

# Gubser flow and hydrodynamical fluctuations

Li Yan (严力)

CNRS, Institut de Physique Théorique, Saclay

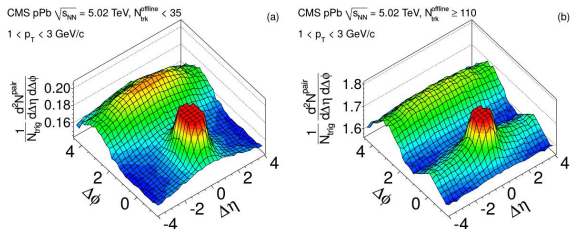


QCD Phase structure III, CCNU, June, 2016

with Hanna Grönqvist, JHEP03(2016)121

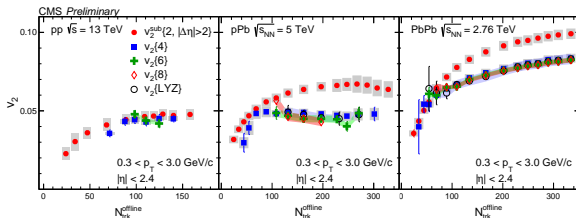
# Medium collectivity in small colliding system: p+A, p+p, etc.

- Long-range multi-particle correlation  $\Leftrightarrow$  medium collective expansion (QGP).



(CMS collaboration)

- Long-range multi-particle correlation  $\Leftrightarrow$  flow  $v_n$ , and flow cumulants  $v_n\{m\}$ .



(CMS collaboration)

- Modeling medium exp. in small colliding systems – **hydro.?** (good in AA)

initial state + hydro. EoM & EoS + freeze-out

- Modeling medium exp. in small colliding systems – **hydro.?** (good in AA)

initial state + hydro. EoM & EoS + freeze-out

- But in small systems, applicability of hydro. is challenged,

- \* convergence of gradient expansion –  $\lambda_{\text{mfp}}/L - \text{Kn}$  (**Niemi and Denicol**)

- \* effects of thermal fluctuations are expected stronger.

- Modeling medium exp. in small colliding systems – **hydro.?** (good in AA)

initial state + hydro. EoM & EoS + freeze-out

- But in small systems, applicability of hydro. is challenged,
  - \* convergence of gradient expansion –  $\lambda_{\text{mfp}}/L - \text{Kn}$  (Niemi and Denicol)
  - \* effects of thermal fluctuations are expected stronger.
- Goal of this work : A preliminary analysis of the effect of thermal noise
  - \* How significant is the effect of thermal fluctuations in heavy-ion collisions?
  - \* In particular, effect of thermal fluctuations in small colliding systems?

- Hydrodynamics – conservation of energy-momentum, etc.

$$T^{\mu\nu} = T_{\text{ideal}}^{\mu\nu} + \Pi^{\mu\nu} \quad \leftrightarrow \quad d_{\mu}T^{\mu\nu} = 0 \quad (\text{Euler and equ. of continuity})$$

especially from hydro EoM, one recognizes that

$$d_{\mu}(su^{\mu}) = -\frac{1}{T}\nabla_{(\mu}u_{\nu)}\Pi^{\mu\nu} \quad \rightarrow \quad \frac{dS}{dt} = -\int d^3x \frac{1}{T}\nabla_{(\mu}u_{\nu)}\Pi^{\mu\nu}$$

Navier-Stokes hydro. (1st order):

$$\Pi^{\mu\nu} = -2\eta \left[ \frac{1}{2}\Delta^{\mu\alpha}\Delta^{\nu\beta}(d_{\alpha}u_{\beta} + d_{\beta}u_{\alpha}) - \frac{1}{3}\Delta^{\mu\nu}\Delta^{\alpha\beta}d_{\alpha}u_{\beta} \right] = -\eta\sigma^{\mu\nu}$$

where

$$\Delta^{\mu\nu} = u^{\mu}u^{\nu} + g^{\mu\nu}$$

- \* We ignore bulk viscosity  $\zeta$  in our work.

# Hydro. and thermal fluctuations in a fluid system

- EoM of a set of physical quantities  $\{x_a\}$  in a thermal system  
(Landau and Lifshitz, J. Kapusta, B. Muller and M. Stephanov)

$$\dot{x}_a = - \underbrace{\sum_b \gamma_{ab} X_b}_{\text{drag}} + \underbrace{y_a}_{\text{fluc.}} \quad \Leftrightarrow \quad \dot{S} = - \sum_a \dot{x}_a X_a$$

$$\text{Maximization of } S \Rightarrow \langle y_a(t_1) y_b(t_2) \rangle = (\gamma_{ab} + \gamma_{ba}) \delta(t_1 - t_2)$$

- Auto-correlations of thermal noise : fluctuation-dissipation
- For hydro.,  $\gamma_{ab}$  is determined then by identifying

$$\dot{x} \rightarrow \Pi^{\mu\nu} \quad \text{and} \quad X \rightarrow \frac{\Delta V}{T} \nabla_{(\mu} u_{\nu)}, \quad \left( \frac{dS}{dt} = - \int d^3x \frac{1}{T} \nabla_{(\mu} u_{\nu)} \Pi^{\mu\nu} \right)$$

- Remarks:

1. White noise – delta function  $\delta() \sim \frac{1}{\Delta V \Delta t}$ .
2. Form of  $\gamma_{ab}$  corresponds to the detailed form of  $\Pi^{\mu\nu}$ ,  $\gamma_{ab} \sim$  dissipations.

- Hydrodynamics with thermal noise (Navier-Stokes hydro):

( Landau and Lifshitz, J. Kapusta et. al. )

$$T^{\mu\nu} = T_{\text{ideal}}^{\mu\nu} + \Pi^{\mu\nu} + S^{\mu\nu}$$

where thermal fluctuation tensor  $S^{\mu\nu}$  are introduced w.r.t.  $\Pi^{\mu\nu}$

$$\langle S^{\mu\nu}(x) \rangle = 0$$

$$\text{Navier-Stokes: } \langle S^{\mu\nu}(x_1) S^{\alpha\beta}(x_2) \rangle = 4T\eta \Delta^{\mu\nu\alpha\beta} \delta^{(4)}(x_1 - x_2)$$

where

$$\Delta^{\mu\nu\alpha\beta} = \frac{1}{2} \left[ \Delta^{\mu\alpha} \Delta^{\nu\beta} + \Delta^{\mu\beta} \Delta^{\nu\alpha} \right] - \frac{1}{3} \Delta^{\mu\nu} \Delta^{\alpha\beta}, \text{ and } \Delta^{\mu\nu\alpha\beta} d_\alpha u_\beta = \sigma^{\mu\nu}$$

- \* One-point functions of physical quantities are not affected.
- \* Thermal noise affects two-point correlations.



## Solving hydro. with thermal noise

- Linearized hydro. EoM with thermal fluctuations around  $d_\mu T_0^{\mu\nu} = 0$

$$T(x) = T_0(x) + \delta T(x)$$

$$\varepsilon(x) = \varepsilon_0(x) + \delta\varepsilon(x)$$

$$\mathcal{P}(x) = \mathcal{P}_0(x) + \delta\mathcal{P}(x)$$

$$u^\mu(x) = u_0^\mu(x) + \delta u^\mu(x)$$

therefore  $d_\mu \delta T^{\mu\nu} = 0 \sim O(\delta)$ ,

$$\delta w D u_\alpha + w \delta u^\mu d_\mu u_\alpha + (D w + w \partial \cdot u) \delta u_\alpha + \nabla_\alpha \delta \mathcal{P} + w D \delta u_\alpha + d_\mu (\delta \Pi_\alpha^\mu + S_\alpha^\mu) = 0$$

$$D \delta \varepsilon + \delta w \partial \cdot u + d_\mu (w \delta u^\mu) + w \delta u^\alpha D u_\alpha - u^\alpha d_\mu (\delta \Pi_\alpha^\mu + S_\alpha^\mu) = 0$$

Note that  $\delta \Pi^{\mu\nu}$  is induced by  $\delta T$ , etc. (C. Young, J. Kapusta et al.)

- Beyond linear order, thermal noise becomes large, e.g., phase transition.
- Simplification can be made when hydro. is analytically solved:

## Solving hydro. with thermal noise

- Linearized hydro. EoM with thermal fluctuations around  $d_\mu T_0^{\mu\nu} = 0$

$$T(x) = T_0(x) + \delta T(x)$$

$$\varepsilon(x) = \varepsilon_0(x) + \delta\varepsilon(x)$$

$$\mathcal{P}(x) = \mathcal{P}_0(x) + \delta\mathcal{P}(x)$$

$$u^\mu(x) = u_0^\mu(x) + \delta u^\mu(x)$$

therefore  $d_\mu \delta T^{\mu\nu} = 0 \sim O(\delta)$ ,

$$\delta w D u_\alpha + w \delta u^\mu d_\mu u_\alpha + (D w + w \partial \cdot u) \delta u_\alpha + \nabla_\alpha \delta \mathcal{P} + w D \delta u_\alpha + d_\mu (\delta \Pi_\alpha^\mu + S_\alpha^\mu) = 0$$

$$D \delta \varepsilon + \delta w \partial \cdot u + d_\mu (w \delta u^\mu) + w \delta u^\alpha D u_\alpha - u^\alpha d_\mu (\delta \Pi_\alpha^\mu + S_\alpha^\mu) = 0$$

Note that  $\delta \Pi^{\mu\nu}$  is induced by  $\delta T$ , etc. (C. Young, J. Kapusta et al.)

- Beyond linear order, thermal noise becomes large, e.g., phase transition.
- Simplification can be made when hydro. is analytically solved:

Bjorken flow and Gubser's flow

- Boost invariance  $\times$  translational invariance in the transverse plane
- Correlation of thermal noise in 1+1D Bjorken hydro,. (J. Kapusta et. al., 2012)

$$\langle S^{\mu\nu}(\tau_1, \xi_1) S^{\alpha\beta}(\tau_2, \xi_2) \rangle = \frac{8}{3\tau_1 A_\perp} T \eta \Delta^{\mu\nu} \Delta^{\alpha\beta} \delta(\tau_1 - \tau_2) \delta(\xi_1 - \xi_2)$$

1.  $\delta(\vec{x}_{1\perp} - \vec{x}_{2\perp}) \rightarrow A_\perp$  characterizes transverse size of the system.
2. Tensor structure of  $S^{\mu\nu}$  is factorized, due to the fact that  $u^\mu = (1, 0)$ .

$$S^{\mu\nu} = w(\tau) f(\tau, \xi) \Delta^{\mu\nu},$$

such that the unknown scalar and dimensionless function  $f(\tau, \xi)$

$$\langle f(\tau_1, \xi_1) f(\tau_2, \xi_2) \rangle = \frac{2\nu}{A_\perp w(\tau_1) \tau_1} \delta(\tau_1 - \tau_2) \delta(\xi_1 - \xi_2) \quad \text{with } \nu = \frac{4}{3} \frac{\eta}{s}$$

3. Magnitude of thermal noise is constrained by (in addition to  $\eta/s$ ):

$$(A_\perp w(\tau) \tau) \sim A_\perp \left( \frac{dE}{\tau d^2 x_\perp d\xi} \right) \tau \sim \frac{dE_\perp}{dy} \sim \text{multiplicity}$$

Multiplicity more crucial than system size.

- Gubser hydro., 2+1D ([Gubser and Yarom, 2010](#))
  - ▶ Bjorken boost – indep. of spatial rapidity  $\xi$
  - ▶ Rotational symmetry w.r.t. to beam axis – for p+A and ultra-central A+A
- Change coordinates  $(\tau, r, \phi, \xi) \leftrightarrow (\rho, \theta, \phi, \xi)$  via the following mapping

$$\sinh \rho = -\frac{1 - q^2 \tau^2 + q^2 r^2}{2q\tau}, \quad \tan \theta = \frac{2qr}{1 + q^2 \tau^2 - q^2 r^2}$$

$\rho$  plays the role of ‘time’ in the new (with a ‘hat’) coordinate system.

$$\hat{u}^\mu = (1, 0, 0, 0)$$

- There are two parameters:  $\hat{T}_0 \rightarrow$  multiplicity,  $q \rightarrow$  transverse size.
- Non-trivial description of the radial expansion.

- Correlation of thermal noise in 2+1D Gubser hydro., ( $X \rightarrow (\rho, \theta, \phi, \xi)$ )

$$\langle \hat{S}^{\mu\nu}(\rho_1, \theta_1, \phi_1, \xi_1) \hat{S}^{\alpha\beta}(\rho_2, \theta_2, \phi_2, \xi_2) \rangle = \frac{2\nu \hat{T} \hat{s}}{\cosh^2 \rho_1 \sin \theta_1} \hat{\mathcal{P}}^{\mu\nu} \hat{\mathcal{P}}^{\alpha\beta} \delta(X_1 - X_2)$$

- Tensor structure of  $\hat{S}^{\mu\nu}$  is factorized, due to  $\hat{u}^\mu = (1, 0, 0, 0)$ .

$$\hat{S}^{\mu\nu}(\rho, \theta, \phi, \xi) = \hat{w}(\rho) \hat{f}(\rho, \theta, \phi, \xi) \hat{\mathcal{P}}^{\mu\nu},$$

and again we have the correlation of scalar function

$$\langle \hat{f}(\rho_1, \theta_1, \phi_1, \xi_1) \hat{f}(\rho_2, \theta_2, \phi_2, \xi_2) \rangle = \frac{2\nu}{\hat{w} \cosh^2 \rho_1 \sin \theta_1} \delta(X_1 - X_2)$$

- For scalar function  $\hat{f}(X)$ , mode decomposition w.r.t. SO(3) symmetry leads to

$$\text{scalar modes: } \hat{f}(\rho, \theta, \phi, \xi) = \sum h(\rho) Y_{lm}(\theta, \phi) e^{ik_\xi \xi}$$

and

$$\langle h(\rho_1) h(\rho_2) \rangle = \frac{2\nu}{\hat{w} \cosh^2 \rho_1} \delta(\rho_1 - \rho_2)$$

- Magnitude of thermal noise is constrained by

$$\hat{w} \sim \hat{T}_0 \sim \text{multiplicity}$$

Multiplicity more crucial than system size.

## Solve Gubser hydro. with thermal noise

- Decompose thermal fluctuations into modes – scalar and vector modes:

$$\delta\hat{T} = \hat{T} \sum \delta_l(\rho) Y_{lm}(\theta, \phi) e^{ik_\xi \xi}$$

$$\delta\hat{u}_i = \sum \left[ v_{ls}(\rho) \partial_i Y_{lm}(\theta, \phi) e^{ik_\xi \xi} + v_{lv}(\rho) \Phi_{i(lm)}(\theta, \phi) e^{ik_\xi \xi} \right]$$

$$\delta\hat{u}_\xi = \sum v_{l\xi}(\rho) Y_{lm}(\theta, \phi) e^{ik_\xi \xi}$$

- EoM of each mode,

$$\tilde{V}'_l(\rho) = -\tilde{\Gamma}(\rho, l, k_\xi) \tilde{V}_l(\rho) + \tilde{\mathcal{K}}(\rho, k_\xi),$$

where prime denotes derivative w.r.t.  $\rho$

$$\tilde{V}_l(\rho) = \begin{pmatrix} \delta_l(\rho) \\ v_{ls}(\rho) \\ v_{l\xi}(\rho) \\ v_{lv}(\rho) \end{pmatrix}, \quad \tilde{\Gamma} \text{ is a } 4 \times 4 \text{ matrix,} \quad \tilde{\mathcal{K}} = \begin{pmatrix} -\frac{2}{3} \tanh \rho h(\rho) \\ \frac{2\hat{T}}{3\hat{T}'} \tanh \rho h(\rho) \\ -\frac{ik_\xi \hat{T}}{\hat{T} + H_0 \tanh \rho} h(\rho) \\ 0 \end{pmatrix}$$

- Coupled EoMs in 3+1D.
- Vector modes are decoupled, and NOT affected by thermal noise.
- Note that thermal fluctuations are indep. of  $m$ .

# Apply noisy Gubser hydro. to ultra-central Pb-Pb, p-Pb and p-p

- $\hat{T}_0$  and  $q$  determine the system.

	PbPb	pPb	pp
$\hat{T}_0$	7.3	3.1	1.7
$q^{-1}(fm)^{-1}$	4.3	1.1	1.1

Note that the strength of thermal fluctuations is fixed once  $\hat{T}_0$  is given.

- Approximates system evolution of first several fm's
- $k_\xi = 0$  mode
  - Long rapidity range correlations, affected also by initial fluctuations.
  - Further simplification with  $v_\xi$  modes decoupled  $\rightarrow$  2 coupled equations.

$$(\delta\hat{T}, \delta\hat{u}_i) \longleftrightarrow (\delta T, \delta u_\tau, \delta u_r, \delta u_\phi)$$

- We will NOT discuss hadronization and freeze-out.

# Apply noisy Gubser hydro. to ultra-central Pb-Pb, p-Pb and p-p

- $\hat{T}_0$  and  $q$  determine the system.

	PbPb	pPb	pp
$\hat{T}_0$	7.3	3.1	1.7
$q^{-1}(fm)^{-1}$	4.3	1.1	1.1

Note that the strength of thermal fluctuations is fixed once  $\hat{T}_0$  is given.

- Approximates system evolution of first several fm's
- $k_\xi = 0$  mode
  - Long rapidity range correlations, affected also by initial fluctuations.
  - Further simplification with  $v_\xi$  modes decoupled  $\rightarrow$  2 coupled equations.

$$(\delta\hat{T}, \delta\hat{u}_i) \longleftrightarrow (\delta T, \delta u_\tau, \delta u_r, \delta u_\phi)$$

- We will NOT discuss hadronization and freeze-out.
- Initial condition ?

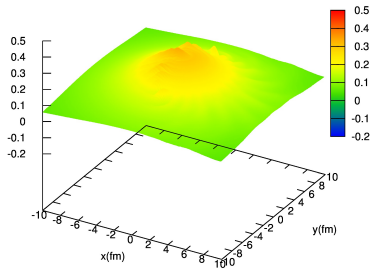
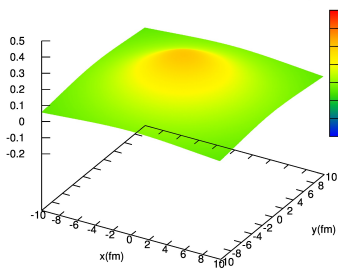


# Evolution of temperature

- $T(\tau, \vec{x}_\perp)$  with thermal noise, one random event

Pb-Pb

$T(\text{GeV}): \tau = 1.7 \text{ fm}$

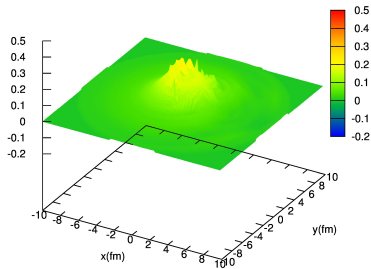
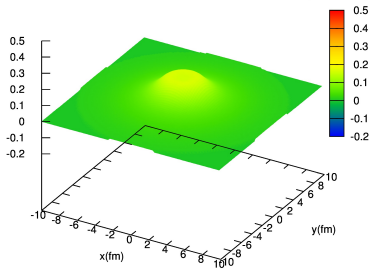


# Evolution of temperature

- $T(\tau, \vec{x}_\perp)$  with thermal noise, one random event

p-Pb

$T(\text{GeV}): \tau = 1.7 \text{ fm}$

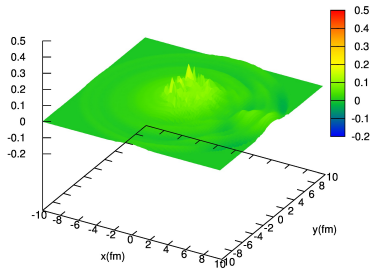
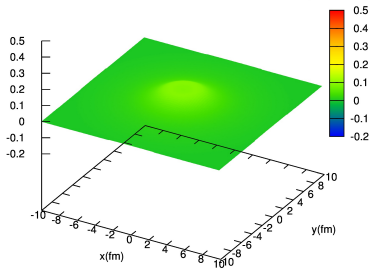


# Evolution of temperature

- $T(\tau, \vec{x}_\perp)$  with thermal noise, one random event

p-p

$T(\text{GeV}): \tau = 1.7 \text{ fm}$



## Two-point correlations – quantitative effects of thermal fluc.

- Two-point correlation of radial flow velocity

$$\begin{aligned} C_{u_r u_r}(\tau, \Delta\phi, r, \phi) &= \langle u_r(\tau, r, \phi) u_r(\tau, r, \phi + \Delta\phi) \rangle - \langle u_r(\tau, r) \rangle^2 \\ &= \underbrace{C_{u_r u_r}^I}_{\text{initial state fluc.}} + \underbrace{C_{u_r u_r}^T}_{\text{thermal fluc.}} \end{aligned}$$

Can also be defined w.r.t. fluctuation modes  $\delta T$ .

- Initial condition : specify  $\delta\hat{T}(\rho_0)$

- ▶ Long-range correlation: initial state eccentricity of order  $m$   $\varepsilon_m$  (single mode)

$$\frac{\delta\hat{T}(\theta, \phi, \rho_0, \xi)}{\hat{T}(\rho_0)} = -\sqrt{\Lambda_{ini}} \left[ (-1)^m \frac{1}{\sqrt{2}} Y_{m,m}(\theta, \phi) + \frac{1}{\sqrt{2}} Y_{m,-m}(\theta, \phi) \right],$$

- ▶ Short-range correlation:  $\delta\hat{T}(\rho_0)$  as a Dirac delta function (all modes)

$$\frac{\delta\hat{T}(\rho_0, \theta, \phi, \xi)}{\hat{T}(\rho_0)} = \sqrt{\Lambda_{ini}} \times \frac{1}{\cosh \rho_0 \sin \theta} \delta(\theta - \theta_0) \delta(\phi - \phi_0).$$

## Two-point correlations – quantitative effects of thermal fluc.

- Two-point correlation of radial flow velocity

$$\begin{aligned} C_{u_r u_r}(\tau, \Delta\phi, r, \phi) &= \langle u_r(\tau, r, \phi) u_r(\tau, r, \phi + \Delta\phi) \rangle - \langle u_{rb}(\tau, r) \rangle^2 \\ &= \underbrace{C_{u_r u_r}^I}_{\text{initial state fluc.}} + \underbrace{C_{u_r u_r}^T}_{\text{thermal fluc.}} \end{aligned}$$

Can also be defined w.r.t. fluctuation modes  $\delta T$ .

- Initial condition : specify  $\delta\hat{T}(\rho_0)$

- ▶ Long-range correlation: initial state eccentricity of order  $m$   $\varepsilon_m$  (single mode)

$$\frac{\delta\hat{T}(\theta, \phi, \rho_0, \xi)}{\hat{T}(\rho_0)} = -\sqrt{\Lambda_{ini}} \left[ (-1)^m \frac{1}{\sqrt{2}} Y_{m,m}(\theta, \phi) + \frac{1}{\sqrt{2}} Y_{m,-m}(\theta, \phi) \right],$$

- ▶ Short-range correlation:  $\delta\hat{T}(\rho_0)$  as a Dirac delta function (all modes)

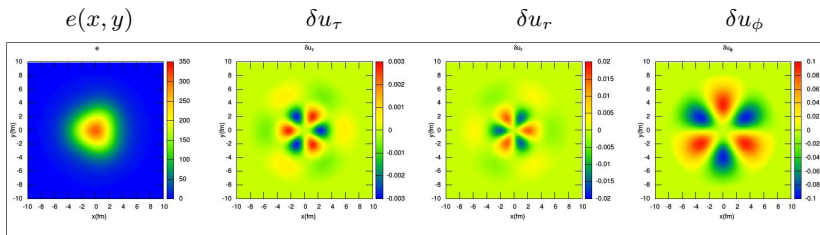
$$\frac{\delta\hat{T}(\rho_0, \theta, \phi, \xi)}{\hat{T}(\rho_0)} = \sqrt{\Lambda_{ini}} \times \frac{1}{\cosh \rho_0 \sin \theta} \delta(\theta - \theta_0) \delta(\phi - \phi_0).$$

$\Lambda_{ini}$  unknown, to be fixed by phenomenology.

- Initial eccentricity is responsible for harmonic flow  $v_n$

$$\varepsilon_m = -\frac{\int d^2\vec{x}_\perp r^m e^{im\phi} e(x, y)}{\int d^2\vec{x}_\perp r^m e(x, y)}$$

- E.g., background with initial  $\varepsilon_3 : Y_{3,3}$  and  $Y_{3,-3}$

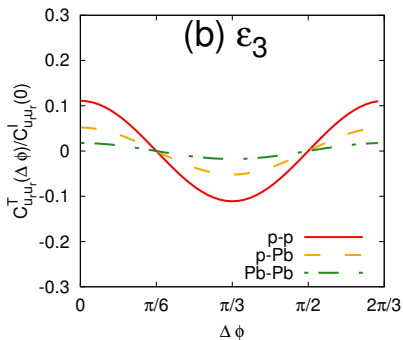
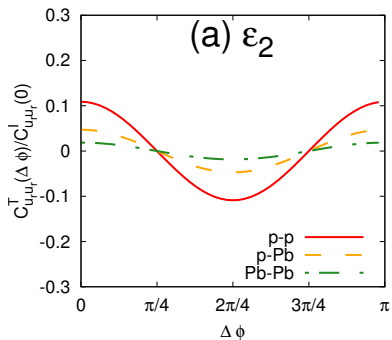


- Fix  $\Lambda_{ini}$  so that at  $\tau = 0.6$  fm: ( $m = 2, 3, 4, 5$ )

$$\varepsilon_m(\text{Pb-Pb}) \approx 0.05, \quad \varepsilon_m(\text{p-Pb}) \approx 0.15, \quad \varepsilon_m(\text{p-p}) \approx 0.2.$$

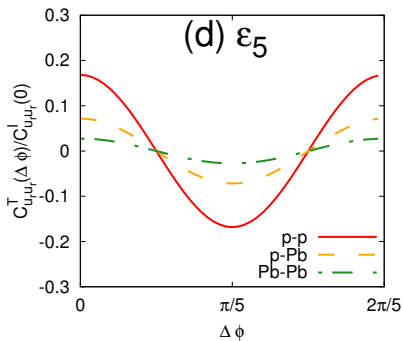
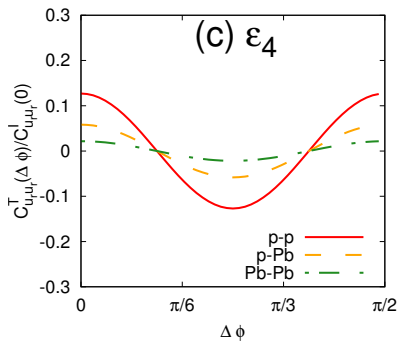
For long-range correlation

- $C_{u_r u_r}^T / C_{u_r u_r}^I$  with an initial  $\varepsilon_m$  vs.  $\Delta\phi$  at  $\tau = 2.5$  fm



For long-range correlation

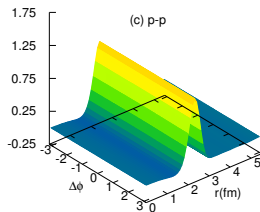
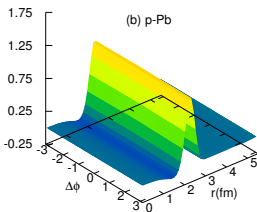
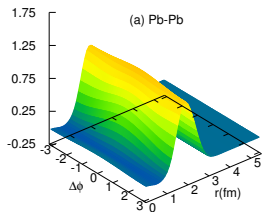
- $C_{u_r u_r}^T / C_{u_r u_r}^I$  with an initial  $\varepsilon_m$  vs.  $\Delta\phi$  at  $\tau = 2.5$  fm





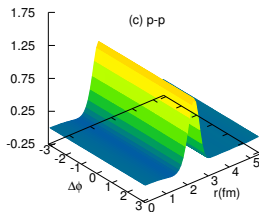
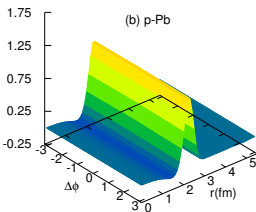
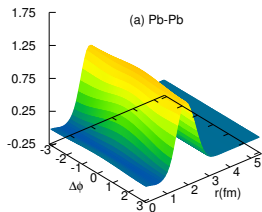
For short-range correlation

- $C_{u_r u_r}^I / C_{u_r u_r}^I(\Delta\phi = \pi)$  with initial Dirac delta vs.  $(r, \Delta\phi)$  at  $\tau = 2.5$  fm

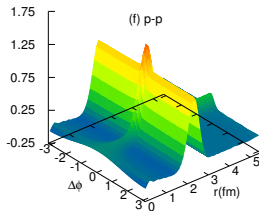
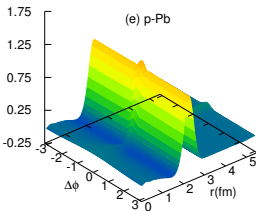
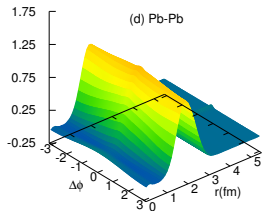


## For short-range correlation

- $C_{u_r u_r}^I / C_{u_r u_r}^I(\Delta\phi = \pi)$  with initial Dirac delta vs.  $(r, \Delta\phi)$  at  $\tau = 2.5$  fm

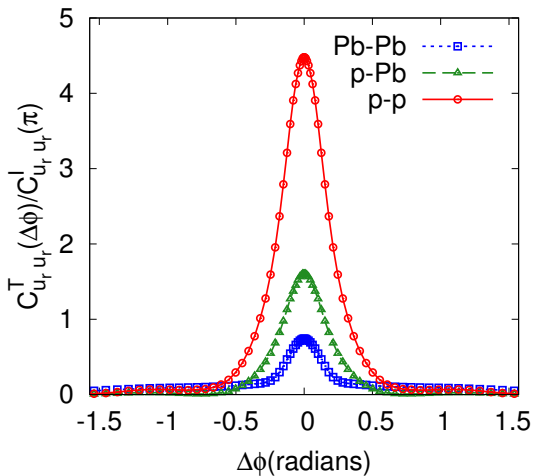


- $(C_{u_r u_r}^I + C_{u_r u_r}^T) / C_{u_r u_r}^I(\Delta\phi = \pi)$  vs.  $(r, \Delta\phi)$



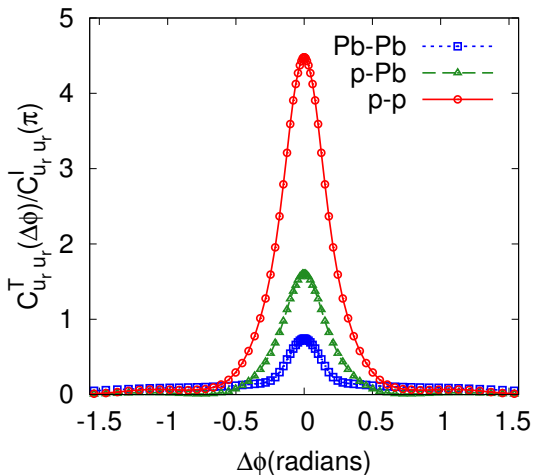
For short-range correlation

- $C_{u_r u_r}^T / C_{u_r u_r}^I(\Delta\phi = \pi)$  with initial Dirac delta vs.  $\Delta\phi$  at  $\tau = 2.5$  fm



For short-range correlation

- $C_{u_r u_r}^T / C_{u_r u_r}^I(\Delta\phi = \pi)$  with initial Dirac delta vs.  $\Delta\phi$  at  $\tau = 2.5$  fm



Related to the near-side peak structure in two-particle correlations.

