

# The impact of the photon PDF and electroweak corrections on $t\bar{t}$ distributions

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## Motivation

### ● EW corrections

- Tension between theory and data at high  $p_T(t)$  region at 8 TeV
- Theory uncertainties decrease  $\rightarrow$  Relevance of EW corrections increase
- Experimental uncertainties will further decrease at LHC13
- $t\bar{t}$  process enters many LHC analyses as signal or background  $\rightarrow$  NNLO QCD and NLO EW predictions are necessary for  $t\bar{t}$  production

**Weak:** *Beenakker et al.*, Nu.Ph.B.411(1994), *Kuhn et al.*, hep-ph/0610335, arXiv:1305.5773, *Bernreuther et al.*, hep-ph/0508091, *Campbell et al.*, arXiv:1608.03356; **QED+ $g\gamma$  LO:** *Hollik et al.*, arXiv:0708.1697; **FB asymmetry:** *Hollik et al.*, arXiv:1107.2606, *Kuhn et al.*, arXiv:1109.6830, *Manohar et al.*, arXiv:1201.3926, *Bernreuther et al.*, arXiv:1205.6580; **NLO+EW+decay (NWA):** *Bernreuther et al.*, arXiv:1003.3926; **EW to  $e^+\mu^-\nu\nu b\bar{b}$ :** *Denner et al.*, arXiv:1607.05571

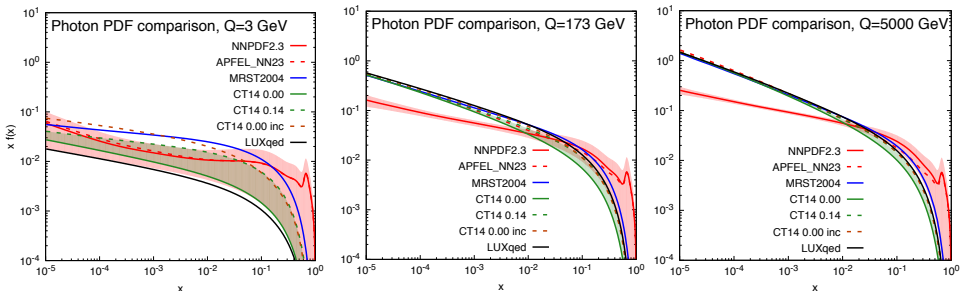
### ● Photon-induced contributions

- The (negative) Sudakov suppression is compensated by the (positive) photon-induced contributions

**PDF sets including  $\gamma(x, Q)$ :** MRST2004QED: *Martin et al. '04*, NNPDF2.3QED: *Ball et al. '13*, CT14QED(inc): *Schmidt et al. '16*, NNPDF3.0QED: *Bertone, Carrazza '16*, LUXqed: *Manohar et al. '16*, MMHTQED (in progress), additional Studies: *Harland-Lang, Khoze, Ryskin '16*

Different PDF sets

LO QED evolution but different assumptions for all PDF sets



- At low  $Q = 3$  GeV there is a similar behaviour
- At high  $Q$  values and low  $x$ , the NNPDF2.3QED is different due to different DGLAP QCD and QED running (not relevant for  $t\bar{t}$ )
- At large  $x \rightarrow \begin{cases} \text{NNPDF2.3QED, large } \gamma(x, Q) \\ \text{CT14QED, LUXqed, small } \gamma(x, Q) \end{cases}$

## Calculation framework

NLO QCD+EW corrections to  $t\bar{t}$  production (MADGRAPH5\_AMC@NLO)

LO

$$\begin{array}{c} q\bar{q}, gg \\ \textcircled{\alpha_s^2} \\ \text{LO QCD} \end{array}$$

$$\begin{array}{c} b\bar{b}, g\gamma \\ \textcircled{\alpha_s \alpha} \\ \text{LO EW} \end{array}$$

$$\textcircled{\alpha^2}$$

$$\Sigma_{\text{QCD}} \equiv \Sigma_{\text{LO QCD}} + \Sigma_{\text{NLO QCD}}$$

$$\Sigma_{\text{EW}} \equiv \Sigma_{\text{LO EW}} + \Sigma_{\text{NLO EW}}$$

NLO

$$\begin{array}{c} qg \\ \textcircled{\alpha_s^3} \\ \text{NLO QCD} \end{array}$$

$$\begin{array}{c} g\gamma, q\gamma \\ \textcircled{\alpha_s^2 \alpha} \\ \text{NLO EW} \end{array}$$

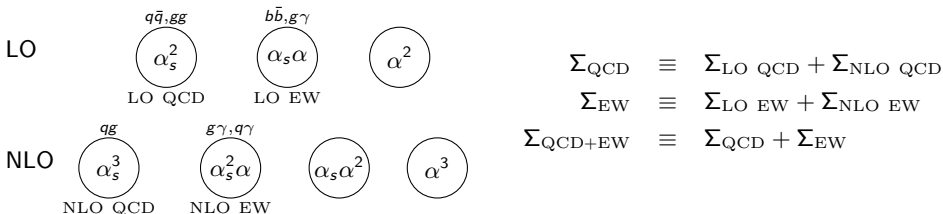
$$\textcircled{\alpha_s \alpha^2}$$

$$\textcircled{\alpha^3}$$

$$\Sigma_{\text{QCD+EW}} \equiv \Sigma_{\text{QCD}} + \Sigma_{\text{EW}}$$

### Calculation framework

NLO QCD+EW corrections to  $t\bar{t}$  production (MADGRAPH5\_AMC@NLO)



- Input parameters

$$\begin{aligned}
 m_t &= 173.3 \text{ GeV}, & m_H &= 125.09 \text{ GeV}, \\
 m_W &= 80.385 \text{ GeV}, & m_Z &= 91.1876 \text{ GeV}, \\
 G_\mu &= 1.1663787 \cdot 10^{-5} \text{ GeV}^{-2}
 \end{aligned}$$

- Scale choice

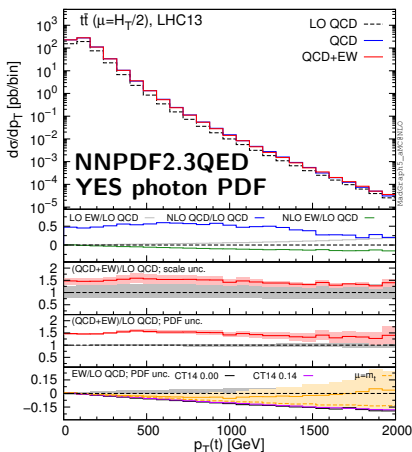
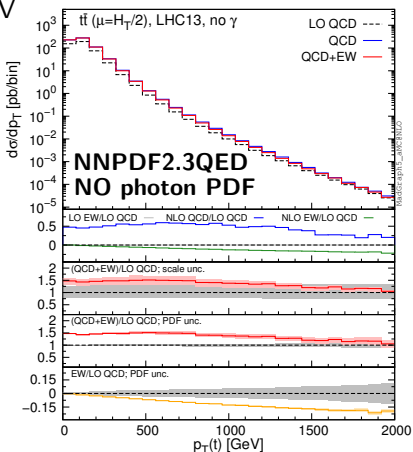
$$\mu = \frac{H_T}{2} = \frac{1}{2} \sum_i m_{T,i}$$

- PDF sets considered

- Main results  $\rightarrow$  NNPDF2.3QED, NNPDF2.3QED (no  $\gamma(x, Q)$ )
- Comparison with CT14QED at 13 TeV

Results at 13 TeV

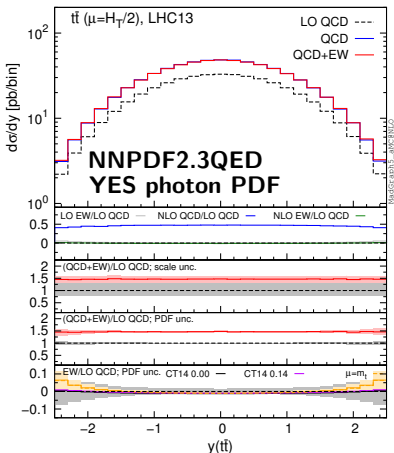
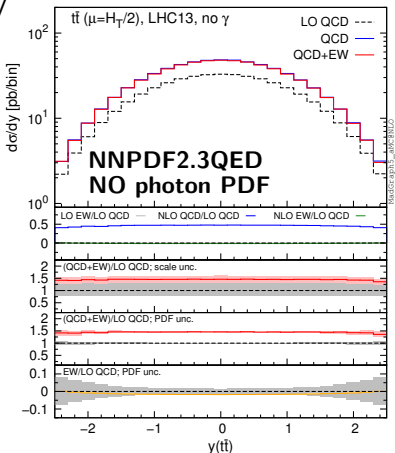
13 TeV



- NNPDF2.3QED: Photon PDF effect is mostly due to LO EW contribution
- NNPDF2.3QED: Large cancellations between Sudakov logs and photon-induced contributions
- photon PDF impact → large in NNPDF2.3QED, negligible in CT14QED

Results at 13 TeV

13 TeV



- NNPDF2.3QED:  $\sim 5\%$  effect of photon PDF at large rapidity regions, which are reachable experimentally
- CT14QED  $\longleftrightarrow$  NNPDF2.3QED (no  $\gamma(x, Q)$ )
- CT14QED and NNPDF2.3QED in agreement within uncertainties



Calculation framework

$t\bar{t}$  distributions at NNLO QCD+NLO EW accuracy (**PRELIMINARY**)

*in collaboration with*  
*Michal Czakon, David Heymes, Alex Mitov, Davide Pagani, Marco Zaro*

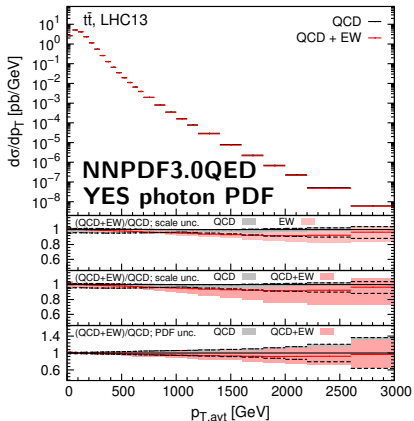
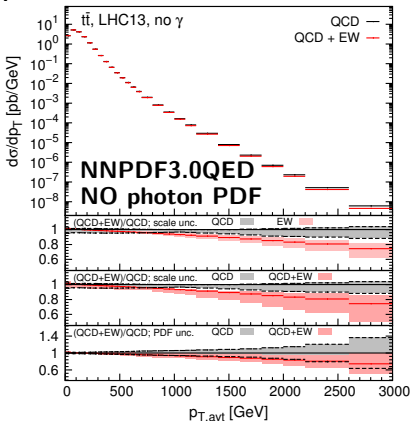
- Scale choice based on arXiv:1606.03350 (Czakon, Heymes, Mitov)

$$\mu = \begin{cases} m_T/2 & \text{for } p_{T,avt} \\ H_T/4 & \text{for } m(t\bar{t}), y_{avt}, y(t\bar{t}) \end{cases}$$

- PDF sets considered
  - Main results  $\rightarrow$  NNPDF3.0QED, NNPDF3.0QED (no  $\gamma(x, Q)$ )
  - Comparison with LUXqed

Preliminary Results at 13 TeV

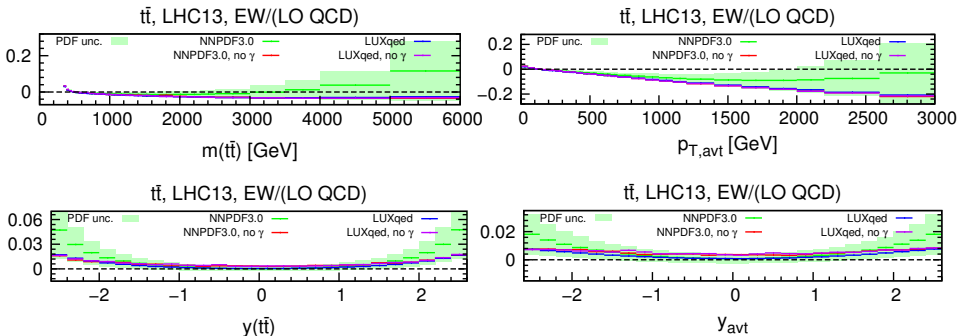
13 TeV



- NNLO QCD corrections reduce the scale dependence significantly
- PDF uncertainties become large at high  $p_T$  region
- For this distribution the photon PDF uncertainty does not increase the total PDF uncertainty

Preliminary Results at 13 TeV

NNPDF3.0QED vs LUXqed



- photon PDF impact  $\rightarrow$  large in NNPDF3.0QED, negligible in LUXqed
- LUXqed  $\leftrightarrow$  NNPDF3.0QED (no  $\gamma(x, Q)$ )
- LUXqed and NNPDF3.0QED in agreement within uncertainties

## Conclusions

- PDF sets
  - NNPDFQED  $\rightarrow$  Large impact of photon-induced contributions accompanied with large uncertainties
  - CT14QED, LUXqed  $\rightarrow$  Negligible impact of photon-induced contributions in  $t\bar{t}$  distributions
- 13 TeV
  - Photon-induced contributions may be visible in rapidity but also in  $p_T$  distributions
  - NNLO QCD are necessary in order to reduce the scale dependence
  - NNLO QCD + NLO EW in progress

## Conclusions

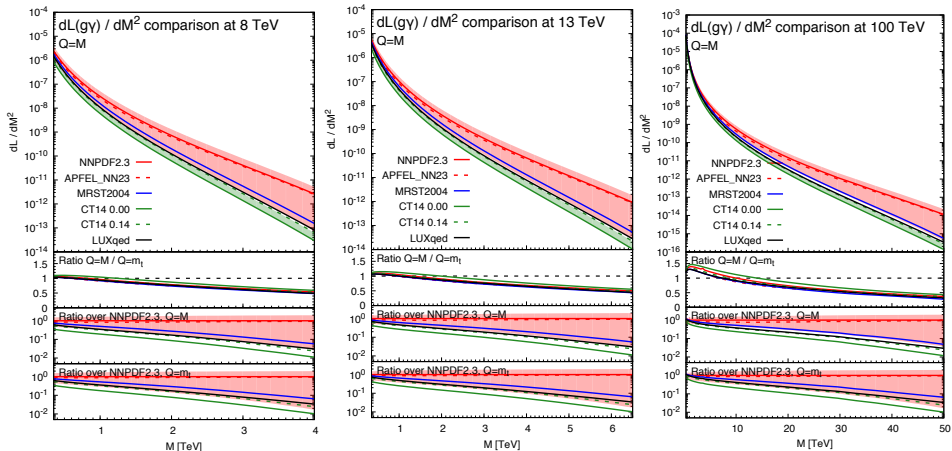
- PDF sets
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*...Thank you*

## The photon PDF

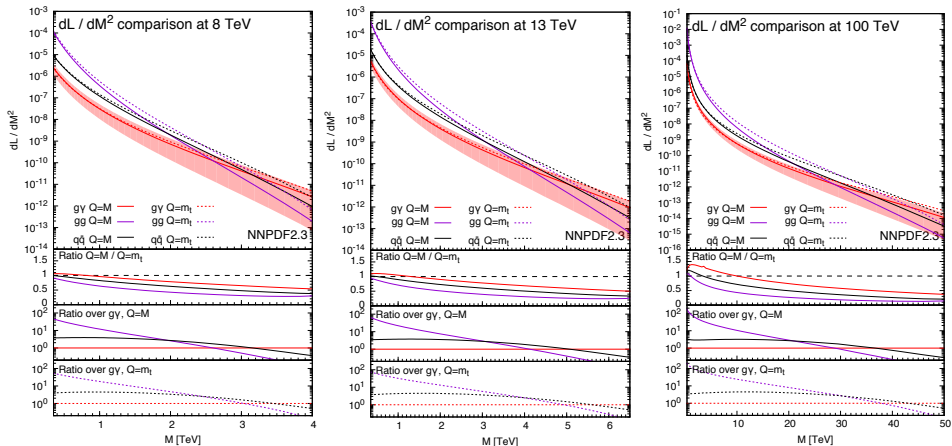
- NNPDF2.3QED
  - No assumption for the  $\gamma(x, Q^0)$  functional form
  - Different scales for QCD/QED evolutions
- CT14QED
  - Uses an ansatz like MRST2004 with one free parameter
  - The momentum fraction carried by the photon is constrained to be  $\leq 0.14\%$  at 90% CL
  - A set including the elastic photon contribution is also provided (CT14QEDinc)
- NNPDF3.0QED
  - Simultaneous evolution of QCD/QED is implemented (also in APFEL\_NN23), which changes the low  $x$  behaviour, but with no effect in  $t\bar{t}$  phenomenology

## The $g\gamma$ Luminosity



- LUXqed lies very close to CT14QED
- Effects due to the different evolution in NNPDF2.3QED are not visible

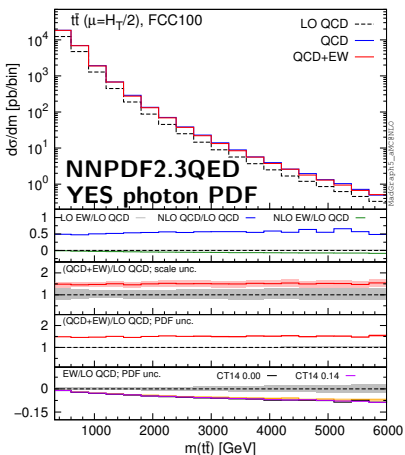
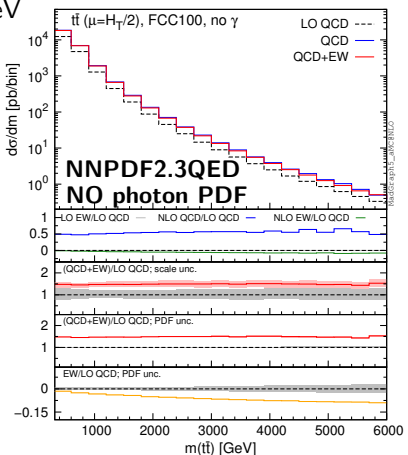
## Parton Luminosities and scale choice



- In both dynamical and fixed scales the  $g\gamma$  luminosity is suppressed with respect to the  $gg$  one at the low  $M$  region

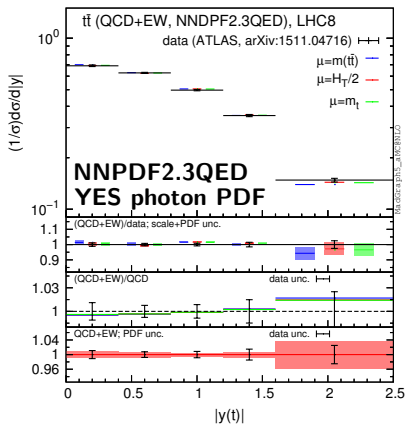
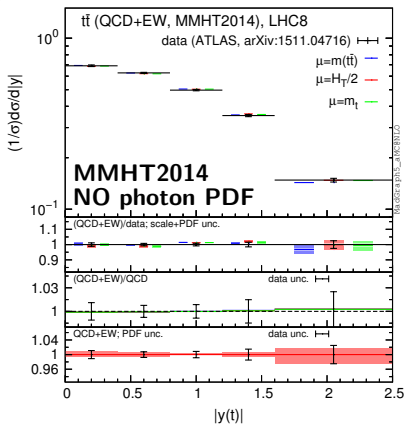


100 TeV



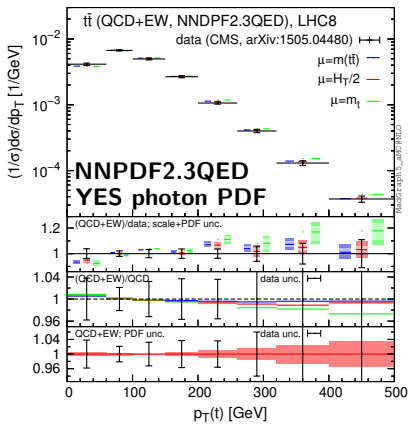
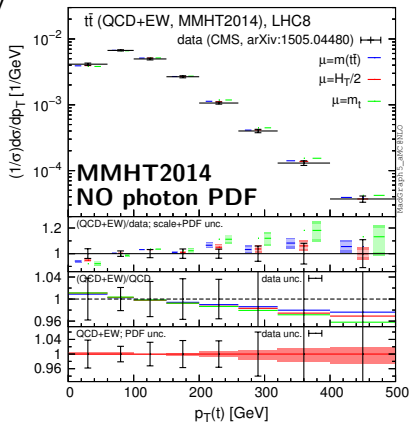
- At 100 TeV  $t\bar{t}$  differential distributions are not sensitive to photon-induced contributions
- $\sqrt{s} \uparrow \implies$  Bjorken  $x$ 's  $\downarrow$
- Larger effects are expected at 8 TeV, where already we have data

# 8 TeV



- Normalised  $(1/\sigma)$  rapidity distributions  $\rightarrow$  Exp. errors reduce at few % level.
- Large PDF uncertainties and visible impact of photon PDF (NNPDF2.3QED)
- Can be used for constraining the photon PDF (NNPDF2.3QED)

# 8 TeV



- In  $p_T$  distributions the impact of the photon PDF is larger at the tail
- Sudakov logs vs  $\gamma(x, Q)$  cancellation depends on the scale definition
- Scale uncertainty still large at NLO QCD  $\rightarrow$  NNLO QCD needed
- For 13 TeV comparisons between theory and experiment EW corrections and photon-induced contributions need to be taken into account