

# Review of Model Predictions for Extensive Air Showers

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**UHECR 2016, Kyoto, Japan**

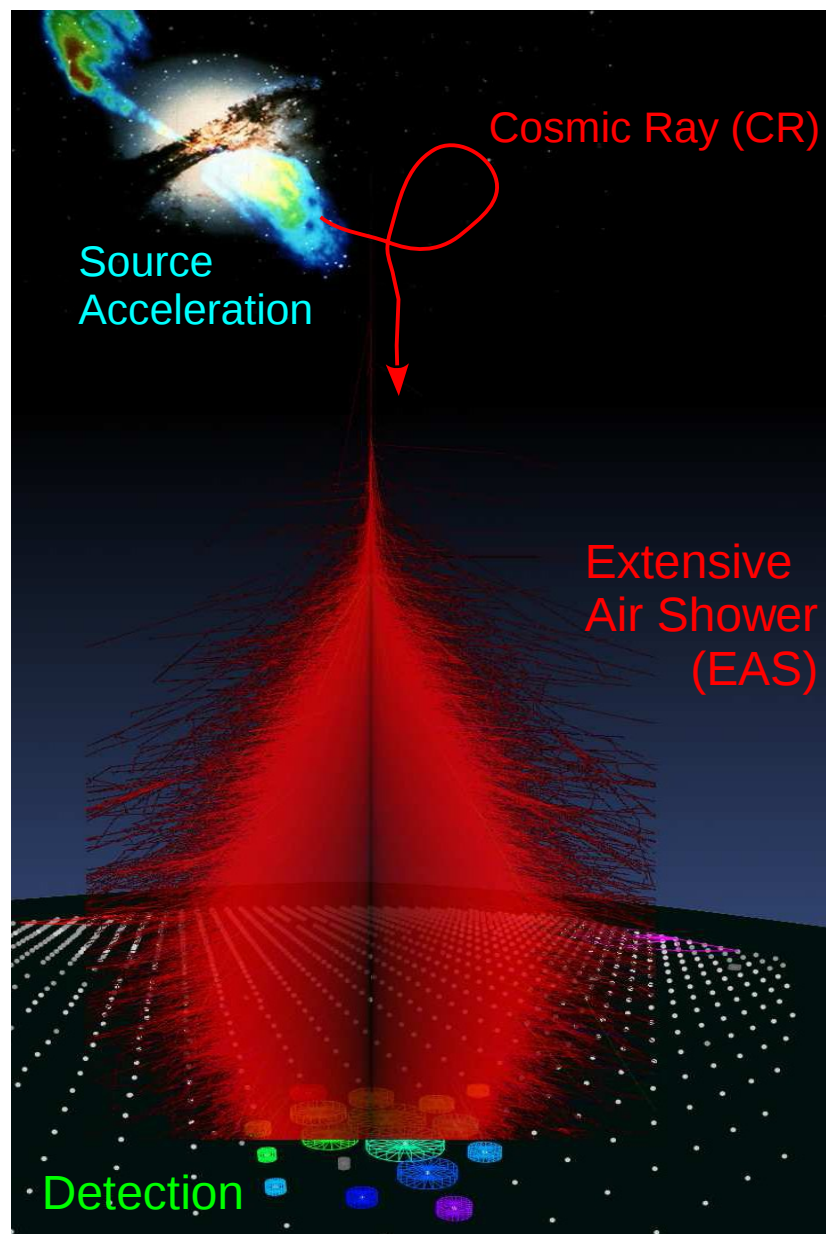
October the 12<sup>th</sup> 2016

# Outline

- Monte-carlo for Cosmic Ray analysis
- MC comparison to accelerator data
- Electromagnetic (EM) signal in extended air showers
  - ➔ source of uncertainties
- Muon signal

LHC data reduced the model uncertainties and **exclude old models** for mass composition of cosmic rays. **Remaining uncertainties** can be further reduced taking into account **forward measurements** AND using (light) **nuclear target**.

# Preamble



From R. Ulrich (KIT)

- **Goal of Astroparticle Physics :**
  - ➔ astronomy with high energy particles
  
- **How to test hadronic interactions ?**
  - ➔ if the source mechanism is well understood we could have a known beam at ultra-high energy ( $10^6$  GeV and more)
    - ➔ improving but not very precise
  - ➔ reasonable minimum limits from CR abundance :
    - ◆ low = hydrogen (proton)
    - ◆ high = iron ( $A=56$ )
  - ➔ test of hadronic interactions in EAS via correlations between observables.

**mass measurements should be consistent and lying between proton and iron simulated showers if physics is correct**

# Hadronic Interaction Models in CORSIKA

(HDPM)

**Old generation :** (SIBYLL 2.1 QGSJET01 DPMJET 2.55 VENUS) (<2001)

**All Glauber based**  
**But differences in hard, remnants, diffraction ...**

**New (!) generation :**

**LHC tuned :**

**LHC inspired :** **SIBYLL 2.3**

**Motivation :**  
- update with latest LHC results in simple model

Engel et al.

semi-hard

(QGSJET II-03)

**QGSJET II-04**

Ostapchenko

**QGSJET III (?)**

**Motivation :**  
- Hard Pomeron-Pomeron connexion

soft

**NEXUS 3.97**

**Attempt to get everything described in a consistent way (energy sharing)**

(EPOS 1.99) (2005-2012)

**EPOS LHC** (2013-)

Pierog & Werner

**EPOS 3** (2016-)

**Motivation :**  
- binary scaling in hard probes

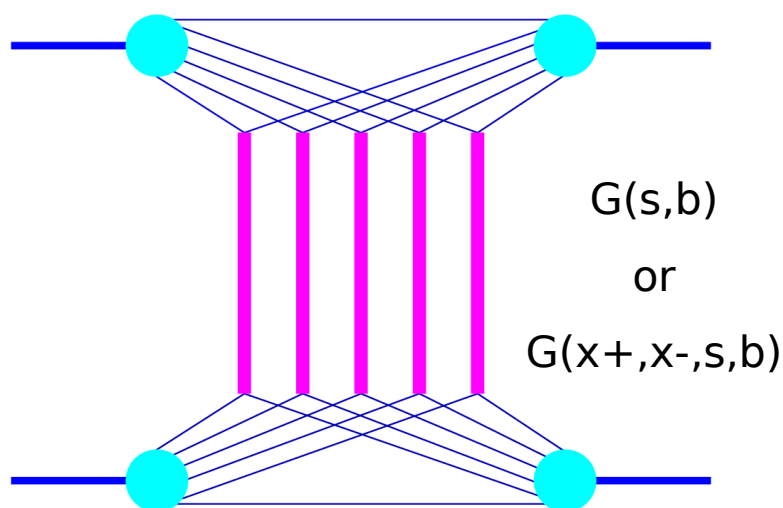
Fedinitch & Engel

**DPMJET III**

**Motivation :**  
- update with LHC results  
- fix high energy

Riehn & Engel

# Cross Section and Multiplicity in Models



## ● Gribov-Regge and optical theorem

- ➔ Basis of all models (multiple scattering) but
  - ◆ Classical approach for QGSJET, SIBYLL and DPMJET (no energy conservation for cross section calculation)
  - ◆ Parton based Gribov-Regge theory for EPOS (**energy conservation at amplitude level**)

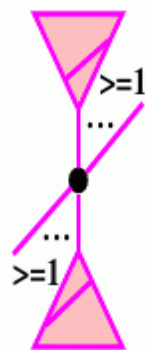
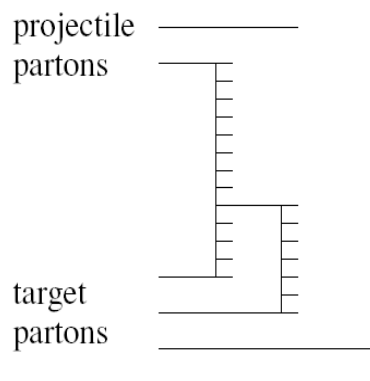
## ● pQCD

- ➔ Minijets with cutoff in SIBYLL and DPMJET
- ➔ Same hard Pomeron (DGLAP convoluted with soft part : no cutoff) in QGSJET and EPOS but
  - ◆ Generalized enhanced diagram in QGSJET-II
  - ◆ Simplified non linear effect in EPOS

- Phenomenological approach

EPOS

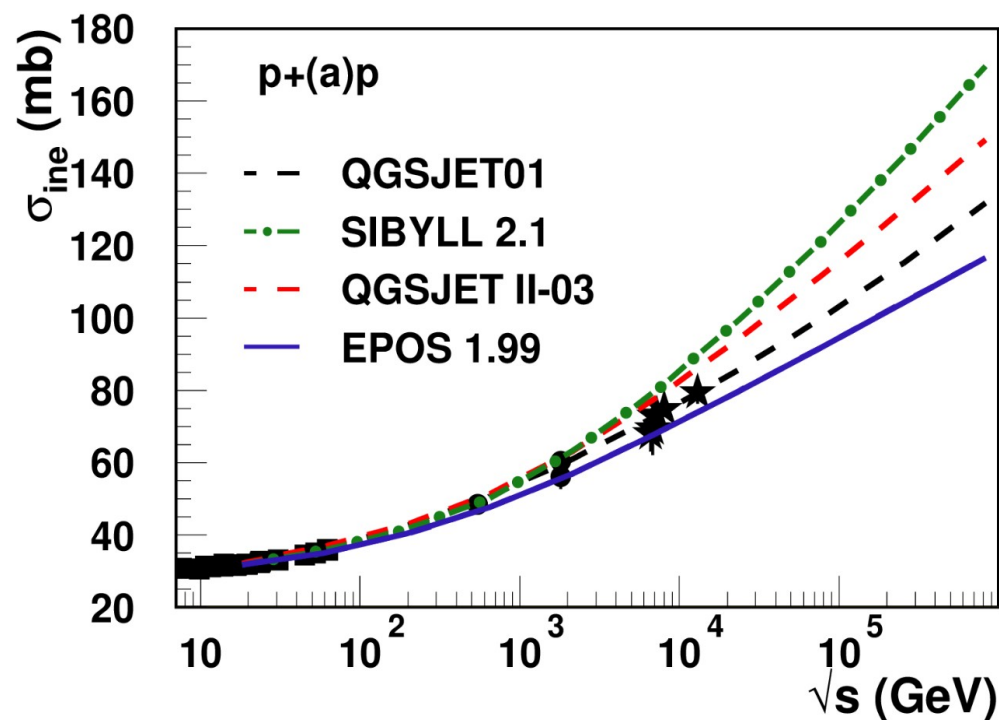
QGSJET II



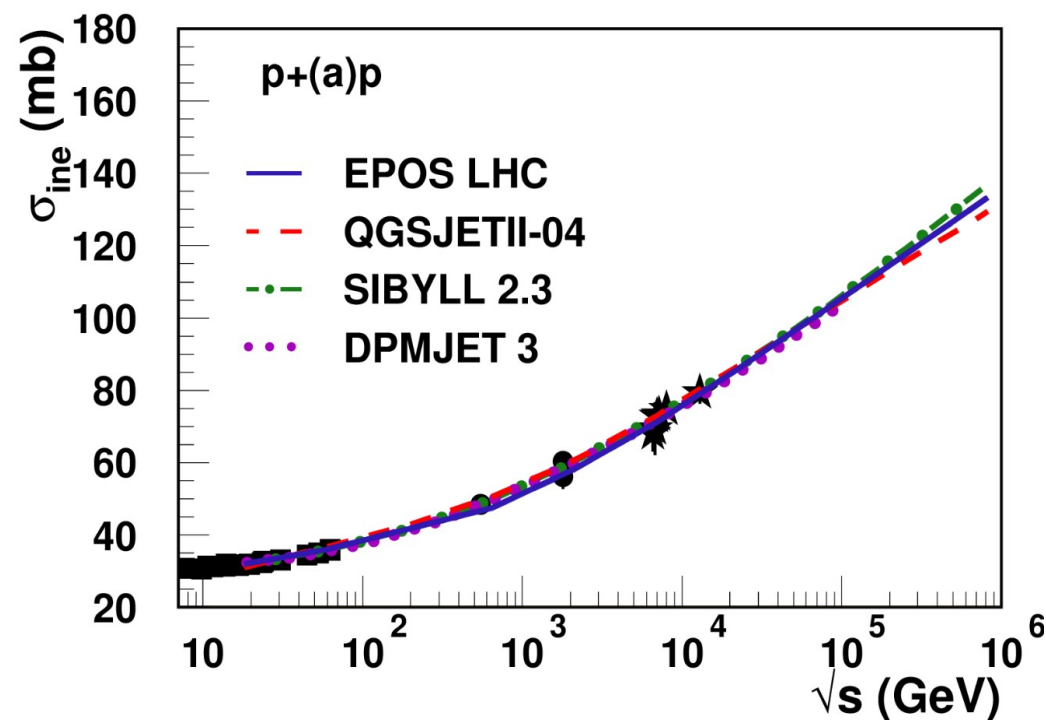
# Cross Sections

- ➔ Same cross section prediction at pp level and low energy (data for tuning)
- ➔ extrapolation to high energy looks settled
  - ◆ different amplitude and scheme
  - ➔ same extrapolations

Pre - LHC



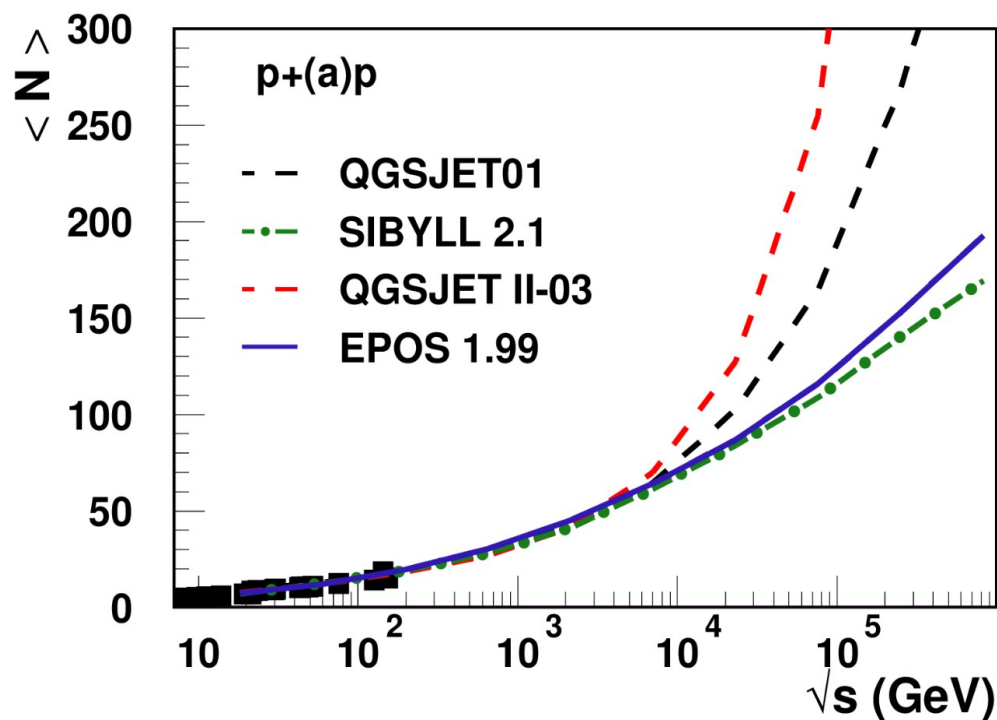
Post - LHC



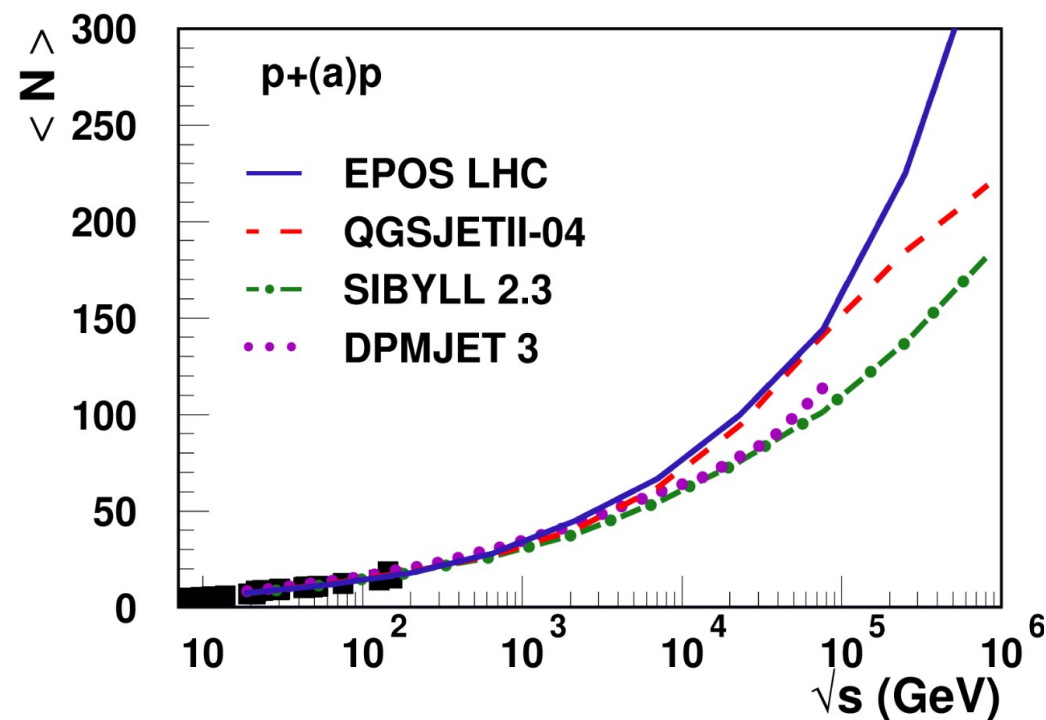
# Multiplicity

- Multiplicity fixed by data up to 900 GeV
- extrapolation to high energy is still model dependent ?

Pre - LHC



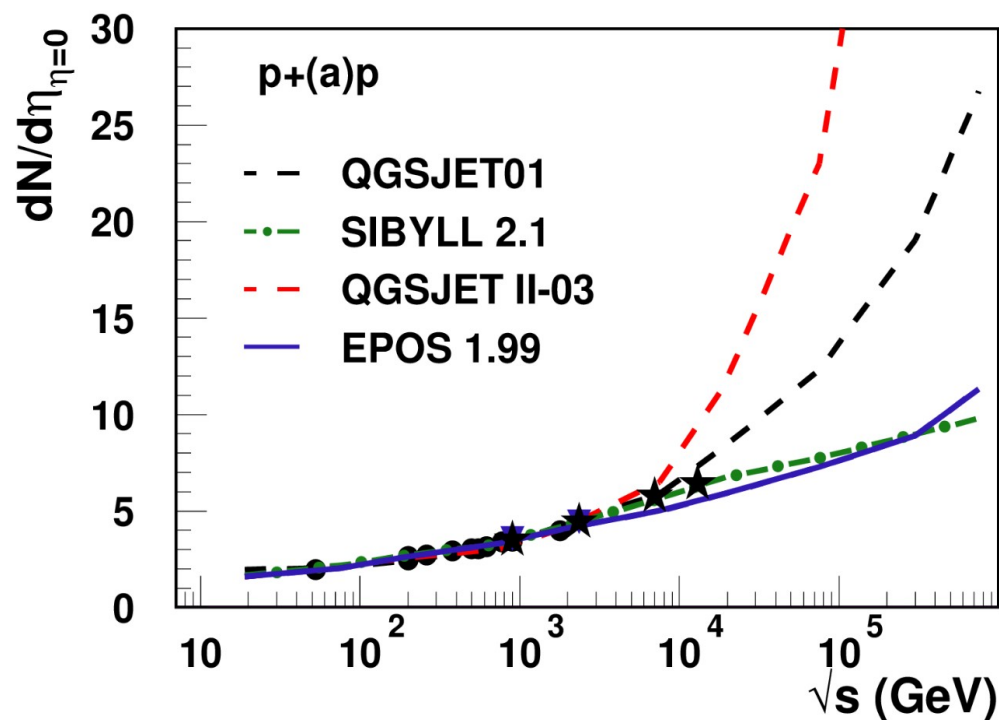
Post - LHC



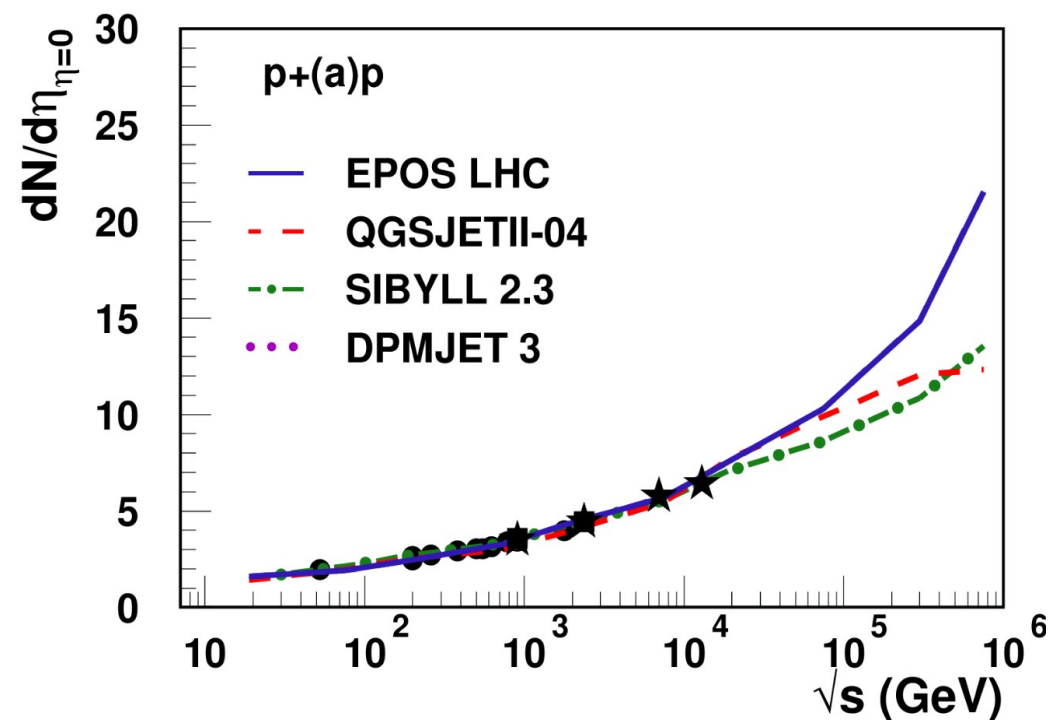
# Multiplicity at mid-rapidity

- Multiplicity fixed by data up to 13 TeV
- extrapolation to high energy less model dependent after LHC
- QGSJET01 and QGSJETII-03 extrapolation excluded

Pre - LHC



Post - LHC

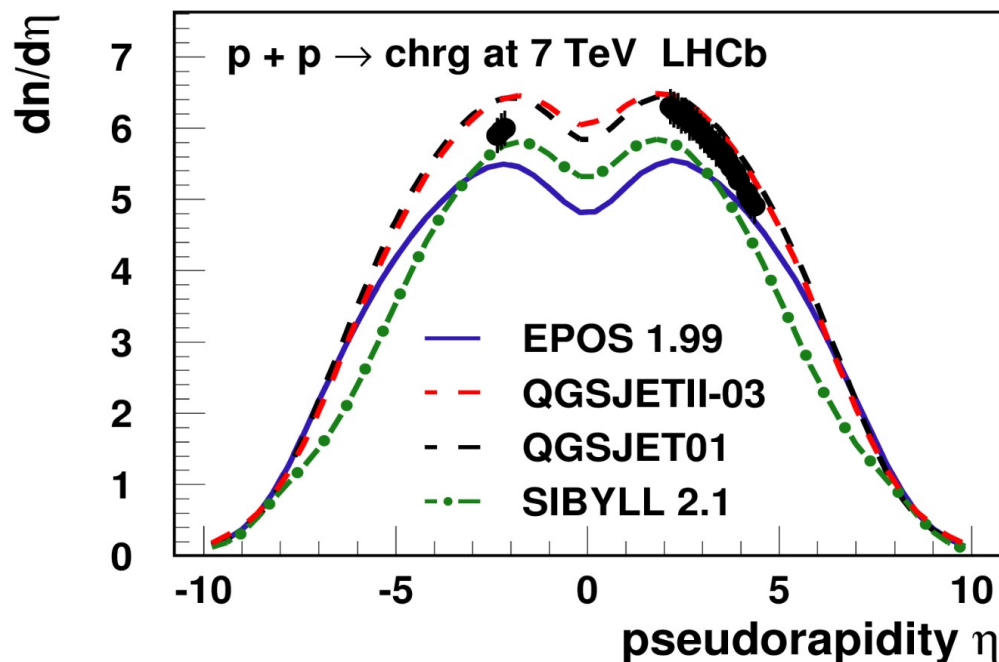




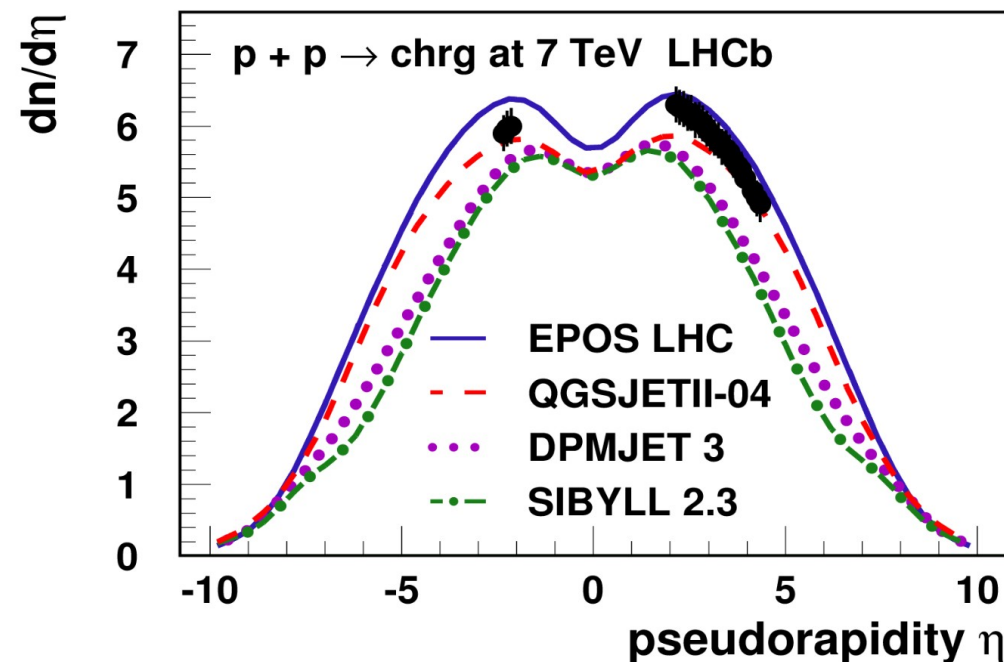
# Pseudorapidity

- ➔ Difference between mid-rapidity and full multiplicity coming from the width of the pseudorapidity distributions
- ➔ From LHC data
  - DPMJET 3 and SIBYLL 2.3 too narrow
  - QGSJETII-04 ~ OK
  - EPOS LHC a bit too large

Pre - LHC



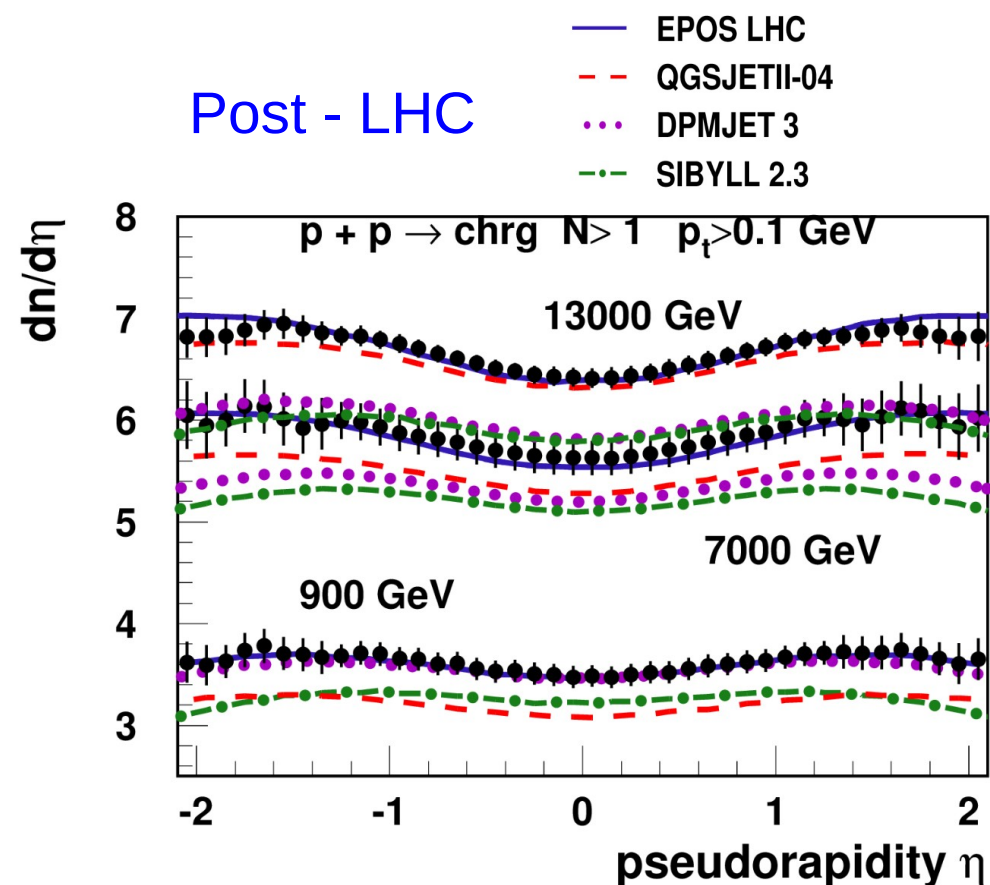
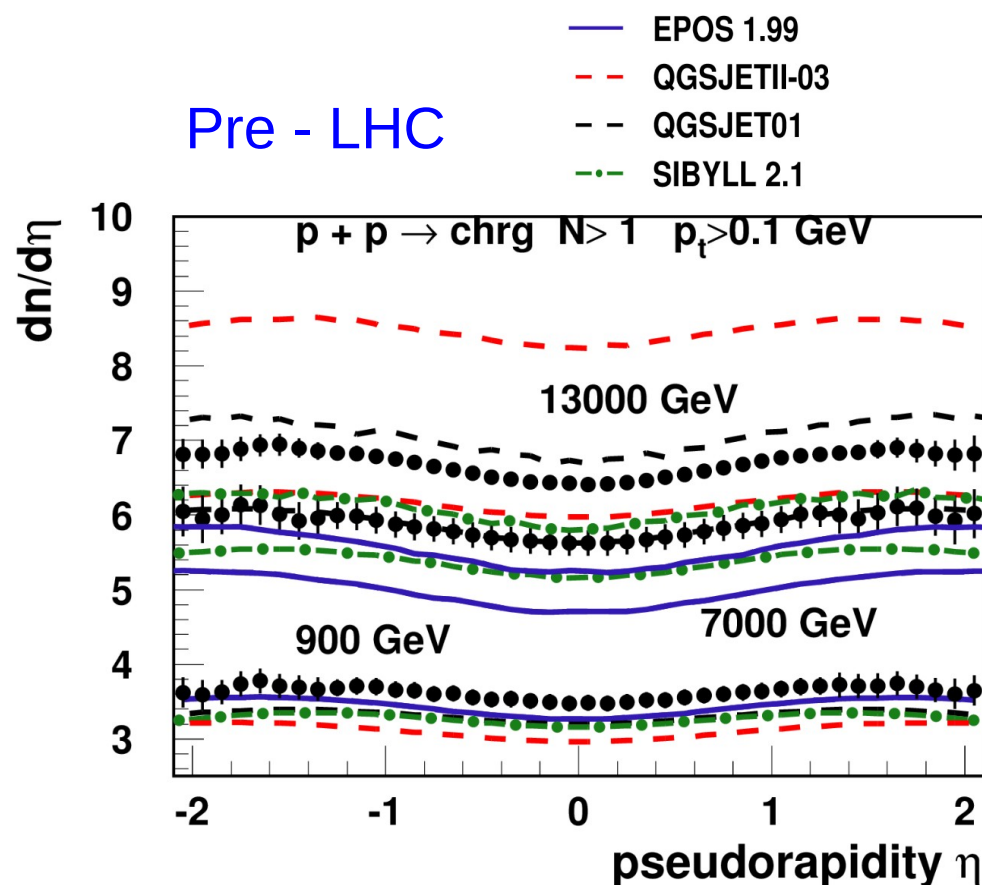
Post - LHC



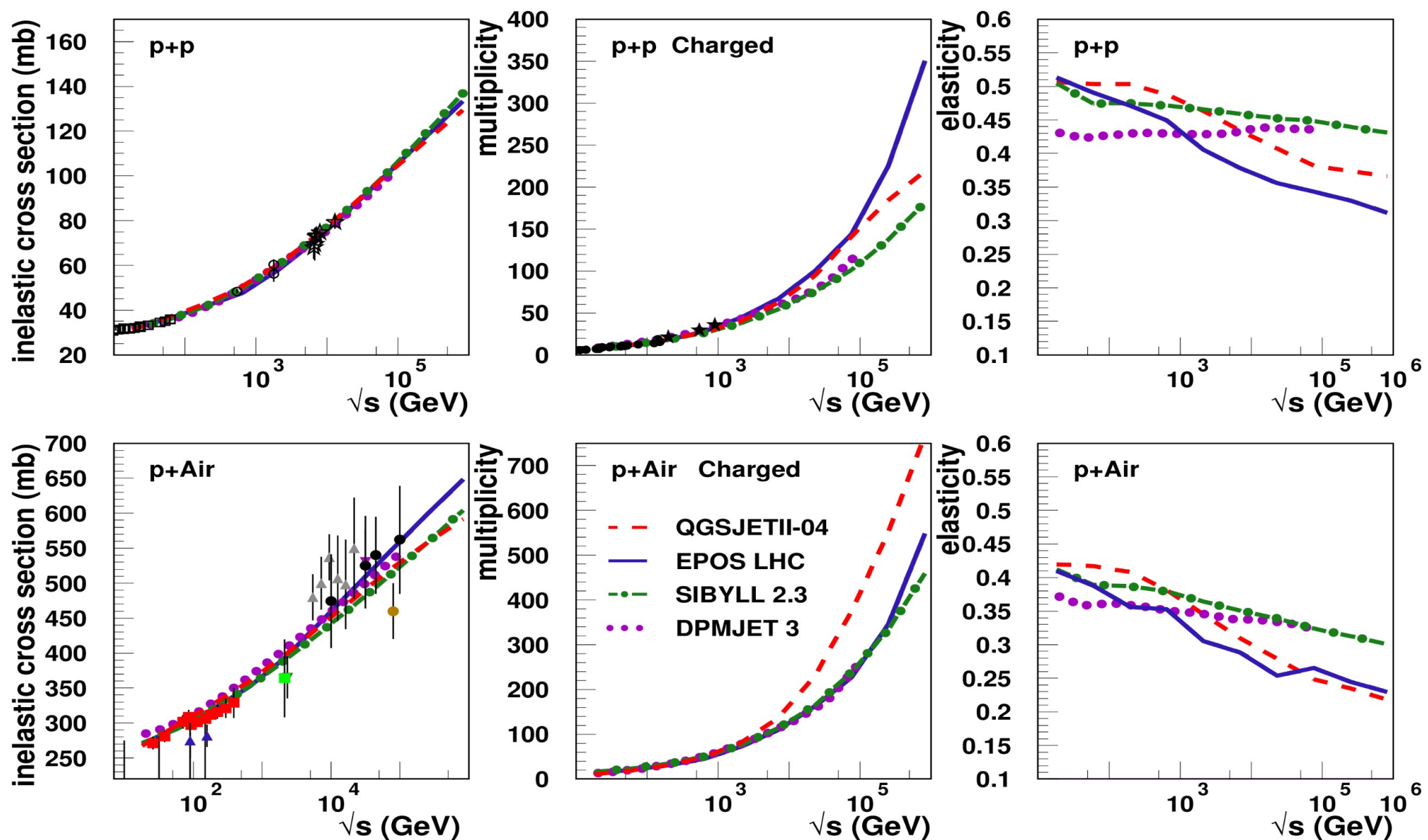
# Test of Models vs Accelerator Data

→ From LHC data

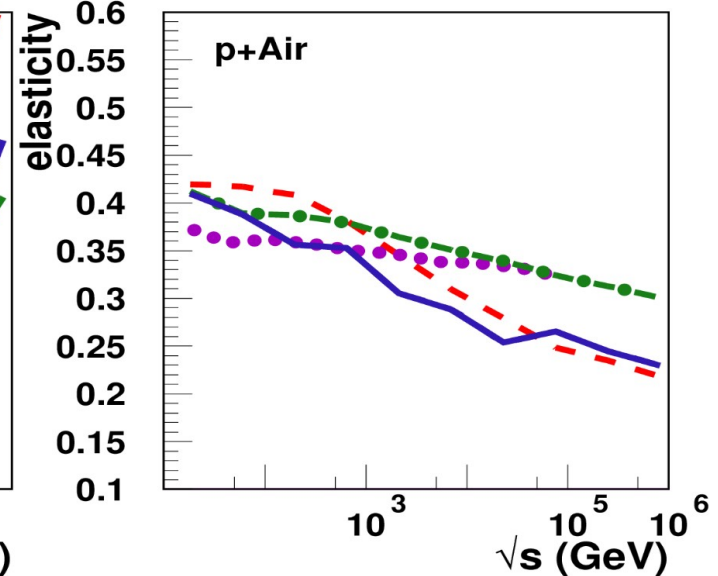
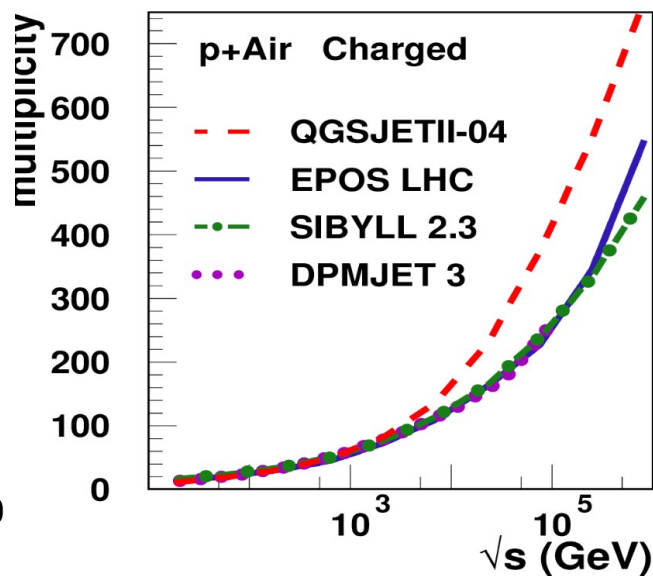
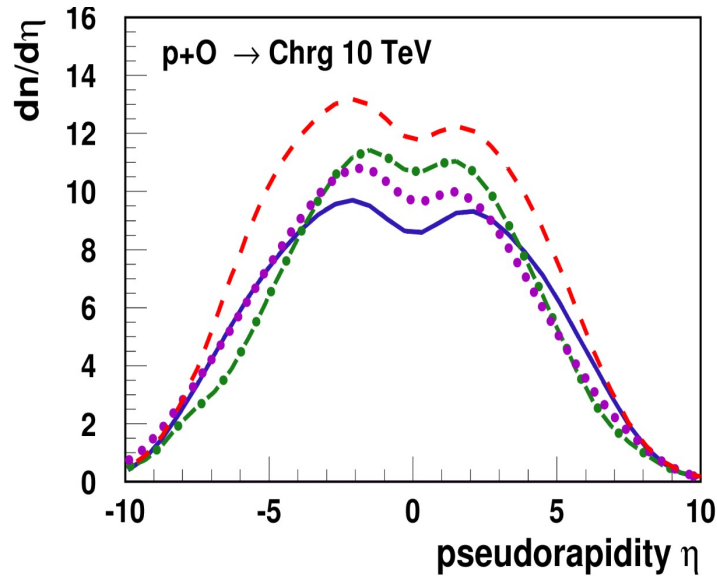
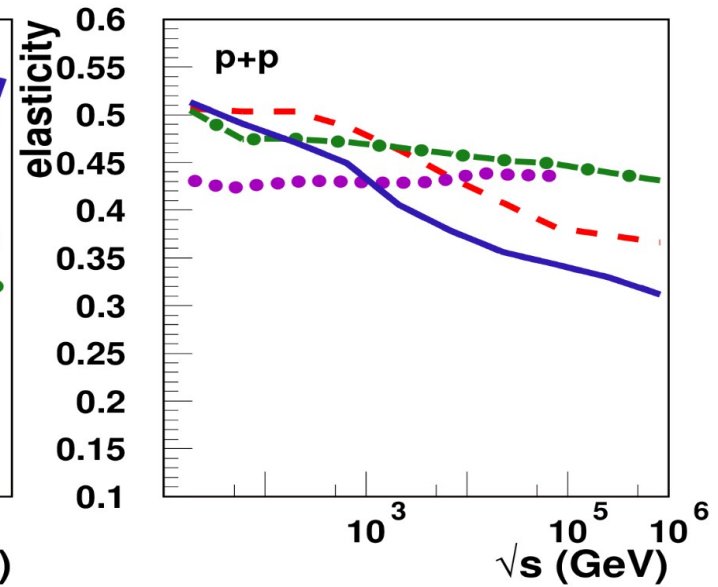
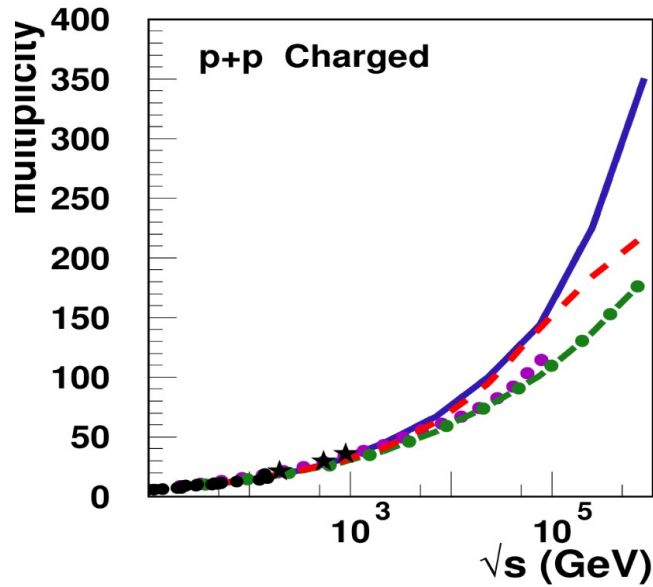
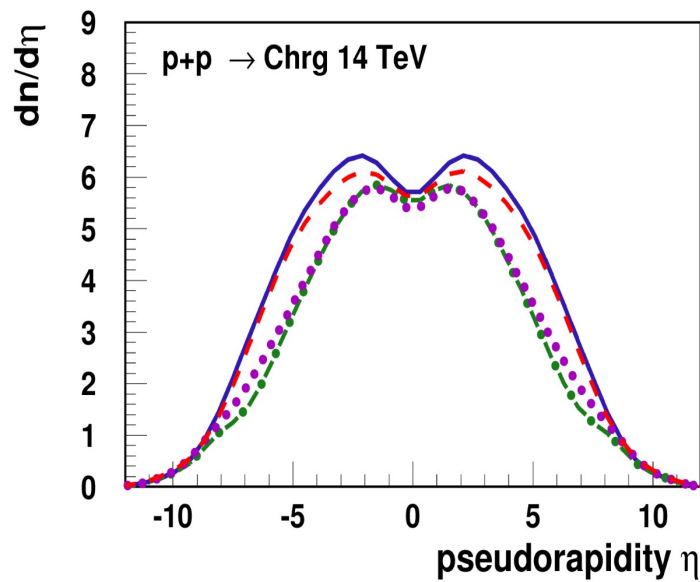
- All pre-LHC models extrapolation excluded
- DPMJET 3 and SIBYLL 2.3 underestimate multiplicity
- QGSJETII-04 and EPOS LHC ~ OK (and similar to Pythia 8)



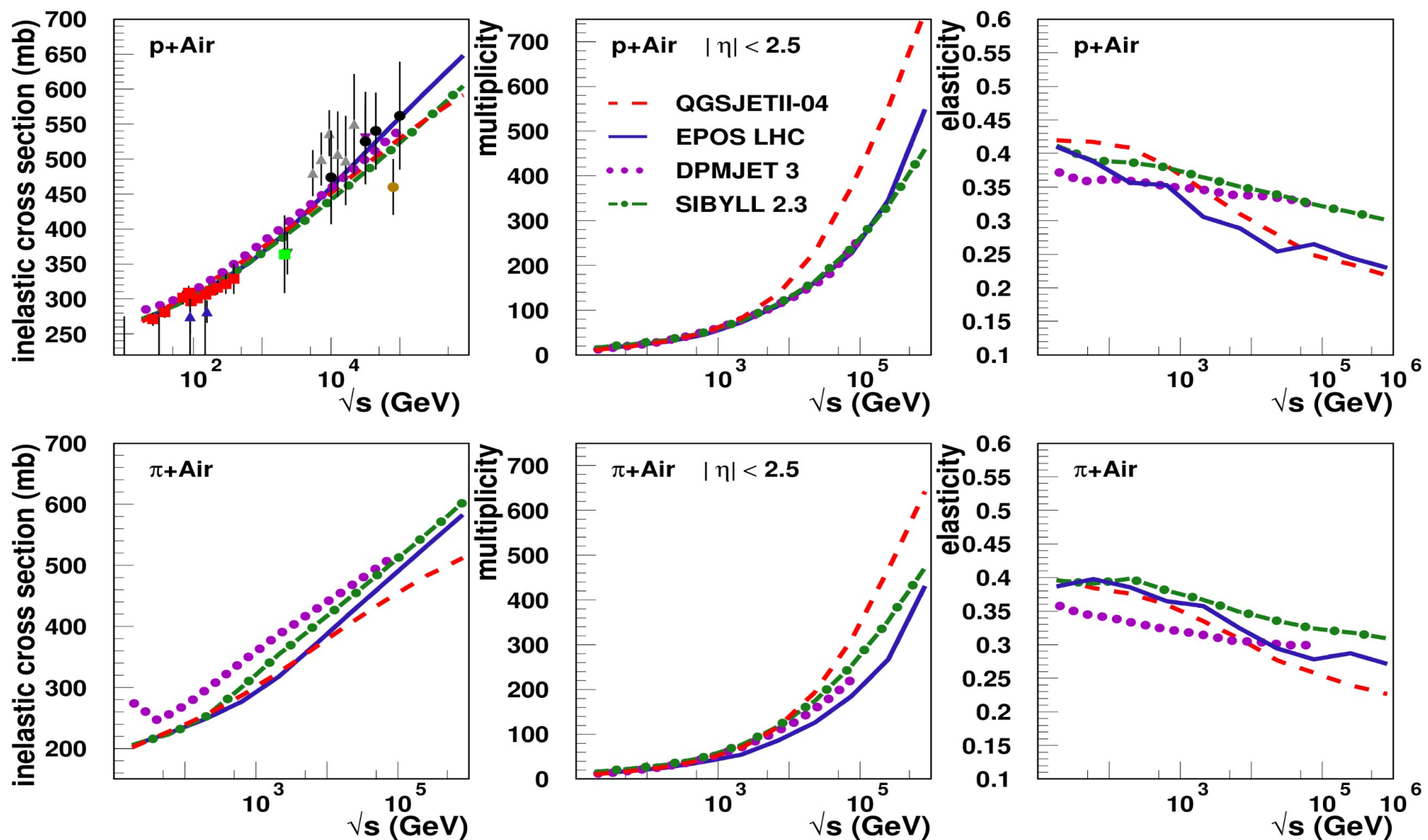
# Ultra-High Energy Hadronic Model Predictions p-Air



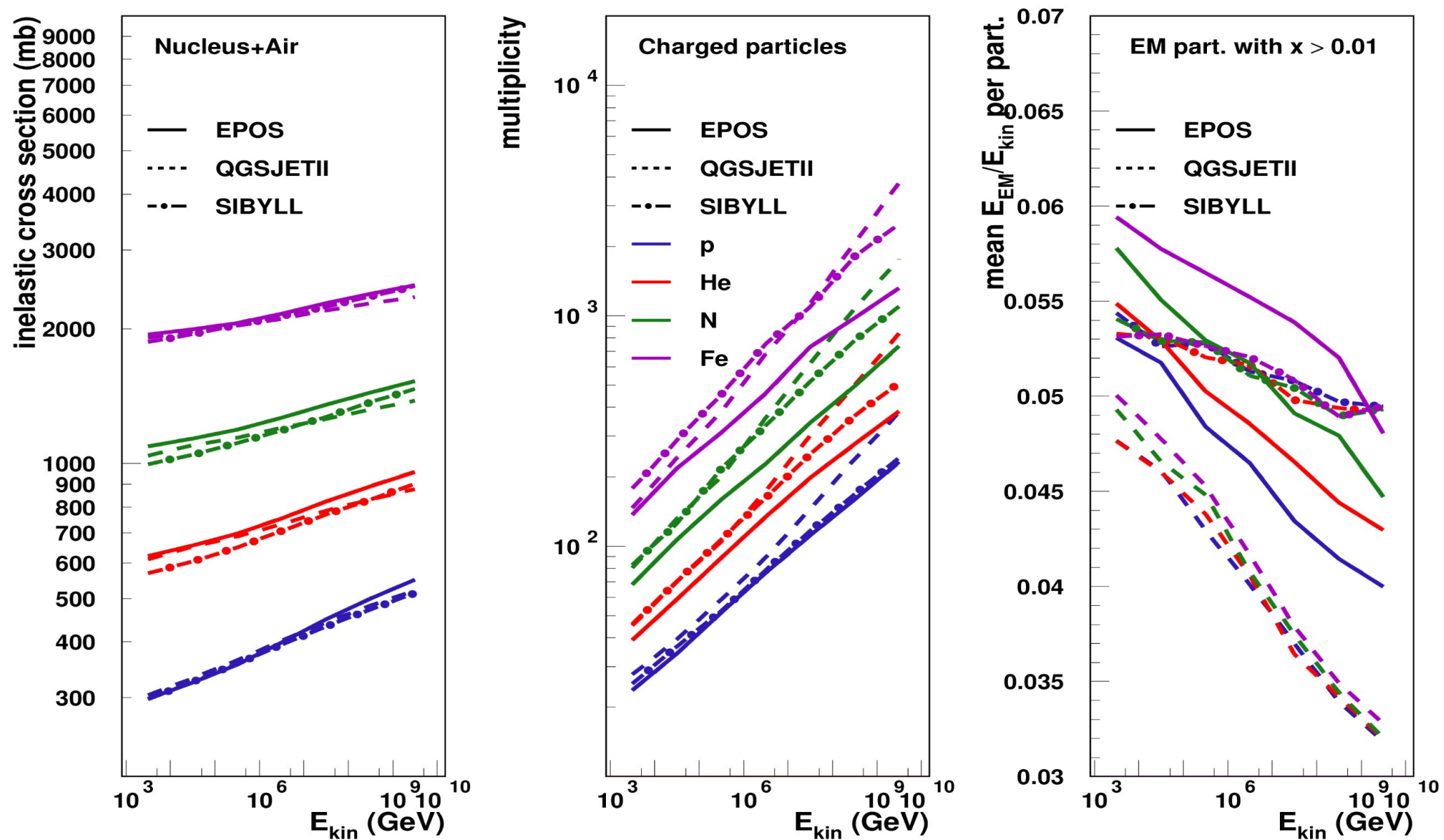
# Ultra-High Energy Hadronic Model Predictions p-Air



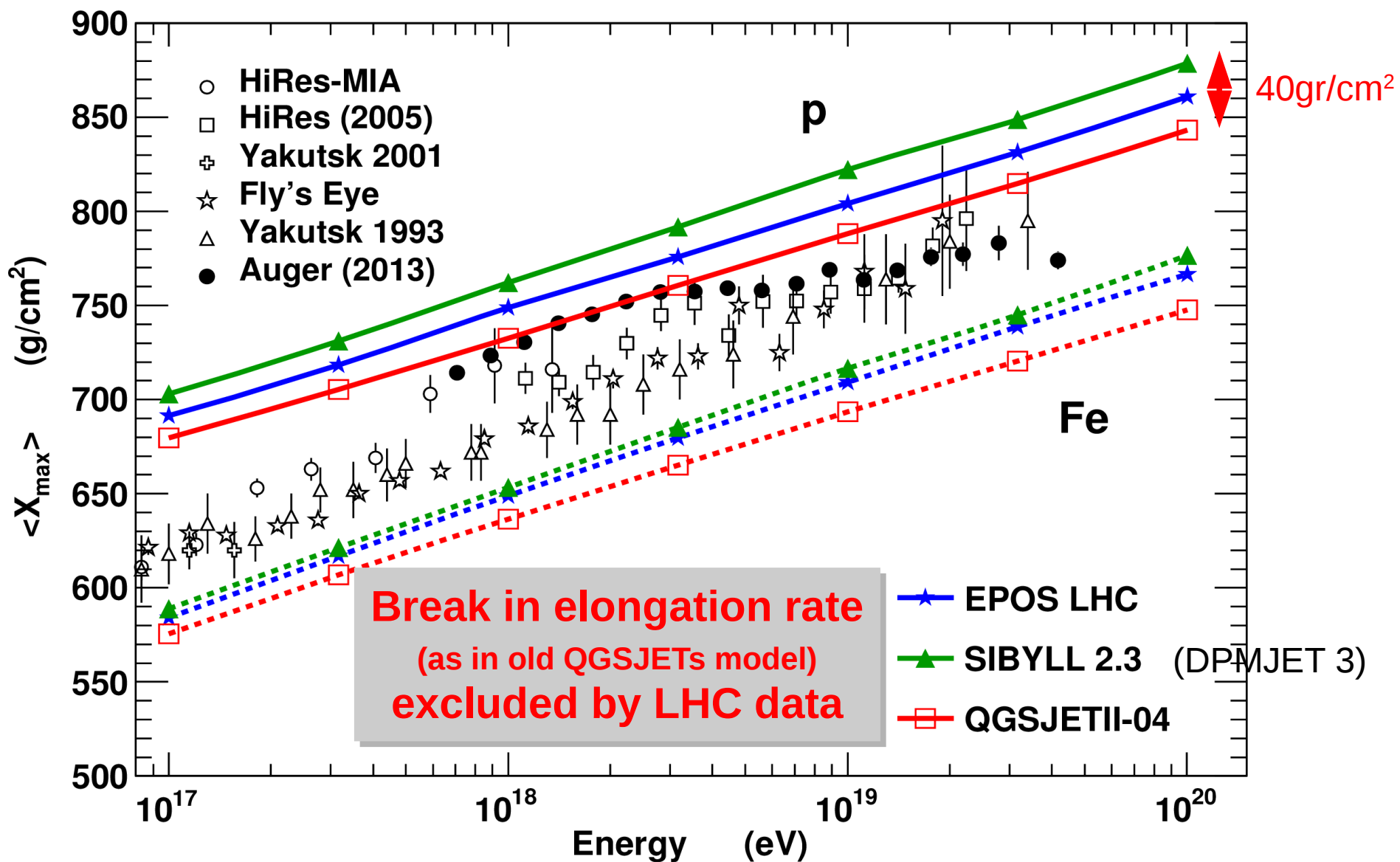
# Ultra-High Energy Hadronic Model Predictions $\pi$ -Air



# Ultra-High Energy Hadronic Model Predictions A-Air



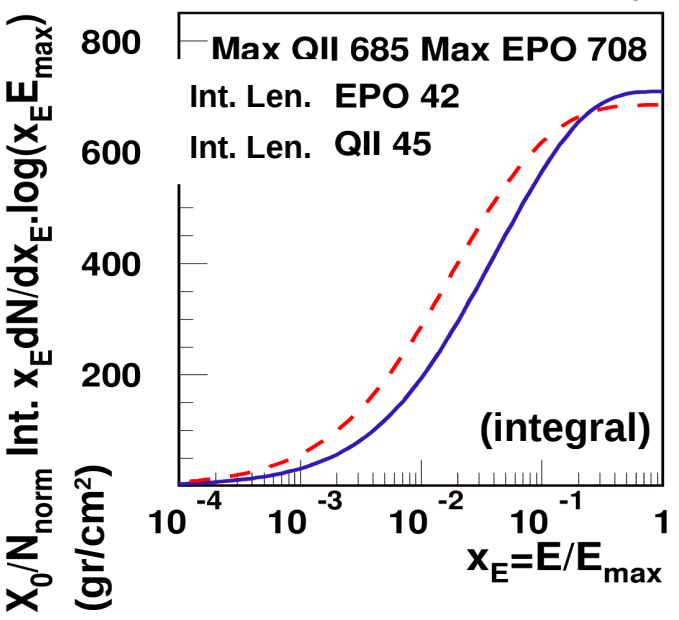
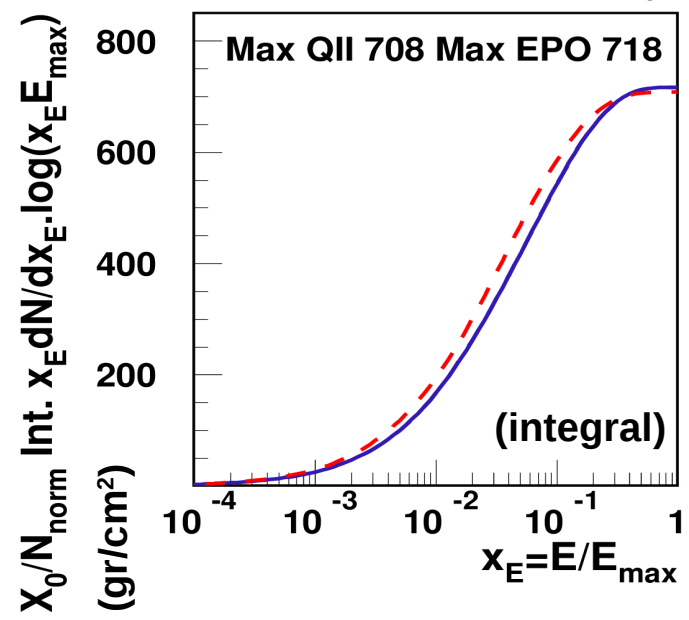
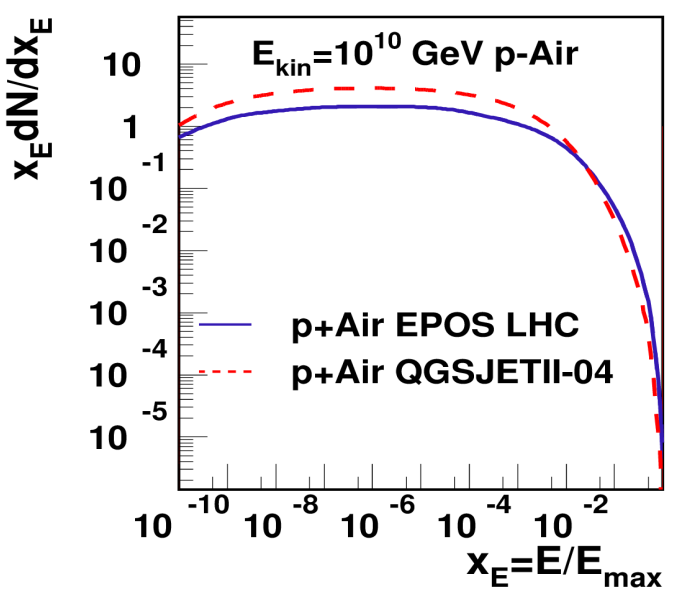
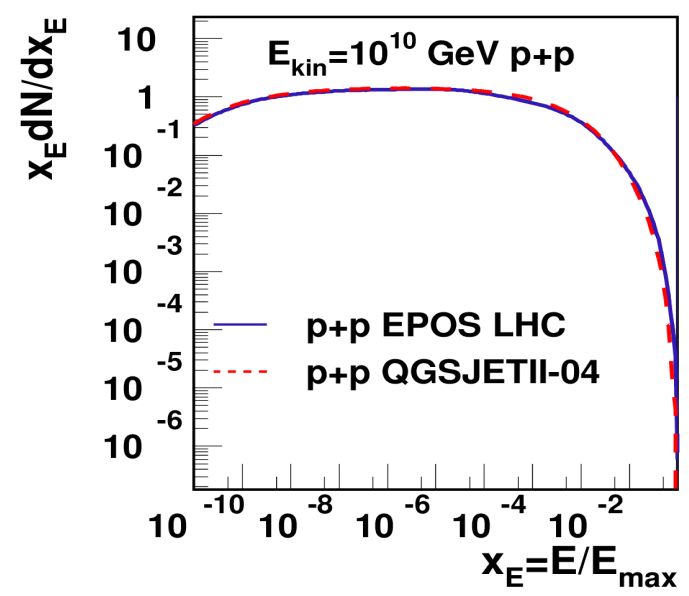
# EAS with Re-tuned CR Models : $X_{\max}$



# Photon Energy Spectra

- **Uncertainties in  $X_{max}$** 
  - ➔ photon energy spectra
  - ➔ elasticity (for 2<sup>nd</sup> interaction)
  - ➔ **extrapolation to nuclear interactions**

- **Use directly energy spectra from first interaction**
  - ➔ which energy is important ?

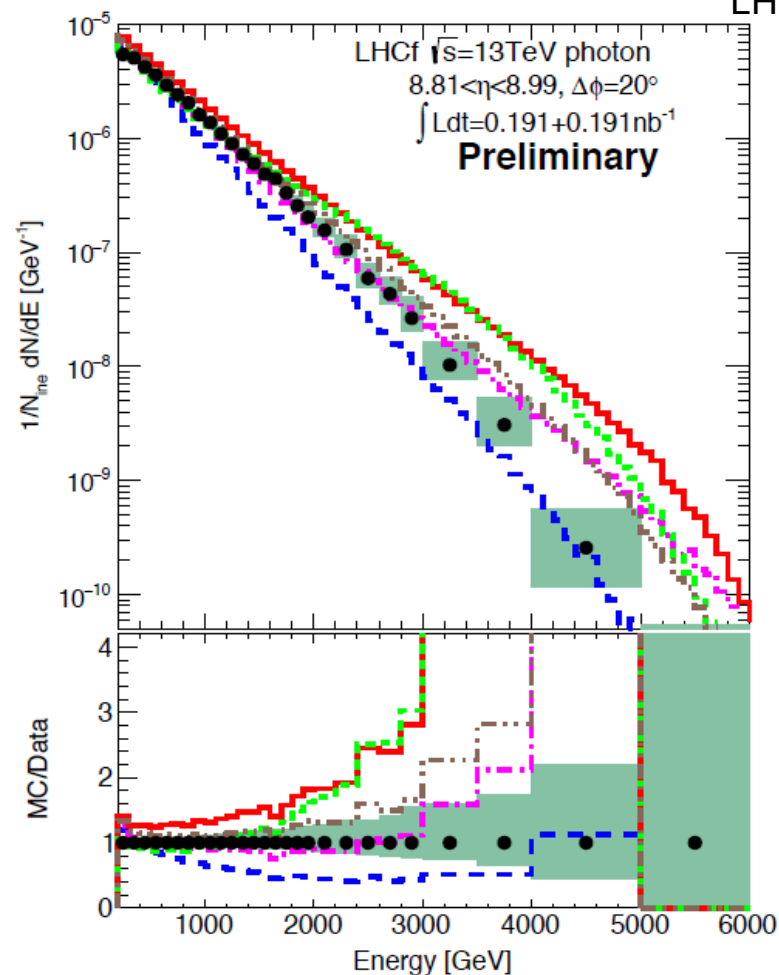
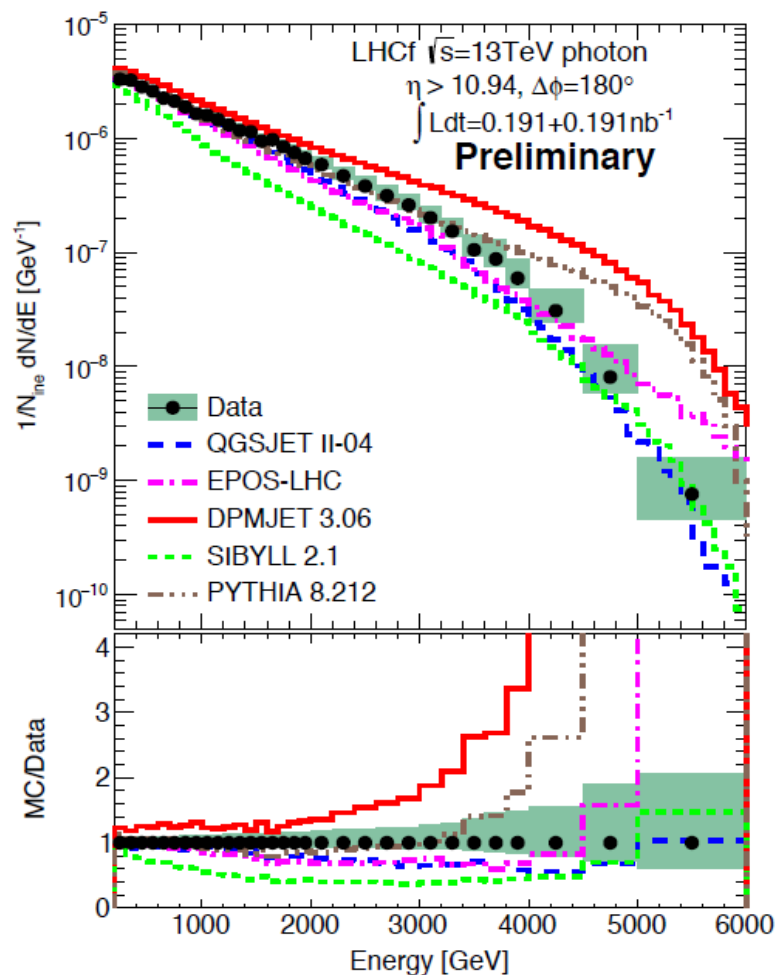




# Comparison with LHCf

- ➔ LHCf favor not too soft photon spectra (EPOS LHC, SIBYLL 2.3) : deep  $X_{\max}$
- ➔ No model compatible with all LHCf measurements : room for improvements !
- ➔ Can p-Pb data be used to mimic light ion (Air) interactions ?

T.Sako for the  
LHCf collaboration

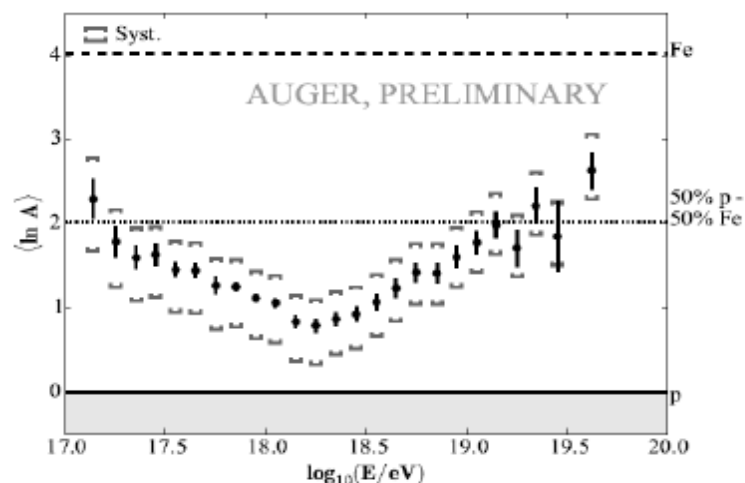


# Model Consistency using Electromagnetic Component

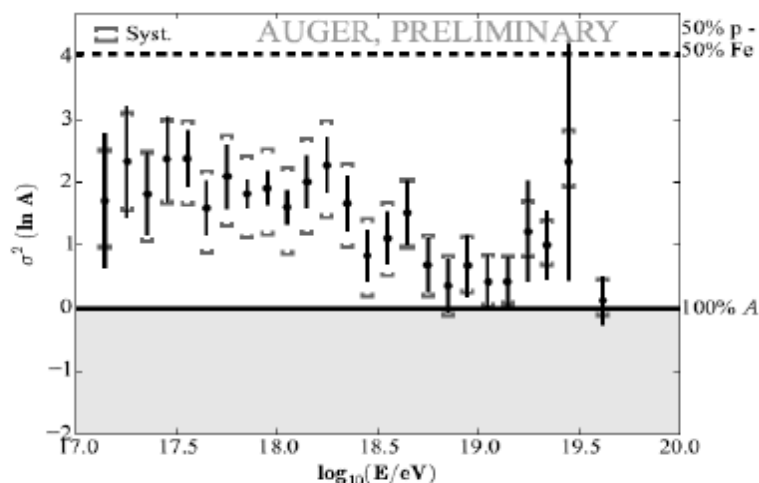
## Study by Pierre Auger Collaboration

→ std deviation of  $\ln A$  allows to test model consistency.

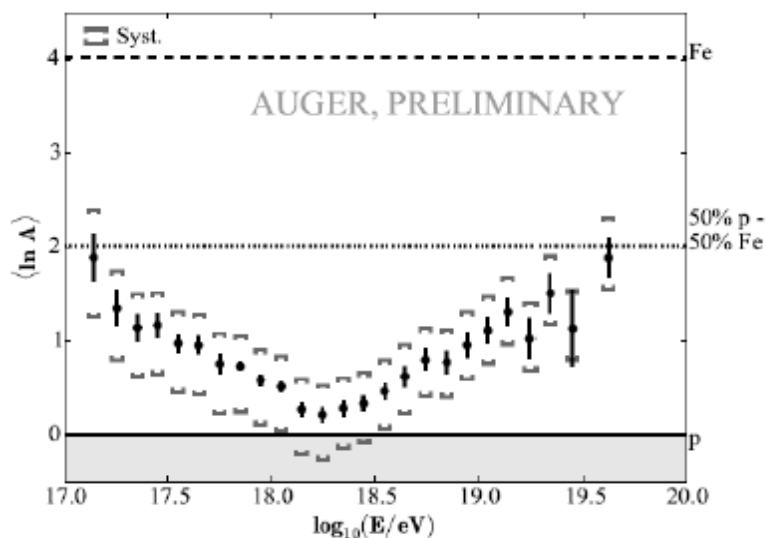
EPOS-LHC (Mean of  $\ln A$ )



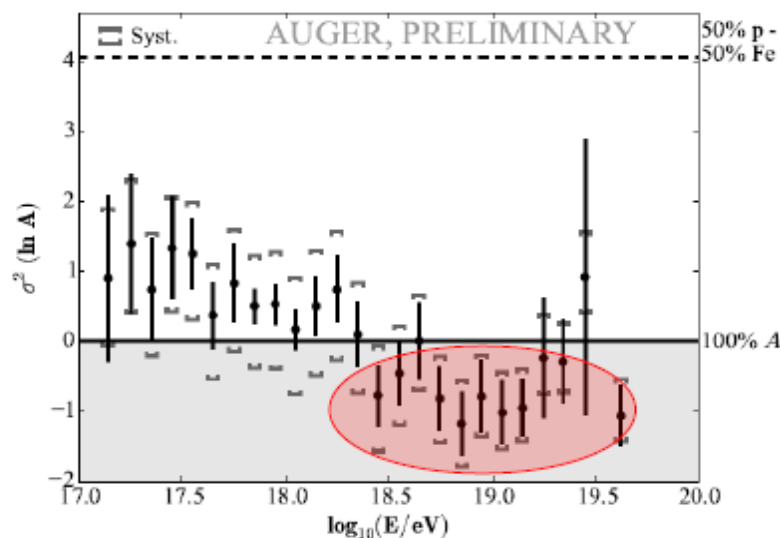
EPOS-LHC (Std. Deviation of  $\ln A$ )



QGSJetII-04 (Mean of  $\ln A$ )



QGSJetII-04 (Std. Deviation of  $\ln A$ )

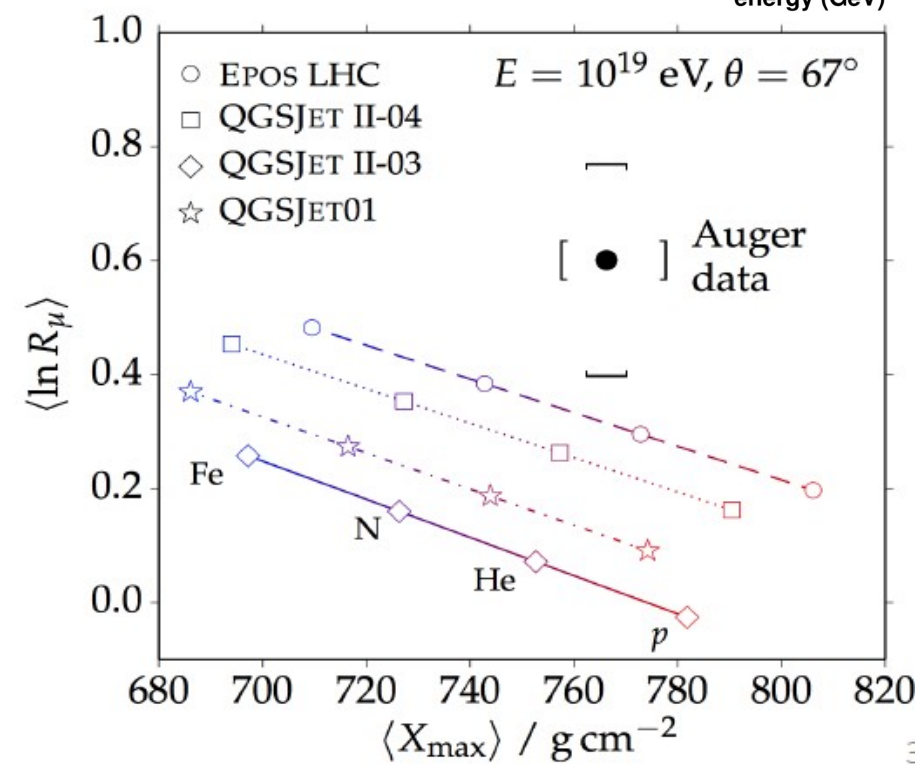
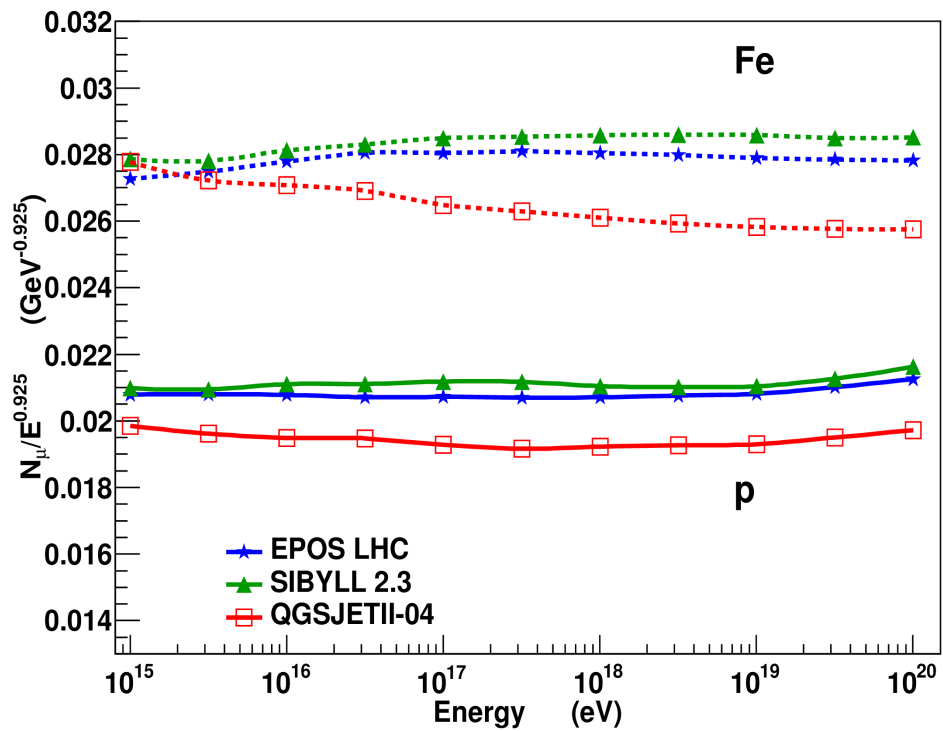
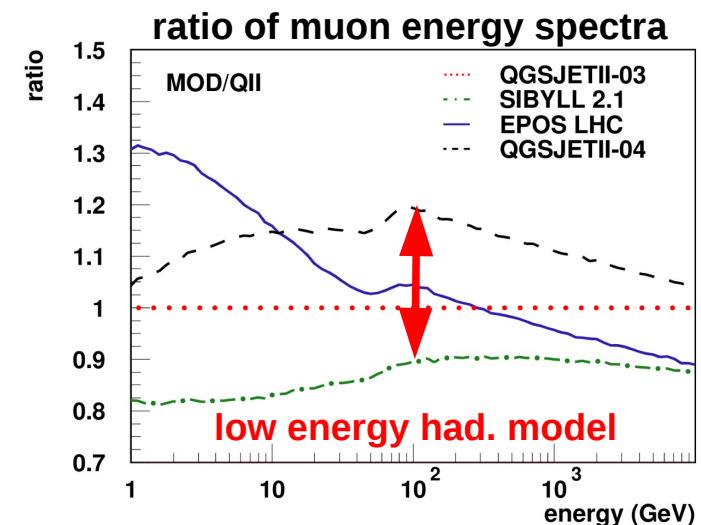


tensions if  $\langle X_{\max} \rangle$   
too small

QGSJETII-04 is a  
lower limit for  $X_{\max}$

# Muons at Ground

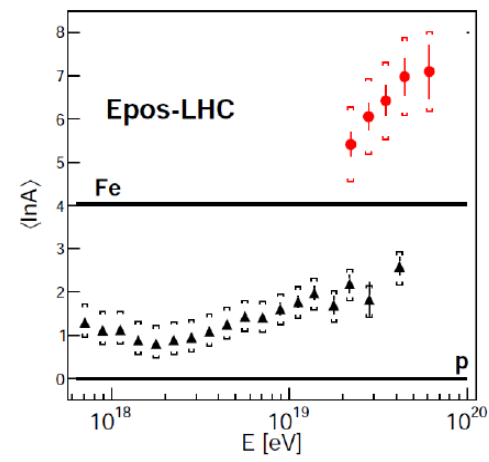
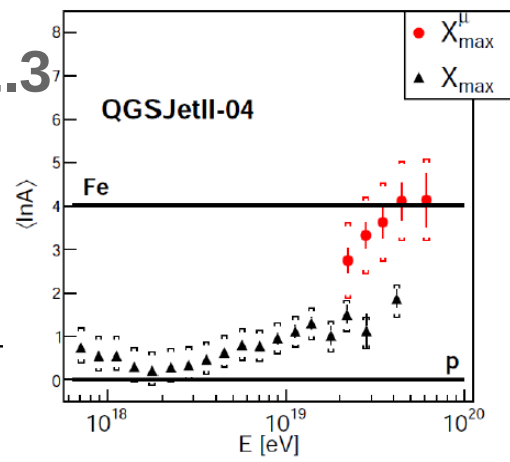
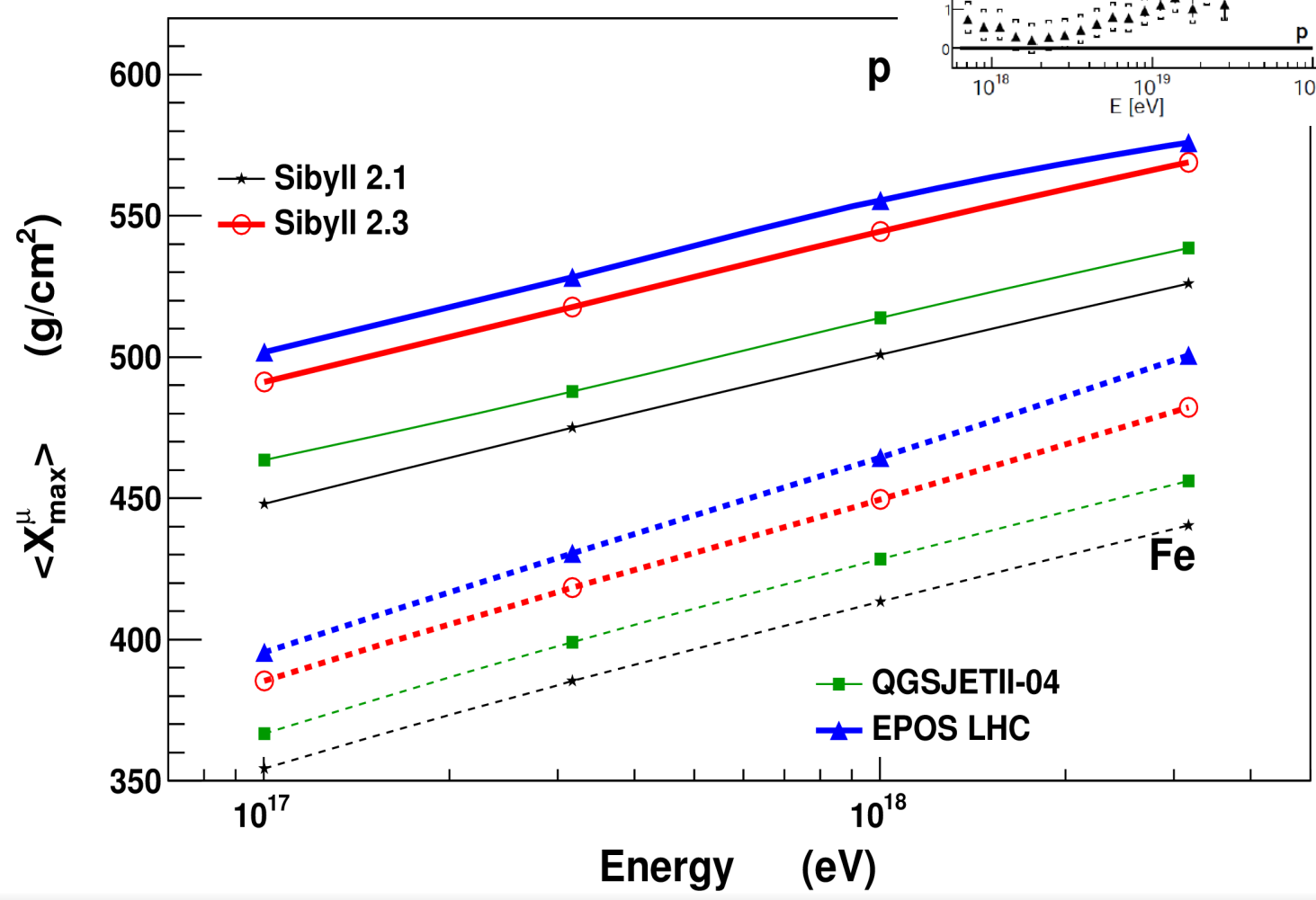
- ➔ Muon production depends on all int. energies
- ➔ Muon production dominated by pion interactions (LHC indirectly important)
- ➔ Resonance and baryon production important
- ➔ Post-LHC Models ~ agrees on numbers but with different production height and spectra



# Muon Production Depth

● Same for EPOS LHC and SIBYLL 2.3

- ➔ low pion-air elasticity: higher  $X_{max}^{\mu}$
- ➔ more forward baryons: higher  $X_{max}^{\mu}$



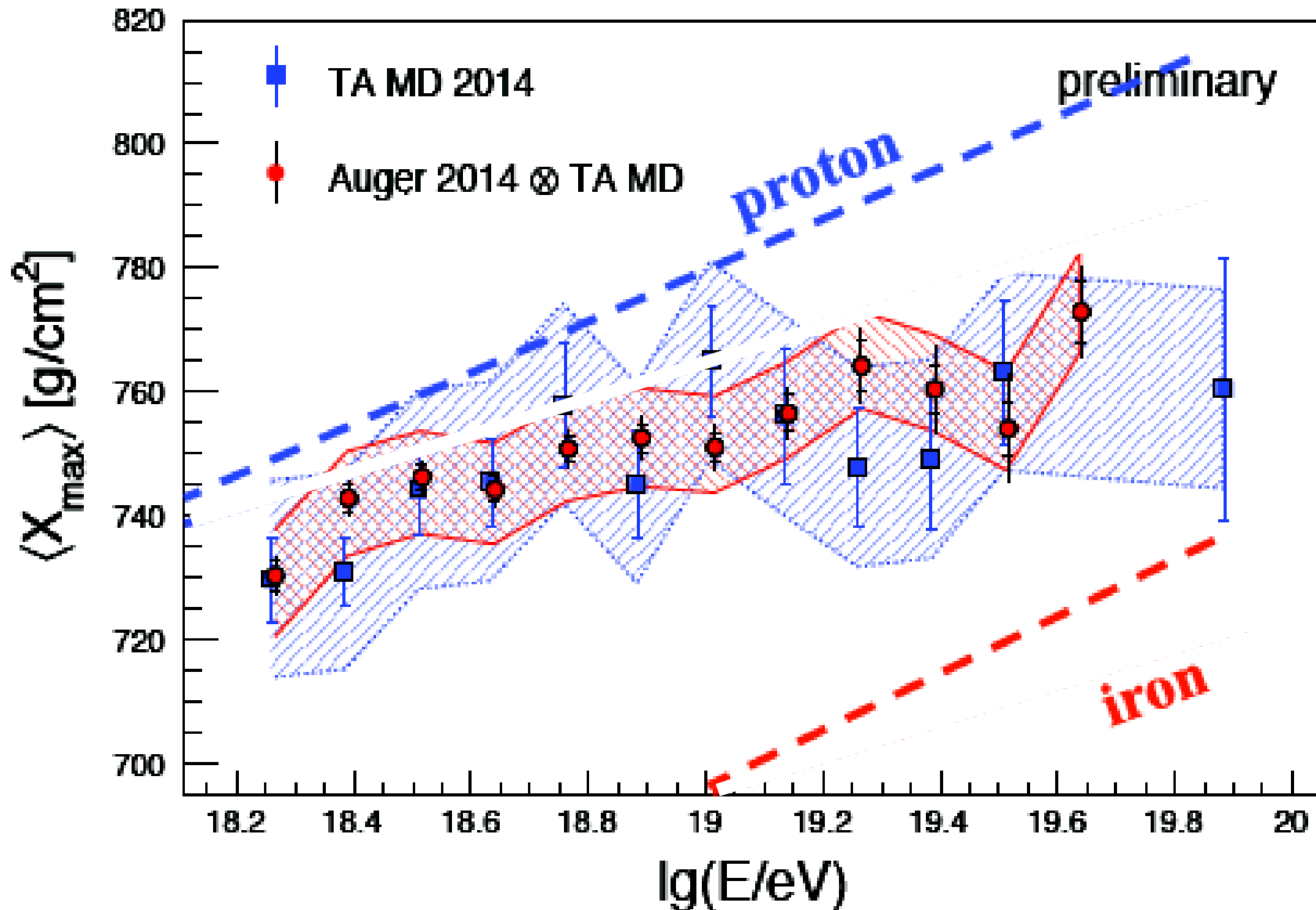
**MPDs sensitive to baryon (less generation) and meson spectra in pion interactions so small effect on  $X_{max}^{\mu}$**

Source of discrepancy understood and will be corrected in EPOS 3.

# Summary

- Auger data (and other low energy cosmic ray experiments) not consistently described by hadronic interaction models (even post LHC)
  - ◆  $\langle X_{\max} \rangle$  and fluctuations, number of muons and muon production depth ...
  - ➔ **but it has never been so good !** only 1 to 2 sigma difference in most of the cases
- Central particle production at LHC reduced model uncertainties in slope of  $X_{\max}$ 
  - ➔ same energy evolution in models important for mass of primary cosmic rays
  - ➔ **all pre-LHC models in contradiction with LHC data** (central and forward prod.)
  - ➔ using latest model version reduce uncertainties and avoid unphysical behavior
- Remaining 20 gr/cm<sup>2</sup> difference for  $X_{\max}$  predictions
  - ➔ **linked to forward physics** (photon spectra and diffraction measured at LHC) not yet taken into account in models used for EAS simulation (coming...)
  - ➔ effect of extrapolation to p-Air interaction
    - ◆ **p-O beam necessary** to check that p-p properly extrapolated
    - ◆ p-Pb measurements can be used but need change in most models (only EPOS reproduces p-Pb data for the moment)

# Conclusion ...

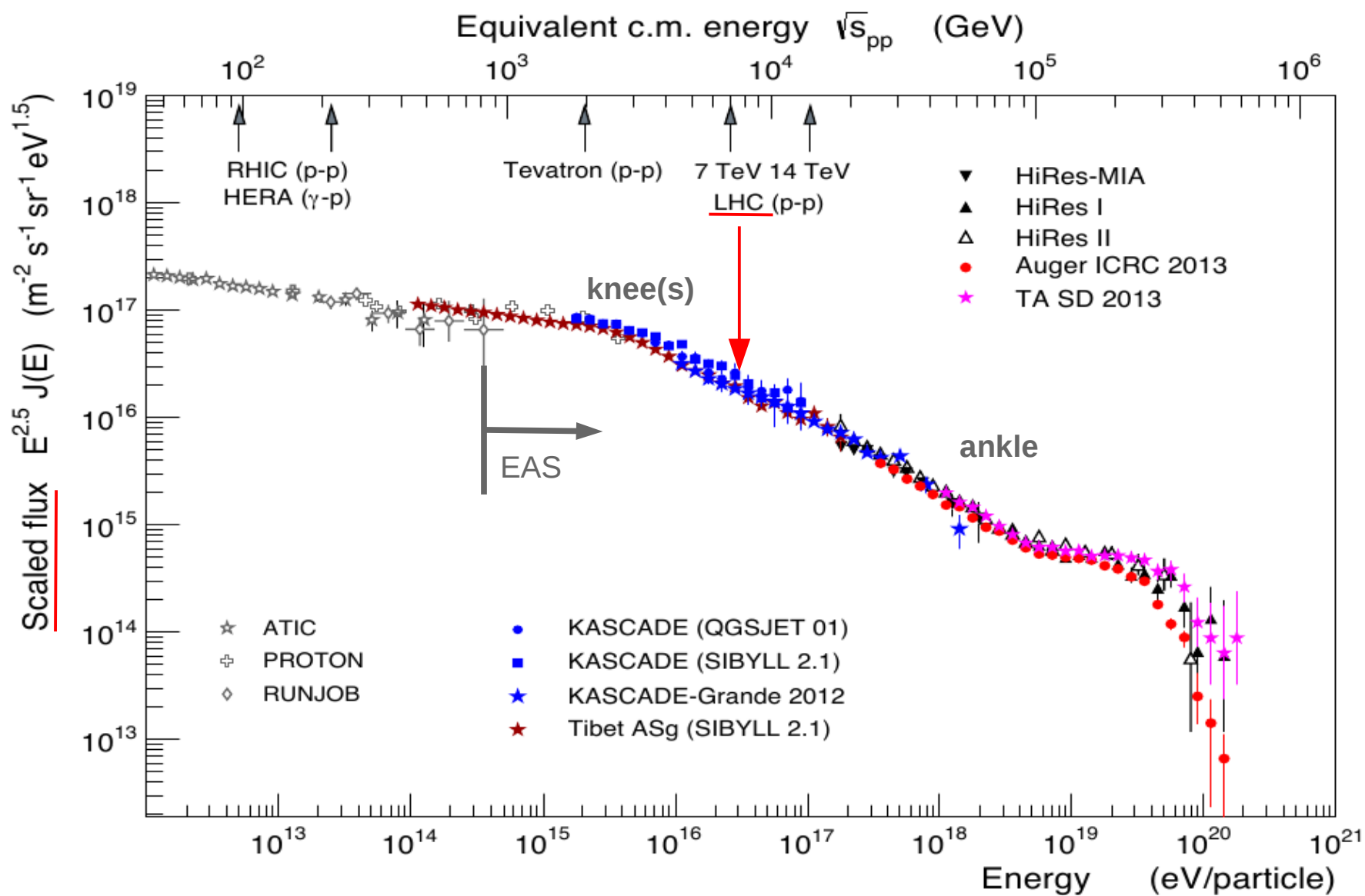


From Roberto Aloiso talk (2015 working group)

# Baryons in Pion-Carbon

- **Very few data for baryon production from meson projectile, but for all :**
    - ➔ strong baryon acceleration (probability  $\sim 20\%$  per string end)
    - ➔ proton/antiproton asymmetry (valence quark effect)
    - ➔ target mass dependence
  - **New data set from NA49 (G. Veres' PhD)**
    - ➔ test  $\pi^+$  and  $\pi^-$  interactions and productions at 158 GeV with C and Pb target
    - ➔ confirm large forward proton production in  $\pi^+$  and  $\pi^-$  interactions but not for anti-protons
      - ◆ forward protons in pion interactions are due to strong baryon stopping (nucleons from the target are accelerated in projectile direction)
      - ◆ strong effect only at low energy
- ➔ EPOS overestimate forward baryon production at high energy

# Cosmic Ray Spectrum



R. Engel (KIT)

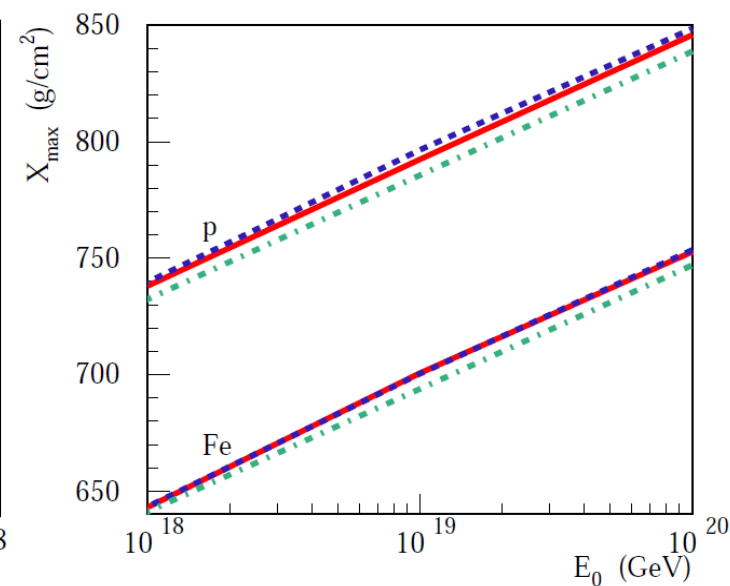
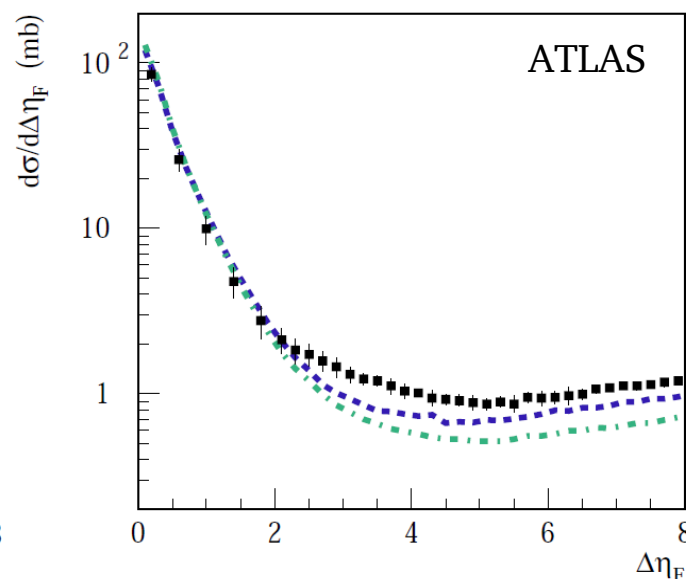
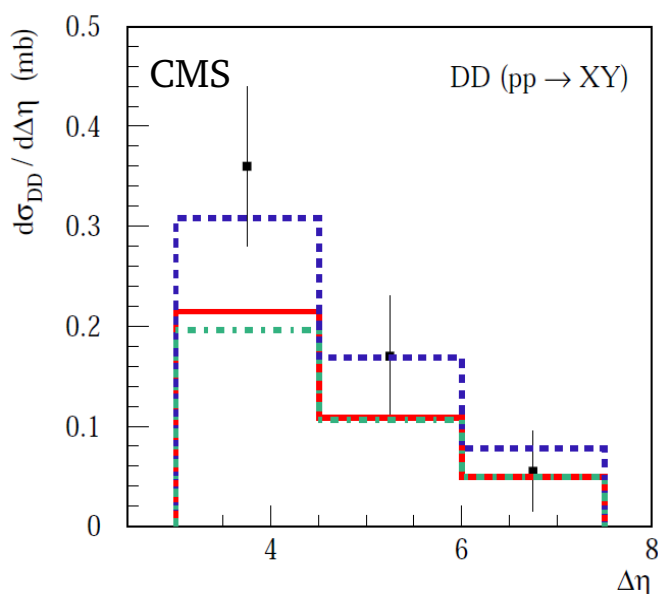


# Diffraction measurements

- TOTEM and CMS diffraction measurement not fully consistent
- Tests by S. Ostapchenko using QGSJETII-04 (PRD89 (2014) no.7, 074009)
  - ➔ SD+ option compatible with CMS
  - ➔ SD- option compatible with TOTEM

$M_X$ range	$< 3.4$ GeV	$3.4 - 1100$ GeV	$3.4 - 7$ GeV	$7 - 350$ GeV	$350 - 1100$ GeV
TOTEM [13, 24]	$2.62 \pm 2.17$	$6.5 \pm 1.3$	$\simeq 1.8$	$\simeq 3.3$	$\simeq 1.4$
QGSJET-II-04	3.9	7.2	1.9	3.9	1.5
option SD+	3.2	8.2	1.8	4.7	1.7
option SD-	2.6	7.2	1.6	3.9	1.7

➔ difference of  $\sim 10$  gr/cm<sup>2</sup> between the 2 options



# Simplified Shower Development

Using generalized Heitler model and superposition model :

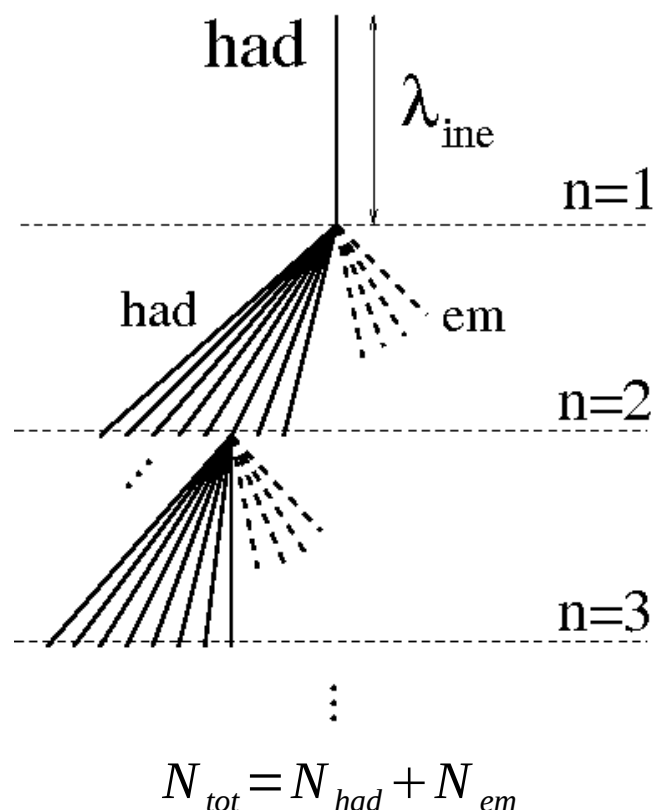
$$X_{max} \sim \lambda_e \ln \left( (1-k) \cdot E_0 / (2 \cdot N_{tot} \cdot A) \right) + \lambda_{ine}$$

➔ Model independent parameters :

- $E_0$  = primary energy
- $A$  = primary mass
- $\lambda_e$  = electromagnetic mean free path

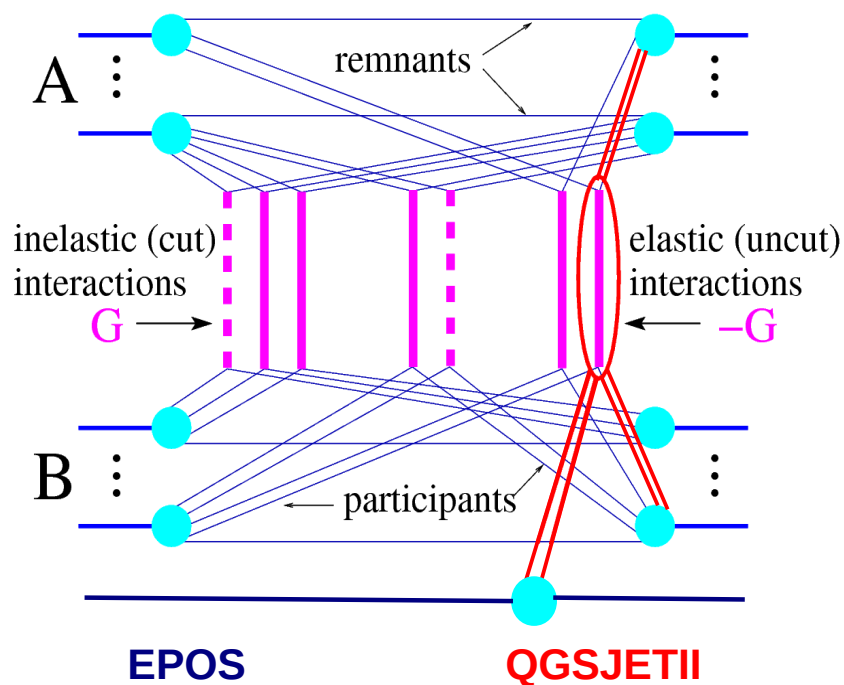
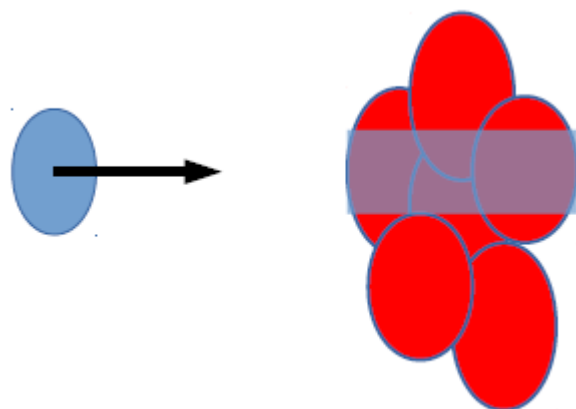
➔ Model dependent parameters :

- $k$  = elasticity
- $N_{tot}$  = total multiplicity
- $\lambda_{ine}$  = hadronic mean free path (cross section)



J. Matthews, Astropart.Phys. 22  
(2005) 387-397

# Nuclear Interactions



## ● Sibyll

→ Glauber for pA

■ with inelastic screening for diffraction in new Sibyll 2.3 (only nuclear effect)

→ superposition model for AA ( $A \times pA$ )

## ● QGSJETII

→ Pomeron configuration based on  $A$  projectiles and  $A$  targets

→ Nuclear effect due to multi-leg Pomerons

## ● EPOS

→ Pomeron configuration based on  $A$  projectiles and  $A$  targets

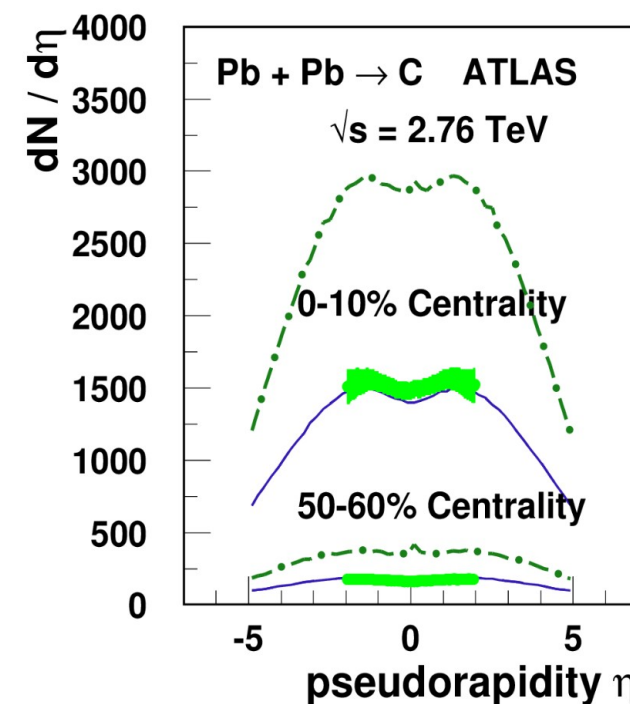
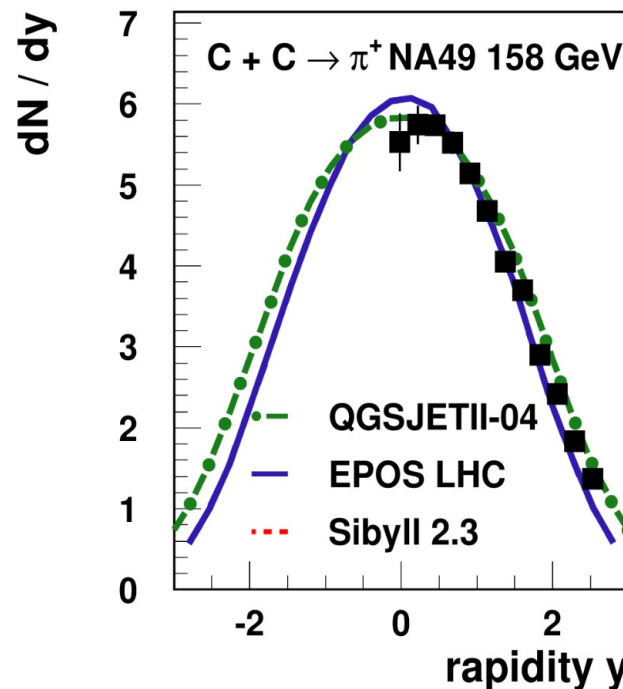
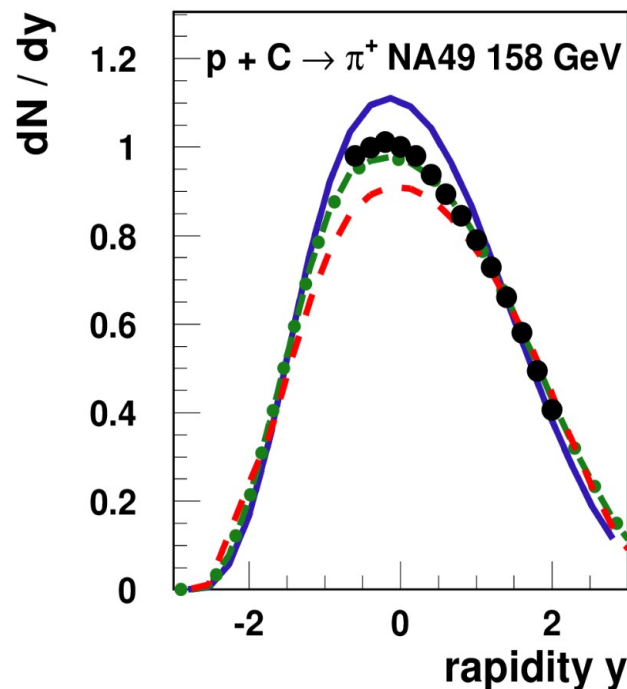
→ screening corrections depend on nuclei

→ final state interactions (core-corona approach and collective hadronization with flow for core)

# Light Ion Data

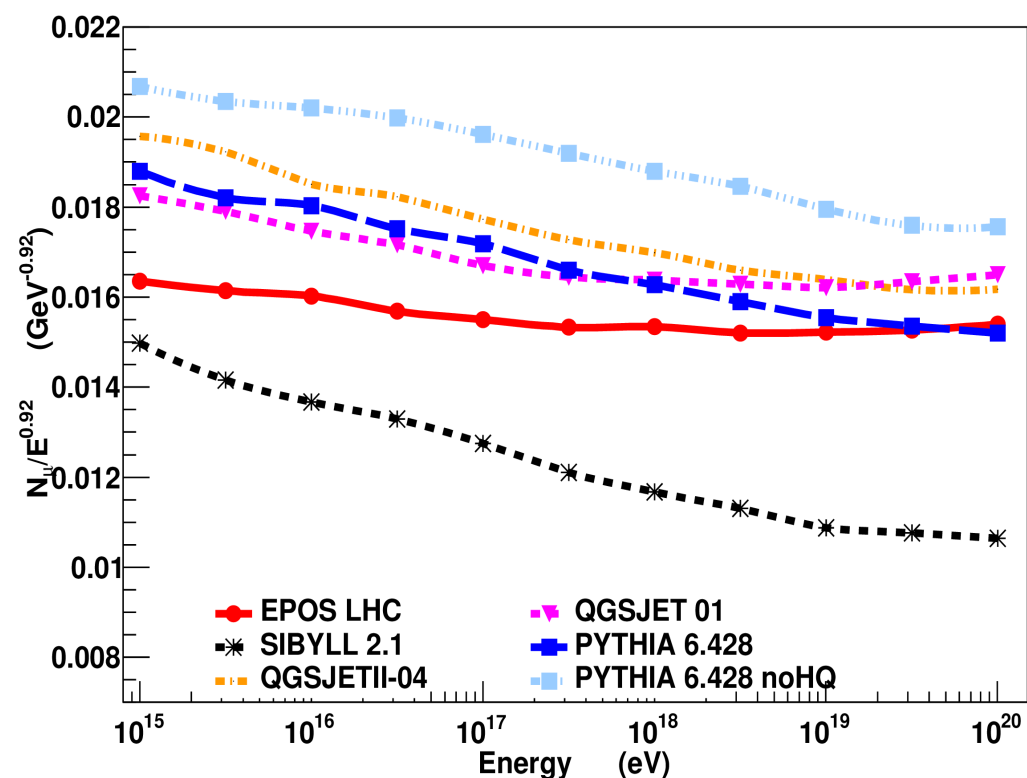
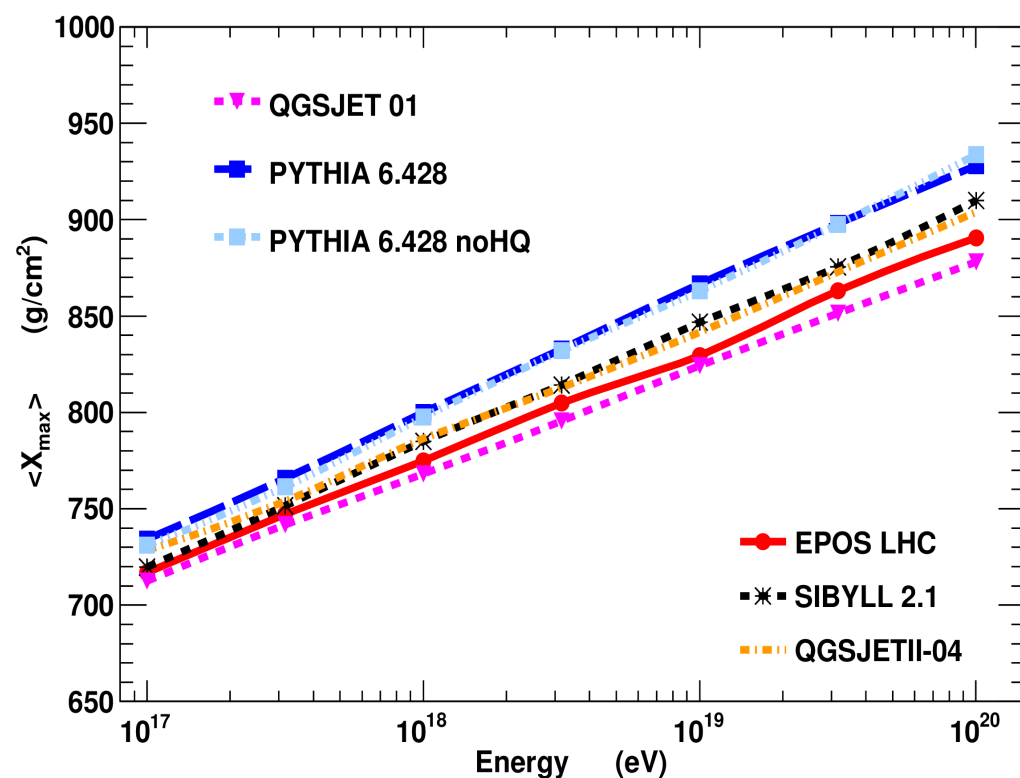
Very few data to compare with all CR models :

- ➔ strong limitations in Sibyll (projectile up to Fe only and target up to O !)
- ➔ no final state interactions exclude heavy nuclei for QGSJETII
- ➔ no light ion data at high energy

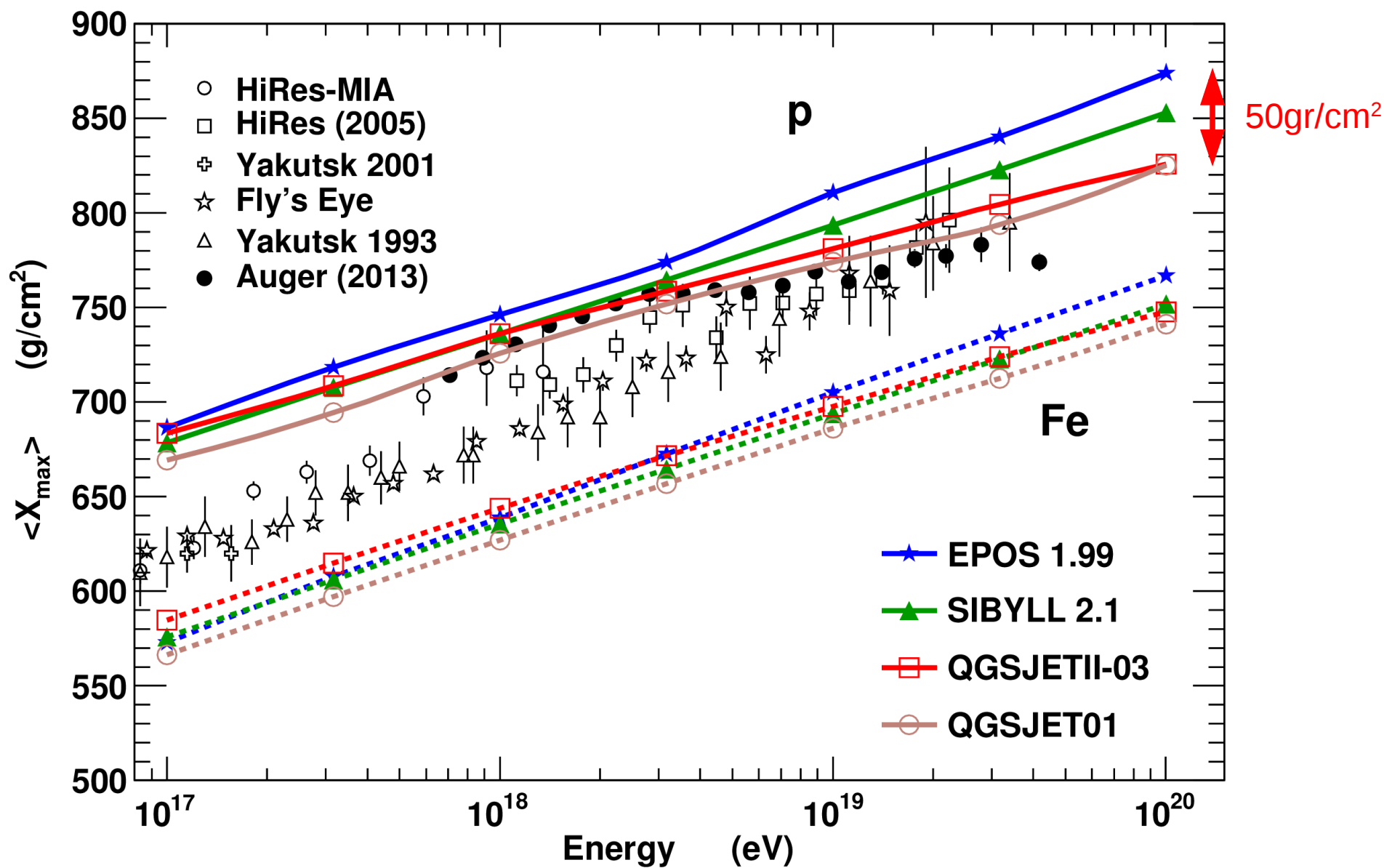


# Tests using hydrogen atmosphere

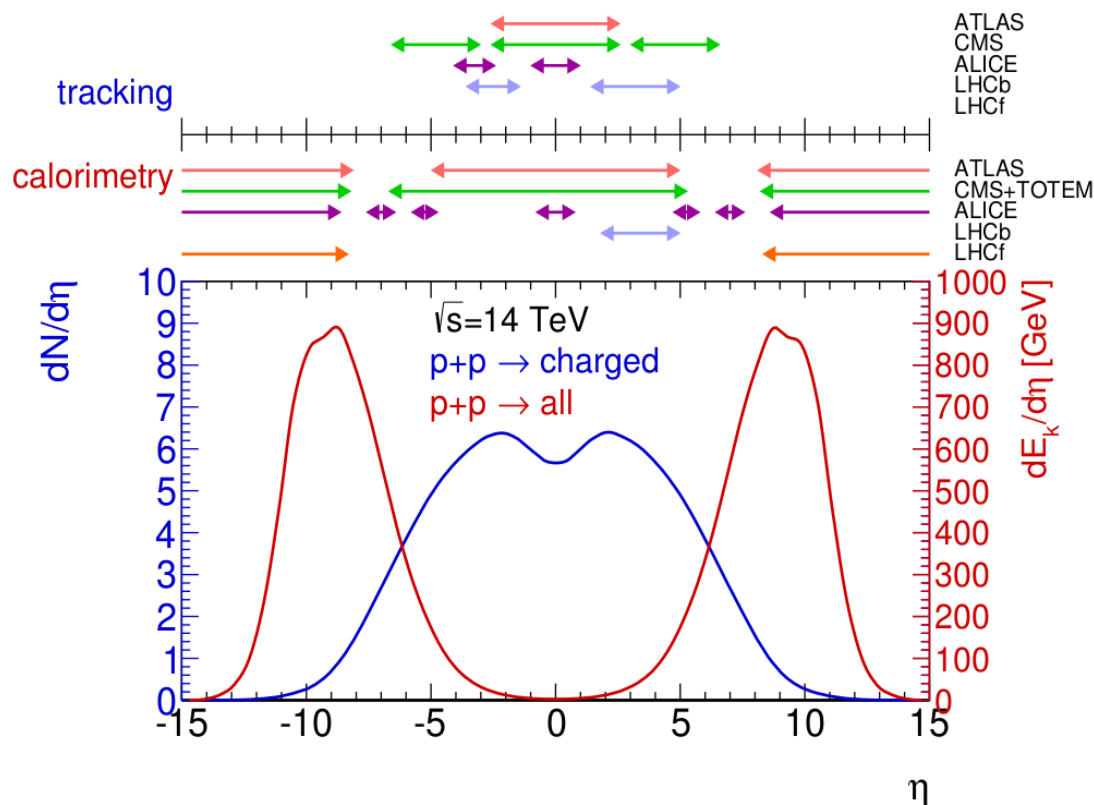
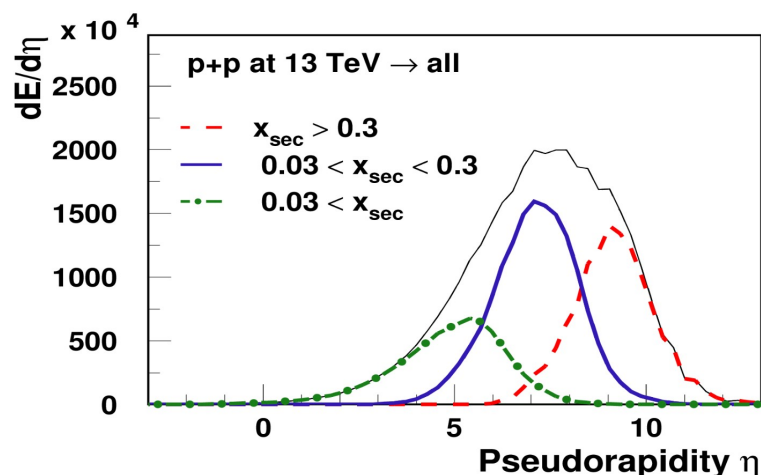
- Work done with David D'Enterria (CERN) and Sun Guanhao
  - ➔ test of Pythia event generator
- Modified air shower simulations with air target replaced by hydrogen
  - ➔ for interactions only (no change in density)
  - ➔ no nuclear effect



# EAS with Old CR Models : $X_{\max}$



# LHC acceptance



- p-p data of central detectors used to reduce uncertainty by factor ~2
  - ➔ p-Pb difficult to compare to CR models (only EPOS)
  - ➔ special centrality selection
  - ➔ pO ?
- Direct photon energy spectra from LHCf
  - ➔ small phase space but relevant for  $X_{\text{max}}$
  - ➔ p-Pb (O) and correlation with ATLAS
- Average elasticity/inelasticity (energy fraction of the leading particle)
  - ➔ all diffraction measurement to be taken into account