

# Photon-Photon Scattering with Intense PW or multi-PW lasers.

**Luis Roso**

**Director**

Apollon Users Meeting  
February 12th, 2016

**CLPU** CENTRO DE  
LASERES  
PULSADOS  
ULTRACORTOS  
ULTRAIINTENSOS

**Centro de Laseres Pulsados**

<http://www.clpu.es/>

**Apollon First Users' Meeting**

Under the auspices of Institut  
Lasers & Plasmas

Feb 11-12, 2016

Venue : Synchrotron Soleil  
Université Paris Saclay

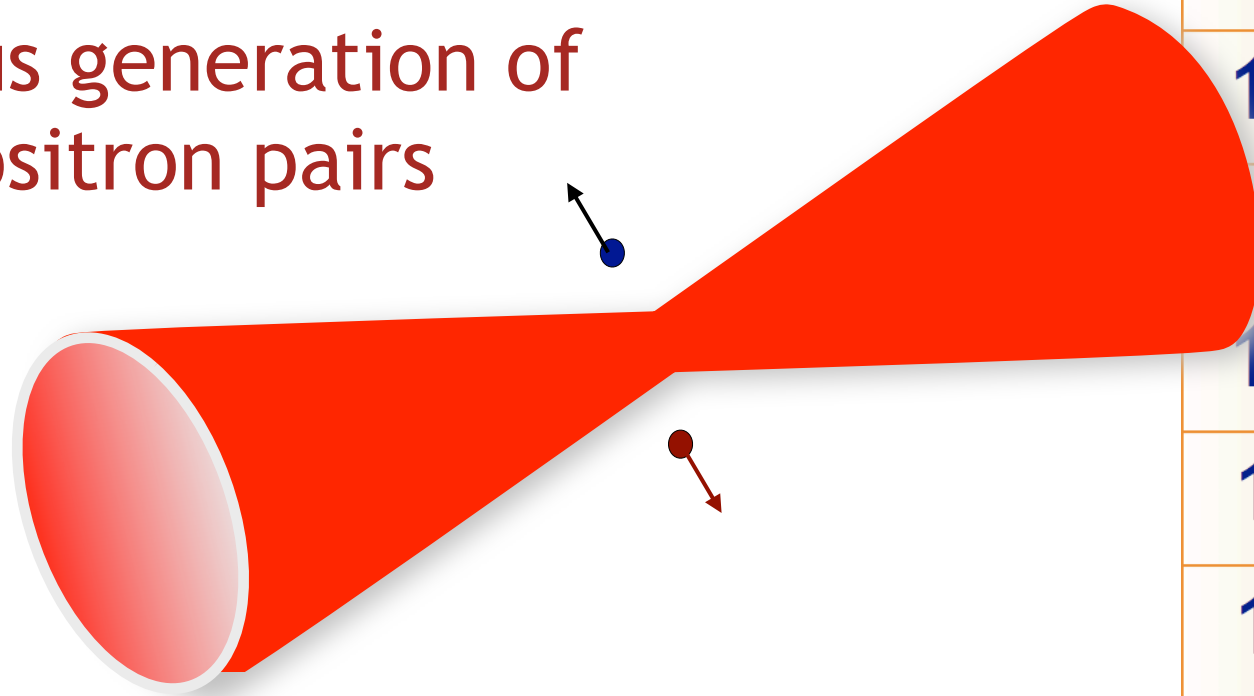


# The Schwinger limit

# Vacuum annihilation: Schwinger



Spontaneous generation of electron-positron pairs



Non-linear QED

Beyond  $10^{29} \text{ W/cm}^2$  vacuum seems to be unstable

$10^{21} \text{ W/cm}^2$

$10^{22} \text{ W/cm}^2$

$10^{23} \text{ W/cm}^2$

$10^{24} \text{ W/cm}^2$

$10^{25} \text{ W/cm}^2$

$10^{26} \text{ W/cm}^2$

$10^{27} \text{ W/cm}^2$

$10^{28} \text{ W/cm}^2$

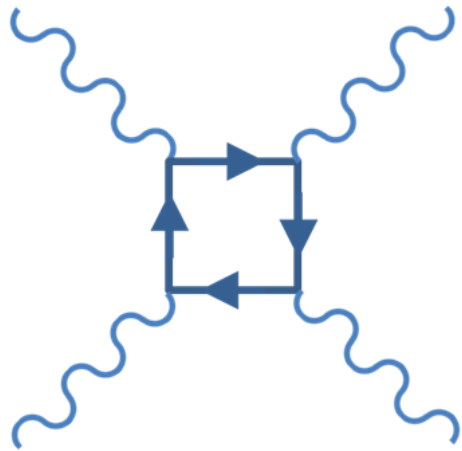
$10^{29} \text{ W/cm}^2$

$10^{30} \text{ W/cm}^2$

$10^{31} \text{ W/cm}^2$

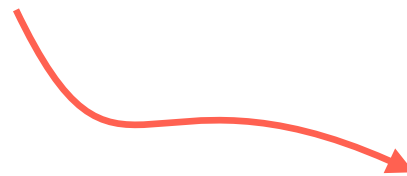


# Quantum vacuum



Intensity where  
virtual pairs start  
to be relevant ???

Generation of real pairs



$10^{18}$ W/cm <sup>2</sup>
$10^{19}$ W/cm <sup>2</sup>
$10^{20}$ W/cm <sup>2</sup>
$10^{21}$ W/cm <sup>2</sup>
$10^{22}$ W/cm <sup>2</sup>
$10^{23}$ W/cm <sup>2</sup>
$10^{24}$ W/cm <sup>2</sup>
$10^{25}$ W/cm <sup>2</sup>
$10^{26}$ W/cm <sup>2</sup>
$10^{27}$ W/cm <sup>2</sup>
$10^{28}$ W/cm <sup>2</sup>



$10^{14}$  W/cm<sup>2</sup>

$10^{15}$  W/cm<sup>2</sup>

Atomic unit of intensity  $10^{16}$  W/cm<sup>2</sup>

$10^{17}$  W/cm<sup>2</sup>

$10^{18}$  W/cm<sup>2</sup>

$10^{19}$  W/cm<sup>2</sup>

$10^{20}$  W/cm<sup>2</sup>

$10^{21}$  W/cm<sup>2</sup>

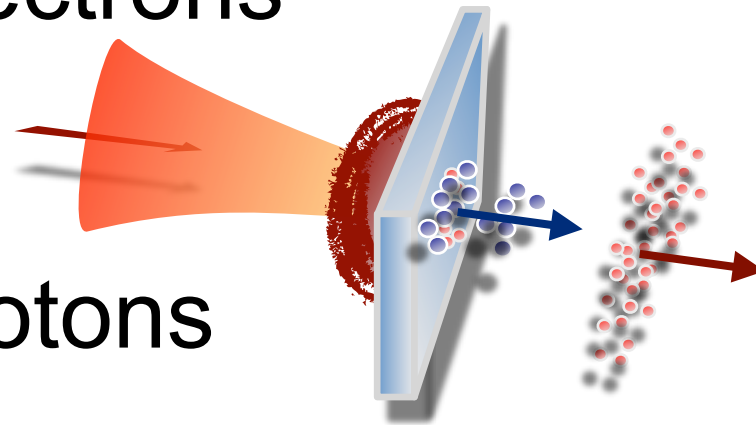
$10^{22}$  W/cm<sup>2</sup>

$10^{23}$  W/cm<sup>2</sup>

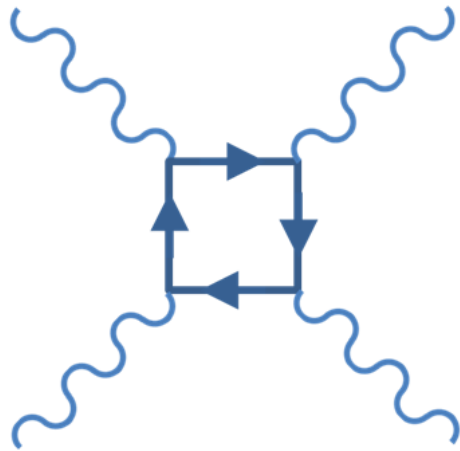
$10^{24}$  W/cm<sup>2</sup>

Relativistic electrons

Relativistic protons

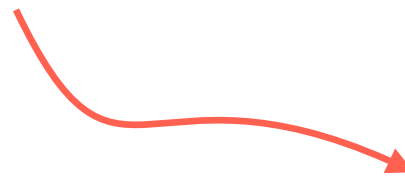


World record



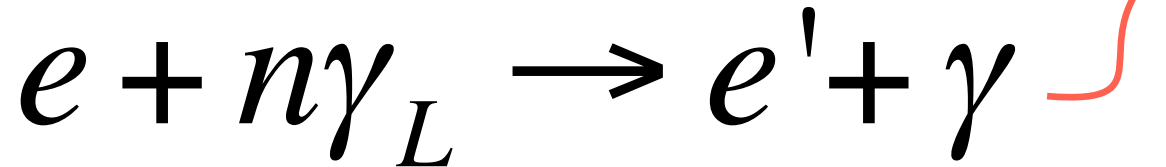
Intensity where  
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to be relevant ???

Generation of real pairs

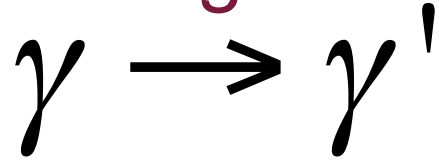


$10^{18}$ W/cm <sup>2</sup>
$10^{19}$ W/cm <sup>2</sup>
$10^{20}$ W/cm <sup>2</sup>
$10^{21}$ W/cm <sup>2</sup>
$10^{22}$ W/cm <sup>2</sup>
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$10^{26}$ W/cm <sup>2</sup>
$10^{27}$ W/cm <sup>2</sup>
$10^{28}$ W/cm <sup>2</sup>

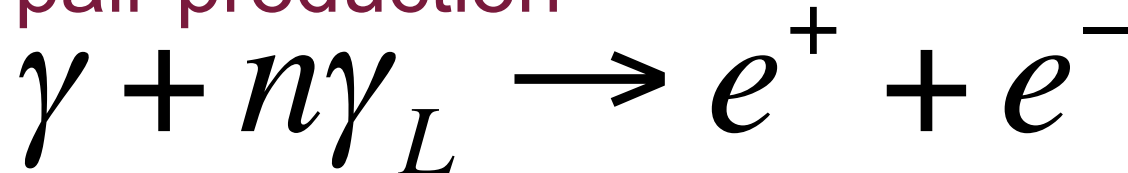
Non-linear Compton-Thomson



Vacuum birefringence



Stimulated pair production



Vacuum pair production



$10^{20}$  W/cm<sup>2</sup>

$10^{21}$  W/cm<sup>2</sup>

$10^{22}$  W/cm<sup>2</sup>

$10^{23}$  W/cm<sup>2</sup>

$10^{24}$  W/cm<sup>2</sup>

$10^{25}$  W/cm<sup>2</sup>

$10^{26}$  W/cm<sup>2</sup>

$10^{27}$  W/cm<sup>2</sup>

$10^{28}$  W/cm<sup>2</sup>

$10^{29}$  W/cm<sup>2</sup>





What QED  
signature can  
be observed  
first?

What  
happens  
here ?

$10^{21} \text{ W/cm}^2$

$10^{22} \text{ W/cm}^2$

$10^{23} \text{ W/cm}^2$

$10^{24} \text{ W/cm}^2$

$10^{25} \text{ W/cm}^2$

$10^{26} \text{ W/cm}^2$

$10^{27} \text{ W/cm}^2$

$10^{28} \text{ W/cm}^2$

Schwinger limit: Generation of real pairs

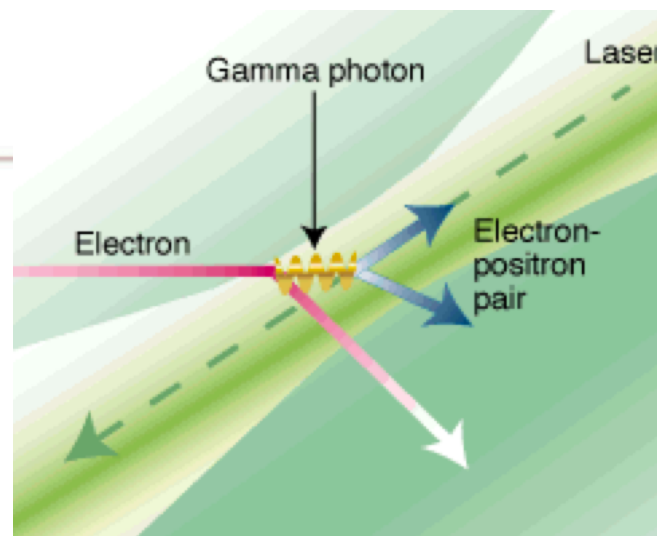
$10^{29} \text{ W/cm}^2$

$10^{30} \text{ W/cm}^2$

$10^{31} \text{ W/cm}^2$



What can be  
observed  
first?



$10^{21}$  W/cm<sup>2</sup>

$10^{22}$  W/cm<sup>2</sup>

$10^{23}$  W/cm<sup>2</sup>

$10^{24}$  W/cm<sup>2</sup>

$10^{25}$  W/cm<sup>2</sup>

$10^{26}$  W/cm<sup>2</sup>

## SLAC-144 experiment, 1998

Studies of nonlinear QED in collisions of 46.6 GeV electrons with intense laser pulses

C. Bamber,<sup>\*</sup> S.J. Boege,<sup>†</sup> T. Koffas, T. Kotseroglou,<sup>‡</sup> A.C. Melissinos, D.D. Meyerhofer,<sup>§</sup> D.A. Reis and W. Ragg<sup>\*\*</sup>

*Department of Physics and Astronomy, University of Rochester, Rochester, New York 14627*

C. Bula,<sup>††</sup> K.T. McDonald and E.J. Prebys

*Joseph Henry Laboratories, Princeton University, Princeton, New Jersey 08544*

D.L. Burke, R.C. Field, G. Horton-Smith,<sup>‡‡</sup> J.E. Spencer and D. Walz

*Stanford Linear Accelerator Center, Stanford University, Stanford, California 94309*

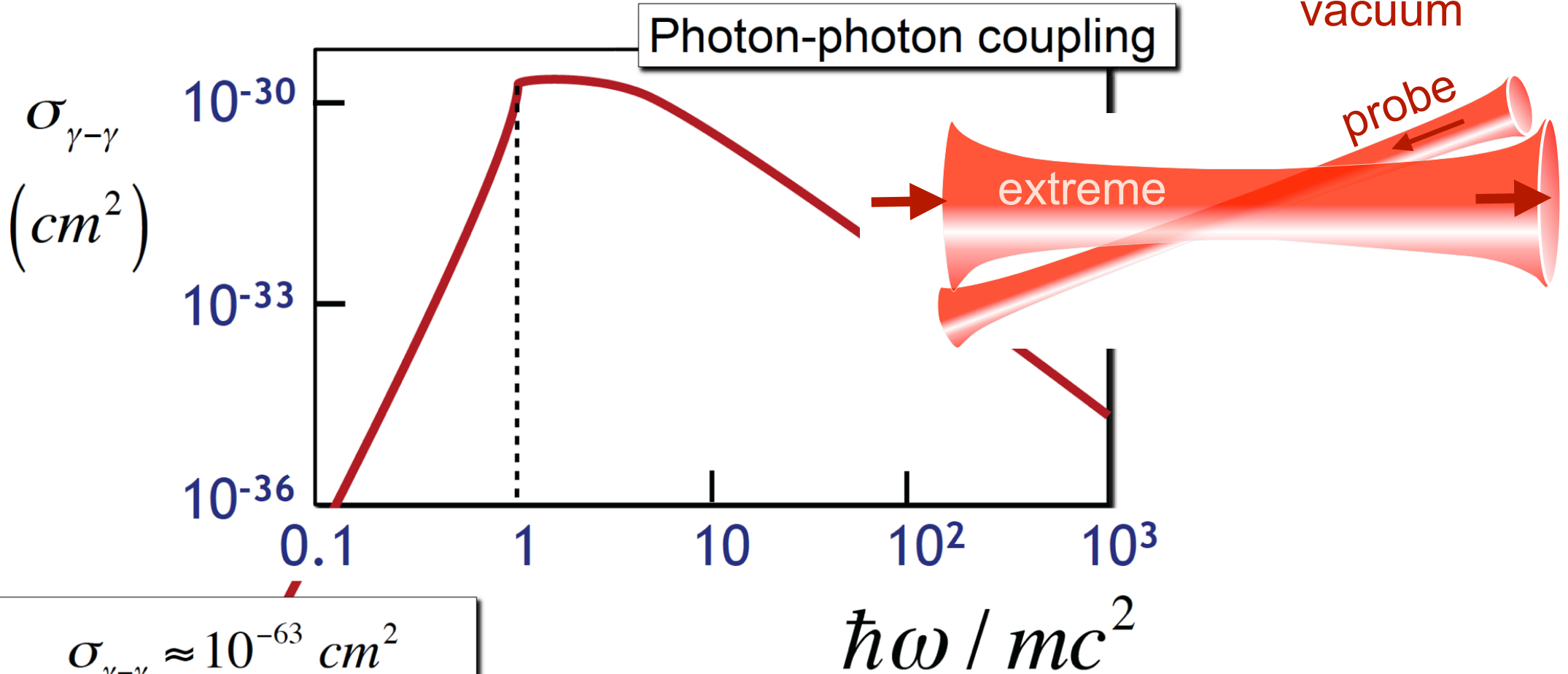
S.C. Berridge, W.M. Bugg, K. Shmakov<sup>§§</sup> and A.W. Weidemann

*Department of Physics and Astronomy, University of Tennessee, Knoxville, Tennessee 37996*



# Photon-photon coupling

birrefringent vacuum



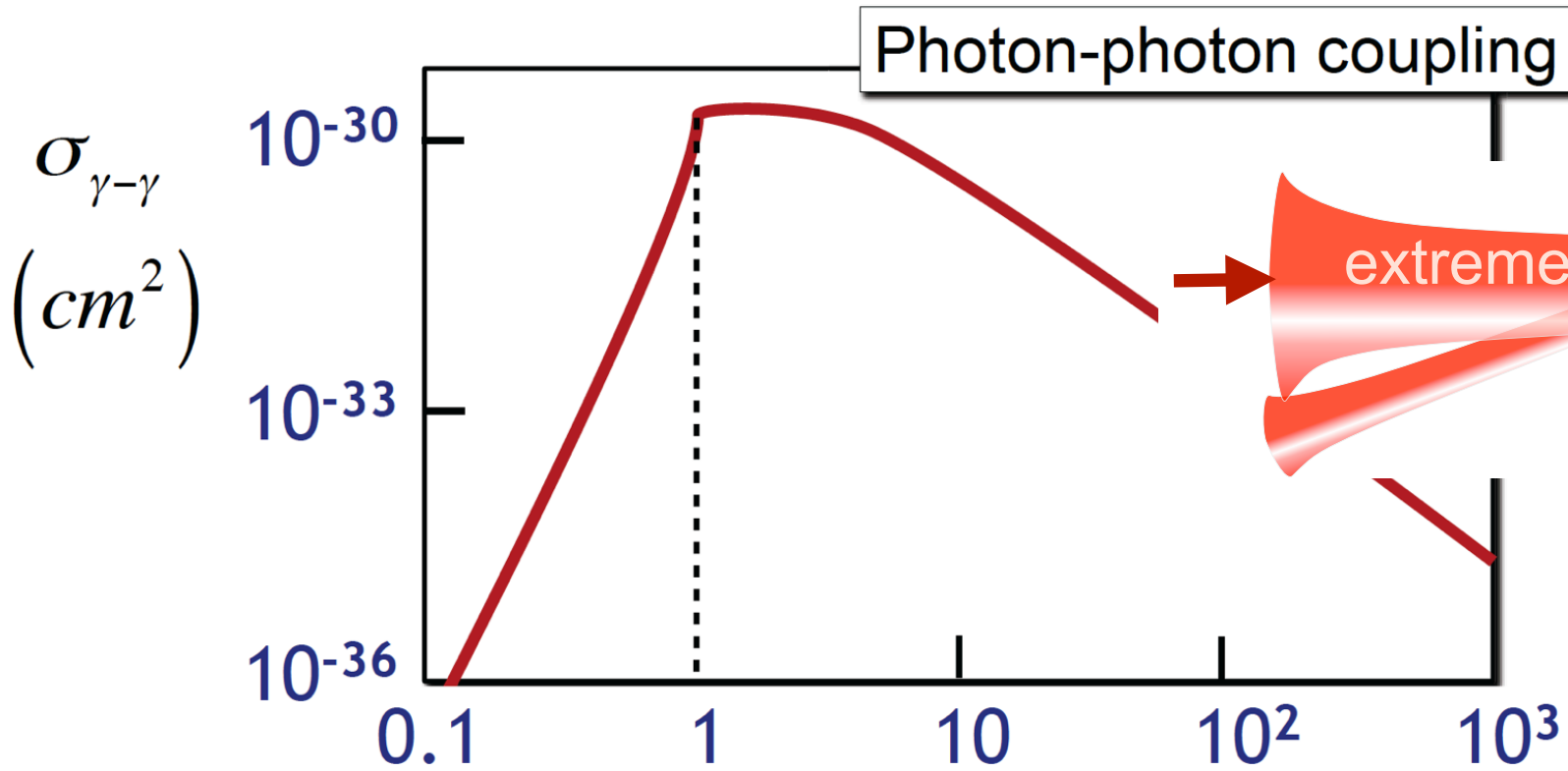
$$\sigma_{\gamma-\gamma} \approx 10^{-63} \text{ cm}^2$$

at optical wavelengths



# Photon-photon coupling

birrefringent vacuum



STAR WARS

$$\sigma_{\gamma-\gamma} \approx 10^{-63} \text{ cm}^2$$

at optical wavelengths



# Lagrangian density, linear

for  $h\nu \ll m_e c^2$

$$\mathcal{L}_0 = \frac{\epsilon_0}{2} \left( \mathbf{E}^2 - c^2 \mathbf{B}^2 \right)$$

For optical photons

# Lagrangian QED non-linear



for  $h\nu \ll m_e c^2$

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

being

$$\mathcal{L}_0 = \frac{\epsilon_0}{2} (\mathbf{E}^2 - c^2 \mathbf{B}^2)$$

$$\mathcal{G} = \epsilon_0 c (\mathbf{E} \cdot \mathbf{B})$$

$\mathcal{L}_0^2$  and  $\mathcal{G}^2$  are the only two Lorentz-covariant terms that can be formed with the electromagnetic fields at the lowest order above  $\mathcal{L}_0$

# Effective lagrangian QED



for  $h\nu \ll m_e c^2$

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

being

$$\mathcal{L}_0 = \frac{\epsilon_0}{2} (\mathbf{E}^2 - c^2 \mathbf{B}^2)$$

$$\mathcal{G} = \epsilon_0 c (\mathbf{E} \cdot \mathbf{B})$$

In QED (Euler-Heisenberg)

$$\xi_L^{QED} = \xi_T^{QED} \equiv \xi = \frac{8\alpha^2 \hbar^3}{45 m_e^4 c^5} = 6.7 \times 10^{-30} \frac{m^3}{J}$$

# Effective lagrangian QED



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$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

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W. Heisenberg and H. Euler, Folgerungen aus der Diracschen Theorie des Positrons Z. Phys. 98, 714 (1936).





# Effective Lagrangian BI

for  $h\nu \ll m_e c^2$

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

In the Born-Infeld model

$$\xi_T^{BI} = 4\xi_L^{BI} / 7$$

Proposed to remove the divergence of the electron's self-energy in classical electrodynamics by introducing an upper bound of the electric field at the origin

M. Born and L. Infeld, Foundations of the New Field Theory, Proc. Roy. Soc. Lond. 144 (1934) 425.



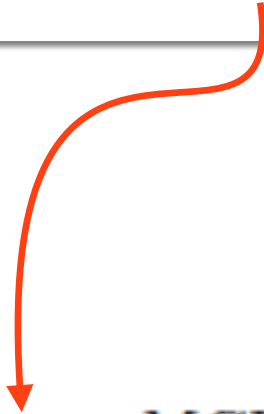



# Other effective Lagrangians

for  $h\nu \ll m_e c^2$

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

Minicharged particles  
new terms in the  
Lagrangian


$$\Delta \xi_L^{\text{MCP}}$$


$$\Delta \xi_T^{\text{MCP}}$$



# Other effective Lagrangians

Minicharged particles

Spin 1/2 fermion

$$\Delta\xi_L^{\text{MCP}} = \Delta\xi_T^{\text{MCP}} = \left(\frac{\epsilon m_e}{m_\epsilon}\right)^4 \xi$$

Spin 0 boson

$$\Delta\xi_L^{\text{MCP0}} = \frac{7}{16} \left(\frac{\epsilon m_e}{m_\epsilon}\right)^4 \xi$$

$$\Delta\xi_T^{\text{MCP0}} = \frac{1}{28} \left(\frac{\epsilon m_e}{m_\epsilon}\right)^4 \xi$$

Spin 1 boson

$$\Delta\xi_L^{\text{MCP1}} = \frac{261}{16} \left(\frac{\epsilon m_e}{m_\epsilon}\right)^4 \xi$$

$$\Delta\xi_T^{\text{MCP1}} = \frac{243}{28} \left(\frac{\epsilon m_e}{m_\epsilon}\right)^4 \xi$$

# Axions ... if they exist

Axions are hypothetical elementary particle postulated by the Peccei-Quinn theory in 1977 to resolve the strong CP problem in QCD.

no electric charge,  
very small mass,  $10^{-6}$  to  $1 \text{ eV}/c^2$

Many world collaborations to look for axions

CAST CERN Axion Solar Telescope

PVLAS Polarizzazione del Vuoto con LASer, Padova

ADMX Axion Dark Matter eXperiment, USA

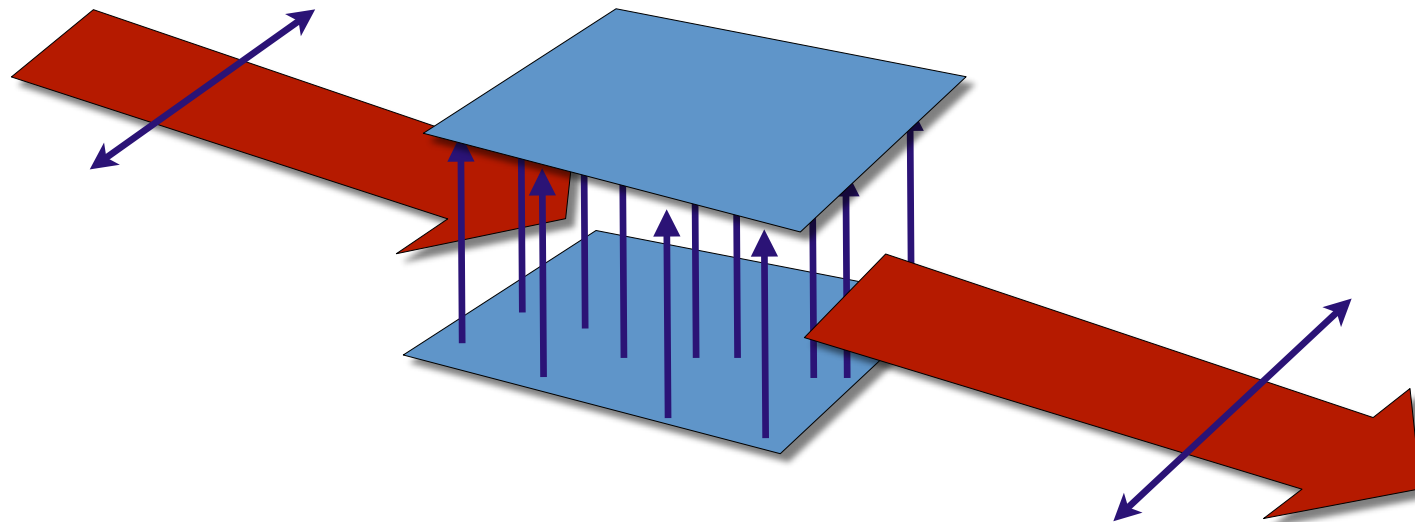
many others ...



# Axions

no electric charge,  
very small mass,  $10^{-6}$  to  $1 \text{ eV}/c^2$

Coupling to magnetic fields ...



small rotation of polarization plane

# Other effective Lagrangians



Axions

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

scalar boson

$$\Delta \xi_T = \frac{2\hbar^3 g_P^2}{7c m_\Phi^2}$$

$$\Delta \xi_L = 0$$

pseudoscalar boson

$$\Delta \xi_T = 0$$

$$\Delta \xi_L = \frac{\hbar^3 g_S^2}{2c m_\Phi^2}$$

R D Peccei, and H R Quinn

CP Conservation in the Presence of Pseudoparticles

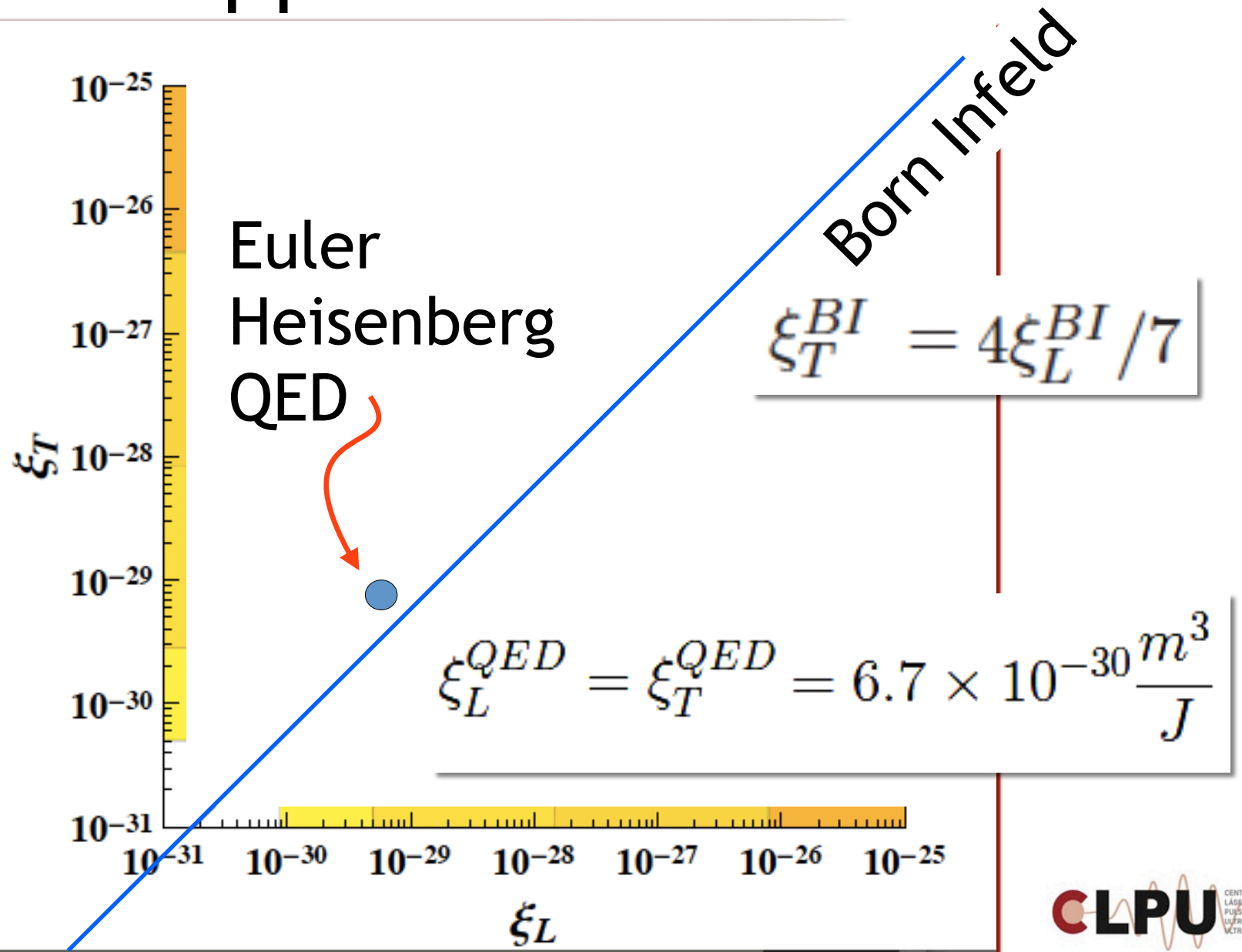
Phys Rev Lett, 1977

# What happens ???



Too-many  
Lagrangians

Two  
parameter  
space



# Too many Lagrangians

of the same kind

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

How to select the correct one ???

With a laser ..

$$\Delta\phi_L = 4\xi_L I k\tau$$

$$\Delta\phi_T = 7\xi_T I k\tau$$

D Tommasini, A Ferrando, H Michinel et M Seco  
Precision tests of QED and non-standard models by  
searching photon-photon scattering in vacuum with  
high power lasers, JHEP, 2009





# Why lasers ...

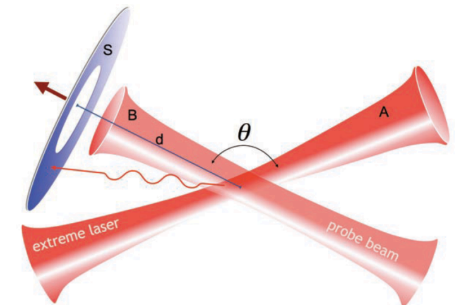
$$\Delta\phi_L = 4\xi_L I k\tau$$

$$\Delta\phi_T = 7\xi_T I k\tau$$

Photon-photon coupling results in a phase shift of the probe

$$N_D^{\mathcal{N}} = \frac{8f\mathcal{N}}{\pi\hbar c} \frac{E_A^2 E_B w_0^2}{\lambda_B w_A^4 w_B^2} \left( e^{-\frac{2r_0^2}{w_D^2}} - e^{-\frac{2R^2}{w_D^2}} \right) (a\xi)_{L,T}^2$$

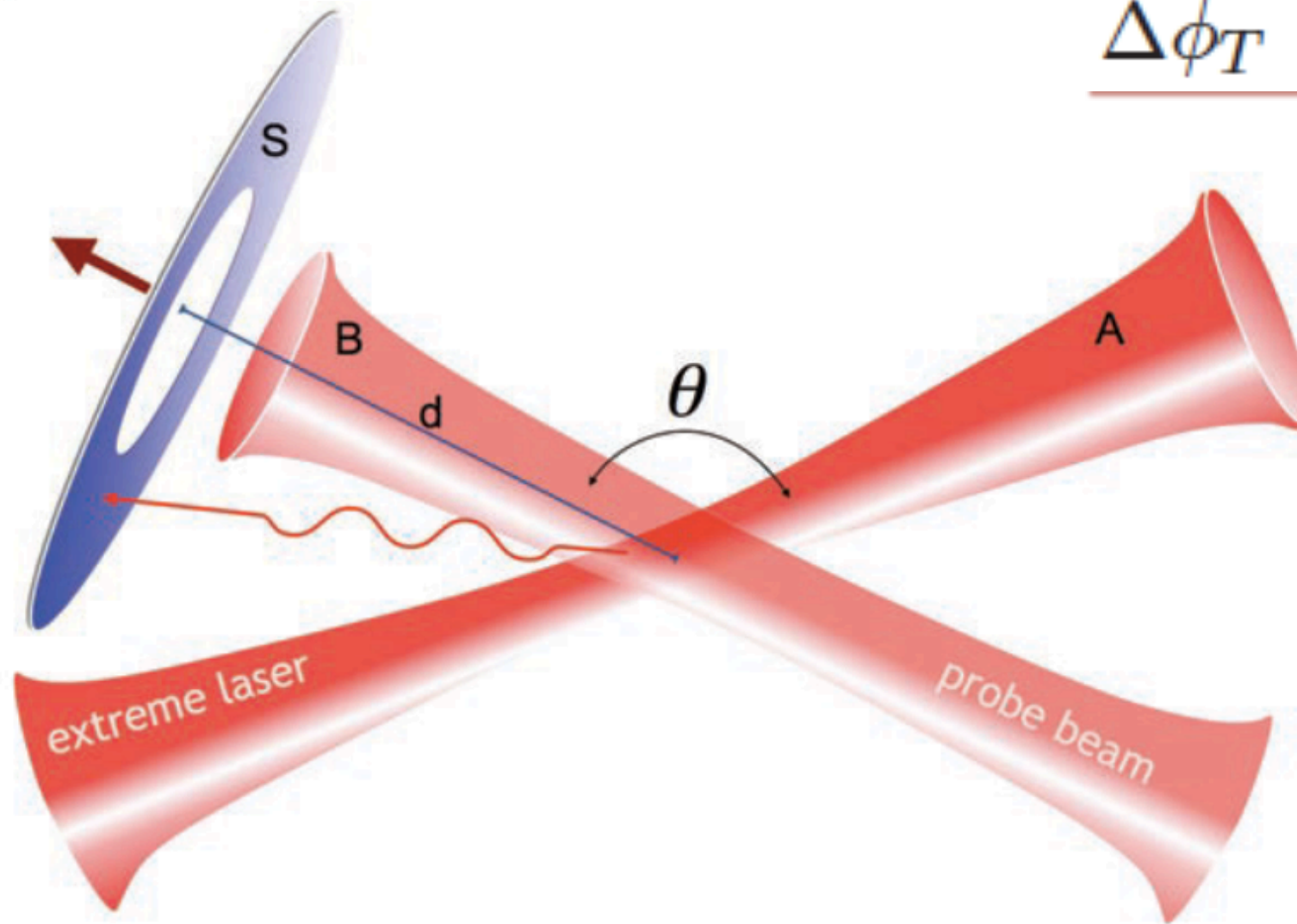
Relation between the number of scattered photons and QED coupling constants



With laser fields ...

$$\Delta\phi_L = 4\xi_L I k\tau$$

$$\Delta\phi_T = 7\xi_T I k\tau$$





# Too many Lagrangians

of the same kind

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

Laser-Lagrangian  
connection

$$\Delta\phi_L = 4\xi_L I k\tau$$

$$\Delta\phi_T = 7\xi_T I k\tau$$

D Tommasini, A Ferrando, H Michinel et M Seco  
Precision tests of QED and non-standard models by  
searching photon-photon scattering in vacuum with  
high power lasers, JHEP, 2009



# Too many Lagrangians

of the same kind

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T \mathcal{G}^2$$

How to select the correct one ???

With a laser ..

Intensity  $I$

Wavenumber  $k$

Pulse duration  $\tau$

$$\Delta\phi_L = 4\xi_L I k \tau$$

$$\Delta\phi_T = 7\xi_T I k \tau$$

D Tommasini, A Ferrando, H Michinel et M Seco

Precision tests of QED and non-standard models by searching photon-photon scattering in vacuum with high power lasers, JHEP, 2009

# Birrefringent vacuum



$$\Delta\phi_L = 4\xi_L I k\tau$$

$$\Delta\phi_T = 7\xi_T I k\tau$$

extreme laser



probe laser



With laser fields ...

D Tommasini,  
A Ferrando,  
H Michinel,  
M Seco  
D Novoa

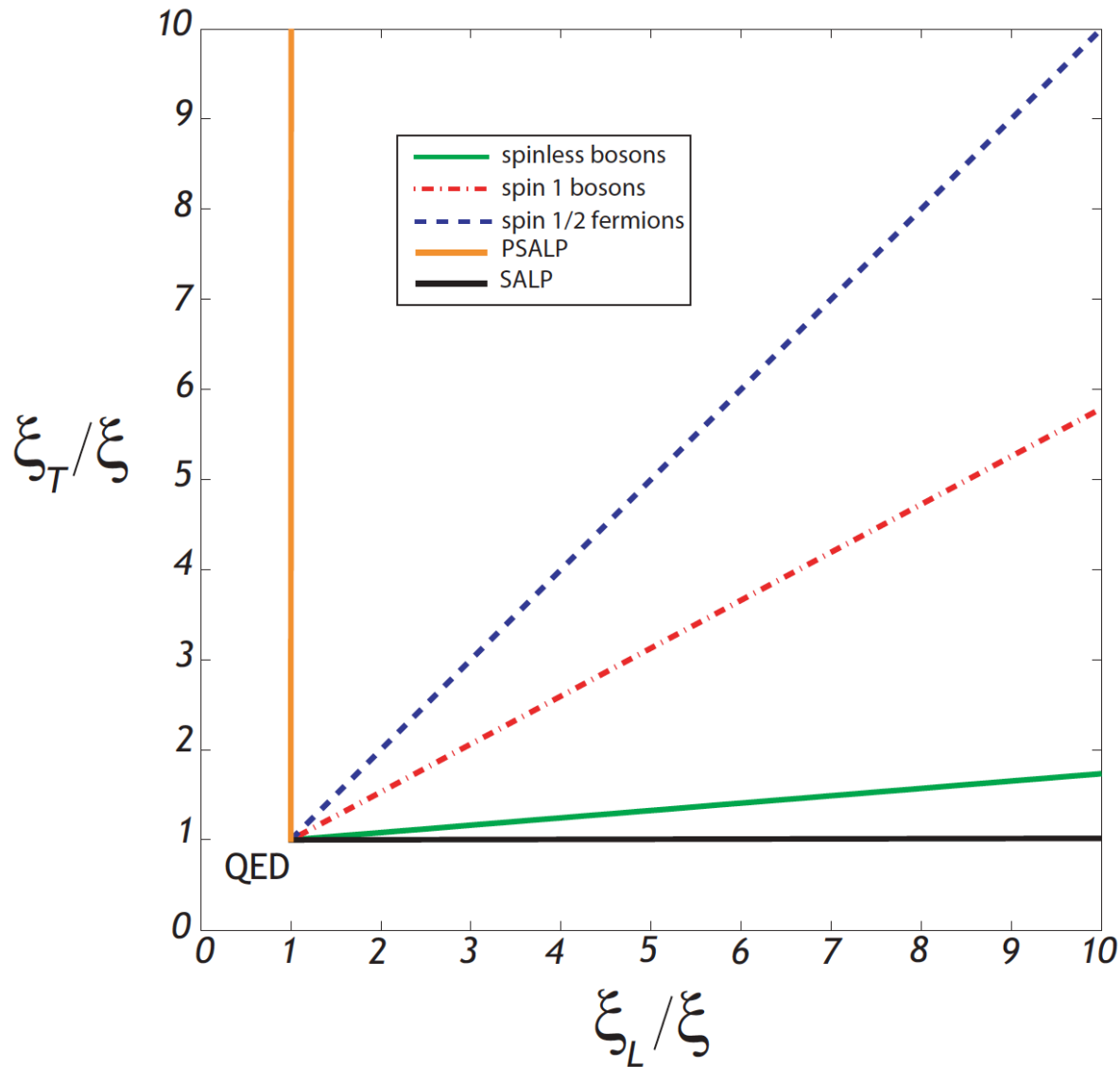
Extreme laser 12  $\mu\text{m}$  waist

59  $\mu\text{m}$  waist Probe laser

$$I_A = 4.6 \times 10^{20} \text{ W/cm}^2 \text{ and } I_B = 3.7 \times 10^{18} \text{ W/cm}^2$$

# With laser fields

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{i}{4} \xi_T \mathcal{G}^2$$

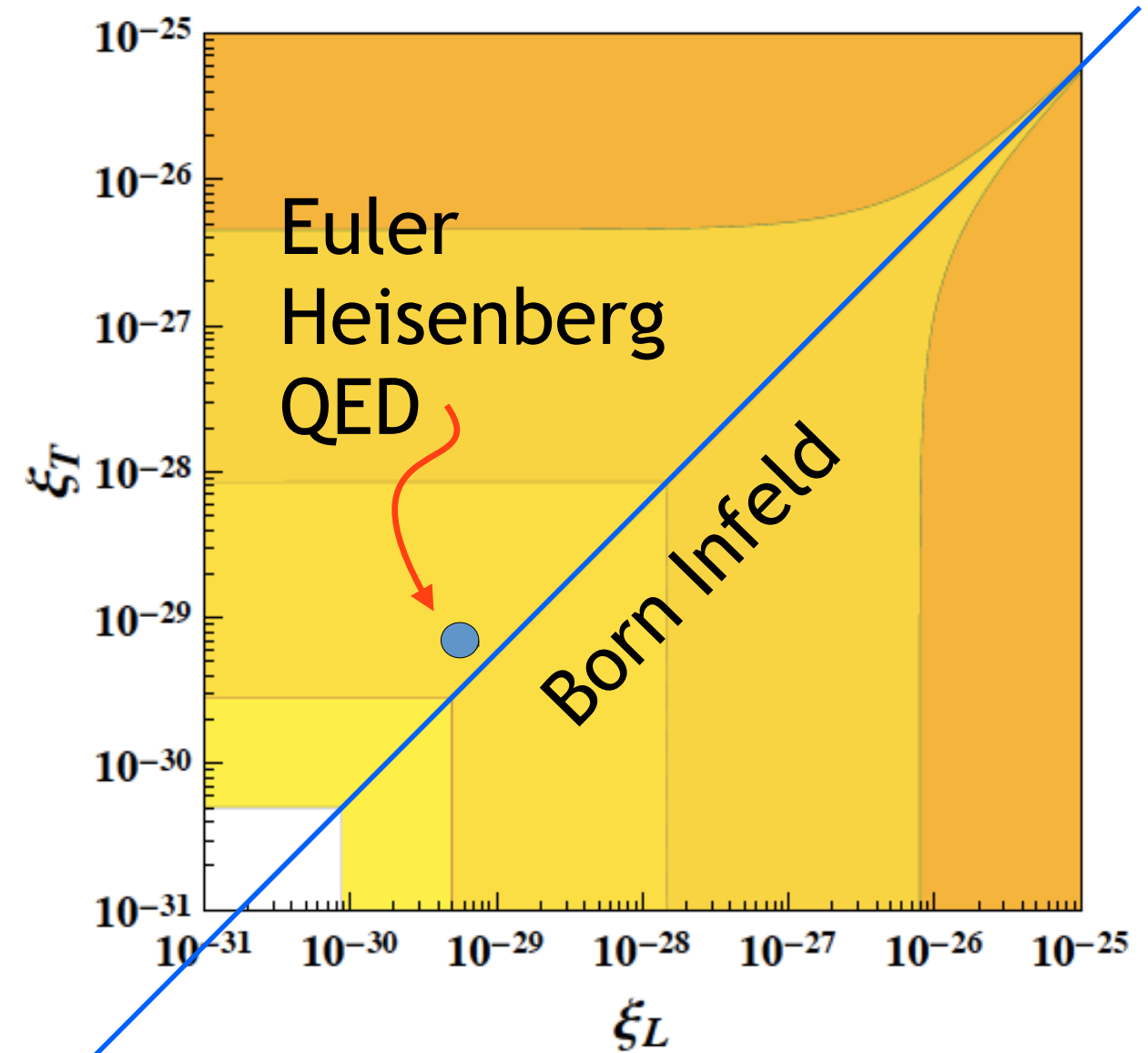




# Effective lagrangian QED

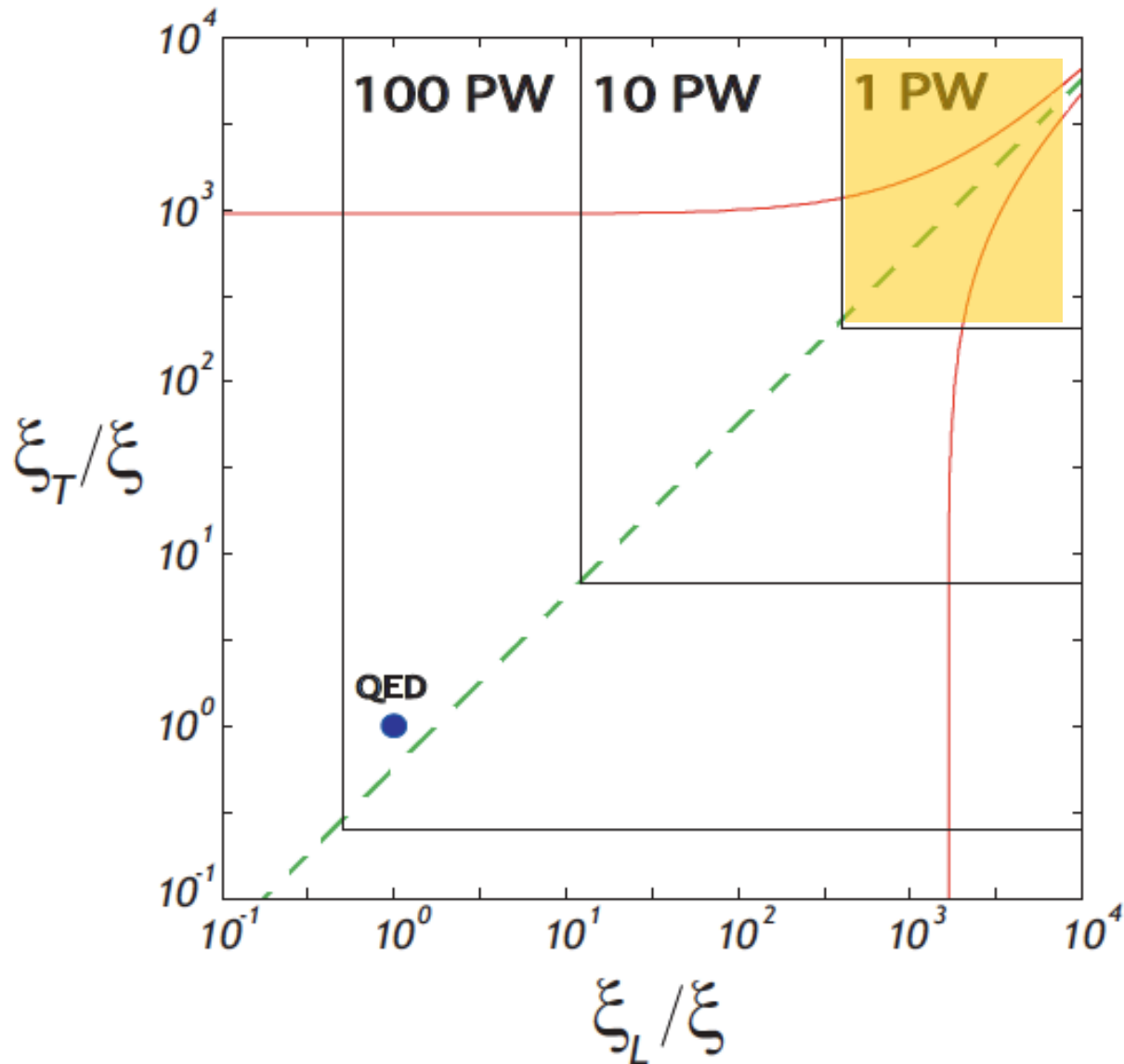
Do axions exist?

I don't know  
but,  
if they exist we  
can see them  
with this kind of  
laser experiment

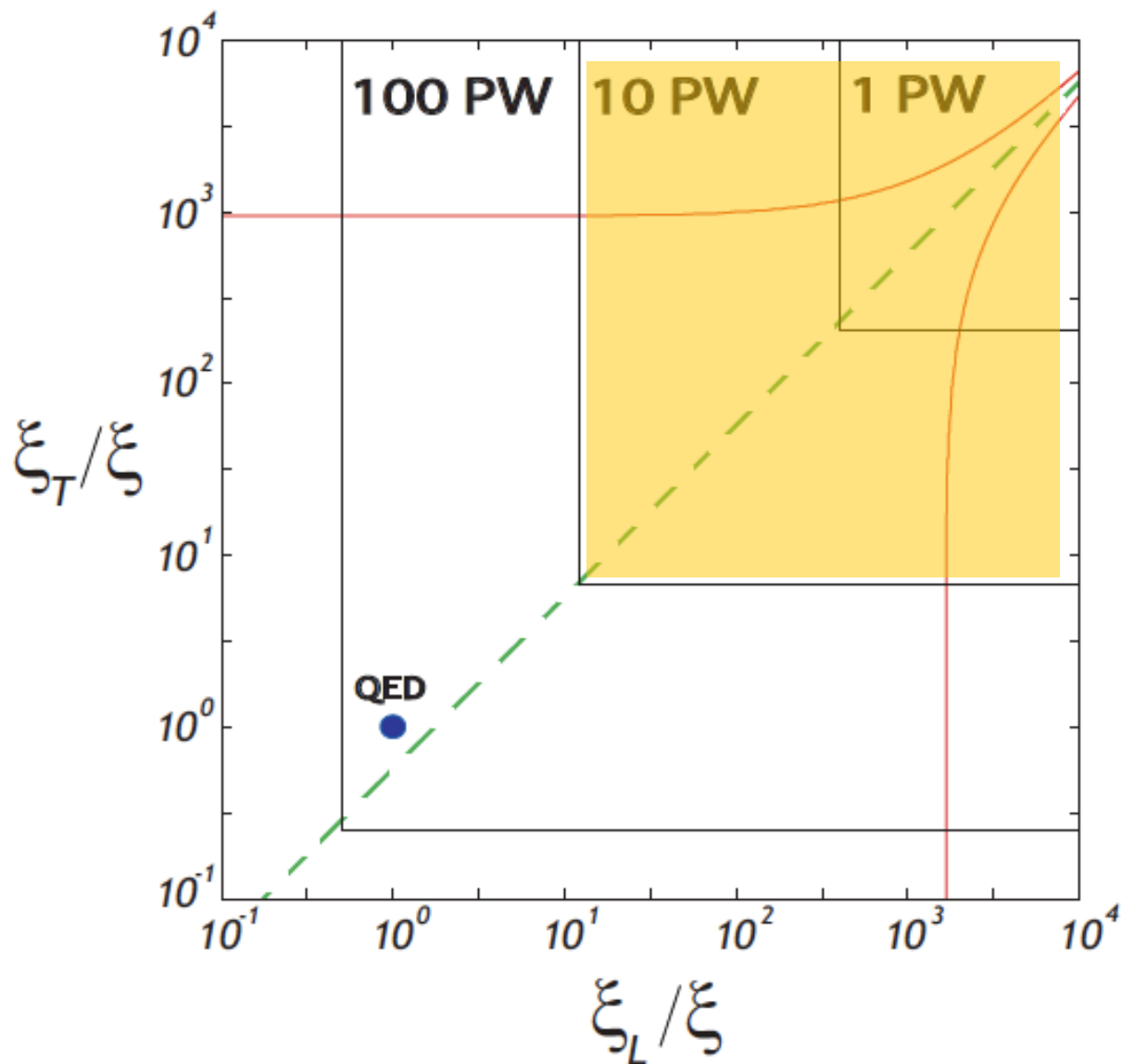




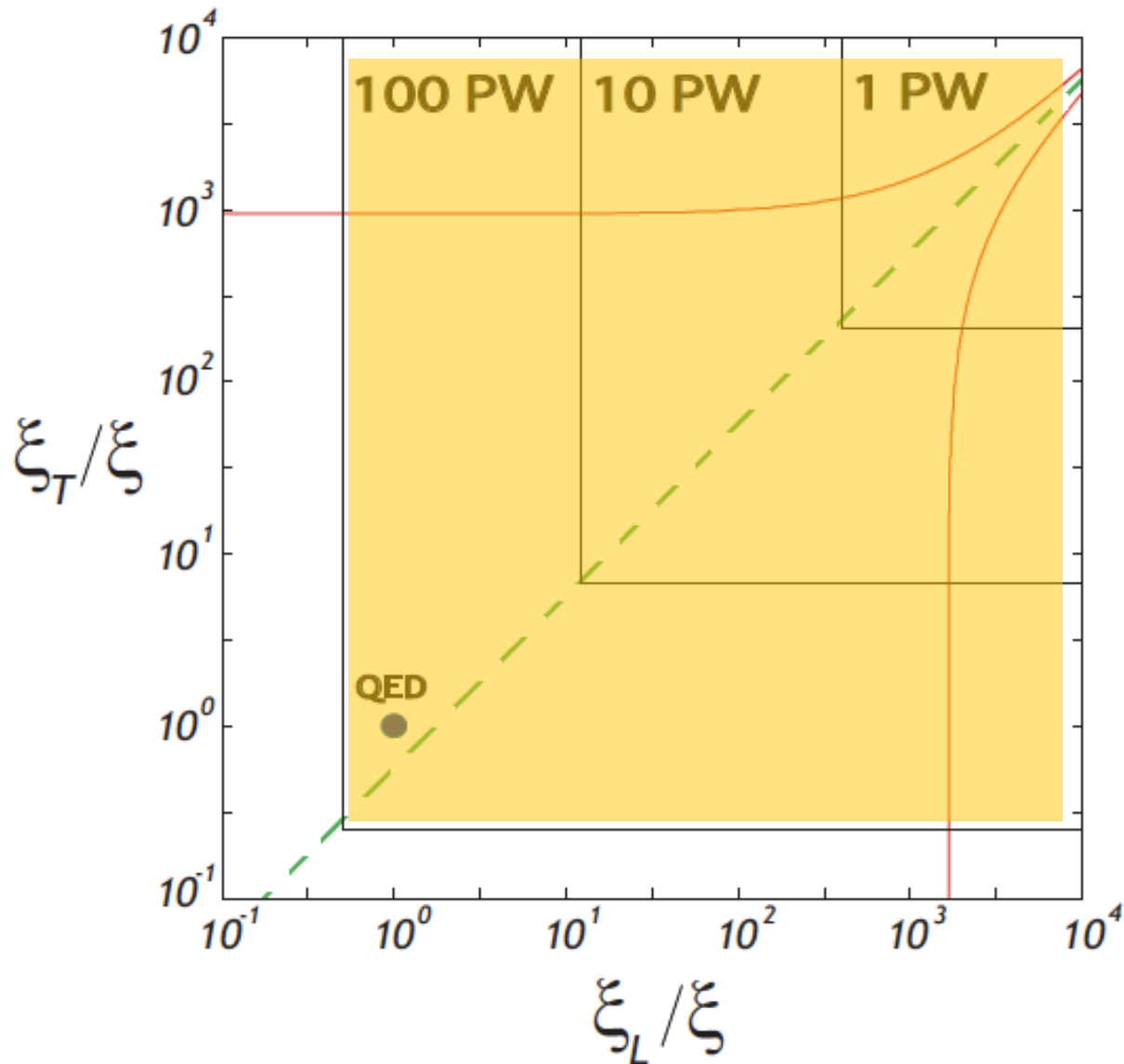
With laser fields . . .  $\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{i}{4} \xi_T \mathcal{G}^2$



With laser fields . . .  $\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{i}{4} \xi_T \mathcal{G}^2$



With laser fields . . .  $\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{i}{4} \xi_T \mathcal{G}^2$





# The VEGA laser

Titanium:sapphire technology

robust and well under control

VEGA	peak power	energy	duration	repetition rate	operation
VEGA 1	20 TW	600 mJ	30 fs	10 / sec	2007
VEGA 2	200 TW	6 J	30 fs	10 / sec	2013
VEGA 3	1 PW	30 J	30 fs	1 /sec	2016





# With laser fields ...



Facility	$P$ (PW)	$\tau$ (fs)	$\lambda$ (nm)	$\xi_T^{\text{lim}} / \xi$	$\xi^{\text{lim}} / \xi^{\text{PVLAS}}$
VEGA	1+0.2	30	800	230	0.1
APOLLON	10	15	800	23	0.01
ELI 10	10	30	800	8	0.003
PETAL	7	500	1053	14	0.006
100 PW	100	30	800	0.24	0.0001

QED 1

$$\mathcal{L} = \mathcal{L}_0 + \xi_L \mathcal{L}_0^2 + \frac{7}{4} \xi_T g^2$$



# Summary

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**Background:** Quantum Electrodynamics predicts that at very high fields vacuum can become unstable and virtual pairs of particles may begin to have measurable effects. **However there is a big controversy on what effect is going to appear first.**

**Objective:** To propose a cutting edge but feasible experiment, using world's most intense lasers, to determine those effects and to decide which one is the right QED lagrangian, and how the relativistic Dirac positron sea can play a role without need to generate real pairs.

**Impact:** This can change our understanding of a big physics facility. A laser, being the most intense field in the world, can trigger unique interactions with such virtual particles. It can also be relevant for neutrino studies and for our understanding of the dark matter of the universe.



# More details

Springer Series in Chemical Physics 106

Kaoru Yamanouchi  
Gerhard G. Paulus  
Deepak Mathur *Editors*

## Progress in Ultrafast Intense Laser Science X

 Springer

PUILS 

JILS 

## Chapter 9.- Quantum Vacuum Polarization Searches with High Power Lasers Below the Pair Production Regime

**Daniele Tommasini,**  
David Novoa and  
Luis Roso







# The collaboration

---

Orense team

Daniele Tommasini

Angel Paredes

Humberto Michinel



CLPU-Salamanca team

Giancarlo Gatti

Cruz Méndez

...

Canada

Robert Fedosejevs,  
Univ Alberta,



US,

Wendell Hill, Univ Maryland

Ronnie Shepherd, Livermore

Roman Sobolewski, Rochester

# Thanks!!!

Luis Roso, Director  
Centro de Láseres Pulsados,  
CLPU, Salamanca