

Coupled Currents in Cosmic Strings with N carrier bosonic fields

M. L. , X. Martin, P. Peter, *Phys. Rev. D* (2009)

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Outline

General considerations

Current carrying strings: microscopic approach

Current carrying strings: worldsheet approach

Mechanical stability for arbitrary N

$N = 2$ case

Analytic approximation

Conclusions

General considerations

Generic production of cosmic (super-) strings

In the SB patterns of *most* supersymmetric grand unified theories

e.g., Jeannerot *et al.*, Phys. Rev. D **68** (2003)

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Cosmic string network evolution affects cosmology

Important to know what kind of strings one is dealing with.

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One or more currents

Microscopic approach

Current-carrying string described by

$$\mathcal{L} = \mathcal{L}_{\text{NG}} - \sum_i^N \left[\frac{1}{2} \partial_\mu \Phi_i^* \partial^\mu \Phi_i - \frac{g_i}{2} (|H|^2 - \eta^2) |\Phi_i|^2 - V(\Phi_i) \right] - \sum_{i \neq j}^N \mathcal{V}(\Phi_i, \Phi_j)$$

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Global U(1) complex scalar fields [Peter, 1992]

$$\left. \begin{array}{l} \text{straight,} \\ \text{static} \\ \text{config.} \end{array} \right\} \Phi_i(x^\mu) = \phi_i(r) e^{i\psi_i} \quad \text{and} \quad \psi_i = \omega_i t - k_i z$$

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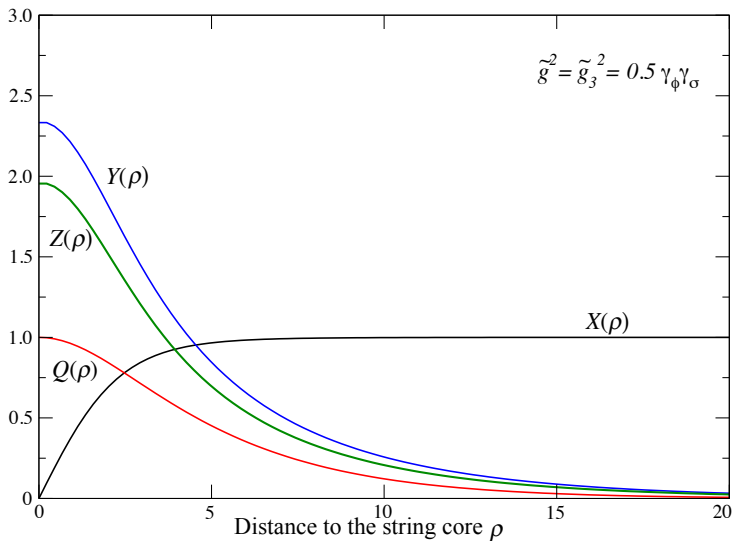
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2 topological invariants

$$\begin{array}{l} \text{EM flux} :: \Phi \\ = \int \mathcal{F}_{\mu\nu} d\sigma^{\mu\nu} \\ = -\frac{2\pi n}{e} \end{array} \quad \left. \begin{array}{l} \text{Phase change} \\ \text{around} \\ \text{closed loop} \\ \text{[Witten, 1985]} \end{array} \right\} \oint \frac{d\psi_i}{d\zeta} d\zeta = 2\pi m_i$$

Field profiles as a function of radial coord. for $N = 2$



$X(\rho)$:: Higgs field

$Q(\rho)$:: Gauge field associated with Higgs field

$Y(\rho)$ and $Z(\rho)$:: global U(1) fields

\tilde{g} :: cross-coupling constant

γ_i :: self-coupling constant

Worksheet approach

Worksheet Lagrangian [Carter, 1989]

$$\tilde{\mathcal{L}}(w_1, \dots, w_N) = 2\pi \int \mathcal{L} r dr \quad \text{with} \quad w_i = \partial^\mu \varphi_i \partial_\mu \varphi_i = k_i^2 - \omega_i^2$$

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Can be used to write down the EOS of the string...

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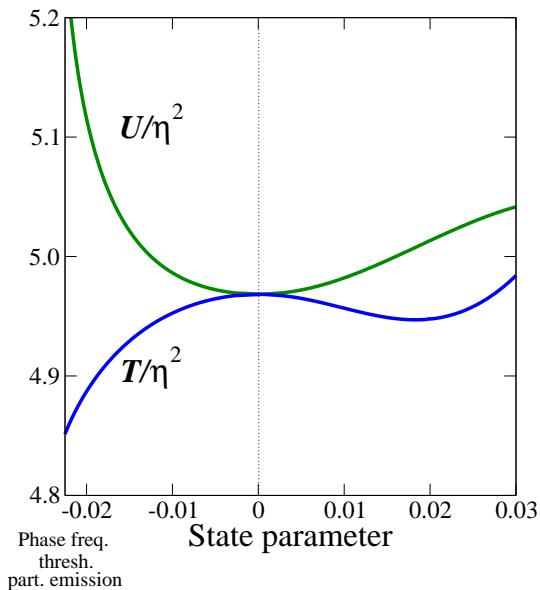
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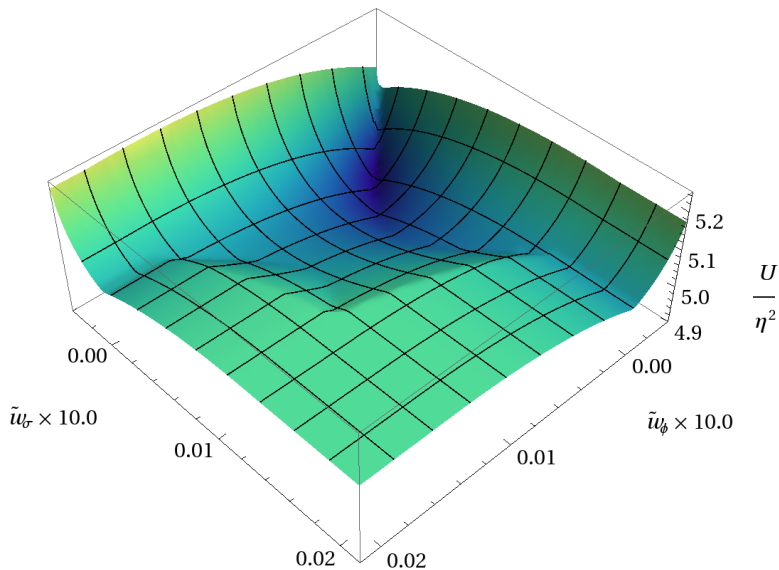
EOS for current-carrying strings: N arbitrary [new]

$$U - T = \sqrt{\mathcal{F} \left[(\mathcal{K}^{ii} \chi_{jj})^2, \chi_{ij}^2 \mathcal{K}_{ii} \mathcal{K}_{jj} \right]} \quad \text{with} \quad \begin{cases} \chi_{ij} = k_i k_j - \omega_i \omega_j \\ \mathcal{K}^{ij} = 2 \frac{\delta \tilde{\mathcal{L}}}{\delta \chi_{ij}} \end{cases}$$

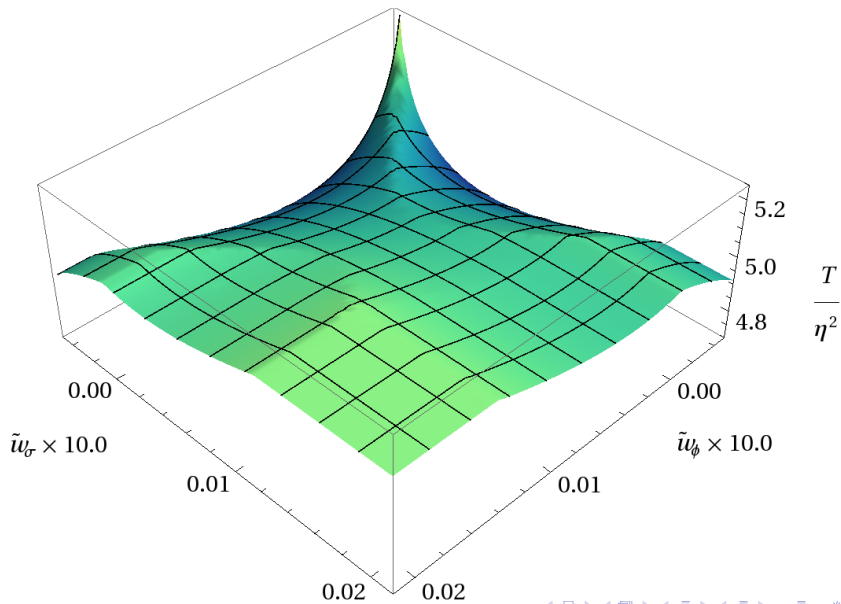
Energy and Tension, $N = 1$



Energy, $N = 2$ and $\chi_{12} = 0$



Tension, $N = 2$ and $\chi_{12} = 0$



Mechanical stability for $N = 1$ [Carter, 1989]

Perturb cons. equations and worldsheet integr. condition

1. Integrability of the worldsheet
2. Conservation of $\tilde{T}^{\mu\nu}$
3. Irrotationality condition of the phase gradients
4. *Current conservation*

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Perturbed quantities ($Q \rightarrow \chi_{ij}, \mathcal{K}_{ij} \dots$)

$$\text{Trans} :: \delta(Q) = e^{i(\omega_t t - k_t z)} \varepsilon(Q) \quad \text{Long} :: \delta(Q) = e^{i(\omega_\ell t - k_\ell z)} \varepsilon(Q)$$

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$$\text{disp. rel.} \in \mathbb{R} \quad :: \quad \frac{\omega_t^2}{k_t^2} = \frac{T}{U} \geq 0 \quad (\text{transverse wave})$$

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Longitudinal stability

$$\text{disp. rel.} \in \mathbb{R} \quad :: \quad \frac{\omega_\ell^2}{k_\ell^2} = -\frac{dT}{dU} \geq 0 \quad (\text{acoustic wave})$$

Mechanical stability for arbitrary N

Transverse stability

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Note: transverse perturbations decouple from perturbations in χ_{ij} and \mathcal{K}_{ij}

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Longitudinal stability

Homogeneous system of $2N$ coupled equations

\Rightarrow Obtain $\frac{\omega_\ell}{k_\ell}$ from $|\dots| = 0$ and impose $\frac{\omega_\ell}{k_\ell} \in \mathbb{R}$

sets bounds on the χ_{ij} 's

$N = 2$ case

Transverse stability

$$\frac{T}{U} \geq 0 \quad \Rightarrow \quad \chi_{12}^2 \leq \frac{1}{\mathcal{K}_{11}\mathcal{K}_{22}} \left[\mathcal{A} - \left(\frac{1}{2}\chi_{11}\mathcal{K}_{11} - \frac{1}{2}\chi_{22}\mathcal{K}_{22} \right) \right]$$

where $\mathcal{A} = 2\pi \int \mathcal{F}(H, A_\mu, \Phi_i, \dots) r dr$.

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Longitudinal stability

$$c_0 + c_1\xi + c_2\xi^2 + c_3\xi^3 + c_4\xi^4 = 0 \quad \text{where} \quad \begin{cases} \xi = \mathcal{G}(\omega_\ell/k_\ell) \\ c_i = \mathcal{H}(\mathcal{K}_{ij}, \chi_{ij}, \dots) \end{cases}$$

Suffices to require $\xi \in \mathbb{R} \Rightarrow$ sets bounds on the χ_{11} , χ_{22} and χ_{12}

compute U , T , \mathcal{K}_{ij} , ξ , etc. from field profiles obtained numerically

Analytic approximation [Carter and Peter, 1995]

$$\tilde{\mathcal{L}}_i(w_i) = -m_h^2 - \frac{m_{\Phi_i}^2}{2} \ln \left(1 + \frac{w_i}{m_{\Phi_i}^2} \right) \quad w_i < 0$$

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Neglects

$$\sum_{i \neq j}^N \mathcal{V}(\Phi_i, \Phi_j)$$

Expected to be a good approx. for $\tilde{g}_{ij} \ll 1$

Mechanical stability of approx. model

Transverse stability

$$T/U \geq 0$$

$$w_1 w_2 \geq 0 \quad \because \quad \chi_{12}^2 \leq w_1 w_2 \frac{(U_1 - m^2 + T_2)(U_2 - m^2 + T_1)}{(U_1 - T_1)(U_2 - T_2)}$$

$$w_1 w_2 \leq 0 \quad \because \quad \chi_{12}^2 \leq |w_1 w_2| \frac{(T_1 + T_2 - m^2)(U_1 + U_2 - m^2)}{(U_1 - T_1)(U_2 - T_2)}$$

Longitudinal stability

All perturbation modes decouple $\Rightarrow -dT/dU \geq 0$

$$-m_i^2 \leq w_i \leq \frac{m_i^2}{3}$$

Has been shown to agree with the numerical results when $\tilde{g} \ll 1$

Conclusions

- ▶ Cosmic (super-)strings seem to be inevitable
 - ▶ They come with various attributes (intercommutation probability, mechanical stability, *etc.*)
 - ▶ Important to develop tools in order to deal with these \neq types of strings
-
- ▶ Focused on mechanical stability of strings
 - ▶ Extended existing formalism to N current carrier fields
 - ▶ Checked against fully analytic approximation