

More on EDM correlations
in SUSY

Oleg Lebedev

CERN

1. S. Abel + OL,
JHEP 0601, 133 (2006)
2. —" —" —" —" —"
WG3 report

Motivation

EDMs of fundamental particles have not been observed,
BUT...

- ① remarkable sensitivity to NEW PHYSICS

$$d_e \approx \left(\frac{300 \text{ GeV}}{M_{\text{susy}}} \right)^2 \sin \Phi_{CP} \times 10^{-25} \text{ e.cm}$$

=> M_{susy} up to 100 TeV
for $d_e \sim 10^{-30} \text{ e.cm}$

- ② experimental progress

$$\left. \begin{array}{l} d_e \rightarrow 10^{-30} \text{ e.cm} \\ d_n \rightarrow 10^{-28} \text{ e.cm} \end{array} \right\} \text{ in a few years}$$

- ③ complementary to collider data

E.g. $\Phi_{CP} \sim 10^{-5}$

- ④ probe fundamental sources of \cancel{CP} , possibly baryogenesis, ...

Relativistic EDMs

$$H_{\text{non-rel.}} = -d \vec{S} \cdot \vec{E}$$



$$\mathcal{L}_{\text{rel.}} = -\frac{i}{2} d \bar{\Psi} (F\sigma) \gamma_5 \Psi$$

$\begin{cases} F_{\mu\nu} = \text{photon field strength} \\ \Psi = \text{fermion} \end{cases}$

For composite objects (n, atoms, ...) also relevant

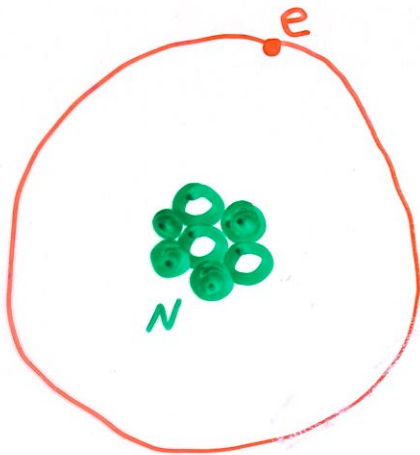
$$\mathcal{L}_{\text{eff}} = \frac{g^2}{32\pi^2} \Theta G_{\mu\nu} \tilde{G}^{\mu\nu} + \frac{w}{3} f^{abc} G_a \tilde{G}_b G_c - \frac{i}{2} \tilde{d} \bar{\Psi} g (G\sigma) \gamma_5 \Psi + \sum_{ij} C_{ij} (\bar{\Psi}_i \Psi_i) (\bar{\Psi}_j i\gamma_5 \Psi_j)$$

$\begin{cases} G_{\mu\nu} = \text{gluon field strength} \\ \Theta = \text{QCD } \Theta\text{-term} \\ w = \text{Weinberg operator} \\ \tilde{d} = \text{colour EDM} \end{cases}$

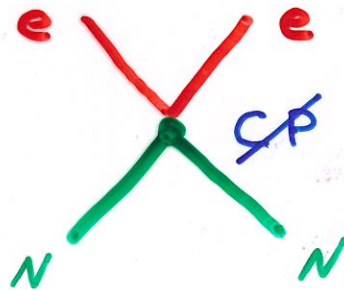
$$EDM = f(\theta, w, d, \tilde{d}, C_{ij})$$

Example:

Tl atom



$$\mathcal{L} \sim \bar{e} i \gamma_5 e \bar{N} N$$



Deforms the atom \rightarrow EDM

$$d_{Tl} \approx -585 d_e - C_s \cdot 43 e \cdot \text{GeV}$$

$$\downarrow \mathcal{L} \sim C_s \bar{e} i \gamma_5 e \bar{N} N$$

C_s can be more important than d_e (e.g. heavy superpartners)

Model - dependence

d_n	Model	Ref.
2.7	MIT bag model	Baluni
3.6	Current algebra	Crewther, ...
3.3	Effective chiral approach	Pich, ...
6.7	HB ch PT	Borasoy
3.0	Chiral bag model	Musakhanov, ...
1.4	Cloudy bag model	Morgan, ...
1.2	Chiral quark-meson model	Mc Govern, ...
2.4	QCD sum rules	Pospelov, ...
1.4	perturbative chiral model	Kuckei, ...

(d_n in units of 10×10^{-16} e.cm)

EDMs in SUSY

Induced by new CP phases:

$$\Delta \tilde{\mathcal{L}} = \mu \bar{\Psi}_{H_1} \Psi_{H_2} + B \mu H_1 H_2 + \text{h.c.}$$

$$+ \frac{1}{2} (m_3 \bar{\lambda}_3 \lambda_3 + m_2 \bar{\lambda}_2 \lambda_2 + m_1 \bar{\lambda}_1 \lambda_1)$$

$$+ A_{ij}^d H_1 \tilde{q}_{L_i} \tilde{q}_{R_j}^* + \dots$$

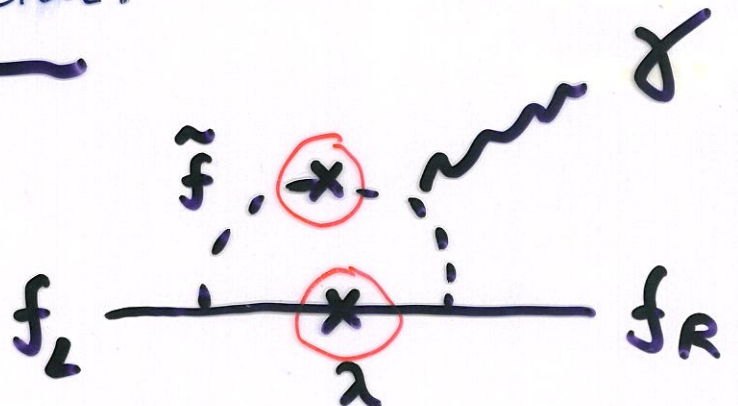
(... = complex

2 phases eliminated by $U(1)_R, U(1)_{PQ}$

Physical phases: $\text{Arg}(m_i^* A), \text{Arg}(B^* A), \dots$

Typical EDM contribution:

Ellis, Ferrara, Nanopoulos '82



$g_{\text{SUSY}} \sim 10^{-2} \Leftarrow$
(CP problem)

The CP problem appears already in the most minimalistic models, e.g.

mSUGRA :

$\mu, A = \text{complex!}$

This is due to holomorphicity of SUSY.

(Unlike the FCNC problem)



EDMs are special observables, probe even minimal scenarios

① Small phases

E.g. phase alignment

$$y_M = y_A = -y_\mu \quad (+ \text{ corrections})$$



$$y(M^* A) = y(\mu A) = \dots = 0 //$$

(dilaton domination + Giudice - Masiero)

② Decoupling

Nath '91
Kizukuri,
Oshimo '92



$$\sim \frac{1}{m^2_{\tilde{q}_{1,2}}} \xrightarrow{\sim 10 \text{ TeV}} 0$$

→ EDMs at 2 loops

③ off diagonal CP

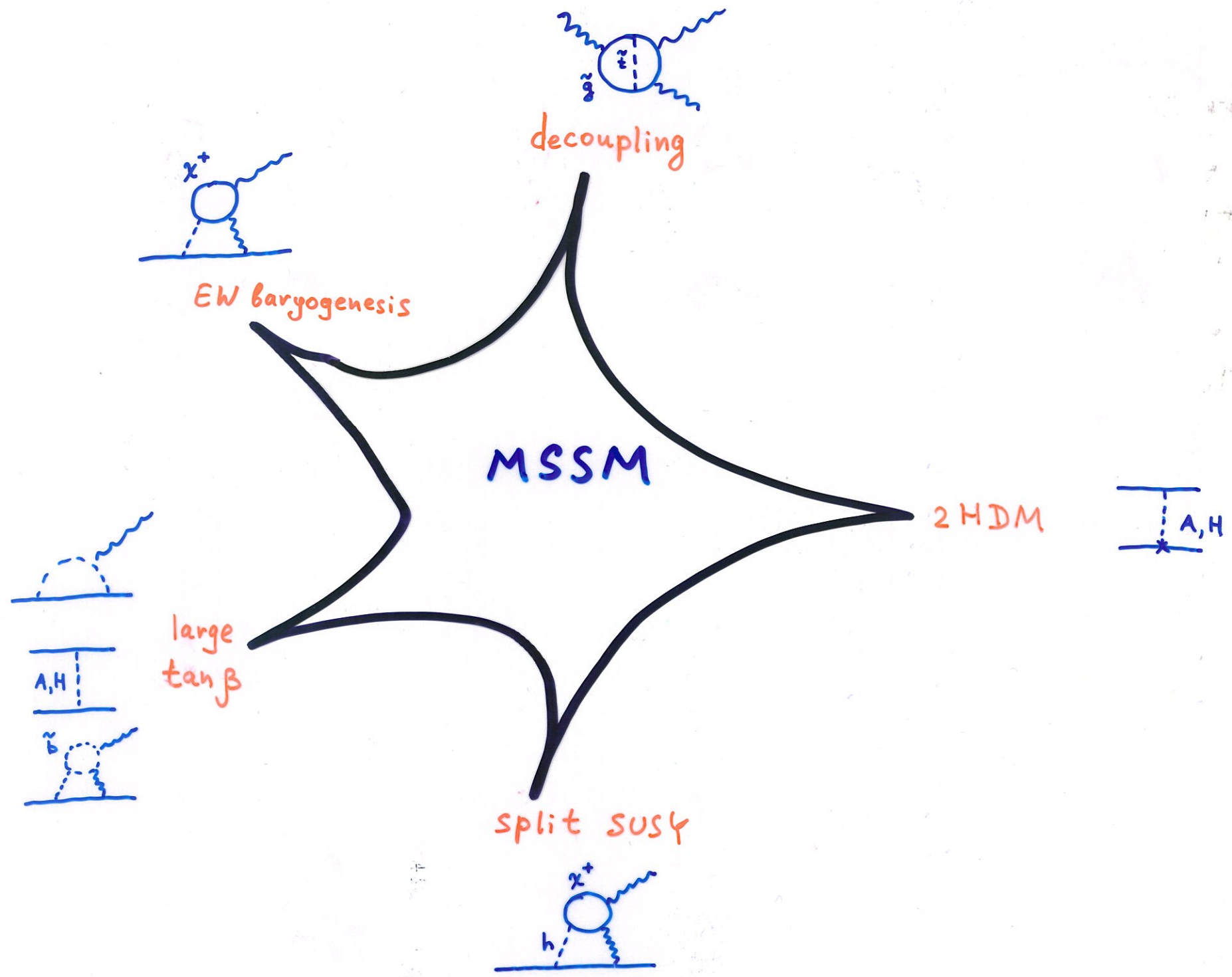
$$\left. \begin{aligned} A &= A^\dagger \\ \mu &= \mu^\dagger \\ \dots \end{aligned} \right\}$$

by symmetry → LR
→ flavour

Mohapatra,
Senjanovich

Abel, Bailin,
Khalil, OL

Like in the SM, small EDMs.



EDM correlations

(10)

SUSY predicts certain correlations between

$$\underline{d_n}, d_e, (d_{Hg}, d_\mu, d_D, \dots)$$



indirect signature of SUSY

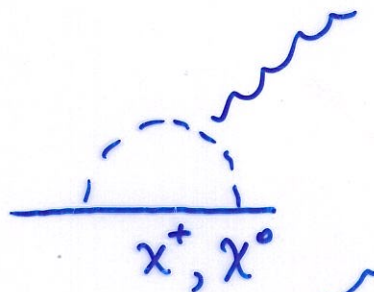
Typically expect:

$$d_n \sim 10 d_e$$

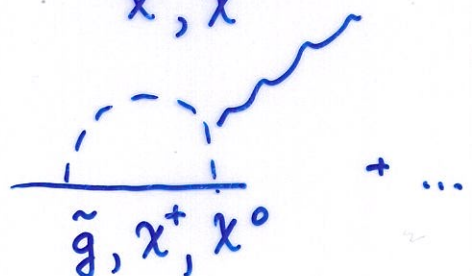
$$\left(\frac{m_q}{m_e} \sim 10 \right)$$

even though

$$d_e \sim$$



$$d_n \sim$$

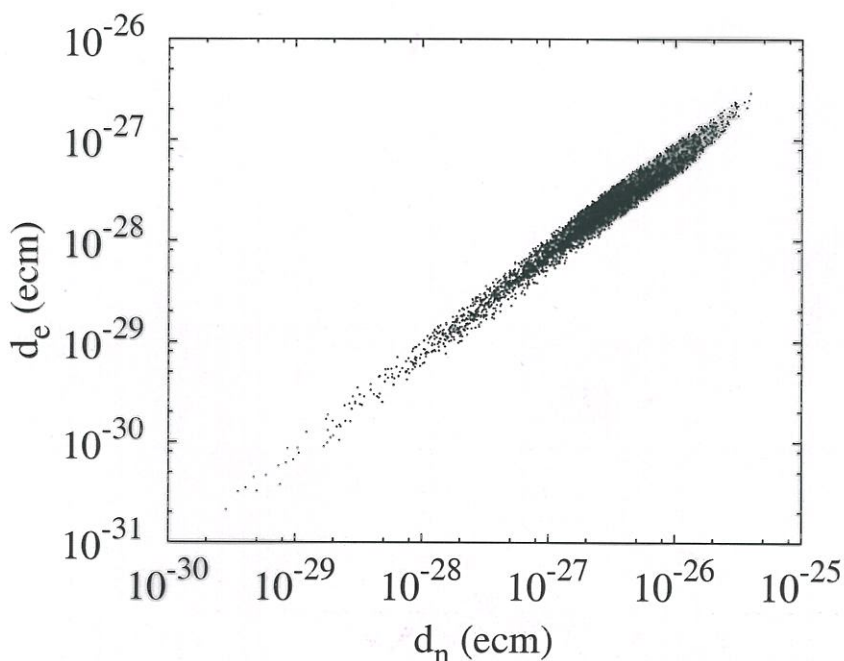


$d_e - d_n$ correlation in mSUGRA

$$m_0, m_{1/2}, |A| \in [200 \text{ GeV}, 1 \text{ TeV}]$$

$$\varphi_\mu \in [-\pi/500, \pi/500]$$

$$\tan \beta = 5$$



$$\underline{d_e \sim 10^{-1} d_n}$$

⇓
indirect
evidence
for
SUSY!

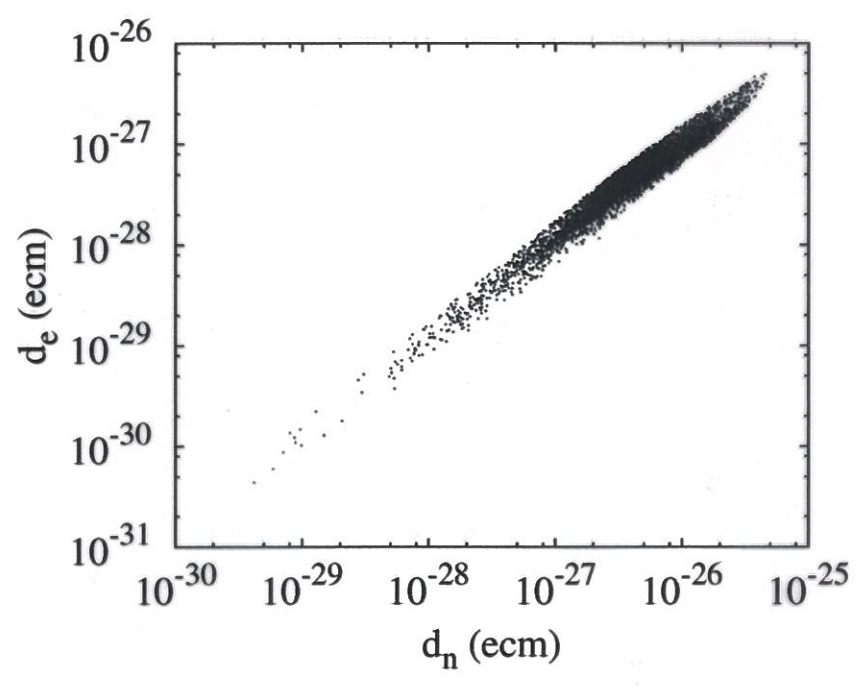
Note: persists for $\tan \beta \sim 35$; also for $m_{\tilde{g}} \neq m_{\tilde{z}}$ (GUT)

Heavy m SUGRA

$$m_0, m_{1/2}, |A| \in [2 \text{ TeV}, 10 \text{ TeV}]$$

$$\Phi_A, \Phi_\mu \in [-\pi, \pi]$$

A very similar picture:



Most points are in the observable range!

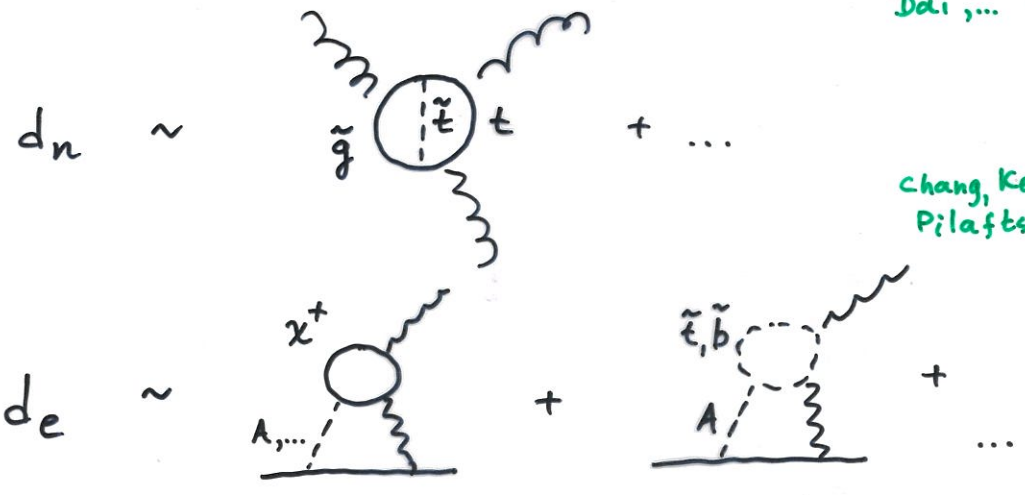
Decoupling

$$M_{\tilde{q}_3}, m_{1/2}, |A| \in [200 \text{ GeV}, 1 \text{ TeV}]$$

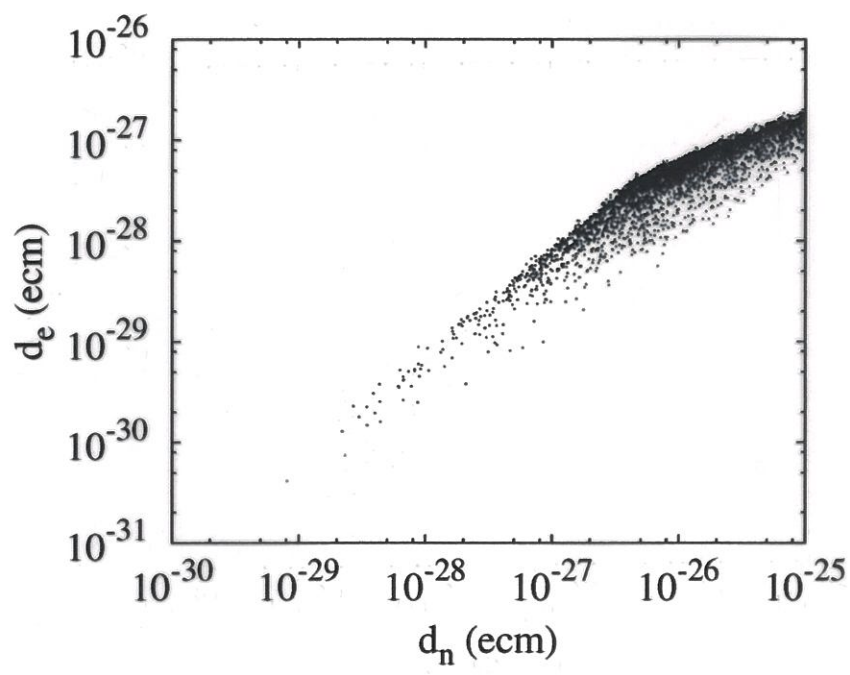
$$y_A, y_\mu \in [-\pi, \pi]$$

$$m_{\tilde{q}_{1,2}} \rightarrow \infty$$

Weinberg '89
Dai, ... '90



Chang, Keung,
Pilaftsis '99



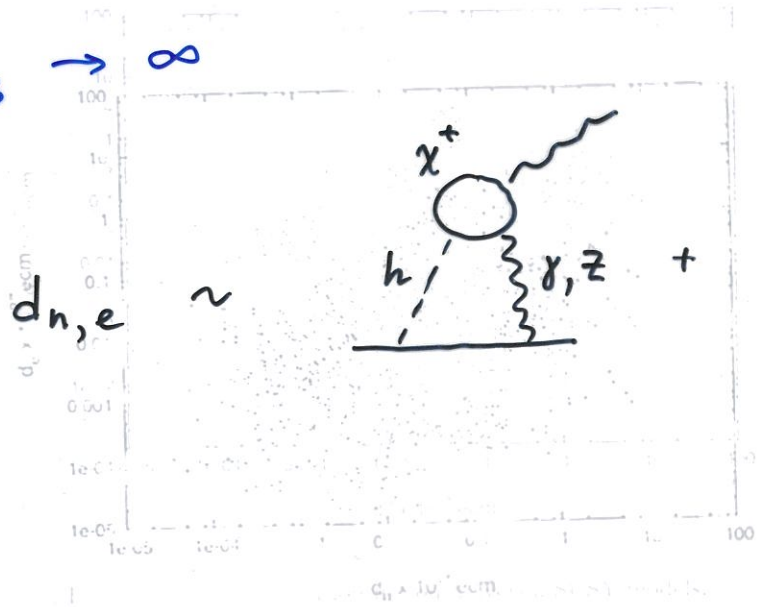
$d_n \sim 10 \div 100 d_e$

Should be
observed
soon!

Split SUSY

$M_{1,2,3} \in [200 \text{ GeV}, 1 \text{ TeV}]$
 $|\mu| \in [200 \text{ GeV}, 1 \text{ TeV}]$
 $m_h \in [100 \text{ GeV}, 300 \text{ GeV}]$

$M_{\text{scalars}} \rightarrow \infty$



Chang, ... '02
Giudice,
Romanino '05
...

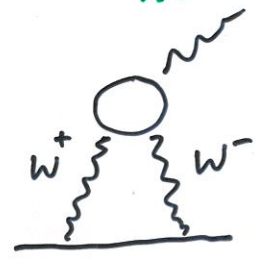
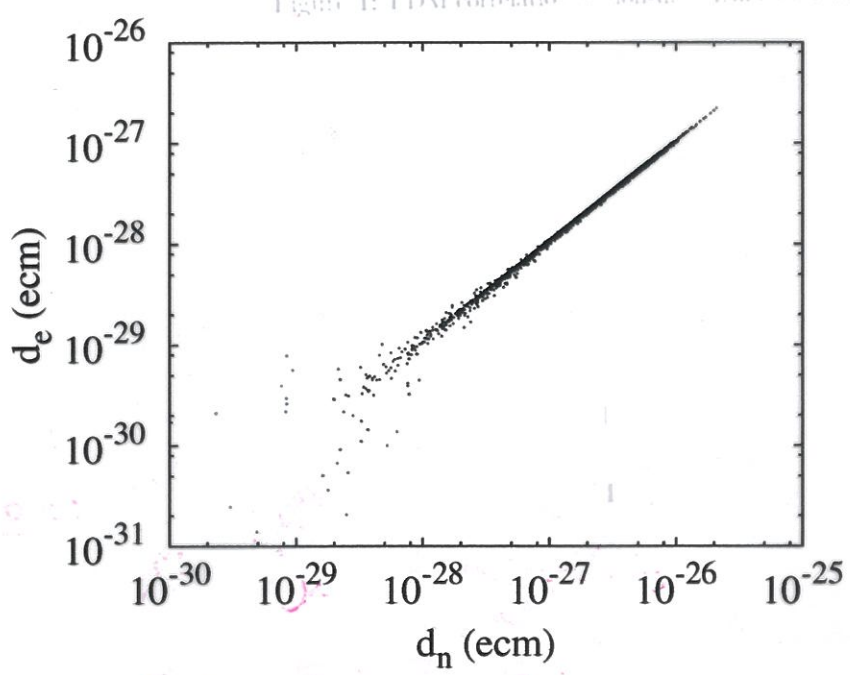


Figure 1: EDM correlator for non-universal SUSY models.



$d_n \sim 10 d_e$

Observable!

Non-universal SUSY

$m_{\text{squark}} \neq m_{\text{slepton}}$

$M_{1,2} \neq M_3$

phases : $\varphi_\mu, \varphi_A, \varphi_{M_3}$

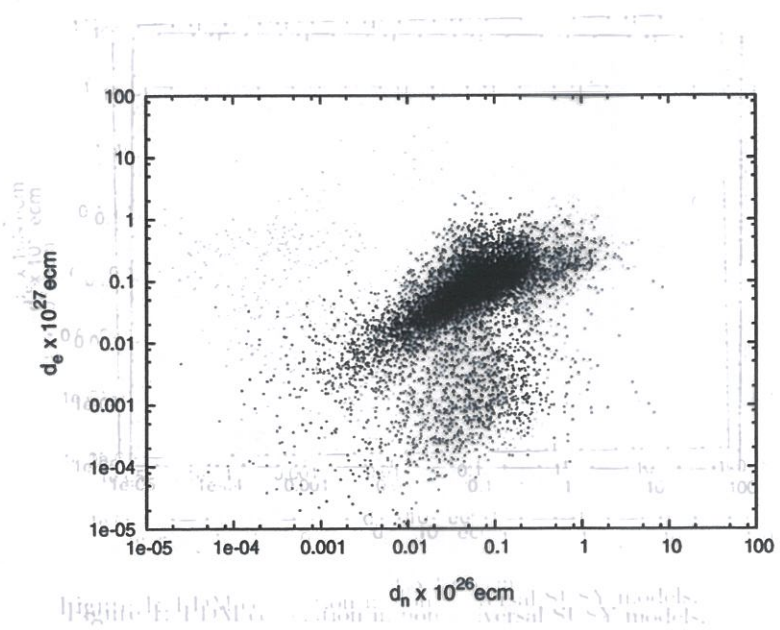


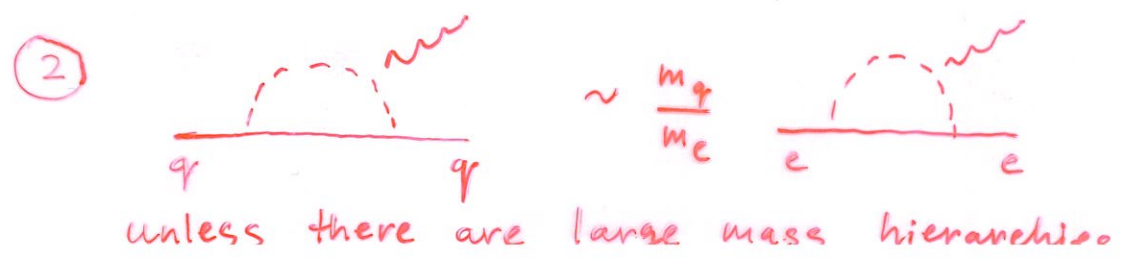
Figure 1: EDM correlation in non-universal SUSY models.

90% : $d_n/d_e \sim 10-100$

10% : $d_n/d_e \sim 10^2-10^4$

Reason:

① φ_μ dominates



Non-universal SUSY

Random scan :

all mass parameters $\in [300 \text{ GeV}, 5 \text{ TeV}]$
complete non-universality
($m_{\tilde{q}_3} \neq m_{\tilde{q}_{1,2}}, \dots$)

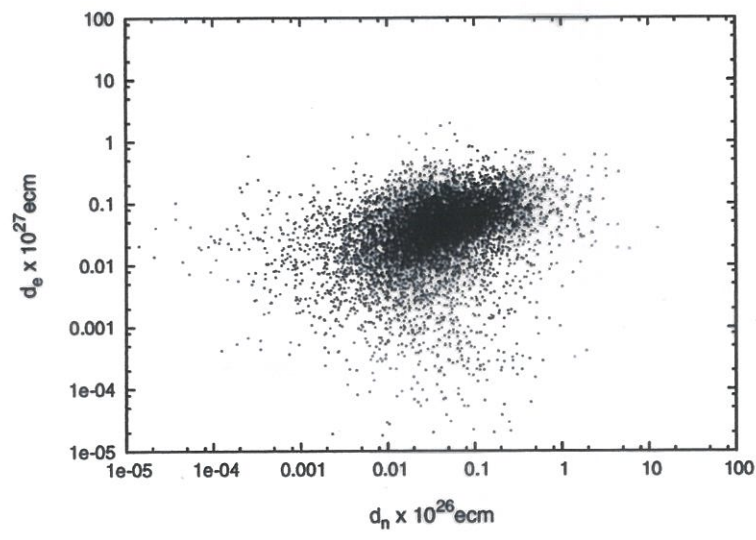


Figure 1: EDM correlation in non-universal SUSY models.

Message:

typically

$d_n \sim 10 \div 100 d_e$

(rather insensitive to $\tan \beta$,
universality assumptions,
...)

Thus,

$d_e \gtrsim d_n$
 $d_e \lll d_n$

\Rightarrow SUSY
disfavored
(at least common models)

Also,

complementary to collider data:

$M_{\text{susy}} \lesssim \text{TeV}, g_{\text{CP}} \sim 10^{-5}$
 $M_{\text{susy}} \gtrsim 5 \text{TeV}, g_{\text{CP}} \sim 1$ } out of colliders' reach

Conclusions

- if there's low energy SUSY, EDMs should be observed in the next ~ 5 years
($d_n > 10^{-28}$ e.cm, $d_e > 10^{-30}$ e.cm)
- SUSY predicts d_e - d_n correlations
(typically $d_n \sim 10 d_e$)
- EDM measurements are complementary to collider data
(e.g. $\mathcal{P}_{CP} \sim 10^{-5}$)