

- EXO-16-005 ttbar+MET 2015
- EXO-16-010 Z(II)+MET 2015 •
- EXO-16-011 monoH(γγ) 2015 •
- EXO-16-012 monoH(bb) 2015 •
- EXO-16-037 monojet 2016 •
- EXO-16-038 Z(II)+MET 2016 •
- EXO-16-039 mono-γ
- EXO-16-040 monotop 2016 •

Analyses/minute=1.05

- 1 8TeV result :19.6 fb⁻¹ •
- 9 2015 results: 2.7 fb⁻¹
- 11 2016 results: 12.9 fb⁻¹•

EXO-16-030 low-mass dijets 2015

EXO-16-032 dijets

- EXO-16-032 Z'→ee/µµ 2016
- EXO-16-043 LQ(ee) 2016
- EXO-16-034 Z(II)gamma 2016
- EXO-16-035 Z(qq)gamma 2016
- 2016 EXO-16-025 Z(qq)gamma 8 TeV
 - EXO-16-022 displaced e-mu 2015
 - EXO-16-036 HSCP 2016
 - B2G-16-020 VW(W→lv) 2016
 - B2G-16-010 VZ (Z→II) 2015
 - B2G-16-017 W'→tb(→lvbb) 2016

B2G-16-009 W'→tb(→qqbb) 2015

- EXO-16-005 ttbar+MET 2015
- EXO-16-010 Z(II)+MET 2015 •
- EXO-16-011 monoH(γγ) 2015
- EXO-16-012 monoH(bb) 2015 •
- EXO-16-037 monojet 201
- EXO-16-038 Z(II)+MET 2016

201

- EXO-16-039 mono-γ
- EXO-16-040 monotop 201

Heavy Mediators

	vviiat5 iii	
5 •	EXO-16-032 dijets	2016
5 •	EXO-16-030 low-mass dijets	s 2015
5 •	EXO-16-032 Z'→ee/µµ	2016
5 •	EXO-16-043 LQ(ee)	2016
6•	EXO-16-034 Z(II)gamma	2016
6 •	EXO-16-035 Z(qq)gamma	2016
6 •	EXO-16-025 Z(qq)gamma	8 Te\
6 •	EXO-16-022 displaced e-mu	2015
•	EXO-16-036 HSCP	2016
•	B2G-16-020 VW(W→Iv)	2016
•	B2G-16-010 VZ (Z→II)	2015
•	B2G-16-017 W'→tb(→lvbb)	2016

B2G-16-009 W'→tb(→qqbb) 2015

- EXO-16-005 ttbar+MET 2015
- EXO-16-010 Z(II)+MET 2015 •
- EXO-16-011 monoH(γγ) 2015 •
- EXO-16-012 monoH(bb) 2015 •
- EXO-16-037 monojet 2016 •
- EXO-16-038 Z(II)+MET 2016 •
- EXO-16-039 mono-γ
- EXO-16-040 monotop

Diphoton stuff

• EXO-16-030 low-mass dijets 2015

EXO-16-032 dijets

- EXO-16-032 Z'→ee/µµ 2016
- EXO-16-043 LQ(ee) 2016
- EXO-16-034 Z(II)gamma 2016
 - EXO-16-035 Z(qq)gamma 2016
- 2016 EXO-16-025 Z(qq)gamma 8 TeV
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- EXO-16-005 ttbar+MET 2015
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- EXO-16-012 monoH(bb) 2015 •
- EXO-16-037 monojet 2016 •
- EXO-16-038 Z(II)+MET 2016 •
- EXO-16-039 mono-γ
- EXO-16-040 monotop

Lifetime Frontier

EXO-16-030 low-mass dijets 2015

EXO-16-032 dijets

- EXO-16-032 Z'→ee/µµ 2016
- EXO-16-043 LQ(ee) 2016
- EXO-16-034 Z(II)gamma 2016
- EXO-16-035 Z(qq)gamma 2016
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- EXO-16-005 ttbar+MET 2015
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- EXO-16-012 monoH(bb) 2015 •
- EXO-16-037 monojet 2016 •
- EXO-16-038 Z(II)+MET 2016 •
- EXO-16-039 mono-γ
- EXO-16-040 monotop 2016 •

Boosted Jet Searches

EXO-16-030 low-mass dijets 2015

EXO-16-032 dijets

- EXO-16-032 Z'→ee/µµ 2016
- EXO-16-043 LQ(ee) 2016
- EXO-16-034 Z(II)gamma 2016
- EXO-16-035 Z(qq)gamma 2016
- 2016 EXO-16-025 Z(qq)gamma 8 TeV
 - EXO-16-022 displaced e-mu 2015
 - EXO-16-036 HSCP 2016
 - B2G-16-020 VW(W→Iv) 2016
 - B2G-16-010 VZ (Z→II) 2015
 - B2G-16-017 W'→tb(→lvbb) 2016
 - B2G-16-009 W'→tb(→qqbb) 2015

EXO-16-005 ttbar+MET 2015

- EXO-16-010 Z(II)+MET 2015
- EXO-16-011 monoH(γγ) 2015
- EXO-16-012 monoH(bb) 2015
- EXO-16-037 monojet 2016
- EXO-16-038 Z(II)+MET 2016
- EXO-16-039 mono-γ
 2016

EXO-16-040 monotop 2016

Dark Matter

- EXO-16-032 dijets 2016
- EXO-16-030 low-mass dijets 2015
- EXO-16-032 Z'→ee/µµ 2016
- EXO-16-043 LQ(ee) 2016
- EXO-16-034 Z(II)gamma 2016
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- B2G-16-017 W'→tb(→lvbb) 2016
- B2G-16-009 W'→tb(→qqbb) 2015

Common Themes

Running at such high luminosity is tough!



- While drops like this are not long term issues
- Challenge of quality results in short time hard

Control of Data



Remarkable control of the *MET* Plot is 1 week old

Same events (less PU sensitivity)

• However we have it mostly under control

- We still have time to push out new techniques (#PUPPI) JME-16-004

The Searches

Making a "Mass" out of of It



Mass (169 ways)







EXO-16-030

Total Luminosity : 2.7 fb⁻¹ (2015) Final State : (qq)+j



Beats UA2 (1988) Goes to new phase space Z' mass < 140 GeV



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EXO-16-030

Total Luminosity : 2.7 fb⁻¹ (2015) Final State : jj+j



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*** Discussion title: Review of EXO-16-030

Dear Authors

I noticed that in your PAS the introduction refers to UA1 and UA2 results at sqrt(s)=300 GeV. These results came out when I was on UA2 in fact, and the SppS ran at sqrt(s)=630GeV.

Best regards Joe[Incandela]



Total Luminosity : 2.7 fb⁻¹ (2016) **Final State** : V I

EXO-16-015



Final State

Total Luminosity : 2.6 fb⁻¹ (2015) Final State : ee+2j Leptoquark(LQ $\rightarrow e+j$)

Leptoquark decaying in electron final state



400

600

800

 M_{IO} (GeV)

1000

1200

1400

Search in tail of mass/ H_{T} No excess and no deficit

8TeV leptoquark search excess is now gone

EXO-16-043

Total Luminosity : 19.6/12.9 fb⁻¹ (8TeV/2016)Final State: $Z+\gamma(Z\rightarrow qq/Z\rightarrow II)$ CMS Preliminary12.9 fb⁻¹ (13 TeV)



V/top-tagged Searches







B2G-16-020





B2G-16-010



Lifetime Frontier



Total Luminosity : 12.9 fb⁻¹ (2016)Heavy & StableFinal State: Long lived track

• Two approaches: dE/dX & Time of Flight



Dark Matter Searches

Search for Dark Matter at LHC

Can split dark matter into two classes of searches



Yukawa coupling to quarks (At the moment no mixing)

Flavor universal to quarks (At the moment no mixing)

All dark matter searches are really a search for Dark Matter + A mediator



Mediator X p_{τ} (13 ways)

Spin 1 DM Searches



Spin 0 DM Searches





EXO-16-037



Small integrated deficit over whole region Mass exclusion out to 750 GeV (2 TeV for monojet/V)

EXO-16-039



Small excess(1σ) over whole region Mass exclusion out to 400 GeV for vector mediator Scalar model BR(Higgs) <0.86 (Guillelmo's talk)

EXO-16-038



EXO-16-040

Total Luminosity : 2.3 fb⁻¹ (2015) Final State : Higgs(bb/ $\gamma\gamma$)+*MET*

Mono-Higgs



EXO-16-011/EXO-16-012



EXO-16-011/EXO-16-012



Sensitivity pushed to exclude Scalar mediator models EXO-16-005

Mediator Arms Race



Who can cover the territory first?

Dark Matter Summary

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In Summary

- No New Physics
 - A plethora of new an interesting results
 - Excess in ZZ(II+j)@650 GeV and γ+j @ 2TeV
- LHC Run II has warmed up
 - Just getting into the interesting part
- Now probing :
 - Spin 1 mediators in the 2-3 TeV range
 - Spin 0 mediators in the hundreds of GeV range
 - Long lived 100 GeV particles in the few cm range

Thanks (750)에 의해 낙심 하지않습니다

*Bonus Heavy Ion QCD talk in back



My Favorite Heavy Ion Result

Softdrop condition

Softdrop = Jet grooming technique removes large-angle soft radiation + remaining background



We use $\beta = 0$ and $z_{cut} = 0.1$ All soft emissions are removed Equivalent to modified Mass Drop

[1] Larkoski, Marzani, Thaler Phys. Rev. D91:111501 (2015) Soft Drop: JHEP 1405 (2014) 146 Soft Drop condition $z > z_{cut} \theta^{\beta}_{\uparrow}_{\downarrow}_{\text{energy angular exponent}}$

$$z_{g} = \frac{\min(p_{T,1}, p_{T,2})}{p_{T,1} + p_{T,2}}$$

Momentum fraction carried by the subleading branch of first splitting

Measurement in pp

Note: For detector resolution reasons require $\Delta R_{12(subjet)} > 0.1$



Pythia and Herwig predictions straddle the pp result



Jet p_T dependence

Modification gets weaker when increasing jet p_T



Dark Matter Searches







Figure 3: Cross section upper limits at 95% CL on various signal models for the tracker-only inalysis (left column) and tracker+TOF analysis (right column) at $\sqrt{s} = 13$ TeV. In the legend, CS' stands for charged suppressed interaction model.

PUPPE_T Performance in Data





Total Luminosity : 2.6 fb⁻¹ (8TeV/2015) ⁶⁰ Final State : VV VCombination



Diboson search now has many different final states

Many different combinations possible

Signal

CRs: γ+jets



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W +



Control: another decay of a Z boson $Z \rightarrow VV$ Remove

Problem is control regions have less events than signal $\sigma_{\mu\mu} = 0.1 \sigma_{\nu\nu}$ -Statistical precision is 4x worse CMS-EXO-16-010 CMS-EXO-12-055



1 Control region 100% uncertainty @ 1 TeV





2 Control regions 60% uncertainty @ 1 TeV





3 Control regions 40% uncertainty @ 1 TeV





4 Control regions 30% uncertainty @ 1 TeV





5 Control regions 15% uncertainty @ 1 TeV





5 Control regions+Signal 15% uncertainty @ 1 TeV All in one big Simultaneous fit



To large extent the γ+jets drives the constraint

- However we need need $Z \rightarrow II$ to constraint γ



A mystery? Understanding $Z/\gamma p_{\tau}$

Can we really use Photons to model Zs?



*See backup

CMS-SMP-14-005

How do we fix this?

Impact of the electroweak corrections



A mystery? The Z p_{τ} spectrum

• These results are missing NLO EWK corrections!



dpT dpT

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However we still have a problem!

Unc.
$$- \frac{d\sigma^{\gamma}}{dp_{\tau}} / \frac{d\sigma^{z}}{dp_{\tau}} = d\sigma^{\gamma}/d\sigma^{z}(\mu)$$



Uncertainty on ratio? How is it done? Scale uncertainty on process #1 Scale uncertainty on process #2 Uncertainty on process #1/process #2 (fully correlated) % unc $\frac{d\sigma^{\gamma}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} = d\sigma^{\gamma}/d\sigma^{Z}(\mu)$ Unc. $\begin{pmatrix} d\sigma^{\gamma}(+\sigma) \\ d\sigma^{Z}(+\sigma) \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \begin{pmatrix} d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0}) \\ d\sigma^{Z}(\mu^{up})/d\sigma^{i}(\mu_{0}) \end{pmatrix}$

Uncertainty on ratio? How is it done? Scale uncertainty on process #1 Scale uncertainty on process #2 Uncertainty on process #1/process #2 (fully correlated)

Unc. $d\sigma^{\gamma} / d\sigma^{z} = d\sigma^{\gamma}/d\sigma^{z}(\mu)$ dp_{T} / dp_{T}

%

 $\begin{pmatrix} d\sigma^{\gamma}(+\sigma) \\ d\sigma^{Z}(+\sigma) \end{pmatrix} = \begin{pmatrix} 1 & C \\ C & 1 \end{pmatrix} \begin{pmatrix} d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0}) \\ d\sigma^{Z}(\mu^{up})/d\sigma^{i}(\mu_{0}) \end{pmatrix}$ Adjust C until uncertainty is

Uncertainty on ratio? How is it done?

Scale uncertainty on process #1 Scale uncertainty on process #2 Uncertainty on process #1/process #2 (fully correlated)

%

Unc. $\frac{d\sigma^{\gamma}}{dp_{T}} / \frac{d\sigma^{Z}}{dp_{T}} = d\sigma^{\gamma}/d\sigma^{Z}(\mu)$ $\begin{pmatrix} d\sigma^{\gamma}(+\sigma) \\ d\sigma^{Z}(+\sigma) \end{pmatrix} = \begin{pmatrix} 1 & C \\ C & 1 \end{pmatrix} \begin{pmatrix} d\sigma^{\gamma}(\mu^{up})/d\sigma^{i}(\mu_{0}) \\ d\sigma^{Z}(\mu^{up})/d\sigma^{i}(\mu_{0}) \end{pmatrix}$

Decorrelate scale unc. until its max of either process

 $d\sigma^{v}/d\sigma^{z}(+\sigma) < \max_{i} (d\sigma^{i}(\mu^{up})/d\sigma^{i}(\mu_{o}))$



What about the EWK uncertainty?

In light of being conservative : Treated <u>full correction as an uncertainty</u> More formal way could be with scale



Additionally <u>de-correlated this per bin</u> Avoids low *MET* to high *MET* constraints Not very logical Other (better) schemes exist

What do the uncertainties look like?



Updated unc still too large

CMS-EXO-12-055

Profiling them in the fit

Constraints after the fit

