## Hydrodynamics at large orders

#### Michał P. Heller

Perimeter Institute for Theoretical Physics, Canada

National Centre for Nuclear Research, Poland

based on **I 302.0697**, **I 409.5087**, **I 503.07514**, **I 603.05344 + work in progress** 

# Hydrodynamization

M. P. Heller, R. A. Janik and P. Witaszczyk, Phys. Rev. Lett. 108, 201602 (2012), **1103.3452** 

## Relativistic hydrodynamics

hydrodynamics is

an EFT of the slow evolution of conserved currents in collective media close to equilibrium

**DOFs**: always local energy density  $\epsilon$  and local flow velocity  $u^{\mu}$   $(u_{\nu}u^{\nu} = -1)$ **EOMs**: conservation eqns  $\nabla_{\mu} \langle T^{\mu\nu} \rangle = 0$  for  $\langle T^{\mu\nu} \rangle \underline{\text{expanded in gradients}}$ 



This talk: behaviour of the gradient expansion at large orders in the number of abla

## Boost-invariant flow [Bjorken 1982]



Boost-invariance: in  $(\tau \equiv \sqrt{x_0^2 - x_1^2}, y \equiv \operatorname{arctanh} \frac{x_1}{x_0}, x_2, x_3)$  coords no y-dep

In a CFT: 
$$\langle T^{\mu}_{\nu} \rangle = \text{diag} \left\{ -\mathcal{E}(\tau), -\mathcal{E} - \tau \dot{\mathcal{E}}, \, \mathcal{E} + \frac{1}{2}\tau \dot{\mathcal{E}}, \, \mathcal{E} + \frac{1}{2}\tau \dot{\mathcal{E}} \right\}$$
  
and via scale-invariance  $\frac{\Delta \mathcal{P}}{\mathcal{E}}$  is a function of  $w \equiv \tau T$   
Gradient expansion: series in  $\frac{1}{w}$ .

# Hydrodynamization 1103.3452 (see also Chesler & Yaffe 0906.4426, 1011.3562)

Ab initio calculation in N=4 SYM at strong coupling:





3/15

# Why hydrodynamization can occur?

M. P. Heller, R. A. Janik and P. Witaszczyk, Phys. Rev. Lett. 110, 211602 (2013), **1302.0697** 

### Excitations in strongly-coupled plasmas

see, e.g. Kovtun & Starinets [hep-th/0506184]



 $\omega(k) \rightarrow 0$  as  $k \rightarrow 0$ : slowly dissipating modes (hydrodynamic sound waves)

all the rest: far from equilibrium (QNM) modes damped over  $t_{therm} = O(1)/T$ 

### Hydrodynamic gradient expansion is divergent

In I302.0697 we computed 
$$f(w) \equiv \frac{2}{3} + \frac{1}{6} \frac{\Delta \mathcal{P}}{\mathcal{E}}$$
 up to  $O(w^{-240})$ :



## Hydrodynamics and QNMs

1302.0697

Analytic continuation of  $f_B(\xi) \approx \sum_{n=0}^{240} \frac{1}{n!} f_n \xi^n$  revealed the following singularities:



Branch cut singularities start at  $\frac{3}{2} i \omega_{QNM_1}!$ 

# Resumming gradient expansion in MIS theory

M. P. Heller, M. Spaliński, Phys. Rev. Lett. 115, 072501 (2015), **1503.07514** 

### The boost-invariant MIS theory

1503.07514

$$(\tau_{\Pi} \mathcal{D} + 1) \Pi_{\mu\nu} = -\eta \sigma_{\mu\nu} + \dots$$

$$\int_{C_{\eta}} C_{\eta} = \frac{\eta}{s} \quad \text{and} \quad C_{\tau_{\Pi}} = \tau_{\Pi} T$$

$$f' = -\frac{1}{C_{\tau_{\Pi}}} + \frac{2}{3C_{\tau_{\Pi}}f} + \frac{16}{3w} - \frac{16}{9wf} + \frac{4C_{\eta}}{9C_{\tau_{\Pi}}wf} - \frac{4f}{w} + \dots$$

$$\int_{W} \text{ gradient expansion and a QNM}$$

$$f = \sum_{n=0}^{\infty} \frac{f_n}{w^n} + \delta f + \mathcal{O}(\delta f^2)$$

$$\exp(-\frac{3}{2C_{\tau_{\Pi}}}w) \times \dots$$

$$general solution has a trans-series form$$

$$f = \sum_{m=0}^{\infty} (c_{amb} + r)^m \left\{ w^{\frac{C_{\eta}}{C_{\tau_{\Pi}}}} \exp\left(-\frac{3}{2C_{\tau_{\Pi}}}w\right) \right\}^m \sum_{n=0}^{\infty} a_{m,n} w^{-n}$$

# Divergent gradient expansion at weak coupling

M. P. Heller, A. Kurkela & M. Spaliński, work-in-progress

### **RTA** kinetic theory

Natural language to talk about weakly coupled media is the Boltzmann equation:

$$p^{\mu}\partial_{\mu}g(x,p) = C[g(x,p)]$$
 with  $\langle T^{\mu\nu}\rangle(x) = \int_{\text{momenta}} g(x,p) p^{\mu}p^{\nu}$ 

LO C[g(x, p)] for gauge theories is complicated. We will use instead

$$C[g(x,p)] = -\frac{p^{\mu}u_{\mu}}{\hat{\mathcal{T}}} \Big\{ g(x,p) - g_0(x,p) \Big\} \text{ with } g_0(x,p) = e^{\frac{u_{\mu}p^{\mu}}{T}}$$

This equation is, typically, highly nonlinear due to  $\langle T^{\mu\nu} \rangle u_{\nu} = -\mathcal{E}(T) u^{\mu\nu}$ 

CFTs: 
$$p^{\mu}p_{\mu} = 0$$
 and  $\hat{\mathcal{T}} = \frac{\gamma}{T}$ . Note that  $\gamma$  can be scaled-away (we set it to I).

### Hydrodynamics in kinetic theory

We search for solution to

$$\partial_{\tau} g(\tau, p^{y}, p^{\perp}) = T(\tau) \Big\{ g(\tau, p^{y}, p^{\perp}) - \exp\left(-\frac{1}{T(\tau)} \sqrt{\tau^{2} (p^{y})^{2} + (p^{\perp})^{2}}\right) \Big\}$$

of the form 
$$g(\tau, p^y, p^{\perp}) = \exp\left(-\frac{1}{T}\sqrt{\tau^2 \left(p^y\right)^2 + \left(p^{\perp}\right)^2}\right) \left\{1 + \sum_{n=1}^{\infty} \frac{1}{(\tau T)^n} g_n\left(\frac{\tau p^\eta}{T(\tau)}, \frac{p^{\perp}}{T(\tau)}\right)\right\}$$

Landau matching allows us to obtain

$$f(w) = \frac{2}{3} + \frac{1}{6} \frac{\Delta \mathcal{P}}{\mathcal{E}} = \frac{2}{3} + \frac{4}{45} \frac{1}{w} + \frac{16}{945} \frac{1}{w^2} = \frac{1}{20} + \dots + \frac{1}{5} \frac{1}{8} = \frac{1}{5} \gamma$$

QNM in kinetic theory



# Gradient expansion in FRLW cosmology

A. Buchel, M. P. Heller & J. Noronha 1603.05344

## Holographic dual to FRLW cosmology

Another symmetric flow: comoving plasma in a FRLW universe:  $ds^2 = -dt^2 + a(t)^2 d\vec{x}^2$ 

conformal For a CFT, equilibrium state in  $\mathbb{R}^{I,3} \xrightarrow{\text{trafo}} \bullet$  equilibrium comoving plasma in FRLW:

$$\epsilon = \frac{3}{8}\pi^2 N_c^2 T^4 + \frac{3N_c^2 \dot{a}^4}{32\pi^2 a^4} \quad \text{and} \quad P = \frac{1}{3}\epsilon - \frac{N_c^2 \dot{a}^2 \ddot{a}}{8\pi^2 a^3} \quad \text{with} \quad T = \frac{\mu}{4\pi a}$$

This is an all-order hydro answer in a CFT. Dual: slicing of AdS-Schwarzschild:

see also Siopsis et al. 0809.3505

$$ds_5^2 = 2dt \ (dr - Adt) + \Sigma^2 \ dx^2$$
 with  $A = \frac{r^2}{8} \left( 1 - \frac{\mu^4}{r^4 a^4} \right) - \frac{\dot{a}}{a} \ r_1 \ \& \ \Sigma = \frac{ar}{2}$ 

No entropy production: 
$$r_{hor} = \frac{\mu}{a} \otimes A_{hor} = \frac{\mu^3}{8}$$
. To introduce it   
- hydro-  
see Rozali et al.  
1505.03901

## Hydrodynamic entropy production in FRLW

To conformal we consider: 
$$S_{EH} + \# \int d^5x \sqrt{-g} \left\{ (\partial \phi)^2 - \frac{\Delta(\Delta - 4)}{L^2} \phi^2 + \ldots \right\}$$
 with

$$\phi = r^{4-\Delta}J + \ldots$$
 and we work to LO in  $\delta = rac{J}{T^{4-\Delta}}$ 

For  $\mathcal{N}=2^*$ :  $\Delta=2$  and  $J\sim m_b^2$  or  $\Delta=3$  and  $J\sim m_f$ 

We get 
$$\frac{dA_{hor}}{dt} \sim (\partial_r \Sigma^3) \times (\partial_t \phi + A \partial_r \phi)^2 \Big|_{r=r_{hor}} + O(\delta^3)$$

In the rest of my talk we focus on  $(\partial_t \phi + A \partial_r \phi) \Big|_{r=r_{hor}}$  alone Key idea behind 1603.05344 : solve EOM<sub> $\phi$ </sub> (to express ) in derivatives of a(t)

12/15

### Numerical holography at the precision frontier

We find 
$$\phi = \delta$$
  $a^{4-\Delta}(4-\Delta) \sum_{n=0}^{\infty} \frac{\mathcal{T}_{n,\Delta}[a]}{\mu^n} F_{\Delta,n}\left(\frac{\mu}{ar}\right)$   
 $\mathcal{T}_{\Delta,n} = a \, \dot{\mathcal{T}}_{\Delta,n-1} + (4-\Delta) \, \dot{a} \, \mathcal{T}_{\Delta,n-1}$   
for  $a = e^{Ht}$ :  
 $\mathcal{T}_{\Delta,n}[a] = \Gamma(n+4-\Delta)a^n H^n$   
 $\mathcal{T}_{\Delta,n}[a] = \Gamma(n+4-\Delta)a^n H^n$ 

divergent gradient expansion provided ( $\checkmark$ )  $F_{\Delta,n}\left(\frac{\mu}{ar}\right)$  does not decay too fast with n

we solve  $M_{ij}F_{\Delta,n}^j = B_i + 2$  bdry cond where ij are indices of pseudospectral representation of  $z = \frac{\mu}{a r}$  coordinate

 $n_{max} = 300 \longrightarrow 150 \text{ points \& 1000 digits}$ 

Key result (for  $\Delta = 3$ )

$$\frac{dA_{hor}}{dt} \sim \left(\partial_r \Sigma^3\right) \times \left(\partial_t \phi + A \,\partial_r \phi\right)^2 \Big|_{r=r_{hor}} + O(\delta^3) \sim \left(\sum_{n=0}^{\infty} c_n \xi^n\right)^2 + O(\delta^3) \text{ with } \xi = \frac{H}{T}$$



# Executive summary



Support

### Evolution equations for relativistic viscous fluids

$$\nabla_{\mu} \{ \epsilon u^{\mu} u^{\nu} + P(\epsilon) \{ g^{\mu\nu} + u^{\mu} u^{\nu} \} - \eta(\epsilon) \sigma^{\mu\nu} \} = 0 \text{ is acausal.}$$

#### Remedy: make $\Pi^{\mu\nu} = \langle T^{\mu\nu} \rangle - (\epsilon u^{\mu}u^{\nu} + P(\epsilon)\{g^{\mu\nu} + u^{\mu}u^{\nu}\})$ a new DOF, e.g. $(\tau_{\Pi}\mathcal{D} + 1)\Pi^{\mu\nu} = -\eta\sigma^{\mu\nu}$

Small perturbations obey Maxwell-Cattaneo equation

$$\partial_t^2 \delta u_z - \frac{\eta}{s} \frac{1}{\tau_{\Pi} T} \partial_x^2 \delta u_z + \frac{1}{\tau_{\Pi}} \partial_t \delta u_z = 0$$

Take it seriously:

$$\label{eq:star} \begin{split} \omega &= -i \frac{\eta}{sT} k^2 + \dots \quad \text{and} \quad \omega = -i \frac{1}{\tau_{\Pi}} + i \frac{\eta}{sT} k^2 + \dots \\ \text{hydrodynamics} \quad \text{purely imaginary "quasinormal mode"} \end{split}$$

Generalization that adds  $\operatorname{Re}(\omega_{QNM})$ :  $\left((\frac{1}{T}\mathcal{D})^2 + 2\omega_I \frac{1}{T}\mathcal{D} + |\omega|^2\right) \Pi^{\mu\nu} = -\eta |\omega|^2 \sigma^{\mu\nu}$ 

1409.5087