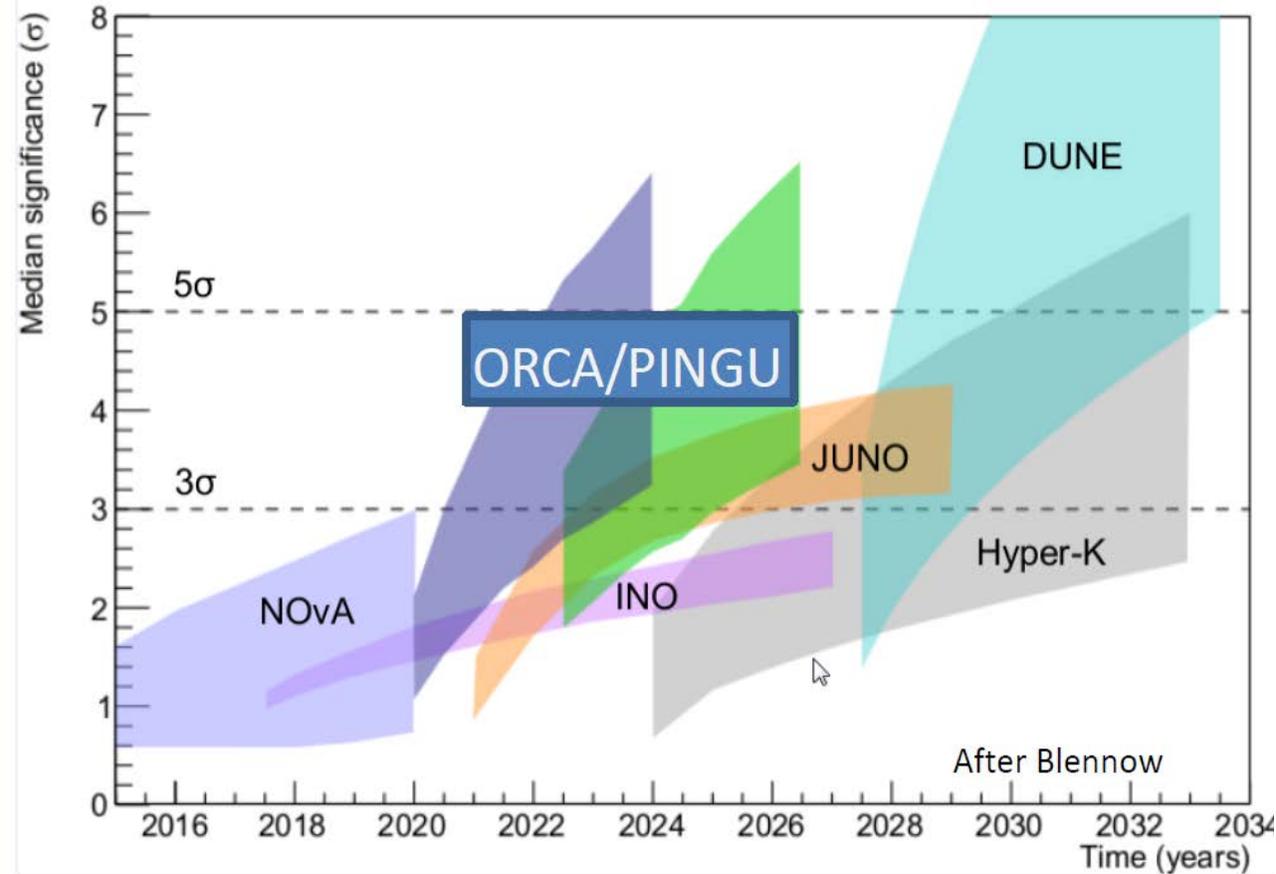


# Timeline: neutrino mass hierarchy

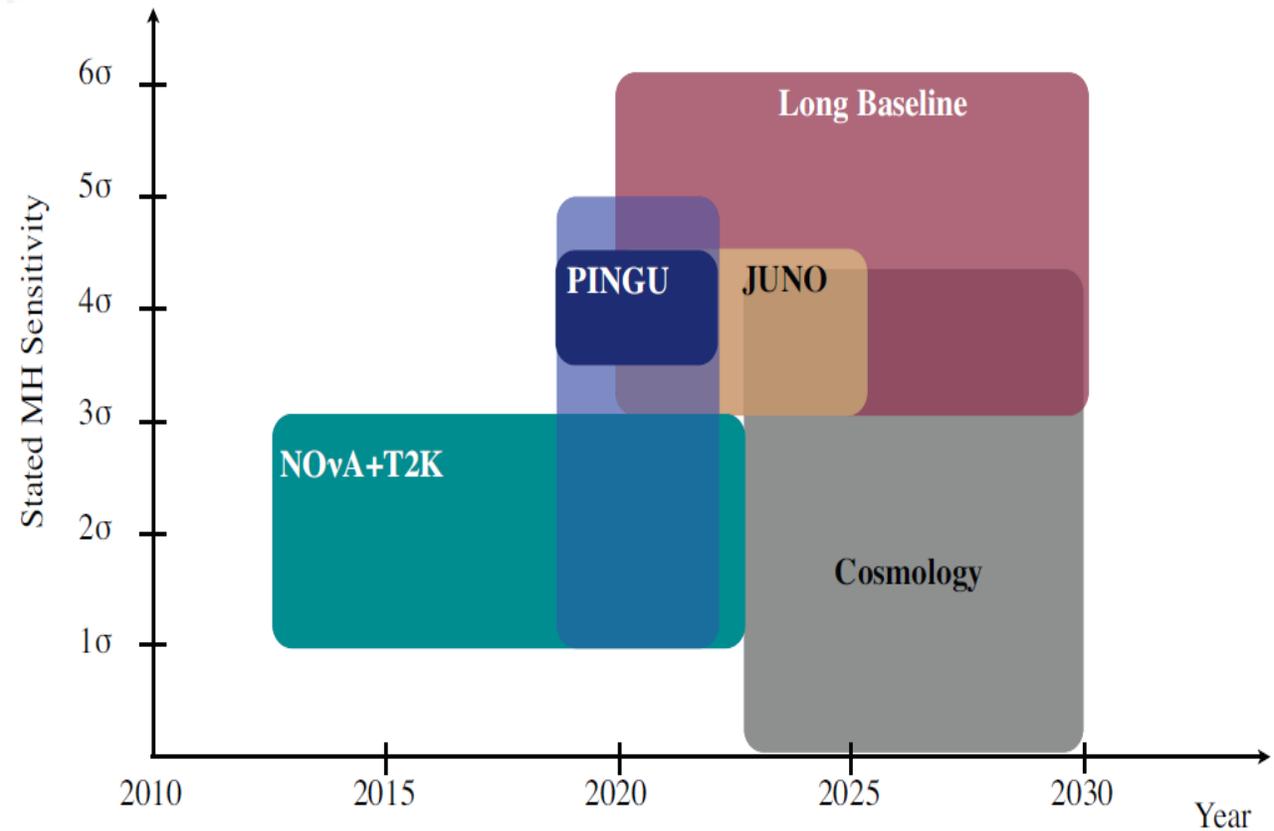
A. Heijboer, NOW2016

M. Blennow et al, JHEP 1403 (2014) 028

### Expected sensitivities vs. Time



P. Huber, NOW2016

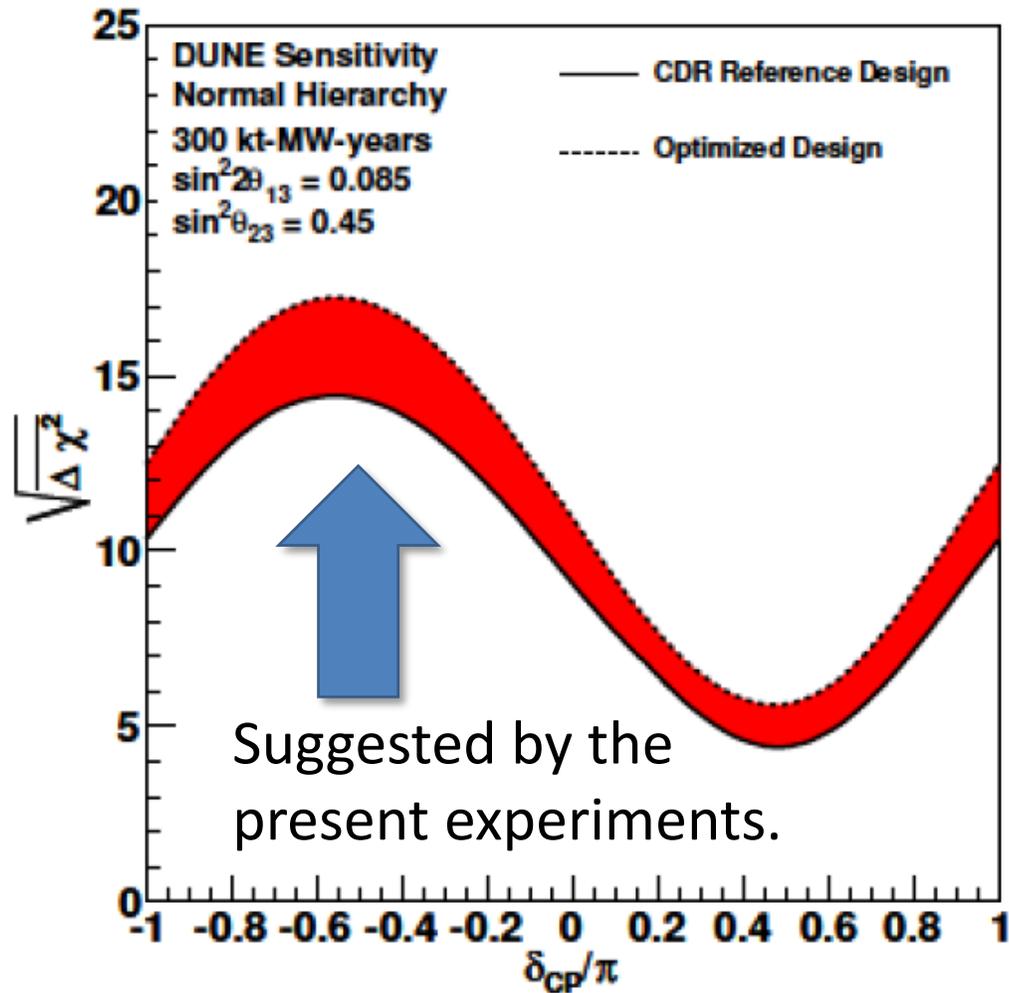


# Sensitivity to mass hierarchy in DUNE

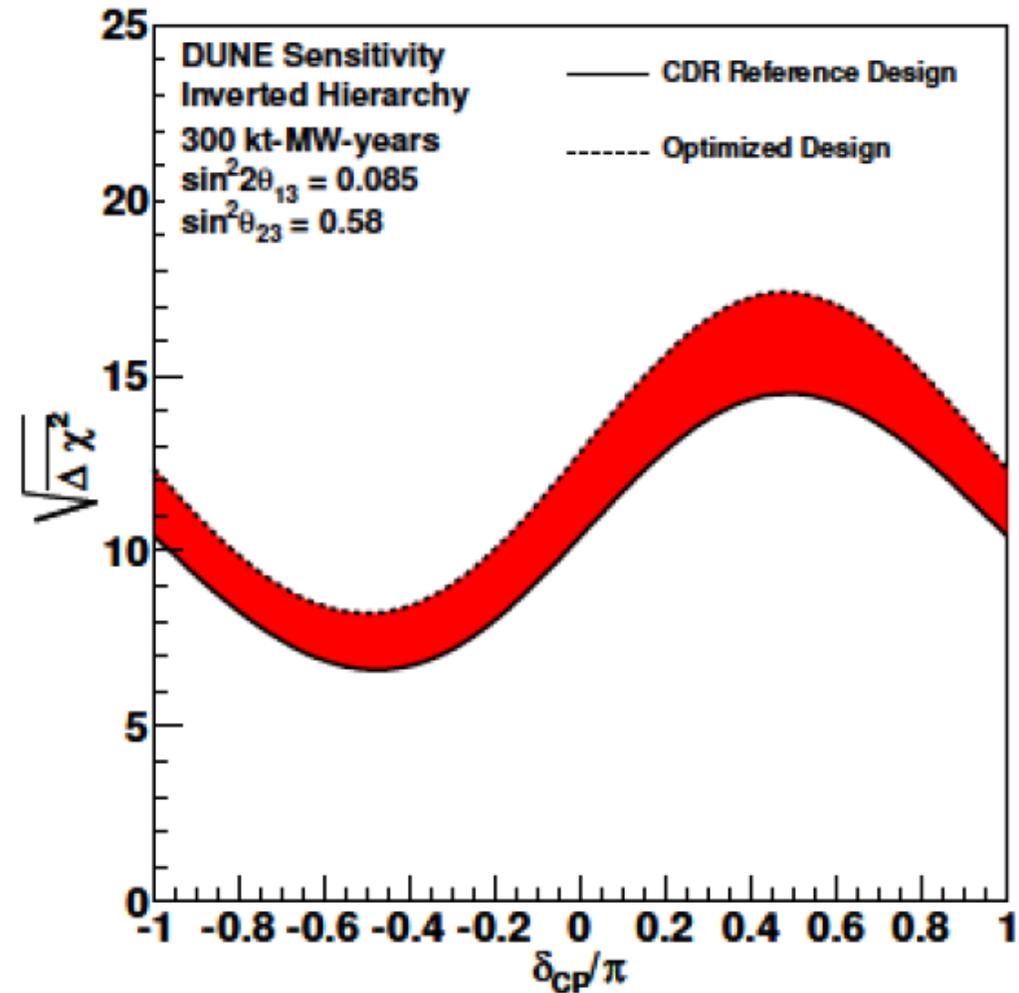
DUNE 300 kt-MW-years ( $\sim 7$  years, 40kton, 1.07MW)

D. Stefan, NOW2016

## Normal hierarchy

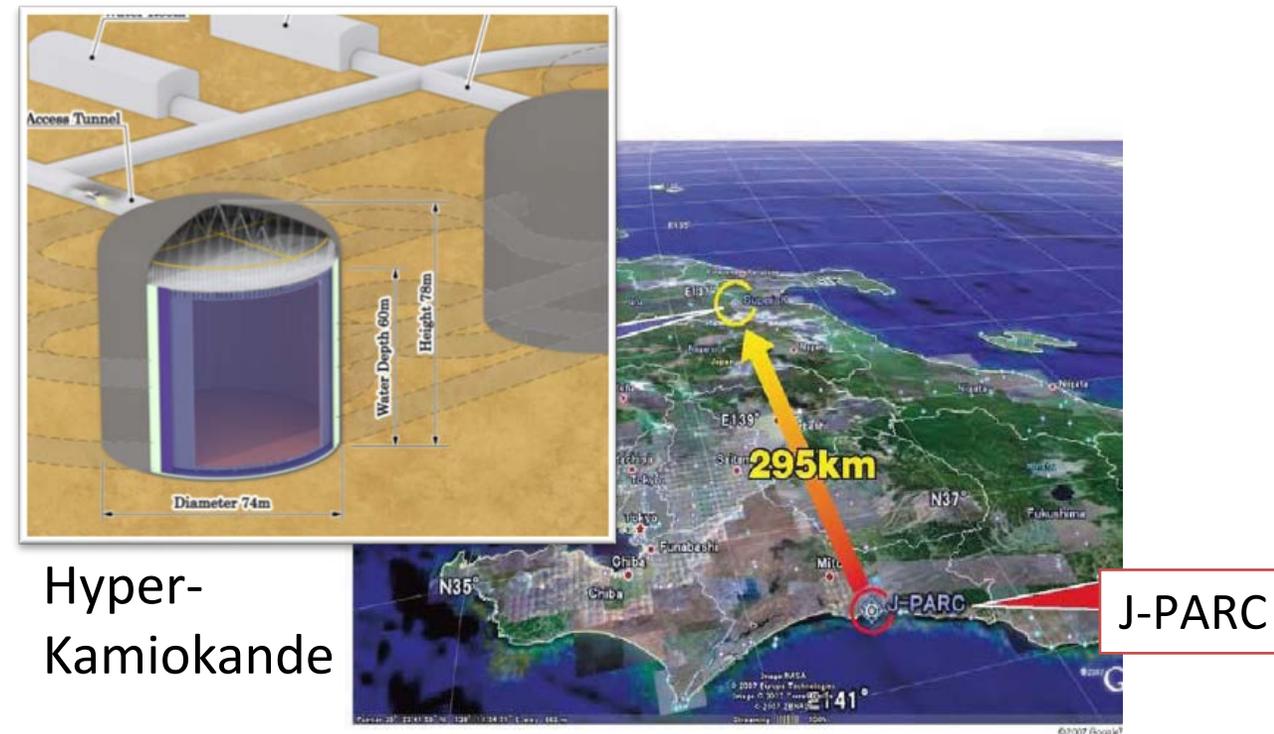
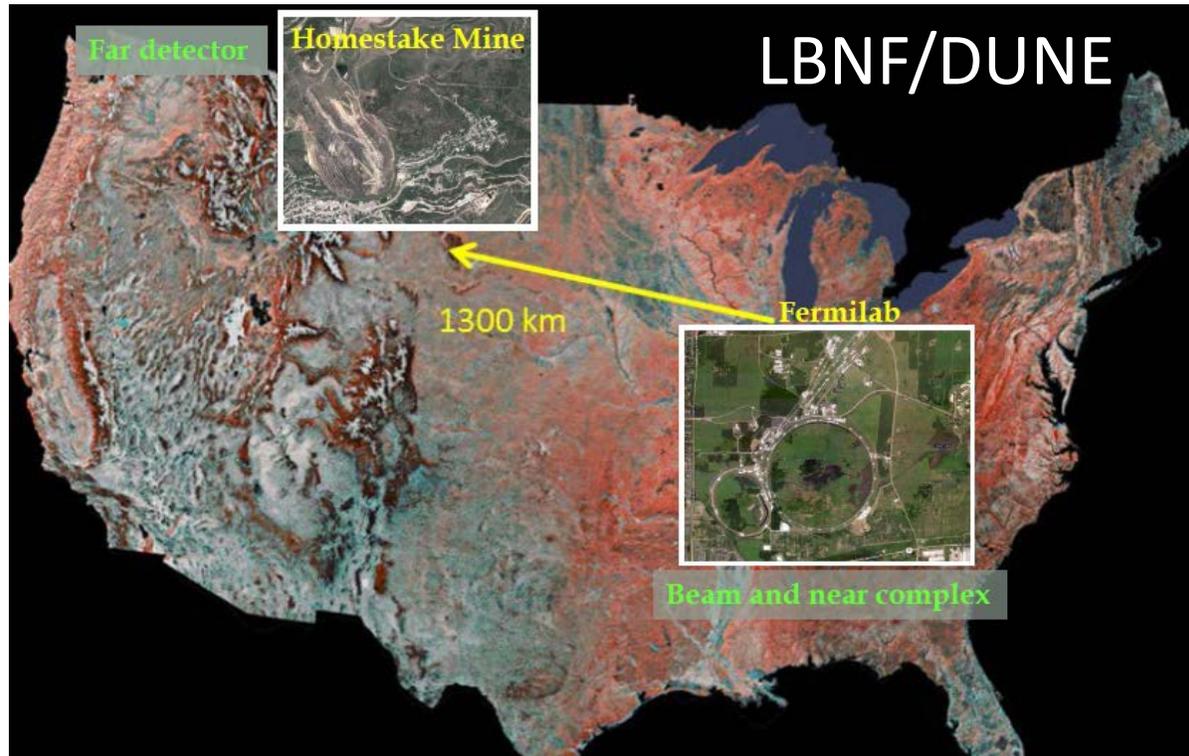


## Inverted hierarchy



# CP violation measurements

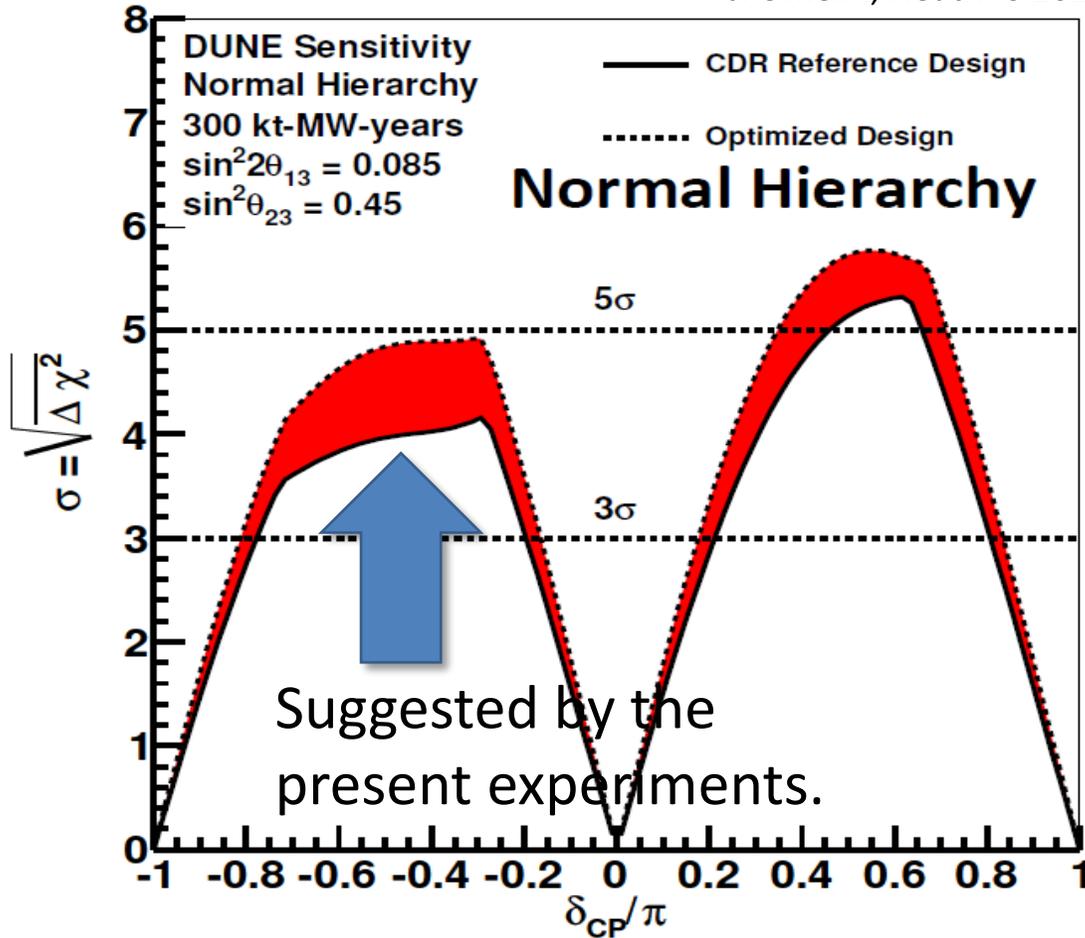
- ✓ We want to observe if  $\nu_{\mu} \rightarrow \nu_e$  oscillation of neutrinos and those of anti-neutrinos are different.
- ✓ This is a difficult experiment. We need the next generation long base line experiments with much higher performance neutrino detectors.



# CP violation sensitivities

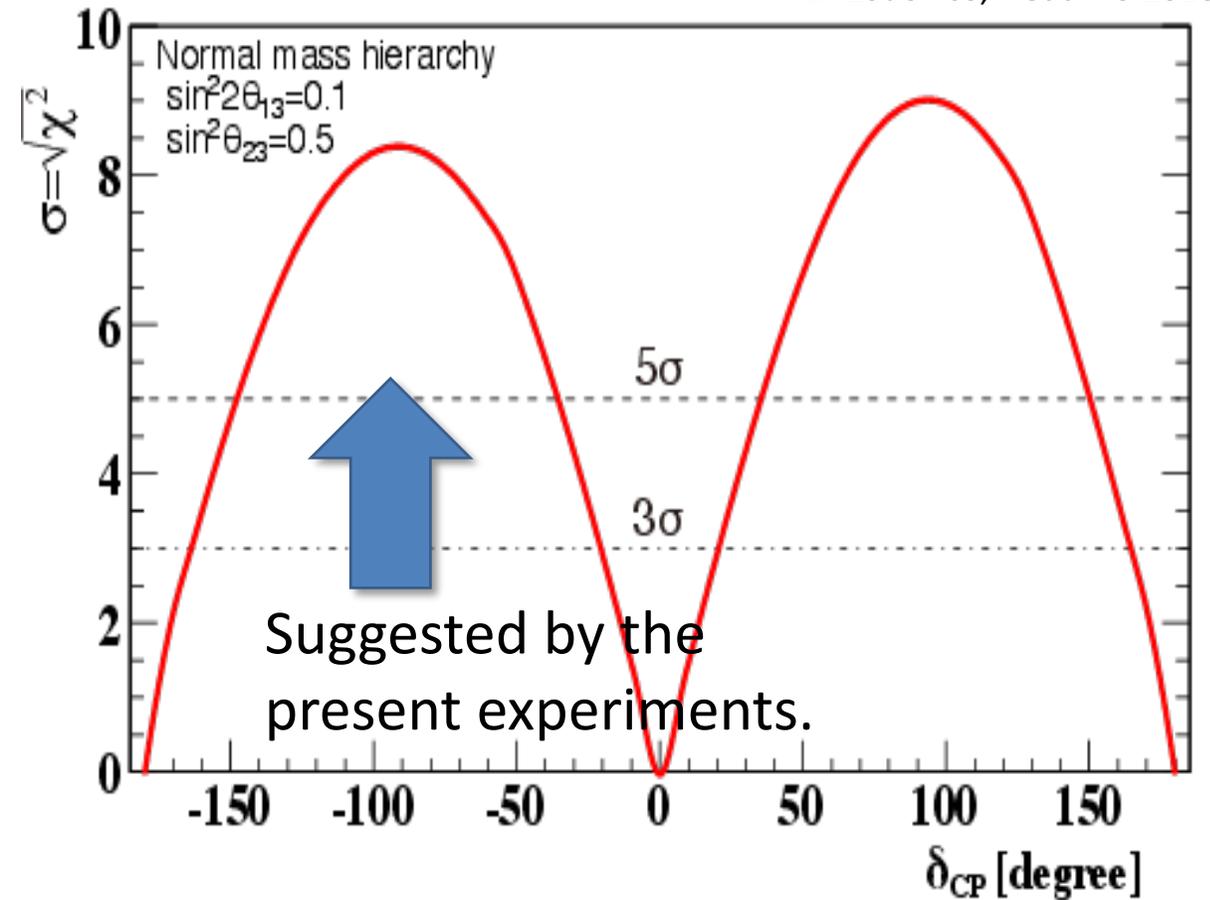
## DUNE (7 years, 1.07MW)

J. Urheim, Neutrino 2016



## Hyper-K (10 years, 1.3MW)

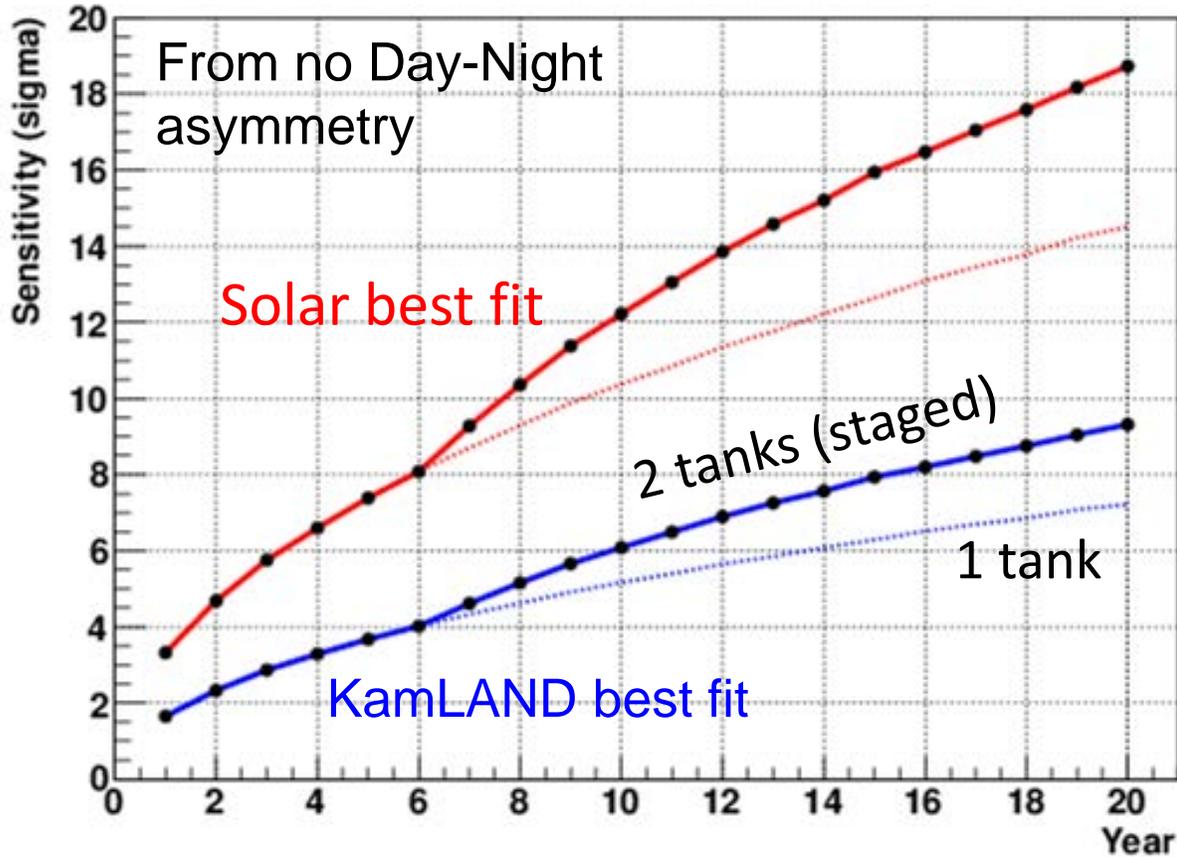
F. Di Lodovico, Neutrino 2016



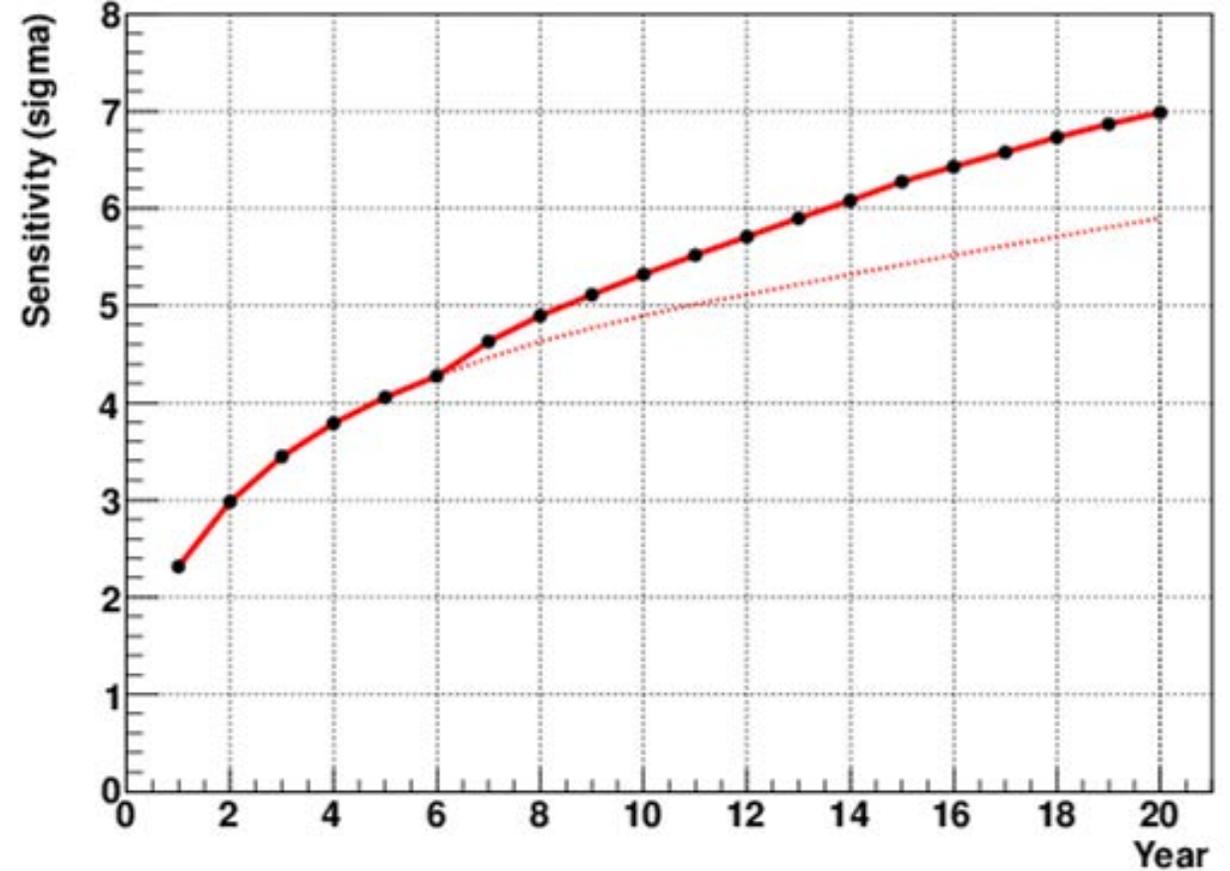
Both projects have high sensitivities to CP violation. (These 2 plots cannot be compared the sensitivities with these figures due to different assumptions)

# Appendix: Solar neutrino measurements (Hyper-K)

## Day-night asymmetry sensitivity

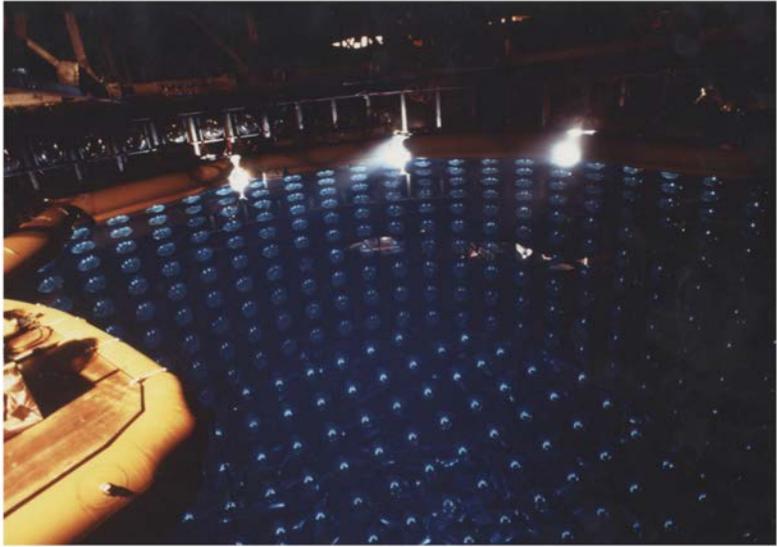


## Spectrum upturn sensitivity



*Other physics with future experiments*

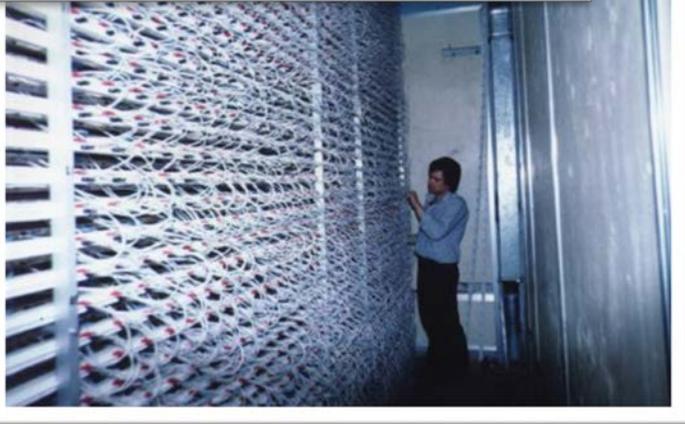
# Proton decay experiments (1980's)



Grand Unified Theories  
(in the 1970's)  
→  $\tau_p = 10^{30 \pm 2}$  years

Kamiokande  
(1000ton)

IMB  
(3300ton)



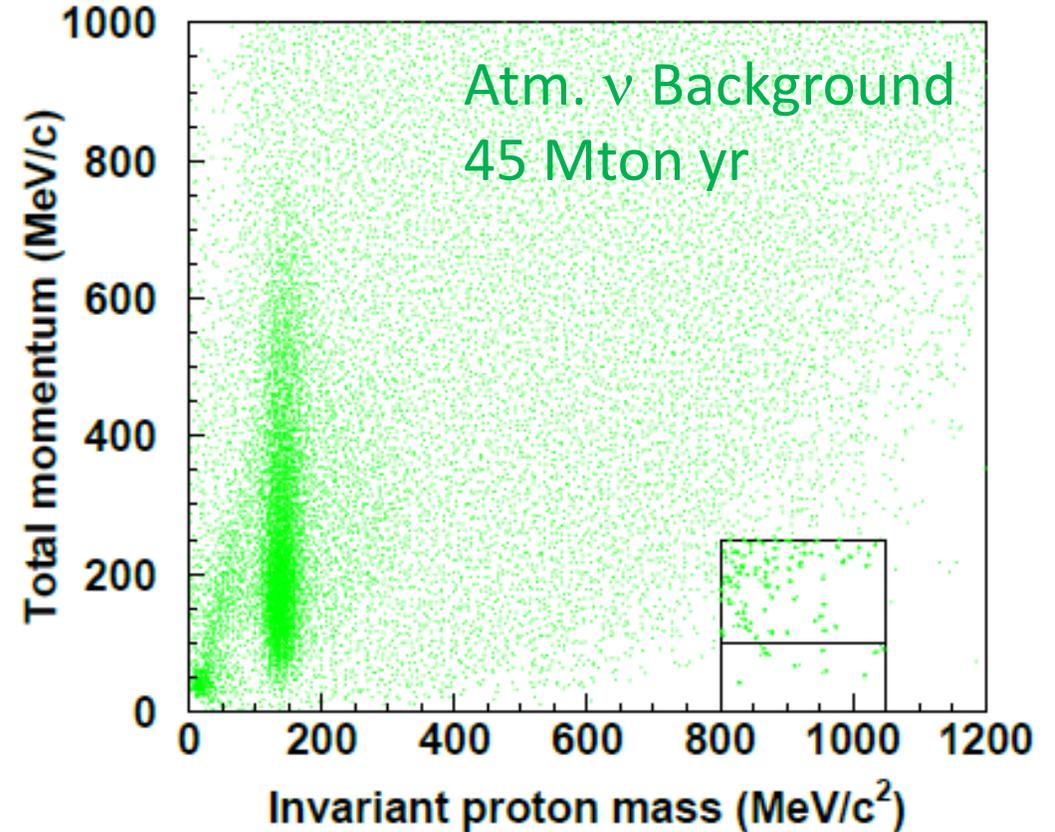
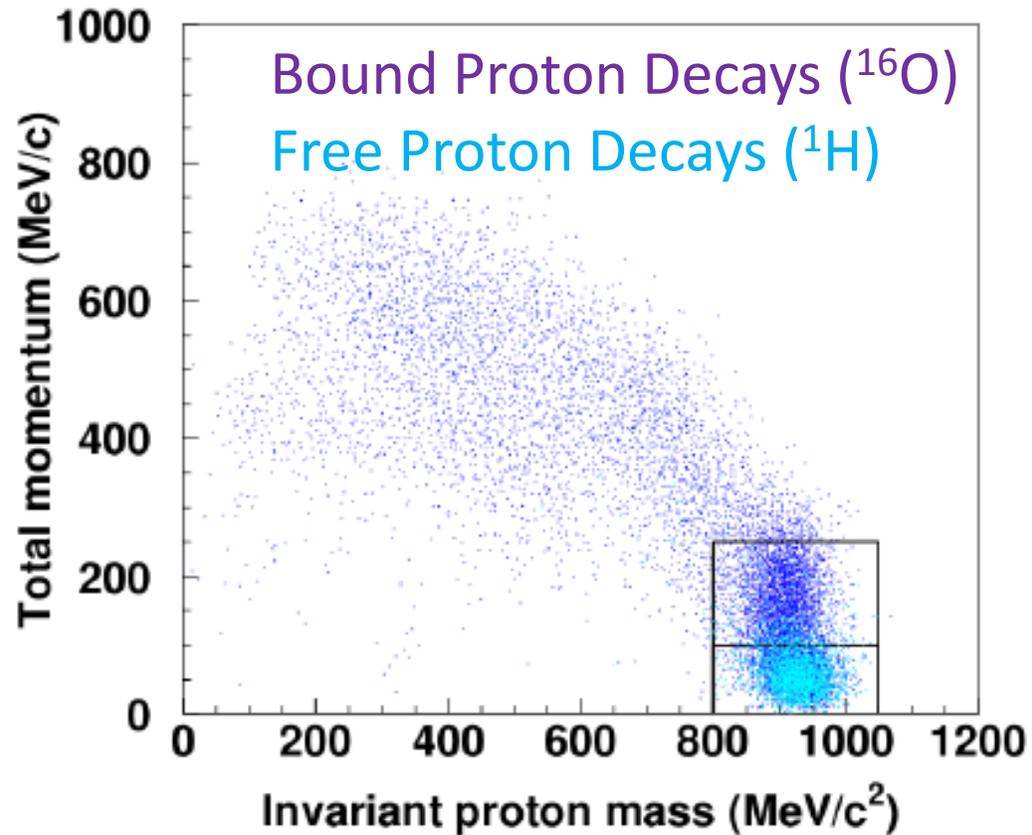
NUSEX  
(130ton)

Frejus  
(700ton)



*These experiments did not observe proton decays and excluded the original GUTs. However, we should not forget about the importance of proton decays!*

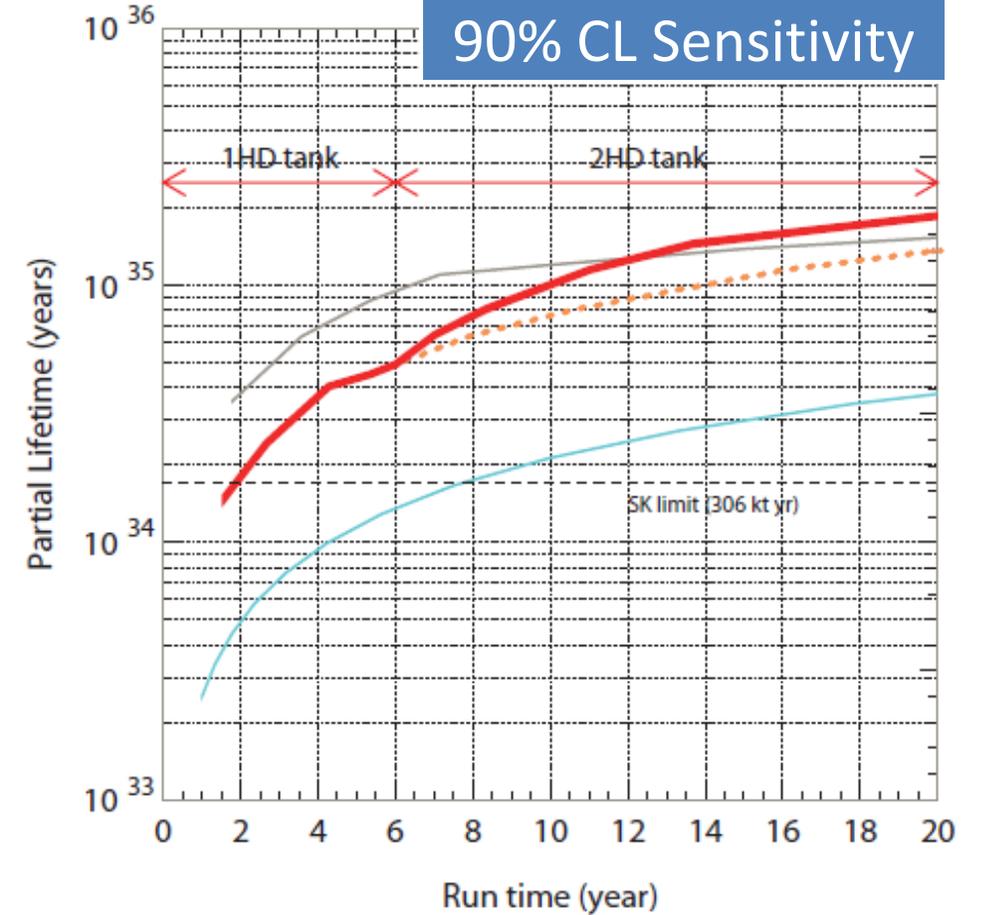
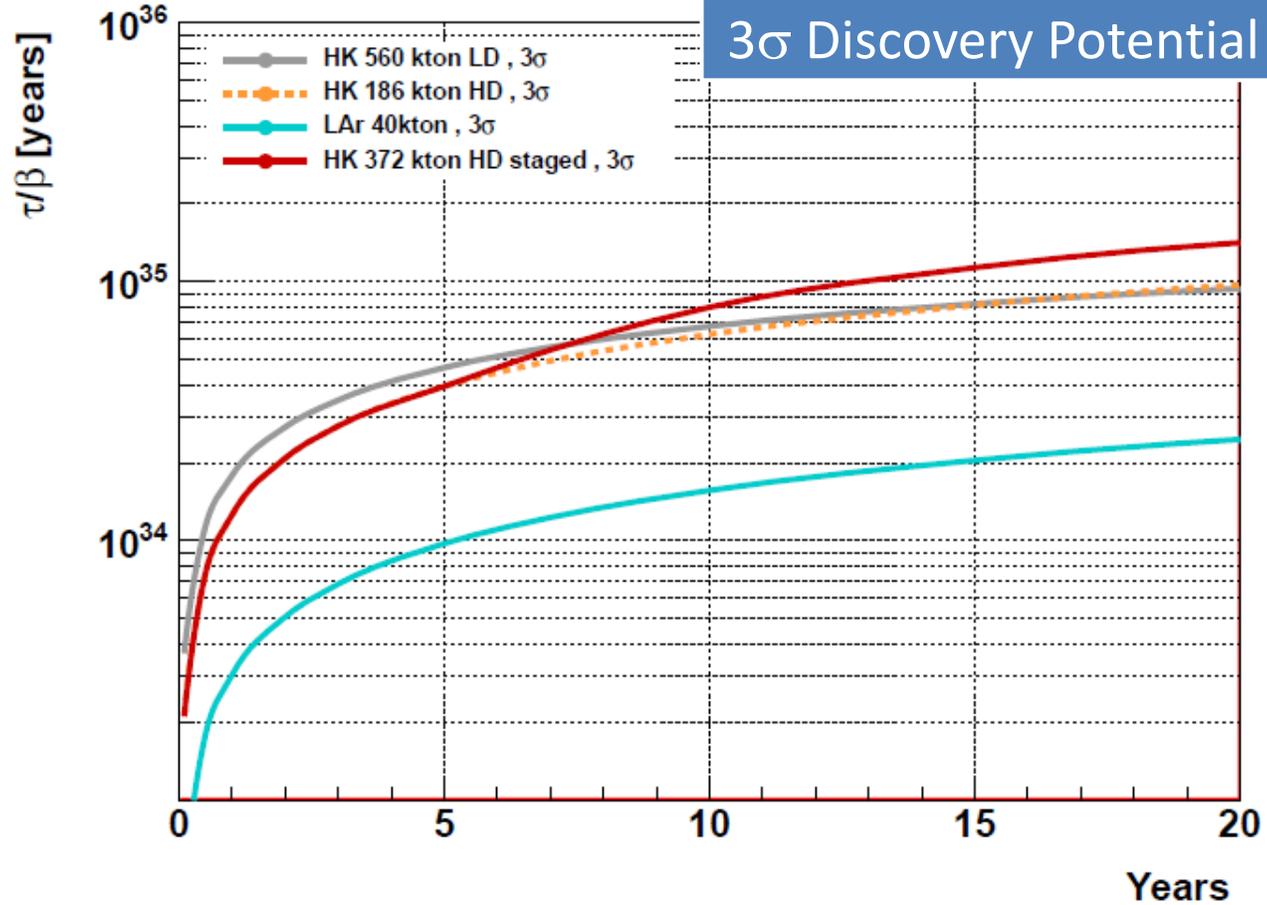
# Proton decay ( $p \rightarrow e^+ \pi^0$ )



	$P_{\text{total}} < 100 \text{ MeV}/c$		$P_{\text{total}} < 250 \text{ MeV}/c$	
	efficiency	Background (/Mtonyr)(*)	Efficiency	Background (/Mtonyr)(*)
Hyper-K	18.7%	0.06	38.1%	0.68

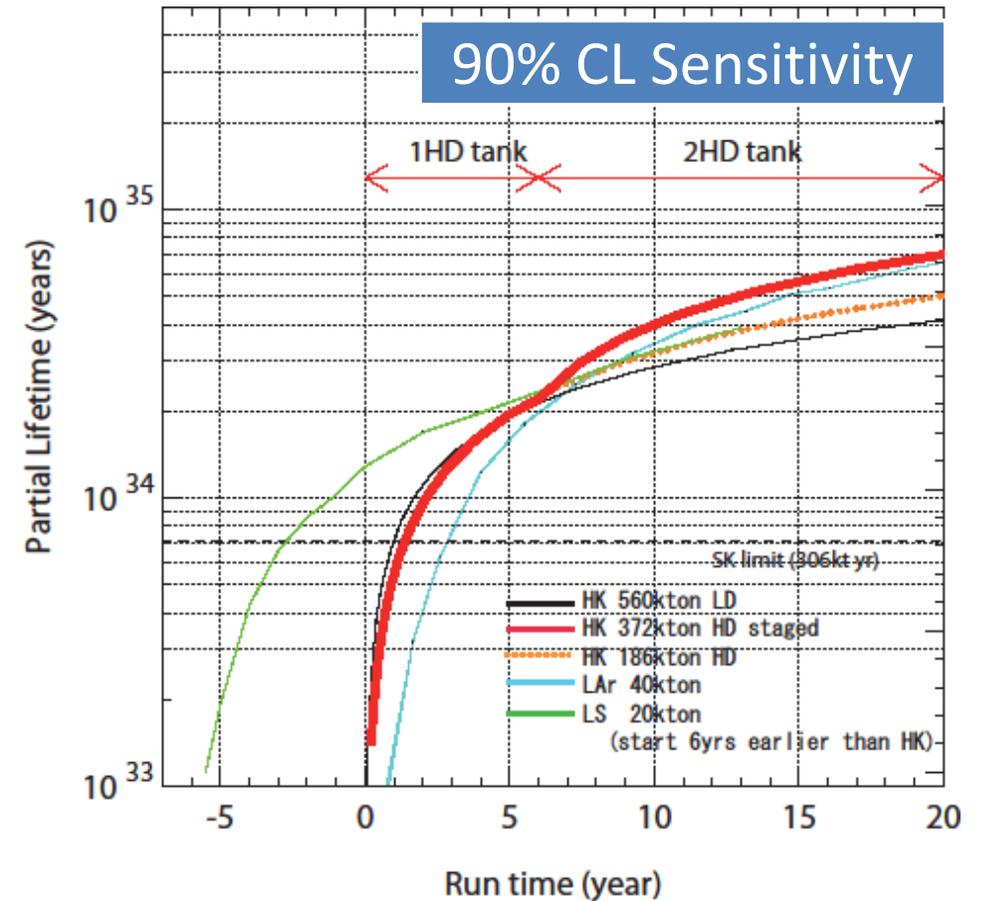
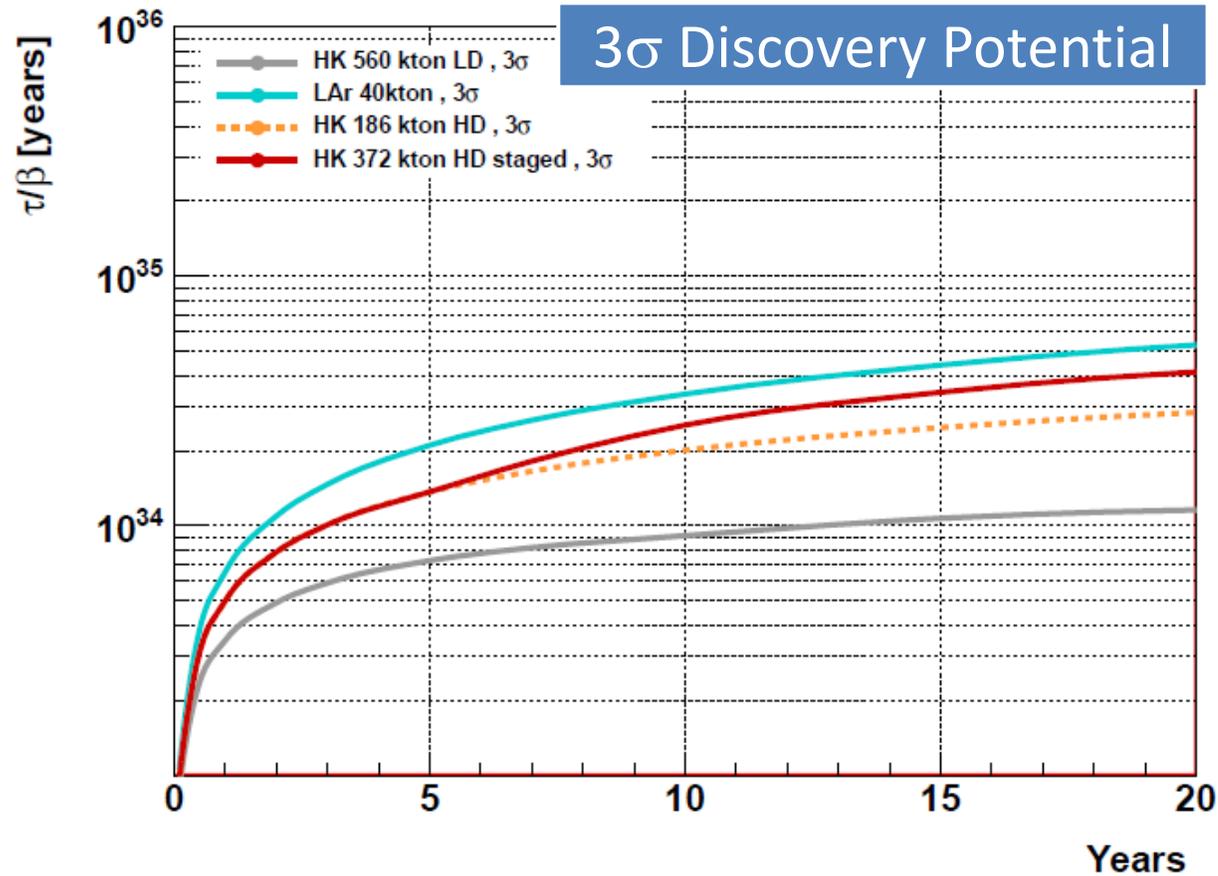
(\*) Neutron tagging included to reduce the background

# $P \rightarrow e^+ \pi^0$ : sensitivity



- ✓  $> 1 \times 10^{35}$  years after 2.7 Mton yr (90%CL) or 3 $\sigma$  discovery with 4.0 Mtonyr.  
(Lines for the liquid argon experiment have been generated based on numbers in the literature (efficiency: 45% bkg: 1 event/Mtonyr ).)

# $P \rightarrow \nu K^+$ : sensitivity



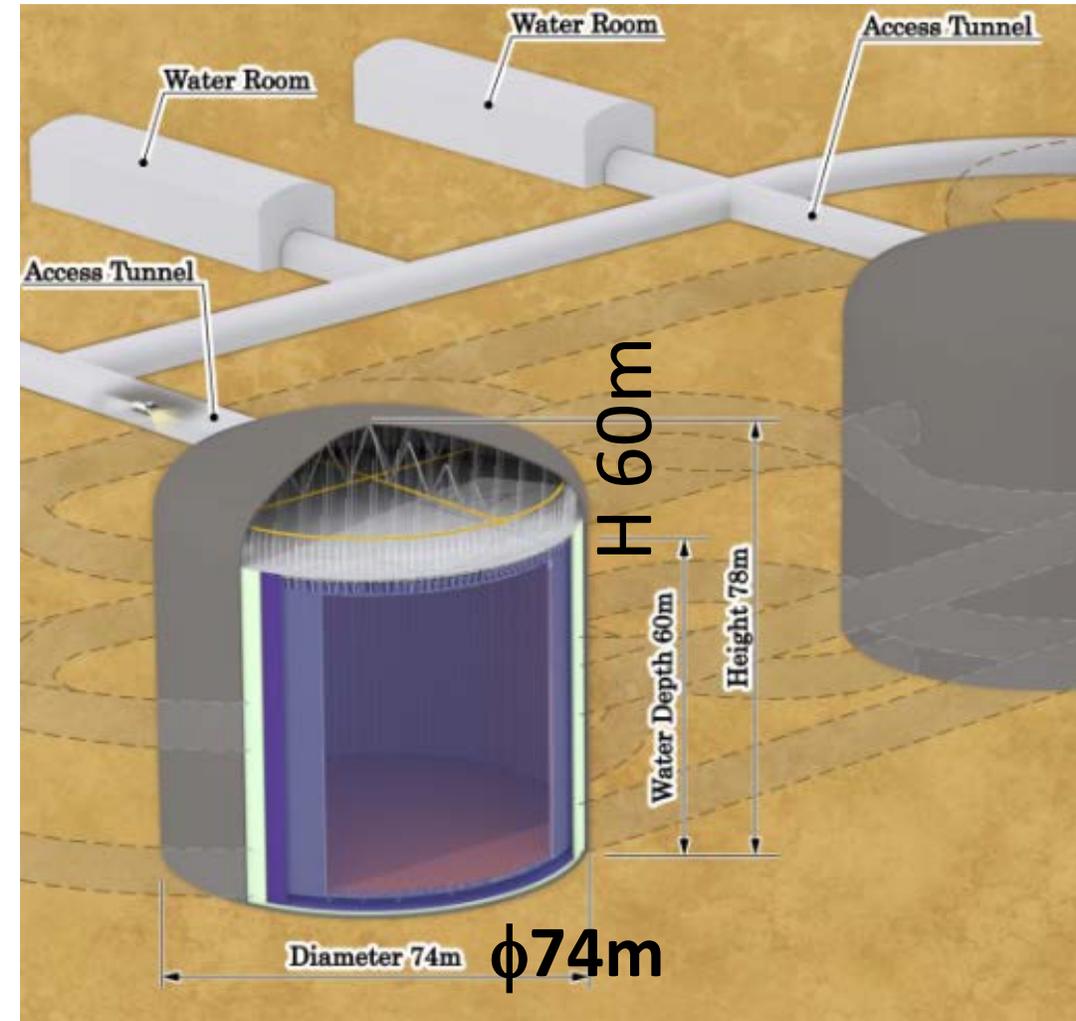
✓ Discovery potential higher in Liq. Ar. (Hyper-K slightly better for 90%CL limit.)

(Discovery potential for the Lar experiment has been generated based on numbers in the literature (efficiency: 97%, bkg: <1 event/Mton year ).

# Appendix: Status of Hyper-K

- ✓ The baseline design was changed to “0.26Mton X 2 tanks (0.38Mton fid. mass) with staging construction” due to the cost issue.
- ✓ However, due to the successful R&D on the PMT improvements, the expected overall performance is not changed substantially.
- ✓ The plan of Hyper-K was submitted to the Japanese Council of Japan. If things goes well, Hyper-K will be listed as one of the “Master Plan Projects”.
- ✓ Hyper-K will be reviewed by the Ministry of the Japanese Ministry of Education, Culture, Sports, Science and Technology (MEXT) in 2017. If it is evaluated highly, it will be listed in the MEXT “Roadmap”.
- ✓ Then, ...

(There is a new possibility to construct the 2<sup>nd</sup> detector in Korea.)



# Summary

- *Since the discovery of neutrino oscillations, various experiments have studied neutrino oscillations. We have almost understood the basic structure of neutrino oscillations.*
- *However, we still do not know some key features of neutrino properties. Neutrinos are one of the keys to understand the physics beyond the Standard Model of particle physics.*
- *It is very good that there are many planned/on-going future neutrino oscillation experiments.*
- *It is also good that these future experiments are sensitive to other important physics, such as proton decay.*
- *We will learn much more on neutrinos in the coming years!*