## A Probabilistic Catalogue of Unresolved High Latitude Fermi LAT Sources

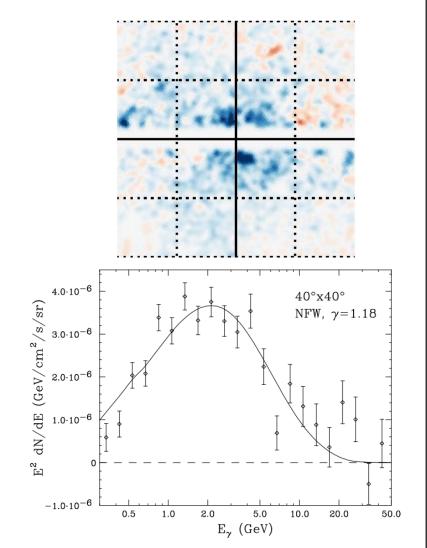
Stephen K N PORTILLO with Tansu DAYLAN and Douglas P FINKBEINER 9 December 2015
Gamma Rays and Dark Matter, Obergurgl



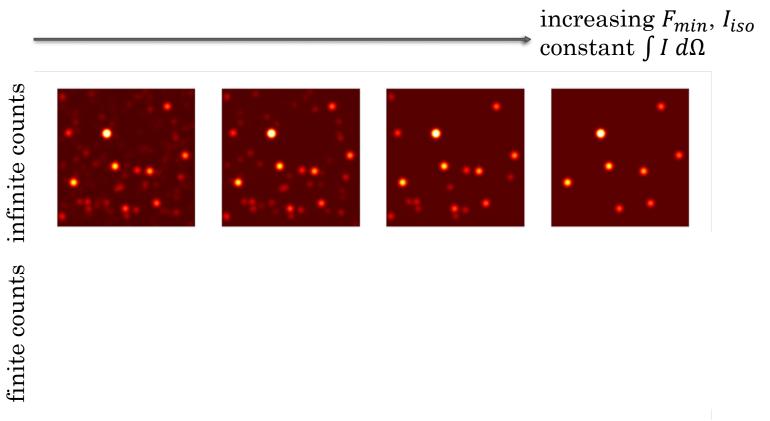


#### GeV Excess Interpretations

- Excess about Galactic Centre in *Fermi* LAT data
- Morphology, intensity, and spectrum compatible with dark matter interpretations
- May also be interpreted as a new population of faint point sources



# Point Source Populations vs Diffuse Sources

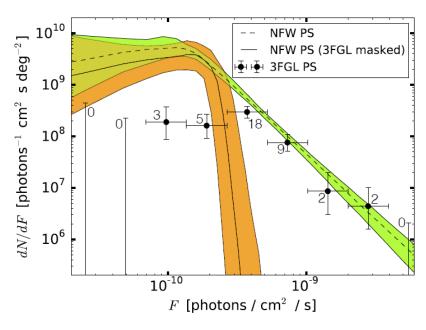


Counting statistics introduce a point source detection threshold But even sources fainter than this threshold can affect image

### Non-Poissonian Template Fitting

- A point source population with a given spatial distribution can be treated as a diffuse source with non-Poissonian statistics
- Lee et al. 2015 greatly favours a point source interpretation of the GeV Excess
- Inferred population's luminosity function peaks just below Fermi LAT detection threshold

(cp. talks by RODD and ZECHLIN)



NFW template point sources NFW template point sources minus 3FGL

#### (Deterministic) Catalogues

- A (deterministic) catalogue is a list of point source candidates above some inclusion threshold  $F_{incl}$   $Data, F_{incl} \rightarrow \left\{\ell_i \pm \sigma_{\ell_i}, \delta_i \pm \sigma_{\delta_i}, F_i \pm \sigma_{F_i}\right\}_{i=1}^N$
- Inclusion threshold = detection threshold: Almost all catalogue sources are true sources But faint true sources are not in the catalogue
- Inclusion threshold < detection threshold:
   <p>More faint true sources are included in the catalogue
   But many catalogue sources are not true sources
   The data is overfitted

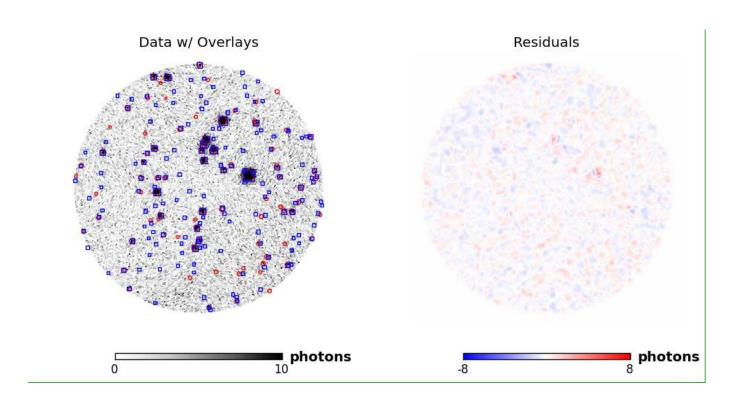
#### Probabilistic Catalogues

- A probabilistic catalogue is a probability distribution over the space of lists of point source candidates  $P(\{\ell_i, \delta_i, F_i\}_{i=1}^N | Data)$
- Sampling the probabilistic catalogue provides an ensemble of catalogues consistent with the data
- This ensemble captures the degeneracies of point source identification
- The reality of a single faint point source candidate will be very uncertain, but the properties of a faint population are constrained
- I currently use Brendon Brewer's DNest3+RJObject to sample
- This technique is a different approach to the GeV Excess that retains more information than non-Poissonian template fitting

#### Validation at High Latitude

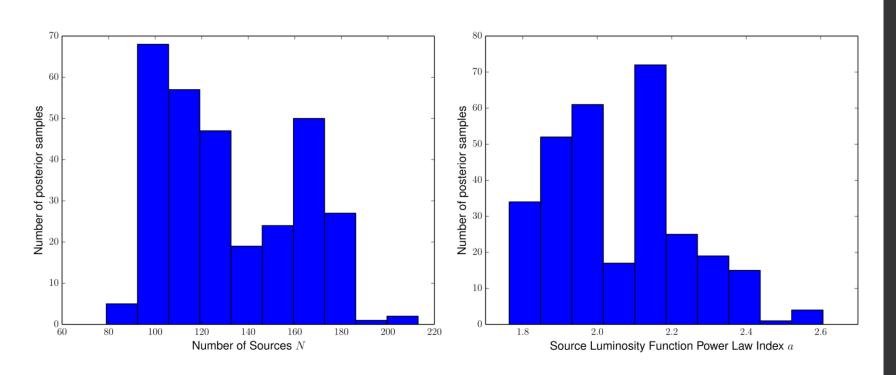
- North Galactic Pole  $b > 70^{\circ} (N_{pix} = 23544)$
- Region includes 84 3FGL sources
- Run with ~900 core-hours
- Diffuse sources:
  - Galactic diffuse emission
  - Isotropic emission
- Point source population:
  - Mostly distant active galaxies
  - Assumed to be isotropically distributed
  - Unknown flux distribution parameterized as power law

#### Probabilistic Catalogue Samples



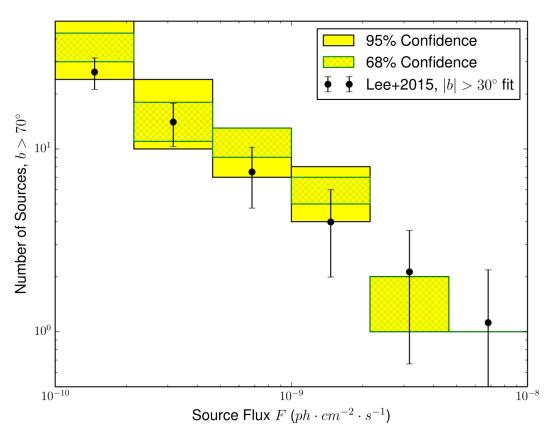
red circles – 3FGL point sources blue squares – point source candidates

#### Point Source Population Constraints



Pass 8 SOURCE 1-3 GeV Weeks 9-217

#### Comparison



Pass 8 ULTRACLEANVETO PSF3 1.893-11.943 GeV Weeks 9-365

#### Conclusion

- Constraining the point source contribution to the GeV Excess essential to its interpretation
- A point source population can be distinguished from a diffuse source, even if the individual sources are below the detection threshold
- Probabilistic catalogues capture the degeneracies of point source identification
- We have constructed a probabilistic catalogue for high latitude *Fermi* LAT sources in reasonable agreement with the 3FGL and non-Poissonian template fitting
- Stay tuned for Galactic Centre results...