Micromegas Based Muon Telescopes for The ScanPyramids Mission - Performance & First Results -

D. Attié, S. Bouteille, D. Calvet, P. Magnier, I. Mandjavidze, S. Procureur, M. Riallot



8th SYMPOSIUM ON LARGE TPCs FOR LOW-ENERGY RARE EVENT DETECTION



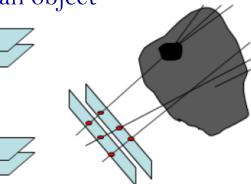
- Muon tomography
- Genetic Multiplexed Resistive Bulk Micromegas
- Demonstrator for a Water Tower (WatTo) imagery
- WatTo to ScanPyramids
- Installation at the Khufu Pyramid
- Results
- Conclusions

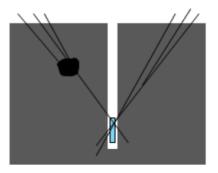


- Muons from cosmic rays interacts with matter and can be either **deviated** by Coulomb scattering or **absorbed**
- Deviation and absorption depend on
 - material length
 - density

Material	Thickness	θ (deg)	P _{absorption}
Air	100 m	0.094	0.78%
Lead	1 cm	0.28	0.49%
Lead	10 cm	1.01	2.9%
Water	1 m	0.35	4.2%
Limestone	1 m	0.74	9.5%
Limestone	100 m		99%
Limestone	1 km		99.999%

• can be used to probe an object

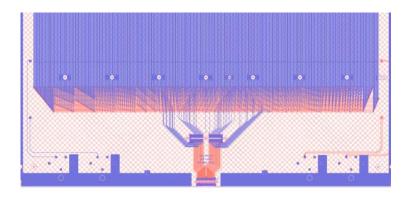


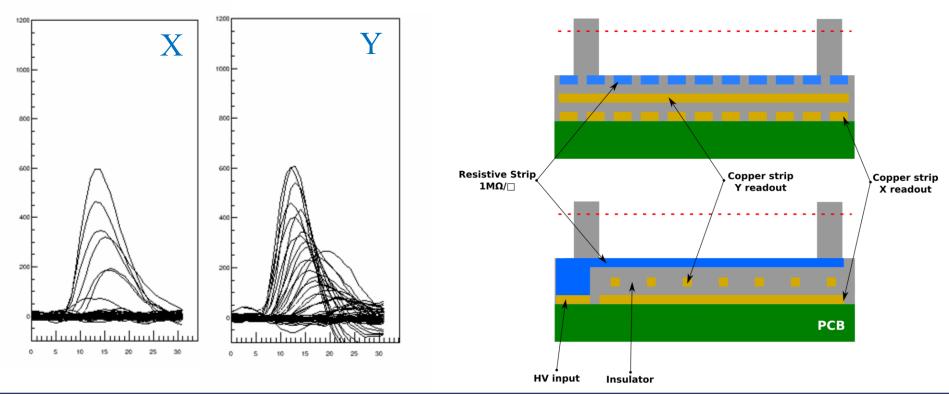


- Many applications in the world: homeland security, volcanology, archeology, mining, ...
- In France: Diaphane, TomuVol, Geo-Azur
- Why another one? Micromegas offer spatial resolution 10-20 times better

Genetic Multiplexed Resistive Bulk Micromegas

- Bulk Micromegas technology (1996, 2005)
- Genetic multiplexing readout (2012)
- Resistive strips X & Y (2010, 2012)
 - capacitive induction
 - spread signal over several strips

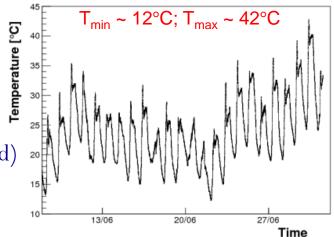


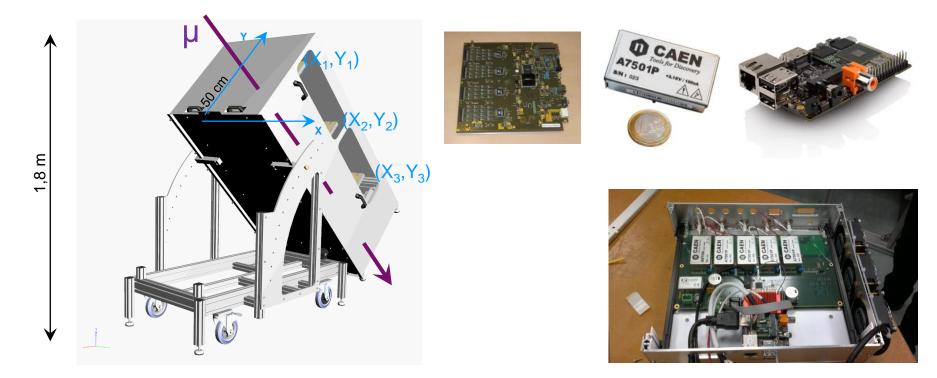


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WatTo (Water Tower): Micromegas Muon Telescope

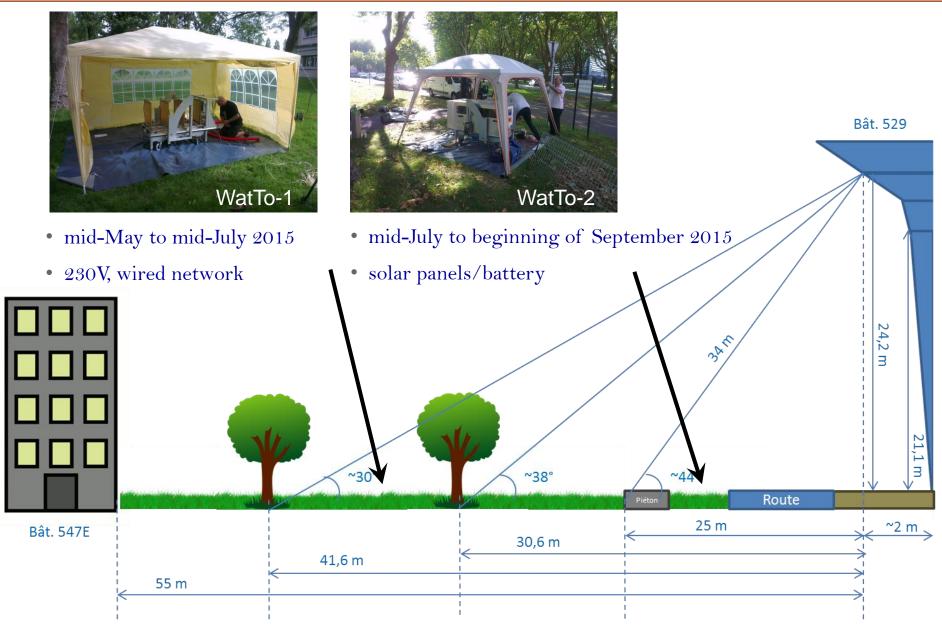
- 1st muon telescope with multiplexed Micromegas
 - 1st use of Micromegas outside: gain stability f(T,P)
 - Self trigger mode of Particle Physics electronics
 - Compact electronic system (nano-PC, specific HV card)
 - 30 W power consumption







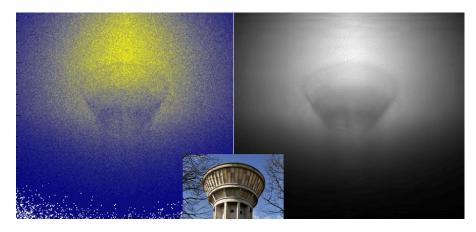
WatTo Configurations



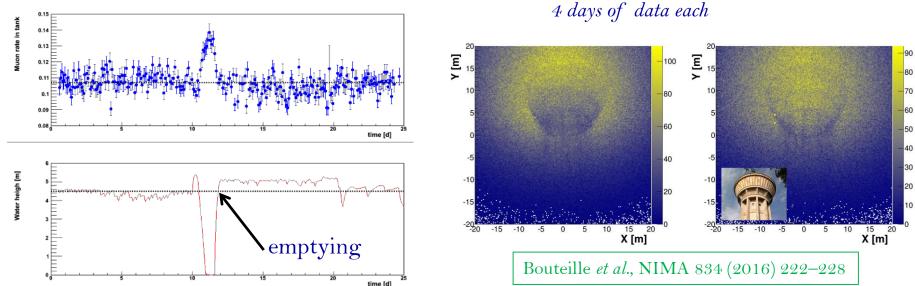


WatTo Results

• WatTo-1: 4 weeks of data

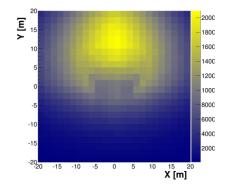


• WatTo-2 (solar boards + dynamics)



Validation of the most precise muon telescope in the world

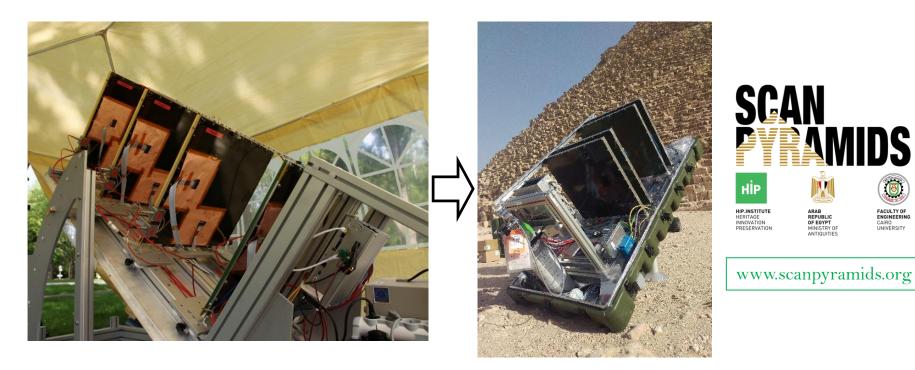
4 weeks of data Scintillators resolution @ same compactness



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- Telescopes: $1 \rightarrow 3$
- Mechanical structure \rightarrow case
- Detectors: prototype (Cern) \rightarrow final version (built by PCB's company Elvia)
- Construction time: 9 monts \rightarrow 3 months
- Weight: ~ 200 kg \rightarrow ~ 130 kg
- Detector power supply: no temperature feedback \rightarrow f(T)
- Data: raw \rightarrow raw + pre-processed

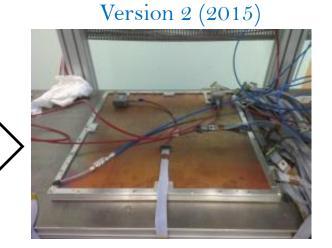




Detection Plane Evolution

Prototype (2014)





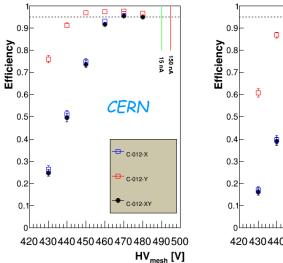
Bouteille et al., NIMA 834 (2016) 187-191

- Better shielding

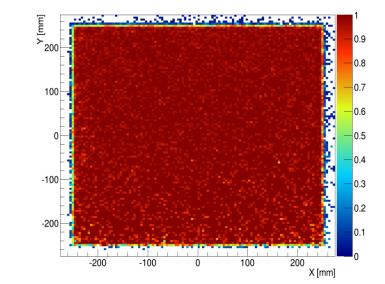
- 13 strips added
- Decreased dead space
- Mechanics improved

• Detector performance

Efficiency vs. HV



2D Efficiency



david.attie@cea.fr

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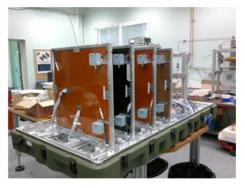


• 3 telescopes installed in front of the Khufu (Cheops) Pyramid end of May 2016







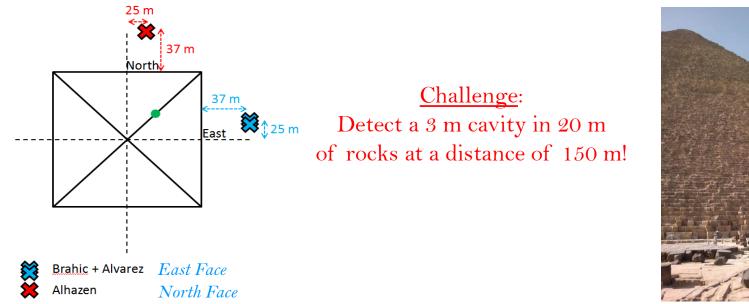


Alhazen (n°1)

Alvarez (n°2)

Brahic (n°3)

• First campaign: performance validation looking to a known cavity



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Installation at Khufu Pyramid



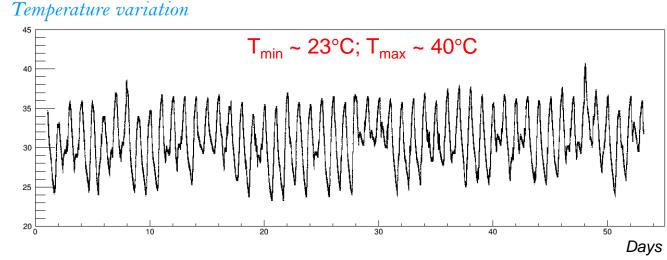
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Data Taking

• Each telescope operate during 2 to 3 months (gas autonomy)



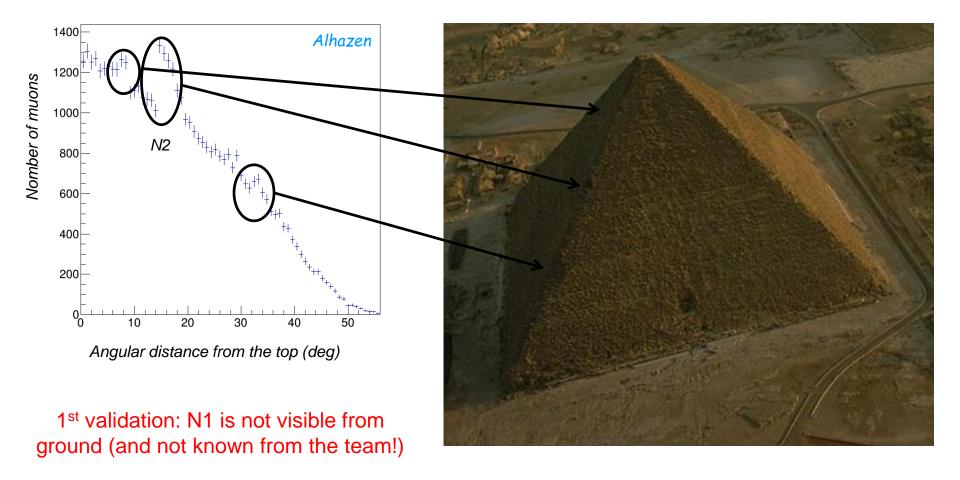
- Stable data taking, thanks to local team (3G, gaz, maintenance)
- Excellent spatial resolution $\rightarrow \leq 1$ m accuracy at a distance of 150 m
- Number of recorded events:
 - Alhazen (Nord): 30.8 millions of triggers (4.5 Hz)
 - Brahic (Est): 24.6 millions (4.2 Hz)
 - Alvarez (Est): 18.7 millions (3.3 Hz)
- Issue with *Alvarez* (July 1st, alignment change)
 - \rightarrow Investigation in process

david.attie@cea.fr

~70% are can be reconstructed



• First slices show notches along the edge

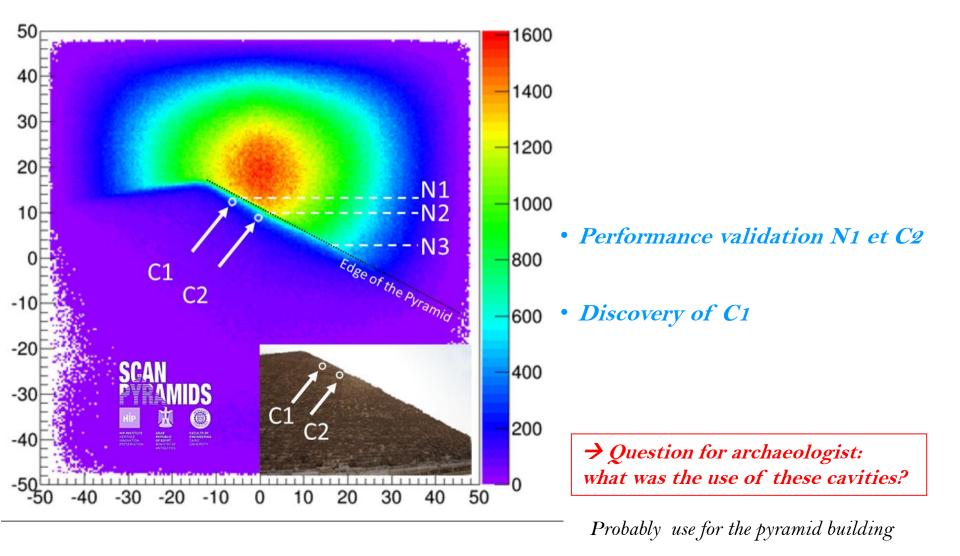


• *Brahic* & *Alvarez* see the same pattern with 3 notches

david.attie@cea.fr



• Press released: October 15th, 2016





- The first ScanPyramids campaign was a success
- Discovery of a new cavity (C1)
- The 3 telescopes are still in Cairo waiting for redeployment
- Next steps:
 - look at other edges
 - scan other Pyramid edges (Khafre, etc.)
 - data analysis is in continuing (add reconstructed events)
- Contact in progress with several private companies (RATP, etc.) for societal applications, but first need simulations or in situ demonstration

