

Compton Scattering and Radiation Reaction Studies

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technical staff*



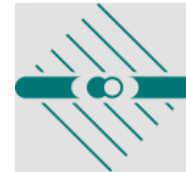
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C. Harvey

Talk Outline

- **RATIONALE**
 - Next generation of multi-MeV γ -rays from PW interactions
 - Quantum and Classical Radiation Reaction
 - Photon-photon scattering cross-section

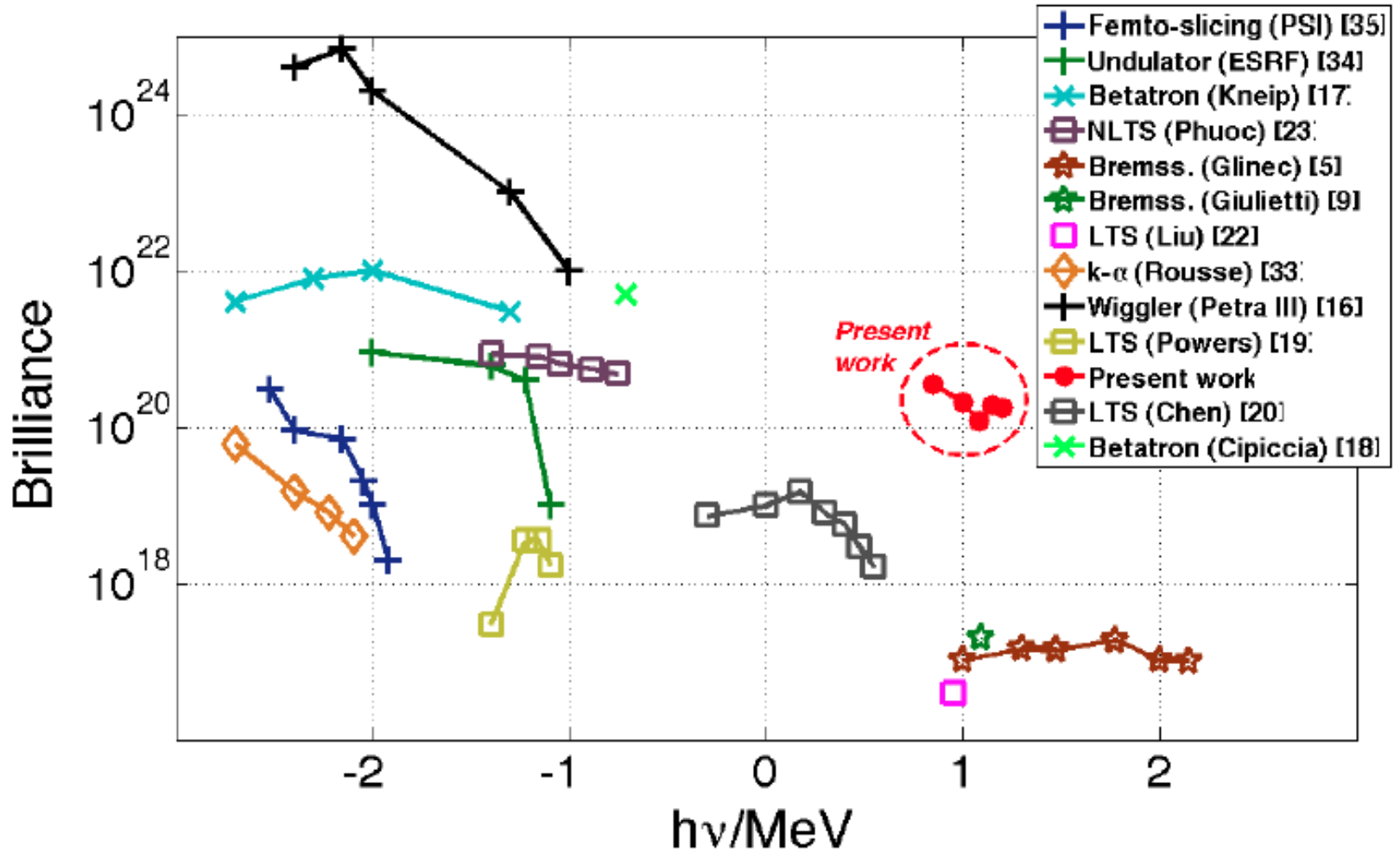
- **THEORETICAL PREDICTIONS**
 - Next generation of multi-MeV γ -rays
 - Quantum and Classical Radiation Reaction
 - Photon-photon scattering cross-section

- **GENERAL EXPERIMENTAL SETUP**

- **EXPERIMENTAL TECHNIQUES**
 - fs-scale synchronisation
 - γ -ray spectrometer
 - Pixelated scintillators
 - Double magnetic spectrometer

- **CONCLUSIONS**

▪ **NEXT GENERATION OF MULTI-MEV γ -RAYS**



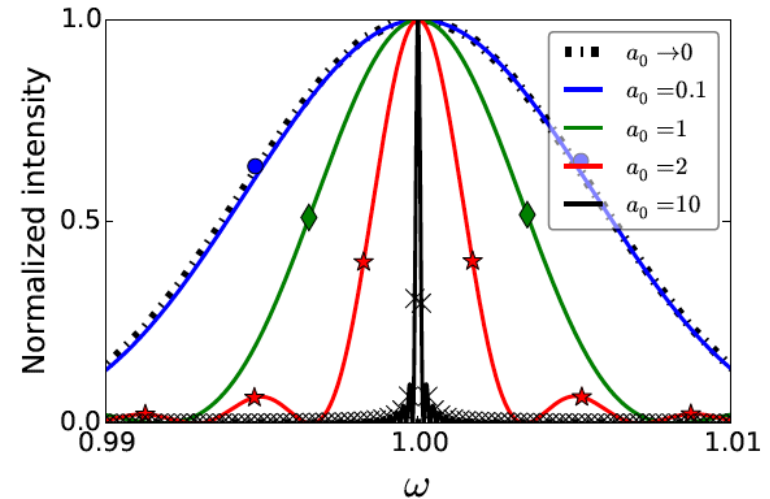
Rationale

MONOCHROMATIC γ -RAYS FROM NON-LINEAR COMPTON SCATTERING

Photon energy:
$$\hbar\omega_X \sim \frac{4\gamma_e^2 \hbar\omega_L N_{ABS}}{1 + (\gamma_e\theta)^2 + \frac{a_0^2}{2} + 2\frac{N_{ABS}\chi}{a_0}}$$

Photon number:
$$\begin{cases} N_X \sim 1.5 \times 10^{-2} a_0^2 & \text{for } a_0 \ll 1 \\ N_X \sim 3.38 \times 10^{-2} a_0 & \text{for } a_0 \geq 1 \end{cases}$$

Bandwidth:
$$\frac{\Delta\omega_X}{\omega_X} \sim \sqrt{\left(\frac{a_0^2}{2}\right)^2 + \left(\frac{\Delta\omega_L}{\omega_L}\right)^2 + \left(\frac{2\Delta\gamma_e}{\gamma_e}\right)^2}$$



S. Rykovanv et al. ArXiv (2014)

INPUT

Electron energy ~ 1 GeV
 Electron bandwidth $\sim 10\%$
 Electron number $\sim 10^9$
 Laser dimensionless intensity ~ 10
 Chirped laser pulse

OUTPUT

Photon energy > 100 MeV
 Photon pulse duration ~ 10 fs
 Photon number $\sim 4 \times 10^8$
 Photon divergence ~ 5 mrad
 Bandwidth $\sim 20\%$
 Brilliance: $> 10^{21}$ ph s $^{-1}$ mm $^{-2}$ mrad $^{-2}$ x0.1%BW

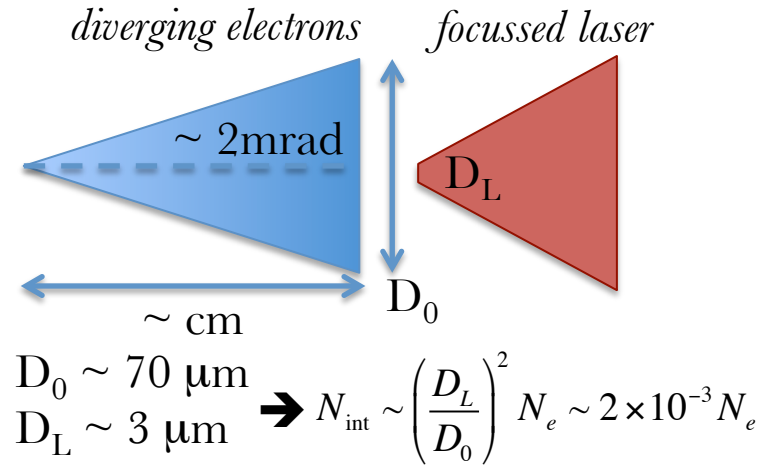
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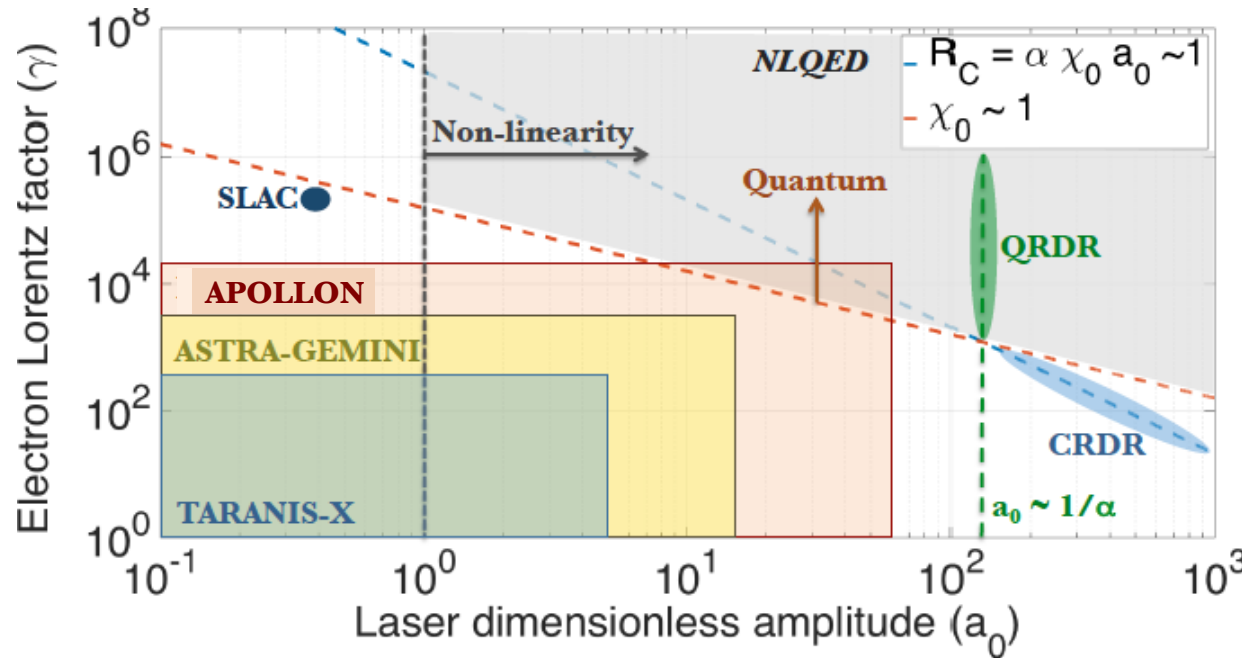
INPUT

Electron energy $\sim 0.6 \text{ GeV}$
 Electron bandwidth $\sim 10\%$
 Electron number $\sim 10^9$
 Laser focal spot $\sim 70 \mu\text{m}$
 Laser dimensionless intensity ~ 6
 Chirped laser pulse

OUTPUT

Photon energy $\sim 60 \text{ MeV}$
 Photon pulse duration $\sim 10 \text{ fs}$
 Photon number $\sim 2 \times 10^8$
 Photon divergence $\sim 3.6 \text{ mrad}$
 Bandwidth $\sim 20\%$
 Brilliance: $\sim 10^{21} \text{ ph s}^{-1} \text{ mm}^{-2} \text{ mrad}^{-2} \times 0.1\% \text{ BW}$

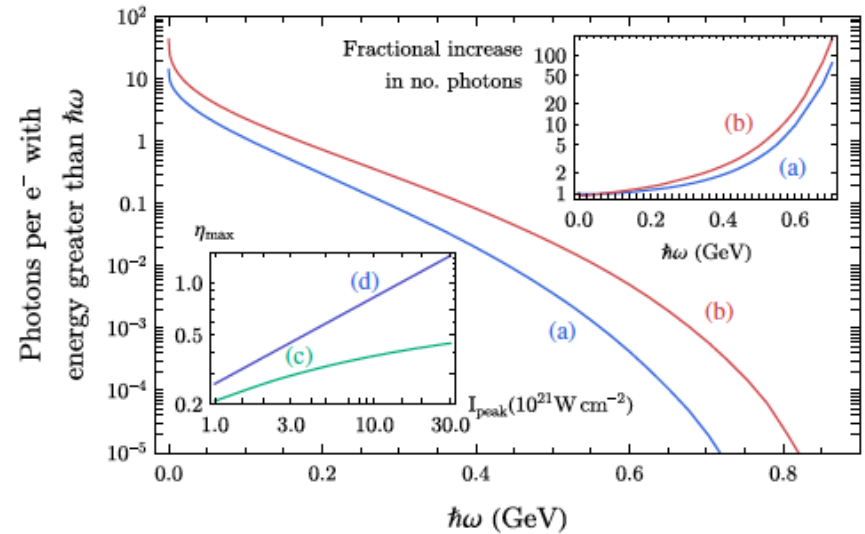
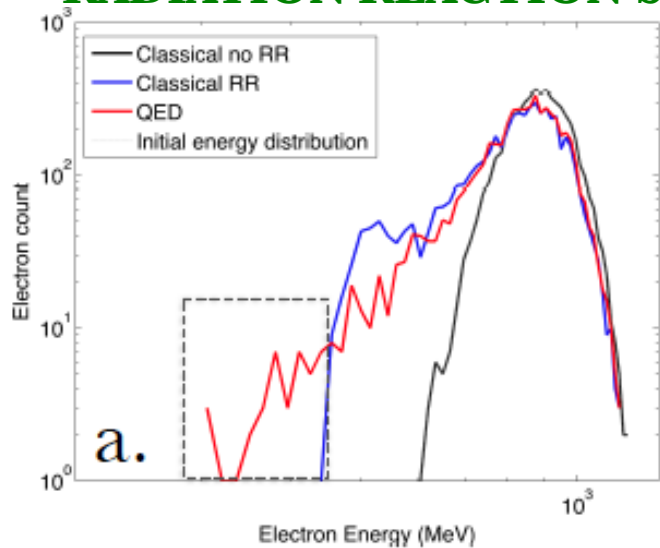
■ RADIATION REACTION STUDIES



- *non-linearity* ($a_0 > 1$)
equivalent to multi-photon absorption
- *quantum effects* ($\chi \sim 1$)
significant electron recoil and pair production
- **CRDR**
electron converting all its energy in photons in a laser period
- **QRDR**
multiple photon emission in one laser period

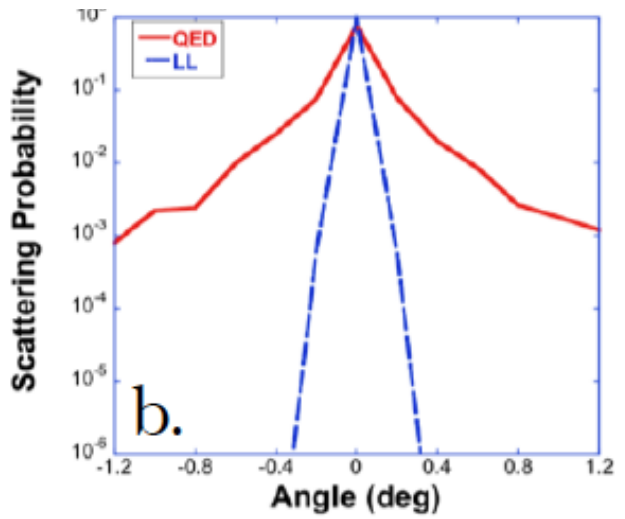
Possibility to access electrodynamics beyond the LL approximation of interest for particle acceleration and astrophysics

■ RADIATION REACTION STUDIES

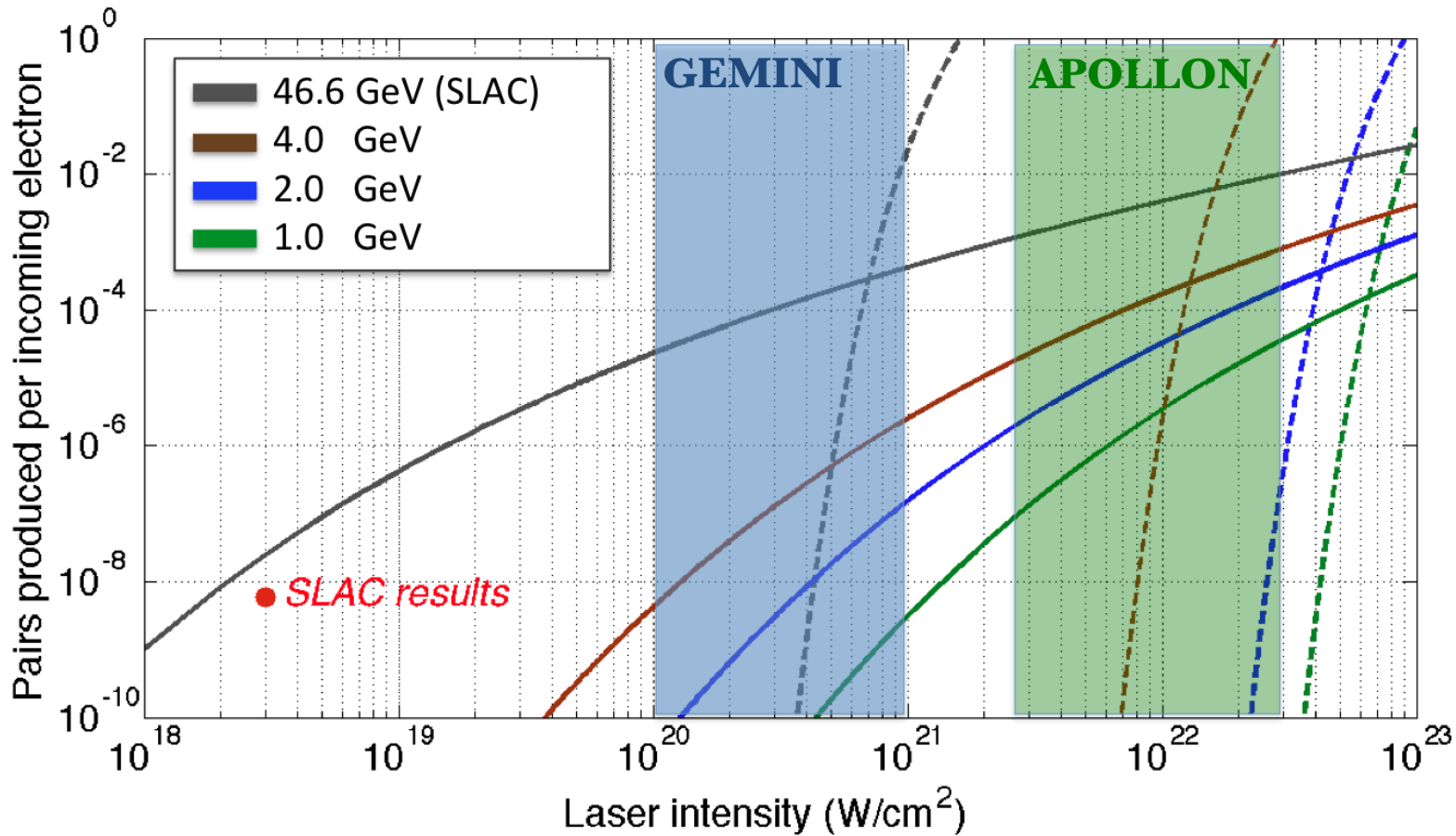


T. Blackburn et al. Phys. Rev. Lett. 112, 015001 (2014)

- Possibility of appreciating stochastic photon emission
- Simultaneous signatures in the electron spectrum and photon beam properties

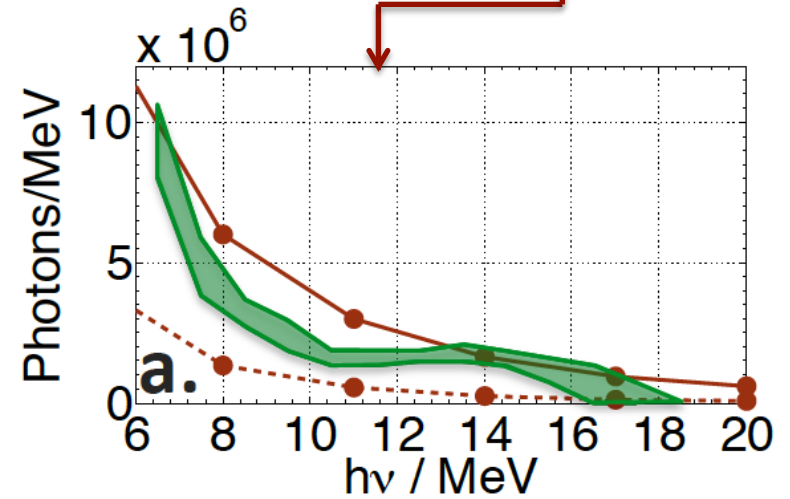
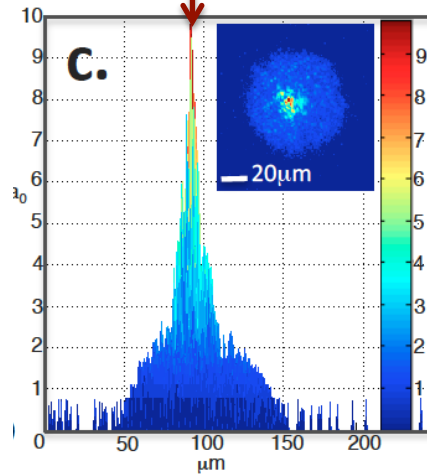
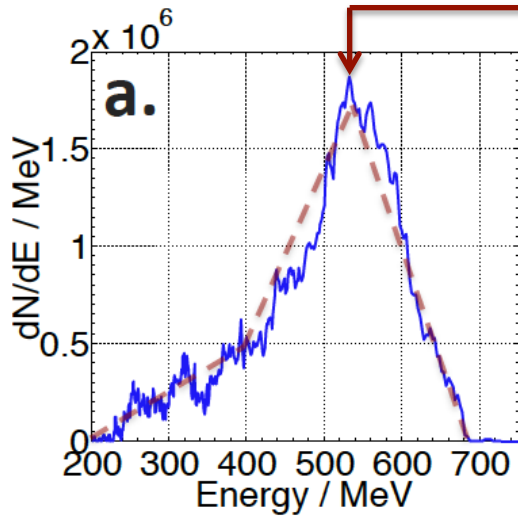
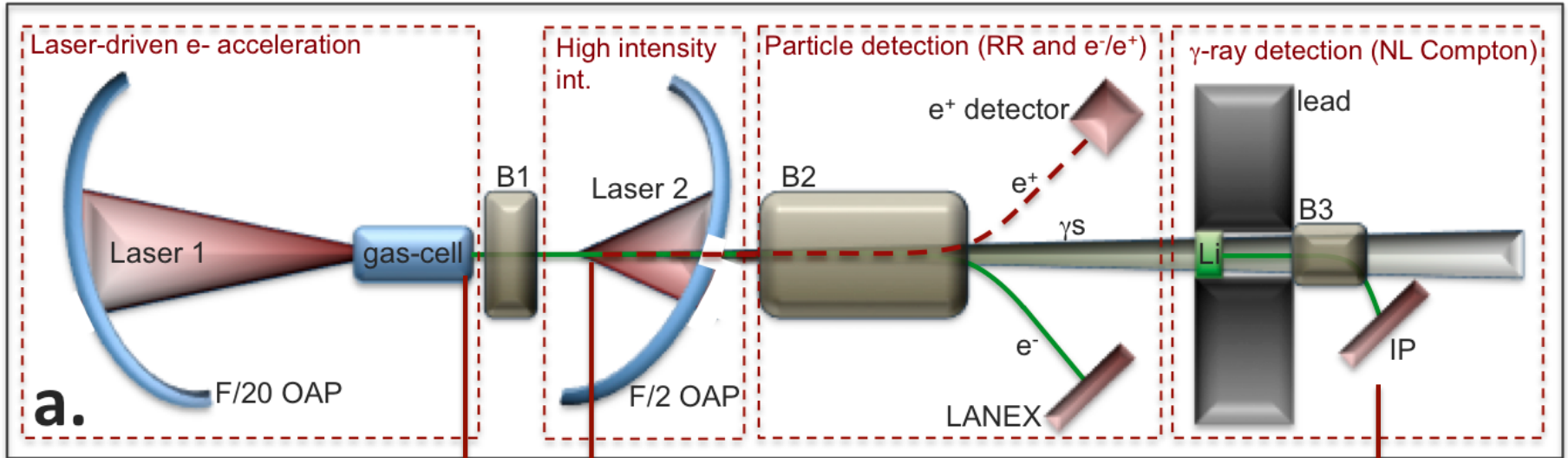


Photon-photon cross section



- Up to 10^{-2} positrons per incoming electron at 4 GeV
- Possibility to access direct electro-production

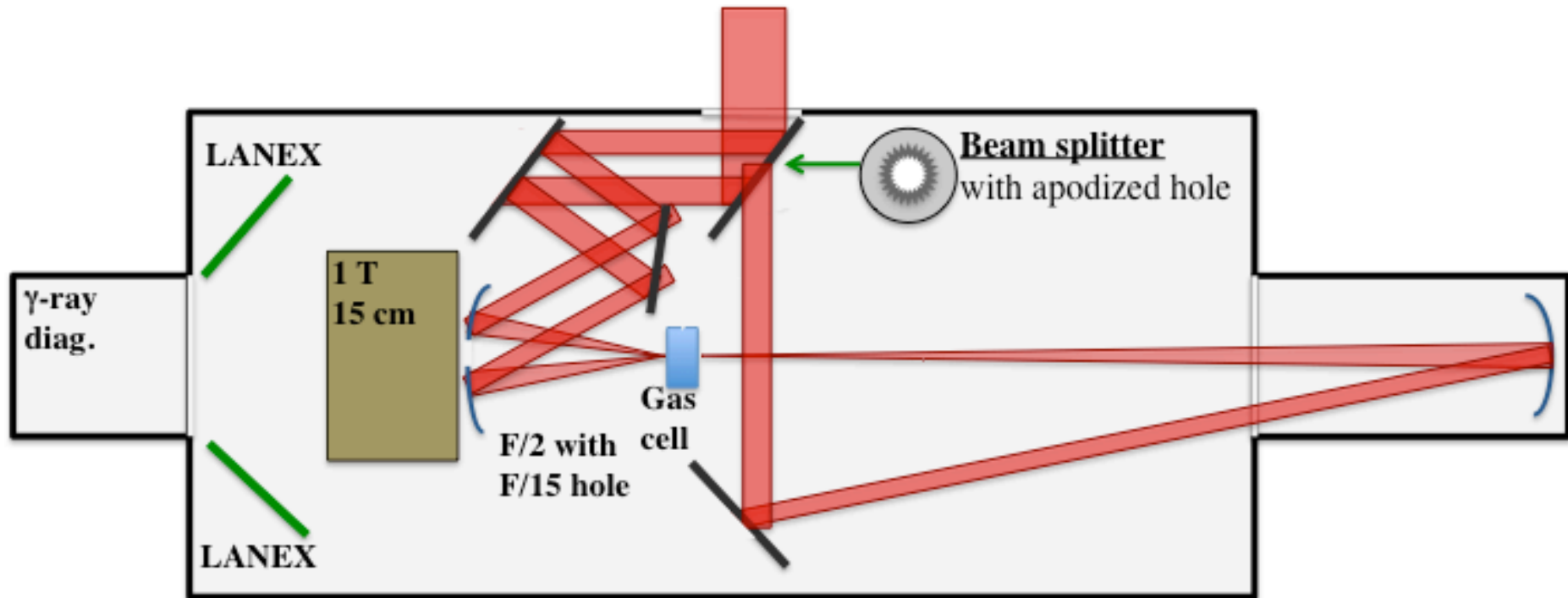
General Setup



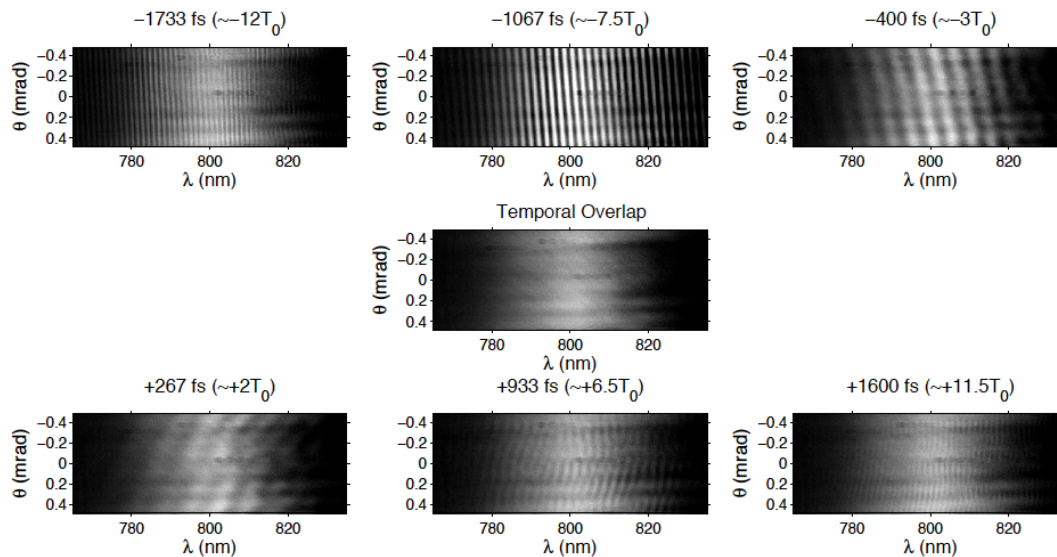
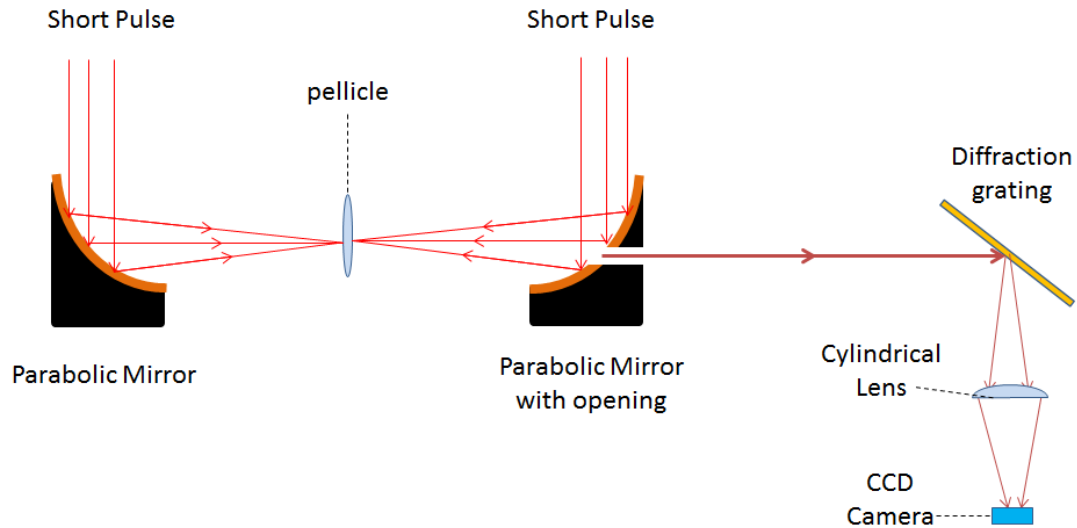
G. Sarri et al. Phys. Rev. Lett. 113, 224801 (2014)

General Setup

▪ SINGLE BEAM POSSIBLE SETUP



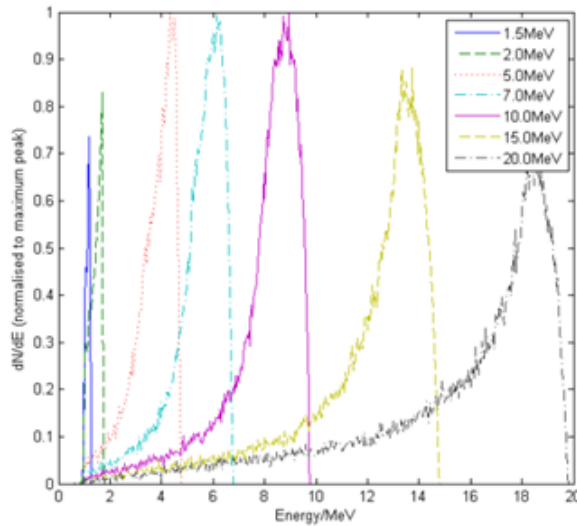
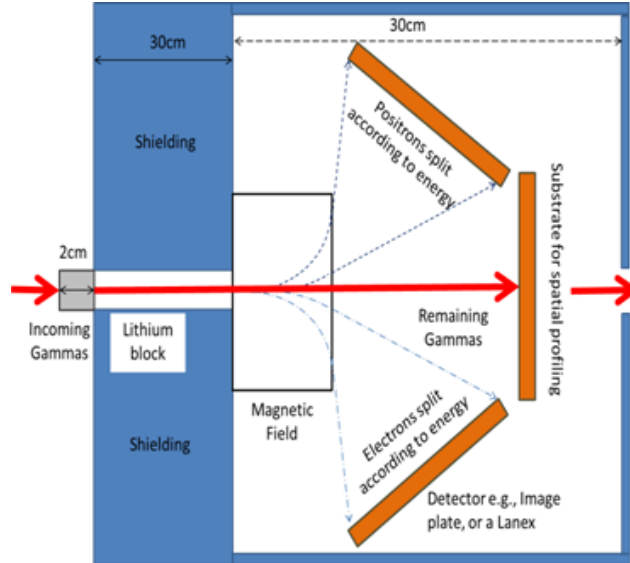
Temporal Synchronisation



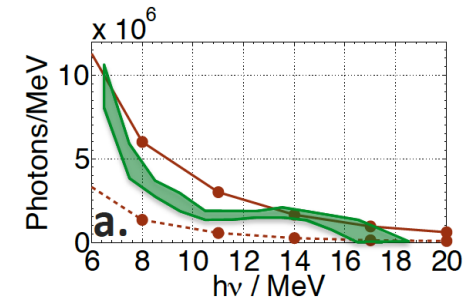
- Both laser beams are sent through the hole of the short focal length parabola onto a diffraction grating.
- Spectral dispersion along one axis provides temporal information, whilst the undispersed axis provides information on the spatial overlap.
- Temporal resolution of the order of 50fs
- Spatial resolution of the order of a few microns

D.J. Corvan et al. Opt. Exp. 24,1852 (2016)

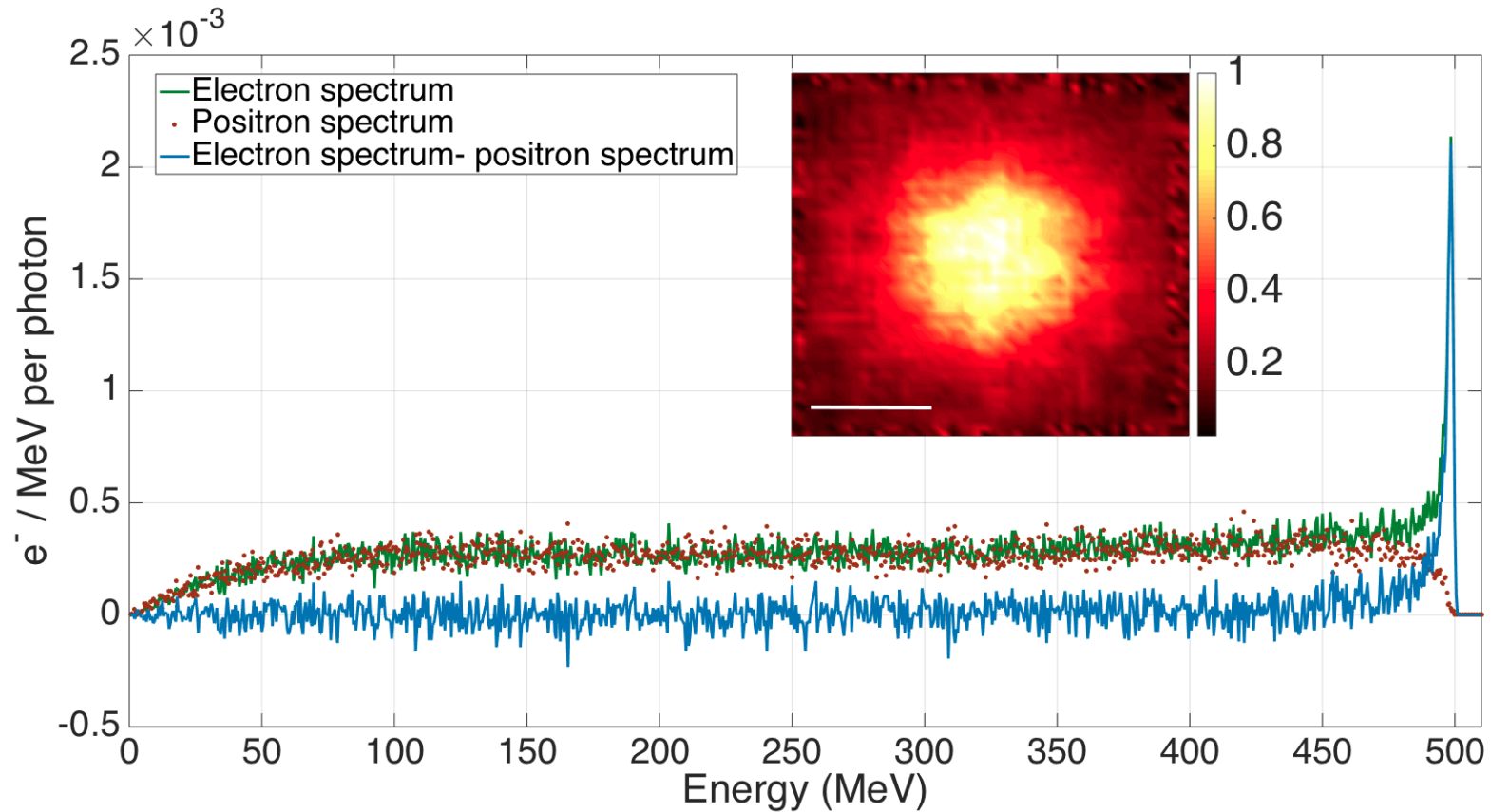
γ -ray spectroscopy: Li



- In its basic configuration, γ -rays undergo inverse Compton scattering, generating a population of electrons with similar spectral distribution on axis.
- Electrons are then spectrally resolved and the resulting spectrum is deconvolved to give the initial γ -ray spectrum
- Energy resolution ~ 1 MeV
- Energy window 3 – 30 MeV
- Conversion into electrons $\sim 3\%$

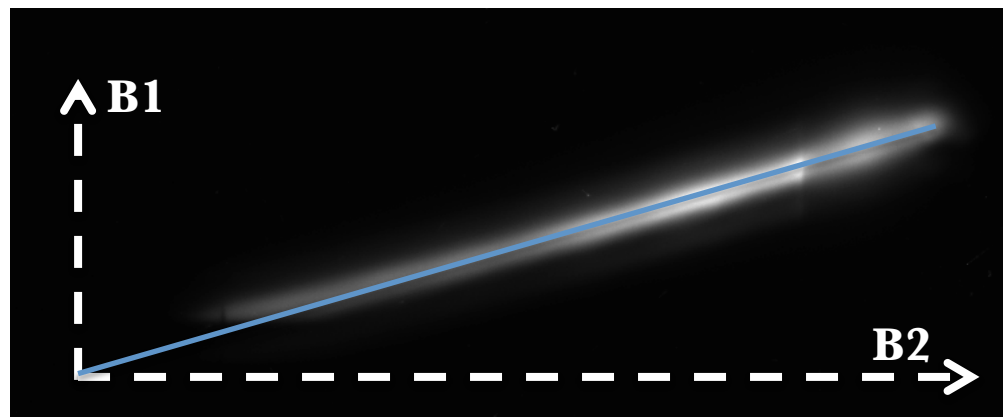
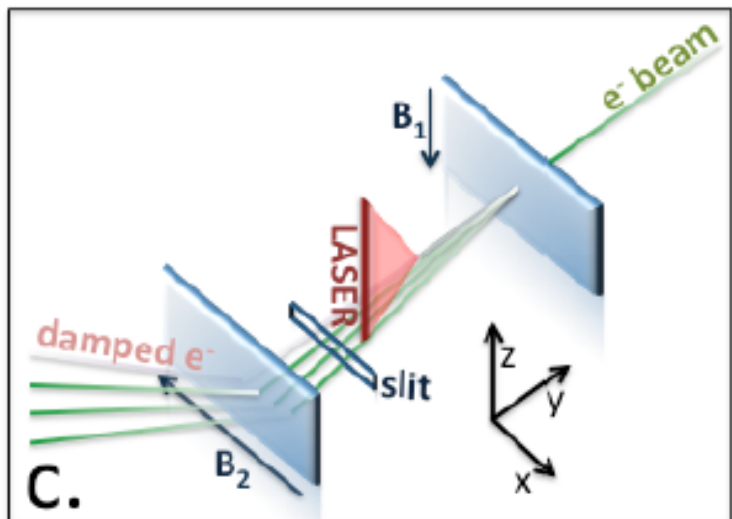
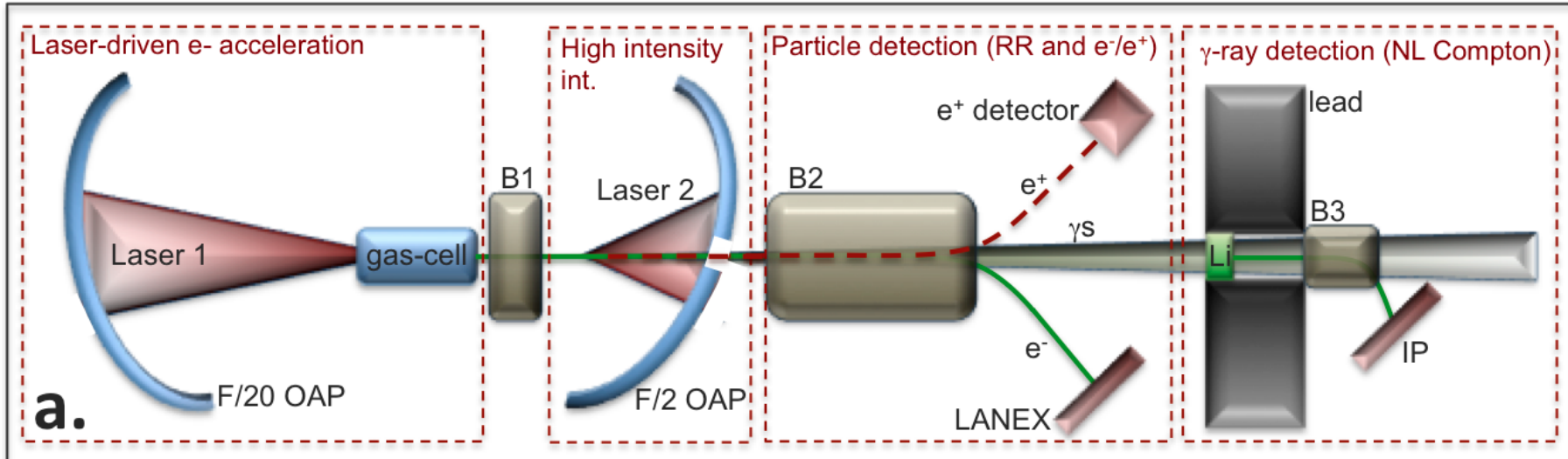


D. J. Corvan et al. Rev. Sci. Instrum. 85, 065119 (2014)

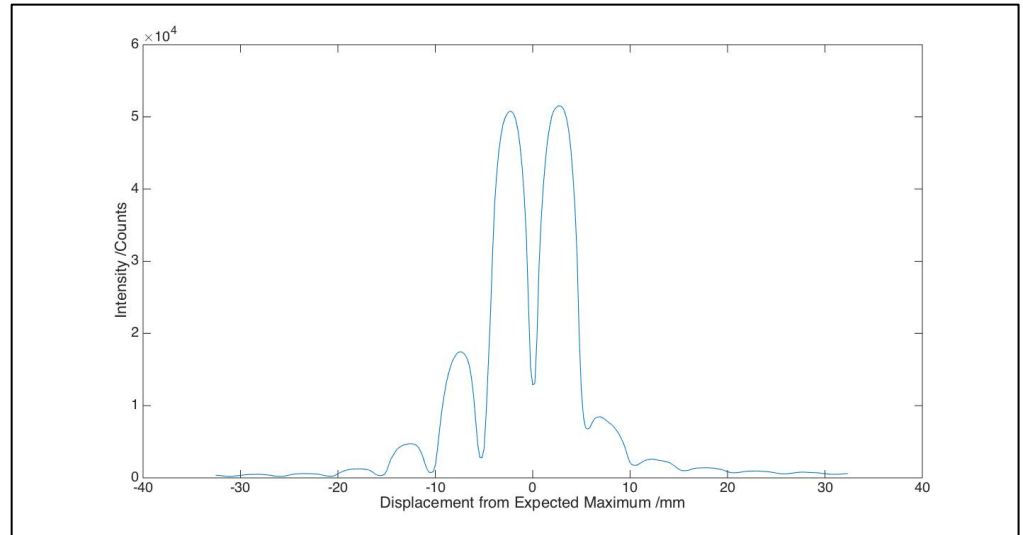
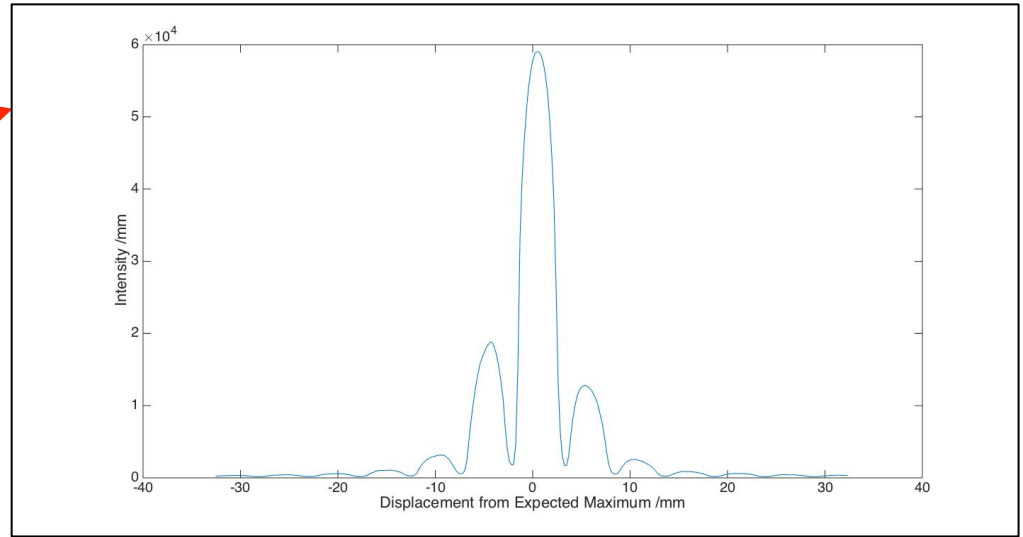
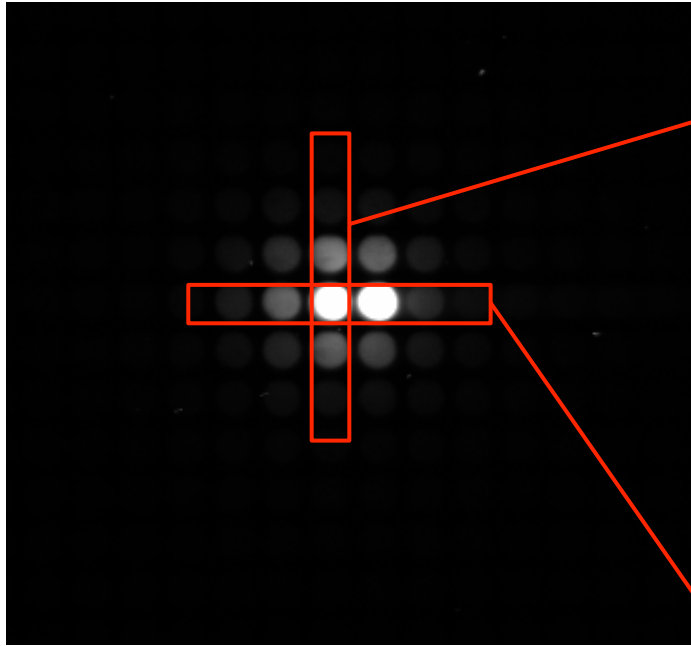


- 20 bar, 50cm long H₂ pipe
- MeV resolution up to the GeV range
- Conversion efficiency: 0.25%
- First prototype built and tested

Double magnetic spectrometer



Pixelated scintillators



- Possibility of directly measuring the laser a_0

Conclusions

- **Unprecedented** possibility of studying physical phenomena that have been thus far out of reach of experiments
- **Radiation Reaction** experiments that can access highly non-linear and quantum regimes. First systematic and direct studies of non-linear quantum electrodynamics
- **Prolific Pair Production** in both Trident and Two-step process, with the possibility of studying the transition between the two. Possibility of directly measuring the photon-photon cross section and its non-linear corrections
- **Next generation of γ -ray beams** with unprecedented short duration, high flux, and high brilliance even in the initial stage of Apollon. Of paramount importance for progressing our knowledge in diverse fields of science.

Thank you for
your attention

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Quantum Radiation Reaction

⇒ Radiation Reaction is one of the oldest and most fundamental problems in electromagnetism: How do we correctly model the electron dynamics if we include radiative losses?

0. Classical Lorentz force

$$m \frac{du^u}{ds} = eF^{uv} u_v$$

✗ No energy loss

1. LAD Equation

$$m \frac{du^u}{ds} = eF^{uv} u_v + \frac{2}{3} e^2 \left(\frac{d^2 u^u}{ds^2} + \frac{du^v}{ds} \frac{du^v}{ds} u^u \right)$$

- ✓ Damping force (radiation reaction term)
- ✗ Classical renormalisation (point-like electron)
- ✗ Runaway solutions! (diverging acceleration even without external field)

2. LL Equation

$$m \frac{du^u}{ds} = eF^{uv} u_v + \frac{2}{3} e^2 \left(\frac{e}{m} (\partial_\alpha F^{uv}) u^\alpha u_v - \frac{e^2}{m^2} F^{uv} F_{\alpha v} u^\alpha + \frac{e^2}{m^2} (F^{\alpha v} u_v)(F_{\alpha \lambda} u^\lambda) u^u \right)$$

- ✓ No runaway solutions
- ✓ Valid in classical relativity