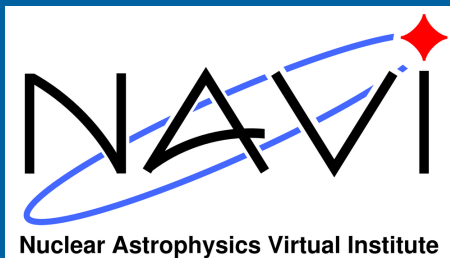


# Static type, windowless gas targets underground

2<sup>nd</sup> Workshop on Nuclear Astrophysics  
at the Canfranc Underground Laboratory

Canfranc/Spain, 01.03.2016

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**TECHNISCHE  
UNIVERSITÄT  
DRESDEN**



# Windowless gas target systems

- ◆ For low-energy, low counting rate experiments, the ion beam induced background may be a problem.
- ◆ This is especially true for experiments with  $^1\text{H}$  beam.

Typical backgrounds:

(a) Direct capture on nuclei with similar atomic charge

- ◆  $^{13}\text{C}(p,\gamma)$  background when  $^{14}\text{N}(p,\gamma)$  is to be studied

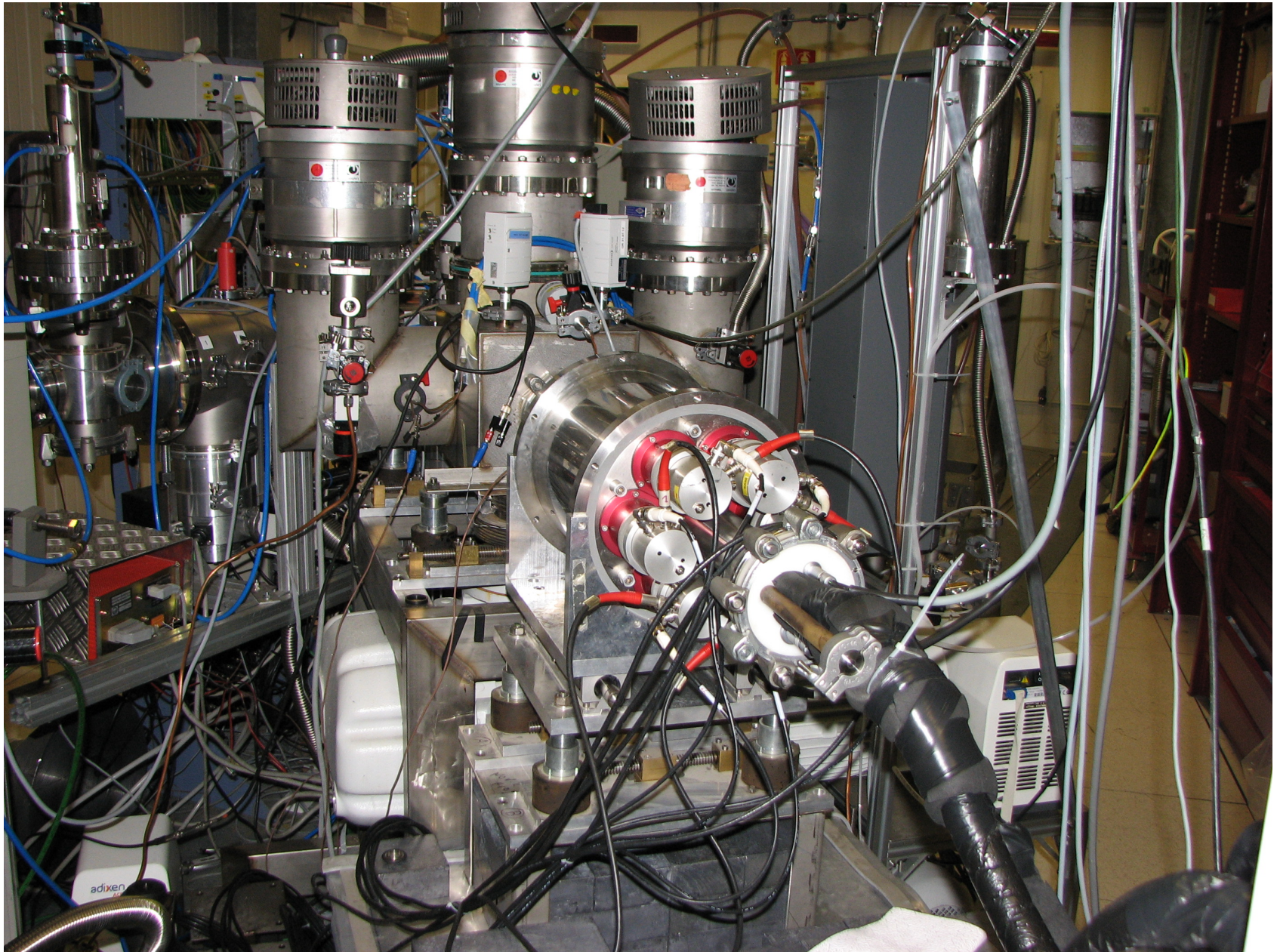
(b) Resonant capture at particular beam energies

- ◆  $^{19}\text{F}(p,\alpha\gamma)$  at 224 keV, 340 keV, and at higher energies
- ◆  $^{11}\text{B}(p,\gamma)$  at 163 keV, 600 keV, and at higher energies
- ◆  $^{15}\text{N}(p,\alpha\gamma)$  at 430 keV and 897 keV, and at higher energies

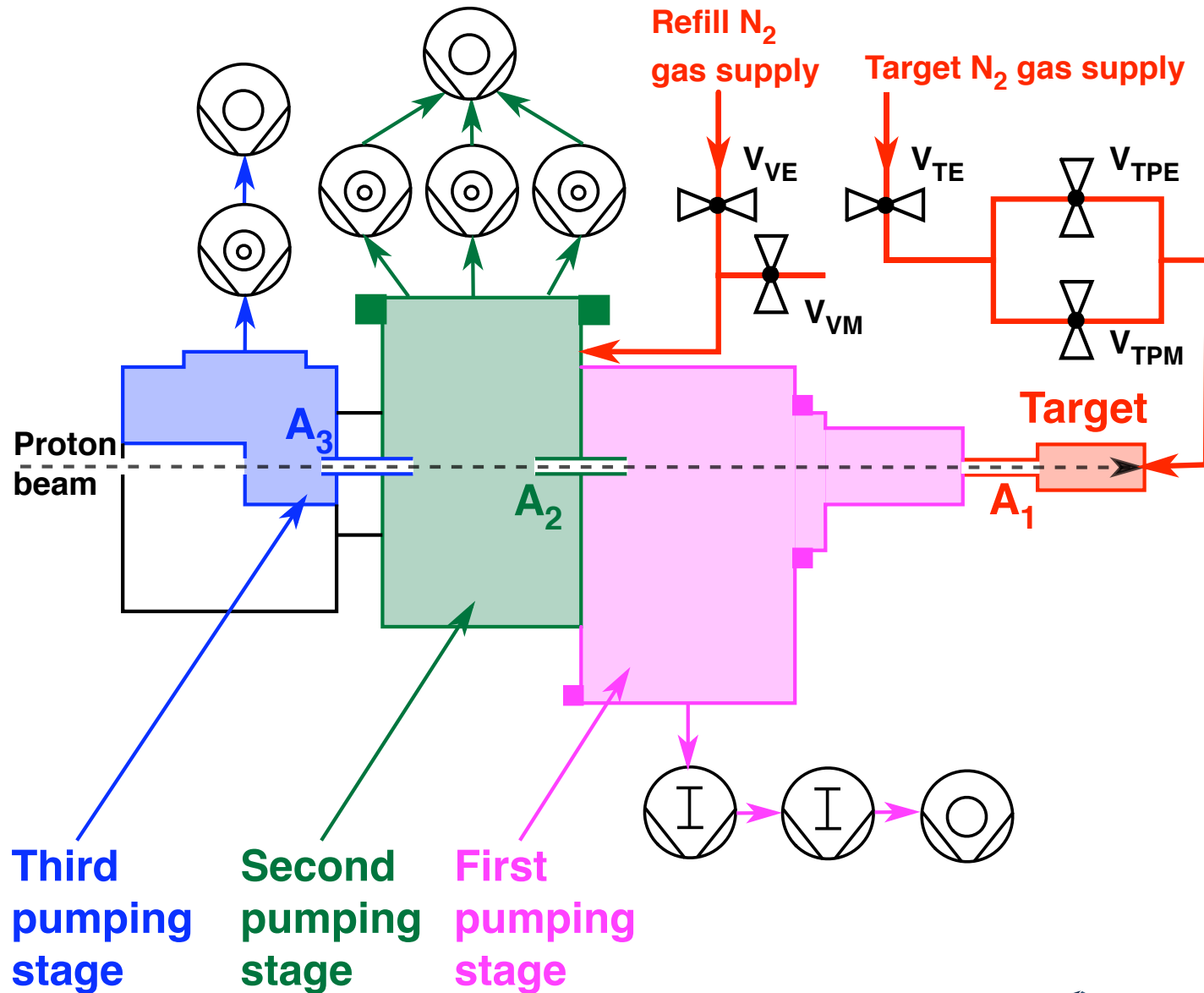
Gas targets are one way to reduce this beam induced background!



# Windowless gas target system at LUNA



# Pumping scheme



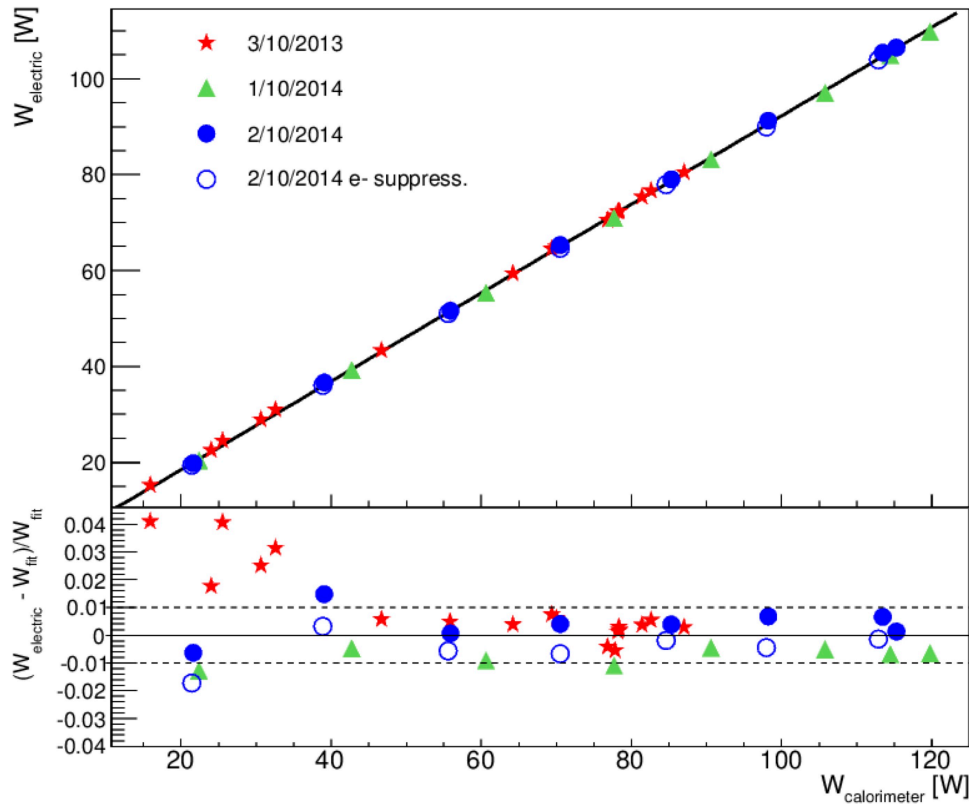
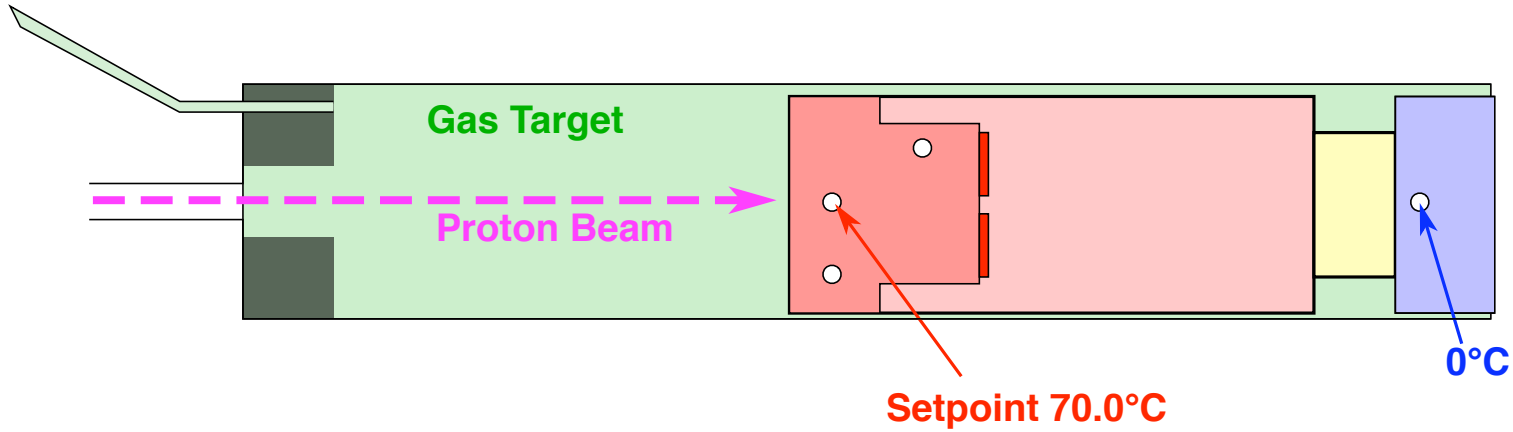
**Third  
pumping  
stage**

**Second  
pumping  
stage**

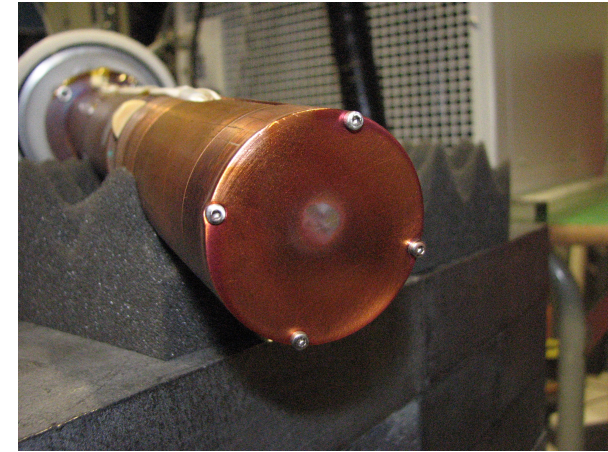
**First  
pumping  
stage**



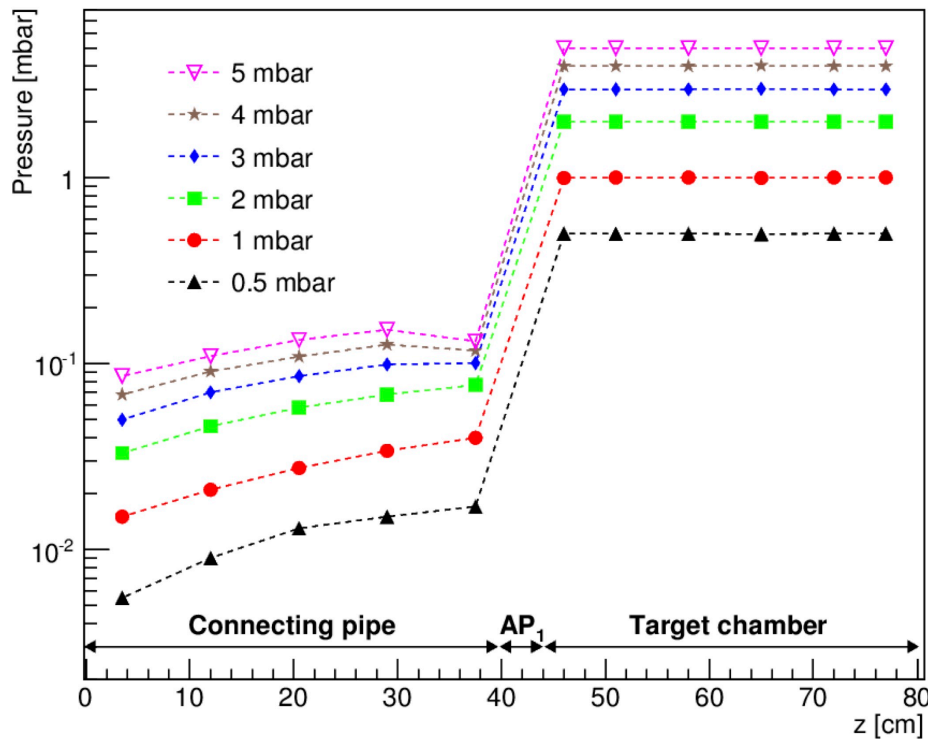
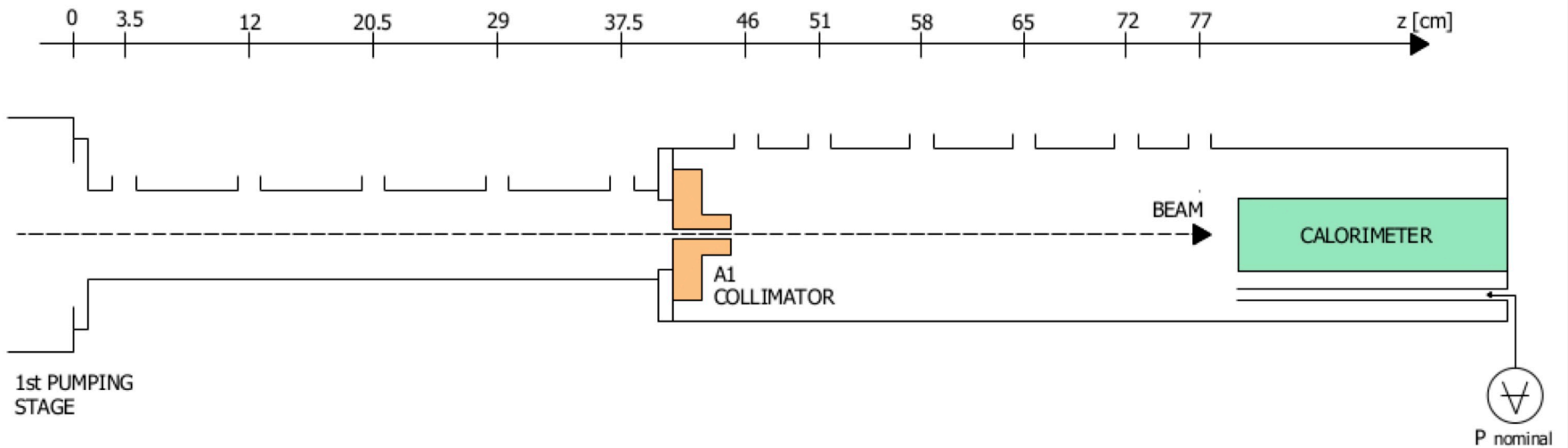
# Beam intensity measurement: Calorimeter with constant temperature gradient



## Electrical calibration

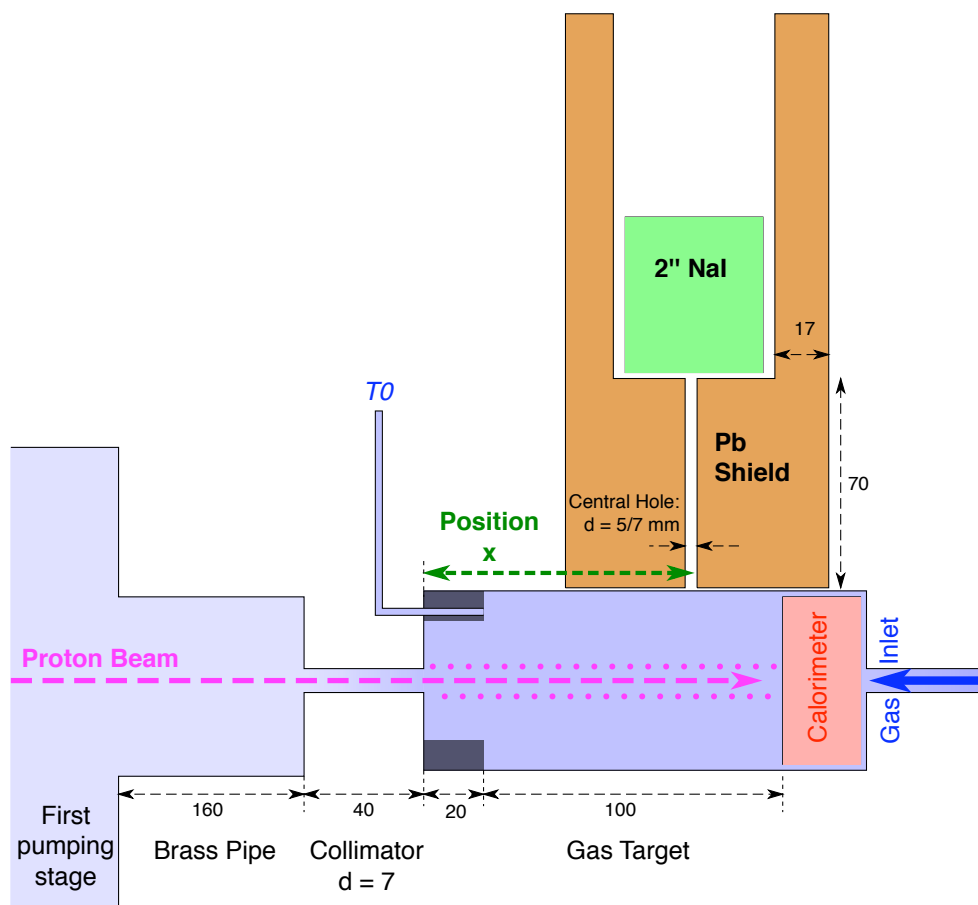


# Target density: Pressure measurement





# Target density: Effective target density (resonance scan technique)

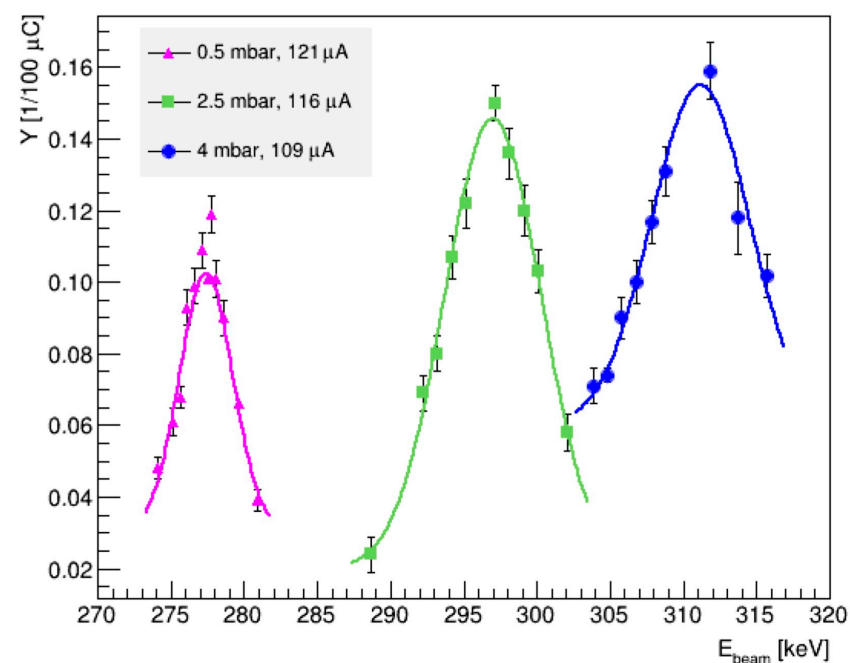


Thinning of the target by local (!) heating of the gas by the ion beam

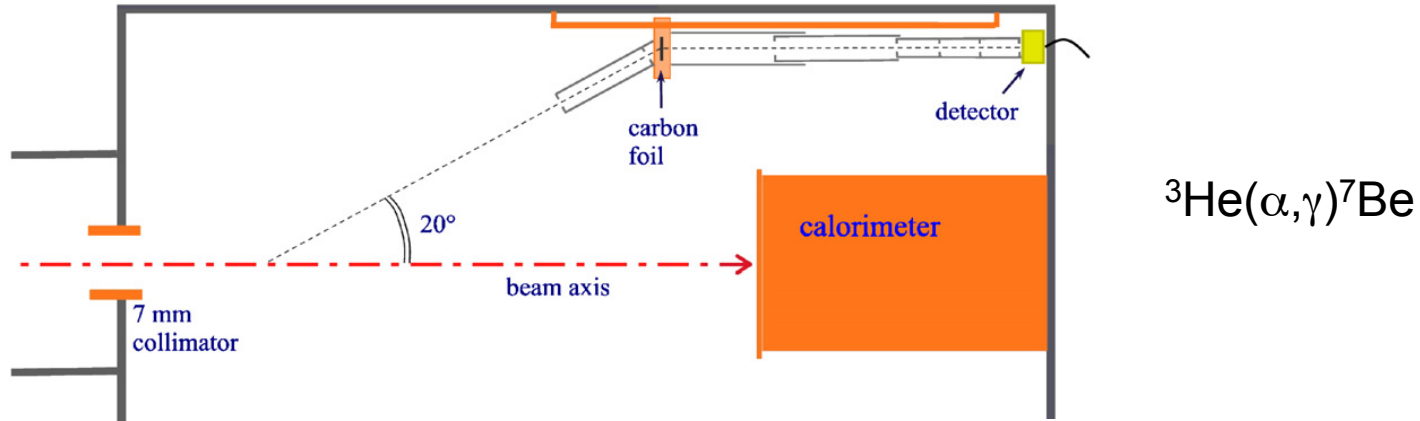
Need narrow resonances:

$^{14}\text{N}(p,\gamma)^{15}\text{O}$  at 278 keV

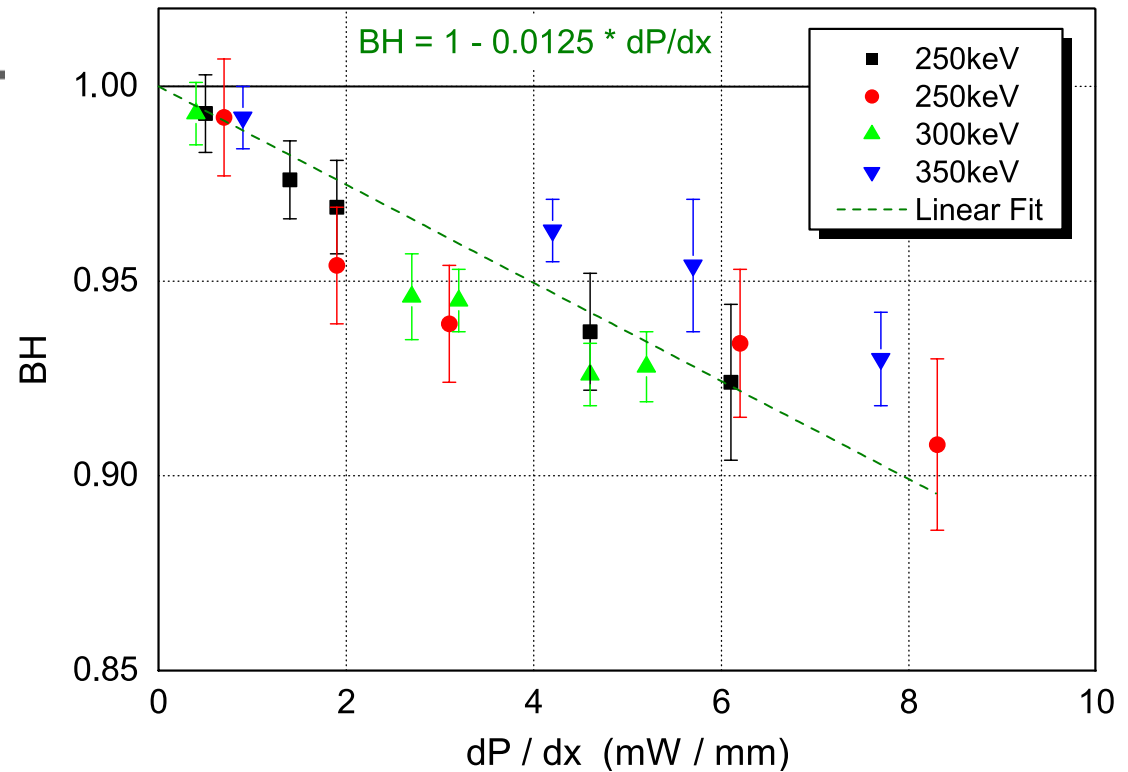
$^{21}\text{Ne}(p,\gamma)^{22}\text{Na}$  at 271 keV



# Target density: Effective target density (elastic scattering)

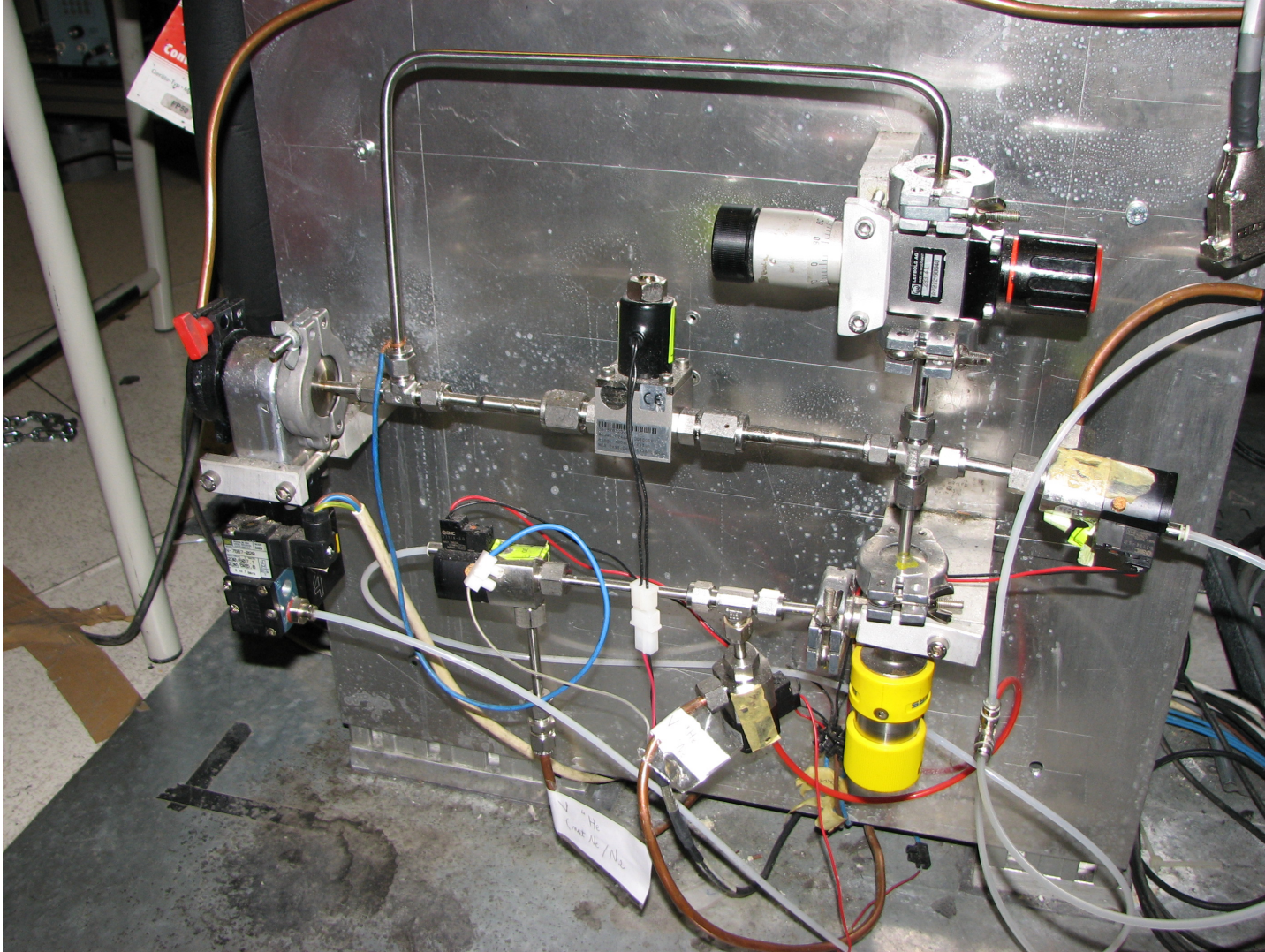


$$\frac{dW}{dx} = n_0 \cdot \frac{dE}{d\tilde{x}} \cdot I_{\text{target}}$$



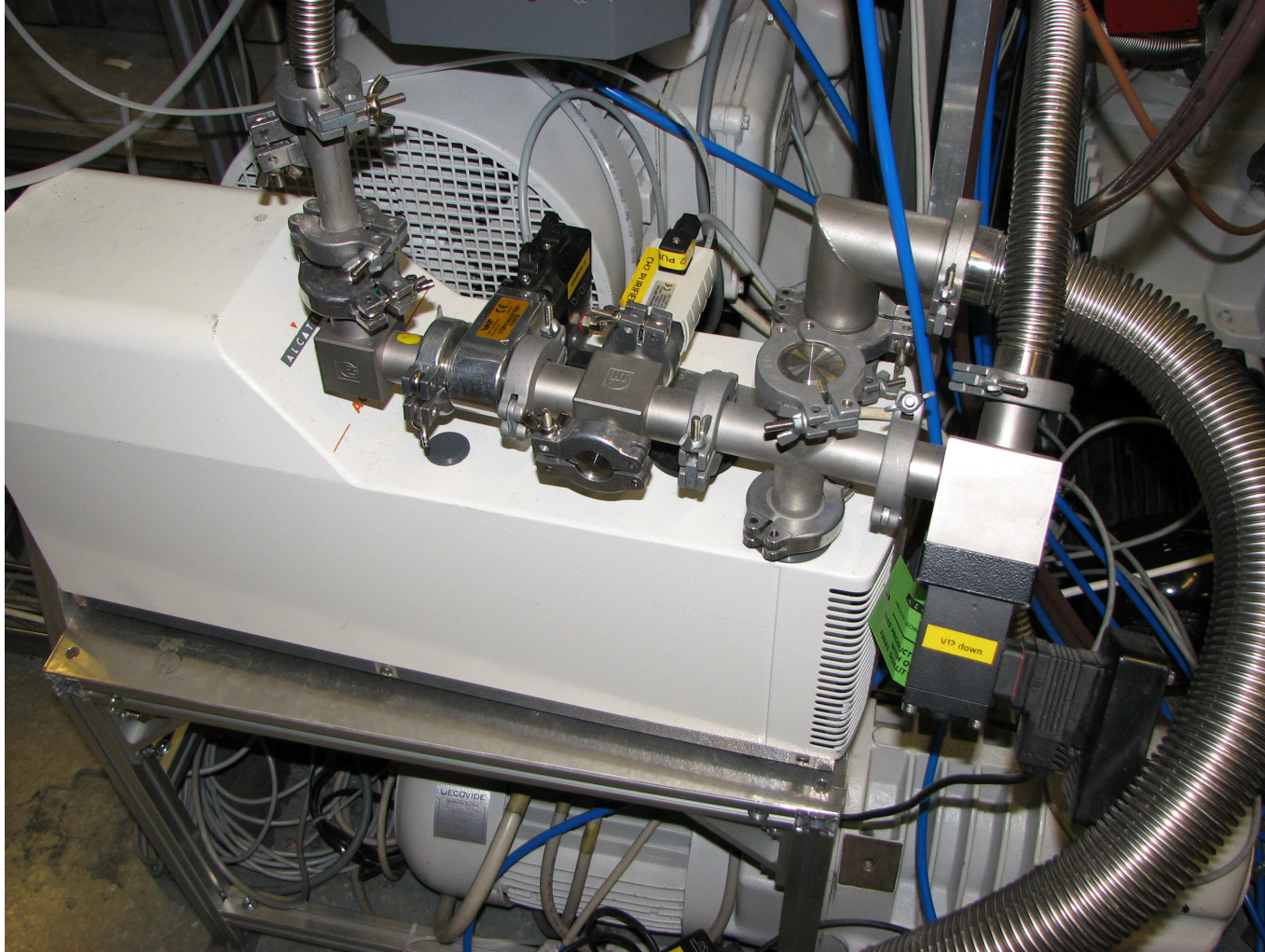


# Gas inlet regulation with manual offset and active feedback



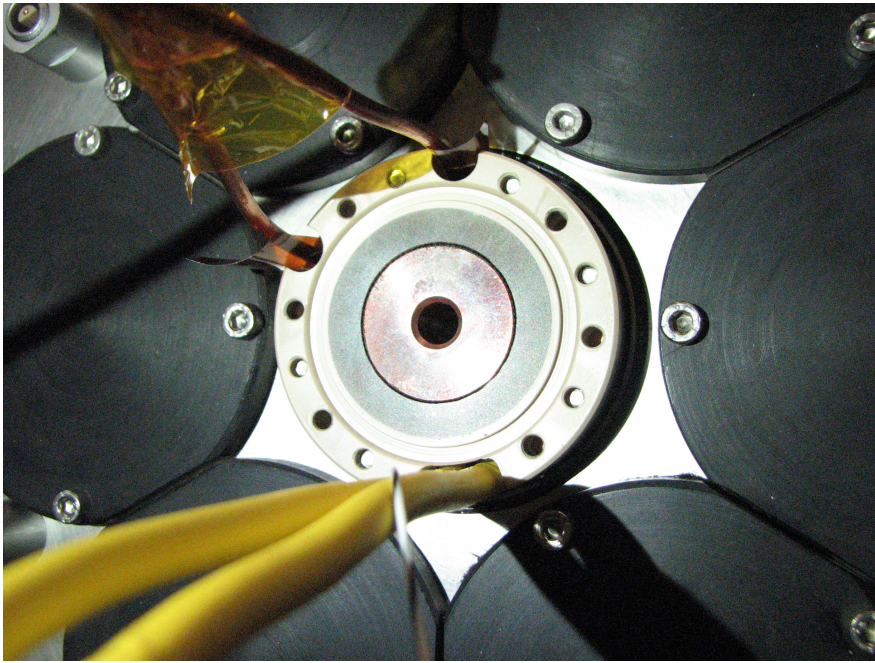


# “Oil-free” forepump

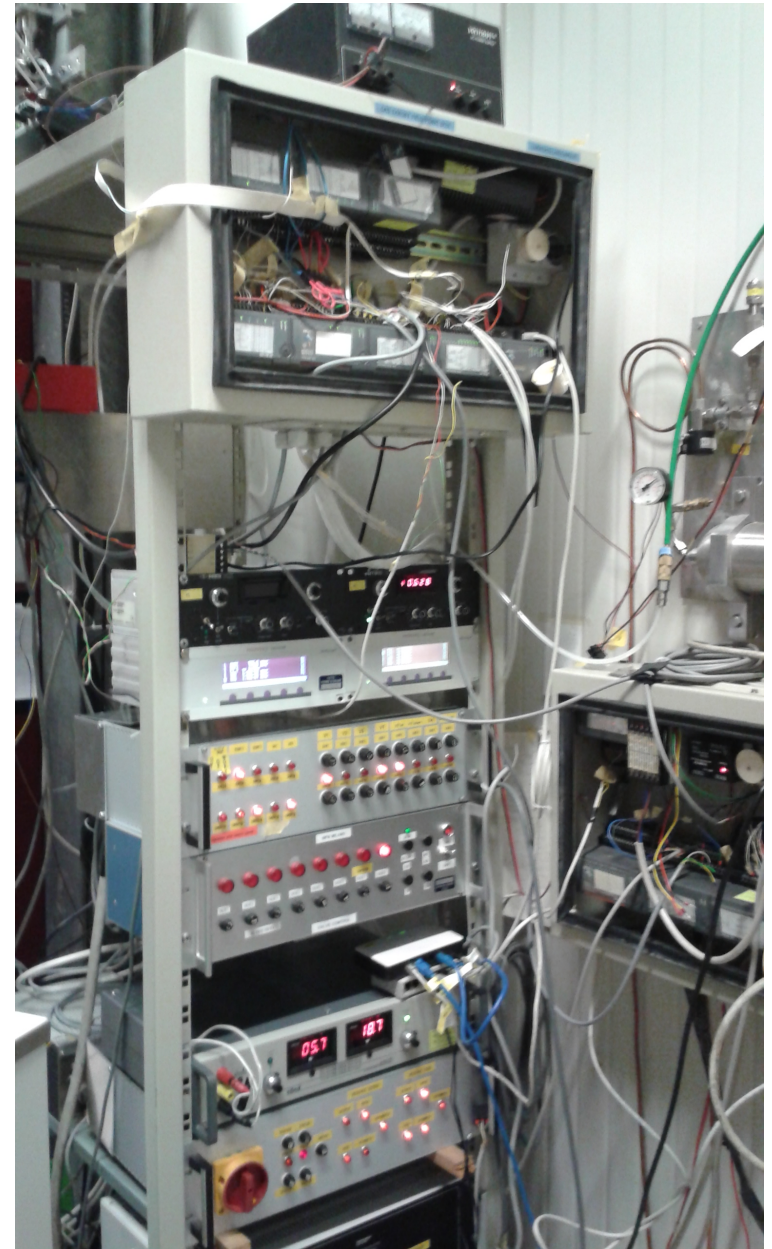




# Watercooled input detector



# Rack with slow control



# Windowless gas target systems, summary and lessons learned

Gas targets were instrumental in several highly successful LUNA measurement campaigns:

- ◆  $^{nat}\text{N}$  for  $^{14}\text{N}(p,\gamma)^{15}\text{O}$  and  $^{15}\text{N}(p,\gamma)^{16}\text{O}$
- ◆  $^3\text{He}$  for  $^3\text{He}(\alpha,\gamma)^7\text{Be}$  and  $^3\text{He}(^3\text{He},2p)^4\text{He}$
- ◆  $^2\text{H}$  for  $^2\text{H}(\alpha,\gamma)^6\text{Li}$  (Davide Trezzi's talk yesterday)
- ◆  $^{22}\text{Ne}$  for  $^{22}\text{Ne}(p,\gamma)^{23}\text{Na}$  (Federico Ferraro's talk yesterday)

## Lessons

- ◆ Try to keep your setup free from water (deuterium), oil ( $^{12}\text{C}$ ,  $^{13}\text{C}$ ), teflon ( $^{19}\text{F}$ )
- ◆ New problem boron, not yet clear where it is coming from

You need many pumps!

- ◆ Sensitive: high pumping speed turbomolecular pumps
- ◆ New turbopumps can handle  $10^{-2}$  mbar pressure over longer periods

You need a recirculation system for expensive gases

- ◆ Chemical getters work fine
- ◆ Monitor either by nuclear reactions, or by mass spectrometer