

Nonabelian Kinetic - Mixing -

At the energy and intensity frontiers

arxiv : 1511.02865

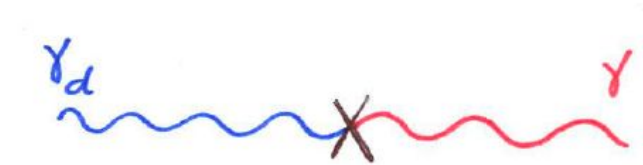
Gabriel Barello

**work with Spencer Chang and Christopher
Newby**



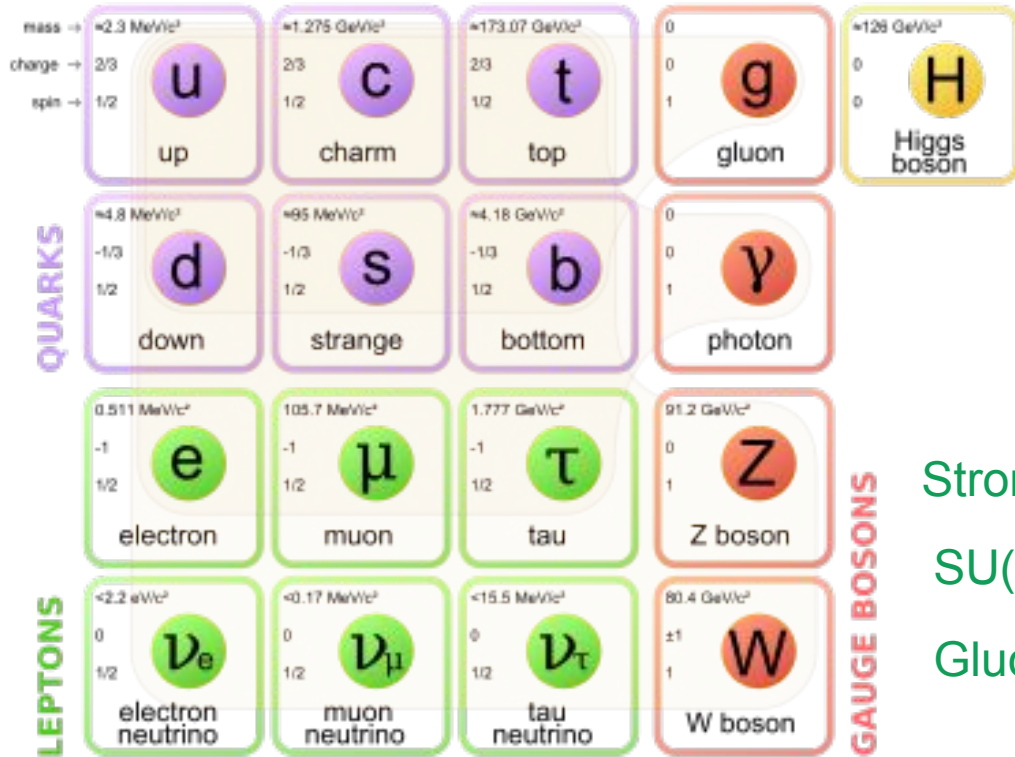
What this talk is about:

Kinetic mixing is a phenomenon in which a new gauge boson (the 'dark photon') can change back and forth into the regular photon



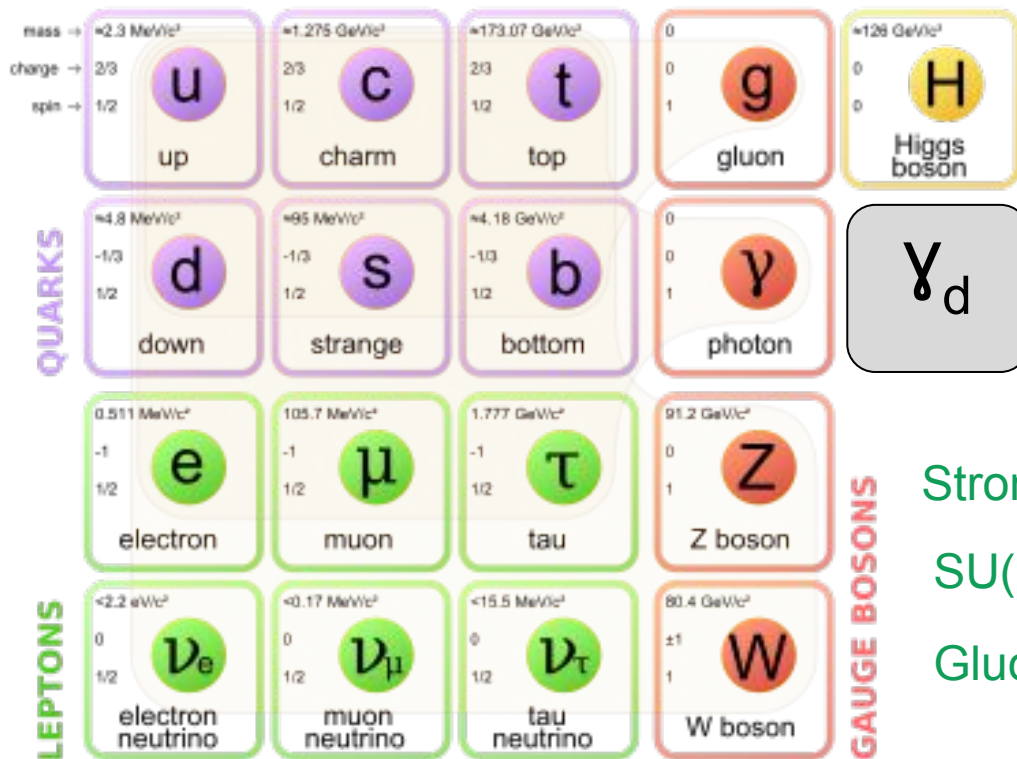
This talk is about nonabelian kinetic mixing

Dark Photons as a Simple Extension:



Strong $SU(3)$ \otimes $SU(2)_L$ \otimes $U(1)$ Y
 Gluons \otimes Electroweak (W/Z/Photon)

Dark Photons as a Simple Extension:



Strong $SU(2)_L$ Y Dark Photon

$SU(3) \otimes SU(2) \otimes U(1) \otimes U(1)$

Gluons Electroweak (W/Z/Photon)

The simplest extension:

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + i\bar{\psi}(\not{D} - m)\psi$$

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

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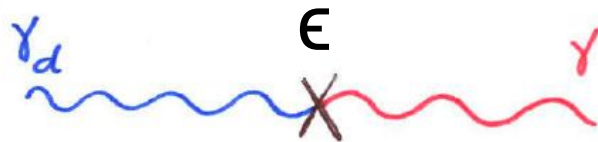
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$$-\frac{1}{4}\tilde{F}_{\mu\nu}\tilde{F}^{\mu\nu} + \frac{\epsilon}{2}\tilde{F}_{\mu\nu}F^{\mu\nu}$$

B. Holdom (1986)

$$F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

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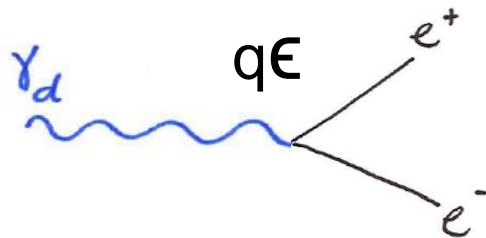
$$-\frac{1}{4}\tilde{F}_{\mu\nu}\tilde{F}^{\mu\nu} + \frac{\epsilon}{2}\tilde{F}_{\mu\nu}F^{\mu\nu} + i\epsilon q\tilde{A}\bar{\psi}\psi$$

field redefinition

B. Holdom (1986)

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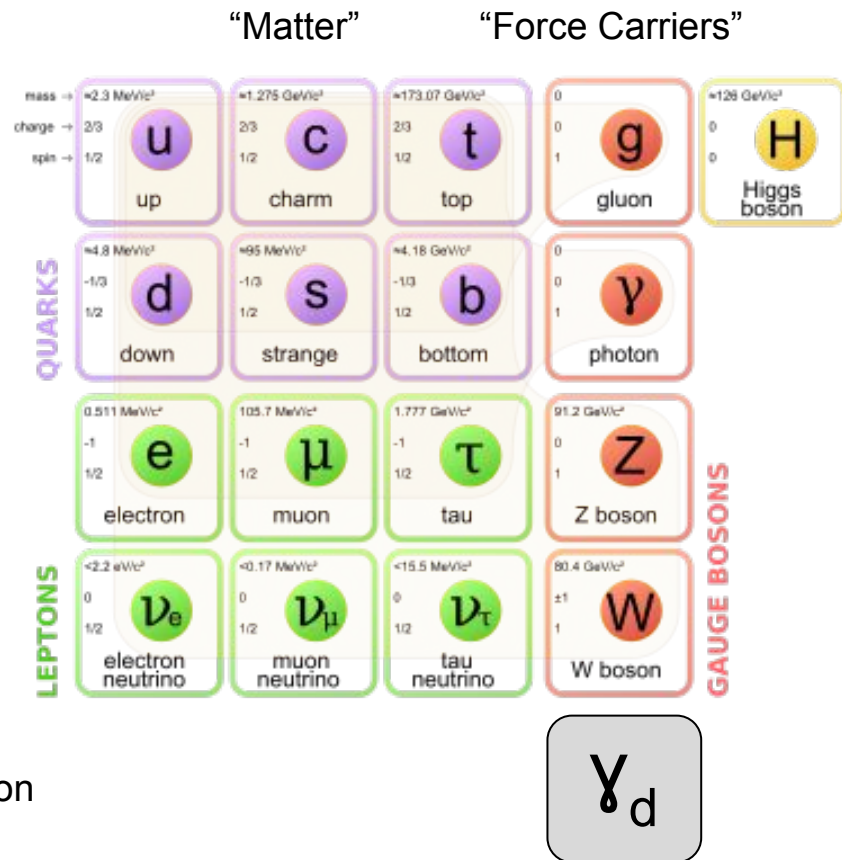


Nonabelian Kinetic Mixing

-“Regular” Kinetic Mixing-

The Dark U(1) and hypercharge mix

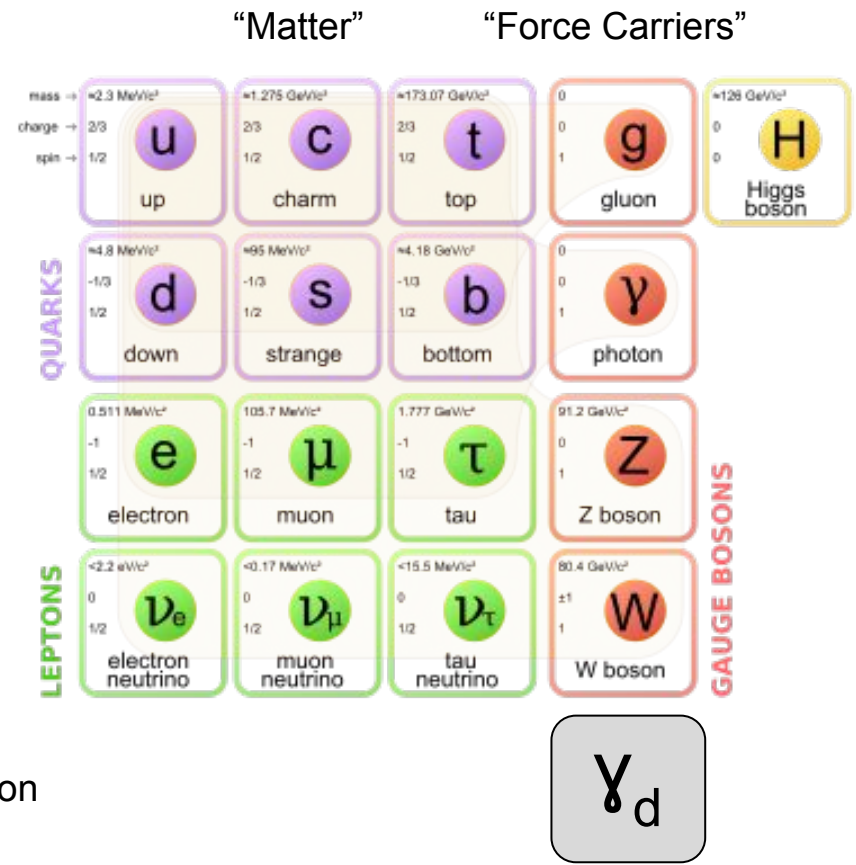
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 $SU(3) \otimes SU(2) \otimes U(1) \otimes U(1)$



Nonabelian Kinetic Mixing

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Strong SU(2)_L Y Dark Photon
 SU(3) ⊗ SU(2) ⊗ U(1) ⊗ U(1)



Originally Arkani-Hamed & Weiner proposed the other way around, dark nonabelian group mixing with SM - arxiv: 0810.0714

PBS NOVA, Fermilab, Office of Science, United States Department of Energy, Particle Data Group

Nonabelian Kinetic Mixing - How can it happen?

$$\mathcal{L}_{\text{NAKM}} \supset \frac{\epsilon}{2} \tilde{F}_{\mu\nu} W_a^{\mu\nu}$$

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Epsilon is a field that transforms under $SU(2)_L$

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The Higgs!

Introduces a new scale
Lambda

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EWSB generates KM

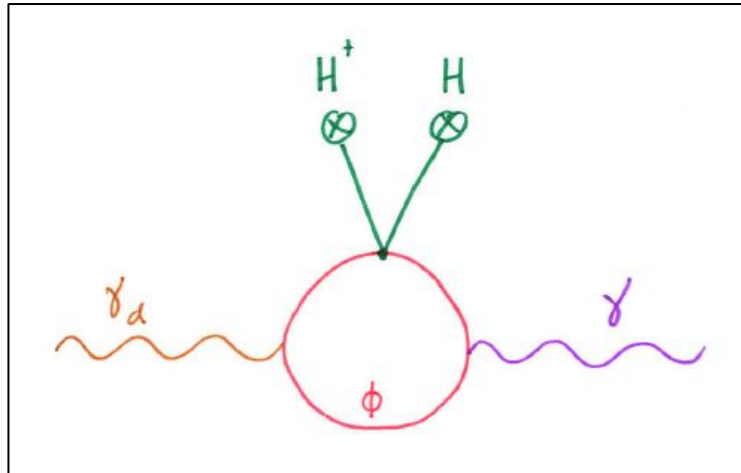
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$$\mathcal{L}_{\text{NAKM}} \supset \frac{v^2}{\Lambda^2} \tilde{F}_{\mu\nu} W_3^{\mu\nu}$$

The origin of weak scales:

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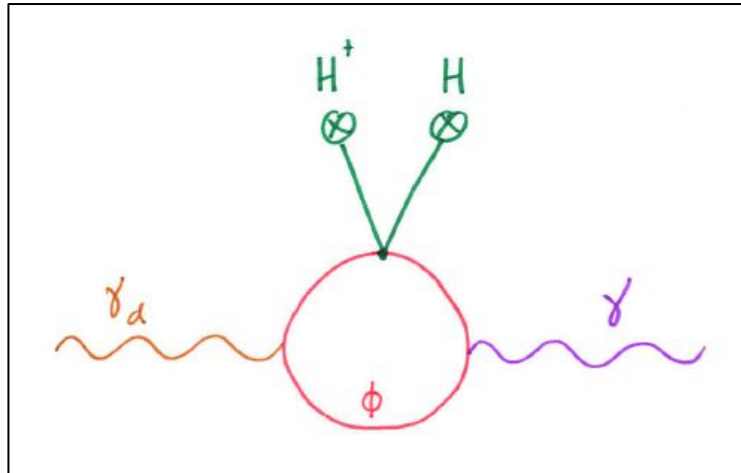


$$\Lambda^2 \sim 16\pi^2 m_\phi^2$$

The origin of weak scales:

$$\mathcal{L}_{\text{NAKM}} \supset \frac{H_i^\dagger \tau_{ij}^a H_j}{\Lambda^2} \tilde{F}_{\mu\nu} W_a^{\mu\nu} \longrightarrow \frac{\epsilon}{2} \tilde{F}_{\mu\nu} F^{\mu\nu}$$

The Photon Field Strength



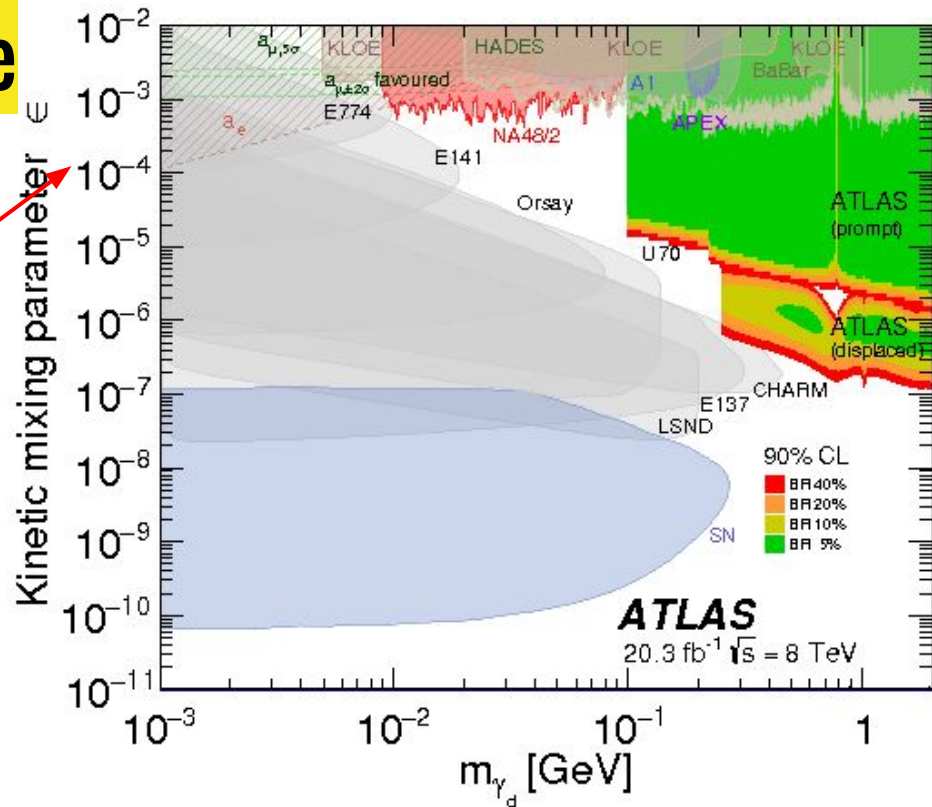
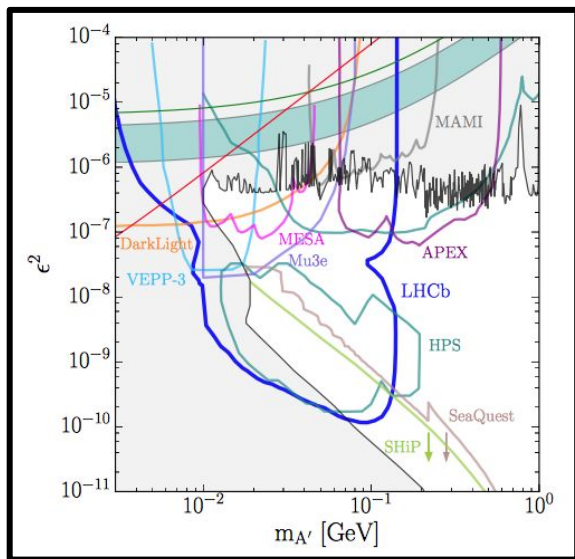
$$\epsilon = \frac{cv^2 s_W}{2\Lambda^2} \quad \Lambda^2 \sim 16\pi^2 m_\phi^2$$

$$m_\phi = \sqrt{\frac{cv^2 s_W}{32\pi^2 \epsilon}} \sim 1\text{TeV} \times \sqrt{\frac{c}{\epsilon/10^{-4}}}$$

A compelling coincidence

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Ilten, Thaler, Williams, Xue: arxiv 1509.06765

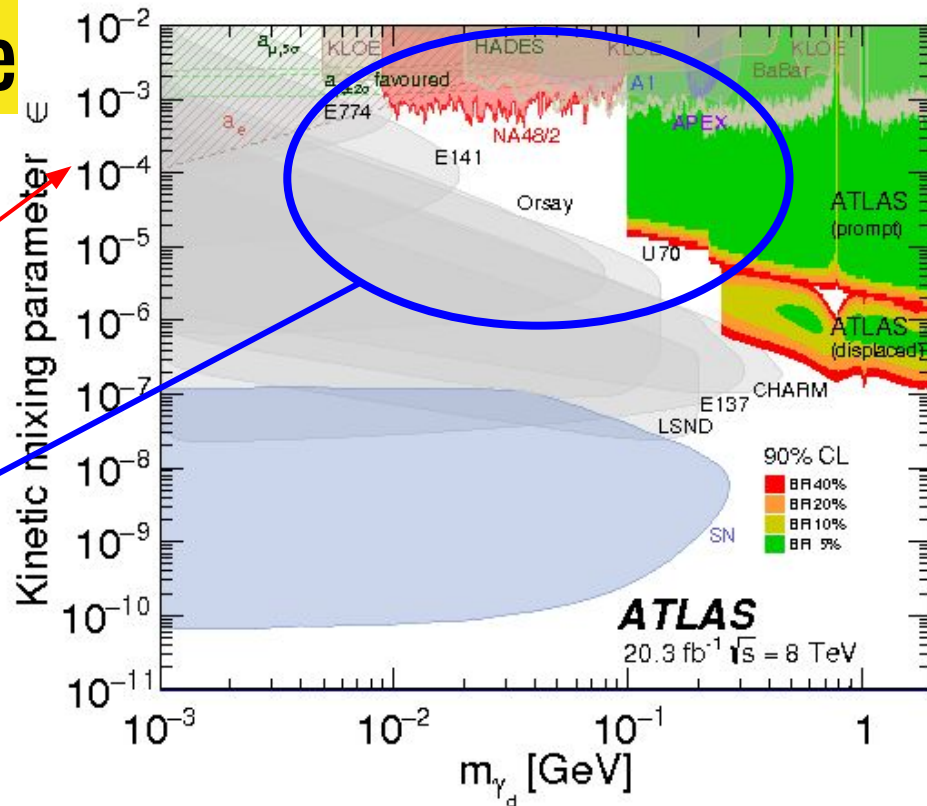
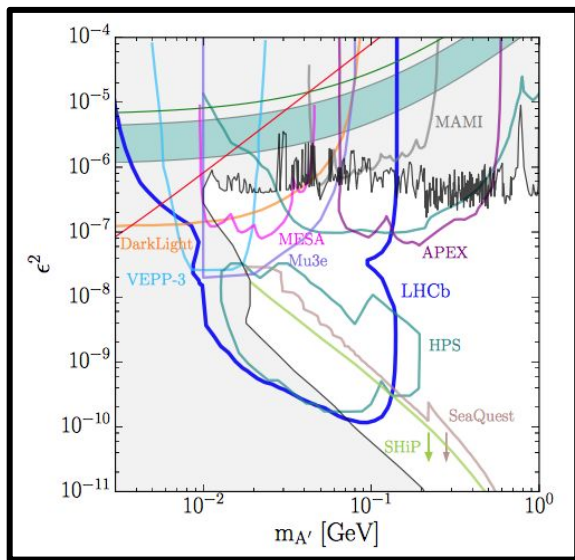


ATLAS Collaboration: arXiv: 1511.05542

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The Connection

“If we discover a dark photon in the next few years, nonabelian kinetic mixing predicts that we’ll produce the mediating particle at the LHC!”

Next:

- Build a model
- Design search strategies
- Explore other related scenarios

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Our Model

- A dark photon γ_d : the force carrier associated with the new force
- The “dark higgs” h_d : Gives the dark photon a mass
- The mediator ϕ : Generates Nonabelian Kinetic Mixing

New Particle Quantum Numbers:					
Name	$SU(2)_L$	$U(1)_d$	$U(1)_Y$	Spin	Mass
Dark Photon (γ_d)	0	0	0	1	0.2 GeV
Dark Higgs (h_d)	0	1	0	0	0.4 GeV
Mediator (ϕ)	3	1	0	0	TBD

NOTE: No particle charged under $U(1)_Y$ means no regular KM!

The Mediator

EWSB splits the mediator into 3 sets of particles:

χ^\pm

Mediator Mass
~ 100s of GeV

ϕ_R^0 / ϕ_I^0

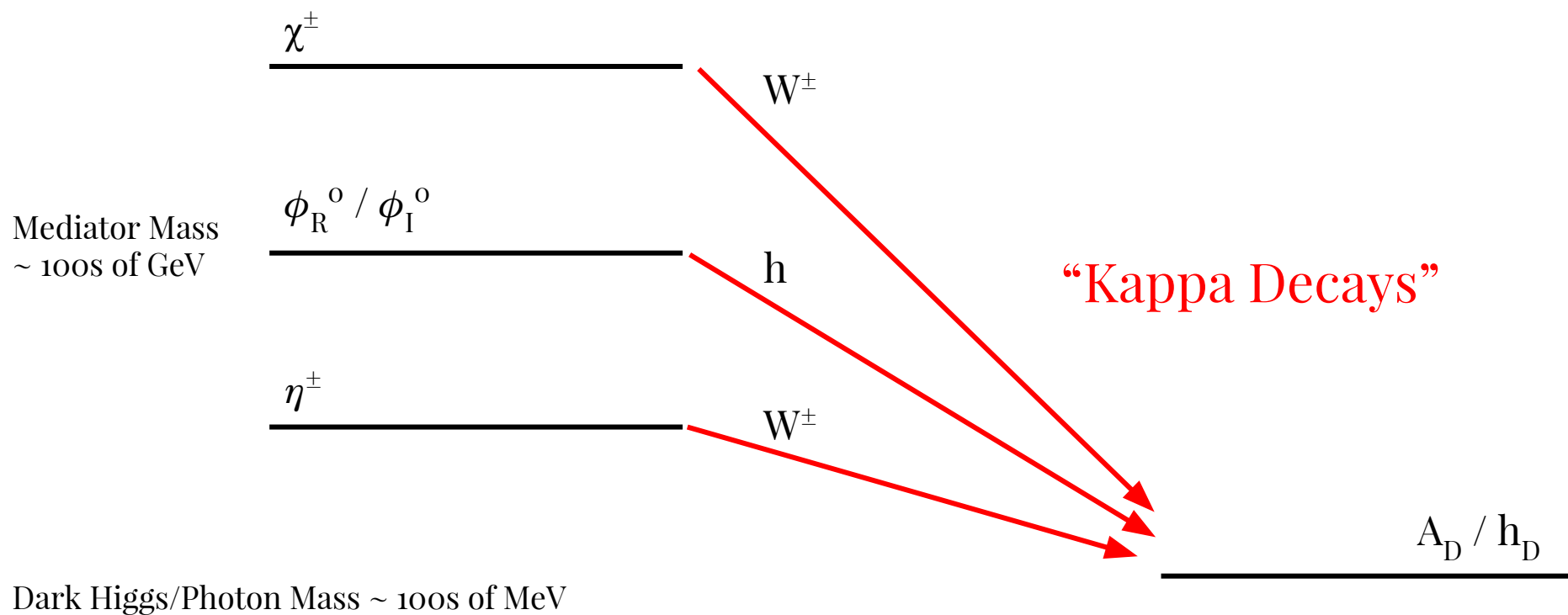
η^\pm

Dark Higgs/Photon Mass ~ 100s of MeV

A_D / h_D

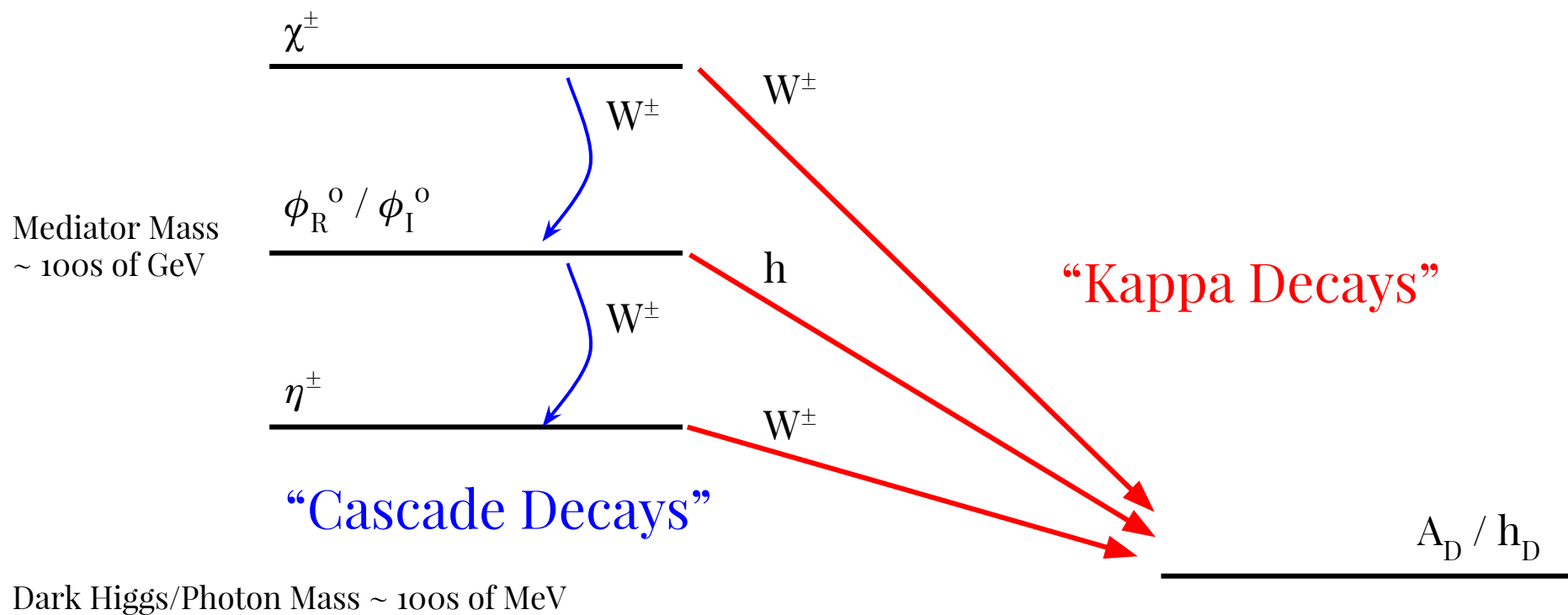
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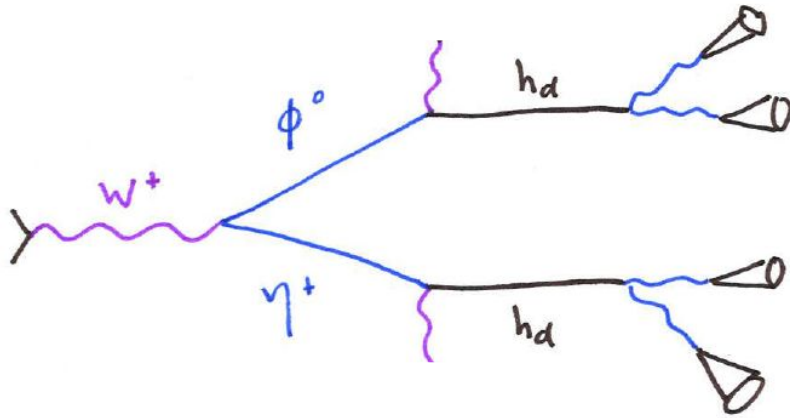


The Mediator

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Search Scenarios:

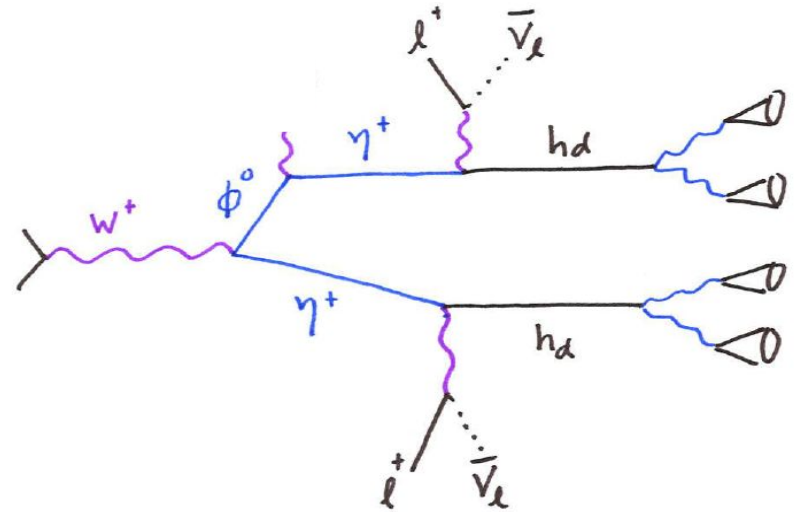


Direct Decay

- Kappa decays dominate

Cascade Decays

- Cascade decays dominate



In the Future:

- We'll finish up this collider study - keep an eye out!
- There is a large space of potential models to explore
 - Change the dark photon mass, larger mass means much different constraints
 - $SU(2)$ of the dark sector mixing with $U(1)$ of the standard model
 - Maybe the mediator could be dark matter?

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Questions?

The lagrangian

$\mathcal{L} = \text{Kinetic Terms} + \text{Mass Terms}$

$$+ \lambda (\phi_a^\dagger T_c^{ab} \phi_b) (H_i^\dagger \tau_{ij}^c H_j) \leftarrow \text{Generates KM}$$

$$+ \kappa \phi^a (H^\dagger \tau^a H) h_d + h.c. \leftarrow \text{Allows the mediator to decay}$$

$$+ \text{Other scalar coupling terms} \leftarrow \text{Unimportant for this talk}$$