## Ahead of the Hunt:

## Small field models

## and GW-rleading candidates

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## Motivation

# The hunt for Primordial Gravitational Waves 

- Alternatives to large field models.
- Fundamental Physics?

Ben-Dayan, Brustein JCAP 1009, 007 (2010)

## Outline

I. Recent Results

- The model
- Covering the plane of interest
- Finding the most probable member
II. Discrepancy between analytics and numerics
III.Summary and outlook
I. Recent Results
- The model
- Covering the plane of interest


## 5 Degree polynomial model

$$
V(\phi)=V_{0}\left(1+\sum_{p=1}^{5} a_{p} \phi^{p}\right)
$$

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$$
V(\phi)=V_{0}\left(1-\sqrt{\frac{r_{0}}{8}} \phi\right.
$$

$$
\left.+\frac{\eta_{0}}{2} \phi^{2}+\frac{\alpha_{0}}{3 \sqrt{2 r_{0}}} \phi^{3}+a_{4} \phi^{4}+a_{5} \phi^{5}\right)
$$

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## Results

Why look at these models?

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## Results

## 5 Degree polynomial model



## $n_{s}$

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## OOPS!... Something's wrong

 Discrepancy between analytics \& numerics.
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Discrepancy between analytics \& numerics


Inflationary perturbations from a potential with a step
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## Summary and outlook


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## Summary and outlook

## where do we go from here?

- Go to higher tensor-to-scalar ratios.
- See what BICEP3 yields, and adjust.
- Find a better analytic expression from first principles?
- Quantify perturbations - give better analytic tools


## Questions?

## Thank you! <br> 

