

# FORBIDDEN WIMPS (AREN'T)

Nirmal Raj




with  
A. Delgado, A. Martin  
160y.xxxxx

# Relic density @ Pheno 2016

$$\Omega_\chi h^2 \sim 0.1$$

- ★ ELDERs : Y-D Tsai
- ★ Co-decays : J. Dror
- ★ Shared asymmetry : L. Necib
- ★ Old-fashioned WIMP freezeout : N. R.

# Thermal cross-section

$$\alpha^2 / M^2$$


WIMP miracle

$$\langle \sigma v \rangle_{\text{ann}} = 3 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$$

# Thermal cross-section

$$\alpha^2 / M^2$$

$$\chi_i \chi_j \rightarrow \chi_k \text{SM}$$

WIMP miracle      WIMPlless miracle      “semi-annihilation”

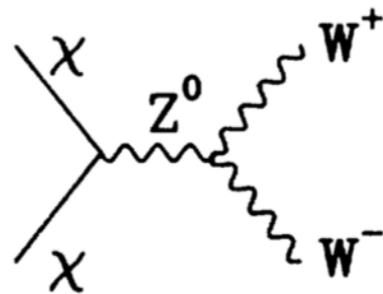
Feng, Kumar: 0803.4196

D' Eramo, Thaler: 1003.5912

$$\langle \sigma v \rangle_{\text{ann}} = 3 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$$



co-annihilation



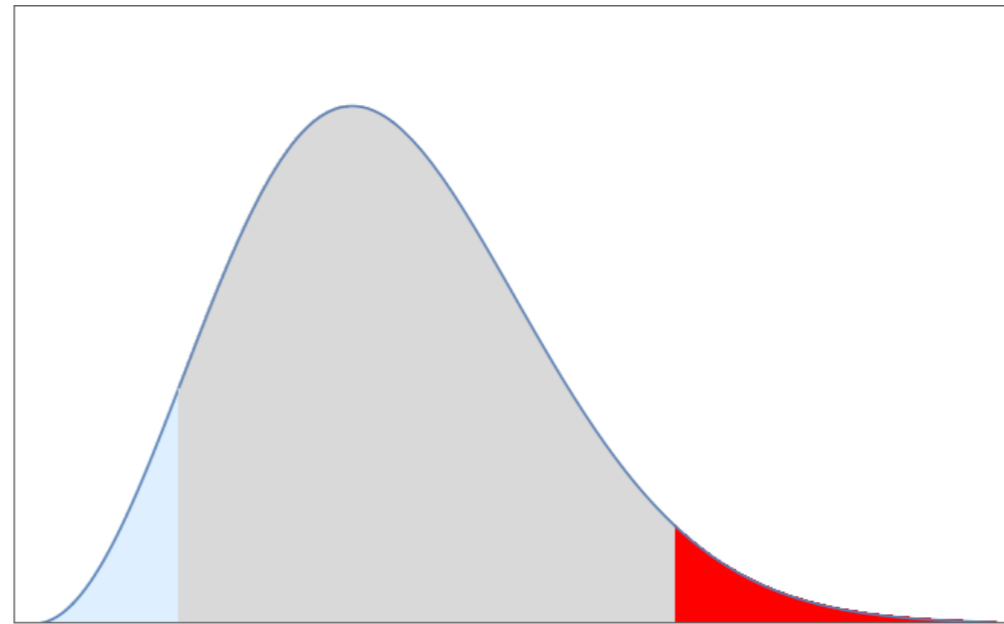
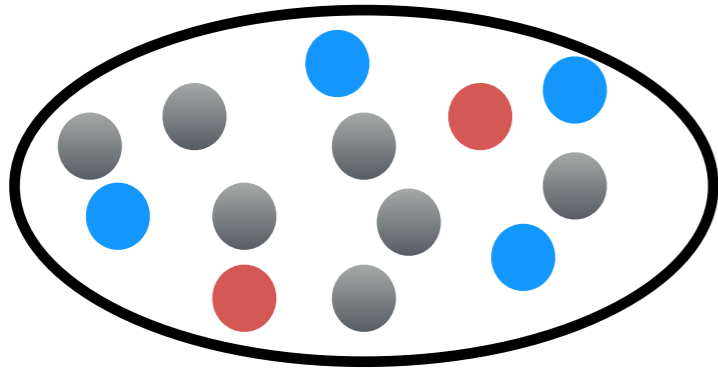
resonant

?

forbidden

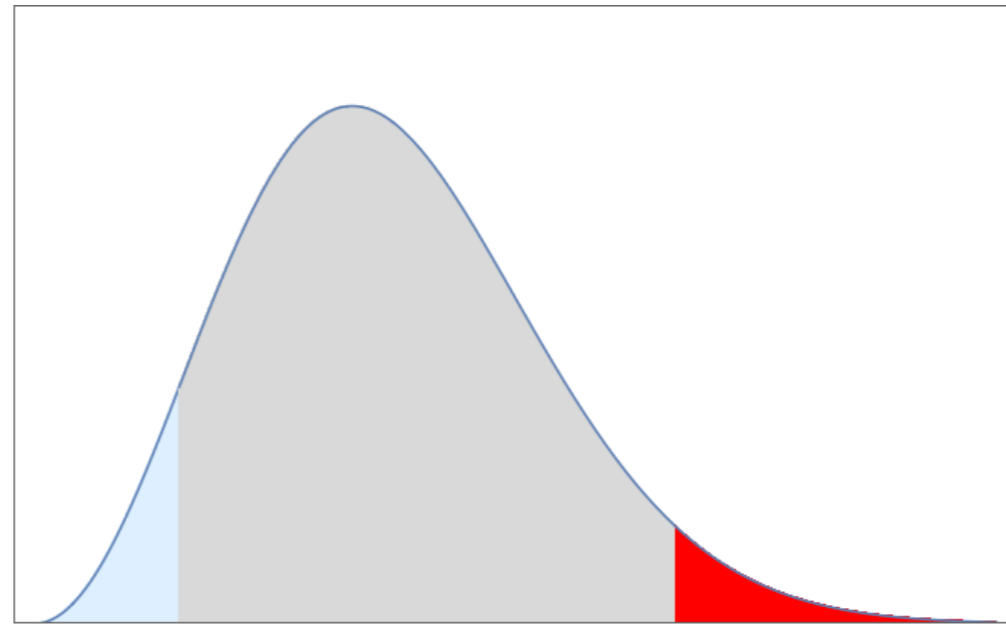
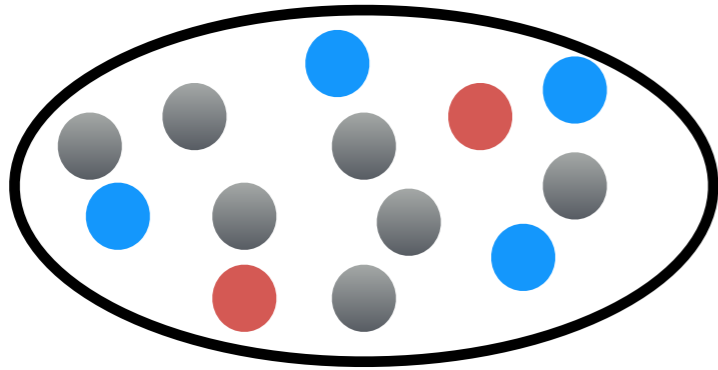
Griest & Seckel: “Three exceptions in the calculation of relic abundances” Phys.Rev. D43 (1991)

# DM velocities



$v$

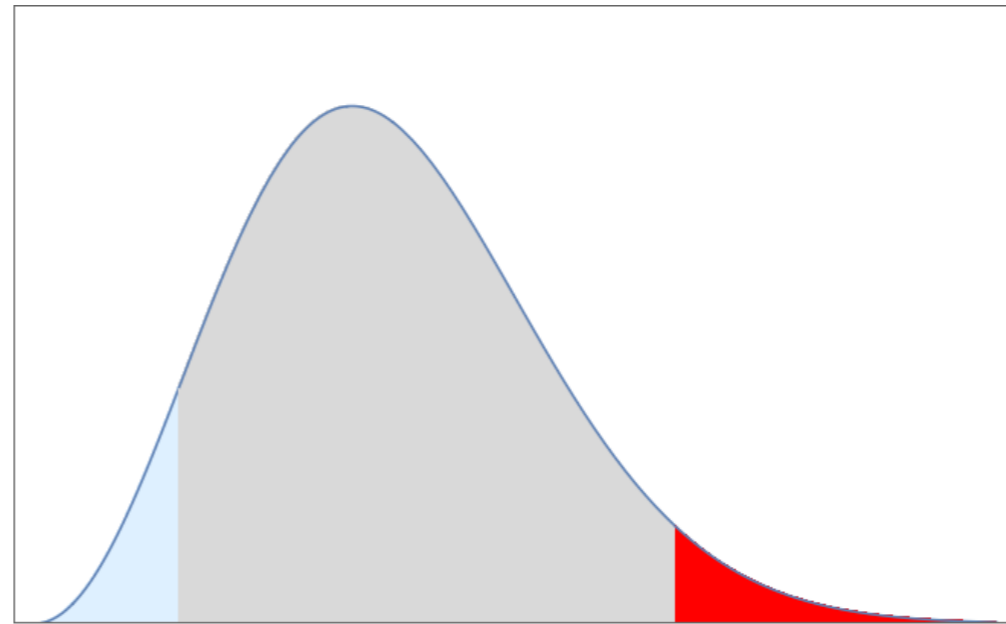
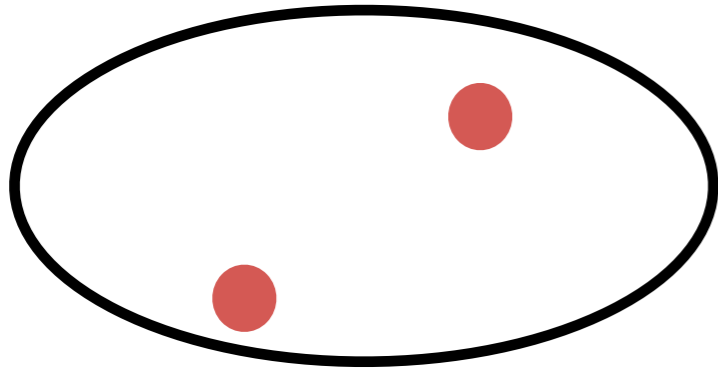
# DM annihilation



$v$

$$\left. \chi\bar{\chi} \rightarrow f\bar{f} \right\} m_\chi > m_f, \langle\sigma v\rangle : \quad \bullet \quad \bullet \quad \bullet$$

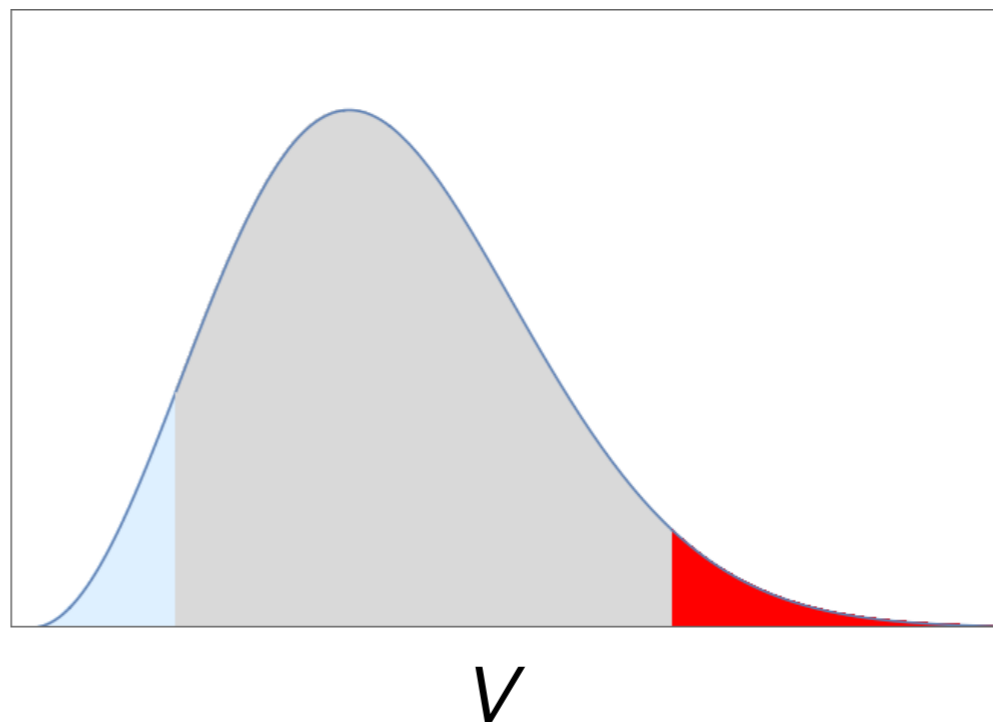
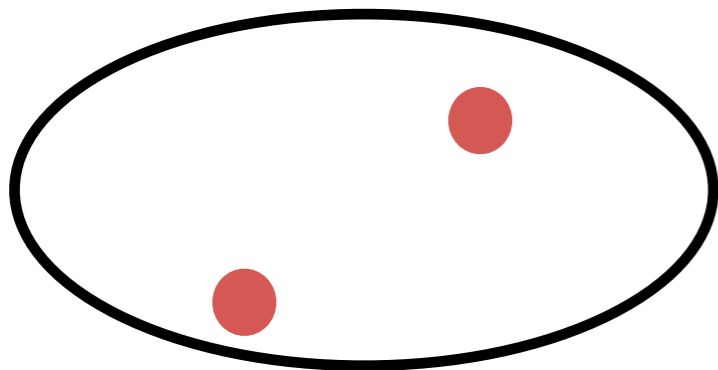
# DM annihilation



v

$$\left. \begin{array}{l} \chi\bar{\chi} \rightarrow f\bar{f} \\ \chi\bar{\chi} \rightarrow f\bar{f} \end{array} \right\} \begin{array}{l} m_\chi > m_f, \langle\sigma v\rangle : \text{blue} \quad \text{gray} \quad \text{red} \\ m_\chi < m_f, \langle\sigma v\rangle : \quad \quad \quad \text{red} \end{array}$$

# DM annihilation



$$\chi\bar{\chi} \rightarrow f\bar{f} \left. \begin{array}{l} m_\chi > m_f, \langle\sigma v\rangle : \text{blue} \quad \text{grey} \quad \text{red} \\ \underline{\underline{m_\chi < m_f}}, \langle\sigma v\rangle : \text{red} \end{array} \right\}$$

$\langle\sigma v\rangle = 0$  at zero temperature, hence “forbidden” DM.

Key ingredient: large  $\sigma \sim |\mathcal{M}|^2$  to begin with, to get  $\langle\sigma v\rangle_{\text{ann}} = 3 \times 10^{-26} \text{ cm}^3 \text{ s}^{-1}$



# Forbidden DM history

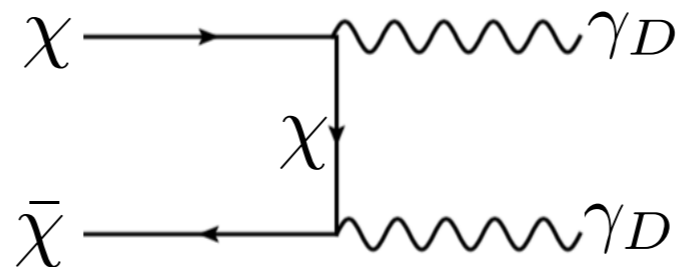
**This morning's plenary talk by Mariangela Lisanti**

R. D'Agnolo, J. Ruderman: 1505.07107

Large  $\sigma \sim |\mathcal{M}|^2$

— light (MeV- GeV) mediators

Relic density from forbidden annihilation



# Forbidden DM history

(a) C. Jackson, et al: 0912.0004, 1302.1802, 1303.471

(b) S. Tulin, H. Yu, K. Zurek: 1208.009

(c) R. D'Agnolo, J. Ruderman: 1505.07107

Large  $\sigma \sim |\mathcal{M}|^2$  — s-channel pole: (a), (b)  
— light (MeV- GeV) mediators: (c)

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Large  $\sigma \sim |\mathcal{M}|^2$  — s-channel pole: (a), (b)  
— light (MeV- GeV) mediators: (c)

Relic density from forbidden ann.: (c)

Exotica final states: (b), (c)

Light final states: (a)

U(1)' gauge: (a), (c)

Main target indirect detection: (a), (b), (c)

# Model wishlist vs Virtues

Weak scale masses

LHC, LUX can now probe all “exceptions”

No gauge group extension

Minimal assumptions (no gauge kinetic mixing, etc.)

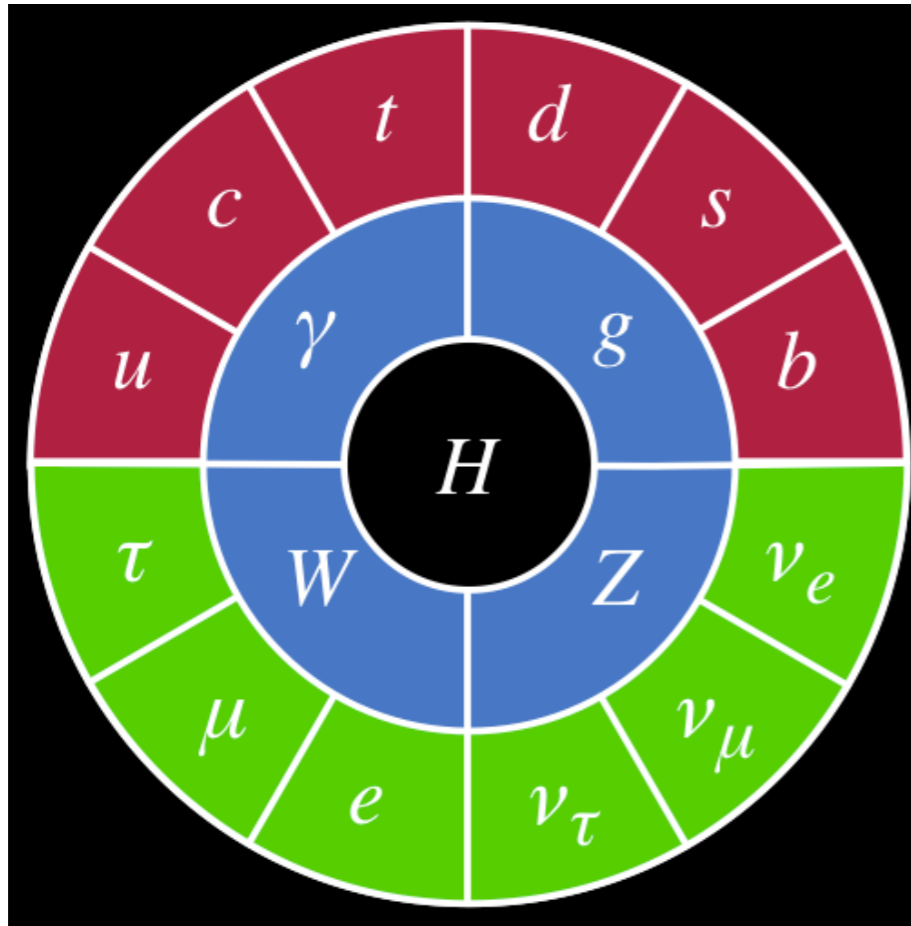
Annihilation to SM states

DM mass roughly known (~90% SM mass)

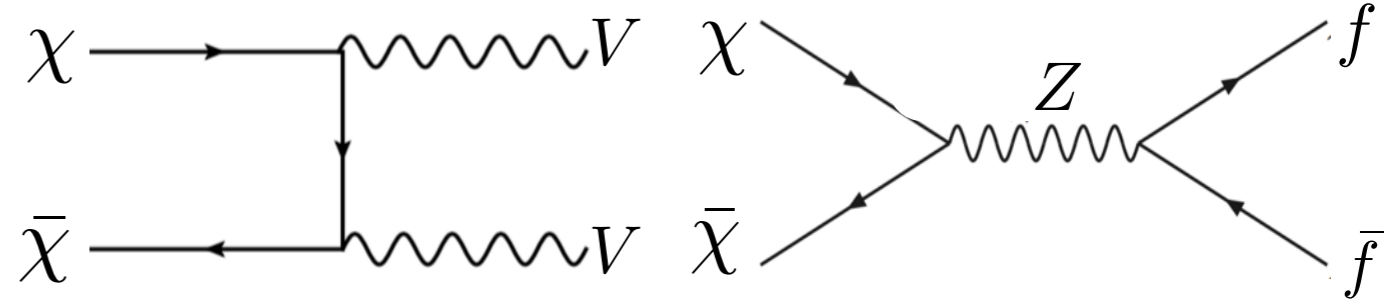
Enforce communication between dark & luminous sectors

MicroOmegas already computes forbidden rates near thresholds.  
But can a WIMP be entirely forbidden?

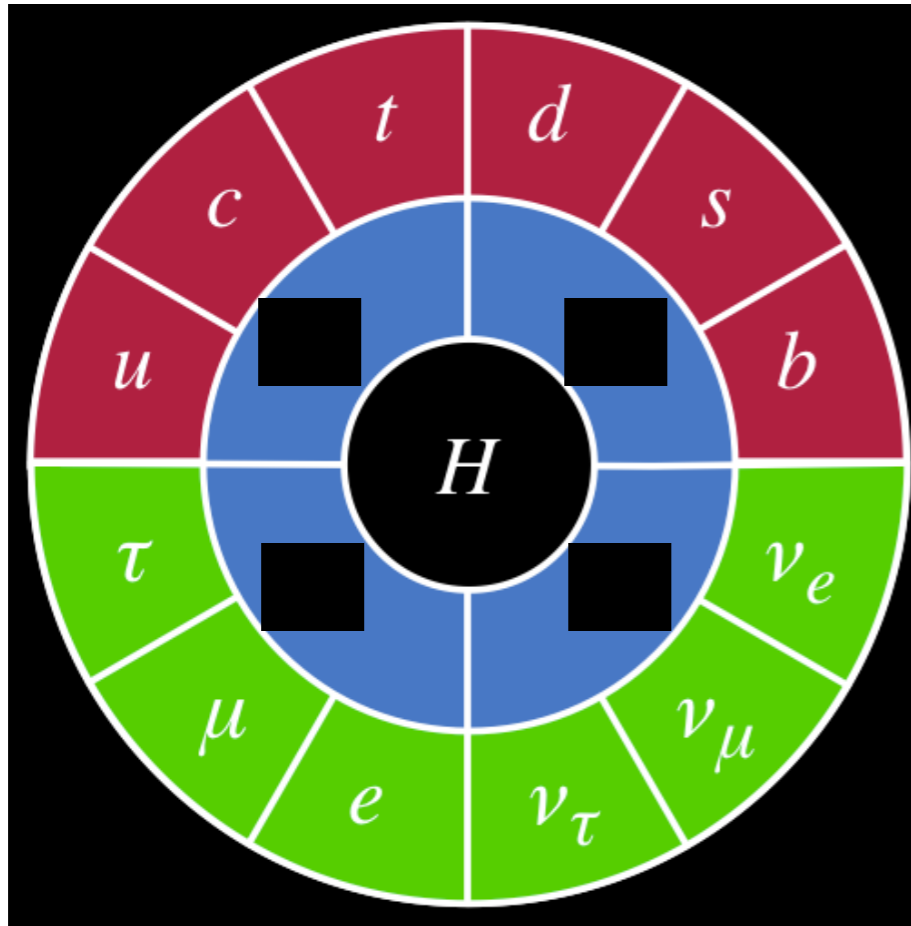
# “Truly” forbidden?



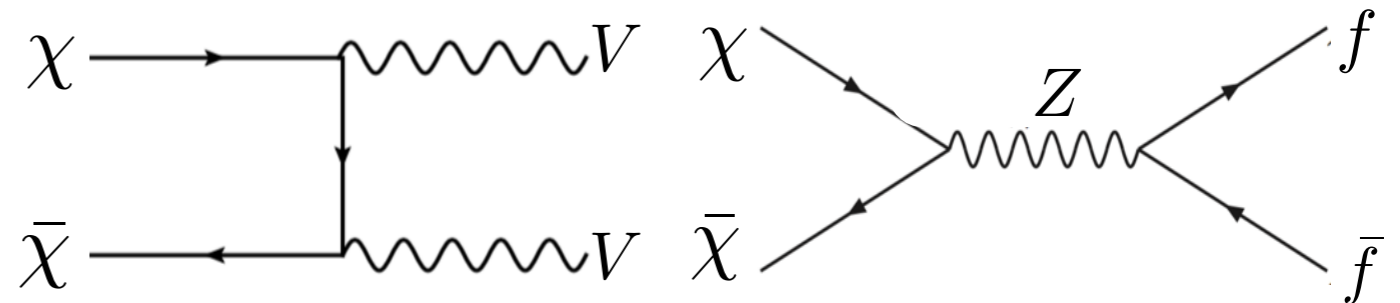
SU(2) charged



# “Truly” forbidden?

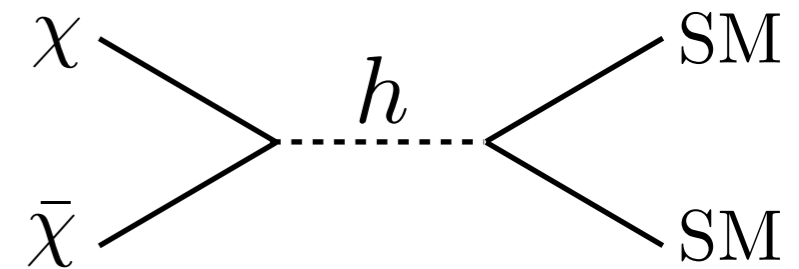


SU(2) charged

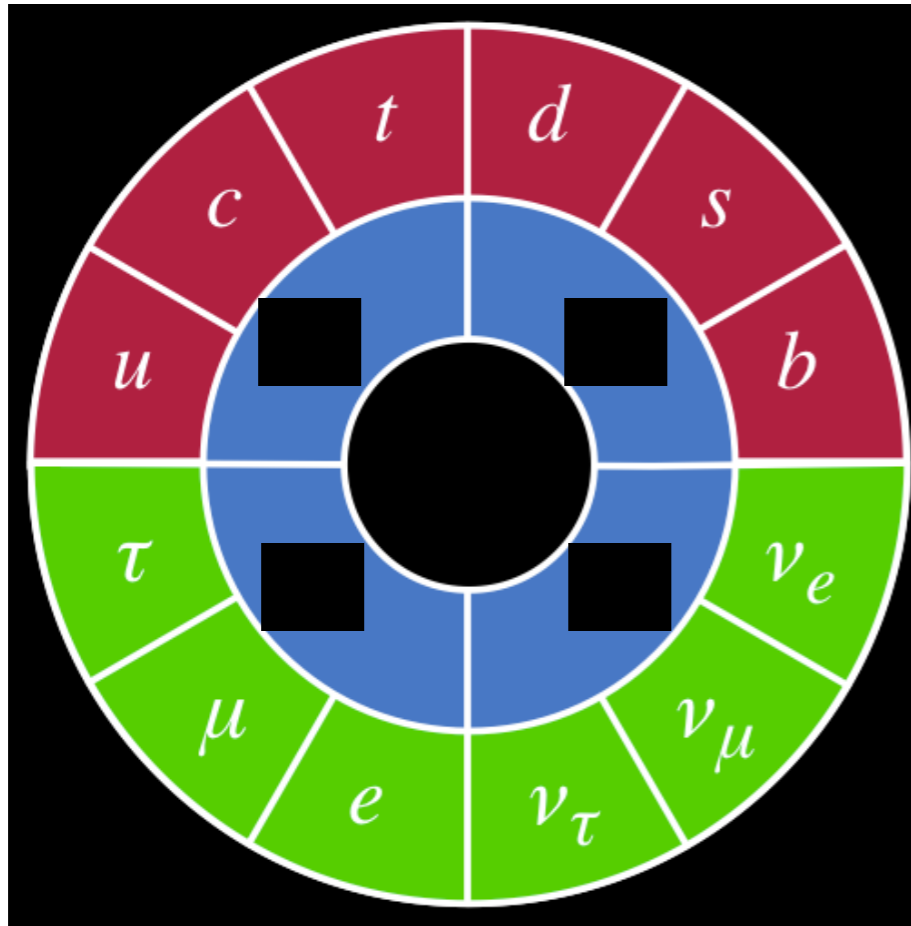


Higgs portal

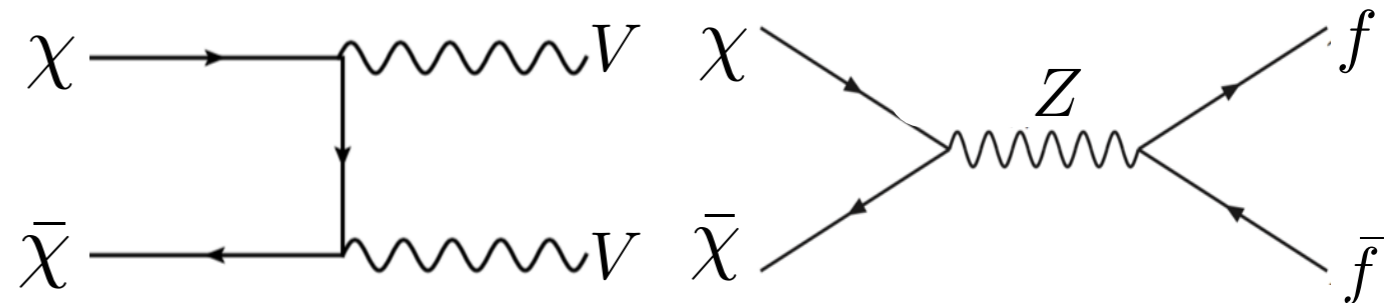
$$\chi\bar{\chi} \rightarrow hh$$



# “Truly” forbidden?

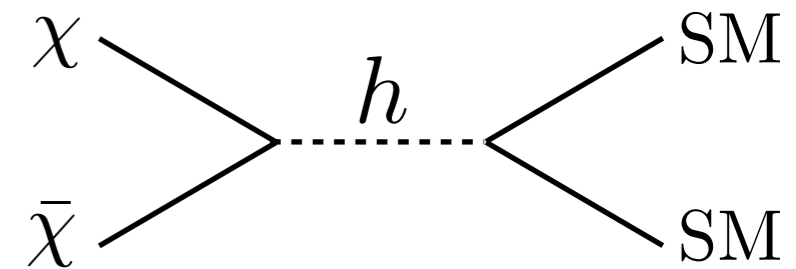


SU(2) charged



Higgs portal

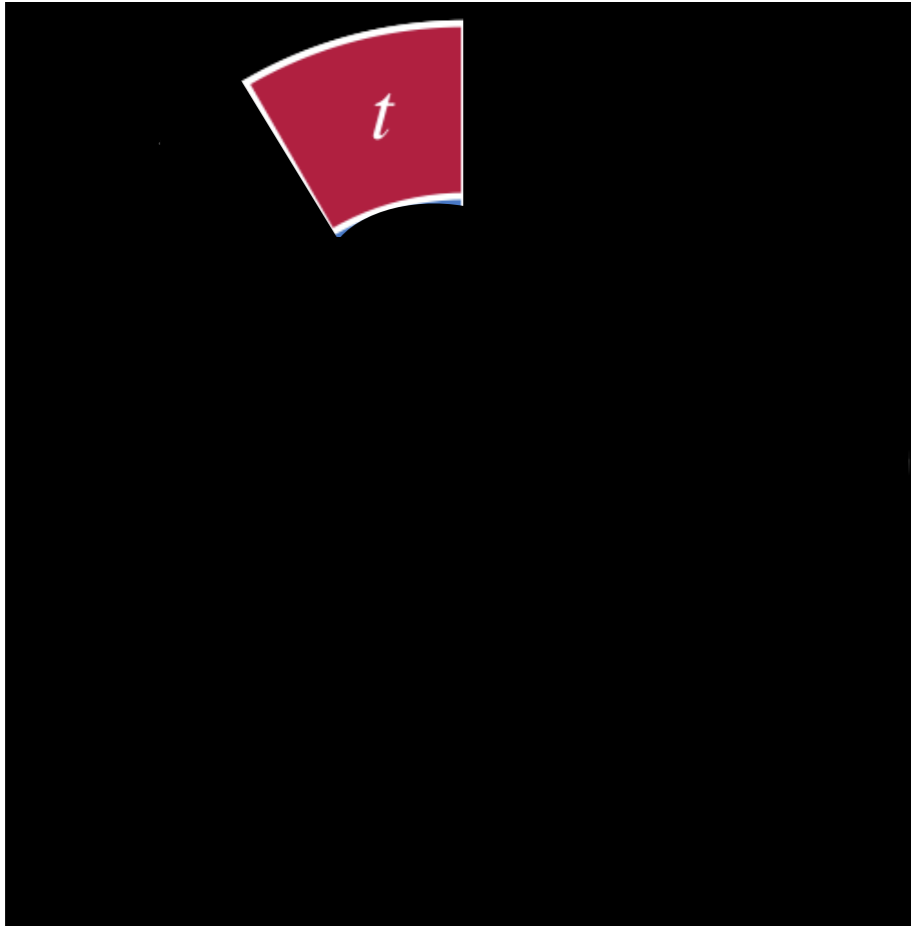
$$\chi\bar{\chi} \rightarrow hh$$



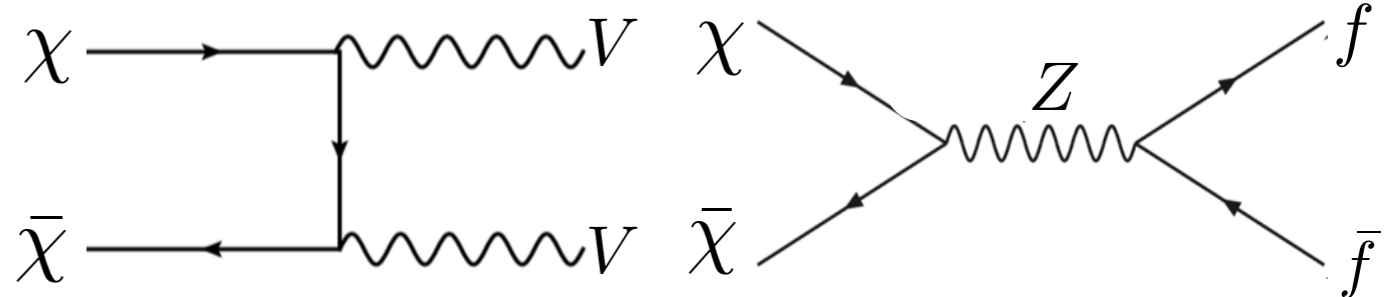
## Fermion portals

Fermion lighter than 100 GeV => mediator mass < O(10 GeV);  
ruled out by LEP or lack of resonances

# “Truly” forbidden?

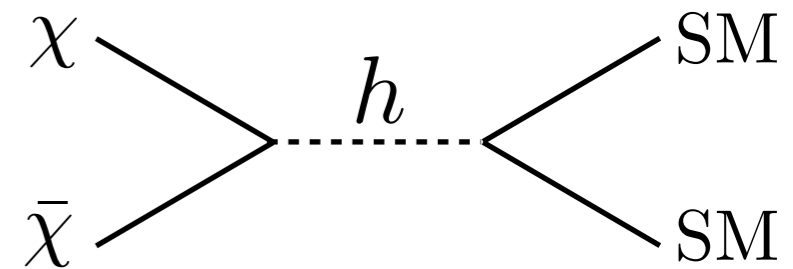


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Higgs portal

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## Fermion portals

Fermion lighter than 100 GeV  $\Rightarrow$  mediator mass  $< O(10 \text{ GeV})$ ;  
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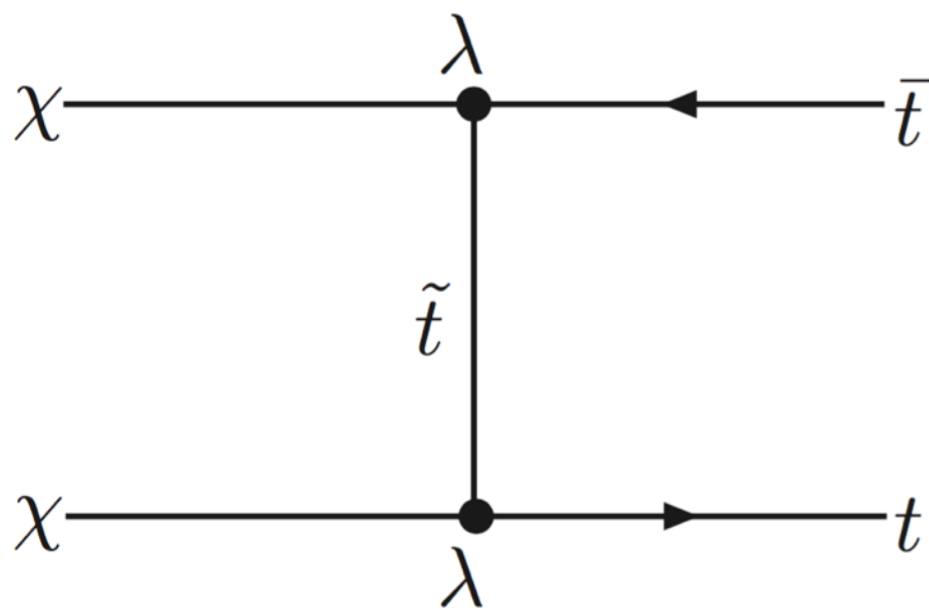


# Top Portal

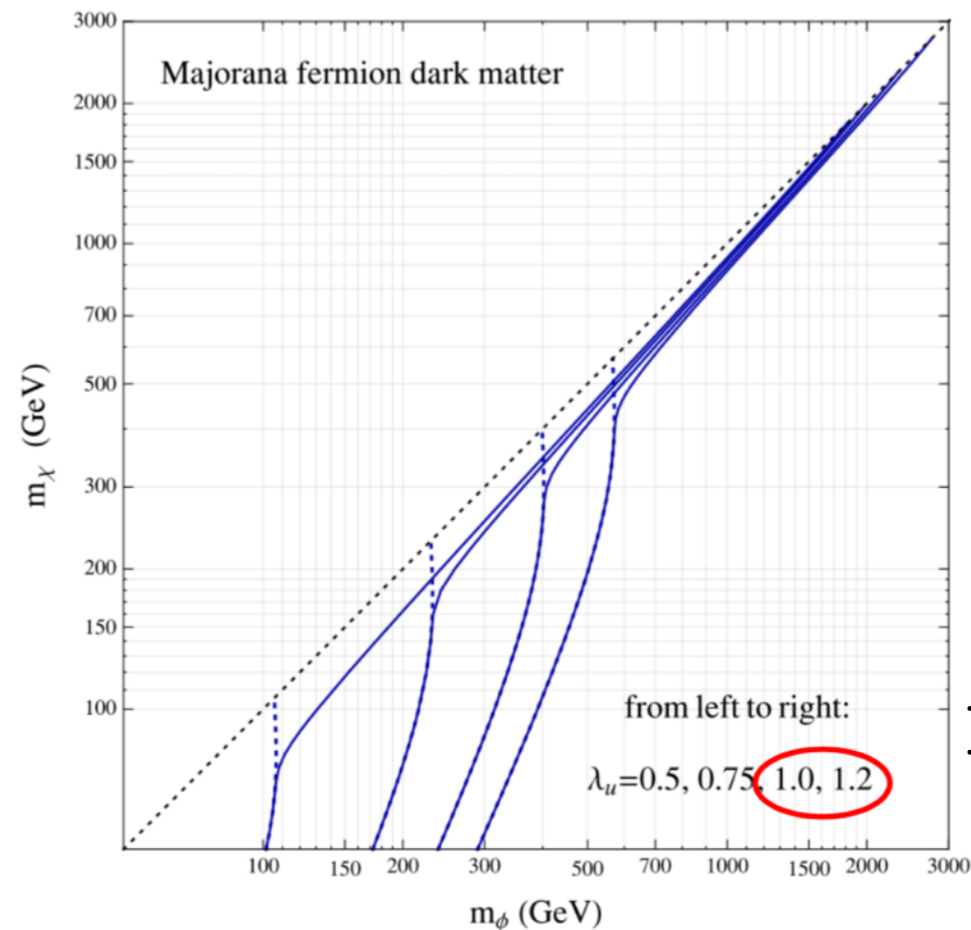
Accomplishing large  $\sigma \sim |\mathcal{M}|^2$

Model		Relic Abundance
$\chi$	Q	
Majorana fermion	Complex scalar	$a \sim m_a^2$ $\lambda \sim 0.5 - 2$
Dirac fermion	Complex scalar	$\lambda \sim 0.2 - 1$

Chang, Luty, et al; 1307.8120



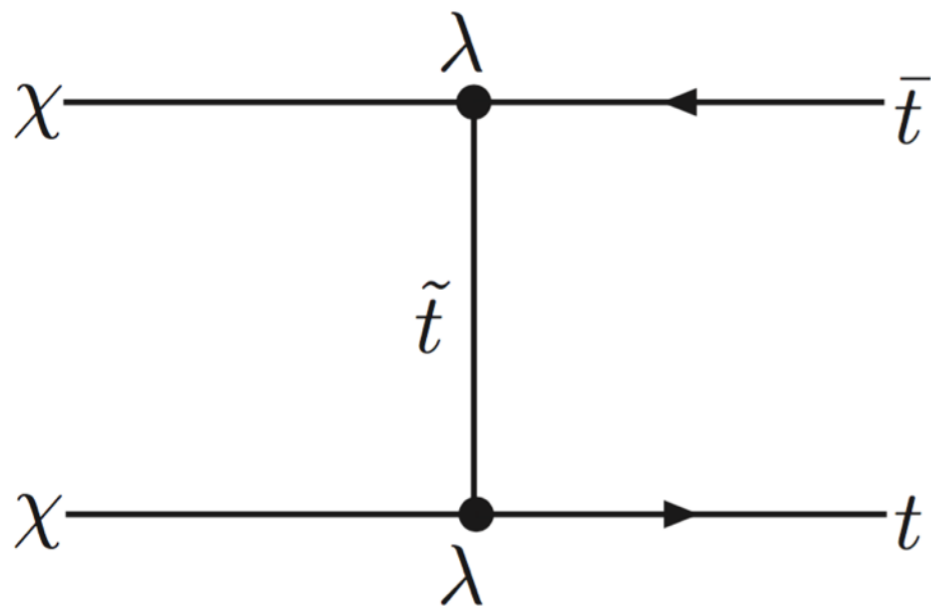
Effective coupling large



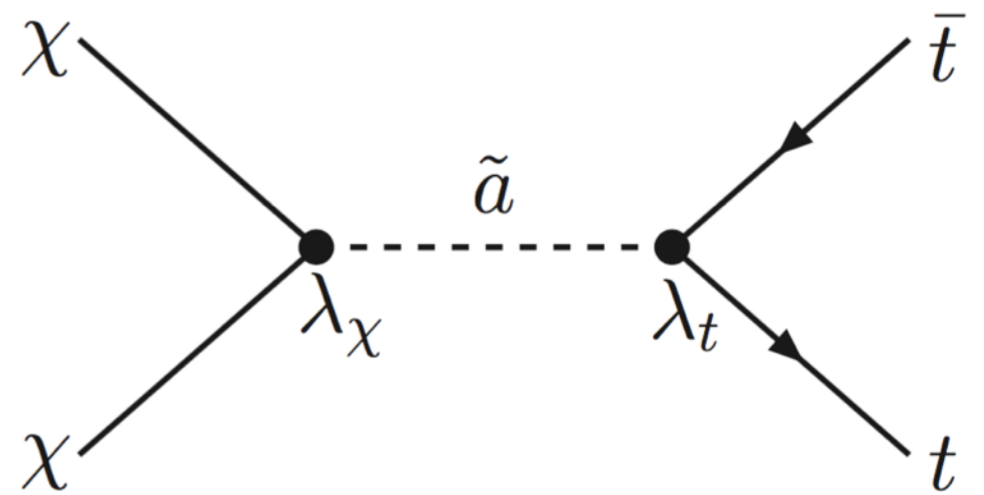
Bai, Berger  
1308.0612

# Top Portal

**Accomplishing large**  $\sigma \sim |\mathcal{M}|^2$



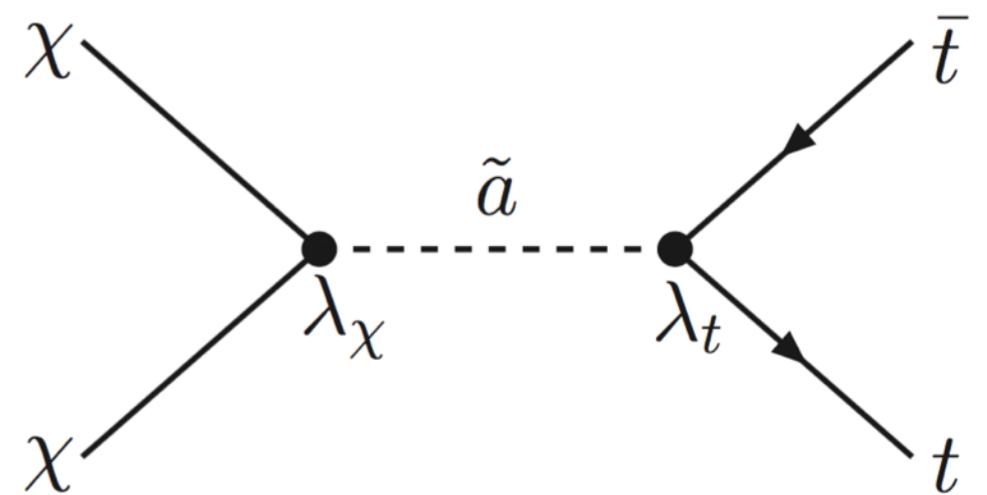
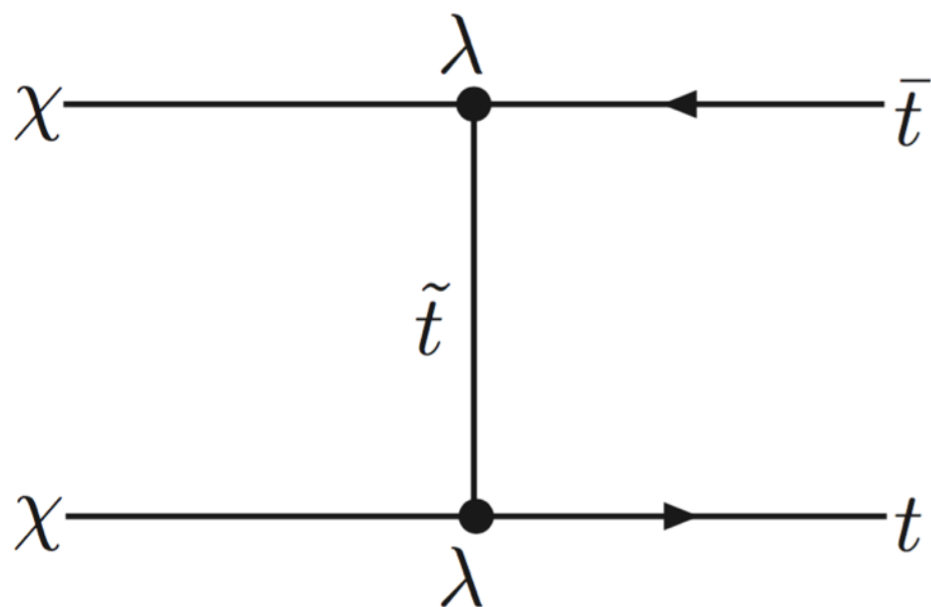
Effective coupling large



Annihilate near pole

# Top Portal Simplified Models

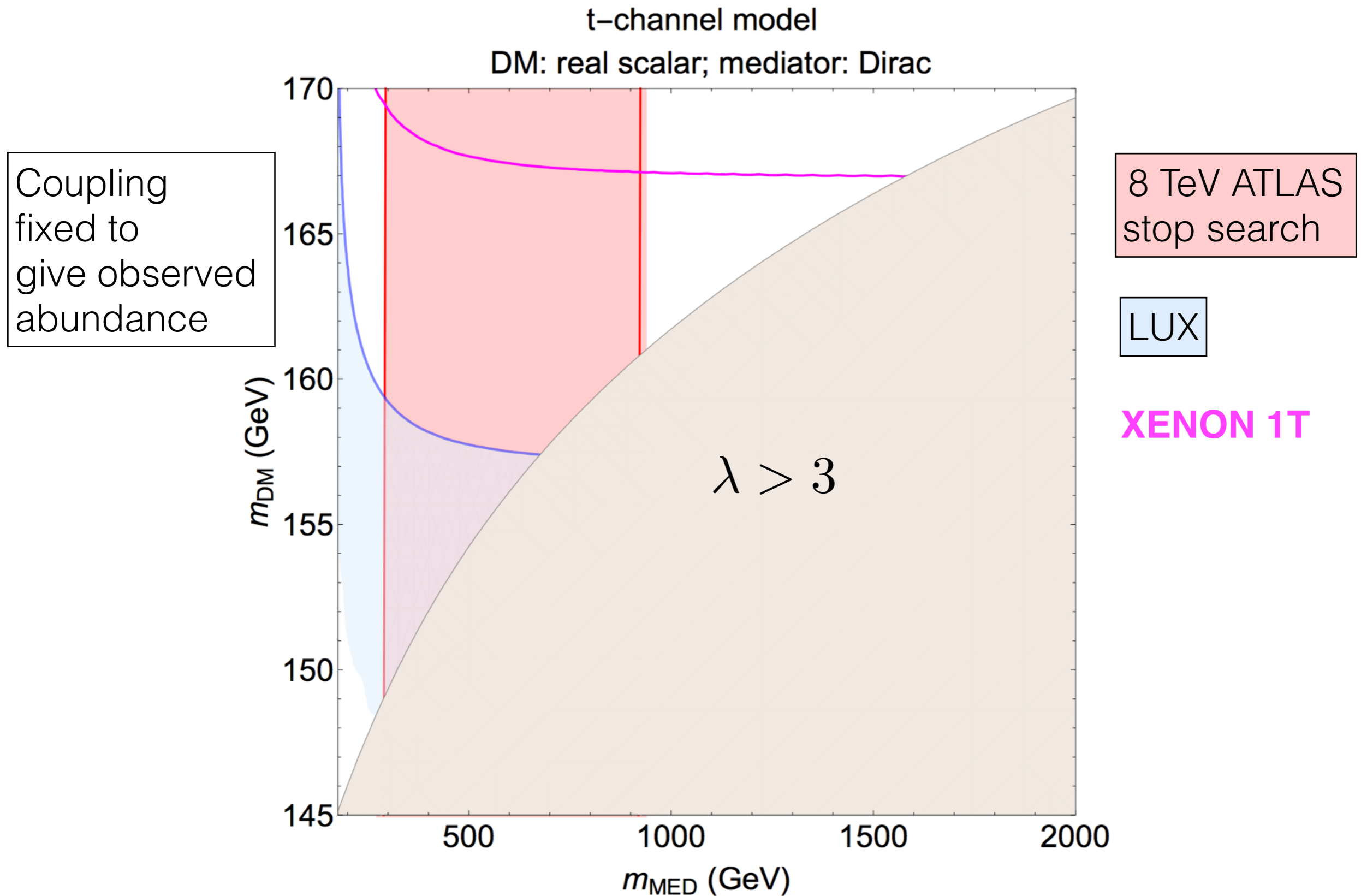
Field	$SU(3)_c \otimes SU(2)_W \otimes U(1)_Y$	$Z_2$	Spins		
			$t$ -channel	$s$ -channel	
$\chi$	$(1,1,0)$	-1	$1/2$	$0$	$1/2$
$\tilde{t}$	$(3,1,2/3)$	-1	$0$	$1/2$	
$\tilde{a}$	$(1,1,0)$	+1			$0$



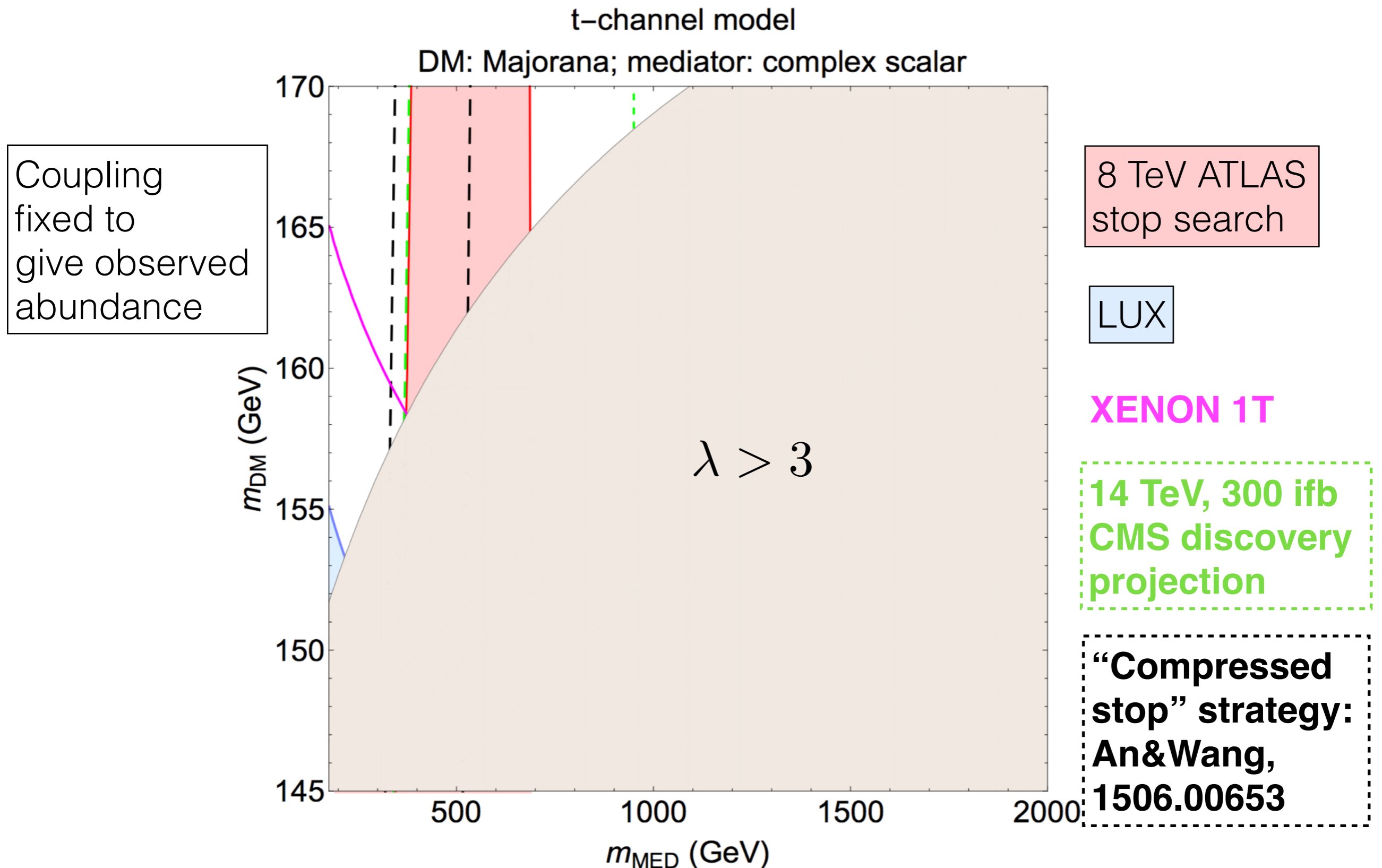
No qualitative difference if fermion DM Majorana/Dirac, scalar DM complex/real

If mediator CP-even, LUX kills model

# Limits & prospects: t-channel



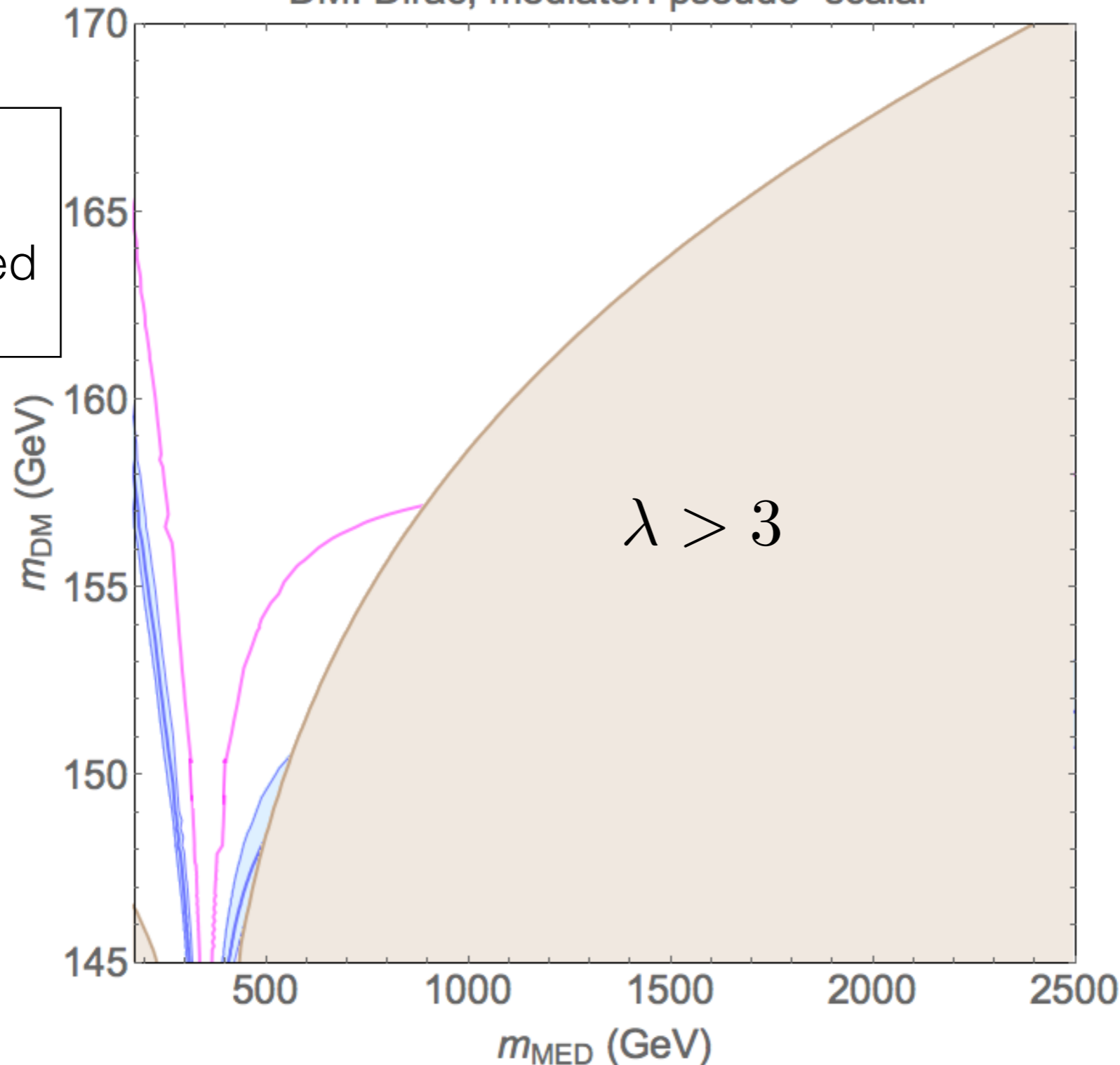
# Limits & prospects: t-channel



# Limits & prospects: s-channel

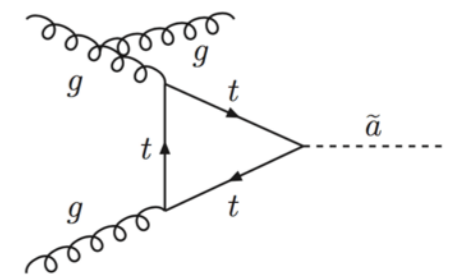
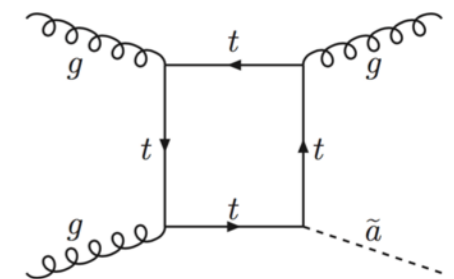
s-channel model

DM: Dirac; mediator: pseudo-scalar



**Monojet @  
8 TeV, 20 ifb**

**Monojet @  
14 TeV, 3 iab**



# Final remarks

Truly forbidden WIMP already possible;

Discovery prospects —

s-channel model: need to get lucky,

t-channel models: high coverage @ LHC and XENON1T

Forbidden WIMP vs “secluded WIMP” [ Pospelov, et al 0711.4866]

good at direct detection & colliders, bad at indirect detection

bad at “ ”, built for “ ”

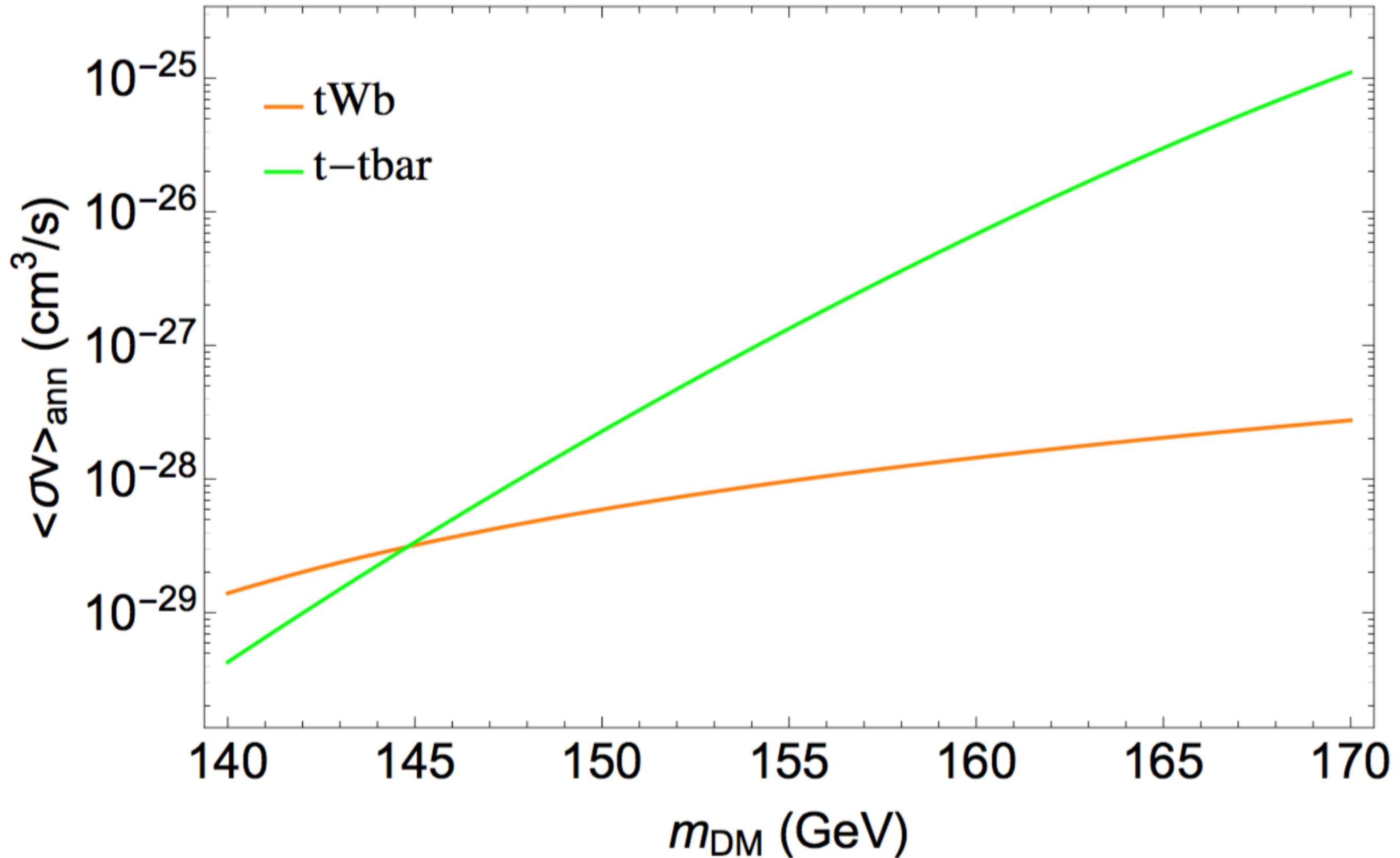
## *Future directions*

- Partially forbidden: multiple SM final states, more pheno
- Mediator final states
- Co-annihilation effects
- Incorporation into SUSY

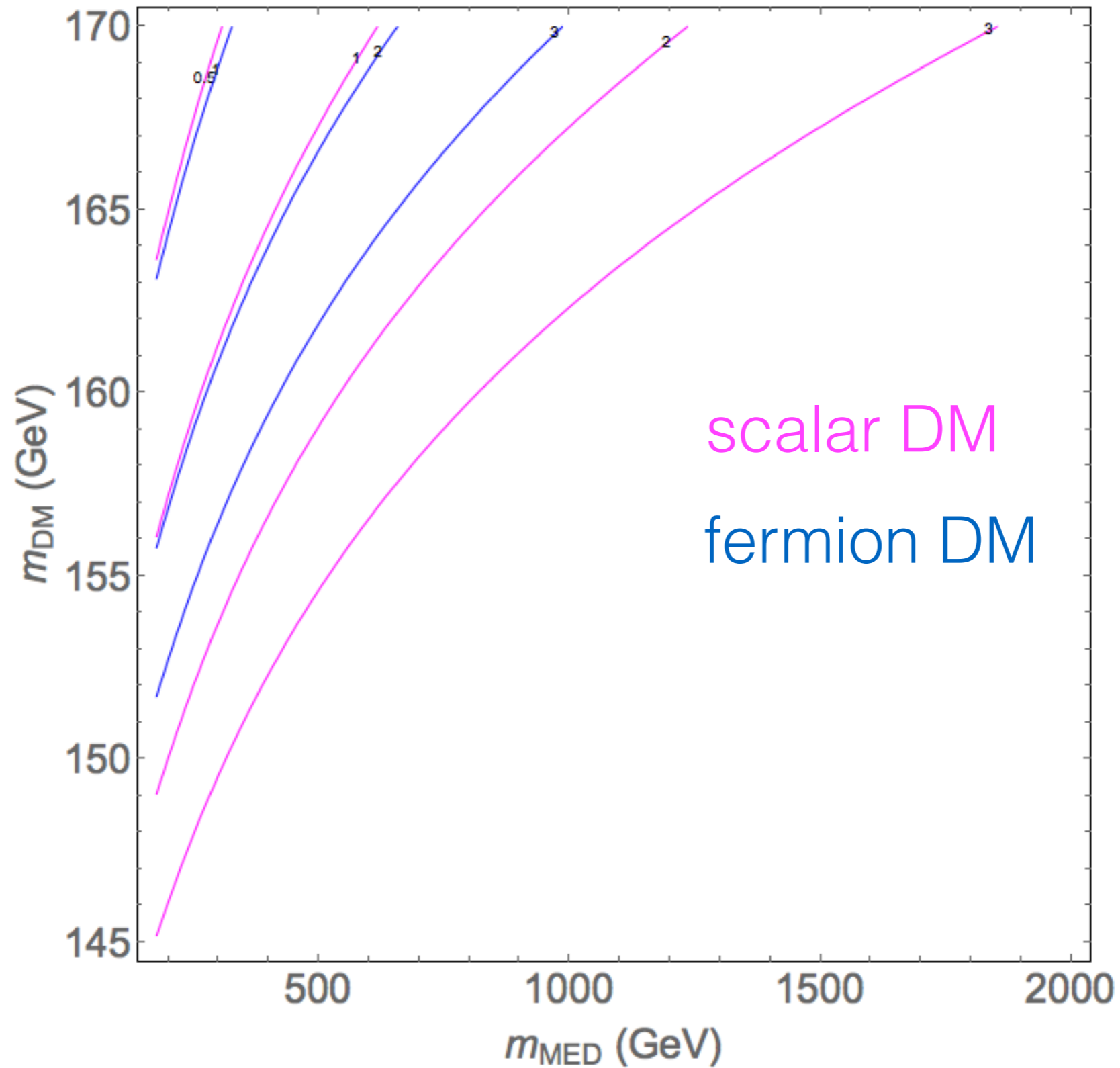
# Back-up Slides



# 3-body annihilations



# Couplings



# Relic abundance

$$\langle\sigma v\rangle_{\chi\chi\rightarrow t\bar{t}} = \left(\frac{n_t^{\text{eq}}}{n_\chi^{\text{eq}}}\right)^2 \langle\sigma v\rangle_{t\bar{t}\rightarrow\chi\chi}$$

$$\langle\sigma v\rangle_{\chi\chi\rightarrow t\bar{t}} = \langle\sigma v\rangle_{t\bar{t}\rightarrow\chi\chi}(1 + \delta)^3 e^{-2\delta x}$$

$$\Omega_\chi h^2 \approx \frac{1.07 \times 10^9 \text{ GeV}^{-1}}{M_{\text{Pl}}} \frac{x_F}{\sqrt{g_*}} \frac{1}{I_a + 3I_b/x_F}$$

$$I_a = x_F(1 + \delta)^3 \int_{x_F}^{\infty} \frac{dx}{x^2} e^{-2\delta x} \quad a$$

$$I_b = 2x_F^2(1 + \delta)^3 \int_{x_F}^{\infty} \frac{dx}{x^3} e^{-2\delta x} \quad b$$

Taylor coefficients of  
 $\langle\sigma v\rangle_{t\bar{t}\rightarrow\chi\chi}$

# Monojet recast

Haisch & Re, 1503.00691

Recast of

CMS collaboration, *Search for dark matter, extra dimensions and unparticles in monojet events in proton-proton collisions at  $\sqrt{s} = 8$  TeV*, *Eur. Phys. J. C* **75** (2015) 235  
[arXiv:1408.3583] [INSPIRE].

$$p_{T,j_1} > 110 \text{ GeV}, \quad |\eta_{j_1}| < 2.4, \quad p_{T,j_2} > 30 \text{ GeV}, \quad |\eta_{j_2}| < 4.5, \quad \Delta\phi_{j_1 j_2} < 2.5, \quad (3.1)$$

where  $\Delta\phi_{j_1 j_2}$  is the azimuthal separation of the two leading jets, which are reconstructed using a radius parameter of  $R = 0.5$ . Another important selection criterion is the imposed jet veto [21], which rejects events if they contain a tertiary jet with  $p_{T,j_3} > 30$  GeV and  $|\eta_{j_3}| < 4.5$ . The CMS measurement is performed in seven distinct  $\cancel{E}_T$  regions, and we find that in the case of the operators  $O_{S,P}^t$  the highest sensitivity is obtained for  $\cancel{E}_T > 450$  GeV. The corresponding 95% confidence level (CL) limit on the fiducial cross section reads

$$\sigma_{\text{fid}}(pp \rightarrow \cancel{E}_T + j) < 7.8 \text{ fb}. \quad (3.2)$$