

BSM&TOP PHYSICS (AFTER LHC)

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OLOMOUC, SEP. 22 2016**



Outline

- * status and prospect for new physics at LHC
- * examples (mostly SUSY) and ideas
- * conclusions

LHC status and ^{flash} perspectives

- * top quark physics at LHC is entering precision age
- * by 2018 LHC will have **0.1/ab=100/fb**
- * in sight: new Higgs bosons and new electroweak states

Menu

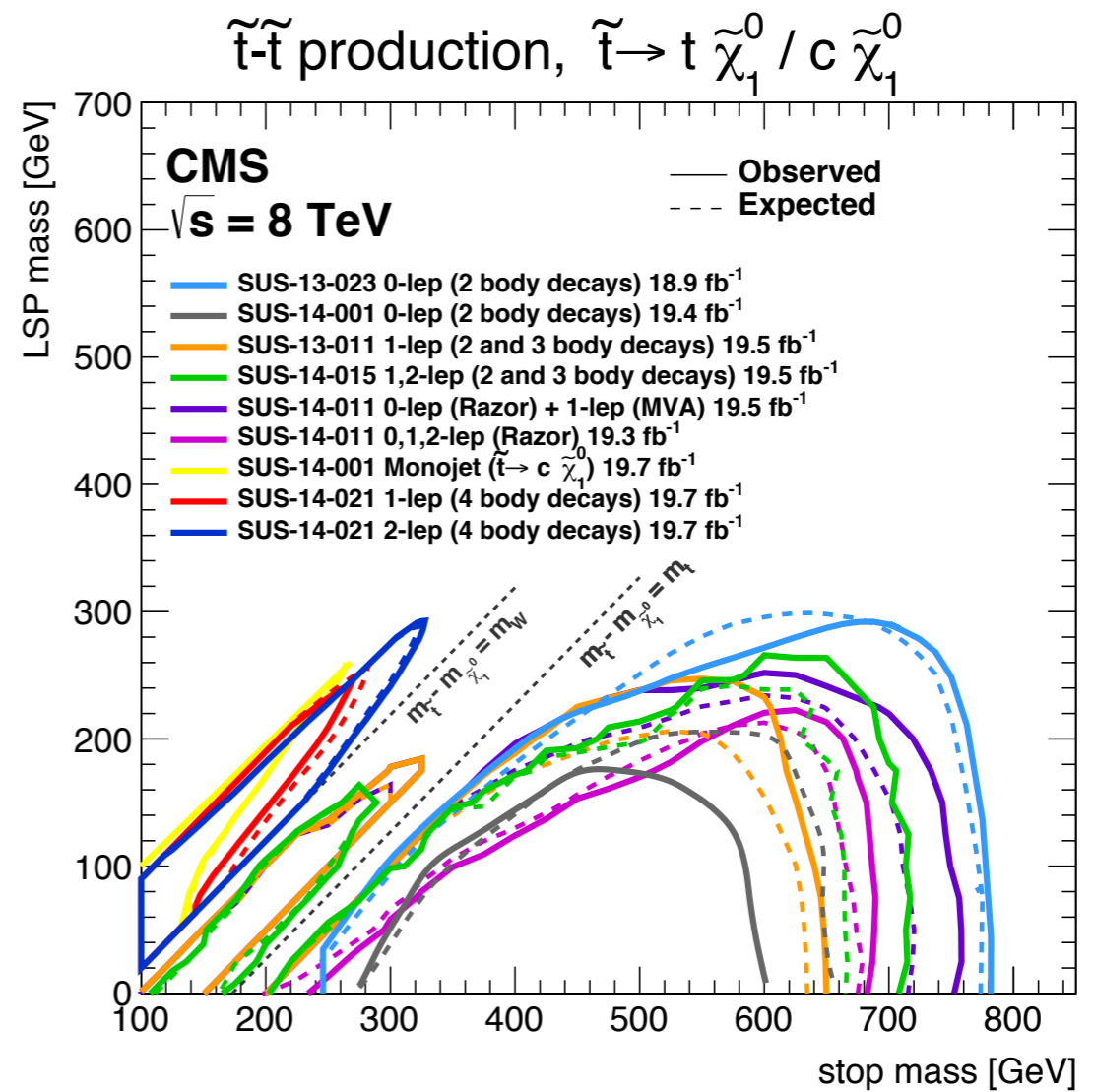
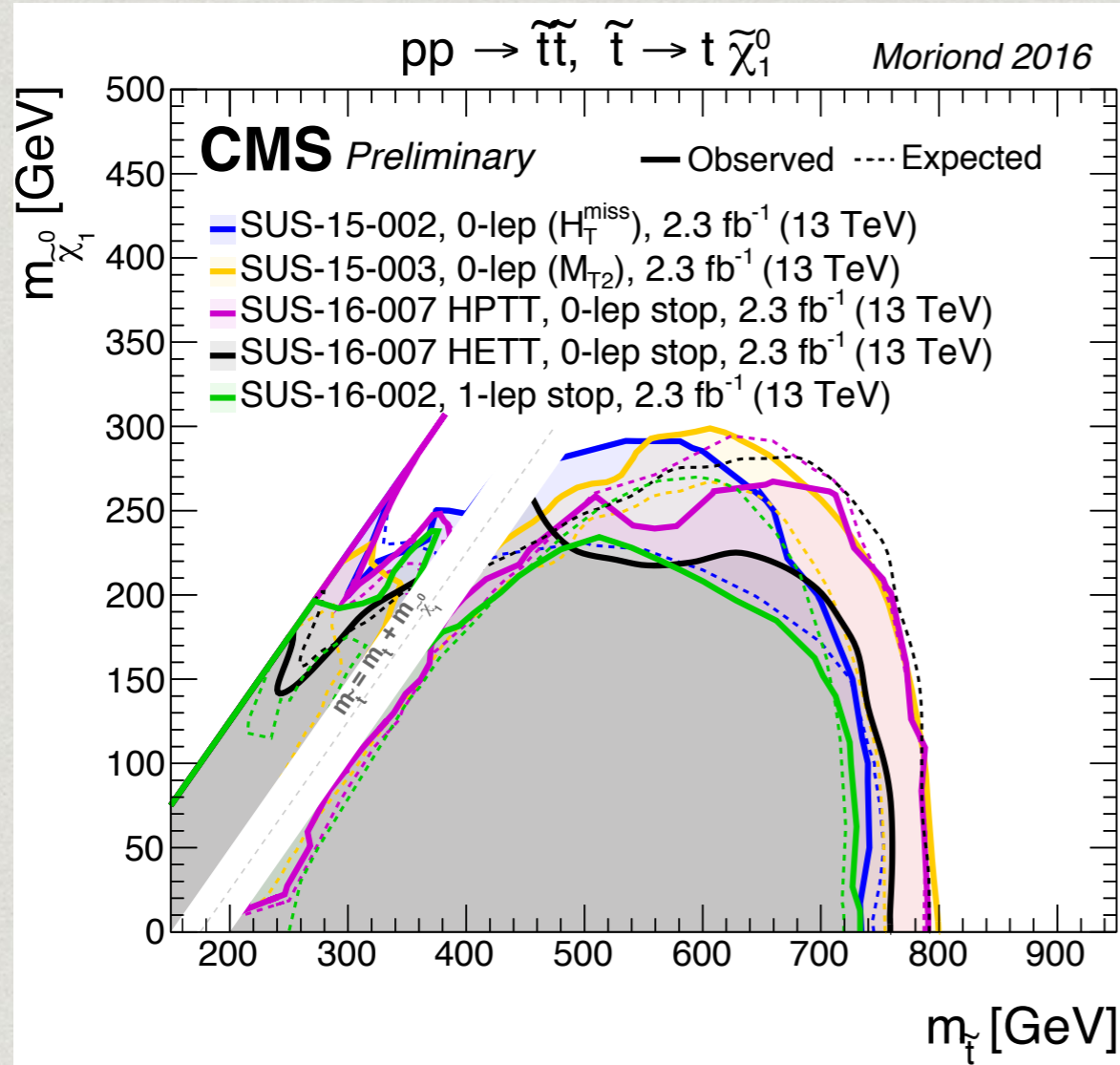
despite the great prospects, some flies on the ointment:

- * squarks in the difficult regions
- * light degenerate chargino (Higgsino-like)
- * light singlets (*e.g.* $pp \rightarrow S \rightarrow \gamma\gamma$)

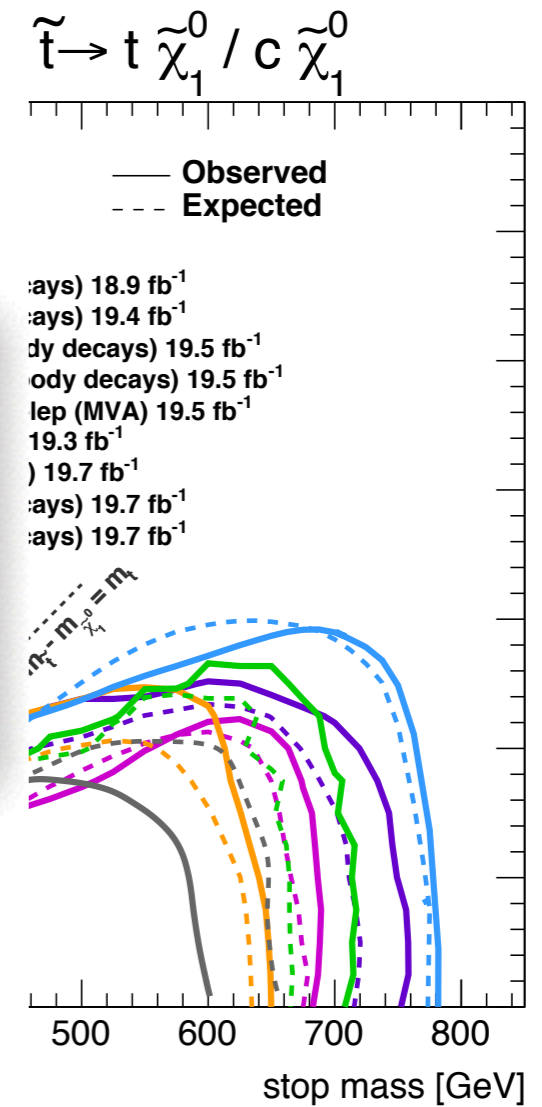
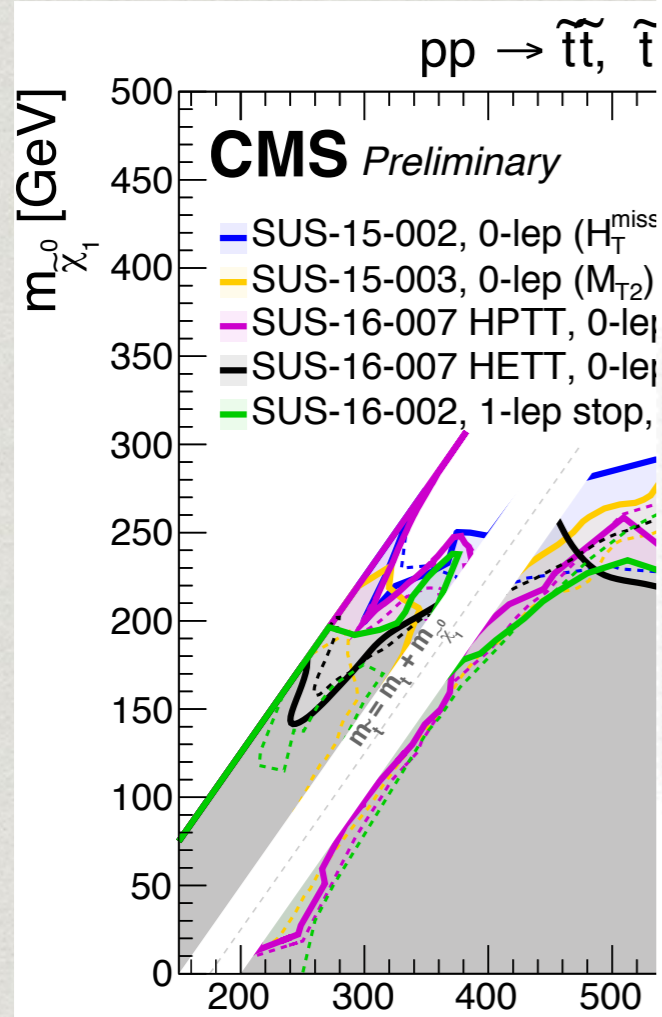
Challenges

- * compressed
- * diluted
- * delayed

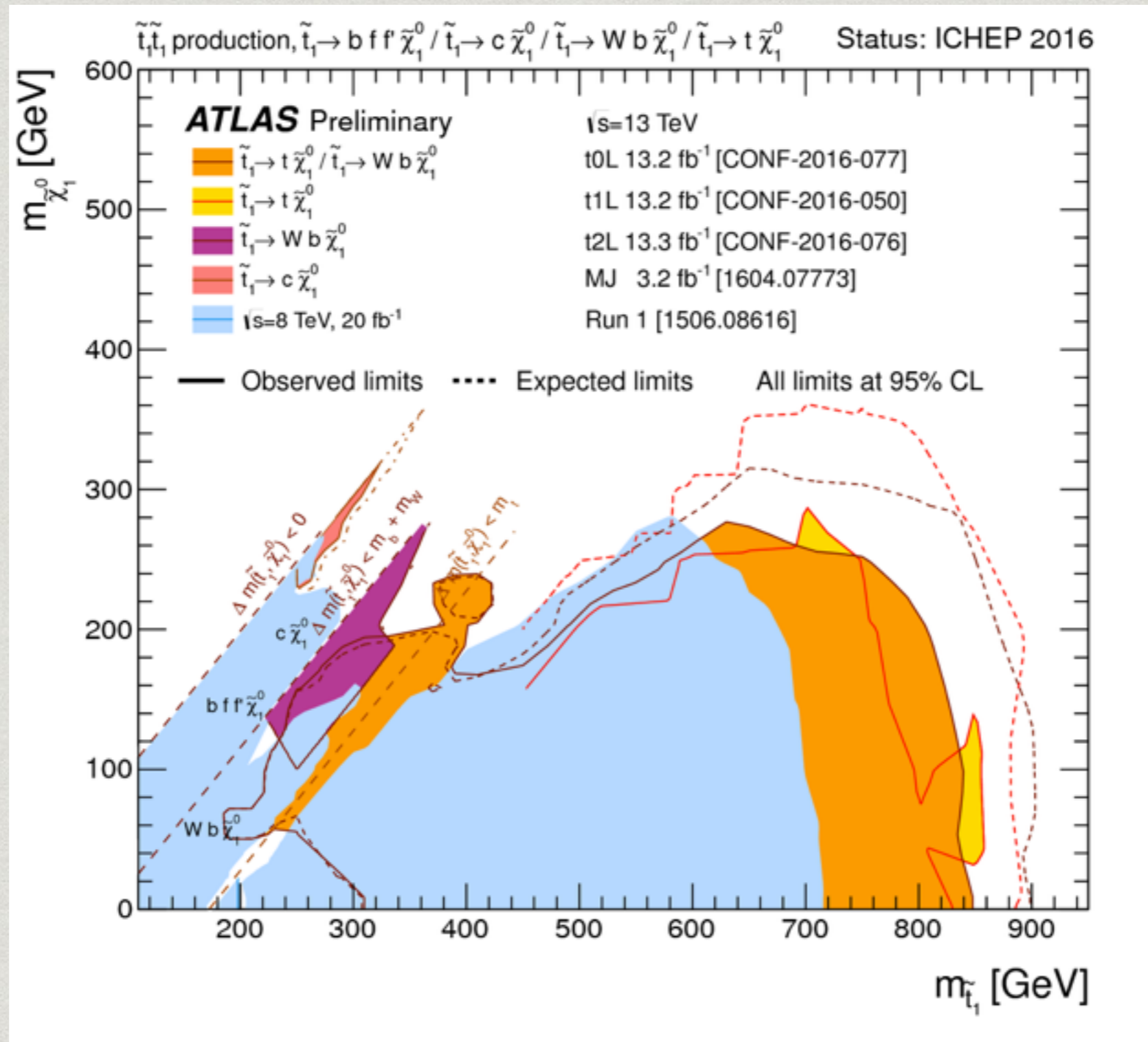
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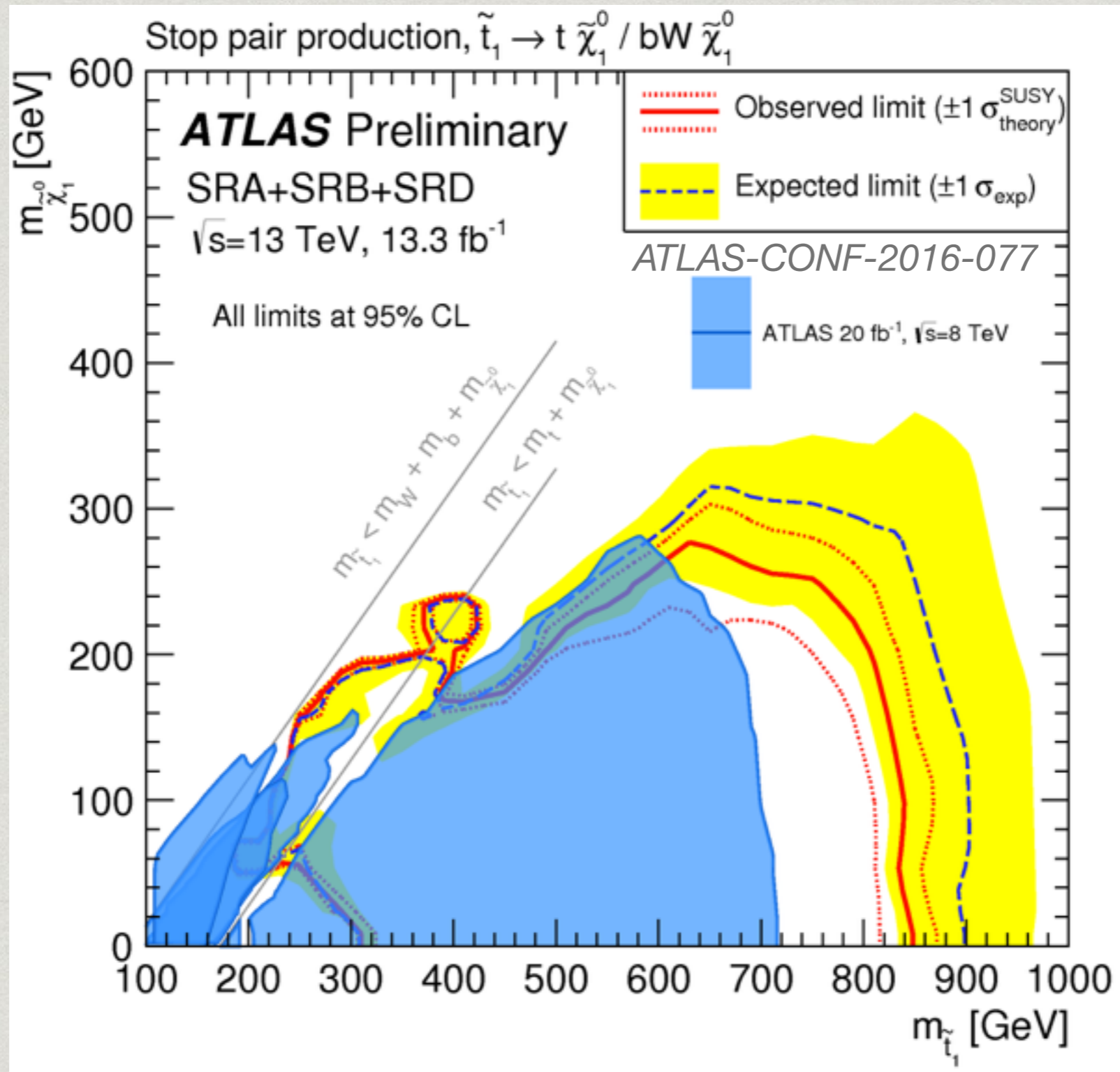


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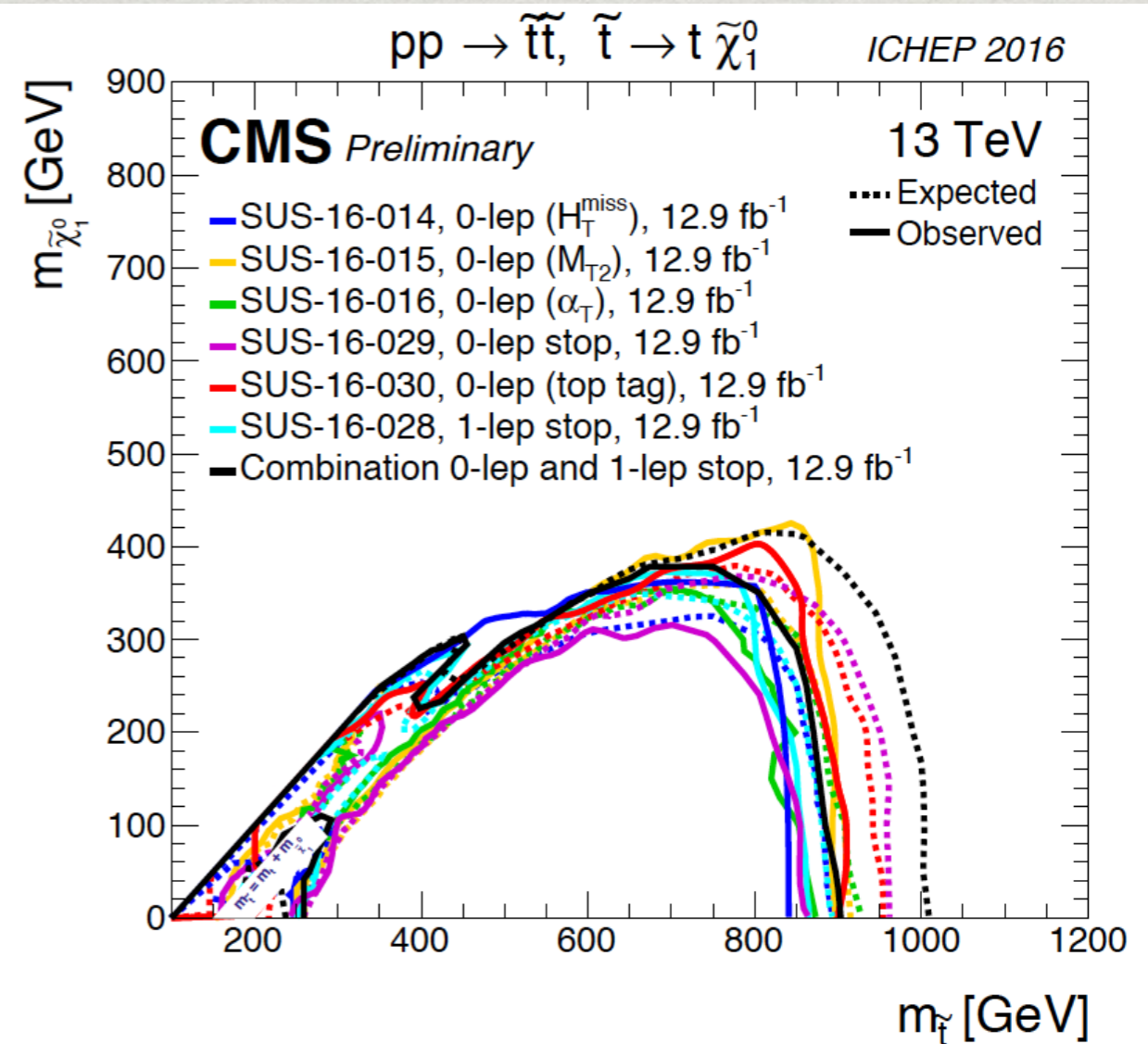
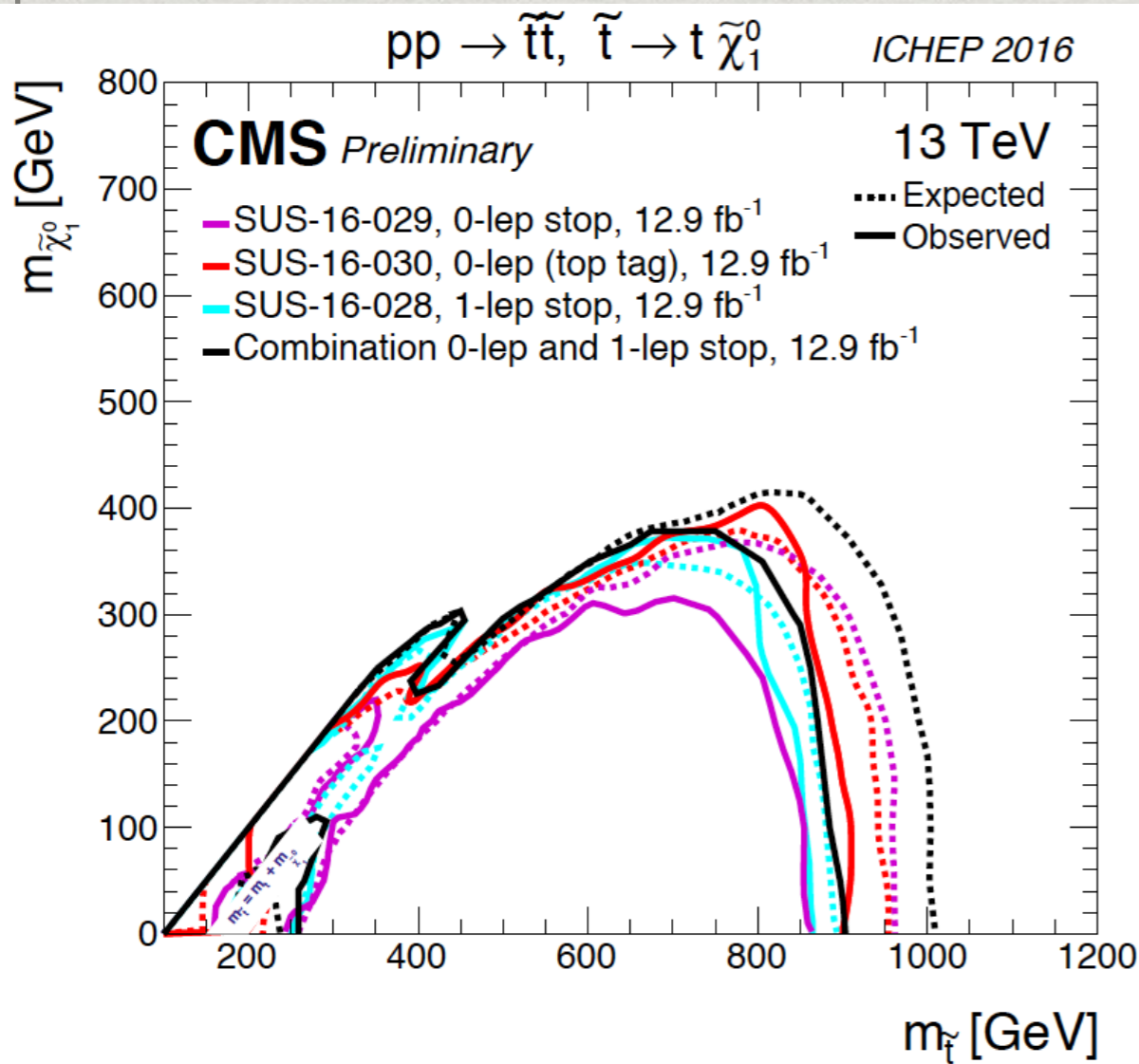


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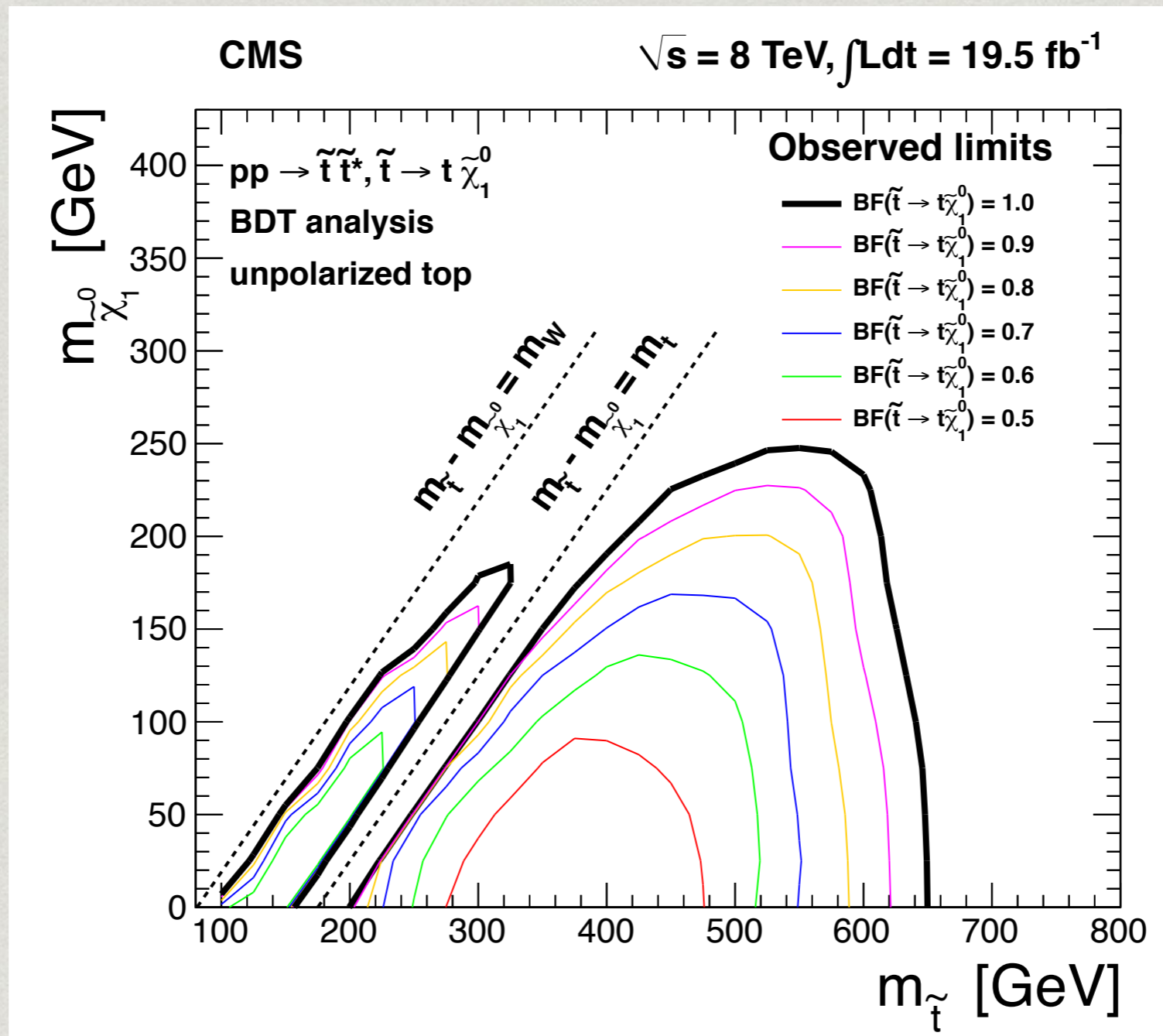
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Challenges

- * compressed \Rightarrow little visible energy
- * diluted
- * delayed

Dilution

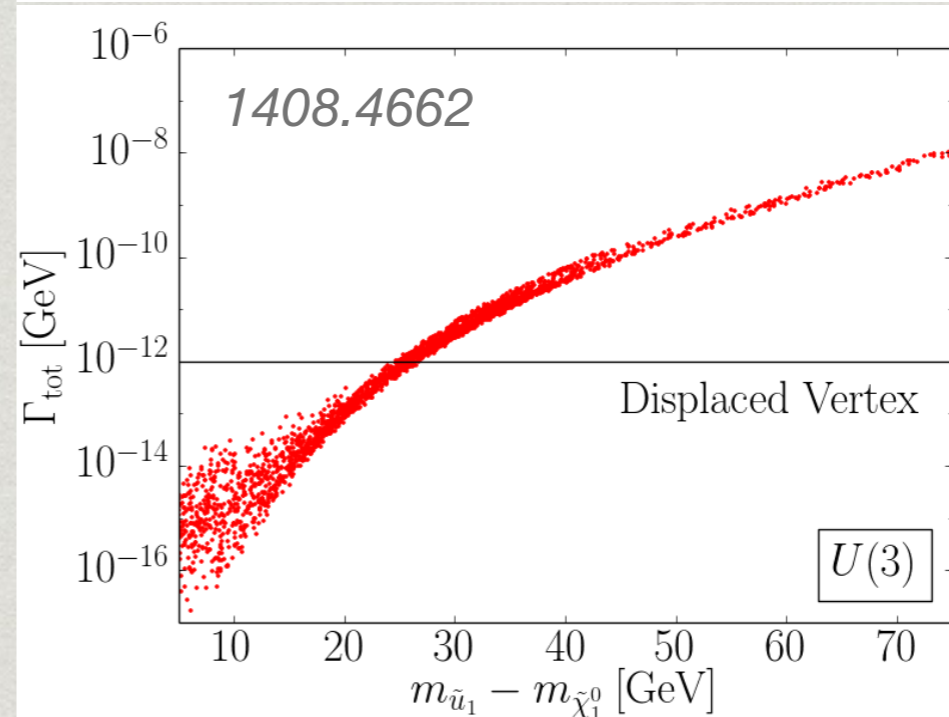
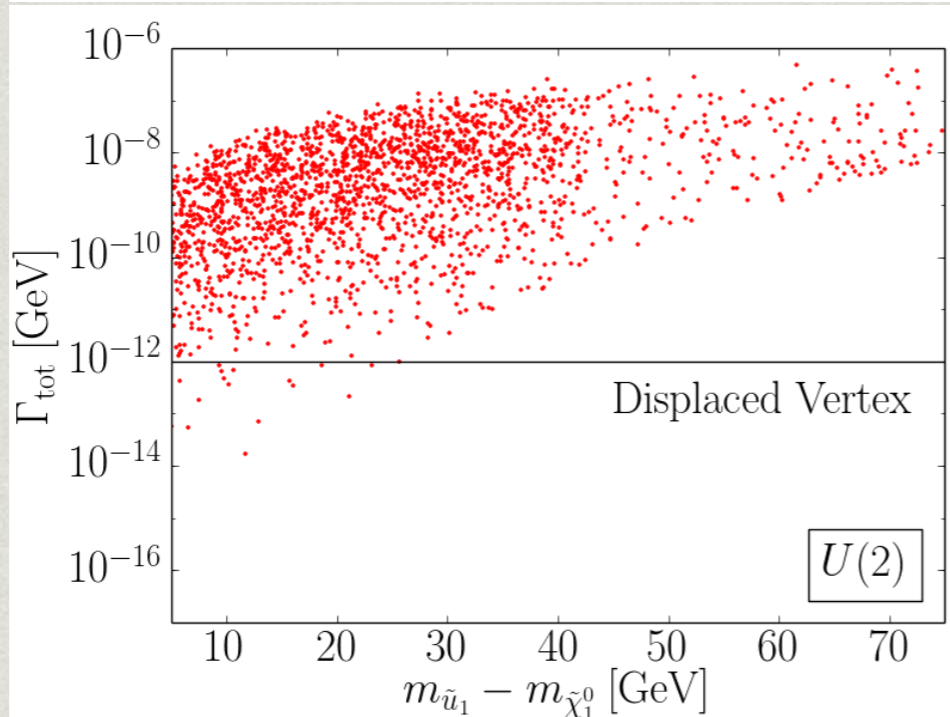
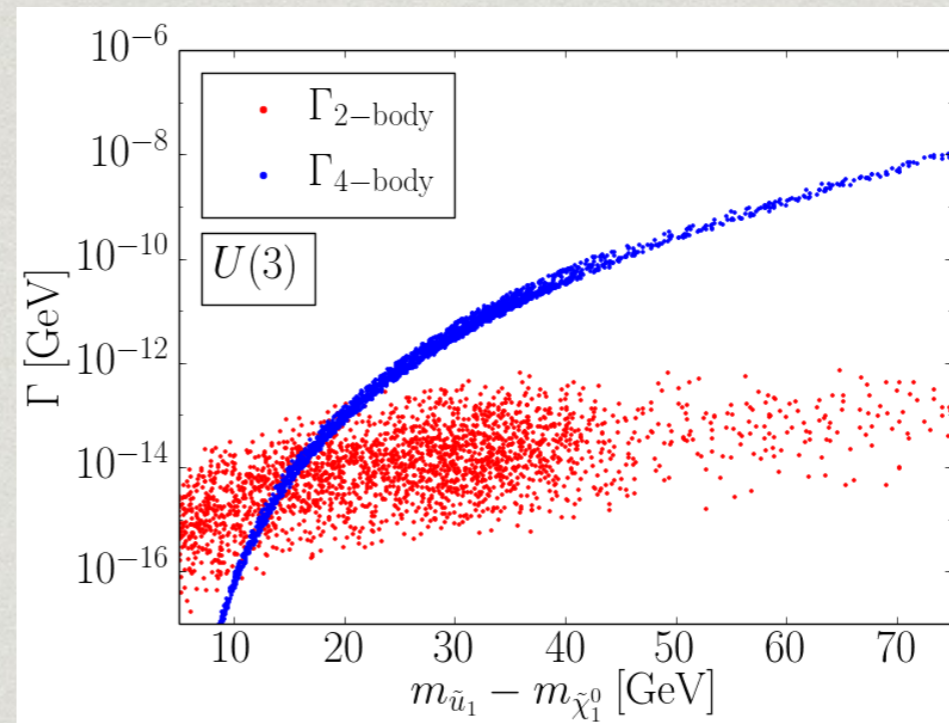
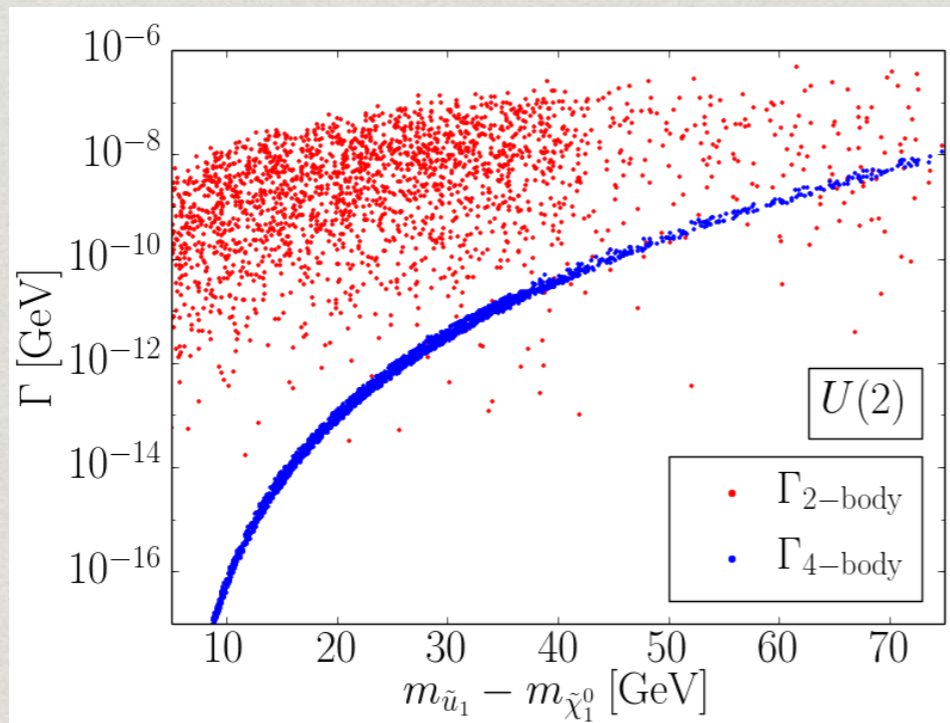


Challenges

- * compressed \Rightarrow little visible energy
- * diluted \Rightarrow spread on many channels
- * delayed

Wait ...

$$\tilde{t} \rightarrow b f f' \chi^0 + c \chi^0$$

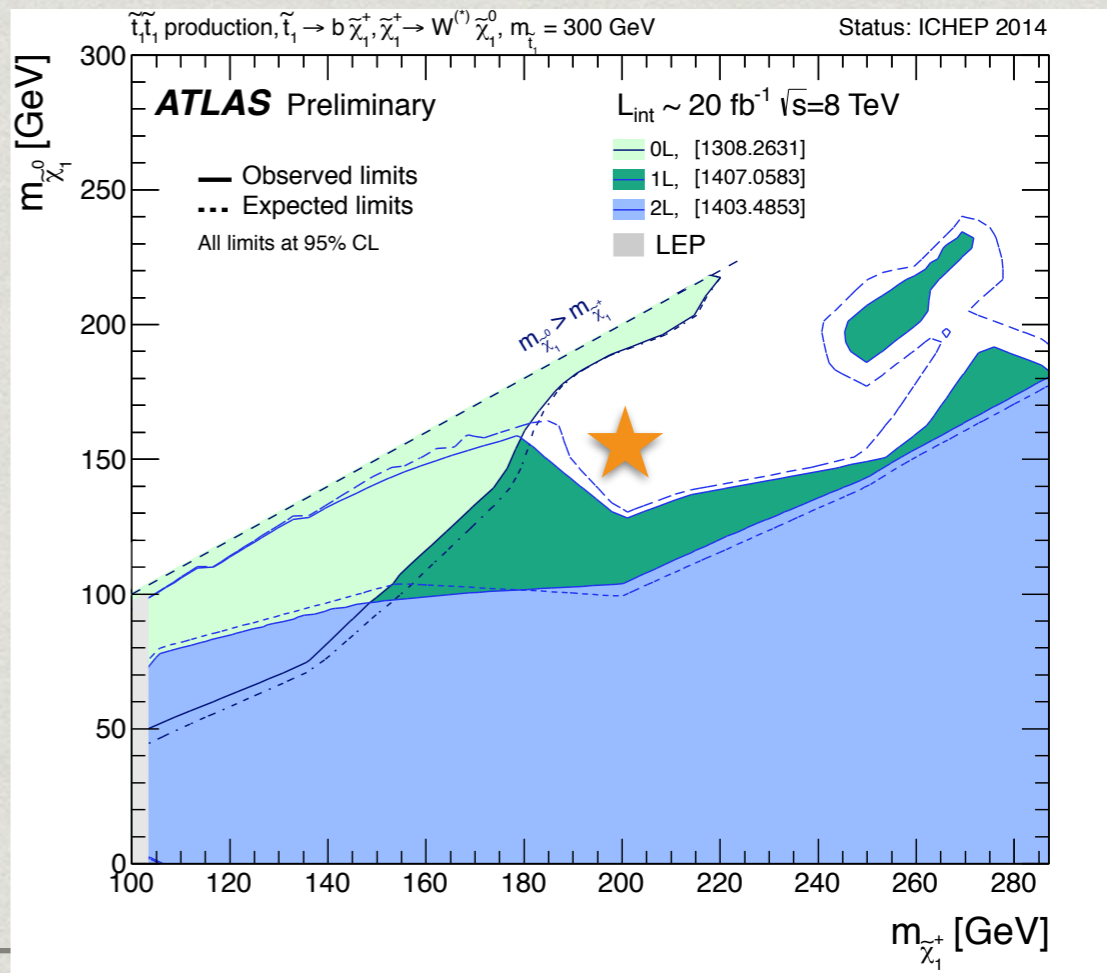
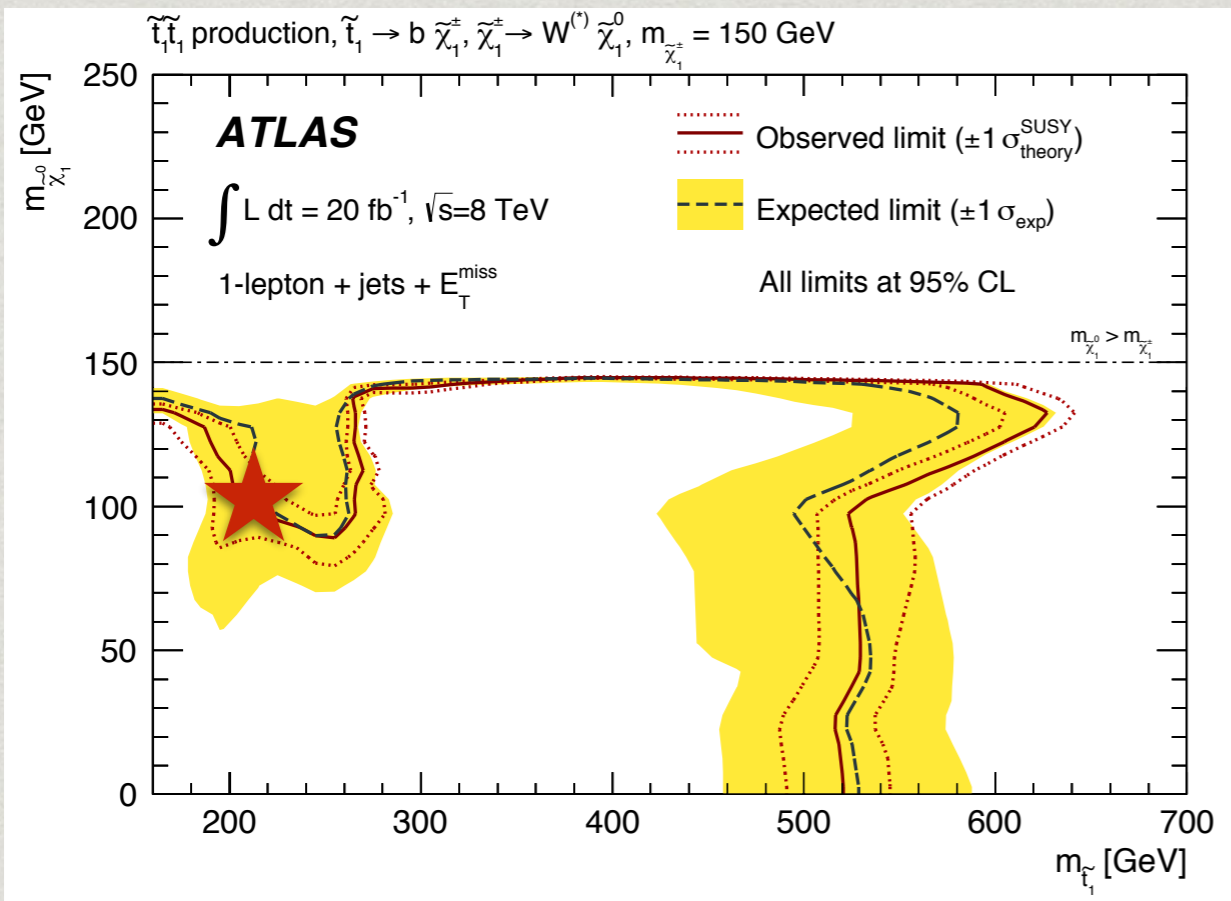
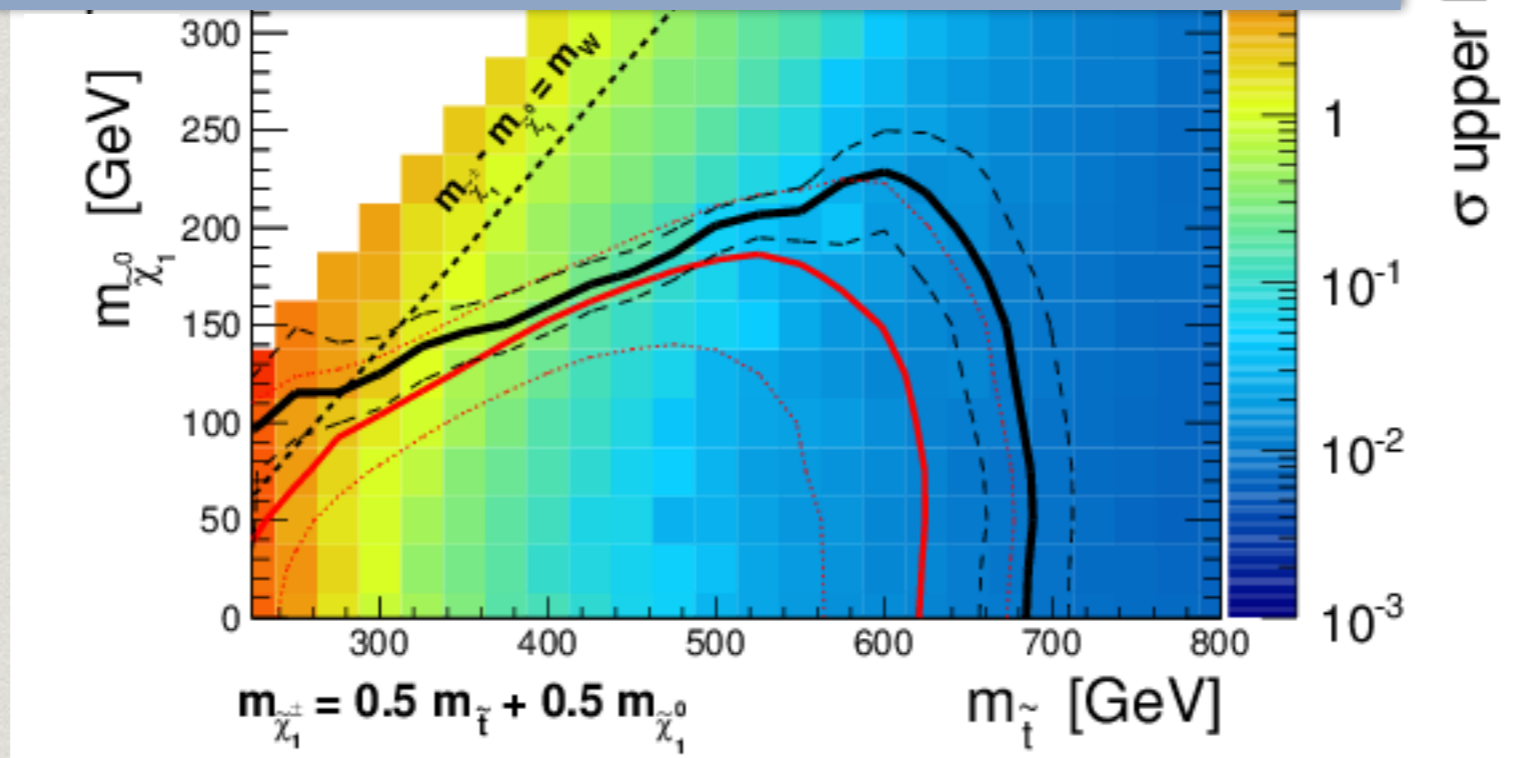


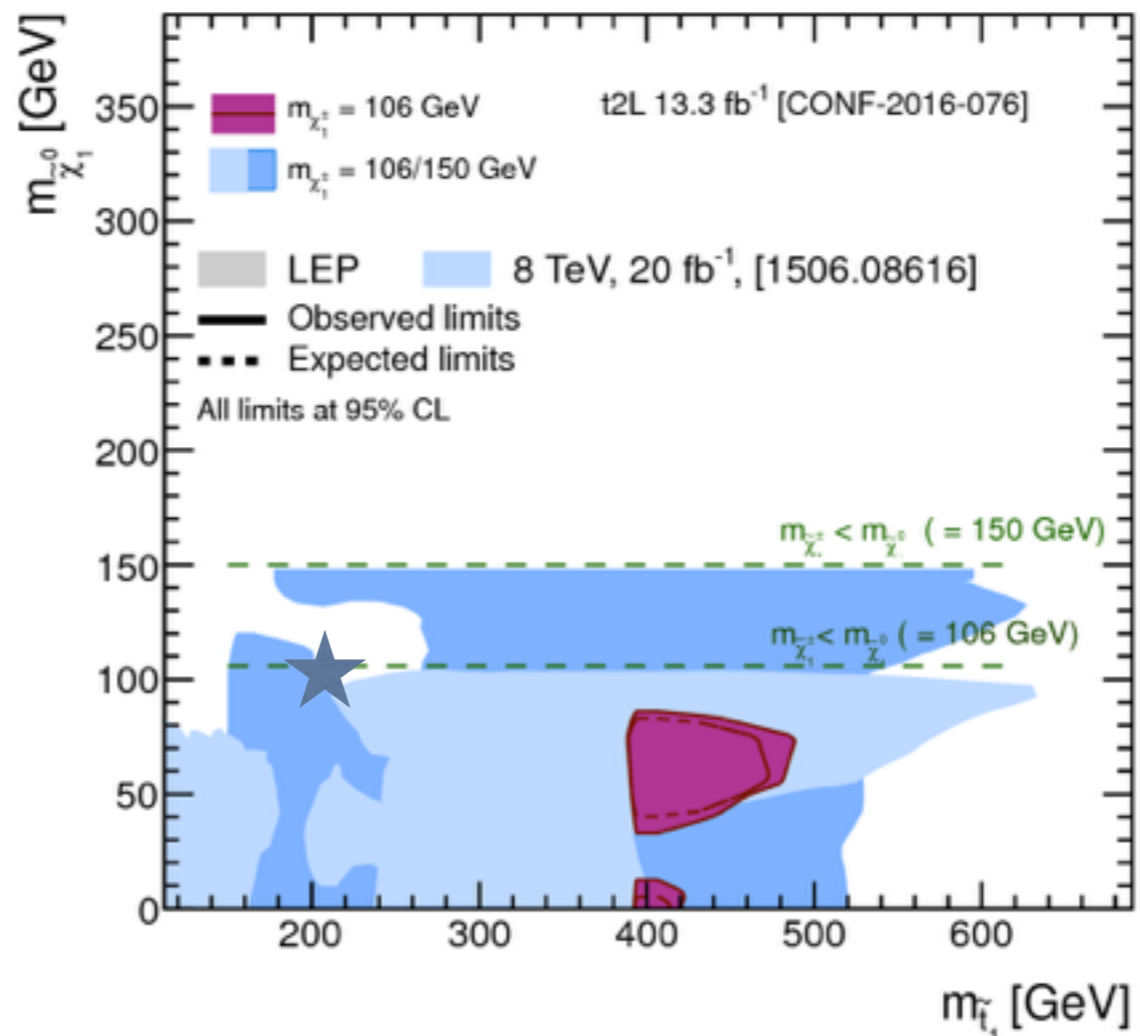
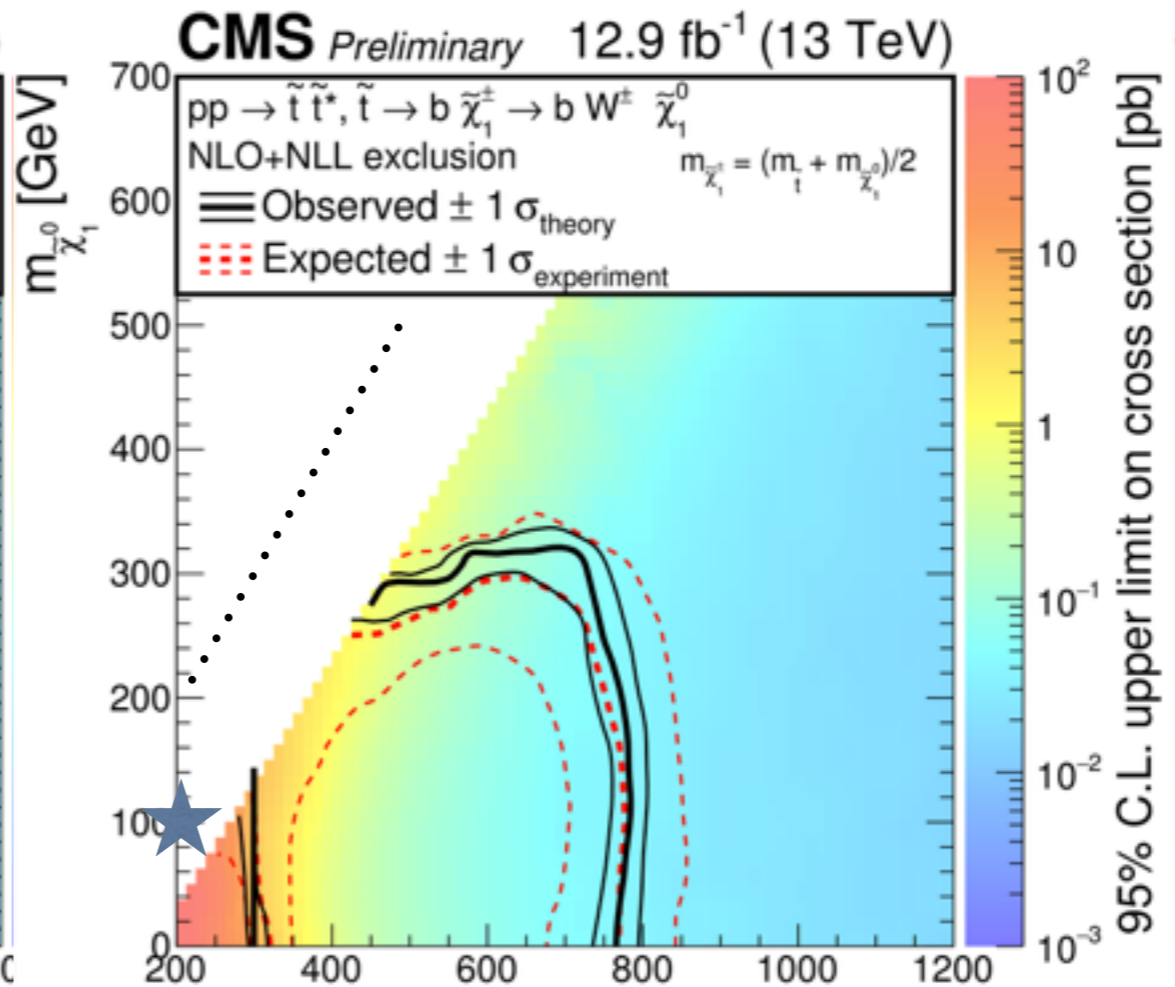
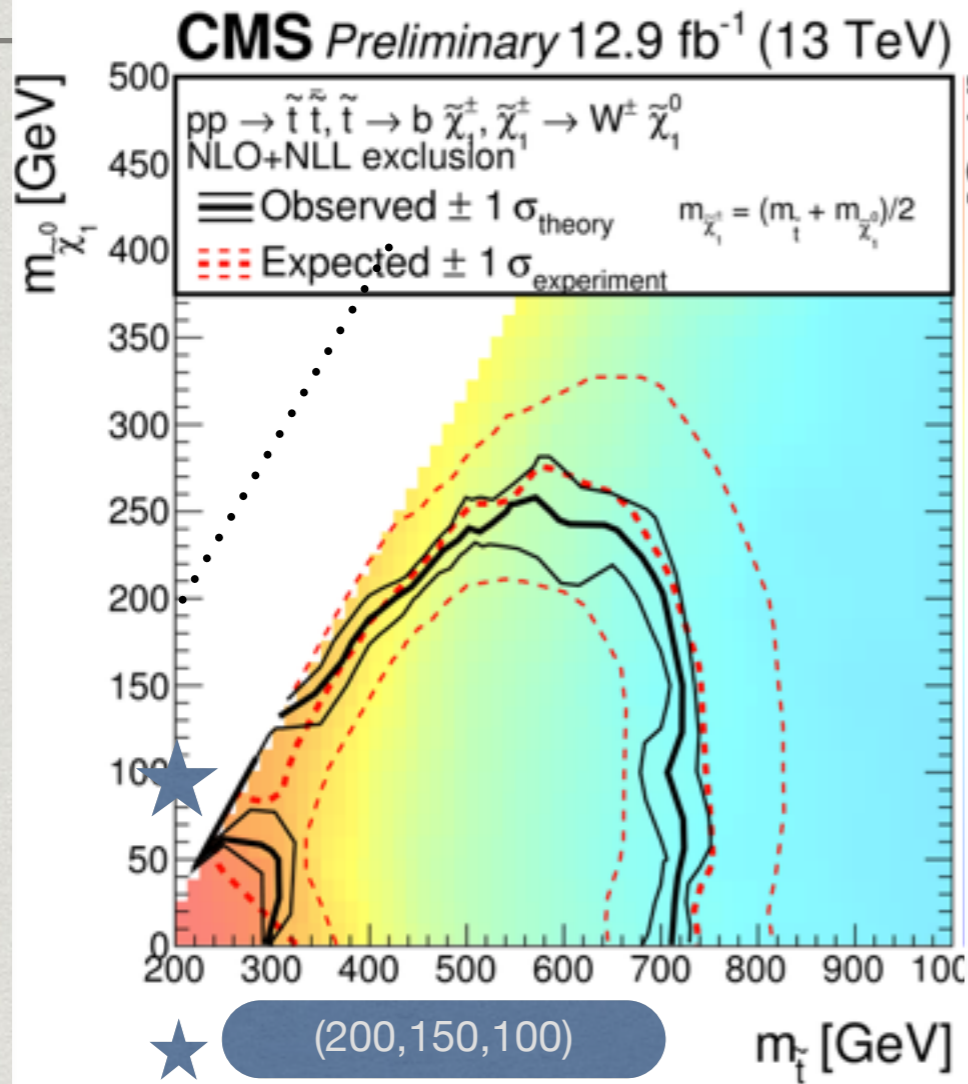
Challenges

- * compressed \Rightarrow little visible energy
- * diluted \Rightarrow spread on many channels
- * delayed \Rightarrow flavor tags may not work, signal is different than what originally thought

THE MORE THE MERRIER

$$\tilde{t} \rightarrow b \chi^+ \rightarrow b \ell \nu \chi^0$$





THE MORE THE MERRIER

INCLUDING CHARGINO IS EVEN MORE CHALLENGING

AFTER SEVERAL YEARS OF LHC

BLINDSPOTS ARE STILL THERE

- * light stop searches seem very challenging
- * situation is even more complex for $\tilde{t} \rightarrow \chi^+ b$



3. Light stop

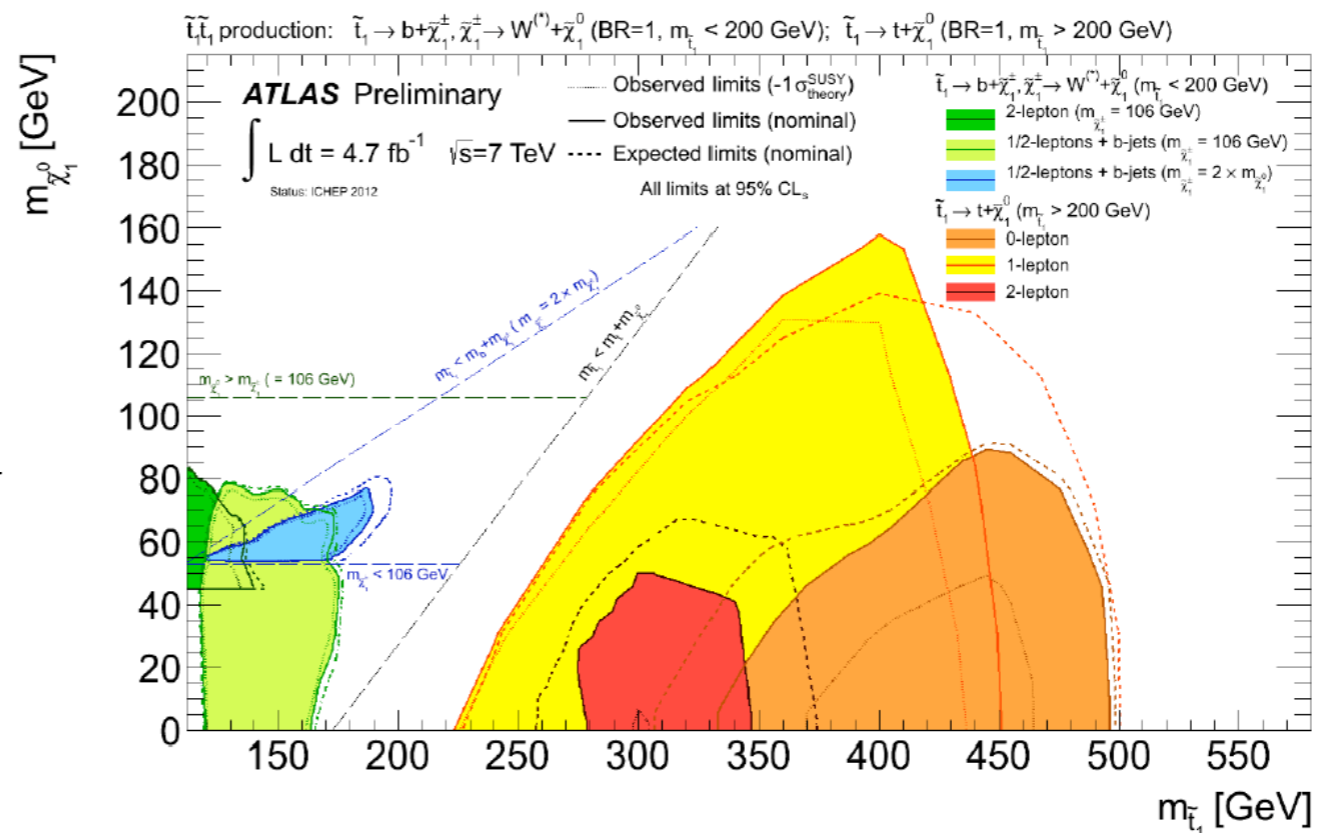
There has been much discussion at this workshop along the lines:

Light stop is highly motivated by naturalness of SUSY.

Light stop is difficult to observe at the LHC. Current limits are weak. Cases such as $m(\tilde{t}) \approx m_t$, $m(\tilde{\chi}) \sim 200$ GeV, competing decay channels subvert these limits.

So, why not

$$m(\tilde{t}) < 250 \text{ GeV} \quad ?$$

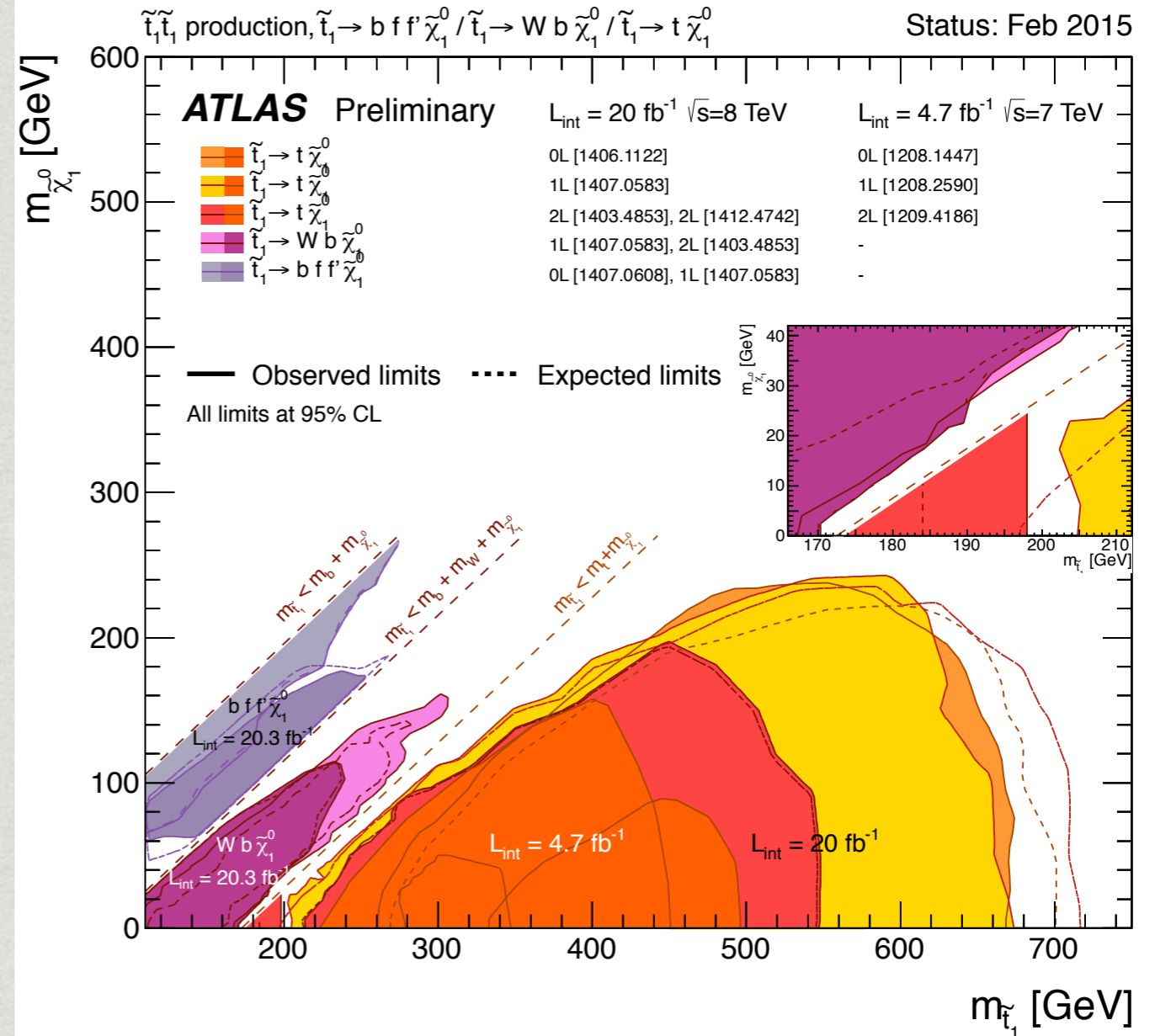
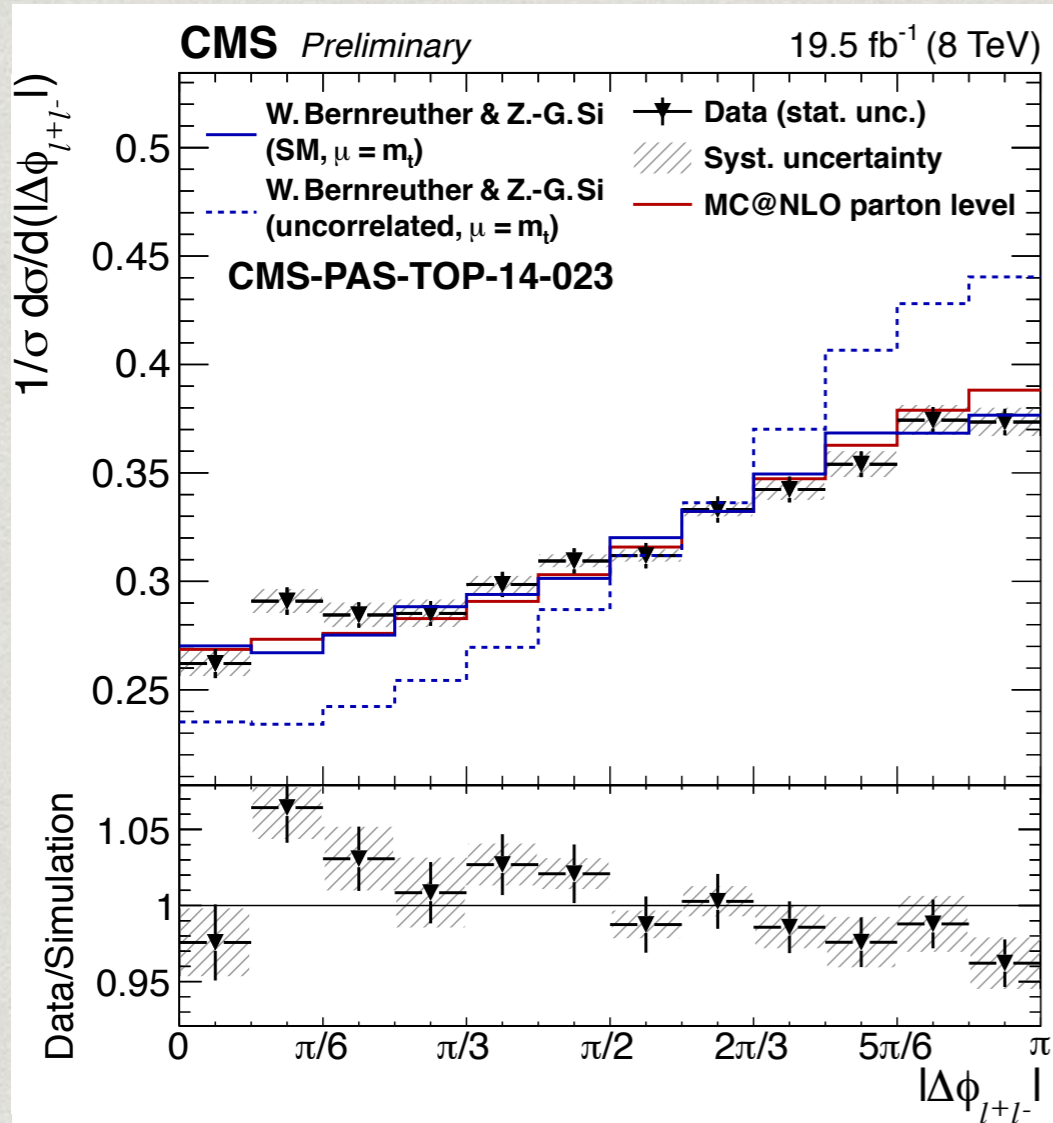


$\Delta\phi(\ell\ell)$

1205.5808

1412.4742+ATLAS-CONF-2014-056

$$\mathcal{L}_{eff} = -\frac{\tilde{\mu}_t}{2} \bar{t} \sigma^{\mu\nu} T^a t G_{\mu\nu}^a - \frac{\tilde{d}_t}{2} i \sigma^{\mu\nu} \gamma_5 T^a t G_{\mu\nu}^a$$

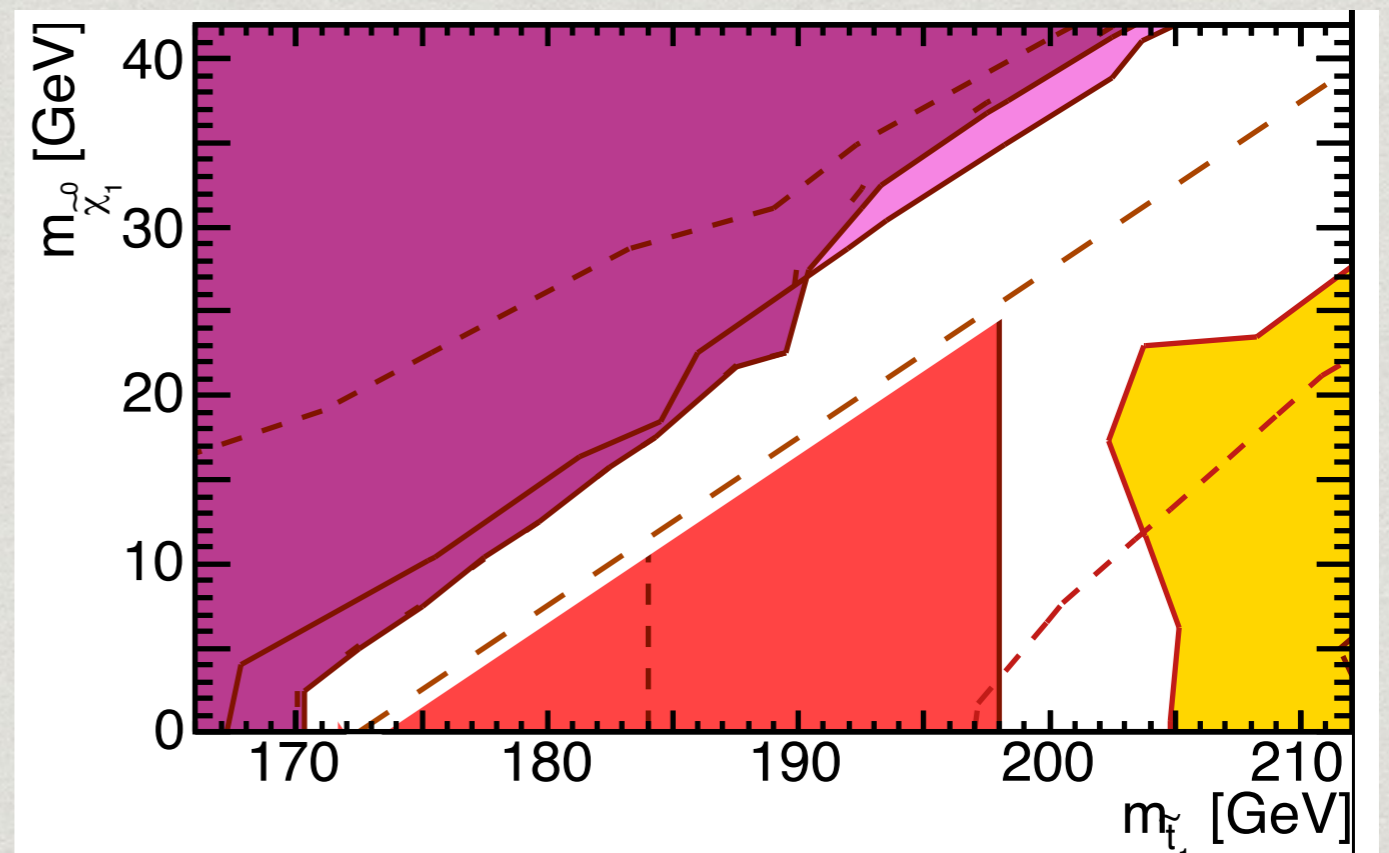
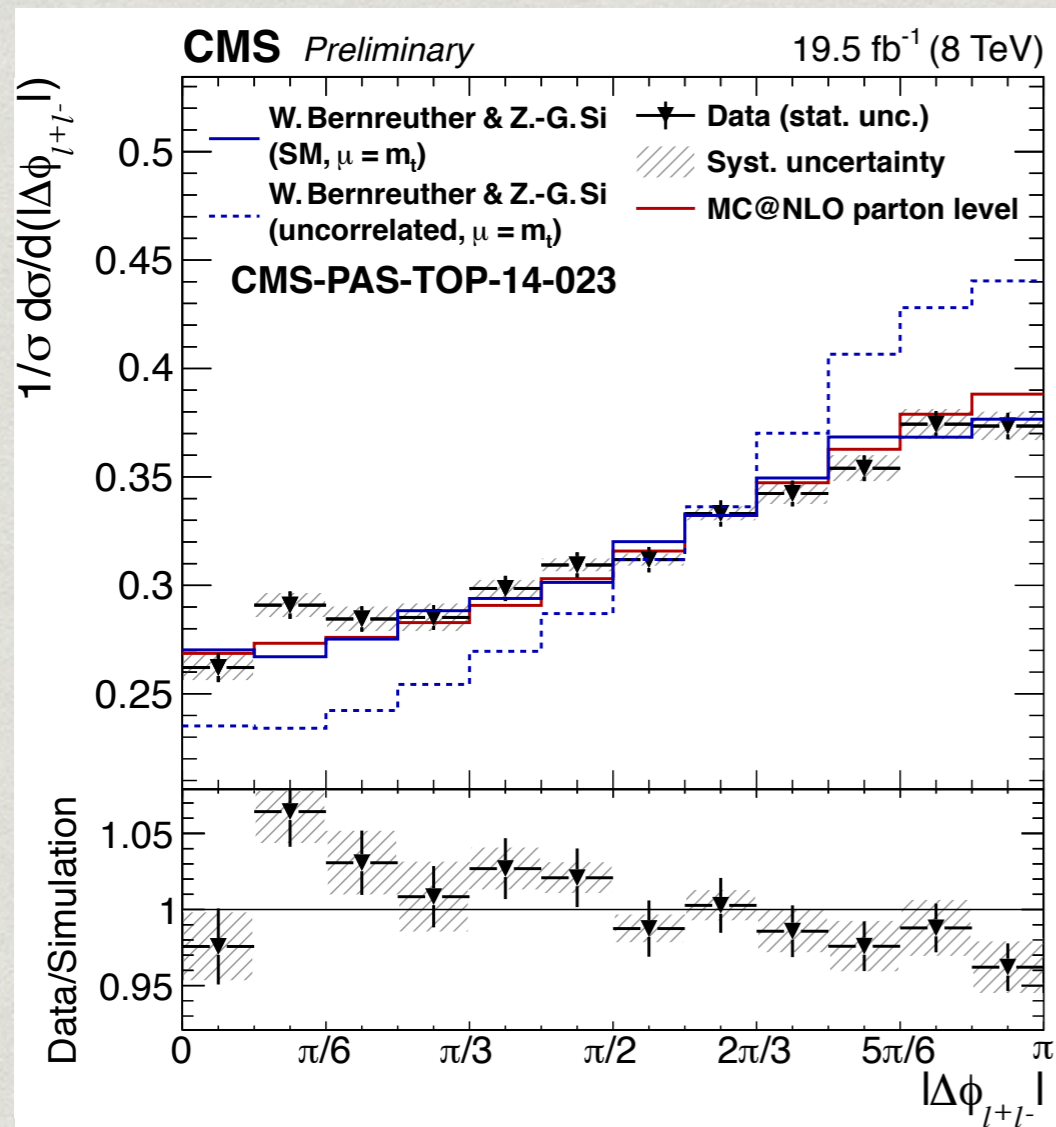


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$$\mathcal{L}_{eff} = -\frac{\tilde{\mu}_t}{2} \bar{t} \sigma^{\mu\nu} T^a t G_{\mu\nu}^a - \frac{\tilde{d}_t}{2} i \sigma^{\mu\nu} \gamma_5 T^a t G_{\mu\nu}^a$$



\tilde{t}_1, \tilde{t}_1 production, $\tilde{t}_1 \rightarrow b f f' \tilde{\chi}_1^0$ / $\tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0$ / $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$

$m_{\tilde{\chi}_1^0}$ [GeV]

ATLAS Preliminary

$L_{int} = 20 \text{ fb}^{-1} \sqrt{s} = 8 \text{ TeV}$

$L_{int} = 4.7 \text{ fb}^{-1} \sqrt{s} = 7 \text{ TeV}$

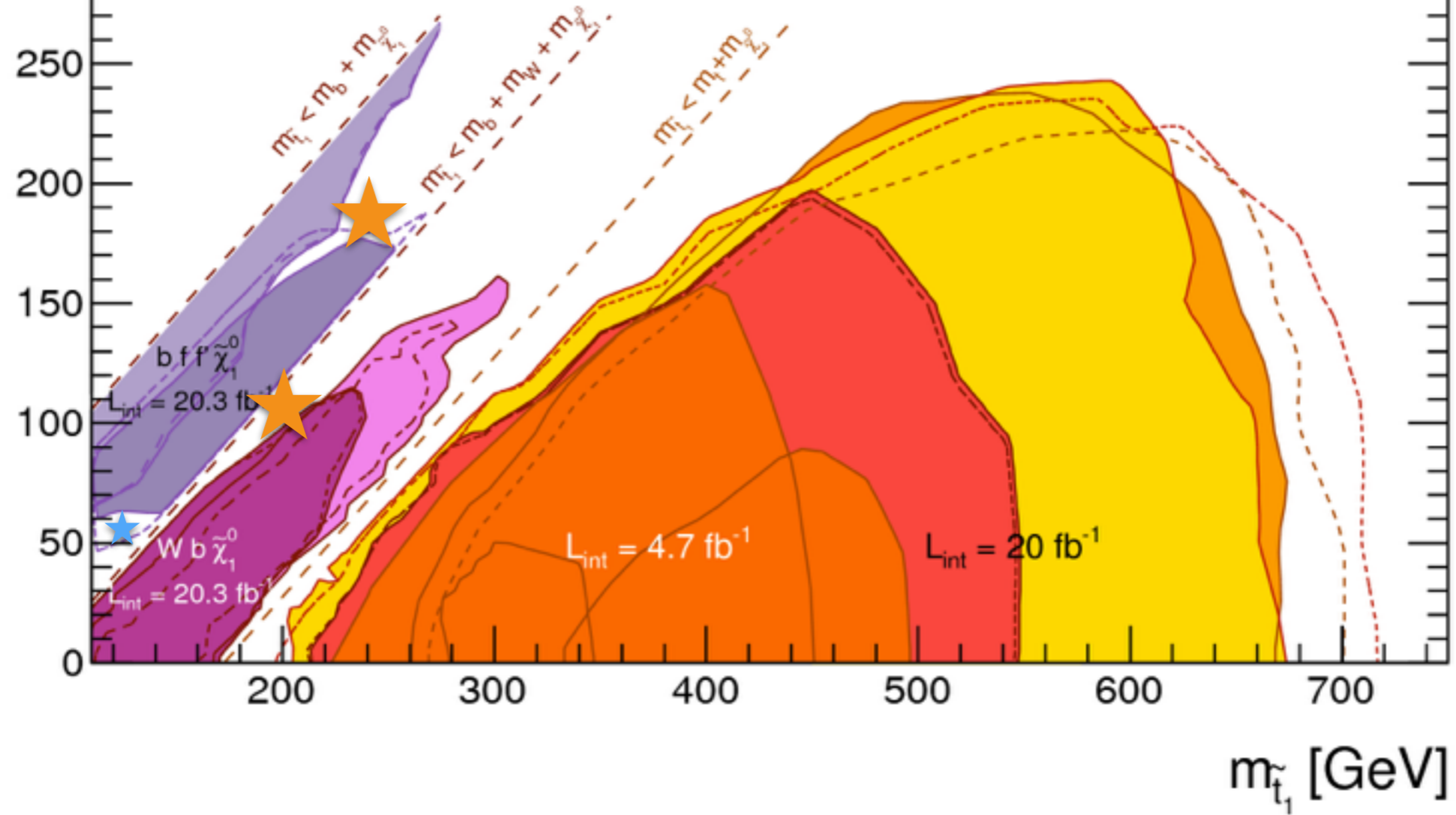
- $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$
- $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$
- $\tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$
- $\tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0$
- $b f f' \tilde{\chi}_1^0$

0L	1406.1122
1L	[1407.0583]
2L	[1403.4853]
1L	[1407.0583], 2L [1403.4853]
0L	[1407.0608], 1L [1407.0583]

0L	[1208.1447]
1L	[1208.2590]
2L	[1209.4186]
-	-
-	-

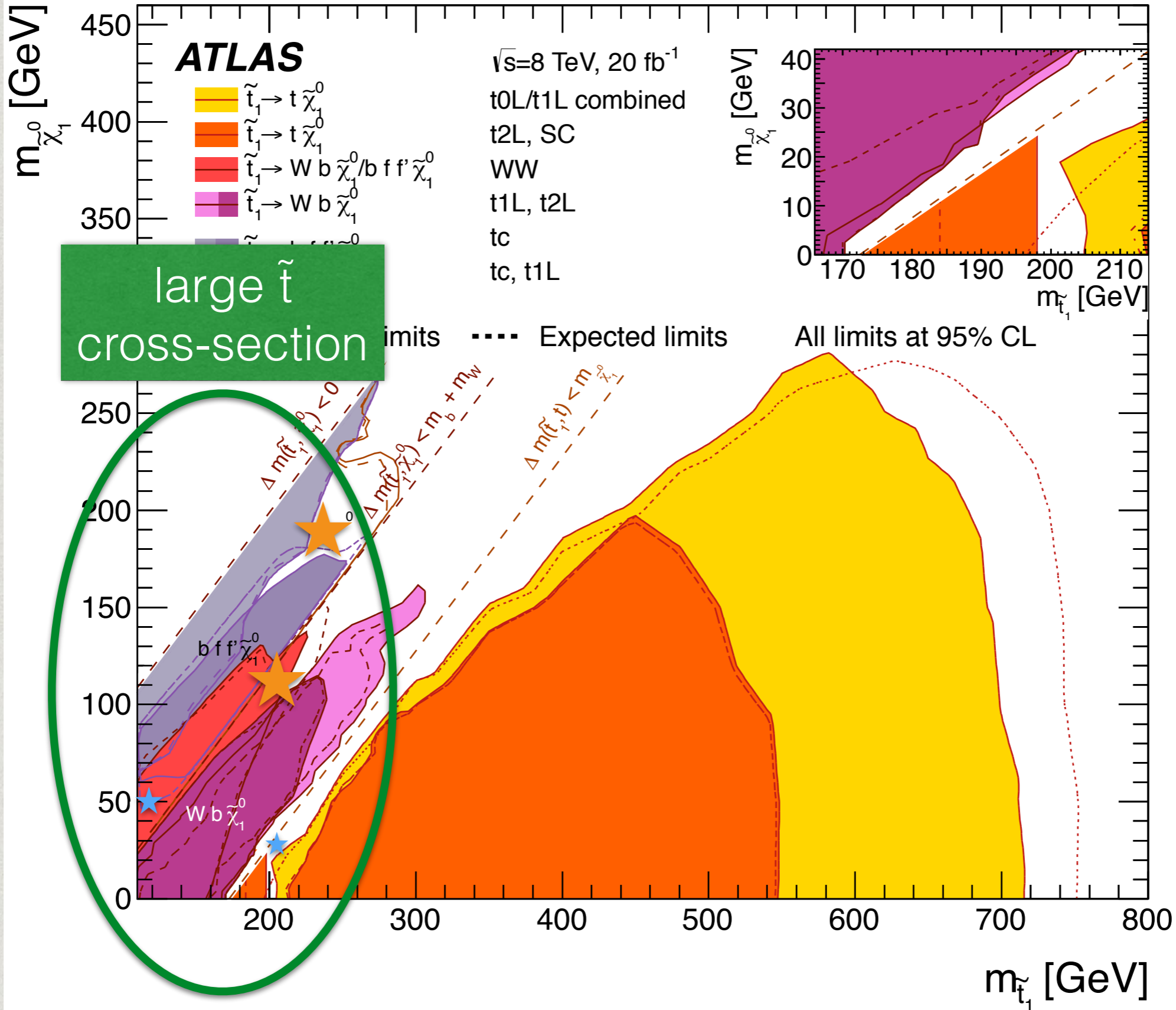
— Observed limits - - - - Expected limits

All limits at 95% CL



$m_{\tilde{t}_1}$ [GeV]

\tilde{t}_1, \tilde{t}_1 production, $\tilde{t}_1 \rightarrow b f' \tilde{\chi}_1^0 / \tilde{t}_1 \rightarrow W b \tilde{\chi}_1^0 / \tilde{t}_1 \rightarrow t \tilde{\chi}_1^0$



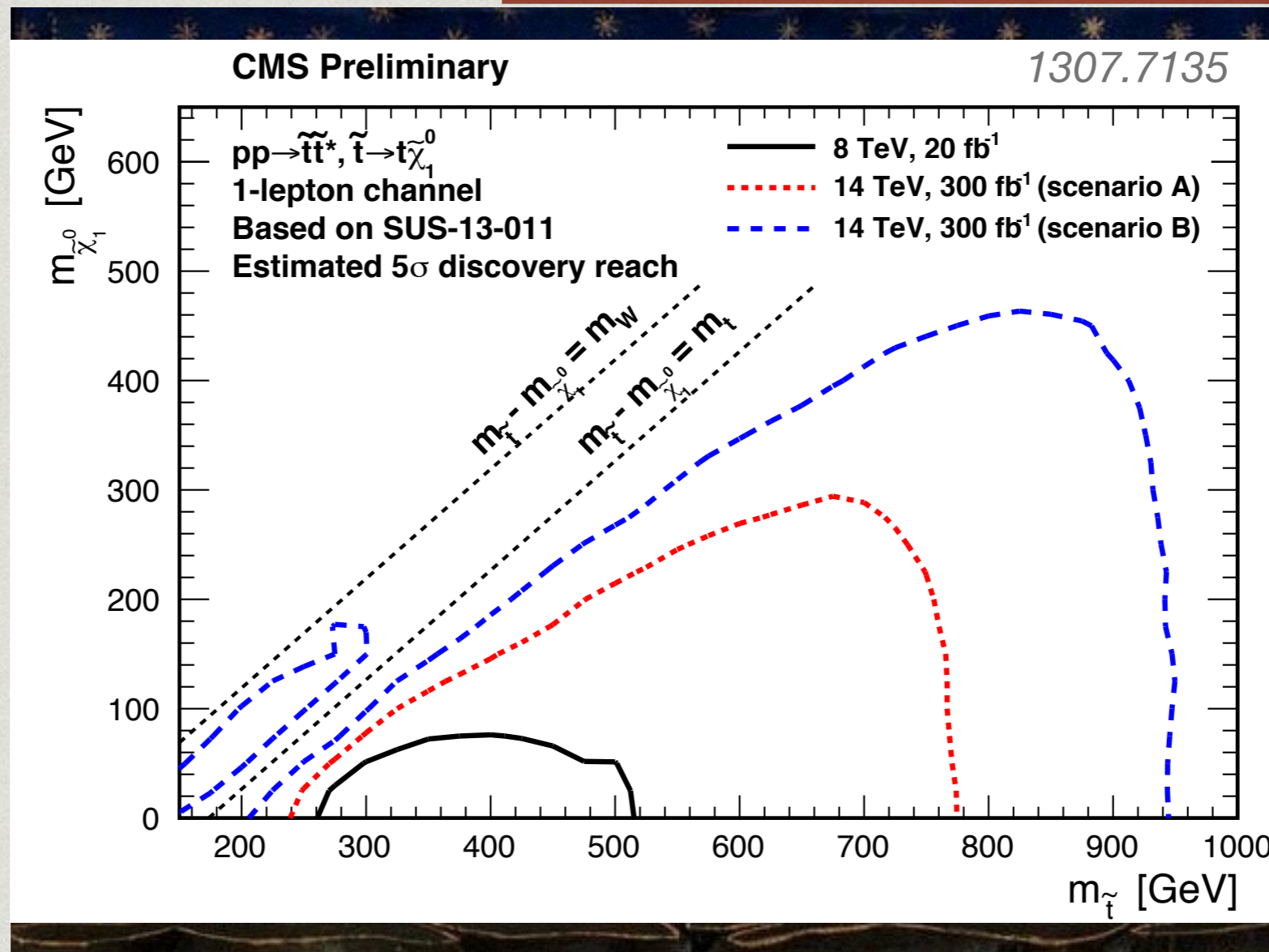
New physics after the end of LHC

Cristoforo de Predis (o Preda) (Milano 1440/1445 - 1486)



New physics after the end of LHC

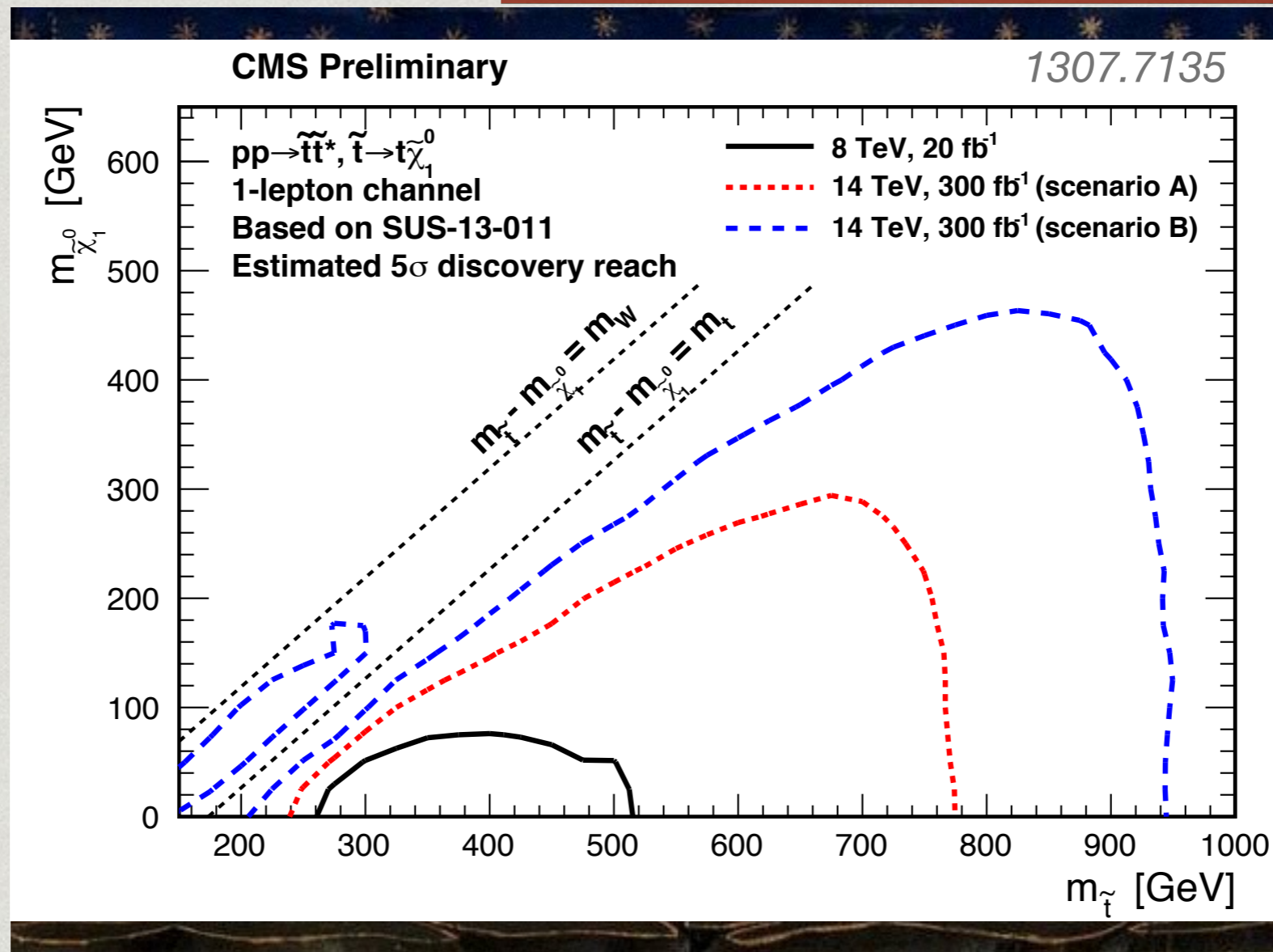
Cristoforo de Predis (o Preda) (Milano 1440/1445 - 1486)



most of the 5σ discovery region is excluded at 95% C.L.

New physics after the end of LHC

Cristoforo de Predis (o Preda) (Milano 1440/1445 - 1486)



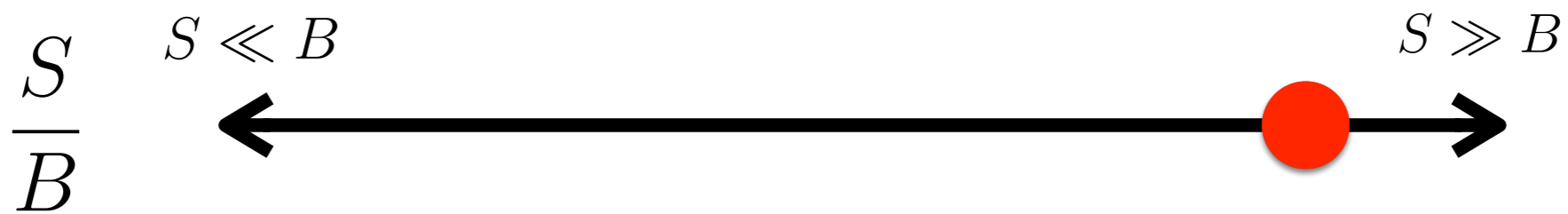
FOCUS ON SUBTLE NEW PHYSICS

We can search well, “hard new physics” (high- p_T)

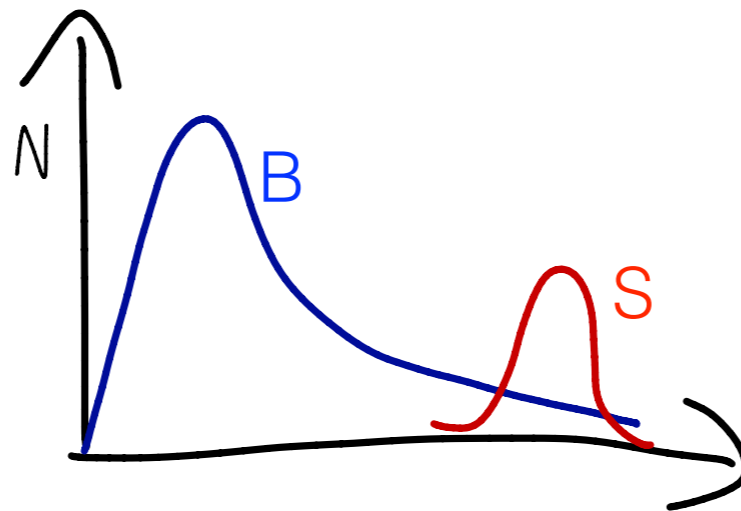
We struggle with the other half of the possibilities (SM-like)

FROM THE KEYNOTE

signal v. background

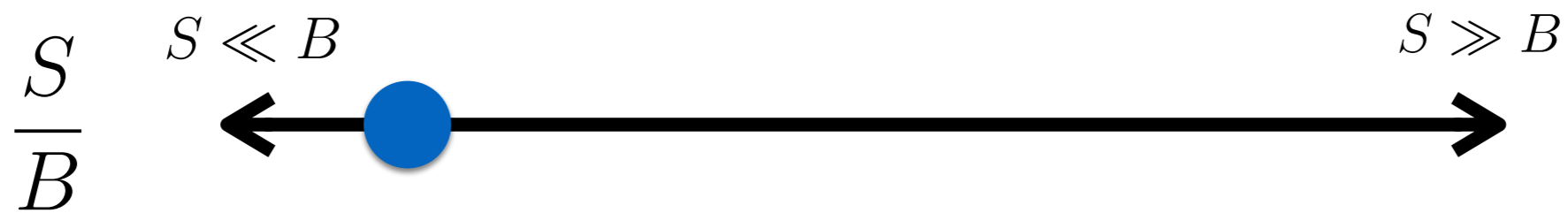


usually we try to separate new physics from the background

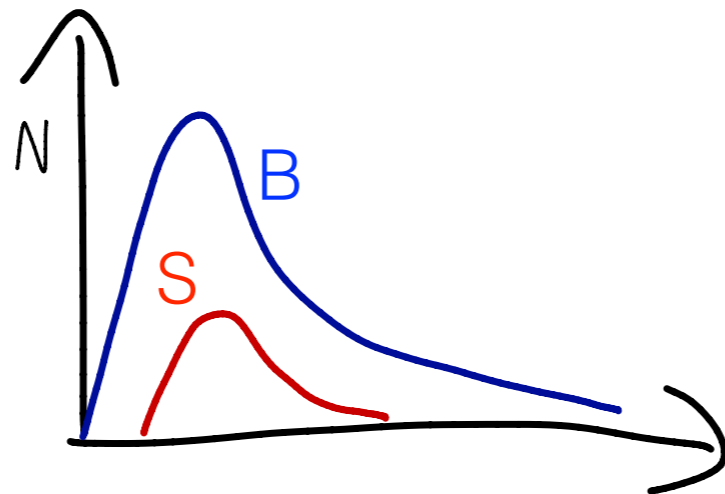


FROM THE KEYNOTE

signal v. background

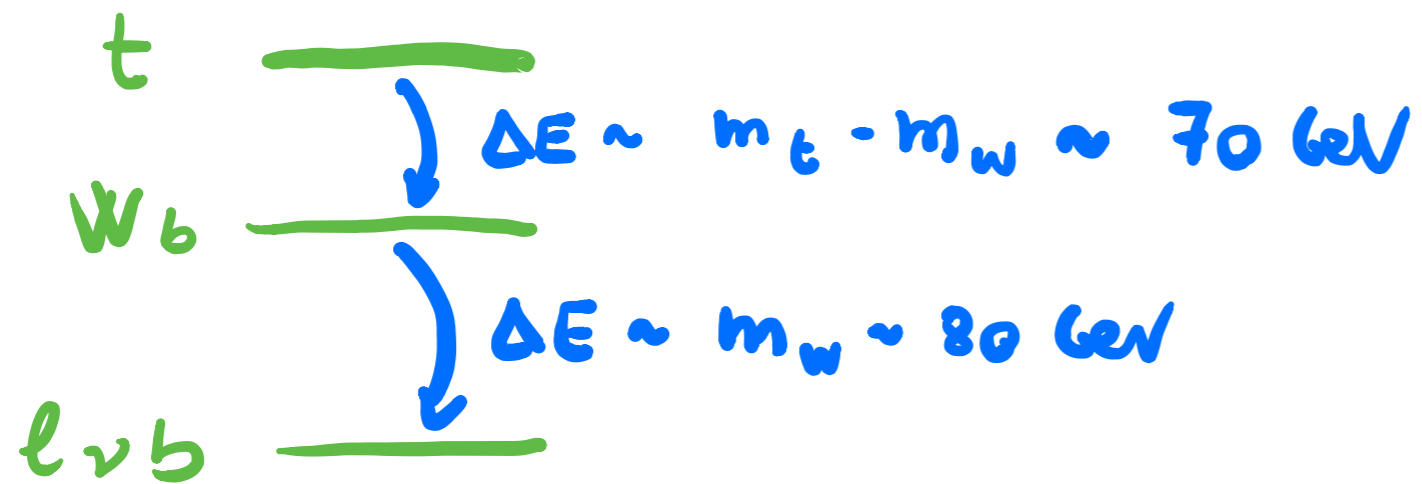


Sometimes we can learn about new physics by precisely studying the background

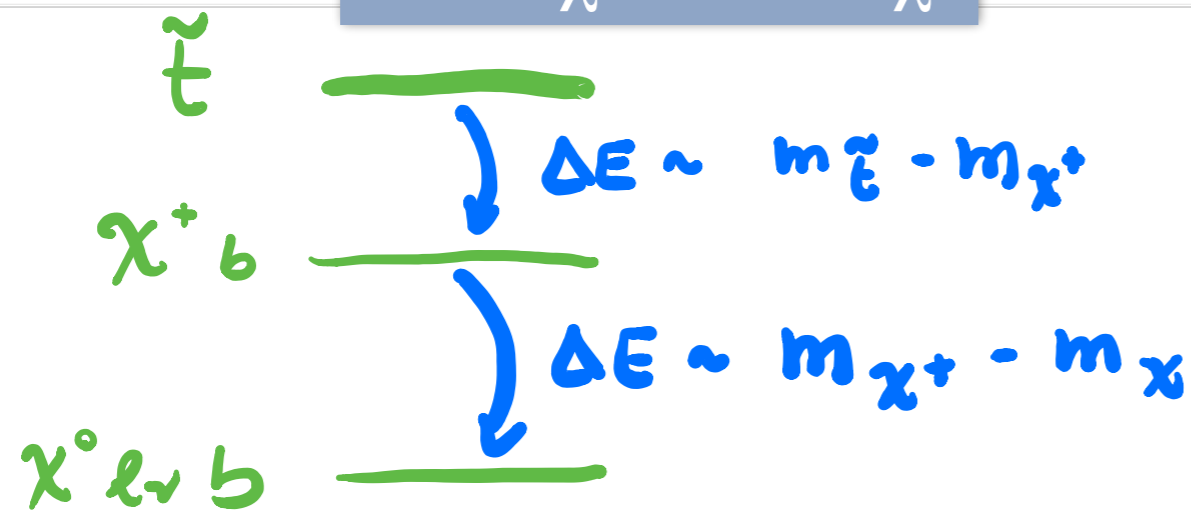


Soft is hard

$t \rightarrow bW \rightarrow b\ell\nu$

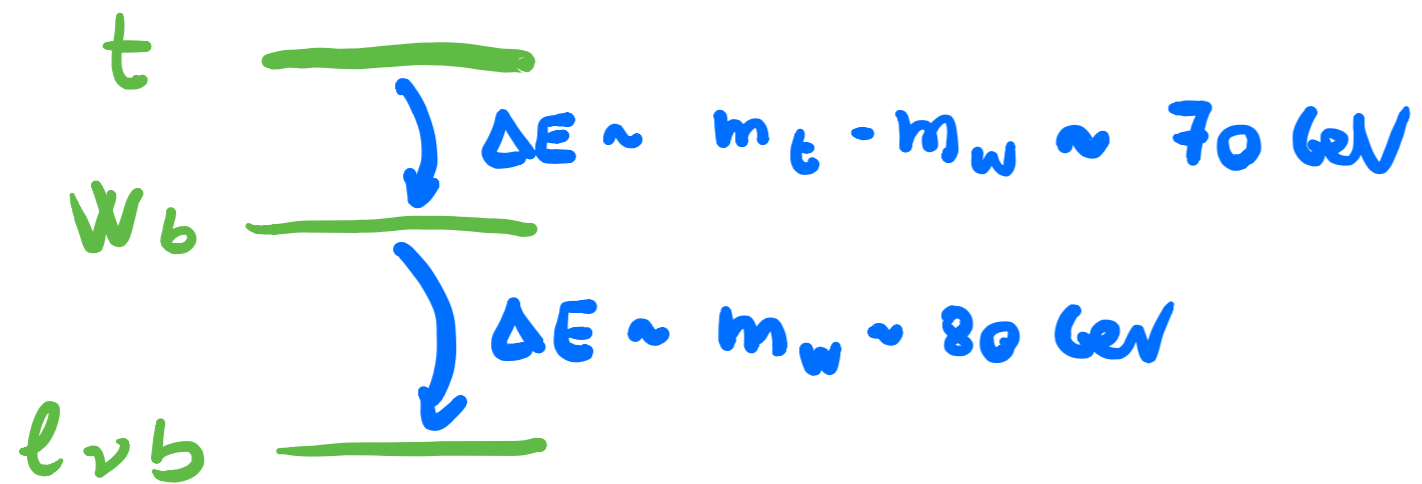


$\tilde{t} \rightarrow b\chi^+ \rightarrow b\ell\nu\chi^0$



Soft is hard

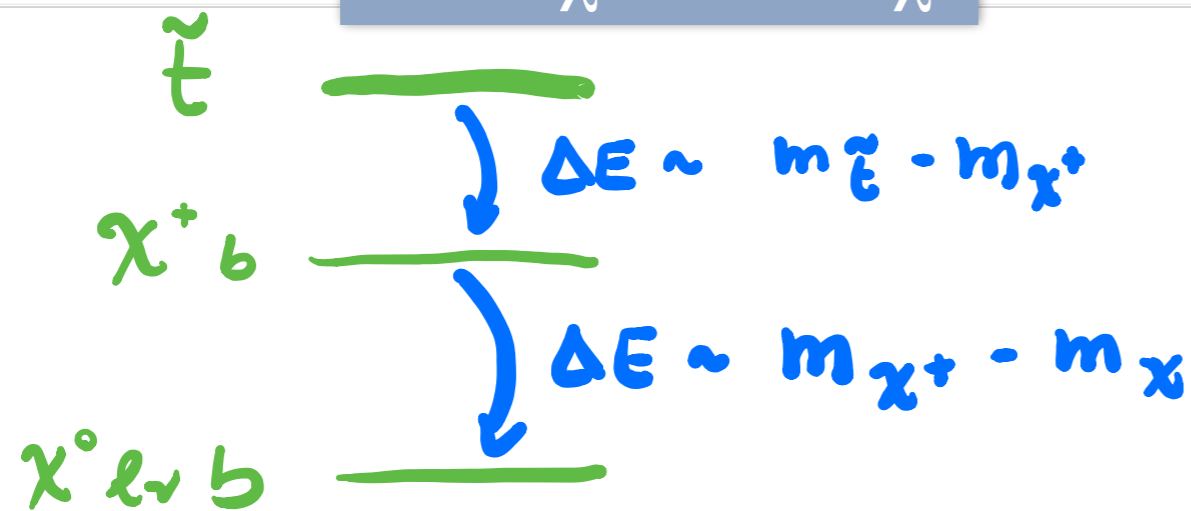
$$t \rightarrow bW \rightarrow b\ell\nu$$



New physics is SM-like



$$\tilde{t} \rightarrow b\chi^+ \rightarrow b\ell\nu\chi^0$$



Better precision

RESOLUTION

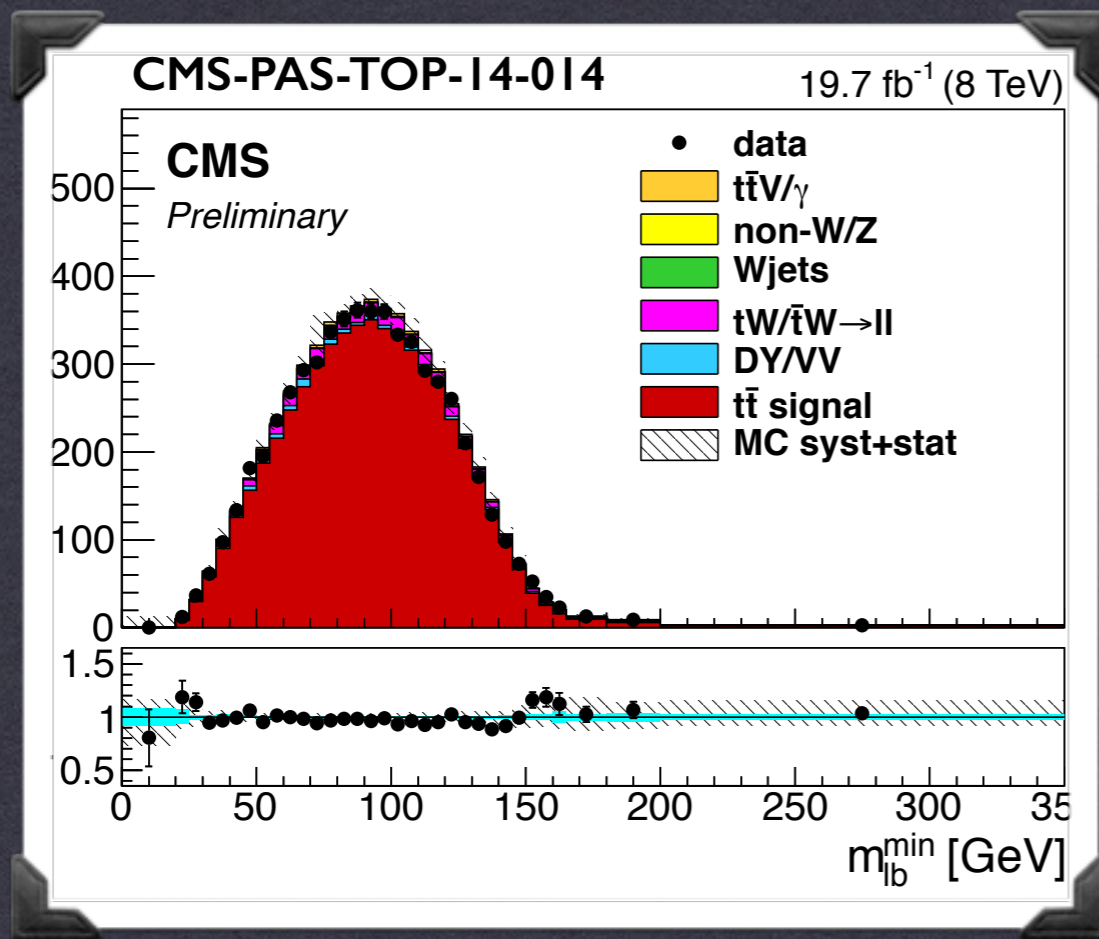
RESOLUTION

CLEAN ENVIRONMENT



RESOLUTION

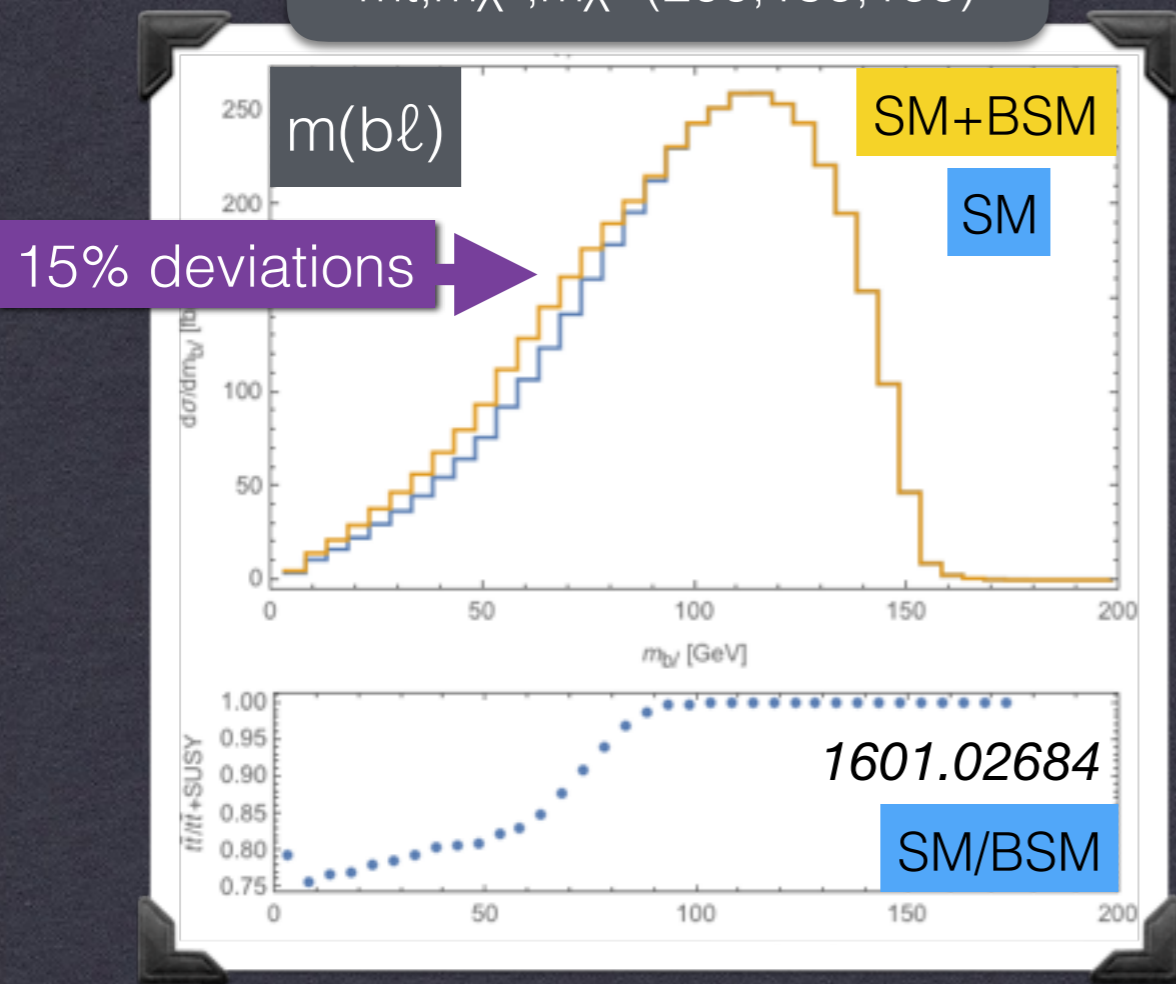
PRECISION AT LHC



RESOLUTION

PRECISION AT LHC

$m_{\tilde{t}}, m_{\chi^+}, m_{\chi^0} = (200, 150, 100)$



$m(b\ell)$

SM+BSM

SM

15% deviations

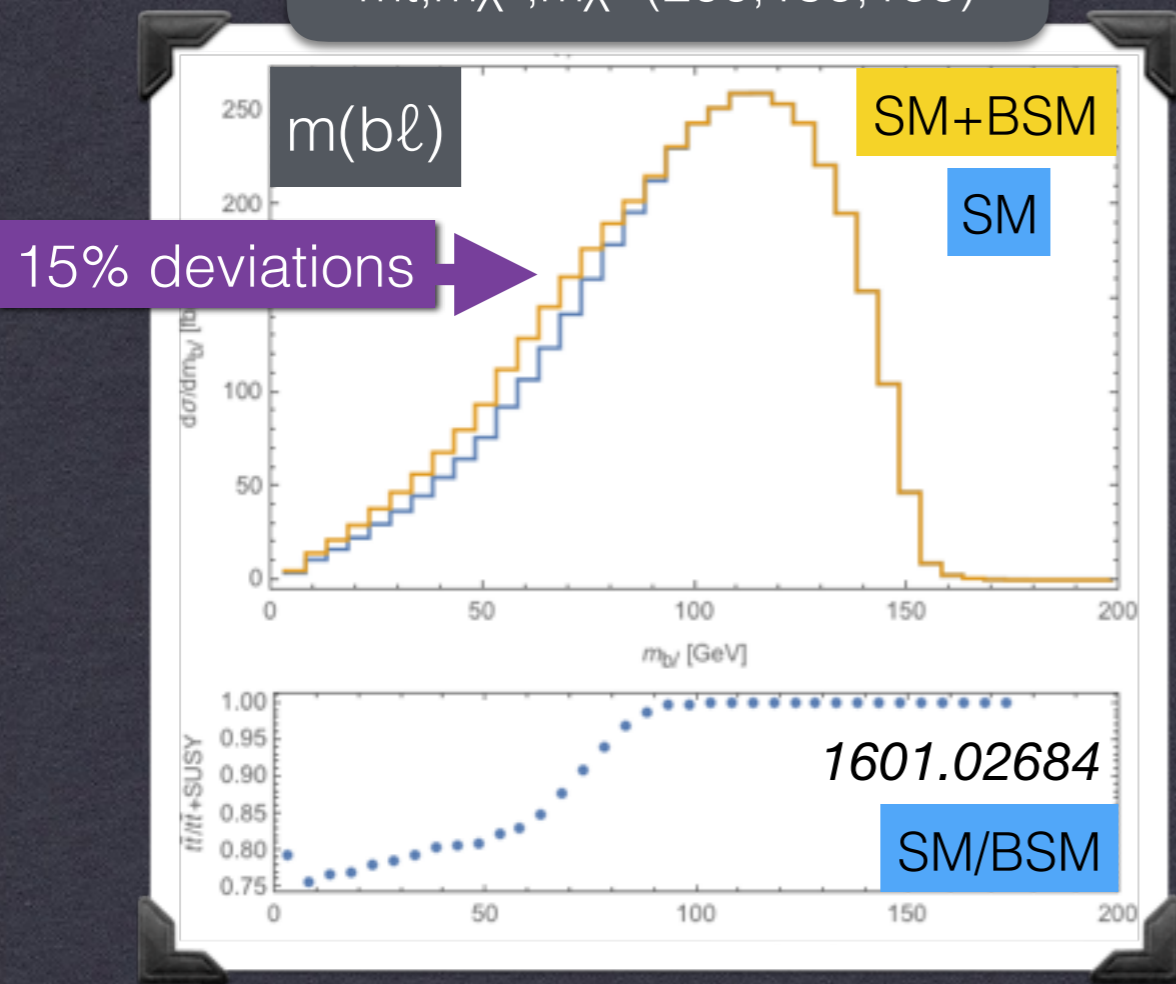
1601.02684

SM/BSM

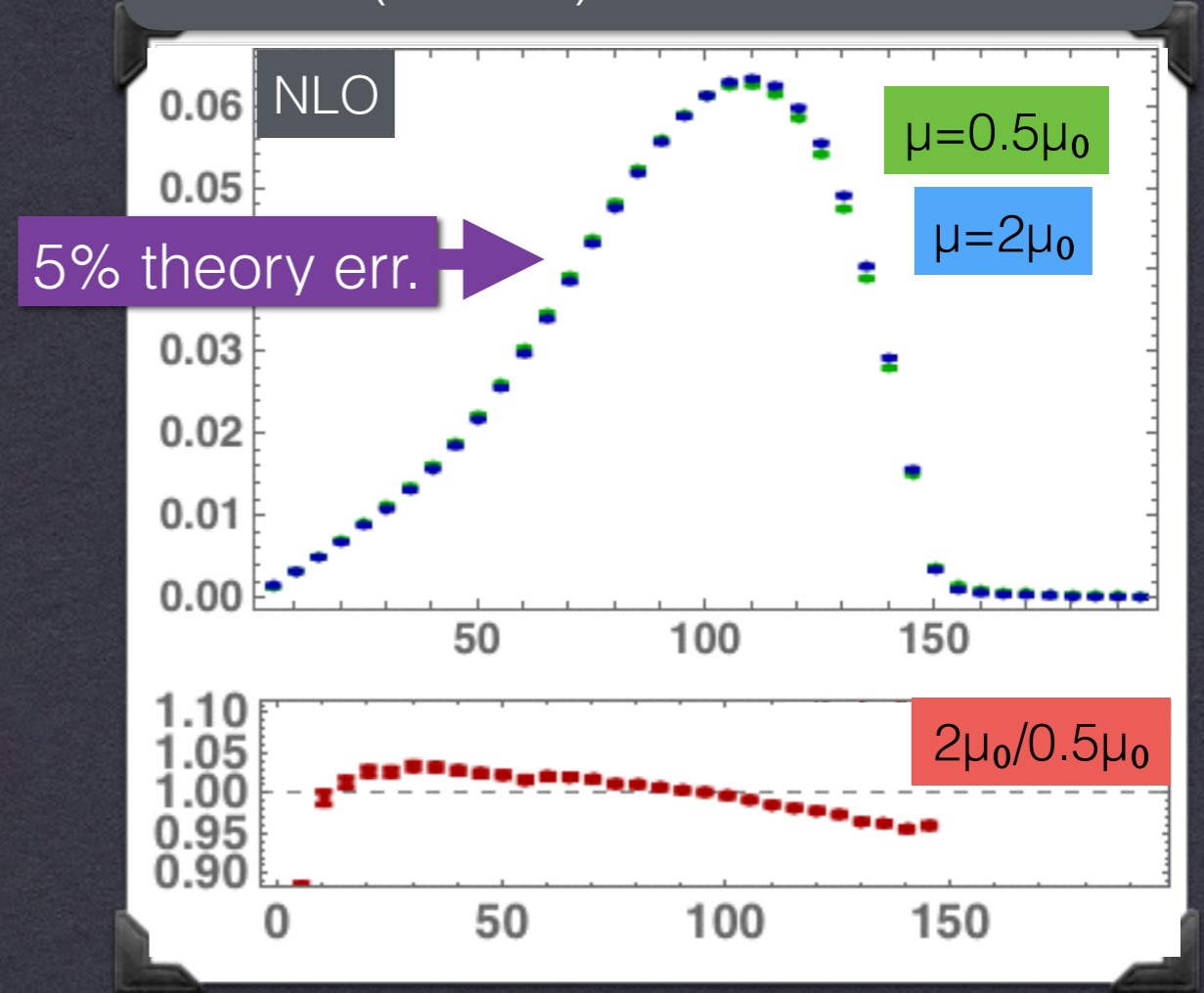
RESOLUTION

PRECISION AT LHC

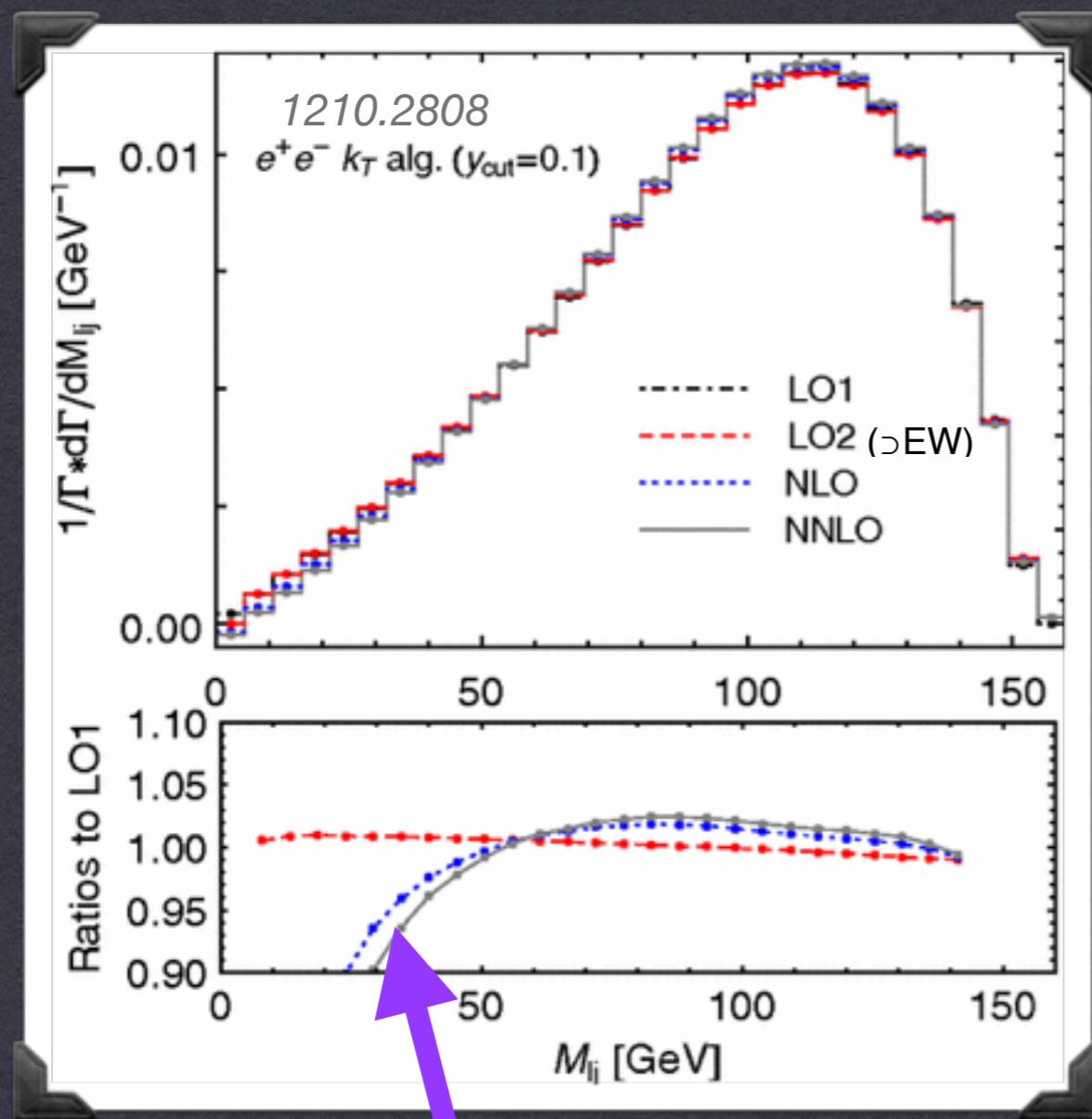
$m_{\tilde{t}}, m_{\chi^+}, m_{\chi^0} = (200, 150, 100)$



SM $t\bar{t}$ (MCFM) as in TOP-14-014



NNLO top decay

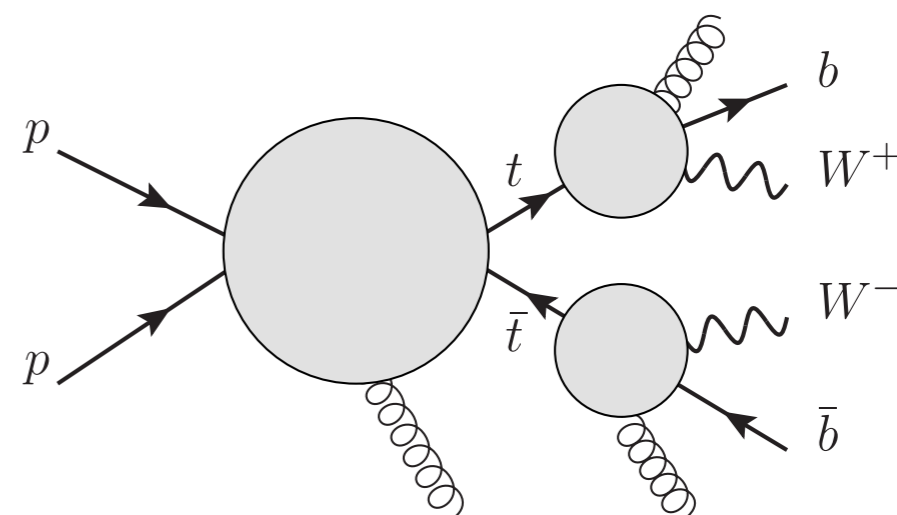


2% NNLO correction

also 1301.7133 NNLO top decay

NLO+PS w/top decay

1607.04538 "bb4l" POWHEG-RES
1412.1828 "tt_dec" POWHEG



S. Pozzorini

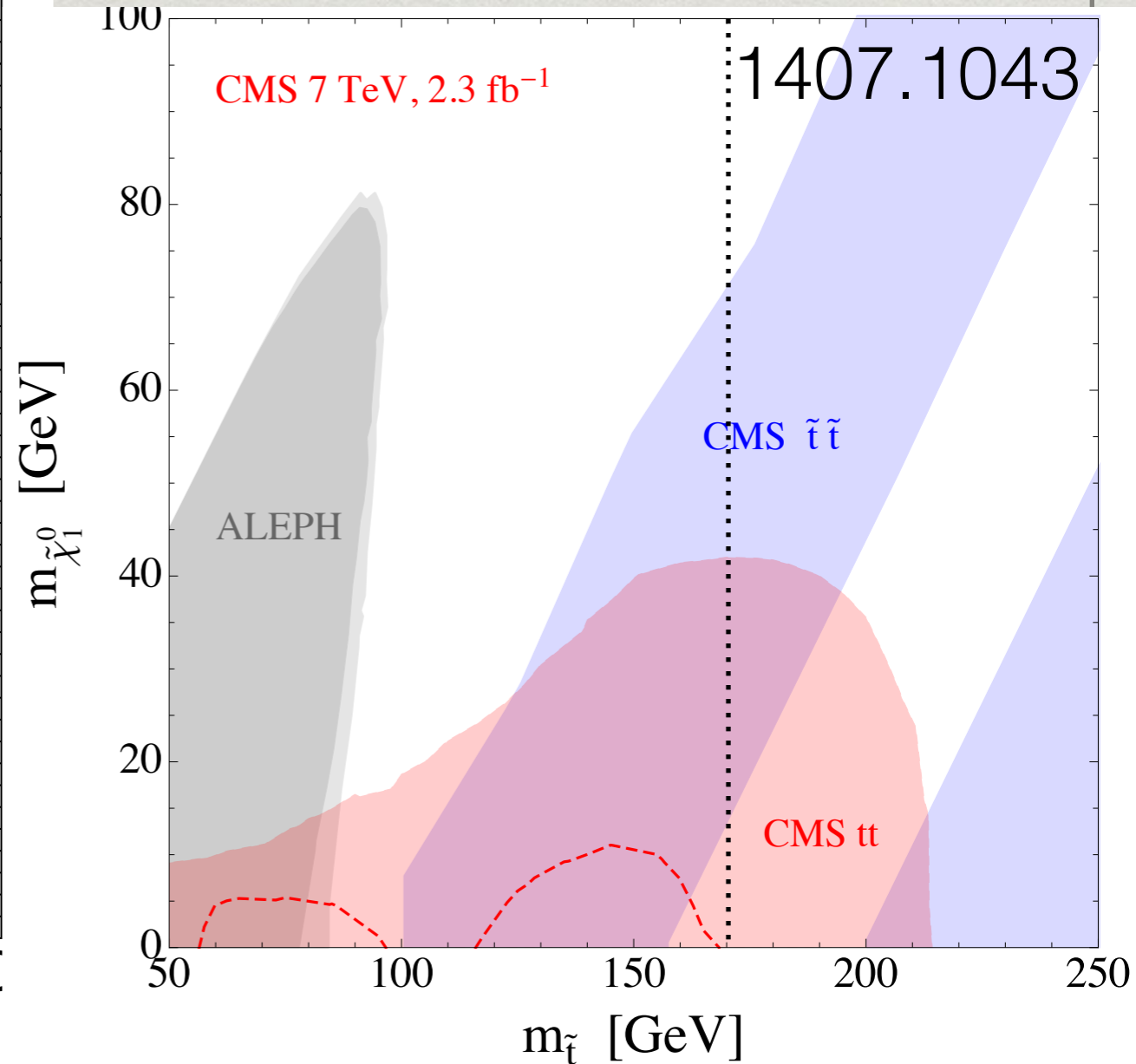
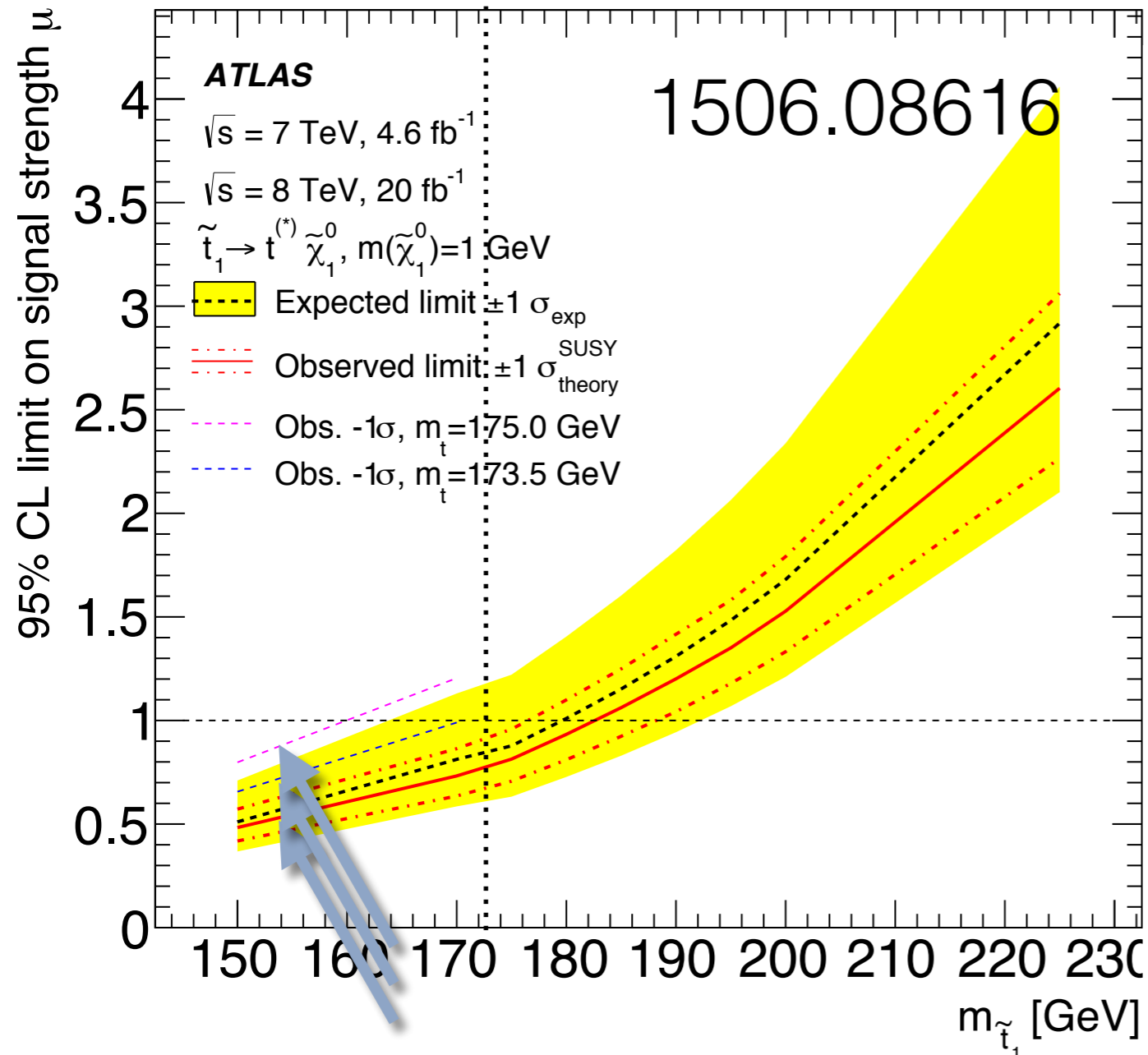
encouraging results
form "tt_dec"

$\sigma(t\bar{t})$

1407.1043 + 1406.5375 + 1506.08616

light stop effects on top cross-section

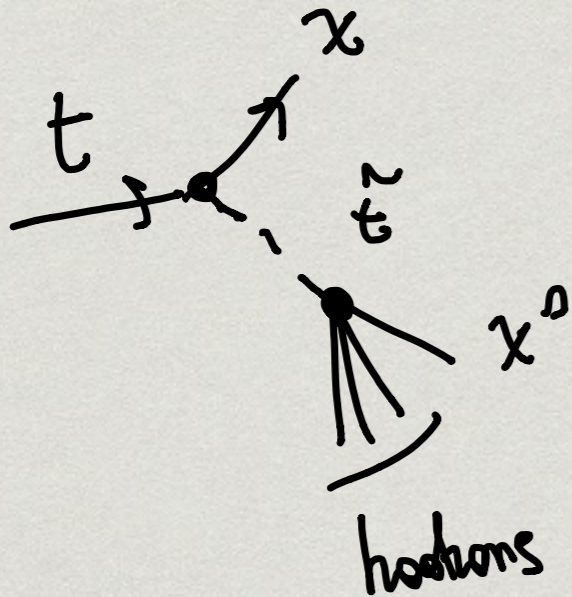
$m_\chi \neq 1$ GeV



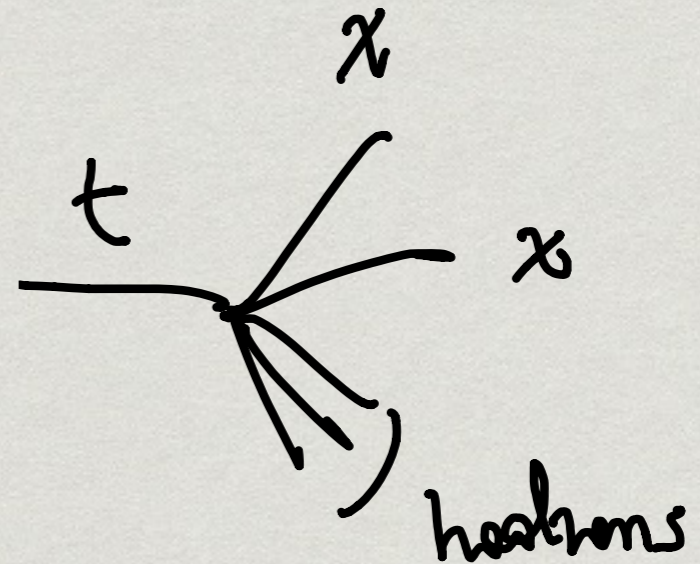
variations of m_{top} reflected on $\sigma(t\bar{t})$

A window on heavier new physics

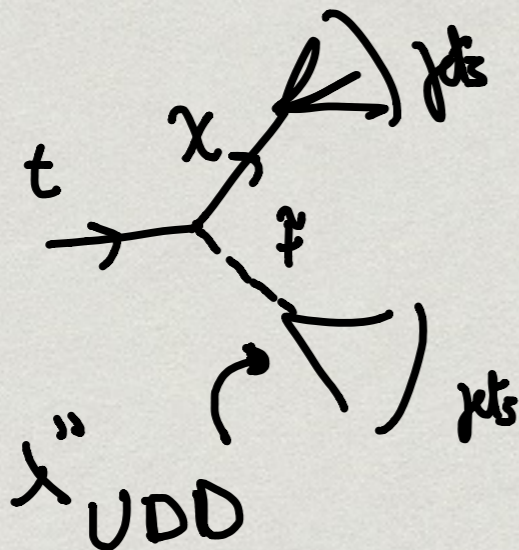
RPC SUSY



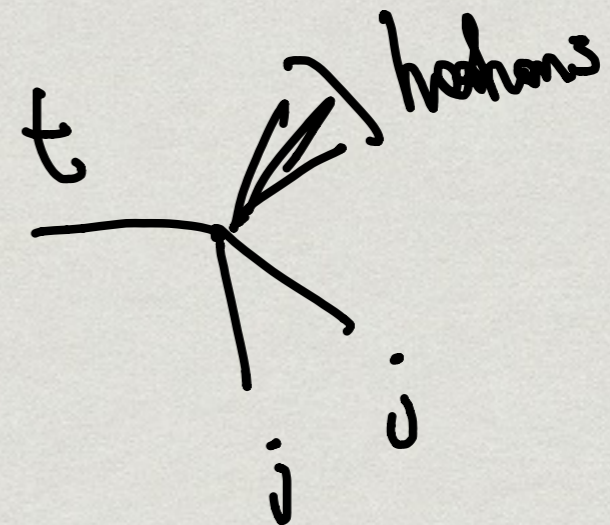
$m_{\tilde{t}} > m_t$



RPV SUSY



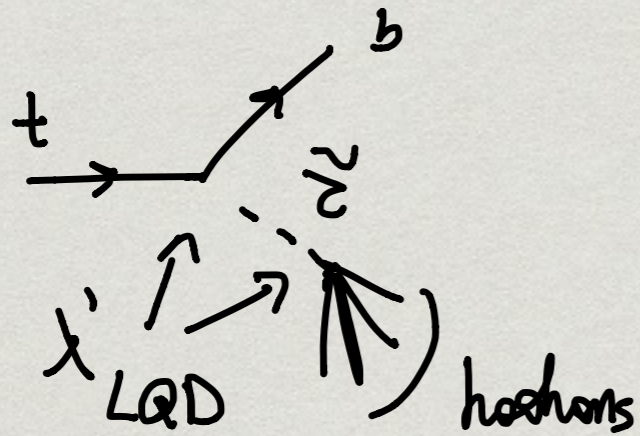
$m_{\tilde{t}} > m_t$



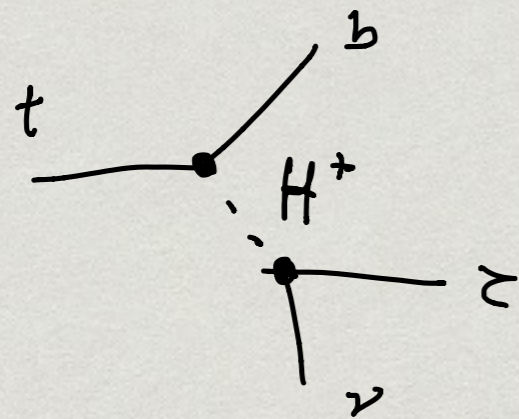
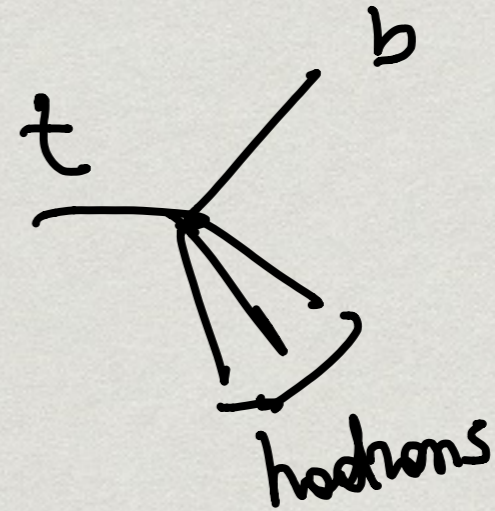
displaced $\tilde{\chi}^0$ decay

A window on heavier new physics

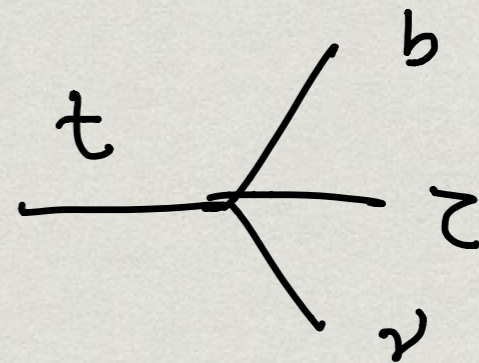
RPV SUSY



$m_{\tilde{t}} > m_t$



$m_{H^+} > m_t$



of course these are just “templates” for thinking, lots of possible signals

top decays to BSM

top as a “portal”

Direct production of light new physics:

- $t \rightarrow \tilde{t} \chi$ (few % BR in the MSSM, $t\text{-}\tilde{t}\text{-}\chi$ coupling)
- $t \rightarrow \tilde{t} b \rightarrow b b c$ (RPV λ' , AFAIK not fully covered at LEP)
- $t \rightarrow b H^+ \rightarrow b \tau \nu$ (CMS-PAS-HIG-12-052)

Indirect test through higher dimensional operators:

- $t \rightarrow c Z, c H$ (and $c \rightarrow q$, 1508.05796, 1312.4194, PAS-TOP-14-020)
- $t \rightarrow b c \ell$ (BNV 1107.3805, 1310.1618)
- $t \rightarrow q e \mu$ (1507.07163)
- $t \rightarrow q W$ (1404.2292)
- $t \rightarrow b b c$ (1407.1724, 1407.1725)
- global BR measurement (1506.05074)

Summary and outlook

- * Accumulating evidence indicates that gaps in searches are difficult to fill at the LHC
- * Projections are difficult, but a breakthrough seems needed to do better with standard searches strategies
- * **Precision in top quark physics has already proven it can be useful for BSM ($\Delta\varphi_{\ell\ell}$, total cross-section, ...)**
- * **more precision studies interpreted for BSM are needed**

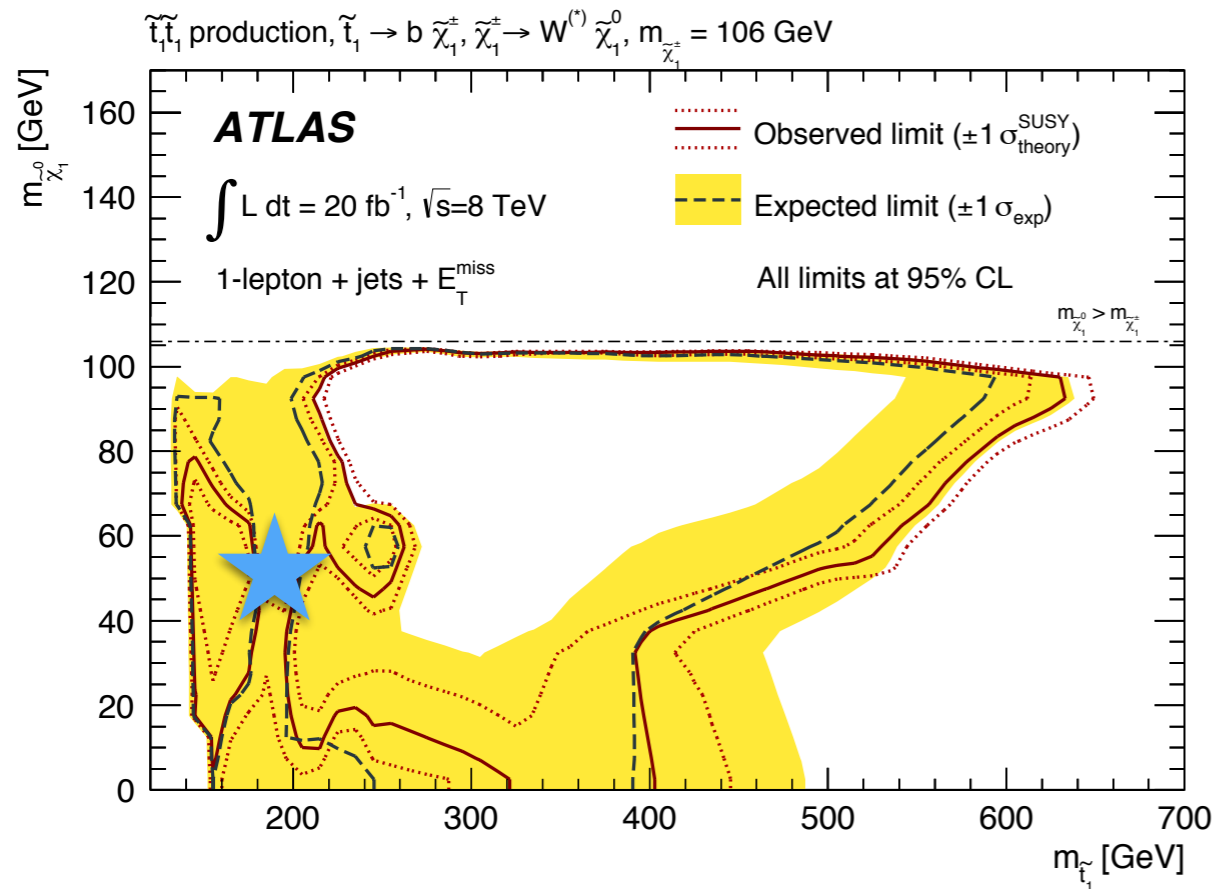
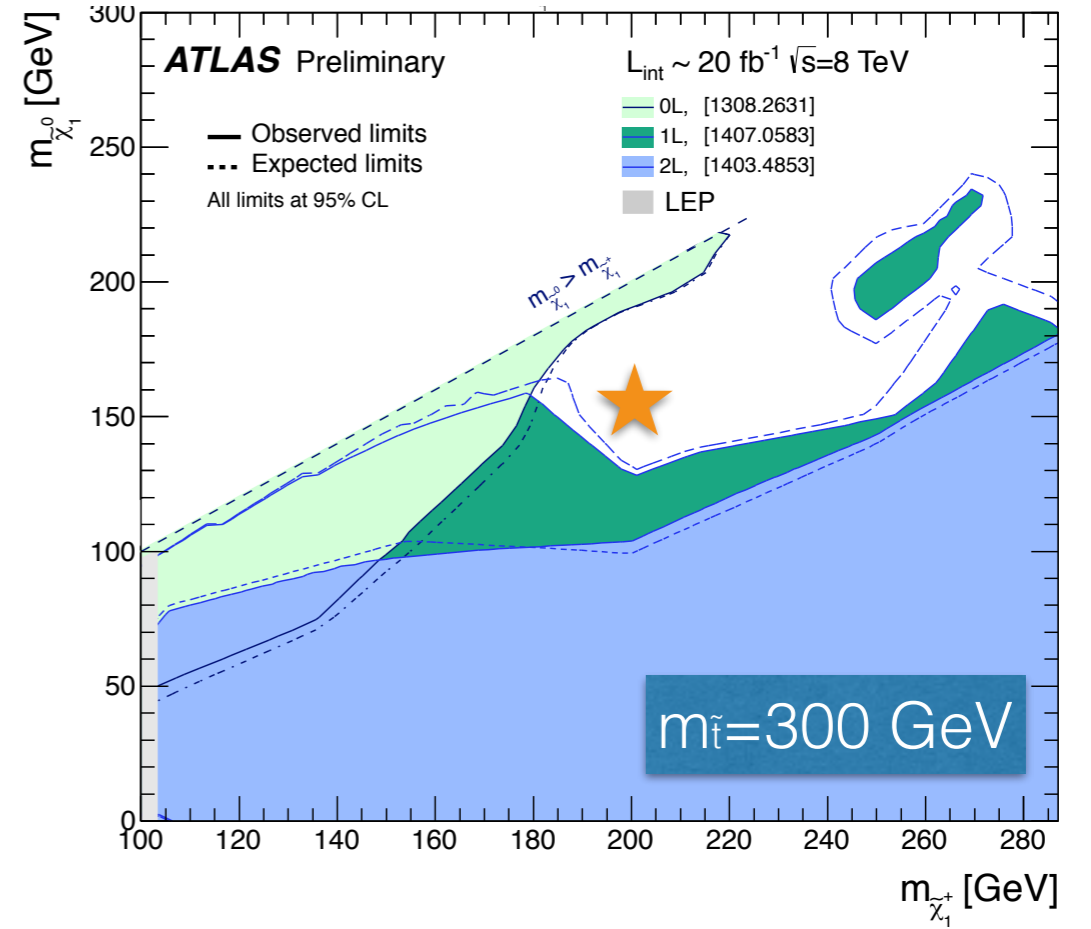
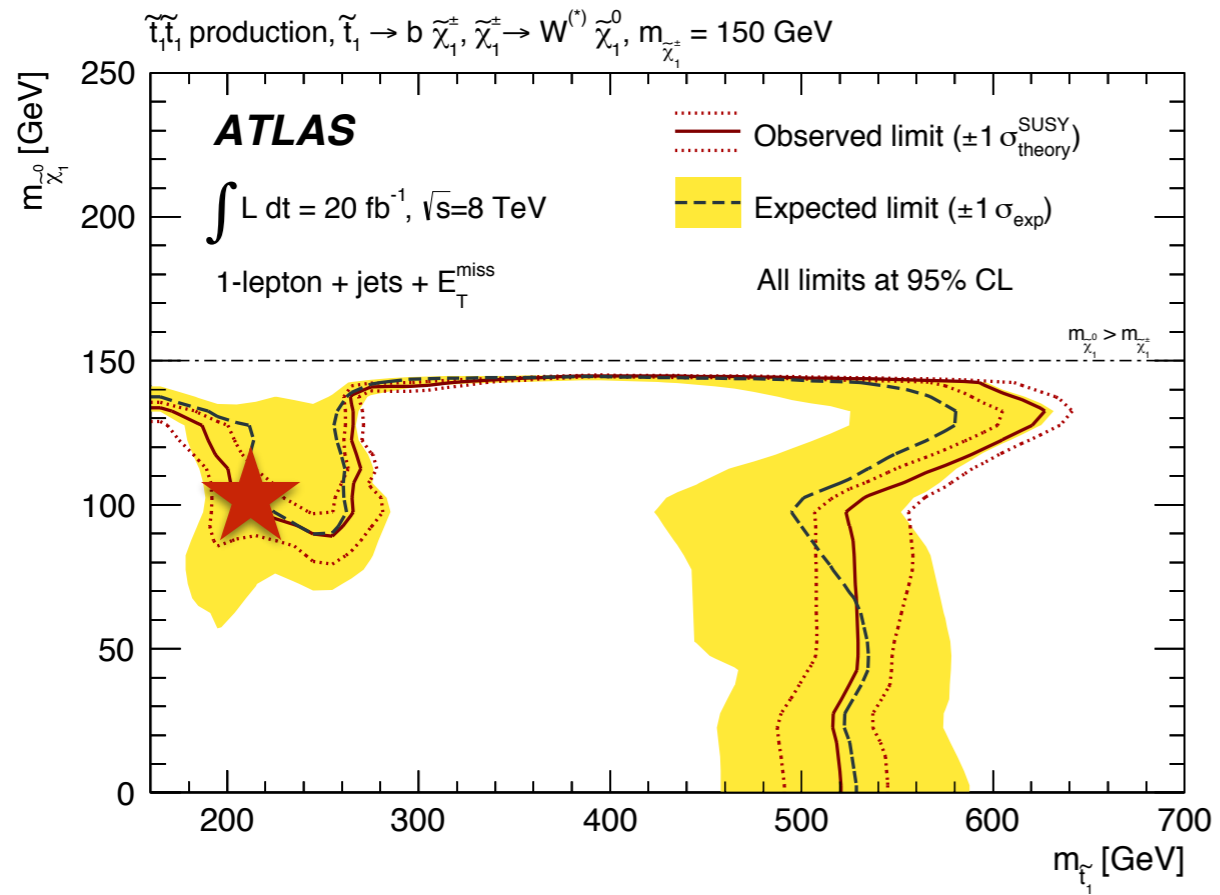
Summary and outlook

- * BSM needs top quark precision physics!
- * two examples discussed:
 - o \tilde{t} production
 - o $t \rightarrow$ BSM (e.g. $t \rightarrow \tilde{t}\chi^0$) and $t \rightarrow$ SM via HDO
- * new colored states in blindspots

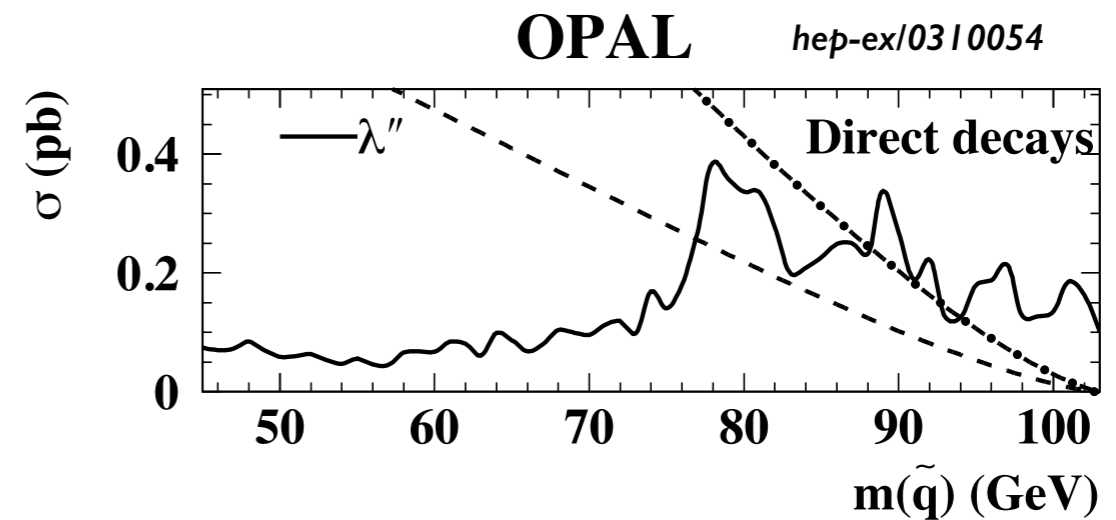
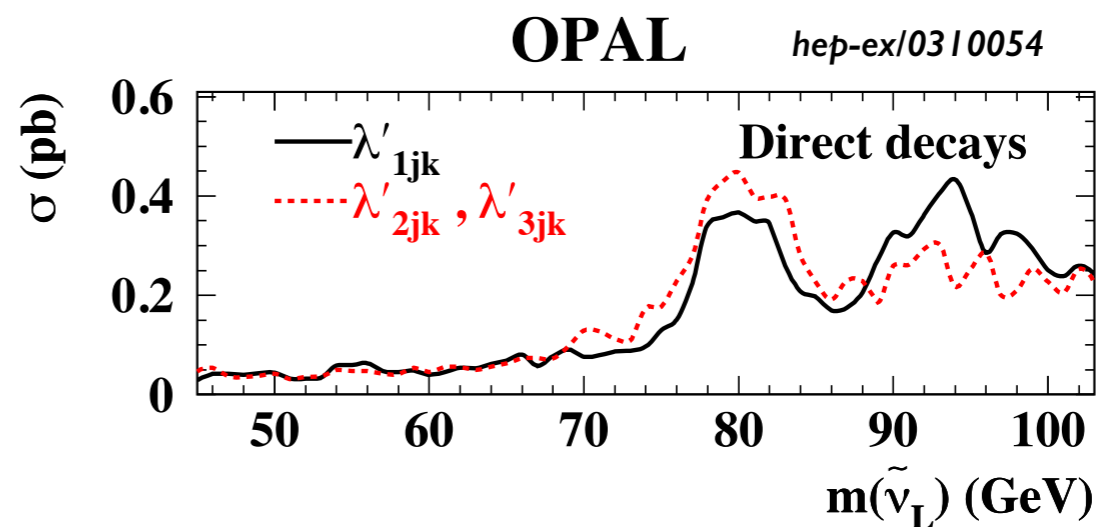
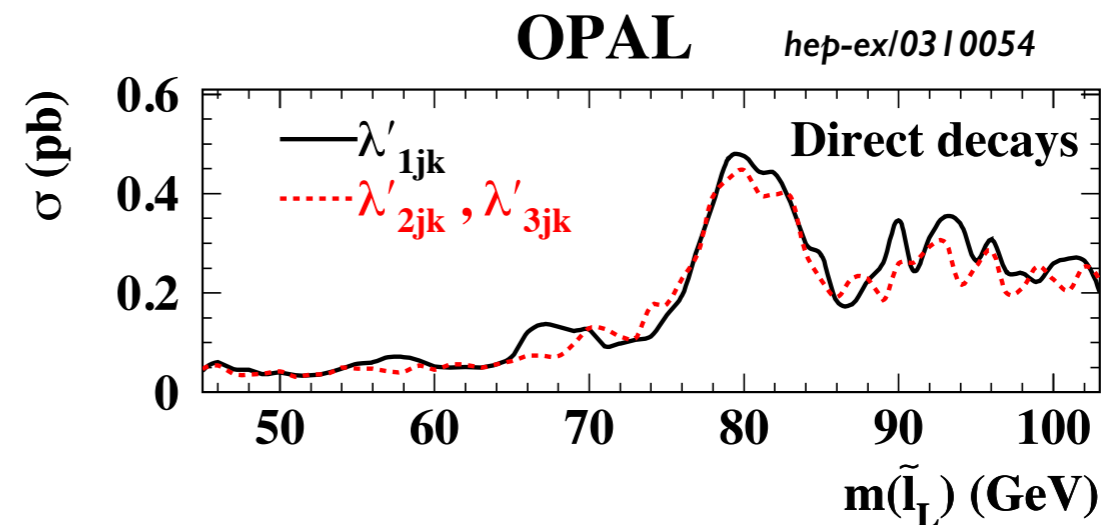
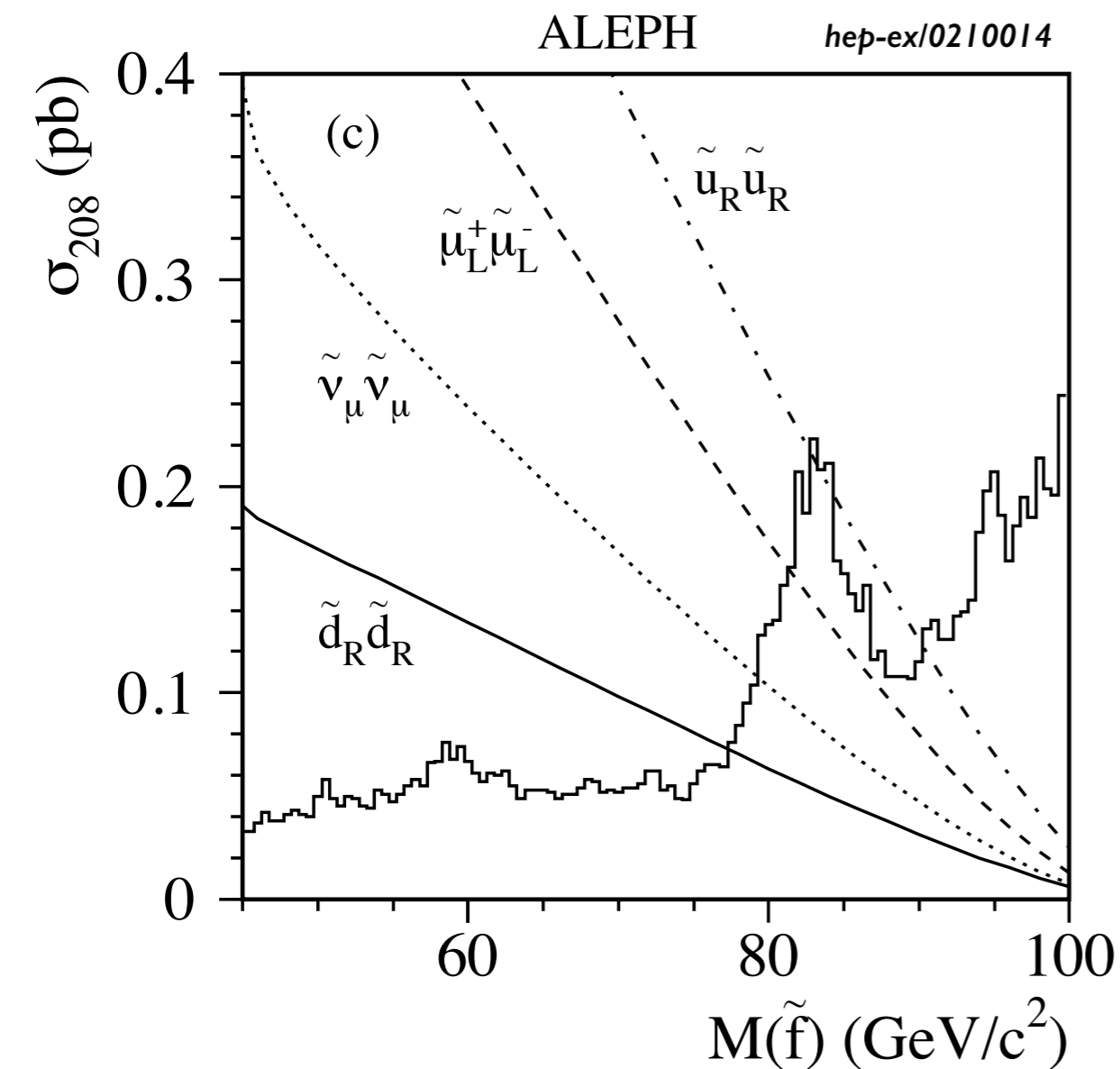
focus on

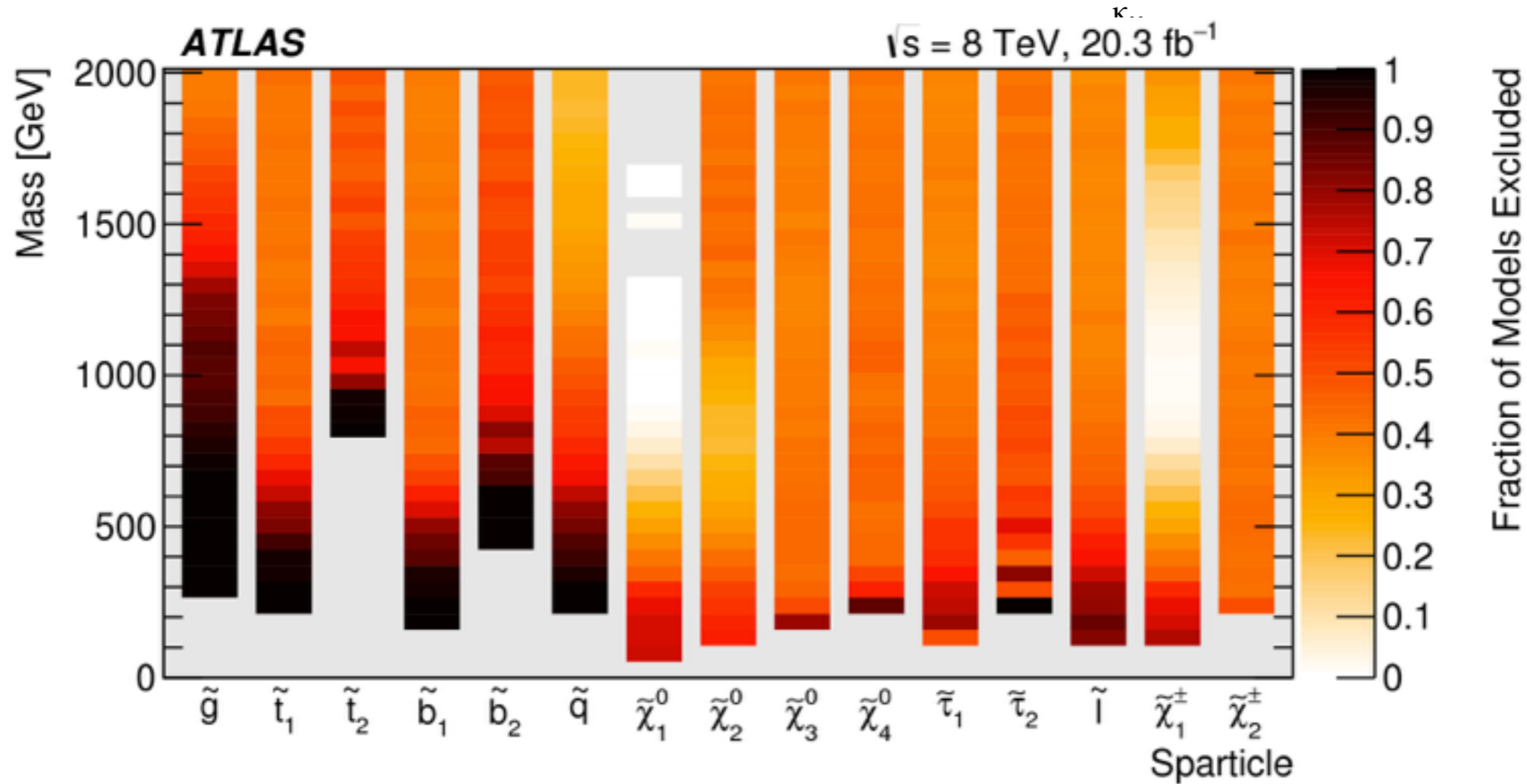
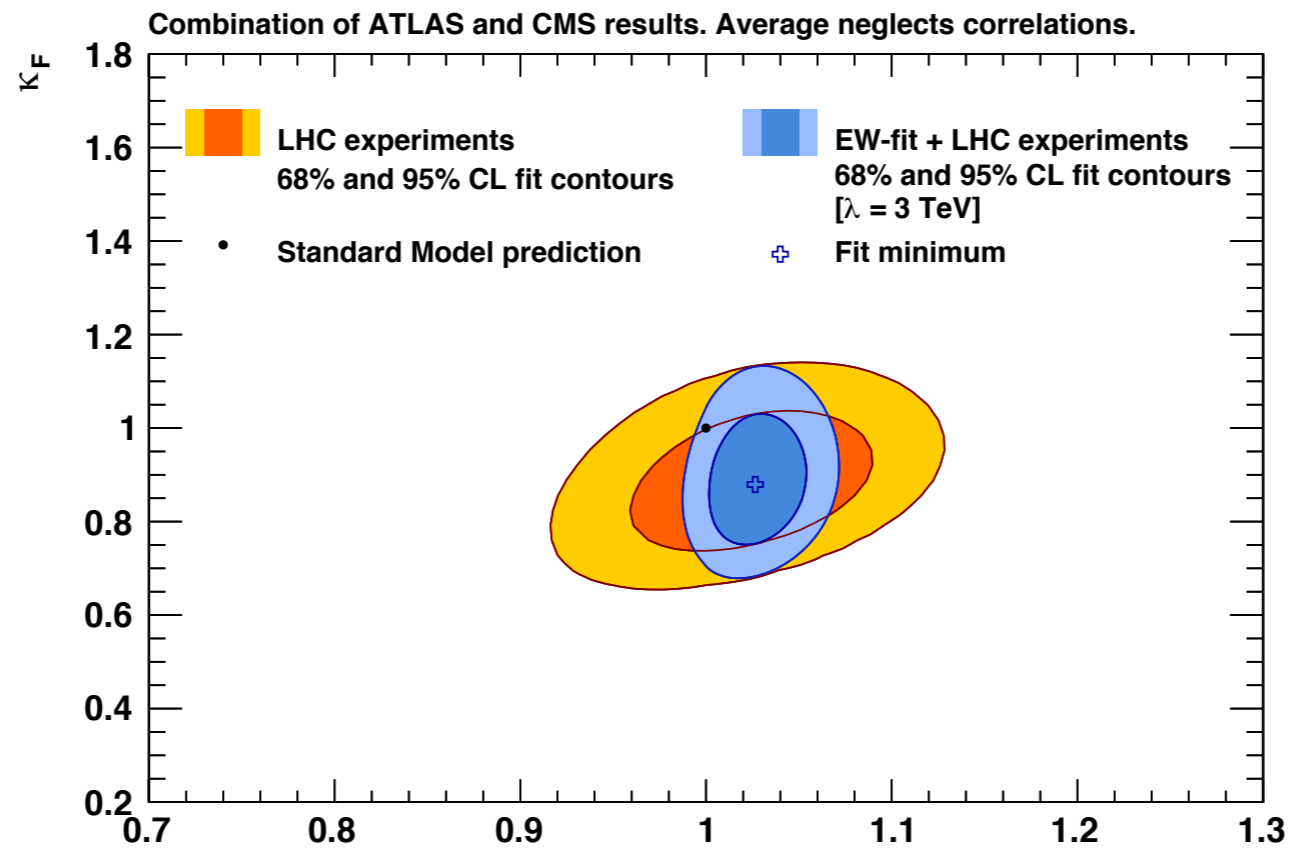
light ($M \lesssim v$) new physics, subtler signatures

Thank you!



$e^+e^- \rightarrow 4j$ limits





BSM Top FCNC overview

1311.2028, ATL-PHYS-PUB-2013-012, CMS-PAS-FTR-13-016

Process	SM	2HDM(FV)	2HDM(FC)	MSSM	RPV	RS
$t \rightarrow Zu$	7×10^{-17}	–	–	$\leq 10^{-7}$	$\leq 10^{-6}$	–
$t \rightarrow Zc$	1×10^{-14}	$\leq 10^{-6}$	$\leq 10^{-10}$	$\leq 10^{-7}$	$\leq 10^{-6}$	$\leq 10^{-5}$
$t \rightarrow gu$	4×10^{-14}	–	–	$\leq 10^{-7}$	$\leq 10^{-6}$	–
$t \rightarrow gc$	5×10^{-12}	$\leq 10^{-4}$	$\leq 10^{-8}$	$\leq 10^{-7}$	$\leq 10^{-6}$	$\leq 10^{-10}$
$t \rightarrow \gamma u$	4×10^{-16}	–	–	$\leq 10^{-8}$	$\leq 10^{-9}$	–
$t \rightarrow \gamma c$	5×10^{-14}	$\leq 10^{-7}$	$\leq 10^{-9}$	$\leq 10^{-8}$	$\leq 10^{-9}$	$\leq 10^{-9}$
$t \rightarrow hu$	2×10^{-17}	6×10^{-6}	–	$\leq 10^{-5}$	$\leq 10^{-9}$	–
$t \rightarrow hc$	3×10^{-15}	2×10^{-3}	$\leq 10^{-5}$	$\leq 10^{-5}$	$\leq 10^{-9}$	$\leq 10^{-4}$

$t \rightarrow cH \rightarrow \gamma\gamma$

