Far from equilibrium particle production and thermalization: From cosmology to heavy-ion collisions



The Big Bang and the little bangs - Non-equilibrium phenomena in cosmology and in heavy-ion collisions, CERN TH, August 2016

Content

- Particle production and thermalization process from extreme conditions:
- Inflaton (p)reheating as a well understood quantum example
 → Bose condensation and fermion production far from equilibrium
- Heavy-ion collisions in the high-energy limit
 → Gluon and quark production using real-time lattice simulations
- Universality far from equilibrium and the role of the nonperturbative infrared regime
- Anomalous fermion production in gauge theories

Space-time evolution of the Universe versus ultra-relativistic nuclear collisions in the laboratory

Schematic space-time evolution histories:



WMAP Science Team

• initial-state quantum fluctuations \rightarrow particle production and thermalization \rightarrow hierarchy of decoupling processes \rightarrow experimentally observed final state

P. Sorensen & C. Shen

Particle production and thermalization process from extreme conditions

Consider general class of models (including lattice gauge theories) with bosons coupled to Dirac fermions:

$$\mathcal{L} = \frac{1}{2} \partial \Phi^* \partial \Phi - V(\Phi) + \sum_{k}^{N_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{pmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{bmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{pmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{pmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (MP_L + M^*P_R) \Psi_k \end{pmatrix} \\ \xrightarrow{\mathcal{N}_f} \begin{bmatrix} i \bar{\Psi}_k \gamma^\mu \partial_\mu$$

- starting from vacuum initial conditions for N_f fermion flavors
- with extreme initial conditions for bosonic fields:

interaction strength field² expectation value
$$\frac{\lambda \cdot \langle \Phi^2 \rangle}{Q^2} \sim 1$$
 characteristic energy/momentum²

→ Extreme conditions can enhance the loss of details about microscopic properties (coupling strengths, initial conditions, ...)

Examples of extreme initial conditions



Preheating after inflation

Parametric resonance preheating: Kofman, Linde, Starobinsky...

Pure state density matrix ρ_0

Large inflaton field amplitude ($\lambda \lesssim 10^{-12}$):

$$\langle \Phi \rangle$$
(*t*) = Tr($ho_0 \Phi$) ~ 0.3 $M_P \gg m/\sqrt{\lambda}$

with vacuum fluctuations





Heavy-ion collisions in the high-energy limit

CGC: Lappi, McLerran, Dusling, Gelis, Venugopalan, Epelbaum...

Large initial gauge fields:

 $\langle A \rangle \sim Q_s/g$

with fluctuations:

$$\langle \delta A^2 \rangle \ \sim \ Q_s^2$$





Preheating and turbulent thermalization: A well understood *quantum* example



Beyond instability regime: Non-perturbative bosonic sector



1) over-occupation at the end of instability regime preempts perturbative power counting

2) development of non-perturbative infrared/Bose condensation during turbulent regime

 \rightarrow 1/N expansion to NLO (2PI) with quantum effective action:

Berges NPA 699 (2002) 847; Aarts et al. PRD 66 (2002) 045008

infinite (geometric) series at NLO in the number of inflaton field components

fermion propagator

Turbulent regime: Self-similar scaling behavior



Fermion production amplification from highly occupied bosons

Semi-classical:
$$iD_{0,ij}^{-1}(x,y) = \left[i\gamma^{\mu}\partial_{\mu} - m_{\psi} - \frac{g}{N_f}\phi(\mathbf{t})\right] \delta^{(4)}(x-y) \delta_{ij}$$

Baacke, Heitmann, Pätzold, PRD 58 (1998) 125013; Greene, Kofman, PLB 448 (1999) 6; Giudice, Peloso, Riotto, Tkachev, JHEP 9908 (1999) 014; Garcia-Bellido, Mollerach, Roulet, JHEP 0002 (2000) 034; ...



Berges, Gelfand, Pruschke, PRL 107 (2011) 061301



 strongly enhanced fermion production due to high boson cccupancies!

 backreaction on bosons controlled by small ~g²

$$\phi=\phi_0\,\sqrt{6N_s/\lambda}\,$$
 , ${\sf m}_\psi$ = 0

Impact of bosons on nonequilibrium fermion spectral function

$$\rho(x,y) = i \left\langle \{\psi(x), \bar{\psi}(y)\} \right\rangle \qquad \checkmark \qquad \rho_V^{\mu} = \frac{1}{4} \operatorname{tr} \left(\gamma^{\mu} \rho\right) \quad \text{vector components} \\ \rho_S = \frac{1}{4} \operatorname{tr} \left(\rho\right) \qquad \text{scalar component} \end{cases}$$

quantum field anti-commutation relation:

$$-i\rho_V^0(t,t;\mathbf{p}) = 1$$



Real-time lattice simulations with fermions

$$\mathcal{L} = \frac{1}{2} \partial \Phi^* \partial \Phi - V(\Phi) + \sum_{k}^{N_f} \left[i \bar{\Psi}_k \gamma^\mu \partial_\mu \Psi_k - \bar{\Psi}_k (M P_L + M^* P_R) \Psi_k \right]$$

Highly-occupied bosons at weak coupling: employ classical-statistical simulations Fermion quantum fluctuations:

$$\int \prod_{k} D\Psi_{k}^{+} D\Psi_{k} e^{i \int \mathcal{L}(\Phi, \Psi^{+}, \Psi)} \rightarrow \qquad \partial_{x}^{2} \Phi(x) + V'(\Phi(x)) + N_{f} J(x) = 0$$
$$J(x) = J^{S}(x) + J^{PS}(x) \qquad \qquad J^{S}(x) = -g \left\langle \bar{\Psi}(x)\Psi(x) \right\rangle = g \operatorname{Tr} D(x, x),$$
$$J^{PS}(x) = -g \left\langle \bar{\Psi}(x)\gamma^{5}\Psi(x) \right\rangle = g \operatorname{Tr} D(x, x)\gamma^{5}$$

For classical $\Phi(x)$ the exact equation for the fermion D(x,y) reads:

$$(i\gamma^{\mu}\partial_{x,\mu} - m + g\operatorname{Re}\Phi(x) - ig\operatorname{Im}\Phi(x)\gamma^{5})D(x,y) = 0$$

Aarts, Smit; Borsanyi, Hindmarsh; Berges, Gelfand, Pruschke; Saffin, Tranberg; Kasper, Hebenstreit; Mace, Mueller, Schlichting, Sharma, Tanji,...

Comparing lattice simulations to NLO quantum results





Heavy-ion collisions in the high-energy limit



JB, Schenke, Schlichting, Venugopalan, NPA931 (2014) 348 for initial spectrum from Epelbaum, Gelis, PRD88 (2013) 085015. Plasma instabilities from wide range of initial conditions:

Mrowczynski; Rebhan, Romatschke, Strickland; Arnold, Moore, Yaffe; Bödecker; Attems, ... Romatschke, Venugopalan; J.B., Scheffler, Schlichting, Sexty; Fukushima, Gelis ...

Overoccupied non-Abelian plasma



To discuss attractor: Initial over-occupation described by family of distributions at τ_0 (read-out in 'Coulomb gauge')



(controls "prolateness" or "oblateness" of initial momentum distribution)

Nonthermal fixed point

Evolution in the `anisotropy-occupancy plane'



Early stage of `bottom-up'* scaling emerges as a consequence of the attractor *Baier et al (BMSS), PLB 502 (2001) 51

Self-similar evolution



Scaling exponents: $\alpha = -2/3$, $\beta = 0$, $\gamma = 1/3$ and scaling distribution function f_s:

$$f(\mathbf{p}_{\mathrm{T}},\mathbf{p}_{\mathrm{z}},\tau) = (Q\tau)^{\alpha} f_{S} \Big((Q\tau)^{\beta} \mathbf{p}_{\mathrm{T}}, (Q\tau)^{\gamma} \mathbf{p}_{\mathrm{z}} \Big)$$

nonthermal fixed-point distribution

Comparing gauge and scalar field theories

with longitudinal expansion



- For gauge & scalar fields: range of *thermal-like transverse spectrum* ~1/p_T even as longitudinal distribution is being `squeezed'
- Strongly enhanced infrared regime for scalars: *inverse particle cascade leading to Bose condensation*, $\sim 1/p^5$ as in *isotropic superfluid turbulence*
- At latest available times for scalars a flat distribution for $p_T\gtrsim Q$ emerges
 - J.B., Boguslavski, Schlichting, Venugopalan, Phys. Rev. Lett. 114 (2015) 6, 061601

Universality far from equilibrium

• Same universal exponents and scaling function in common $1/p_{\tau}$ range

$$\alpha = -2/3$$
 , $\beta = 0$, $\gamma = 1/3$



 \rightarrow Remarkably large universality class far from equilibrium!

Some puzzles and challenges

• The scaling solution seems well understood for the gauge theory (BMSS) - but the corresponding kinetic theory arguments fail for the scalar theory

Consider the Boltzmann equation

$$[\partial_{\tau} - \frac{p_Z}{\tau} \partial_{p_Z}]f(p_T, p_Z, \tau) = C[f](p_T, p_Z, \tau)$$

with a self-similar evolution

$$f(p_T, p_Z, \tau) = (Q\tau)^{\alpha} f_S((Q\tau)^{\beta} p_T, (Q\tau)^{\gamma} p_Z)$$

→ Non-thermal fixed point solution $(f \gg 1)$

$$[\alpha + \beta p_T \partial_{p_T} + (\gamma - 1) p_Z \partial_{p_Z}] f_S(p_T, p_Z) = Q^{-1} C[f_S](p_T, p_Z)$$

→ Scaling exponents determined by scaling relations for

- Small angle elastic scattering
- Energy conservation
- Particle number conservation $(\alpha 2\beta \gamma = -1)$

tering $(2\alpha - 2\beta + \gamma = -1)$ for gauge theory $(\alpha - 3\beta - \gamma = -1)$ vation $(\alpha - 2\beta - \gamma = -1)$

 $\rightarrow \alpha = -2/3, \beta = 0, \gamma = 1/3$ in excellent agreement with lattice data!

However: No dominance for small angle scattering in scalar theory expected! More general principle underlying nonthermal fixed point? In the weak-coupling limit, kinetic theory is expected to have an overlapping range of validity with classical-statistical simulations



kinetic theory

However, kinetic theory seems not to reproduce important quantities such as P_L / P_T characterizing isotropization of the longitudinally expanding plasma:



- Scaling behavior of P_L / P_T very similar for scalar and gauge field simulations!
- In scalar theory behavior of P_L / P_T known to arise from infrared contributions (Bose condensation)

→ Nonperturbative despite (weak) coupling parameter

Perturbative estimates extrapolated beyond the weak-coupling regime suggest absence of transient universal scaling behavior

Gauge: Kurkela, Zhu, PRL 115 (2015) 182301: Scalar: Epelbaum, Gelis, Jeon, Moore, Wu, JHEP 09 (2015) 117



However, no such indications (yet no expansion) from nonperturbative estimates in scalar quantum field theory JB, Wallisch, arXiv:1607.02160 (and holographic superfluids?) Ewerz et al., JHEP 18 (2015) 1505 (cf. also Adams, Chesler, Liu, Science 341 (2013) 368)

By now, detailed understanding of the infrared dynamics (Bose condensation) in scalar QFT (NLO 1/M) 02504 Boguslavski

condensation time:

$$t_f \simeq t_0 \left(\frac{|\psi_0|^2(t_f)}{f(t_0,0)}\right)^{1/\alpha} V^{1/\alpha}$$



(2015)

92

PRD

ЛВ,

Piñeiro Orioli

However, no such understanding in gauge theories – despite indications for significant infrared contributions to gauge invariant quantities (P_1 / P_T)

Quark production from over-occupied gauge fields



Anomalous fermion production with real-time lattice simulations



Anomaly induced axial charge can persist even beyond decoherence time

Application: 'Chiral magnetic effect' in QED

Mueller, Hebenstreit, JB, PRL 117 (2016) 061601



Dynamical tracking behavior of $\varphi = \measuredangle(\mathbf{E}, \mathbf{B})$ \rightarrow collinear configs with max. anom. current



\rightarrow see also talk by Florian Hebenstreit

(cf. Kharzeev et al. Nature Physics (2016), Mueller, Schlichting, Sharma arXiv:1606.00342)



Conclusions

Extreme conditions \rightarrow loss of details about microscopic model & initial conditions:

• New dynamical universality classes:



• Anomalous fermion production with tracking behavior $(E > E_c = m^2/e)$:



Macroscopic fields, condensates and fluctuations

• In a quantum theory the field amplitude corresponds to the expectation value of a (here relativistic, real) Heisenberg field operator $\Phi(t, \mathbf{x})$:

$$\phi(t) = \langle \Phi(t, \mathbf{x}) \rangle \equiv \operatorname{Tr} \left\{ \varrho_0 \Phi(t, \mathbf{x}) \right\}$$

time-dependent expectation value density operator at some `initial' time t = 0

• Fluctuations derive from correlation functions, e.g. spatially homogeneous:

Preheating: Insensitivity to initial condition details

Example: *Inflaton'* $\lambda \phi^4$ theory ($\lambda \ll 1$), $\phi = \phi_0 + \delta \phi$

1. Large initial field:

2. High occupancy:



Schematic behavior near nonthermal fixed point: dual cascade



Particle versus energy transport

 $K \ll Q \ll \Lambda$

number conservation:

energy conservation:

$$\dot{n}_Q = \dot{n}_K + \dot{n}_\Lambda \qquad \qquad Q\dot{n}_Q = K\dot{n}_K + \Lambda\dot{n}_\Lambda$$
$$\dot{n}_K = \frac{\Lambda - Q}{\Lambda - K}\dot{n}_Q \simeq \dot{n}_Q \qquad \qquad \dot{n}_\Lambda = \frac{Q - K}{\Lambda - K}\dot{n}_Q \simeq \frac{Q}{\Lambda}\dot{n}_Q$$
$$\Rightarrow \Lambda \dot{n}_\Lambda \simeq Q\dot{n}_Q$$

Particles are transported towards lower scales, energy towards higher scales

Self-similarity

$$f(t, \mathbf{p}) = s^{\alpha/\beta} f(s^{-1/\beta} t, s\mathbf{p})$$

$$f(t,\mathbf{p}) = t^{\alpha} f_S(t^{\beta}\mathbf{p})$$

Time-independent scaling function:

$$f_S(t^{\beta}\mathbf{p}) \equiv f(1, t^{\beta}\mathbf{p})$$

Scaling exponents α and β determine rate and direction of transport:

$$K(t_1) = K_1 \qquad \Rightarrow \qquad K(t) = K_1 (t/t_1)^{-\beta}$$
$$f(t, K(t)) \sim t^{\alpha}$$

e.g. $\alpha > 0$, $\beta > 0$: particle transport towards lower momentum scales

Self-similar dynamics: infrared scaling

 $f(t, \mathbf{p}) = t^{\alpha} f_S(t^{\beta} \mathbf{p})$, $\alpha = 1.51 \pm 0.13$, $\beta = 0.51 \pm 0.04$



Mass scale separating non-relativistic infrared regime



- non-relativistic infrared dynamics expected because of the generation of a mass gap (condensate + medium)
- \rightarrow relativistic & non-relativistic field theories have same infrared scaling

Piñeiro Orioli, Boguslavski, Berges, PRD 92 (2015) 025041

Universal scaling form of the distribution function



Estimating scaling properties

$$\frac{\partial f(t, \mathbf{p})}{\partial t} = C[f](t, \mathbf{p}) \quad \text{`collision integral'}$$

 $C[f](t, \mathbf{p}) = s^{-\mu} C[f](s^{-1/\beta}t, s\mathbf{p}) = t^{-\beta\mu} C[f_S](1, t^{\beta}\mathbf{p})$ $f(t, \mathbf{p}) = s^{\alpha/\beta} f(s^{-1/\beta}t, s\mathbf{p}) \qquad s^{-1/\beta}t = 1$

$$\frac{\partial}{\partial t} \left[t^{\alpha} f_{S}(t^{\beta} \mathbf{p}) \right] = t^{\alpha - 1} \left[\alpha + \beta \, \mathbf{q} \cdot \nabla_{\mathbf{q}} \right] f_{S}(\mathbf{q}) |_{\mathbf{q} = t^{\beta} \mathbf{p}}$$

Time-independent *fixed point equation:*

$$[\alpha + \beta \mathbf{p} \cdot \nabla_{\mathbf{p}}] f_S(\mathbf{p}) = C[f_S](1, \mathbf{p})$$

+ scaling relation:

$$\alpha - 1 = -\beta\mu$$

Micha, Tkachev, PRD 70 (2004) 043538

Conservation laws

$$n = \int \frac{d^d p}{(2\pi)^d} f(t, \mathbf{p}) = t^{\alpha - \beta d} \int \frac{d^d q}{(2\pi)^d} f_S(\mathbf{q})$$

 $\Rightarrow \quad particle \ conservation: \ \alpha = \beta d$

$$\epsilon = \int \frac{d^d p}{(2\pi)^d} \,\omega(\mathbf{p}) f(t, \mathbf{p}) = t^{\alpha - \beta(d+z)} \int \frac{d^d q}{(2\pi)^d} \,\omega(\mathbf{q}) f_S(\mathbf{q})$$

energy conservation: $\alpha = \beta(d+z)$

 \Rightarrow

$$\omega(\mathbf{p}) = s^{-z}\omega(s\mathbf{p})$$

Perturbative estimate (Gross-Pitaevskii)

gives scaling relation: $\alpha - 1 = -\beta \left[3d - (d+2) - 3\alpha/\beta \right]$

$$\begin{array}{c} \alpha = \beta d \\ \Rightarrow \quad nonrel. \ particle \ transport: \ \alpha = -\frac{d}{2}, \ \beta = -\frac{1}{2} \end{array}$$

Negative perturbative exponents do not account for inverse particle cascade!

Beyond perturbation theory: large-N expansion to NLO

$$\begin{split} \int \mathrm{d}\Omega^{2\leftrightarrow 2}(\mathbf{p},\mathbf{l},\mathbf{q},\mathbf{r}) &\longrightarrow \int d\Omega^{\mathrm{NLO}}[f](t,\mathbf{p},\mathbf{l},\mathbf{q},\mathbf{r}) \sim \int \frac{d^d l}{(2\pi)^d} \frac{d^d q}{(2\pi)^d} \frac{d^d r}{(2\pi)^d} \\ &\times (2\pi)^{d+1} \,\delta^{(d)}(\mathbf{p}+\mathbf{l}-\mathbf{q}-\mathbf{r}) \,\delta(\omega_{\mathbf{p}}+\omega_{\mathbf{l}}-\omega_{\mathbf{q}}-\omega_{\mathbf{r}}) \\ &\times g_{\mathrm{eff}}^2[f](t,\mathbf{p},\mathbf{q}) \,, \end{split}$$



FIG. 11. Illustration of different scattering channels. The vertex correction at NLO may be viewed as an effective interaction, which involves the exchange of an intermediate particle.

based on Berges Nucl. Phys. A 699 (2002) 847; Aarts et al. Phys. Rev. D 66 (2002) 045008

Vertex correction (NLO 1/N)

$$g_{\text{eff}}^2(t, \mathbf{p}, \mathbf{q}) \equiv \frac{g^2}{|1 + \Pi_{\text{nr}}^R(t, \omega_{\mathbf{p}} - \omega_{\mathbf{q}}, \mathbf{p} - \mathbf{q})|^2}$$

one-loop retarded self-energy:

$$\Pi_{\mathrm{nr}}^{R}(t,\omega,\mathbf{p}) = g \int \frac{\mathrm{d}^{d}q}{(2\pi)^{d}} f(t,\mathbf{p}-\mathbf{q}) \\ \times \left[\frac{1}{\omega_{\mathbf{q}}-\omega_{\mathbf{p}-\mathbf{q}}-\omega-i\epsilon} + \frac{1}{\omega_{\mathbf{q}}-\omega_{\mathbf{p}-\mathbf{q}}+\omega+i\epsilon}\right]$$

scaling behavior:

$$\Pi_{\mathrm{nr}}^{R}(t,\omega_{\mathbf{p}},\mathbf{p}) = s^{\alpha/\beta-d+2} \Pi_{\mathrm{nr}}^{R}(s^{-1/\beta}t,\omega_{s\mathbf{p}},s\mathbf{p})$$

 $\alpha/\beta \geq d\;$, $\Pi^R_{\rm nr}(t,\omega_{\bf p},{\bf p})\gg 1\;$ in the infrared:

$$\Rightarrow \qquad g_{\text{eff}}^2(t,\mathbf{p},\mathbf{q},\mathbf{r}) = s^{-2(\alpha/\beta - d + 2)} g_{\text{eff}}^2(s^{-1/\beta}t,s\mathbf{p},s\mathbf{q},s\mathbf{r})$$

Scaling solution at NLO 1/N

$$C_{\mathrm{nr}}^{\mathrm{NLO}}[f](t,\mathbf{p}) = s^{-(2-\alpha/\beta)} C_{\mathrm{nr}}^{\mathrm{NLO}}[f](s^{-1/\beta}t,s\mathbf{p})$$
$$= t^{\alpha-2\beta} C_{\mathrm{nr}}^{\mathrm{NLO}}[f_S](1,t^{\beta}\mathbf{p})$$

gives scaling relation: $\alpha - 1 = \alpha - 2\beta$

nonrel. transport:
$$\beta = \frac{1}{2}$$
 of $\begin{cases} particles: \alpha = d/2 \\ energy: \alpha = (d+2)/2 \end{cases}$

Positive nonperturbative exponents can describe inverse cascade! NLO result in good agreement with full numerical simulation

Piñeiro Orioli, Boguslavski, Berges, PRD 92 (2015) 025041

Self-similar dynamics from classical-statistical simulations

 $f(t, \mathbf{p}) = t^{\alpha} f_S(t^{\beta} \mathbf{p})$, $\alpha = 1.66 \pm 0.12$, $\beta = 0.55 \pm 0.03$



Condensation far from equilibrium

$$F(t, t', \mathbf{x} - \mathbf{x}') = \frac{1}{2} \langle \psi(t, \mathbf{x}) \psi^*(t', \mathbf{x}') + \psi(t', \mathbf{x}') \psi^*(t, \mathbf{x}) \rangle$$
$$f(t, \mathbf{p}) + (2\pi)^3 \delta^{(3)}(\mathbf{p}) |\psi_0|^2(t) \equiv \int d^3 x \, e^{-i\mathbf{p}\mathbf{x}} F(t, t, \mathbf{x})$$
$$\bigvee_{\text{volume: } (2\pi)^3 \delta(\mathbf{0}) \to V}$$



Time: t

Condensation time



Nonthermal fixed point in non-Abelian gauge theory

Interpret scaling condition with **energy/number conserving*** Fokker-Planck-type dynamics for the collisional broadening of longitudinal momentum distribution:

$$C^{(\text{elast})}[p_T, p_z; f] = \hat{q} \; \partial_{p_z}^2 f(p_T, p_z, t)$$

with momentum diffusion parameter: $\hat{q} \sim \alpha_S^2 \int \frac{d^2 p_T}{(2\pi)^2} \int \frac{dp_z}{2\pi} f^2(p_T, p_z, t)$

1)
$$\mu = 3\alpha - 2\beta + \gamma$$
 $\xrightarrow{\alpha - 1 = \mu(\alpha, \beta, \gamma)}$ $2\alpha - 2\beta + \gamma + 1 = 0$
2) number conservation \longrightarrow $\alpha - 2\beta - \gamma + 1 = 0$
3) energy conservation $\xrightarrow{remarkable agreement}} \alpha - 3\beta - \gamma + 1 = 0$
 $\alpha = -2/3$, $\beta = 0$, $\gamma = 1/3$ with lattice data!

*cf. early stages of Baier, Mueller, Schiff, Son, PLB 502 (2001) 51 ('BMSS')

A closer look at scalar N-component field theory

with longitudinal expansion



• Isotropic scaling properties and condensation in the infrared regime