Jet energy loss in small systems with finite-size effects and running coupling

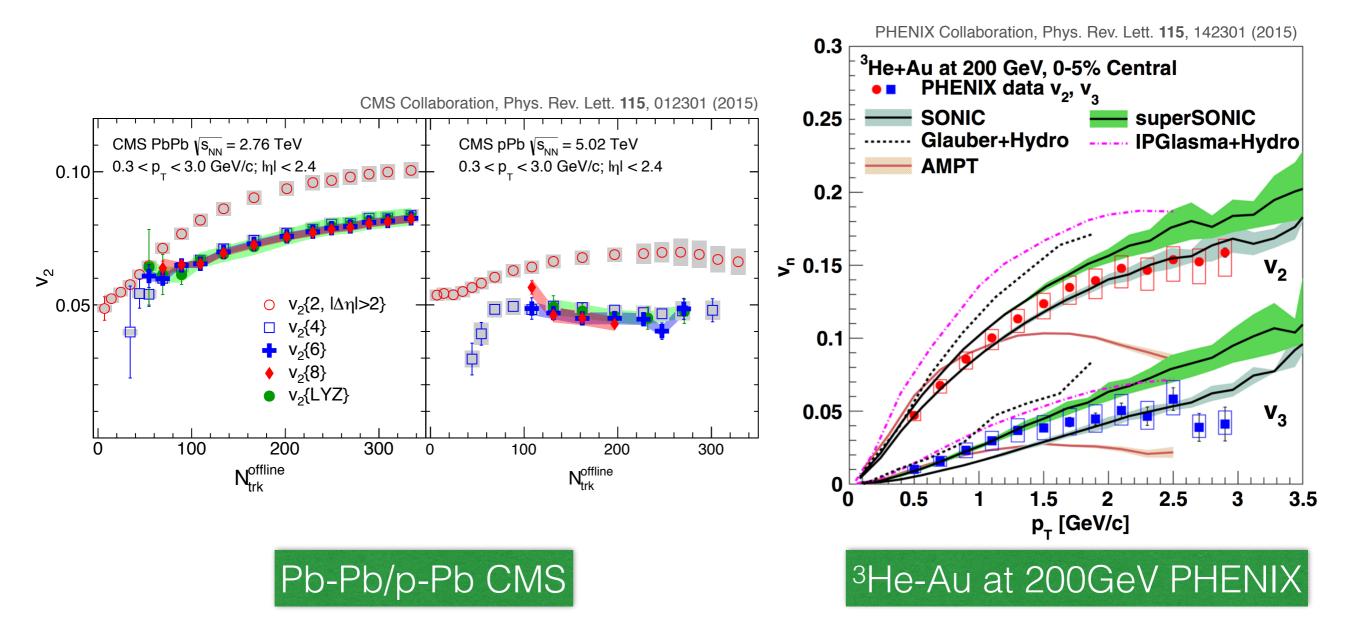
Chanwook Park

Collaboration with C. Shen, S. Jeon, C. Gale





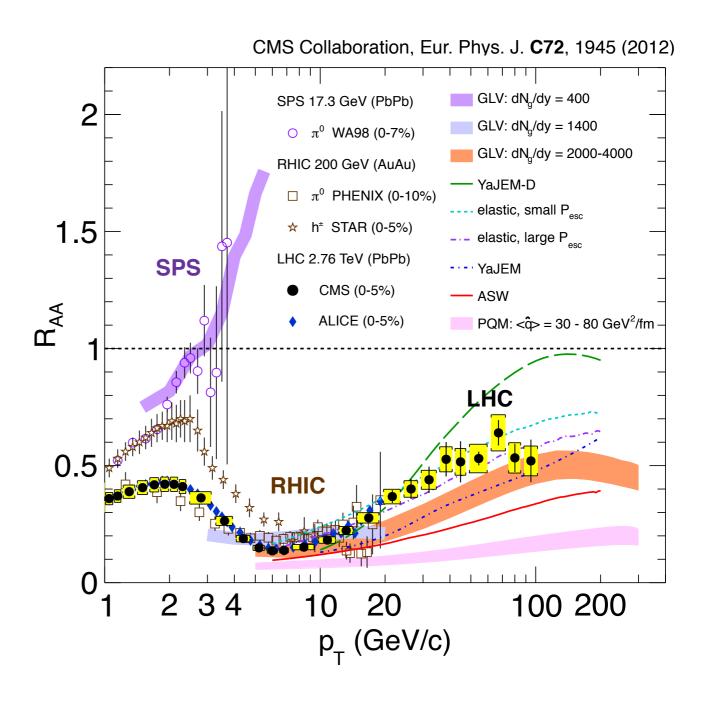
Collectivity in small systems



- Collective behaviour was measured in central collisions of small systems
- This suggests strongly coupled medium in small systems

A Tool for studying QGP: Jets

- Jet quenching : due to colour opacity of a strongly coupled medium
- Strong suppression in AA collisions was measured at RHIC and the LHC
- Similar behaviour in small systems?



Jet-medium interaction

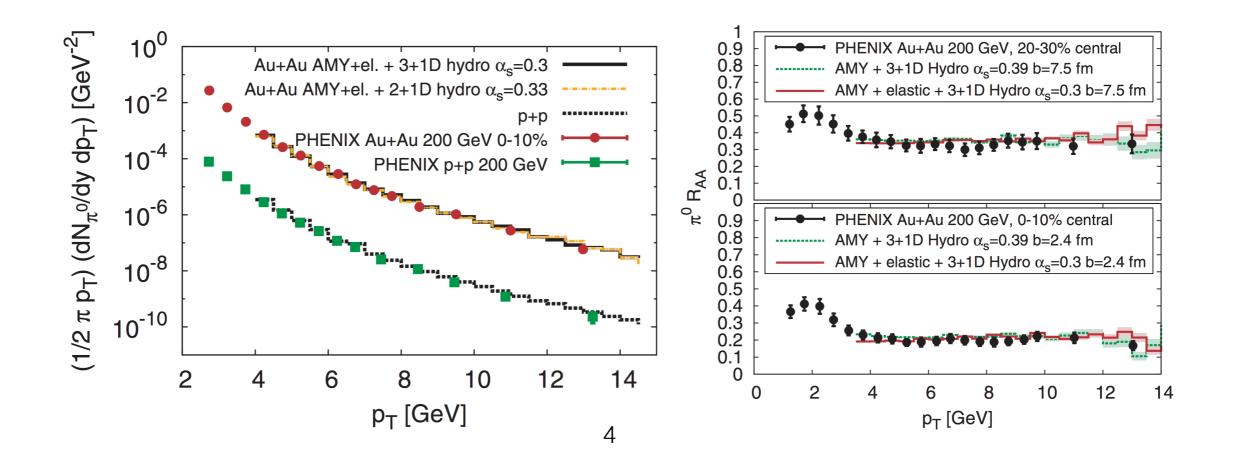
MARTINI B. Schenke, S. Jeon, C. Gale, Phys. Rev. C80, 054913 (2009)

- Nucleus MARTINI

 QGP medium

 Hard

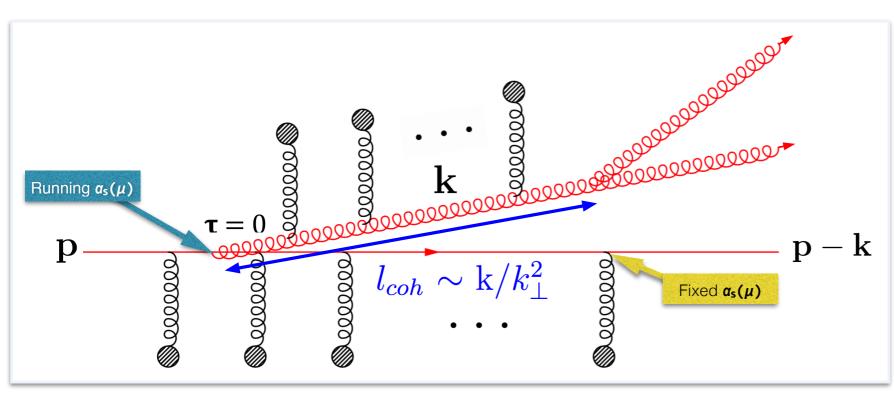
 Process
- Event generator for hard jets in heavy ion collisions
- Based on PYTHIA 8.2
- AMY(Arnold-Moore-Yaffe) radiation scheme as well as collisional processes
- E-by-E hydrodynamics background is available

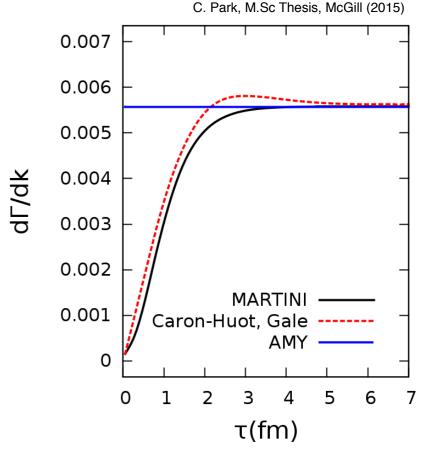


New developments in MARTINI

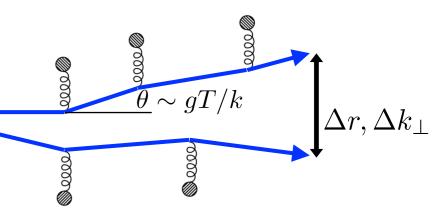
Finite-size effects on radiation

S. Caron-Huot and C. Gale, Phys. Rev. C82, 064902 (2010)





- Finite formation time of radiation
 : Time dependence is applied in AMY rate
- Implementation : Random walk (separation condition $\Delta r \Delta k_{\perp} > \frac{1}{2}$



New developments in MARTINI

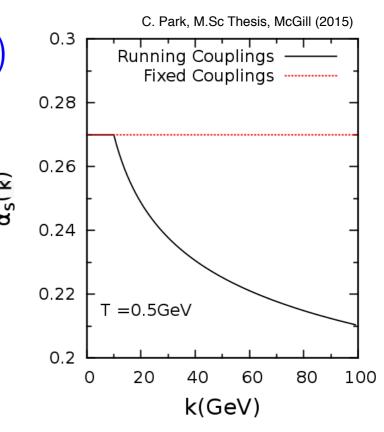
Running coupling

C. Young, B. Schenke, S. Jeon, and C. Gale, Nucl. Phys. A910-911, 494 (2013)

Origin of g_s in the AMY rate

$$\frac{d\Gamma(p,k)}{dk} = \frac{C_s g_s^2}{16\pi p^7} \frac{1}{1 \pm e^{-k/T}} \frac{1}{1 \pm e^{-(p-k)/T}} \times \begin{cases} \frac{\frac{1+(1-x)^2}{x^3(1-x)^2}}{N_f \frac{x^2+(1-x)^2}{x^2(1-x)^2}} & q \to qg \\ N_f \frac{x^2+(1-x)^2}{x^2(1-x)^2} & g \to qq \\ \frac{1+x^4+(1-x)^4}{x^3(1-x)^3} & g \to gg \end{cases} \\
\times \int \frac{d^2 \mathbf{h}}{(2\pi)^2} 2\mathbf{h} \cdot \text{Re } \mathbf{F}(\mathbf{h}, p, k),$$

- g_s^2 : Gluon radiation vertex $\to g_s^2 = 4\pi\alpha_s(Q)$
- In MARTINI, $Q \sim \langle |k_{\perp}| \rangle = (\hat{q}k)^{\frac{1}{4}}$ $\langle k_{\perp}^2 \rangle = \hat{q}l_{coh}$ $l_{coh} = k/\langle k_{\perp}^2 \rangle$
- g_s^2 in the integral equation : elastic scattering (soft) \rightarrow fixed

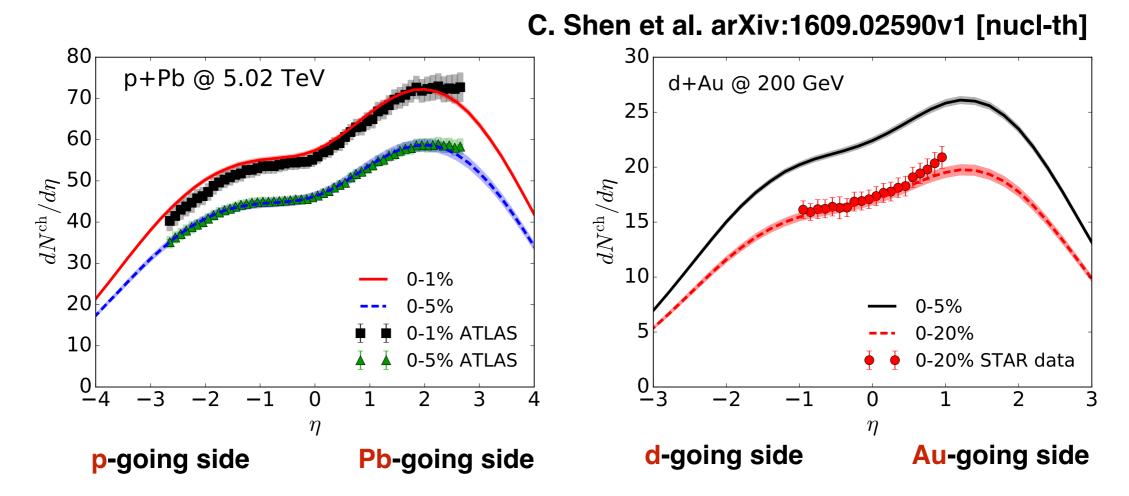


3+1D hydrodynamic simulations

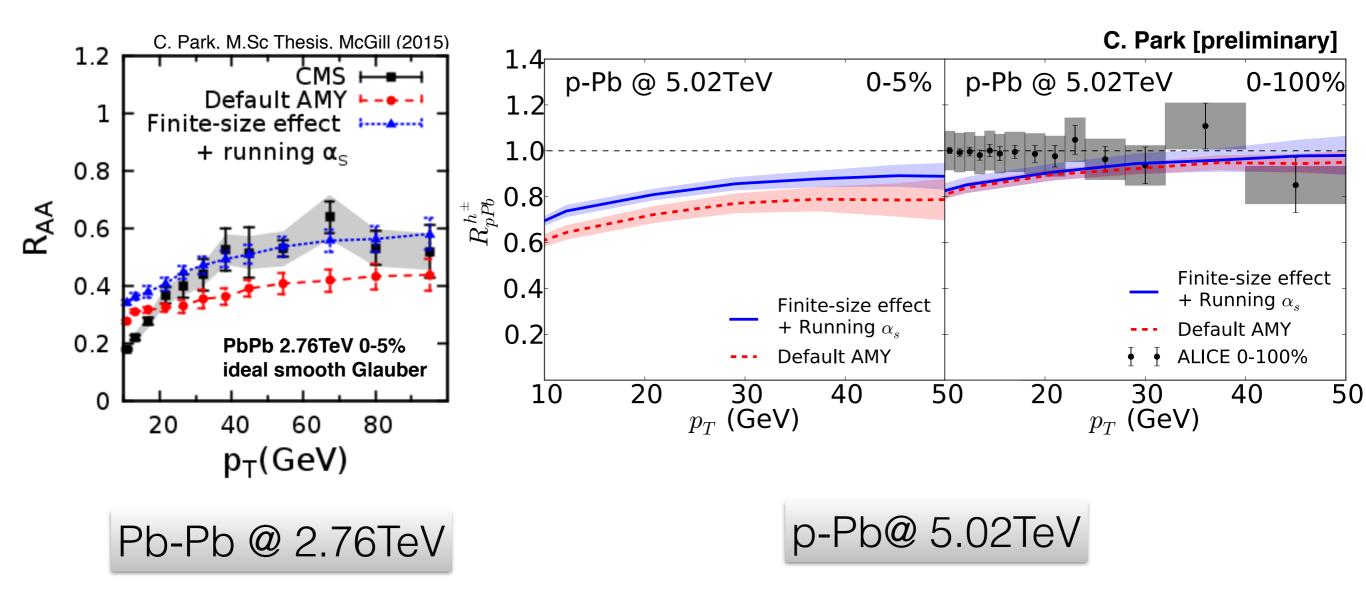
MUSIC

B. Schenke, S. Jeon, C. Gale, Phys. Rev. C82, 014903 (2009)

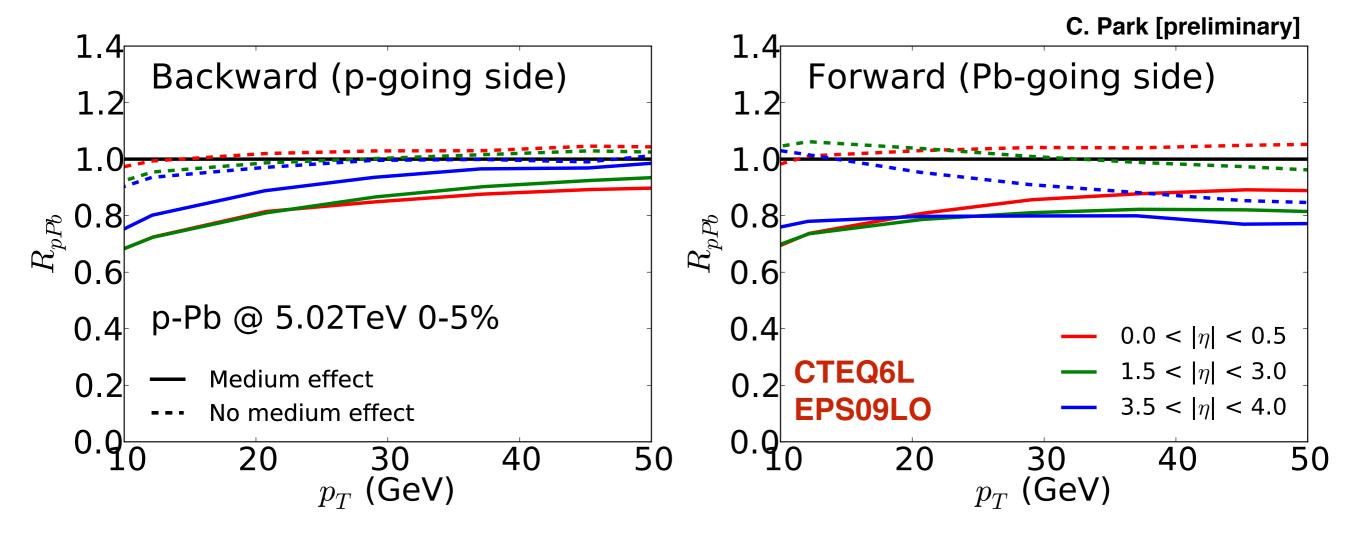
- 3+1D: Suitable for studying non-trivial longitudinal dynamics and rapidity dependence of jet observables
- Event-by-event simulation + MC Glauber initial condition



Finite-size effects & running coupling

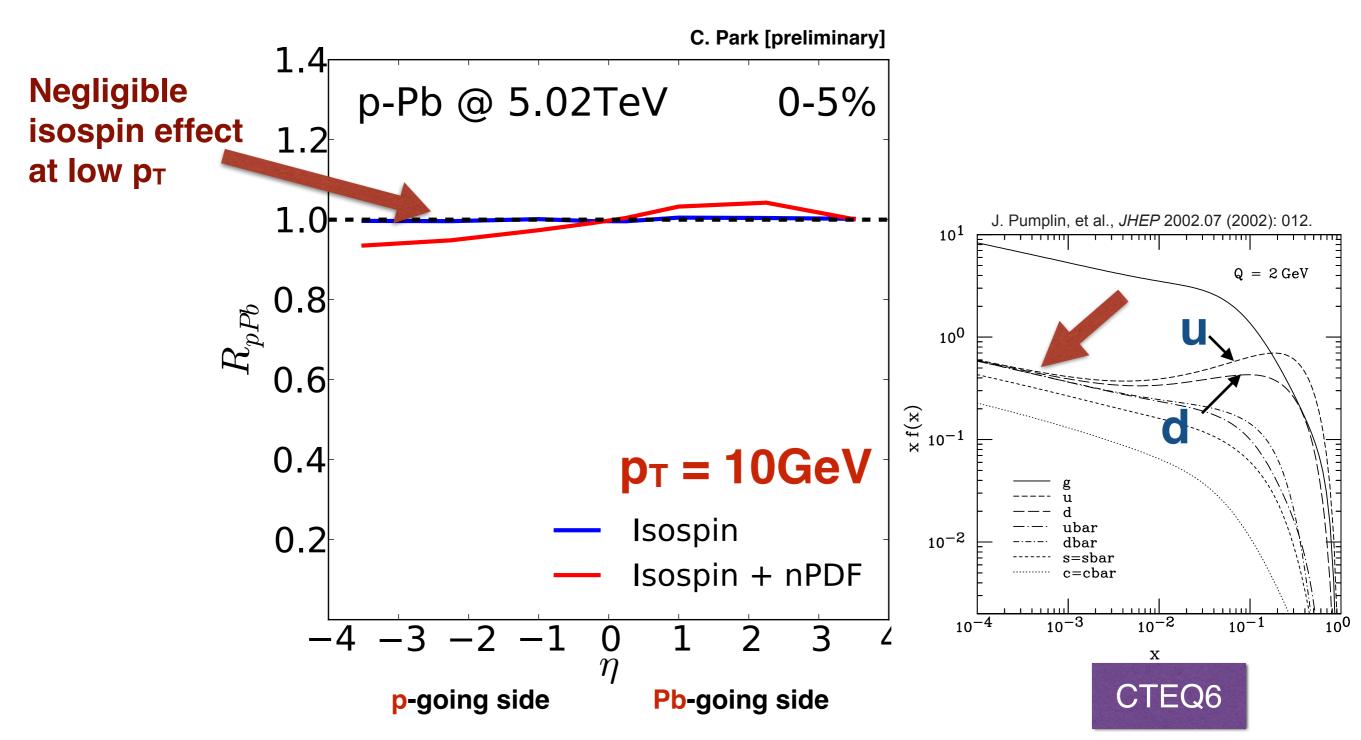


- Two effects give constant shifts in RAA & RpA
- For large systems, the new prescriptions were favoured, but underestimated ALICE data @ 10-20GeV in small systems.

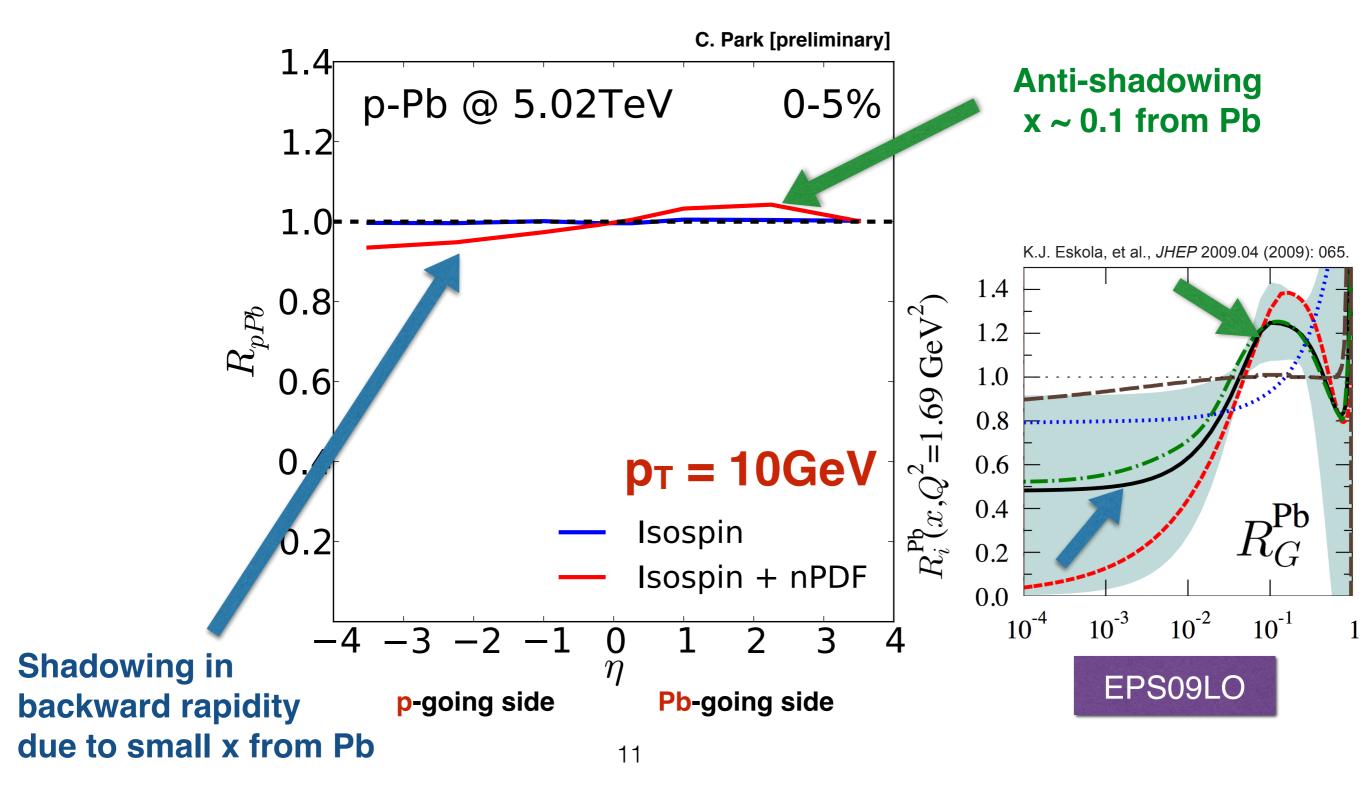


- Net effect : Convolution of medium effect + nuclear PDF + isospin
- Better understanding of their effects in rapidity directions is required.

Dominant nuclear PDF effect in low p_T R_{pPb}

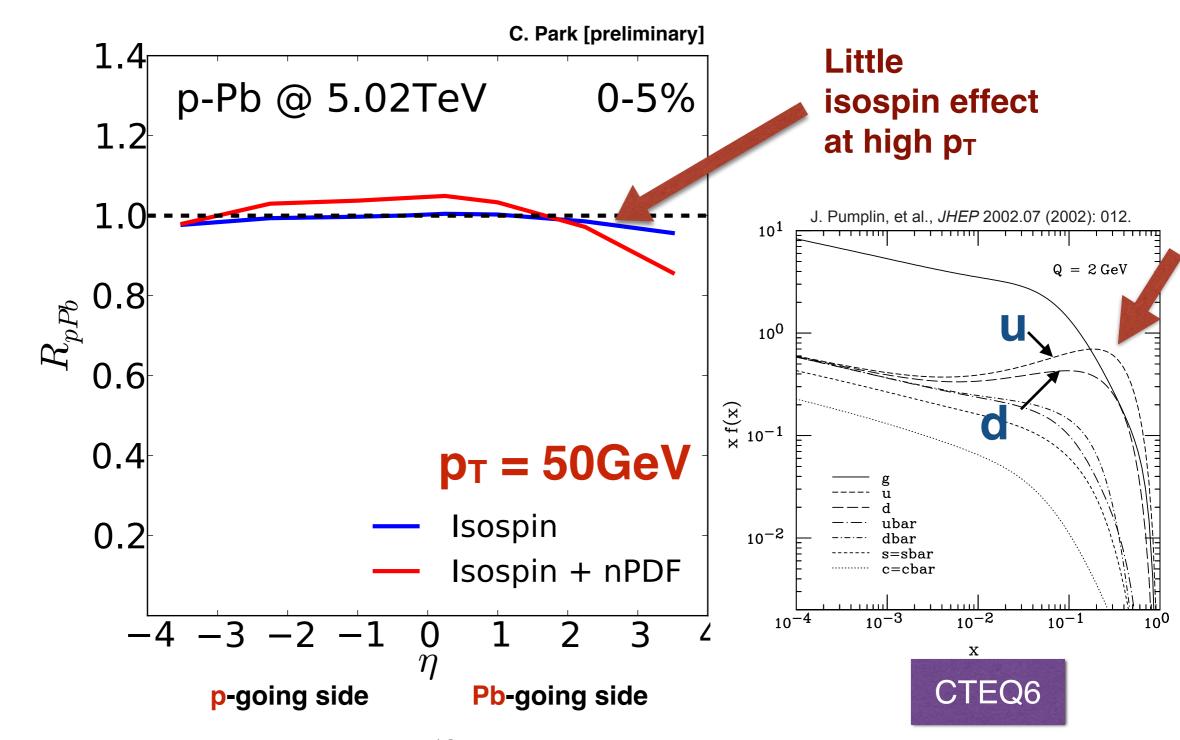


Dominant nuclear PDF effect in low p_T R_{pPb}



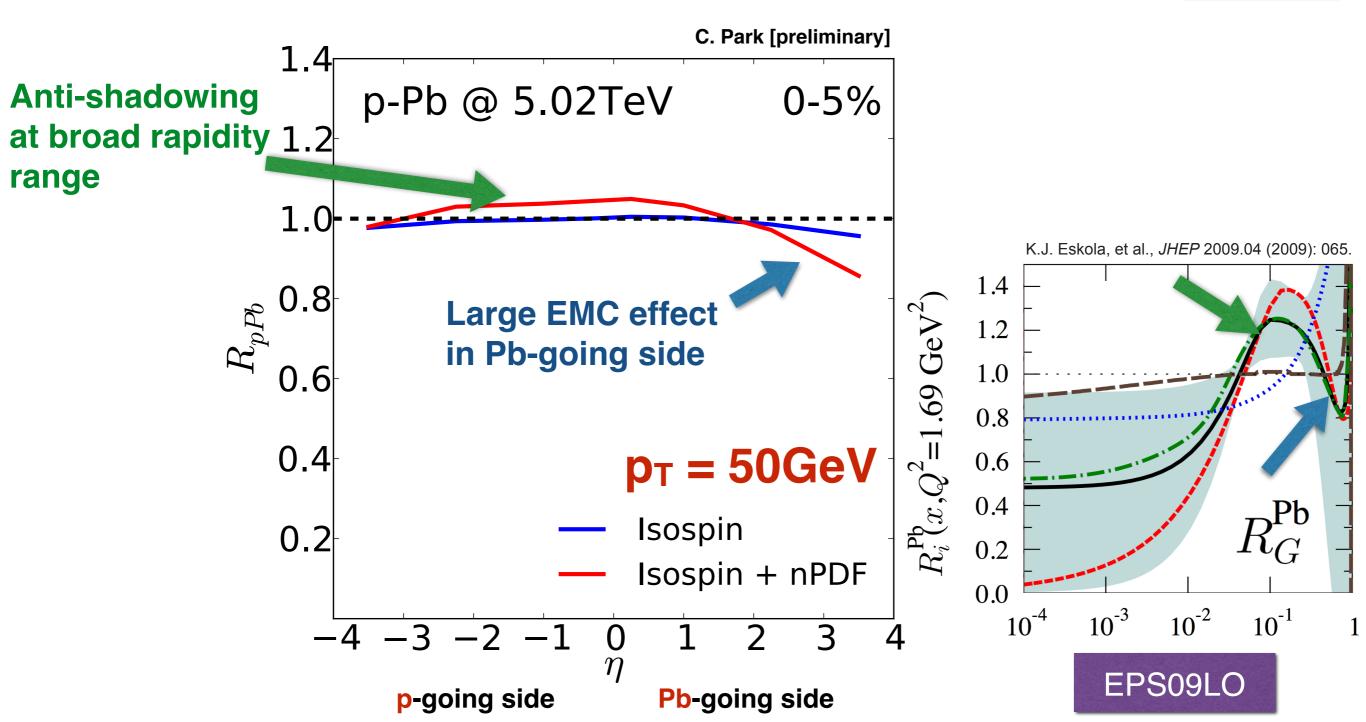
Isospin effect and anti-shadowing in high p_T R_{pPb}

high p_⊤ = larger x

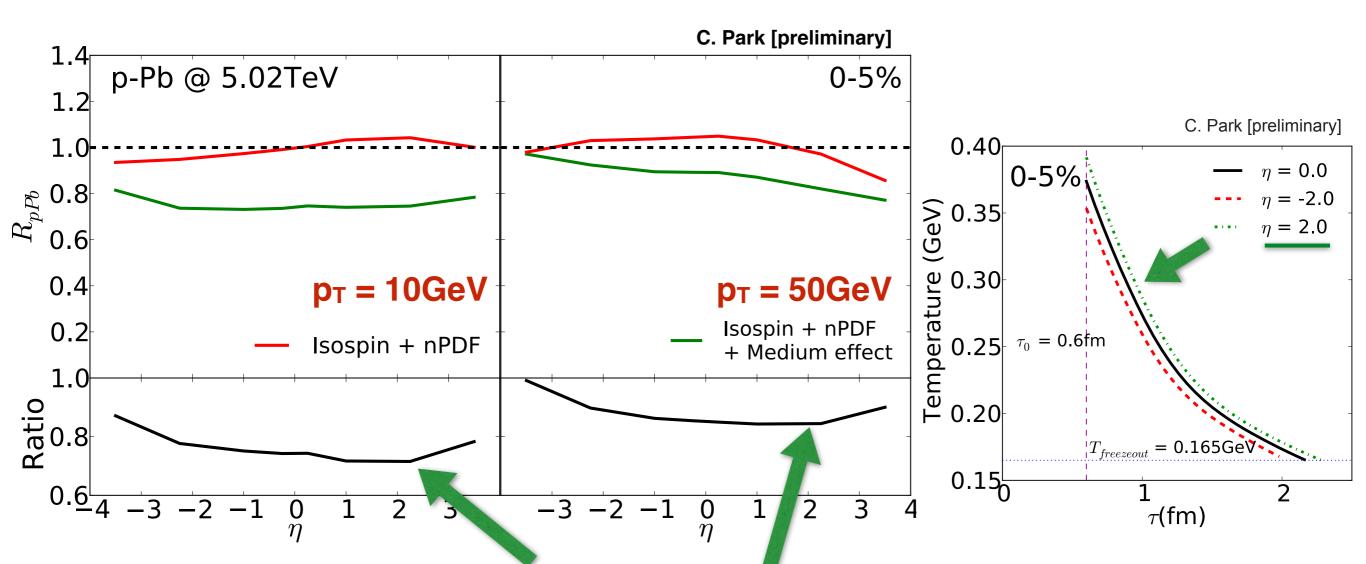


Isospin effect and anti-shadowing/EMC in high p_T R_{pPb}

high p_⊤ = larger x

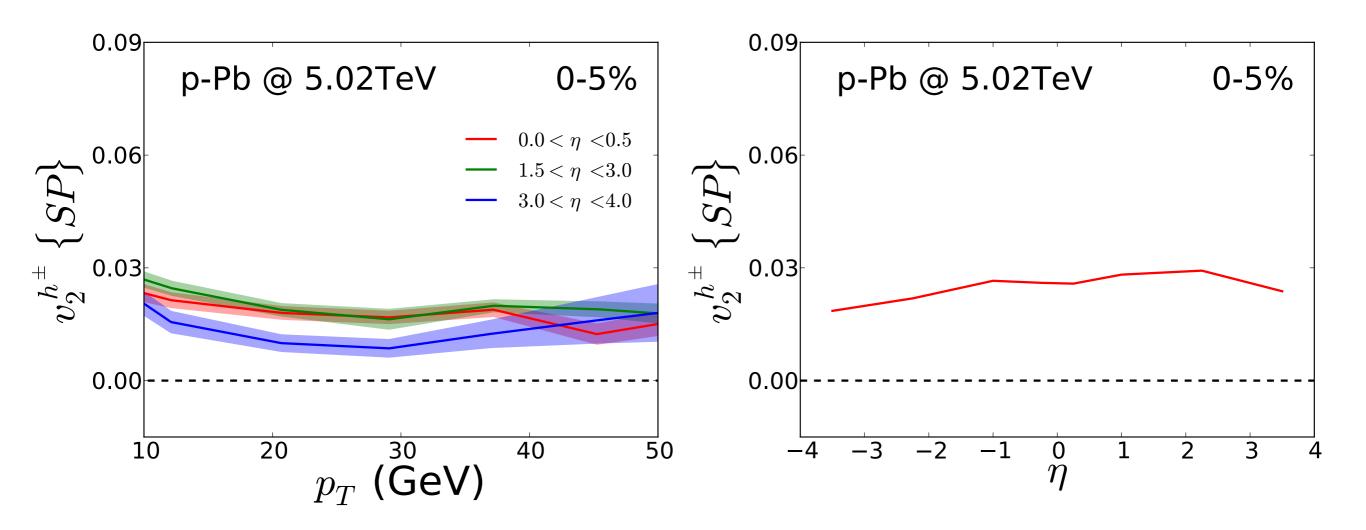


- Net QGP medium energy loss related to temperature and energy
 - a clear signature of QGP droplet in central collisions is predicted



Sizable medium induced energy loss due to high T

Charged hadron v₂ at high p_T



- Scalar product method; reference flow integrated from 0.3 3GeV
- 1~3% v₂ in central p-Pb collisions
- Medium induced energy loss dependency on integrated v₂

Conclusion

- Finite-size effect and running coupling for the radiation are essential to describe R_{AA} in large systems.
- A 3+1D medium in the small system offers the opportunity to study the key physics that governs different rapidity regimes.
 e.g. nPDF, isospin, and medium induced energy loss
- We predict sizable energy loss in central collisions of small systems due to the formation of the QGP droplet.
- We predict non-zero v₂ in central collisions of small systems having energy loss dependency.
 - Next work: jet reconstruction to find missing energy in small systems

Backup

AMY integral equation

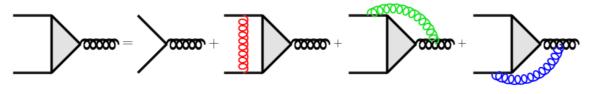


Image by G. Qin

Integral equation

$$2\mathbf{h} = i\delta E(\mathbf{h}, p, k)\mathbf{F}(\mathbf{h}) + g_s^2 \int \frac{d^2\mathbf{q}_{\perp}}{(2\pi)^2} C(\mathbf{q}_{\perp}) \left\{ (C_s - C_A/2)[\mathbf{F}(\mathbf{h}) - \mathbf{F}(\mathbf{h} - k\mathbf{q}_{\perp})] + (C_A/2)[\mathbf{F}(\mathbf{h}) - \mathbf{F}(\mathbf{h} + P\mathbf{q}_{\perp}) + (C_s - C_A/2)[\mathbf{F}(\mathbf{h}) - \mathbf{F}(\mathbf{h} - (p - k)\mathbf{q}_{\perp})] \right\}$$

Differential rate to exchange q⊥

Energy difference δE

$$C(\mathbf{q}_{\perp}) = \frac{m_D^2}{\mathbf{q}_{\perp}^2(\mathbf{q}_{\perp}^2 + m_D^2)}, \quad m_D^2 = \frac{g_s^2 T^2}{6} (2N_c + N_f).$$

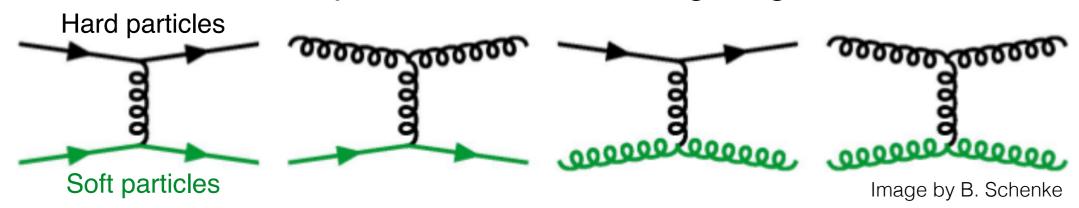
$$\delta E(\mathbf{h},p,k) = rac{\mathbf{h}^2}{2pk(p-k)} + rac{m_k^2}{2k} + rac{m_{p-k}^2}{2(p-k)} - rac{m_p^2}{2p}$$

AMY radiative
$$\frac{d\Gamma(p,k)}{dk} = \frac{C_s g_s^2}{16\pi p^7} \frac{1}{1 \pm e^{-k/T}} \frac{1}{1 \pm e^{-(p-k)/T}} \times \begin{cases} \frac{\frac{1+(1-x)^2}{x^3(1-x)^2}}{x^3(1-x)^2} & q \to qg \\ N_f \frac{x^2+(1-x)^2}{x^2(1-x)^2} & g \to qq \\ \frac{1+x^4+(1-x)^4}{x^3(1-x)^3} & g \to gg \end{cases}$$

$$\times \int \frac{d^2\mathbf{h}}{(2\pi)^2} 2\mathbf{h} \cdot \operatorname{Re} \; \mathbf{F}(\mathbf{h}, p, k),$$

Collisional process

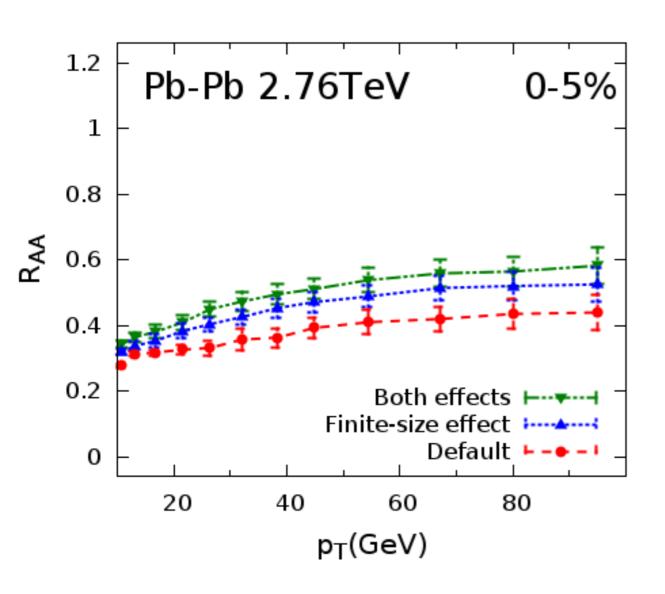
All the possible 2-2 scattering diagrams

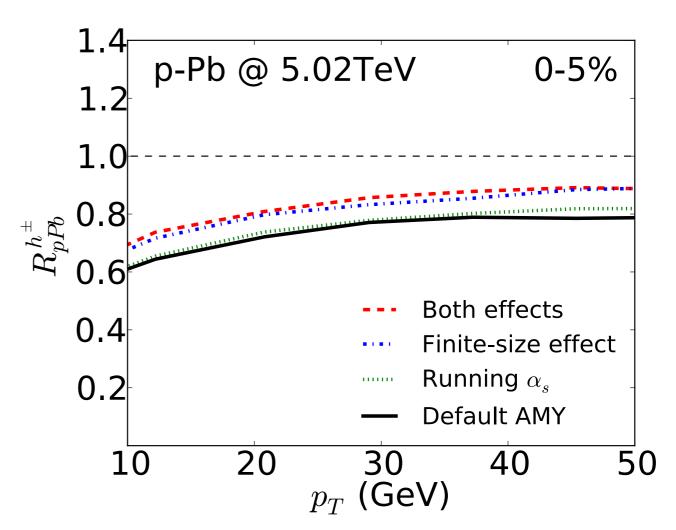


Matrix elements

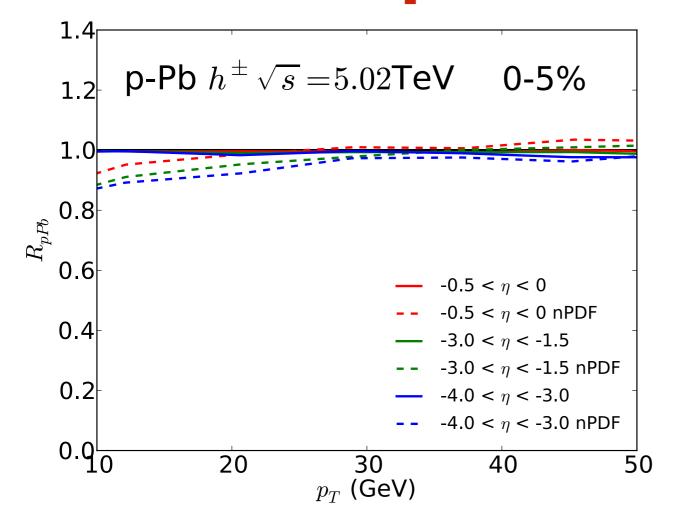
$$|\mathcal{M}|_{qq}^{2} = \frac{4}{9}g^{4}\frac{s^{2} + u^{2}}{t^{2}}, \qquad |\mathcal{M}|_{qg}^{2} = 2g^{4}\left(1 - \frac{su}{t^{2}}\right), |\mathcal{M}|_{gq}^{2} = 2g^{4}\left(1 - \frac{su}{t^{2}}\right), \qquad |\mathcal{M}|_{gg}^{2} = \frac{9}{2}g^{4}\left(\frac{17}{8} - \frac{su}{t^{2}}\right)$$

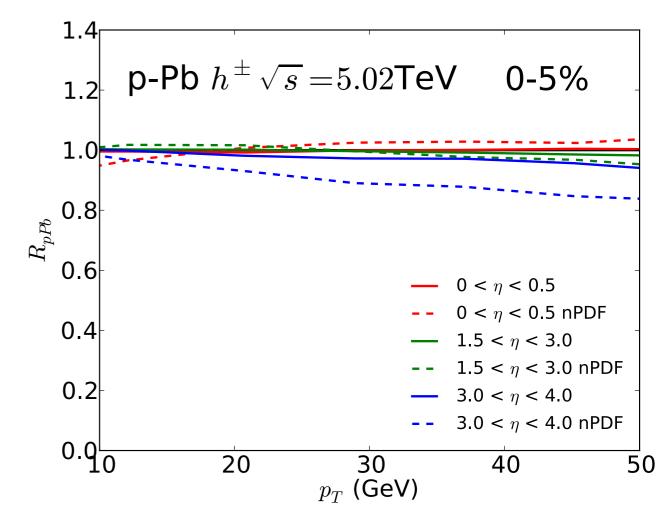
Finite-size effects & running coupling





Isospin & nuclear PDF





p-going side

Pb-going side