



Contribution ID: 208

Type: **Parallel Talk**

Precision Measurement of ^3He -to- ^4He ratio in Cosmic Rays with the AMS Detector on the Space Station

Thursday, 6 July 2017 12:06 (18 minutes)

The knowledge of the energy dependence of the ^3He -to- ^4He flux ratio ($^3\text{He}/^4\text{He}$) is one of the most important sources for studying and testing cosmic ray propagation models.

Similar to the B/C measurement, where B is assumed to originate from interactions of primary Carbon and Oxygen in Cosmic Rays, in the $^3\text{He}/^4\text{He}$ measurement ^3He is assumed to be produced by interactions of ^4He and heavier nuclei with the interstellar matter.

The $^3\text{He}/^4\text{He}$ ratio is a powerful tool for determining the amount of interstellar material traversed by cosmic rays and, since ^4He has smaller cross section compared to C and O, this provides testing of the propagation models over larger travelled distances.

The AMS results are unique and distinct from all the previous data and they are presented here for the first time. The AMS $^3\text{He}/^4\text{He}$ flux ratio is based on 9 million ^3He events and 56 million ^4He events and it extends from 0.7 GeV/n to 10 GeV/n in a region where previous measurements have large errors and are not consistent with each other. This prevents accurate comparison with models. We will present comparison of our results with theoretical models.

Experimental Collaboration

AMS collaboration

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Session Classification: Astroparticle physics

Track Classification: Astroparticle Physics