

Three Pieces in Closed Time Path

Sangyong Jeon

*Department of Physics
McGill University
Montréal, QC, CANADA*

CERN
August 24, 2016

McGill is in Montréal, Québec, Canada



Montreal
QC
Canada

McGill is in Montréal, Québec, Canada



McGill is in Montréal, Québec, Canada



McGill is in Montréal, Québec, Canada



Mr. McGill going home after a hard day's work.



Rutherford carried out his Nobel Prize (1908) winning work at McGill (1898-1907).

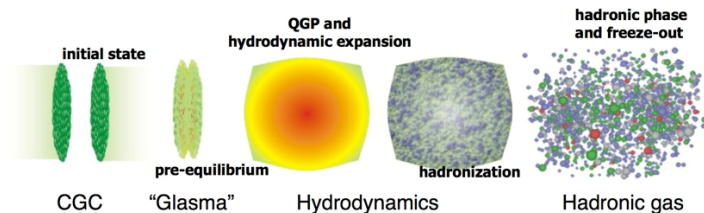
His *original* equipment on display

- Charles Gale
- Sangyong Jeon
- Chun Shen → BNL
- Alina Czajka
- Li Yan ← Saclay
- Sangwook Ryu → Frankfurt
- Michael Richard
- Igor Kozlov
- Chanwook Park
- Scott McDonald
- Mayank Singh
- Sigtryggur Hauksson

MUSIC (Hydro), MARTINI (Jets), AdS/QCD, FTFT, Many-body QCD, ...

Former members: Björn Schenke, Clint Young, Gabriel Denicol, Matt Luzum, Gojko Vujanovic, Jean-Francois Paquet, Abhijit Majumder, Mohammed Mia, Abhee Kanti Dutt-Mazumder, Prashanth Jaikumar, Champak B. Das, Thomas Epelbaum, ...

Stages in heavy ion collisions



- CGC (Color Glass Condensate) & Glasma stages: Dominated by Classical Yang-Mills field
- What is Classical Yang-Mills?
- Quantum corrections?

Want to use Closed Time Path + Keldysh Rotation:

- Physics Motivation: Ultra-relativistic Heavy Ion Collisions
 - *Dense* gluon dynamics (i.e. classical Yang-Mills) dominate the initial state as well as the formation of Quark-Gluon Plasma
 - (Local) Thermalization mechanism – Coherent classical field to the maximum entropy state?
- Technical Motivation: *How to do classical field theory in QFT*
 - CTP + Keldysh Rotation – Clean separation of Q and C d.o.f.

Transition Amplitude vs Expectation Value

- Transition amplitude in QM

$$\begin{aligned} T_{fi} &= \langle f | \hat{U}(t_f, t_i) | i \rangle \\ &= \langle f | q_f \rangle \langle q_f | \hat{U}(t_f, t_i) | q_i \rangle \langle q_i | i \rangle \\ &= \langle f | q_f \rangle \int_{q_i}^{q_f} \mathcal{D}q e^{i \int_{t_i}^{t_f} dt L(q, \dot{q})} \langle q_i | i \rangle \end{aligned}$$

- If a classical path dominates (for instance, large energy), shift $q \rightarrow q_{cl} + q$ to get

$$\int_{q_i}^{q_f} \mathcal{D}q e^{i \int_{t_i}^{t_f} dt L(q, \dot{q})} = e^{i S_{cl}(q_f, q_i)} \int_0^0 \mathcal{D}q e^{i \int_{t_i}^{t_f} dt \delta L(q, \dot{q} | q_{cl})}$$

where $\delta L(q, \dot{q} | q_{cl})$ is at least quadratic in q and \dot{q}

- The problem here is one of the **boundary value problem**
- This is *not* what we want

What we want: Initial value problem

- Expectation value in QM

$$\langle O(t) \rangle = \langle i | \hat{U}(t_j, t) \hat{O}_S \hat{U}(t, t_j) | i \rangle$$

is an initial value problem

- It is going to involve *two* transition amplitudes or two path integrals. One for $\hat{U}(t_f, t_j)$ and another for $\hat{U}(t_j, t_f)$
- How does one formulate classical initial value problem from this?

Expectation Value

- Expectation value

$$\langle O(t) \rangle_i = \langle i | \hat{U}(t_{\text{init}}, t) \hat{O}_S \hat{U}(t, t_{\text{init}}) | i \rangle$$

Time flows from t_{init} to t and then back to t_{in} . Only the in-state is specified. \implies *Initial value problem*

- Natural to define the initial density operator

$$\hat{\rho}_i = |i\rangle\langle i|$$

or more generally

$$\hat{\rho}_{\text{init}} = \sum_n P_n |n\rangle\langle n| \quad \text{with} \quad \sum_n P_n = 1$$

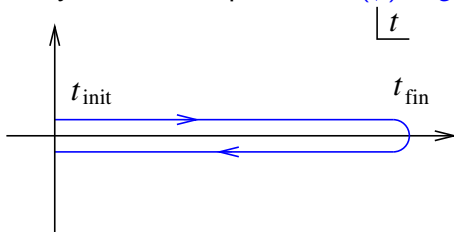
so that

$$\langle O(t) \rangle = \text{Tr} \hat{\rho}(t) \hat{O}_S \quad \text{with} \quad \hat{\rho}(t) = \hat{U}(t, t_i) \hat{\rho}_{\text{init}} \hat{U}(t_i, t)$$

Formulating the initial value problem

Closed Time Path (Schwinger-Keldysh, Keldysh-Schwinger, in-in ...)

Using the scalar theory as an example with $V(\phi) = g^2 \phi^4 / 4!$



- Upper branch (ϕ_1) represents $\langle \text{fin} | \hat{U}(t_{\text{fin}}, t_{\text{init}}) | \text{init} \rangle$
- Lower branch (ϕ_2) represents $\langle \text{init} | \hat{U}(t_{\text{init}}, t_{\text{fin}}) | \text{fin} \rangle$
- Generating functional

$$Z[J_1, J_2]$$

$$= \int [d\phi_f] \langle \phi_f | \hat{\rho}(t_{\text{fin}}) | \phi_f \rangle$$

$$= \int [d\phi_f][d\phi_1][d\phi_2] \langle \phi_f | \hat{U}_{J_1}(t_{\text{fin}}, t_{\text{init}}) | \phi_1 \rangle \langle \phi_1 | \hat{\rho}_{\text{init}} | \phi_2 \rangle \langle \phi_2 | \hat{U}_{J_2}(t_{\text{init}}, t_{\text{fin}}) | \phi_f \rangle$$

Closed Time Path – Cont.

- Generating functional

$$Z[J_1, J_2] = \int [d\phi_f][d\phi_1^i][d\phi_2^i] \int_{\phi_1^i}^{\phi_f} \mathcal{D}\phi_1 \int_{\phi_2^i}^{\phi_f} \mathcal{D}\phi_2 \rho_{\text{init}}[\phi_1^1, \phi_2^2] \exp\left(i \int_{t_i}^{t_f} (L(\phi_1) - L(\phi_2) + J_1\phi_1 - J_2\phi_2)\right)$$

- t_f can be taken to be at $t = \infty$
- Propagator takes on the 2×2 matrix structure

$$\mathbf{G}_{\text{CTP}} = \begin{pmatrix} G_{11} & G_{12} \\ G_{21} & G_{22} \end{pmatrix}$$

$$G_{11} = G_{22}^* = G_F = \langle T\phi(x)\phi(y) \rangle$$

$$G_{12} = \langle \phi_1(x)\phi_2(y) \rangle = \langle \phi(x)\phi(y) \rangle$$

$$G_{21} = \langle \phi_2(x)\phi_1(y) \rangle = \langle \phi(y)\phi(x) \rangle$$

Keldysh rotation, AKA r-a formalism

- Introduce a change of variables

$$\phi_c (= \phi_r) = \frac{\phi_1 + \phi_2}{2} \quad \text{and} \quad J_c = J_1 - J_2$$

$$\phi_q (= \phi_a) = \phi_1 - \phi_2 \quad \text{and} \quad J_q = \frac{J_1 + J_2}{2}$$

- The time derivative terms in the Lagrangian:

$$\frac{\dot{\phi}_1^2}{2} - \frac{\dot{\phi}_2^2}{2} = \dot{\phi}_c \dot{\phi}_q = \partial_t(\dot{\phi}_c \phi_q) - \ddot{\phi}_c \phi_q$$

Upon integrations by part ($\phi_q^f = 0$ because $\phi_1^f = \phi_2^f$)

$$\int (L(\phi_1) - L(\phi_2)) = \int_{t_{\text{init}}}^{t_{\text{fin}}} \left(\phi_q E[\phi_c] - \frac{g^2}{4!} \phi_q^3 \phi_c \right) - \dot{\phi}_c^i \phi_q^i$$

where $E[\phi_c] = \frac{\delta S[\phi_c]}{\delta \phi_c}$ is *the sourceless classical field equation*

Keldysh rotation – Cont.

- Carrying out $[d\phi_q^i]$ integral with the surface term $-\dot{\phi}_c^i \phi_q^i$ transforms the initial density matrix to the Wigner form

$$\int [d\phi_q^i] e^{-i\dot{\phi}_c^i \phi_q^i} \rho[\phi_c^i + \phi_q^i/2, \phi_c^i - \phi_q^i/2] = \rho_W[\phi_c^i, \dot{\phi}_c^i]$$

and the Generating functional now becomes

$$\begin{aligned} Z[J_c, J_q] &= \int \mathcal{D}\phi_c \int \mathcal{D}\phi_q \rho_W[\phi_c^i, \dot{\phi}_c^i] \\ &\quad \exp\left(i \int_{t_i}^{t_f} \left(\phi_q (E[\phi_c] + J_q) - \frac{g^2}{4!} \phi_q^3 \phi_c + J_c \phi_c \right)\right) \end{aligned}$$

with no functional integration over ϕ_q at t_{init} and t_{fin}

- If the $\phi_q^3 \phi_c$ term is ignored, the evolution of the field is entirely *classical* with J_q as the source and with the initial data $(\phi_c^i, \dot{\phi}_c^i)$

Free field theory and the Propagators

- Free field theory

$$Z[J_c, J_q] = \int \mathcal{D}\phi_c \int \mathcal{D}\phi_q \rho_W[\phi_c^i, \dot{\phi}_c^i] \exp\left(i \int_{t_i}^{t_f} \left(\phi_q(-\partial^2 - m^2)\phi_c + J_q\phi_q + J_c\phi_c\right)\right)$$

- Integrating over ϕ_q yields

$$Z[J_c, J_q] = \int \mathcal{D}\phi_c \rho_W[\phi_c^i, \dot{\phi}_c^i] \delta[(\partial^2 + m^2)\phi_c - J_q] \exp\left(i \int_{t_i}^{t_f} J_c\phi_c\right)$$

- Solve the classical EoM in the (t, \mathbf{k}) space

$$\phi_c(t, \mathbf{k}) = \phi_h(t, \mathbf{k}) + \int_{t_i}^{t_f} dt' G_R(t - t', \mathbf{k}) J_q(t', -\mathbf{k})$$

where $G_R(t - t', \mathbf{k}) = \theta(t - t') \sin(E_k(t - t'))/E_k$ and $\phi_h(t, \mathbf{k}) = \phi_c^i(\mathbf{k}) \cos(E_k(t - t_i)) + \dot{\phi}_c^i(\mathbf{k}) \sin(E_k(t - t_i))/E_k$

Free field theory and the Propagators

$$Z[J_c, J_q] = \int \mathcal{D}\phi_c \rho_W[\phi_c^i, \dot{\phi}_c^i] \delta[(\partial^2 + m^2)\phi_c - J_q] \exp\left(i \int_{t_i}^{t_f} J_c \phi_c\right)$$

- Integrating over ϕ_c produces the Jacobian $|\text{Det}(\partial^2 + m^2)|^{-1}$
- Changing variables from ϕ_f to $\dot{\phi}_i$ produces $\left| \text{Det} \left(\frac{\delta\phi_f}{\delta\dot{\phi}_i} \right) \right|$

Since $\dot{\phi}_i = \frac{\delta\mathcal{S}}{\delta\phi_i}$, this is the inverse of the van Vleck determinant

$|\text{Det}(\partial^2 + m^2)|$ which cancels the Jacobian

$$Z[J_c, J_q] = \int [d\phi_c^i][d\pi_c^i] \rho_W[\phi_c^i, \pi_c^i] \exp\left(i \int_{t_i}^{t_f} J_c \phi_c[\phi_c^i, \pi_c^i, J_q]\right)$$

with $\pi_c^i = \dot{\phi}_c^i$

Propagators

Since $\int J\phi_c = \int J_c G_R J_q + \int J_c \phi_h[\phi_c^i, \phi_c^i]$,

- These two are obvious
 - $\langle \phi_c(t)\phi_q(t') \rangle = iG_R(t-t')$ (time always flows from ϕ_q to ϕ_c)
 - $\langle \phi_q(t)\phi_q(t') \rangle = 0$
- The symmetric propagator $G_S = \langle \phi_c \phi_c \rangle$ depends on $\rho_W[\phi_c^i, \pi_c^i]$
 - Classical vacuum: $\rho_W = \delta[\phi_c^i]\delta[\pi_c^i]$ which gives $\langle \phi_c(t)\phi_c(t') \rangle = 0$
 - Quantum *perturbative* vacuum:

$$\rho_W[\phi, \pi] = \exp\left(-\int \frac{d^2k}{(2\pi)^3 E_k} (E_k^2 \phi(\mathbf{k})\phi(-\mathbf{k}) + \pi(\mathbf{k})\pi(-\mathbf{k}))\right)$$

which gives $\langle \phi_c(t)\phi_c(t') \rangle = FT [\pi\delta(p^2 - m^2)]$

- Thermal medium: $\langle \phi_c(t)\phi_c(t') \rangle = FT [(1/2 + n_B(p^0))2\pi\delta(p^2 - m^2)]$
- *Quantum effect*: Non-vanishing $G_S = \langle \phi_c \phi_c \rangle$

Almost classical interpretation

$$Z[J_c, J_q] = \int \mathcal{D}\phi_c \int \mathcal{D}\phi_q \rho_W[\phi_c^i, \dot{\phi}_c^i] \exp \left(i \int \left(\phi_q (E[\phi_c] + J_q) - \frac{g^2}{4!} \phi_q^3 \phi_c + J_c \phi_c \right) \right)$$

- $\rho_W[\phi_c^i, \dot{\phi}_c^i] \sim$ Probability distribution of the initial data (Not strictly, since it's a Wigner transform)
- If for some reason $\phi_q \ll \phi_c$, then drop the ϕ_q^3 term to get

$$\int \mathcal{D}\phi_q e^{i \int \phi_q (E[\phi_c] + J_q)} = \delta[E[\phi_c] + J_q]$$

which enforces the classical equation of motion

- Origin of quantum effects
 - $\rho_W[\phi_c^i, \dot{\phi}_c^i]$: Includes quantum effects. Especially the zero-point motions.
 - Quantum vertex $g^2 \phi_q^3 \phi_c / 4!$: Provides correlations absent in the classical theory

When the classical field dominates

- Let $V = g^2 \phi^4 / 4!$. The EoM is

$$(\partial^2 + m^2)\phi_c + \frac{g^2}{3!}\phi_c^3 = J_q$$

- Suppose we have a *physical* source $J_q = J_{\text{phys}}$.
- If $J_{\text{phys}} = O(1/g)$, then

$$\phi_c = O(1/g)$$

and the interaction term is as big as the free field terms.

The Lagrangian

- Let φ be the solution of the classical EoM and let $\phi_c \rightarrow \varphi + \phi_c$
- The Lagrangian

$$\begin{aligned} L &= \phi_q (E[\varphi + \phi_c] + J_{\text{phys}}) + \frac{g^2}{4!} \phi_q^3 \varphi + \frac{g^2}{4!} \phi_q^3 \phi_c \\ &= \phi_q \left((\partial^2 + m^2 + \underbrace{\frac{g^2}{2} \varphi^2}_{O(1)}) \phi_c + \underbrace{\frac{g^2}{2} \varphi \phi_c^2}_{O(g)} + \underbrace{\frac{g^2}{3!} \phi_c^3}_{O(g^2)} \right) \\ &\quad + \underbrace{\frac{g^2}{4!} \varphi \phi_q^3}_{O(g)} + \underbrace{\frac{g^2}{4!} \phi_q^3 \phi_c}_{O(g^2)} \end{aligned}$$

- One can do perturbation theory if one knows $G_R = 1/(\partial^2 + m^2 + g^2 \varphi^2/2)$
- If interested in only the leading order corrections, just ignore ϕ_q^3 terms and solve classical field equations with the fluctuating initial condition.

- Carrying out $\int \mathcal{D}\phi_q$ integrals results in $\delta[E[\phi_c] + J_q]$
- Carrying out $\int \mathcal{D}\phi_c$ integrals results in $\text{Det}^{-1} \left(\frac{\delta E[\phi_c]}{\delta \phi_c} \right)$
- Swapping the boundary value problem (with ϕ_i, ϕ_f) with the initial value problem (with $\phi_i, \dot{\phi}_i$) results in

$$\text{Det} \left(\frac{\delta \phi_f}{\delta \dot{\phi}_i} \right) = \text{Det} \left(\frac{\delta^2 \mathcal{S}}{\delta \phi_i \delta \phi_f} \right)^{-1} = \text{Det} \left(\frac{\delta E[\phi_c]}{\delta \phi_c} \right)$$

- Any observable up to LO + NLO (with $\pi = \dot{\phi}$)

$$\langle \mathcal{O}(t) \rangle = \int [d\phi_c^i][d\pi_c^i] \rho_W[\phi_c^i, \pi_c^i] \mathcal{O}[\phi_{\text{cl}}(t), \pi_{\text{cl}}(t)]$$

- In principle

$$\begin{aligned}
 \langle \mathcal{O}(t) \rangle &= \int \mathcal{D}\phi_c \int \mathcal{D}\phi_q \rho_W[\phi_c^i, \dot{\phi}_c^i] \exp\left(i \int \left(\phi_q(E[\phi_c] + J_q) - \frac{\phi_q^3}{4!} V''''(\phi_c) \right)\right) \mathcal{O}[\phi_c, \pi_c] \\
 &= \int \mathcal{D}\phi_c \int \mathcal{D}\phi_q \rho_W[\phi_c^i, \dot{\phi}_c^i] \exp\left(i \int (\phi_q(E[\phi_c] + J_q))\right) \\
 &\quad \times \left(1 - i \int d^4x \frac{\phi_q^3}{4!} V''''(\phi_c) + \dots \right) \mathcal{O}[\phi_c, \dot{\phi}_c] \\
 &= \int \mathcal{D}\phi_c \rho_W[\phi_c^i, \dot{\phi}_c^i] \delta[E[\phi_c] + J_q] \mathcal{O}[\phi_c, \dot{\phi}_c] \\
 &\quad - i \int d^4x \frac{\delta^3}{\delta J_q(x)^3} \int \mathcal{D}\phi_c \rho_W[\phi_c^i, \dot{\phi}_c^i] \delta[E[\phi_c] + J_q] \mathcal{O}[\phi_c, \dot{\phi}_c] \frac{1}{4!} V''''(\phi_c(x)) + \dots
 \end{aligned}$$

provide that $\rho_W[\phi_c^i, \dot{\phi}_c^i]$ is also accurate up to the first order quantum correction

- In practice, not so easy

Vacuum Initial State Density

- The vacuum functional satisfies the Schrödinger equation

$$\mathcal{H}|\Psi\rangle = 0$$

where

$$\mathcal{H} = \int d^3x \left(\frac{\pi^2}{2} + \frac{(\nabla\phi)^2}{2} + \frac{m^2}{2}\phi^2 + V(\phi) \right)$$

and

$$\pi(\mathbf{x}) = -i \frac{\delta}{\delta\phi(\mathbf{x})}$$

- Perturbative vacuum (products of SHO ground states)

$$\langle\phi|\Psi\rangle = \exp\left(-\frac{1}{2} \int \frac{d^3k}{(2\pi)^3} E_k \phi(\mathbf{k})\phi(-\mathbf{k})\right)$$

- Vacuum Wigner functional

$$\rho_W[\phi, \pi] = \exp\left(-\int \frac{d^3k}{(2\pi)^3} E_k \left(E_k^2 \phi(\mathbf{k})\phi(-\mathbf{k}) + \pi(\mathbf{k})\pi(-\mathbf{k}) \right)\right)$$

Coherent State

Coherent state: Eigenfunctional of the annihilation operator – Minimum uncertainty state \sim Classical field

- Creation operator in the ϕ representation

$$\mathcal{A}^\dagger(\mathbf{k}) = \left(E_k \phi(-\mathbf{k}) - (2\pi)^3 \frac{\delta}{\delta \phi(\mathbf{k})} \right)$$

- Annihilation operator in the ϕ representation

$$\mathcal{A}(\mathbf{k}) = \left(E_k \phi(\mathbf{k}) + (2\pi)^3 \frac{\delta}{\delta \phi(-\mathbf{k})} \right)$$

- Commutator

$$[\mathcal{A}(\mathbf{k}), \mathcal{A}^\dagger(\mathbf{k}')] = 2E_k (2\pi)^3 \delta(\mathbf{k} - \mathbf{k}')$$

Coherent States

- Ground state: Solving

$$\langle \phi | \mathcal{A}(\mathbf{k}) | \Psi \rangle = 0$$

gives

$$\langle \phi | \Psi \rangle = \mathcal{N} \exp \left(- \int \frac{d^3 q}{(2\pi)^3} \frac{E_q}{2} \phi(\mathbf{q}) \phi(-\mathbf{q}) \right)$$

- Coherent state: Solving

$$\langle \phi | \mathcal{A}(\mathbf{k}) | \varphi + i\Pi \rangle = \varphi(\mathbf{k}) \langle \phi | \varphi + i\Pi \rangle$$

yields

$$\langle \phi | \varphi + i\Pi \rangle = \exp \left(i \int \frac{d^3 k}{(2\pi)^3} \Pi(\mathbf{k}) \phi(-\mathbf{k}) - \frac{1}{2} \int \frac{d^3 k}{(2\pi)^3} E_k |\phi(\mathbf{k}) - \varphi(\mathbf{k})|^2 \right)$$

Coherent State Initial Density Matrix

- Wigner transform of the coherent state functional

$$\rho_W[\phi, \pi] = \exp \left[- \int \frac{d^2k}{(2\pi)^3 E_k} \left(E_k^2 |\phi(\mathbf{k}) - \varphi(\mathbf{k})|^2 + |\pi(\mathbf{k}) - \Pi(\mathbf{k})|^2 \right) \right]$$

- Since this is just a shifted vacuum functional, in practice:

$$\begin{aligned}\phi_c^i &= \varphi + \delta\phi \\ \pi_c^i &= \Pi + \delta\pi\end{aligned}$$

where $\delta\phi$ and $\delta\pi$ follows the usual Gaussian vacuum distributions.

Application 0

Scattering Amplitude

Expansion of Self-Energy

Kadanoff Baym Equation

- Standard application of the CTP formalism
- KB Eq in the more-or-less standard form:

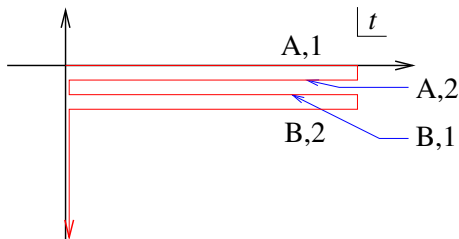
$$(p \cdot \partial) G^{<, >} = \frac{1}{2} (\Pi^> G^< - \Pi^< G^>)$$

- In Quasi-particle approximation with

$$G^>(X, p) = 2\pi\delta(p^2 - m^2) \left[\theta(p^0)(1 + f_+(X, p)) + \theta(-p_0)f_-(X, -p) \right]$$

this can become Kinetic theory equation (e.g. the Boltzmann eq)
provided that the self-energy is expanded in scattering amplitudes

Two-sweep CTP



Simon Caron-Huot's Masters Thesis

By separating the A fields and B fields, one can show

$$\begin{aligned} \Pi^>(P) &= \sum_{n, \{Q\}} \frac{1}{n!} |\mathcal{M}_{ar\dots r}(P; Q_1, \dots, Q_n)|^2 \\ &\quad \times G^>(Q_1) \cdots G^>(Q_n) (2\pi)^4 \delta^{(4)}(Q_1 + \cdots + Q_n - P) \end{aligned}$$

where

- r, a index = c, q index
- $\mathcal{M}_{ar\dots r}(P; Q_1, \dots, Q_n)$: Fully retarded and 1-PI correlation function
- $G^>(Q)$: Full Wightman function

Scattering Matrix Expansion

$$\begin{aligned}\Pi^>(P) &= \sum_n \frac{1}{n!} |\mathcal{M}_{ar\dots r}(P; Q_1, \dots, Q_n)|^2 \\ &\quad \times G^>(Q_1) \cdots G^>(Q_n) (2\pi)^4 \delta^{(4)}(Q_1 + \dots + Q_n - P)\end{aligned}$$

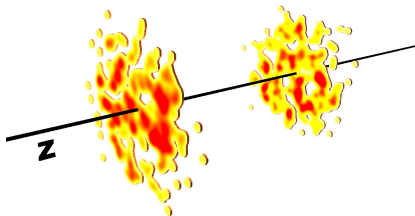
- This appears in the Kadanoff-Baym equation
- Becomes the collision terms in the kinetic theory
- Tells you *what* to calculate for the *in-medium* scattering amplitude – It's *not* the usual Feynman (time-ordered) amplitude in vacuum

Application 1

Color Glass Condensate and the JIMWLK RG Equation

Application – Color Glass Condensate

[Venugopalan, McLerran, JIMWLK, Gelis, Hatta, Fukushima, Dumitru, Kovchegov, Itakura, Lappi, Nara, ...]



- Main idea: Highly accelerated hadrons are composed of
 - Large x partons: 2D frozen-in-time (Color Glass) color current
 - Small x gluons: Weizsäcker-Williams field generated by large x partons (Condensate) \Rightarrow Classical field
- Small x part of the gluon PDF $\sim \langle A_{cl} A_{cl} \rangle$

Yang-Mills with an external color current

Try first

$$L = -\frac{1}{4} G_a^{\mu\nu} G_{\mu\nu}^a - J_a^\mu A_\mu^a$$

- Classical EoM OK: $[D_\mu, G^{\mu\nu}] = J^\nu$ with $[D_\mu, J^\mu] = 0$
- Trouble: Gauge transformation
 - $A' = UAU^\dagger - \frac{1}{ig} U\partial U^\dagger$
 - $J' = UJU^\dagger$
- In QED, $J^\mu A_\mu$ is gauge invariant as long as $\partial_\mu J^\mu = 0$
- In the full QCD, the color current is a part of $\bar{\psi}\gamma^\mu D_\mu\psi$. Without the $\bar{\psi}\gamma^\mu\partial_\mu\psi$ term, however, the L above is *not* gauge invariant even if $[D_\mu, J^\mu] = 0$
- Way out: Non-local interaction. Use $\text{Tr} \ln(\gamma^\mu D_\mu)$ or $\text{Tr} \rho W$ or $\text{Tr} \rho \ln W$ where W is the Wilson line along $u^\mu = J^\mu / \rho$ with $U_i = U_f$ [Jalilian-Marian, Jeon, Venugopalan, Phys.Rev.D63:036004,2001]

- CTP Lagrangian for pure glue with a color current

$$\mathcal{L} = \eta_\nu^a ([D_\mu, G^{\mu\nu}] - J^\nu)_a + \frac{ig}{4} [D_\mu, \eta_\nu]^a [\eta^\mu, \eta^\nu]_a$$

where D_μ and $G^{\mu\nu}$ contains only

$$A_\mu = \frac{A_{1,\mu} + A_{2,\mu}}{2}$$

whereas $\eta_\mu = A_{1,\mu} - A_{2,\mu}$

- The source $J^\mu = J_1^\mu = J_2^\mu$ is the physical external source
- In principle,

$$A'_{1,2} = U_{1,2} A_{1,2} U_{1,2}^\dagger + \frac{1}{ig} U_{1,2} \partial U_{1,2}^\dagger$$

where U_1 and U_2 are not necessarily the same.

- If $U_1 \neq U_2$, $\mathcal{L} = L_1 - L_2$ is not gauge invariant since neither L_1 nor L_2 is gauge invariant
- If $U_1 = U_2 = U$,

$$A' = UAU^\dagger + \frac{1}{ig}U\partial U^\dagger$$

$$\eta' = U\eta U^\dagger$$

and the $J_a^\mu \eta_\mu^a$ term is gauge invariant.

- When one is given a color current J^μ without the corresponding kinetic energy term, this is the results in a gauge invariant *local* theory
- Let

$$A = A_{\text{cl}} + a$$

then systematic perturbative study is possible.

An *exact* solution of Classical YM equation

- There aren't too many exact solution of classical Yang-Mills equation even in static situations
- When $J^\mu = \delta^{\mu\pm} \rho(x^\mp, \mathbf{x}_\perp)$, an *exact* solution of CYM can be found – *MV (McLerran-Venugopalan) model*
- Abelian subgroup solution: Suppose

$$J^\mu = J_3^\mu t_3 + J_8^\mu t^8$$

and

$$A^\mu = A_3^\mu t^3 + A_8^\mu t^8$$

then since $[t^3, t^8] = 0$ (both are diagonal), the classical YM equations reduces to two sets of Maxwell equations – *Of limited utility*.

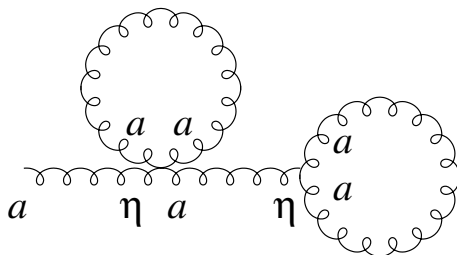
JIMWLK equation

– Quantum correction (RG equation) on top of the MV solution

- Formulated in the light-cone coordinate system $x^\pm = \frac{x^0 \pm x^3}{\sqrt{2}}$
- Main idea: The properties of small x gluons are determined by the underlying color charge distribution.
- Vacuum fluctuation introduces *all* x scales even though the color charge density ρ itself is soft
- Where is the dividing line between small x and large x ? \implies RG approach is necessary
- Main point: This quantum correction to the MV model can be still treated within the classical theory
- CTP calc: [Jeon, Annals Phys. 340 (2014) 119-170]

JIMWLK = Jalilian-Marian, Iancu, McLerran, Weigert, Leonidov, Kovner

Tadpole diagrams



Caution: a here is the c field or the “ r ” field fluctuation and η here is the q field or the “ a ” field

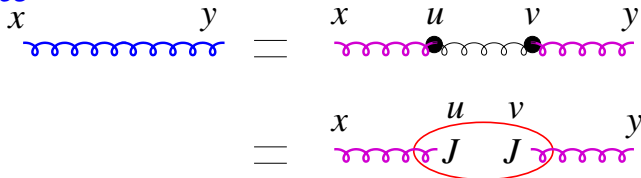
- Leading order quantum corrections
- This is $O(1/g)$ if the UV regulated tadpole contribution is $O(1/g^2)$
- The same as the size of the classical source $J_{\text{phys}} = O(1/g)$
- Large x gluons can act like an *additional* $O(1/g)$ source

Tadpole = Source Correlation at the initial time

- Furthermore, since

$$G_S(x, y) \sim \int_{u, v} \partial_{u^+} G_R(x, u) G_S^0(u - v) \partial_{v^+} G_R(v, y)$$

tadpoles can be generated by *2-point correlation of classical sources*



- Schematically, $\langle J(u)J(v) \rangle \sim \partial_{u^+} \partial_{v^+} G_S^0(u - v)$
- The symmetric propagator is just the transformed Minkowski one

$$G_S^0(x) = \int_k e^{-ik^+ x^- - ik^- x^+ + i\mathbf{k}_\perp \cdot \mathbf{x}_\perp} \pi \delta(2k^+ k^- - \mathbf{k}_\perp^2)$$

Basic Idea

- Construct the source correction $Y[\lambda, \rho]$ so that for any observable $\mathcal{O}[A]$

$$\begin{aligned} & \int \mathcal{D}\rho W_\rho[\rho] \int [da_i] \rho W[a_i] \mathcal{O}[A[\rho, a_i]] \\ &= \int \mathcal{D}\rho W_\rho[\rho] \int \mathcal{D}\lambda Y[\lambda, \rho] \mathcal{O}[A[\rho + \lambda]] \\ &= \int \mathcal{D}\rho W'_\rho[\rho'] \mathcal{O}[A[\rho']] \end{aligned}$$

including the leading quantum corrections

- $W[\rho]$: Geometric color charge distribution
- $Y[\lambda, \rho]$: Gives the same $\langle a_i \rangle$ and $\langle a_i(x) a_i(y) \rangle$ (calculating these correctly in CTP-YM is the main task)
- The combined density $W'_\rho[\rho'] = \int \mathcal{D}\lambda W[\rho - \lambda] Y[\lambda, \rho - \lambda]$

JIMWLK equation

The combined density can be shown to satisfy

$$\frac{\partial W}{\partial Y} = \mathcal{H}W$$

where

$$\mathcal{H} = \frac{1}{2\pi} \int_{\mathbf{u}_\perp, \mathbf{v}_\perp} \frac{\delta}{\delta \alpha_a(\mathbf{u}_\perp)} \eta^{ab}(\mathbf{u}_\perp | \mathbf{v}_\perp) \frac{\delta}{\delta \alpha_b(\mathbf{v}_\perp)}$$

with

$$\begin{aligned} & \eta(\mathbf{x}_\perp | \mathbf{y}_\perp) \\ &= - \int_{\mathbf{u}_\perp} (1 - V^\dagger(\mathbf{u}_\perp)V(\mathbf{y}_\perp) - V^\dagger(\mathbf{x}_\perp)V(\mathbf{u}_\perp) + V^\dagger(\mathbf{x}_\perp)V(\mathbf{y}_\perp)) \\ & \quad \partial_x^i G_T(\mathbf{x}_\perp - \mathbf{u}_\perp) \partial_i^y G_T(\mathbf{u}_\perp - \mathbf{y}_\perp) \end{aligned}$$

where

$$V(u_1^-, u_2^-; \mathbf{u}_\perp) = \mathcal{P} \exp \left(ig \int_{u_2^-}^{u_1^-} dz^- A_-(z^-, \mathbf{u}_\perp) \right)$$

is the color rotation factor when crossing the current and $\nabla_\perp^2 G_T(\mathbf{x}) = \delta^{(2)}(\mathbf{x})$

To get the rapidity evolution of the gluon PDF,

- Solve

$$\frac{\partial W}{\partial Y} = \mathcal{H}W$$

- At a given Y , sample ρ from W
- Solve CYM with the sampled ρ
- Calculate the gluon number density
- Average over many configurations

Thermalization

Initial Conditions in Heavy Ion Collisions

- Initial condition *before* the collision – Composed of classical particles (source) and the classical field they generate
- Initial moments *right after* the collision – Interaction of two classical Yang-Mills fields from the two nuclei \implies **Glasma**
- Around $\tau \approx 0.5$ fm, hydrodynamics starts to apply – Local equilibrium (or some semblance of it) is necessary
- In particular, in the local rest frame

$$T^{\mu\nu} \approx \begin{pmatrix} \varepsilon & 0 & 0 & 0 \\ 0 & P_x & 0 & 0 \\ 0 & 0 & P_y & 0 \\ 0 & 0 & 0 & P_z \end{pmatrix}$$

with $P_x \approx P_y \approx P_z \approx \varepsilon/3$

How do we get here?

Glasmata: [Lappi, McLerran, Romatschke, Gelis, Fukushima, Venugopalan, Jeon, Itakura, ...]

Scalar theory example

- Simple Simulation: Build up scalar field with

$$-\partial^2\phi - \frac{g^2}{3!}\phi^3 = J$$

where $J = O(1/g)$ is the source for $t < 0$

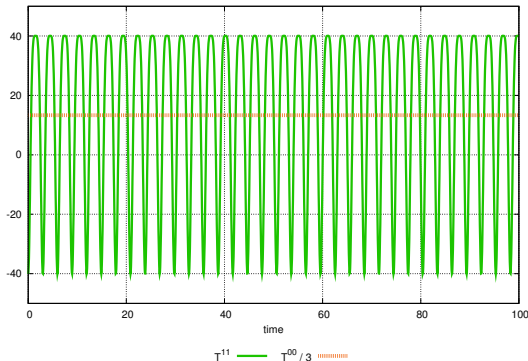
- See what happens at $t > 0$

Thermalization in CTP

Scalar theory example

[Dusling, Epelbaum, Gelis, Venugopalan (DEGS)]

Spatially homogeneous case without vacuum fluctuations:



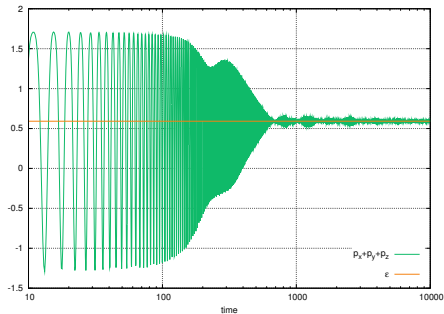
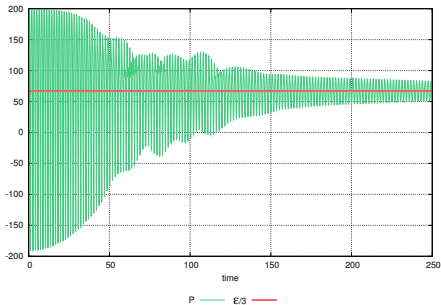
Energy is conserved, but pressure oscillates wildly
(Analytic solution possible in terms of Jacobi elliptic function)

Thermalization in CTP

Scalar theory example

Add vacuum noise

$$\langle O \rangle = \int [d\phi_i][d\pi_i] \rho_V[\phi_i, \pi_i] \delta[E[\varphi + \phi]] \delta[E[\varpi + \pi]] O(\varphi + \phi, \varpi + \pi)$$



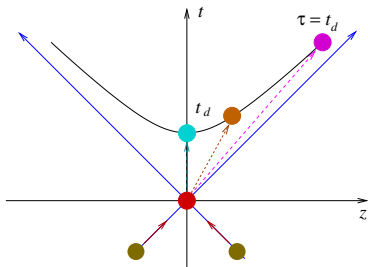
$g = 0.5$, DEGS, NPA 850, 69, 2011

$g = 1.0$, EG, NPA 872, 210, 2011

Miline Space Vacuum Functionals

Initial condition in τ

- Why τ ? – Time dilation



- If the three fireballs all start out from $t = 0, z = 0$ and evolve exactly the same way (e.g. thermalization), the state of the cyan at $t = t_d$ is the same as the state of the brown and magenta at $\tau = t_d$
- Appropriate “time” variable
 - Relativistic case: $\tau = \sqrt{t^2 - z^2} = t\sqrt{1 - v_z^2}$ is the most natural time variable – Local time at z
 - Non-relativistic case: $z = v_z t \ll t \implies t$ is the the most natural time variable.

Milne space

- Milne space

$$\tau = \sqrt{t^2 - z^2}$$
$$\eta = \tanh^{-1}(z/t)$$

or

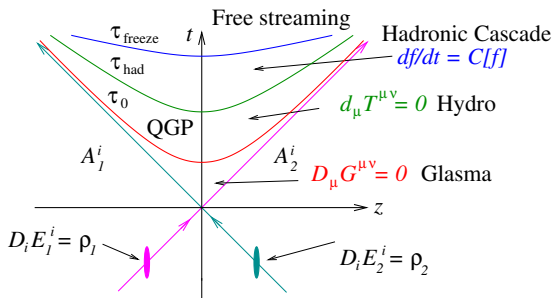
$$t = \tau \cosh \eta$$
$$z = \tau \sinh \eta$$

- Lorentz boost by $v_z = z/t = \tanh \eta$ yields

$$t' = t / \cosh \eta = \tau$$
$$z' = 0$$

- If the collective velocity $v_z = z/t \implies$ Boost invariant Bjorken expansion

Glasma Initial Condition



- Glasma initial condition set at $\tau = 0^+$
- Calculate

$$\langle \mathcal{O} \rangle = \int [d\phi_i][d\pi_i] \rho_V[\phi_i, \pi_i] \delta[E[\varphi + \phi]] \delta[E[\varpi + \pi]] \mathcal{O}(\varphi + \phi, \varpi + \pi)$$

in the τ - η coordinate system \implies Need to know the vacuum Wigner functional in the τ - η coordinate system in the forward light cone

Milne space vacuum functional

Scalar theory [Long and Shore, 1996]

- Schrödinger Equation for the perturbative vacuum

$$\int d\eta d^2x_{\perp} \tau \left(-\frac{1}{2\tau^2} \frac{\delta^2}{\delta\phi^2} - \phi \left(\nabla_{\perp}^2 - \frac{\partial_{\eta}^2}{\tau^2} \right) \phi \right) \langle \phi | \text{vac} \rangle = i \frac{\partial}{\partial \tau} \langle \phi | \text{vac} \rangle$$

- Can't let the RHS vanish because of the $\frac{1}{\tau^2}$ term in the LHS
- Gaussian ansatz

$$\langle \phi | \text{vac} \rangle = \mathcal{N}(\tau) \exp \left(-\frac{\tau^2}{2} \int d\eta_x d^2x_{\perp} \int d\eta_y d^2y_{\perp} G(\tau, x, y) \phi(x) \phi(y) \right)$$

- In the momentum space, $G_T(\tau, \tilde{k})$ satisfies

$$i\partial_{\tau}(\tau^2 G_T) = \tau^3 G_T^2 - \left(\tau k_{\perp}^2 + \frac{k_{\eta}^2}{\tau} \right)$$

with $\tilde{k} = (\mathbf{k}_{\perp}, k_{\eta})$

Milne space vacuum functional

Scalar theory

- Solution

$$G(\tau, x, y) = \int \frac{d^2 q_{\perp} dq_{\eta}}{(2\pi)^3} e^{i\mathbf{q}_{\perp} \cdot (\mathbf{x}_{\perp} - \mathbf{y}_{\perp}) + iq_{\eta}(\eta_x - \eta_y)} \frac{-i\partial_{\tau} H_{iq_{\eta}}^{(1)}(m_T^q \tau)}{\tau H_{iq_{\eta}}^{(1)}(m_T^q \tau)}$$

- Wigner functional [Jeon & Epelbaum Annals of Phys. 364, 1, 2016]

$$\rho_W[\tau, \phi, \pi] = \mathcal{N} \exp \left(-2 \int \frac{d^3 \tilde{\mathbf{k}}}{(2\pi)^3} \left| \pi(\tilde{\mathbf{k}}) a^*(\tau, \tilde{\mathbf{k}}) - \phi(\tilde{\mathbf{k}}) e^*(\tau, \tilde{\mathbf{k}}) \right|^2 \right)$$

where $\pi = \tau \partial_{\tau} \phi$ and

$$a(\tau, \tilde{\mathbf{k}}) = \sqrt{\frac{\pi}{4}} e^{\pi k_{\eta}/2} H_{ik_{\eta}}^{(2)}(k_{\perp} \tau)$$

$$e(\tau, \tilde{\mathbf{k}}) = \tau \partial_{\tau} a(\tau, \tilde{\mathbf{k}})$$

Gauge theory vacuum

- Abelian gauge theory Lagrangian

$$L = \frac{\tau}{2}(\partial_\tau \mathbf{A}_\perp)^2 - \frac{\tau}{2}(\nabla \times \mathbf{A}_\perp)^2 + \frac{1}{2\tau} \mathbf{A}_\perp \cdot \partial_\eta^2 \mathbf{A}_\perp \\ + \frac{1}{2}(\partial_\tau A_\eta)^2 + \frac{1}{2\tau} A_\eta \nabla_\perp^2 A_\eta - \frac{1}{\tau} A_\eta \partial_\eta \nabla_\perp \cdot \mathbf{A}_\perp$$

- Gauss law

$$0 = \frac{\partial_\eta \partial_\tau A_\eta}{\tau^2} + \nabla_\perp \cdot \partial_\tau \mathbf{A}_\perp$$

- Decompose $\mathbf{A}_\perp = \mathbf{A}_T + \mathbf{A}_L$ with

$$\mathbf{A}_T = i \int \frac{d^2 k_\perp}{(2\pi)^2} e^{i\mathbf{k}_\perp \cdot \mathbf{x}_\perp} A_T(\hat{\mathbf{k}}_\perp \times \mathbf{e}_z)$$

and

$$\mathbf{A}_L = \nabla_\perp \varphi$$

Wigner functional

[Jeon & Epelbaum Annals of Phys. 364, 1, 2016]

- Equations for A_T are identical to the scalar case
- Equations for A_η is much more complicated because A_η and φ couple, but solvable.
- Longitudinal Wigner functional

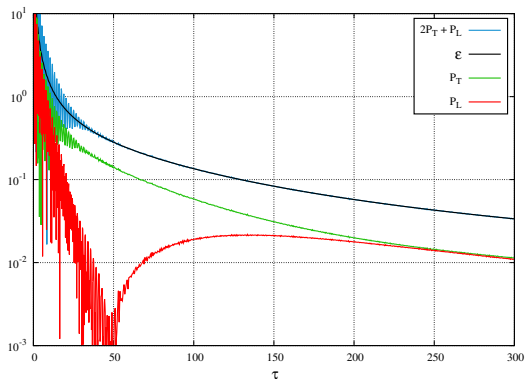
$$\begin{aligned} & \rho_L[\tau, A_\eta, \pi_\eta] \\ &= \mathcal{N} \exp \left(-2 \int \frac{d^3 \tilde{k}}{(2\pi)^3} \frac{1}{k_\perp^2} \left| e^*(\tau, -\tilde{k}) \pi_\eta + k_\perp^2 a^*(\tau, -\tilde{k}) A_\eta \right|^2 \right) \end{aligned}$$

with $\pi_\eta = \frac{1}{\tau} \partial_\tau A_\eta$ and

$$a(\tau, \tilde{k}) = \sqrt{\frac{\pi}{4}} e^{\pi k_\eta / 2} H_{ik_\eta}^{(2)}(k_\perp \tau)$$

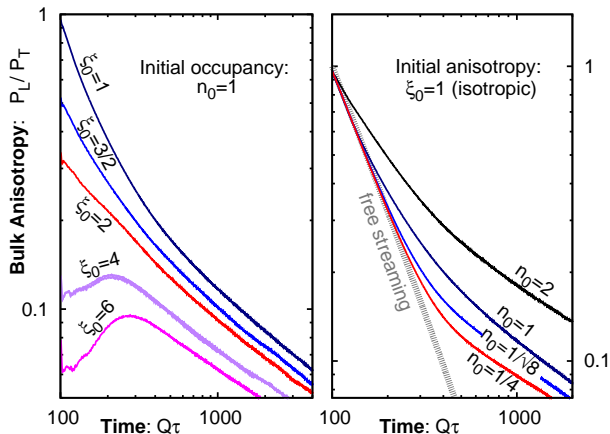
$$e(\tau, \tilde{k}) = \tau \partial_\tau a(\tau, \tilde{k})$$

Scalar theory in Milne space



$g = 4$, DEGV, PRD 86, 085040, 2012

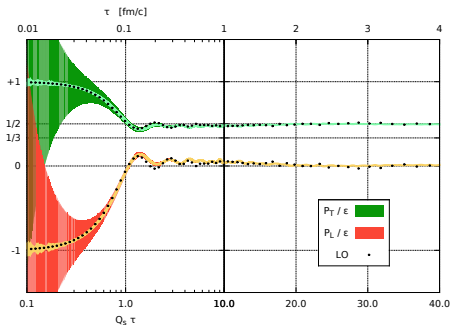
Yang-Mills in Milne space



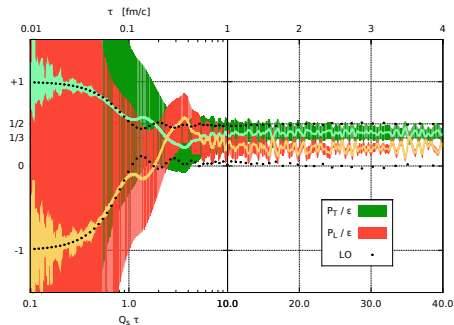
Berges, Boguslavski, Schlichting, Venugopalan, PRD 89, 074011 (2014) (arXiv 1303.5650)

$N_T = 256 - 512$, $N_\eta = 1024 - 4096$

Yang-Mills in Milne space



$g = 0.1$



$g = 0.5$

EG, PRL 111, 232301, 2013 (arXiv 1307.2214)

$N_T = 64$, $N_\eta = 128$

Conclusions, Problems & Perspectives

- CTP is useful in thinking about *initial value problems*
 - Conceptually
 - Practically
- Strong classical + 1st order quantum correction can be done within classical physics – Perfect way to simulate QFT in real time
- Difficulties: Higher order corrections
 - Feynman diagrams are useful conceptually but not computationally – Fourier transform does not produce $\delta(E_i - E_f)$ when $t_i < t < t_f$
 - $\phi_q^3 \phi_c$ terms are intrinsically quantum – Hard to include in purely classical evolution
 - Need to calculate $\delta\rho_W$
- Lots of interesting problems still to be considered:
 - Glasma evolution with quantum corrections
 - Field to Particle transition
 - NNLO JIMWLK
 - Fermions? ...