

A TeV scale messenger (e.g. 750 GeV) of Dark Matter

Andi Hektor
(NICPB, Tallinn, Estonia)

[in collaboration with L. Marzola, M. Raidal, S. Di
Chiara, K. Kannike]

CERN, TeVPA
Sept 16, 2016

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- From the effective WIMP to the simplified WIMP
- Two ways to escape from Direct Detection
 - (i) ‘Coy DM’
 - (ii) ‘Resonance Portal’
- ‘Coy DM’
 - Gamma-ray line signal from Coy DM
 - A pseudoscalar case of TeV-scale mediator
- ‘Resonance portal’
 - Life in the resonance zone is interesting — velocity dependent annihilation signal!

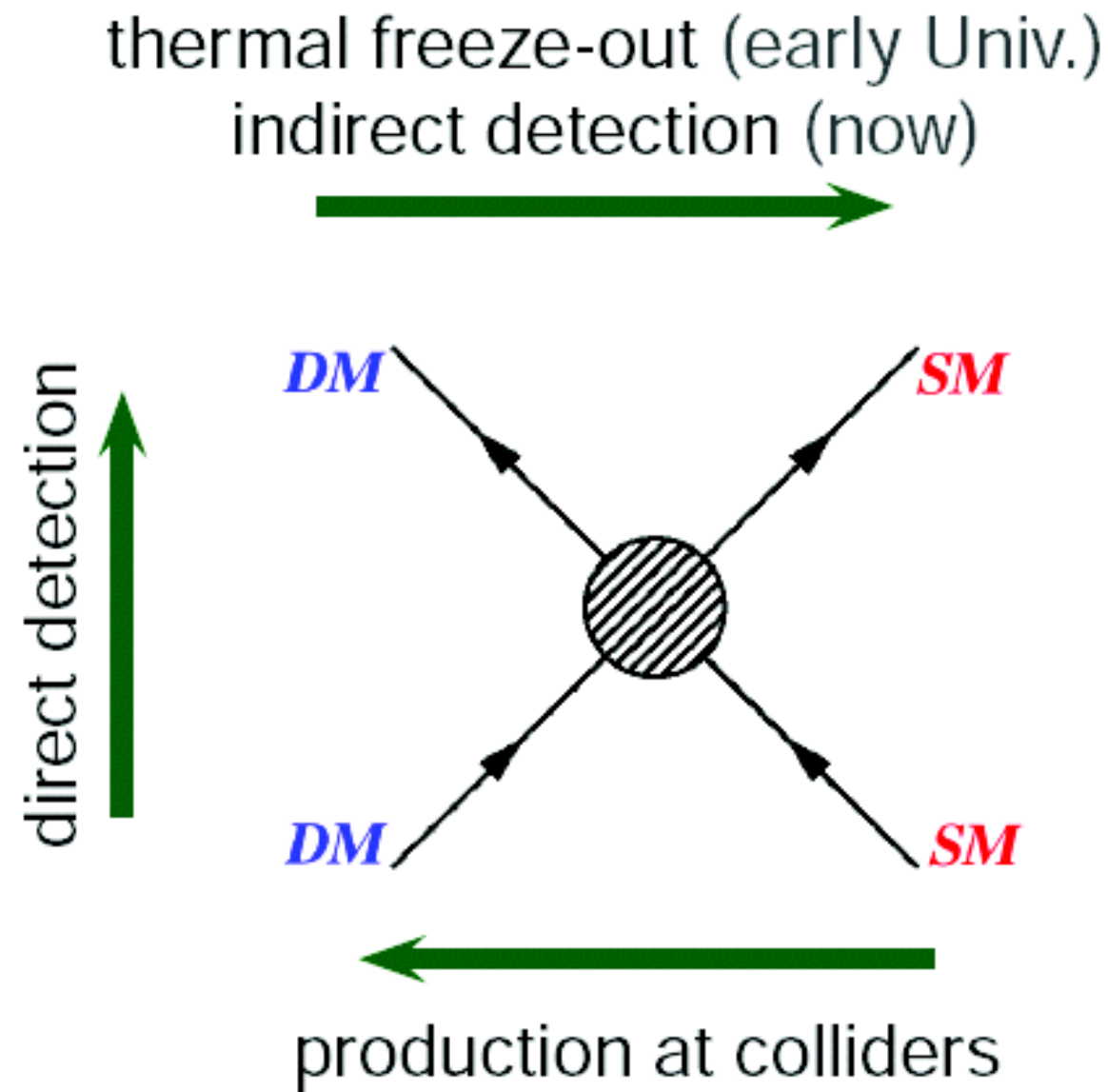
Based on:

1403.3401, 1507.05096, 1602.00004, 1603.07263

& two appearing in some next weeks



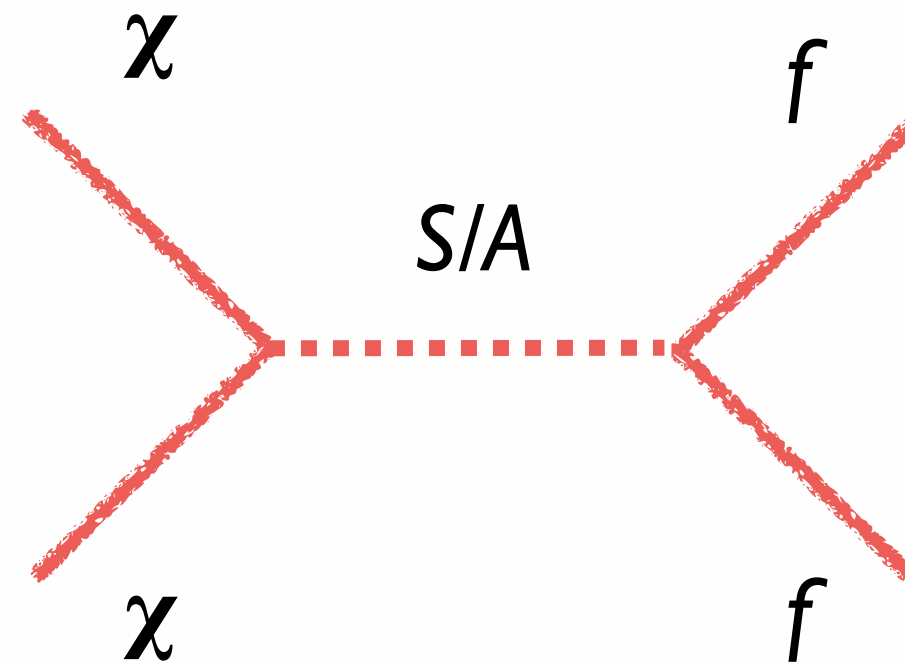
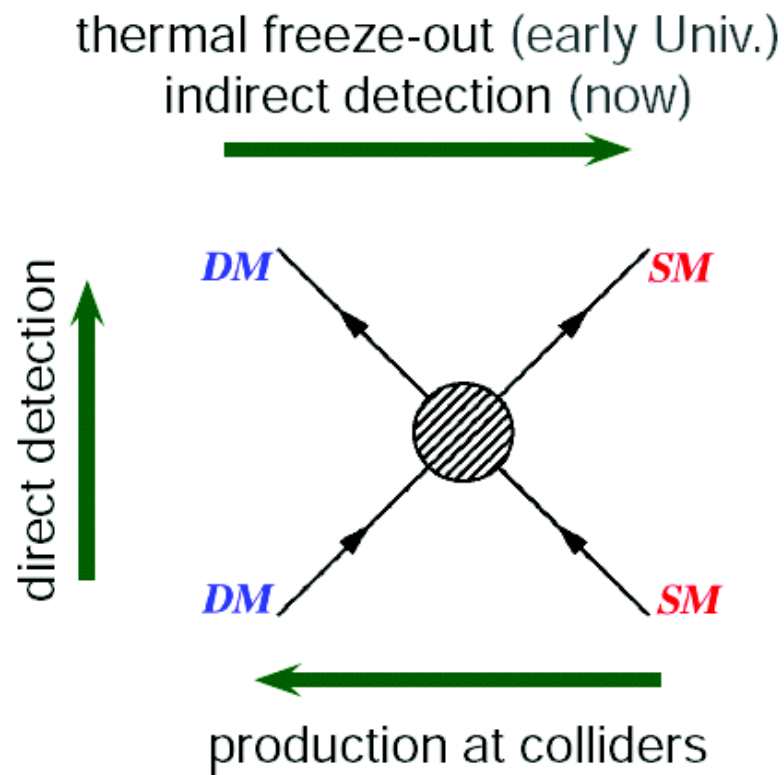
Effective WIMP



$$\mathcal{L}_{\text{EFT}} = \frac{1}{M_*^2} (\bar{q}q) (\bar{\chi}\chi)$$



From the effective to the simplified DM

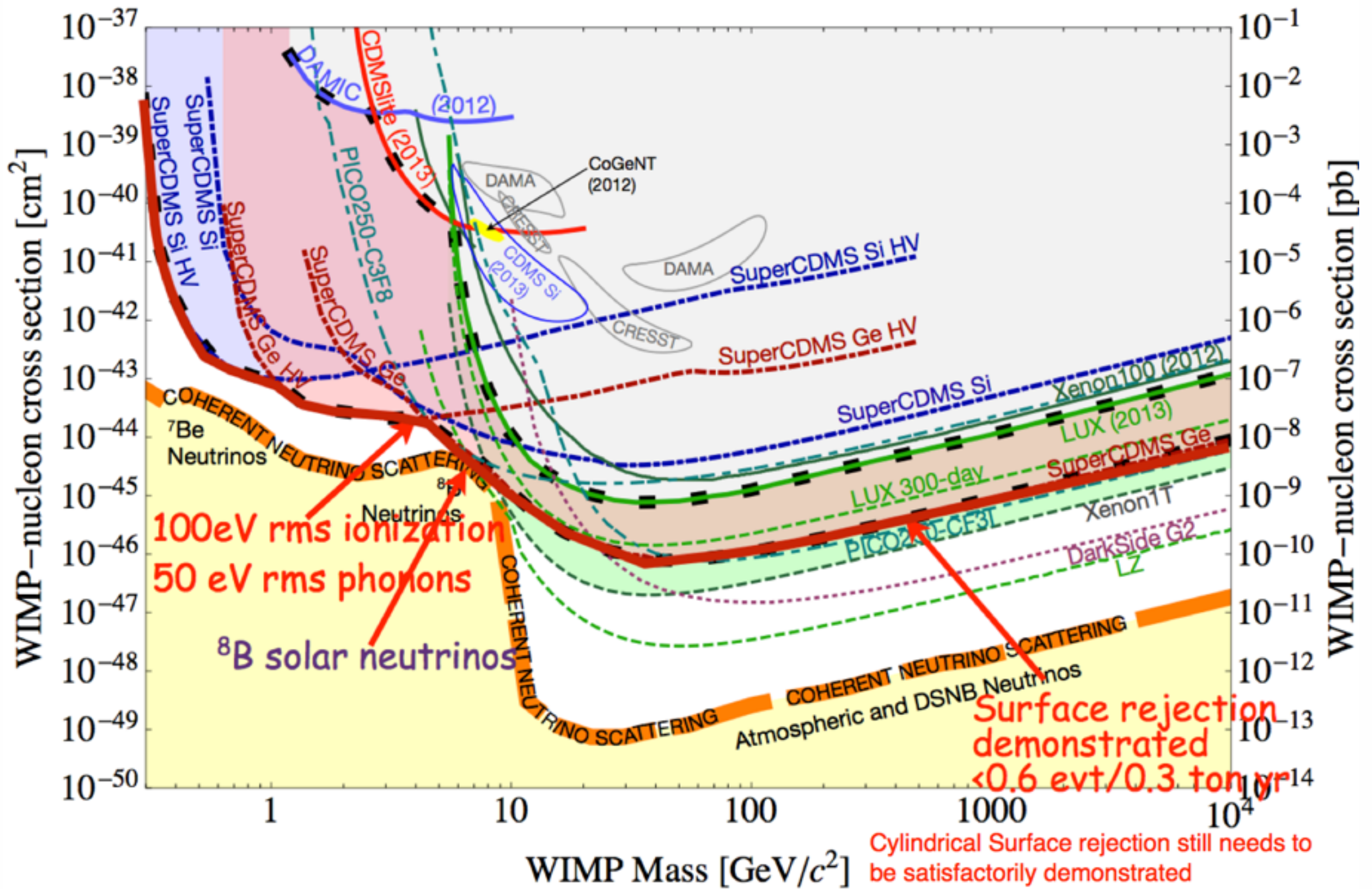


$$\mathcal{L}_{\text{EFT}} = \frac{1}{M_*^2} (\bar{q}q) (\bar{\chi}\chi)$$

$$\mathcal{L}_{0_S s \frac{1}{2}} = \frac{1}{2} (\partial_\mu S)^2 - \frac{1}{2} m_S^2 S^2 + \bar{\chi} (i \not{\partial} - m_\chi) \chi - g_\chi S \bar{\chi} \chi - g_{\text{SM}} S \sum_f \frac{y_f}{\sqrt{2}} \bar{f} f$$

$$\mathcal{L}_{0_A s \frac{1}{2}} = \frac{1}{2} (\partial_\mu A)^2 - \frac{1}{2} m_A^2 A^2 + \bar{\chi} (i \not{\partial} - m_\chi) \chi - i g_\chi A \bar{\chi} \gamma^5 \chi - i g_{\text{SM}} A \sum_f \frac{y_f}{\sqrt{2}} \bar{f} \gamma^5 f$$

The model killer, Direct Detection





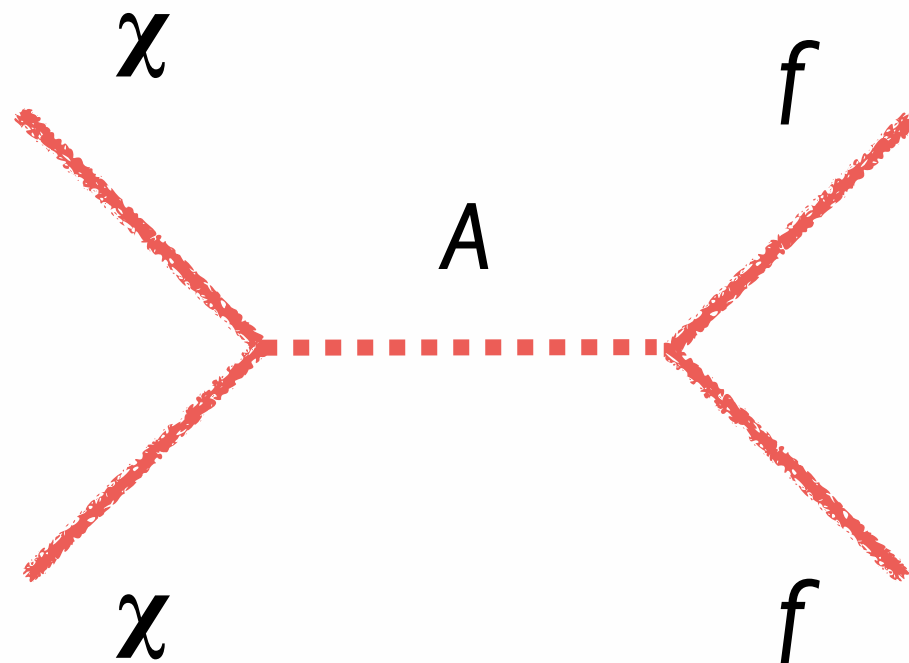
Escaping from the killer Direct Detection:

A. ‘Coy DM’

B. ‘Resonance Portal’



Coy DM



$$\sigma_{SI} \propto v^2$$

$$\mathcal{L}_{0As\frac{1}{2}} = \frac{1}{2}(\partial_\mu A)^2 - \frac{1}{2}m_A^2 A^2 + \bar{\chi}(i\not{\partial} - m_\chi)\chi - ig_\chi A \bar{\chi}\gamma^5\chi - ig_{SM}A \sum_f \frac{y_f}{\sqrt{2}} \bar{f}\gamma^5 f$$

Bohm et al, [1401.6458](#); Berlin et al, [1404.0022](#); Dolan et al, [1412.5174](#)

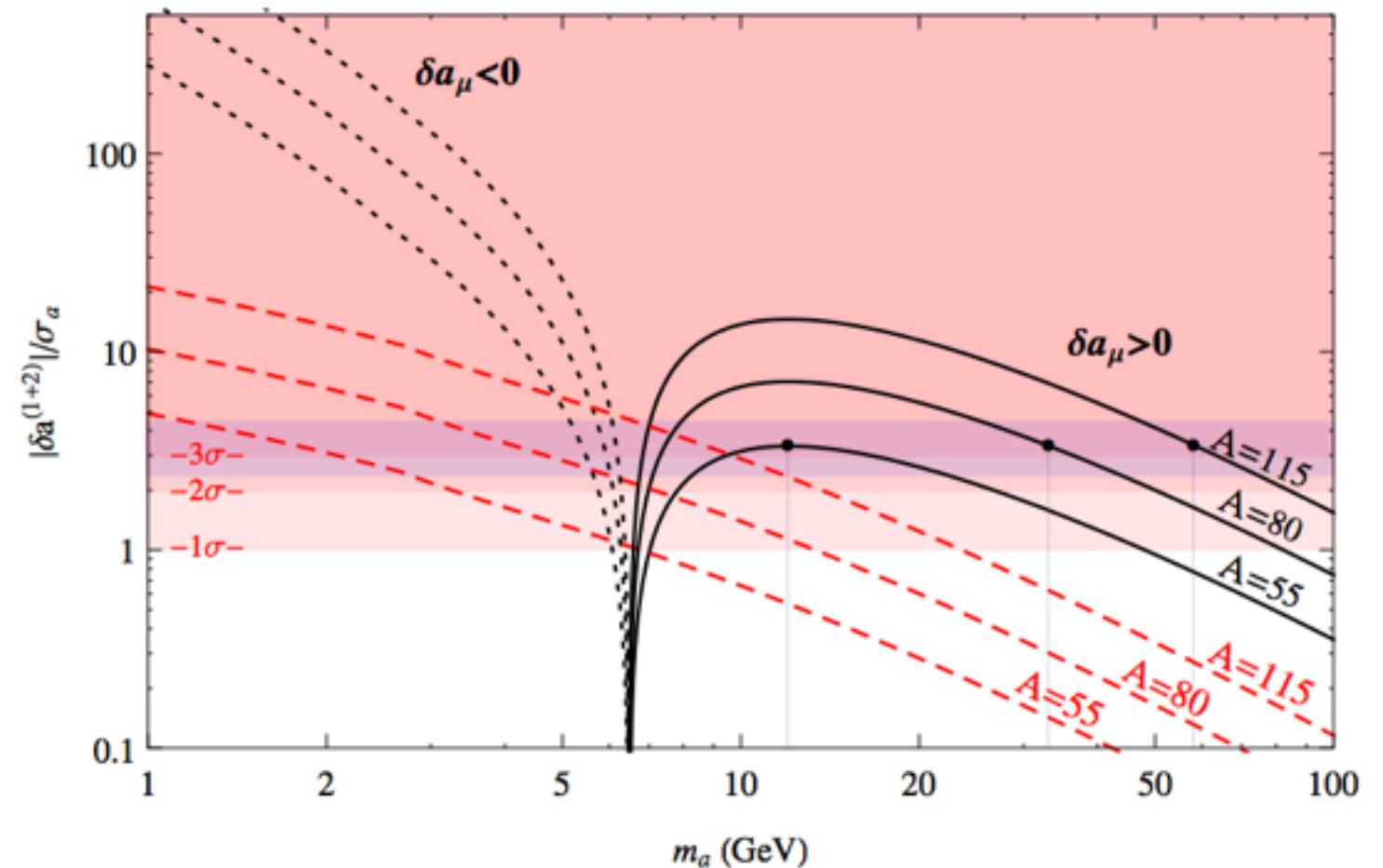
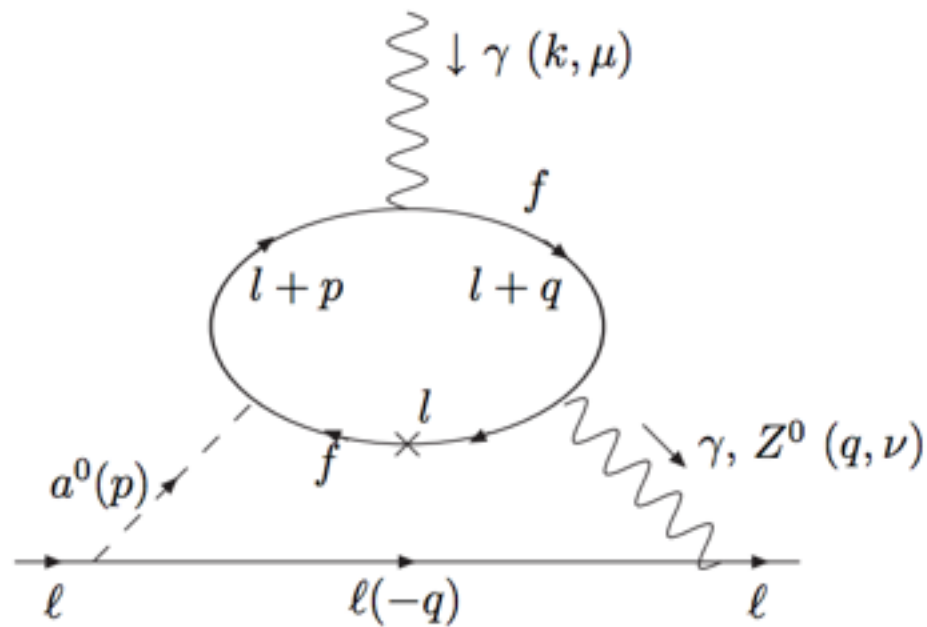


Coy DM has surprisingly many features...

- Self-interacting DM, DAMA/LIBRA etc
- Muon $g-2$
- γ -ray line



Coy DM & muon g-2



Chang et al, hep-ph/0009292

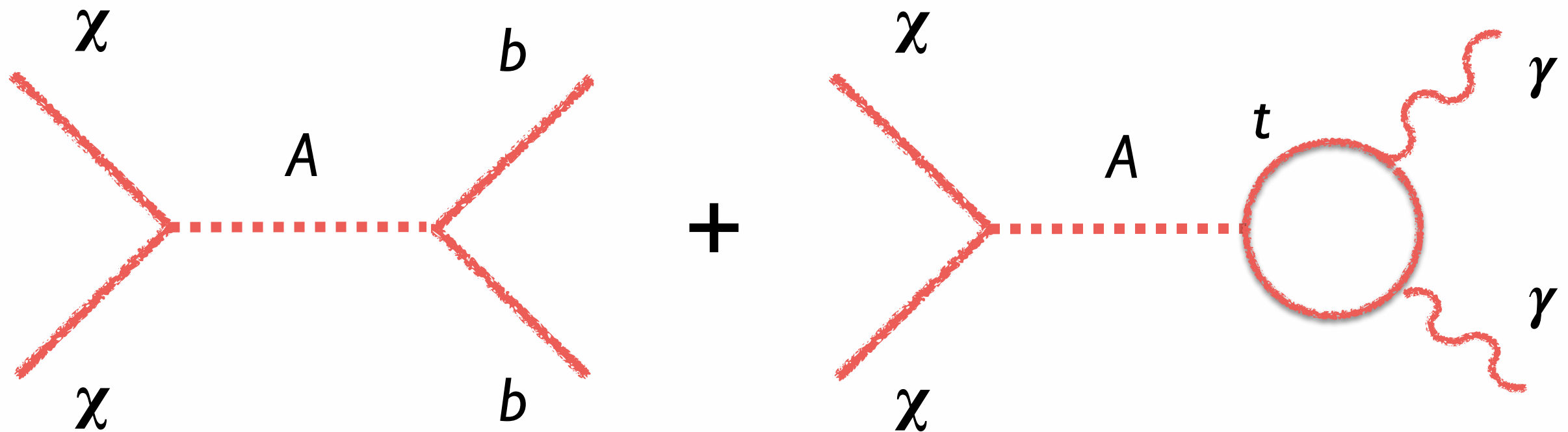
‘Leptophilic’ Coy DM,
Hektor & Marzola, 1403.3401



Coy DM & γ -ray line



Coy DM & γ -ray line

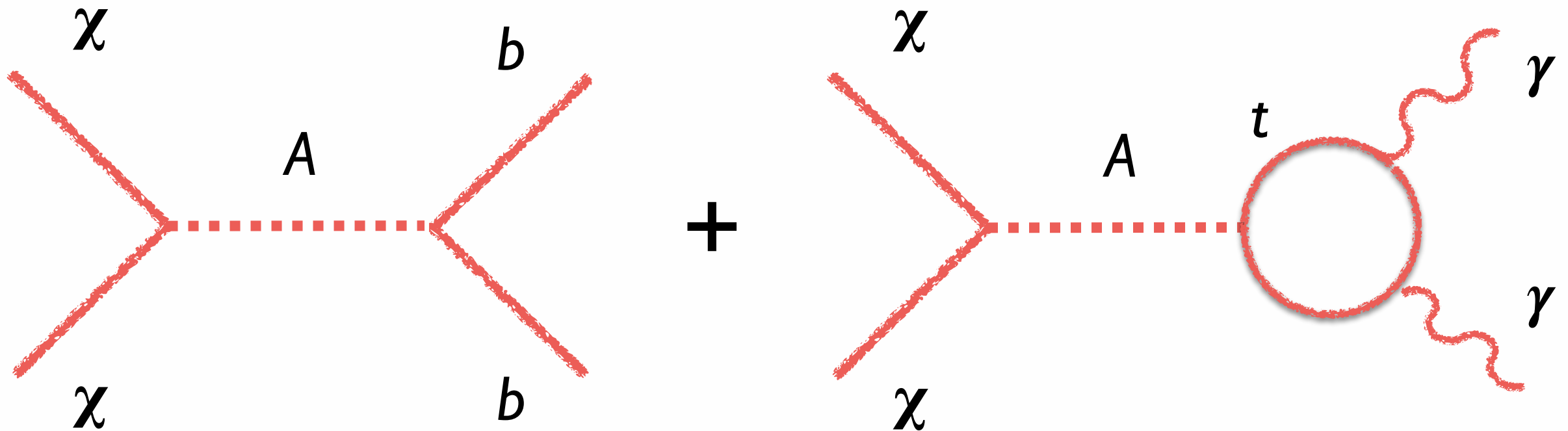


$$m_\chi < m_t$$

$$\mathcal{L}_{0As\frac{1}{2}} = \frac{1}{2}(\partial_\mu A)^2 - \frac{1}{2}m_A^2 A^2 + \bar{\chi}(i\not{\partial} - m_\chi)\chi - ig_\chi A \bar{\chi}\gamma^5\chi - ig_{\text{SM}}A \sum_f \frac{y_f}{\sqrt{2}} \bar{f}\gamma^5 f$$



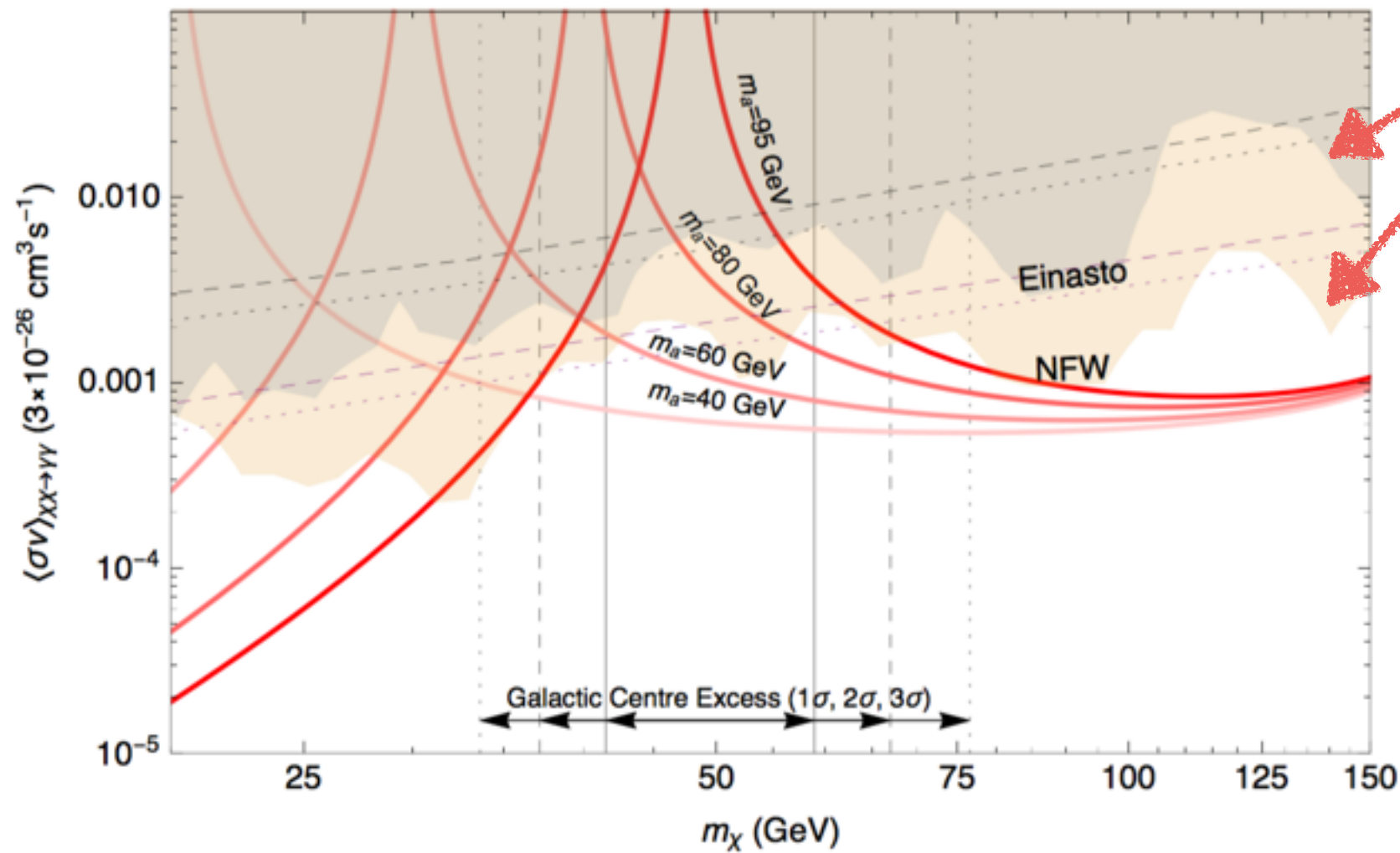
Coy DM & γ -ray line



- Loop suppression, $\sim 10^{-2} \dots 10^{-3}$
- Sensitivity of line/broader distribution search, $\sim 10^2$



γ -ray line signal alone

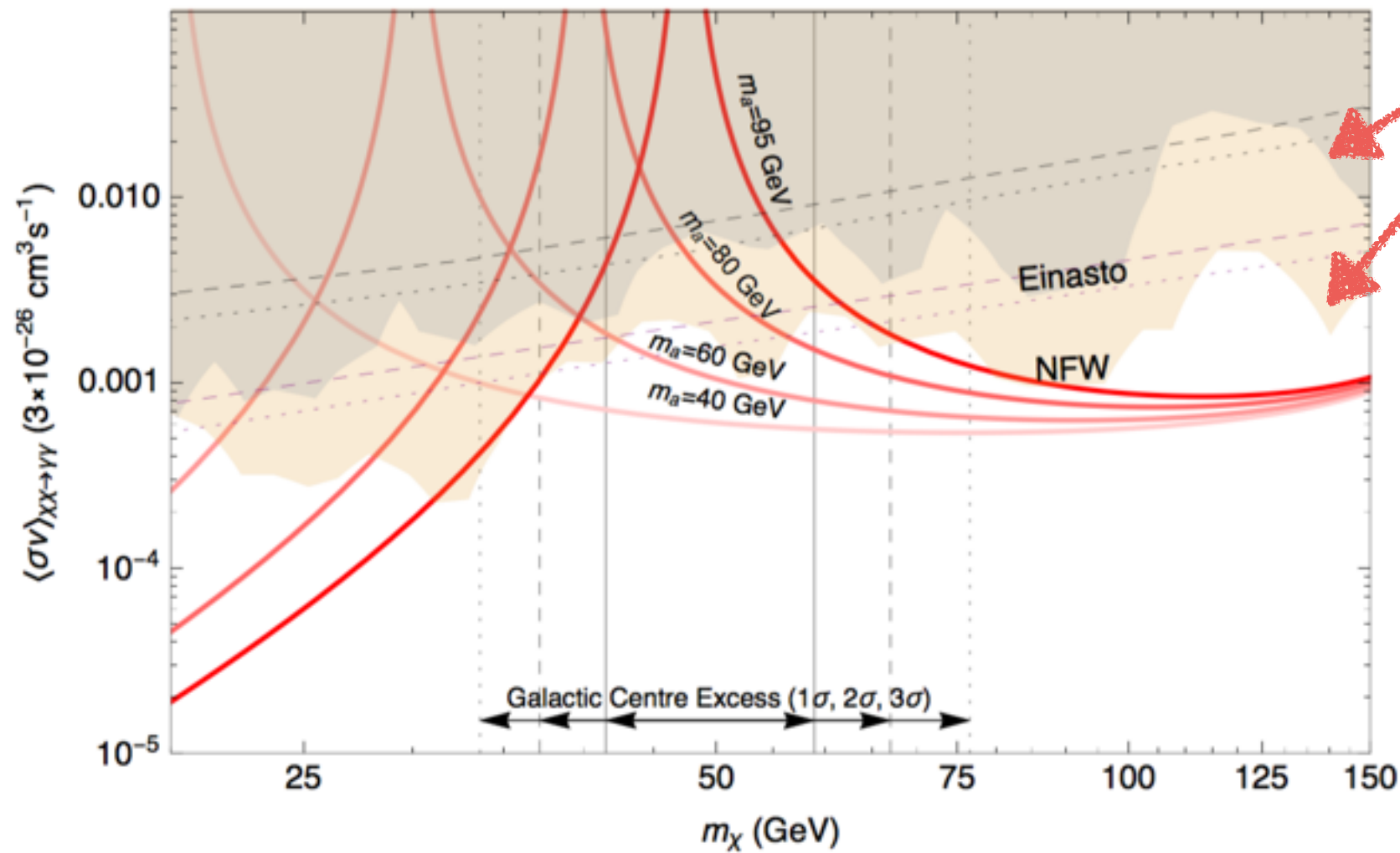


Fermi-LAT
1506.00013

$$\sigma_{\chi\chi\rightarrow\gamma\gamma} = \frac{1}{512\pi^3} \frac{(g_\chi g_f \alpha Q_f N_c^f)^2 (2m_\chi)^4}{(m_a^2 - 4m_\chi^2)^2 + m_a^2 \Gamma_a^2} \left| \frac{A_1(\tau_f)}{m_f} \right|^2$$



γ -ray line signal alone



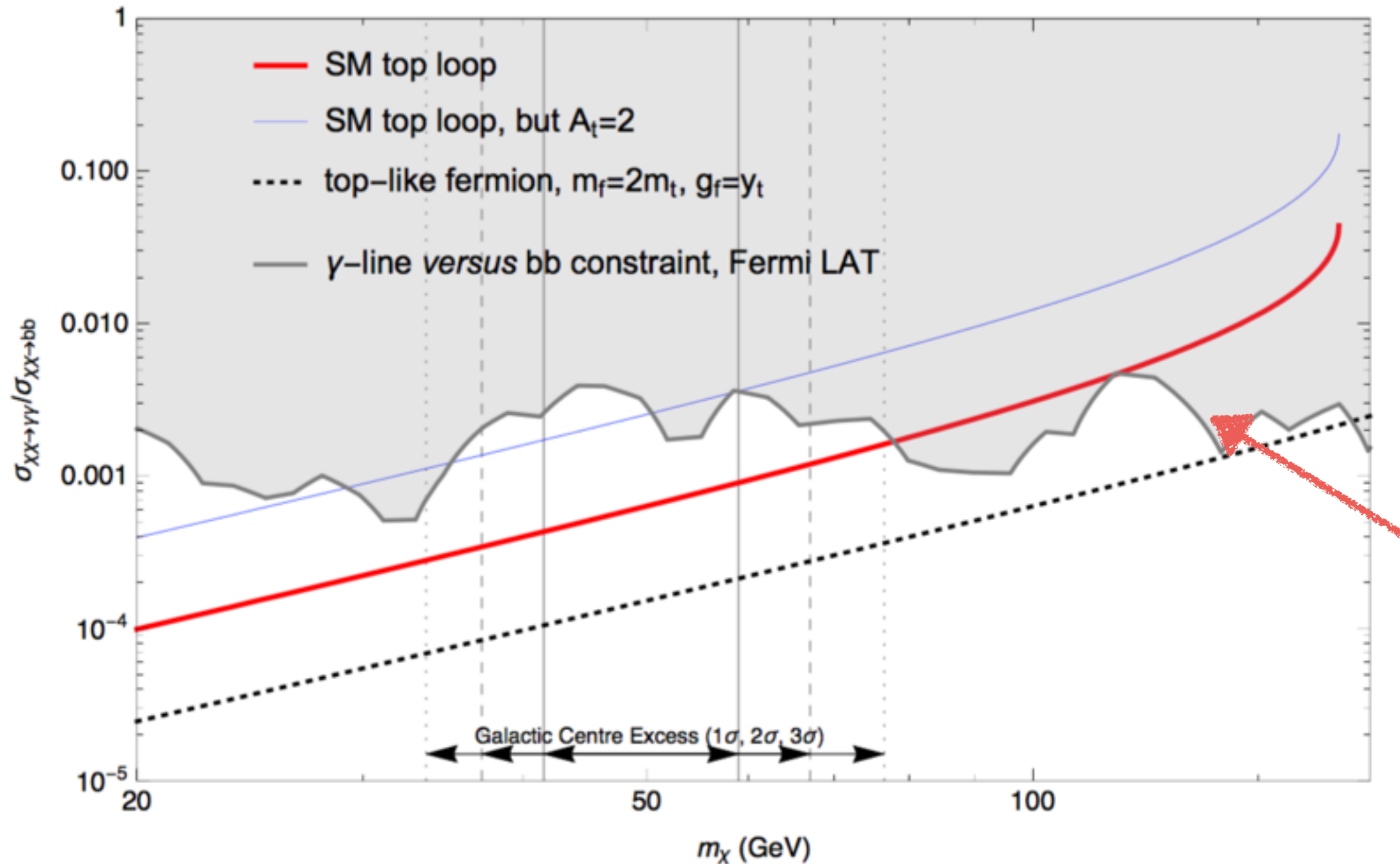
Fermi-LAT
1506.00013

$$\sigma_{\chi\chi\rightarrow\gamma\gamma} = \frac{1}{512\pi^3} \frac{(g_\chi g_f \alpha Q_f N_c^f)^2 (2m_\chi)^4}{(m_a^2 - 4m_\chi^2)^2 + m_a^2 \Gamma_a^2} \left| \frac{A_1(\tau_f)}{m_f} \right|^2$$

$$\sigma_{\chi\chi\rightarrow b\bar{b}} = \frac{N_c (g_\chi g_b)^2}{8\pi} \frac{m_\chi^2 \sqrt{1 - \frac{m_b^2}{m_\chi^2}}}{(m_a^2 - 4m_\chi^2)^2 + m_a^2 \Gamma_a^2}$$



Ratio of the γ -ray line and bb signal



Fermi-LAT
1506.00013,
1503.02641

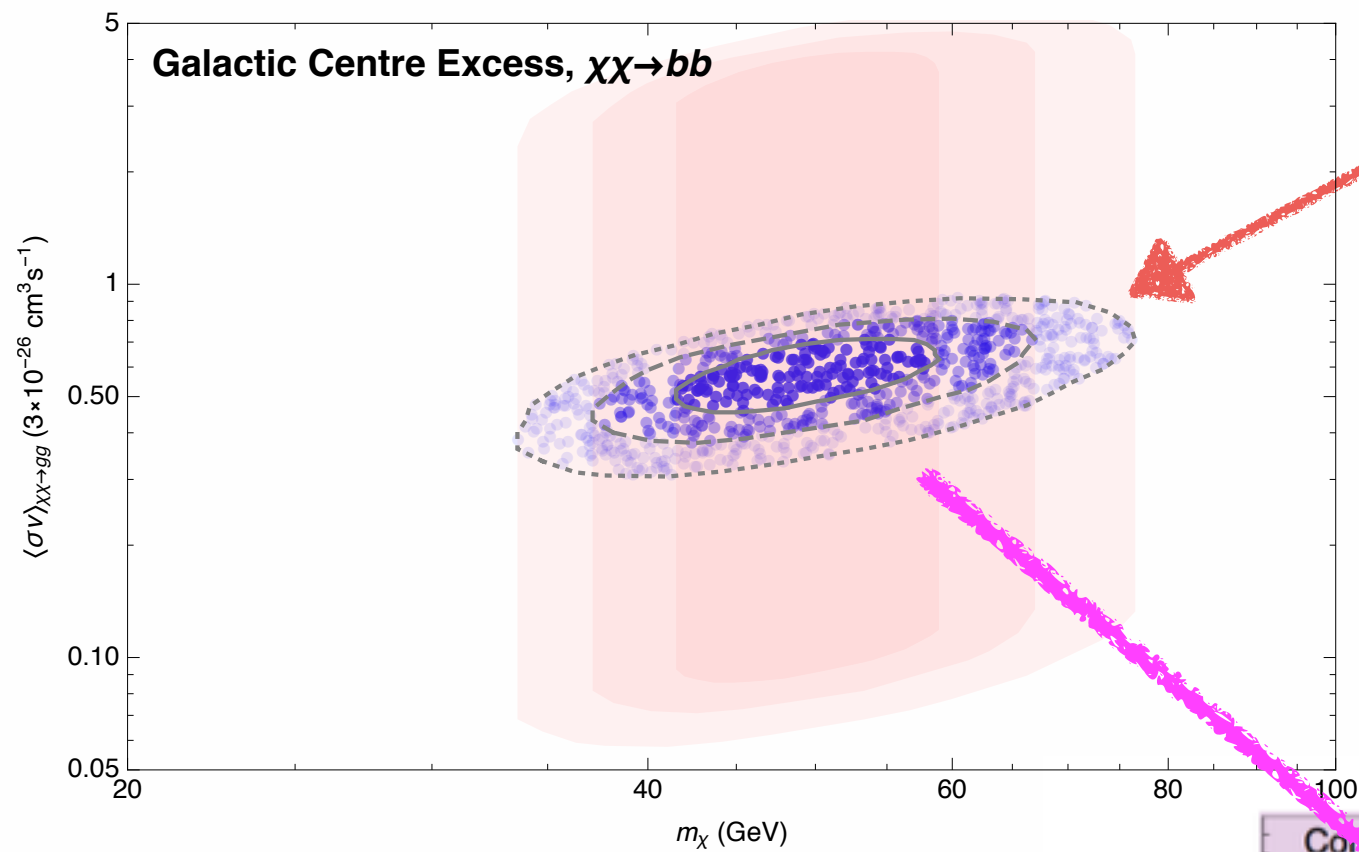
$$R = \frac{(N_c^t \alpha Q_t y_t)^2}{2\pi^2 N_c^b y_b^2} \frac{m_t^2}{m_\chi^2 \sqrt{1 - \frac{m_t^2}{m_\chi^2}}} \left| \arcsin^2\left(\frac{m_\chi}{m_t}\right) \right|^2$$



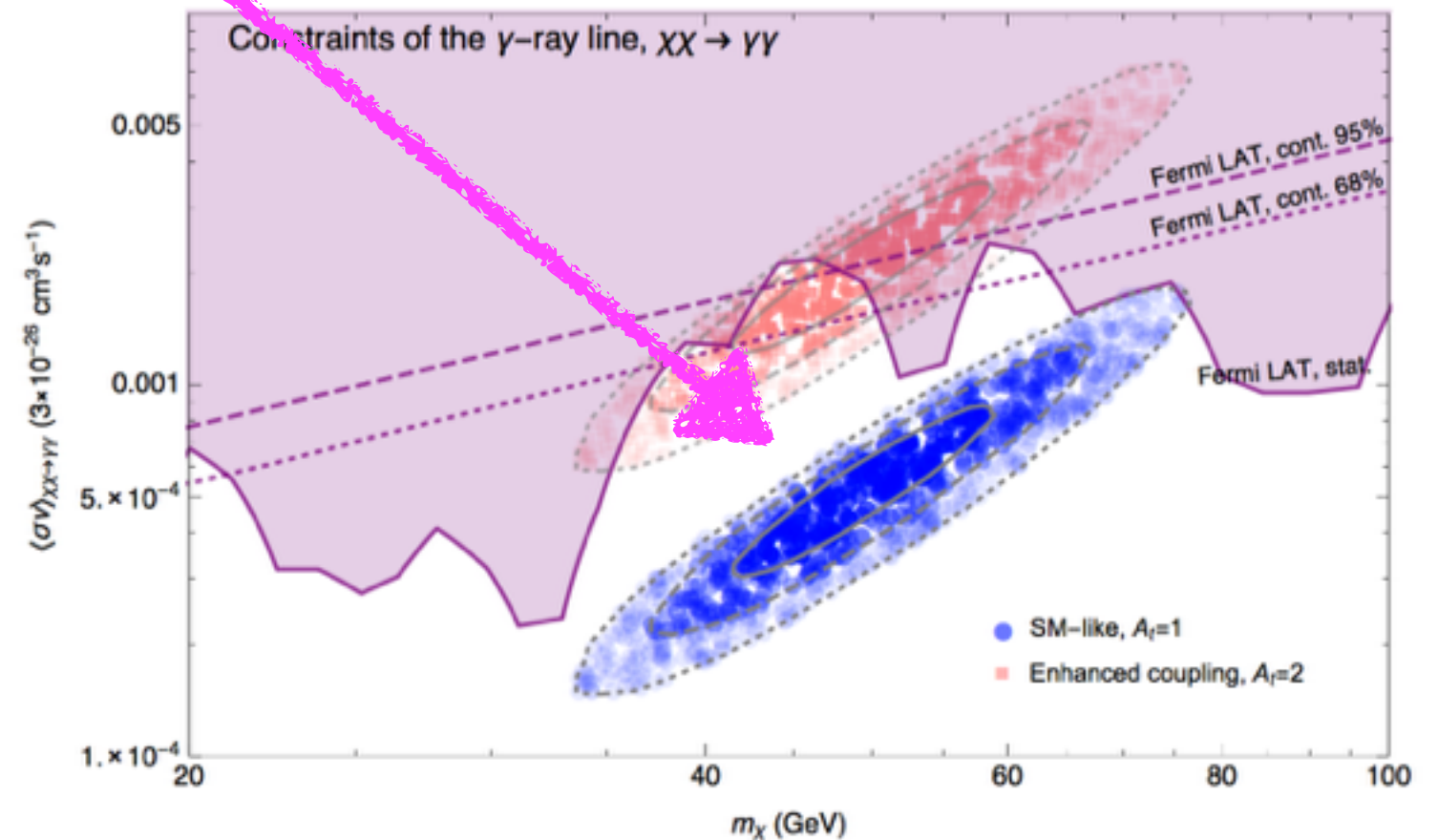
The γ -ray line and the Galactic Centre Excess

Calore et al,

1411.4647

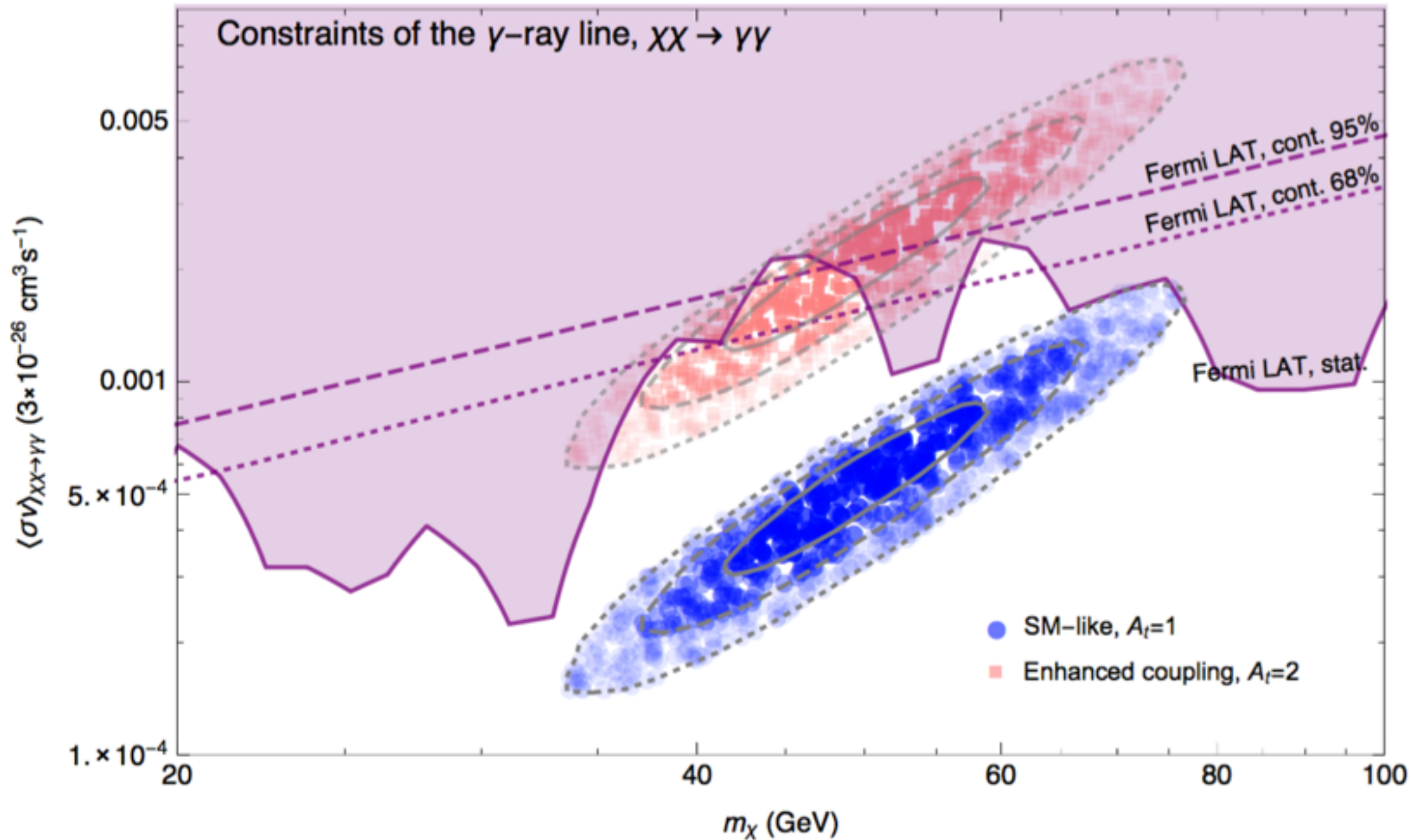


$$R = \frac{(N_c^t \alpha Q_t y_t)^2}{2\pi^2 N_c^b y_b^2} \frac{m_t^2}{m_\chi^2 \sqrt{1 - \frac{m_t^2}{m_\chi^2}}} \left| \arcsin^2\left(\frac{m_\chi}{m_t}\right) \right|^2$$





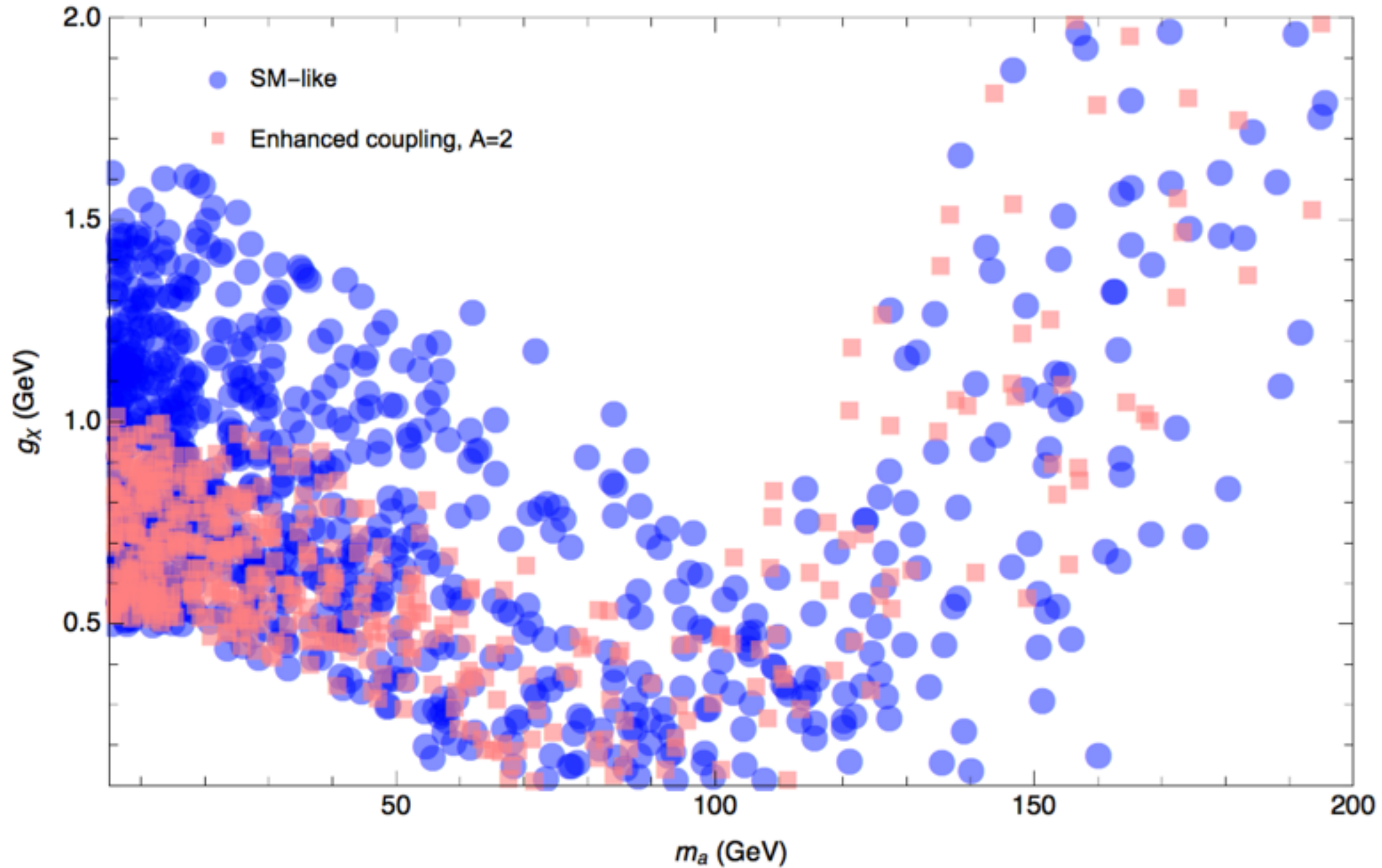
The γ -ray line and the Galactic Centre Excess



$$R = \frac{(N_c^t \alpha Q_t y_t)^2}{2\pi^2 N_c^b y_b^2} \frac{m_t^2}{m_\chi^2 \sqrt{1 - \frac{m_t^2}{m_\chi^2}}} \left| \arcsin^2\left(\frac{m_\chi}{m_t}\right) \right|^2$$



The γ -ray line and the Galactic Centre Excess



$$R = \frac{(N_c^t \alpha Q_t y_t)^2}{2\pi^2 N_c^b y_b^2} \frac{m_t^2}{m_\chi^2 \sqrt{1 - \frac{m_t^2}{m_\chi^2}}} \left| \arcsin^2\left(\frac{m_\chi}{m_t}\right) \right|^2$$



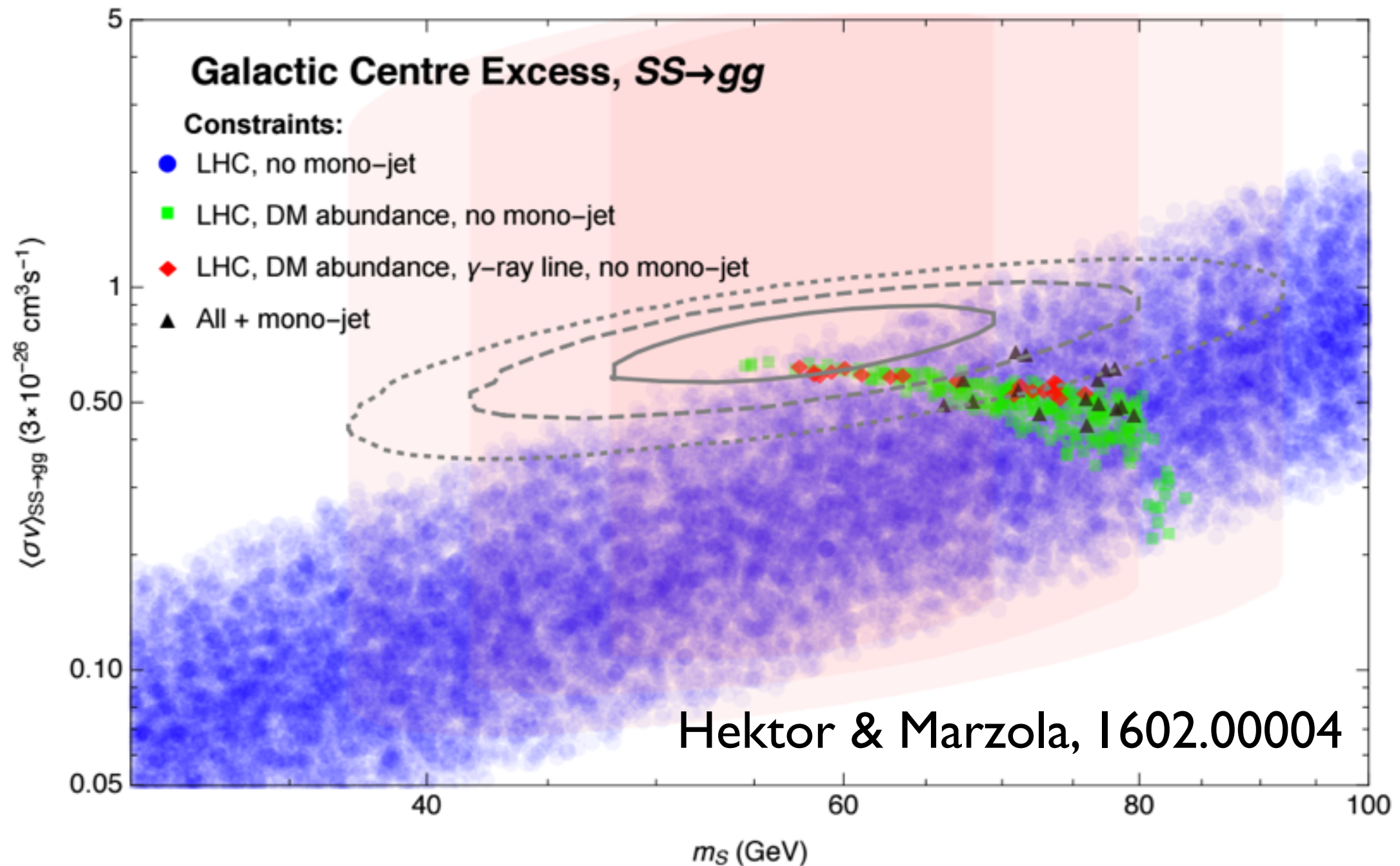
Galactic Centre Excess & TeV-scale pseudoscalar mediator

or

What did we learn from the 750 GeV pseudo-particle?



Coy DM & Galactic Centre Excess & TeV-scale mediator



- Pseudoscalar mediator, $m_A = 750 \text{ GeV}$
- Scalar DM particle



Life in 'Resonance portal'



Escaping from the killer Direct Detection:

- A. 'Coy DM'
- B. 'Resonance Portal'

Is the Resonance Portal fine tuning?

1. Yes
2. No. There can be theoretical motivation, e.g. hep-ph/9207234, hep-ph/9704403, hep-ph/9804231



RP offers interesting phenomenology for indirect section

$$\sigma_{\chi\chi \rightarrow S \rightarrow ff} = \frac{N_c (g_\chi A y_f)^2}{16 \pi} \frac{m^2 \left(1 - \frac{m_f^2}{m^2}\right)^{3/2}}{(M^2 - s[m^2, v_{\text{rel}}])^2 + M^2 \Gamma^2} v_{\text{rel}}$$

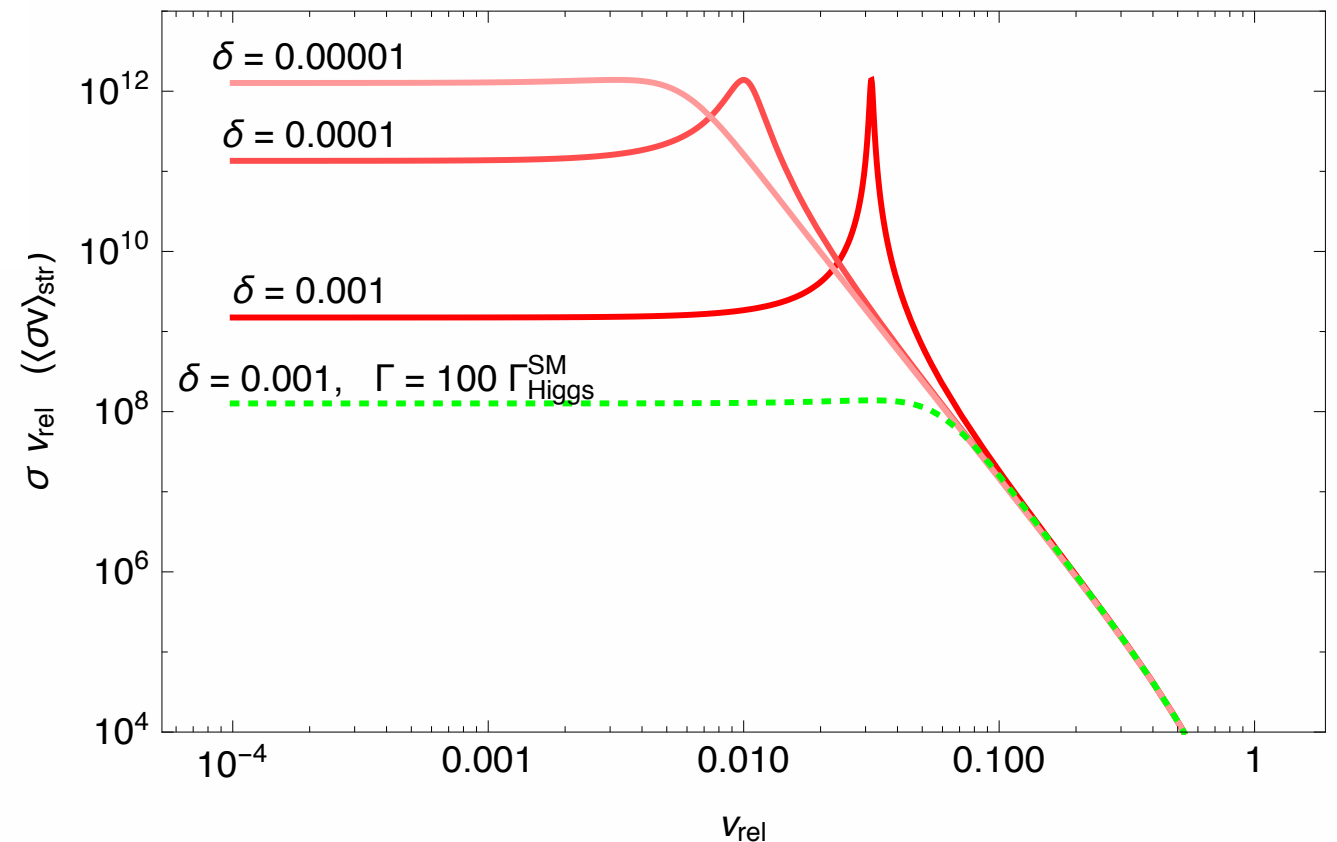
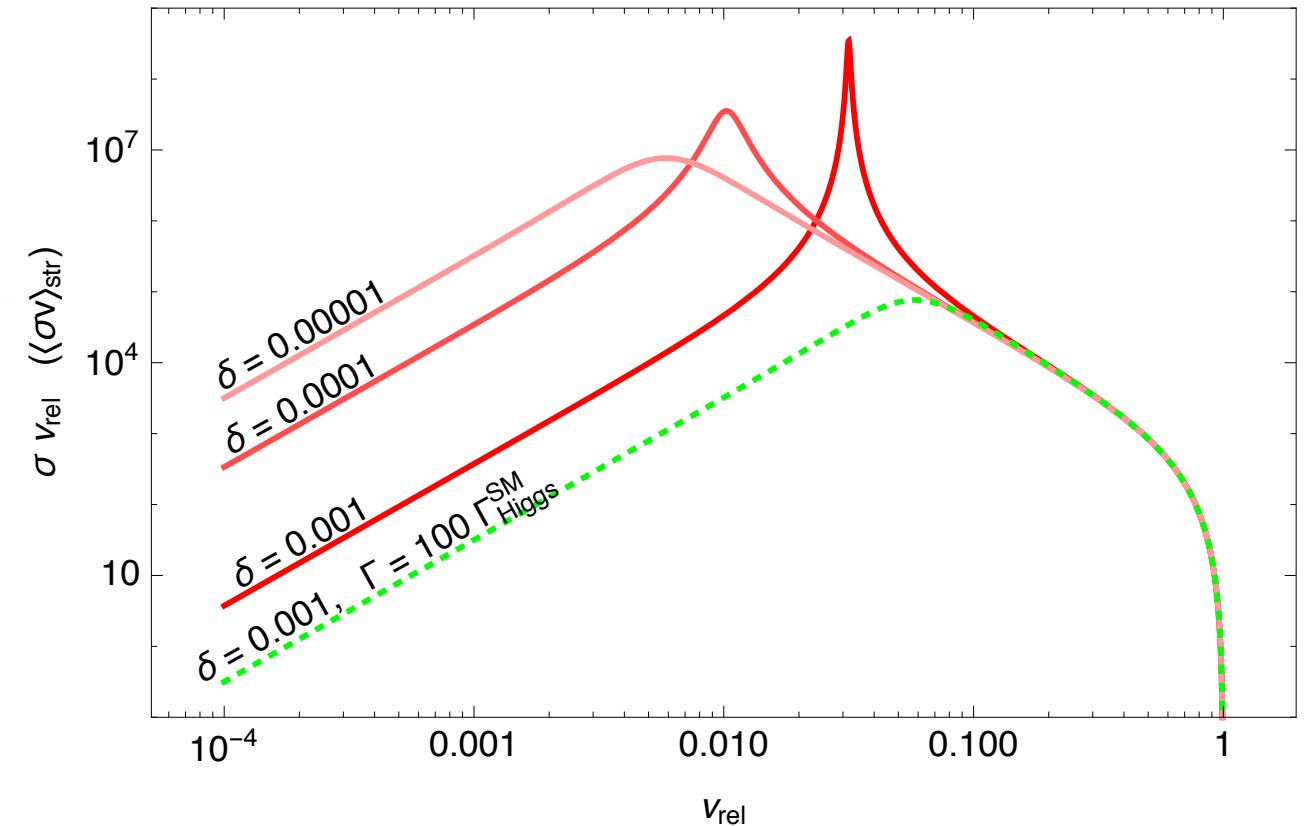
$$\sigma_{\chi\chi \rightarrow A \rightarrow ff} = \frac{N_c (g_\chi A y_f)^2}{16 \pi} \frac{m^2 \left(1 - \frac{m_f^2}{m^2}\right)^{1/2}}{(M^2 - s[m^2, v_{\text{rel}}])^2 + M^2 \Gamma^2} v_{\text{rel}}^{-1}$$

$$m^2 = \frac{M^2}{4(1 + \delta)}, \quad s[m^2, v_{\text{rel}}] = \frac{4 m^2}{1 - v_{\text{rel}}^2}$$

$$M = m_{\text{higgs}}, \quad g_\chi = 1$$

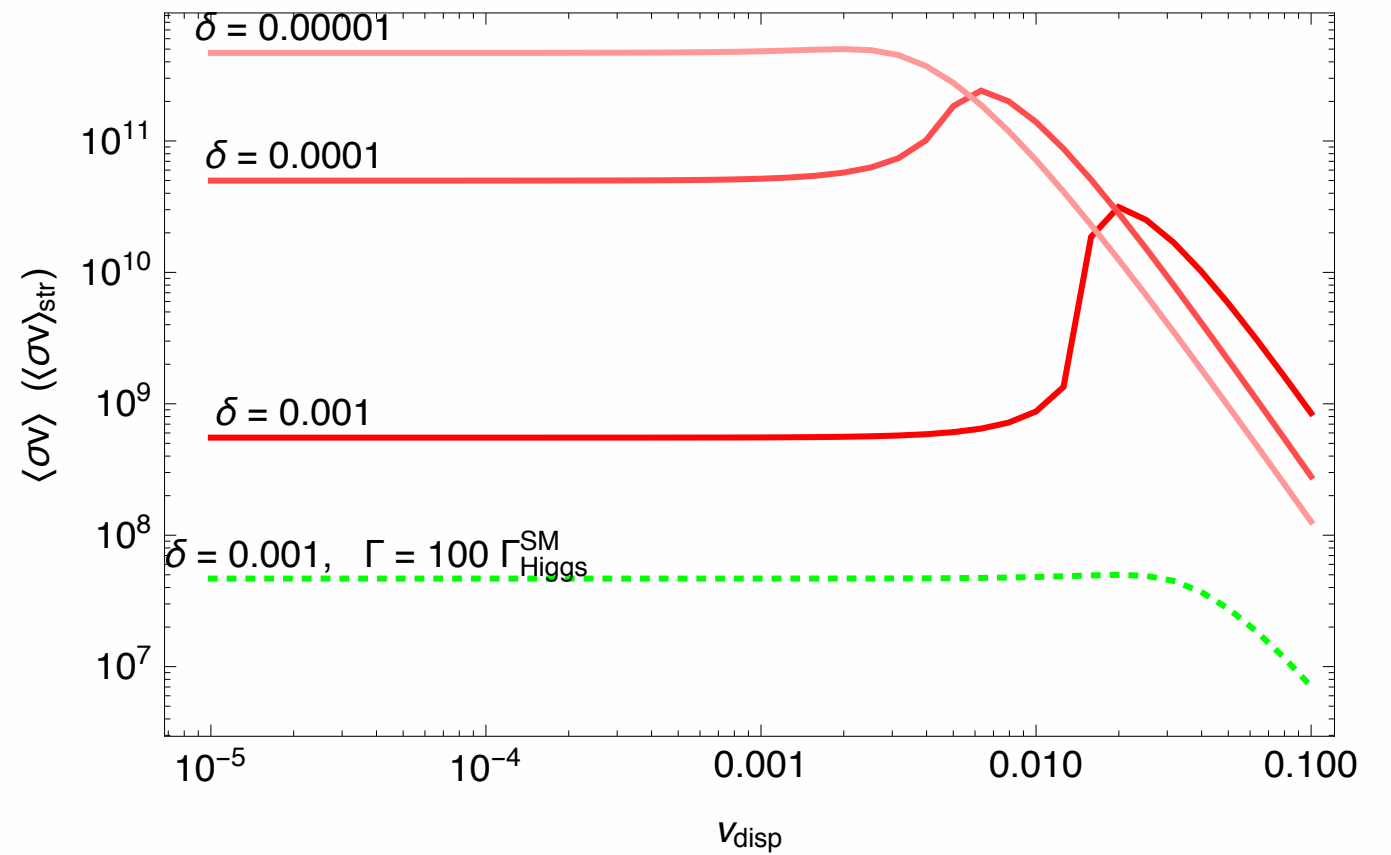
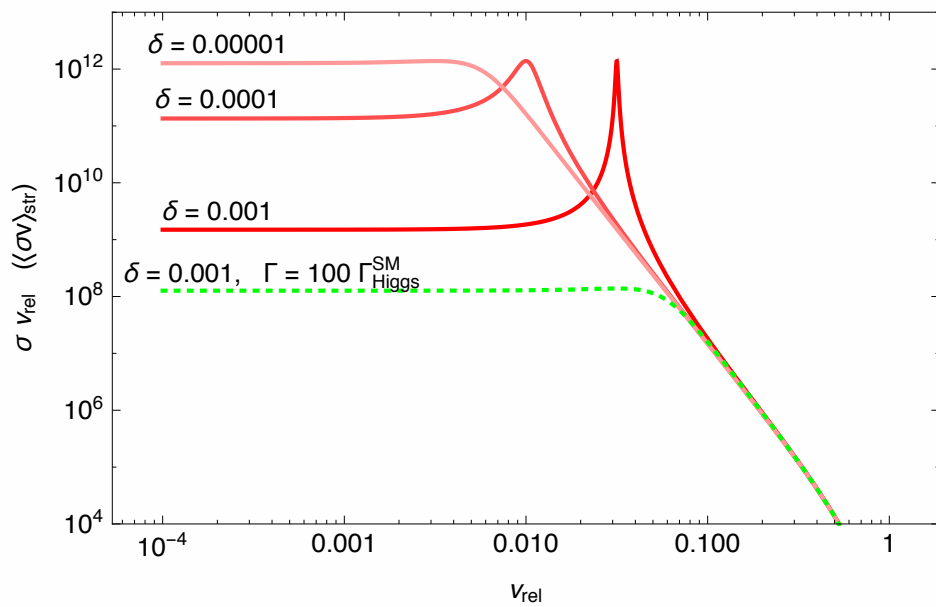
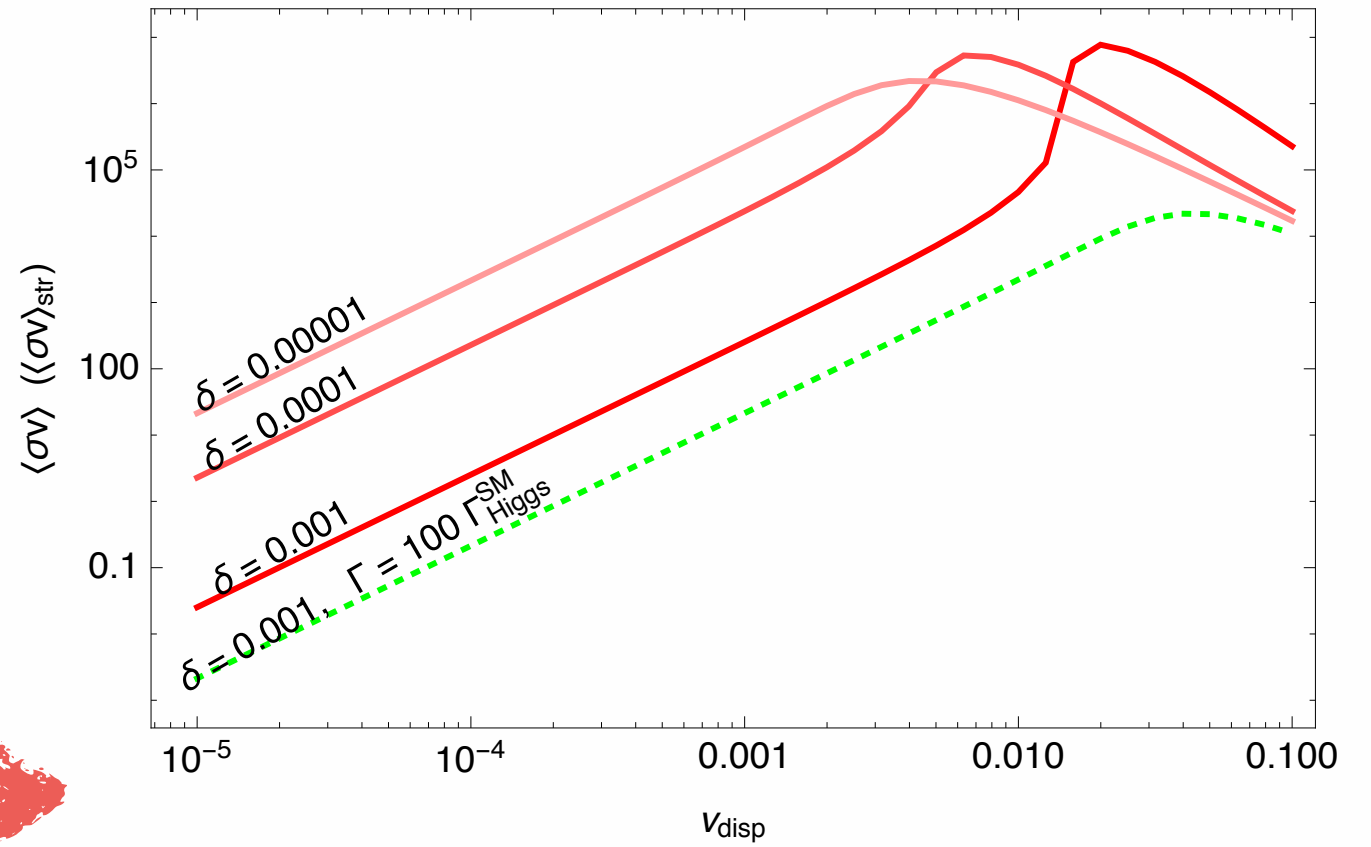
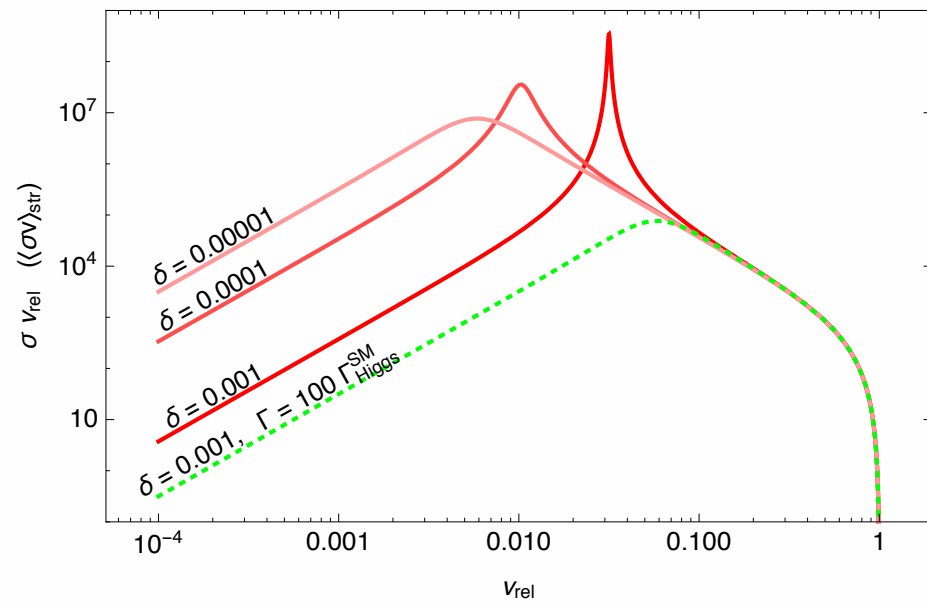
$$A = 1, \quad y_f = 1$$

$$\Gamma = \Gamma_{\text{higgs}}$$





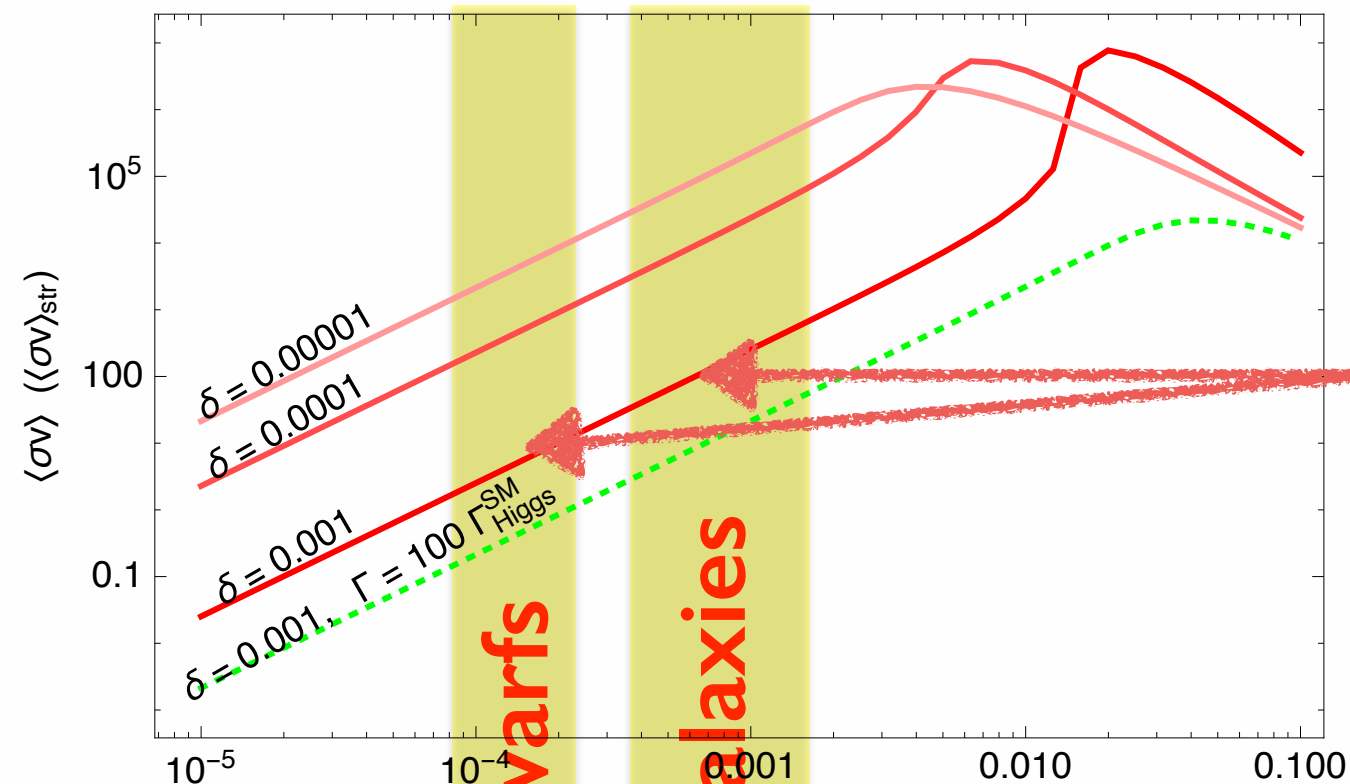
RP offers interesting phenomenology for indirect section





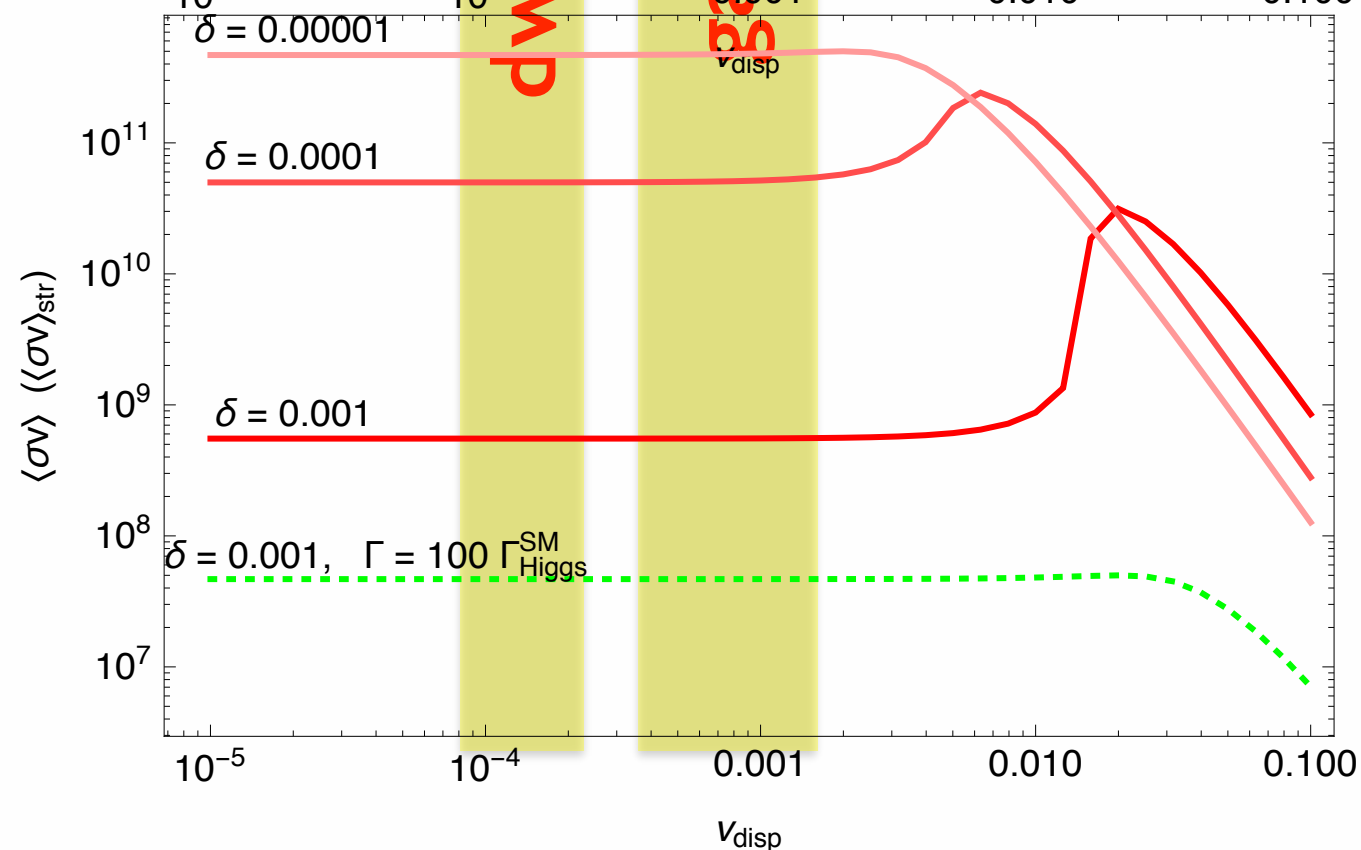
RP offers interesting phenomenology for indirect section

scalar



Difference of
two orders magnitude!

pseudo-
scalar



However, the real life is
more complicated!



Takeaway messages from my talk

- Coy DM offers rich phenomenology — be careful with the γ -ray line!
- ‘Resonance Portal’ — the annihilation cross-section constraints from galaxy clusters, galaxies, dwarfs etc are different!



Thank you!