MSSM dark matter and the muon g-2

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Introduction

The LHC has resumed its search for Supersymmetry (SUSY) in run 2, but direct evidence is still yet to be found. The search continues on!



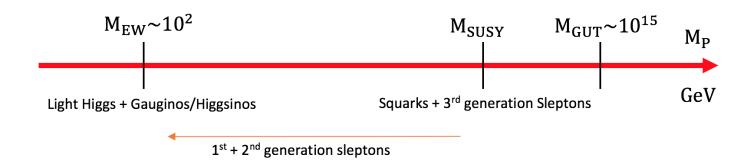
Introduction

The observation of a Higgs at 125 GeV has strengthened the need for SUSY to appear at the weak-scale.

- Tree-level higgs mass $\sim m_Z$.
- Existence of electroweakinos (partners of EW gauge bosons).
- A light neutralino great for DM!
- Composition of DM is important (Wino, Bino, Higgsino).

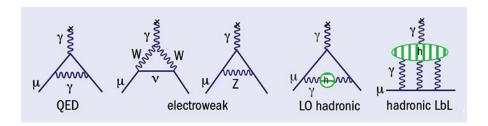
Minimal SUSY mass hierarchy

- Universal squark and 3rd gen slepton masses decoupled
- 125 GeV higgs finely-tuned
- Gauginos/higgsinos at weak scale, protected by chiral symmetry
- ◆ Light 1st and 2nd generation sleptons allowed by FCNC constraints
 → muon g-2



The muon g-2

Contributions to the SM:



 Main theoretical uncertainty comes from LO Hadronic loop contributions (quarks and gluons)

$$20.6 \times 10^{-10} < \Delta a_{\mu} < 36.6 \times 10^{-10}$$
 (1 σ)

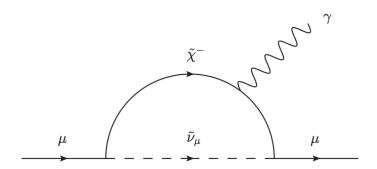
$$20.6 \times 10^{-10} < \Delta a_{\mu} < 36.6 \times 10^{-10}$$
 (1 σ)
 $12.6 \times 10^{-10} < \Delta a_{\mu} < 44.6 \times 10^{-10}$ (2 σ)

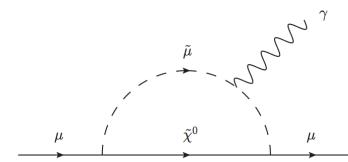
where

$$\Delta a_{\mu} \equiv a_{\mu}^{ ext{exp}} - a_{\mu}^{ ext{SM}}$$

The muon g - 2 in SUSY

Contributions come from sneutrino-chargino and smuon-neutralino loop diagrams





The muon g-2 in SUSY

Contribution from the MSSM:

$$\Delta a_{\mu} = \frac{\alpha m_{\mu}^{2} \mu \tan(\beta)}{4\pi} \left[\frac{M_{2}}{\sin^{2} \theta_{W} m_{\tilde{\mu}_{L}}^{2}} \left(\frac{f_{\chi}(M_{2}^{2}/m_{\tilde{\mu}_{L}}^{2}) - f_{\chi}(\mu^{2}/m_{\tilde{\mu}_{L}}^{2})}{M_{2}^{2} - \mu^{2}} \right) + \frac{M_{1}}{\cos^{2} \theta_{W}(m_{\tilde{\mu}_{R}}^{2} - m_{\tilde{\mu}_{L}}^{2})} \left(\frac{f_{N}(M_{1}^{2}/m_{\tilde{\mu}_{R}}^{2})}{m_{\tilde{\mu}_{R}}^{2}} - \frac{f_{N}(M_{1}^{2}/m_{\tilde{\mu}_{L}}^{2})}{m_{\tilde{\mu}_{L}}^{2}} \right) \right]$$

 f_{χ} and f_{N} are loop functions:

$$f_{\chi}(x) = \frac{x^2 - 4x + 3 + 2\ln(x)}{(1 - x)^3}, \quad f_{\chi}(1) = -2/3$$

$$f_{N}(x) = \frac{x^2 - 1 - 2x\ln(x)}{(1 - x)^3}, \quad f_{N}(1) = -1/3$$

Explaining the muon g-2 in the MSSM

• The following particles are important in analyzing the $(g-2)_{\mu}$ in the MSSM:

$$\tilde{\mu}, \tilde{\nu}_{\mu}, \tilde{\chi}^0, \tilde{\chi}^{\pm}$$
 (1)

- We can heavily constrain the muon g-2 through slepton and chargino searches at colliders
- Smuons should be kept light (less than around 500 GeV) to increase contribution to the $(g-2)_{\mu}$
- Dark Matter (Direct/Indirect) searches can constrain neutralino LSPs in R-Parity conserving SUSY

Constraints from Experiment

• LEP constraints on chargino and Slepton masses:

$$m_{ ilde{l}_L}, m_{ ilde{l}_R} > 100 \, ext{GeV} \quad (I=e,\mu) \ m_{ ilde{\chi}_1^\pm} > 105 \, ext{GeV}$$

Constraints on neutralino LSP as a DM candidate:

$$m_{\tilde{\chi}_1^0} > 30 \,\mathrm{GeV}$$

Higgs mass from ATLAS/CMS:

$$123 < m_{h^0} < 127 \, \text{GeV}$$

- Higgs precision constraints (LEP, Tevatron and LHC)
- Dark matter relic density (PLANCK 2013)

$$\Omega h^2 = 0.112 \pm 0.006 \quad (1\sigma)$$

WIMP-nucleon Spin-Independent Cross Section (LUX 2016)

MSSM Parameter Scan

We calculate the $(g-2)_{\mu}$ and mass spectrum in the MSSM using FeynHiggs-1.12.0:

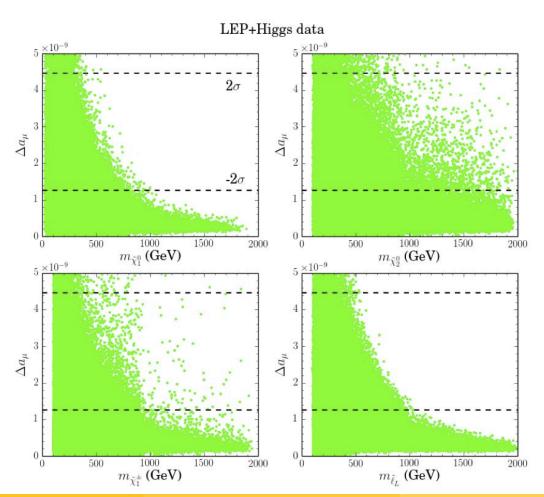
- Decoupled Squarks at 5 TeV (Ignore B-Physics constraints)
- Stau sleptons $m_{\tilde{\tau}_I} = m_{\tilde{\tau}_R} = 5 \, \text{TeV}$
- Gluino mass $M_3 \sim 3$ TeV
- Trilinear coupling A_t in range $|A_t| < 5 \, \text{TeV}$ (We keep $|X_t/M_S| < 2$ to avoid charge/colour-breaking minima)
- Rest of higgs sector decoupled by setting $m_{A^0} = 2 \text{ TeV}$

Parameter scan range:

$$egin{aligned} 10 < an(eta) < 50, \ |M_1|, |M_2|, |\mu| < 2 \, ext{TeV}, \ 0.1 < m_{\widetilde{l}_l}, m_{\widetilde{l}_P} < 2 \, ext{TeV}, \end{aligned} \qquad (\emph{l} = \emph{e}, \mu)$$

Higgs mass calculated in FeynHiggs, precision constraints in HiggsBounds-4.2.1. SUSY spectrum calculated in SPheno, MicrOmegas to calculate DM relic density and SI WIMP-nucleon CS.

Limits on neutralinos, charginos and smuons



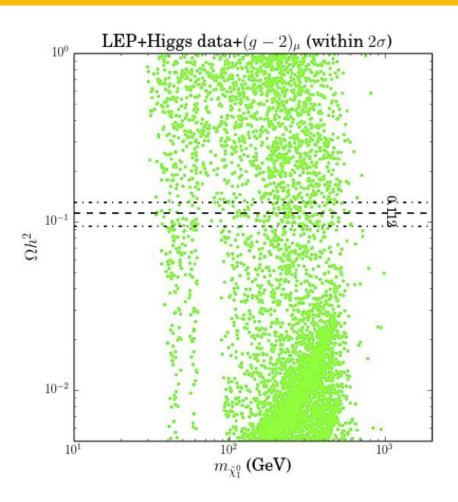
(N)LSP component

Parameters	LSP	NLSP
$M_1 > M_2 > \mu$	Higgsino	Wino
$M_1 > \mu > M_2$	Wino	Higgsino
$M_2 > \mu > M_1$	Bino	Higgsino
$\mu > M_2 > M_1$	Bino	Wino

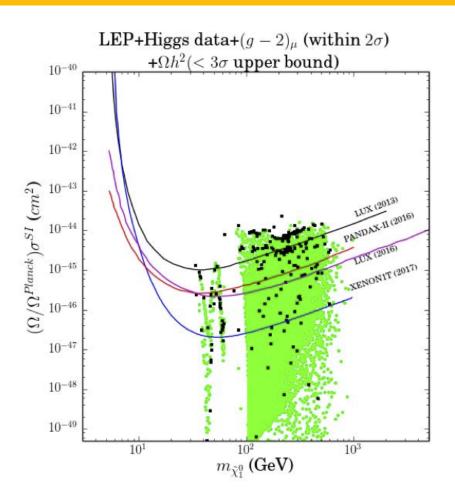
Constraints on χ_1^0 vary for different compositions of Bino, Wino and Higgsinos

- It is well known that pure Bino-like DM relics are typically overabundant (suppressed annihilation cross section), except in the case where the bino co-annihilates with other sparticles
- We can enhance the annihilation rate with a wino or higgsino component in χ_1^0
- To avoid significant constraint, for any LSP abundance less than the relic density, we assume additional DM component (possibly axion-like DM)

Relic Density, Ωh^2



WIMP-nucleon SI Cross Section



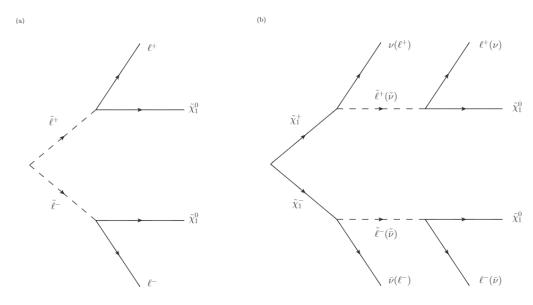
Collider Simulation

We study constraints from multilepton + MET searches at the LHC.

- We study electroweakinos at $\sqrt{s}=8$ TeV LHC from slepton/sneutrino and W/Z decays
- Parameter sets that pass the previous collider and direct/indirect dark matter searches are considered
- ullet Points are considered within the 2σ limit of Δa_{μ}
- We also present the prospects for electroweakino searches with a 100 TeV collider
- NLO events are simulated using MadGraph 5 interfaced with Pythia
- These are passed to CheckMATE-1.2.2 to check exclusion limits at 95% CL

Electroweakinos and sleptons at colliders

$$2\ell + \cancel{E}_T$$
 (2 leptons + missing energy) ¹



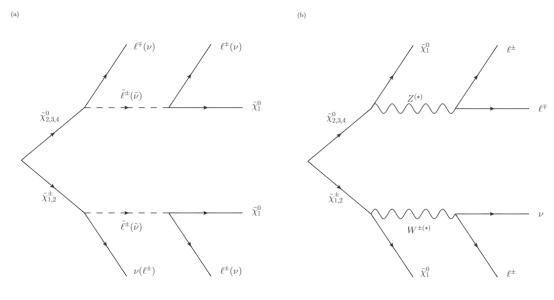
(a) via direct slepton decays

(b) via sleptons/sneutrinos

¹atlas_conf_2013_049

Electroweakinos and sleptons at colliders

$$3\ell + \mathcal{E}_T$$
 (3 leptons + missing energy) ²

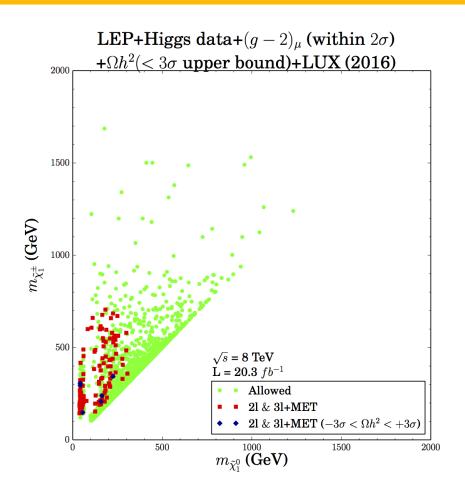


(a) via sleptons/sneutrinos

(b) via gauge bosons

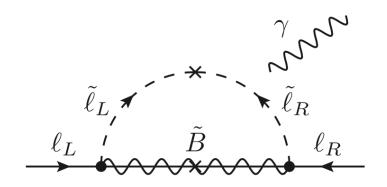
²atlas_1402_7029

Results in $m_{\chi_1^0}-m_{\chi_1^\pm}$ plane



Large μ case

It has been noted that one can explain the $(g-2)_{\mu}$ can be explained with a dominant bino-smuon loop contribution.



This is enhanced with a large smuon left-right mixing.

Too large, and this can spoil the electroweak vacuum stability.

Large μ case

We scan the extended region:

$$egin{aligned} 10 < an(eta) < 50, \ |\mathcal{M}_1|, |\mathcal{M}_2| < 3 \, ext{TeV}, \ 10 < \mu < 100 \, ext{TeV}, \ 0.1 < m_{ ilde{I}_I}, m_{ ilde{I}_R} < 2 \, ext{TeV}, \end{aligned} \qquad (\emph{I} = \emph{e}, \mu)$$

with staus decoupled at $m_{ ilde{ au}_L}=m_{ ilde{ au}_R}=10\,\mathrm{TeV}$ and $A_{ au}=0.$

To explain $(g-2)_\mu$ within 2σ , we find upper limits of $m_{\tilde{\chi}_1^0} < 2.4\,\text{TeV}$ and $m_{\tilde{\ell}_1} < 1.1\,\text{TeV}$.

The previous DM constraints severely limit this case, and so is **not the preferred scenario**.

100 TeV Analysis

The 3 lepton + MET events at 100 TeV are expected to have the largest reach over the MSSM parameter space.

We scale the signal (S) and background (B) events for the 8 TeV analysis by the ratio:

$$N^{100\,\text{TeV}} = (\sigma^{100\,\text{TeV}}/\sigma^{8\,\text{TeV}})(3000\,\text{fb}^{-1}/20.3\,\text{fb}^{-1})N^{8\,\text{TeV}}$$

Sources of background (B):

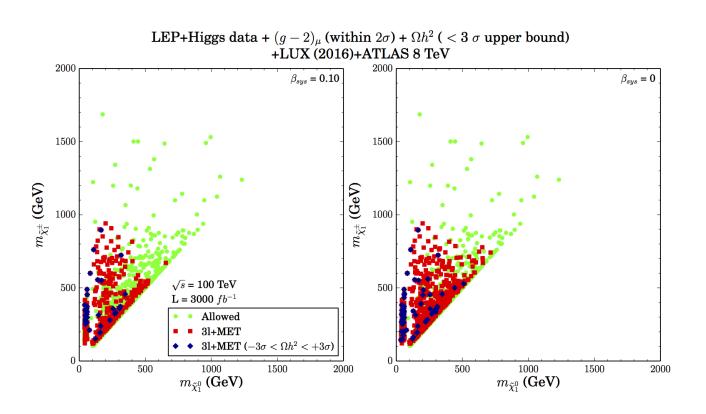
- WZ, ZZ, H
- ttV + ttZ
- VVV
- Reducible (t single/pair, WW, single W/Z with jets or photons)

We exclude events corresponding to:

$$\frac{S}{\sqrt{B + (\beta_{sys}B)^2}} \ge 2$$

where β_{sys} parameterizes the systematic uncertainty.

Results for 100 TeV Analysis



Conclusions

- We studied constraints from direct/indirect measurements on the MSSM with heavy squarks and light sleptons.
- A 100 TeV collider could potentially probe almost the entire mass range for electroweakinos in this model as an explanation for the muon $(g-2)_{\mu}$ and dark matter.
- One can further the analysis using monojet-like signals with greater sensitivity to the degenerate mass region.
- Our 100 TeV analysis can be considered a preliminary one, that can be improved once the collider environment details are known (and/or a public code is released).