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CERN

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In the previous episode...

- We've discussed data formats
- And especially data structures
- We've talked about organising our data





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- We've discussed data formats
- And especially data structures
- We've talked about organising our data

Today

Let's see how we store our data !





Outline



- Existing devices
- Hierarchical storage



Distributed storage

- Data distribution
- Data federation

3 Parallelizing files' storage

- Striping
- Introduction to Map/Reduce









Storage devices

Storage devices

- Existing devices
- Hierarchical storage

2 Distributed storage

- 3 Parallelizing files' storage
- 4 Conclusion



A variety of storage devices

Main differences

- Capacities from 1 GB to 10 TB per unit
- Prices from 1 to 300 for the same capacity
- Very different reliability
- Very different speeds



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Typical numbers in 2015

	Capacity per unit	Latency	\$/TB	Speed	reliability
RAM	16 GB	5 ns	7500 \$	$10\mathrm{GBs^{-1}}$	volatile
SSD	500 GB	10 µ s	400 \$	$550\mathrm{MBs^{-1}}$	poor
HD	6 TB	3 ms	30 \$	$150\mathrm{MBs^{-1}}$	average
Tape	10 T B	100 s	25 \$	$300\mathrm{MBs^{-1}}$	good

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A variety of storage devices





Reliability in real world (CERN)

For disks

- $\bullet\,$ probability of losing a disk per year : few %, up to $10\%\,$
 - with 60K disks, it's around 10 per day
 - and all files are lost
- \bullet one unrecoverable bit error in $10^{14}\ {\rm bits}\ {\rm read}/{\rm written}$
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For tapes

- probability of losing a tape per year : 10^{-4}
 - and you recover most of the data on it
 - net result is 10^{-7} file loss per year
- one unrecoverable bit error in 10^{19} bits read/written
 - for 1GB files, that's one file corrupted per 1G files written



Practical Mass Storage - Real Big Data

when you count in 100s of PetaBytes...

The constraints

- disks or tapes are the only possible solutions
- disks are unreliable at that scale, and need redundancy
 - we'll see that extensively
- tapes are cheaper long term storage by factor 2-2.5
- tape latency imposes data access on disk



Specificities of tape storage

Key points

- 300MB/s in sequential read/write
 - 3x the speed of a disk
 - who said tape is slow ?
- latency/seek time in the order of minutes !
 - due to mount time and robot arm moving
 - due to positionning
- $\bullet\,$ storage is cheap, I/O is not
 - 20\$/TB for storage capacity
 - 25K\$ for each drive





Tape efficiency

Computation

$$efficiency = \frac{I/O \ time}{mount \ time + I/O \ time}$$
$$mount \ size = mount \ time * drive \ speed$$
$$efficiency = \frac{1}{1 + \frac{mount \ size}{data \ size}}$$
$$mount \ size \simeq 50 \ \text{GB}$$

zoo HSM



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Hierarchical storage

Layers

- tape as primary storage
- disks used as a cache in front of the tape system
- SSD used as cache in front of disks (or inside them)

CERN's case (2015)

- tape capacity : 200 PB
- tape usage : 120 PB
- disk raw capacity : 150 PB
- disk usable space : 80 PB



Some consequences

Disk cache management is needed

- the disk cache needs garbage collection
- different algorithm used depending on usage
 - FIFO First In First Out
 - LRU Least Recently Used

User need to adapt their usage

- data need to be prefetched from tape before access
- preferably in bulk



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Distributed storage

1 Storage devices



- Data distribution
- Data federation

3 Parallelizing files' storage

4 Conclusion



Handling large distributed storage

Standard issues

- failures are very common
- data distribution is hard to balance
- congestions are frequent



Keeping balanced data distribution

sounds initially easy

- standard load balancing
- taking benefit of the numerous clients

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Growing the cluster

- when you add disks, their are empty
- data need to be rebalanced





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Data congestions

The problem of popular files

- on writes, you can send a piece of data where you like
- on reads, you have to send the client where the data is
- but for popular pieces of data, the device holding it may become congested



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- temperature is data popularity
- "hot" data is accessed more often
- data getting hotter than given threshold get replicated
- on cool down, replicas should be dropped



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The global view

A World distributed storage

- large experiments' data are distributed around the world
- usually replicating the data at least once
- networks are now fast and you can access remote data through wide area network when local data in not available
- but you do not want to do that by default
- and you want to try the "closest" replica first



The global view

A World distributed storage

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Data geolocalization

- a global central catalog lists all available replicas
- and their "physical" location
- with information on the topology of the network to reach them





Data federation

Hiding the underlying complexity

- some interfaces that hide this world wide distribution of data
- by redirecting clients dynamically to the closest data available

Xrootd federation

• widely used by LHC experiments across the world



Typical federation in an LHC experiment










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Parallelizing files' storage

Storage devices

2 Distributed storage

3 Parallelizing files' storage

- Striping
- Introduction to Map/Reduce

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Why to parallelize storage ?

to work around limitations

- individual device speed (think disk)
 - a file is typically stored on a single device
- network cards' speed
 - 1 Gbit network still present
 - network congestion on a node reduces bandwidth per stream
- core network throughput
 - switches / routers are expensive
 - machines may have less throughput than their card(s) allow(s)
- hot data congestions
 - and the black hole it can generate
 - as slower tranfers allow to accumulate more transfers

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Parallelizing through striping

Main idea

- use several devices in parallel for a single stream
- moving the limitations up by summing performances

Basic striping : Divide and conquer for storage

- split data into chunks aka stripes on different devices
- access in parallel



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RAID

- stands to "Redundant Array of Inexpensive Disks"
- set of configurations that employ the techniques of striping, mirroring, or parity to create large reliable data stores from multiple general-purpose computer hard disk drives (Wikipedia)

Useful RAID levels

RAID 0 striping

- RAID 1 mirroring
- RAID 5 parity

RAID 6 double parity

Can be implemented in hardware or software



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RAID 0

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Useful RAID levels

RAID 0 striping

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See data preservation talk

Can be implemented in hardware or software

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RAID versus **RAIN**

RAIN

- Redundant Array of Inexpensive Nodes
- similar to RAID but across nodes

Main interest

- tackle also the network limitations
- when used for redundancy, improves reliability
- more on this in subsequent lecture

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Practical striping - the stripe size

Desired picture







Practical striping - the stripe size

Stripes too big



RAID has practically not effect



Practical striping - the stripe size

Striped too small



RAID will only kill performance by forcing disk to seek far too often

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How to choose the stripe size

size of the stripe

- must be as small as possible to let small reads benefit from parallelization
- must not be too small
 - to avoid having to deal with too much metadata
 - to avoid too much disk seeking

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A generic solution for the stripe size

Idea

- disentangle "stripe size" from "object size"
- "stripe size" is the size of one slice of data
- "object size" is the size of one block of data on disk
- several stripes are put together into one bigger object



Ceph striping

Object 0	Object 1	Object 1 Object 2	Object 3	Object 4
Unit 0	Unit 1	Unit 1 Unit 2	Unit 3	Unit 4
Unit 5	Unit 6	Unit 6 Unit 7	Unit 8	Unit 9
Unit 10	Unit 11	Unit 11 Unit 12	Unit 13	Unit 14
Object 0	Object 1	Object 1 Object 2	Object 3	Object 4
Object 5	Object 6	Object 6 Object 7	Object 8	Object 9
Unit 15	Unit 16	Unit 16 Unit 17	Unit 18	Unit 19
Unit 20	Unit 21	Unit 21 Unit 22	Unit 23	Unit 24
Unit 25	Unit 26	Unit 26 Unit 27	Unit 28	Unit 29
Object 5	Object 6	Object 6 Object 7	Object 8	Object 9



Ceph striping

Object 0	Object 1	Object 2	Object 3	Object 4	
Unit 0	Unit 1	Unit 2	Unit 3	Unit 4	Stripe Unit 4
Unit 5	Unit 6	Unit 7	Unit 8	Unit 9	
Unit 10	Unit 11	Unit 12	Unit 13	Unit 14	
Object 0	Object 1	Object 2	Object 3	Object 4	
Object 5	Object 6	Object 7	Object 8	Object 9	
Unit 15	Unit 16	Unit 17	Unit 18	Unit 19	
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Ceph striping

Object 0	Object 1	Object 2	Object 3	Object 4	
Unit 0	Unit 1	Unit 2	Unit 3	Unit 4	Stripe Unit 4
Unit 5	Unit 6	Unit 7	Unit 8	Unit 9	
Unit 10	Unit 11	Unit 12	Unit 13	Unit 14	Stripe 2
Object 0	Object 1	Object 2	Object 3	Object 4	
Object 5	Object 6	Object 7	Object 8	Object 9	
Unit 15	Unit 16	Unit 17	Unit 18	Unit 19	
Unit 20	Unit 21	Unit 22	Unit 23	Unit 24	
Unit 25	Unit 26	Unit 27	Unit 28	Unit 29	
Object 5	Object 6	Object 7	Object 8	Object 9	

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Ceph striping

Object 0					
Object 0	Object 1	Object 2	Object 3	Object 4	
Unit 0	Unit 1	Unit 2	Unit 3	Unit 4	Stripe Unit 4
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Object 0	Object 1	Object 2	Object 3	Object 4	
	4				
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Ceph striping

Object 0					
Object 0	Object 1	Object 2	Object 3	Object 4	
Unit 0	Unit 1	Unit 2	Unit 3	Unit 4	Stripe Unit 4
Unit 5	Unit 6	Unit 7	Unit 8	Unit 9	
Unit 10	Unit 11	Unit 12	Unit 13	Unit 14	Stripe 2
Object 0	Object 1	Object 2	Object 3	Object 4	
Object 5	Object 6	Object 7	Object 8	Object 9	
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Object 5	Object 6	Object 7	Object 8	Object 9	
	S. Pon	re - CERN	Disk 3		striping map

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Ceph striping

Object 0						
Object 0	Object 1	Object 2	Object 3	Object 4		
Unit 0	Unit 1	Unit 2	Unit 3	Unit 4	Stripe Unit 4	
Unit 5	Unit 6	Unit 7	Unit 8	Unit 9		
Unit 10	Unit 11	Unit 12	Unit 13	Unit 14	Stripe 2	
Object 0	Object 1	Object 2	Object 3	Object 4		
Object 5	Object 6	Object 7	Object 8	Object 9		
Unit 15	Unit 16	Unit 17	Unit 18	Unit 19		
Unit 20	Unit 21	Unit 22	Unit 23	Unit 24	Object Set 1	
Unit 25	Unit 26	Unit 27	Unit 28	Unit 29		
Object 5	Object 6	Object 7	Object 8	Object 9		
S. Ponce - CERN			Disk 3		striping m	apreduc



Practical striping - number of disks

Why to have many

- to increase parallelism
- to get better performances

Why to have few

- to limit the risk of losing files
- as losing a disk now means losing all files of all disks
- if p is the probability to lose a disk the probability to lose one in n is $p_n = np(1-p)^{n-1} \sim np$

A generic solution for the number of disks

Idea

- disentangle "nb disks" from "nb stripes"
- do not use all disks for all files
- adapt your number of disks to each file
 - more disks for high performance files
 - less disks for more safety

A generic solution for the number of disks

Idea

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- do not use all disks for all files
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Going further : Map/Reduce

What do we have with striping

- striping allows to distribute server I/O on several devices
- but client still faces the total I/O
- and CPU is not distributed

Map/Reduce Idea

- send computation to the data nodes
- "the most efficient network I/O is the one you don't do"



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Map/Reduce Introduction

Schema

Map processes data locally and returns key/value pairs Shuffle sends output to *Reduce* step based on output key Reduce merges partial results to final output



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Example : usage of protocols from logs

open protoA close open protoB close stat protoB protoA



Example : usage of protocols from logs

split logs on different nodes





Example : usage of protocols from logs

- split logs on different nodes
- Parse and extract key/value (Protocol/Date) in Map





Example : usage of protocols from logs

- split logs on different nodes
- **2** parse and extract key/value (*Protocol/Date*) in Map
- Sount accesses and output *Date/nb accesses* in Reduce
- o build graph





The case of Hadoop, HDFS

HDFS : the Hadoop Distributed FileSystem

- Distributed
 - files split in blocks spread over the cluster
 - blocks are typically 128MB
- Redundant
 - mirroring, 3 copies by default
 - erasure coding coming in 3.0
- hardware aware


















striping mapreduce





striping mapreduce









Many ways to store data

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Conclusion

- Storage devices
- 2 Distributed storage
- 3 Parallelizing files' storage



Many ways to store data

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Conclusion

Key messages of the day

- Take benefit of all existing devices to optimize your efficiency
- Do not underestimate the problem of data distribution
- Striping solves many problems. Use it but think twice
- $\bullet~\mbox{Map}/\mbox{Reduce}$ avoids a lot of I/O when results are small