

DAMPE: first data from space

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Purple Mountain Observatory
(on behalf of DAMPE collaboration)

The physics goals

High energy particle detection in space

- Study of the cosmic electron and photon spectra
- Study of cosmic ray protons and nuclei: spectrum and composition
- High energy gamma ray astronomy
- Search for dark matter signatures in e/gamma spectra

Detection of
GeV - 10 TeV e/ γ
50 GeV - 500 TeV protons and nuclei
with excellent energy resolution , tracking precision
and particle identification capabilities

- Exotica and “unexpected” , e.g. GW e.m. counterpart in the FoV

The collaboration

- **CHINA**

- Purple Mountain Observatory, CAS, Nanjing
- Institute of High Energy Physics, CAS, Beijing
- National Space Science Center, CAS, Beijing
- University of Science and Technology of China, Hefei
- Institute of Modern Physics, CAS, Lanzhou



- **ITALY**

- INFN Perugia and University of Perugia
- INFN Bari and University of Bari
- INFN Lecce and University of Salento



- **SWITZERLAND**

- University of Geneva



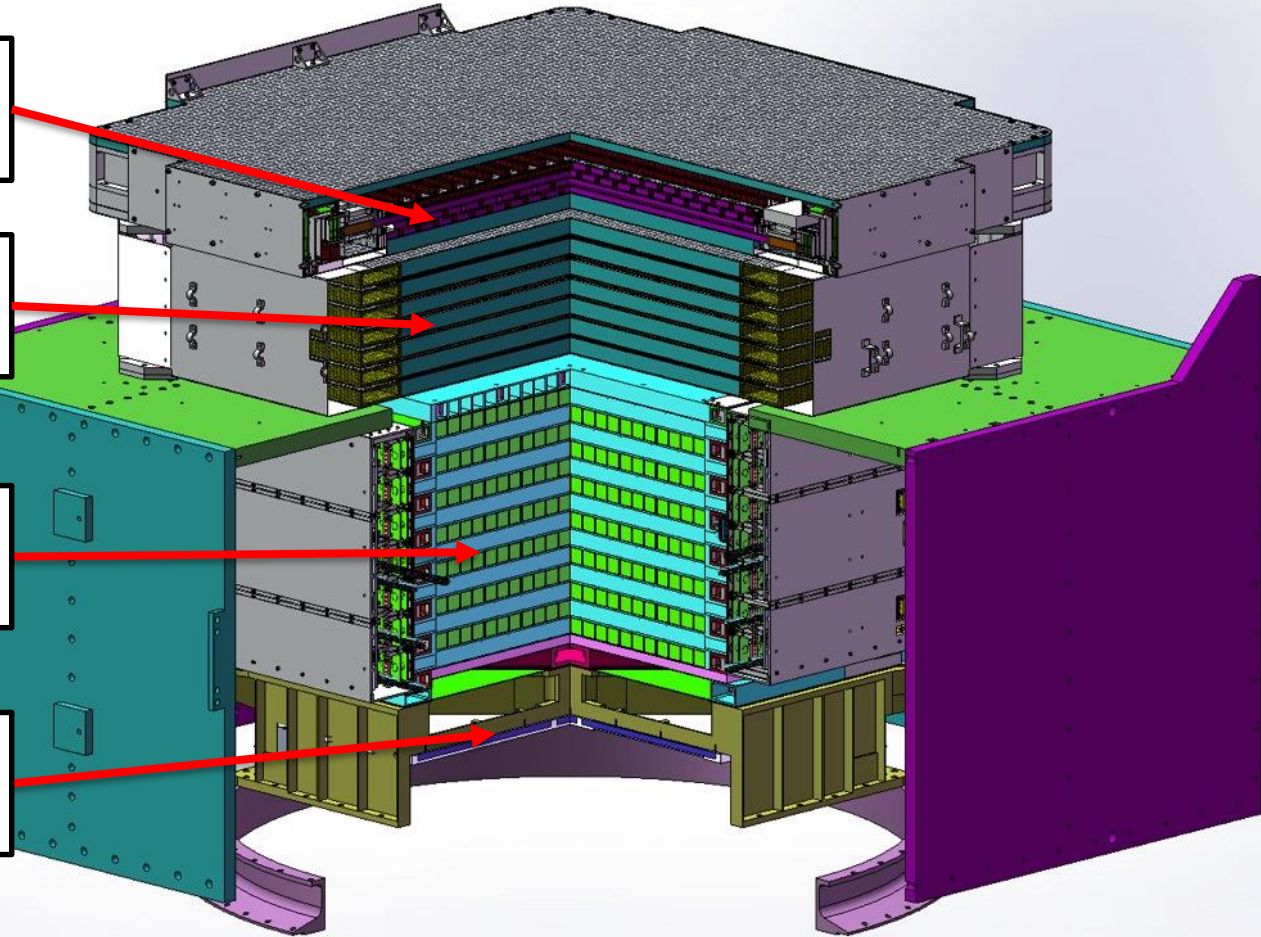
The detector

Plastic Scintillator Detector
(PSD)

Silicon-Tungsten Tracker
(STK)

BGO Calorimeter
(CALO)

Neutron Detector
(NUD)

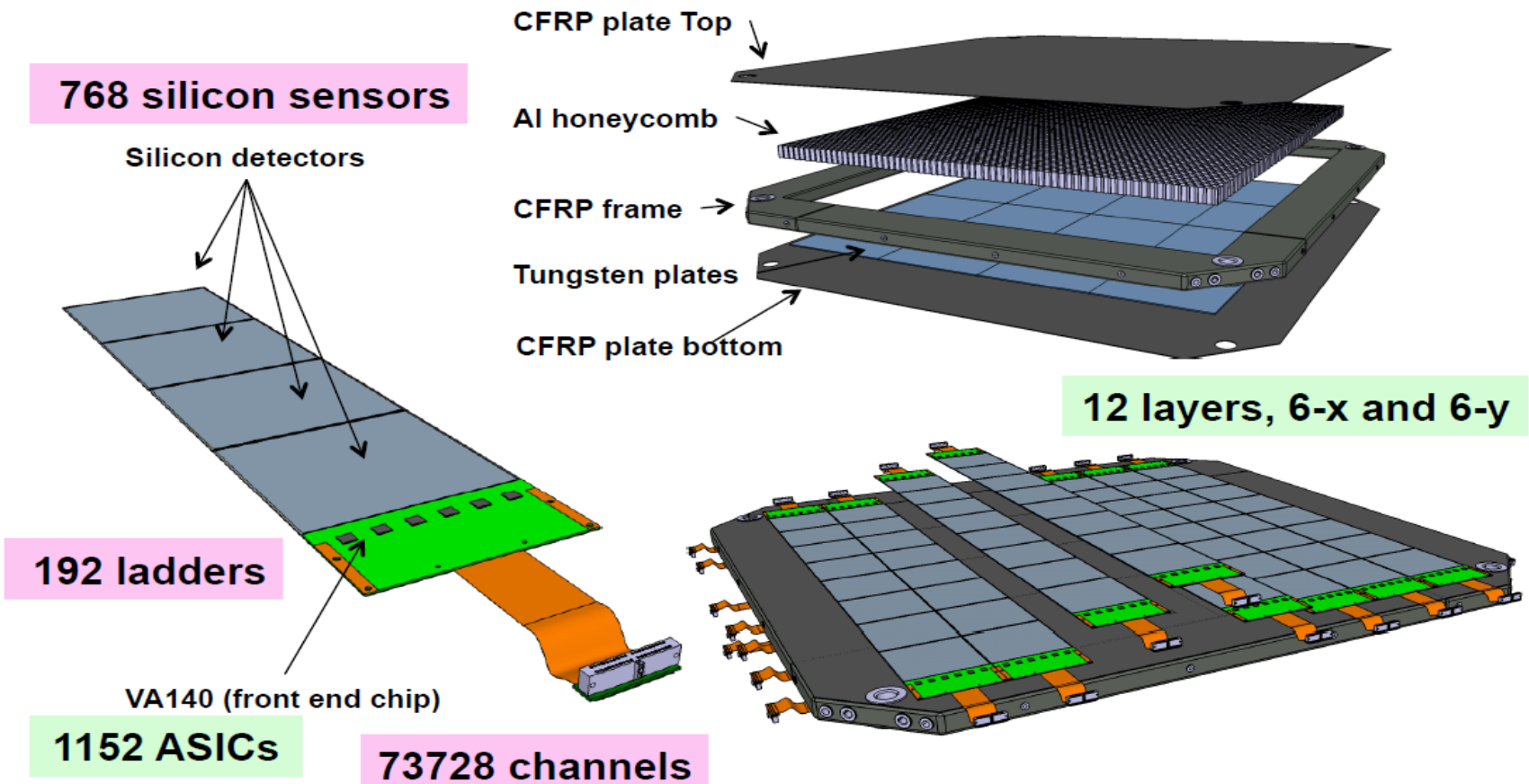


- Charge measurement (dE/dx in PSD , STK and BGO)
- Tungsten converter (pair production)
- Precise tracking (silicon strips)
- Thick calorimeter (BGO bars)
- Hadron rejection (neutron detector)



high energy
 γ -ray, electron and cosmic ray
telescope

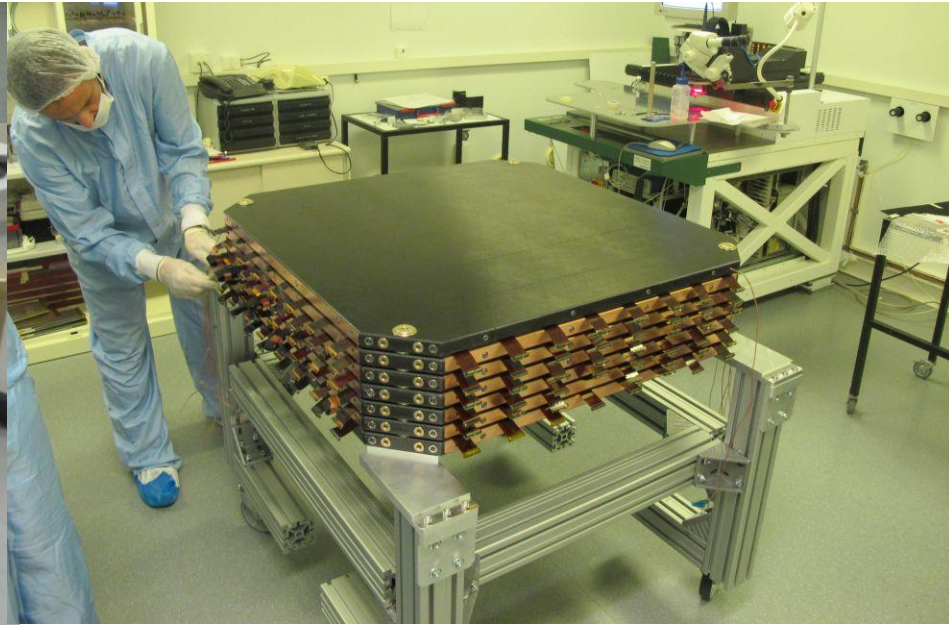
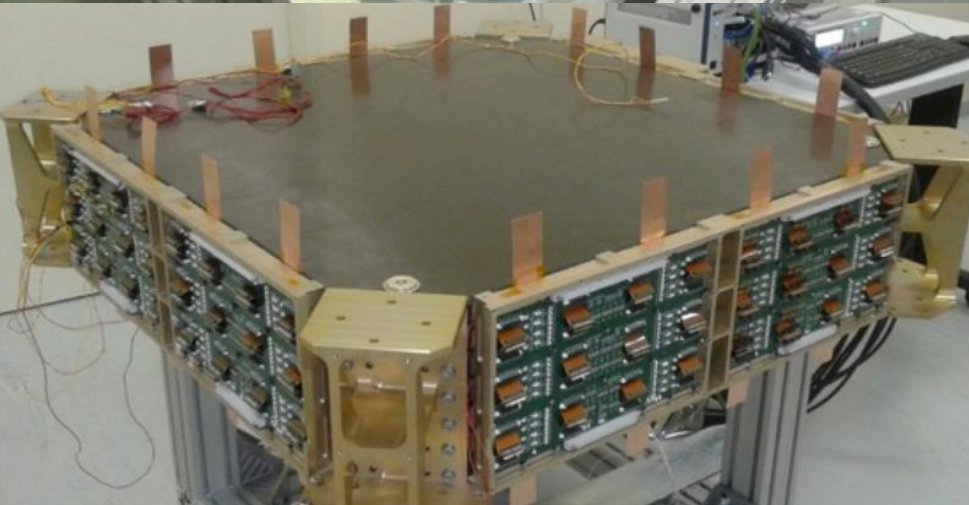
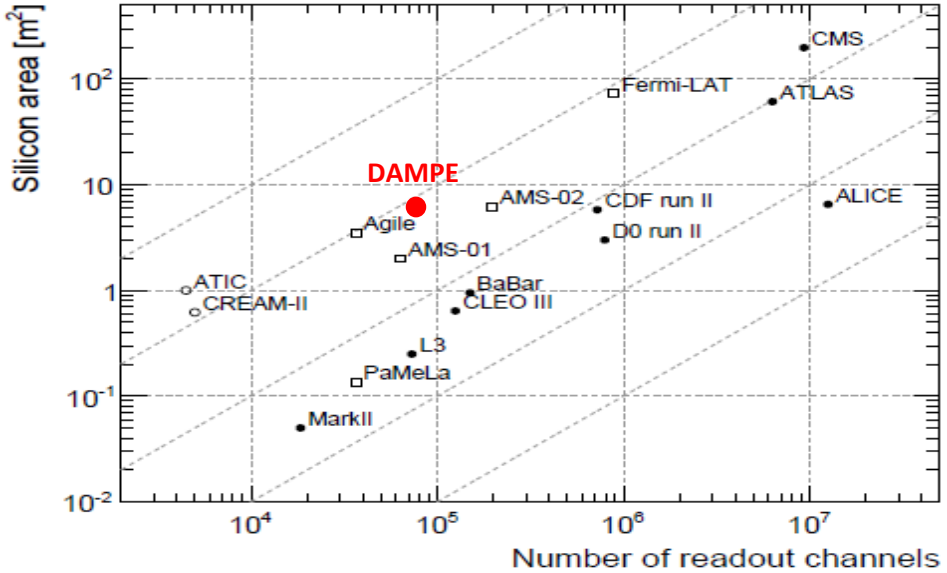
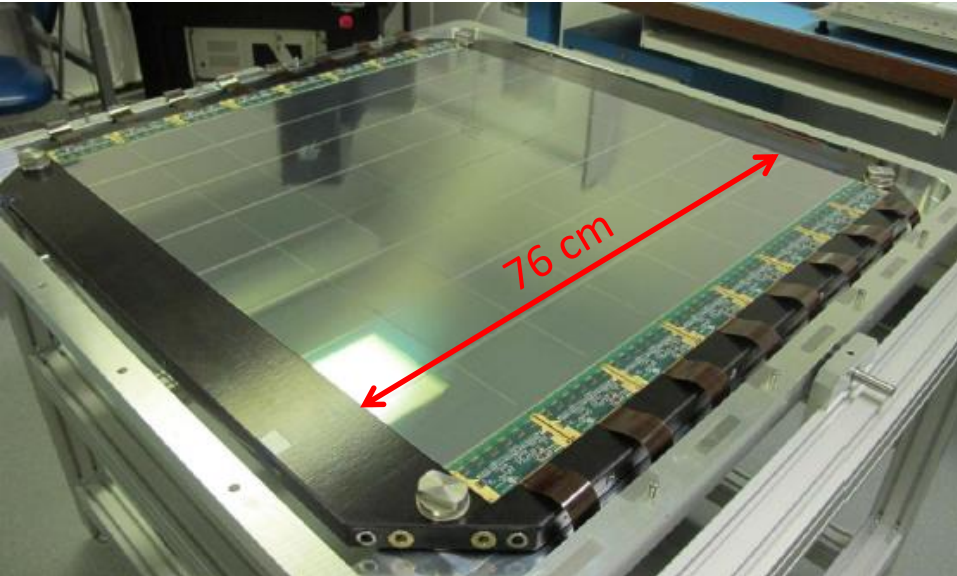
The Silicon Tracker (STK)



- 48 μm wide Si strips with 121 μm pitch
- (95 × 95 × 0.32 mm³) Silicon Strip Detectors (SSD) with 768 strips
- One ladder composed by 4 Silicon Strip Detectors (SSD)
- 16 Ladders per layer (76 cm × 76 cm)
- 12 layers (6x + 6y)

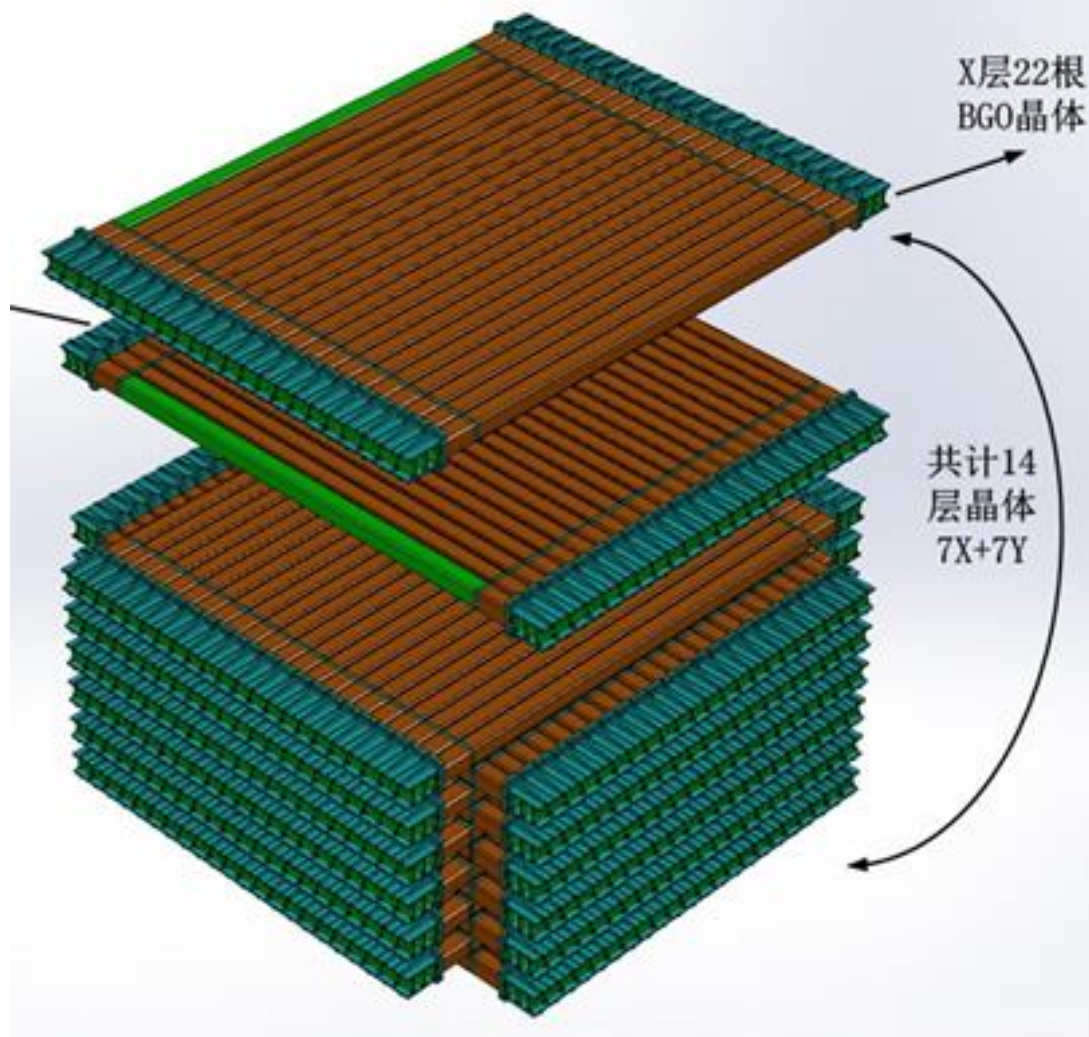
Analog Readout of each second strip:
384 channels / SSD- Ladder
Charge sharing

The Silicon Tracker (STK) - 2

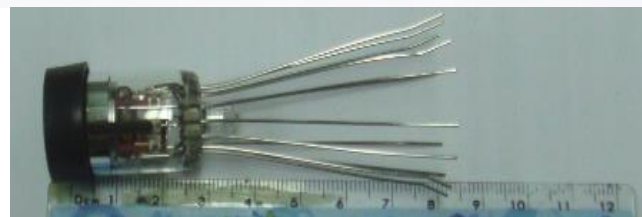


The CALOrimeter

- 14 layers of 22 BGO bars
 - Dimension of BGO bar: $2.5 \times 2.5 \times 60 \text{ cm}^3$
 - 14 hodoscopic stacking alternating orthogonal layers
 - depth $\sim 32X_0$
- Two PMTs coupled with each BGO crystal bar in two ends
- Electronics boards attached to each side of module



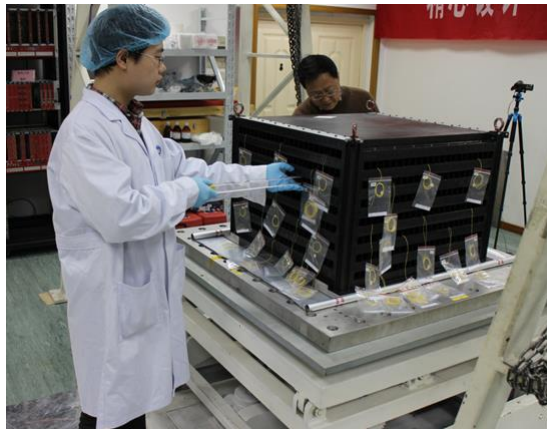
308 bars
616 PMTs



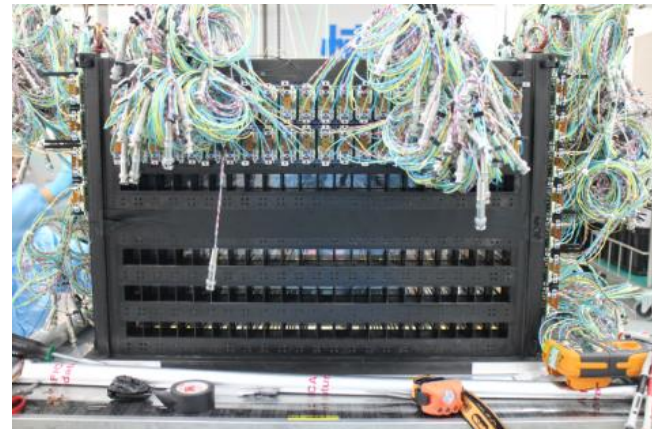
The CALOrimeter -2



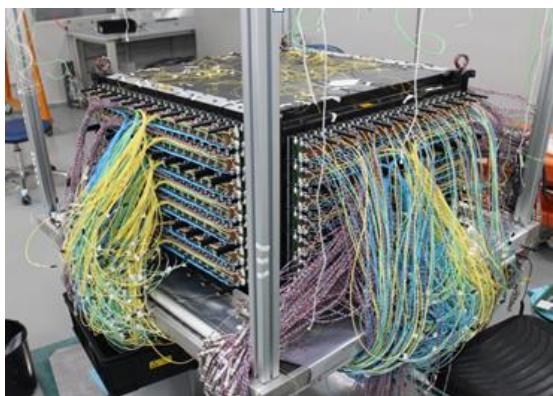
Carbon Fiber Structure



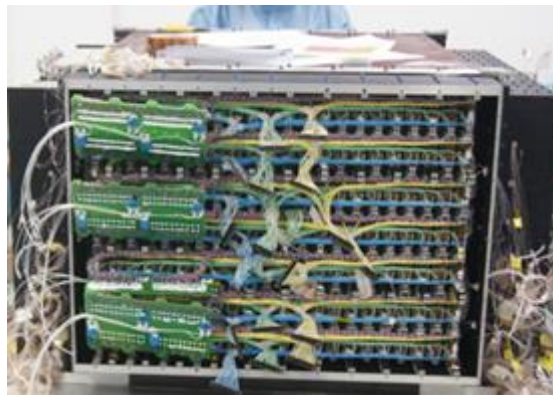
BGO crystal installation



PMT installation



Cable arranging



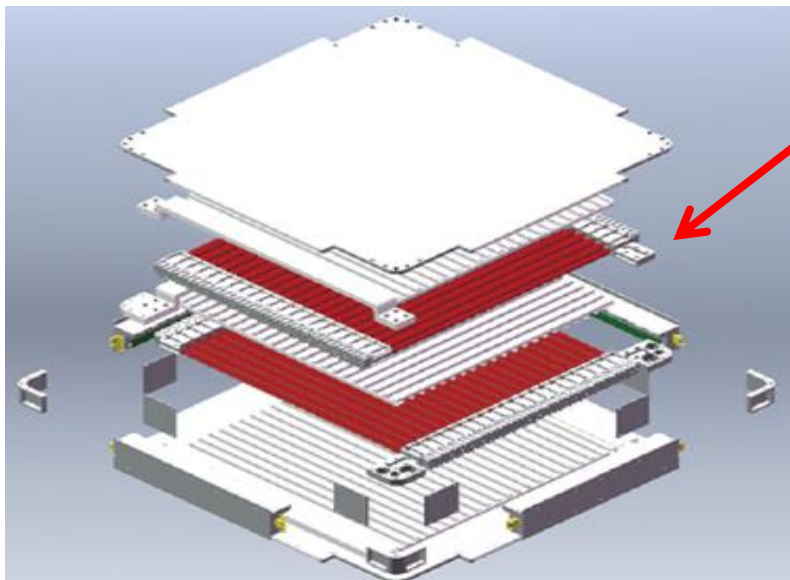
Cable connector



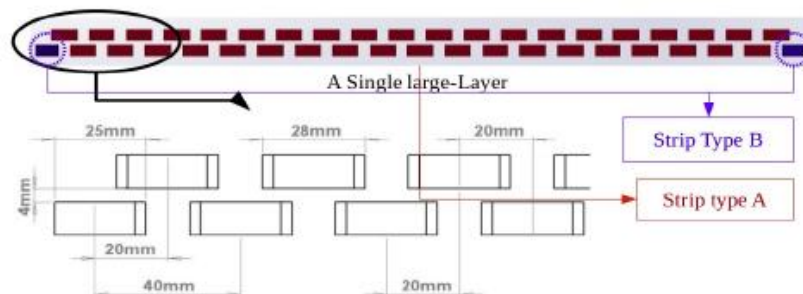
BGO Cal

The PSD and the NUD

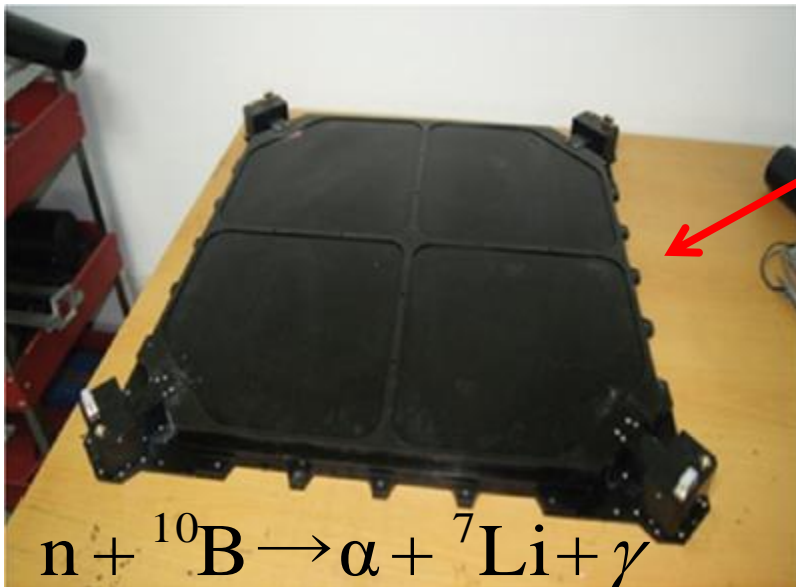
PSD



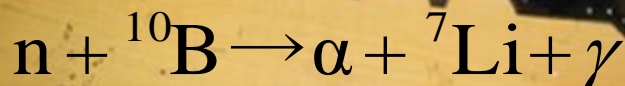
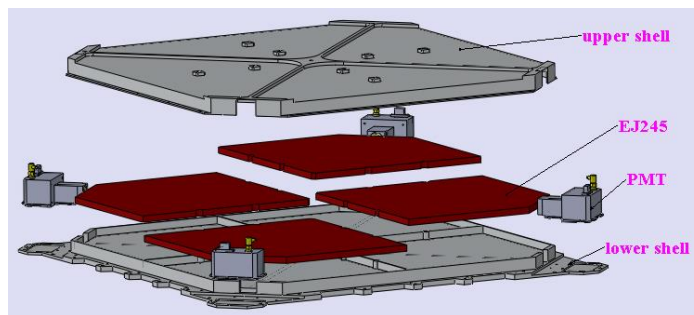
- 1.0 cm thick ,2.8cm wide and 82.0 cm long scintillator strips
- staggered by 0.8 cm in a layer
- 82 cm × 82 cm layers
- 2 layers (x and y)



NUD

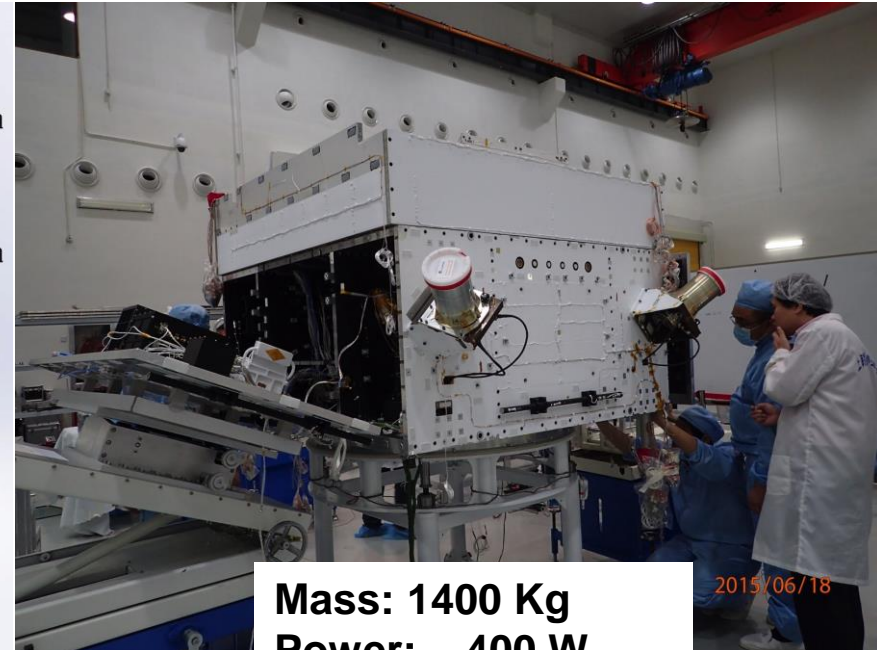
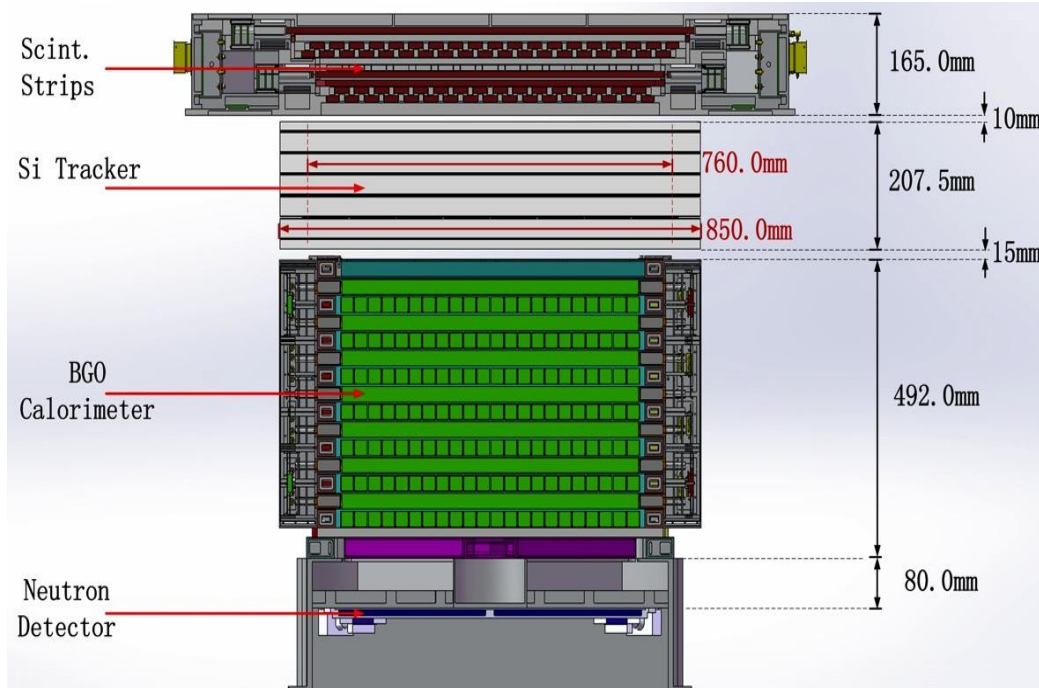


- 4 large area boron-doped plastic scintillators (30 cm × 30 cm × 1 cm)



Comparison with AMS-02 and FERMI

	DAMPE	AMS-02	Fermi LAT
e/ γ Energy res.@100 GeV (%)	1.5	3	10
e/ γ Angular res.@100 GeV ($^{\circ}$)	0.1	0.3	0.1
e/p discrimination	10^5	$10^5 - 10^6$	10^3
Calorimeter thickness (X_0)	32	17	8.6
Geometrical accep. (m^2sr)	0.3	0.09	1

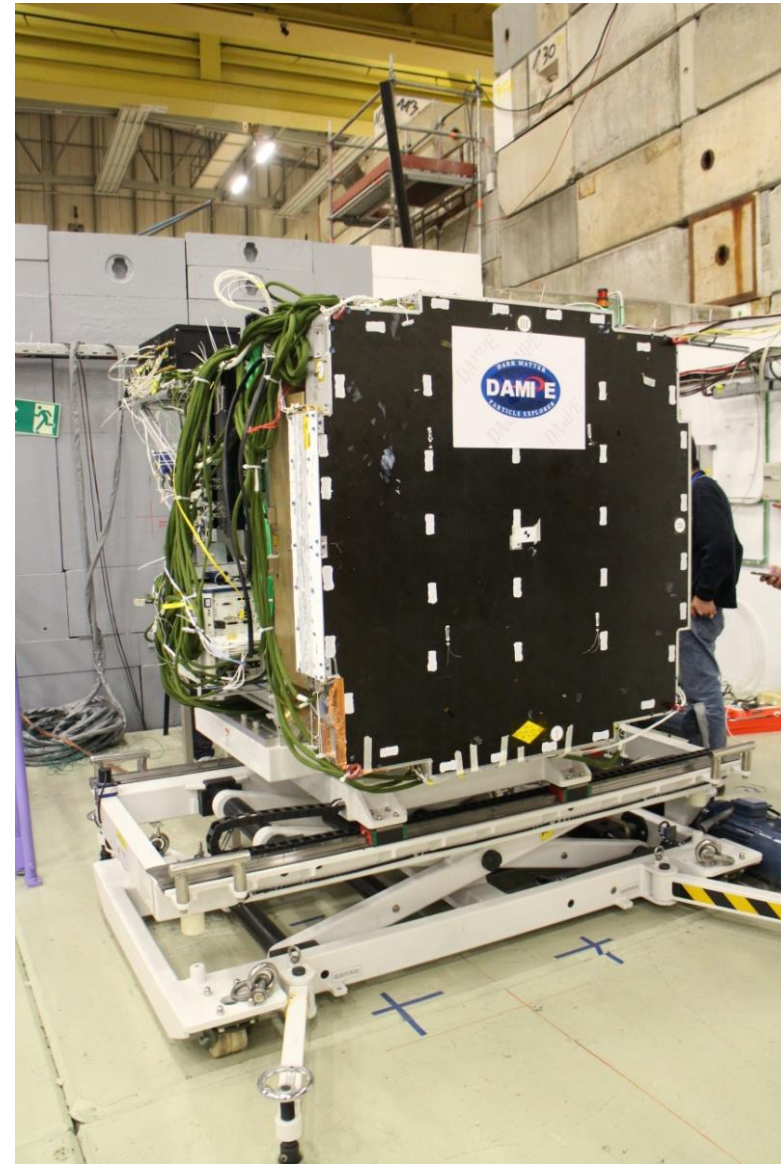


Mass: 1400 Kg
Power: ~ 400 W
Lifetime: > 3 years

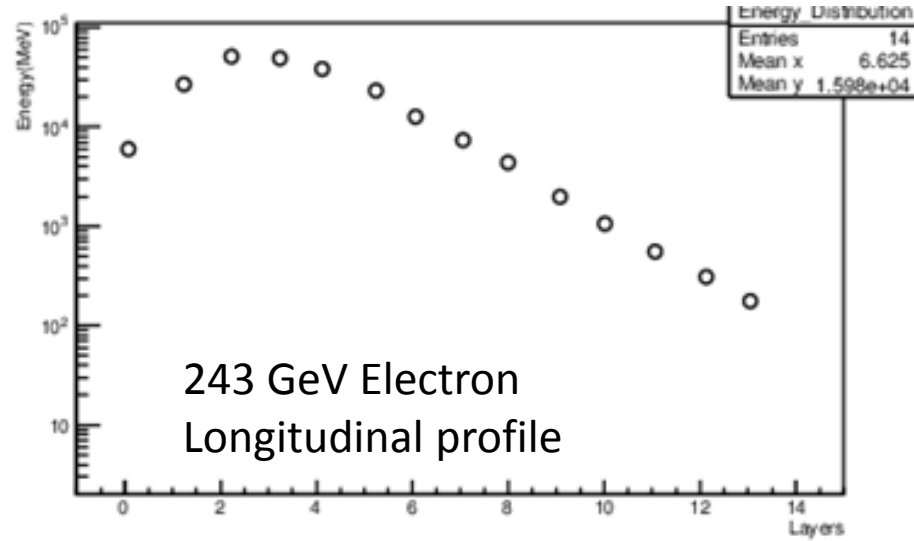
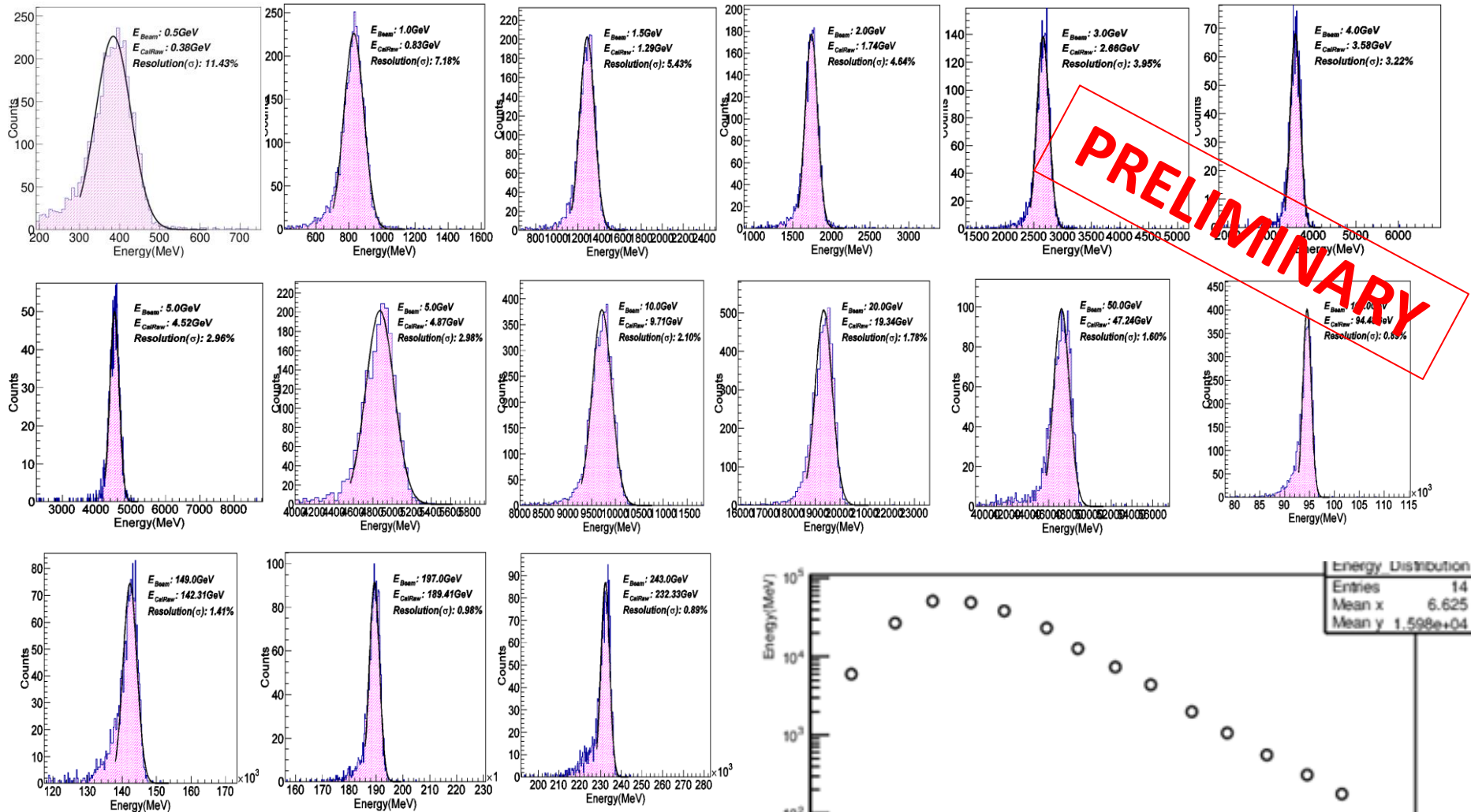
2015/06/18

Test beam activity at CERN

- **14days@PS, 29/10-11/11 2014**
 - e @ 0.5GeV/c, 1GeV/c, 2GeV/c, 3GeV/c, 4GeV/c, 5GeV/c
 - p @ 3.5GeV/c, 4GeV/c, 5GeV/c, 6GeV/c, 8GeV/c, 10GeV/c
 - π^- @ 3GeV/c, 10GeV/c
 - γ @ 0.5-3GeV/c
- **8days@SPS, 12/11-19/11 2014**
 - e @ 5GeV/c, 10GeV/c, 20GeV/c, 50GeV/c, 100GeV/c, 150GeV/c, 200GeV/c, 250GeV/c
 - p @ 400GeV/c (SPS primary beam)
 - γ @ 3-20GeV/c
 - μ @ 150GeV/c,
- **17days@SPS, 16/3-1/4 2015**
 - Fragments: 66.67-88.89-166.67GeV/c
 - Argon: 30A- 40A- 75AGeV/c
 - Proton: 30GeV/c, 40GeV/c
- **21days@SPS, 10/6-1/7 2015**
 - Primary Proton: 400GeV/c
 - Electrons @ 20, 100, 150 GeV/c
 - γ @ 50, 75 , 150 GeV/c
 - μ @ 150 GeV /c
 - π^+ @10, 20, 50, 100 GeV/c
- **10days@SPS, 11/11-20/11 2015**
 - Pb 30AGeV/c (and fragments) (HERD)
- **6days@SPS, 20/11-25/11 2015**
 - Pb 030 AGeV/c (and fragments)



Test beam activity at CERN: electrons

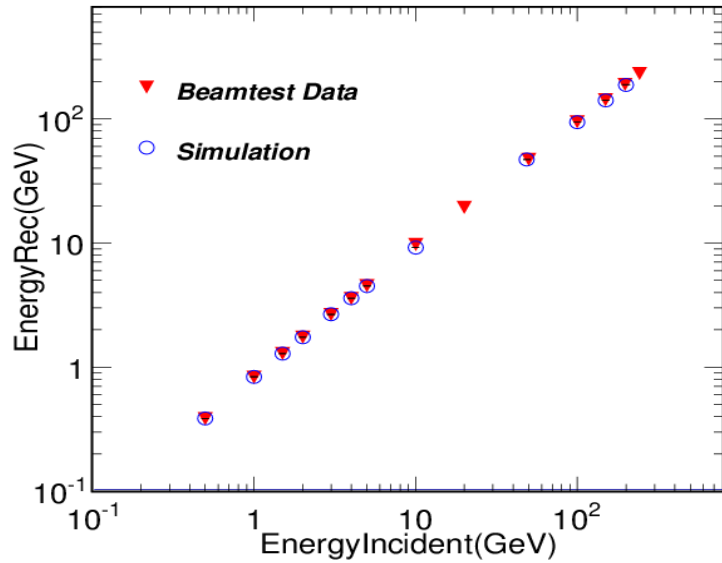


Electrons

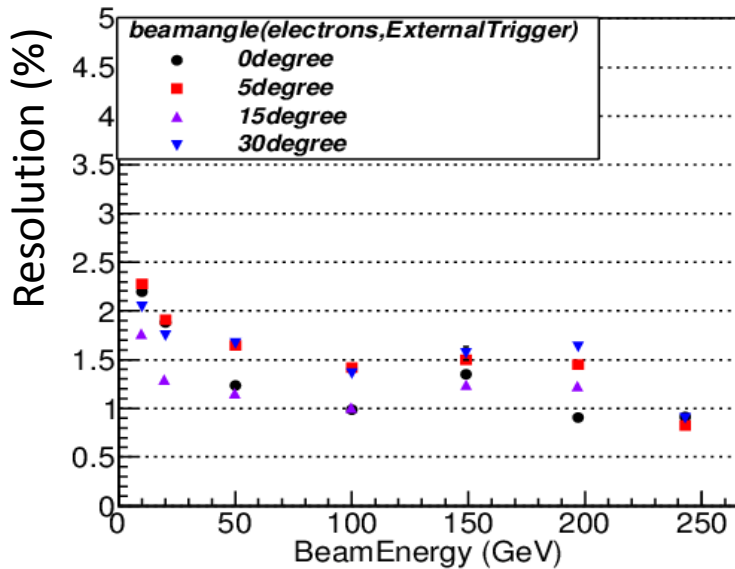
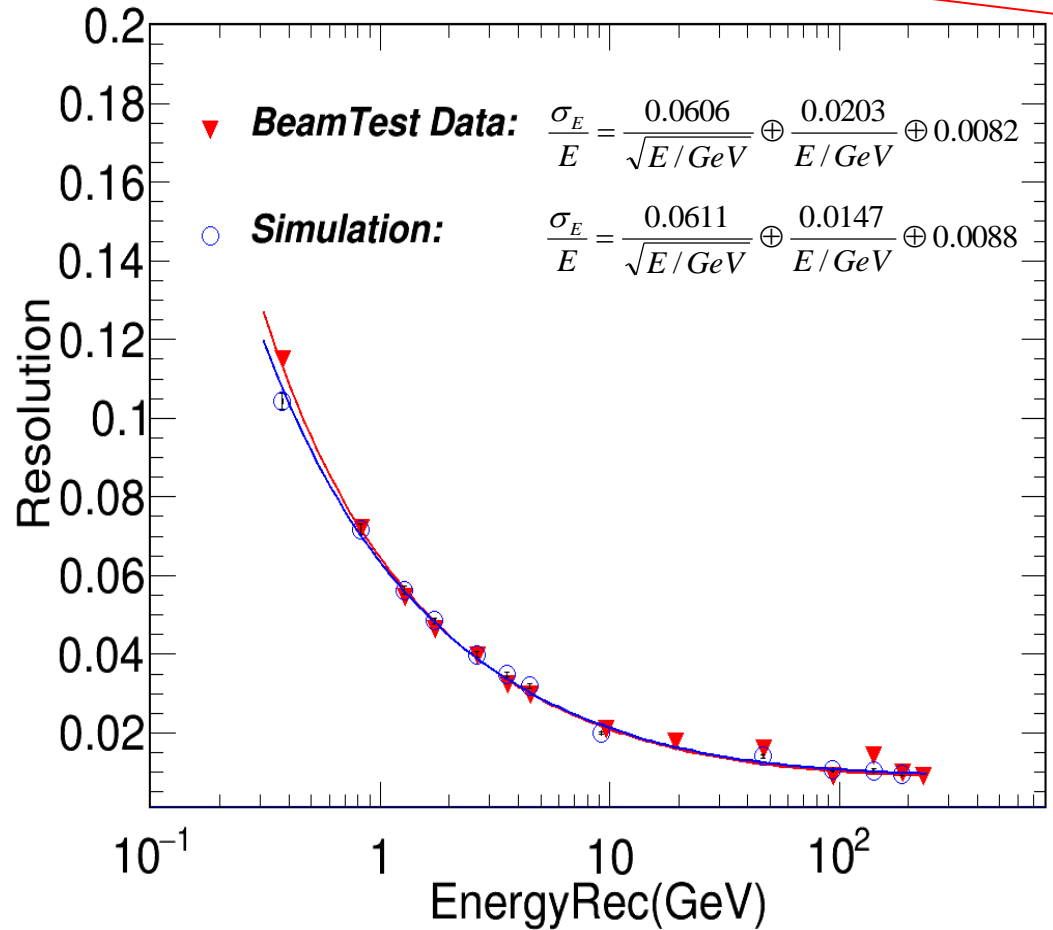
0.5, 1, 1.5, 2, 3, 4, 5 GeV @ PS

5, 10, 20, 50, 100, 149, 197, 243 GeV @ SPS

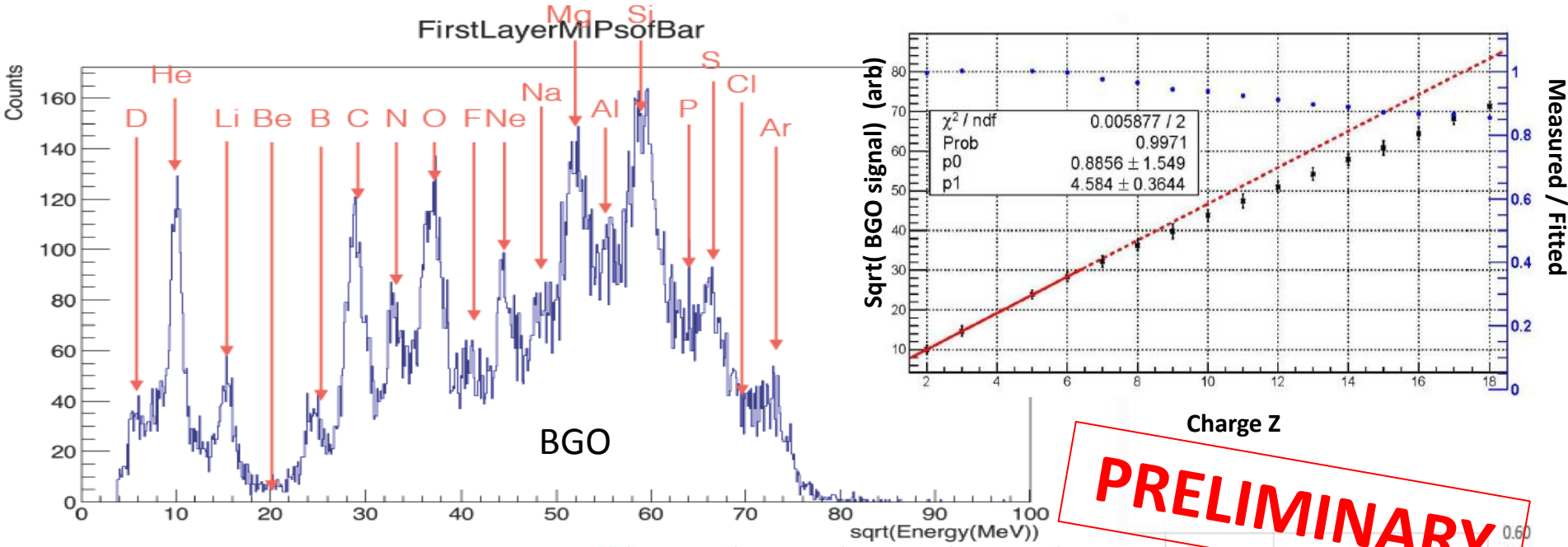
Test beam activity at CERN: electrons



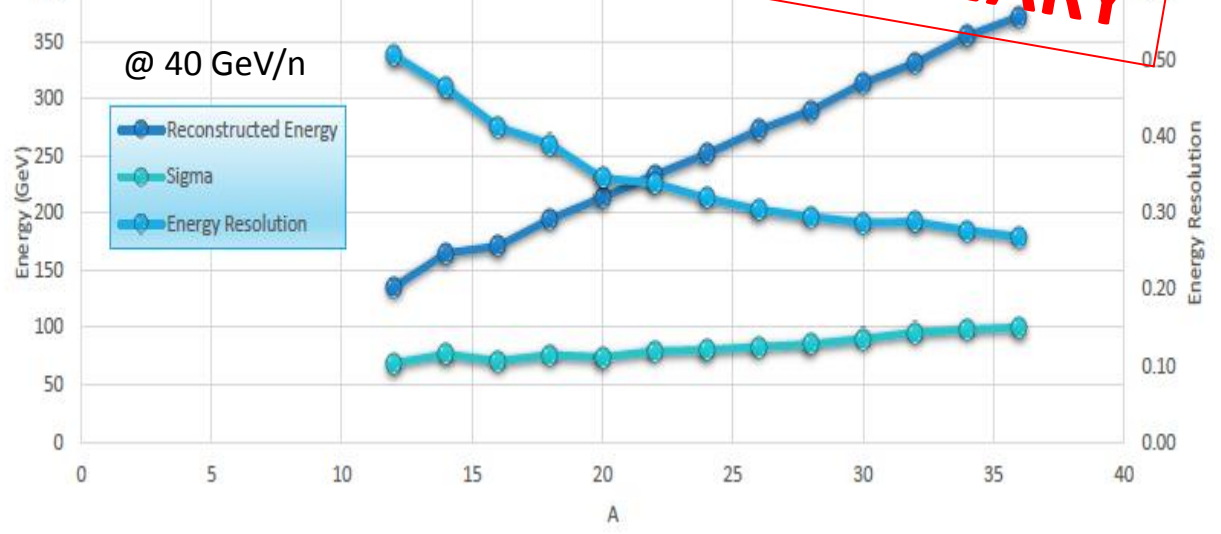
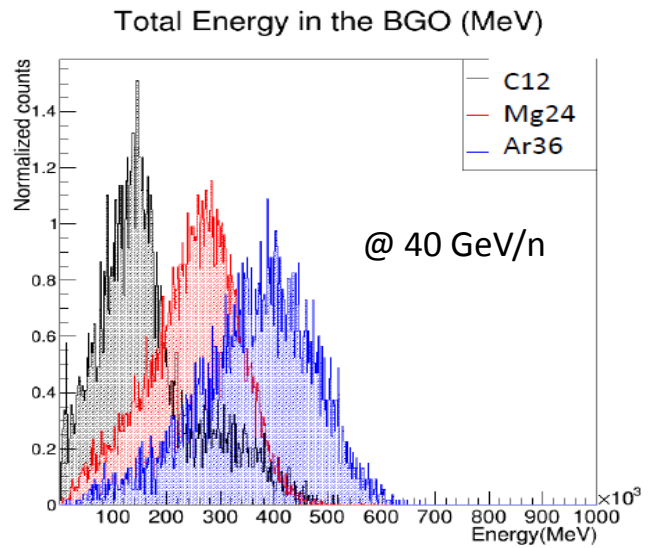
PRELIMINARY



Test beam activity at CERN: ions



PRELIMINARY



The launch: Dec 17th 2015, 0:12 UTC



Jiuquan Satellite Launch Center
Gobi desert

Orbit: sun synchronous , 500km

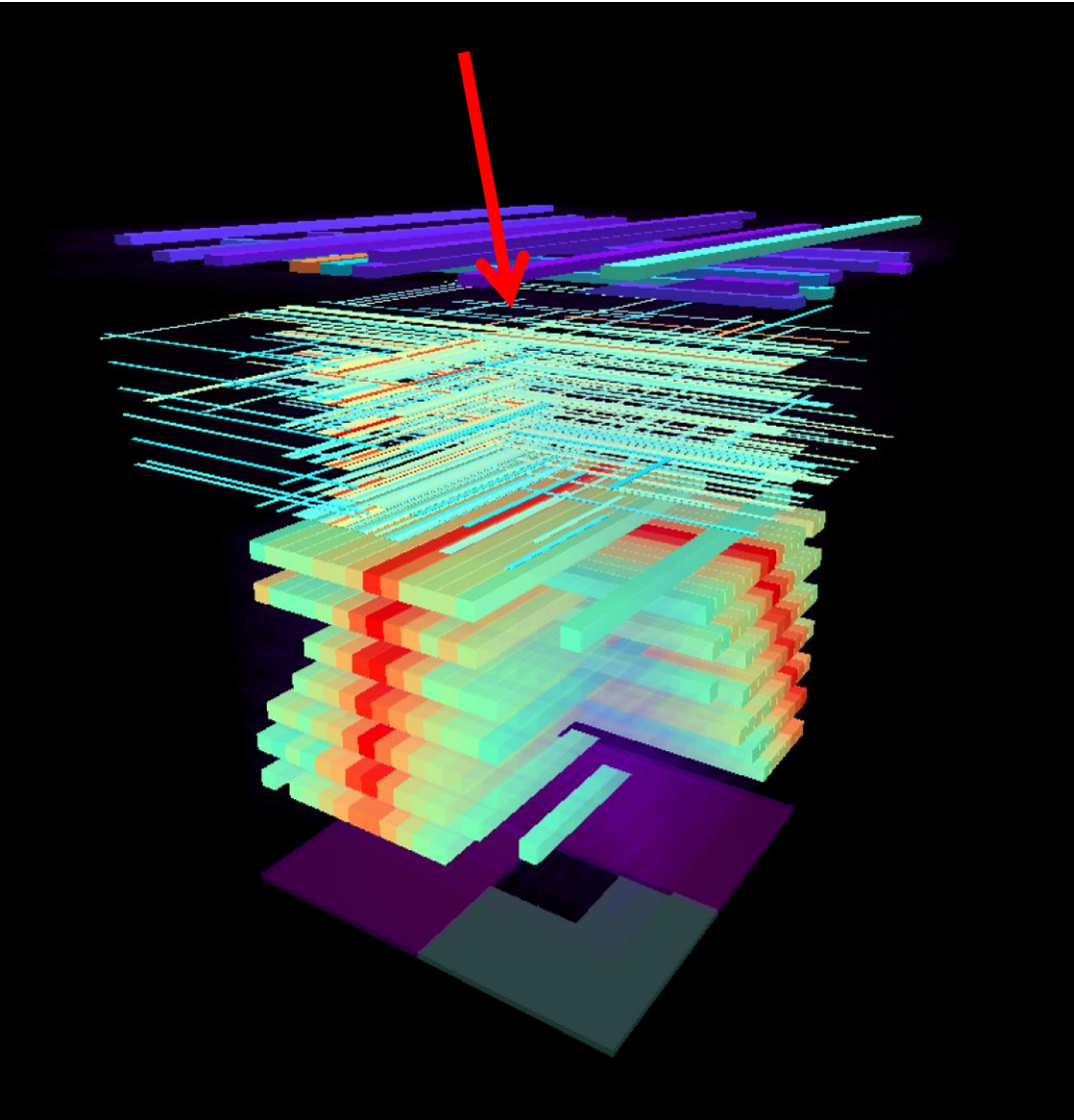


DAMPE → WUKONG

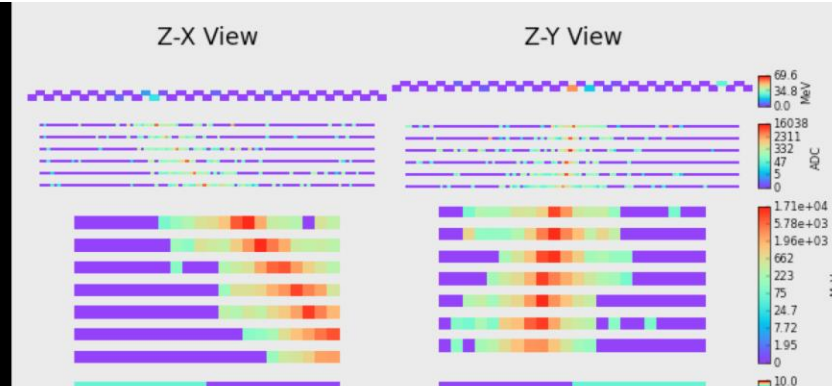


Dec 24th 2015: HV on

330 GeV electron



Z-X View Z-Y View



69.6
34.8
0.0 MeV

16038
2311
332
47
5
0 ADC

1.71e+04
5.78e+03
1.96e+03
662
223
75
24.7
7.72
1.95
0 MeV

<< First < Previous 4911 Next > Last >>

Goto

Colors: 01 02 03 04 05 06 07 08

Stereo Effects: Red Cyan Red Blue Active Passive No Stereo

Advanced Show: Show Trajectory Start Animation Continuous Animation

File Name(s):
../display/20151224_2A/DAMPE_OBS_20151224B012559_RECO5000.root

Event Number:
4911

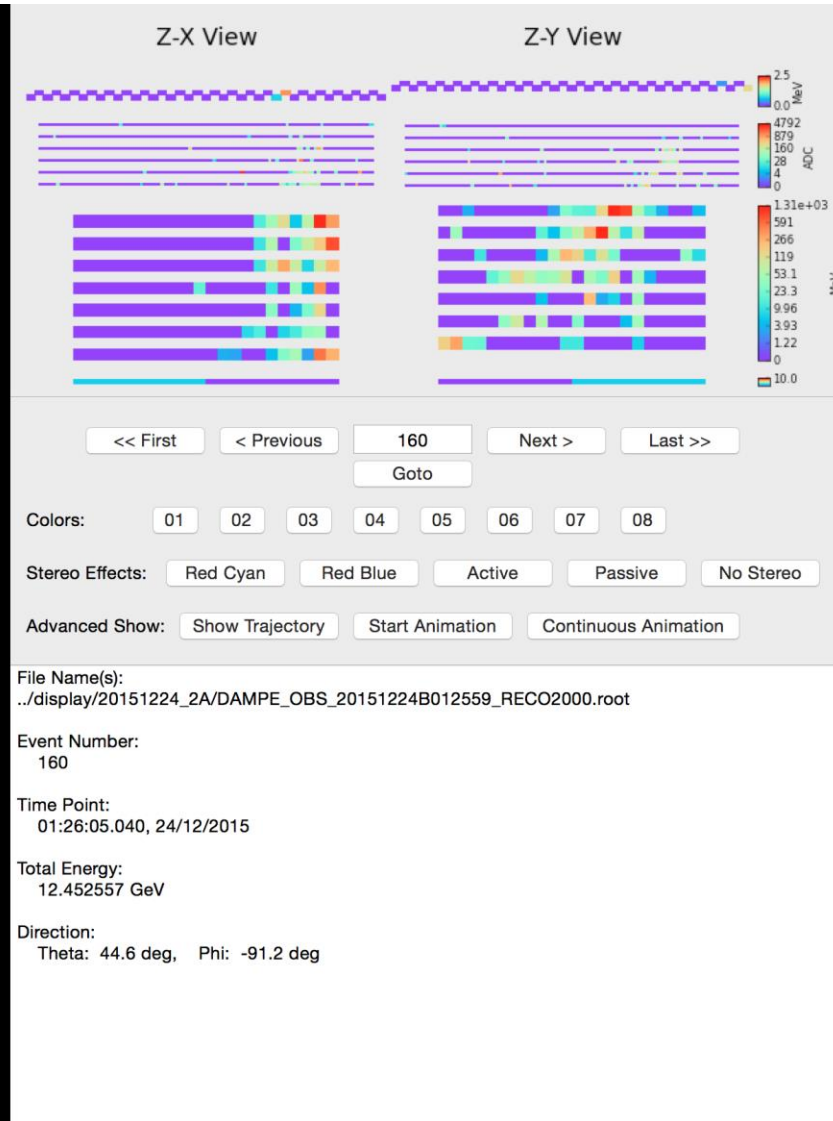
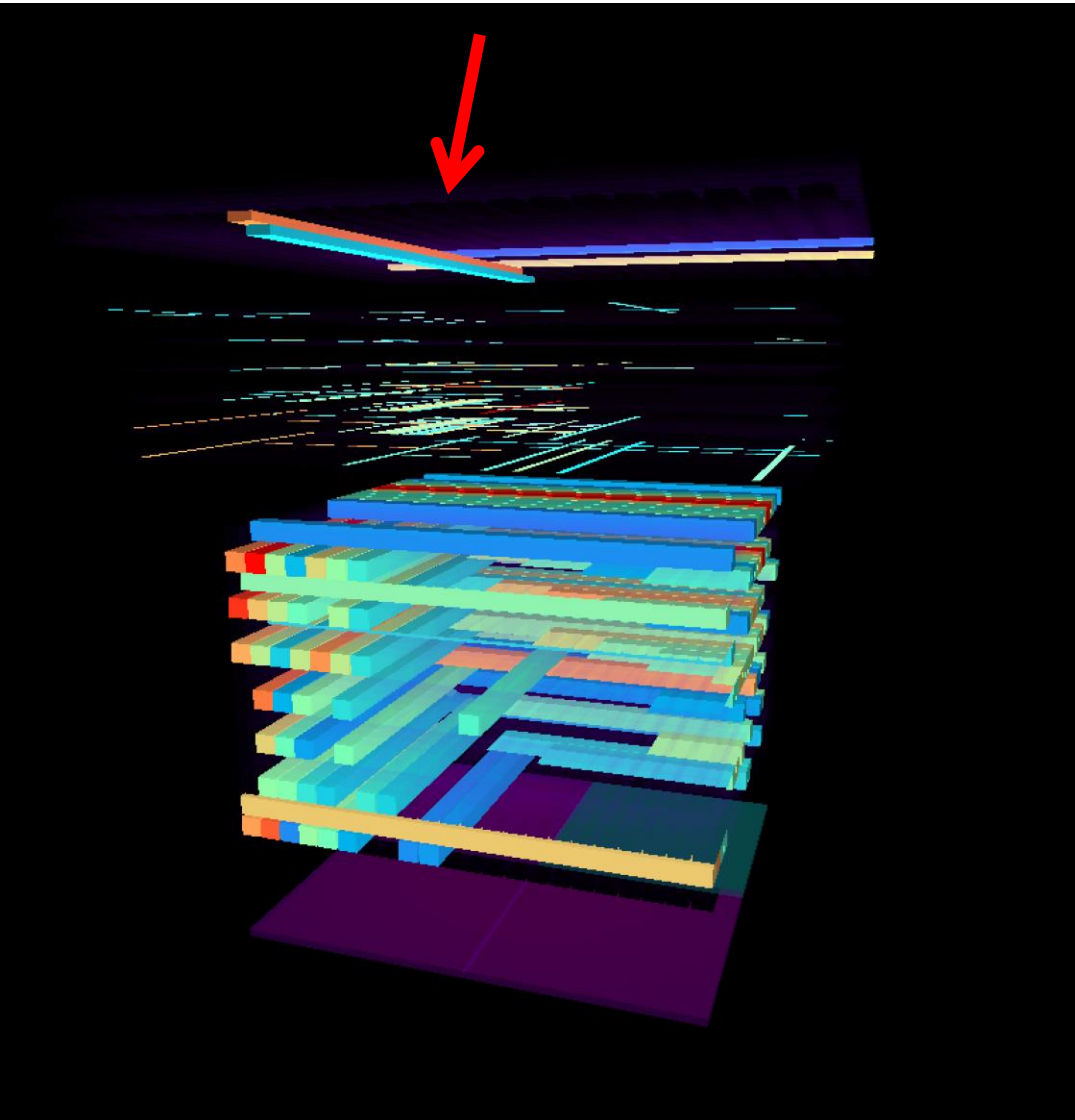
Time Point:
01:28:17.746, 24/12/2015

Total Energy:
328.276531 GeV

Direction:
Theta: 31.3 deg, Phi: -13.0 deg

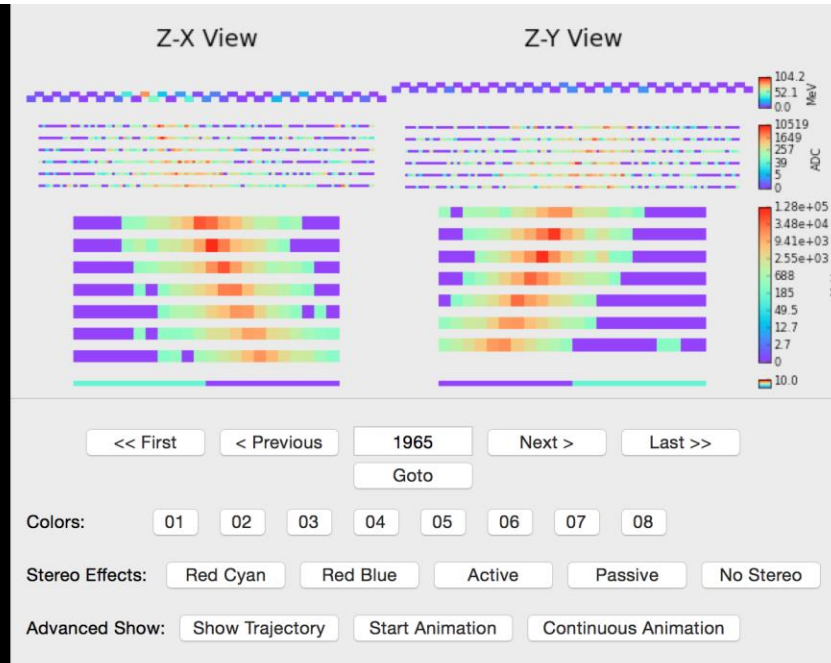
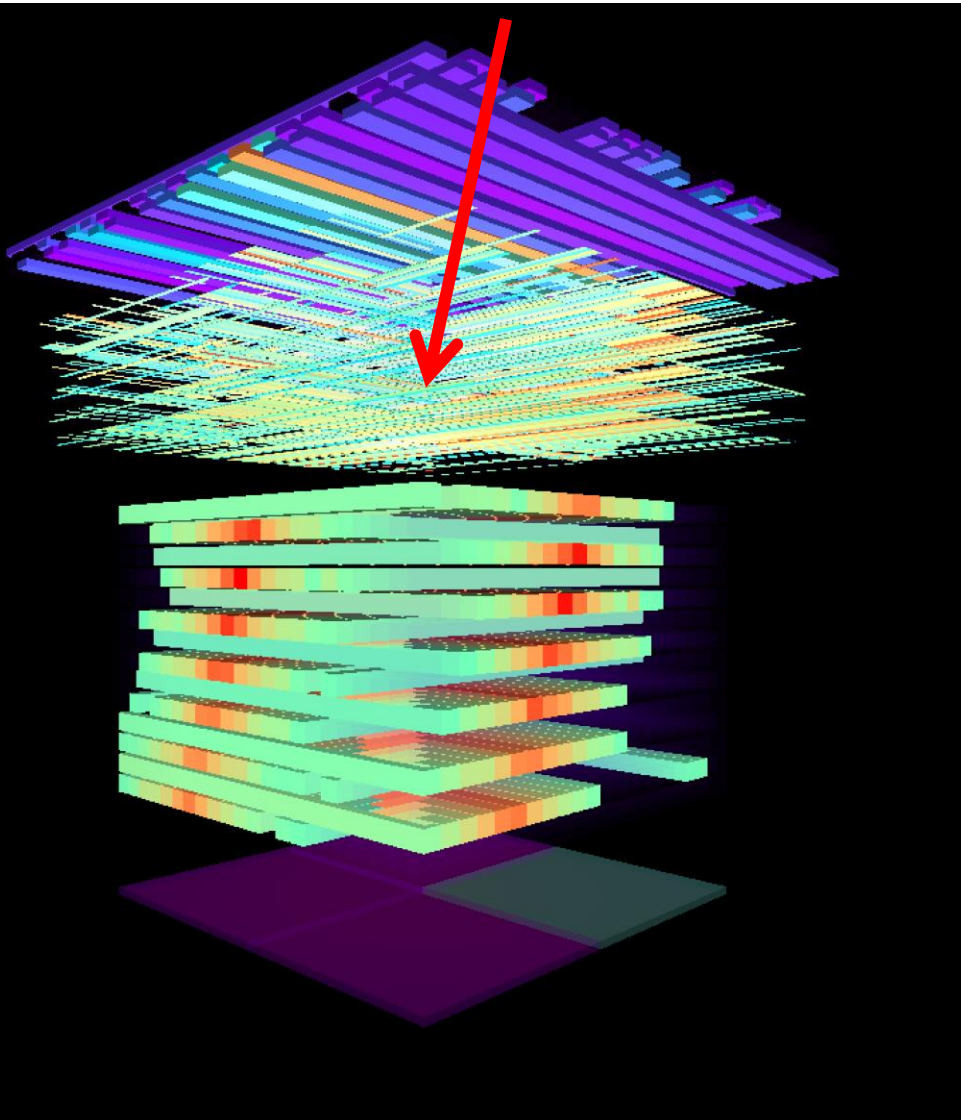
Dec 24th 2015: HV on

12 GeV proton



Dec 24th 2015: HV on

1.3 TeV carbon



File Name(s):
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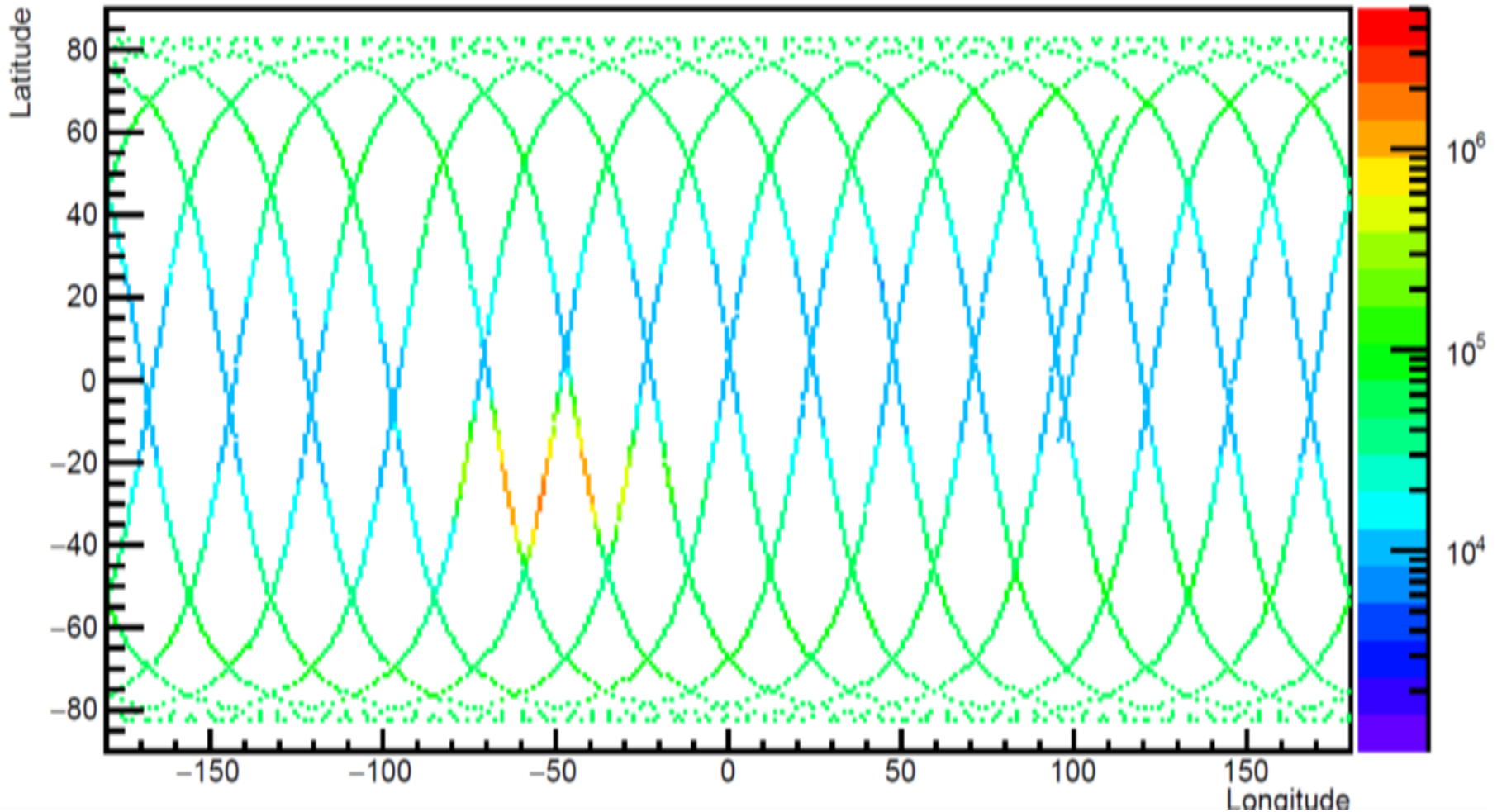
Event Number:
1965

Time Point:
01:26:56.936, 24/12/2015

Total Energy:
1306.882750 GeV

Direction:
Theta: 26.8 deg, Phi: -45.9 deg

Trigger rate in orbit

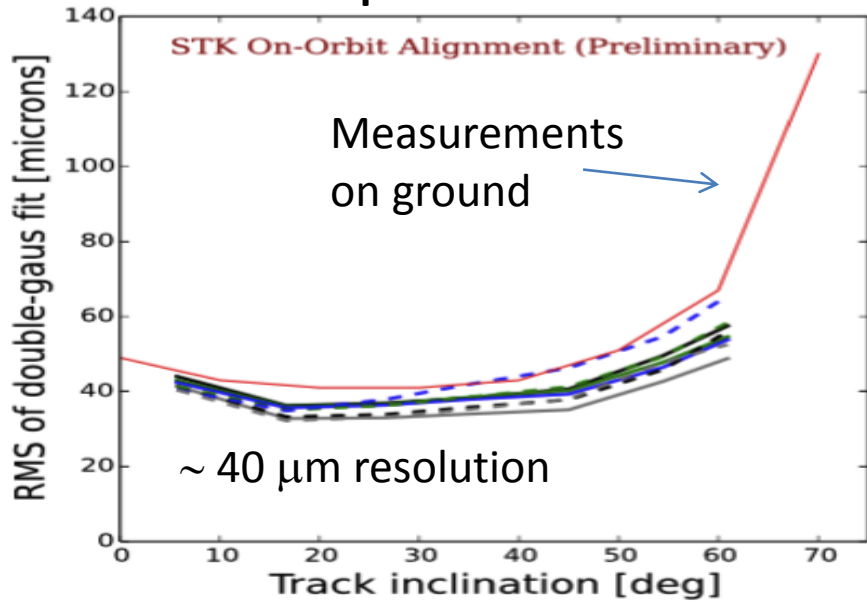


~50 Hz average trigger rate

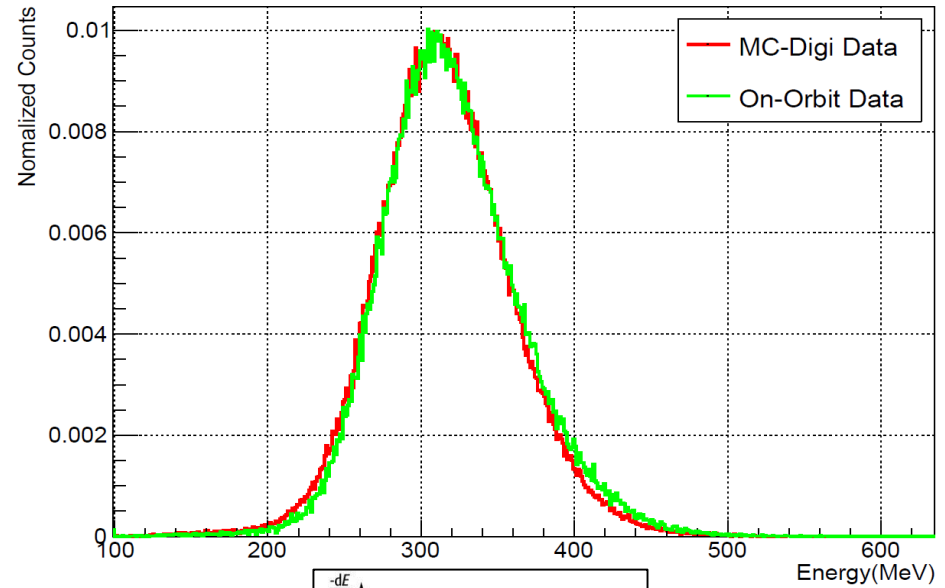
→ 100GB/day on ground (about 4 M events)

Some on-orbit performance plot

STK space resolution

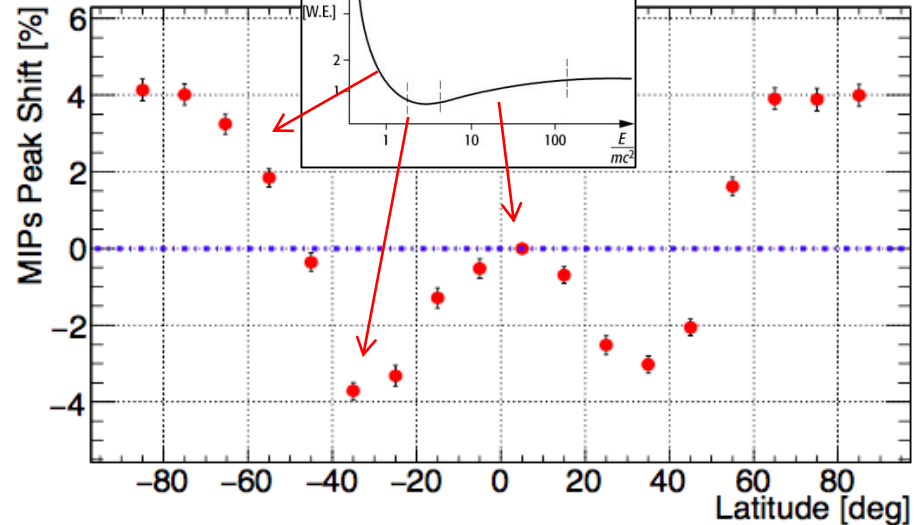
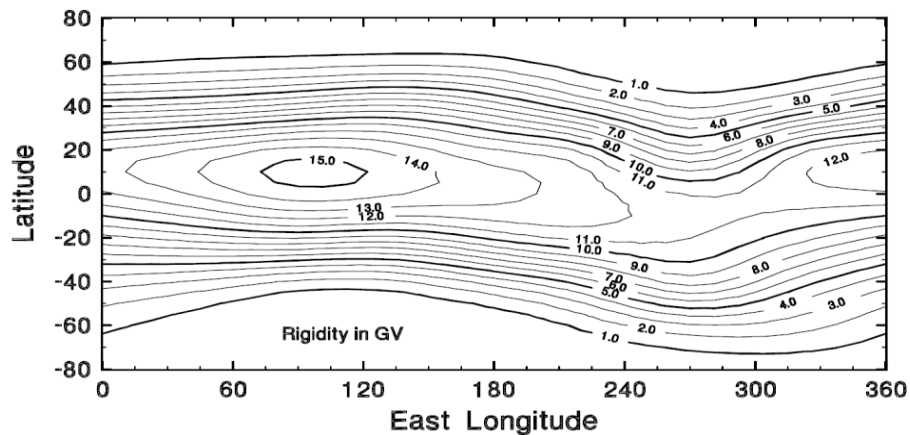


BGO "mip" peak



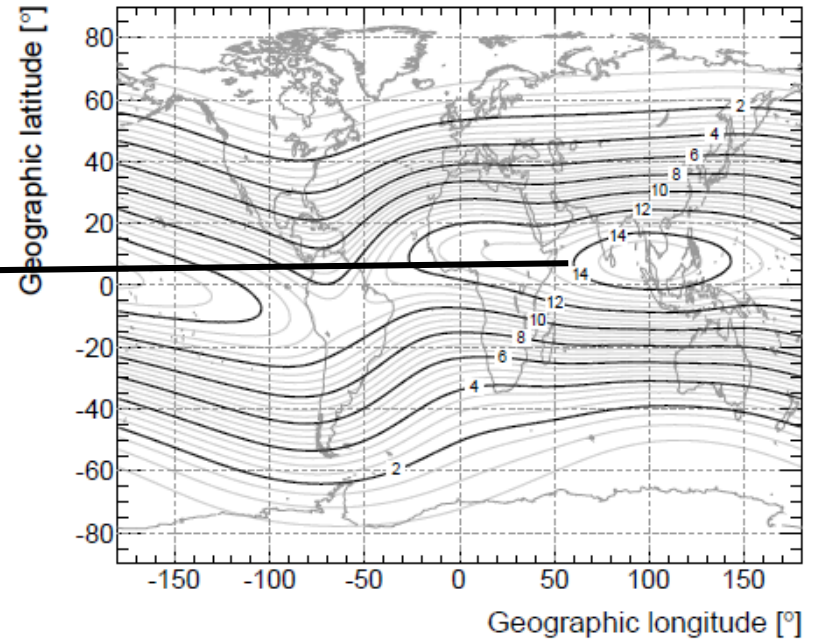
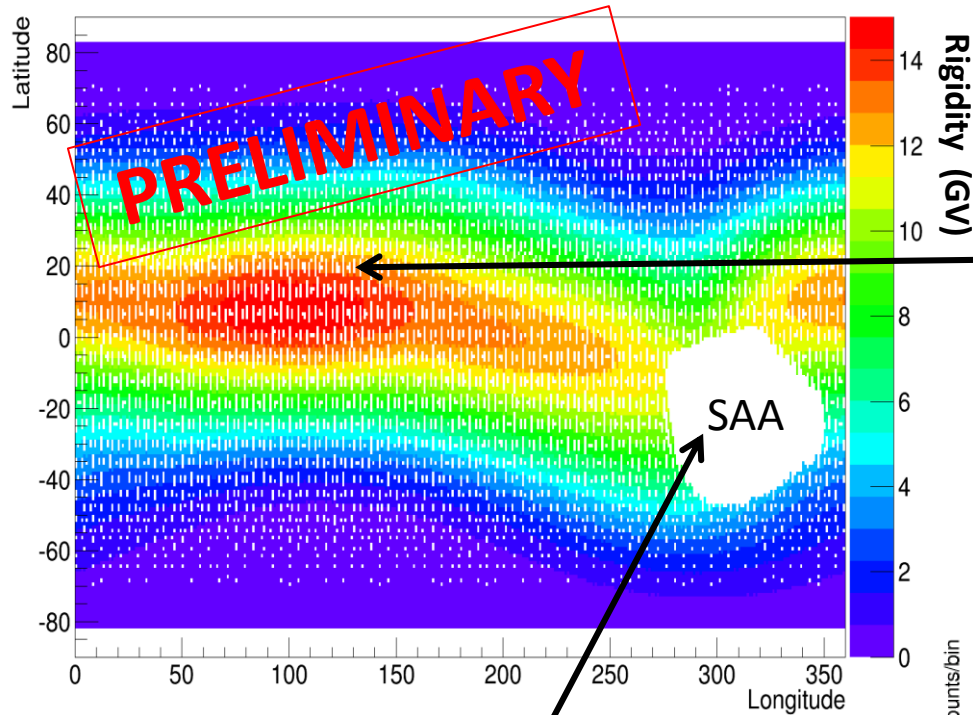
VERTICAL CUTOFF RIGIDITIES AT 450 KM

Tsyganenko model $K_p = 0$

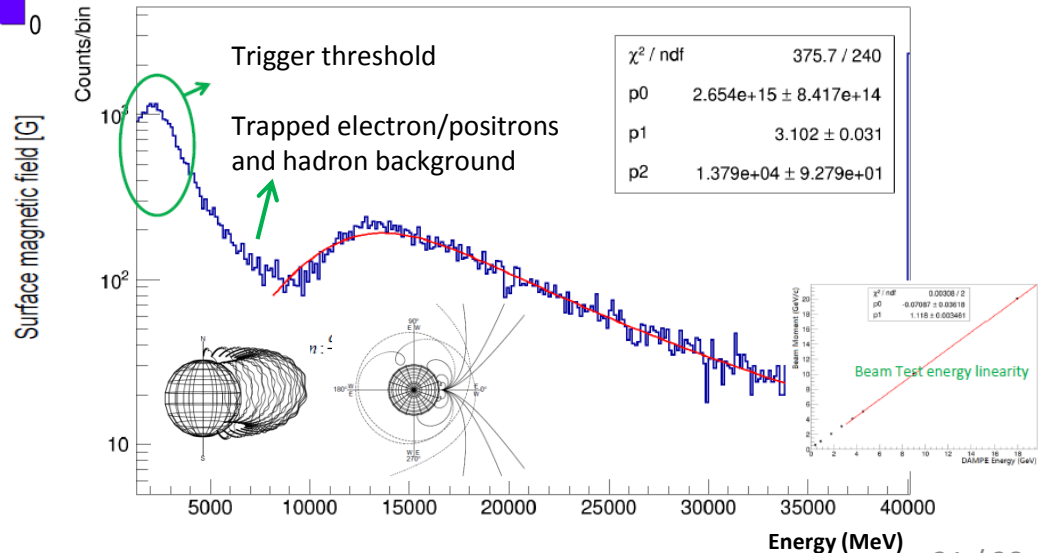
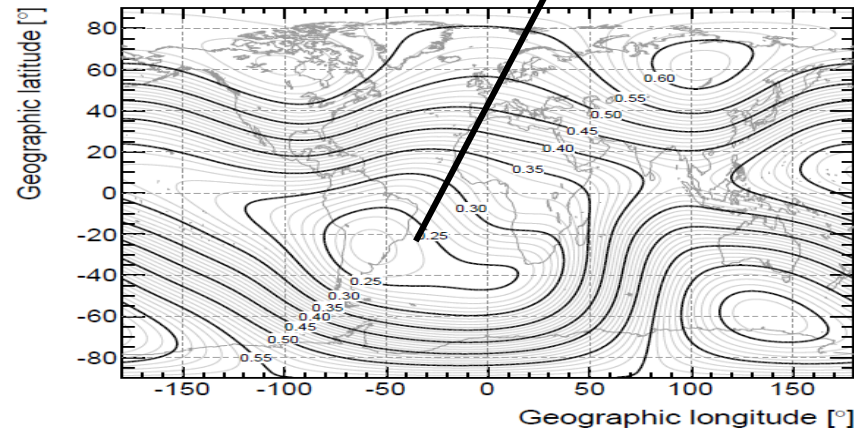


On-orbit energy calibration

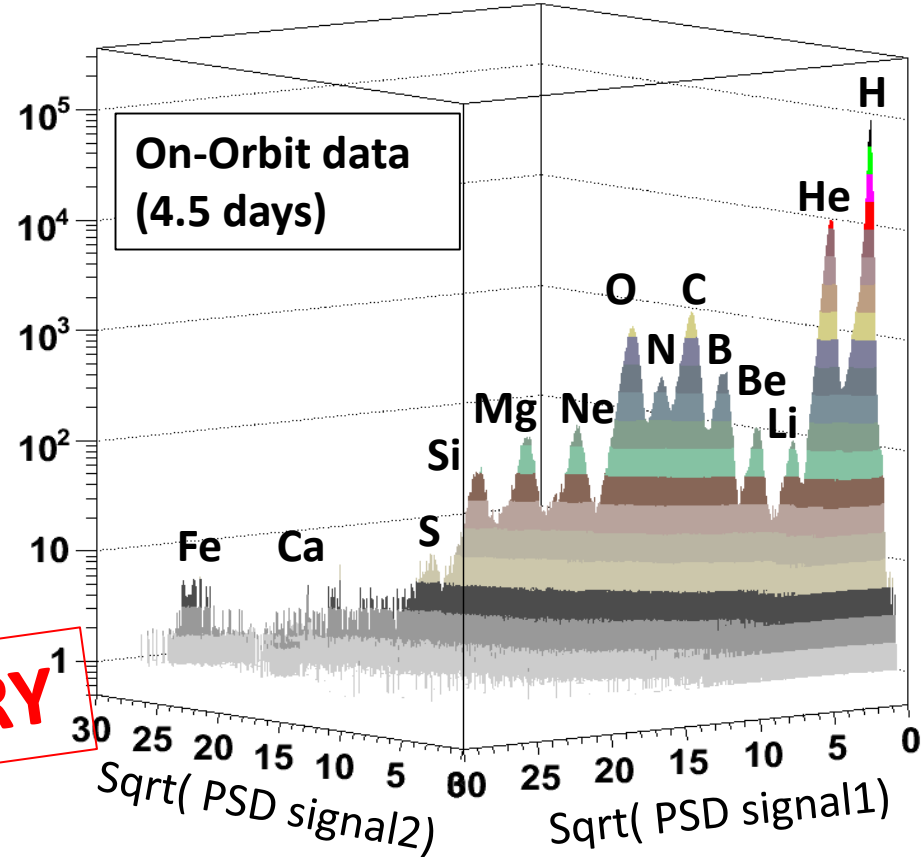
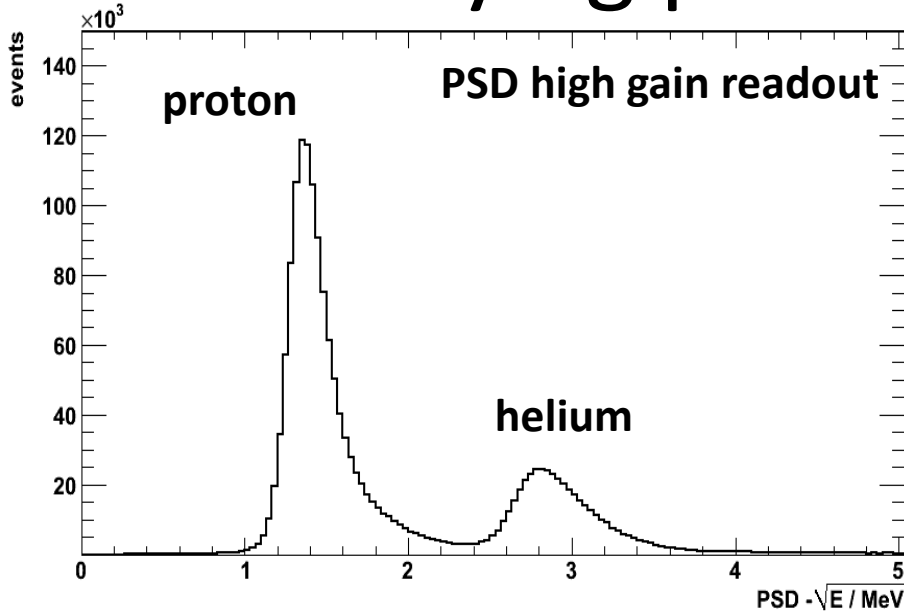
e^\pm rigidity cutoff with 1 month statistics



Vertical rigidity cutoff @ $h = 500$ km [GV]



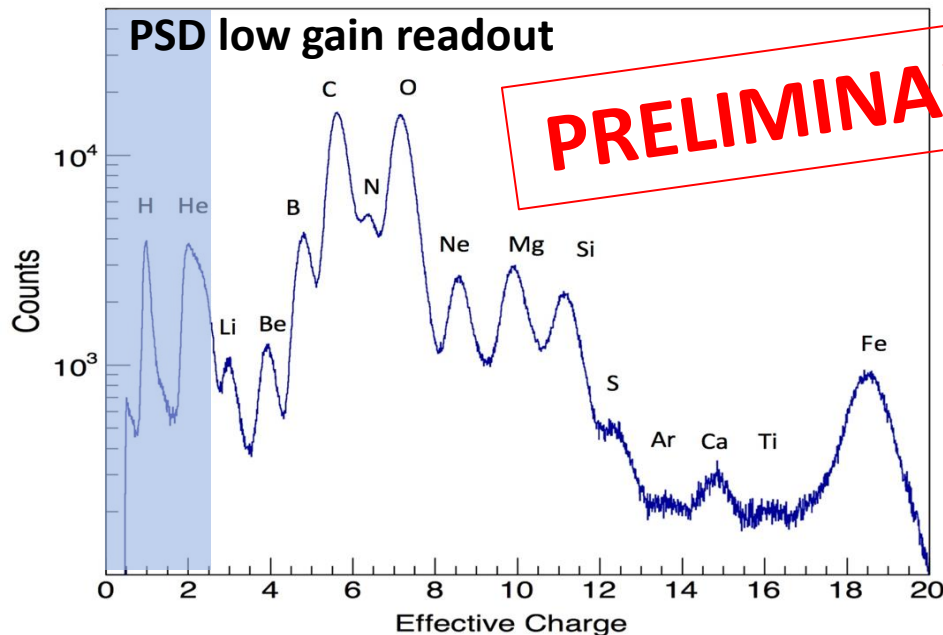
Identifying protons and nuclei



PRELIMINARY

Charge resolution $\sim 0.2-0.3 e$ depending on Z

Charge measurement also given by STK and (with lower precision) by the BGO bars

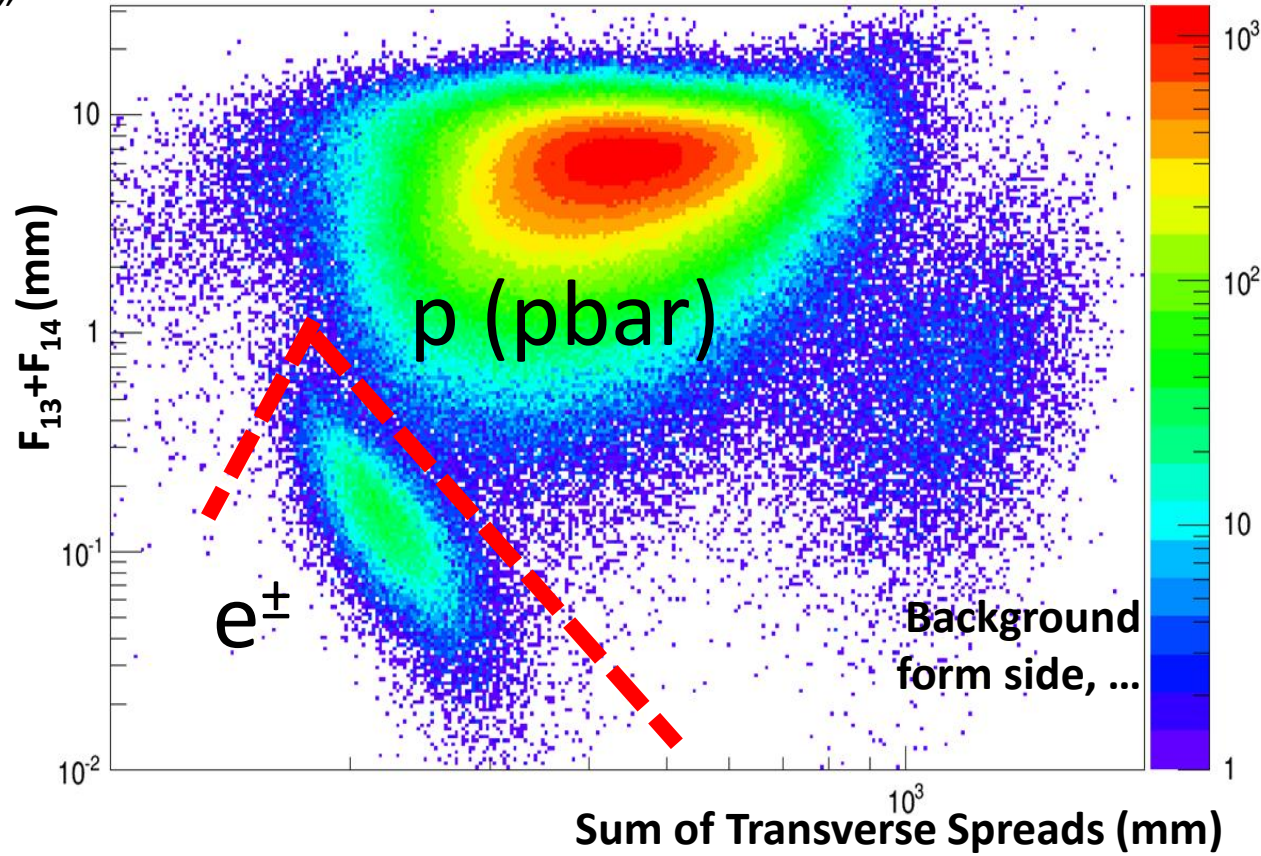
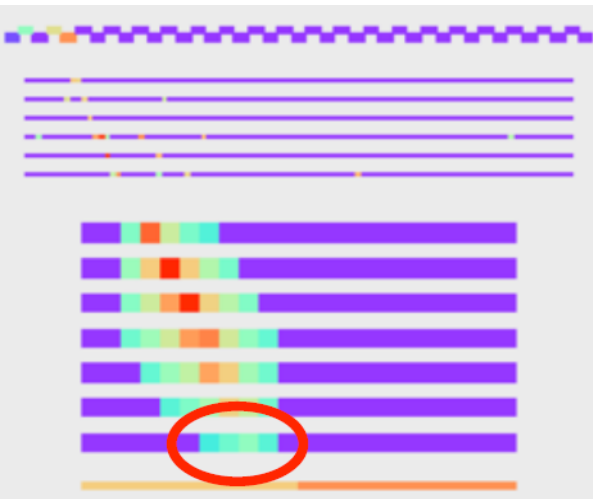


Electron identification

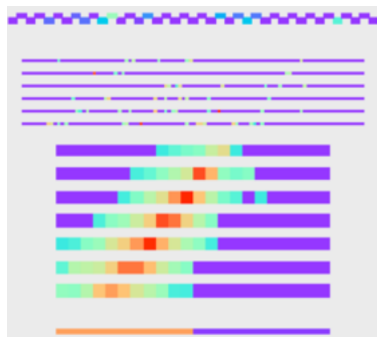
One possible “shape parameter”

$$F_i = Spread_i \times \frac{E_i}{E_{tot}}$$

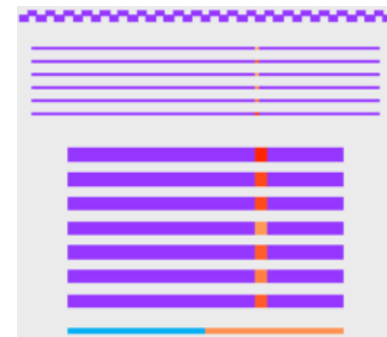
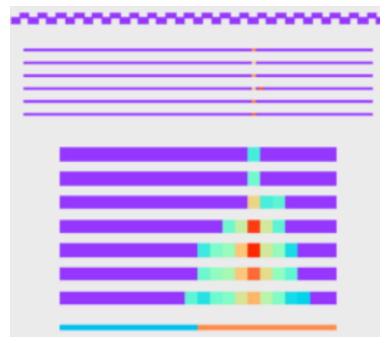
Rejection power $> 10^5$



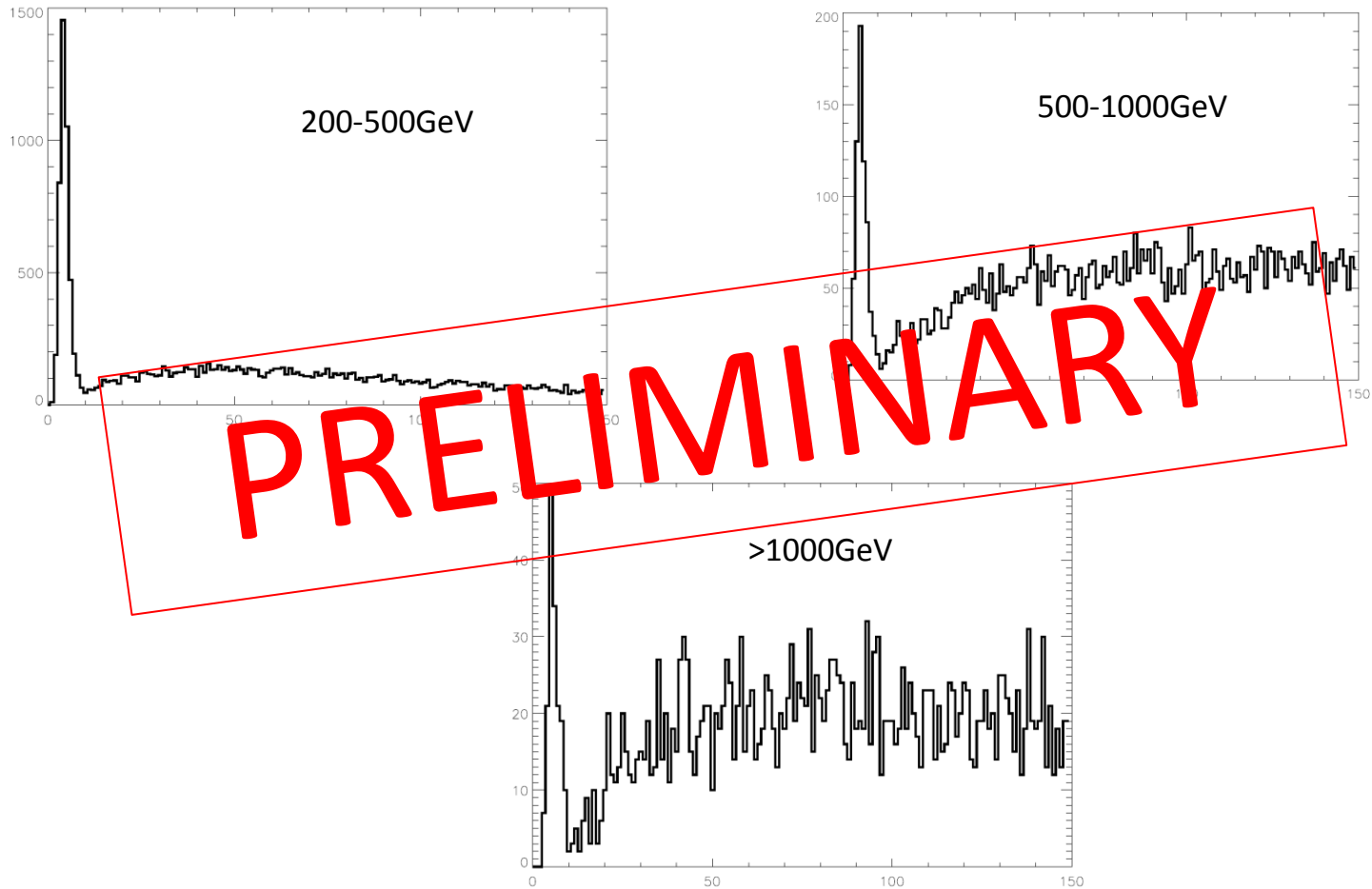
Electrons and positrons



Protons and nuclei



Electron identification



Photons

PRELIMINARY

DAMPE 165 days

$E > 1\text{ GeV}$

Counts / $(0.5^\circ)^2$ pixel

$\sigma_\theta \approx 0.2^\circ$ @ 3 GeV

Vela

Geminga

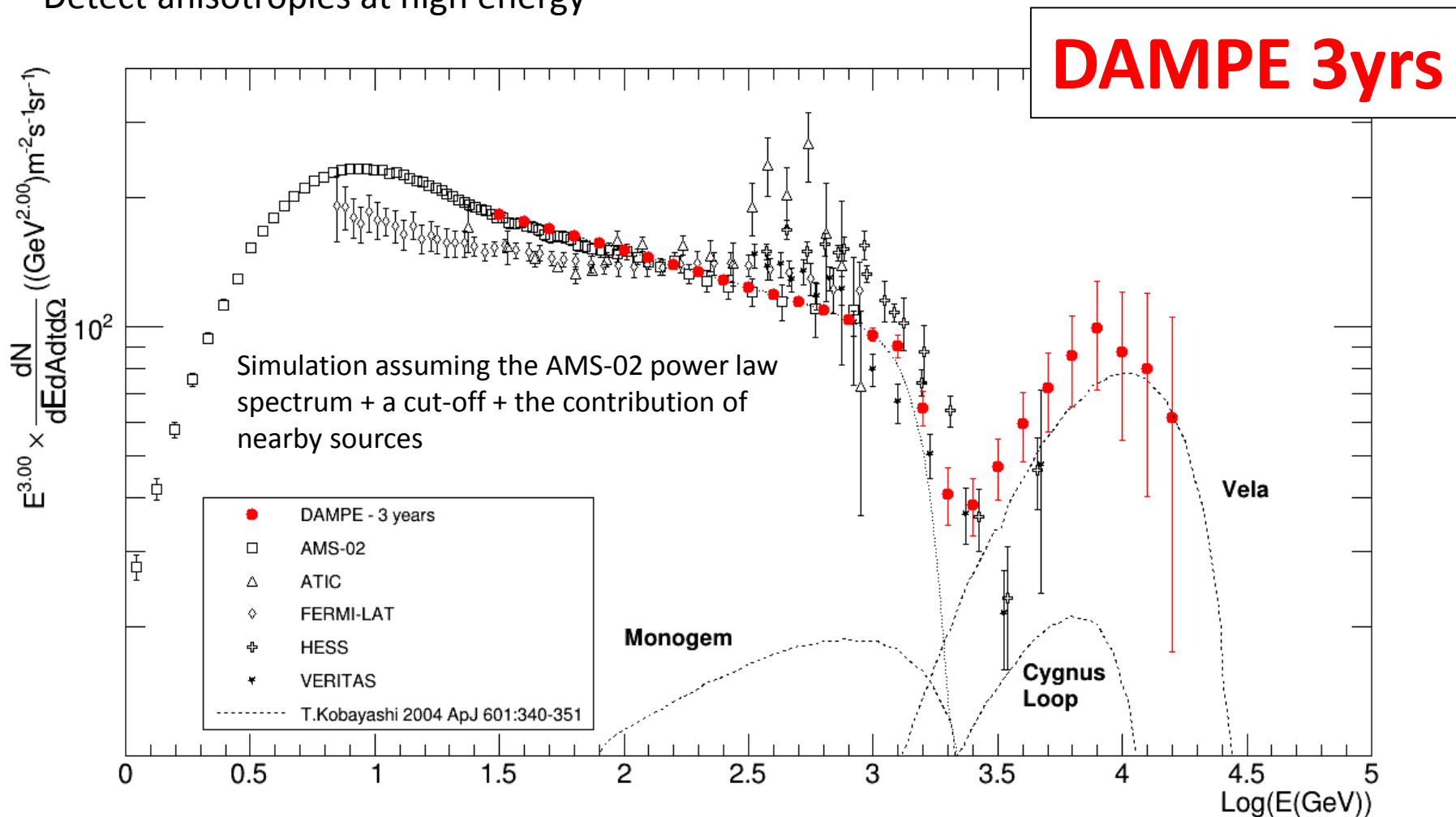
Crab

FERMI 5 years

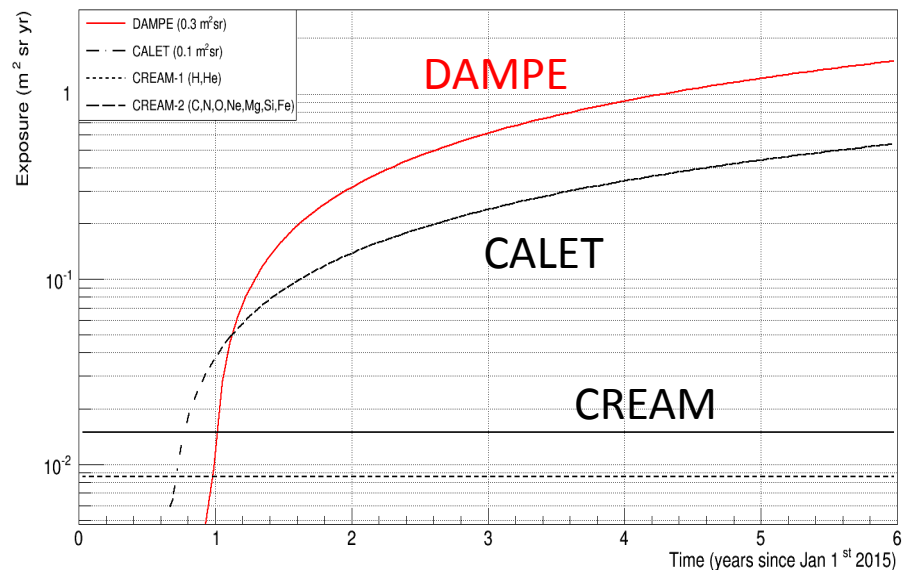
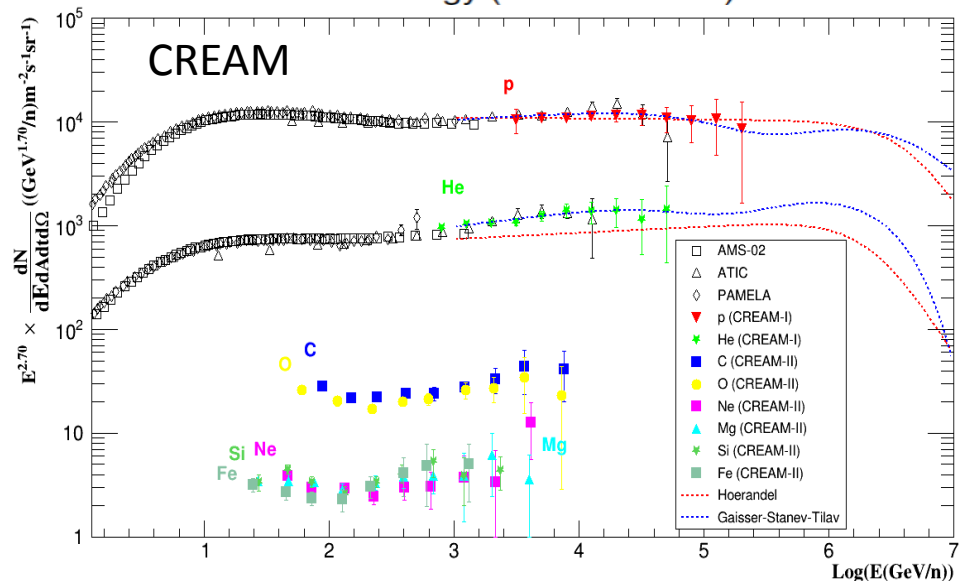
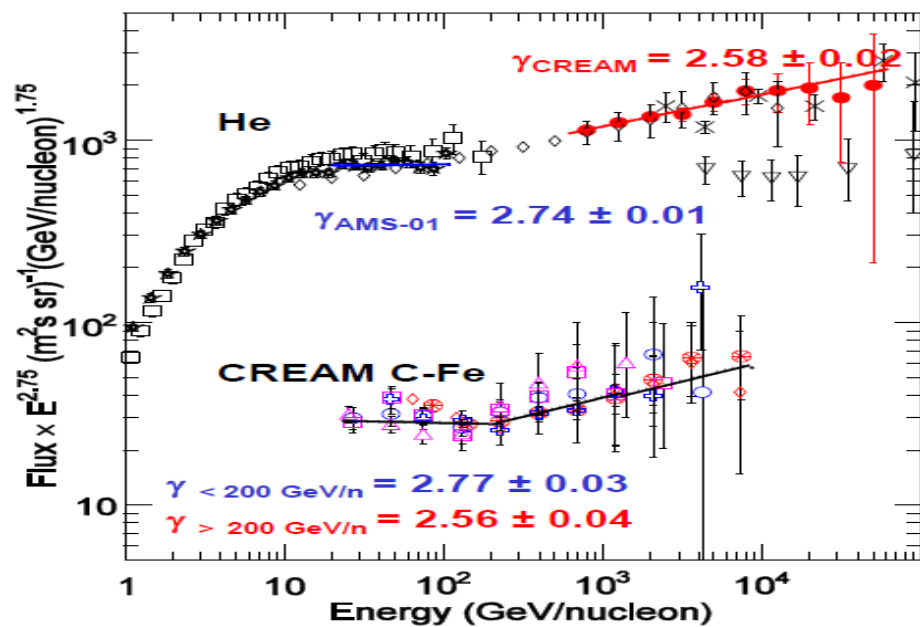
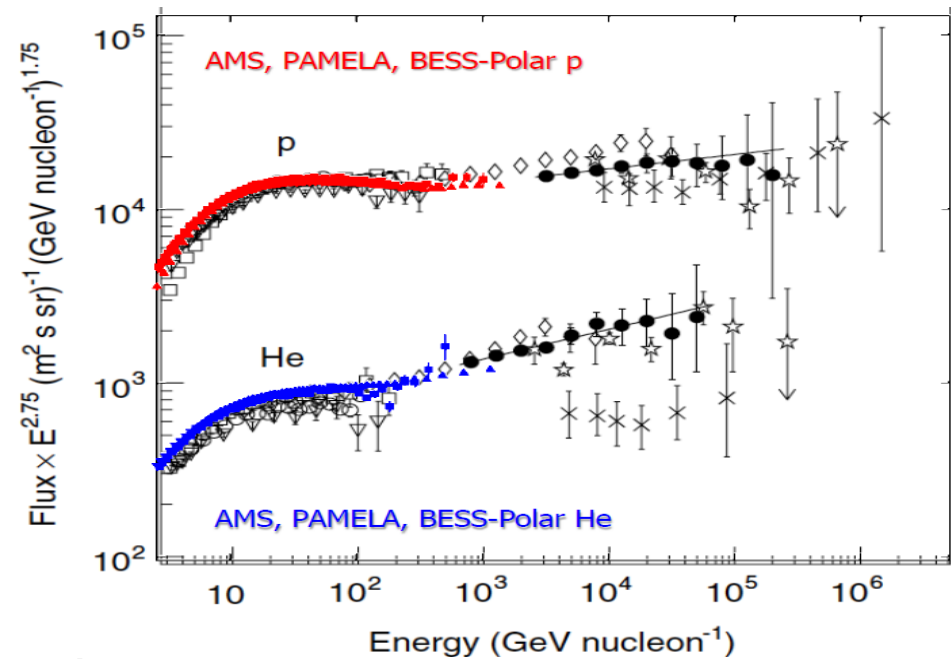
$E > 1\text{ GeV}$

All-electron spectrum

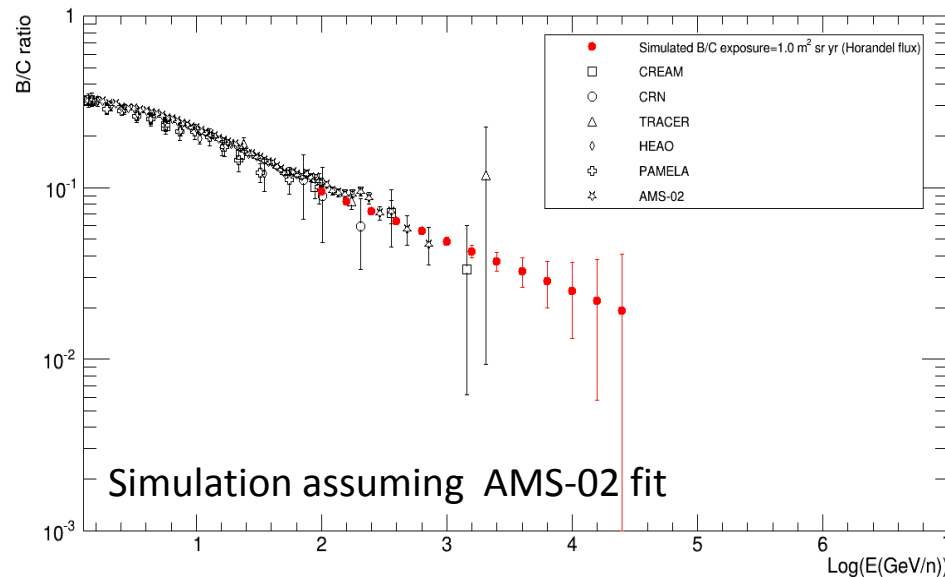
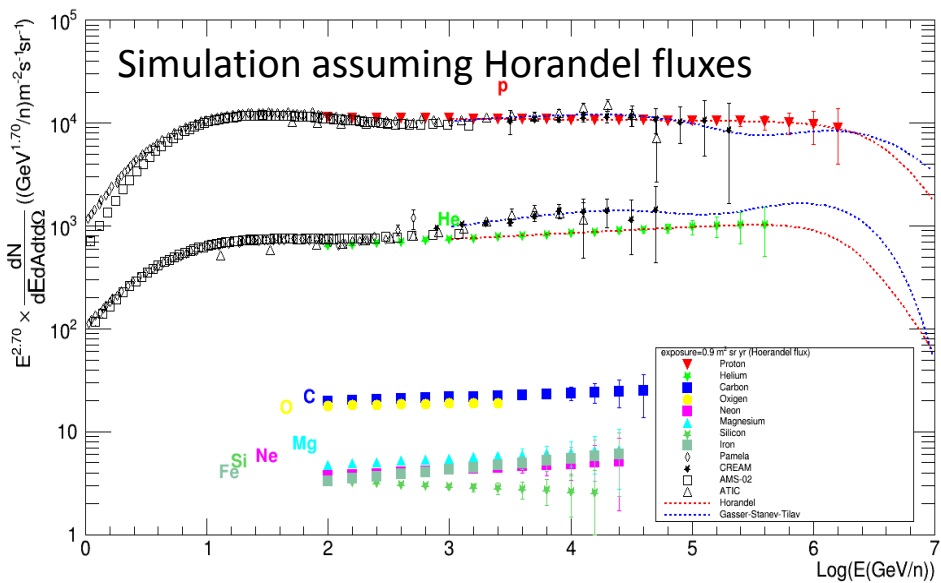
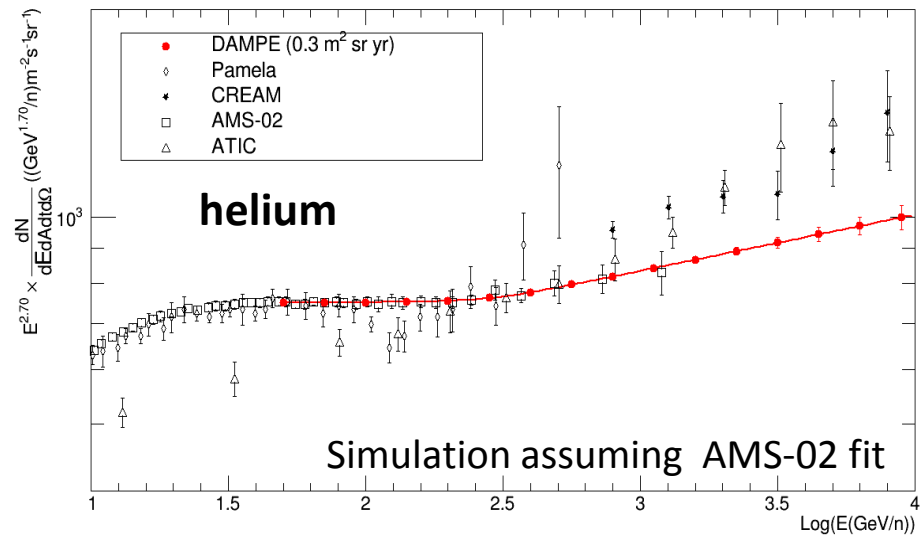
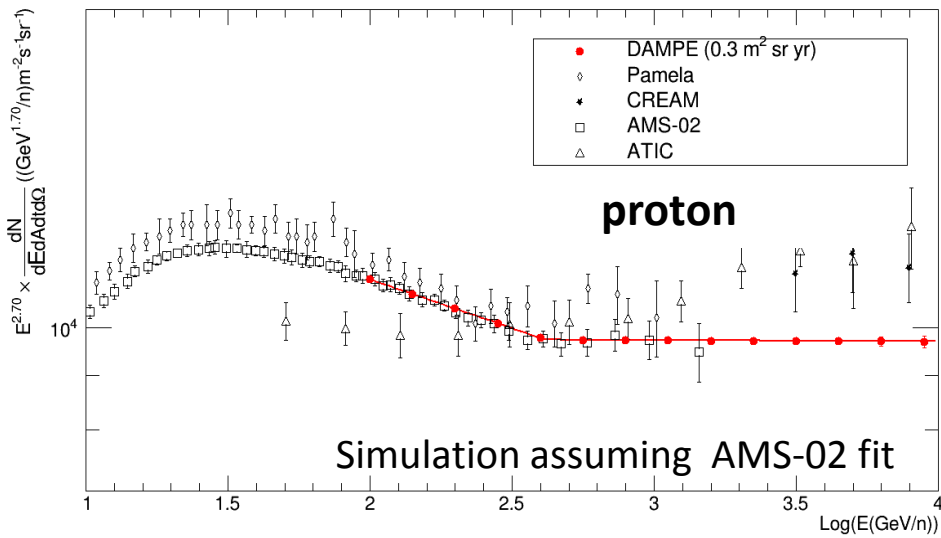
- Measure the all-electron flux up to about 10TeV
- Measure with high accuracy the sub-TeV region and the possible cut-off around one TeV
- Detect structures in the spectrum due to nearby sources and/or DM induced excesses
- Detect anisotropies at high energy



Protons and nuclei spectra



Protons and nuclei: DAMPE 3years



Summary

The detector

- Large geometric factor instrument ($0.3 \text{ m}^2 \text{ sr}$ for p and nuclei)
- Precision Si-W tracker ($40 \mu\text{m}$, 0.2°)
- Thick calorimeter ($32 X_0$, σ_E/E better than 1% above 50 GeV for e/γ , $\sim 35\%$ for hadrons)
- “Mutiple” charge measurements (0.2-0.3 e resolution)
- e/p rejection power $> 10^5$ (topology alone, plus neutron detector)

Launch and performances

- Succesfull launch on dec 17, 2015
- On orbit operation steady and with high efficiencies
- Absolute energy calibration by using the geomagnetic cut-off
- Absolute pointing cross check by use of the photon map

Physics goals

- Study of the cosmic electron and photon spectra
- Study of electron anisotropy and nearby sources contribution
- Study of cosmic ray protons and nuclei: spectrum and composition
- Precise measurement of CR discrepant hardenings and spectral indexes
- High energy gamma ray astronomy
- Search for dark matter signatures in lepton spectra
- The “unexpected”: GW electromagnetic follow up in FoV

