



Single Top & Dark Matter

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LPTHE - CNRS - UPMC

3rd CMS Single Top Workshop

Strasbourg - 02-03 June 2016

Heavy rains in central France region...



Targeting dark matter and top quarks at the LHC

 There is strong (indirect Cosmic microwave back Baryonic acoustic oscillational anomalies Gravitational anomalies etc. 	5
Underground nuclear re	ct dark matter and measure its properties ecoil experiments (DM-nucleon scattering) n in the galaxy, gamma rays missing energy
The top mass is close to	believed to be a sensitive probe for new physics o the electroweak scale ary for stabilizing the Higgs mass Signals of models with a top quark connected to dark matter?





Simplified models for top-philic dark matter

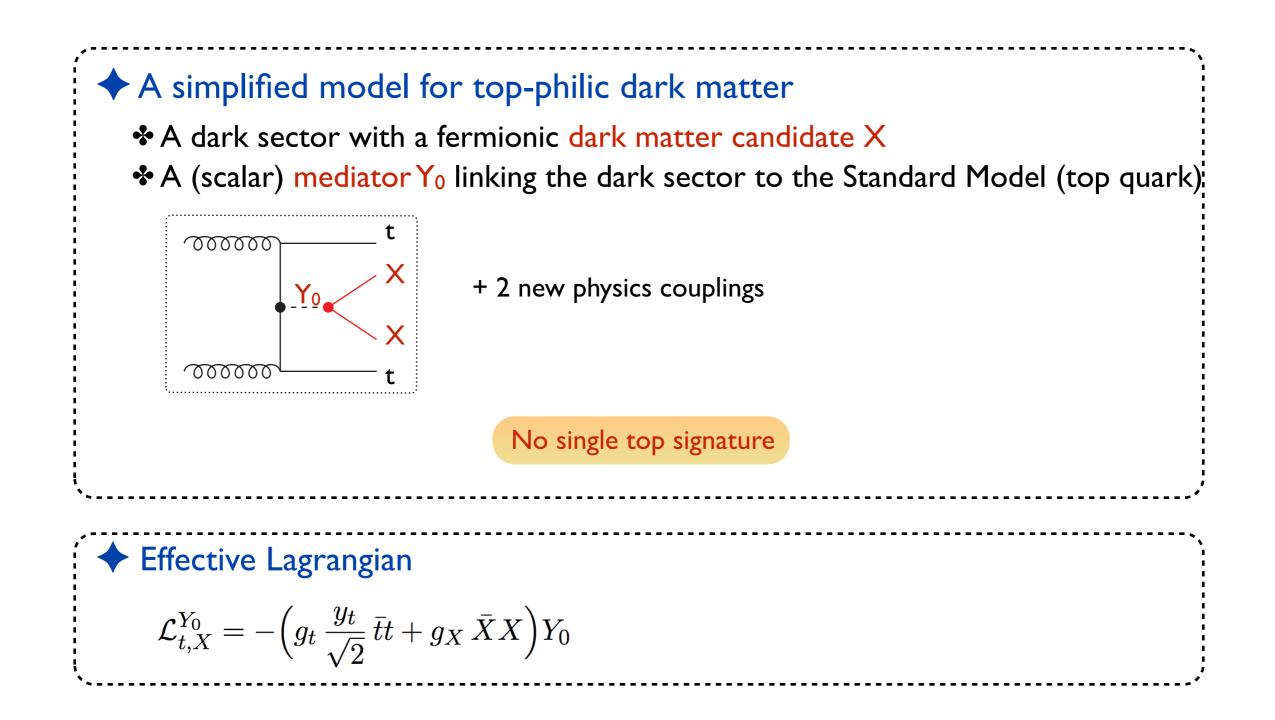


2. Monotop models and their constraints

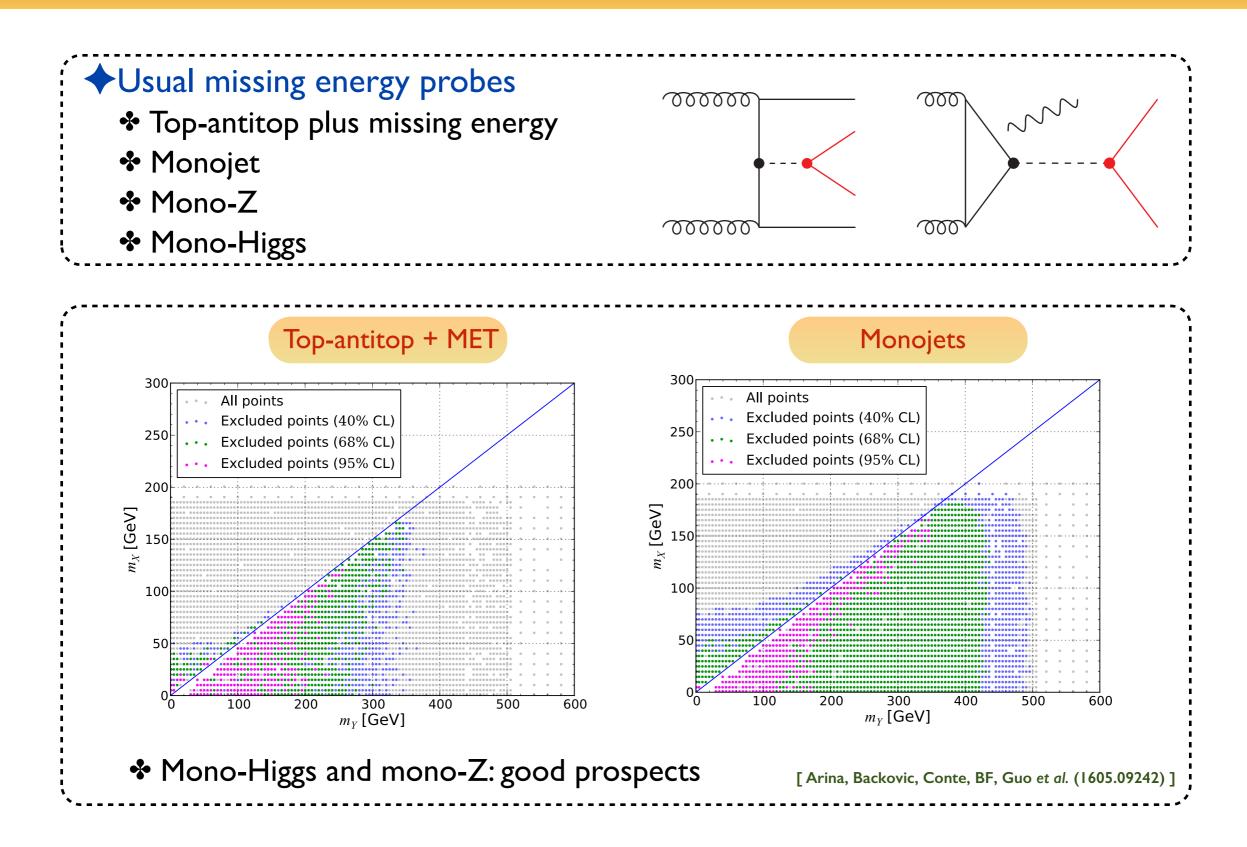


Effects of next-to-leading-order predictions

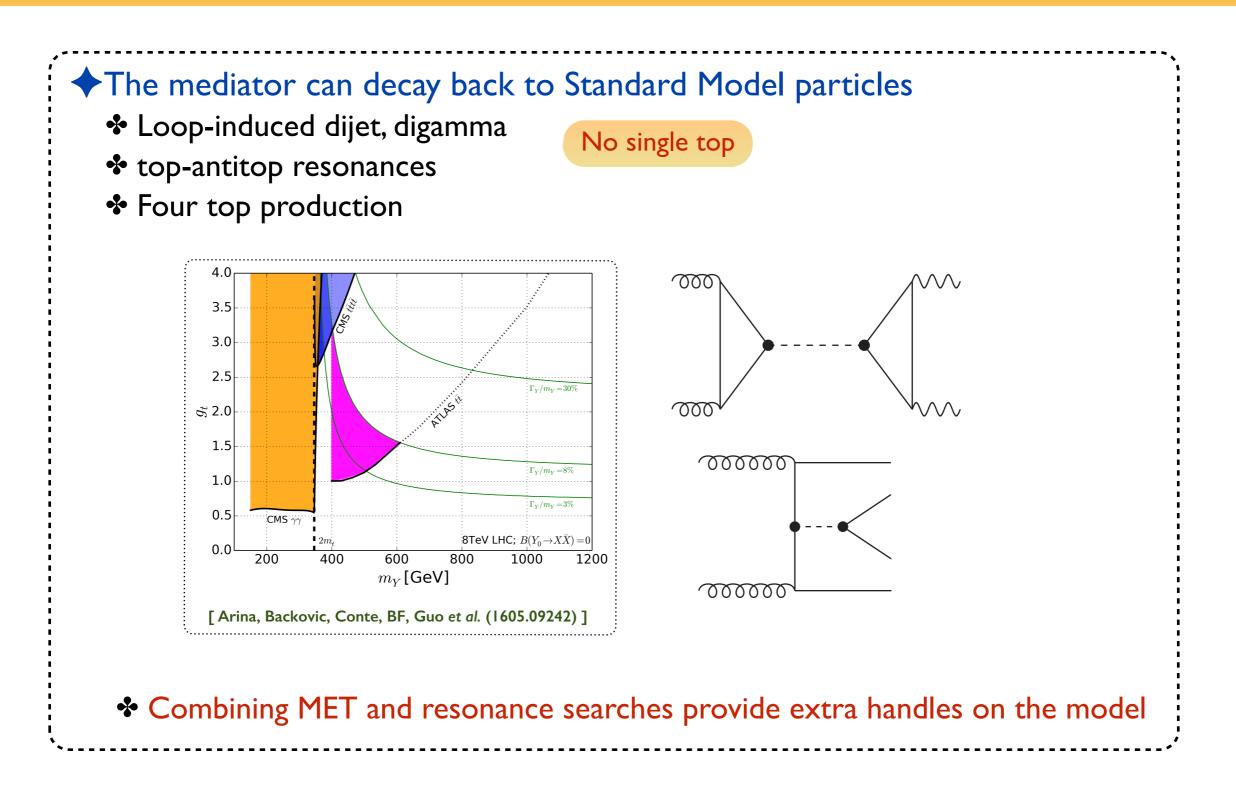
Top-philic dark matter



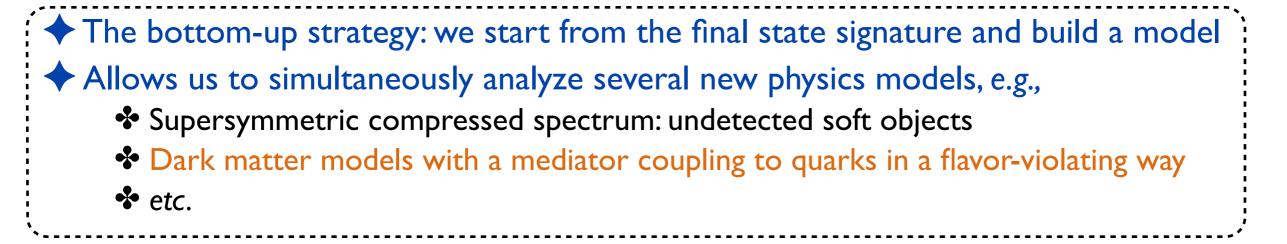
Probing top philic dark matter at the LHC



Non-MET searches are important

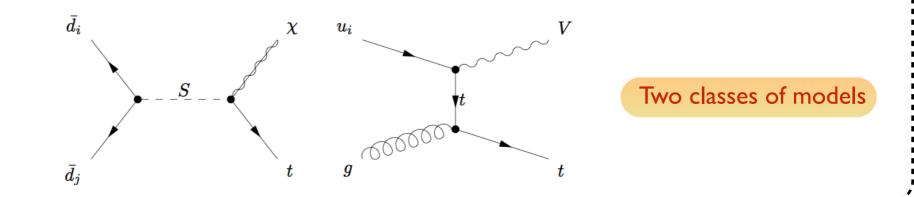


Monotops at hadron colliders: the general case



Generic monotop production

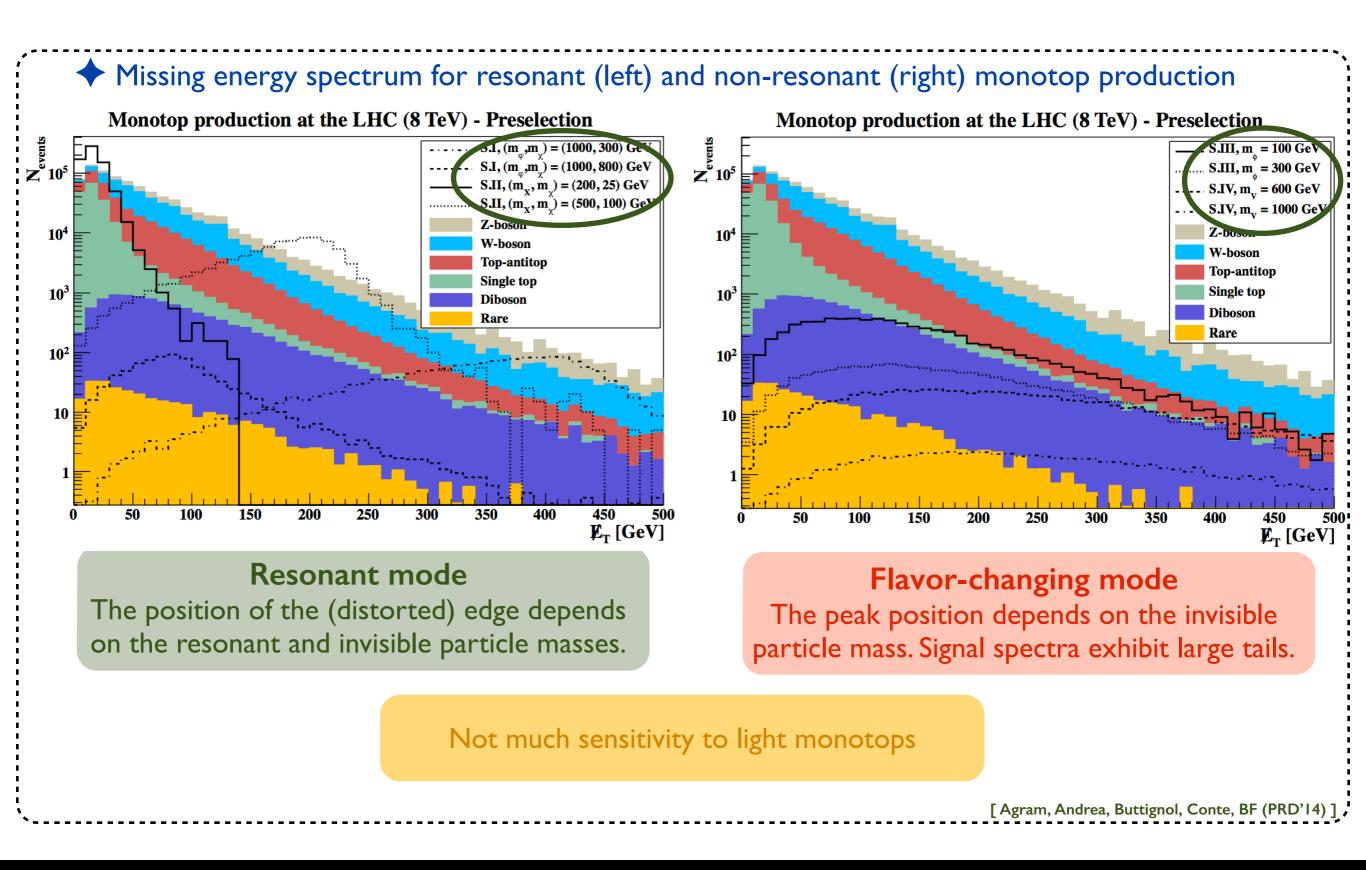
- Missing energy (dark matter candidate or mediator decaying to dark matter particles)
 - \star Bosonic or fermonic state
 - \star One-particle or n-particle state
 - * Neutral, weakly-interacting, long-lived/stable/invisible
- Initial state: two possibilities
 - * A down-type (anti)quark pair \rightarrow baryon-number-violating process
 - \star An up-type quark / gluon associated pair \rightarrow flavor-changing neutral interactions



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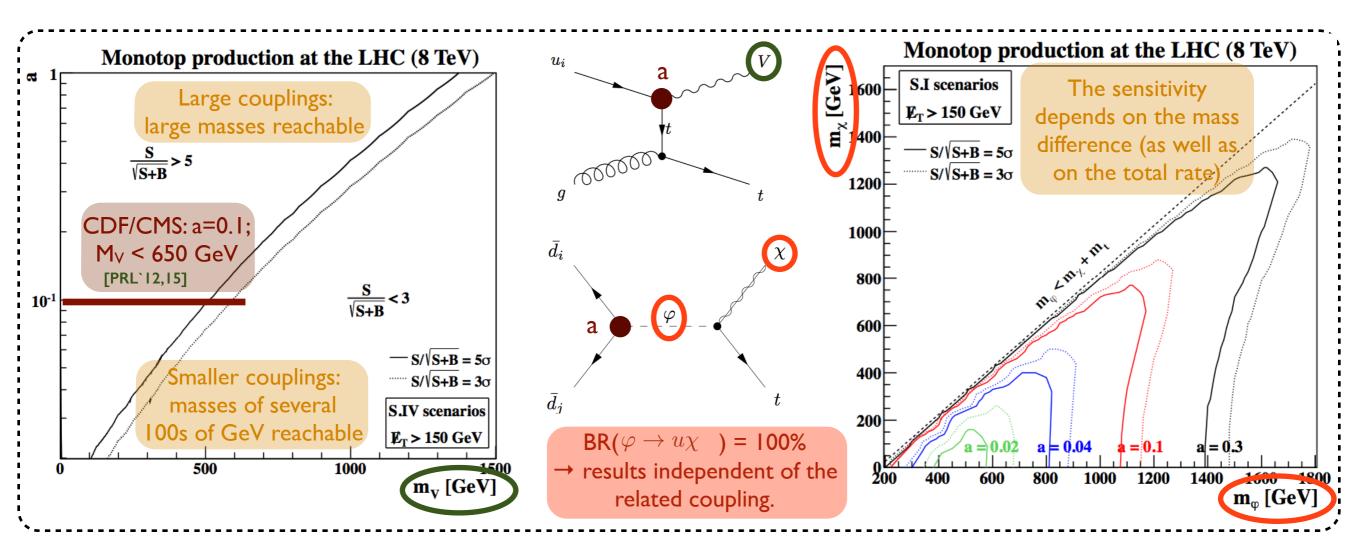
[Andrea, BF, Maltoni (PRD '11)]

Monotops with 20 fb⁻¹ of 8 TeV LHC data (1)



Monotops with 20 fb⁻¹ of 8 TeV LHC data (2)

[Agram, Andrea, Buttignol, Conte, BF (PRD'14)]



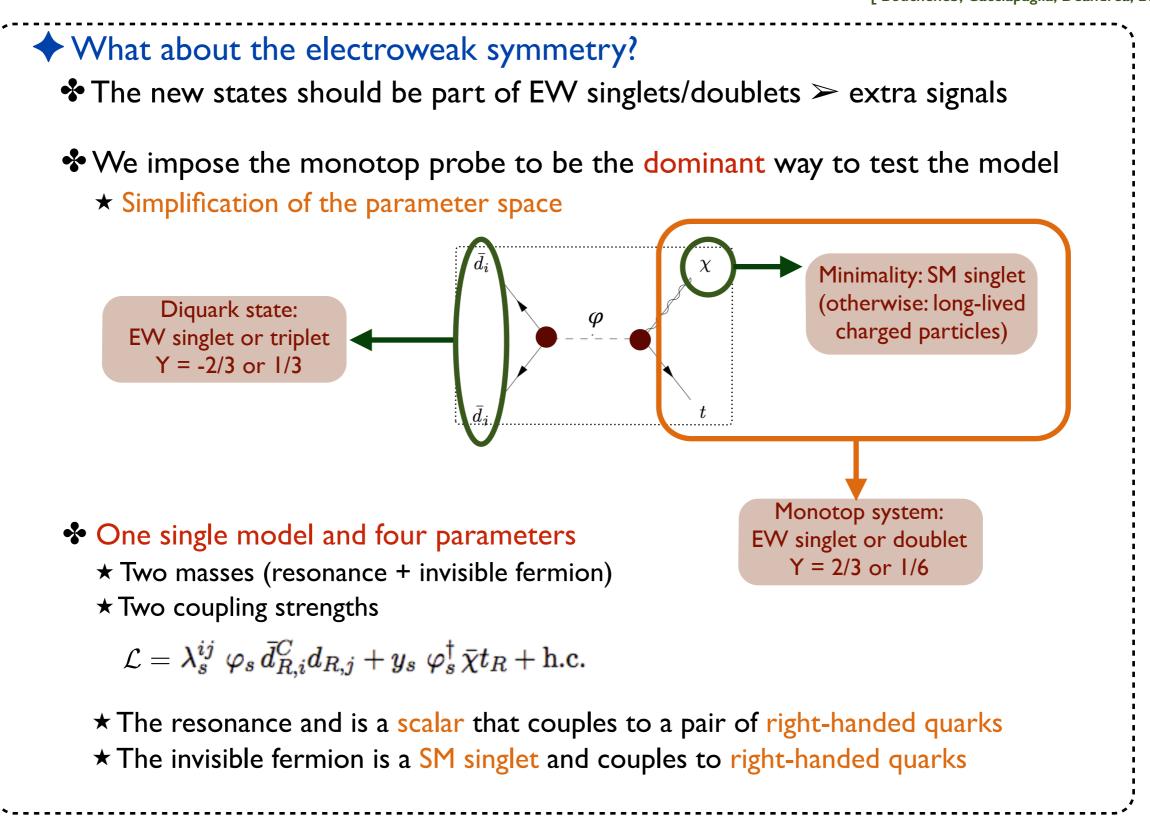
Hadronic monotop selection (similar results for the leptonic case)

- Lepton veto
- ✤ 2 light jets and I b-tagged jet
- ✤ M_{jj} and M_W compatible
- \clubsuit M_{bjj} and M_{t} compatible
- Non-collinear top and missing energy

The LHC can largely access the monotop mot general parameter space

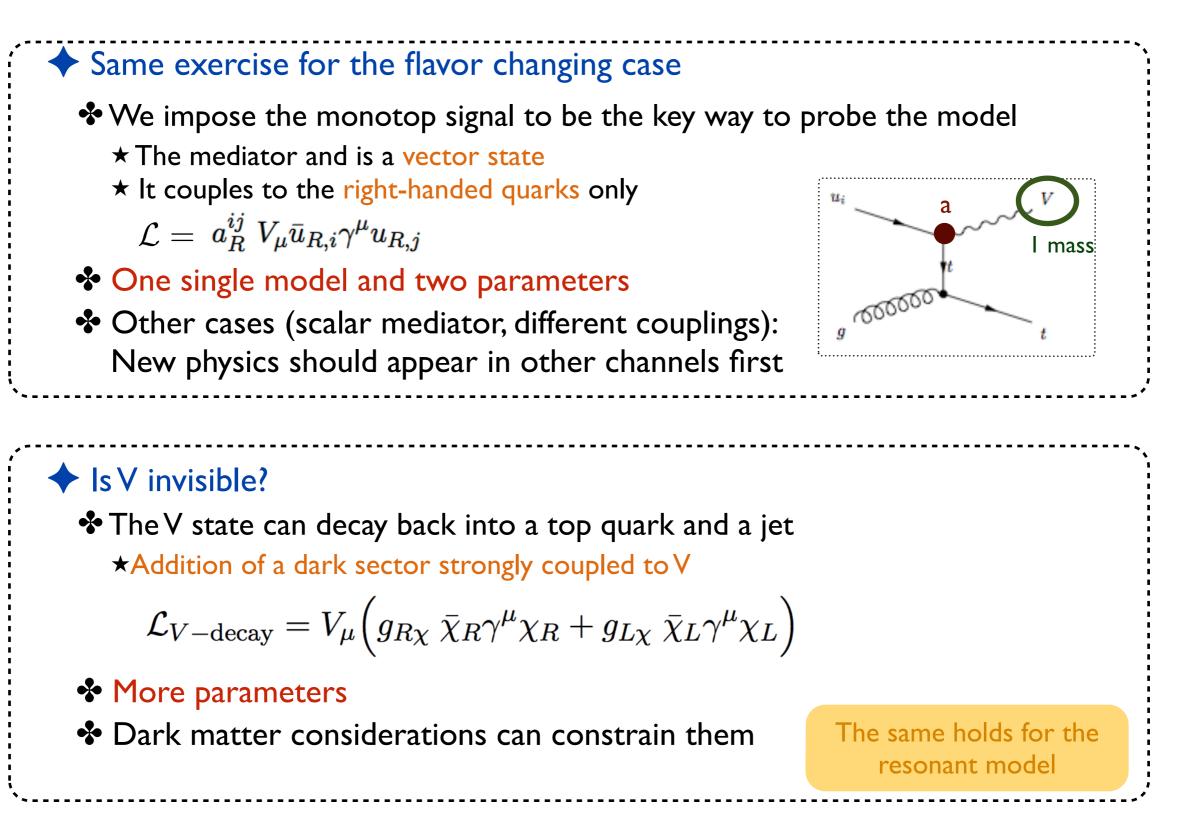
Towards UV completion: the EW symmetry (1)

[Boucheneb, Cacciapaglia, Deandrea, BF (JHEP '15)]



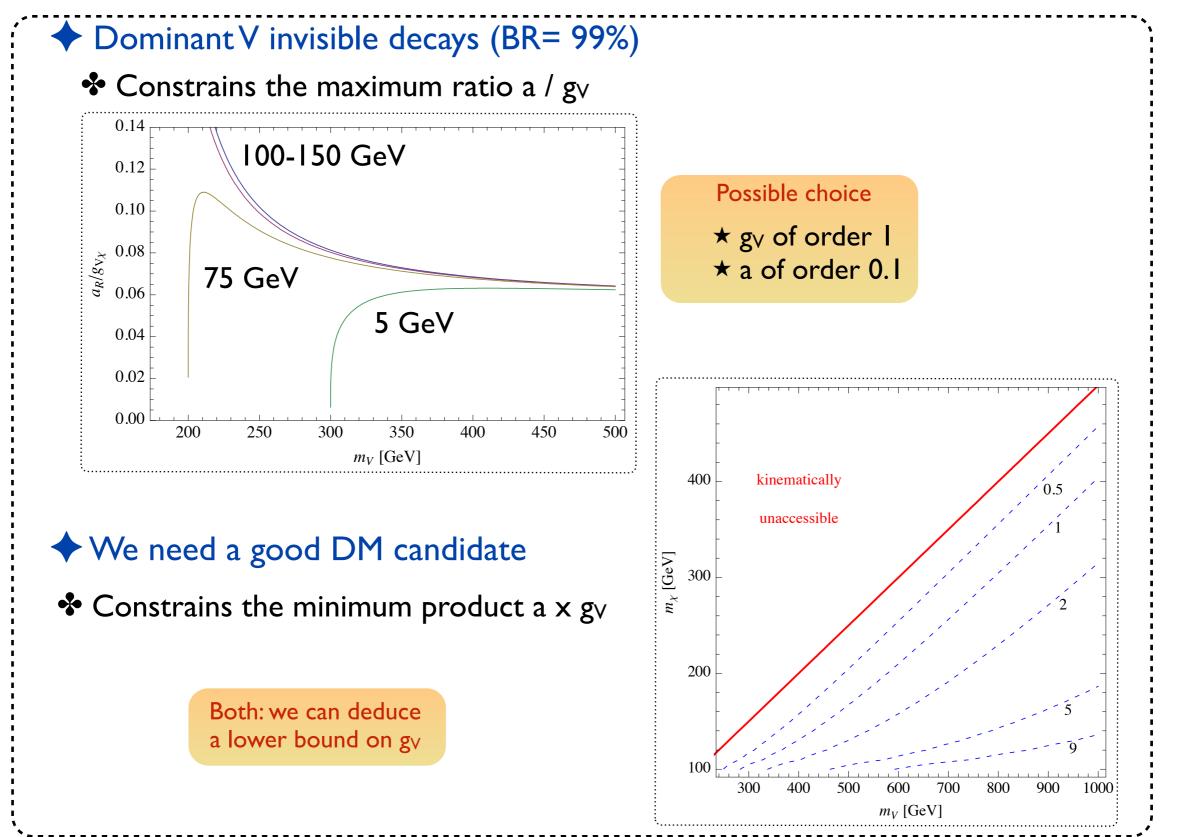
Towards UV completion: the EW symmetry (2)

[Boucheneb, Cacciapaglia, Deandrea, BF (JHEP '15)]



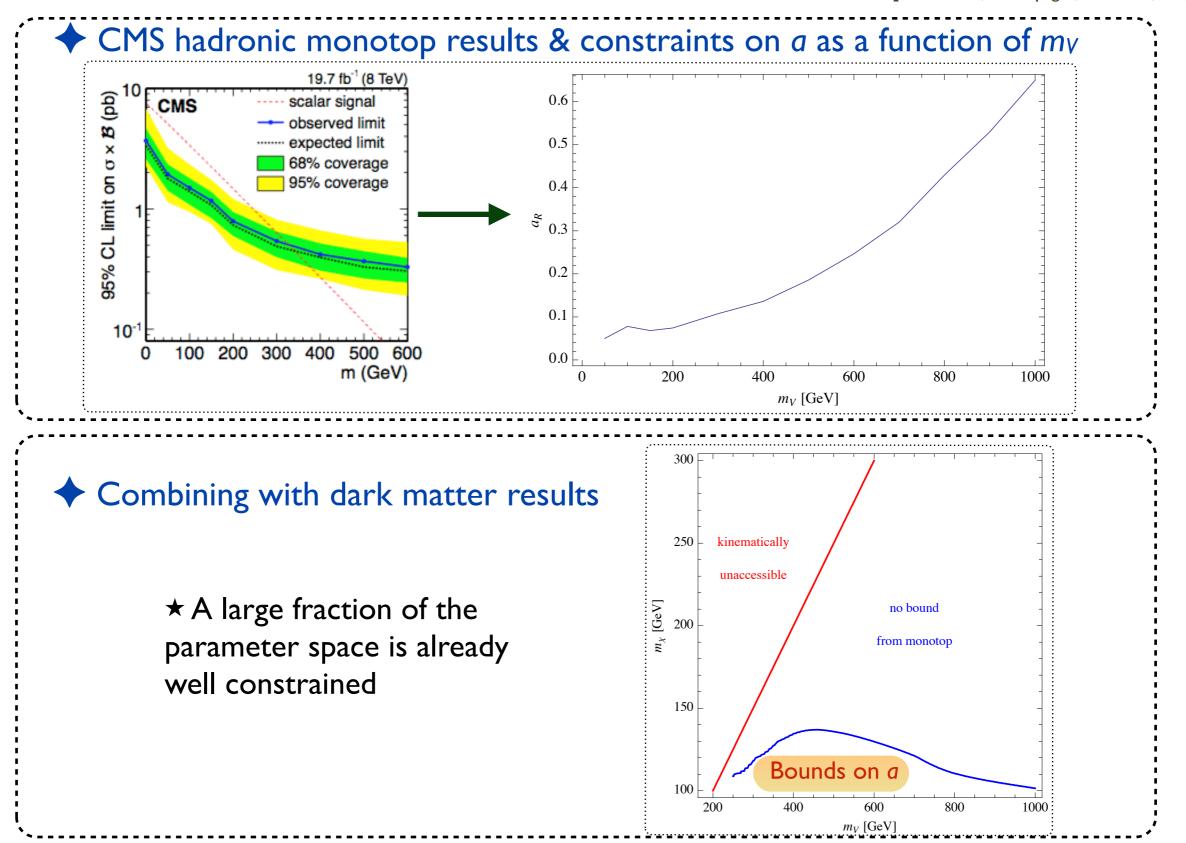
Pinning down the flavor changing monotop model

[Boucheneb, Cacciapaglia, Deandrea, BF (JHEP '15)]



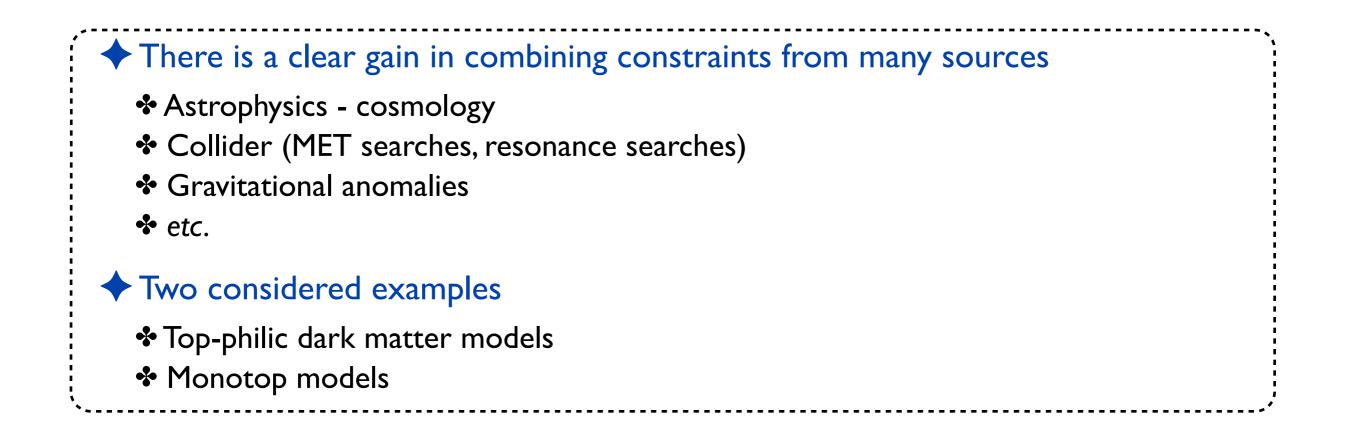
Combining collider and dark matter constraints

[Boucheneb, Cacciapaglia, Deandrea, BF (JHEP '15)]



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Conclusions (before the next topic)







Simplified models for top-philic dark matter

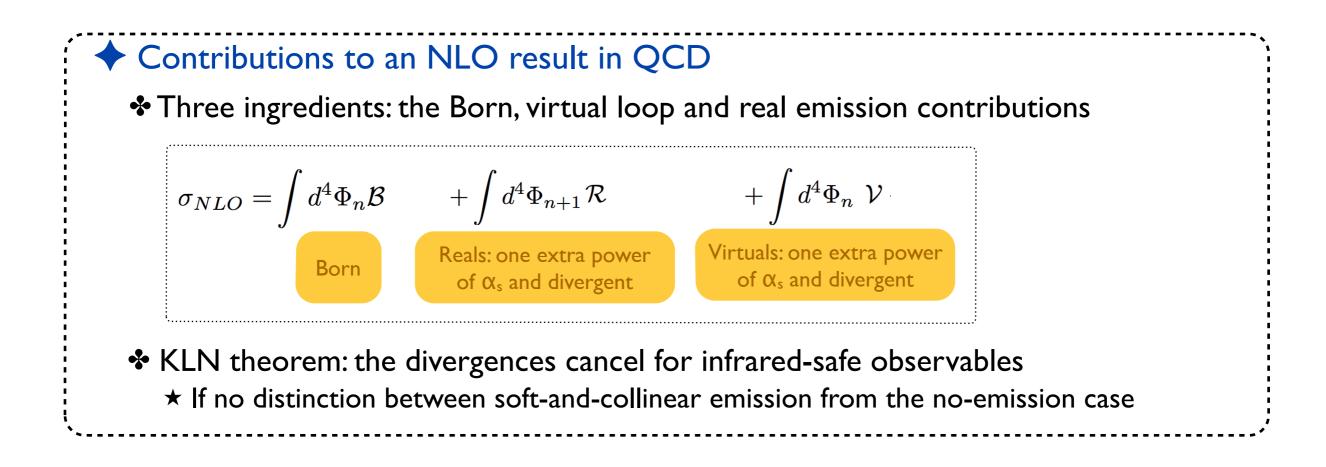


2. Monotop models and their constraints

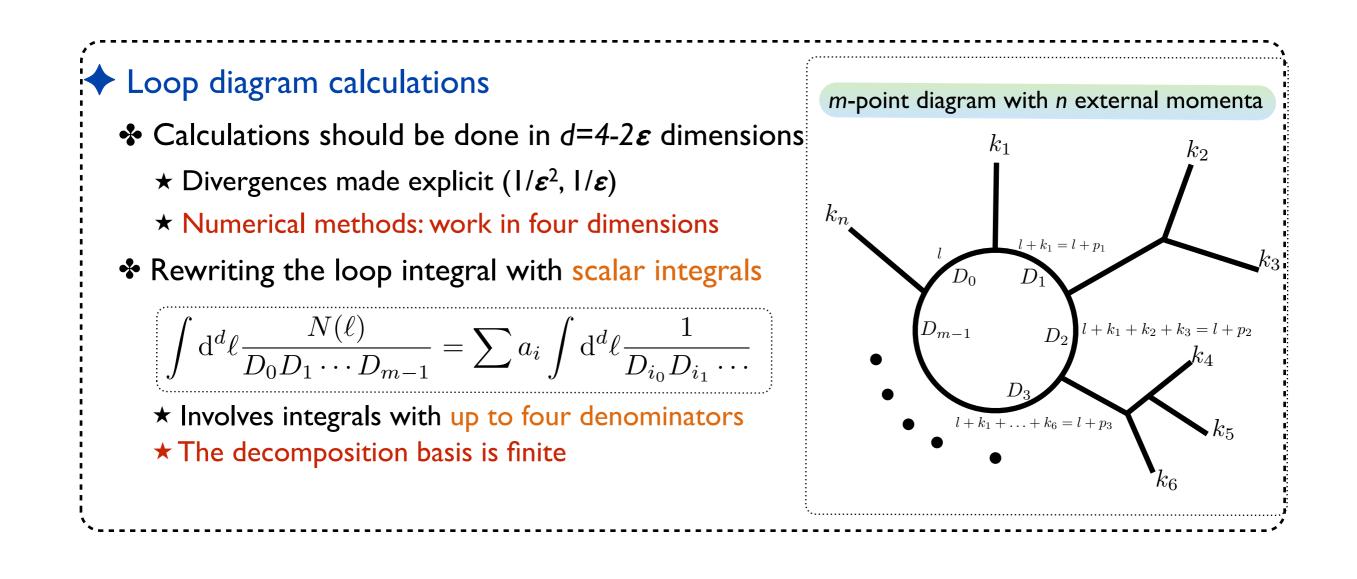


Effects of next-to-leading-order predictions

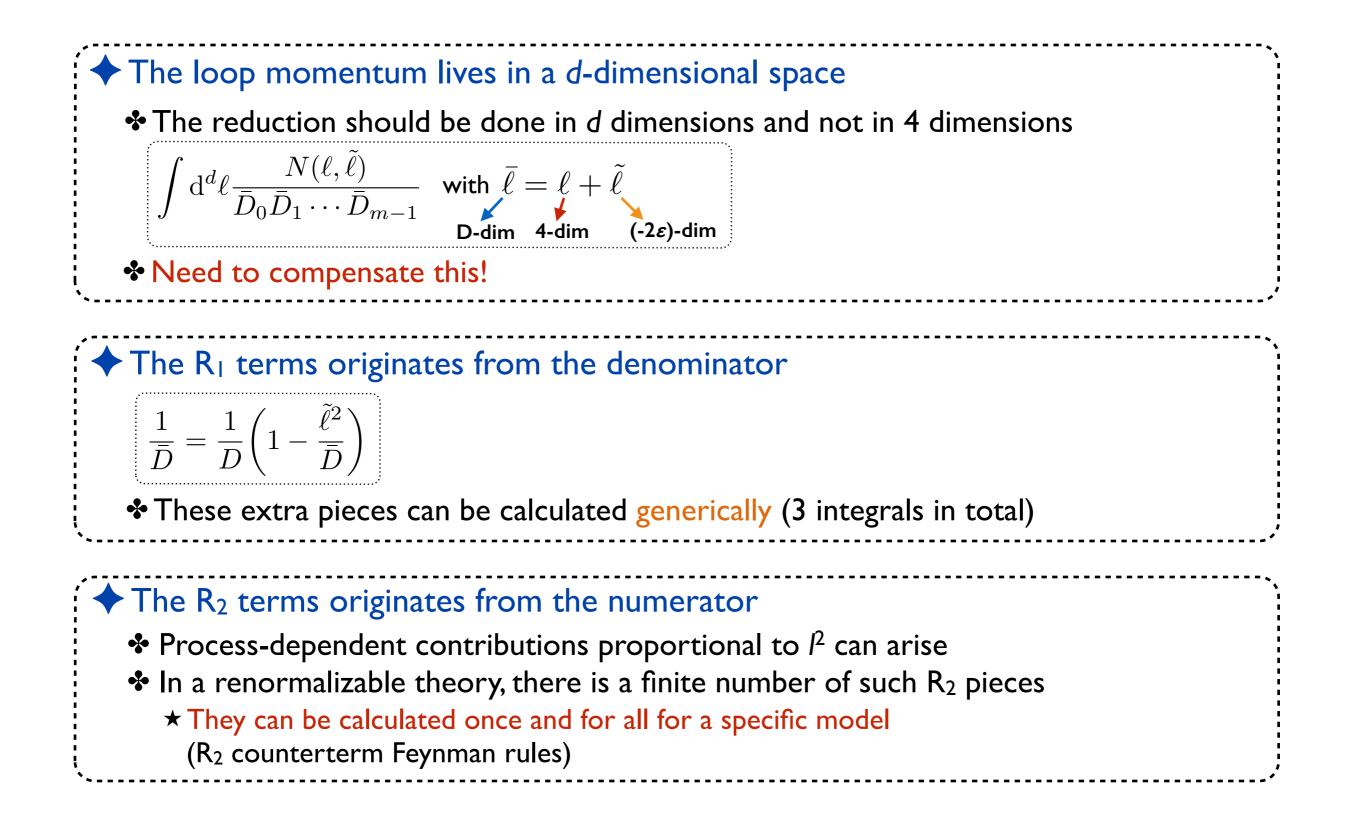
NLO calculations in a nutshell



The virtual contributions



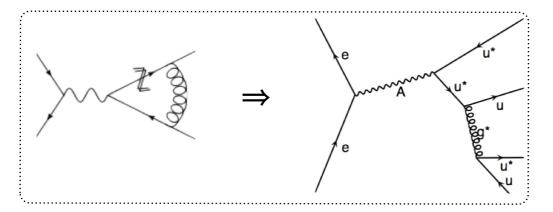
Reduction in four dimensions - the rational terms



Loop calculations with MADLOOP

MADLOOP uses MADGRAPH tree-level capabilities for loop calculations

* Loop-diagrams with n external legs are cut: tree-level diagrams with n+2 external legs



All diagrams with two extra partons in the final states are generated
 A first filter removes the non-necessary ones (including permutations, mirror graphs, etc.)
 A second filter removes the external-line tadpoles and bubbles graphs

MADLOOP then calculates the virtual contributions

- Contraction with Born diagrams, color traces calculations, etc.
- Internal propagator denominators are removed
 - \star We have the loop integrand numerator
 - \star It can be reduced and evaluated

 \clubsuit UV and R₂ counterterm diagrams added

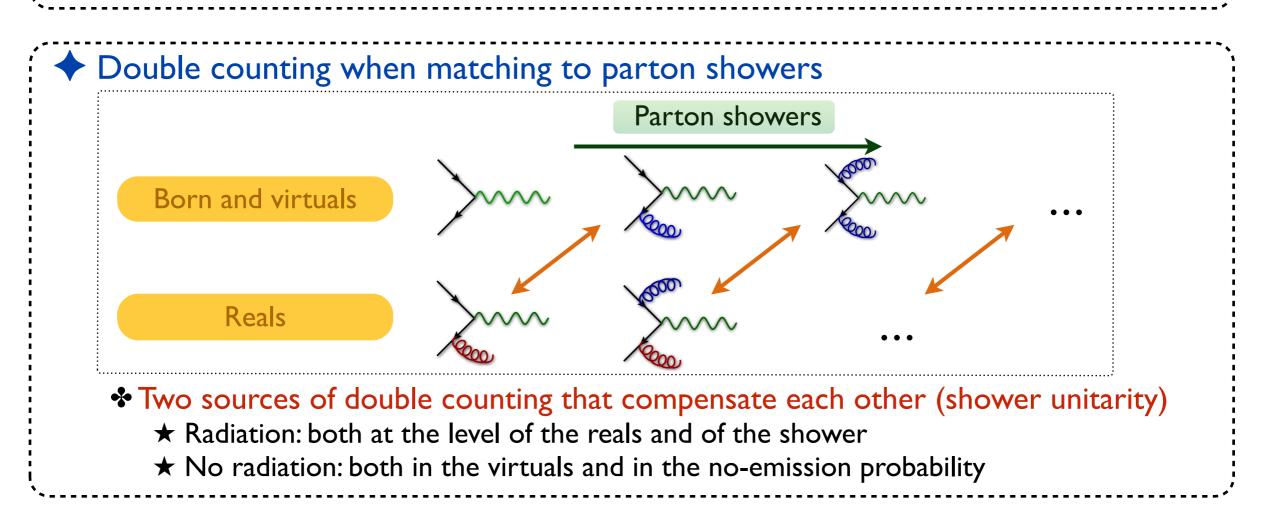
Matching NLO calculations to parton showers

Subtracting the poles

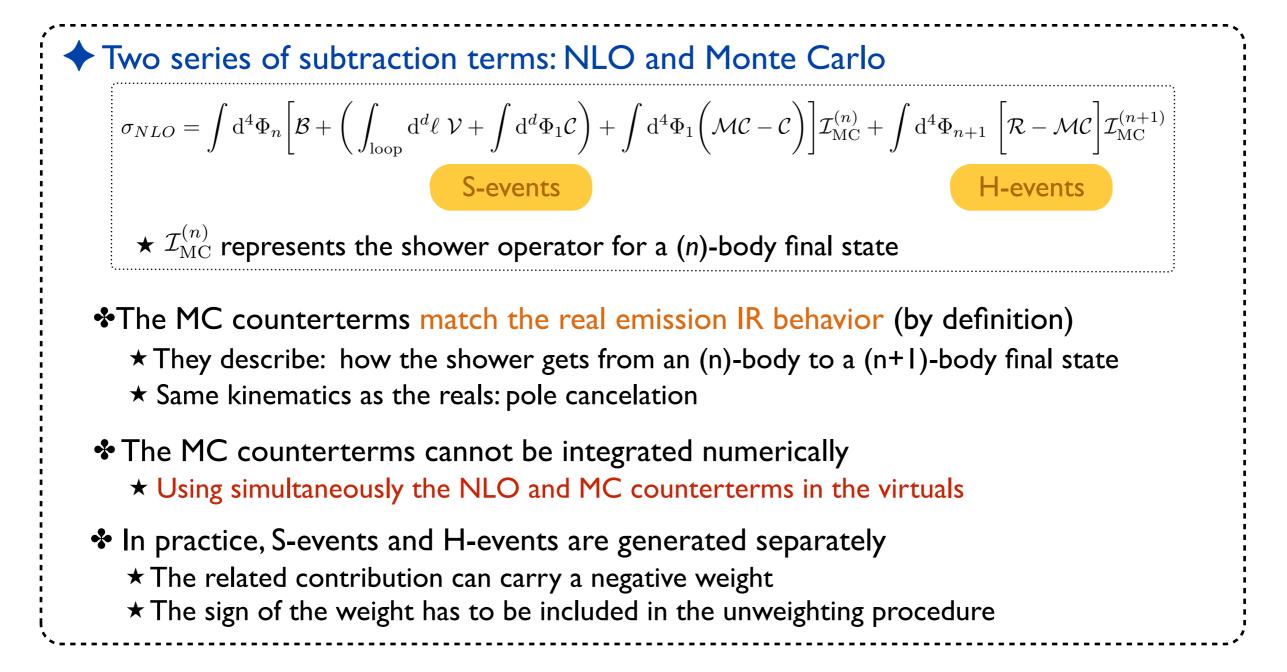
* The structure of the poles appearing at NLO is known \succ subtraction methods

- $\star \mathcal{C}$ subtracted from the reals \succ makes them finite
- $\star \mathcal{C}$ integrated and added back to the virtuals \succ makes them finite
- * Integrals can be made numerically in four dimensions

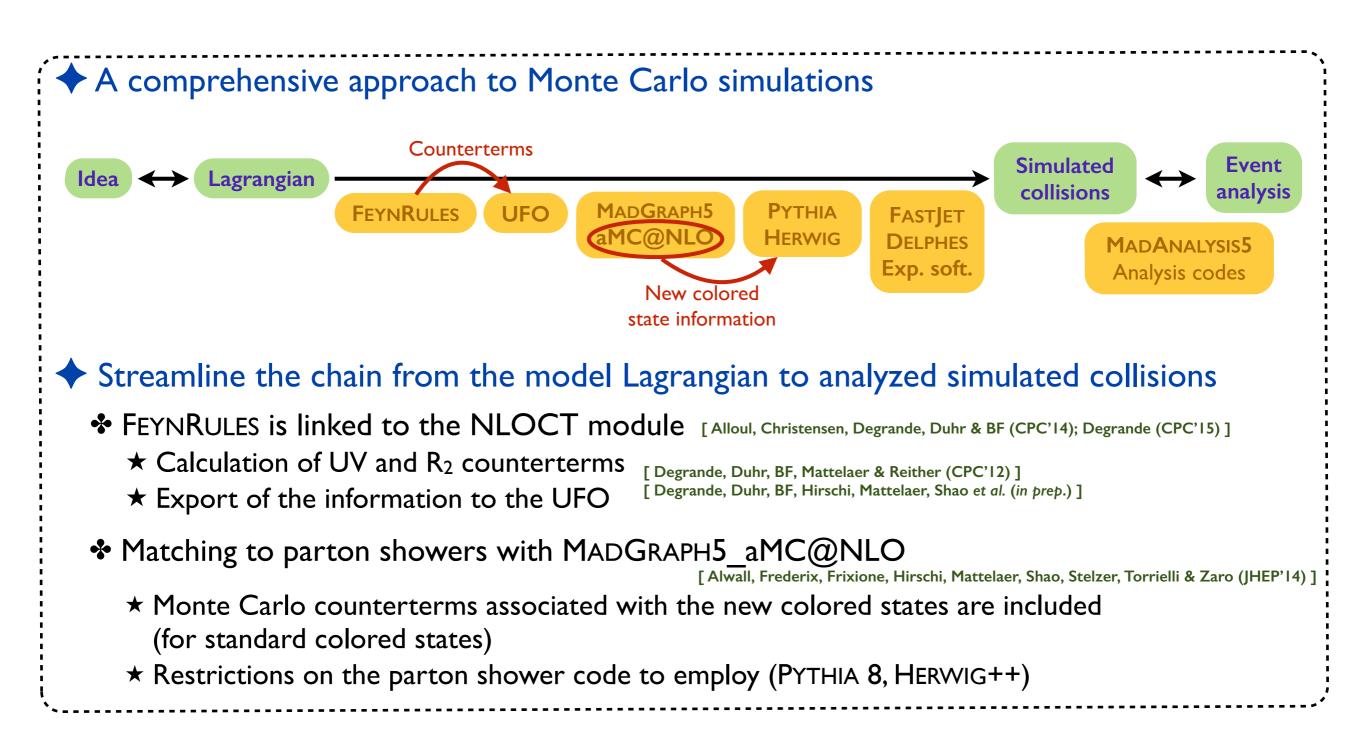
$$\sigma_{NLO} = \int d^4 \Phi_n \ \mathcal{B} + \int d^4 \Phi_{n+1} \left[\mathcal{R} - \mathcal{C} \right] + \int d^4 \Phi_n \left[\int_{\text{loop}} d^d \ell \ \mathcal{V} + \int d^d \Phi_1 \mathcal{C} \right]$$



Counterterms and master formula (MC@NLO)



Automating NLO event generation

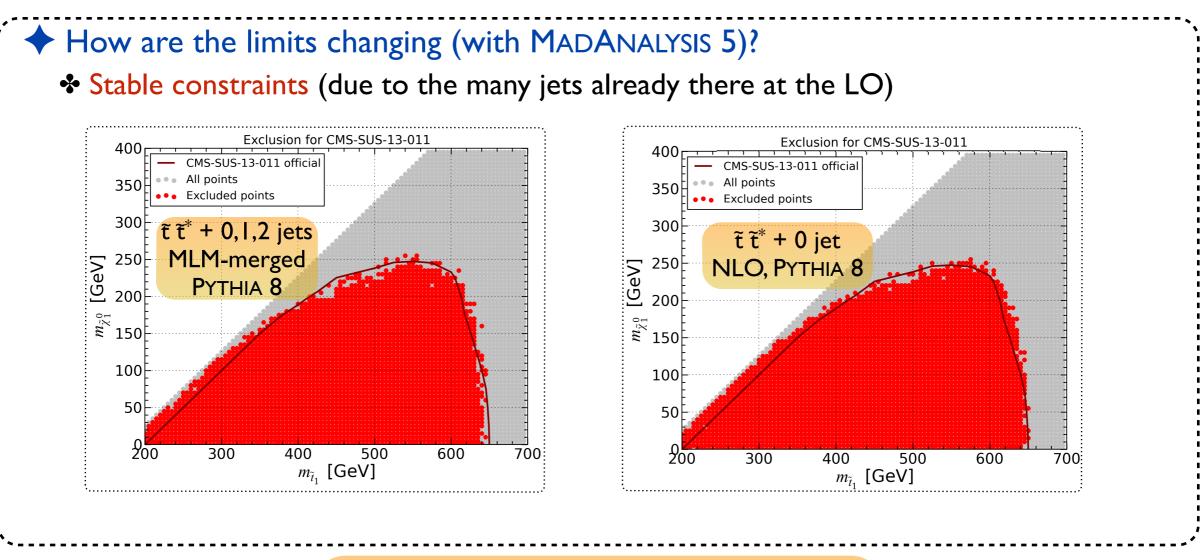


SUS-13-011: NLO effects on an exclusion limit

[Ambrogi, Conte, BF, Kulkarni & Molter (in preparation)]

♦ LO and NLO

- I. Simulated signal: $p p \rightarrow \tilde{t} \tilde{t}^* + 0, 1, 2$ jets @LO ; PYTHIA 8 with the MONASH tune
- **2.** Simulated signal: $p p \rightarrow \tilde{t} \tilde{t}^* + 0$ jet @NLO ; PYTHIA 8 with the MONASH tune



This is an analysis dependent statement NLO effects could be crucial for some analyses!!! Better control of the uncertainties in all cases!

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B2G-I2-002: top-antitop + missing energy search

[Arina, Backovic, Conte, BF, Guo et al. (1605.09242)]

	(m_Y,m_X)	$\sigma_{ m LO}~[{ m pb}]$	CL_{LO} [%]	$\sigma_{ m NLO}$ [pb]	CL_{NLO} [%]
Ι	$(150, 25) { m GeV}$	$0.658^{+34.9\%}_{-24.0\%}$	$98.7^{+0.8\%}_{-13.0\%}$	$0.773^{+6.1\%}_{-10.1\%}$	$95.0^{+2.7\%}_{-0.4\%}$
II	$(40, 30) { m GeV}$	$0.776^{+34.2\%}_{-24.1\%}$	$74.7^{+19.7\%}_{-17.7\%}$	$0.926^{+5.7\%}_{-10.4\%}$	$84.2^{+0.4\%}_{-14.4\%}$
III	$(240, 100) { m GeV}$	$0.187^{+37.1\%}_{-24.4\%}$	$91.6^{+6.4\%}_{-18.1\%}$	$0.216^{+6.7\%}_{-11.4\%}$	$86.5^{+8.6\%}_{-5.5\%}$

(Final) conclusions

There is a clear gain in combining constraints from many sources

- Astrophysics cosmology
- Collider (MET searches, resonance searches)
- Gravitational anomalies

setc.

Two considered examples

- Top-philic dark matter models
- Monotop models

NLO simulations become available

- Many effects: cross sections, differential distributions, efficiencies, CLs
- Number of available models increasing with time

[http://feynrules.irmp.ucl.ac.be/wiki/NLOModels]