

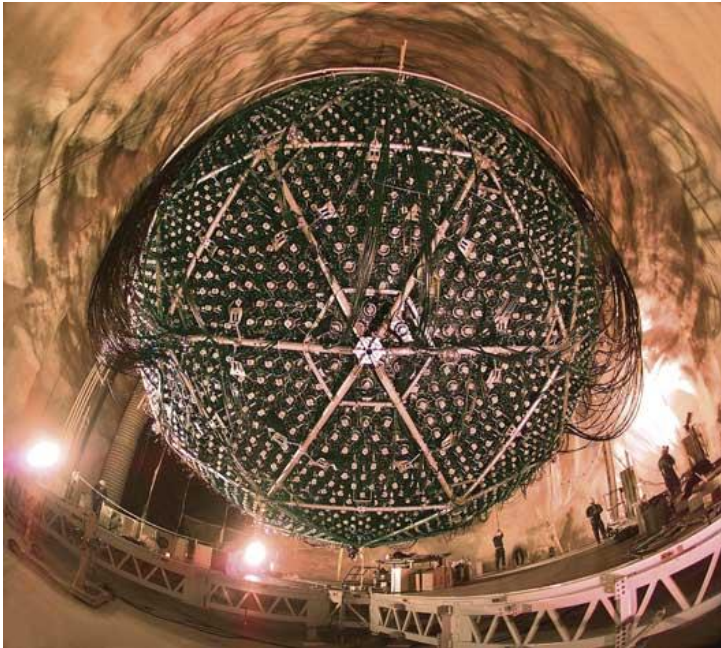


Searching for $0\nu\beta\beta$ with EXO-200 and nEXO

- Motivation for $\beta\beta$ search
- The EXO-200 and nEXO experiments

Thomas Brunner for the nEXO collaboration
CAP2016 – June 13, 2016

Neutrino oscillations



SNO, picture taken from <http://www.oit.on.ca>

Relative mass scale

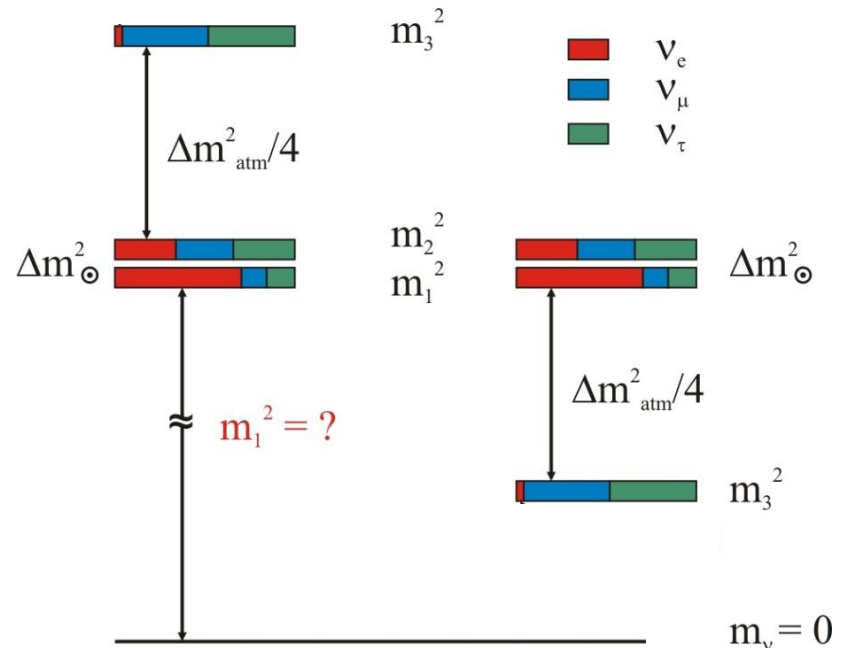
- Indicate a neutrino mass
- Determination of mixing angle θ_{ij}
- Indicate mass hierarchy
- Determination of δm^2

Pontecorvo–Maki–Nakagawa–Sakata matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_{m1} \\ \nu_{m2} \\ \nu_{m3} \end{pmatrix}$$

Normal Hierarchy

Inverted Hierarchy
(only if $m_1^2 \geq \Delta m_{\text{atm}}^2$)

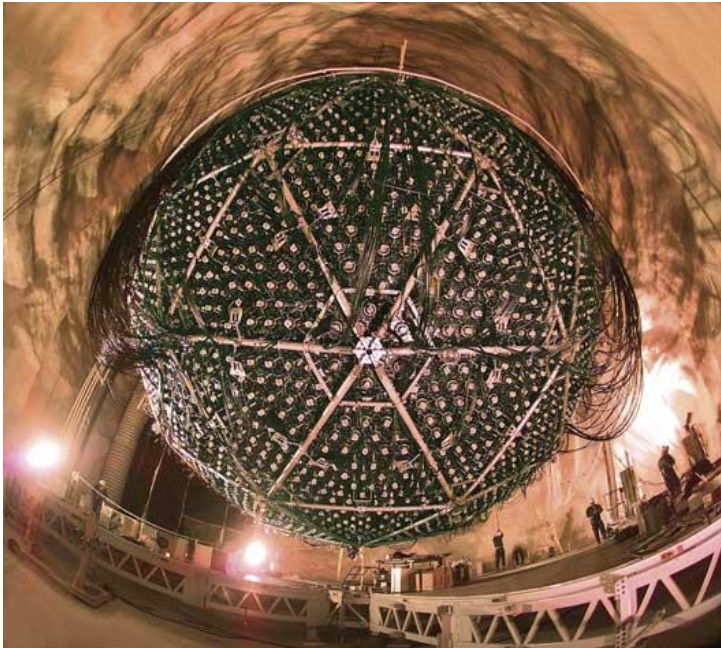


Neutrino oscillations



Pontecorvo–Maki–Nakagawa–Sakata matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_{m1} \\ \nu_{m2} \\ \nu_{m3} \end{pmatrix}$$



SNO, picture taken from <http://www.oit.on.ca>

Relative mass scale

- Indicate a neutrino mass
- Determination of mixing angle θ_{ij}
- Indicate mass hierarchy
- Determination of δm^2

What oscillation experiments cannot tell us about ν 's

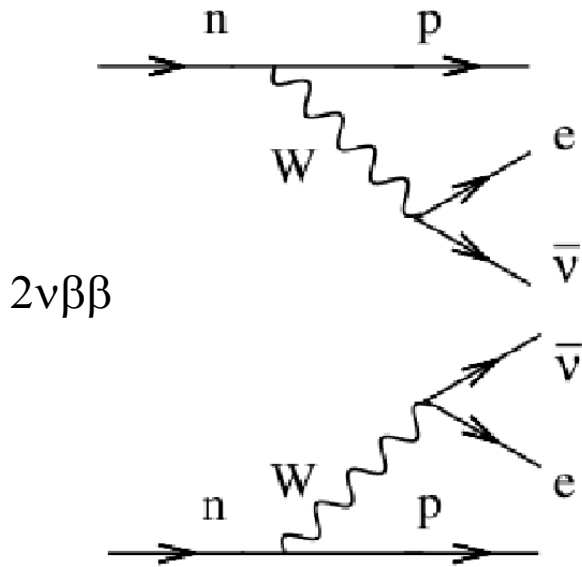
- What is the absolute mass scale
- Why is the neutrino mass so small?
- What is the nature of the ν : Dirac or Majorana?

→ Search for $0\nu\beta\beta$ decay

Double beta decay



M. Goeppert-Mayer, Phys. Rev. 48 (1935) 512



Double beta decay

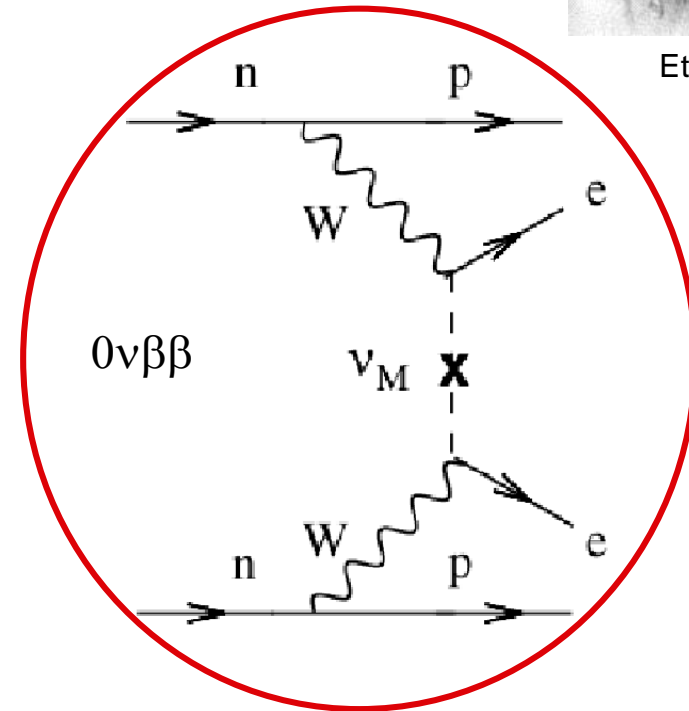
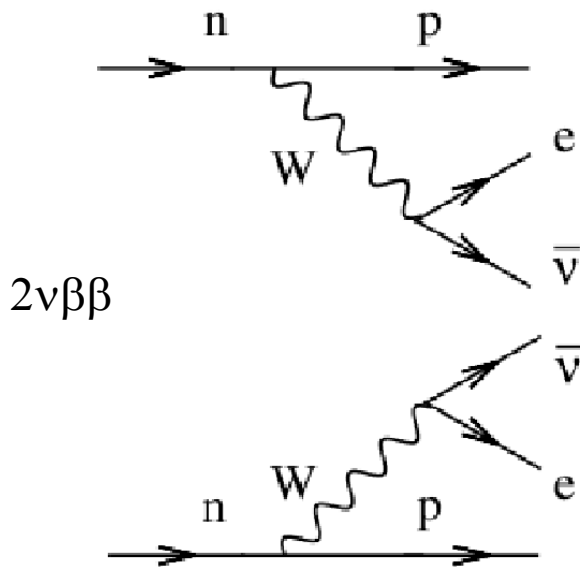


M. Goeppert-Mayer, Phys. Rev. 48 (1935) 512

The most promising approach to determine the nature of the neutrino!
Lepton number is violated in this decay!



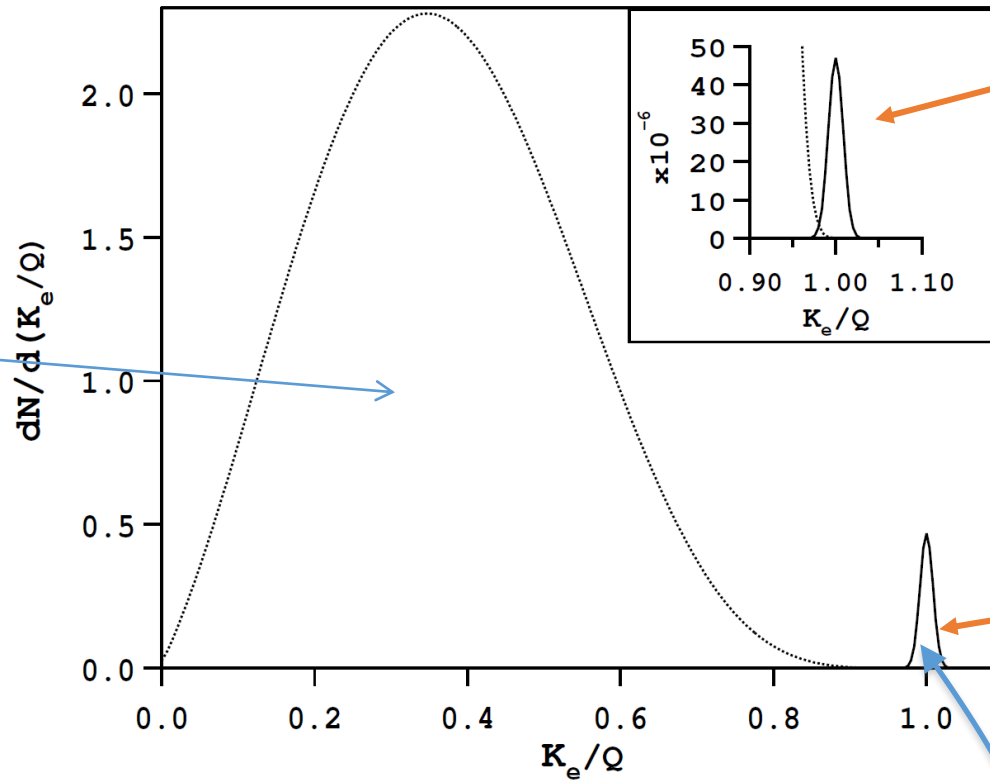
Ettore Majorana



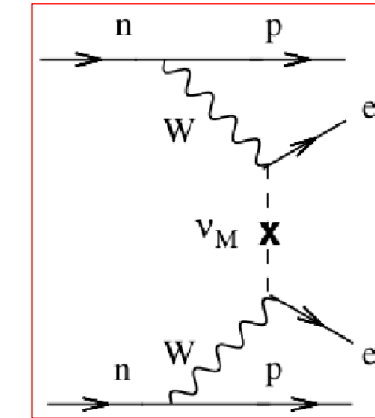
This process can only occur for a Majorana neutrino!

Neutrinoless double beta decay

[arXiv:hep-ph/0611243]



$0\nu\beta\beta$ peak
(normalized to 10^{-6})



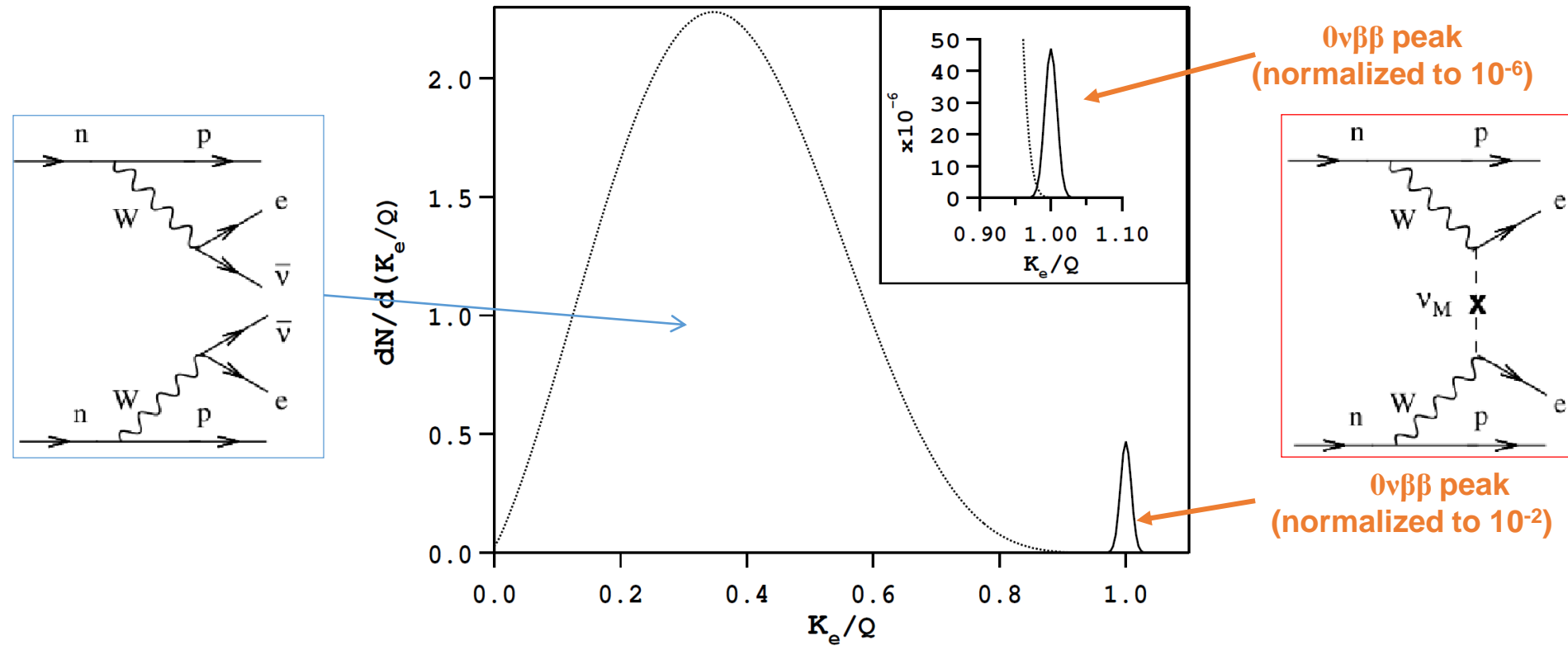
$0\nu\beta\beta$ peak
(normalized to 10^{-2})

kinetic energy K_e of the two electrons
in units of kinematic endpoint (Q)

Smearred by the energy resolution
of the hypothetical detector

Neutrinoless double beta decay

[arXiv:hep-ph/0611243]

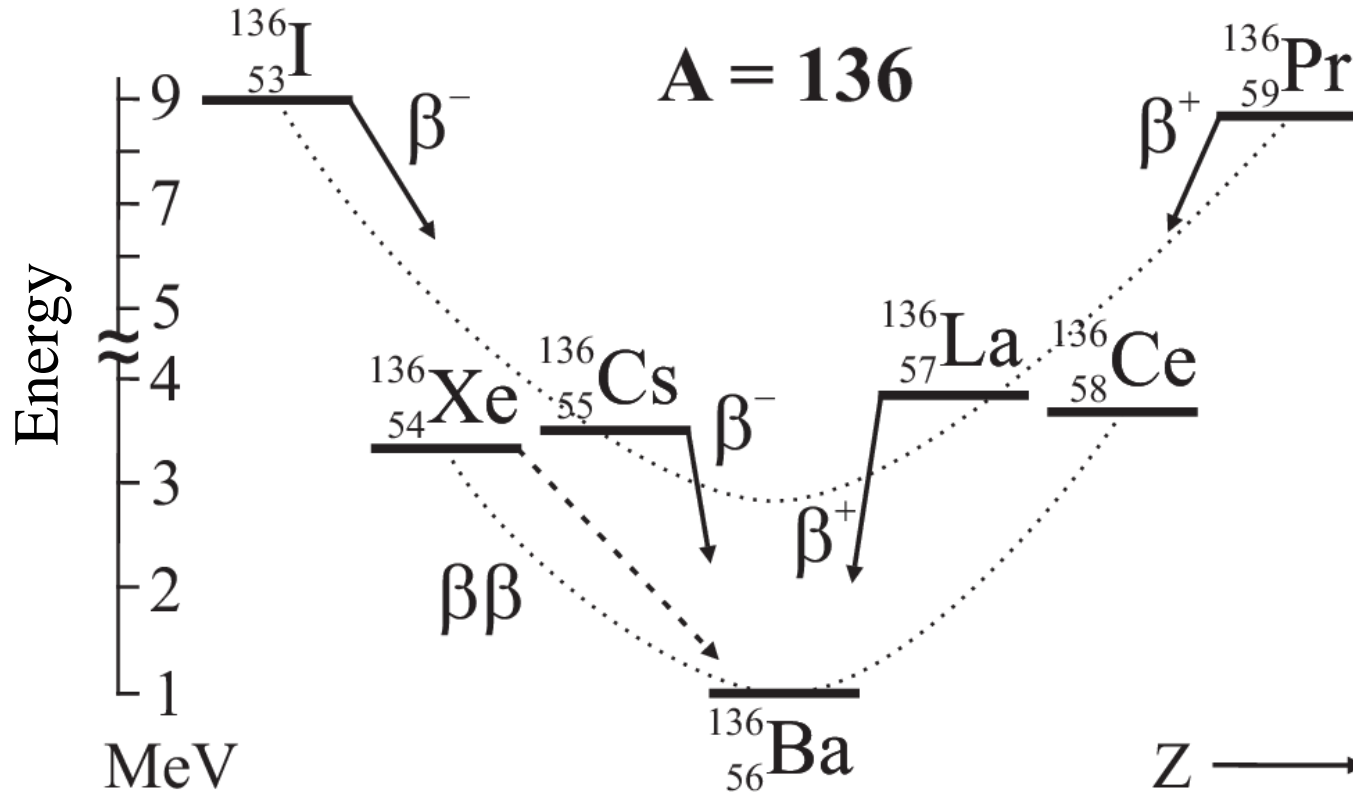


$$\left[T_{1/2}^{0\nu} \right]^{-1} = G^{0\nu} \left| M^{0\nu} \right|^2 \langle m_\nu \rangle^2$$

$G^{0\nu}$ is a phase space factor
 $M^{0\nu}$ is the nuclear matrix element

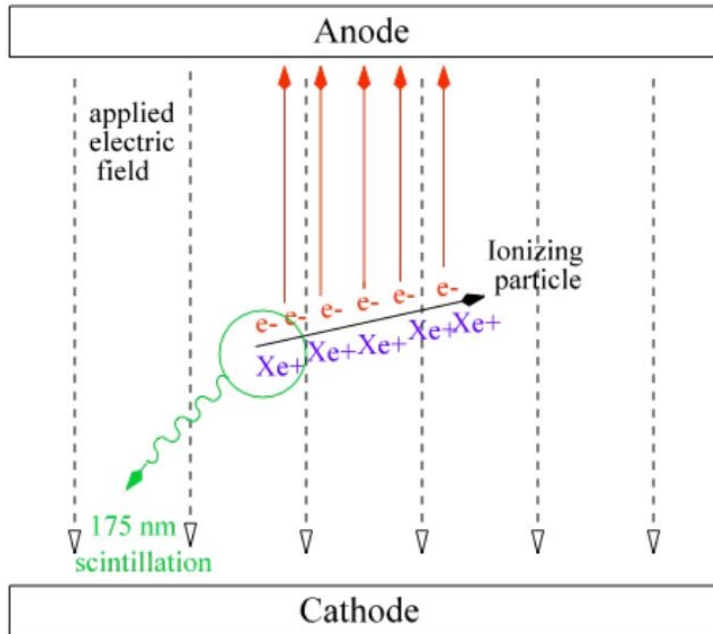
Effective Majorana mass: $\langle m_\nu \rangle = \left| \sum_i U_{ei}^2 m_i \varepsilon_i \right|$ (light neutrino exchange mechanism only)

Double Beta Decay



- If first-order beta decay is forbidden energetically or by spin, second-order double beta decay (a weak nuclear process) can be observed
- True for several isotopes such as: ^{48}Ca , ^{76}Ge , ^{130}Te , ^{136}Xe

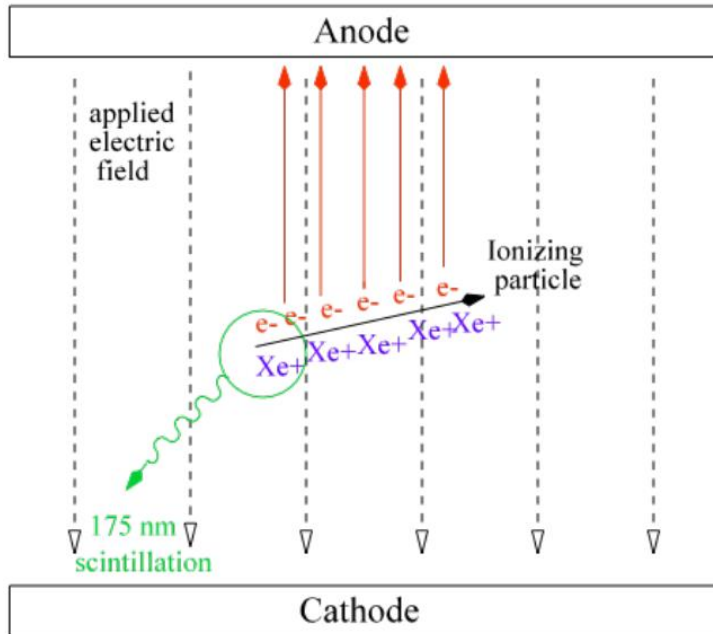
Searching for $0\nu\beta\beta$ in ^{136}Xe with EXO



Liquid-Xe Time Projection Chamber

- Liquid Xe at 168K
- Cryogenic electronics in LXe
- Detection of scintillation light and secondary charges
- 2D read out of secondary charges at segmented anode
- Full 3D event reconstruction:
 1. Energy reconstruction
 2. Position reconstruction
 3. Event Multiplicity

Searching for $0\nu\beta\beta$ in ^{136}Xe with EXO



Liquid-Xe Time Projection Chamber

- Liquid Xe at 168K
- Cryogenic electronics in LXe
- Detection of scintillation light and secondary charges
- 2D read out of secondary charges at segmented anode
- Full 3D event reconstruction:
 1. Energy reconstruction
 2. Position reconstruction
 3. Event Multiplicity



$T_{1/2}^{0\nu} > 10^{25}$ years !!

→ Need:

- high target mass
- high exposure
- low background rate
- good energy resolution

Natural radiation decay rates

A banana	~10 decays/s
A bicycle tire	~0.3 decays/s
1 l outdoor air	~1 decay/min
100 kg of ^{136}Xe (2ν)	~1 decay/10 min

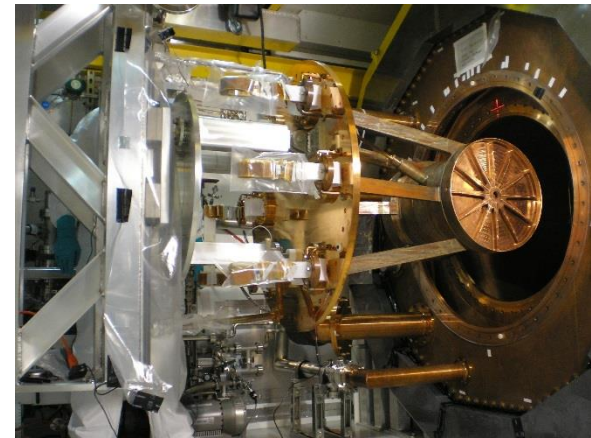
$0\nu\beta\beta$ decay	>1000 x rarer than $2\nu\beta\beta$
Age of universe	1.4×10^{10} years

Advantages of ^{136}Xe

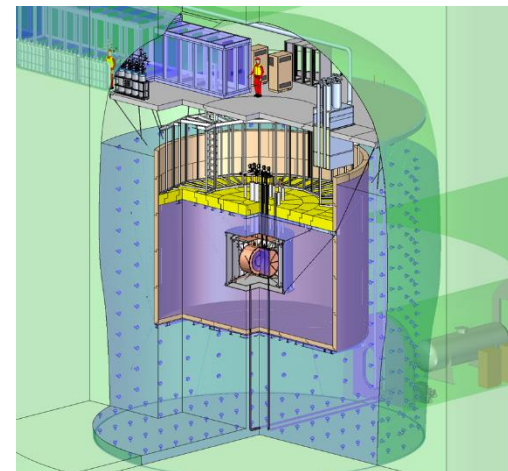
- **Easy to enrich**: 8.9% natural abundance but can be enriched relatively easily (better than growing crystals)
- **Can be purified** continuously, and reused
- **High $Q_{\beta\beta}$** (2458 keV): higher than most naturally occurring backgrounds
- **Minimal cosmogenic activation**: no long-life radioactive isotopes
- **Energy resolution**: improves using scintillation and charge anti-correlation
- **LXe self shielding**
- Background can be potentially reduced by **Ba⁺⁺ tagging**

Phased approach:

1. EXO-200: 200kg liquid-Xe TPC



2. nEXO: 5-ton liquid Xe TPC with Ba tagging option (SNO lab cryopit)

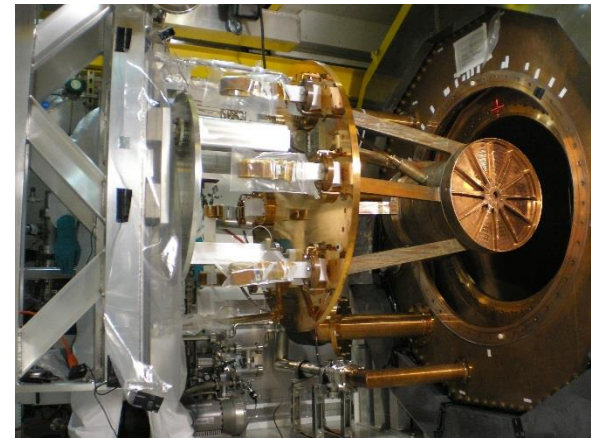


Advantages of ^{136}Xe

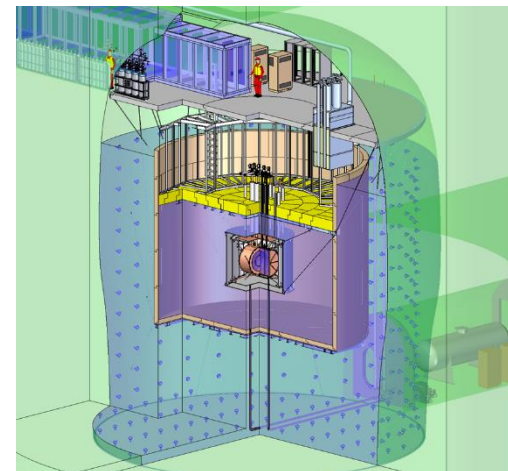
- **Easy to enrich**: 8.9% natural abundance but can be enriched relatively easily (better than growing crystals)
- **Can be purified** continuously, and reused
- **High $Q_{\beta\beta}$** (2458 keV): higher than most naturally occurring backgrounds
- **Minimal cosmogenic activation**: no long-life radioactive isotopes
- **Energy resolution**: improves using scintillation and charge anti-correlation
- **LXe self shielding**
- Background can be potentially reduced by **Ba⁺⁺ tagging**

Phased approach:

1. EXO-200: 200kg liquid-Xe TPC



2. nEXO: 5-ton liquid Xe TPC with Ba tagging option (SNO lab cryopit)

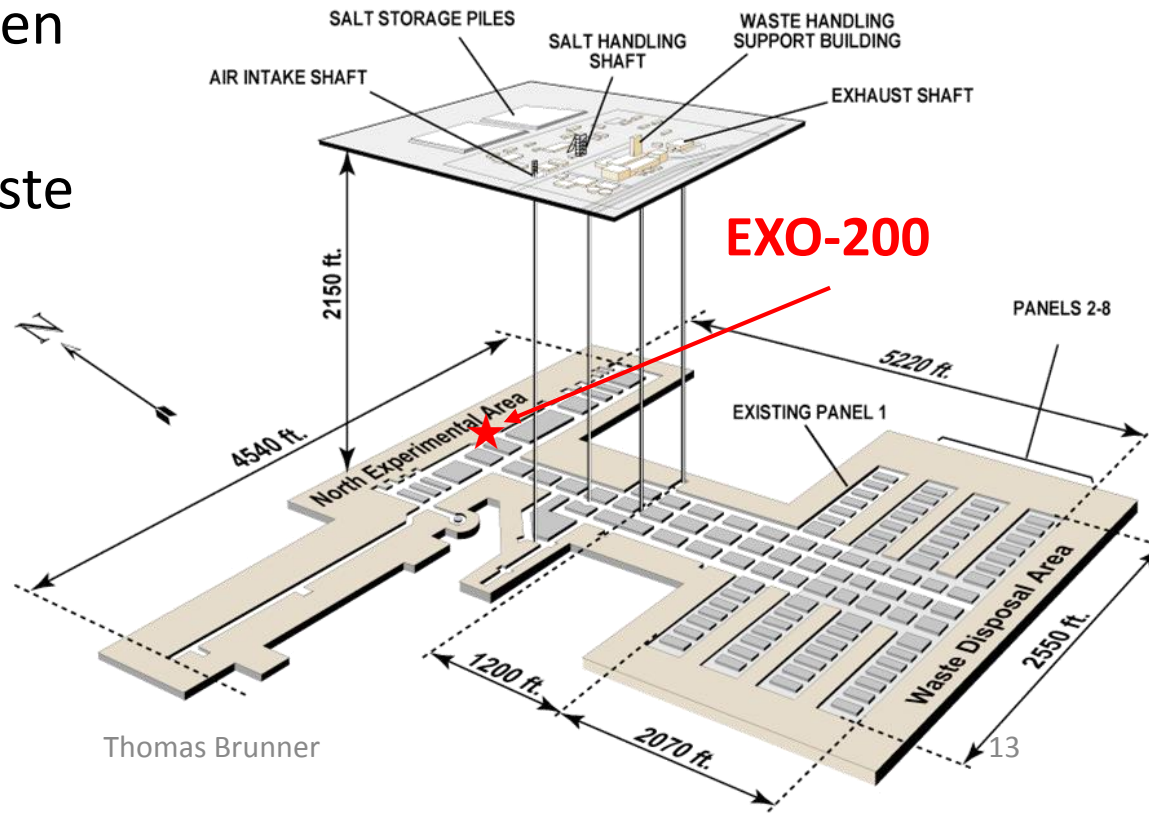


See talks by

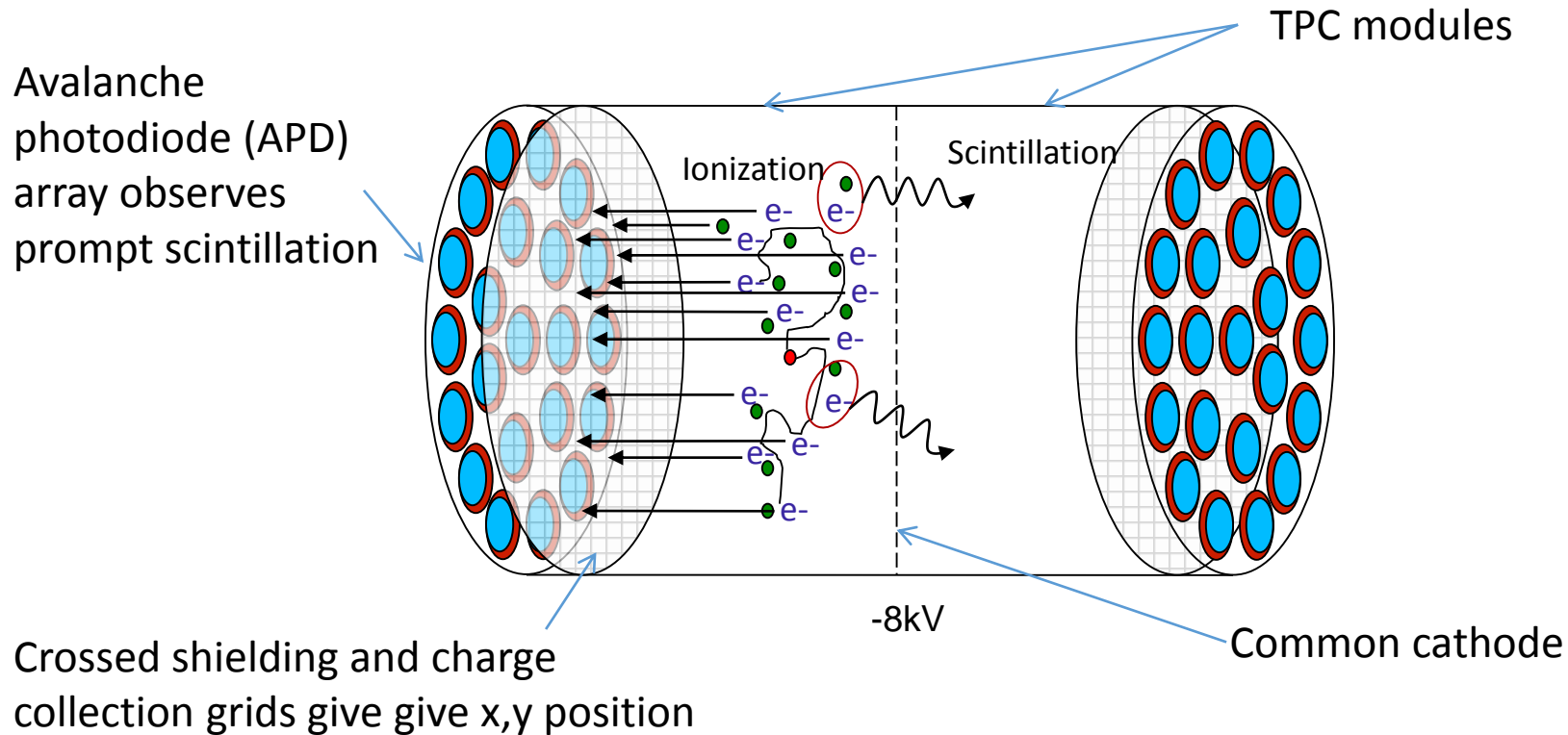
- Y. Lan M1-4
- R. Gornea T1-5

EXO-200

- Located at the Waste Isolation Pilot Plant at $32^{\circ}22'30''\text{N}$ $103^{\circ}47'34''\text{W}$ (Carlsbad, NM).
- 2150 feet depth ($\sim 655\text{m}$), ≈ 1585 mwe flat overburden
- U.S. DOE permanent repository for nuclear waste
- Low radioactivity levels:
 - U, Th $< 100\text{ppb}$
 - Radon background $< 10\text{ Bq/m}^3$

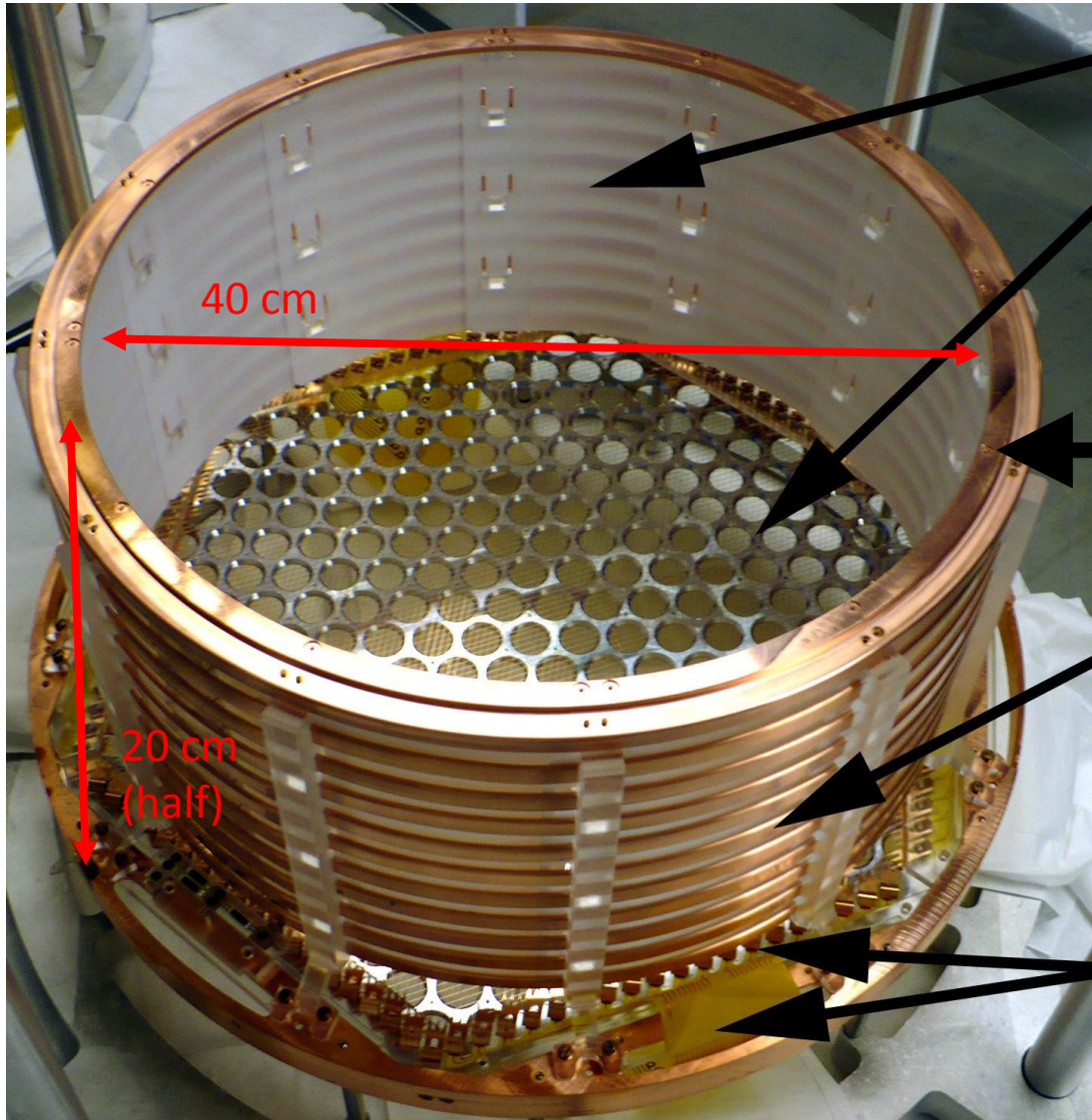


EXO-200 Time Projection Chamber (TPC) Basics

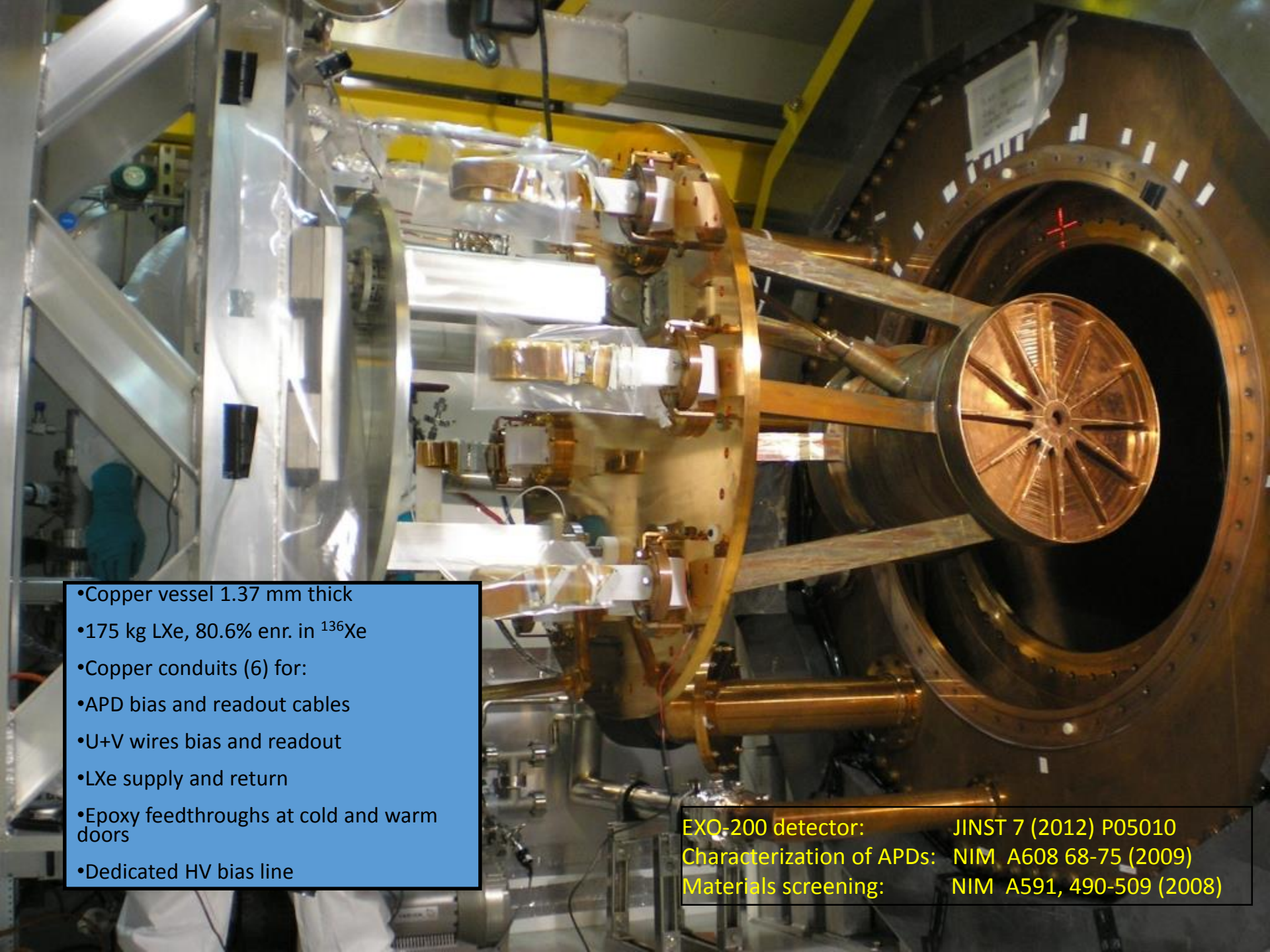


- Z-position from the time difference between scintillation and ionization
- Event energy from the combination of ionization and scintillation
- TPC allows rejection of some gamma backgrounds because Compton scattering results in multiple energy deposits

EXO-200 TPC



- Teflon Reflectors
(increase light collection)
- APD plane and wire planes
(wires are photo-etched)
- Central HV plane
(photo-etched phosphor bronze)
- Acrylic supports
and field shaping
rings
- Kapton flex cables
(spring connections
eliminate solder joints
and glue)

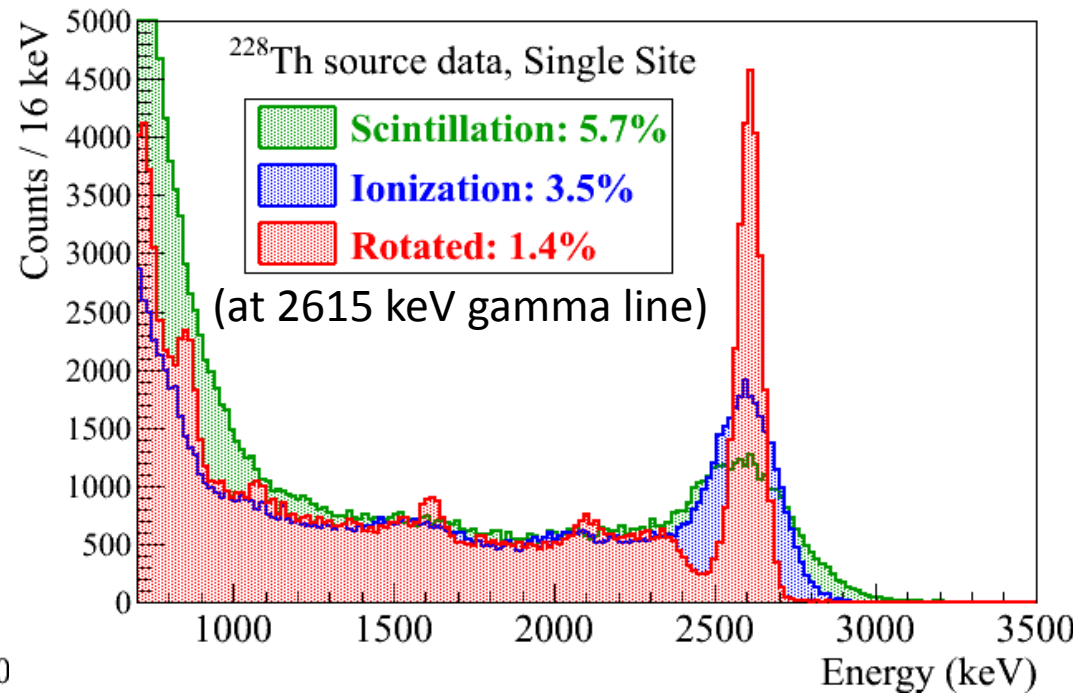
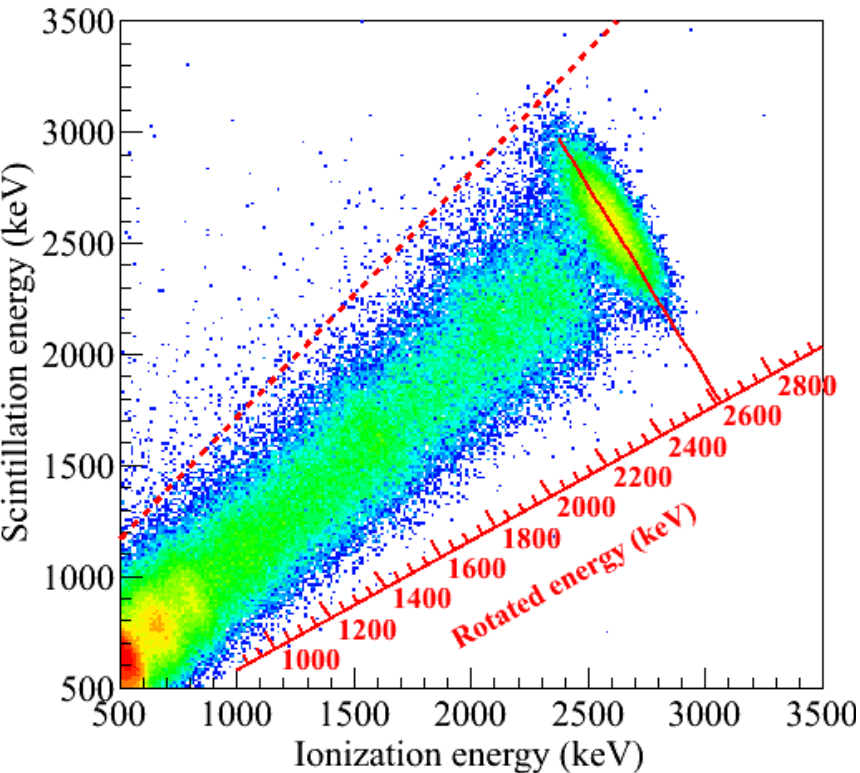


- Copper vessel 1.37 mm thick
- 175 kg LXe, 80.6% enr. in ^{136}Xe
- Copper conduits (6) for:
 - APD bias and readout cables
 - U+V wires bias and readout
 - LXe supply and return
- Epoxy feedthroughs at cold and warm doors
- Dedicated HV bias line

EXO-200 detector: JINST 7 (2012) P05010
Characterization of APDs: NIM A608 68-75 (2009)
Materials screening: NIM A591, 490-509 (2008)

Energy measurement

Combination of charge and light

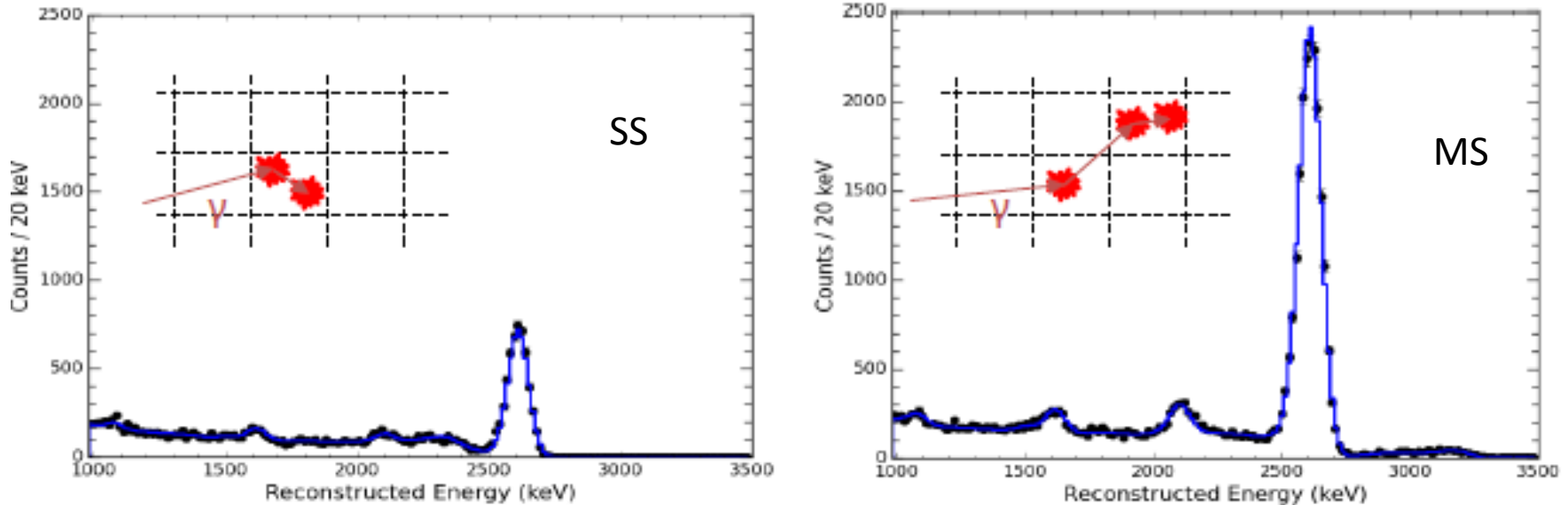


‘Rotation angle’ determined weekly using ^{228}Th source data, defined as angle which gives best ‘rotated’ resolution

Position/multiplicity reconstruction

Background measurement/reduction

^{228}Th calibration source in EXO-200 detector



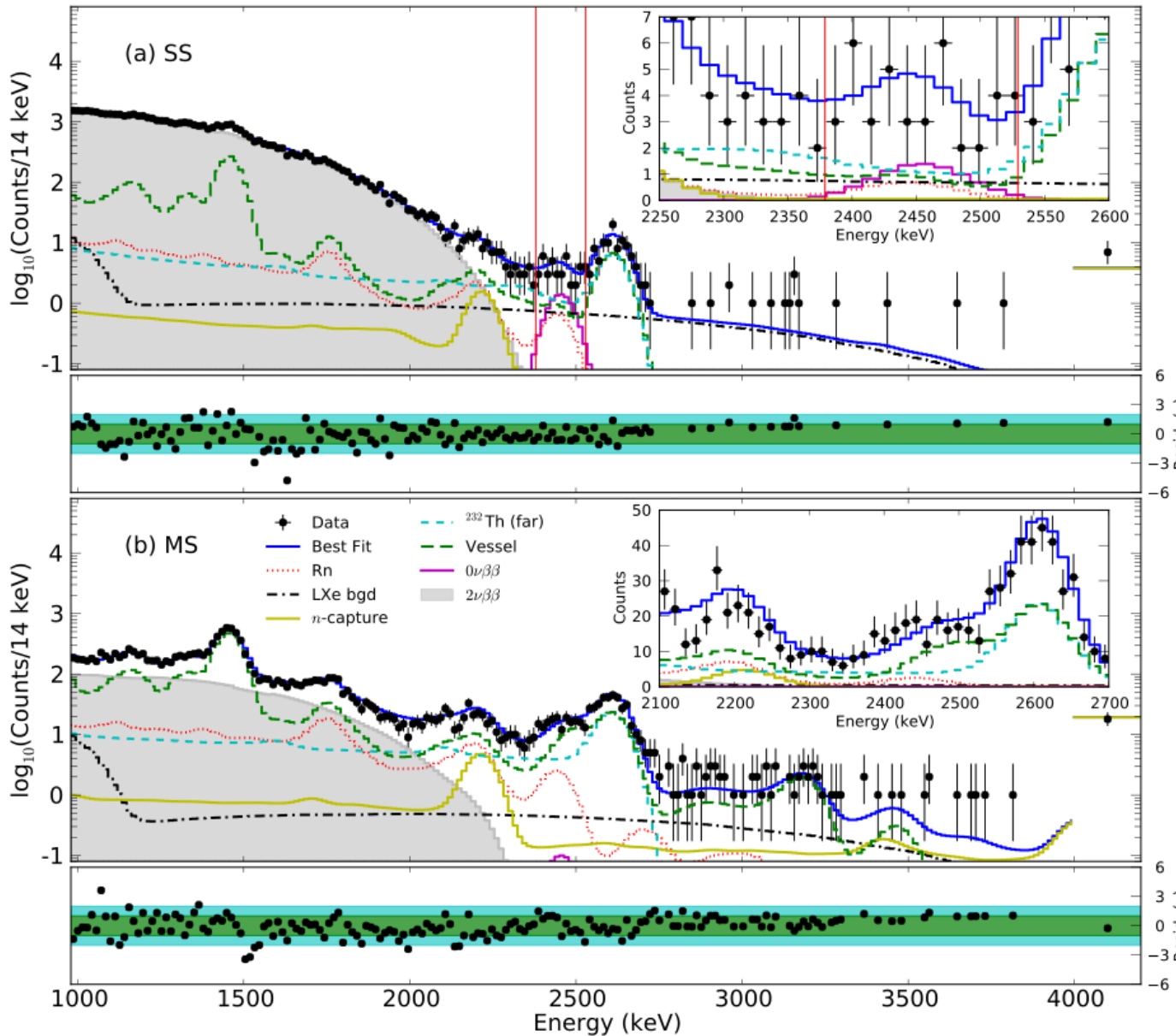
Events with > 1 charge cluster: multi-site (MS) events

Event with 1 charge cluster: single-site (SS) events

$0\nu\beta\beta$: $\sim 90\%$ SS

γ s: $\sim 30\%$ SS at $0\nu\beta\beta$ energy

Recent $0\nu\beta\beta$ decay result

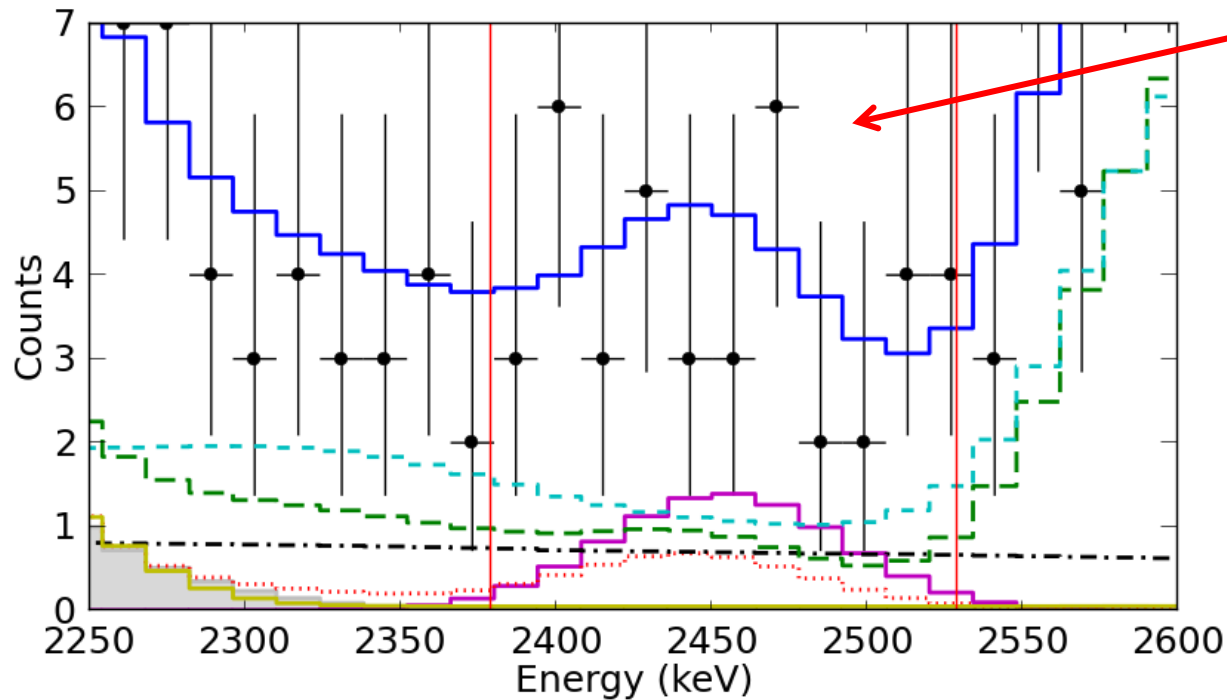


Run 2 data consists of:
 Run 2a already used for
 PRC 2014 and PRL 2012
 09/22/2011 – 04/15/2012
 Runs 2b and 2c
 04/16/2012 – 09/01/2013
 477.60 \pm 0.01 days of data

**¹³⁶Xe exposure:
 99.8 kg yr**

Simultaneous fit to
 energy and standoff
 distance for SS and MS

Recent $0\nu\beta\beta$ decay result



39 counts in
 $\pm 2\sigma$ ROI

Background fit in
 $\pm 2\sigma$ ROI

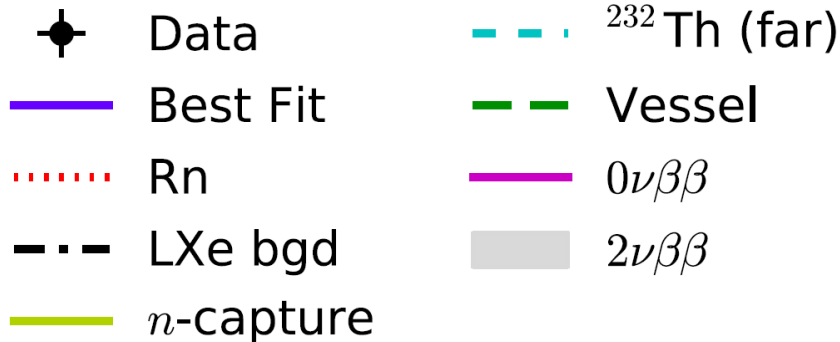
^{232}Th	16.0
^{238}U	8.1
^{137}Xe	7.0
Total	31.1 ± 3.8

From profile likelihood:

$$T_{1/2}^{0\nu\beta\beta} > 1.1 \cdot 10^{25} \text{ yr}$$

$$\langle m_{\beta\beta} \rangle < 190 - 450 \text{ meV}$$

(90% C.L.)

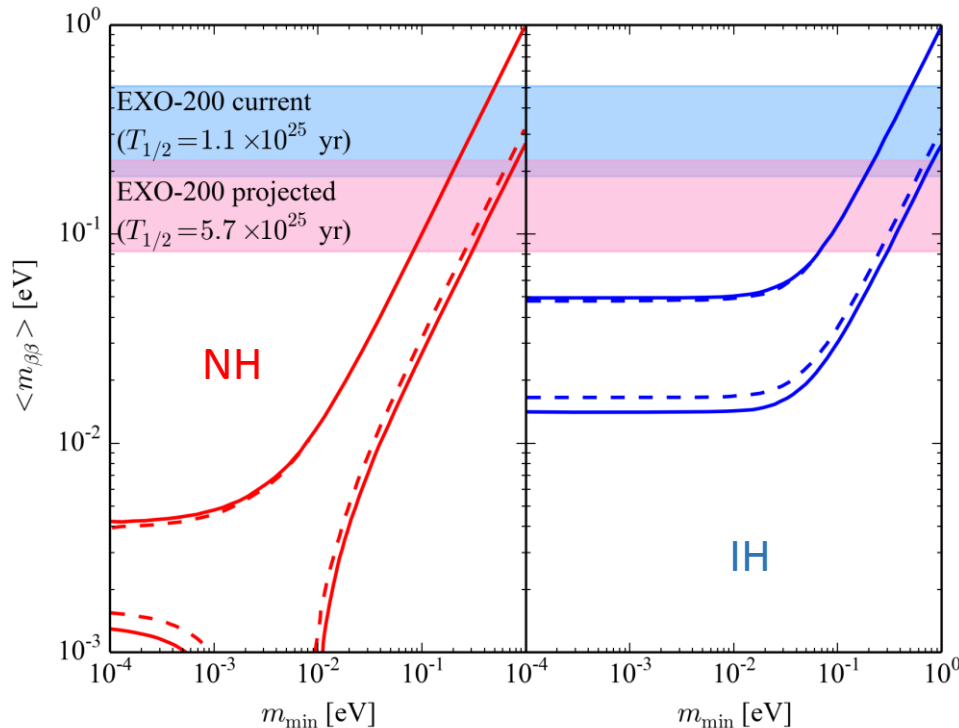


EXO-200 $(0\nu)\beta\beta$ search

- 2011 First measurement of $2\nu\beta\beta$ in ^{136}Xe [PRL 107, 212501 (2011)]
- 2012 First $0\nu\beta\beta$ result, best $m_{\beta\beta}$ limit [PRL 109, 032505 (2012)]
- 2013 Most precisely measured $2\nu\beta\beta$ rate — and the lowest
→ slowest process ever directly measured in nature! [PRC 89, 015502 (2014)]
- 2014 Improved sensitivity to $m_{\beta\beta}$ [Nature 510, 229 (2014)]

$$T_{1/2}^{2\nu\beta\beta} = 2.165 \pm 0.016(\text{stat}) \pm 0.059(\text{syst}) \times 10^{21} \text{ yr}$$

$$T_{1/2}^{0\nu\beta\beta} > 1.1 \times 10^{25} \text{ yr @ 90\% C.L.}$$

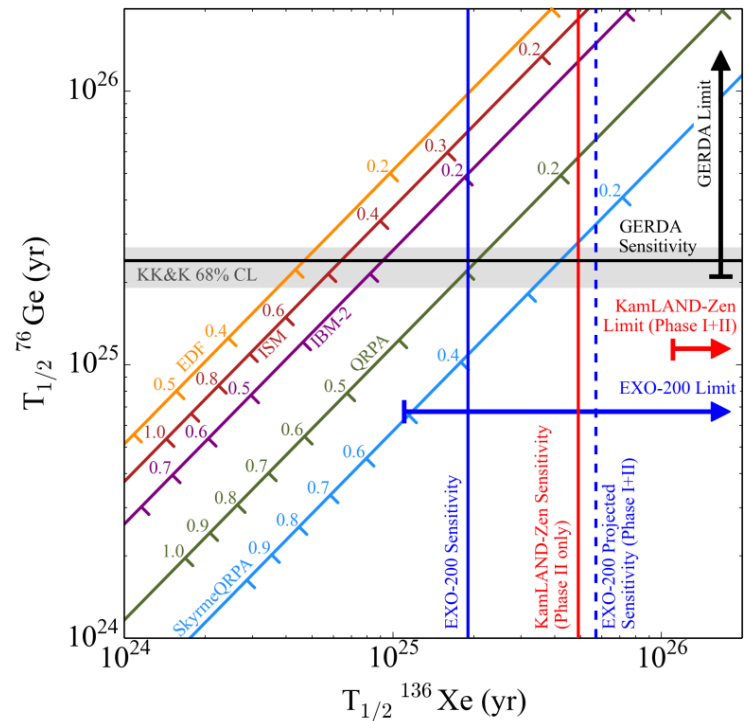
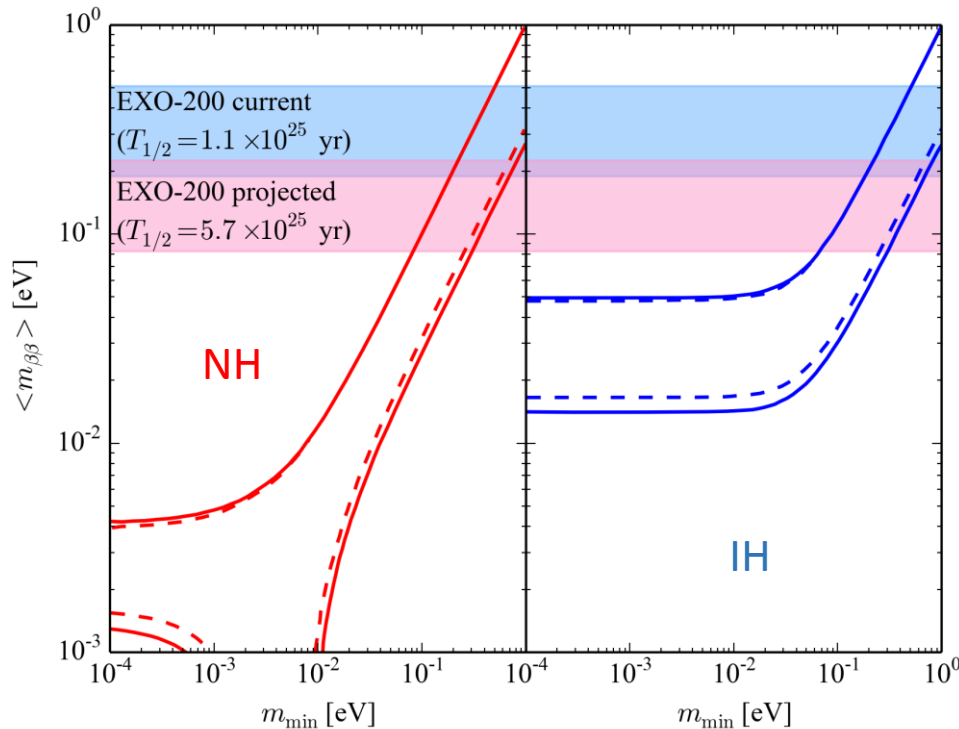


EXO-200 (0ν) $\beta\beta$ search

- 2011 First measurement of $2\nu\beta\beta$ in ^{136}Xe [PRL 107, 212501 (2011)]
- 2012 First $0\nu\beta\beta$ result, best $m_{\beta\beta}$ limit [PRL 109, 032505 (2012)]
- 2013 Most precisely measured $2\nu\beta\beta$ rate — and the lowest
 \rightarrow slowest process ever directly measured in nature! [PRC 89, 015502 (2014)]
- 2014 Improved sensitivity to $m_{\beta\beta}$ [Nature 510, 229 (2014)]

$$T_{1/2}^{2\nu\beta\beta} = 2.165 \pm 0.016(\text{stat}) \pm 0.059(\text{syst}) \times 10^{21} \text{ yr}$$

$$T_{1/2}^{0\nu\beta\beta} > 1.1 \times 10^{25} \text{ yr @ 90\% C.L.}$$

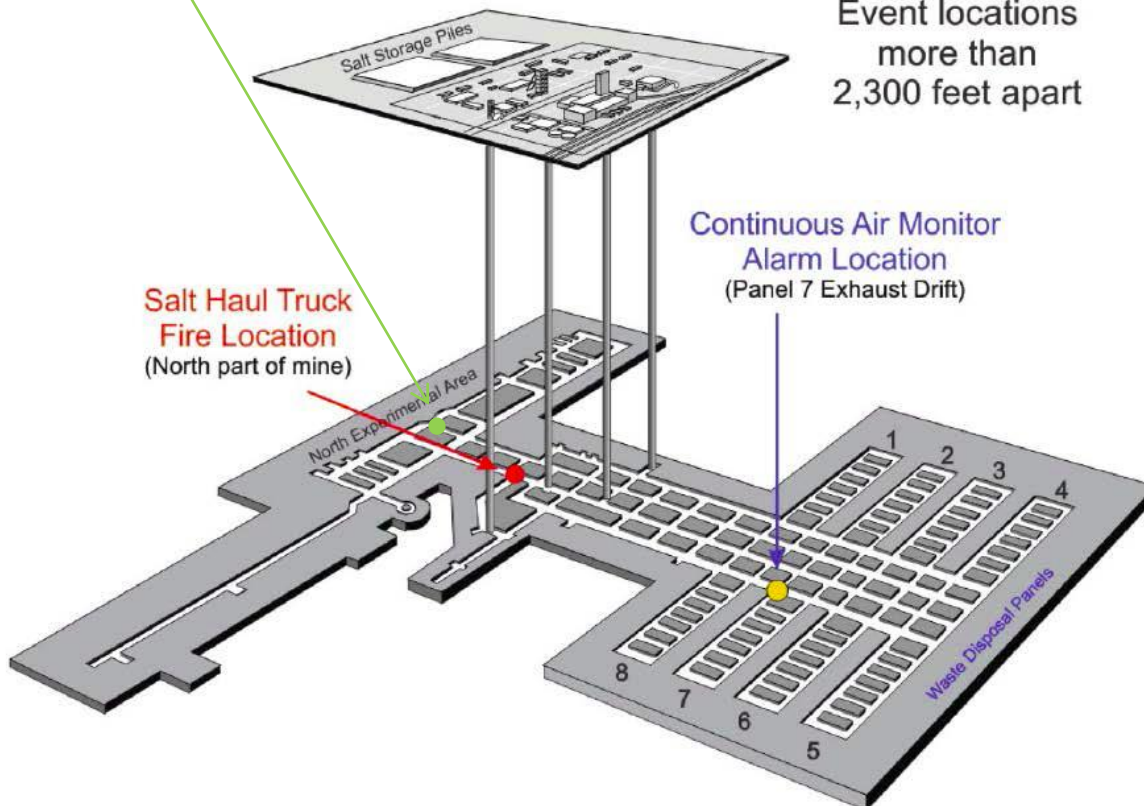


The future of EXO-200

EXO-200 is about 1.2 km from the radiation event



Event locations more than 2,300 feet apart

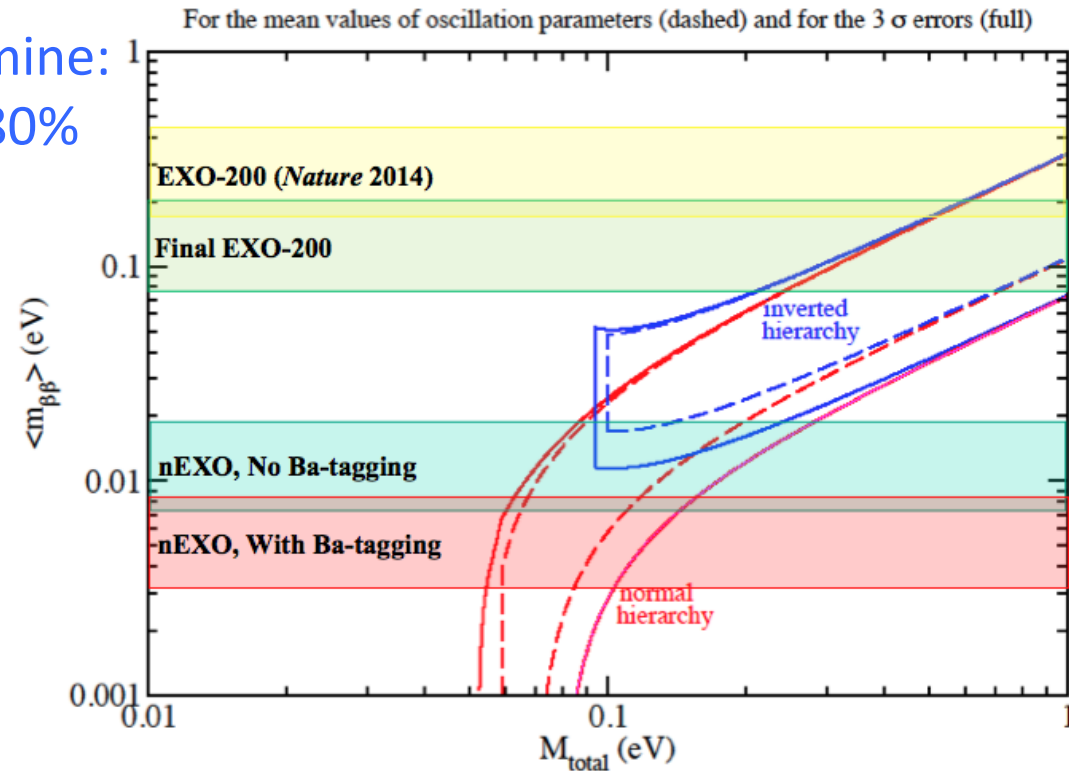


- **Feb. 5 2014:** Fire in WIPP underground
- **Feb. 14, 2014:** Radiation release event
- So far no radioactivity has been measured at EXO-200
- EXO clean up finished
- Low background data taking resumed in early 2016
- Stay tuned for new results

$0\nu\beta\beta$ search with EXO

Multi-phase program :

- **EXO-200** – operational at WIPP mine:
 - ~175kg xenon enriched at ~80%
 - Current limit on $0\nu\beta\beta$:
 1.1×10^{25} years (EXO-200)
 - Continue data taking for 2 more years
 - Sensitivity: 100-200 meV
- **nEXO** - R&D underway:
 - 5T xenon enriched at ~90%
 - Sensitivity: 5-30 meV
 - Improved techniques for background suppression and possibly Ba tagging



→ **Development of nEXO is well advanced**

$0\nu\beta\beta$ search with EXO

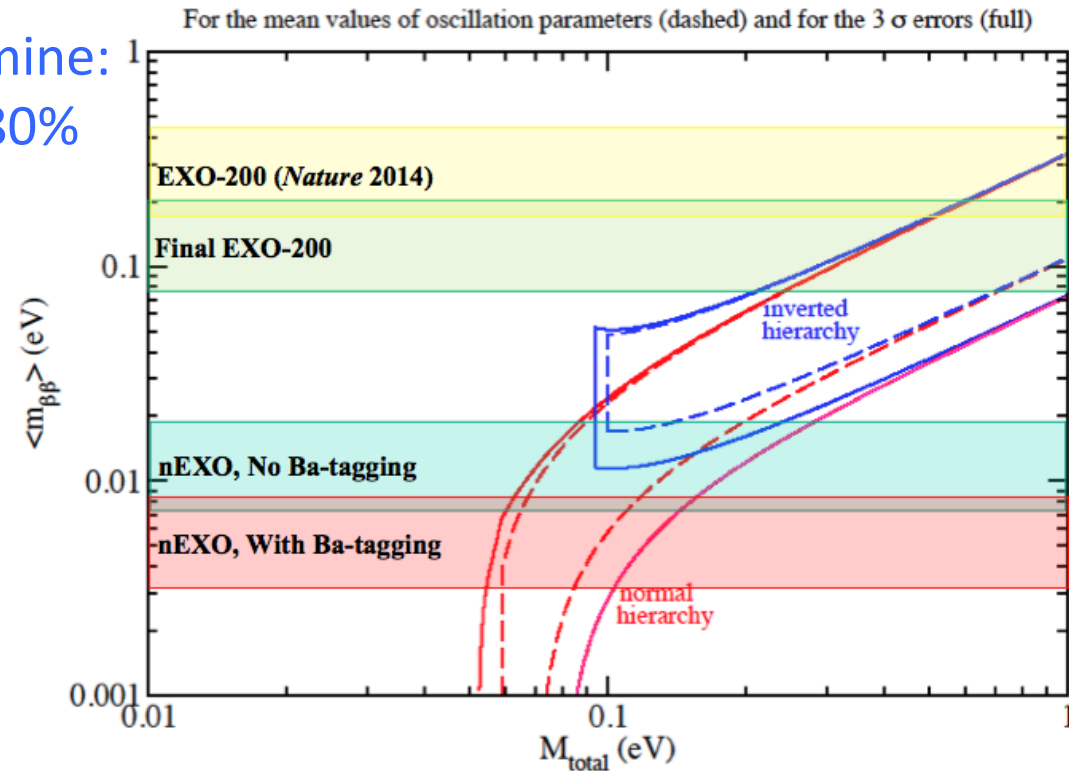
Multi-phase program :

- **EXO-200** – operational at WIPP mine:

- ~175kg xenon enriched at ~80%
- Current limit on $0\nu\beta\beta$:
 1.1×10^{25} years (EXO-200)
- Continue data taking for 2 more years
- Sensitivity: 100-200 meV

- **nEXO** - R&D underway:

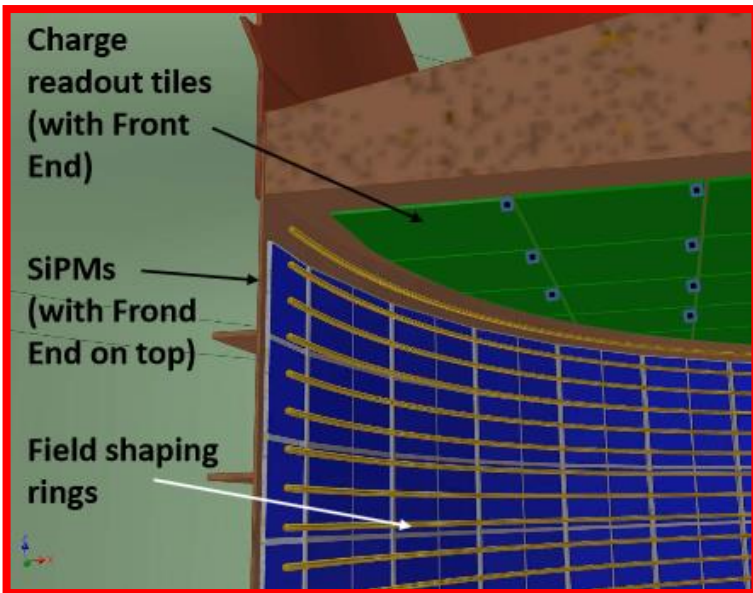
- 5T xenon enriched at ~90%
- Sensitivity: 5-30 meV
- Improved techniques for background suppression and possibly Ba tagging



For more information on Ba tagging:

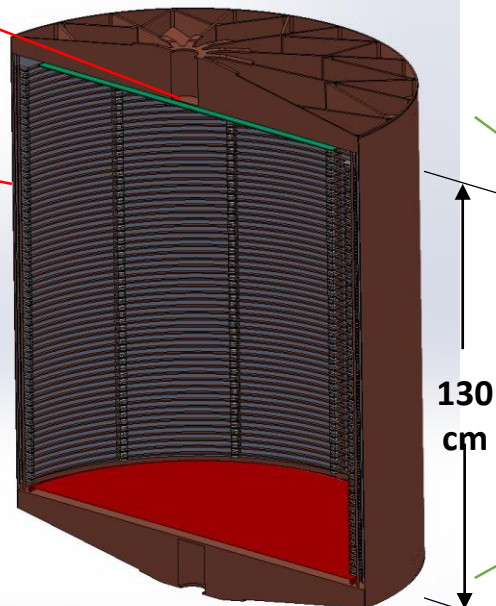
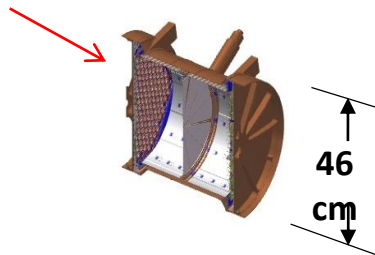
- Y. Lan M1-4
- R. Gornea T1-5

Searching for $0\nu\beta\beta$ with nEXO

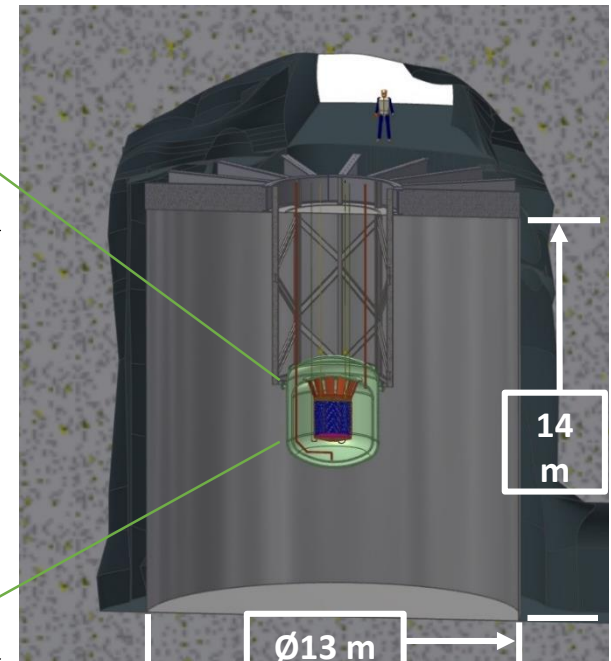


- Next-generation neutrinoless double beta decay detector
- 5 t liquid xenon TPC similar to EXO-200 (50x the size)
- Possible location in SNOLab Cryo Pit (6010 mwe)
- SiPM for light detection
- Tiles for charge read out
- 3D event reconstruction
- Expected σ/E of 1% at Q-value
- Possible addition of Ba-tagging after 5 years

EXO-200 for size comparison



nEXO TPC



nEXO at the SNOLab Cryopit

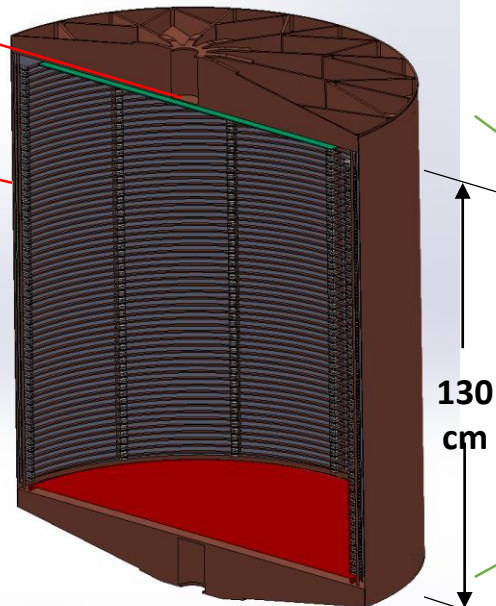
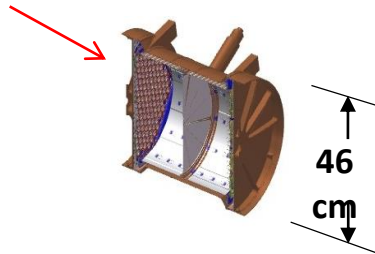
Searching for $0\nu\beta\beta$ with nEXO

4m² of VUV sensitive SiPMs

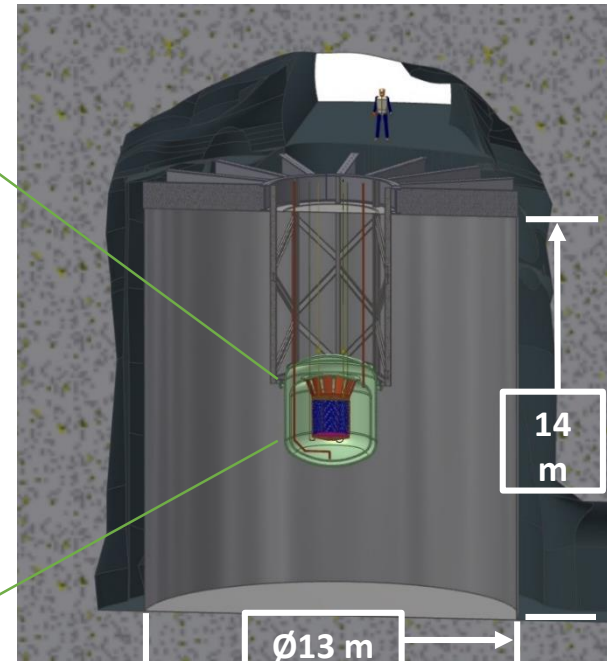


- Next-generation neutrinoless double beta decay detector
- 5 t liquid xenon TPC similar to EXO-200 (50x the size)
- Possible location in SNOLab Cryo Pit (6010 mwe)
- SiPM for light detection
- Tiles for charge read out
- 3D event reconstruction
- Expected σ/E of 1% at Q-value
- Possible addition of Ba-tagging after 5 years

EXO-200 for size comparison



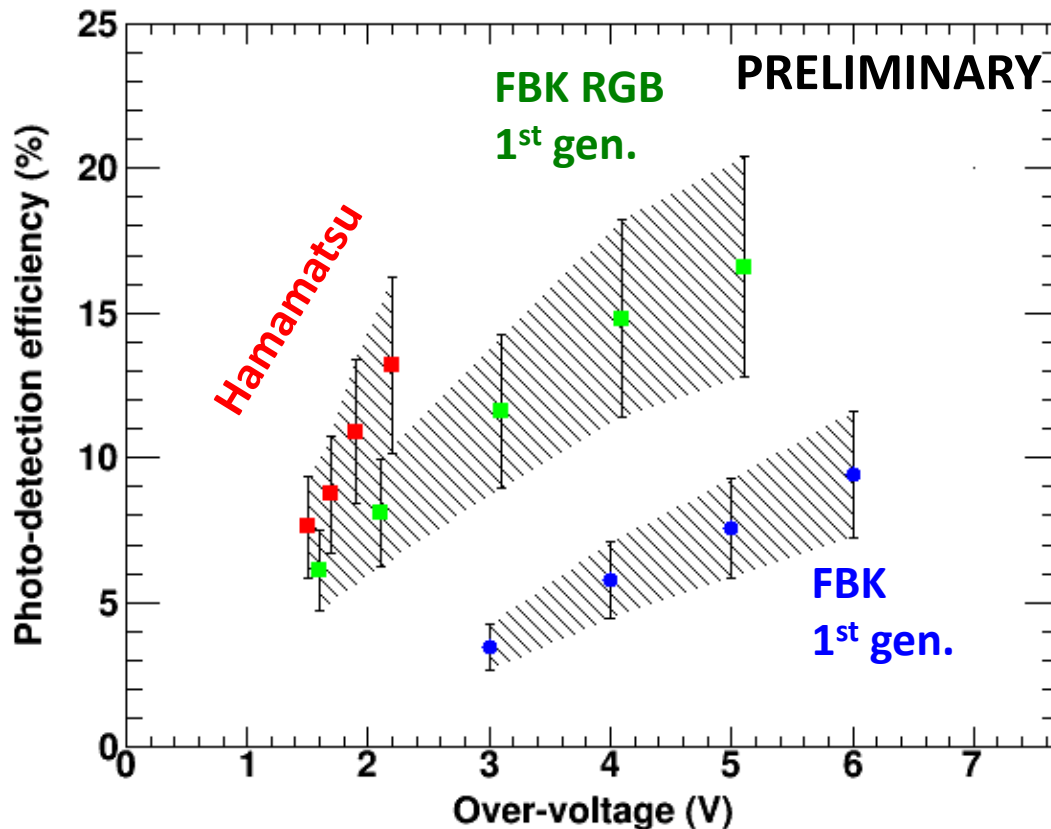
nEXO TPC



nEXO at the SNOLab Cryopit

SiPM Photodetector

- Hamamatsu produces devices with QE= ~12% @ 175nm but encapsulation is too radioactive → trying to procure un-encapsulated devices
- First nEXO-specific run at FBK (Italy) provided ~10% QE [I.Ostrovskiy et al. IEEE TNS 62 (2015) 1825.]
- New FBK “RGB” devices reach 15% QE with 7.7x7.7mm².



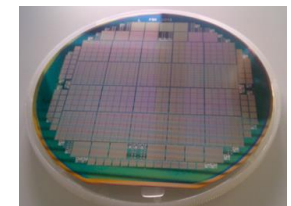
June 13, 2016

Thomas Brunner

- Working closely with manufacturers to develop SiPMs to reach >15% QE at 175nm
- Radioassay of SiPMs to determine radioactivity
- Development of integration of 1x1cm² SiPMs into 10x10cm² tiles
- Tests in liquid Xe planned



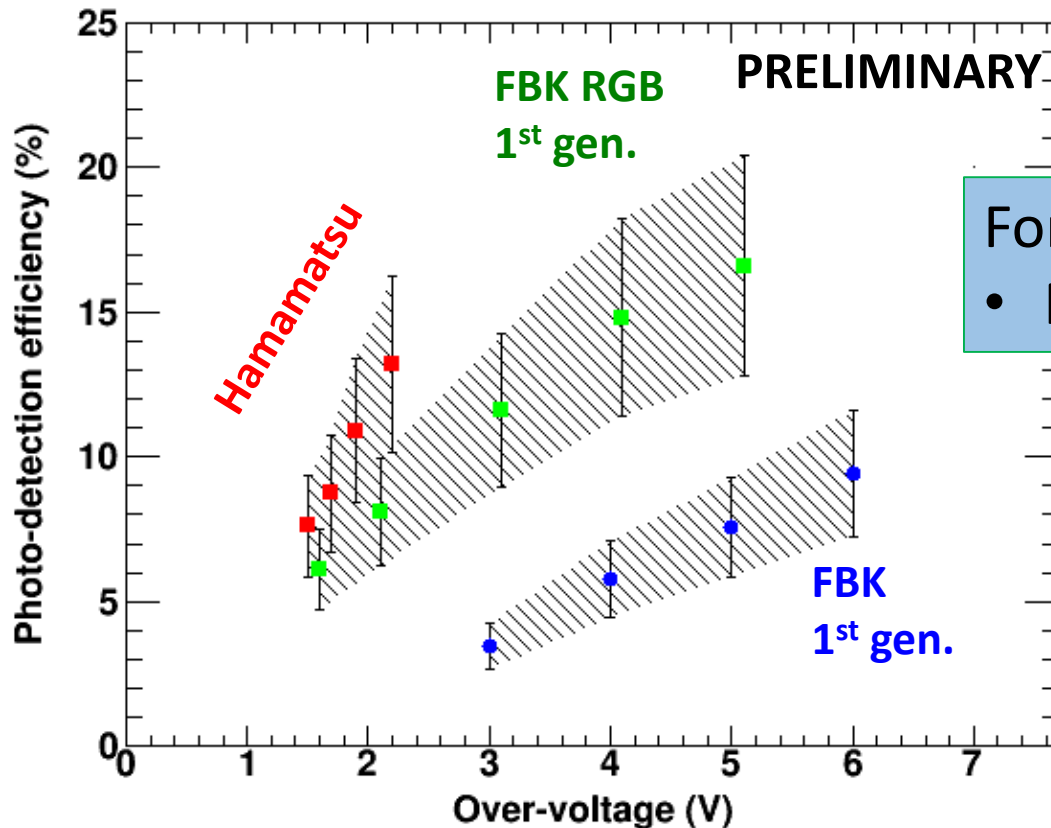
Hamamatsu MEG MPPC



FBK SiPM

SiPM Photodetector

- Hamamatsu produces devices with QE= ~12% @ 175nm but encapsulation is too radioactive → trying to procure un-encapsulated devices
- First nEXO-specific run at FBK (Italy) provided ~10% QE [I.Ostrovskiy et al. IEEE TNS 62 (2015) 1825.]
- New FBK “RGB” devices reach 15% QE with 7.7x7.7mm².



June 13, 2016

Thomas Brunner

- Working closely with manufacturers to develop SiPMs to reach >15% QE at

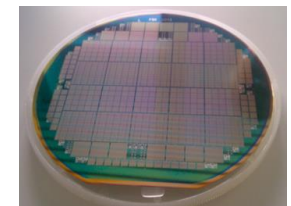
For more information on SiPM:

- F. Retiere W2-4

- Development of integration of 1x1cm² SiPMs into 10x10cm² tiles
- Tests in liquid Xe planned



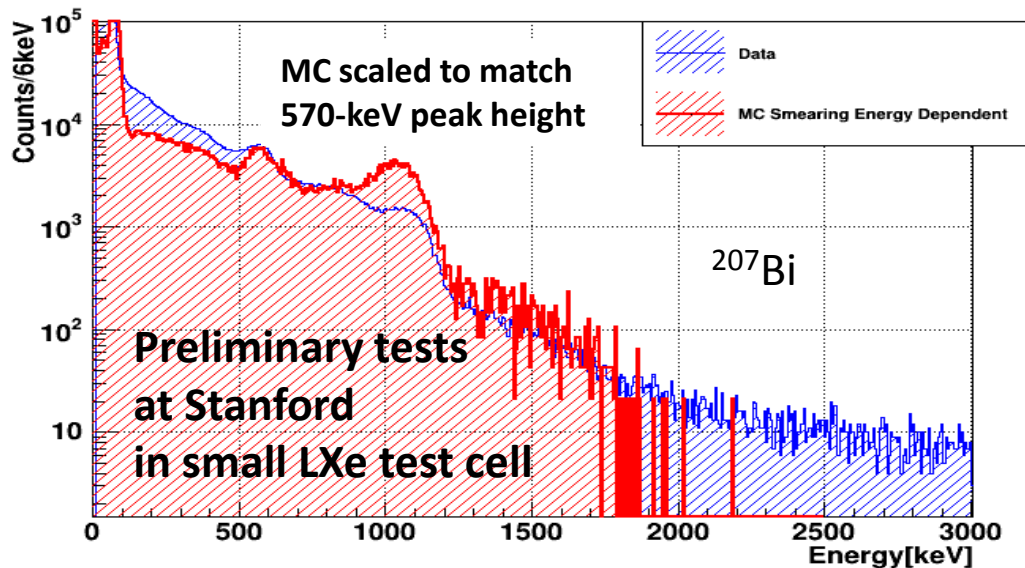
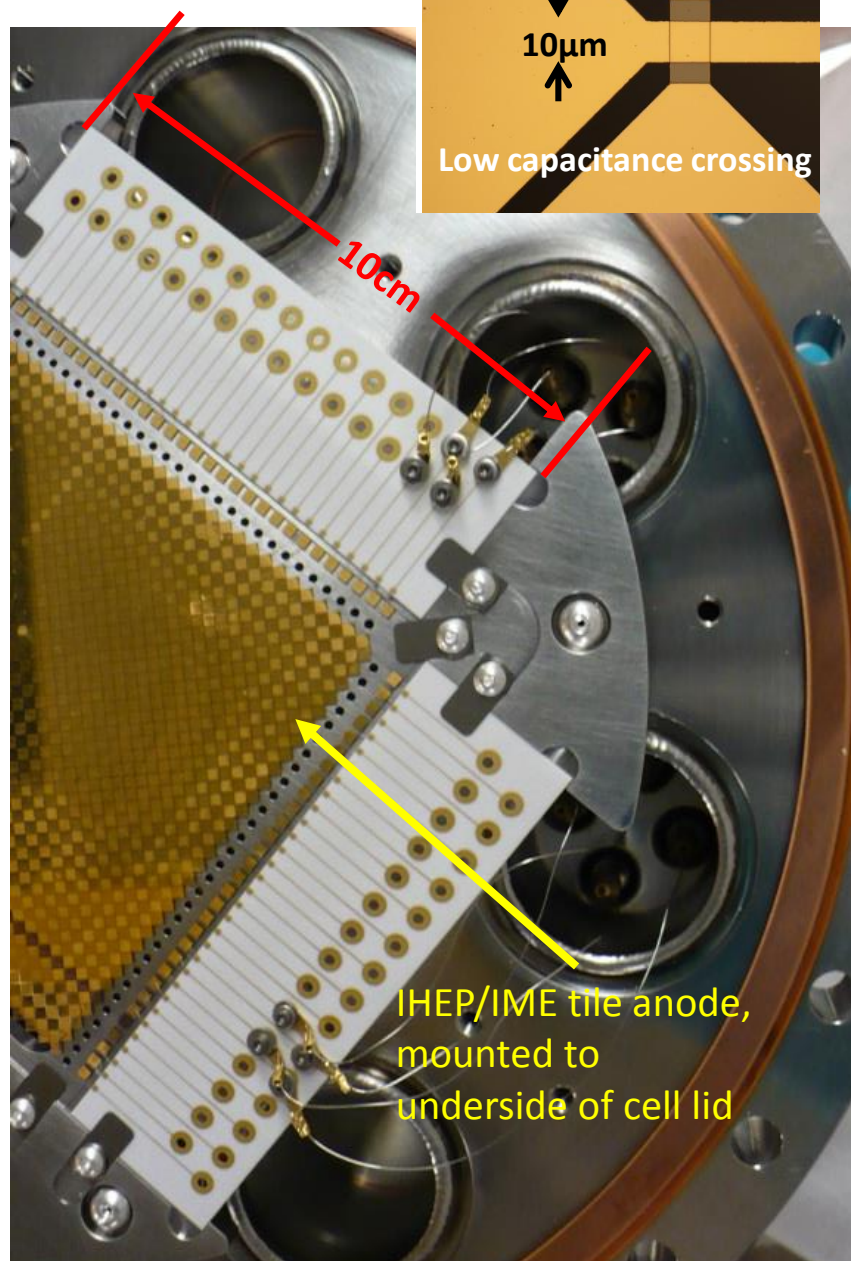
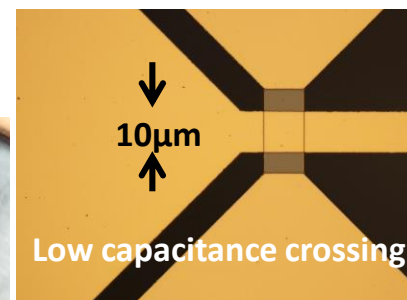
Hamamatsu MEG MPPC



FBK SiPM

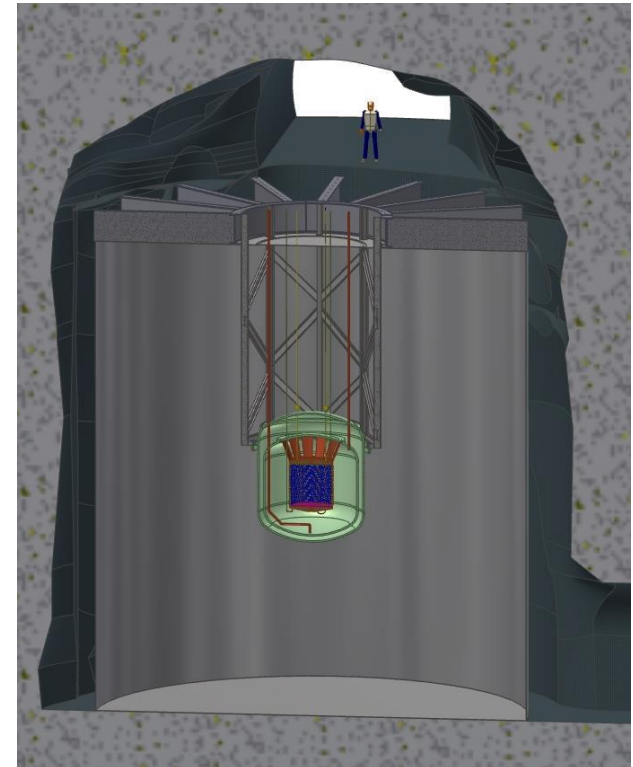
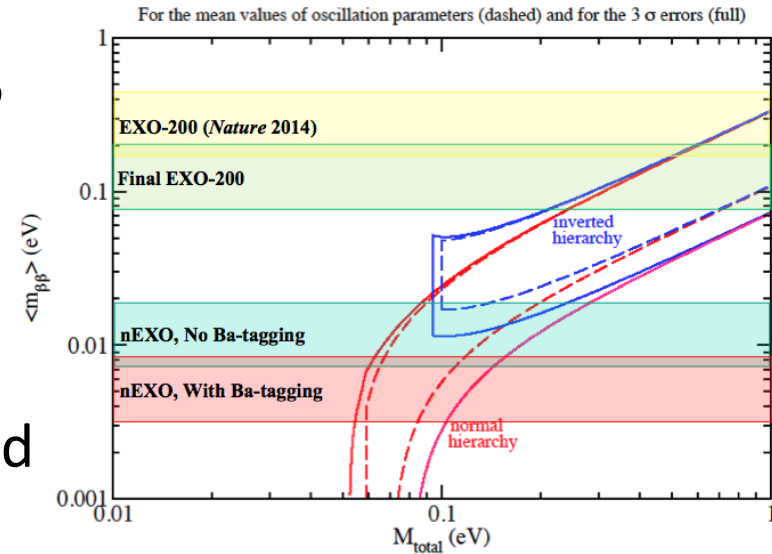
Charge Readout Tiles

- EXO-200 used wires for charge-readout
- Produced by IHEP/IME; functional testing in LXe in the US.
- 10 x 10cm² Prototype Tile
- Metallized strips on fused silica substrate
- 60 orthogonal channels (30 x 30)
- 3mm strip pitch
- Strip intersections isolated with SiO₂ layer
- Currently testing in LXe with a ²⁰⁷Bi source



Summary & Plans

- EXO-200 is operational and taking low background data
- nEXO is the next generation $0\nu\beta\beta$ experiment with 5 T isotopically enriched LXe
- nEXO expands on the success of EXO-200 and improves performance via R&D efforts
- nEXO will have many handles on background
- nEXO has discovery potential in Inverted Hierarchy pushing the lower bound of $\langle m_{\beta\beta} \rangle$
- The 10meV region is within reach
- Strong Canadian contribution to EXO-200 and nEXO





University of Alabama, Tuscaloosa AL, USA — T Didberidze, M Hughes, A Piepke, R Tsang

University of Bern, Switzerland — J-L Vuilleumier

Brookhaven National Laboratory, Upton NY, USA — M Chiu, G De Geronimo, S Li, V Radeka, T Rao, G Smith, T Tsang, B Yu

California Institute of Technology, Pasadena CA, USA — P Vogel

Carleton University, Ottawa ON, Canada — I Badhrees, Y Baribeau, M Bowcock, M Dunford, M Facina,

R Gornea, K Graham, P Gravelle, R Killick, T Koffas, C Licciardi, K McFarlane, R Schnarr, D Sinclair

Colorado State University, Fort Collins CO, USA — C Chambers, A Craycraft, W Fairbank Jr, T Walton

Drexel University, Philadelphia PA, USA — E Callaghan, MJ Dolinski, YH Lin, E Smith, Y-R Yen

Duke University, Durham NC, USA — PS Barbeau, G Swift

University of Erlangen-Nuremberg, Erlangen, Germany — G Anton, R Bayerlein, J Hoessl, P Hufschmidt, A Jamil, T Michel, T Ziegler

IBS Center for Underground Physics, Daejeon, South Korea — DS Leonard

IHEP Beijing, People's Republic of China — G Cao, W Cen, X Jiang, H Li, Z Ning, X Sun, T Tolba, W Wei, L Wen, W Wu, J Zhao

ITEP Moscow, Russia — V Belov, A Burenkov, A Karelin, A Kobayakin, A Kuchenkov, V Stekhanov, O Zeldovich

University of Illinois, Urbana-Champaign IL, USA — D Beck, M Coon, S Li, L Yang

Indiana University, Bloomington IN, USA — JB Albert, S Daugherty, TN Johnson, LJ Kaufman, G Visser, J Zettlemoyer

University of California, Irvine, Irvine CA, USA — M Moe

Laurentian University, Sudbury ON, Canada — B Cleveland, A Der Mesrobian-Kabakian, J Farine, U Wichoski

Lawrence Livermore National Laboratory, Livermore CA, USA — O Alford, J Brodsky,

M Heffner, G Holtmeier, A House, M Johnson, S Sangiorgio

University of Massachusetts, Amherst MA, USA — S Feyzbakhsh, S Johnston, M Negus, A Pocar

McGill University, Montreal QC, Canada — T Brunner, K Murray

Oak Ridge National Laboratory, Oak Ridge TN, USA — L Fabris, D Hornback, RJ Newby, K Zioc

Pacific Northwest National Laboratory, Richland, WA, USA — EW Hoppe, JL Orrell

Rensselaer Polytechnic Institute, Troy NY, USA — E Brown, K Odgers

Université de Sherbrooke — S Charlebois, D Danovitch, R Fontaine, JF Pratte, J Sylvestre

SLAC National Accelerator Laboratory, Menlo Park CA, USA — J Dalmasson, T Daniels, S Delaquis,

G Haller, R Herbst, M Kwiatkowski, A Odian, M Oriunno, B Mong, PC Rowson, K Skarpaas

University of South Dakota, Vermillion SD, USA — J Daughhettee, R MacLellan

Stanford University, Stanford CA, USA — R DeVoe, D Fudenberg, G Gratta, M Jewell, S Kravitz,

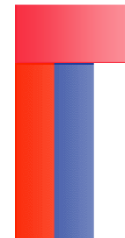
D Moore, I Ostrovskiy, A Schubert, M Weber

Stony Brook University, SUNY, Stony Brook, NY, USA — K Kumar, O Njoya, M Tarka

Technical University of Munich, Garching, Germany — P Fierlinger, M Marino

TRIUMF, Vancouver BC, Canada — J Dilling, P Gumplinger, R Krücken, Y. Lan, F Retière, V Strickland

The nEXO Collaboration





University of Alabama, Tuscaloosa AL, USA — T Didberidze, M Hughes, A Piepke, R Tsang

University of Bern, Switzerland — J-L Vuilleumier

Brookhaven National Laboratory, Upton NY, USA — M Chiu, G De Geronimo, S Li, V Radeka, T Rao, G Smith, T Tsang, B Yu

California Institute of Technology, Pasadena CA, USA — P Vogel

Carleton University, Ottawa

Colorado State University,

Drexel University, Philadel

Duke University, Durham I

University of Erlangen-Nur

IBS Center for Undergroun

IHEP Beijing, People's Rep

IITP Moscow, Russia — V

University of Illinois, Urban

Indiana University, Bloomi

University of California, Irv

Laurentian University, Sud

Lawrence Livermore Nation

University of Massachusett

McGill University, Montrea

Oak Ridge National Labora

Pacific Northwest National

Rensselaer Polytechnic Inst

Université de Sherbrooke —

SLAC National Accelerator

University of South Dakota

Stanford University, Stanfo

Stony Brook University, SUNY, Stony Brook, NY, USA — K Kumar, O Njoya, M Tarka

Technical University of Munich, Garching, Germany — P Fierlinger, M Marino

TRIUMF, Vancouver BC, Canada — J Dilling, P Gumplinger, R Krücken, Y. Lan, F Retière, V Strickland

The nEXO Collaboration

Thanks to my Canadian collaborators:

Carleton University



Laurentian University



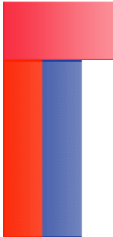
McGill University



Université de Sherbrooke



TRIUMF





University of Alabama, Tuscaloosa AL, USA — T Didberidze, M Hughes, A Piepke, R Tsang

University of Bern, Switzerland — J-L Vuilleumier

Brookhaven National Laboratory, Upton NY, USA — M Chiu, G De Geronimo, S Li, V Radeka, T Rao, G Smith, T Tsang, B V

California Institute of Technology, Pasadena CA, USA — P Vogel

Carleton University, Ottawa ON, Canada — I Badhrees, Y Baribeau, M Bowcock, M Dunford, M Facina,
R Gornea, K Graham, P Gravelle, R Killick, T Koffas, C Licciardi, K McFarlane

Colorado State University, Fort Collins CO, USA — C Chambers, A Craycraft, W Fairbank

Drexel University, Philadelphia PA, USA — E Callaghan, MJ Dolinski, YH Lin, E S

Duke University, Durham NC, USA — PS Barbeau, G Swift

University of Erlangen-Nuremberg, Erlangen, Germany — G Anton, T B

IBS Center for Underground Physics, Daejeon, South Korea

IHEP Beijing, People's Republic of China — G Cao, W

ITEP Moscow, Russia — V Belov, A Burenkov, v Stekhanov, O Zeldovich

University of Illinois, Urbana-Champaign, IL, USA — L Yang

Indiana University, Bloomington IN, USA — TN Johnson, LJ Kaufman, G Visser, J Zettlemoyer

University of California

Laurentian University, Sudbury ON, Canada — Cleveland, A Der Mesrobian-Kabakian, J Farine, U Wichoski

Lawrence Livermore National Laboratory, Livermore CA, USA — O Alford, J Brodsky,

M Heffner, G Holtmeier, A House, M Johnson, S Sangiorgio

University of Massachusetts Lowell, Lowell MA, USA — S Feyzbakhsh, S Johnston, M Negus, A Pocar

McGill University, Montreal QC, Canada — T Brunner, K Murray

Oak Ridge National Laboratory, Oak Ridge TN, USA — L Fabris, D Hornback, RJ Newby, K Ziock

Pacific Northwest National Laboratory, Richland, WA, USA — EW Hoppe, JL Orrell

Rensselaer Polytechnic Institute, Troy NY, USA — E Brown, K Odgers

Université de Sherbrooke — S Charlebois, D Danovitch, R Fontaine, JF Pratte, J Sylvestre

SLAC National Accelerator Laboratory, Menlo Park CA, USA — J Dalmasson, T Daniels, S Delaquis,

G Haller, R Herbst, M Kwiatkowski, A Odian, M Oriunno, B Mong, PC Rowson, K Skarpaas

University of South Dakota, Vermillion SD, USA — J Daughhete, R MacLellan

Stanford University, Stanford CA, USA — R DeVoe, D Fudenberg, G Gratta, M Jewell, S Kravitz,

D Moore, I Ostrovskiy, A Schubert, M Weber

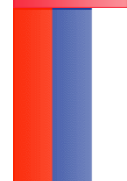
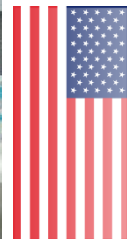
Stony Brook University, SUNY, Stony Brook, NY, USA — K Kumar, O Njoya, M Tarka

Technical University of Munich, Garching, Germany — P Fierlinger, M Marino

TRIUMF, Vancouver BC, Canada — J Dilling, P Gumplinger, R Krücken, F Retière, V Strickland

Thank you for your attention

The nEXO Collaboration



Backup

nEXO - Homogeneity is Crucial

- Increased mass
- Taking full advantage of self shielding
 - more effective
 - Improve Compton tag efficiency by double-hit recognition

LXe mass (kg)	Diameter or length (cm)
5000	130
150	40
5	13



→ **Benefits of monolithic detector compared to segmented detectors**

Physics searches with EXO-200

An Optimal Energy Estimator to Reduce Correlated Noise for the EXO-200 Light Readout

C.G. Davis et al. Submitted to JINST (May 2016). [arxiv:1605.06552](https://arxiv.org/abs/1605.06552) [[physics.ins-det](https://arxiv.org/archive/physics)]

Cosmogenic Backgrounds to $0\nu\beta\beta$ in EXO-200

J.B. Albert et al., J. Cosmol. & Astropart. Phys. (JCAP) 2016 4 (2016) 029

First Search for Lorentz and CPT Violation in Double Beta Decay with EXO-200

J.B. Albert et al., Phys. Rev. D 93, 072001

Search for $2\nu\beta\beta$ decay of ^{136}Xe to the 0_1^+ excited state of ^{136}Ba with EXO-200

J.B. Albert et al., Phys. Rev. C 93, 035501

Measurements of the ion fraction and mobility of alpha and beta decay products in liquid xenon using EXO-200

J.B. Albert et al., Phys. Rev. C 92, 045504 (2015).

Investigation of radioactivity-induced backgrounds in EXO-200

J.B. Albert et al. Phys. Rev. C 92, 015503 (2015).

Search for Majoron-emitting modes of double-beta decay of ^{136}Xe with EXO-200

J.B. Albert, et al. Phys. Rev. D 90, 092004 (2014).

Search for Majorana neutrinos with the first two years of EXO-200 data

J.B. Albert, et al. Nature 510 (2014) 229-234

An improved measurement of the $2\nu\beta\beta$ half-life of Xe-136 with EXO-200

J.B. Albert, et al. Phys. Rev. C 89, 015502 (2014)

Search for Neutrinoless Double-Beta Decay in ^{136}Xe with EXO-200

M. Auger, et al. Phys. Rev. Lett. 109, 032505 (2012)

Observation of Two-Neutrino Double-Beta Decay in Xe-136 with EXO-200

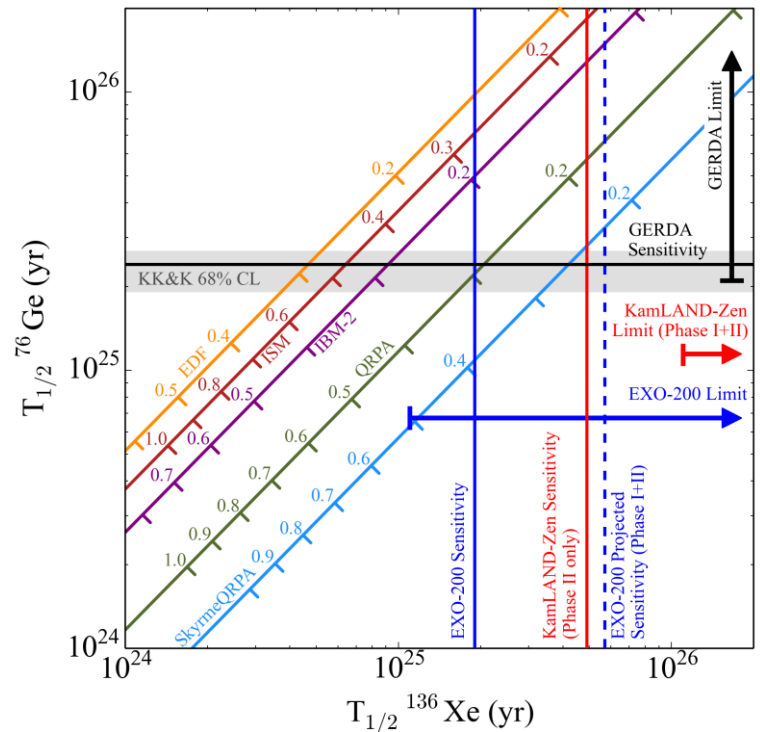
N. Ackerman, et al. Phys. Rev. Lett. 107, 212501 (2011)

EXO-200 $0\nu\beta\beta$ search

- 2011 First measurement of $2\nu\beta\beta$ in ^{136}Xe [PRL 107, 212501 (2011)]
- 2012 First $0\nu\beta\beta$ result, best $m_{\beta\beta}$ limit [PRL 109, 032505 (2012)]
- 2013 Most precisely measured $2\nu\beta\beta$ rate — and the lowest
→ slowest process ever directly measured in nature! [PRC 89, 015502 (2014)]
- 2014 Improved sensitivity to $m_{\beta\beta}$ [Nature 510, 229 (2014)]

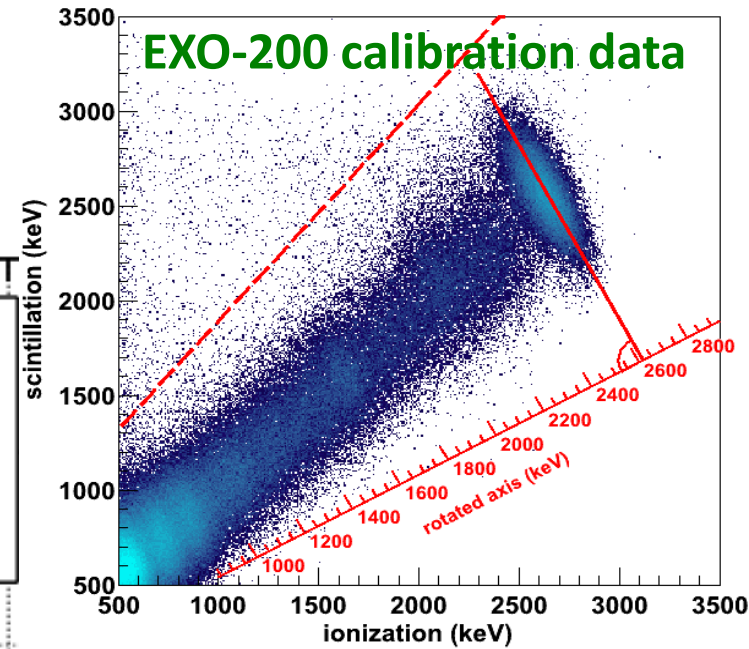
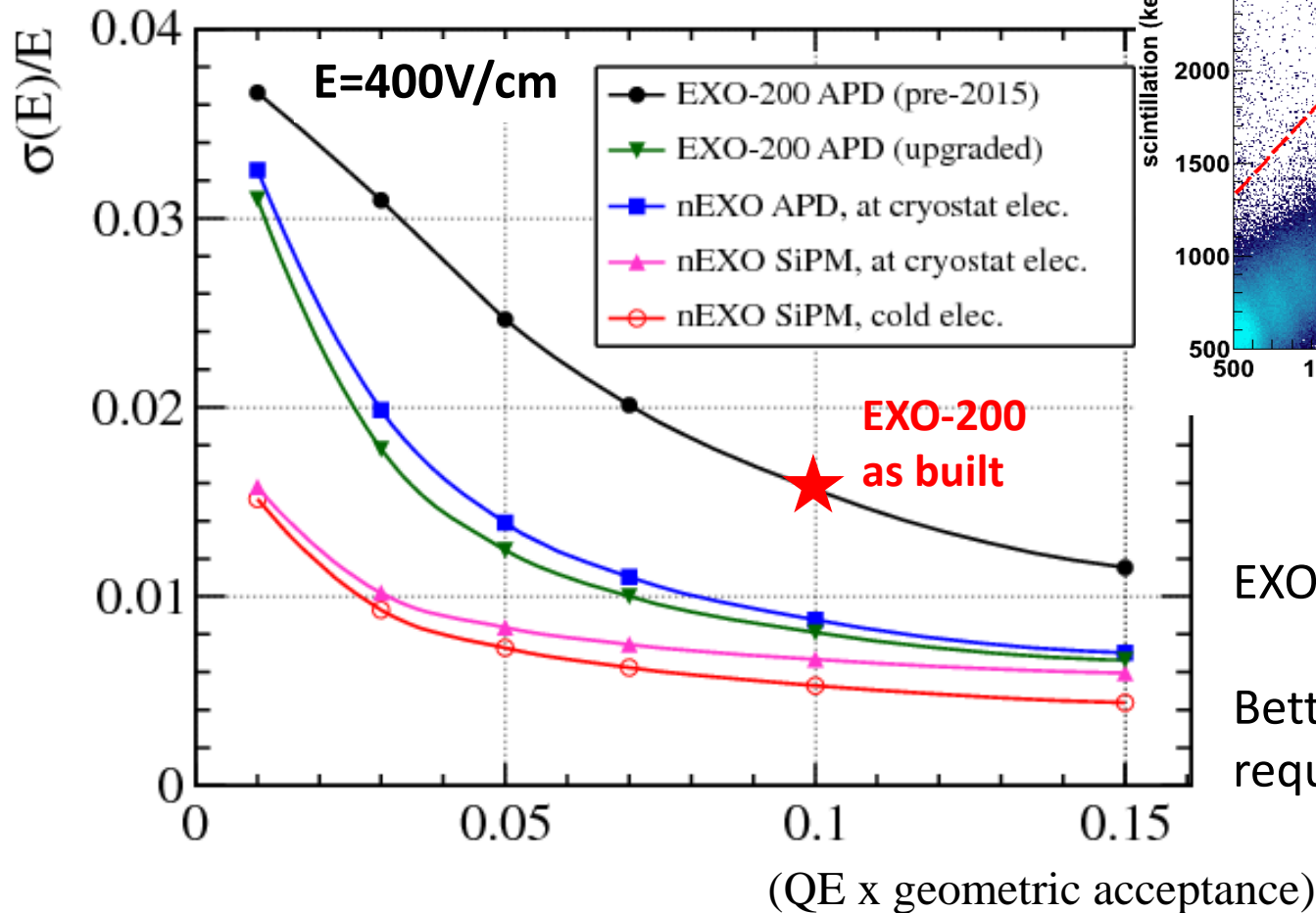
NMEs used:

Model:	$M_{0\nu}$:	Reference:
EDF	4.20	PRL 105, 252503 (2010)
ISM	2.19	Nucl Phys A 818, 139 (2009)
IBM-2	3.05	PRC 91, 034304 (2015)
Skyrme QRPA	1.55	PRC 87 064302 (2013)
QRPA	2.02	PRC 89, 064308 (2014)



Photon Detection

Good energy resolution requires efficient readout of the scintillation light to be combined with the ionization signal

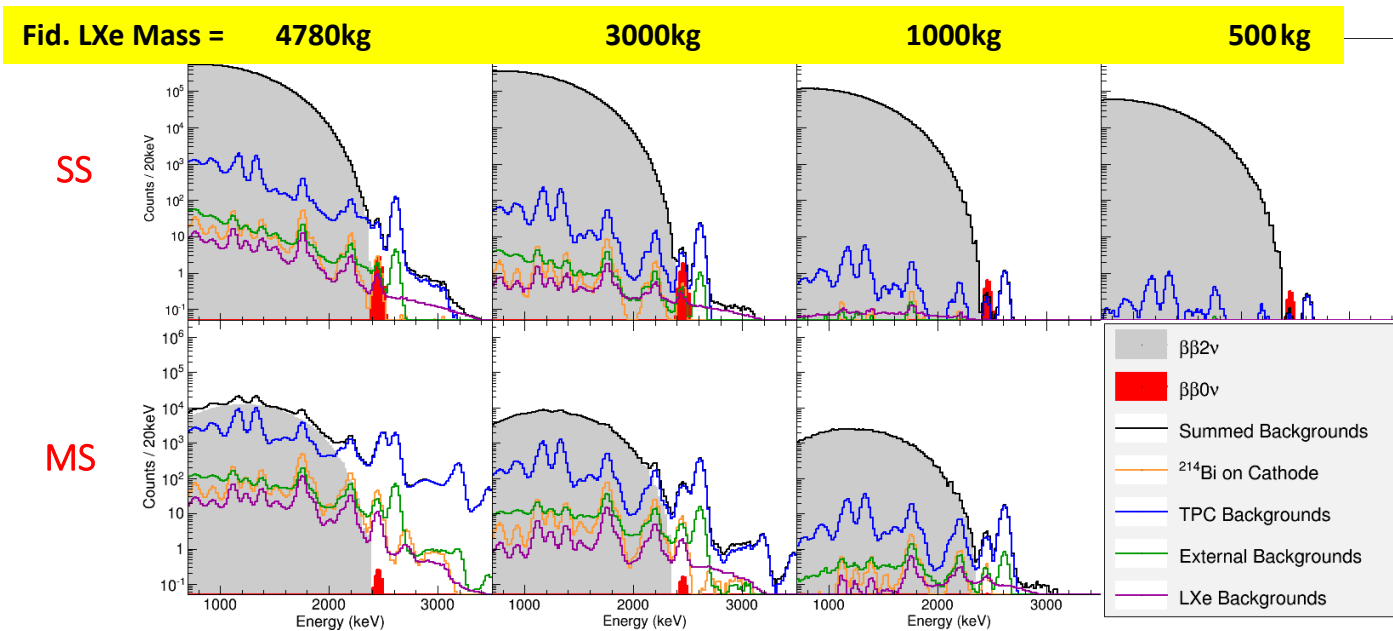


EXO-200 reaches 1.4% at $Q_{\beta\beta}$

Better than 1% resolution required for nEXO

The role of the standoff distance in background identification and suppression

Example: nEXO, 5 yr data, $0\nu\beta\beta$ @ $T_{1/2}=6.6\times 10^{27}$ yr,
projected backgrounds from subsets of the total volume



The fit gets to see all this information and use it in the optimal way