

# Relic Abundance in Secluded Dark Matter Scenario with Massive Mediator

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The relic abundance of the dark matter (DM) particle  $d$  is studied in a secluded DM scenario, in which the  $d$  number decreasing process dominantly occurs not through the pair annihilation of  $d$  into the standard model particles, but via the  $dd \rightarrow mm$  scattering process with a subsequently decaying mediator particle  $m$ . It is pointed out that the cosmologically observed relic abundance of DM can be accomplished even with a massive mediator having a mass  $m_m$  non-negligibly heavy compared with the DM particle mass  $m_d$ .

In the degenerated  $d$ - $m$  case

( $m_d = m_m$ ),

the DM relic abundance is realized by adjusting the  $dd \rightarrow mm$  scattering amplitude large enough and by choosing an appropriate mediator particle life-time.

The DM evolution in the early universe exhibits characteristic “terrace” behavior, or two-step number density decreasing behavior, having a “fake” freeze-out at the first step.

Based on these observations, a novel possibility of the DM model buildings is introduced

in which the mediator particle  $m$  is unified with the DM particle  $d$  in an approximate dark symmetry multiplet.

A pionic DM model is proposed to illustrate this idea in a renormalizable field theory framework.

S. Okawa, M. Tanabashi and M. Yamanaka,

“Relic Abundance in Secluded Dark Matter Scenario with Massive Mediator,”  
arXiv:1607.08520 [hep-ph].

## Summary

The relic abundance of the dark matter (DM) particle  $d$  is studied in a secluded DM scenario, in which the  $d$  number decreasing process dominantly occurs not through the pair annihilation of  $d$  into the standard model particles, but via the  $dd \rightarrow mm$  scattering process with a subsequently decaying mediator particle  $m$ . It is pointed out that the cosmologically observed relic abundance of DM can be accomplished even with a massive mediator having a mass  $m_m$  non-negligibly heavy compared with the DM particle mass  $m_d$ .

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