

Streamer phenomenology in streamer-mode RPCs

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Two or three things I learned about streamer-operated RPCs.... (personal pick of past results)

Outline:

- 1) OPERA RPC system**
- 2) Streamer formation time**
- 3) Streamer and particle parameters**
- 4) Streamer and gas mixtures**

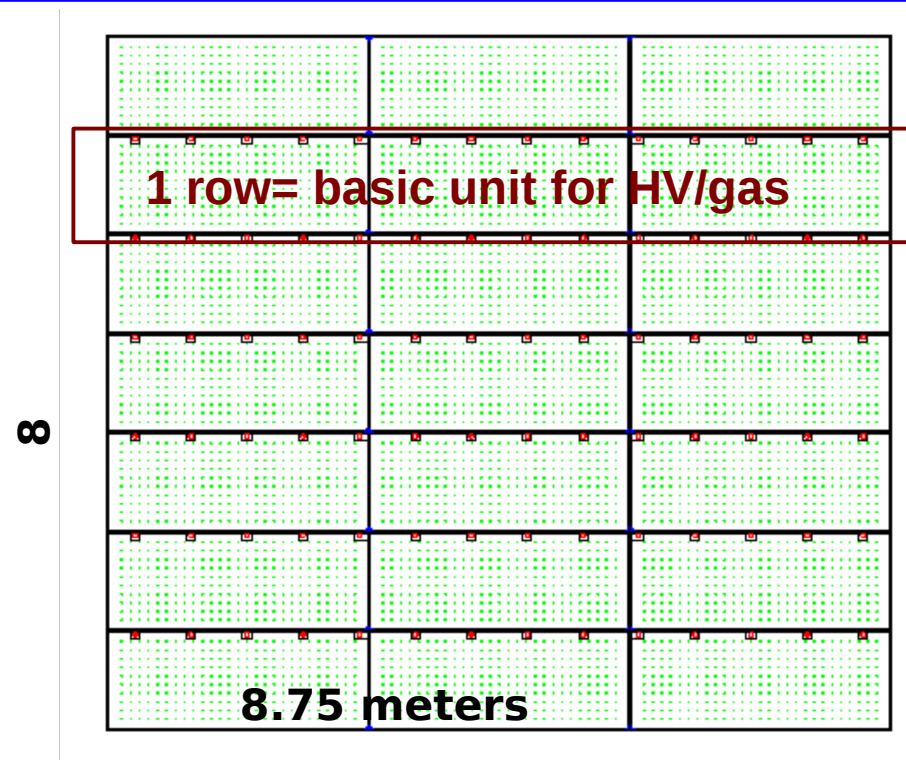
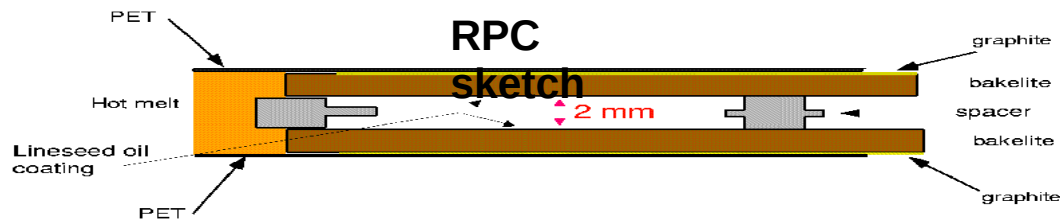
OPERA RPCs

Located in 2 cm gaps between iron slabs:

- Track reconstruction inside magnet
- Shower leakage measurement
- Trigger & Timing for the drift tubes

1 layer = 21 RPCs of size (2.9*1.1) m²
 1 spectrometer = 504 RPCs/XPCs

- High resistivity bakelite electrodes (low rate expected): $\rho > 5 \cdot 10^{11} \Omega \text{ cm}$ @20°C
- Special curved contour chambers
- Streamer mode operation (large signals)
- Read-out by means of ~8 m strips with 2.6 (3.5) cm pitch for bending (orthogonal) view



Total surface of the system ~ 3200 m²

Number of digital electronics channels ~ 28000

Streamer operation with $\text{Ar}/\text{C}_2\text{H}_2\text{F}_4/\text{isoC}_4\text{H}_{10}/\text{SF}_6 = 75.4/20/4/0.6$ (5 refills/day, open-flow system).

See JINST4 (2009) P04018 for Front-End discriminators and DAQ.

Streamer formation time measurement

At some point in OPERA we needed a precise time calibration....

RPC time response calibration

Reference
scintillator

Test RPC

Trigger RPC

Cosmic rays.

OPERA gas mixture.

Trigger = scintillator and Trigger RPC.

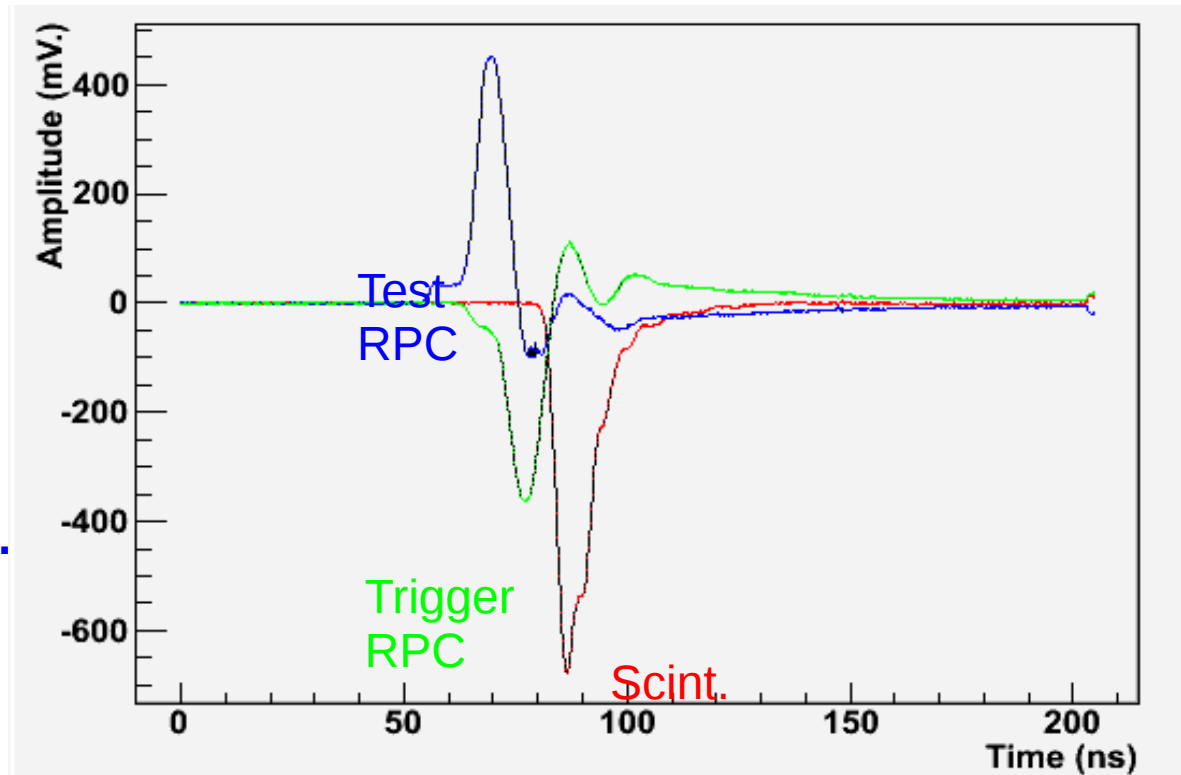
RPCs read-out by means of 3.5*3.5 cm² pads.

Scintillator time response calibrated with ps laser.

All other set-up delays calibrated with scope.

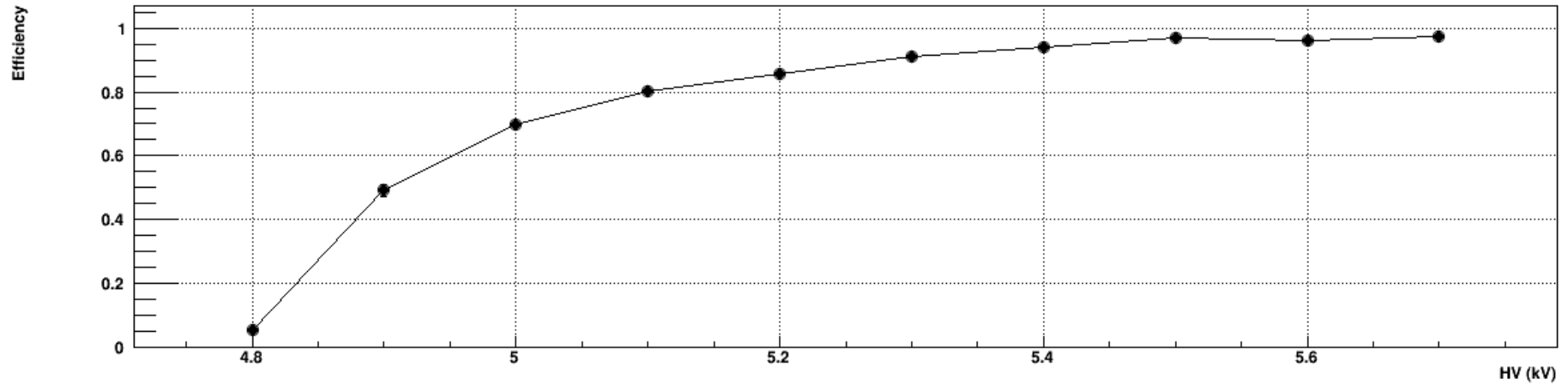
Analogic signals acquired with a waveform digitizer
(5 Gsamples/s).

Analysis refinements:
Cuts on amplitude of scintillator
and trigger RPC signals.
Threshold on test RPC: 26 mV/50 Ohm.

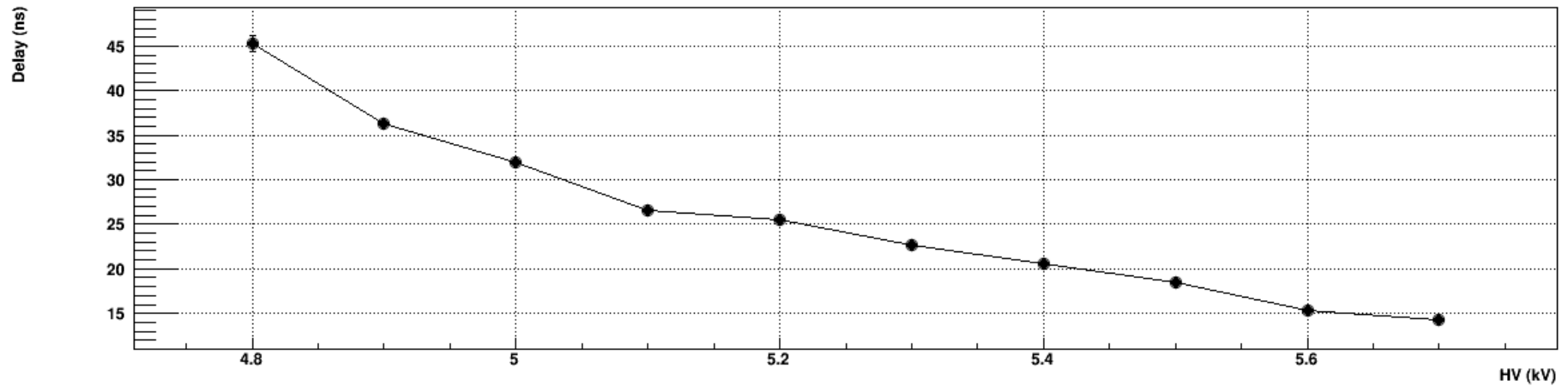


Results

Efficiency vs HV



Delay vs HV



For reasonable efficiency values, there is a 20 ns delay between the passage of the particle and the streamer formation.

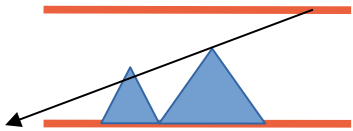
Streamer and particle properties

OPERA data (cosmic and CNGS produced muons) in the first two slides and one test-beam at LNS (parasitic mode).

Angular dependence

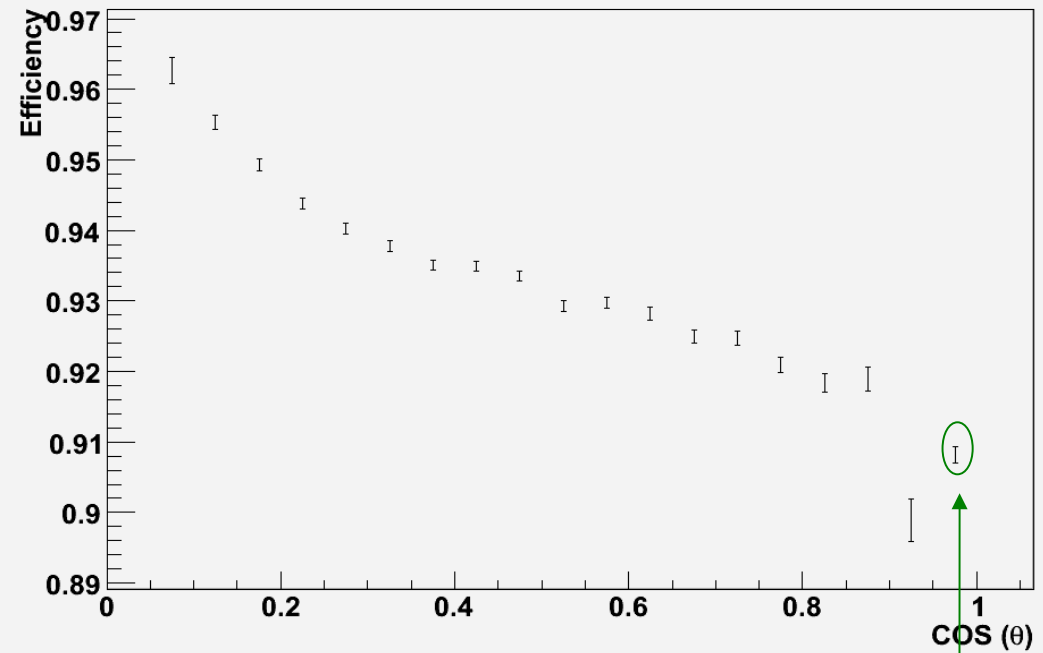
Higher efficiency and cluster size for inclined tracks:

- Geometry effect (more streamers in parallel for inclined tracks)

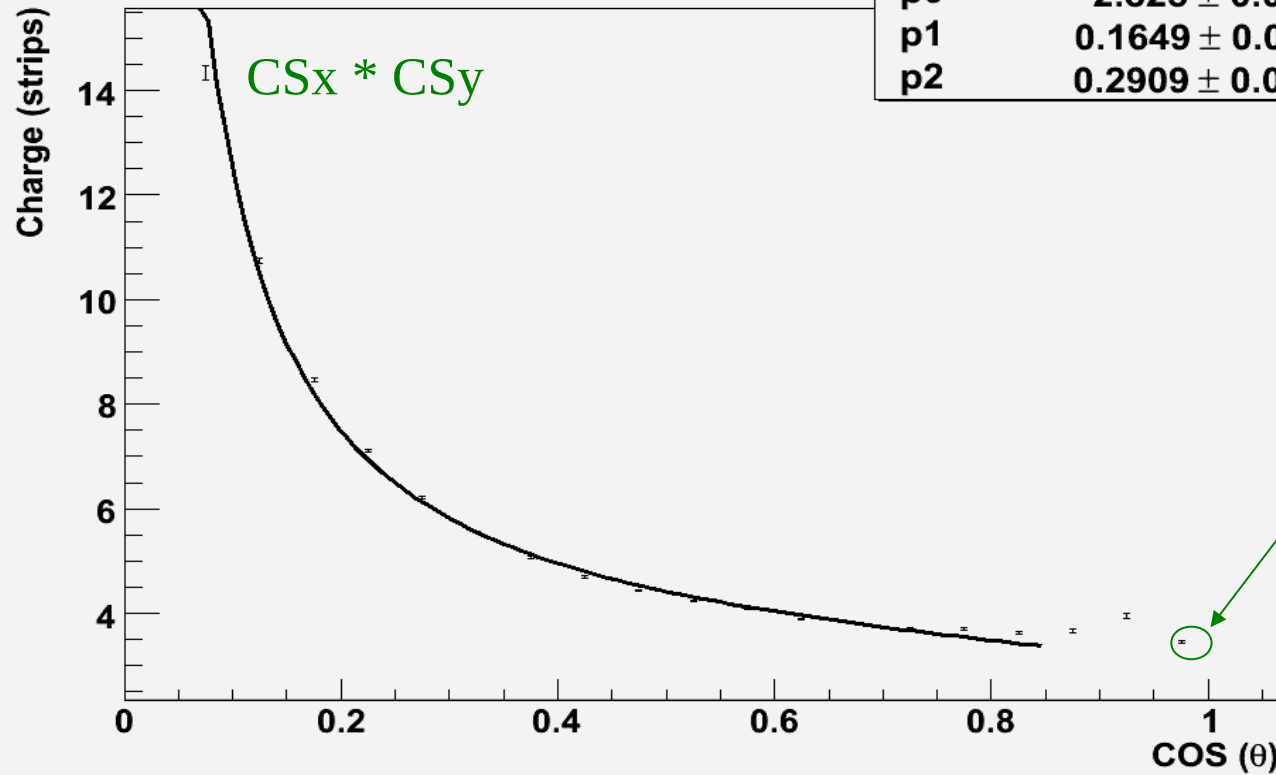


θ = angle wrt to z-axis, normal to RPC layers

Efficiency



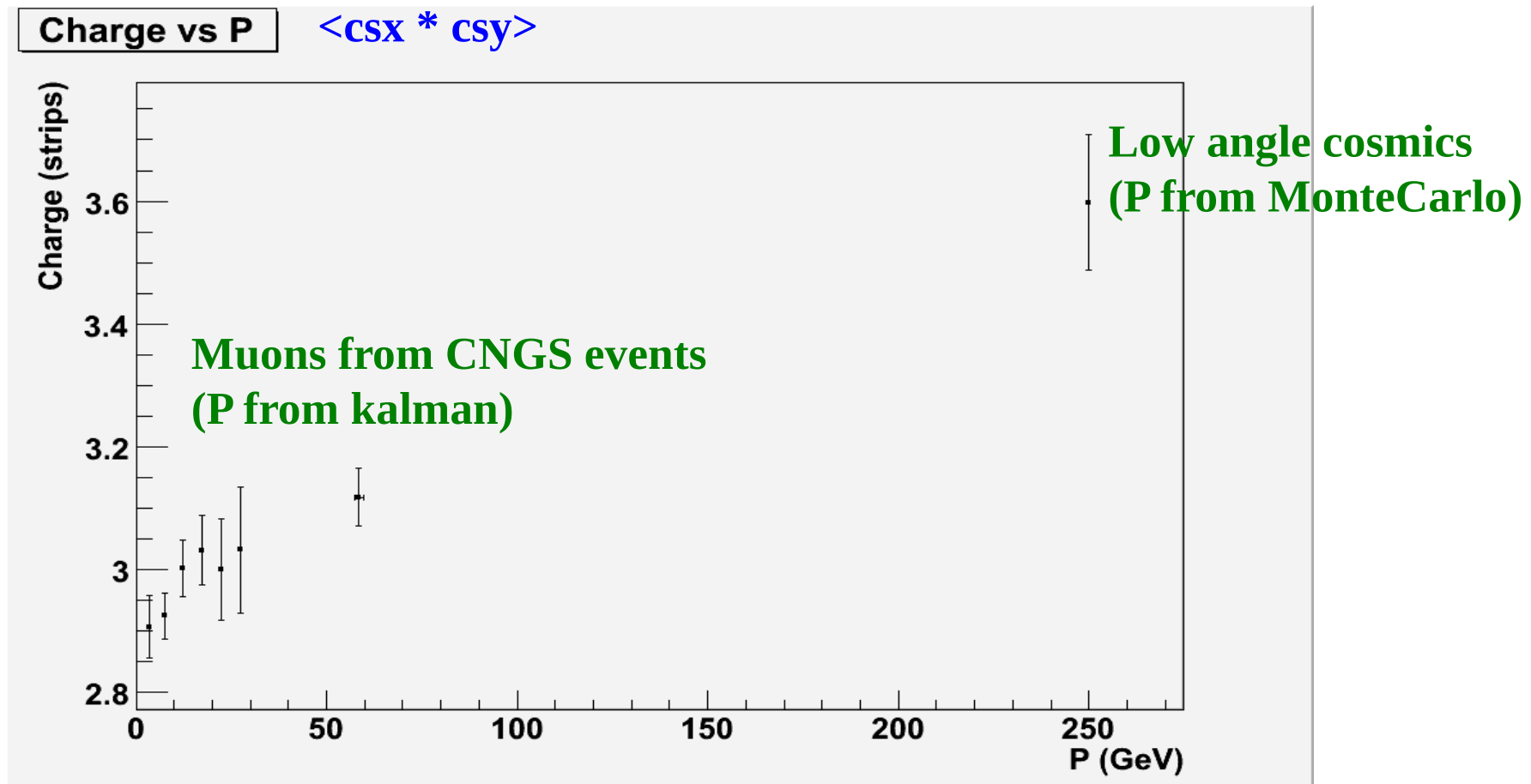
Total Cluster Size



χ^2 / ndf	534.9 / 13
p0	2.825 ± 0.03064
p1	0.1649 ± 0.07672
p2	0.2909 ± 0.02465

Beam events
($p \sim 20 \text{ GeV} \ll \text{cosmic muons}$)

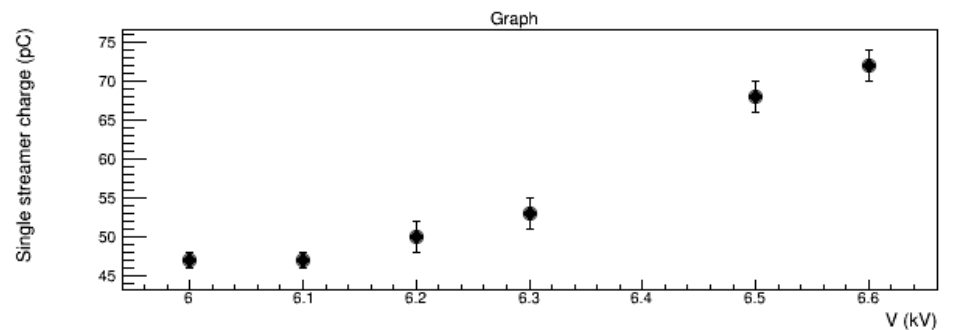
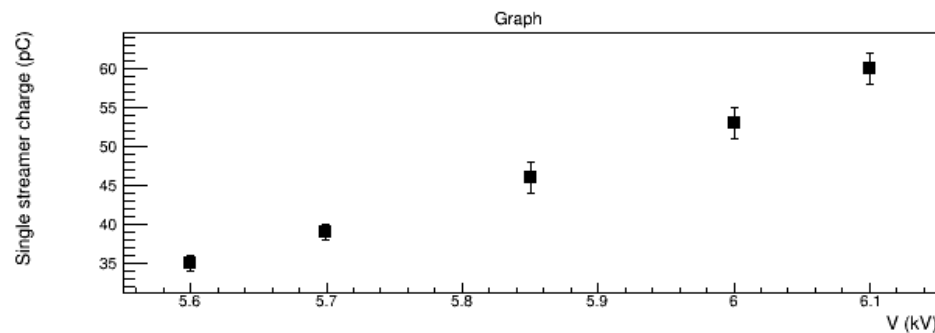
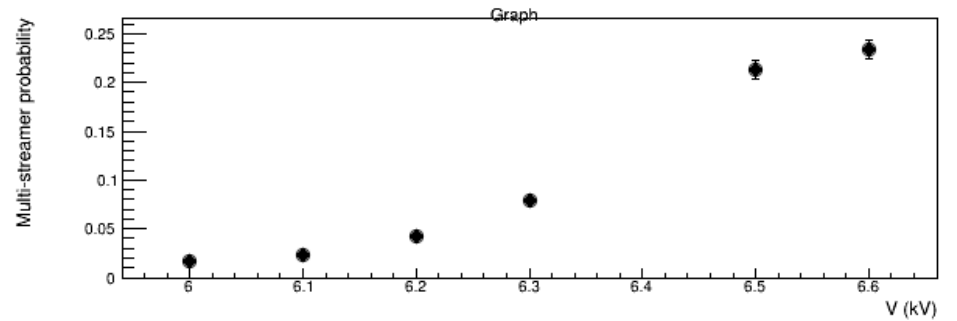
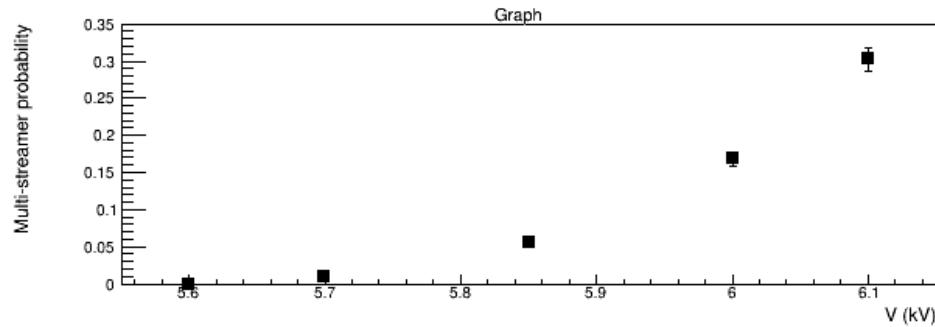
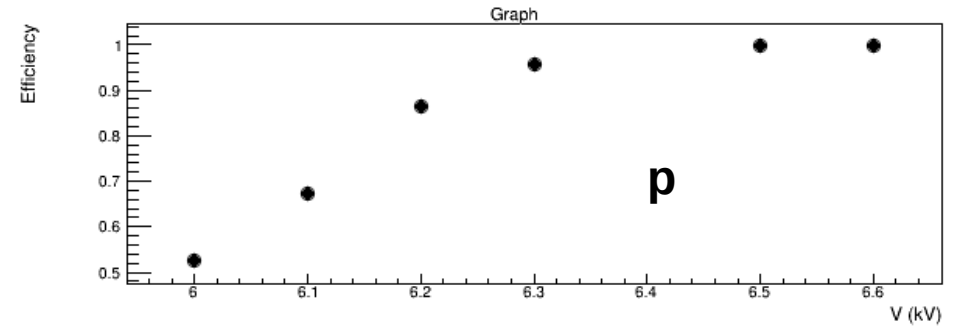
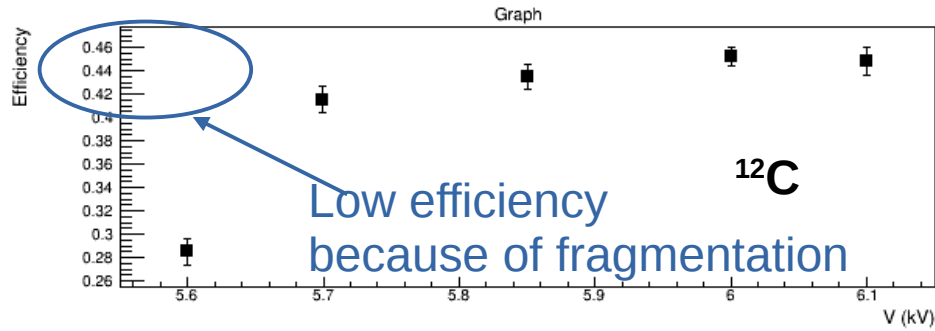
Momentum dependence for perpendicular tracks



For particles impinging with the same inclination, the cluster size increases with the momentum (suggesting a streamer charge dependence on the primary ionization ?)

Highly ionising particles

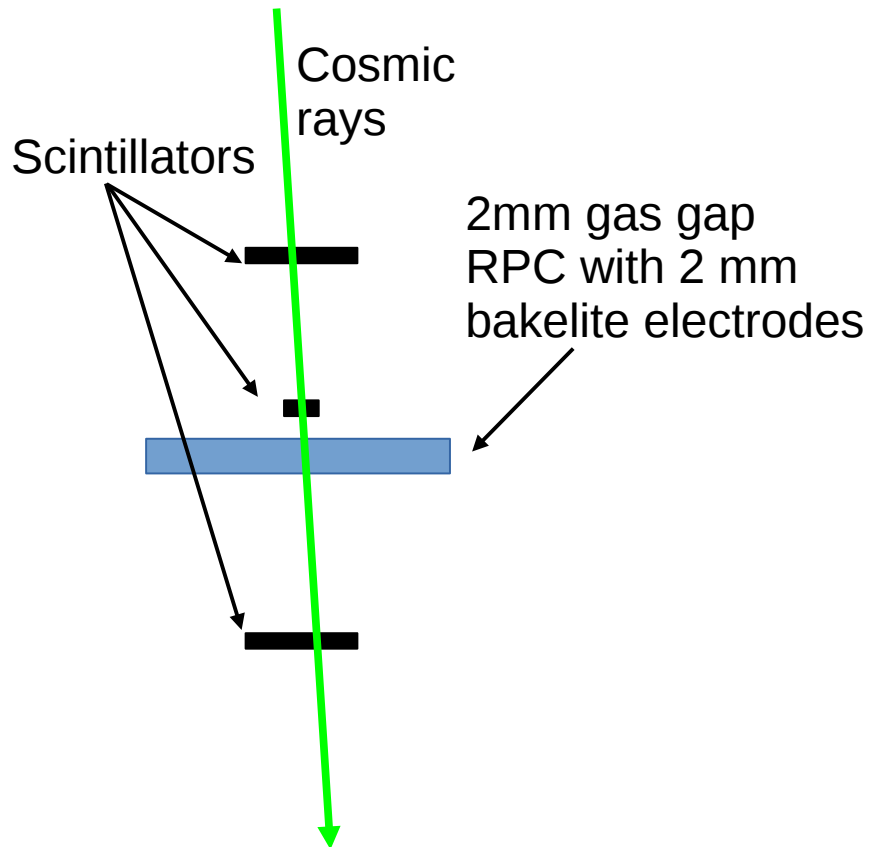
Test-beam in parasitic mode at LNS with 80 MeV/u p and ^{12}C



On ^{12}C ions (36 times more ionizing) streamers develop 500 V lower than on protons.

Streamer and gas mixture

Experimental set-up



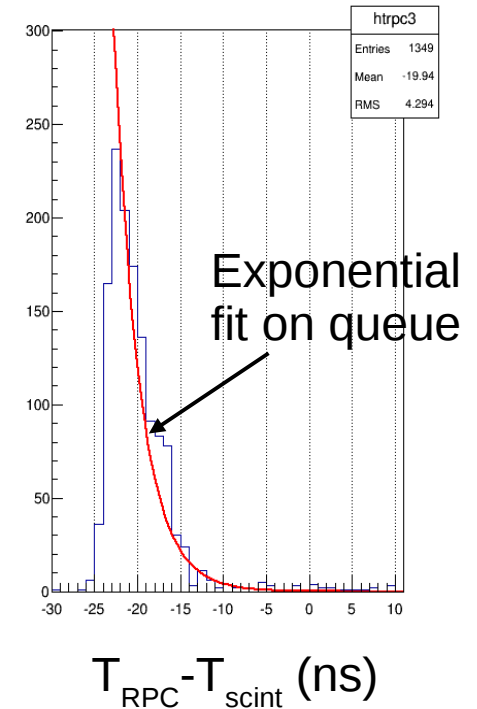
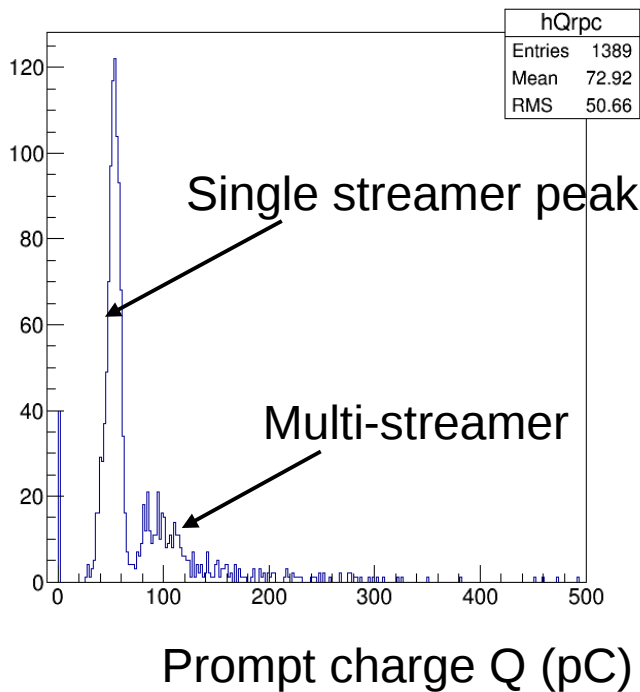
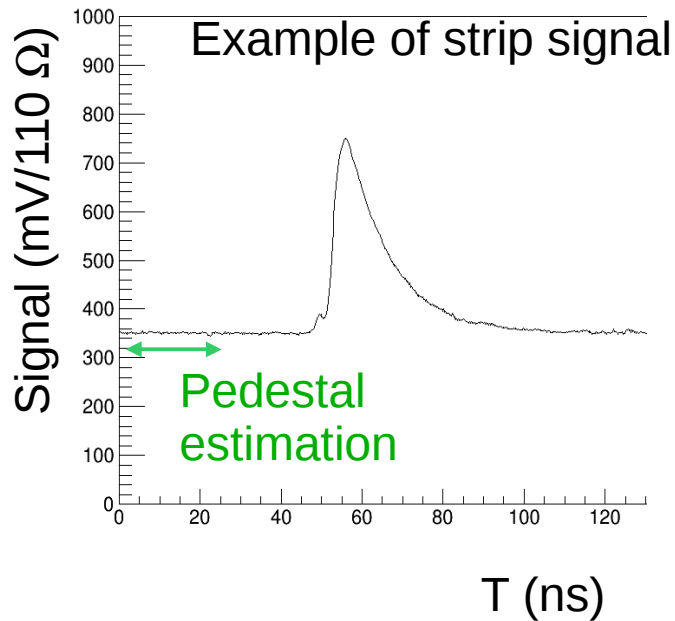
RPC under test read-out by 3.5 cm wide copper strips of 70 cm length. Terminated on 110Ω on the read-out side, on 25Ω on the opposite side.

Cosmic ray trigger made by three scintillators coincidence (the central one used also as time reference).

Signals from 11 strips acquired by means of N6742 CAEN digitizer (5 Gsamples/sec).

Strip signals discriminated at 50 mV/ 110Ω .

Analysis techniques description



Signal treatment:

Pedestal subtraction channel by channels using first 100 samples.

Common mode noise estimation using strips without signals.

RPC and streamer properties from measured distributions:

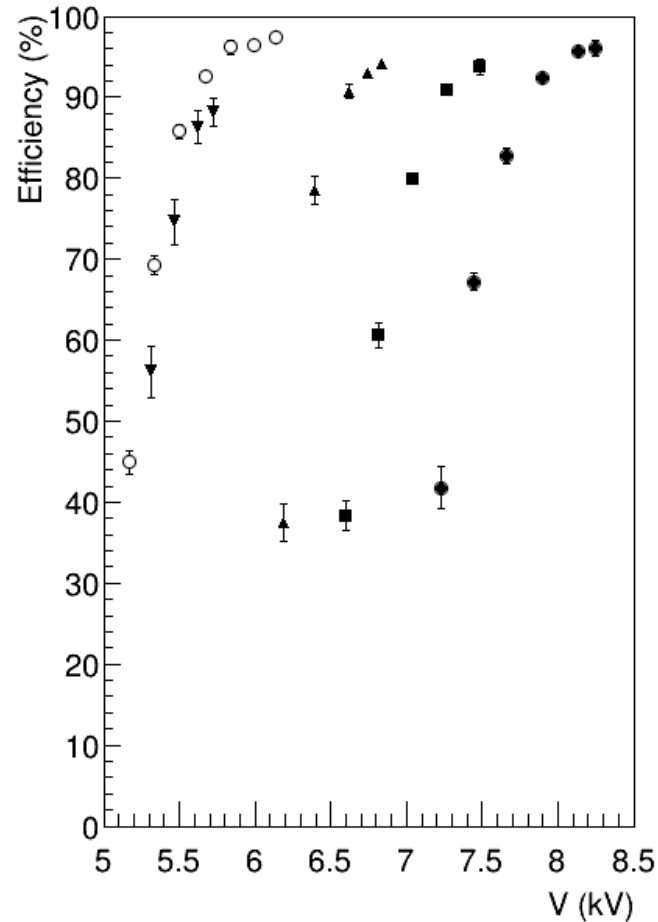
Prompt charge: efficiency, multistreamer probability, single streamer charge.

Single strip events: streamer amplitude, FWHM, risetime (10% - 90% of amplitude).

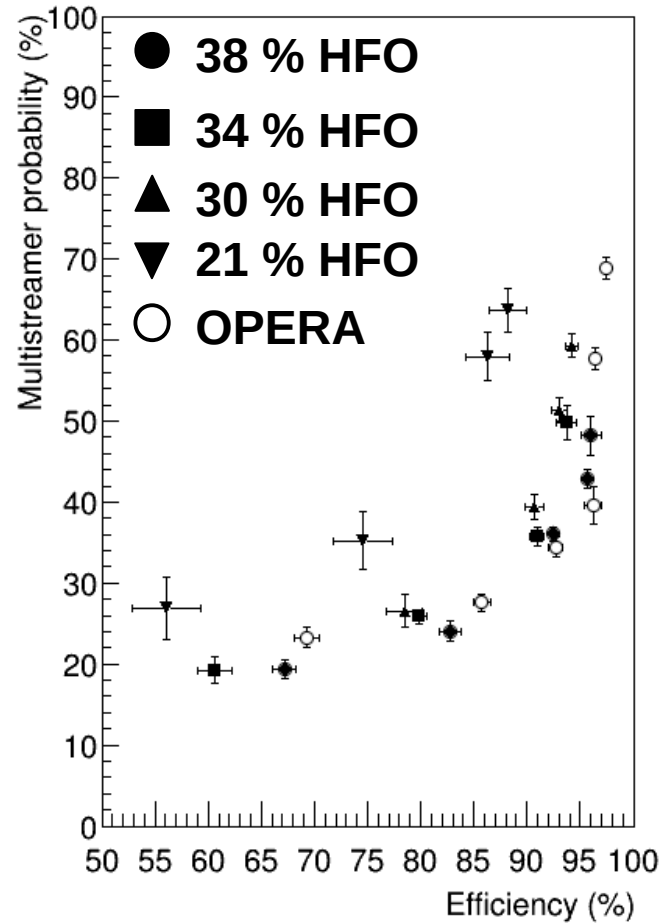
$T_{RPC} - T_{scint}$: streamer arrival time (relative to the scint.), time resolution (exp fit on the queue).

Ar/HFO-1234ze binary mixtures

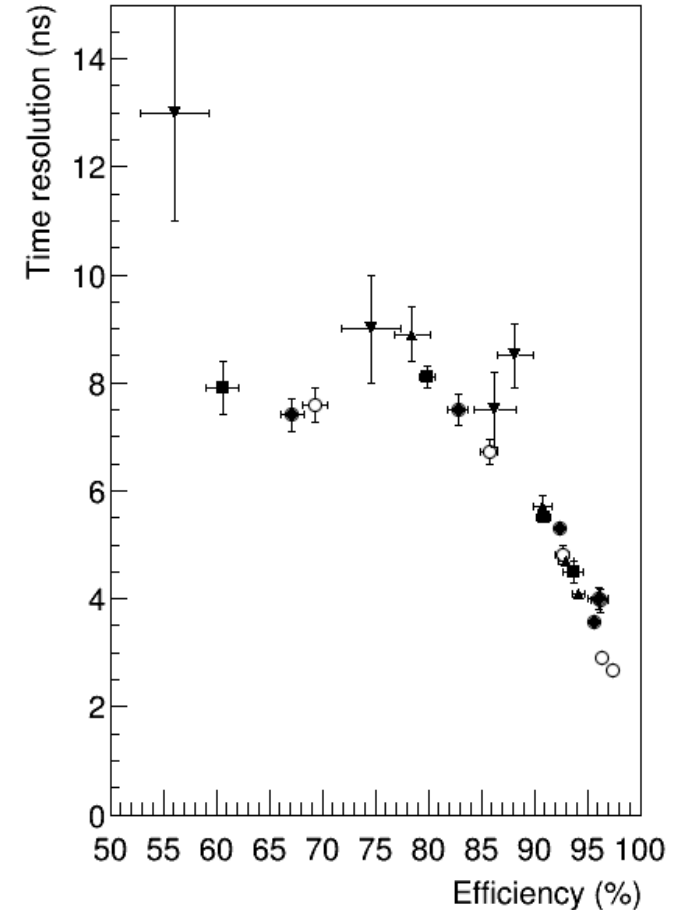
Efficiency



Multistreamer probability



Time resolution (ns)



The higher the HFO concentration, the higher the operating voltage and the quenching power.

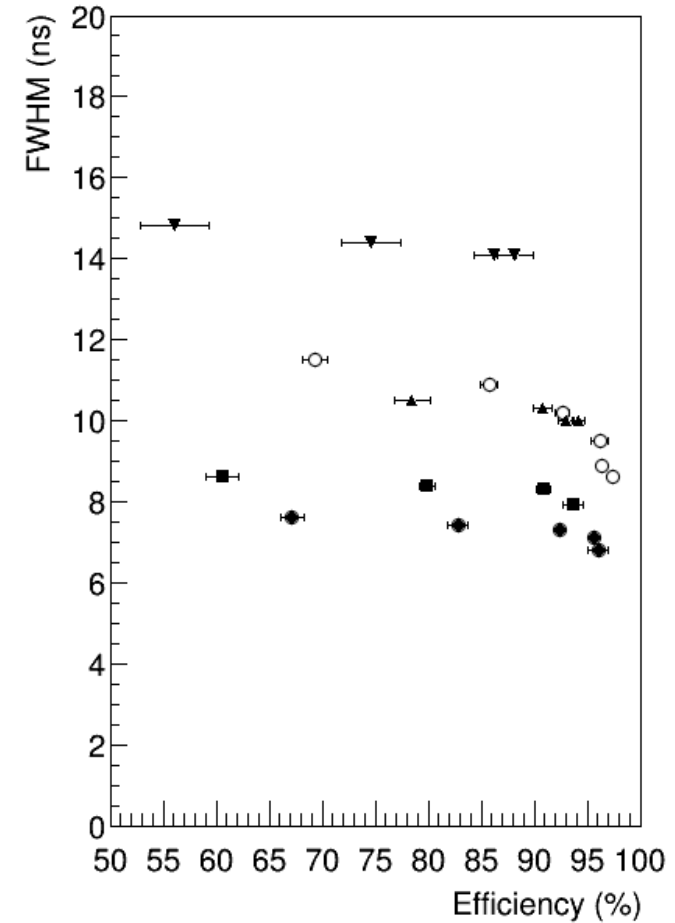
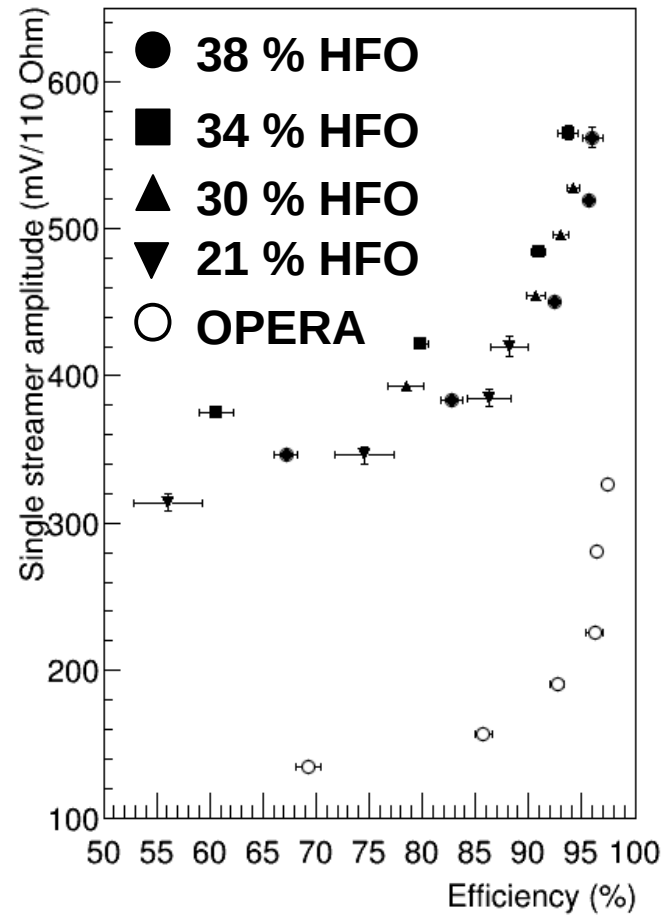
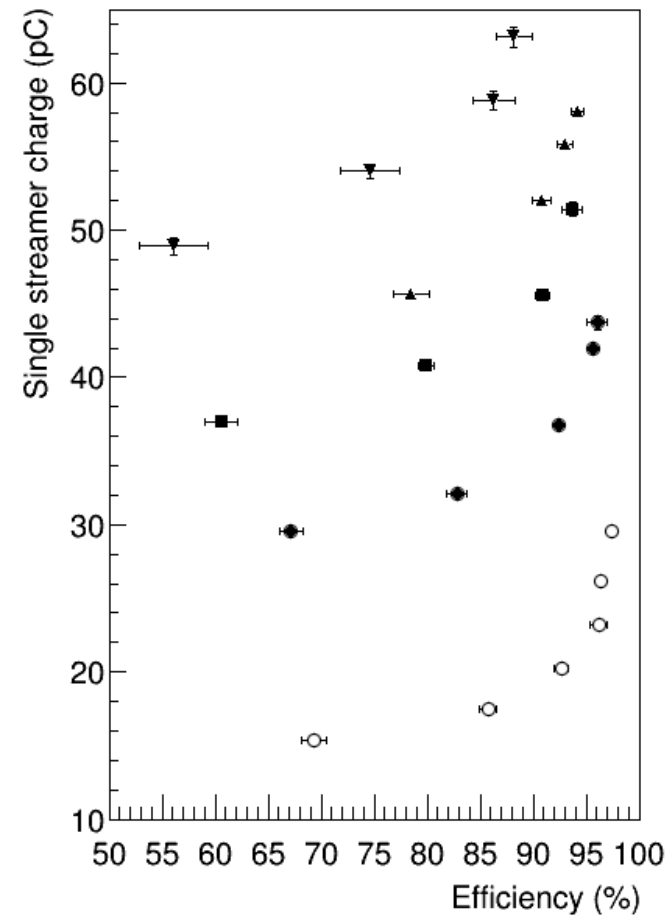
Better to keep HFO concentration to 30%, to limit multistreamers.

Ar/HFO-1234ze binary mixtures

Single streamer charge (pC)

Single streamer amplitude (mV)

FWHM (ns)



The higher the HFO concentration, the higher the operating voltage and the quenching power.
Better to keep HFO concentration to 30%, to limit multistreamers.
The higher the HFO concentration, the lower the prompt charge (FWHM smaller).

Gas mixtures for streamer RPCs

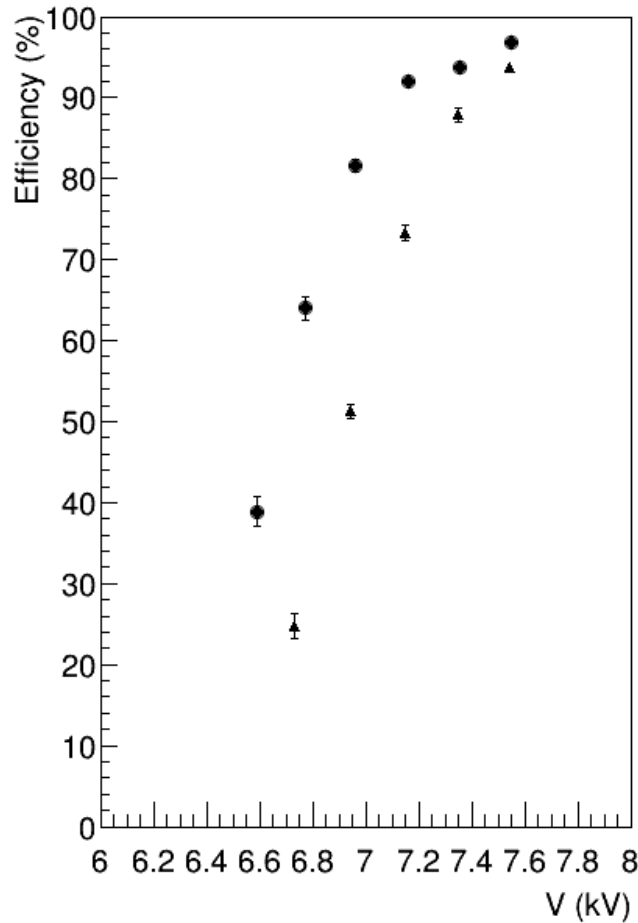
Typical mixture is composed of Argon/Quencher.

Quencher can be: R134a+Iso-butane, HFO-1234ze, HFO-1234yf, Iso-butane

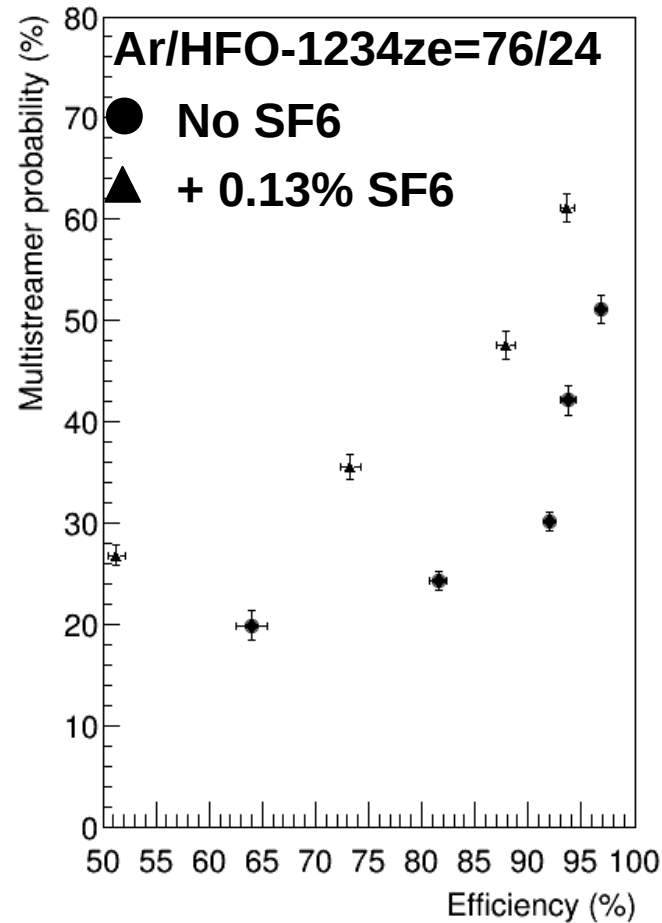
Quencher cannot be: CO_2 , CF_4 .

SF₆ addition

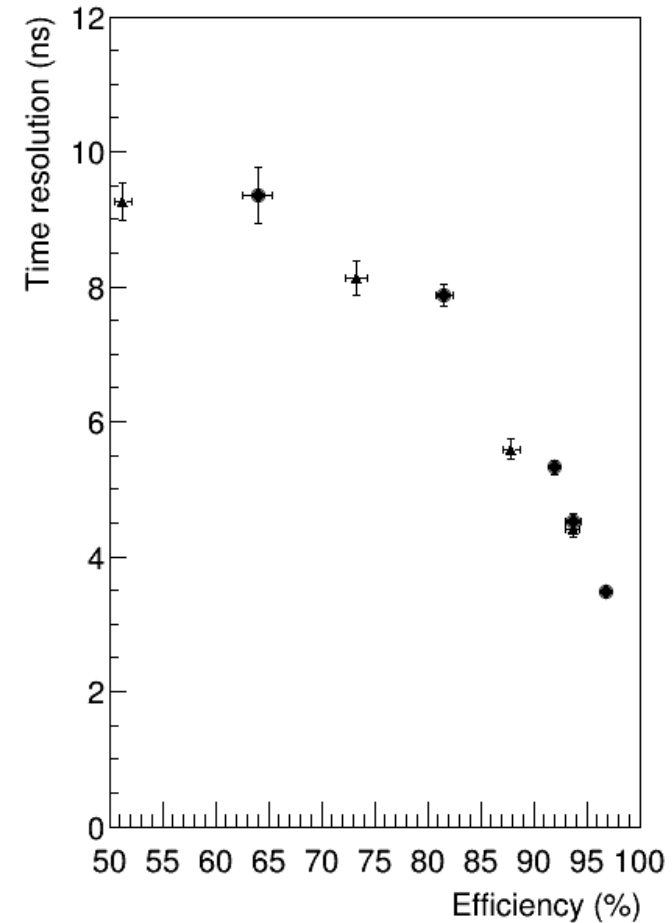
Efficiency



Multistreamer probability



Time resolution (ns)



Similar to “standard” mixtures. Even with small additions

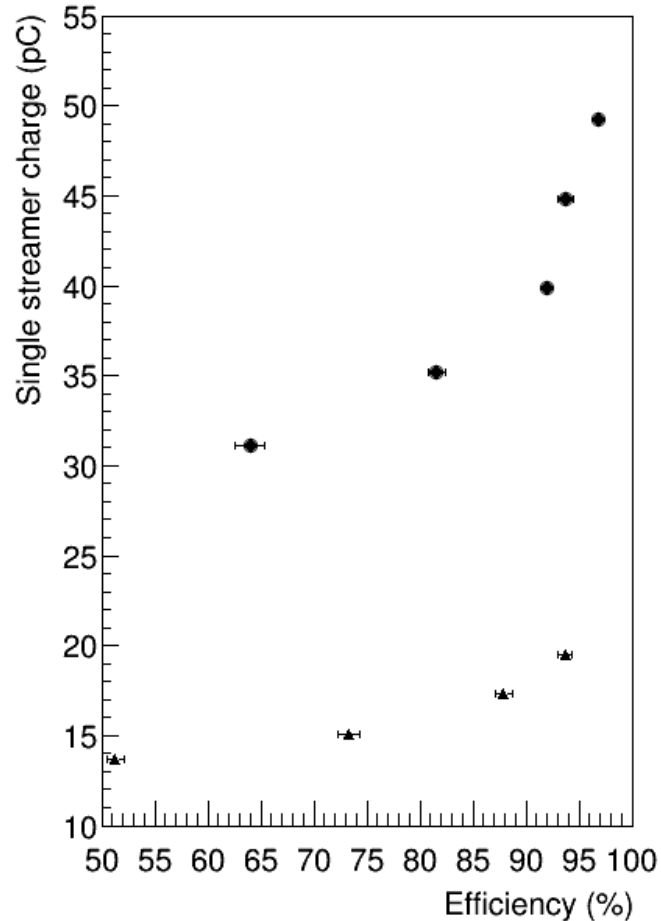
Strong charge suppression (both in signal amplitude and FWHM).

Observed also an increase of multistreamer probability.

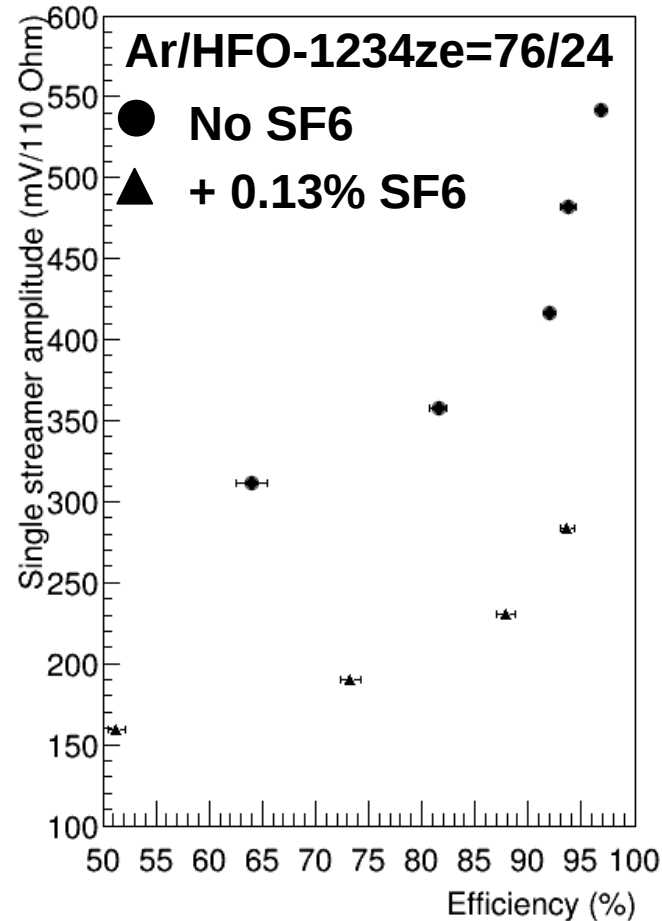
Tested also on Ar/HFOyf=74/26,70/30 Ar/HFOze=70/30 Ar/HFOyf/ibut=63/34/3.

SF₆ addition

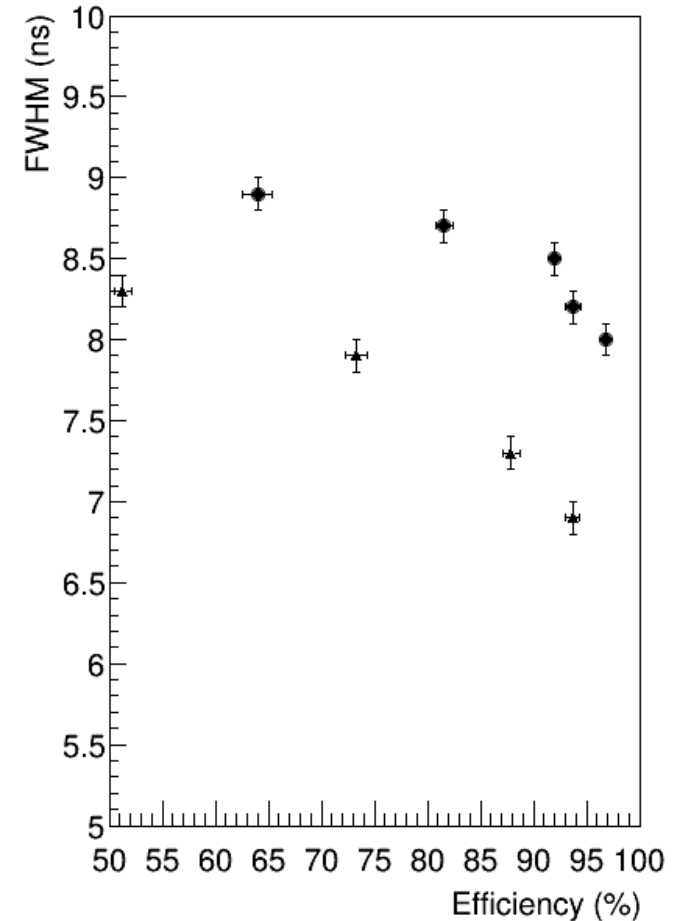
Single streamer charge (pC)



Single streamer amplitude (mV)



FWHM (ns)



Similar to “standard” mixtures. Even with small additions
Strong charge suppression (both in signal amplitude and FWHM).

SF₆ addition

SF6 is used in avalanche operated RPCs as a streamer suppressor.

In streamer mode it suppresses the streamer charge, making stable economic mixtures with low concentrations of quenching gases.

The streamer charge suppression with SF6 is believed to be due to its high attachment coefficient at small electron velocities.