

Learning about neutron stars from pulsar precession observations

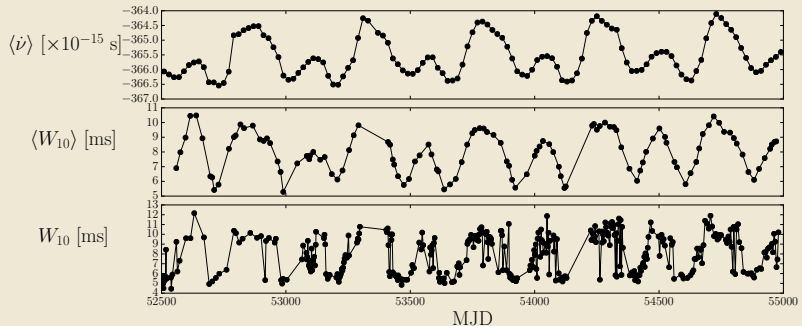
NewCompStar, Istanbul, 2016

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Data courtesy of Lyne et al. (2010): *Switched Magnetospheric Regulation of Pulsar Spin-Down*

Potential explanations:

- ▶ **Precession:** conflict with vortex-pinning of superfluid core
- ▶ **Magnetospheric switching**

We would like to quantify how well the two models fit the data. To do this we will use Bayes theorem:

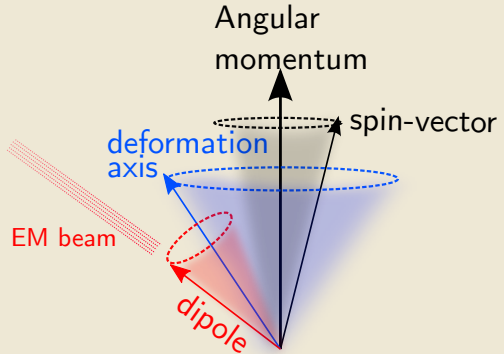
$$P(\mathcal{M}|\mathbf{y}_{\text{obs}}) = P(\mathbf{y}_{\text{obs}}|\mathcal{M}) \frac{P(\mathcal{M})}{P(\mathbf{y}_{\text{obs}})}.$$

The odds ratio:

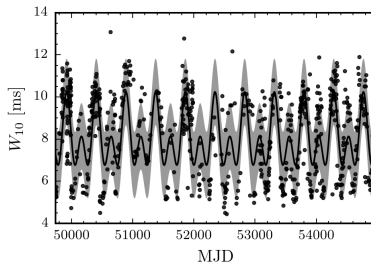
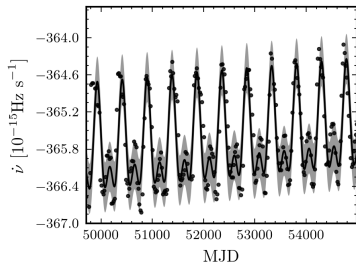
$$\mathcal{O} = \frac{P(\mathcal{M}_A|\mathbf{y}_{\text{obs}})}{P(\mathcal{M}_B|\mathbf{y}_{\text{obs}})} = \frac{P(\mathbf{y}_{\text{obs}}|\mathcal{M}_A)}{P(\mathbf{y}_{\text{obs}}|\mathcal{M}_B)} \underbrace{\frac{P(\mathcal{M}_A)}{P(\mathcal{M}_B)}}_{=1}$$

Calculate the *marginal-likelihood* $P(\mathbf{y}_{\text{obs}}|\mathcal{M}_A)$ using Markov chain Monte-Carlo method.

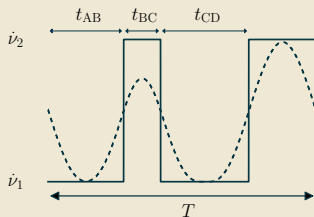
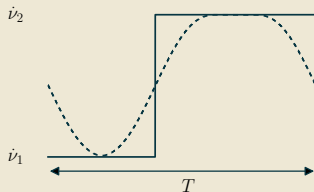
- ▶ Precession is a geometric effect in non-spherical bodies where the spin-vector is misaligned from the angular momentum
- ▶ It will produce periodic modulations of:
 - ▶ the beam width
 - ▶ the spin-down rate
- ▶ Complicated interaction with the EM torque can amplify the spin-down modulations

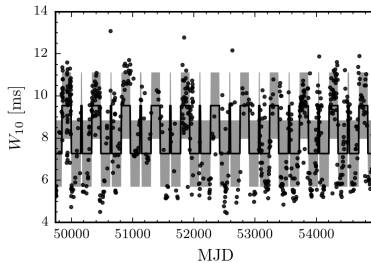
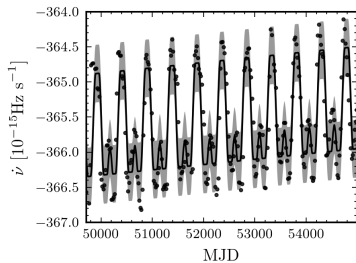


See for example: *Jones & Andersson (2001)*, *Link & Epstein (2001)*, *Akgun et al. (2006)* *Zanazzi & Lai (2015)*, *Arzamasskiy et al. (2015)*



- ▶ Lyne et al. (2010): the magnetosphere undergoes periodic switching between two states
- ▶ The smooth modulation in the spin-down is due to time-averaging of this underlying spin-down model
- ▶ To explain the *double-peak*, Perera (2015) suggested four times were required





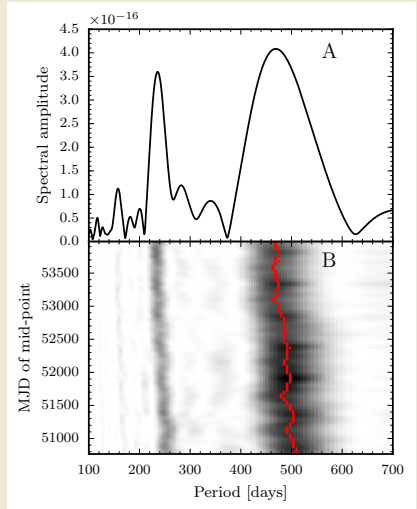
- ▶ Results published in *arXiv:1510.03579* with a conclusion

$$\frac{P(\text{precession}|\text{data})}{P(\text{switching}|\text{data})} = 10^{2.7 \pm 0.5},$$

favouring the precession interpretation

- ▶ This odds-ratio is for *simple* models with unbiased priors
- ▶ Questions for the precession model:
 - ▶ Connection with the pinning of the superfluid core
 - ▶ Presence of a glitch just after our data ends
- ▶ We can extend the models. . .

- ▶ We noticed different sections of data gave different modulation periods
- ▶ Studied with a *Lomb-Scargle* periodogram
- ▶ Finds the expected two peaks
- ▶ Precession period decays from 503 to 467 days over a period of 3211 days



- ▶ Modulation period in the precession model is given by

$$\tau_p = \left| \frac{P}{\epsilon} \right| \cos \theta$$

- ▶ Can rule out variation in P due to spin-down: not large enough and makes the precession period longer not shorter
- ▶ Can rule out variation in θ as there is no corresponding change in the amplitude of modulations

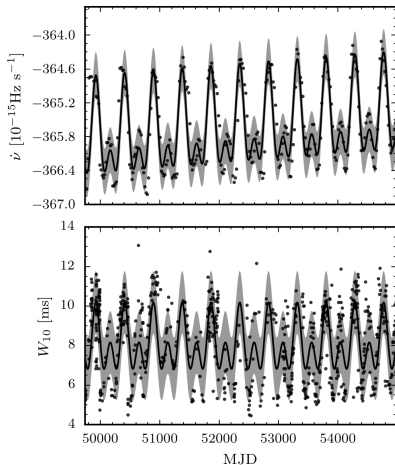
Secular evolution of $\epsilon(t)$

$$\epsilon(t) = \epsilon_0 + \dot{\epsilon}t$$

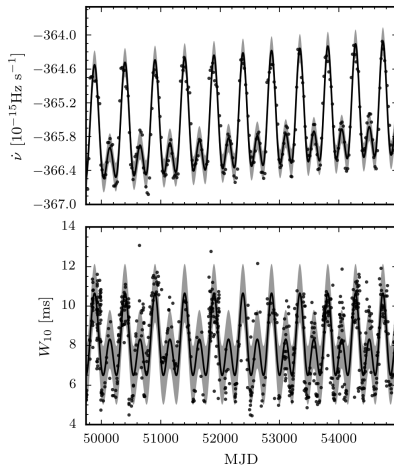
Discreet jumps in $\epsilon(t)$

$$\epsilon(t) = \epsilon_0 \left(1 + \sum_j^N H(t, t_j) \Delta_j \right)$$

Basic precession

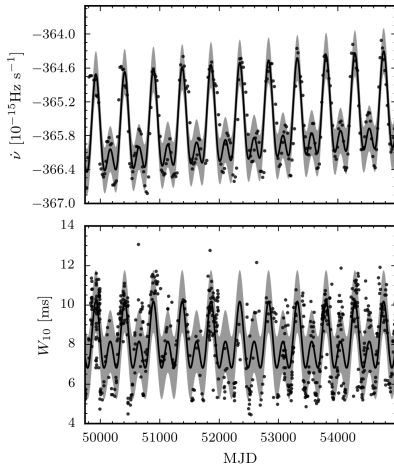


Secular evolution

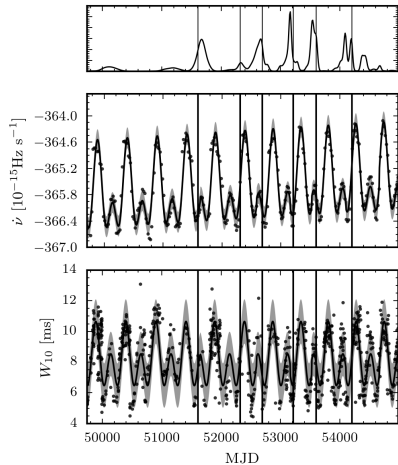


$$\frac{P(\text{secular evolution of } \epsilon(t)|\text{data})}{P(\text{basic precession}|\text{data})} = 10^{74.55 \pm 0.8}$$

Basic precession



N=6 discreet jumps



$$\frac{P(6 \text{ discreet jumps in } \epsilon(t)|\text{data})}{P(\text{basic precession}|\text{data})} = 10^{74.57 \pm 1.1}$$

- ▶ Discreet jumps have a 'preference' for the point in the precessional phase with which they occur: rules out external models.
- ▶ Fractional size of the jumps is $\sim 10^{-2}$
- ▶ The odds-ratio between these models is

$$\frac{P(\text{secular evolution of } \epsilon(t) \mid \text{data})}{P(\text{6 discreet jumps in } \epsilon(t) \mid \text{data})} \approx 1$$

We need a physical model to explain why ϵ changes on a timescale of 200 yrs.

- ▶ **Accretion:** from back of the envelope calculation would require $\dot{M} \approx 10^{-11} M_{\odot}/\text{yr}$.
- ▶ **Evolution of the magnetic field:** requires internal magnetic field to vary on a timescale of ~ 400 yr.
- ▶ **Evolution of the pinned superfluid:** requires

$$\frac{I_{\text{Pinned superfluid}}}{I_{\text{total}}} \leq 10^{-8}$$

and the amount of pinned superfluid to *increase* on a timescale of 200 yrs.

Conclusion

Can rule out some models, but *decreasing* modulation period is difficult to understand in the context of precession.

- ▶ We have found strong evidence in support of a increasing modulation frequency in PSR B1828-11
- ▶ Under the precession interpretation this corresponds to an increase in the deformation $\epsilon(t)$
- ▶ Unclear exactly *how* $\epsilon(t)$ is changing
- ▶ New physical ideas needed?