The AMS02 Spectrometer: first data and detector performance



G. Ambrosi

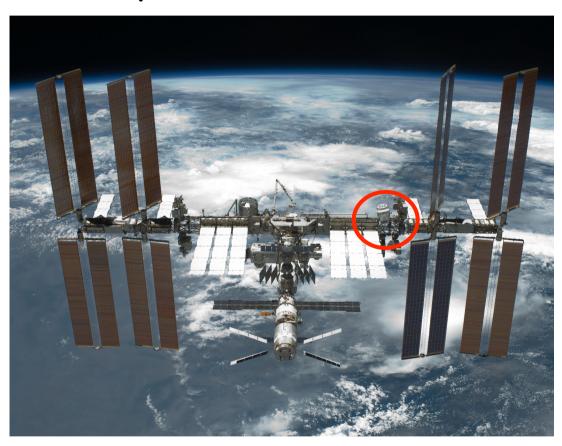
on behalf of the Tracker group (Perugia INFN and University, Geneva University)



AMS on the International Space Station



- Cosmic Antimatter search with 10⁻⁹ sensitivity
- Indirect Dark Matter search (e⁺, p
 , γ)
- Relative abundance of nuclei and isotopes in primary cosmic rays
- γ ray astrophysics

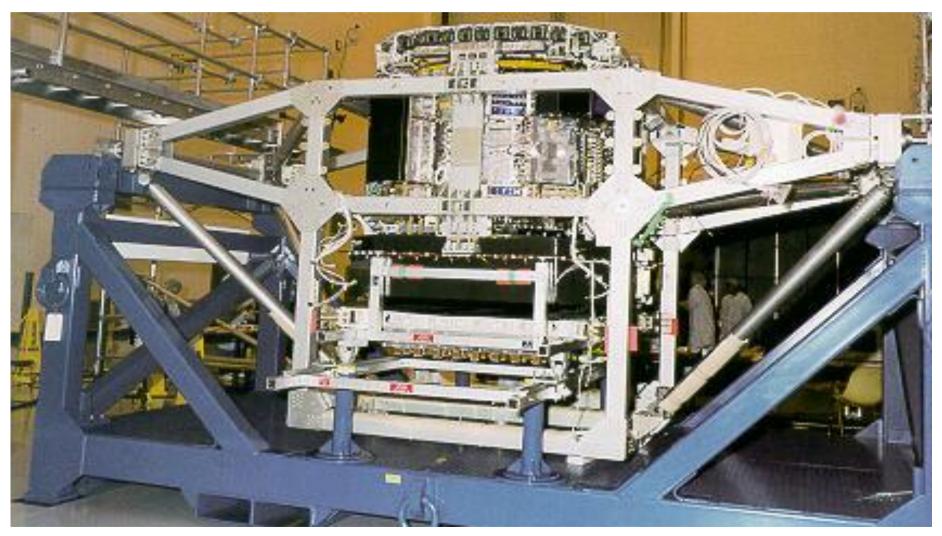


The purpose of the AMS experiment is to perform accurate, high statistics, long measurements of charged cosmic rays (0.5 GV - 1 TV) and γ rays (E>1GeV)





AMS01 at KSC (Florida) in 1998





AMS-01 pilot experiment: STS91, June 2nd - 12th 1998



- 10 days of data taking in orbit:
 - 400 Km altitude
 - latitudes +51.7°
 - all longitudes
- 10⁸ events recorded
- Physics results (Phys. Rep. 366 (2002) 331)
 - precise measurements of primary fluxes
 - detection of secondary fluxes (quasi trapped)
 - antimatter limit at 10⁻⁶

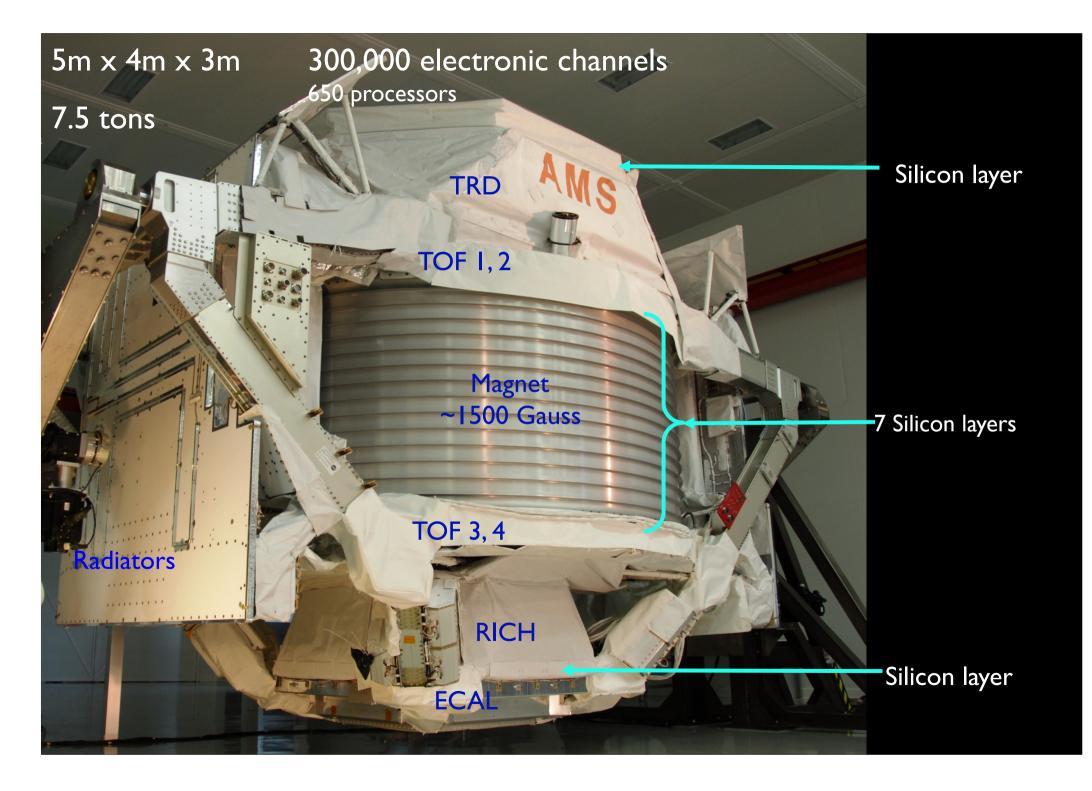


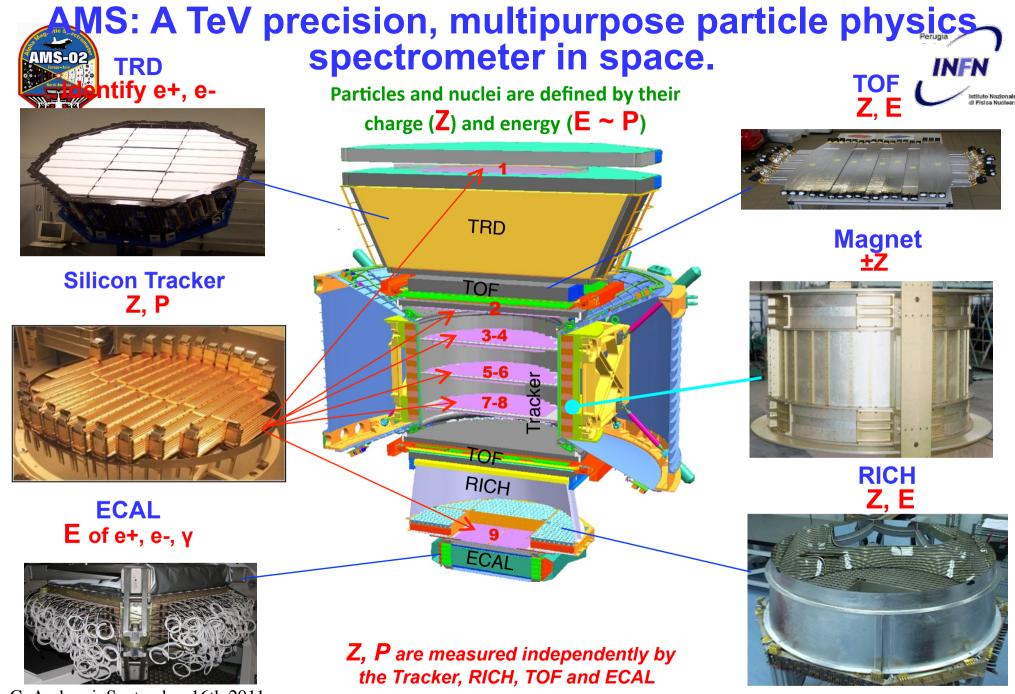




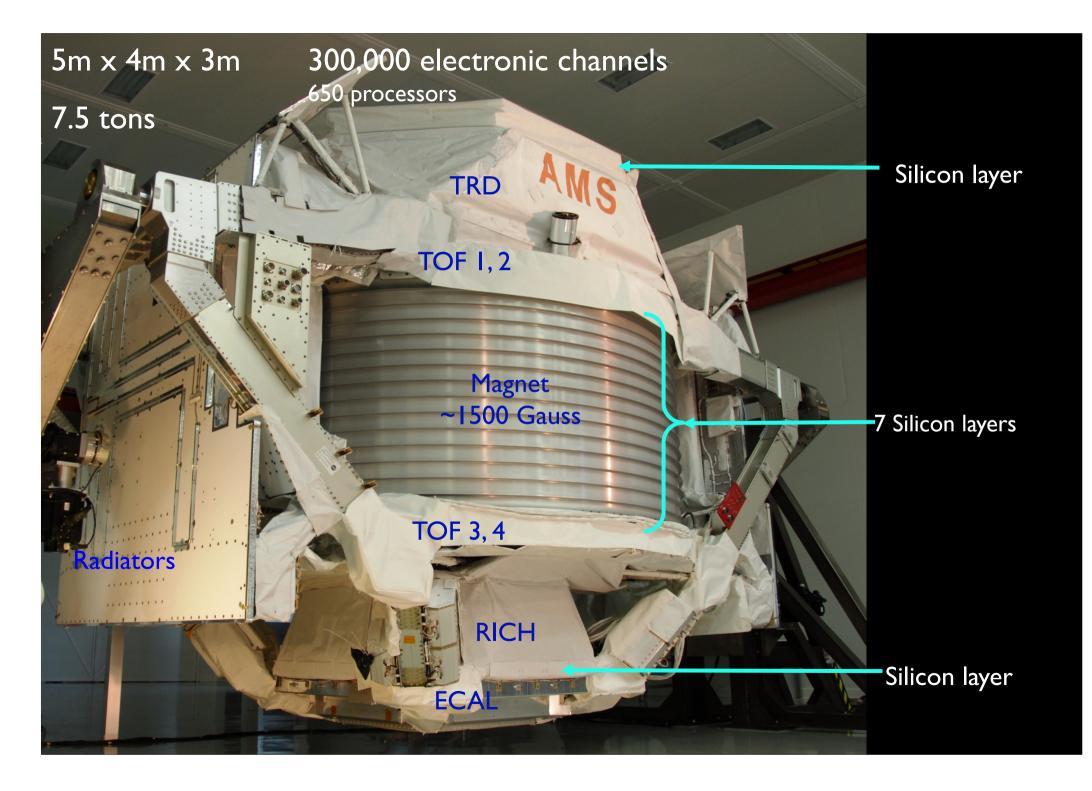


- performance a la `particle physics':
 - high resolution measurements of momentum, velocity, charge and energy
- characteristics to properly work in the space environment:
 - Vibration (6.8 G rms) and acceleration (17 G)
 - Temperature variation (day/night $\Delta T = 100^{\circ}C$)
 - Vacuum (10⁻¹⁰ Torr)
 - Orbital debris and micrometeorites
 - Radiation (Single Event Effect)
- limitation in weight (15000 lb), power (~2KW), bandwidth and maintenance
- Compliant with Electromagnetic Interference and Electromagnetic Compatibility specs





G. Ambrosi, September 16th 2011







	e -	Ρ	He,Li,Be,Fe		Y	e+	Ē, D	He, C
TRD		r	7				•	r
TOF	Ŧ	T	ř	•		*	ţ	ř
Tracker	J			八				ノ
RICH	······.		$\langle \cdots \rangle \longrightarrow \langle \cdots \rangle$					
ECAL		******	Ŧ				****	¥
Physics example	Cosmic Ray Physics					Dark matter Antimatter		





- fall 2009: integration at CERN
- February 2010: test beam at CERN
- spring 2010: EMI and TV test at ESTEC (ESA)
- late spring 2010: magnet replacement at CERN
- August 2010: test beam at CERN
- fall/winter/spring 2010-2011 integration at KSC (Florida)
- May 16th 2011: launch!
- May 19th 2011: first activation in space: everything is working!!





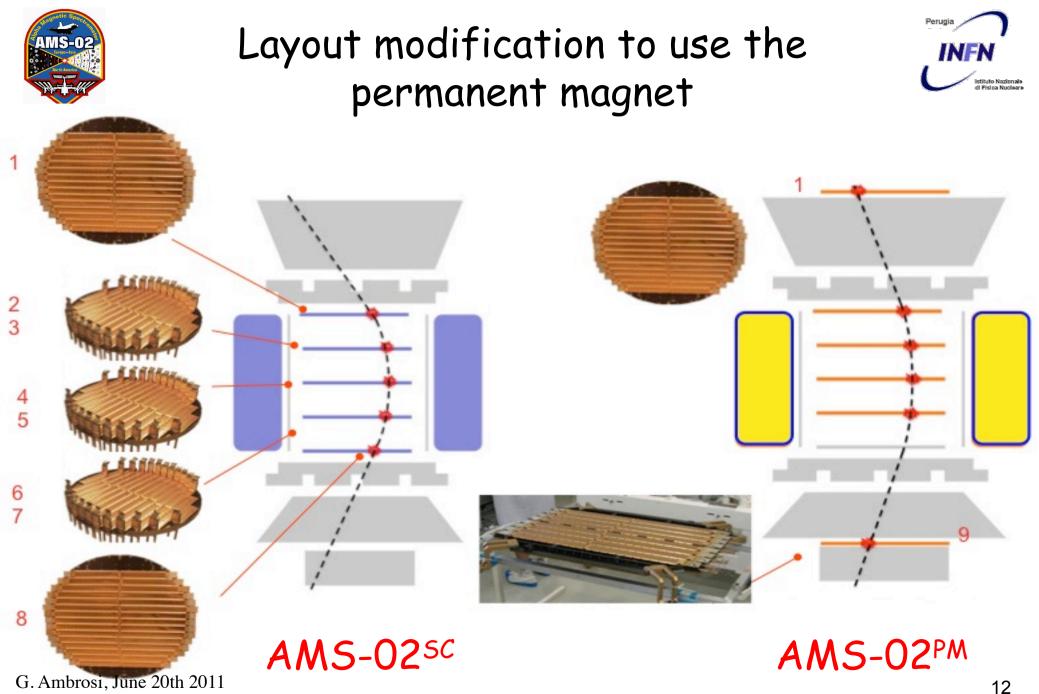
recent AMS02 history in short

- fall 2009: integra
- February 2010: t
- spring 2010: EMI
- late spring 2010:
- August 2010: tes
- fall/winter/spring (Florida)
- May 16th 2011: la
- May 19th 2011: f everything is wor



(ESA) at CERN

n at KSC



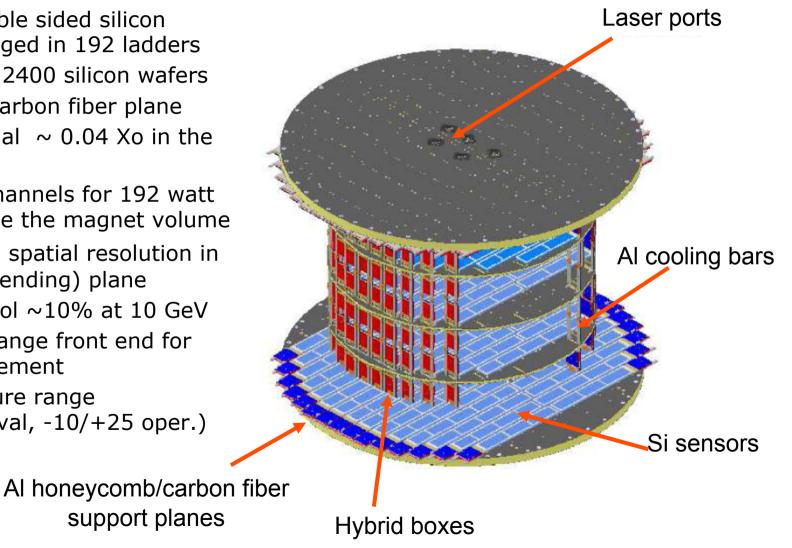


Silicon Tracker



- 9 layers of double sided silicon detectors arranged in 192 ladders for a total of ~ 2400 silicon wafers
- 6 honeycomb carbon fiber plane
- detector material ~ 0.04 Xo in the inner tracker
- total of 200 kchannels for 192 watt dissipated inside the magnet volume
- $10 \ \mu m$ (30 μm) spatial resolution in bending (non bending) plane
- momentum resol ~10% at 10 GeV
- high dynamic range front end for charge measurement
- wide temperature range (-20/+40 survival, -10/+25 oper.)

support planes

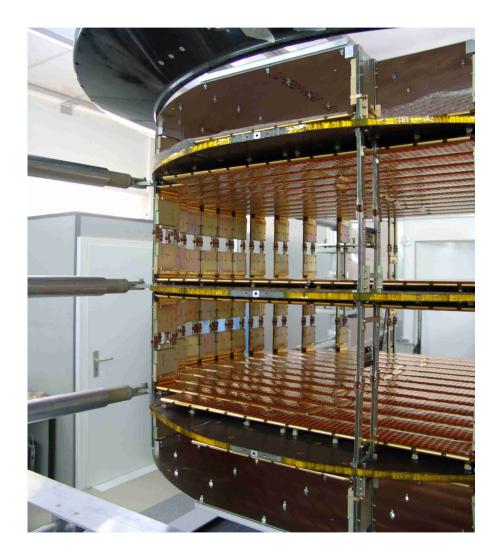


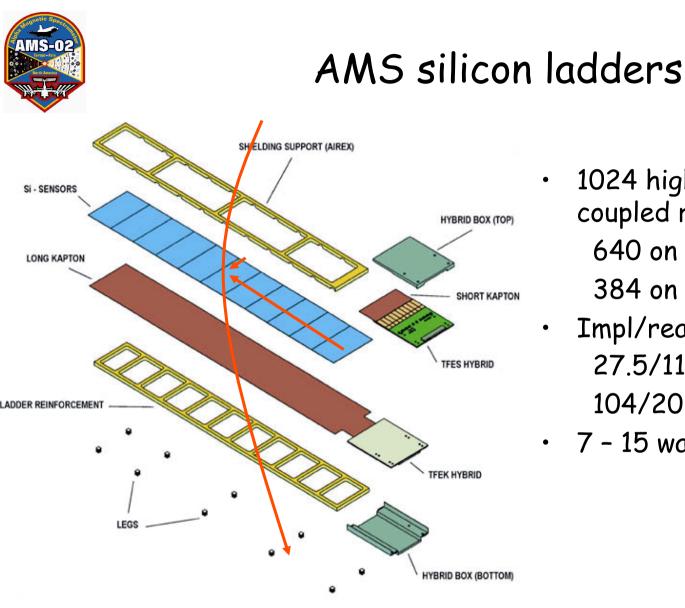


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- 6 honeycomb carbon fiber plane
- detector material ~ 0.04 Xo in the inner tracker
- total of 200 kchannels for 192 watt dissipated inside the magnet volume
- 10 μm (30 μm) spatial resolution in bending (non bending) plane
- momentum resol ~10% at 10 GeV
- high dynamic range front end for charge measurement
- wide temperature range (-20/+40 survival, -10/+25 oper.)





- 1024 high dynamic range, AC coupled readout channels:
 640 on junction (S) side
 384 on ohmic (K) side
- Impl/readout pitch:
 27.5/110 μm (S side)
 104/208 μm (K side)

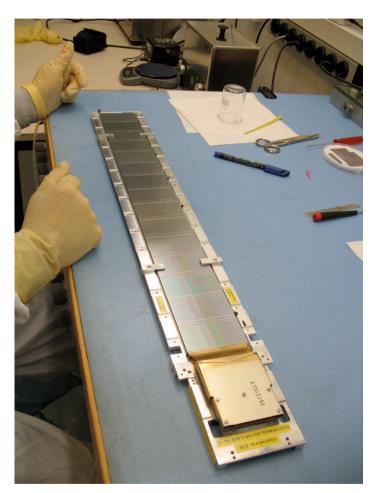
192 flight units, 210 assembled in 3 lines: Perugia (I), Geneva-ETHZ (CH), G&A (Carsoli, I)

Peruni



AMS silicon ladders





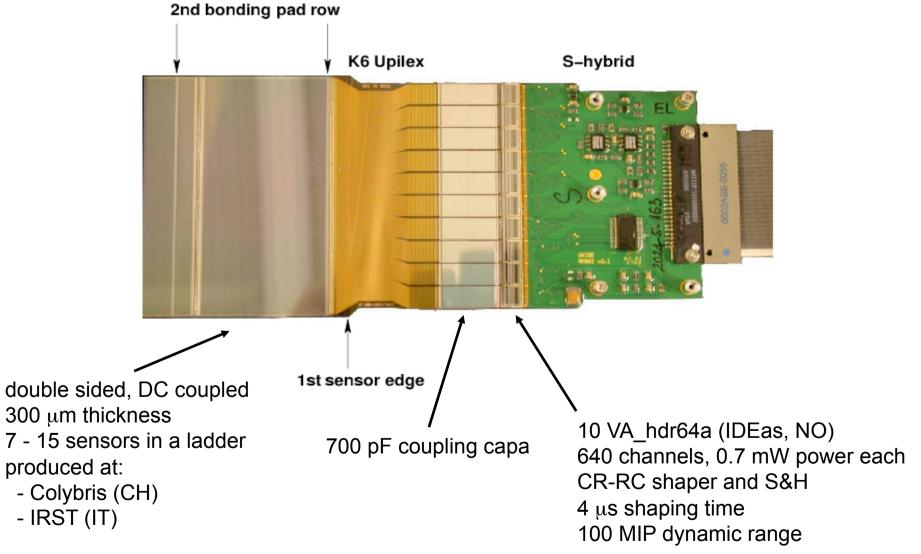
- 1024 high dynamic range, AC coupled readout channels:
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- Impl/readout pitch: 27.5/110 μm (S side) 104/208 μm (K side)
- 7 15 wafers (28 60 cm)

192 flight units, 210 assembled in 3 lines: Perugia (I), Geneva-ETHZ (CH), G&A (Carsoli, I)

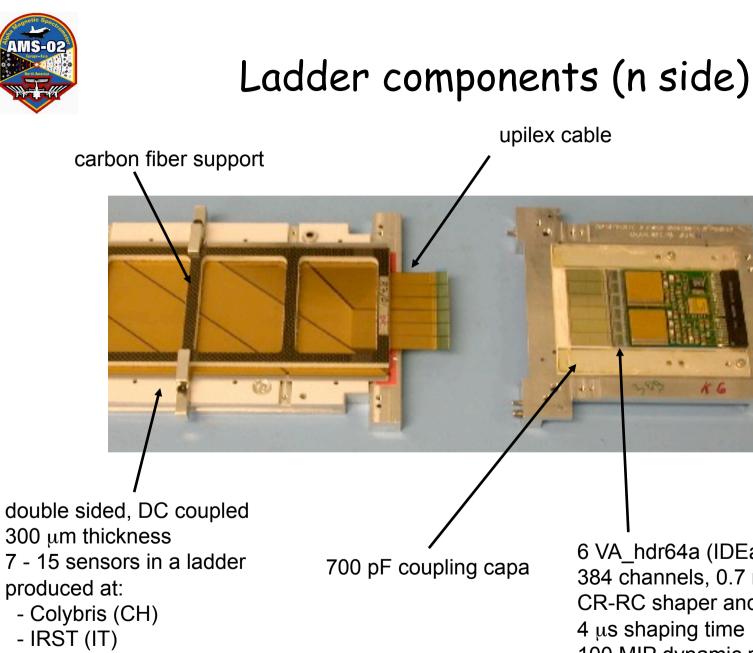




Ladder components (p side)



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6 VA_hdr64a (IDEas, NO) 384 channels, 0.7 mW power each CR-RC shaper and S&H 4 μs shaping time 100 MIP dynamic range

Perugi

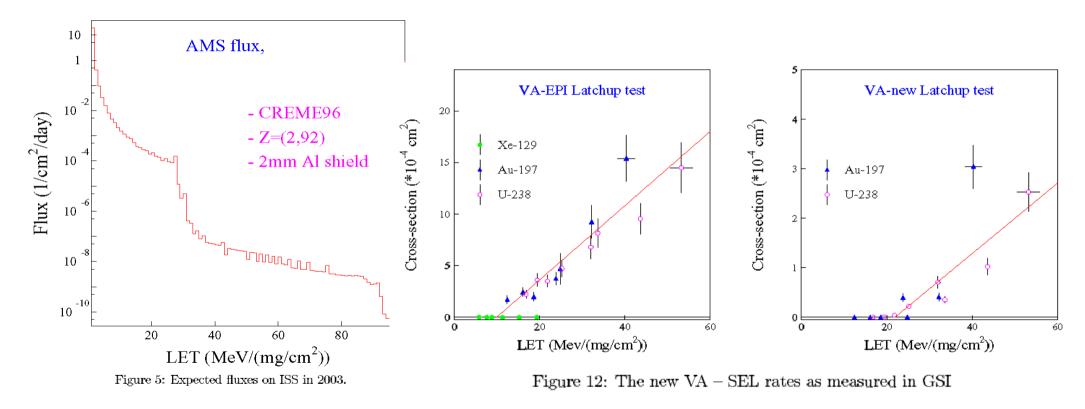
INEN





Radiation 'hard' electronics

The problem are the SEE (Single Event Effect)



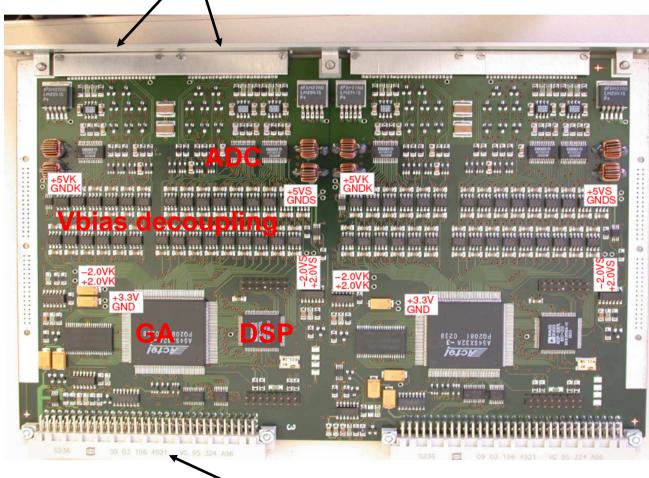
current limit protection is present for all active components



Data Reduction Board (TDR2)



analog signal in



compressed digital out

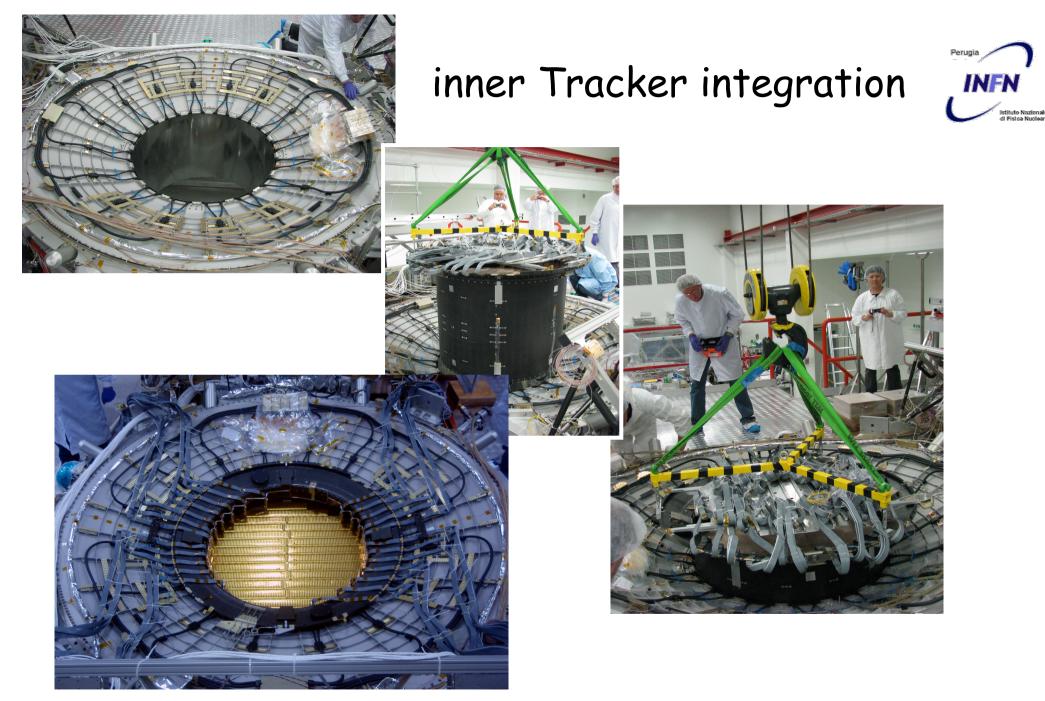
Collect analog data and digitize it (100 µs irred. dead time)

Perform online data compression

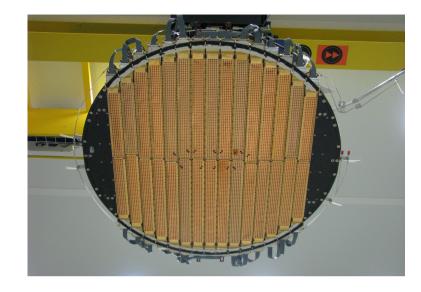
- Remove Pedestals
- Calculate and Remove Common Noise
- Search Clusters

Up to 5 KHz trigger rate in compressed mode

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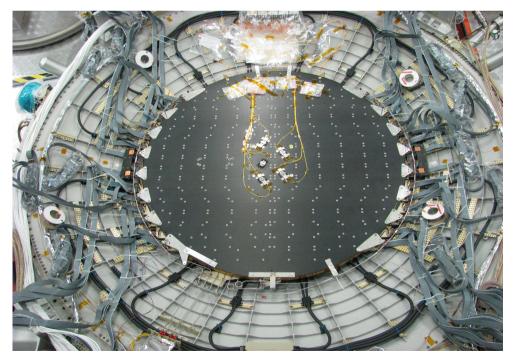














Tracker integration (2)

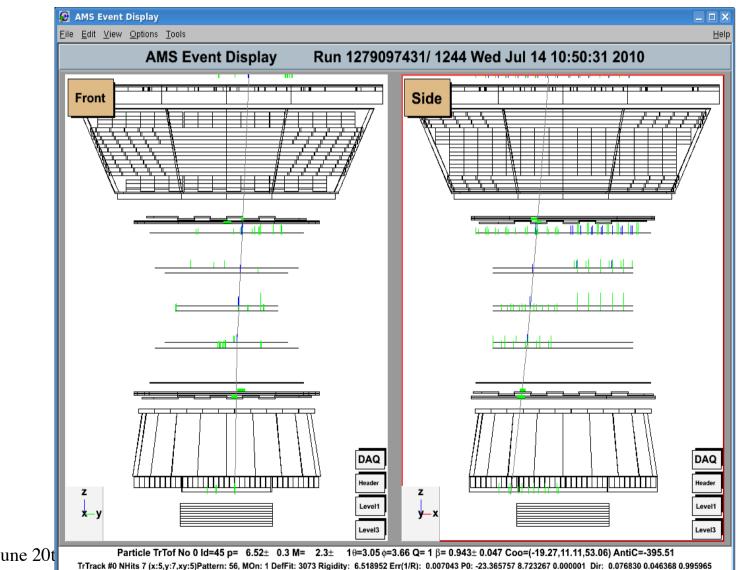






the first muon with the new Tracker

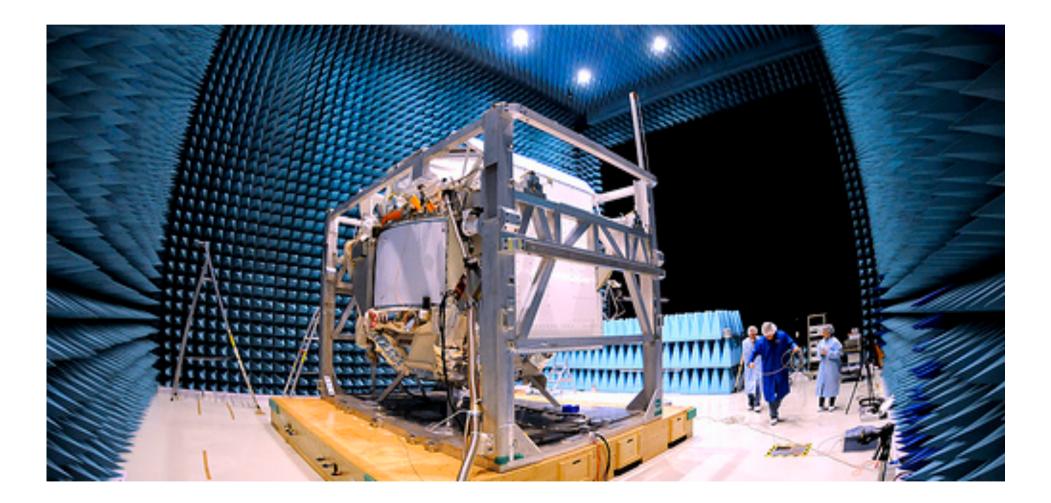




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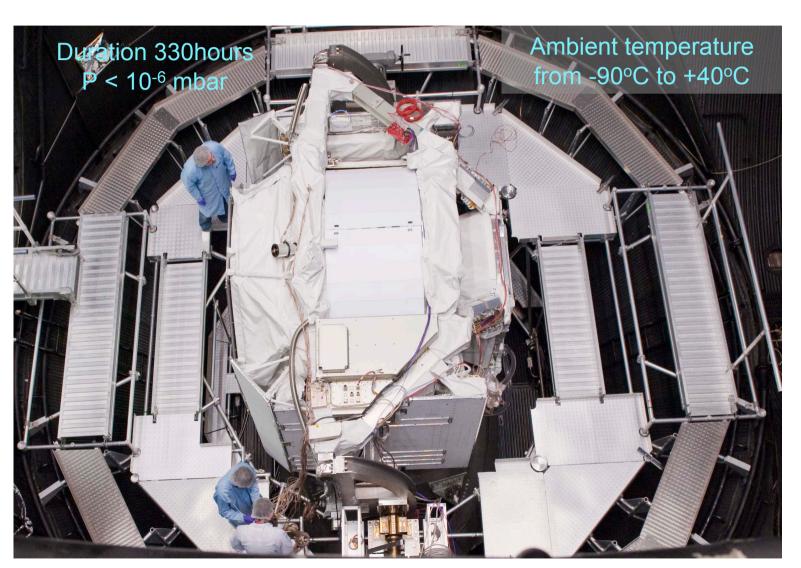


AMS in the ESA EMI Chamber, March 2010, ESTEC, Noordwijk, Netherlands





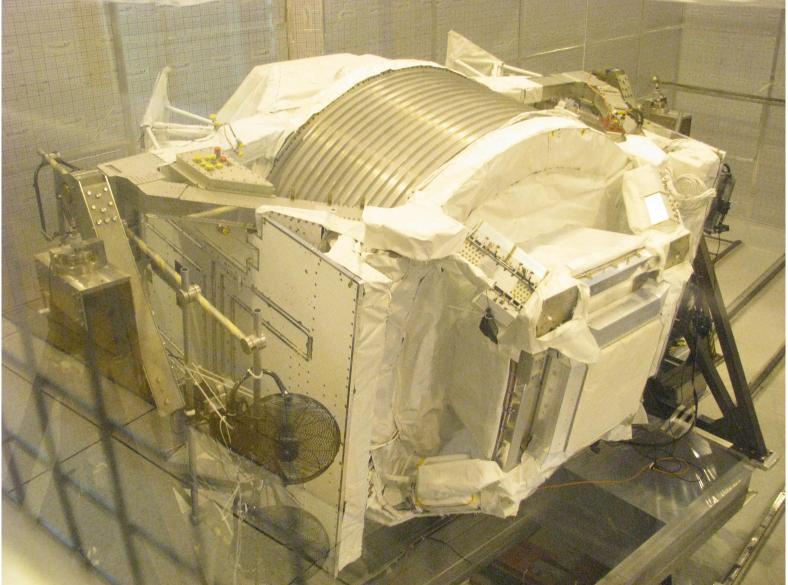
AMS in the ESA TVT Chamber, April 2010, ESTEC



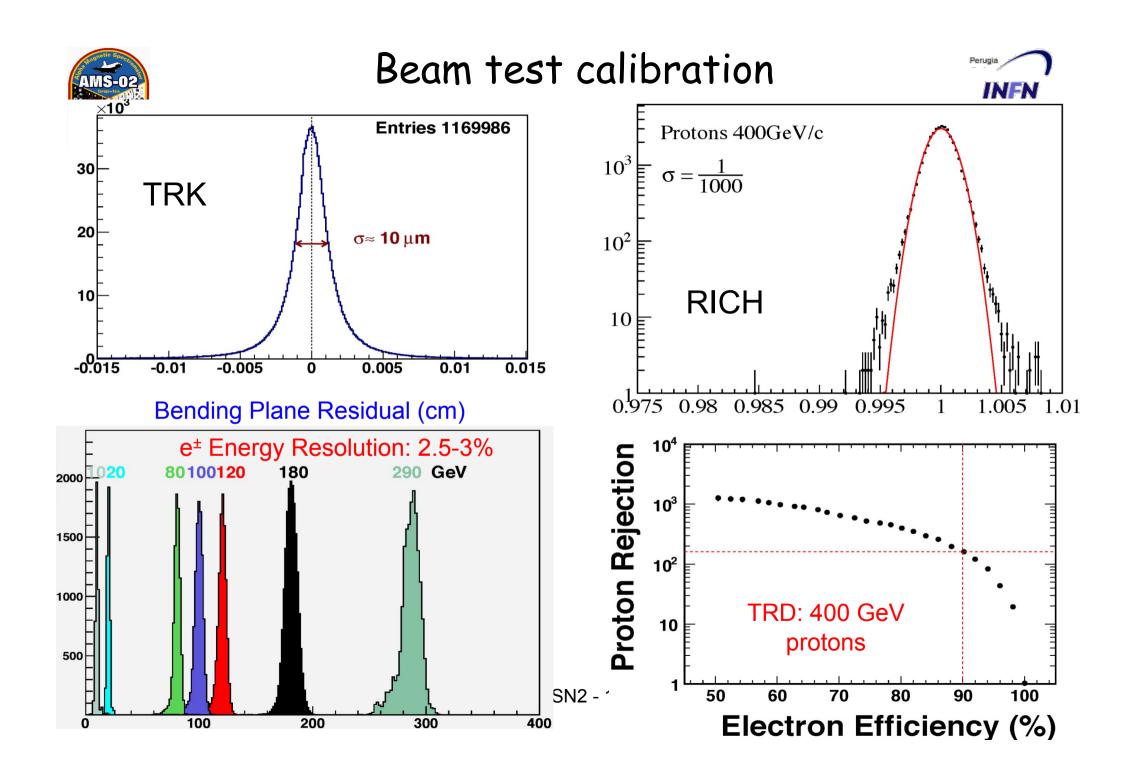
Istituto Nazionale



AMS in Test Beam - 8-20 Aug 2010, CERN



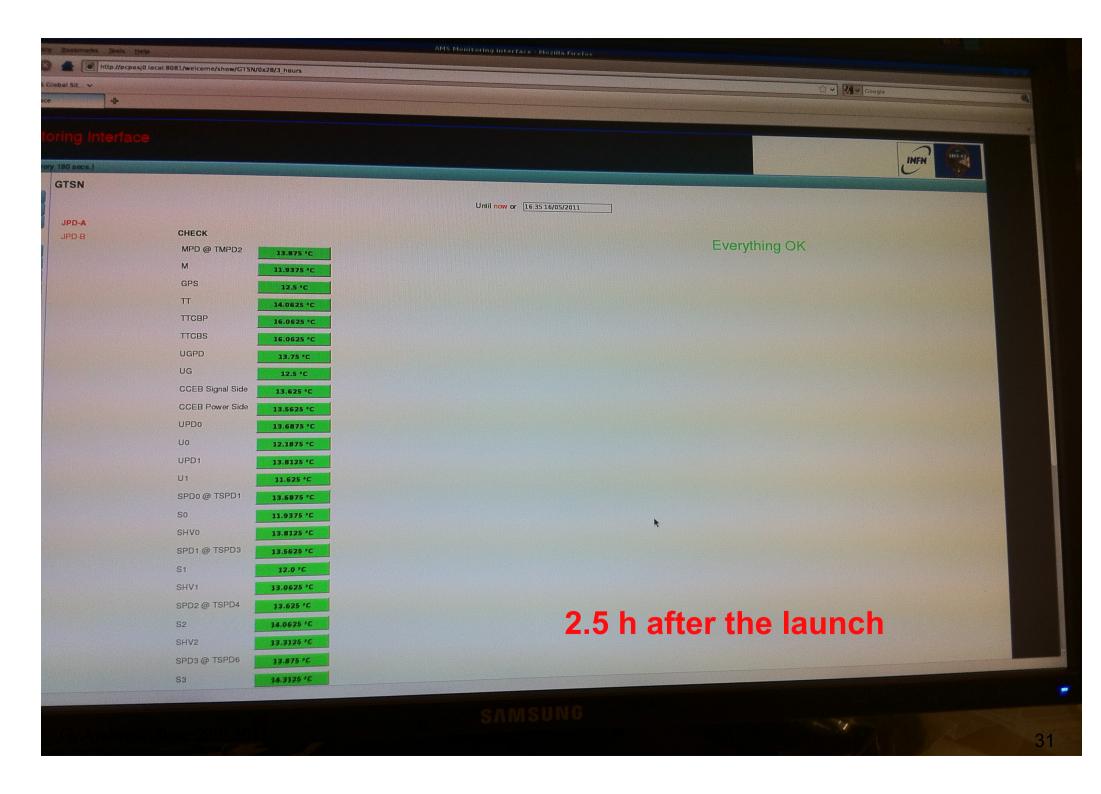
Istituto Nazionale







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AMS-02 : the Collaboration @ NASA-JSC

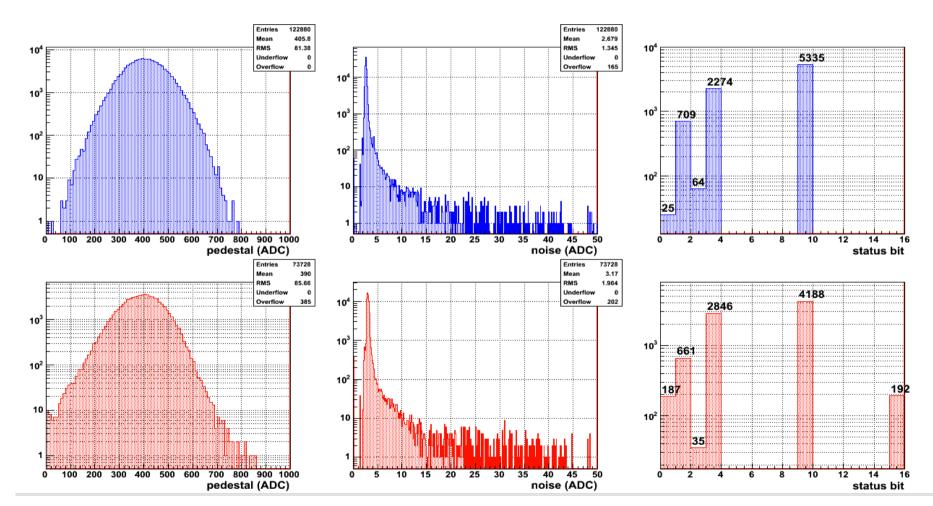






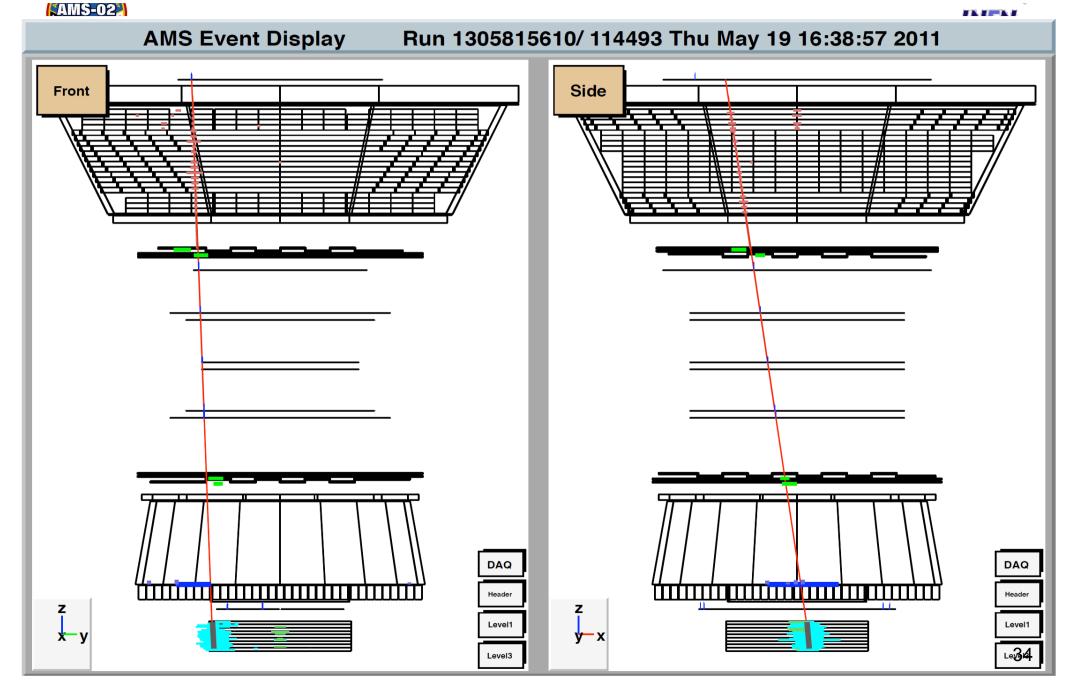
First Tracker calibration in space



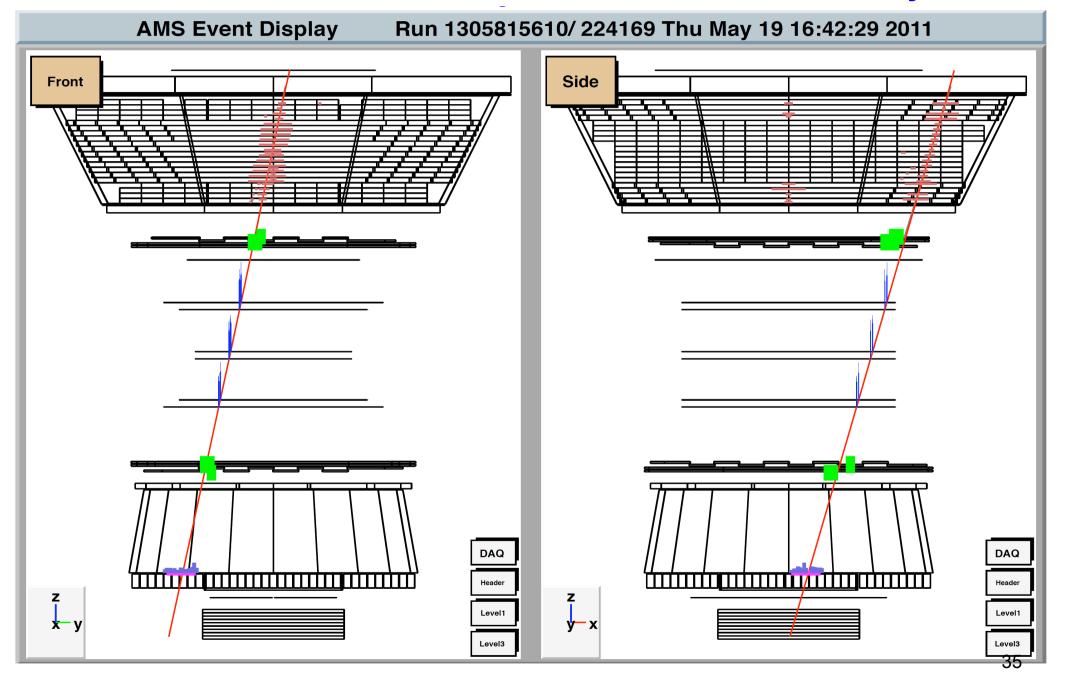


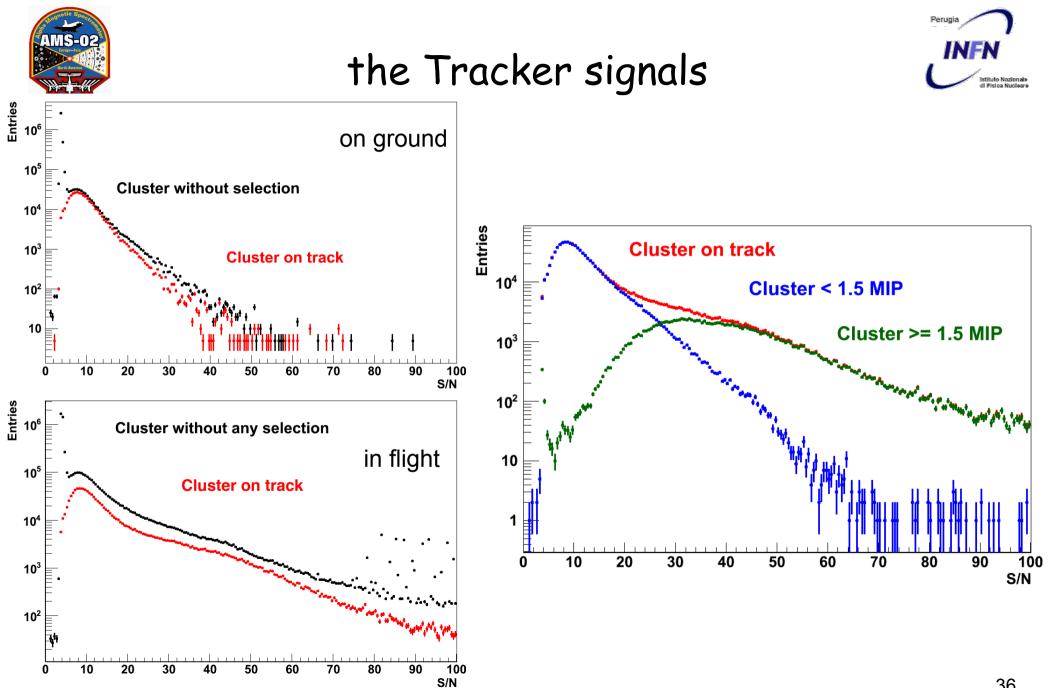
G. Ambrosi, September 16th 2011

Data from the 1st few minutes – 20 GeV Electron, 19 May 2011



Data from the 1st few minutes – 42 GeV/c Carbon, 19 May 2011

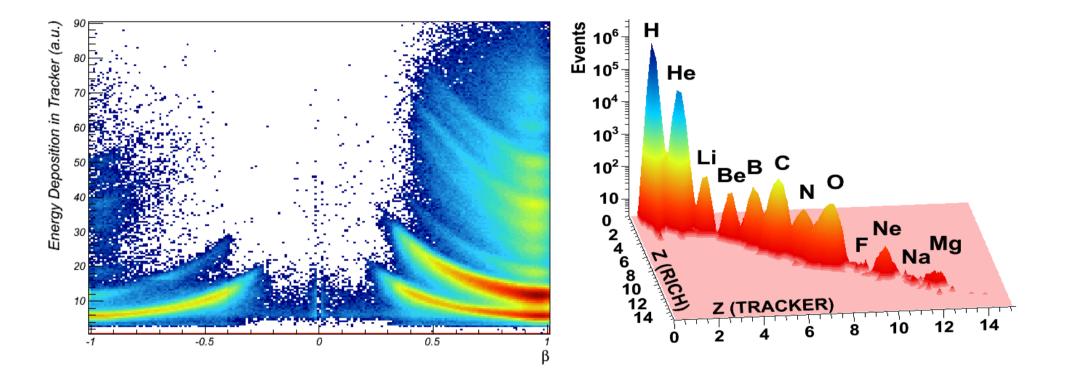






the particles we see

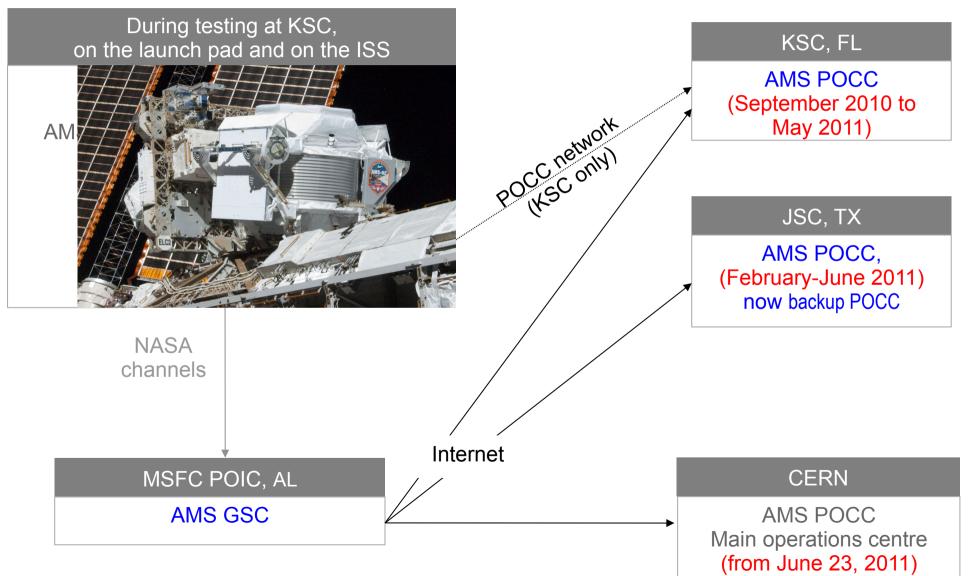






AMS-02 : the flight experience

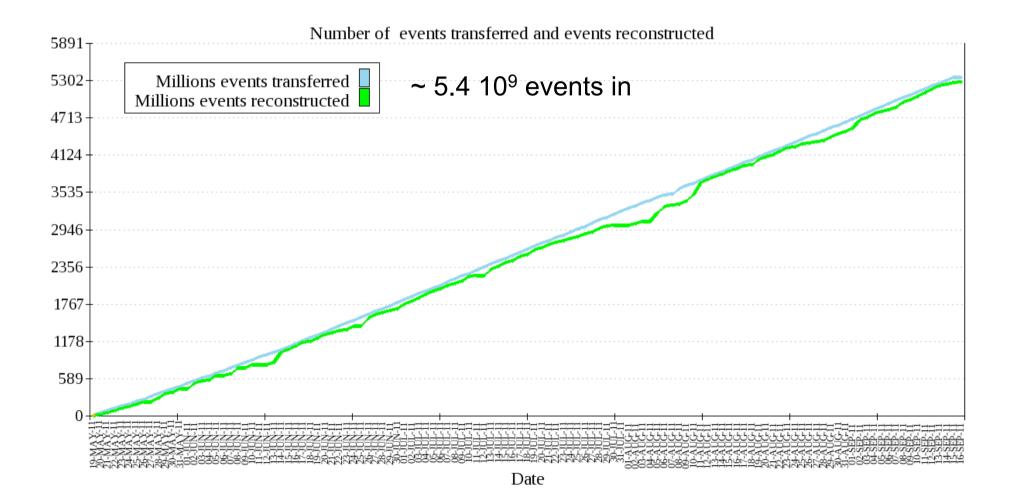






DAQ and data processing: routine operations

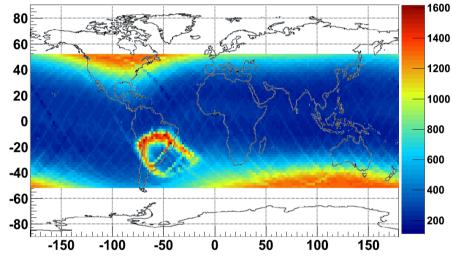


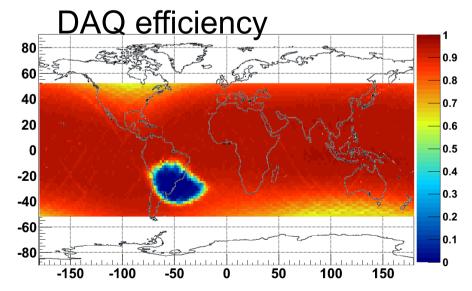




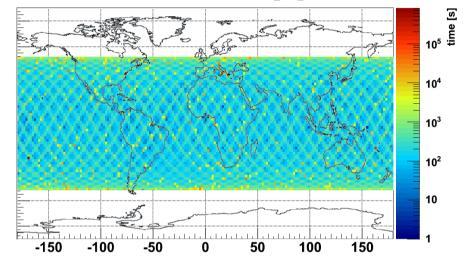
DAQ orbital parameter







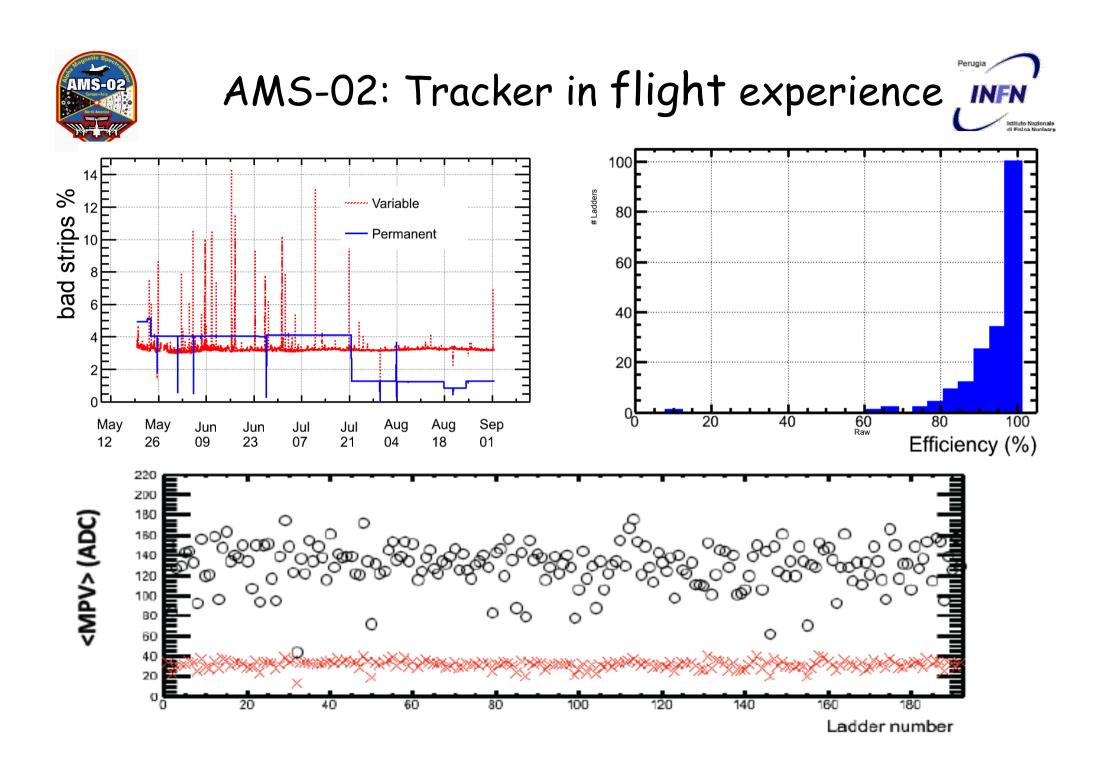
Time at location [s]



Particle rates vary from 200 to 2000 Hz per orbit

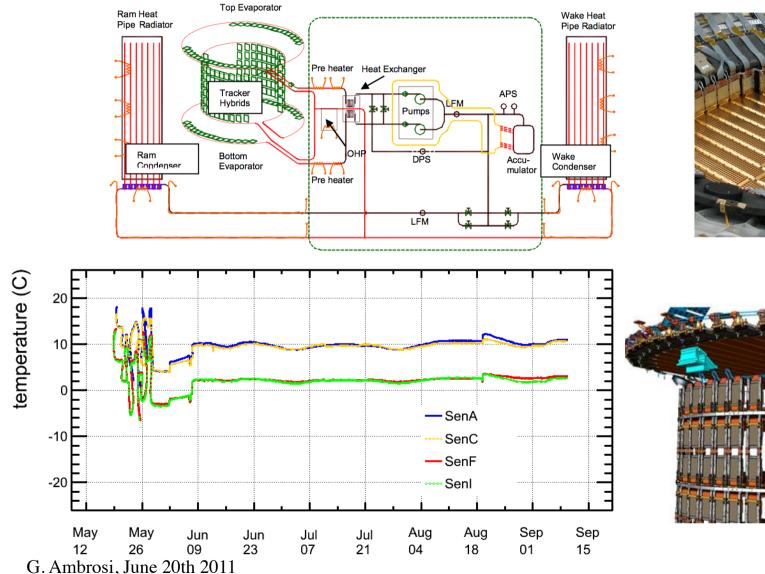
On average: DAQ efficiency 85% DAQ rate ~700Hz Perugia

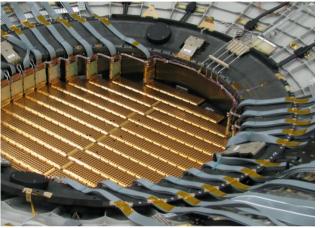
INFN



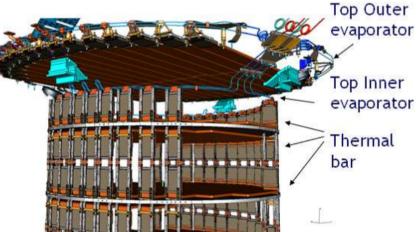


Tracker cooling: 2 phases CO2 pumped loop



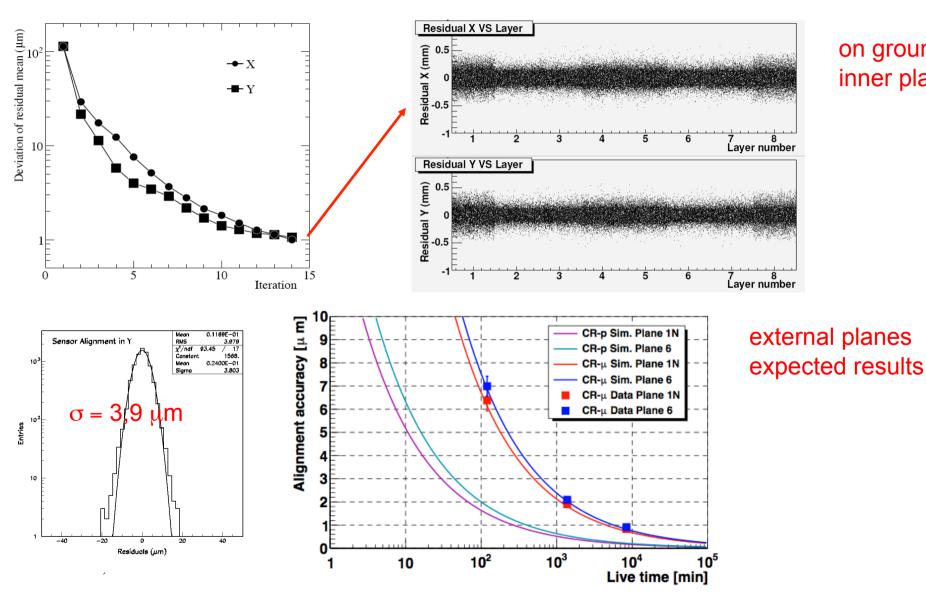


Perugia





experience to come: alignment

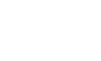


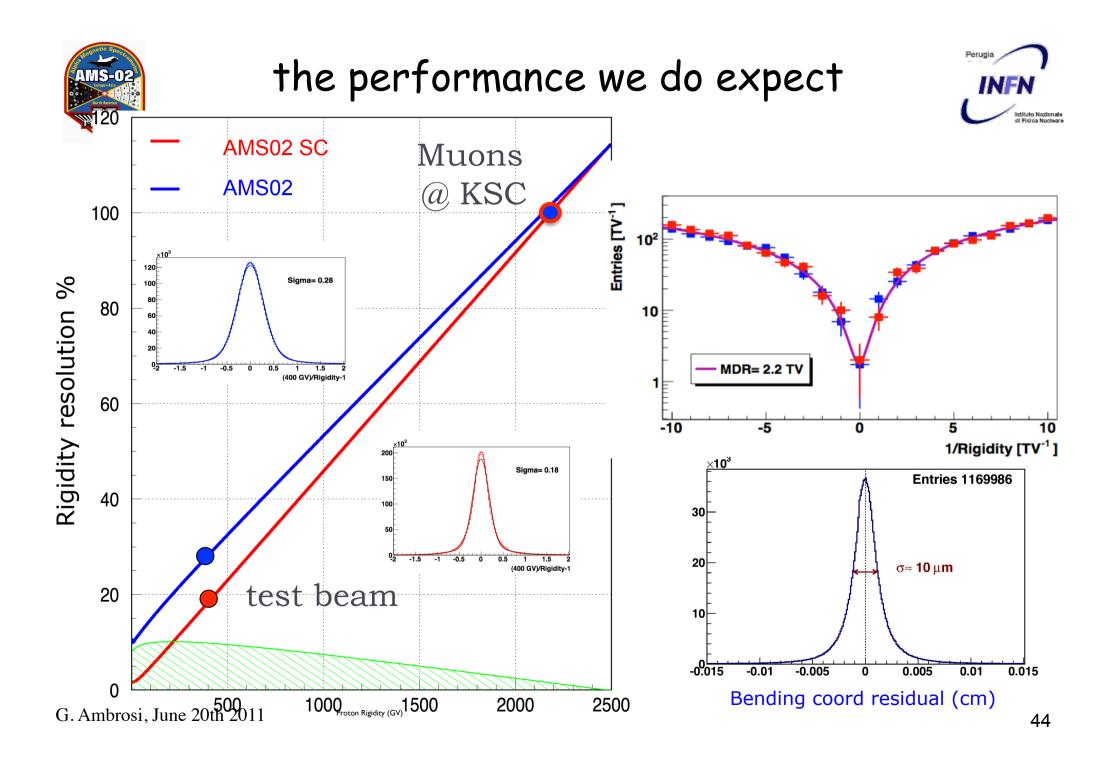
on ground results inner planes

Perugia

INFN

Intitute National



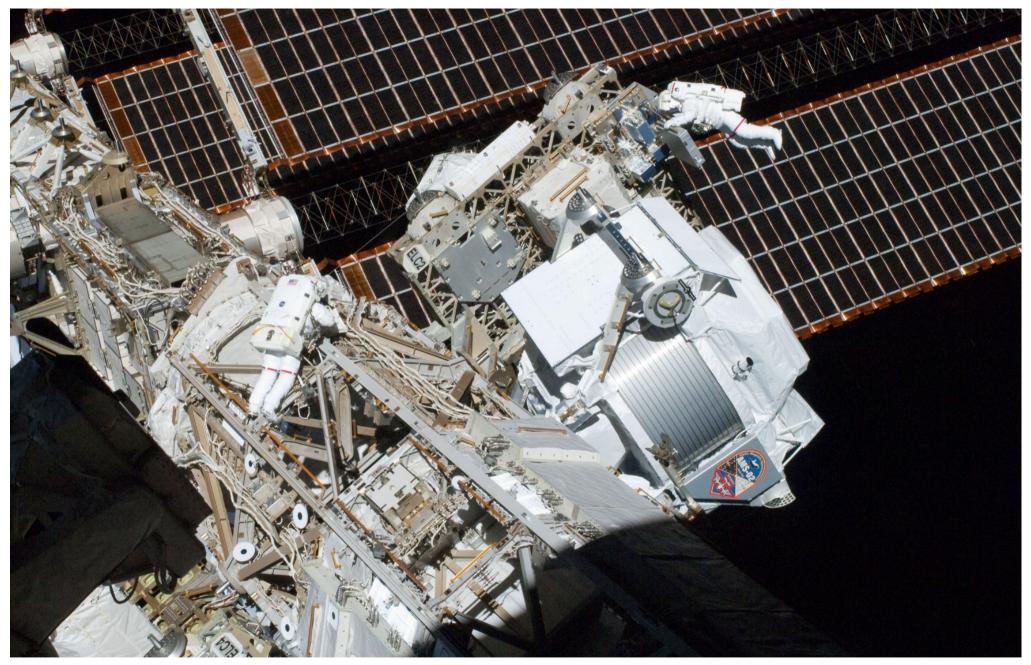




Conclusions



- AMS02 is in orbit since May 16th 2011
- No damage due to the launch stress or to the space environment, all the system are working in both the primary and redundant part
- All the detectors are properly functioning with DAQ in nominal conditions since May 19th 2011
- Tracker behavior is as expected in term of signal and noise levels
- 10+ years on board the ISS: great discovery potential, lot of work ongoing (alignment!)



Science will come soon!

G. Ambrosi, June 20th 2011



Science will come soon!

G. Ambrosi, June 20th 2011