



## Standard Model of Particle Physics

Gauge Theory based on the group:

$$SU(3) \times SU(2) \times U(1)$$

$SU(3) \Rightarrow$  Quantum Chromodynamics

Strong Force (Quarks and Gluons)

$SU_L(2) \times U(1) \Rightarrow$  ElectroWeak Interactions broken to  $U_{EM}(1)$

by HIGGS

$$\underline{SU_L(2) \times U_Y(1) \Rightarrow U_{EM}(1)}$$

Force Carriers:  $W^\pm, Z^0$  and  $\gamma$  masses: 80, 91 and 0 GeV

quark, SU(2) doublets:  $\begin{pmatrix} u \\ d \end{pmatrix}_L, \begin{pmatrix} c \\ s \end{pmatrix}_L, \begin{pmatrix} t \\ b \end{pmatrix}_L$

up-quark, SU(2) singlets:  $u_R, c_R, t_R$

down-quark, SU(2) singlets:  $d_R, s_R, b_R$

lepton, SU(2) doublets:  $\begin{pmatrix} \nu_e \\ e \end{pmatrix}_L, \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L, \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L$

neutrino, SU(2) singlets: — — —

charge lepton, SU(2) singlets:  $e_R, \mu_R, \tau_R$

### Electron mass

comes from a term of the form

$$\bar{L}\phi e_R$$

Absence of  $\nu_R$

forbids such a mass term (dim 4)

for the Neutrino

Therefore in the SM neutrinos are massless  
and hence travel at speed of light.

### Interactions:

Charge Current (CC)

$W^- \rightarrow l_\alpha^- + \bar{\nu}_\alpha$

Neutral Current (NC)

$Z^0 \rightarrow \nu_\alpha + \bar{\nu}_\alpha$

$Z^0 \rightarrow l_\alpha^- + l_\alpha^+$

$$\Gamma(Z^0 \rightarrow f + \bar{f}) = K \frac{g_Z^2 M_Z}{48\pi} [|c_V^f|^2 + |c_A^f|^2]$$

$\alpha = e, \mu, \text{ or } \tau$

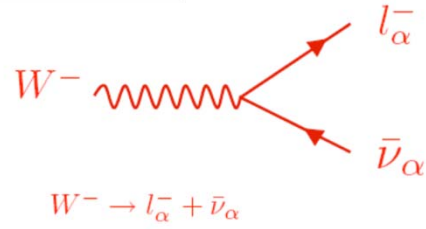
Invisible width of Z plus other data from LEP:

$Z^0 \rightarrow \nu\bar{\nu}$

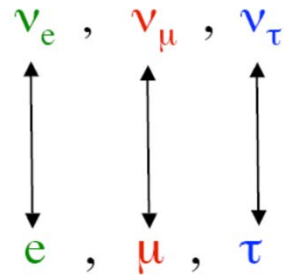
Implies  $N_\nu = 2.99 \pm 0.01$

Three Active Neutrinos!!!      Sterile Neutrinos don't couple to  $Z^0$

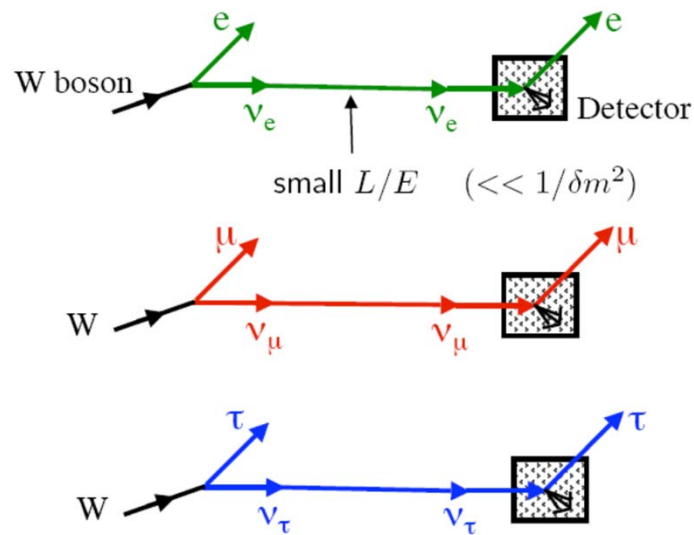
Note That

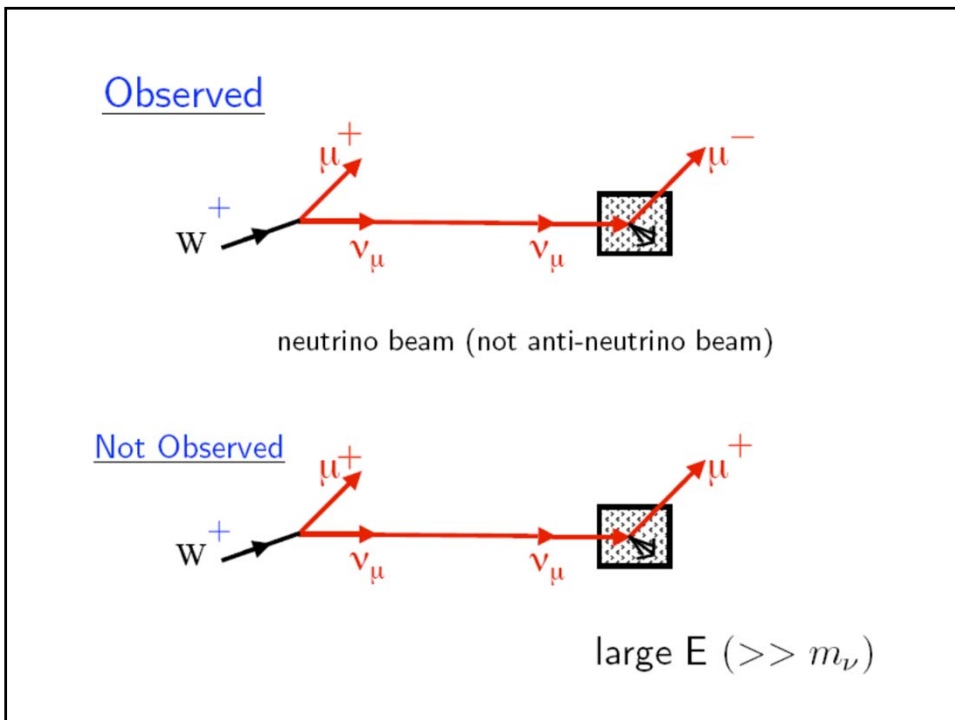
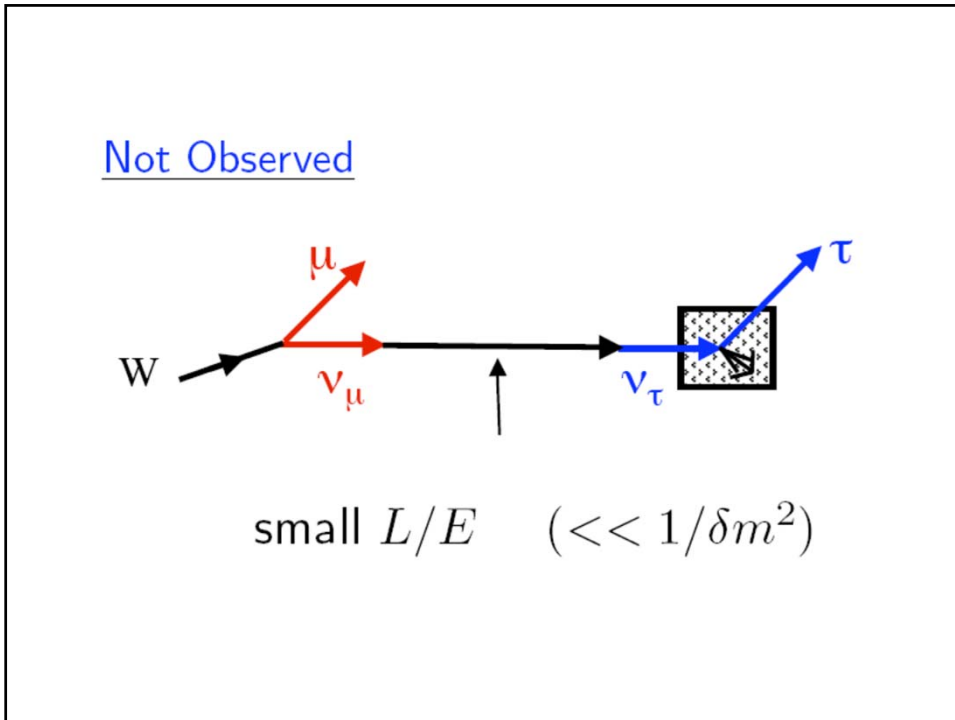


Implies

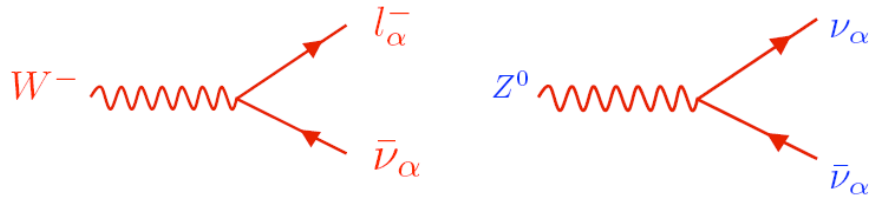


Observed





### Standard Model

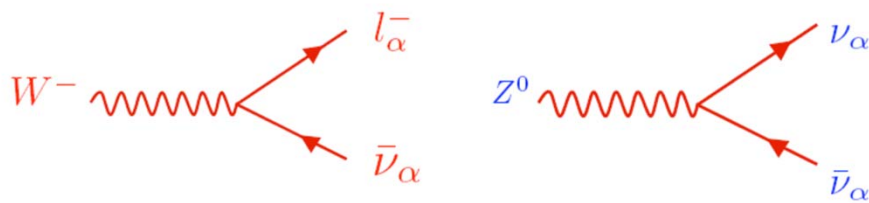


couplings conserve the **Lepton Number L**  
defined by—

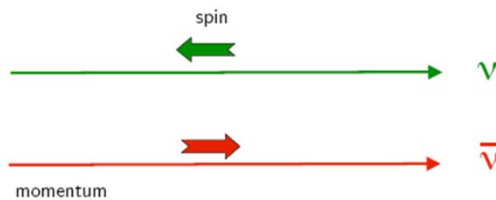
$$L(\nu) = L(\ell^-) = -L(\bar{\nu}) = -L(\ell^+) = 1.$$

Actually  $L_e$ ,  $L_\mu$ , and  $L_\tau$   
separately

### Left Handed Nature of The Neutrino



Produce Left-Handed Neutrinos  
and Right-Handed Anti-Neutrinos



What about the RH neutrinos and LH anti-neutrino ????

$\pi^+$  decay

Pion Rest Frame

Left Handed anti-fermion

Suppressed by powers of  $m_f$

Why  $\pi^+ \rightarrow \mu^+ \nu_\mu > 99\%$  and  $\pi^+ \rightarrow e^+ \nu_e$  is 0.01%.

There exist three fundamental and discrete transformations in nature:

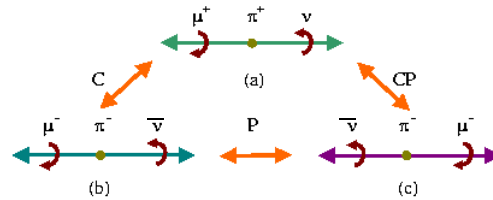
- Parity  $\mathcal{P}$   $\vec{x} \rightarrow -\vec{x}$
- Time reversal  $\mathcal{T}$   $t \rightarrow -t$
- Charge conjugation  $\mathcal{C}$   $q \rightarrow -q$

$\mathcal{P}$ ,  $\mathcal{T}$  and  $\mathcal{C}$  are conserved in the classical theories of mechanics and electrodynamics!

$CPT \leftrightarrow$  Lorentz invariance  $\oplus$  unitarity: is an essential building block of field theory

$CPT$  : L particle  $\leftrightarrow$  R antiparticle

Neutrinos in the MSM are massless and exist only in two states: particle with negative helicity and antiparticle with positive one: **Weyl fermion**



$\mathcal{P}$ : L particle  $\leftrightarrow$  R particle

Parity violation is nowhere more obvious than in the neutrino sector: the reflection of a left-handed neutrino in a mirror is nothing !

### Summary of $\nu$ 's in SM:

Three flavors of massless neutrinos

$$W^- \rightarrow l_{\alpha}^- + \bar{\nu}_{\alpha}$$

$$W^+ \rightarrow l_{\alpha}^+ + \nu_{\alpha}$$

$$\alpha = e, \mu, \text{ or } \tau$$

Anti-neutrino,  $\bar{\nu}_{\alpha}$ , has +ve helicity, Right Handed

Neutrino,  $\nu_{\alpha}$ , has -ve helicity, Left Handed

$\nu_L$  and  $\bar{\nu}_R$  are CPT conjugates

massless implies helicity = chirality



## Beyond the SM

What if Neutrino have a MASS?

speed is less than c therefore time can pass

and

Neutrinos can change character!!!

What are the stationary states?

How are they related to the interaction states?

## NEUTRINO OSCILLATIONS:

Two Flavors

flavor eigenstates  $\neq$  mass eigenstates

$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

W's produce  $\nu_\mu$  and/or  $\nu_\tau$ 's

but  $\nu_1$  and  $\nu_2$  are the states

that change by a phase over time, mass eigenstates.

$$|\nu_j\rangle \rightarrow e^{-ip_j \cdot x} |\nu_j\rangle \quad p_j^2 = m_j^2$$

$\alpha, \beta \dots$  flavor index

$i, j \dots$  mass index

Production:

$$|\nu_\mu\rangle = \cos\theta|\nu_1\rangle + \sin\theta|\nu_2\rangle$$

Propagation:

$$\cos\theta e^{-ip_1 \cdot x}|\nu_1\rangle + \sin\theta e^{-ip_2 \cdot x}|\nu_2\rangle$$

Detection:

$$|\nu_1\rangle = \cos\theta|\nu_\mu\rangle - \sin\theta|\nu_\tau\rangle$$

$$|\nu_2\rangle = \sin\theta|\nu_\mu\rangle + \cos\theta|\nu_\tau\rangle$$

$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = |\cos\theta(e^{-ip_1 \cdot x})(-\sin\theta) + \sin\theta(e^{-ip_2 \cdot x})\cos\theta|^2$$

$$P(\nu_\mu \rightarrow \nu_\tau) = |\cos\theta(e^{-ip_1 \cdot x})(-\sin\theta) + \sin\theta(e^{-ip_2 \cdot x})\cos\theta|^2$$

$$\text{Same } E, \text{ therefore } p_j = \sqrt{E^2 - m_j^2} \approx E - \frac{m_j^2}{2E}$$

$$e^{-ip_j \cdot x} = e^{-iEt} e^{-ip_j L} \approx e^{-i(Et - EL)} e^{-im_j^2 L/2E}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2\theta \cos^2\theta |e^{-im_2^2 L/2E} - e^{-im_1^2 L/2E}|^2$$

$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

$$\delta m^2 = m_2^2 - m_1^2 \text{ and } \frac{\delta m^2 L}{4E} \equiv \Delta \text{ kinematic phase:}$$

$$P(\nu_\mu \rightarrow \nu_\tau) = |\cos\theta(e^{-ip_1 \cdot x})(-\sin\theta) + \sin\theta(e^{-ip_2 \cdot x})\cos\theta|^2$$

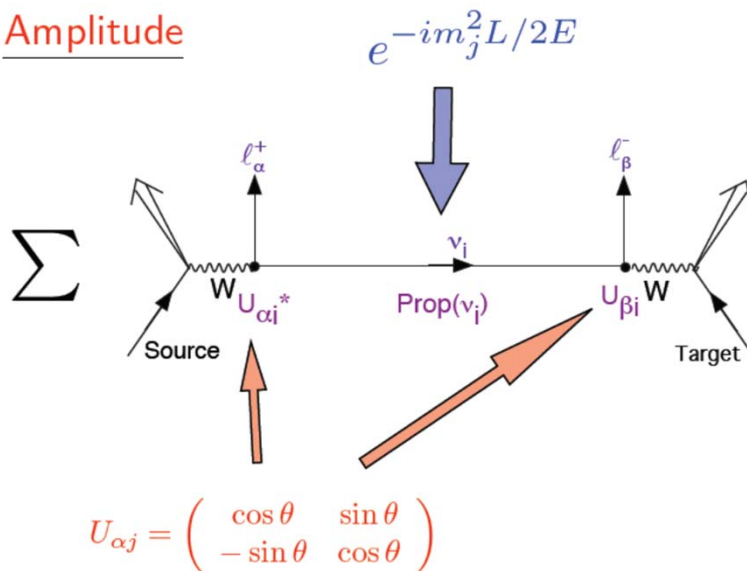
Same E, therefore  $p_j = \sqrt{E^2 - m_j^2} \approx E - \frac{m_j^2}{2E}$

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$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E} \frac{c^4}{\hbar c}$$

### Amplitude



Appearance:

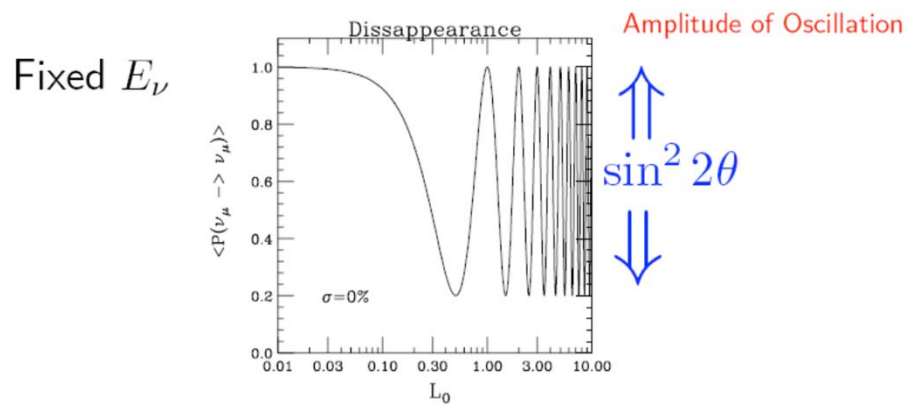
$$P(\nu_\mu \rightarrow \nu_\tau) = \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

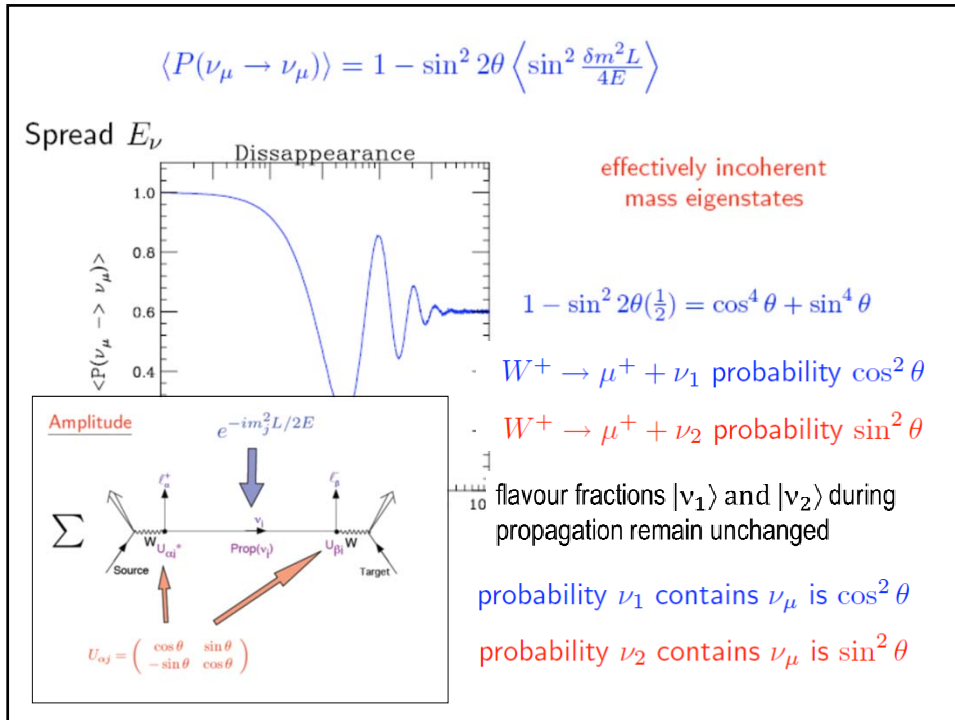
Disappearance:

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

$$P(\nu_\mu \rightarrow \nu_\mu) = 1 - \sin^2 2\theta \sin^2 \frac{\delta m^2 L}{4E}$$

Oscillation Length  $L_0 = 4\pi E / \delta m^2$





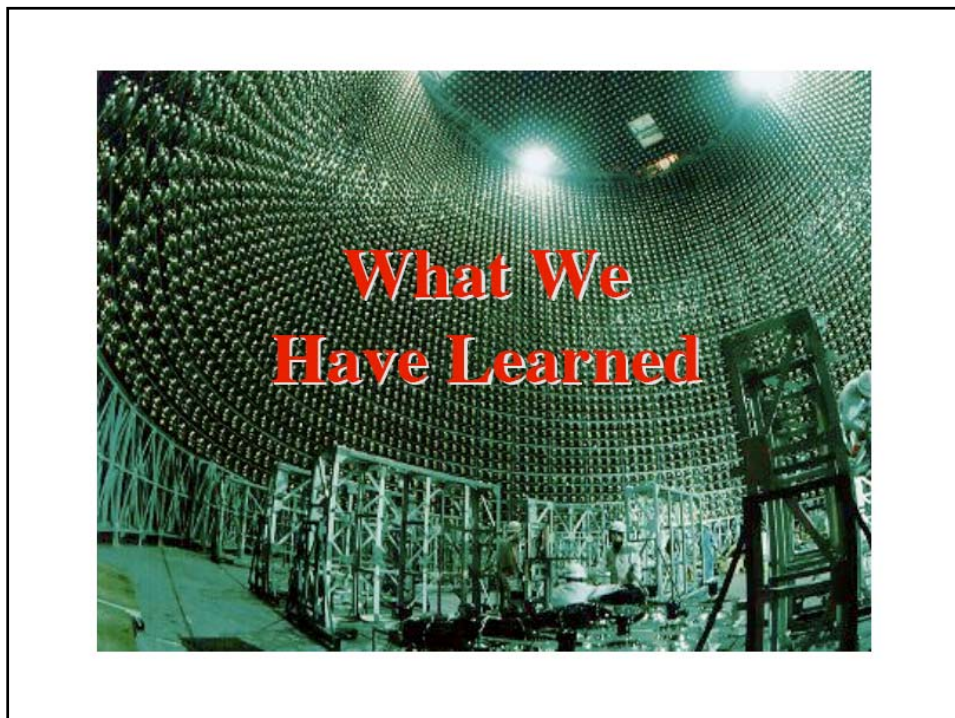
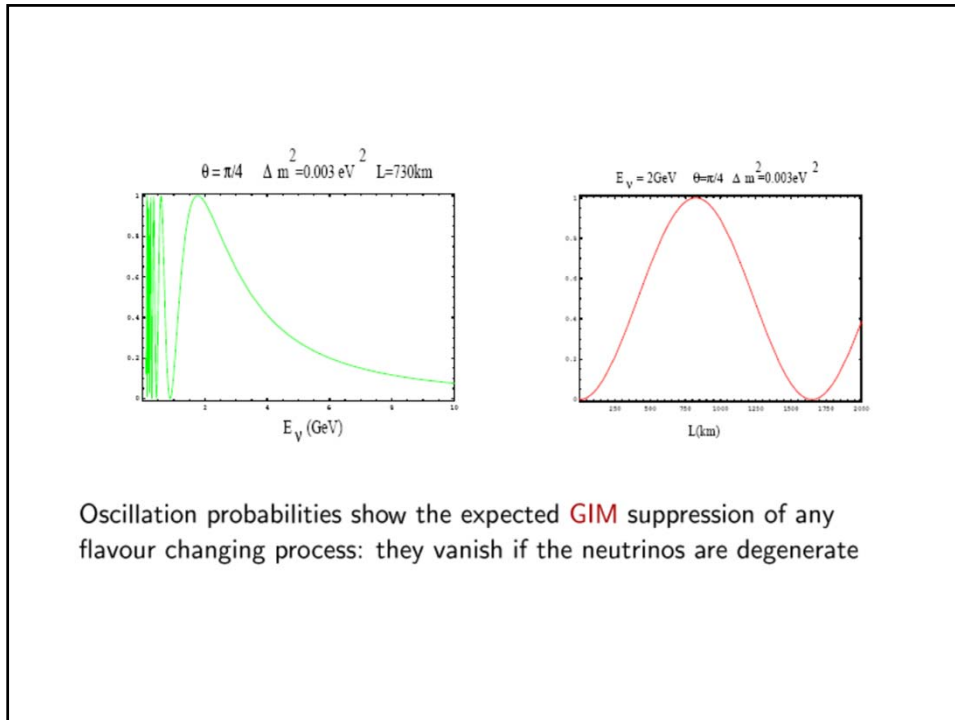
Using the unitarity of the mixing matrix: ( $W_{\alpha\beta}^{jk} \equiv [V_{\alpha j} V_{\beta j}^* V_{\alpha k} V_{\beta k}]$ )

$$P(\nu_\alpha \rightarrow \nu_\beta) = \delta_{\alpha\beta} - 4 \sum_{k>j} \text{Re}[W_{\alpha\beta}^{jk}] \sin^2 \left( \frac{\Delta m_{jk}^2 L}{4E_\nu} \right) \pm 2 \sum_{k>j} \text{Im}[W_{\alpha\beta}^{jk}] \sin \left( \frac{\Delta m_{jk}^2 L}{2E_\nu} \right)$$

For 2 families:  $V_{MNS} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix}$

$$P_{\alpha\beta} = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E_\nu} \right) \rightarrow \text{appearance}$$

$$P_{\alpha\alpha} = 1 - P_{\alpha\beta} < 1 \rightarrow \text{disappearance}$$



## Probability for Neutrino Oscillation in Vacuum

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\text{Amp}(\nu_\alpha \rightarrow \nu_\beta)|^2 =$$

$$P_{\alpha\beta} = \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E\nu} \right) \rightarrow \text{appearance}$$

$$P_{\alpha\alpha} = 1 - P_{\alpha\beta} < 1 \rightarrow \text{disappearance}$$

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$$\left( 1.27 \frac{\Delta m^2 (eV^2) L(km)}{E(GeV)} \right)$$

L/E becomes crucial !!!

## Evidence for Flavor Change:

\*\*\* Atmospheric and Accelerator Neutrinos with  $L/E = 500 \text{ km/GeV}$

\*\*\* Solar and Reactor Neutrinos with  $L/E = 15 \text{ km/MeV}$

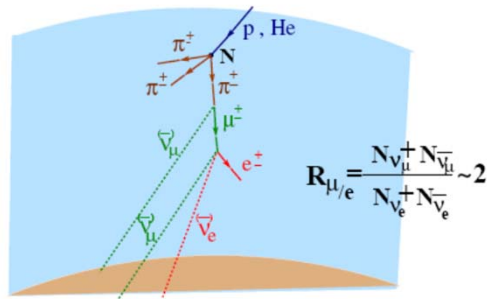
Neutrinos from Stopped muons  $L/E = 2 \text{ m/MeV}$  (Unconfirmed)

### Atmospheric neutrinos

- Atmospheric neutrinos are produced by the interaction of *cosmic rays* ( $p, \text{He}, \dots$ ) with the Earth's atmosphere:

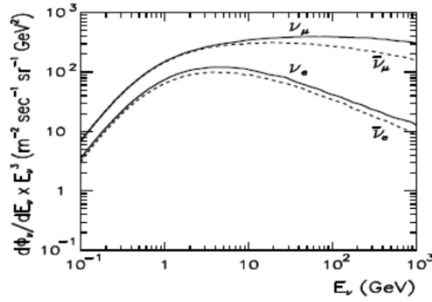
- 1  $A_{\text{cr}} + A_{\text{air}} \rightarrow \pi^{\pm}, K^{\pm}, K^0, \dots$
- 2  $\pi^{\pm} \rightarrow \mu^{\pm} + \nu_{\mu}$ ,
- 3  $\mu^{\pm} \rightarrow e^{\pm} + \nu_e + \nu_{\mu}$ ;

- at the detector, some  $\nu$  interacts and produces a **charged lepton**, which is observed.



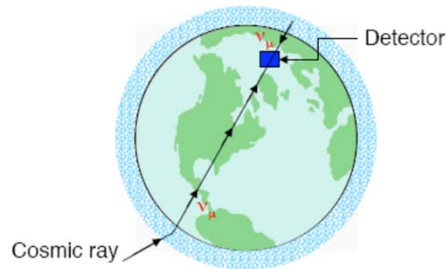


$\nu$  are produced in the atmosphere when primary cosmic rays impinge on it producing  $K, \pi$  which subsequently decay:



A deficit was observed in the ratio  $\mu/e$  events: **Soudan2, IMB, Kamiokande**

### Atmospheric Neutrinos



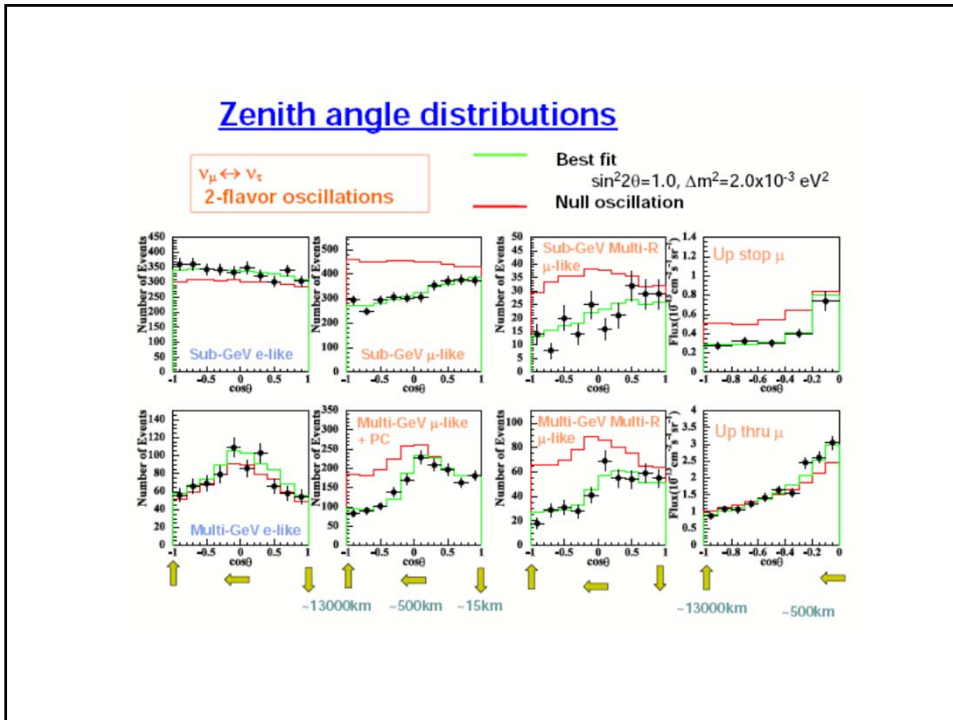
Isotropy of the  $\geq 2$  GeV cosmic rays + Gauss' Law + No  $\nu_\mu$  disappearance

$$\Rightarrow \frac{\phi_{\nu_\mu}(\text{Up})}{\phi_{\nu_\mu}(\text{Down})} = 1 .$$

But Super-Kamiokande finds for  $E_\nu > 1.3$  GeV

$$\frac{\phi_{\nu_\mu}(\text{Up})}{\phi_{\nu_\mu}(\text{Down})} = 0.54 \pm 0.04 .$$





Half of the upward-going, long-distance-traveling  $\nu_\mu$  are disappearing.

Voluminous atmospheric neutrino data are well described by —

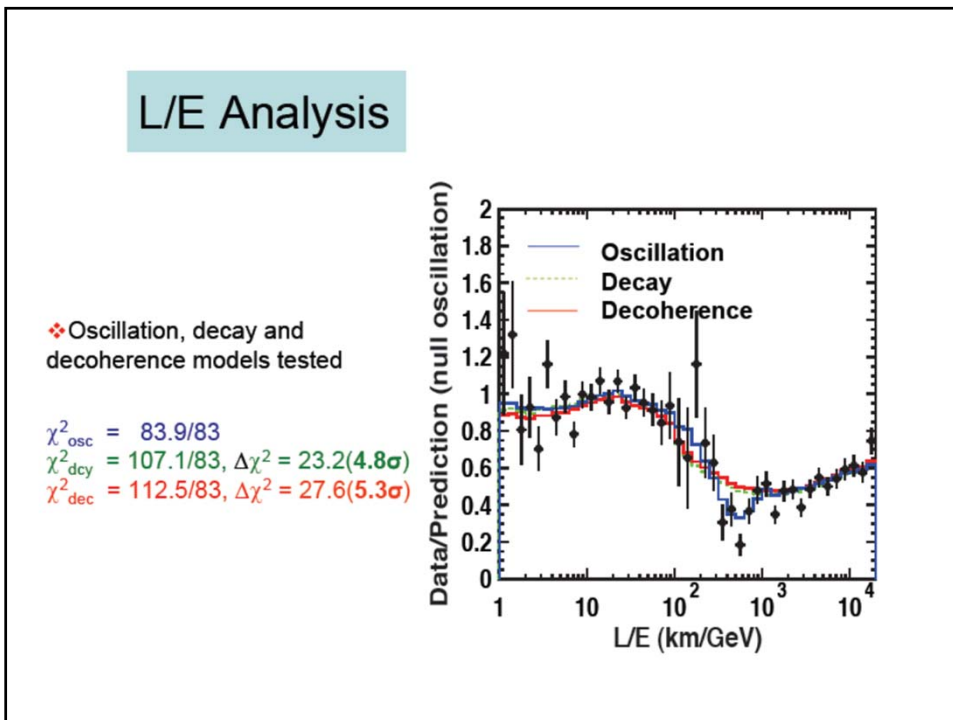
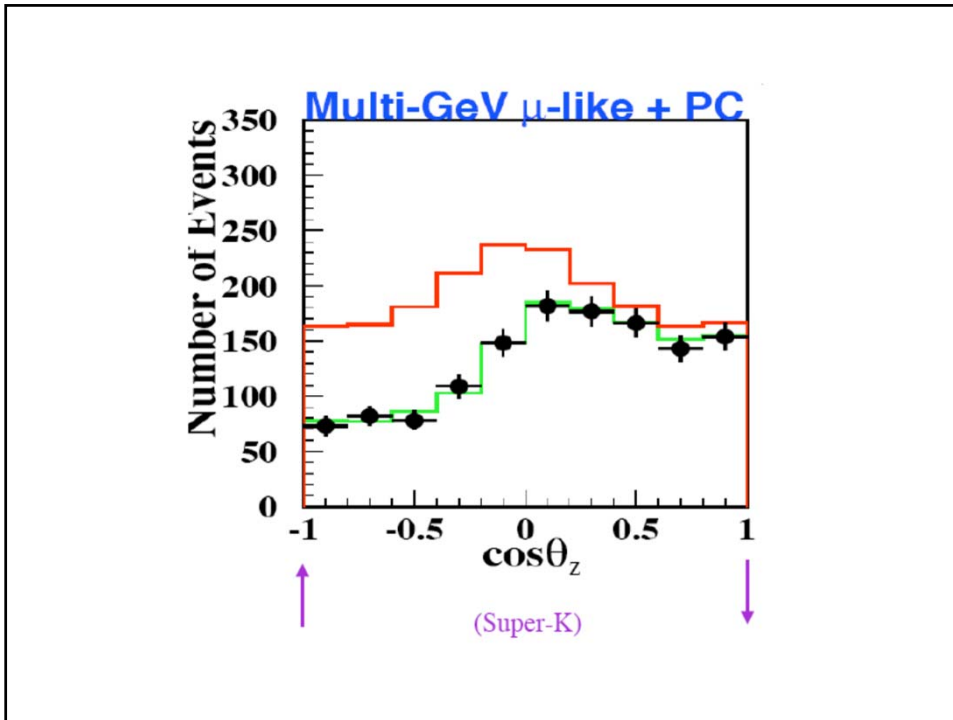
$$\nu_\mu \longrightarrow \nu_\tau$$

with —

$$\Delta m_{\text{atm}}^2 \cong 2.4 \cdot 10^{-3} \text{ eV}^2$$

and —

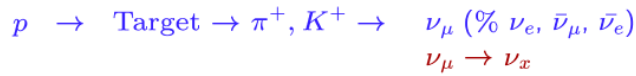
$$\sin^2 2\theta_{\text{atm}} \cong 1$$



Confirmation of the oscillation hypothesis in "man-made"  $\nu$  sources

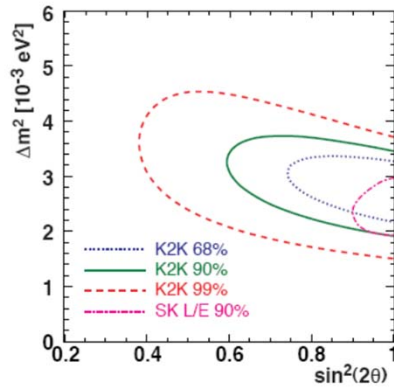
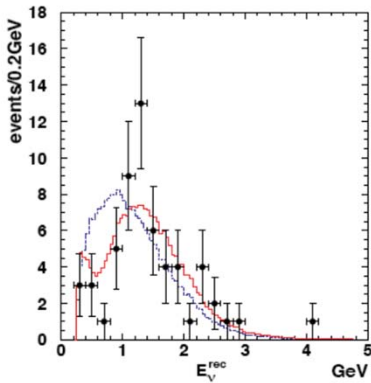
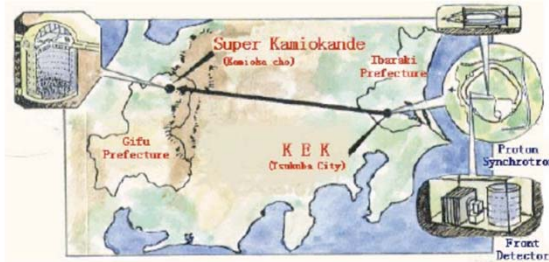
$$|\Delta m_{\text{atmos}}^2| \sim \frac{E_\nu(1 - 10\text{GeV})}{L(10^2 - 10^3\text{km})}$$

$\nu$  beams in this energy range can easily be produced at accelerators



Three such conventional beams **KEK-Kamioka (235km)**, **Fermilab-Soudan (730km)**, **CERN-Gran Sasso (730km)** are looking for the disappearance of  $\nu_\mu$  or appearance of  $\nu_\tau$ :

K2K

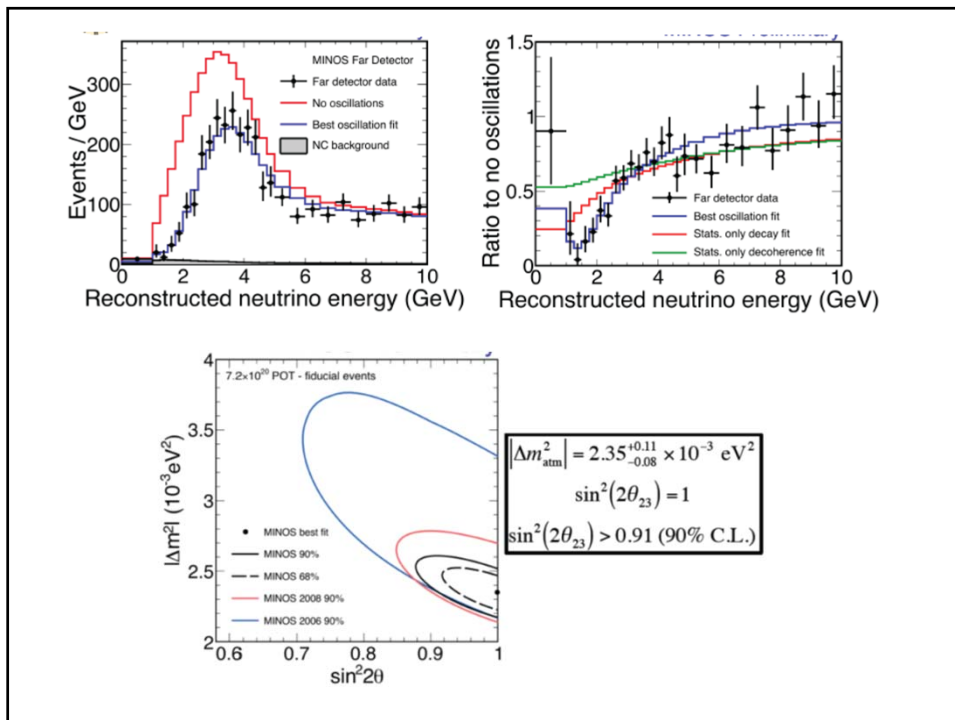


### The MINOS Experiment

**Near Detector**  
 1 kton  
 $3.8 \times 4.8 \times 15 \text{ m}^3$   
 282 steel planes  
 153 scintillator planes

**Far Detector**  
 5.4 kton  
 $8 \times 8 \times 30 \text{ m}^3$   
 484 planes

Fermilab 10 km Soudan  
735 km



### "Atmospheric" Neutrino Summary

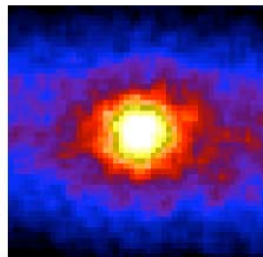
$$\nu_\mu \rightarrow \nu_\tau$$

no evidence of  $\nu_e$  involvement:

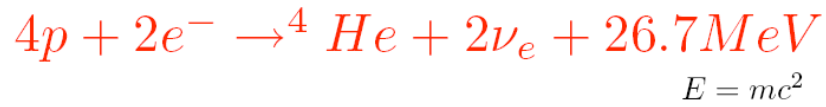
$$\delta m_{atm}^2 = 2.7_{-0.3}^{+0.4} \times 10^{-3} eV^2 \quad L/E = 500 \text{ km/GeV}$$

$$\sin^2 2\theta_{atm} > 0.92 \quad \Rightarrow 0.35 < \sin^2 \theta_{atm} < 0.65$$

## Solar $\delta m^2$



Solar Engine:

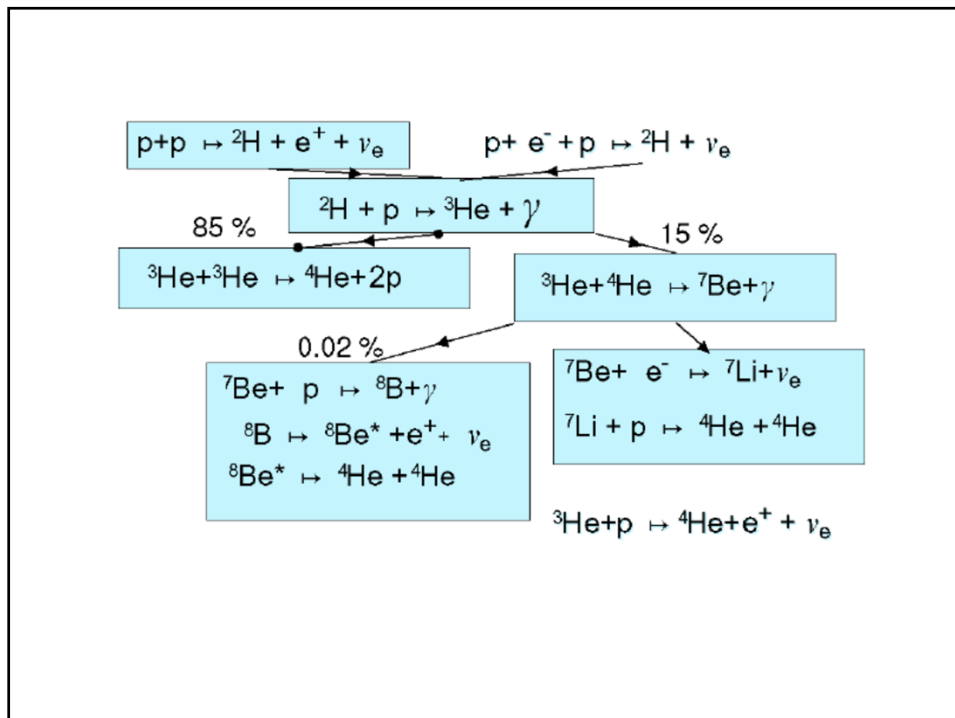


1  $\nu_e$  for every 13.4 MeV ( $=2.1 \times 10^{-12}$  J)

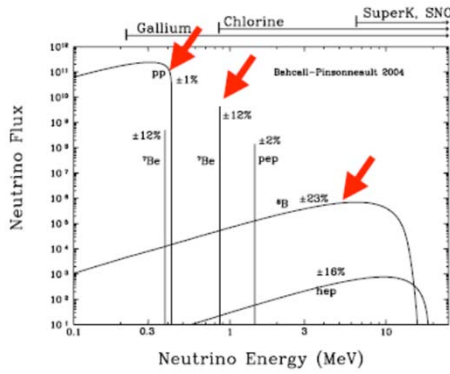
$\mathcal{L}_\odot$  at earth's surface 0.13 watts/cm<sup>2</sup>

$$\phi_\nu = \frac{0.13}{2.1 \times 10^{-12}} = 6 \times 10^{10} / \text{cm}^2 / \text{sec}$$

This corresponds to an average of 2  $\nu$ 's per cm<sup>3</sup>  
since they are going at speed c.



### Solar Spectrum:



$$p + p \rightarrow {}^2\text{H} + e^+ + \nu_e$$

$$\phi_{pp} = 5.94(1 \pm 0.01) \times 10^{10} \text{cm}^{-2} \text{sec}^{-1}$$

$${}^7\text{Be} + e^- \rightarrow {}^7\text{Li} + \nu_e$$

$$\phi_{{}^7\text{Be}} = 4.86(1 \pm 0.12) \times 10^9 \text{cm}^{-2} \text{sec}^{-1}$$

$${}^7\text{Be} + p \rightarrow {}^8\text{B} \rightarrow {}^8\text{Be}^* + e^- + \nu_e$$

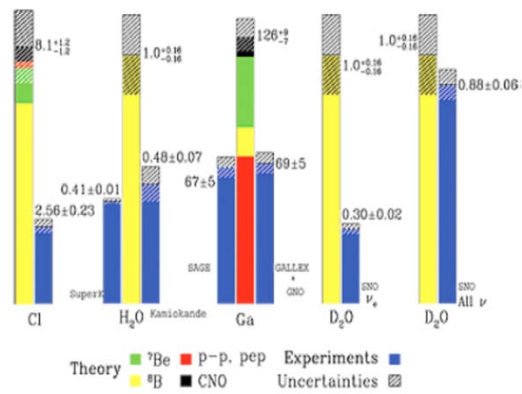
$$\phi_{{}^8\text{B}} = 5.82(1 \pm 0.23) \times 10^6 \text{cm}^{-2} \text{sec}^{-1}$$

**Figure 1.** The predicted solar neutrino energy spectrum. The figure shows the energy spectrum of solar neutrinos predicted by the BP04 solar model [22]. For continuum sources, the neutrino fluxes are given in number of neutrinos  $\text{cm}^{-2} \text{s}^{-1} \text{MeV}^{-1}$  at the Earth's surface. For line sources, the units are number of neutrinos  $\text{cm}^{-2} \text{s}^{-1}$ . Total theoretical uncertainties taken from column 2 of table 1 are shown for each source. To avoid complication in the figure, we have omitted the difficult-to-detect CNO neutrino fluxes (see table 1).



Ray Davis & John Bahcall

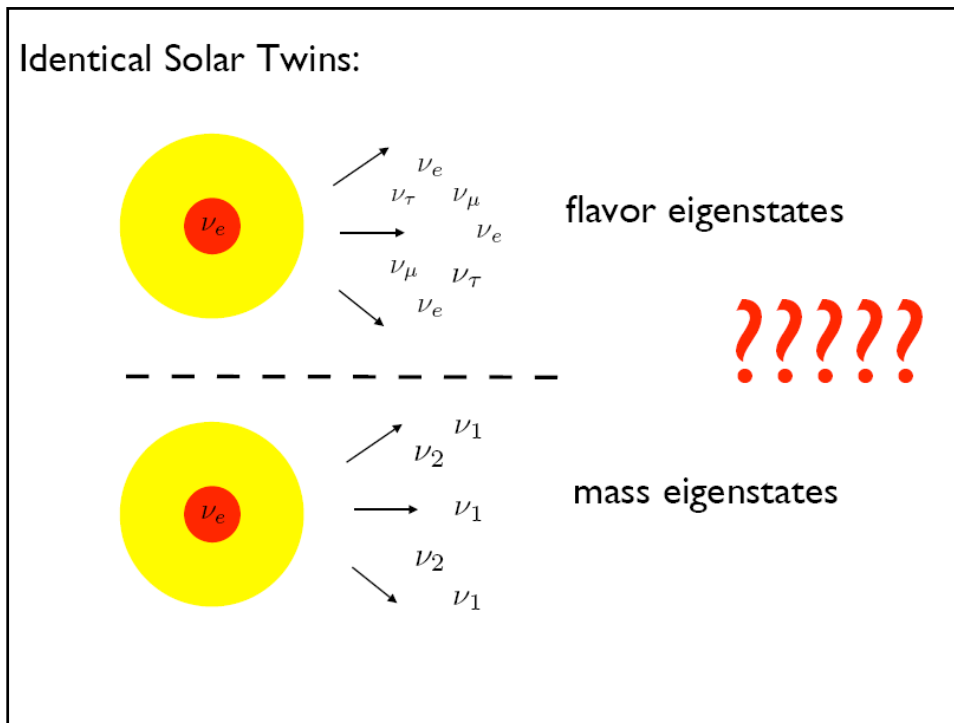
Total Rates: Standard Model vs. Experiment  
Bahcall-Serenelli 2005 [BS05(OP)]



Theory v Exp.

**Neutrino Flavor Transitions!!!**





Kinematical Phase:

$$\delta m_{\odot}^2 = 8.0 \times 10^{-5} eV^2$$

$$\sin^2 \theta_{\odot} = 0.31$$

$$\Delta_{\odot} = \frac{\delta m_{\odot}^2 L}{4E} = 1.27 \frac{8 \times 10^{-5} eV^2 \cdot 1.5 \times 10^{11} m}{0.1-10 MeV}$$

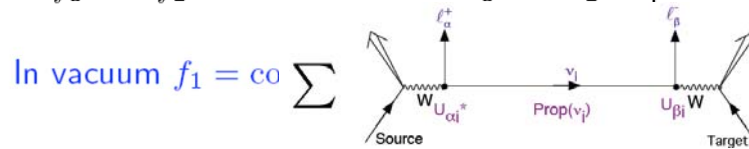
$$\Delta_{\odot} \approx 10^{7 \pm 1}$$

**Effectively Incoherent !!!**

Vacuum  $\nu_e$  Survival Probability:

$$\langle P_{ee} \rangle = f_1 \cos^2 \theta_\odot + f_2 \sin^2 \theta_\odot$$

where  $f_1$  and  $f_2$  are the fraction of  $\nu_1$  and  $\nu_2$  at production.

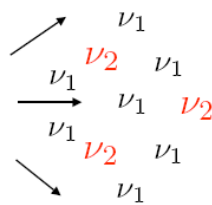
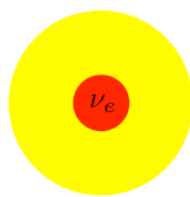


$$\langle P_{ee} \rangle = \cos^4 \theta_\odot + \sin^4 \theta_\odot = 1 - \frac{1}{2} \sin^2 2\theta_\odot$$

for pp and  ${}^7\text{Be}$  this is approximately THE ANSWER.

$$f_1 \sim 69\% \text{ and } f_2 \sim 31\% \text{ and } \langle P_{ee} \rangle \approx 0.6$$

pp and  ${}^7\text{Be}$



$$f_1 \sim 69\%$$

$$f_2 \sim 31\%$$

$$\langle P_{ee} \rangle \approx 0.6$$

$$f_3 = \sin^2 \theta_{13} < 4\%$$

What about  $^8B$  ?

SNO's CC/NC

CC:  $\nu_e + d \rightarrow e^- + p + p$

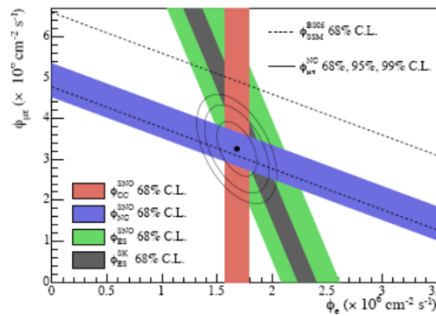
NC :  $\nu_x + d \rightarrow \nu_x + p + n$

ES:  $\nu_\alpha + e^- \rightarrow \nu_\alpha + e^-$

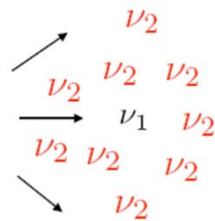
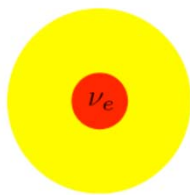
$$\frac{CC}{NC} = \langle P_{ee} \rangle = f_1 \cos^2 \theta_\odot + f_2 \sin^2 \theta_\odot$$

$$f_1 = \left( \frac{CC}{NC} - \sin^2 \theta_\odot \right) / \cos 2\theta_\odot$$

$$= (0.35 - 0.31) / 0.4 \approx 10 \%$$



$^8B$



$$f_2 \sim 90\%$$

$$f_1 \sim 10\%$$

$$\langle P_{ee} \rangle = \sin^2 \theta + f_1 \cos 2\theta_\odot \approx \sin^2 \theta_\odot = 0.31$$

Wow!!! How did that happen???

energy dependence!!!

