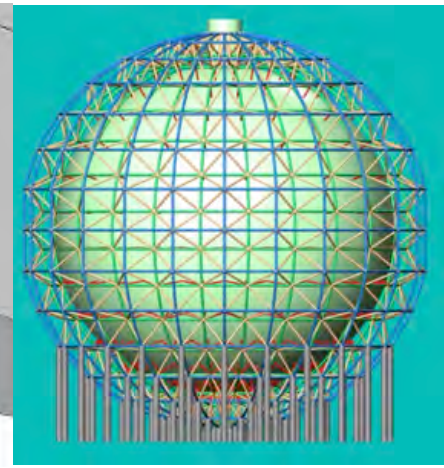
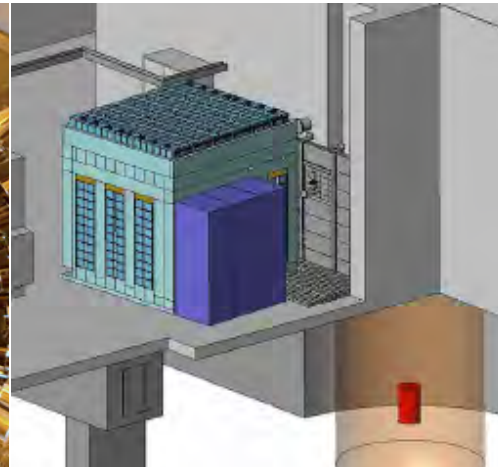
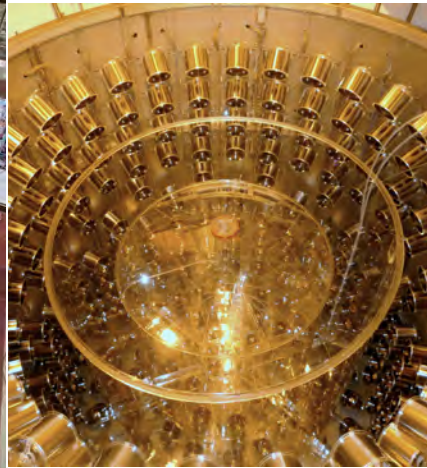
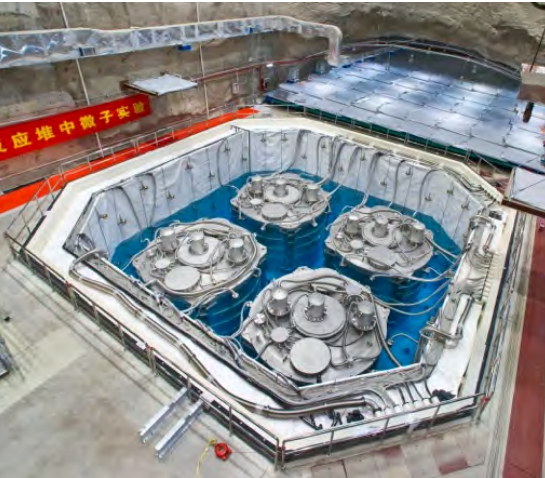


Reactor Neutrino Experiments



Karsten M. Heeger
Yale University

ACP, January 14, 2015

Reactor Antineutrinos

A Tool for Discovery

2003 - First observation of reactor antineutrino disappearance

1995 - Nobel Prize to Fred Reines at UC Irvine

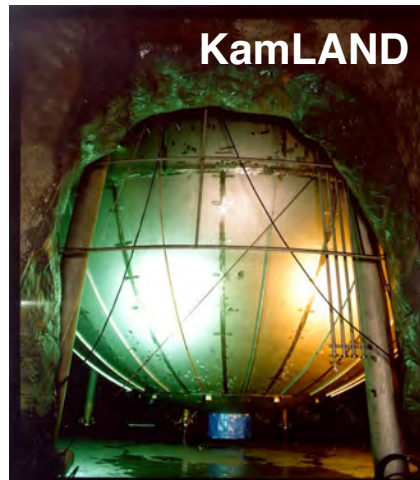


1956 - First observation of (anti)neutrinos



Savannah River

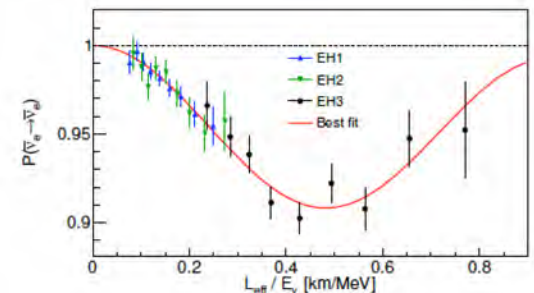
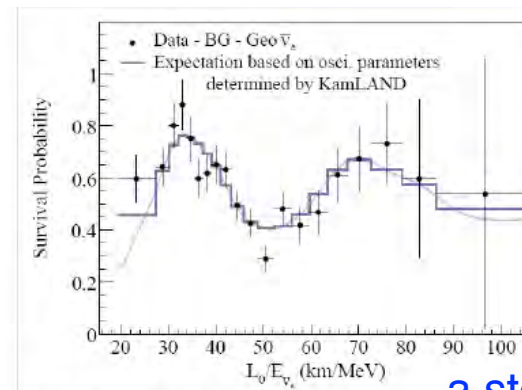
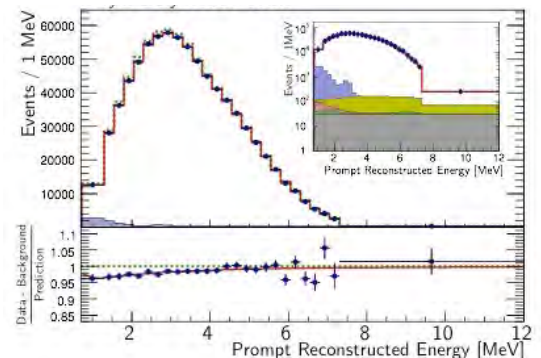
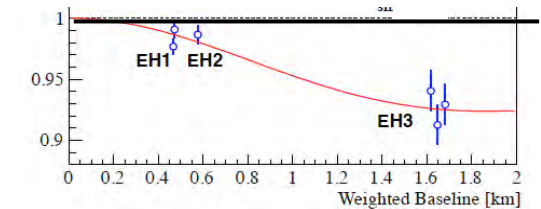
2012 - Measurement of θ_{13} with Reactor Neutrinos



KamLAND



Daya Bay, Double Chooz, RENO



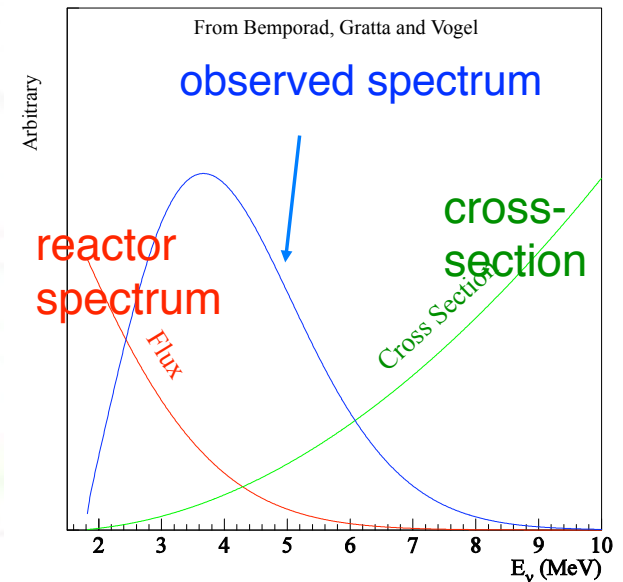
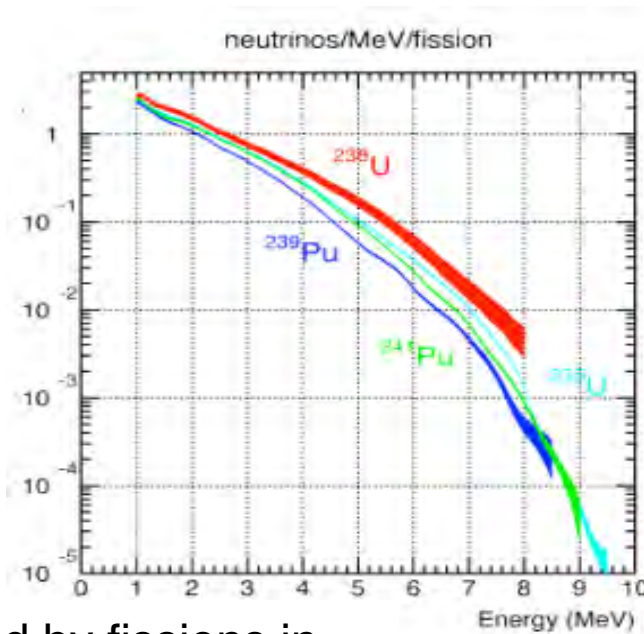
a story of varying baselines... 2

Reactor Antineutrinos

$\bar{\nu}_e$ from β -decays, pure $\bar{\nu}_e$ source

of n-rich fission products

on average ~ 6 beta decays until stable



$> 99.9\%$ of $\bar{\nu}_e$ are produced by fissions in
 ^{235}U , ^{238}U , ^{239}Pu , ^{241}Pu

$$\frac{d^2N(E,t)}{dEdt} \equiv \sum_i \frac{W_{ih}(t)}{\sum_j f_j(t)e_j} f_i(t) S_i(E) c_i^{ne}(E,t) + S_{SNF}(E,t) \quad (1)$$

mean energy of $\bar{\nu}_e$: 3.6 MeV

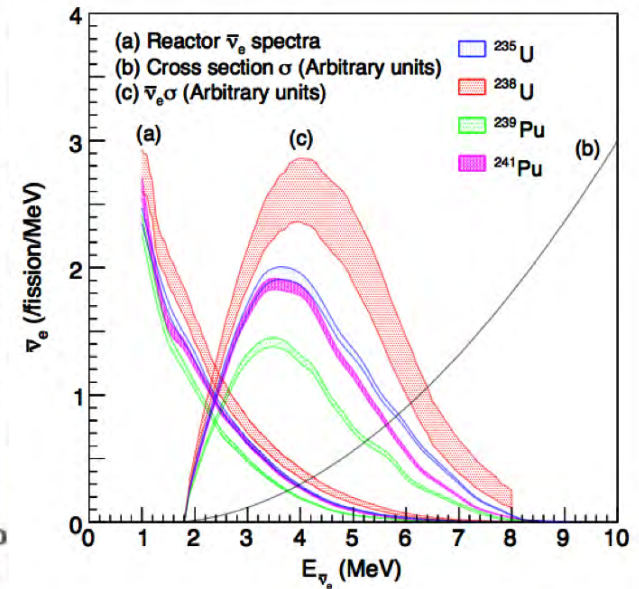
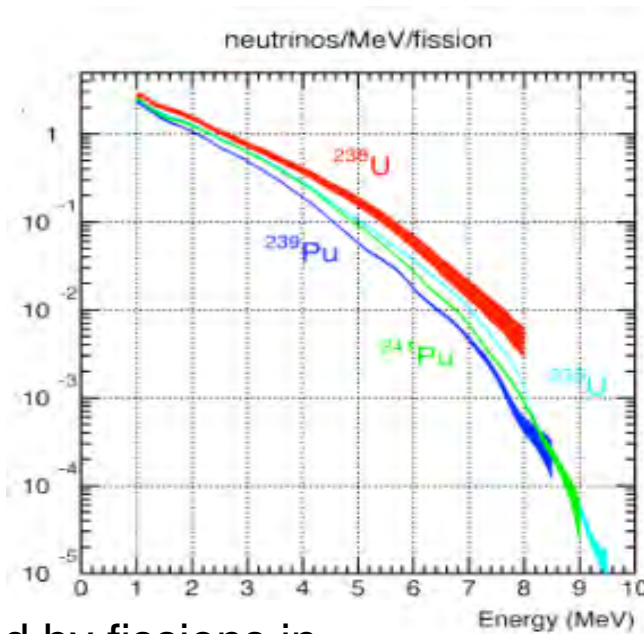
only disappearance
experiments possible

Reactor Antineutrinos

$\bar{\nu}_e$ from β -decays, pure $\bar{\nu}_e$ source

of n-rich fission products

on average ~ 6 beta decays until stable



> 99.9% of ν_e are produced by fissions in

^{235}U , ^{238}U , ^{239}Pu , ^{241}Pu

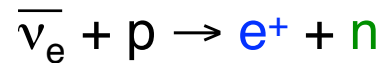
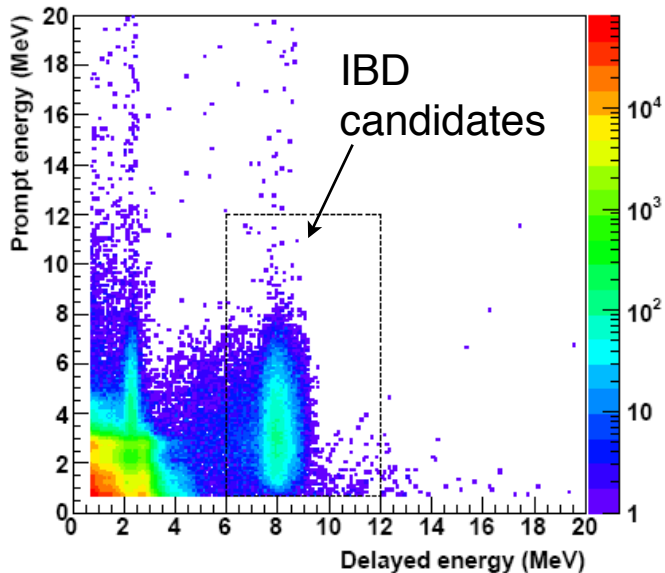
$$\frac{d^2N(E,t)}{dEdt} \equiv \sum_i \frac{W_{ih}(t)}{\sum_j f_j(t)e_j} f_i(t) S_i(E) c_i^{ne}(E,t) + S_{SNF}(E,t) \quad (1)$$

mean energy of $\bar{\nu}_e$: 3.6 MeV

only disappearance experiments possible

Antineutrino Candidates (Inverse Beta Decay)

Prompt + Delayed Coincidence



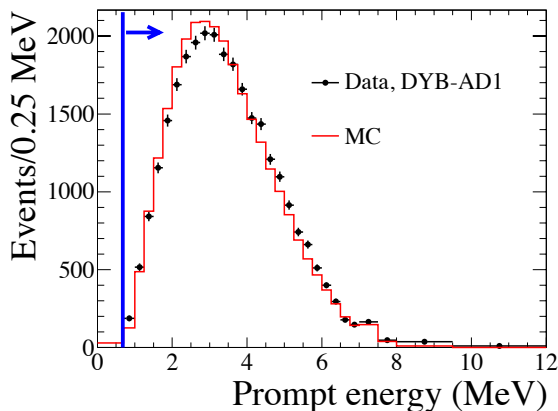
prompt event:

positron deposits energy and annihilates (\sim ns)

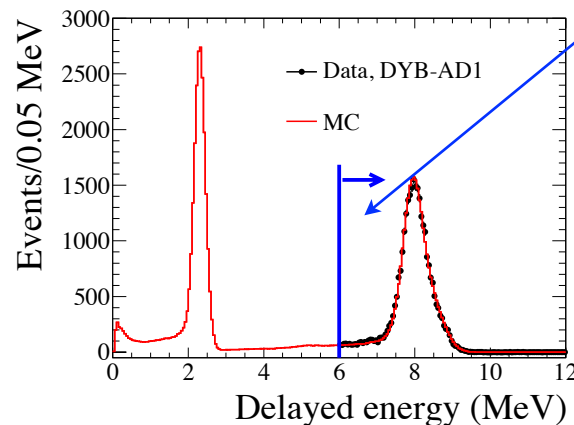
delayed event:

neutron thermalizes and captures on **Gd**

Prompt Energy Signal

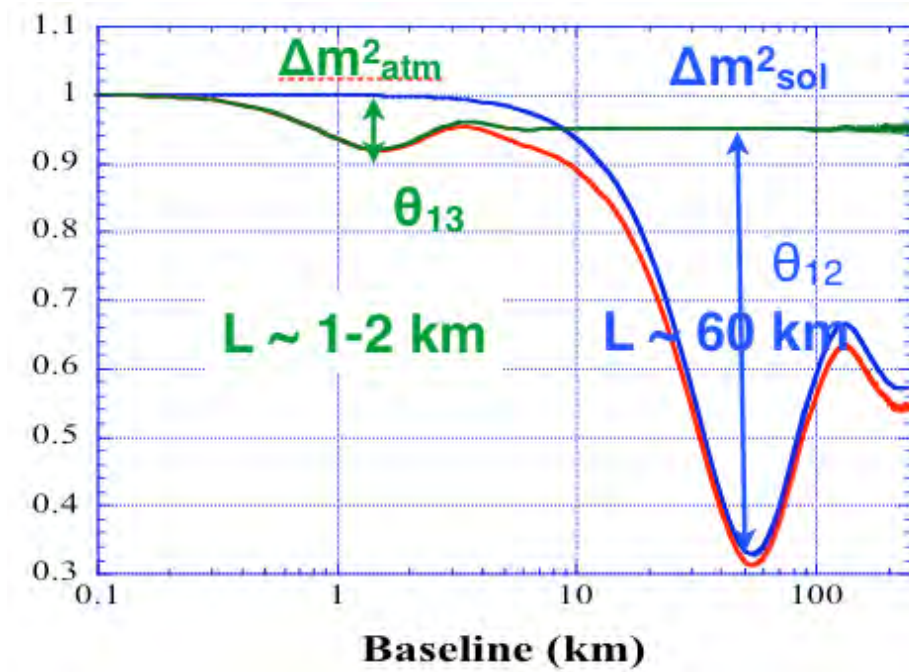


Delayed Energy Signal



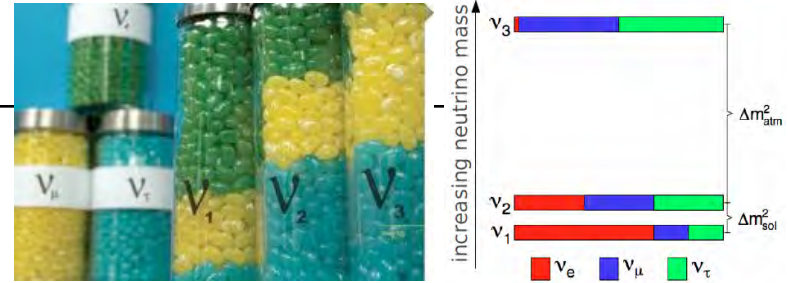
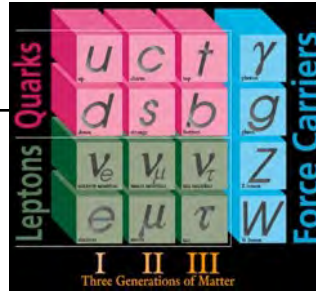
Uncertainty in relative E_d efficiency (0.12%) between detectors is largest systematic.

Oscillation Measurements



Neutrino Mixing

Mixing Angles



$$U = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} = \begin{pmatrix} 0.8 & 0.5 & U_{e3} \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix} \quad \mathbf{U}_{\text{MNSP Matrix}}$$

Maki, Nakagawa, Sakata, Pontecorvo

$$= \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}}_{\text{atmospheric, K2K}} \times \underbrace{\begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta_{CP}} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta_{CP}} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix}}_{\text{reactor and accelerator}} \times \underbrace{\begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{SNO, solar SK, KamLAND}} \times \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha/2} & 0 \\ 0 & 0 & e^{i\alpha/2+i\beta} \end{pmatrix}}_{0\nu\beta\beta}$$

$$\sin^2 \theta_{23}$$

$$0.50^{+0.07}_{-0.06}$$

maximal?

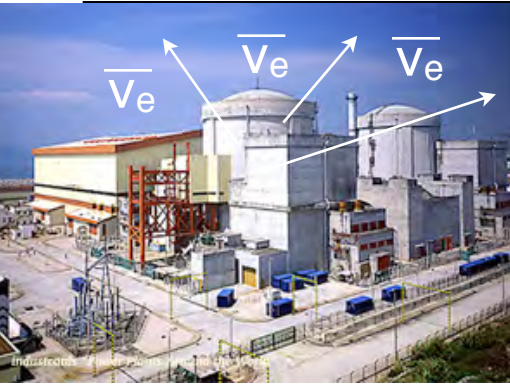
$$\sin^2 \theta_{13}$$

$$\sin^2 \theta_{12}$$

$$0.318^{+0.019}_{-0.016}$$

large, but not maximal!

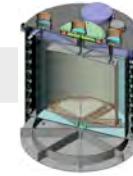
Reactor Neutrino Oscillations



$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E_\nu}\right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2\left(\frac{\Delta m_{21}^2 L}{4E_\nu}\right)$$

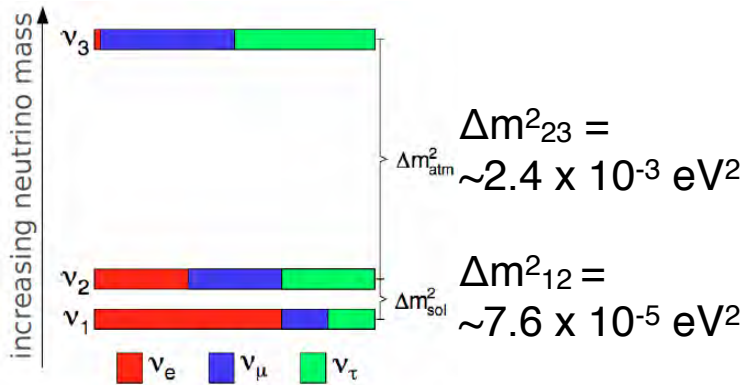
$\bar{\nu}_e$

$\bar{\nu}_{e,x}$

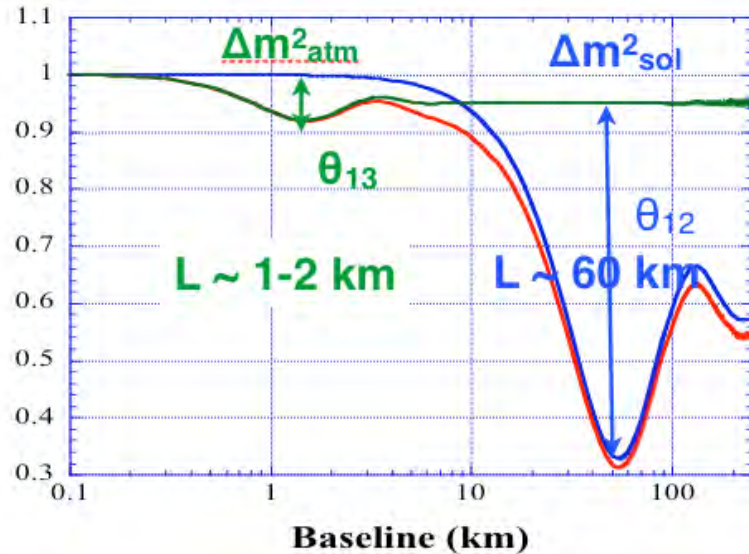


$\bar{\nu}_{e,x}$

for 3 active ν , two different oscillation length scales: $\Delta m_{12}^2, \Delta m_{23}^2$

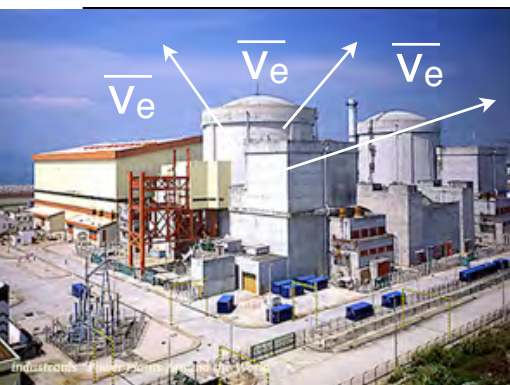


amplitude of oscillation θ

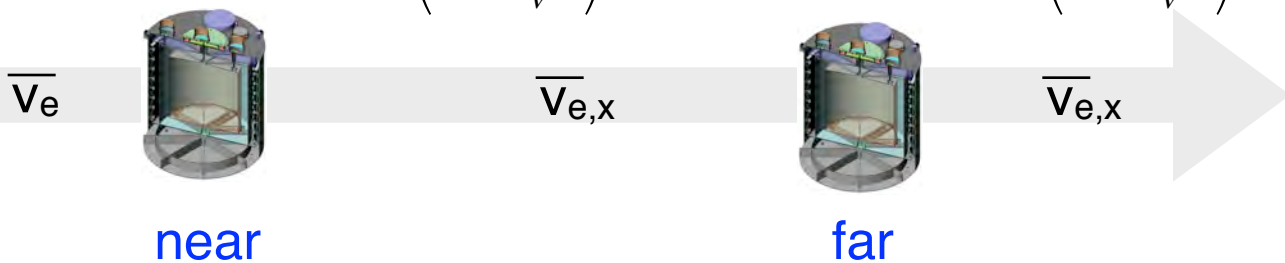


oscillation frequency $L/E \rightarrow \Delta m^2$

Reactor Neutrino Oscillations

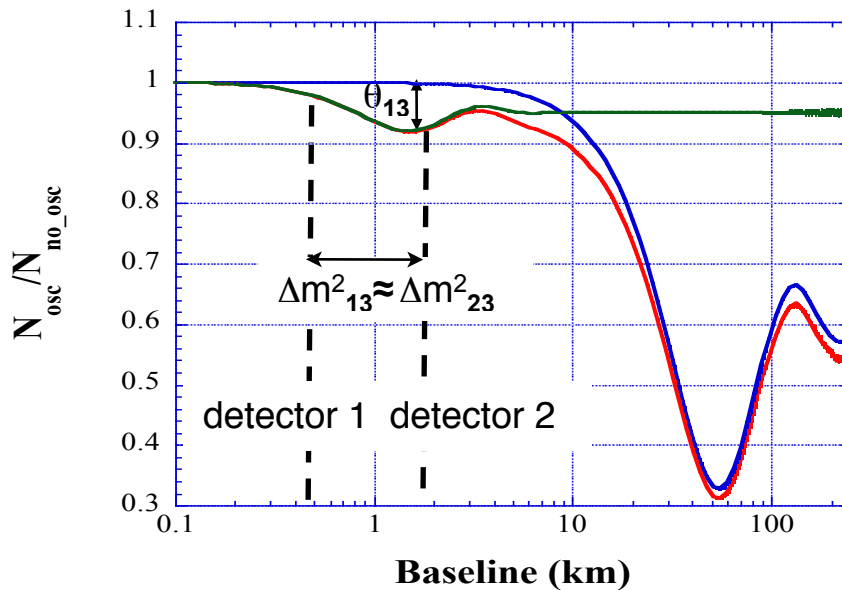


$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2\left(\frac{\Delta m_{31}^2 L}{4E_\nu}\right) - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2\left(\frac{\Delta m_{21}^2 L}{4E_\nu}\right)$$



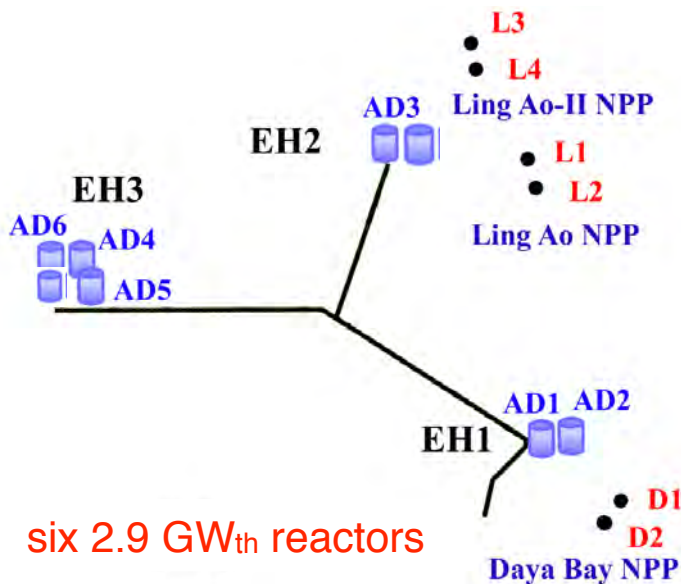
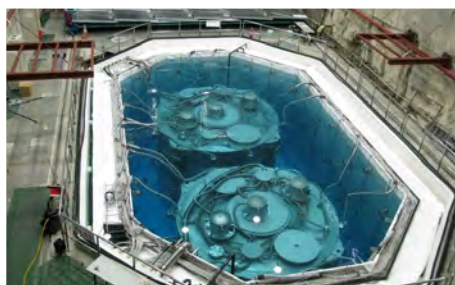
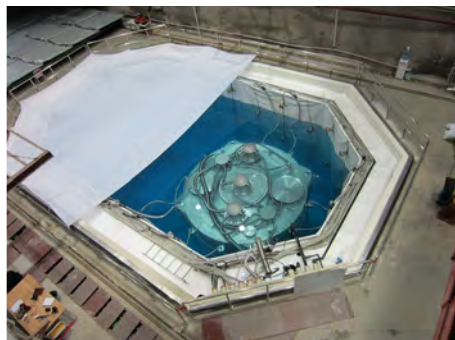
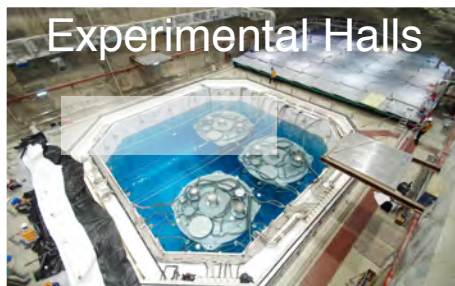
Absolute Reactor Flux
Largest uncertainty in previous measurements

Relative Measurement
Removes absolute uncertainties!



relative measurement (largely) cancels reactor systematics

Daya Bay Reactor Experiment

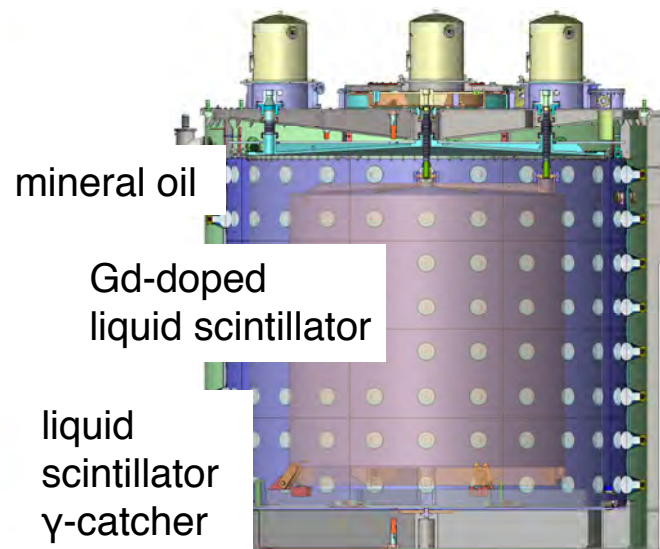


six 2.9 GW_{th} reactors

6 detectors, Dec 2011- Jul 2012
217 days

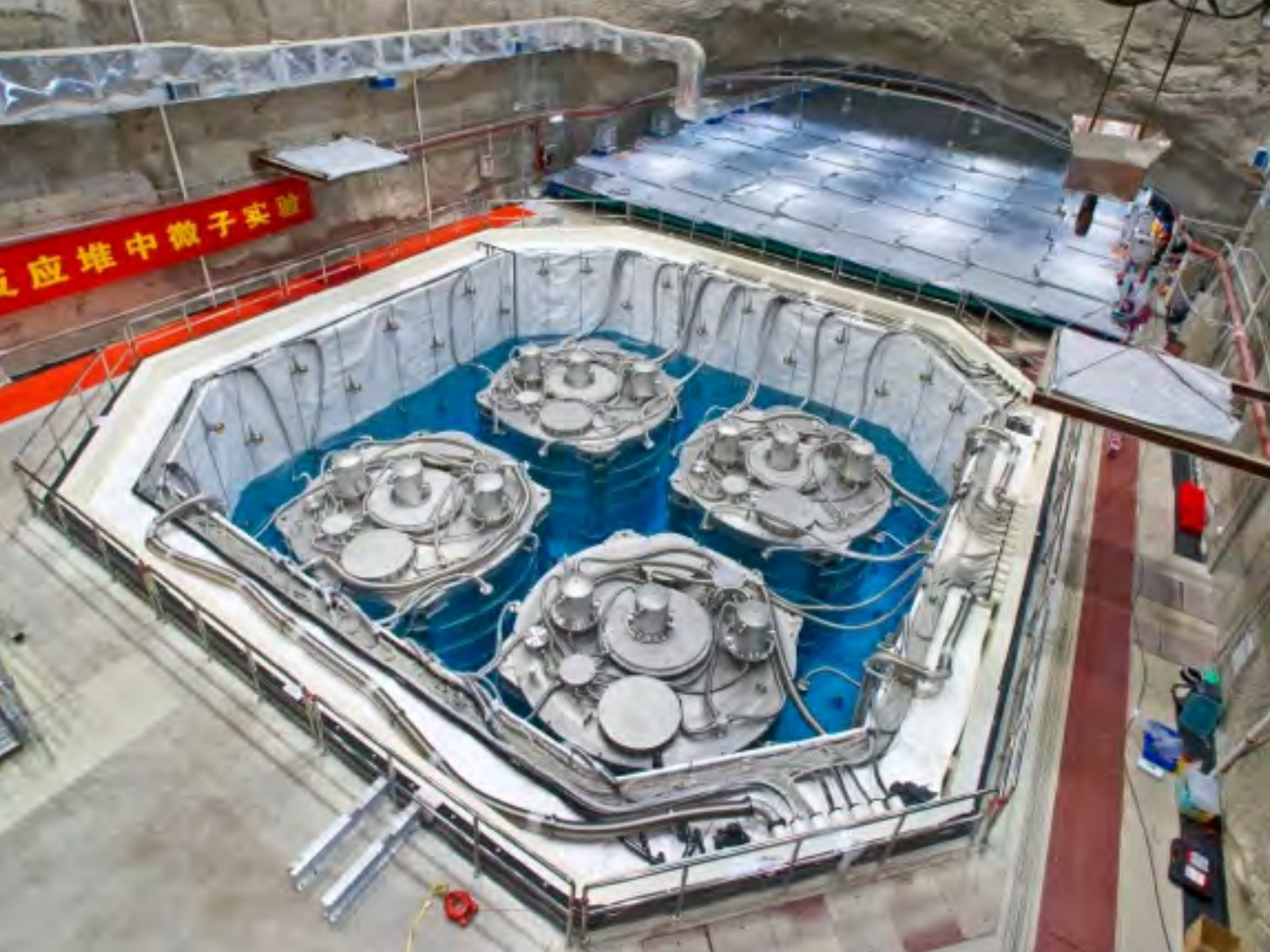
now running with 8 detectors

Antineutrino Detector



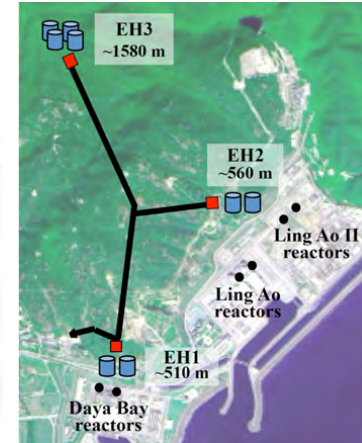
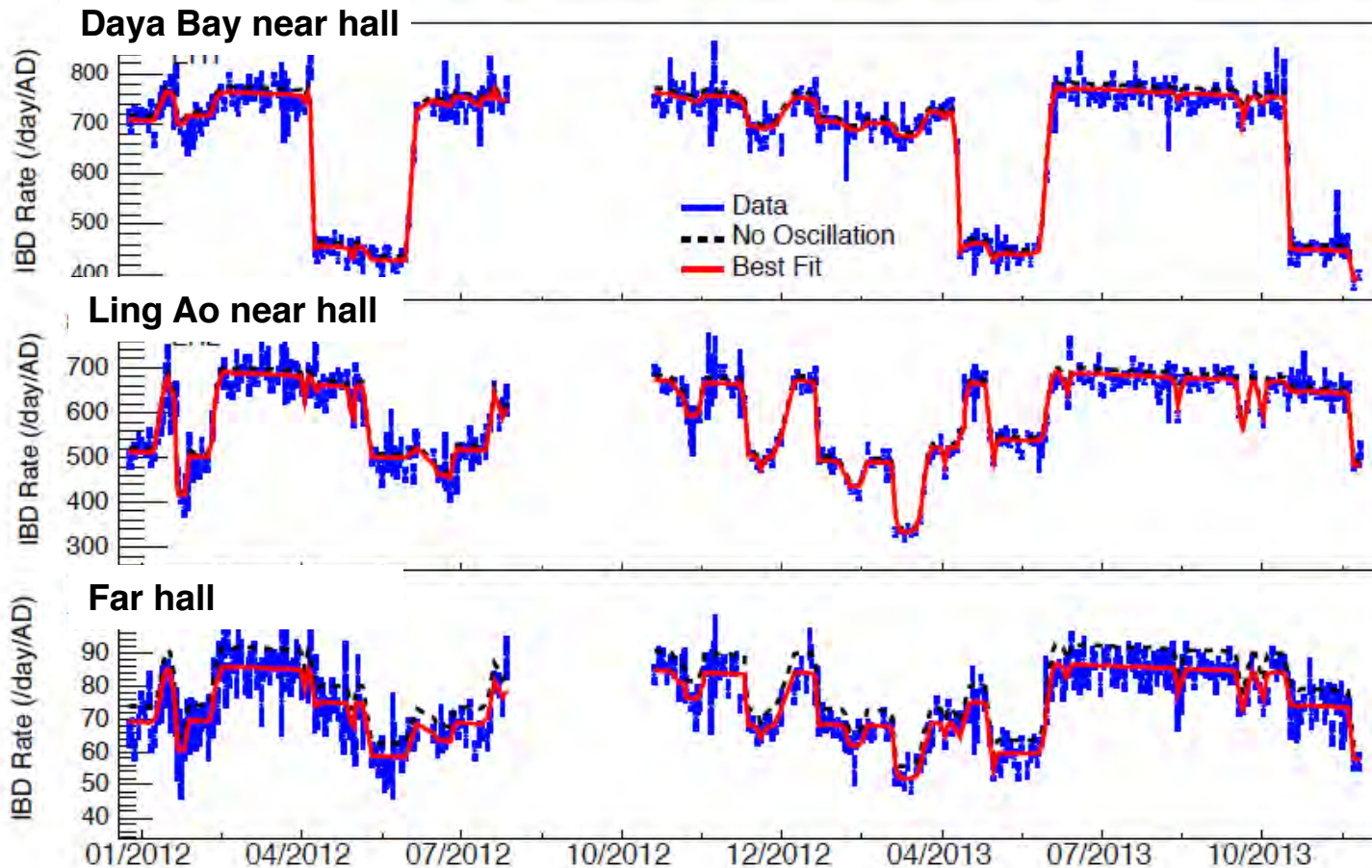
target mass: 20 ton per AD
 photosensors: 192 8"-PMTs
 energy resolution: $(7.5 / \sqrt{E} + 0.9)\%$

反应堆中微子实验



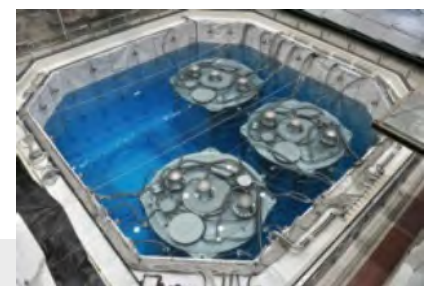
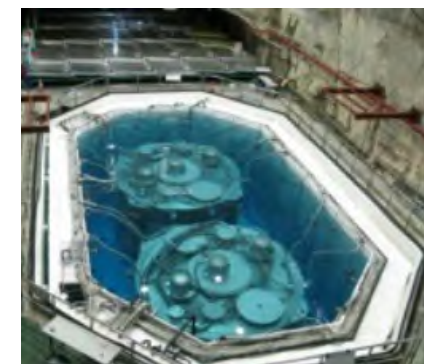
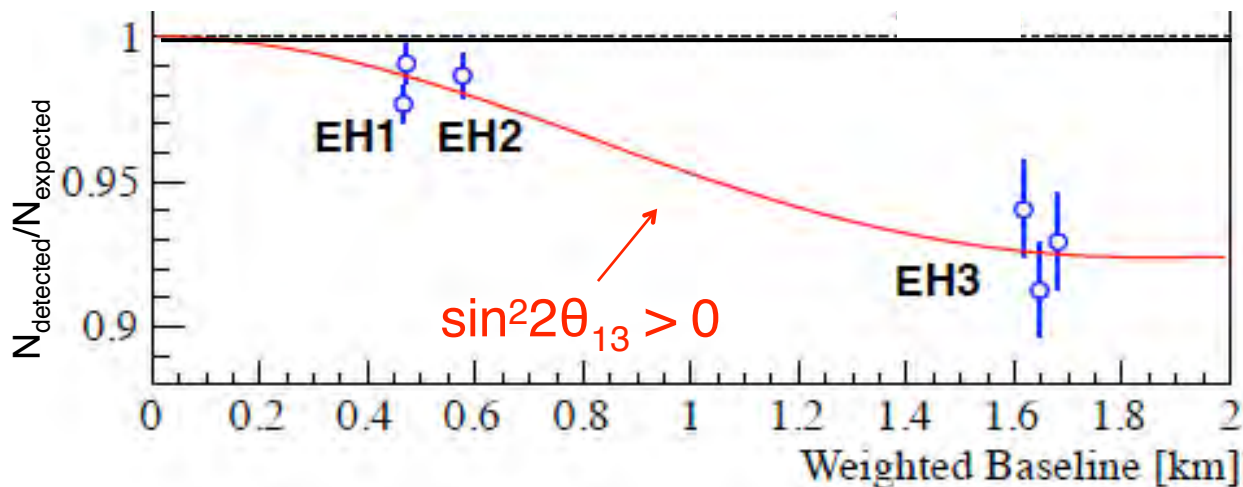
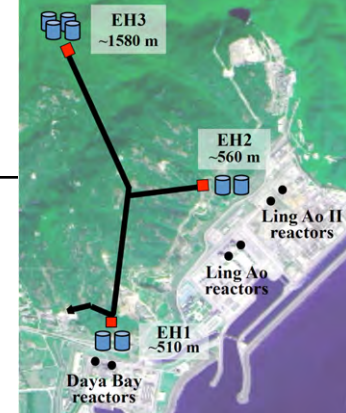
Daya Bay Antineutrino Rate vs Time

Over 1 Million Antineutrino Interactions Detected



Observation of $\bar{\nu}_e$ Disappearance

Based on 55 days of data with 6 ADs, discovered disappearance of reactor $\bar{\nu}_e$ at short baseline. [PRL **108**, 171803]



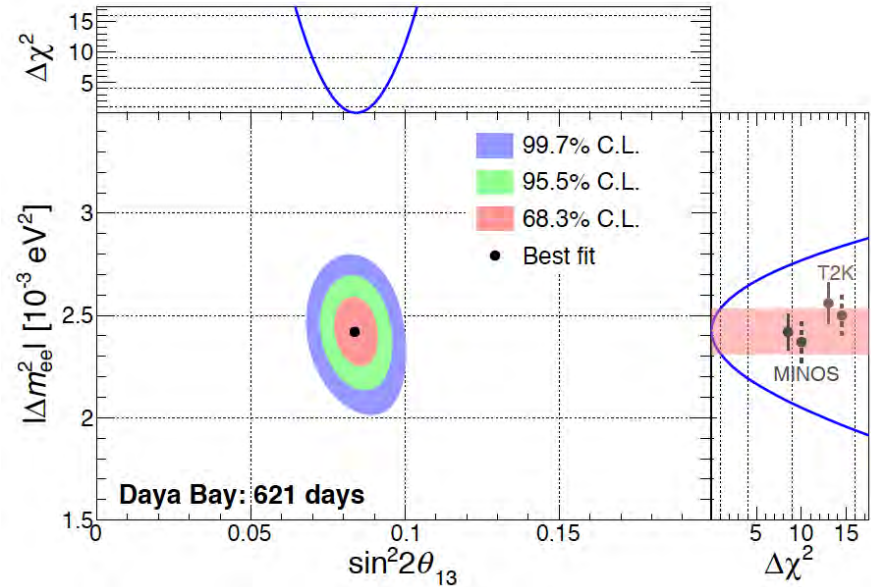
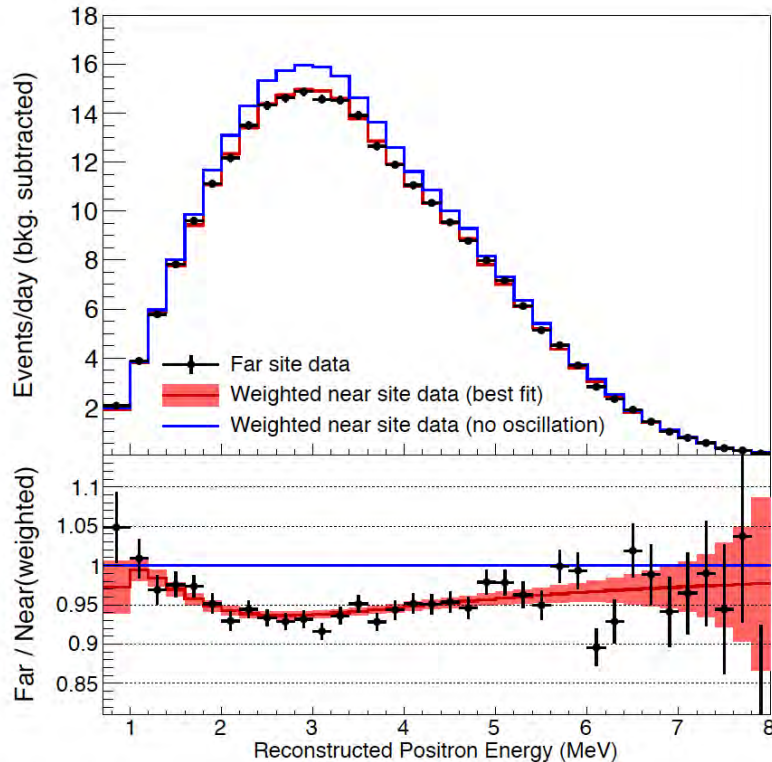
Obtained the most precise value of θ_{13} :

$$\sin^2 2\theta_{13} = 0.089 \pm 0.010 \pm 0.005 \quad [\text{CPC } \mathbf{37}, 011001]$$

One of Science's breakthroughs of year 2012

Daya Bay Neutrino Oscillation

621 days of data, n+Gd



$$\sin^2 2\theta_{13} = 0.084^{+0.005}_{-0.005}$$

$$|\Delta m_{ee}^2| = 2.44^{+0.10}_{-0.11} \times 10^{-3} \text{ eV}^2$$

$$\chi^2/NDF = 134.7/146$$

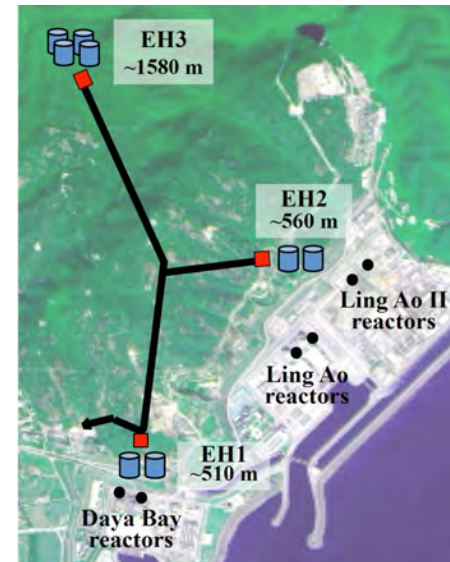
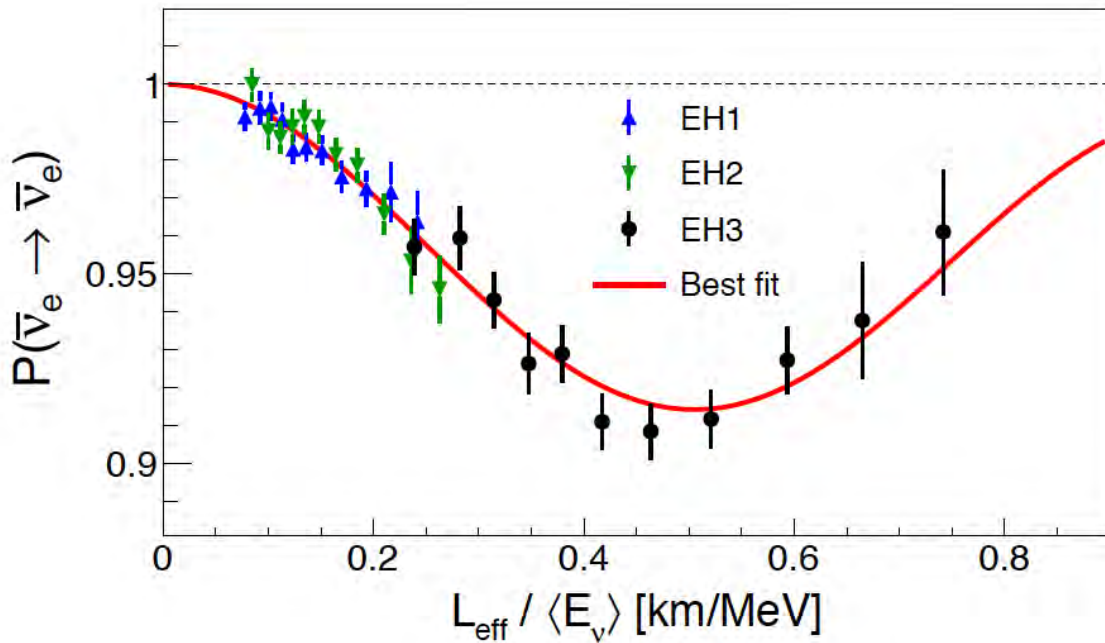
most precise measurement of $\sin^2 2\theta_{13}$ (6%), and Δm_{ee}^2 in the electron neutrino disappearance channel (4%)

Phys.Rev.Lett. 115 (2015) 11, 111802

Daya Bay Neutrino Oscillation

Neutrino oscillation is energy and baseline dependent

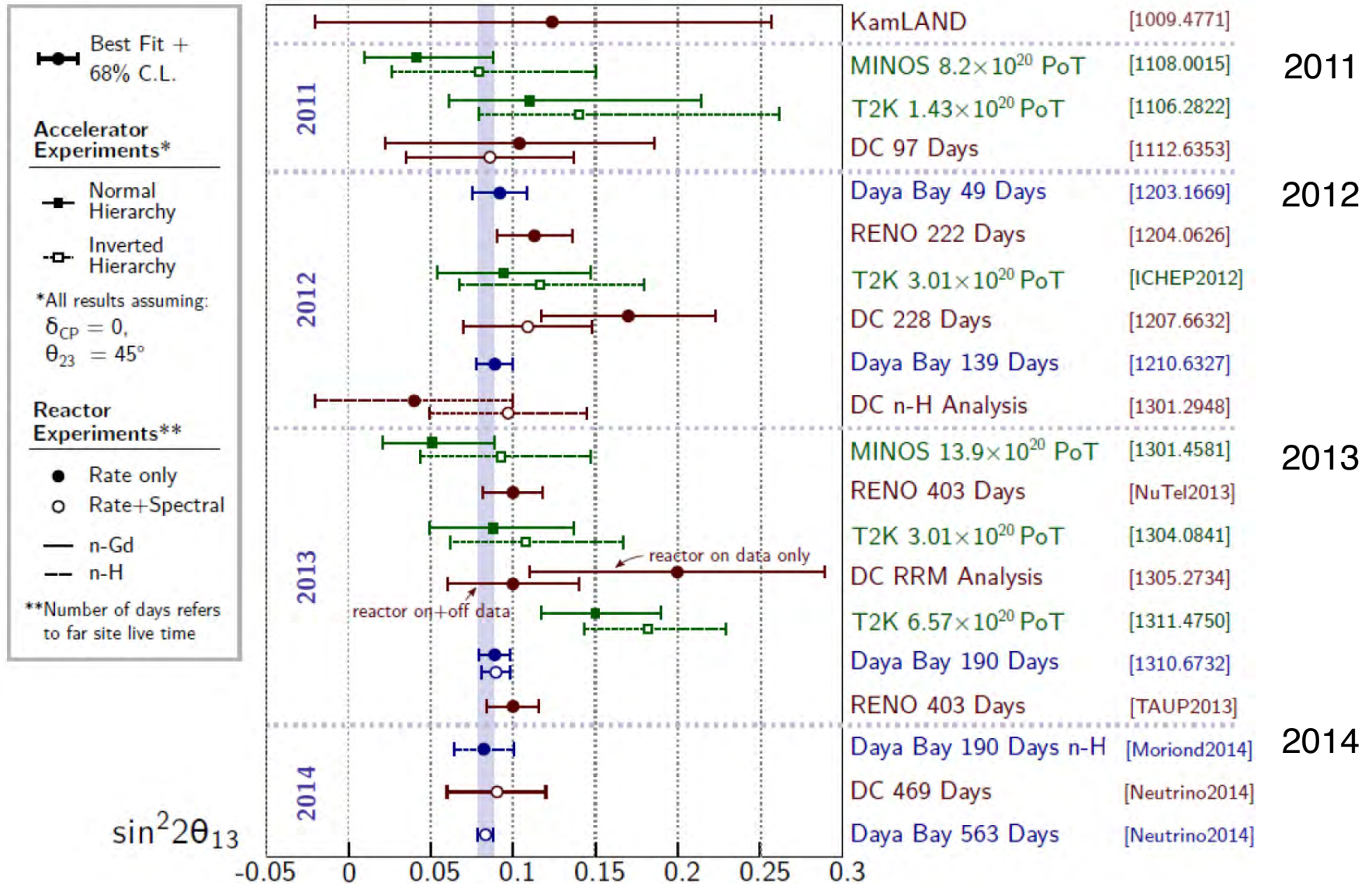
$$P_{i \rightarrow j} = \sin^2 2\theta \sin^2 \left(1.27 \Delta m^2 \frac{L}{E} \right)$$



Daya Bay demonstrates L/E oscillation

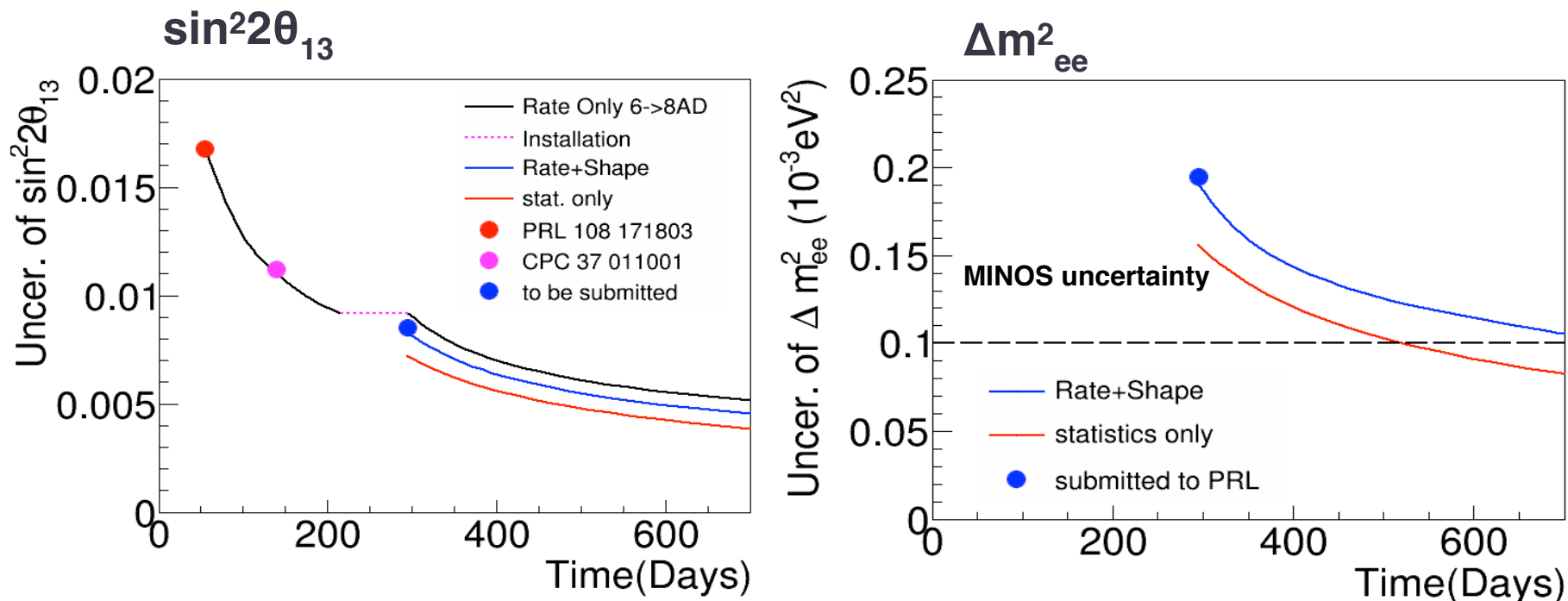
Phys.Rev.Lett. 115 (2015) 11, 111802

Daya Bay Precision Measurement of θ_{13}



Daya Bay Sensitivity Projections

Precision Measurements in $\sin^2 2\theta_{13}$ and Δm^2_{ee}



Daya Bay remains statistically limited through 2015. Will also improve systematics.

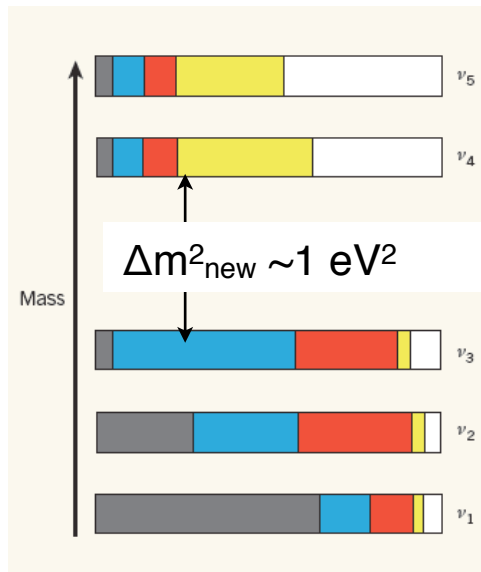
Major systematics:

θ_{13} : Relative + absolute energy, and relative efficiencies

$|\Delta m^2_{ee}|$: Relative energy model, relative efficiencies, and backgrounds

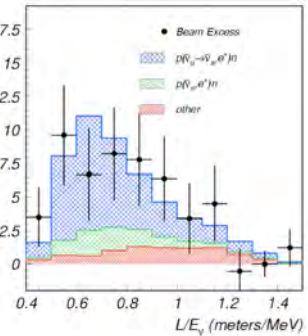
Aim to improve precision of $\sin^2 2\theta_{13}$ and Δm^2_{ee} to 3% by 2017.

ν Anomalies Beyond 3 Neutrinos?

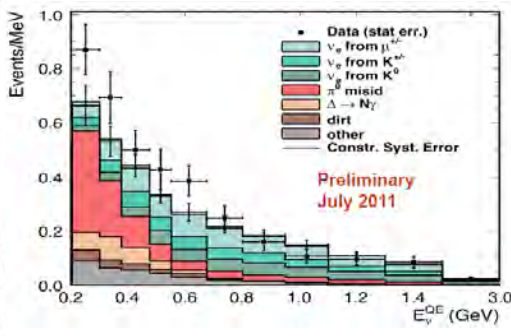


Neutrino Anomalies - More than 3 ν ?

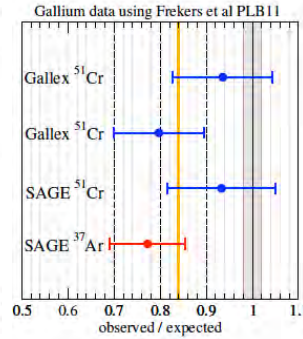
LSND



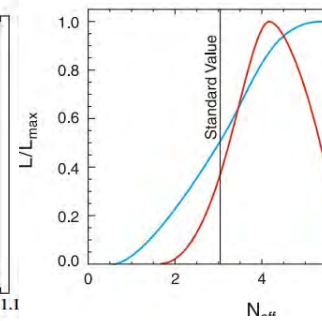
MiniBoone



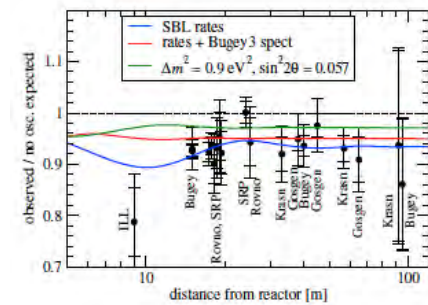
Ga Source



Cosmology (WMAP)



Reactor



Anomalies in 3- ν global oscillation data

LSND ($\bar{\nu}_e$ appearance)

MiniBoone (ν_e appearance)

Ga anomaly

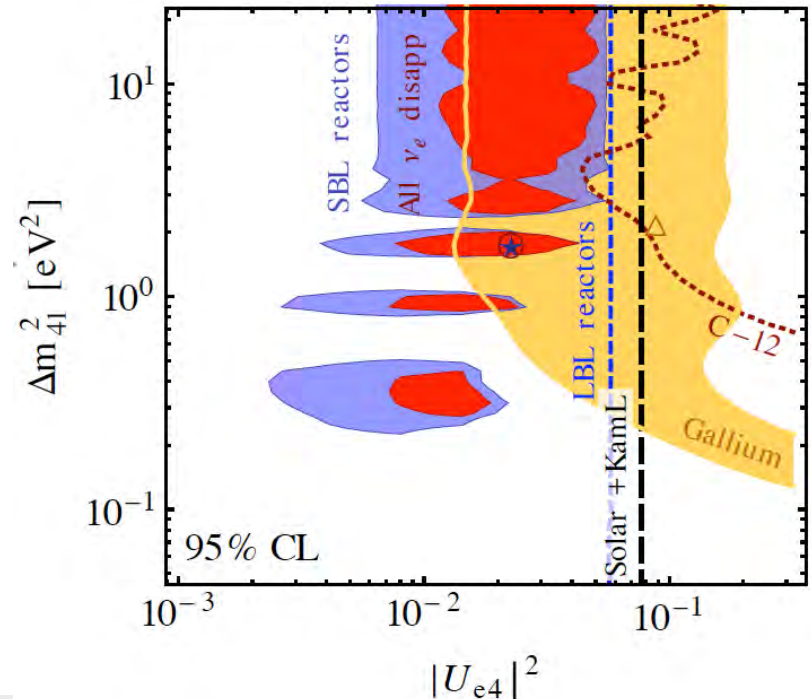
N_{eff} in cosmology

Reactor anomaly and spectrum ($\bar{\nu}_e$ disappearance)

new oscillation signal requires:

$\Delta m^2 \sim O(1\text{eV}^2)$ and $\sin^2 2\theta > 10^{-3}$

“sterile” neutrino states

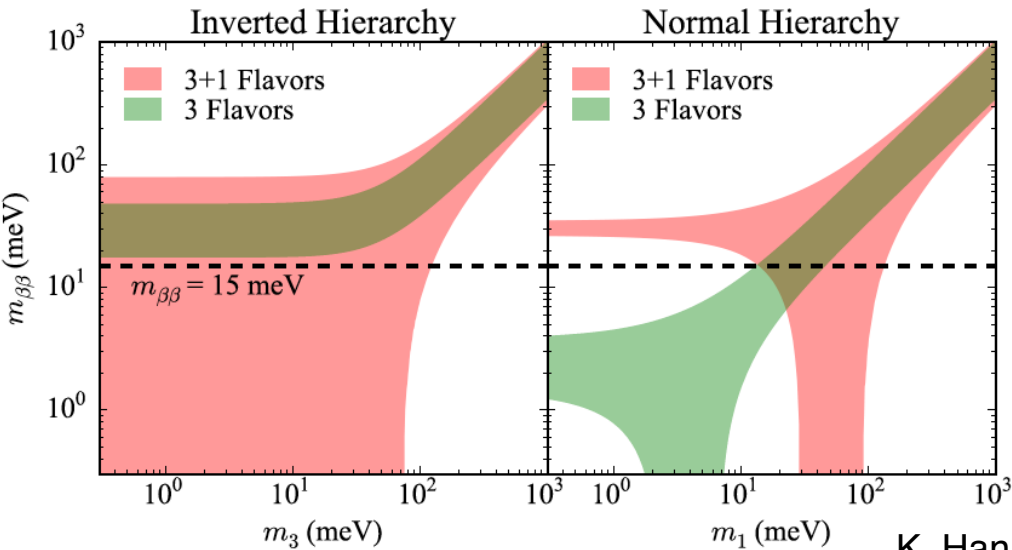


Implications for Future Neutrino Program

Discovery of eV-scale sterile neutrinos would be a paradigm change for particle physics.

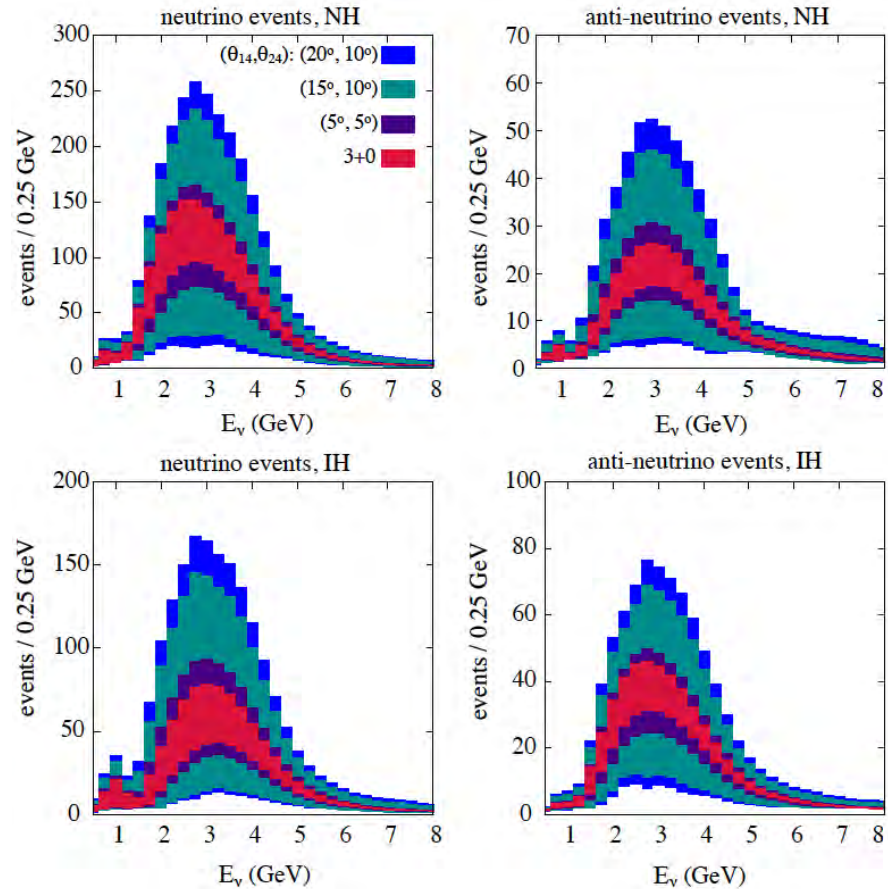
- Expected neutrino spectrum and sensitivity to CP violation for long-baseline neutrino program
- Effective neutrino mass measured by $0\nu\beta\beta$

Neutrinoless Double Beta Decay



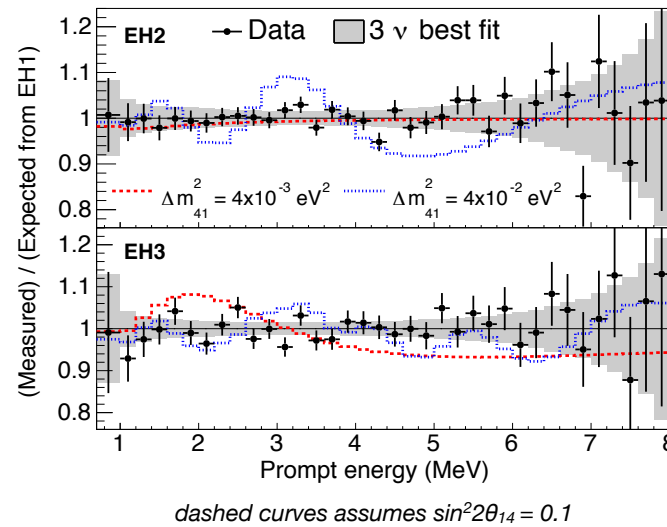
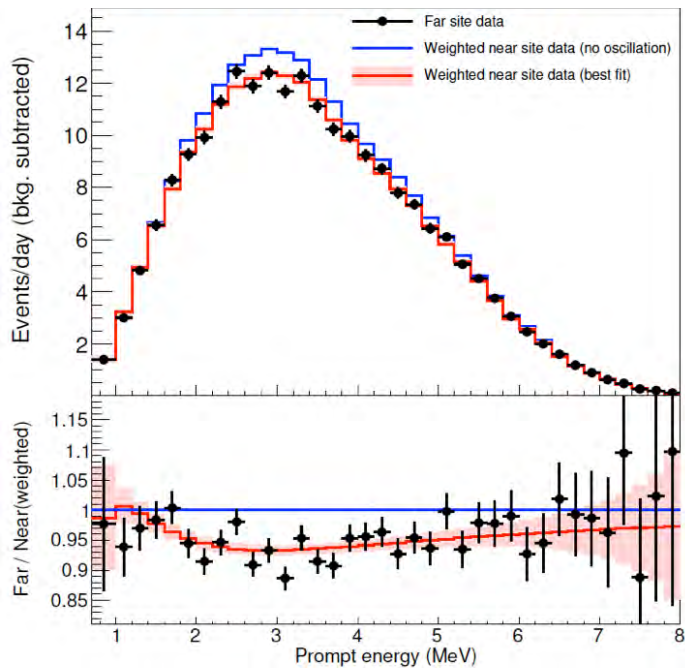
K. Han

DUNE



Gandhi, Kayser, Masud, Prakash arXiv: 1508.06275

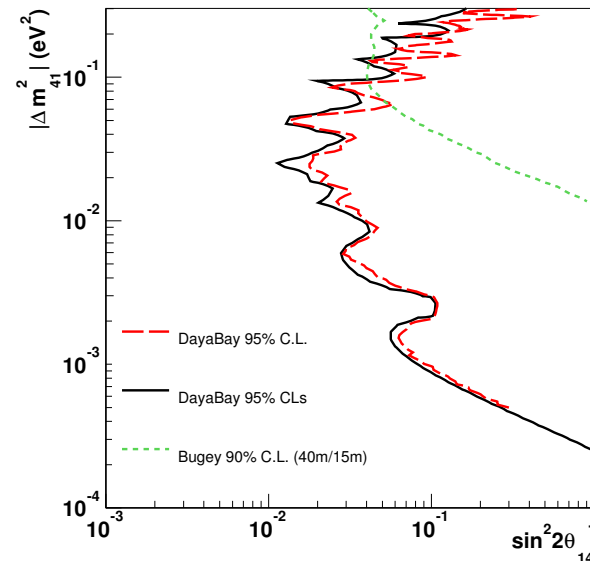
Search for Sterile Neutrinos at Daya Bay



Phys. Rev. Lett. 113, 141802 (2014)

sterile neutrinos would appear as additional spectral distortion and overall rate deficit

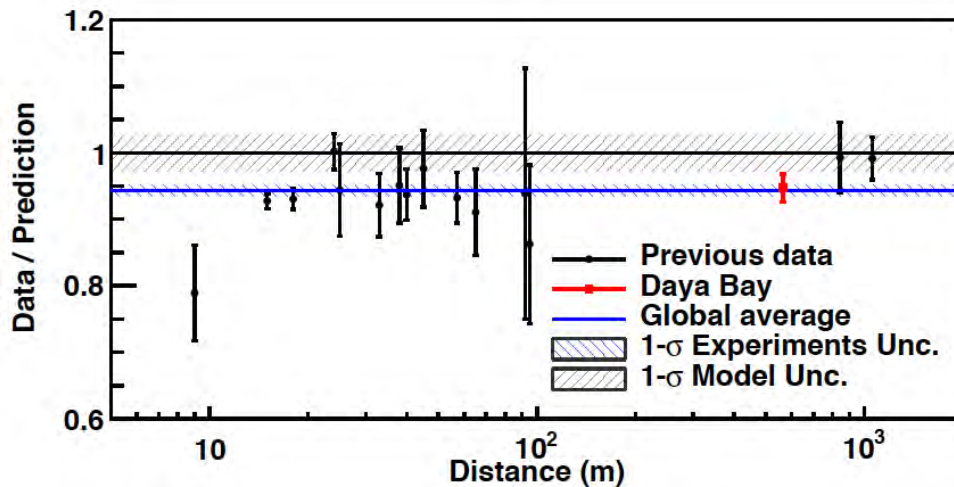
Daya Bay probes largely unexplored region at $\Delta m_{41}^2 < 0.1 \text{ eV}^2$



Reactor Flux and Spectrum “Anomalies”

Flux Deficit

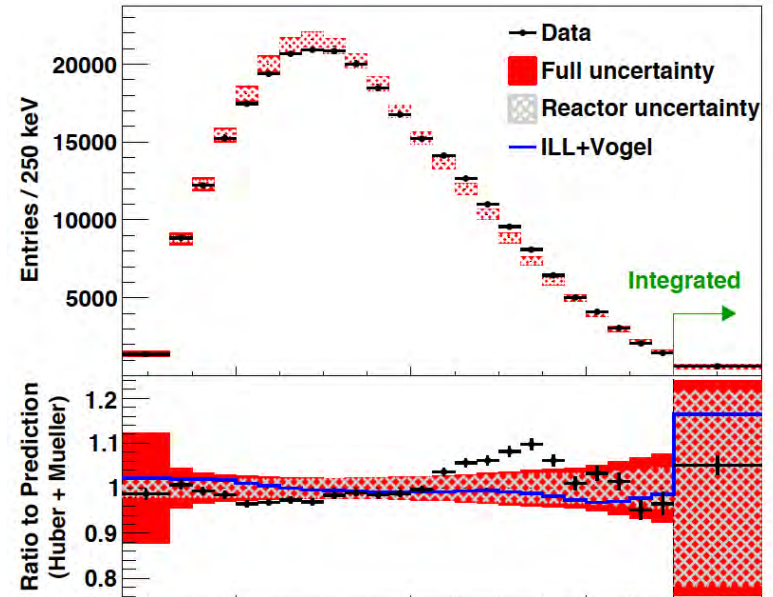
Consistent with previous experiments



Extra neutrino oscillations or artifact of flux predictions?

Understanding reactor flux and spectrum anomalies requires reactor measurements

Spectral Deviation



New feature in 4-6 MeV region of spectrum.

arXiv:1508.04233, accepted by PRL
Daya Bay collaboration

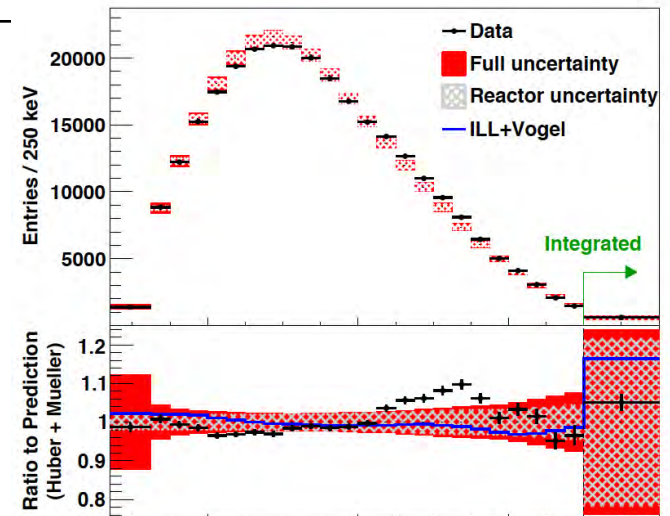
Reactor Spectrum Anomaly

Spectral deviation

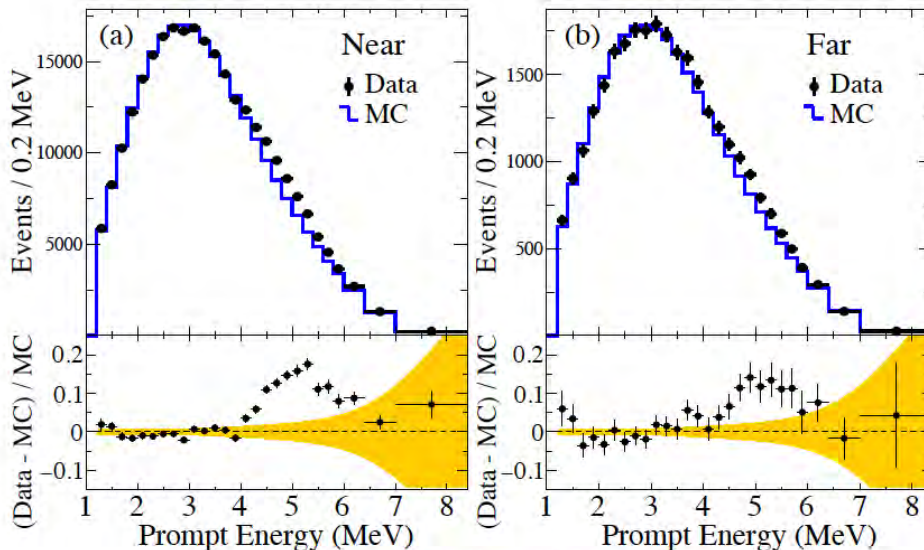
- 10% excess in 4-6 MeV region when compared to model calculations

Observed in all 3 θ_{13} experiments

Daya Bay

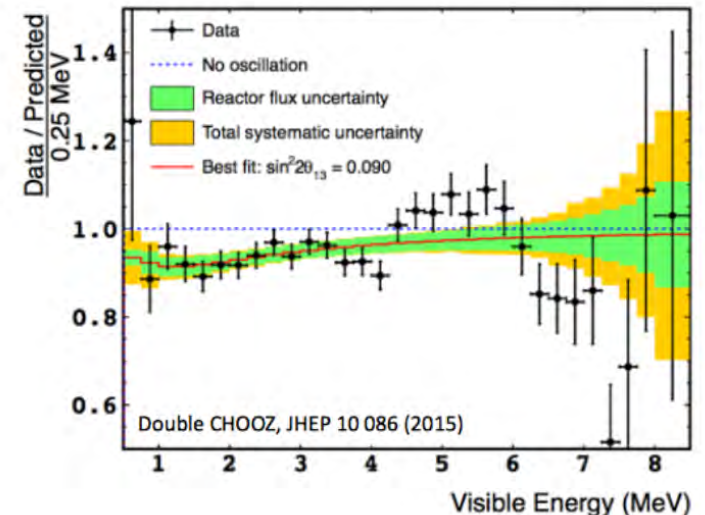


RENO



RENO: arXiv: 1511.05849

Double Chooz



Modeling the Reactor Spectrum

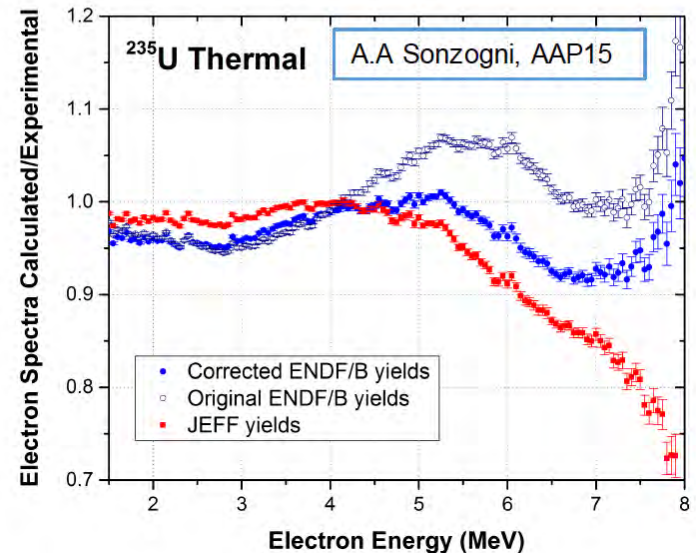
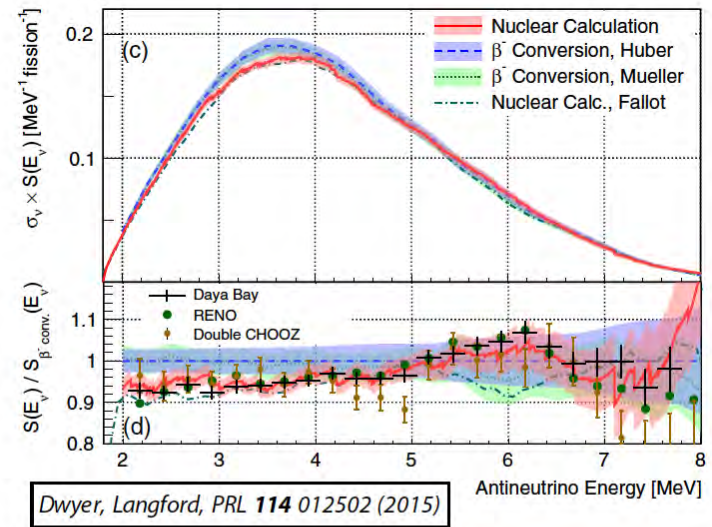
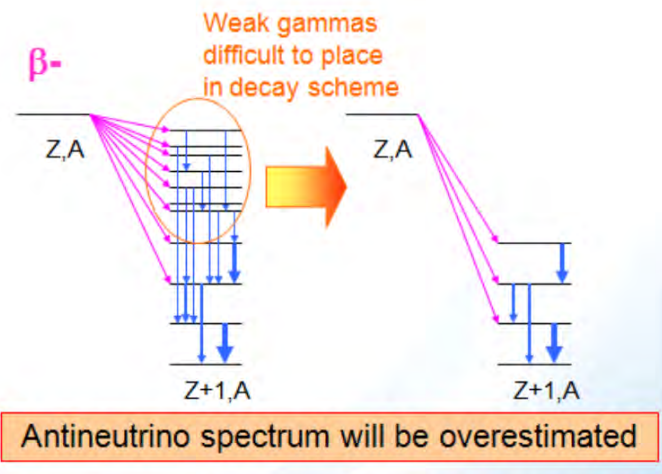
Challenges

Reactor neutrino spectrum is an admixture of thousands beta branches from fission products of ^{235}U , ^{238}U , ^{239}Pu and ^{241}Pu

Conversion method: Cumulative neutrino spectrum from measured beta spectrum

Summation method: Combine fission yields with decay data in databases

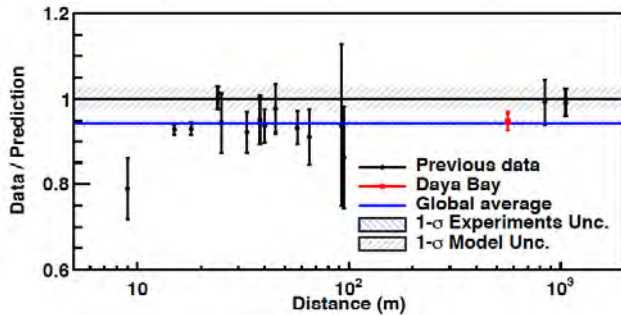
- discrepancies between databases
- decay schemes



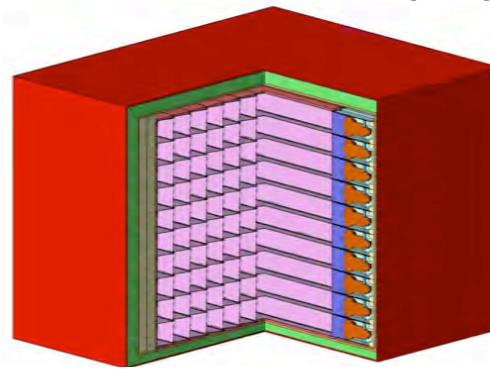
Short-Baseline Reactor Neutrino Experiments

Search for sterile neutrinos through neutrino oscillations

Test reactor anomaly



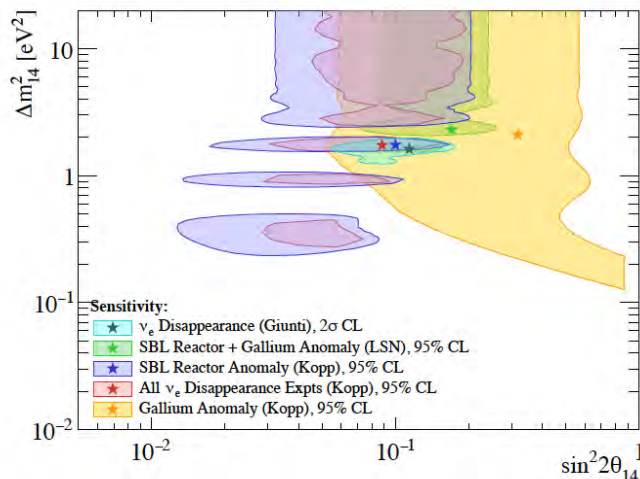
Measurement of the Relative Reactor Flux and Spectrum at Different Baselines independent of reactor models/predictions



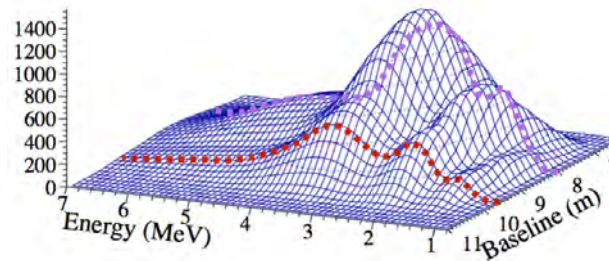
Segmented detector

Relative measurement within detector

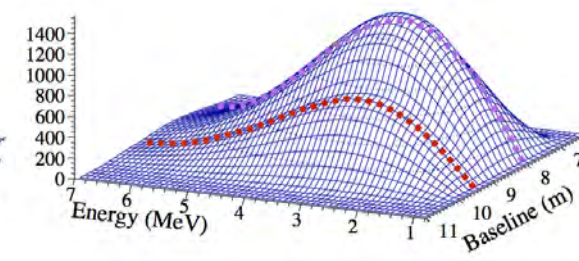
Test allowed oscillation parameter space



oscillated spectrum



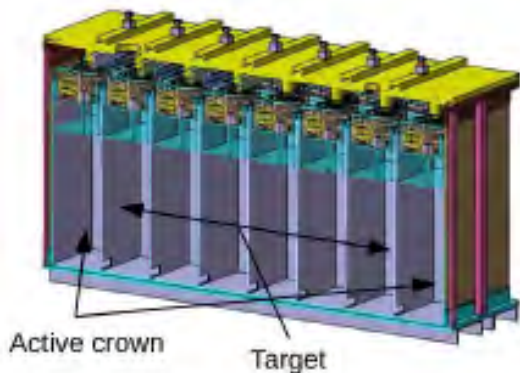
unoscillated spectrum



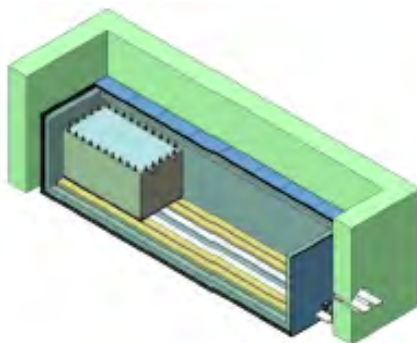
each segmented measures L/E

Short-Baseline Reactor Experiments Worldwide

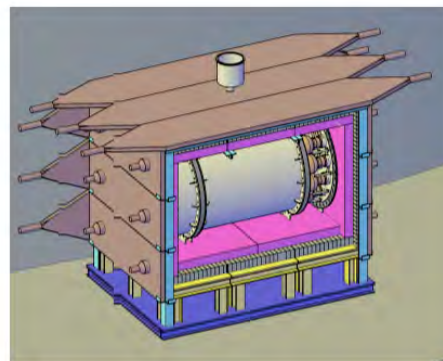
STEREO: Gd-LS detector at 10m from ILL, France



Neutrino-4: Gd-LS detector at 6-12m from SM-3, Russia



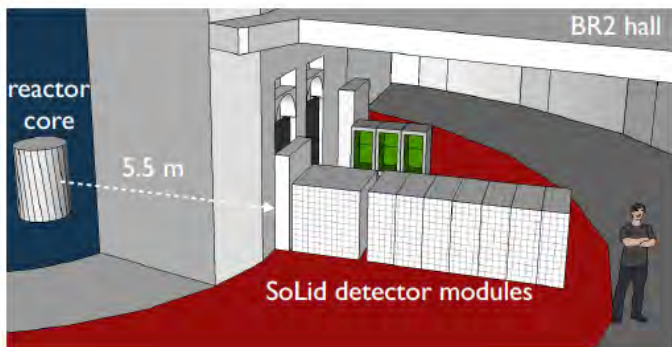
NEOS: Gd-LS detector at ~30m from Hanbit, Korea



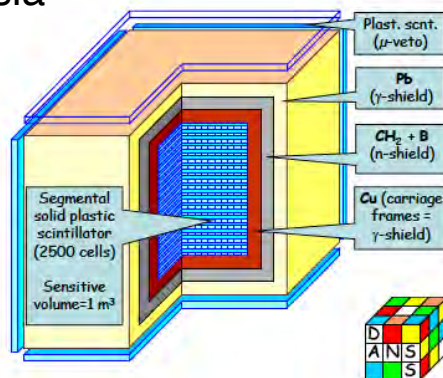
NuLAT: Boron-loaded plastic scintillator cubes



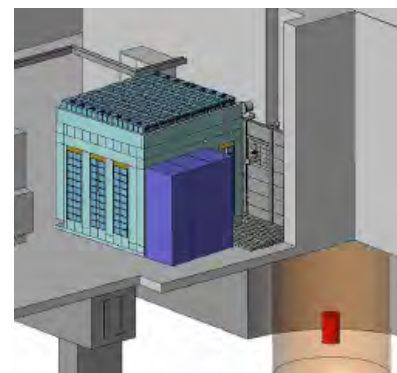
SoLid/CHANDLER: segmented composite scintillator cubes at 5.5m from BR2, Belgium



DANSS: Segmented plastic scintillator at ~10m from KNPP, Russia



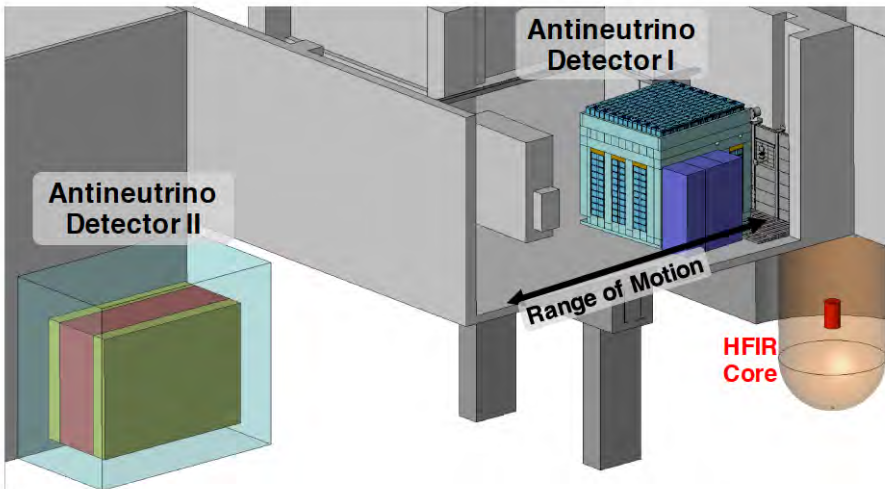
PROSPECT: Segmented 6Li liquid scintillator at 7-12m from HFIR, US



Precision Oscillation and Spectrum Experiment PROSPECT

Search for short-baseline oscillation at distances $<10\text{m}$
Precision measurement of ^{235}U reactor $\bar{\nu}_e$ spectrum

2 detectors, movable baseline, research reactor



Phase I

one movable detector AD-I, $\sim 7\text{-}12\text{ m}$ baseline

Phase II

two detectors,
movable AD-I, $\sim 7\text{-}12\text{m}$ baseline
stationary AD-II, $\sim 15\text{-}19\text{m}$ baseline

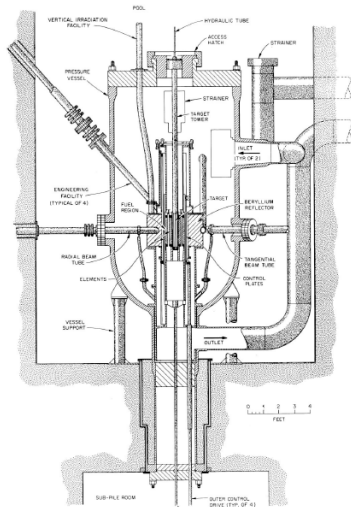
power: 85 MW (research)
fuel: highly enriched uranium (^{235}U)
core shape: cylindrical, compact
duty-cycle: 41%

physics program, [arXiv: 1512.02202](https://arxiv.org/abs/1512.02202)
test detector studies, JINST 10 P11004 (2015)
background measurements, NIM A806 (2016) 401
whitepaper, [arXiv: 1309.7647](https://arxiv.org/abs/1309.7647)

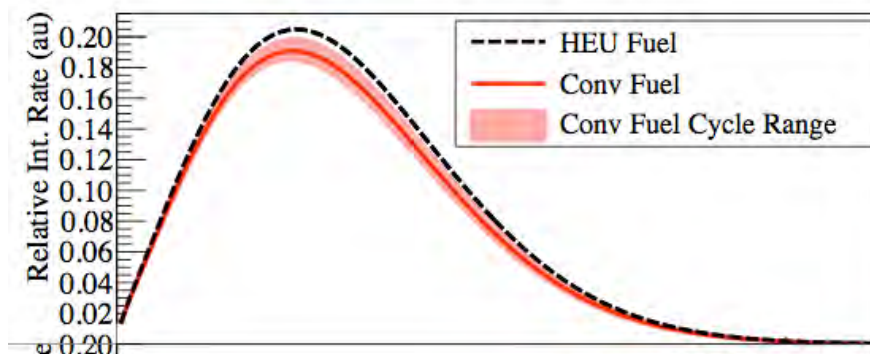
prospect.yale.edu

High Flux Isotope Reactor, Oak Ridge National Lab

US Research Reactor



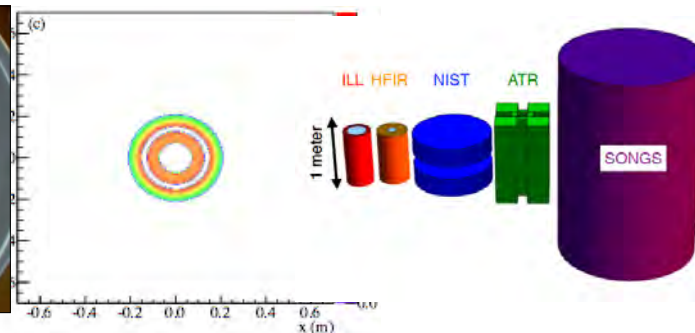
Research Reactor Spectrum



HEU core provides static spectrum of mainly ^{235}U .

power: 85 MW (research)
fuel: highly enriched uranium (^{235}U)
core shape: cylindrical
size: $h=0.5\text{m}$ $r=0.2\text{m}$ (compact)
duty-cycle: 41%

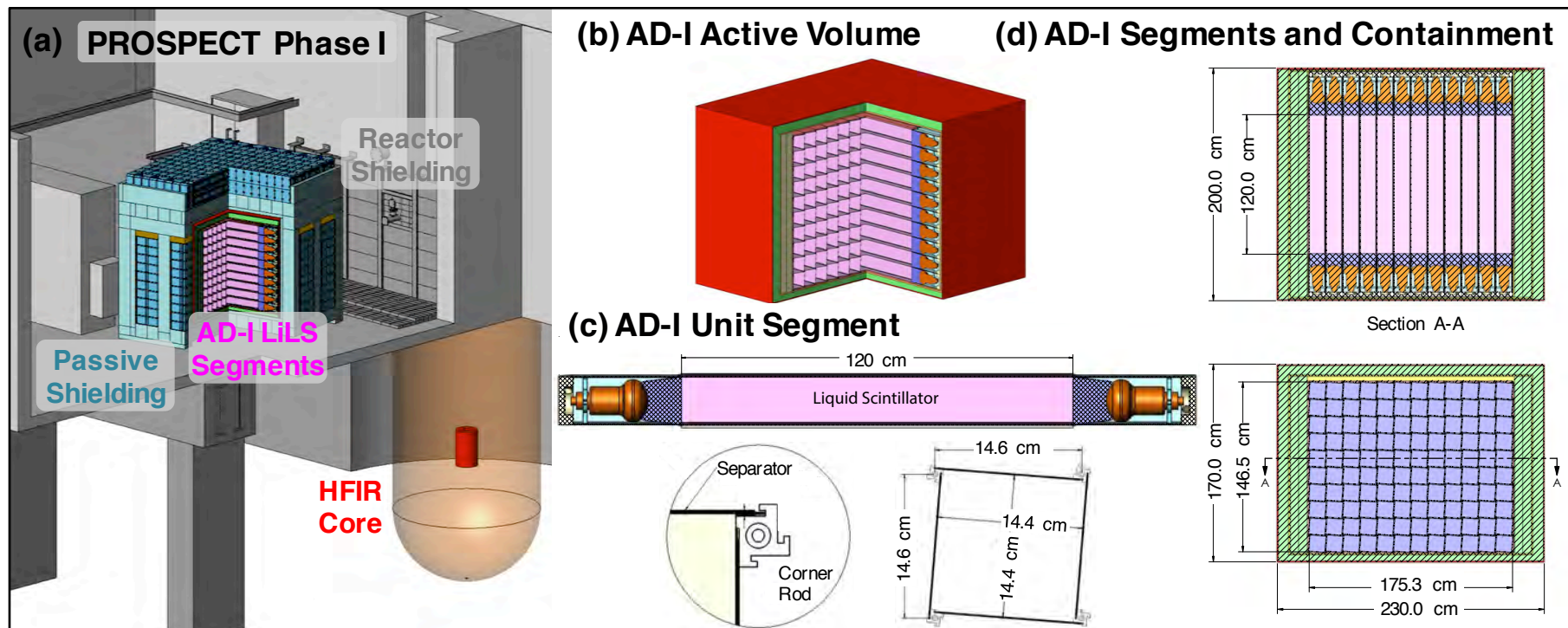
Compact reactor core



Nucl. Instrum. Meth. A806 (2016) 401–419,
 arXiv:1506.03547,
 PROSPECT collaboration

Compact core ($< 1\text{m}$) avoids oscillation washout

Antineutrino Detector



- 3000L of ${}^6\text{Li}$ liquid scintillator
- 120 scintillator loaded cells, $\sim 15 \times 15 \times 120 \text{ cm}$
- double ended PMT readout, light guides, $< 4\text{-}5\% / \sqrt{E}$ resolutions
- thin optical separators, minimal dead material
- containment vessel, filled in place

PROSPECT Physics

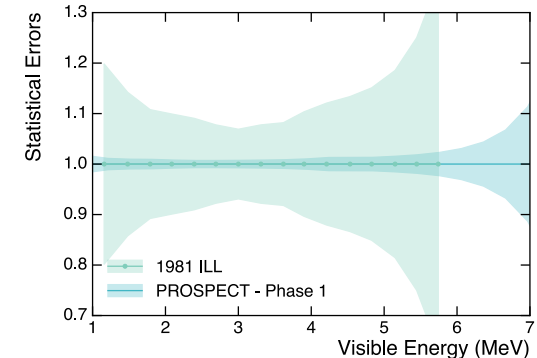
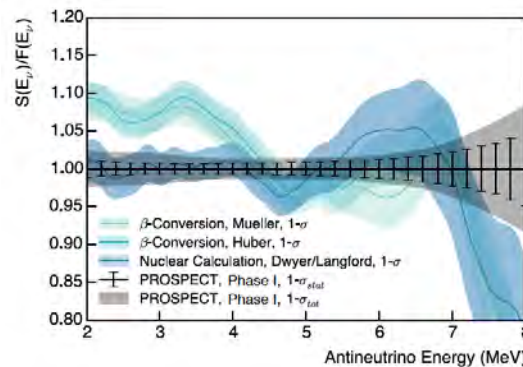
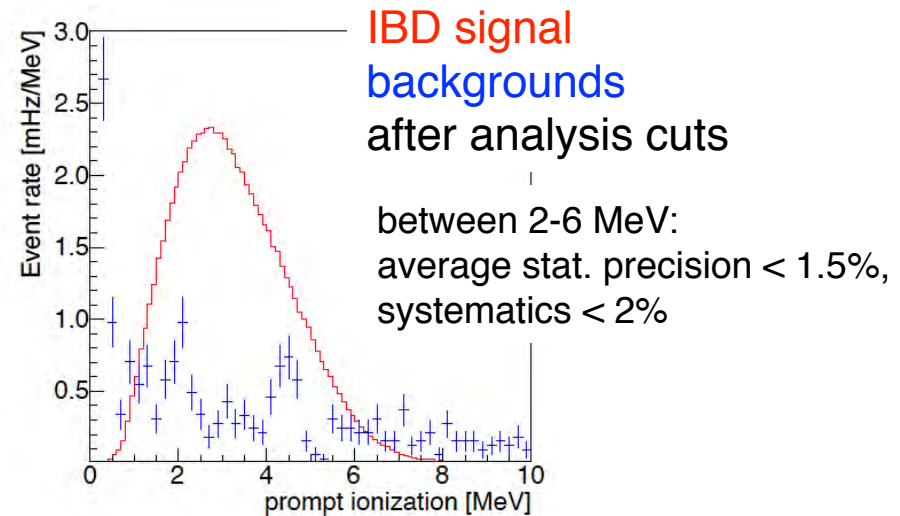
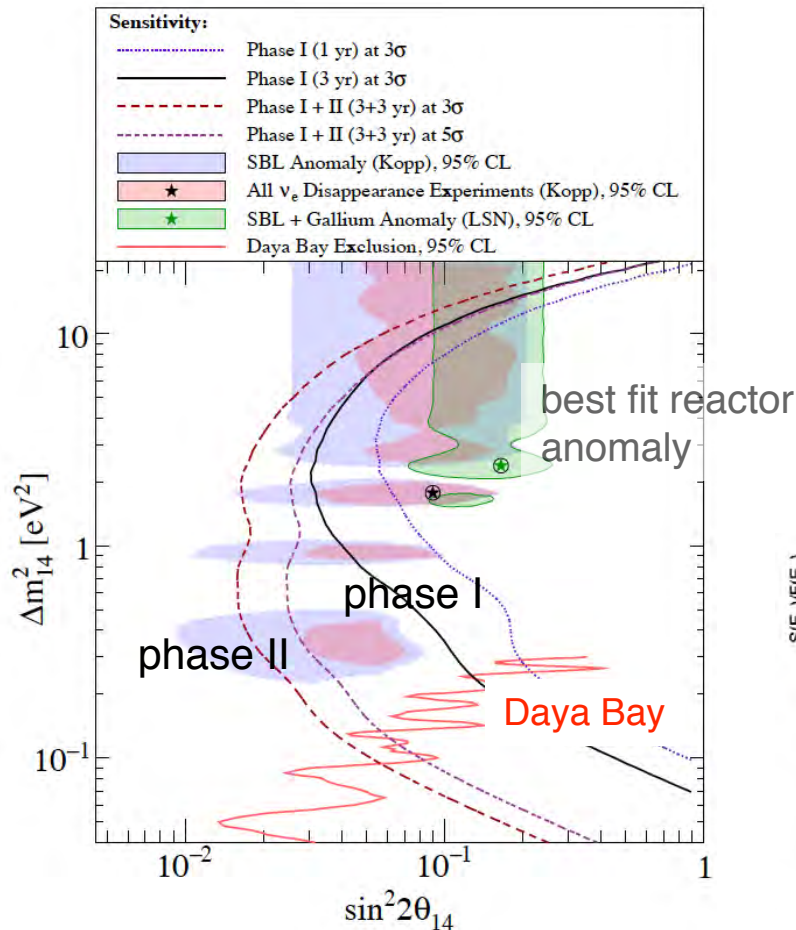


A Precision Oscillation Experiment

- 4 σ test of best fit after 1 year
- >3 σ test of favored region after 3 years
- 5 σ test of allowed region after 3+3 years

A Precision Spectrum Experiment

- Measurement of ^{235}U spectrum
- Compare different reactor models
- Opportunity to compare different reactor cores



PROSPECT Detector and Shielding Development



PROSPECT-0.1

Characterize LS

Aug 2014-Spring 2015

5cm length

0.1 liters

LS, ${}^6\text{LiLS}$



PROSPECT-2

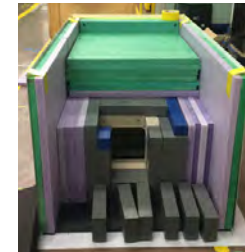
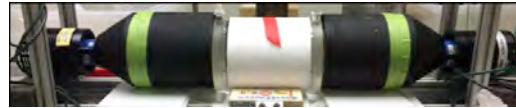
Background studies

Dec 2014 - Aug 2015

12.5 cm length

1.7 liters

${}^6\text{LiLS}$



multi-layer shielding

PROSPECT-20

Segment characterization

Scintillator studies

Background studies

Spring/Summer 2015

1m length

23 liters

LS, ${}^6\text{LiLS}$



PROSPECT-50

Baseline design prototype

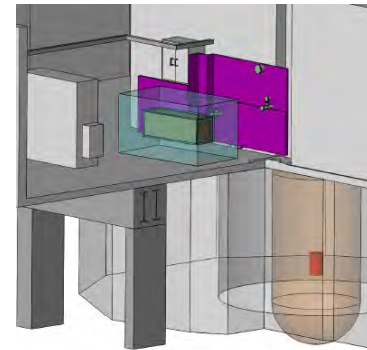
Winter 2015

1x2 segments

1.2m length

50 liters

${}^6\text{LiLS}$



local reactor shielding

PROSPECT-400*

Fiducialization and background studies

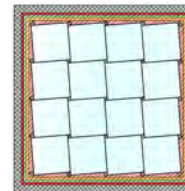
Mid 2016

4x4 segments

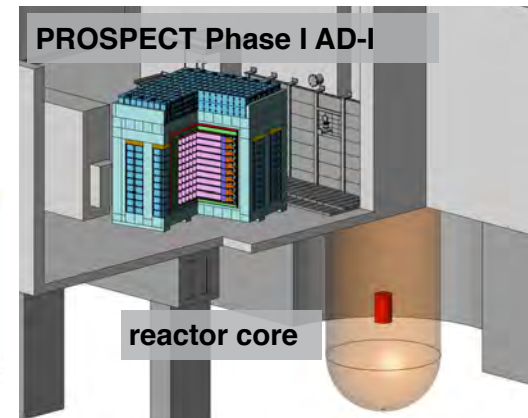
1.2m length

400 liters

${}^6\text{LiLS}$



PROSPECT Phase I AD-I



reactor core

PROSPECT AD-I

Physics measurement

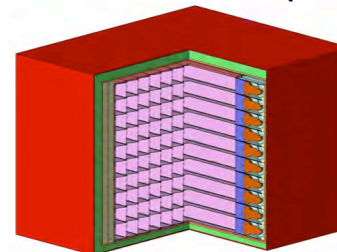
Late 2016

10x12 segments

1.2m length

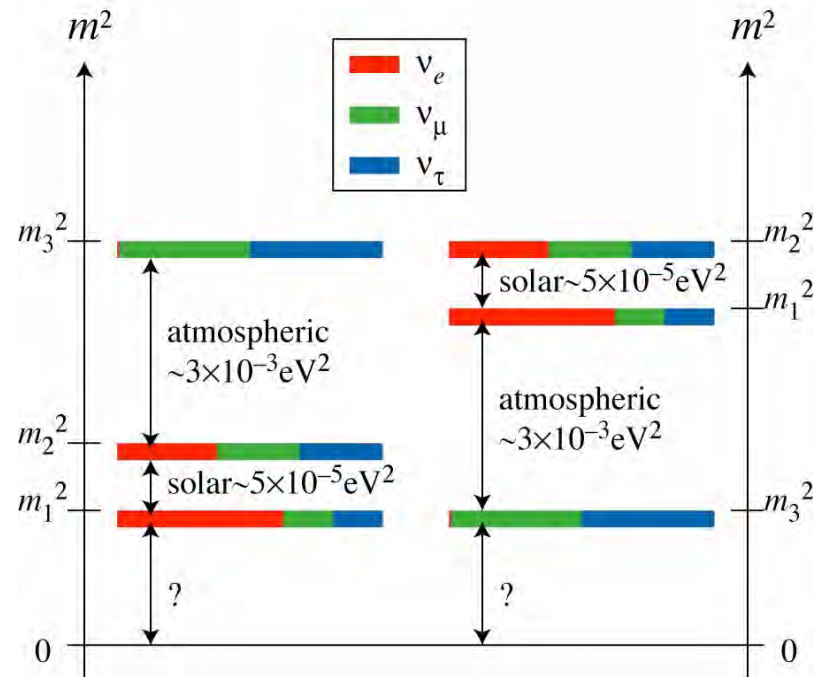
~3 tons

${}^6\text{LiLS}$



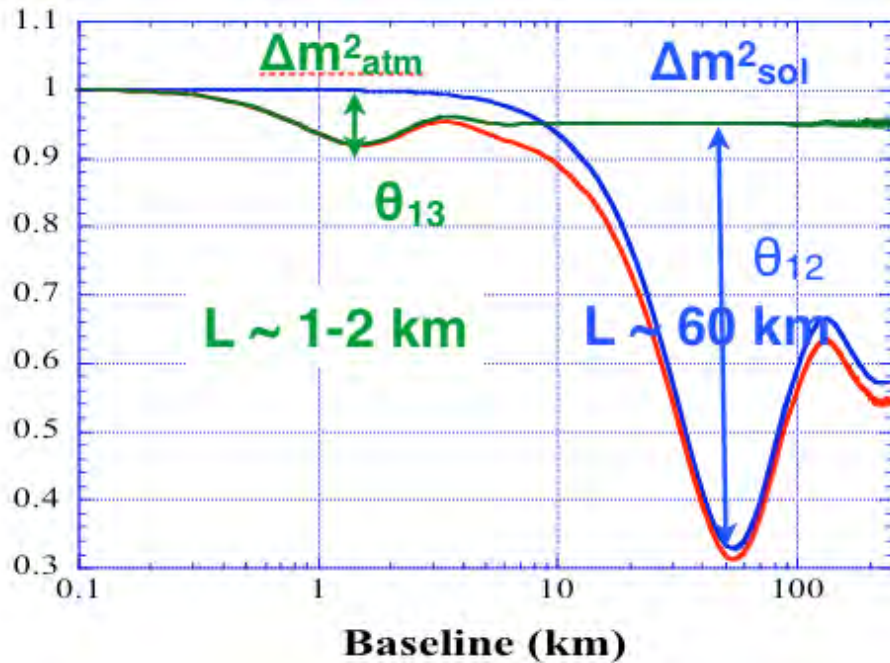
*Technically ready to proceed directly to AD-1 with available funding

Mass Hierarchy?



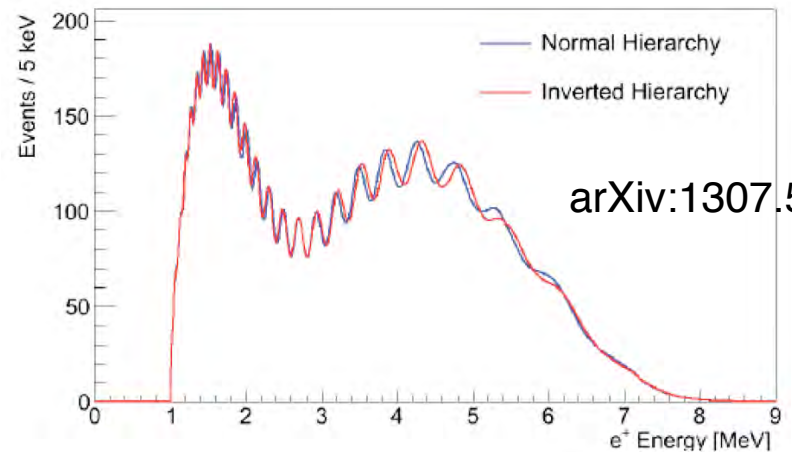
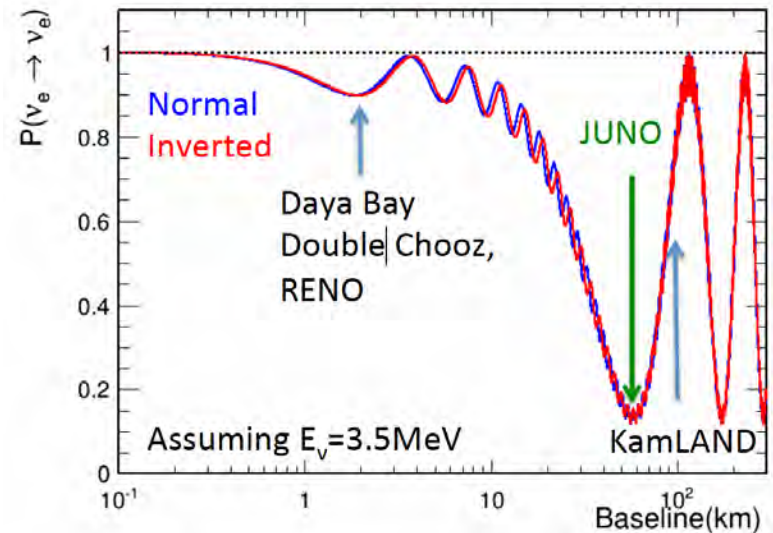
Mass Hierarchy and Reactor Neutrinos

Precision Measurement at ~ 58km



$$P_{\bar{\nu}_e \rightarrow \bar{\nu}_e} = 1 - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \Delta_{21} - \sin^2 2\theta_{13} (\cos^2 \theta_{12} \sin^2 \Delta_{31} + \sin^2 \theta_{12} \sin^2 \Delta_{32})$$

determine mass hierarchy from precision measurements of $|\Delta m^2_{31}|$ and $|\Delta m^2_{32}|$

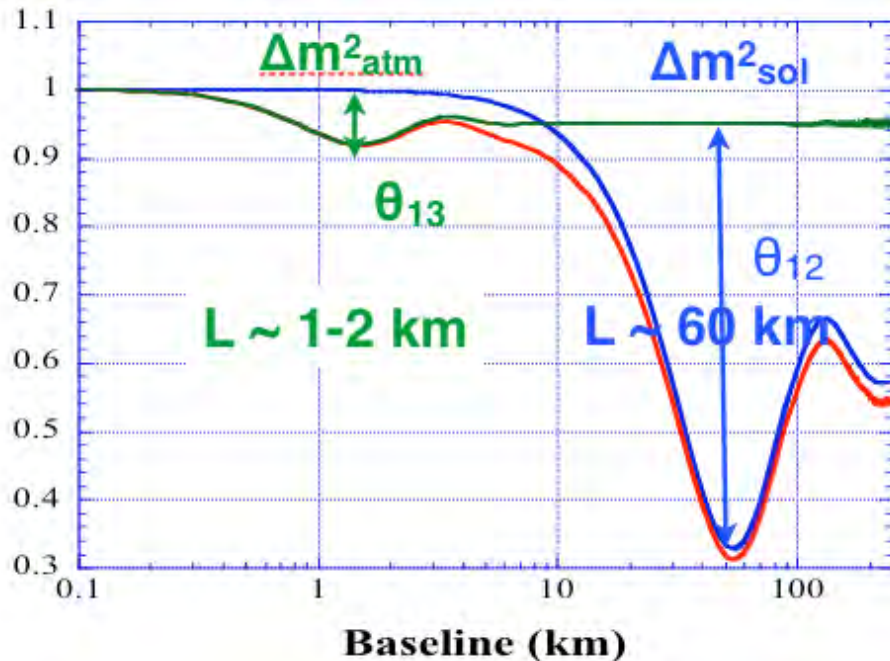


arXiv:1307.5487

mass hierarchy is contained in the spectrum independent of the unknown CP phase

Mass Hierarchy and Reactor Neutrinos

Precision Measurement at ~ 58km

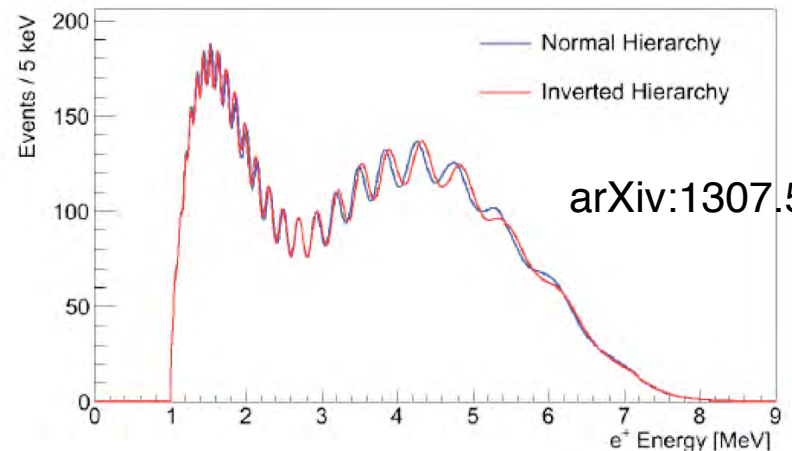
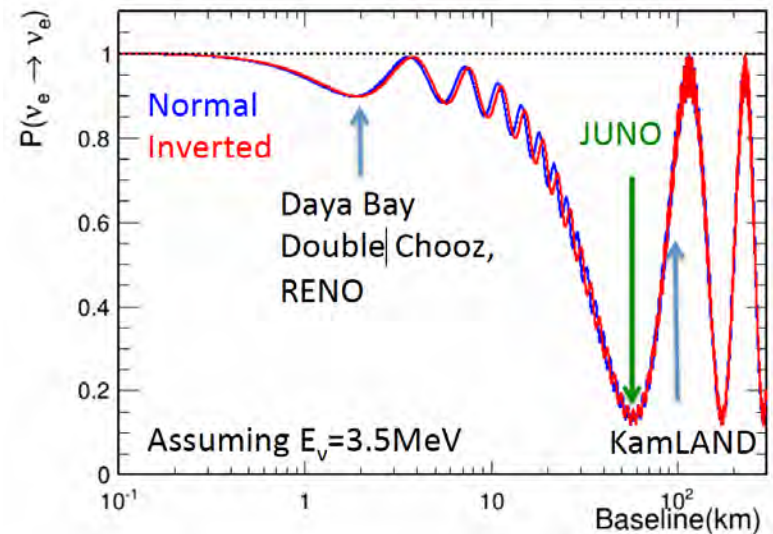


$$\Delta m_{31}^2 = \Delta m_{32}^2 + \Delta m_{21}^2$$

$$\text{NH: } |\Delta m_{31}^2| = |\Delta m_{32}^2| + |\Delta m_{21}^2|$$

$$\text{IH: } |\Delta m_{31}^2| = |\Delta m_{32}^2| - |\Delta m_{21}^2|$$

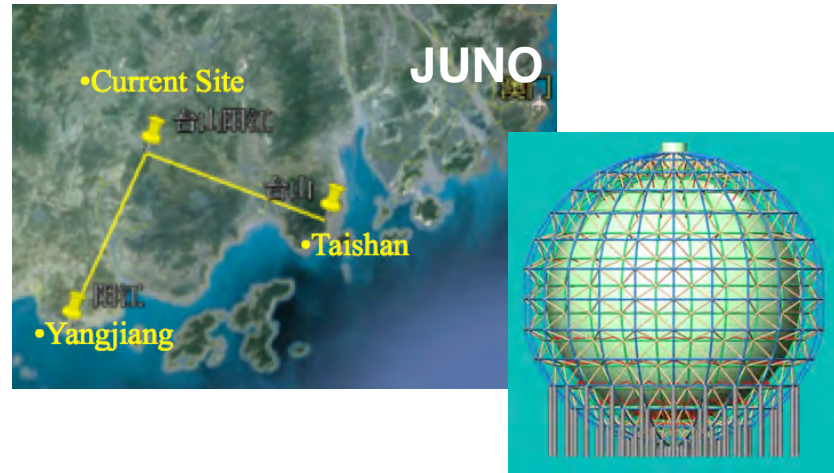
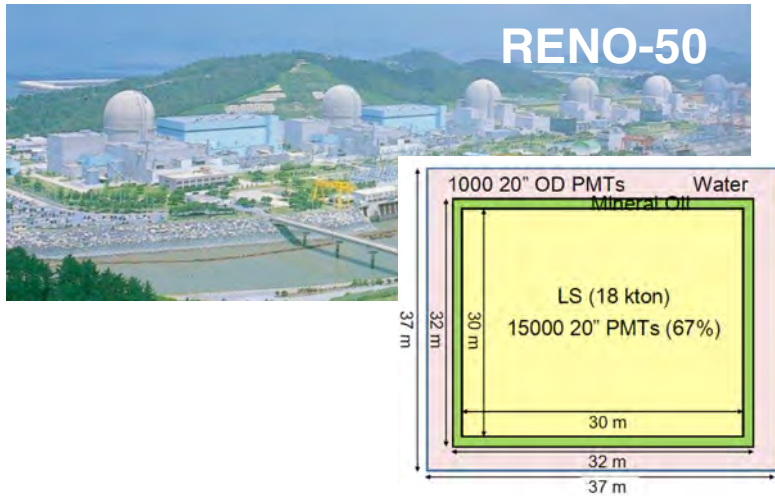
determine mass hierarchy from precision measurements of $|\Delta m_{31}^2|$ and $|\Delta m_{32}^2|$



mass hierarchy is contained in the spectrum independent of the unknown CP phase

Mass Hierarchy and Reactor Neutrinos

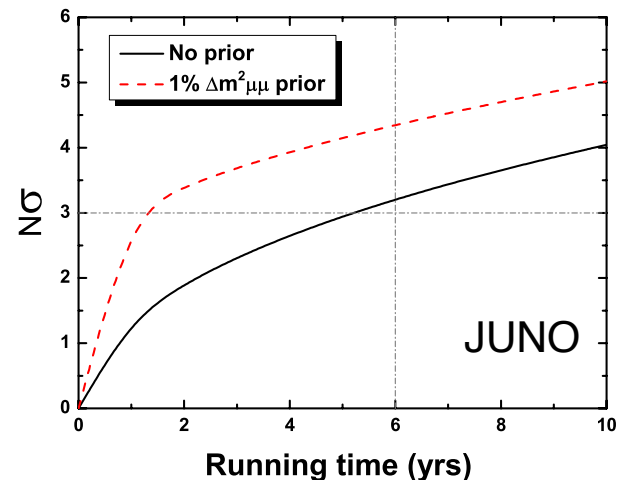
Proposed Projects: JUNO and RENO-50



Precision 3-ν Oscillation Physics

	Current	JUNO
Δm^2_{12}	3%	0.6%
Δm^2_{23}	5%	0.6%
$\sin^2\theta_{12}$	6%	0.7%
$\sin^2\theta_{23}$	20%	N/A
$\sin^2\theta_{13}$	10% (~4% in 3 yrs)	15%

Mass Hierarchy Sensitivity



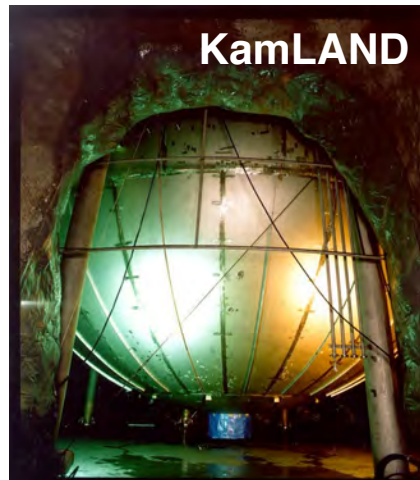
Reactor Antineutrinos in History

A Tool for Discovery



2012 - Measurement of θ_{13} with Reactor Neutrinos

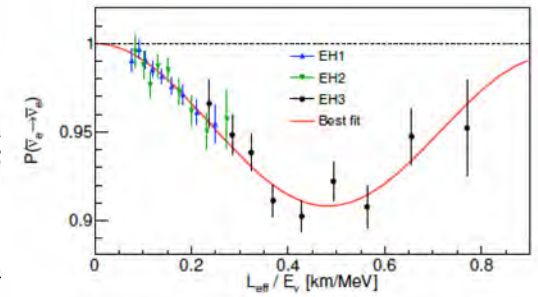
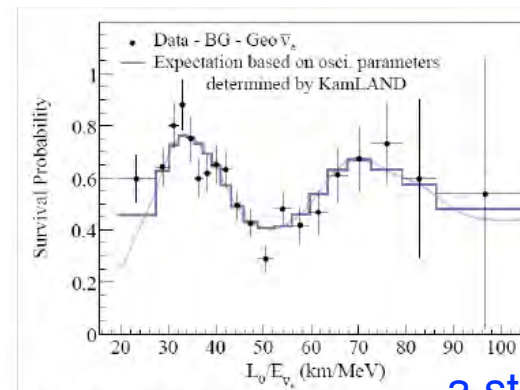
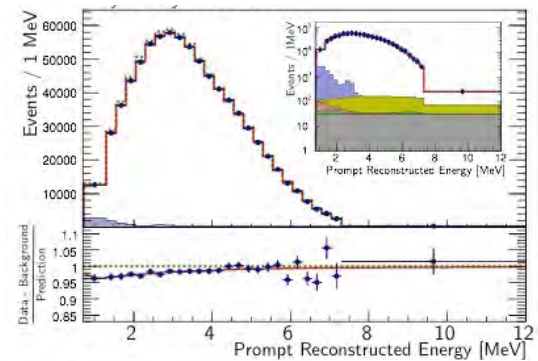
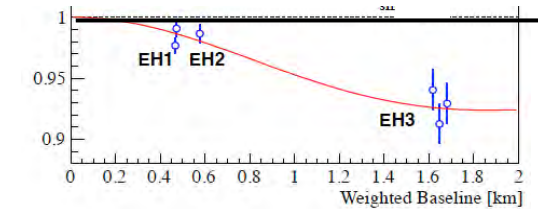
2003 - First observation of reactor antineutrino disappearance



1995 - Nobel Prize to Fred Reines at UC Irvine



1956 - First observation of (anti)neutrinos



a story of varying baselines... 36

Summary & Outlook



Current reactor experiments (**L~1-2km**) provide precision data on θ_{13} , and **reactor antineutrino flux and spectra**, and **complementary limits on sterile neutrinos**. Flux measurement is consistent with previous short-baseline measurements (~6% deficit). Positron spectrum appears inconsistent with current predictions in 4-6 MeV region.

Short-baseline (**L~10m**) experiments (e.g. PROSPECT) offer opportunities for **precision studies of reactor spectrum** and a **definitive search for short-baseline oscillation** and **sterile neutrinos**.

Reactor experiments may inform nuclear modeling of reactors. Detectors may find applications in reactor monitoring.

Medium-baseline experiments (**L~60km**) (e.g JUNO, RENO-50) are technically demanding but may offer <1% **precision oscillation physics** and a **window to the mass hierarchy**.

*After 60 years of reactor neutrino experiments, future is bright.
Active field with ongoing and planned experiments.*