

Neutrino Landscape 2016

José W F Valle



<https://www.facebook.com/ific.ahep/>

Humboldt Kolleg "From the Vacuum to the Universe", Kitzbühel Austria, June-2016

Why neutrinos

Can not do without neutrinos

Basic cosmological and astro probe

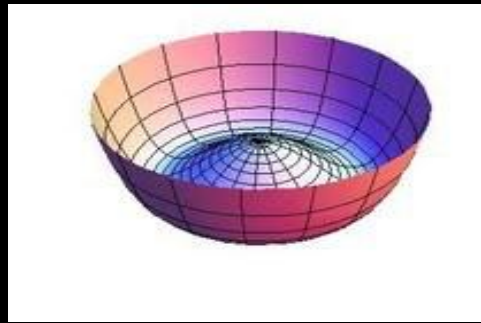
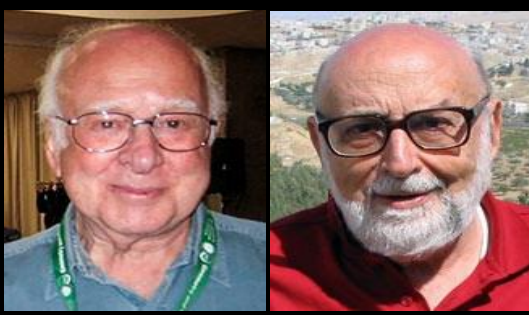
key building block of the Standard Model

Standard model

Three Generations of Matter (Fermions) spin $\frac{1}{2}$

	I	II	III	
mass →	2.4 MeV	1.27 GeV	171.2 GeV	0
charge →	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0
name →	u up	c charm	t top	g gluon
	Left Right	Left Right	Left Right	0
	4.8 MeV	104 MeV	4.2 GeV	0
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0
Quarks	d down	s strange	b bottom	γ photon
	Left Right	Left Right	Left Right	91.2 GeV
	0 eV	0 eV	0 eV	0
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z⁰ weak force
	Left Right	Left Right	Left Right	80.4 GeV
	0.511 MeV	105.7 MeV	1.777 GeV	± 1
Leptons	e electron	μ muon	τ tau	W[±] weak force
	Left Right	Left Right	Left Right	

Bosons (Forces) spin 1



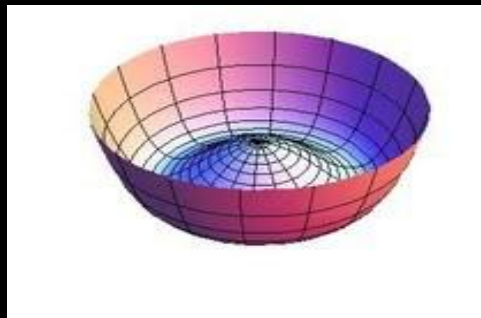
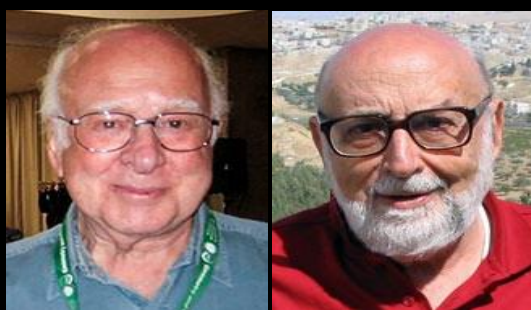
Standard model

125 GeV
H
Higgs boson
spin 0

Three Generations of Matter (Fermions) spin 1/2

	I	II	III	
mass →	2.4 MeV	1.27 GeV	171.2 GeV	0
charge →	2/3	2/3	2/3	0
name →	u up	c charm	t top	g gluon
	Left Right	Left Right	Left Right	0
	d down	s strange	b bottom	γ photon
Quarks	Left Right	Left Right	Left Right	0
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	91.2 GeV
	0 eV	0 eV	0 eV	0
	e electron	μ muon	τ tau	Z⁰ weak force
Leptons	Left Right	Left Right	Left Right	80.4 GeV
	-1	-1	-1	W[±] weak force
	0.511 MeV	105.7 MeV	1.777 GeV	

Bosons (Forces) spin 1



Standard model

125 GeV

H

Higgs boson

spin 0

Three Generations of Matter (Fermions) spin 1/2

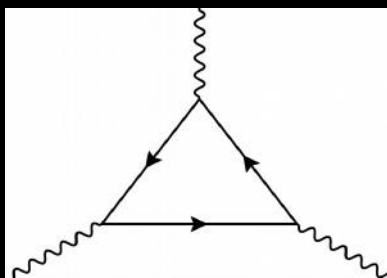
	I	II	III	
mass →	2.4 MeV	1.27 GeV	171.2 GeV	0
charge →	2/3	2/3	2/3	0
name →	u Left up Right	c Left charm Right	t Left top Right	g gluon
	4.8 MeV	104 MeV	4.2 GeV	0
Quarks	d Left down Right	s Left strange Right	b Left bottom Right	0 γ photon
	0 eV	0 eV	0 eV	91.2 GeV
	ν_e Left electron neutrino Right	ν_μ Left muon neutrino Right	ν_τ Left tau neutrino Right	0 Z weak force
Leptons	0.511 MeV	105.7 MeV	1.777 GeV	80.4 GeV
	e Left electron Right	μ Left muon Right	τ Left tau Right	+1 W weak force

Bosons (Forces) spin 1

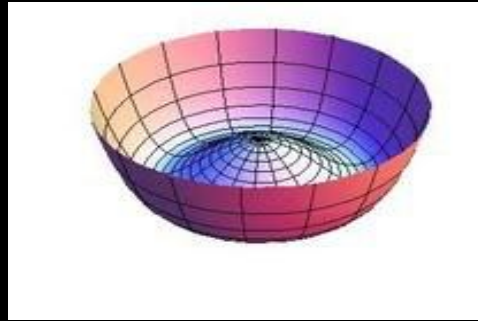
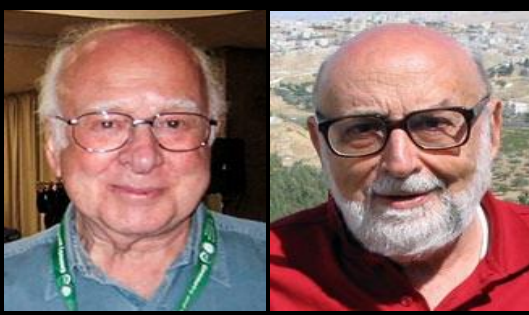
Last brick ? ...

$$-\frac{1}{2} \quad -\frac{1}{2} \quad \frac{1}{6} \quad \frac{1}{6} \quad -1 \quad \frac{2}{3} \quad -\frac{1}{3}$$

anomalies



Charge quantization



Standard model

125 GeV
H
 Higgs boson
 spin 0

Last brick? ...

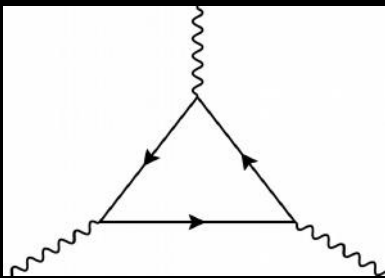
Three Generations of Matter (Fermions) spin 1/2

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charge →	2/3	2/3	2/3	0
name →	u Left up Right	c Left charm Right	t Left top Right	g gluon
	4.8 MeV	104 MeV	4.2 GeV	0
Quarks	d Left down Right	s Left strange Right	b Left bottom Right	γ photon
	0 eV	0 eV	0 eV	91.2 GeV
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z⁰ weak force
Leptons	0.511 MeV	105.7 MeV	1.777 GeV	80.4 GeV
	e Left electron Right	μ Left muon Right	τ Left tau Right	W[±] weak force
				Bosons (Forces) spin 1

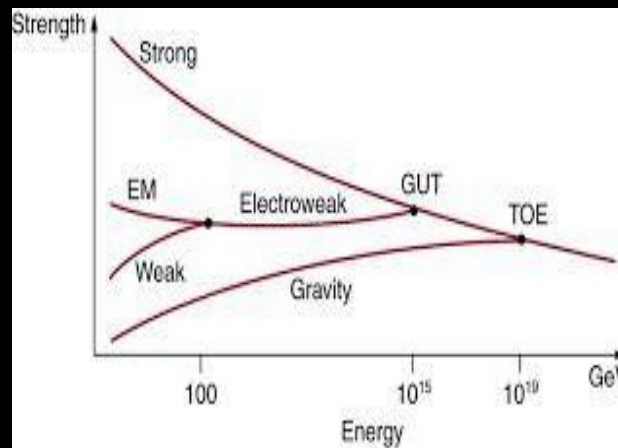
$$-\frac{1}{2} \quad -\frac{1}{2} \quad \frac{1}{6} \quad \frac{1}{6} \quad -1 \quad \frac{2}{3} \quad -\frac{1}{3}$$

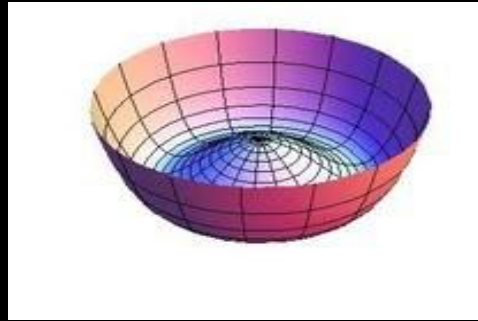
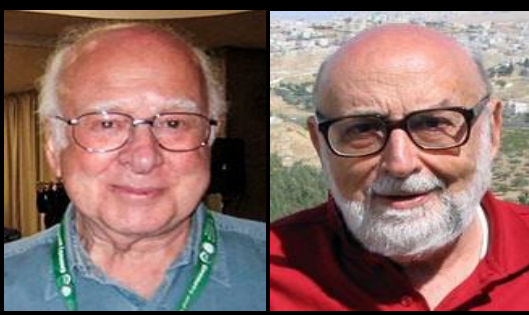
anomalies

coupling unification



Charge quantization





Standard model

125 GeV
H
 Higgs boson
 spin 0

Last brick? ...

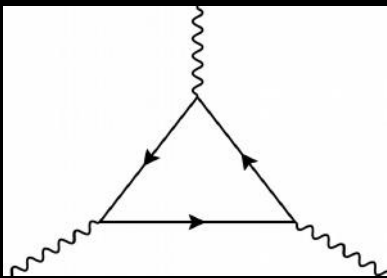
Three Generations of Matter (Fermions) spin 1/2

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charge →	2/3	2/3	2/3	0
name →	u Left up Right	c Left charm Right	t Left top Right	g gluon
Quarks	4.8 MeV -1/3 d Left down Right	104 MeV -1/3 s Left strange Right	4.2 GeV -1/3 b Left bottom Right	0 0 γ photon
Leptons	0 eV 0 ν_e Left electron neutrino Right	0 eV 0 ν_μ Left muon neutrino Right	0 eV 0 ν_τ Left tau neutrino Right	91.2 GeV 0 Z⁰ weak force
	0.511 MeV -1 e Left electron Right	105.7 MeV -1 μ Left muon Right	1.777 GeV -1 τ Left tau Right	80.4 GeV ±1 W[±] weak force

Bosons (Forces) spin 1

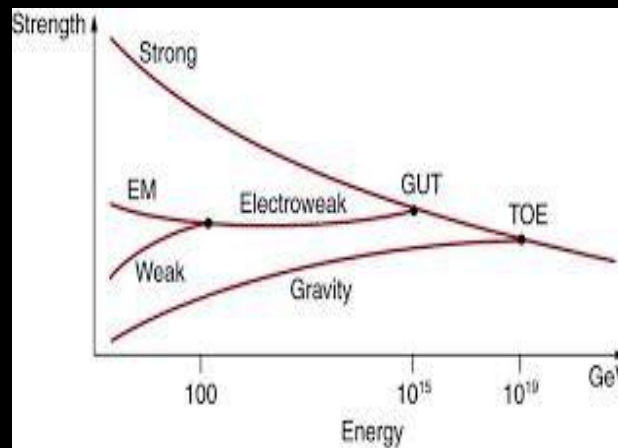
$$-\frac{1}{2} \quad -\frac{1}{2} \quad \frac{1}{6} \quad \frac{1}{6} \quad -1 \quad \frac{2}{3} \quad -\frac{1}{3}$$

anomalies

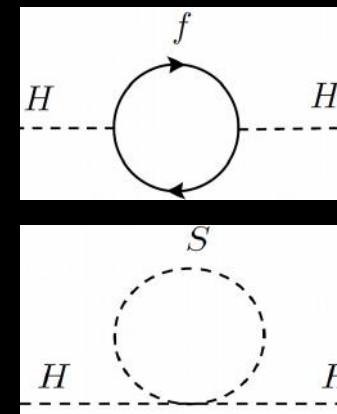


Charge quantization

coupling unification



Consistency of SSB

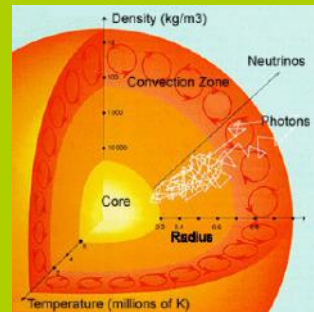


Gravity ...

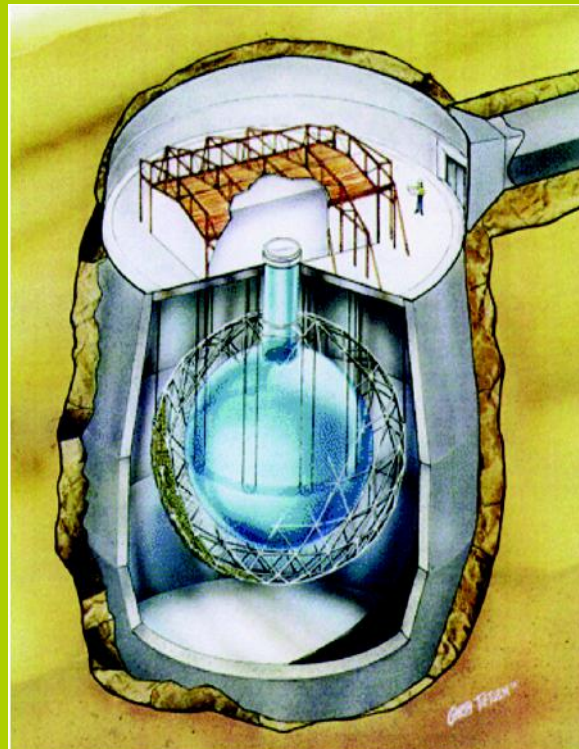
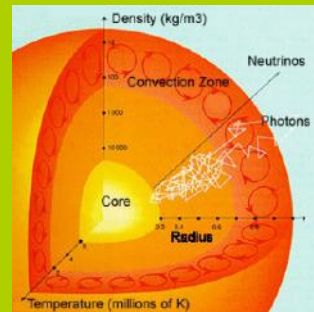
Why 3 families ...

Neutrino mass

Neutrino oscillations

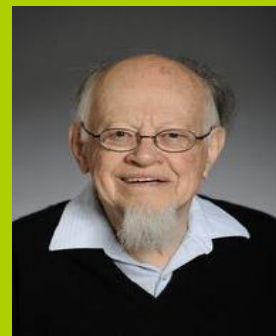
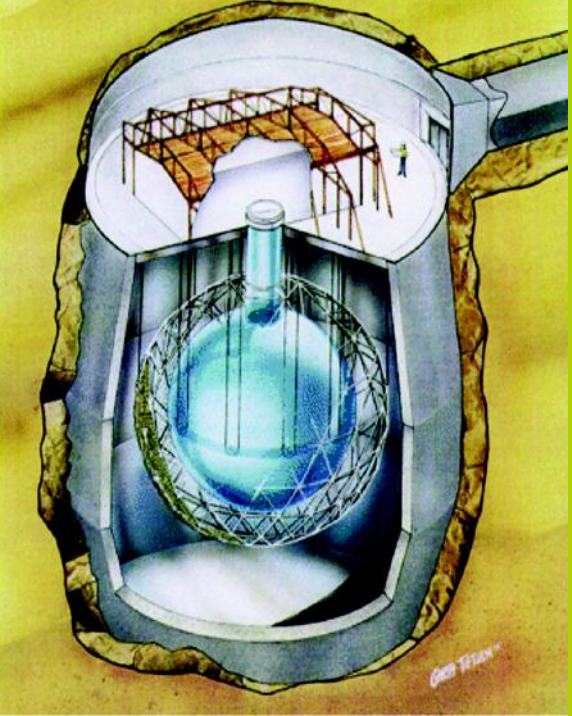
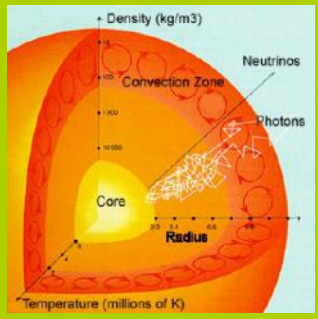
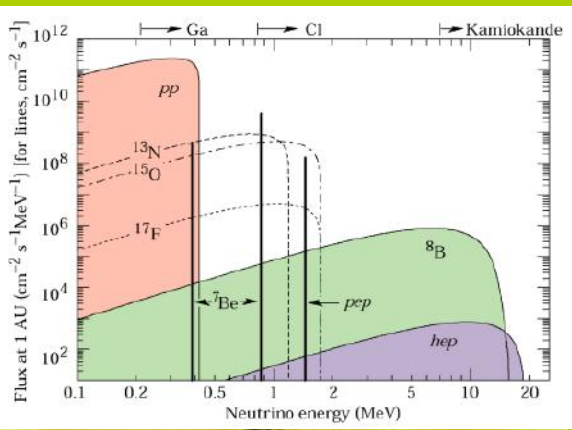


Neutrino oscillations

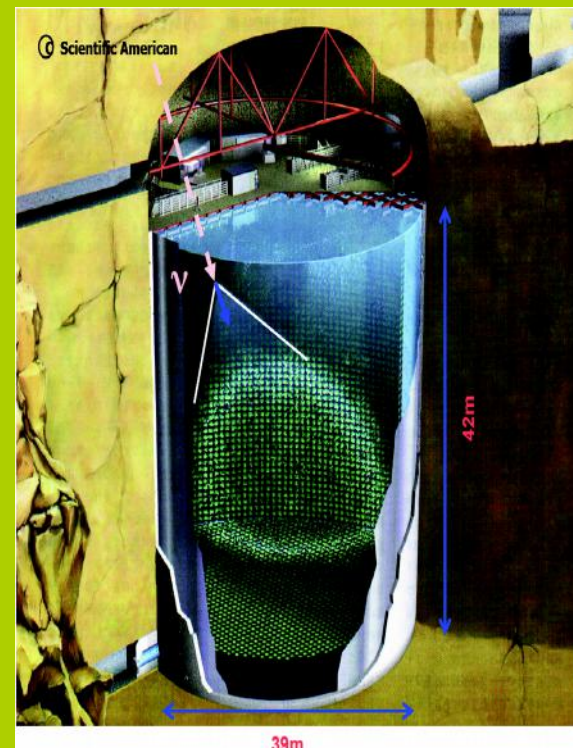
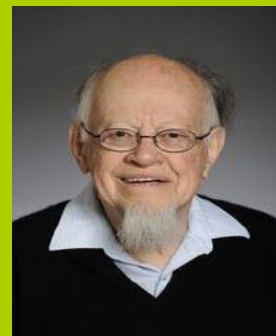
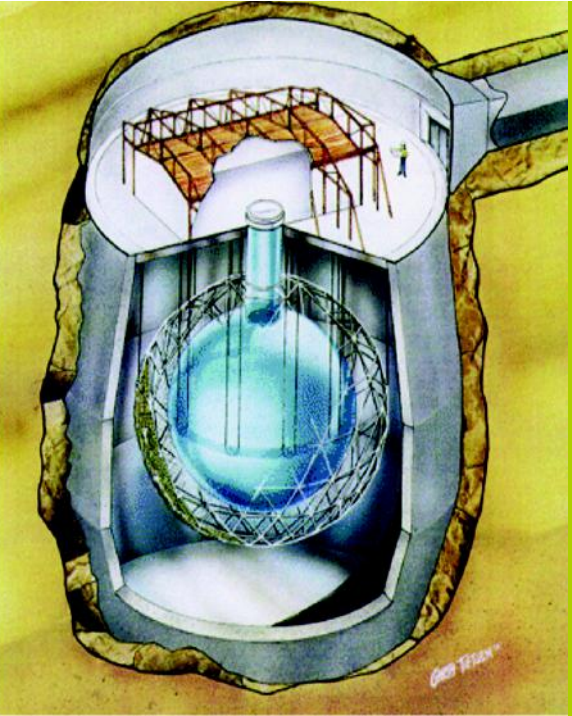
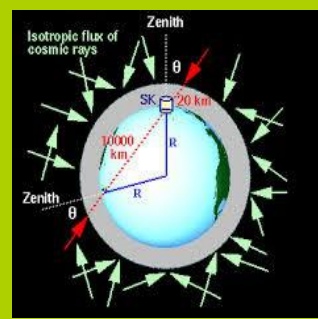
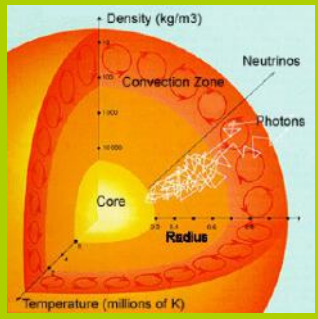
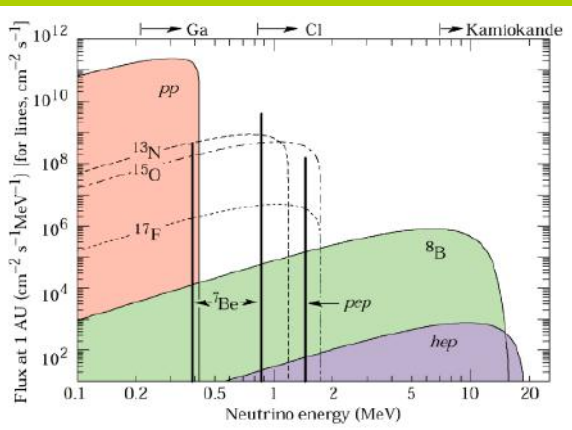


Valle

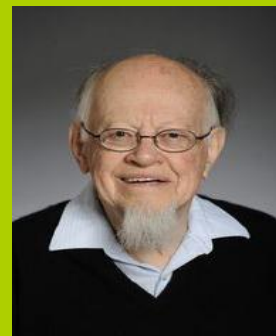
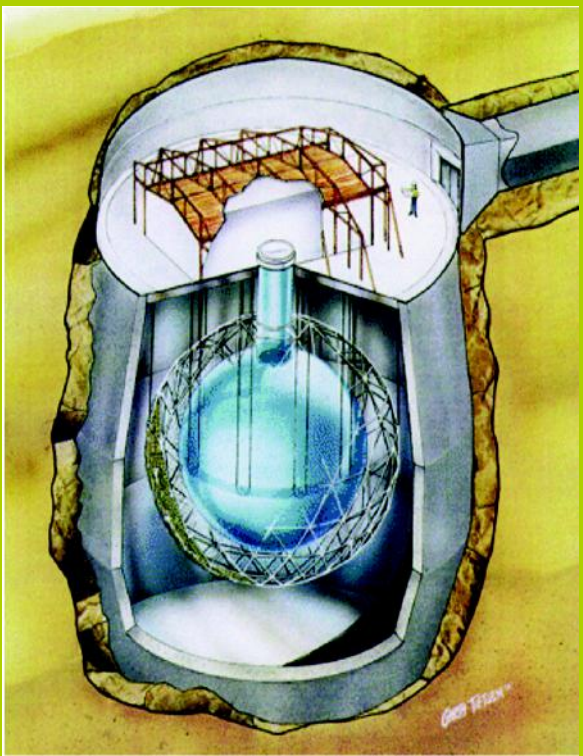
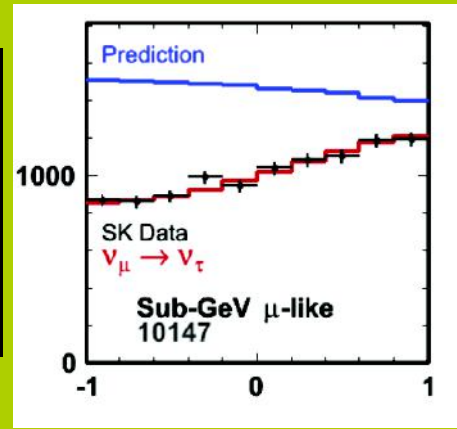
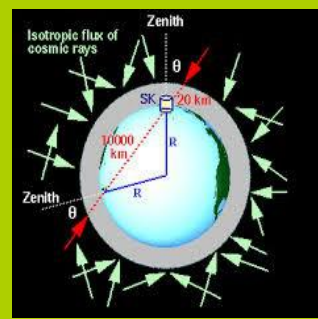
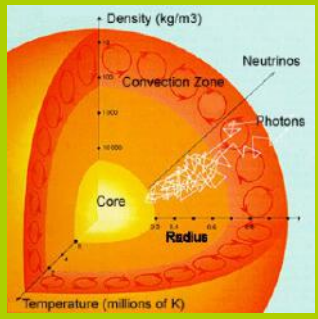
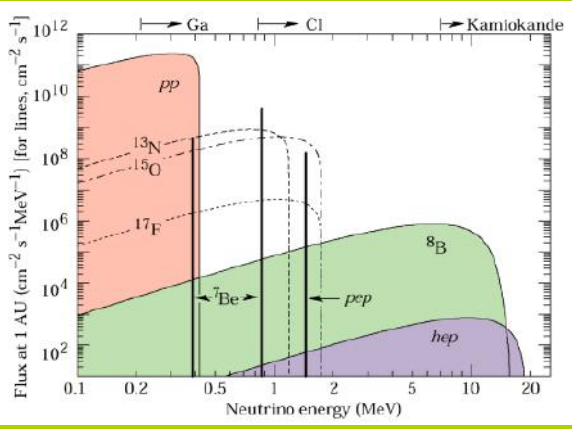
Neutrino oscillations



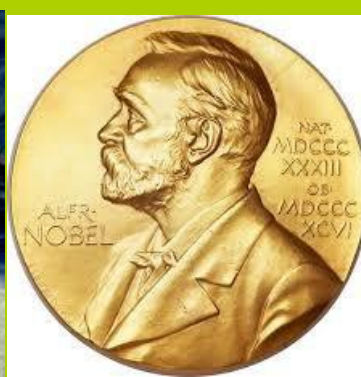
Neutrino oscillations



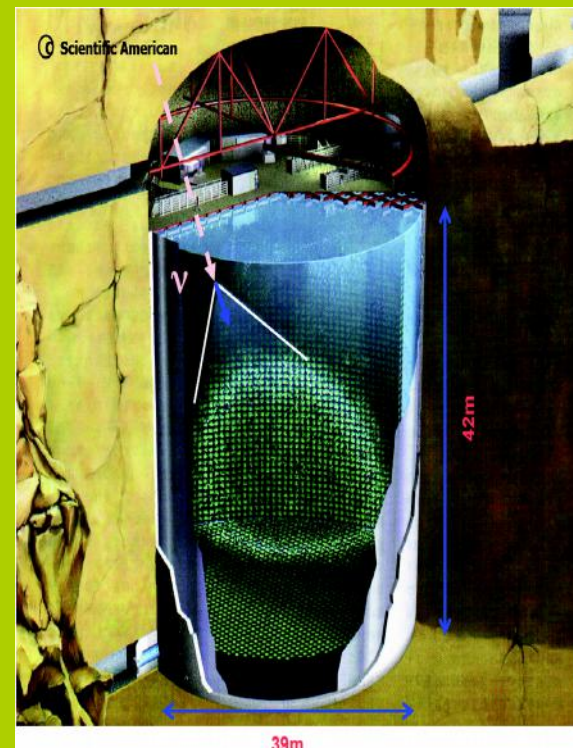
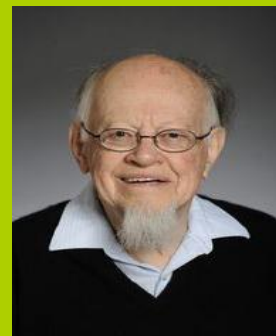
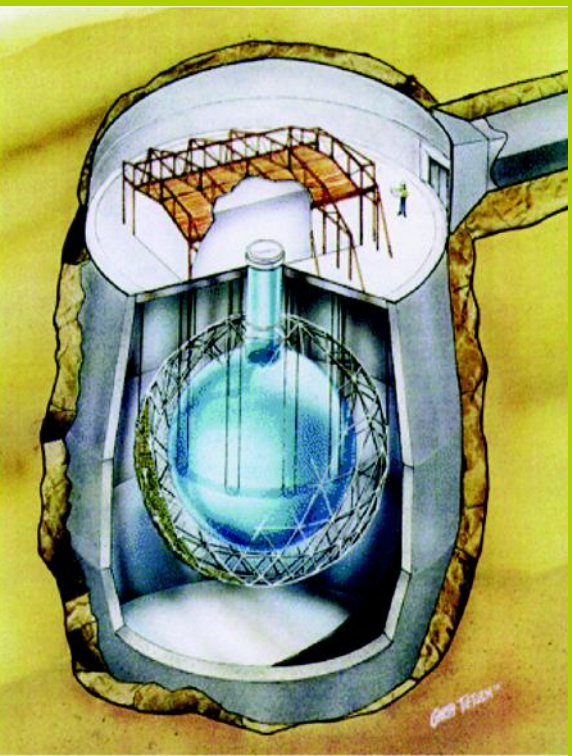
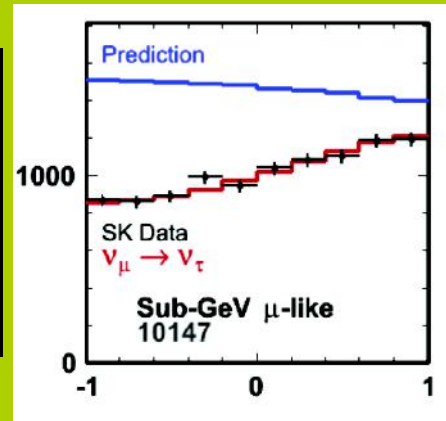
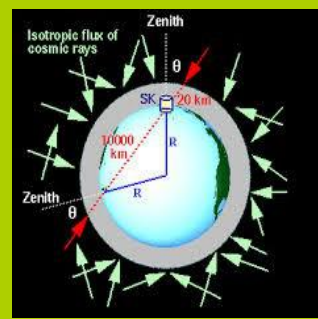
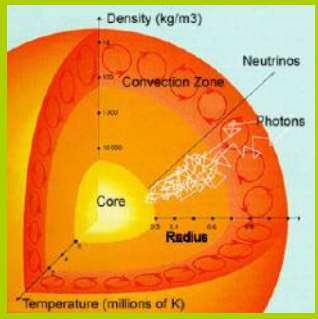
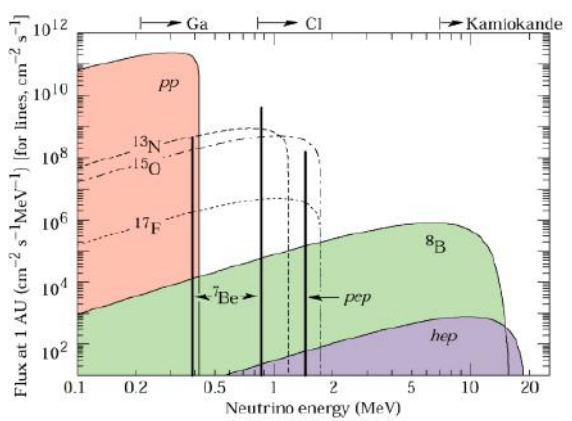
Neutrino oscillations



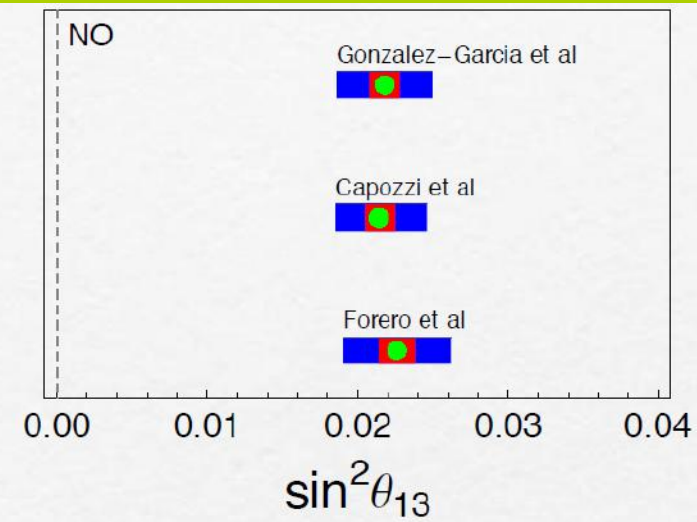
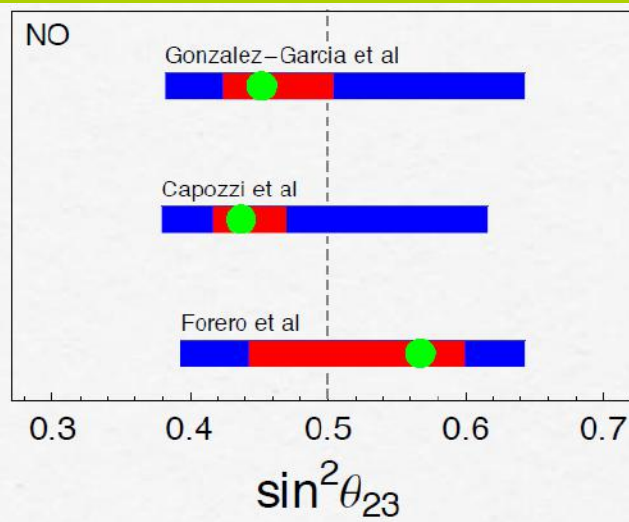
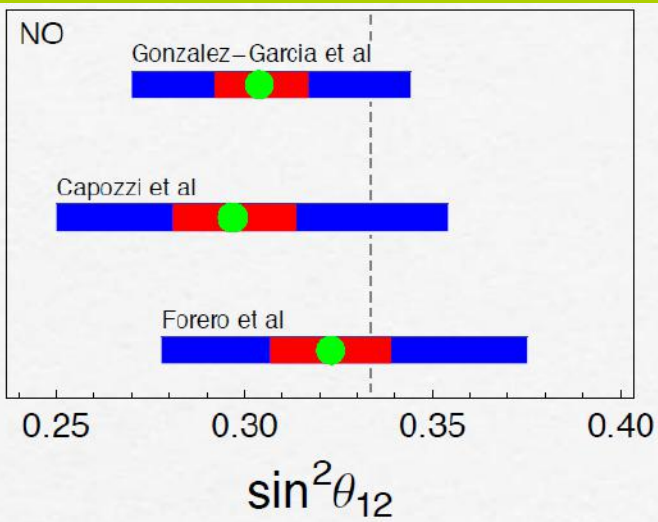
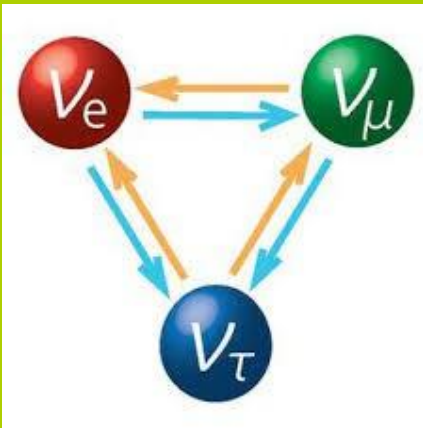
Neutrino oscillations



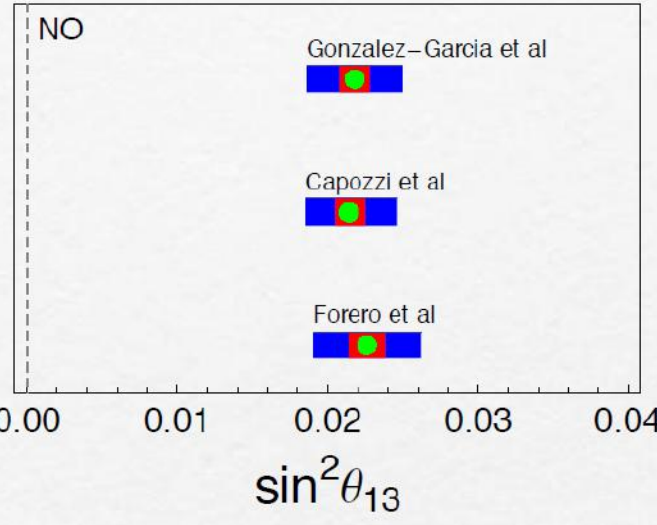
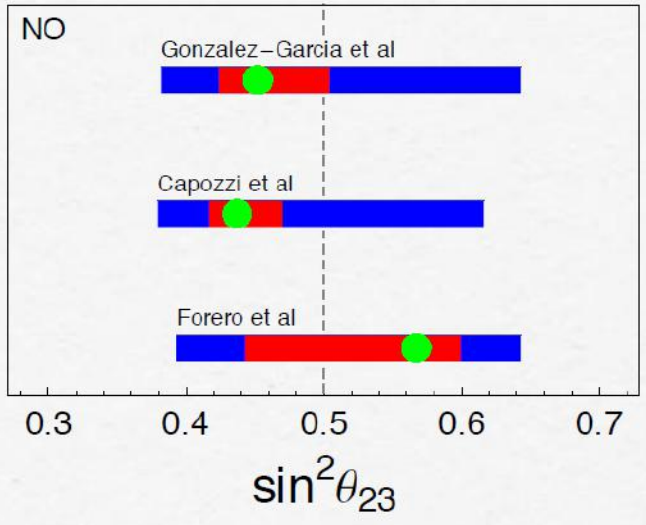
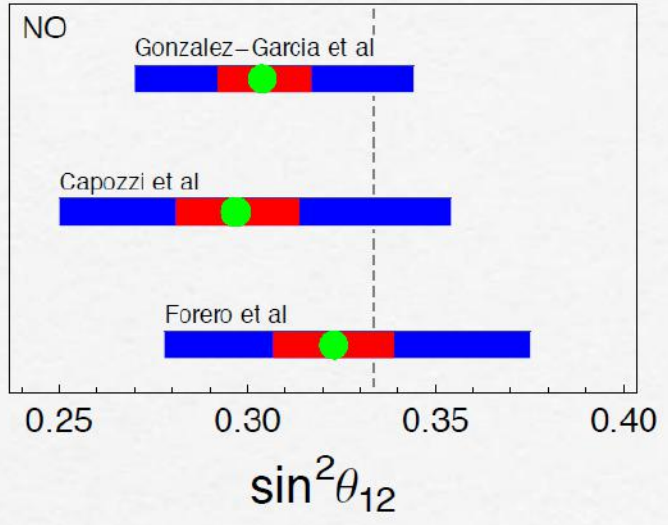
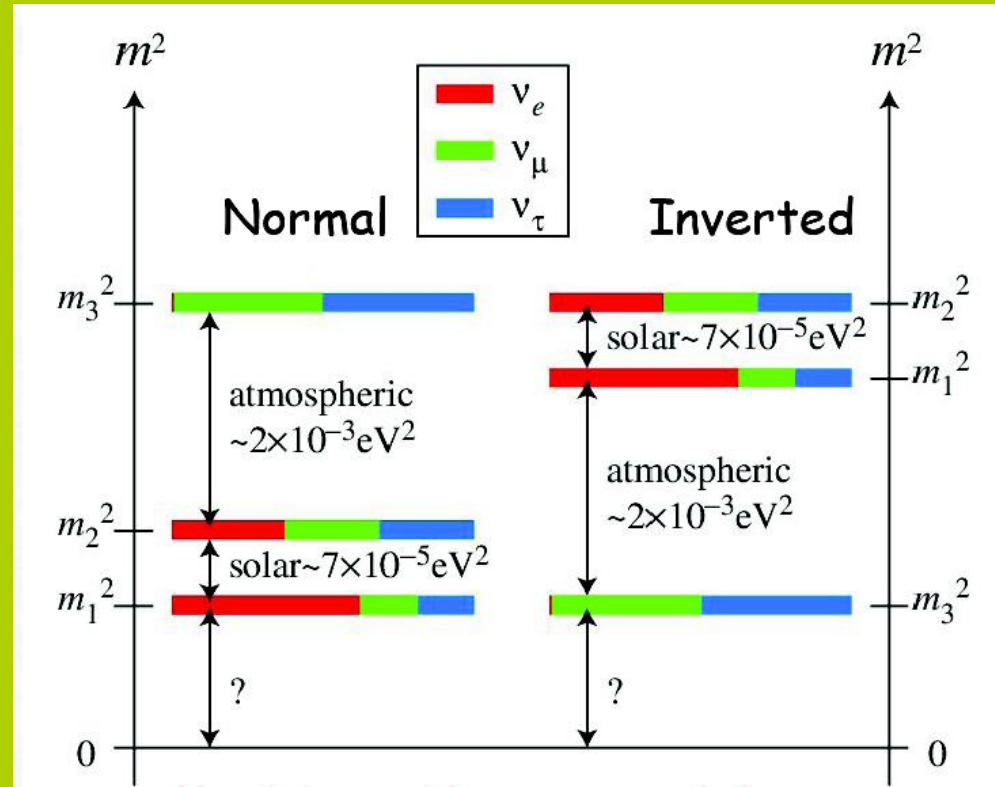
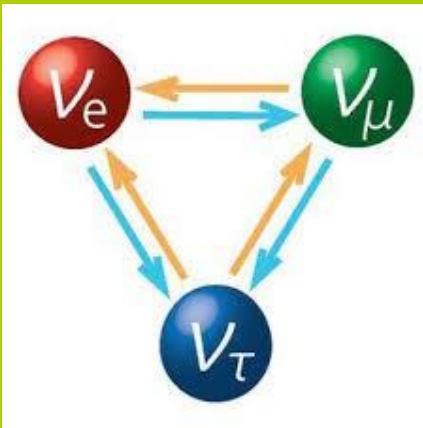
The Nobel Prize in Physics 2015



Oscillation parameters



Oscillation parameters

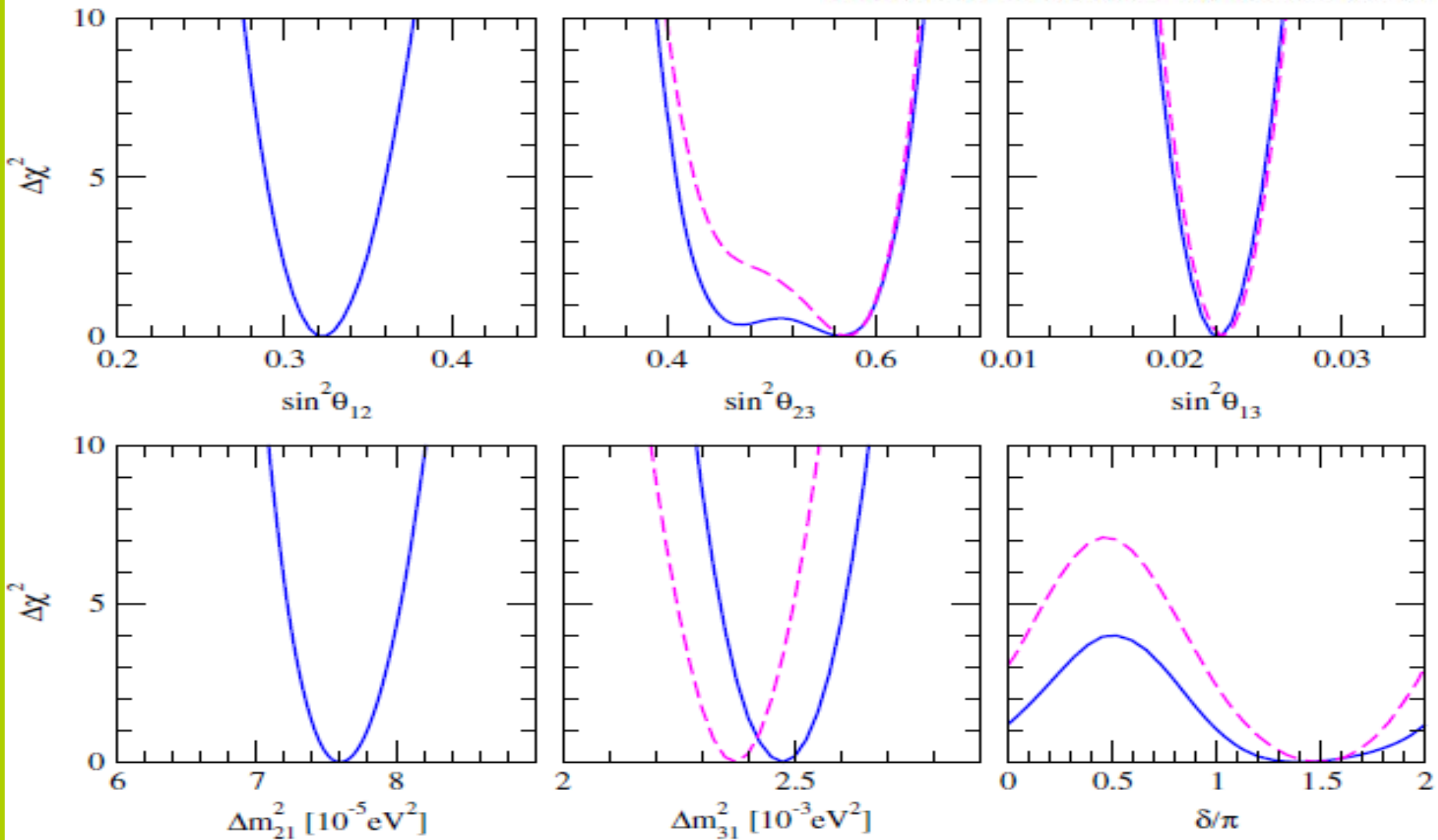




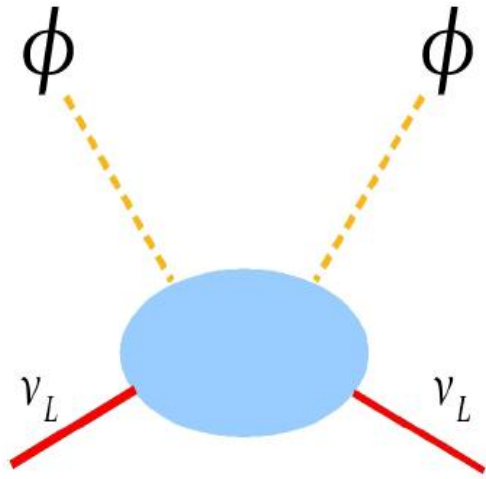
Oscillation parameters

Precision era starts

PHYSICAL REVIEW D 90, 093006 (2014)



The origin of neutrino mass

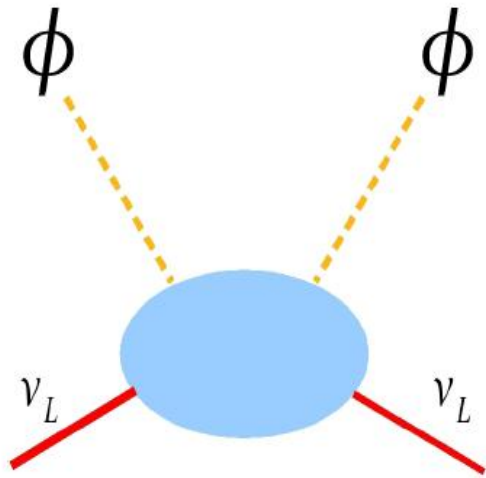


MECHANISM

SCALE

FLAVOR STRUCTURE

The origin of neutrino mass



MECHANISM

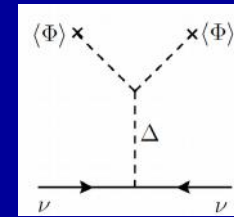
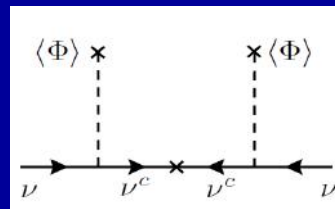
SCALE

FLAVOR STRUCTURE

Seesaw

$$v_3 v_1 \sim v_2^2$$

The origin of neutrino mass

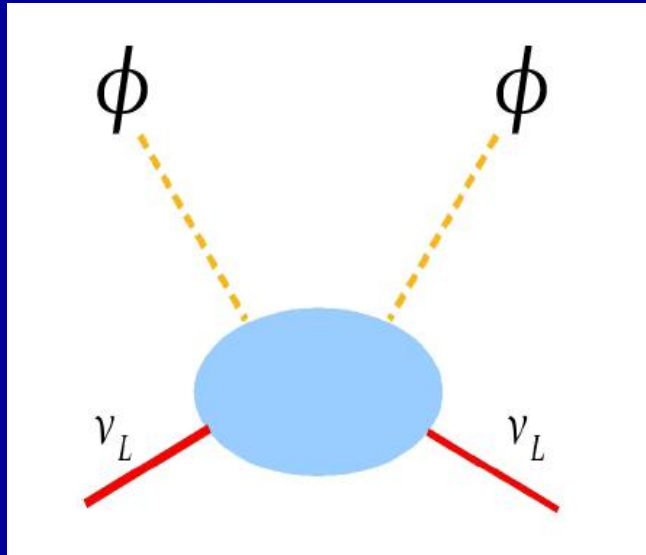


TYPE I

Minkowski 77
 Gellman Ramond Slansky 80
 Glashow, Yanagida 79
 Mohapatra Senjanovic 80
 Lazarides Shafi Weterrich 81
 Schechter-Valle, 80 & 82

TYPE II

Schechter-Valle 80/82



Seesaw

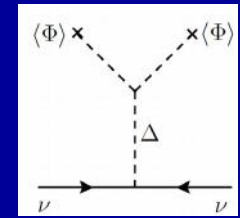
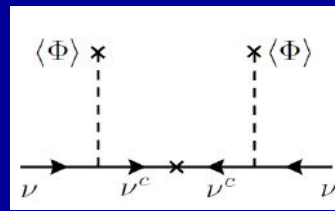
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MECHANISM

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FLAVOR STRUCTURE

The origin of neutrino mass

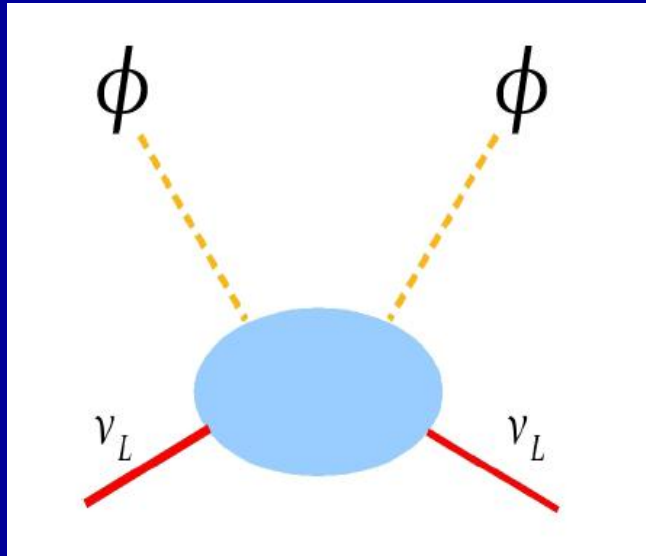


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MECHANISM

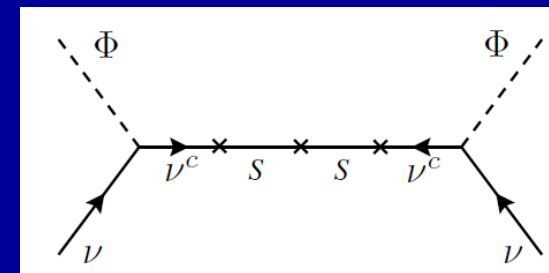
SCALE

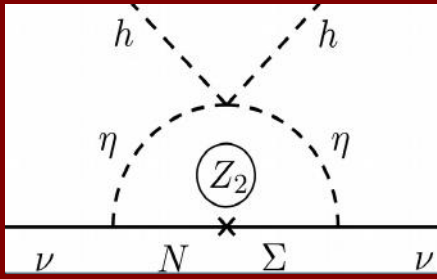
FLAVOR STRUCTURE

Number & properties of messengers

LOW-SCALE SEESAW

Mohapatra-Valle 86
 Akhmedov et al PRD53 (1996) 2752
 Malinsky et al PRL95(2005)161801
 Bazzocchi et al, PRD81 (2010) 051701

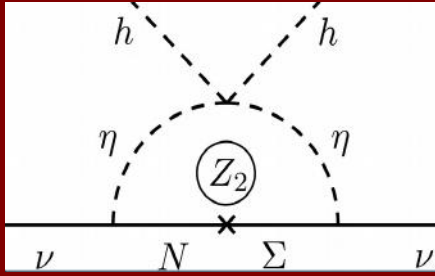




Radiative neutrino mass

many low-scale neutrino mass schemes ...

arXiv:1404.3751



Radiative neutrino mass

many low-scale neutrino mass schemes ...

arXiv:1404.3751

331 electroweak theory # generations = # colours

Singer, Valle, Schechter, Phys.Rev. D22 (1980) 738

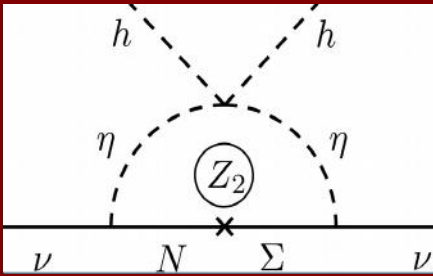
Radiative neutrino mass

many low-scale neutrino mass schemes ...

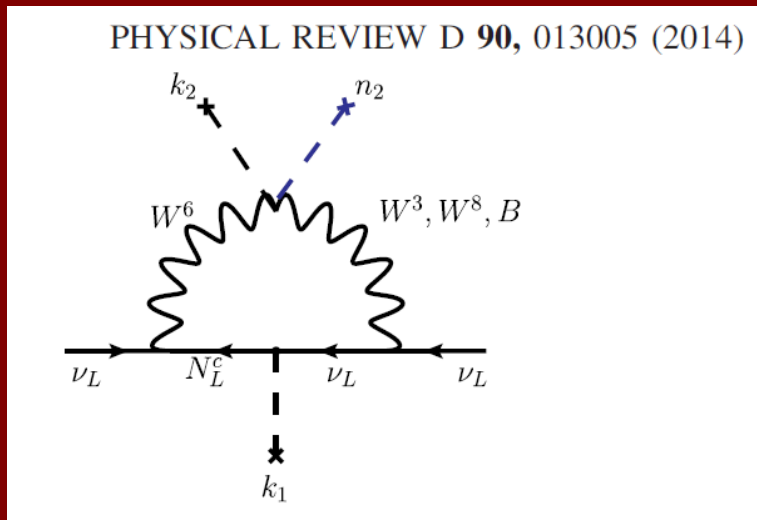
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331 electroweak theory # generations = # colours

Singer, Valle, Schechter, Phys.Rev. D22 (1980) 738



Gauge vs Higgs



BOUCENNA, MORISI, AND VALLE

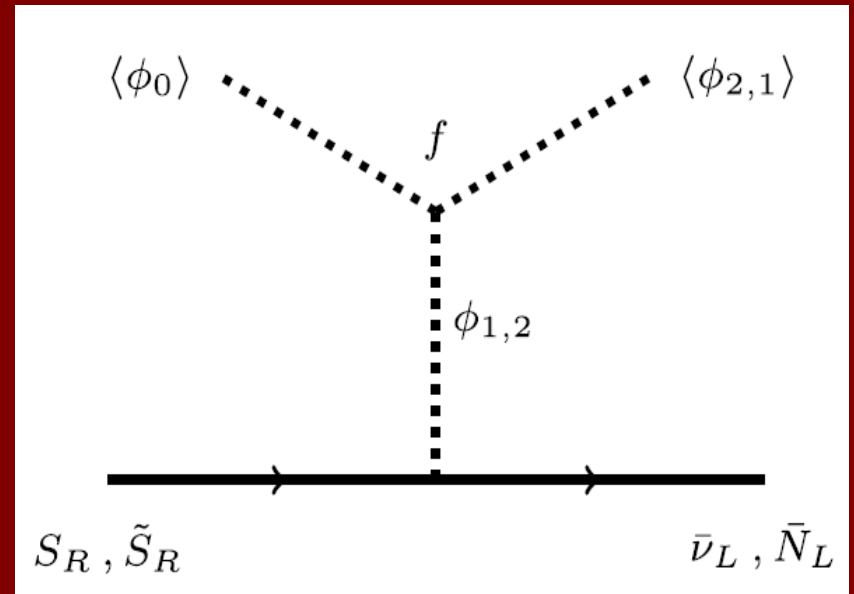
TABLE I. Matter content of the model, where $\hat{u}_R \equiv (u_R, c_R, t_R, t'_R)$ and $\hat{d}_R \equiv (d_R, s_R, b_R, d'_R, s'_R)$ (see text).

	ψ_L^c	ℓ_R	$Q_L^{1,2}$	Q_L^3	\hat{u}_R	\hat{d}_R	S	ϕ_1	ϕ_2	ϕ_3
$SU(3)_c$	1	1	3	3	3	3	1	1	1	1
$SU(3)_L$	3*	1	3	3*	1	1	1	3*	3*	3*
$U(1)_X$	$-\frac{1}{3}$	-1	0	$+\frac{1}{3}$	$+\frac{2}{3}$	$-\frac{1}{3}$	0	$+\frac{2}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$
\mathcal{L}	$-\frac{1}{3}$	-1	$-\frac{2}{3}$	$+\frac{2}{3}$	0	0	1	$+\frac{2}{3}$	$-\frac{4}{3}$	$+\frac{2}{3}$

$$Q = T_3 + \frac{1}{\sqrt{3}}T_8 + X, \quad (2)$$

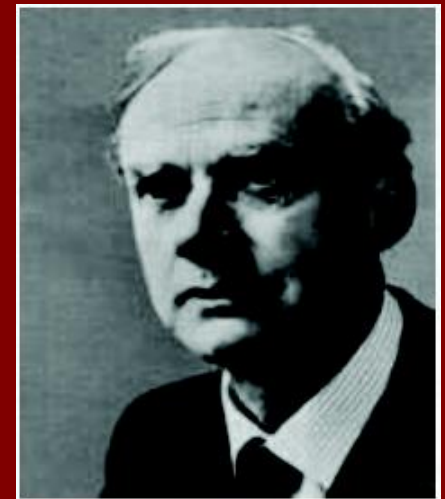
$$L = \frac{4}{\sqrt{3}}T_8 + \mathcal{L}. \quad (3)$$

Type-II Dirac seesaw



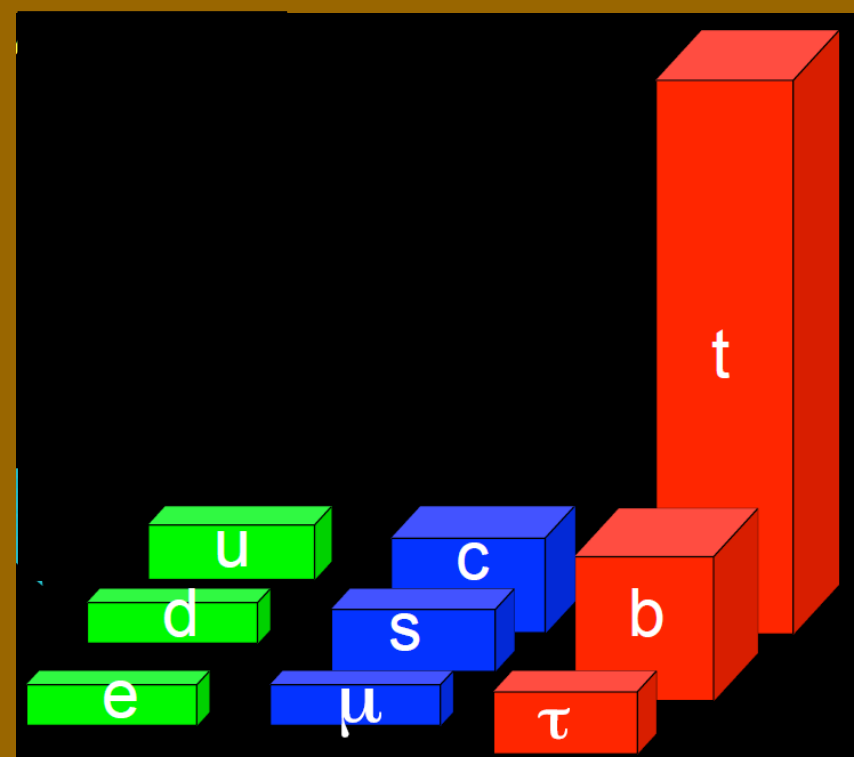
Physics Letters B 755 (2016) 363–366

arXiv:1605.08362



Flavor puzzle

masses



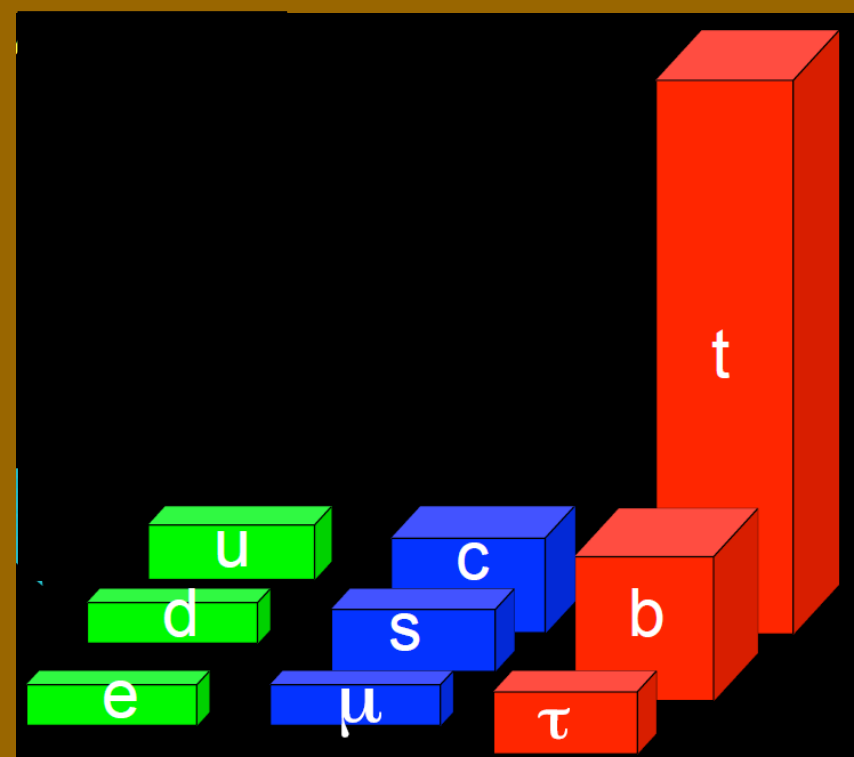
Neutrinos : Lepton number?

mixings

Flavor puzzle

masses

b-tau unification



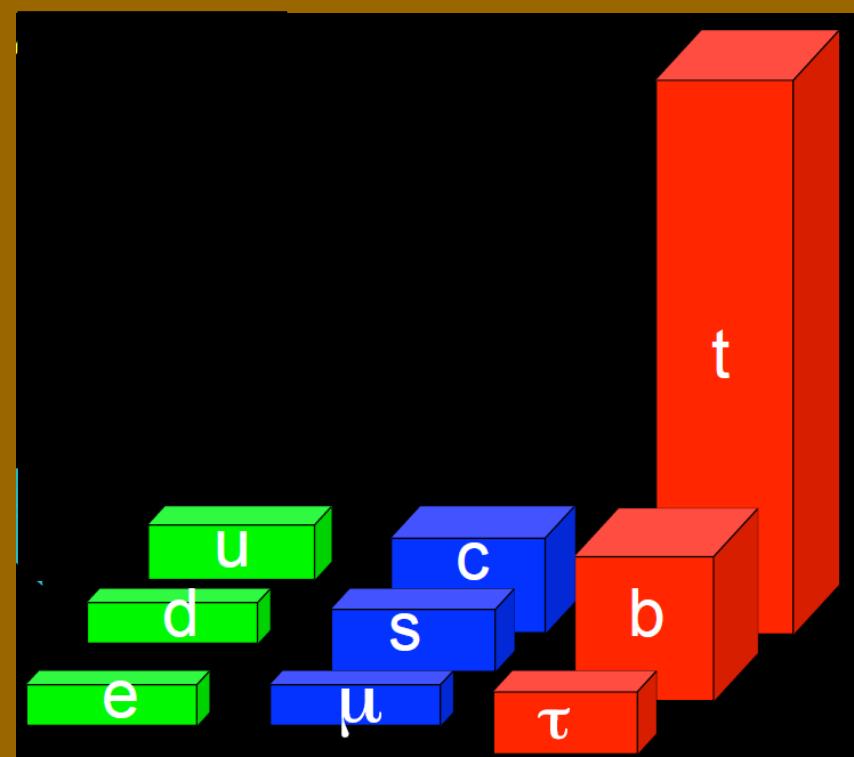
Neutrinos : Lepton number?

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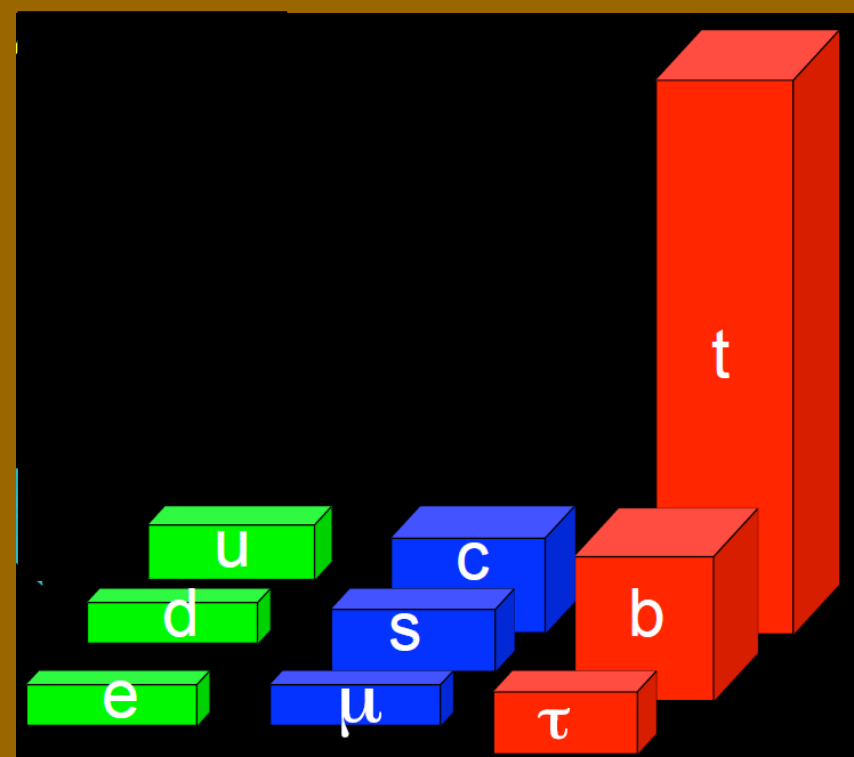
Neutrinos : Lepton number?

mixings

Flavor puzzle

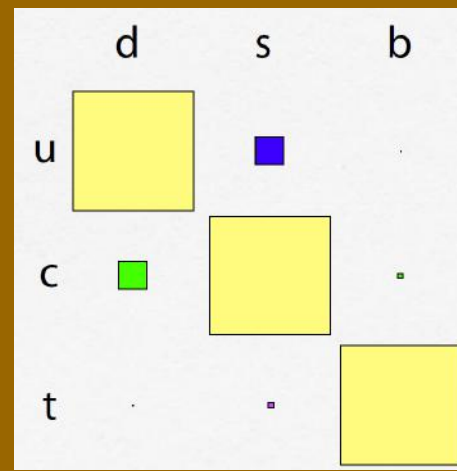
masses

b-tau unification

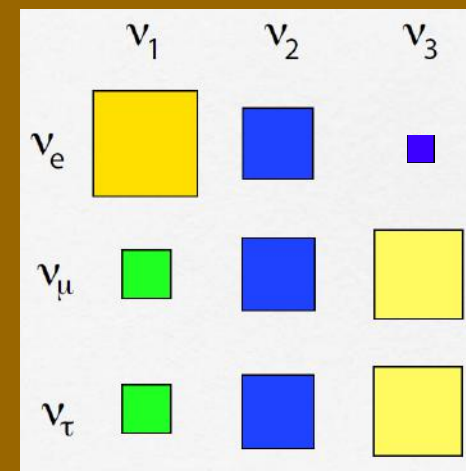


Neutrinos : Lepton number?

mixings

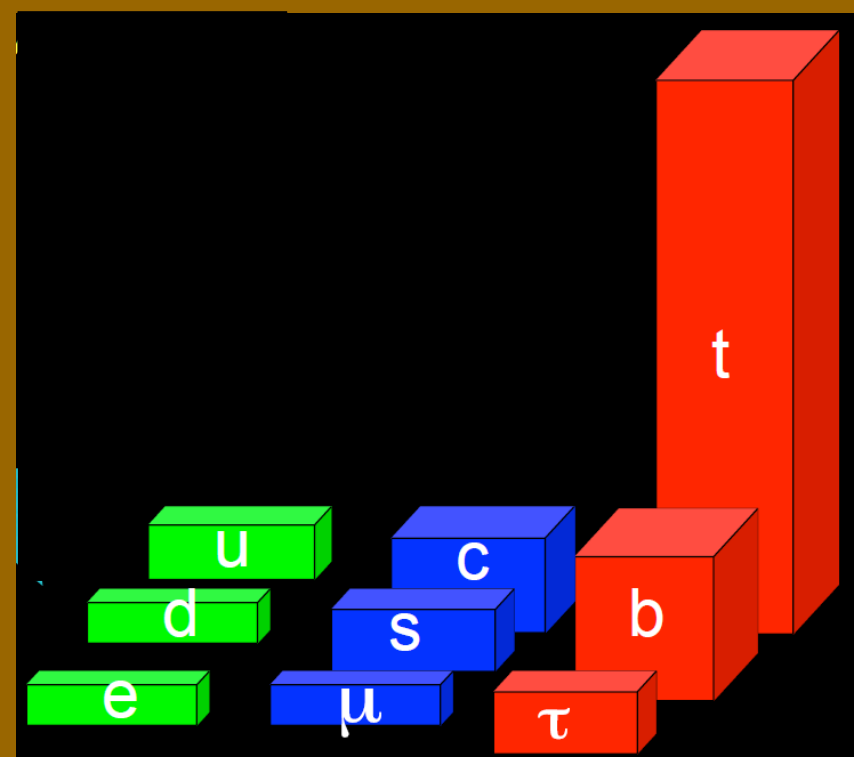


VS



Flavor puzzle

masses



b-tau unification without GUTS

Neutrinos : Lepton number?

$$\frac{m_\tau}{\sqrt{m_e m_\mu}} \approx \frac{m_b}{\sqrt{m_d m_s}}$$

mixings

Morisi et al Phys.Rev. D84 (2011) 036003

King et al

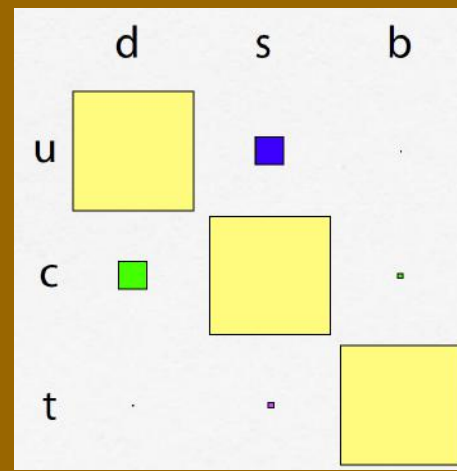
Phys. Lett. B 724 (2013) 68

Morisi et al

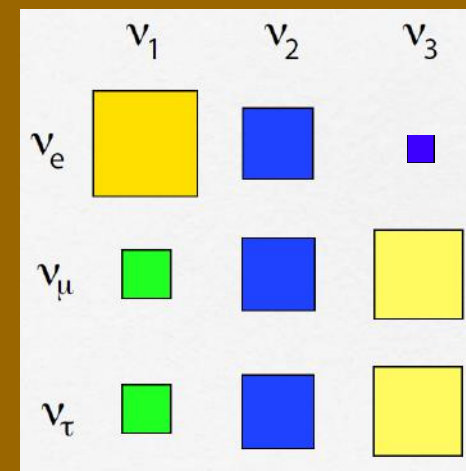
Phys.Rev. D88 (2013) 036001

Bonilla et al

Phys.Lett. B742 (2015) 99

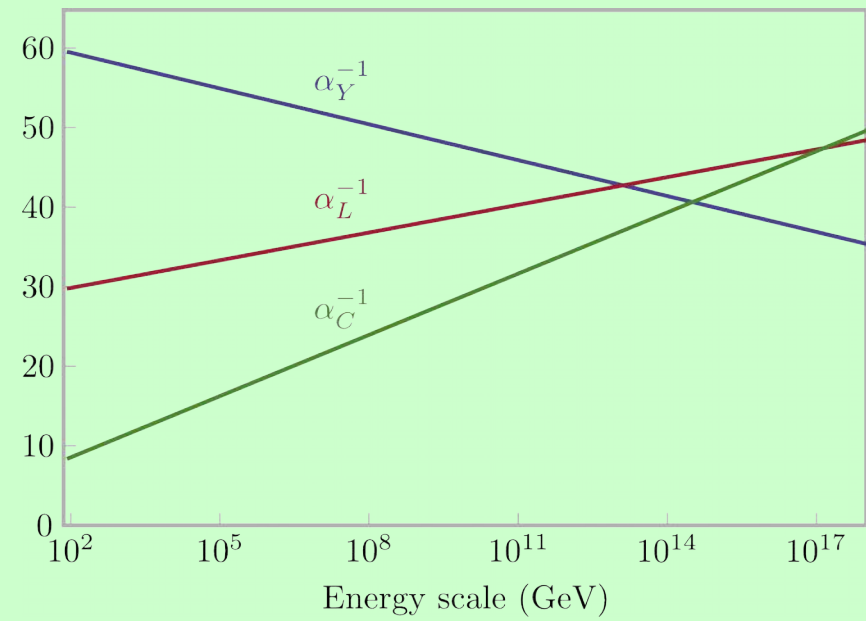


νs



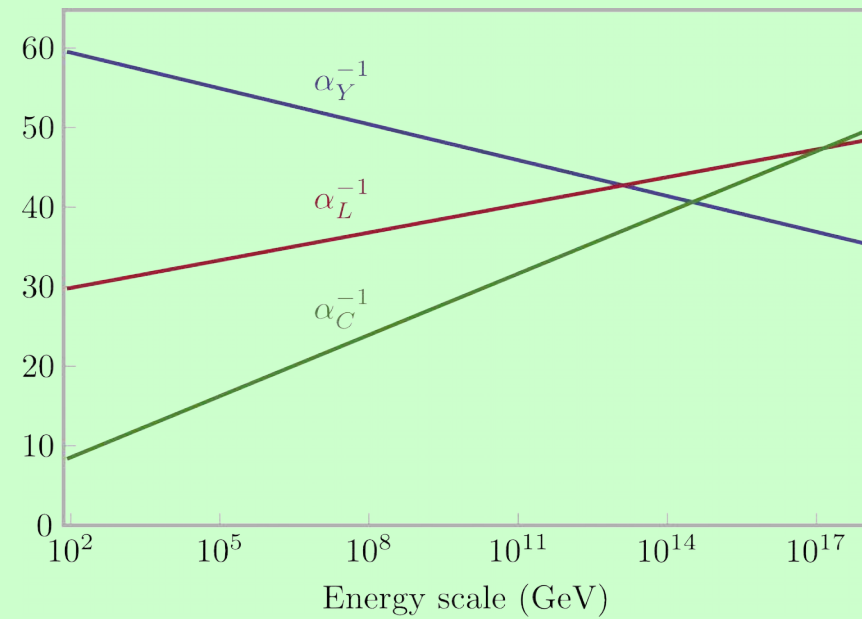
Gauge coupling unification

a near miss ...



Gauge coupling unification

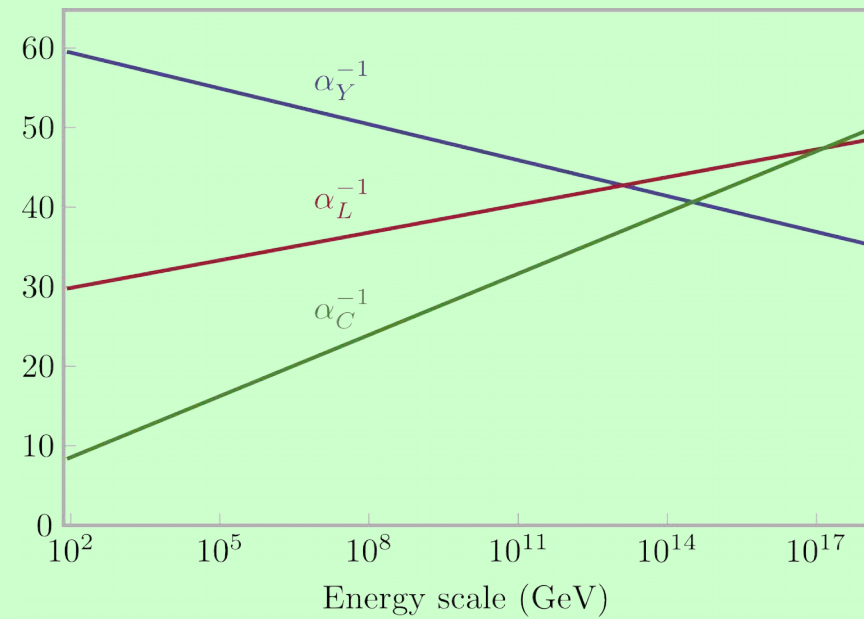
a near miss ...



What makes the gauge couplings unify? **GUT (p decay)**
SUSY

Gauge coupling unification

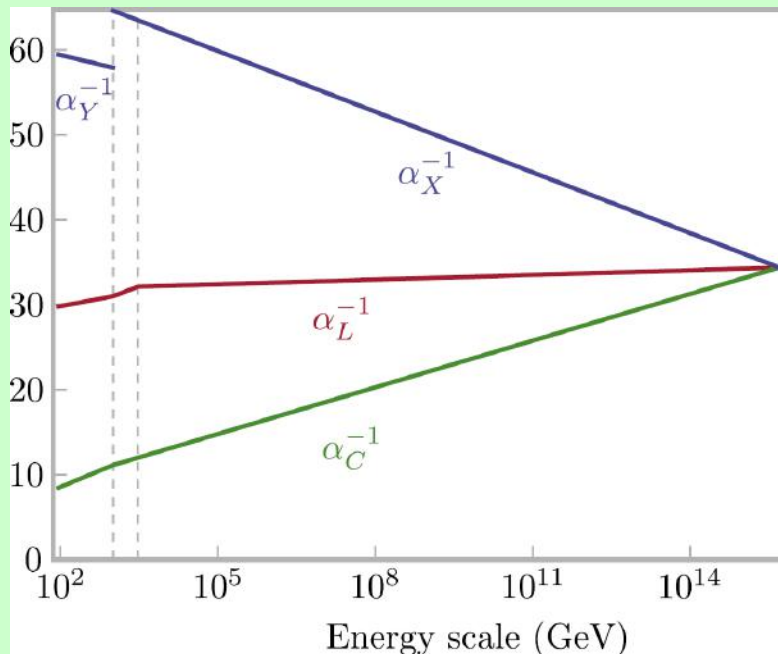
a near miss ...



What makes the gauge couplings unify? GUT (p decay)

SUSY

neutrino



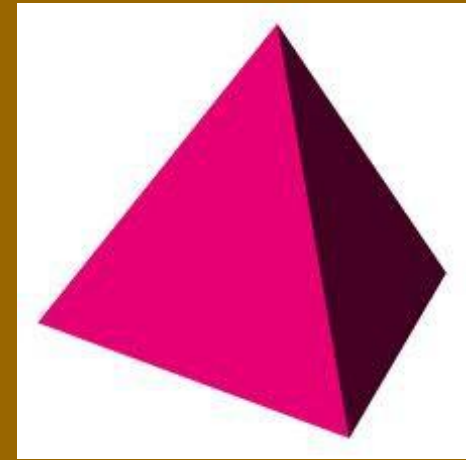
The physics responsible for gauge coupling unification may also induce small neutrino masses

Phys. Rev. D 91, 031702 (2015)

Boucenna, Fonseca, Gonzalez-Canales, JV

Valle

Flavor Symmetry



$\begin{pmatrix} \nu_e \\ e \\ e_R \end{pmatrix}_L$	$\begin{pmatrix} \nu_\mu \\ \mu \\ \mu_R \end{pmatrix}_L$	$\begin{pmatrix} \nu_\tau \\ \tau \\ \tau_R \end{pmatrix}_L$
$\begin{pmatrix} u \\ d \\ u_R \\ d_R \end{pmatrix}_L$	$\begin{pmatrix} c \\ s \\ c_R \\ s_R \end{pmatrix}_L$	$\begin{pmatrix} t \\ b \\ t_R \\ b_R \end{pmatrix}_L$

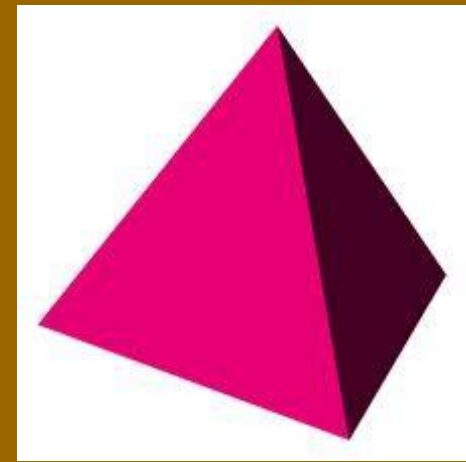
A4

Babu-Ma-Valle PLB552 (2003) 207
Hirsch et al PRD69 (2004) 093006

$$\sin^2 \theta_{23} = 0.5$$

$$\sin^2 \theta_{13} = 0$$

Flavor Symmetry



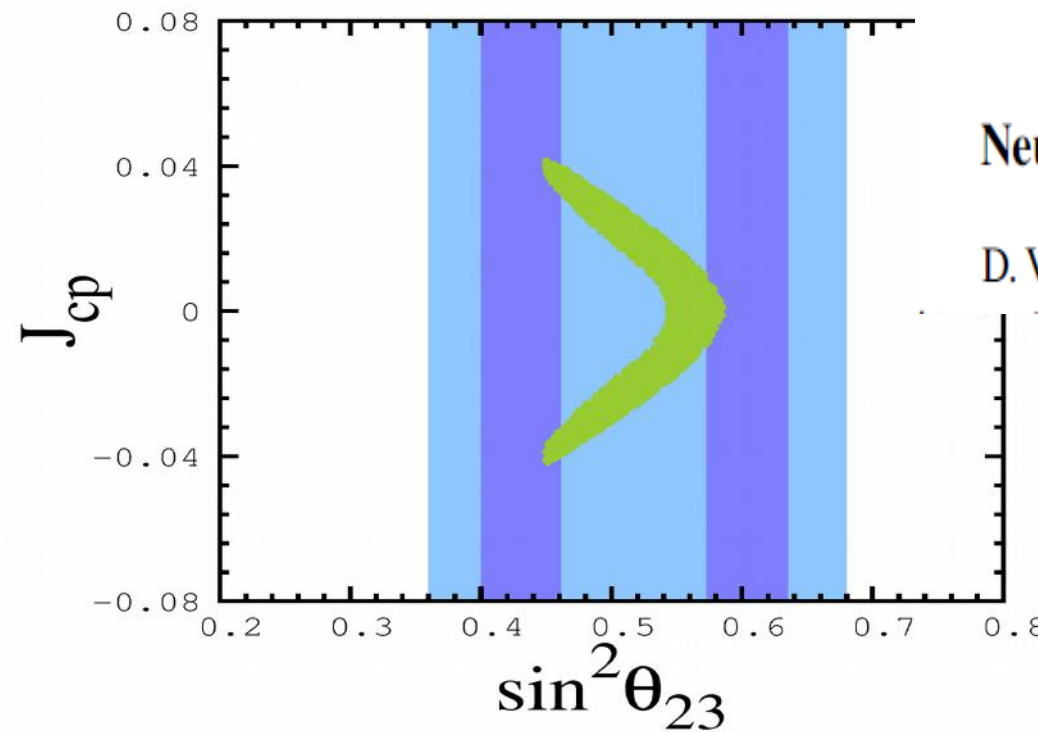
$\begin{pmatrix} \nu_e \\ e \\ e_R \end{pmatrix}_L$	$\begin{pmatrix} \nu_\mu \\ \mu \\ \mu_R \end{pmatrix}_L$	$\begin{pmatrix} \nu_\tau \\ \tau \\ \tau_R \end{pmatrix}_L$
$\begin{pmatrix} u \\ d \\ u_R \\ d_R \end{pmatrix}_L$	$\begin{pmatrix} c \\ s \\ c_R \\ s_R \end{pmatrix}_L$	$\begin{pmatrix} t \\ b \\ t_R \\ b_R \end{pmatrix}_L$

A4

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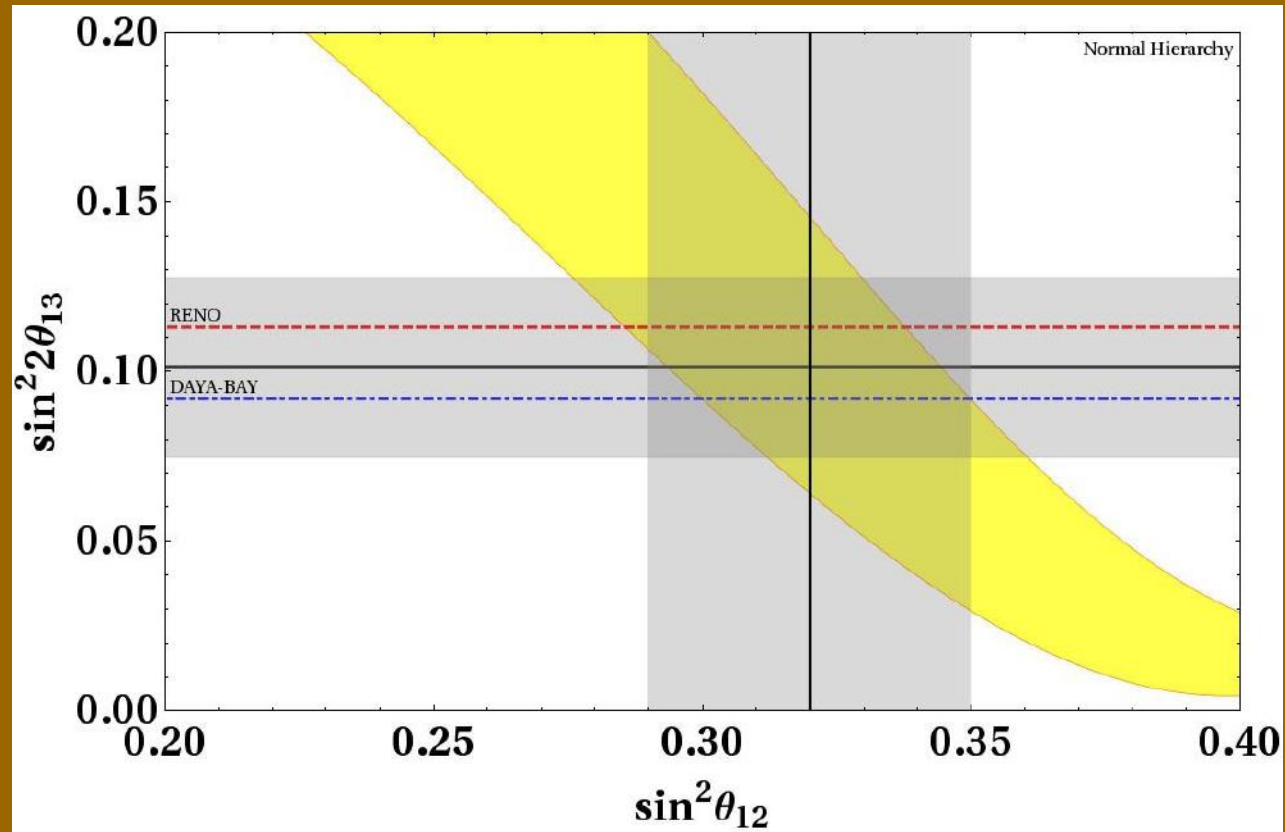
PHYSICAL REVIEW D 88, 016003 (2013)

Neutrino mixing with revamped A4 flavor symmetry

D. V. Forero,^{1,2,*} S. Morisi,^{3,†} J. C. Romão,^{1,‡} and J. W. F. Valle^{2,§}

Flavor correlations

Boucenna et al
PhysRevD.86.073008



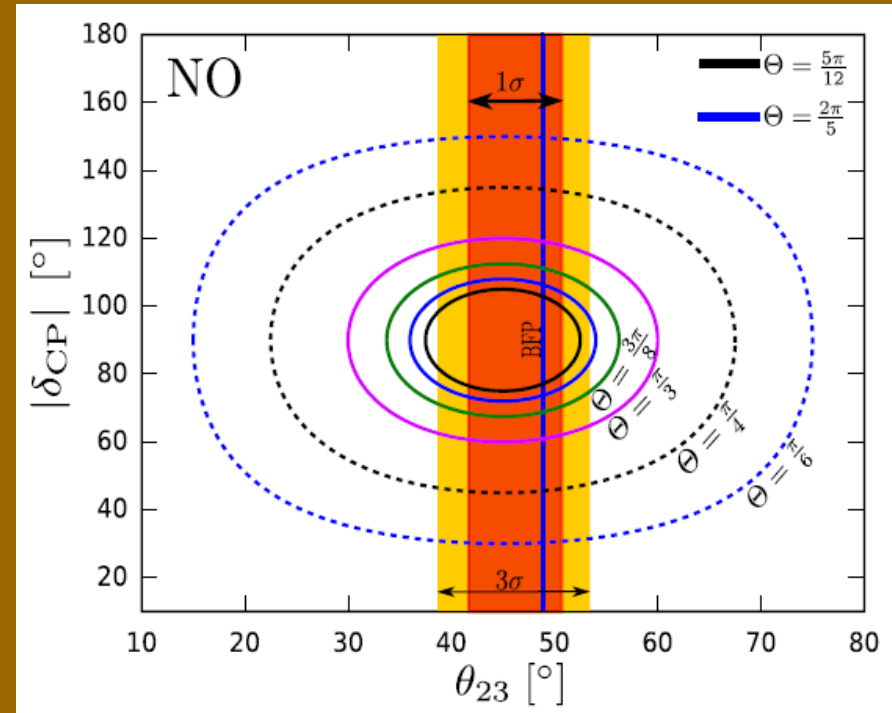
Valle

Model-independent flavor approach

$$\mathbf{X}^T \mathbf{m}_\nu \mathbf{X} = \mathbf{m}_\nu^*$$

*Predicting neutrino mixing
from residual CP symmetries*

P. Chen et al. / Physics Letters B 753 (2016) 644–652



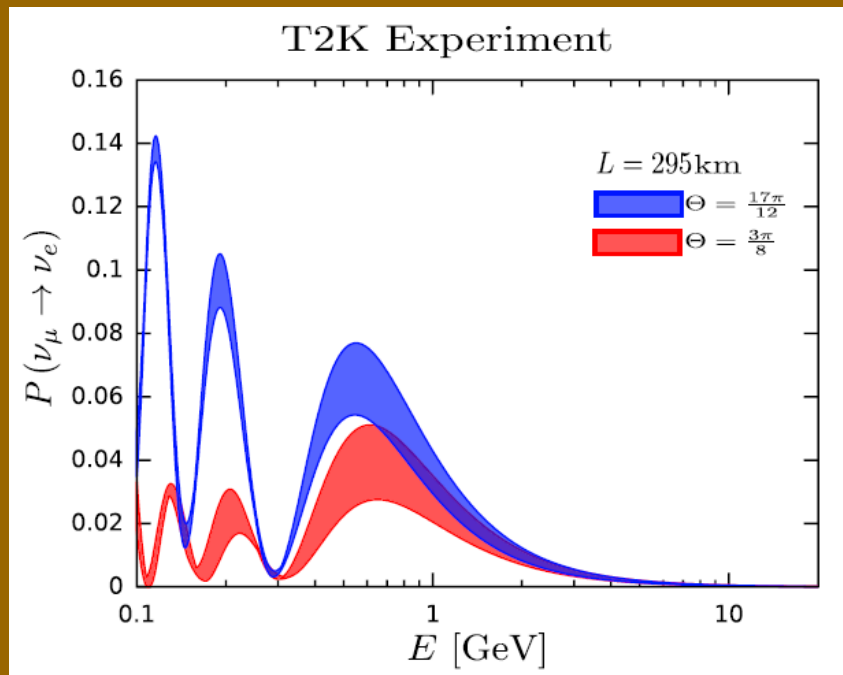
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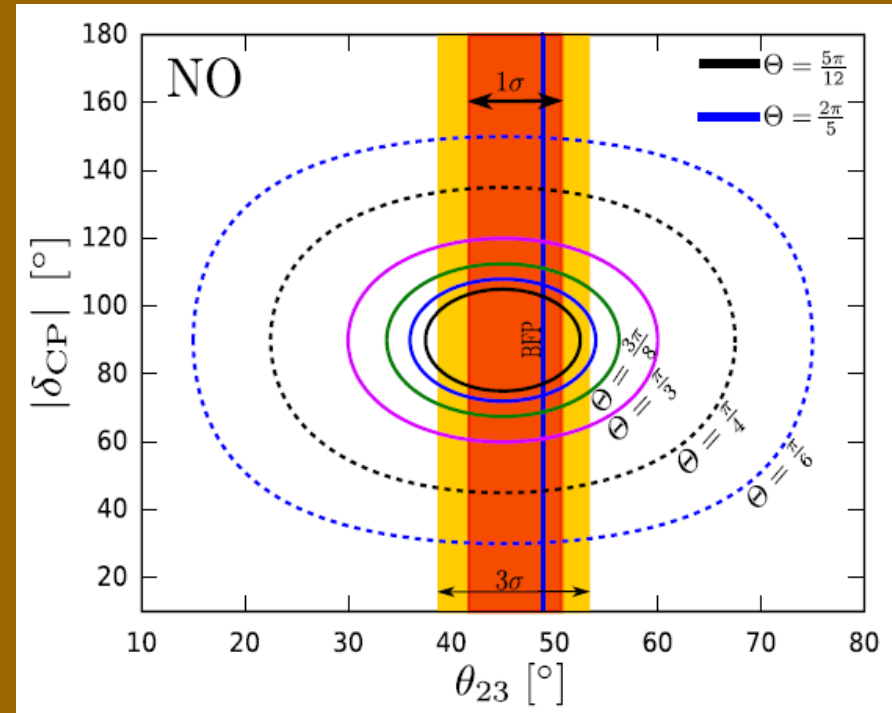


*Predicting neutrino mixing
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P. Chen et al. / Physics Letters B 753 (2016) 644–652

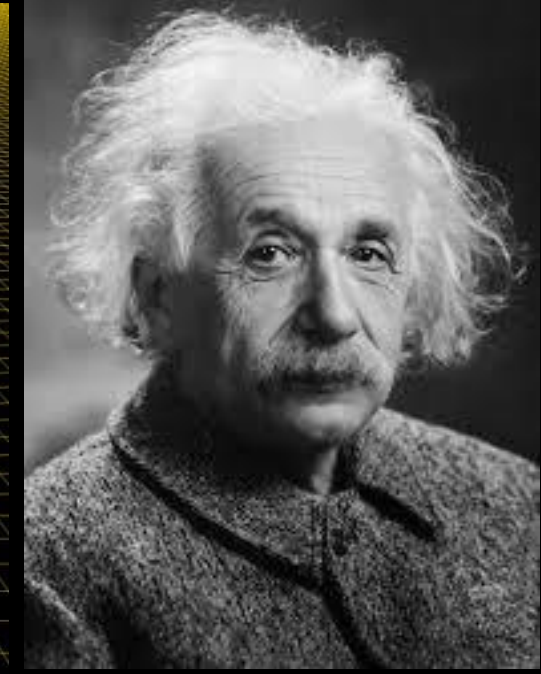
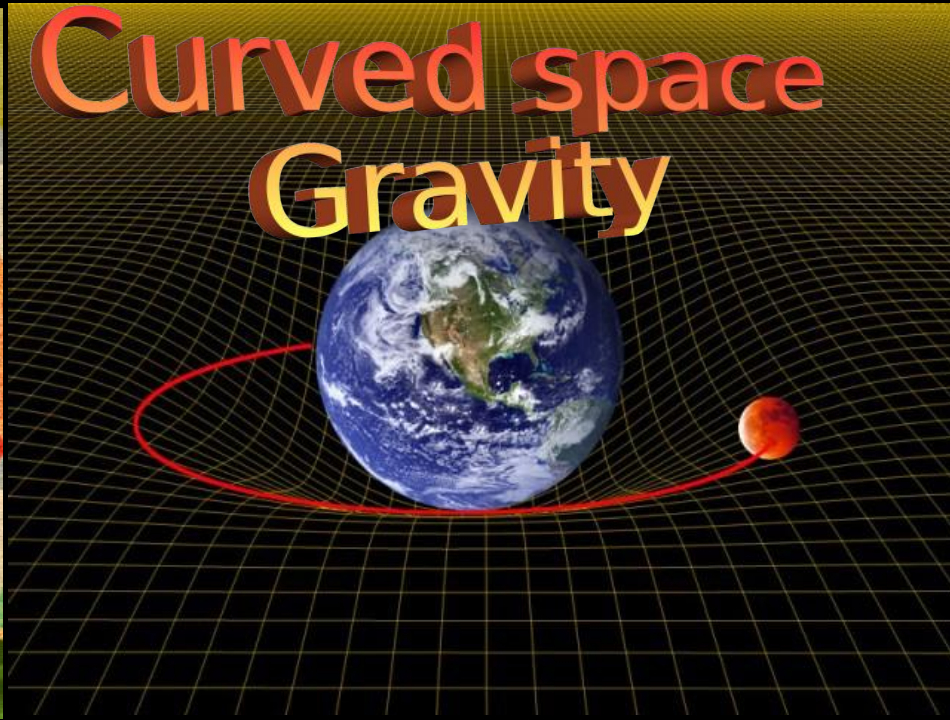
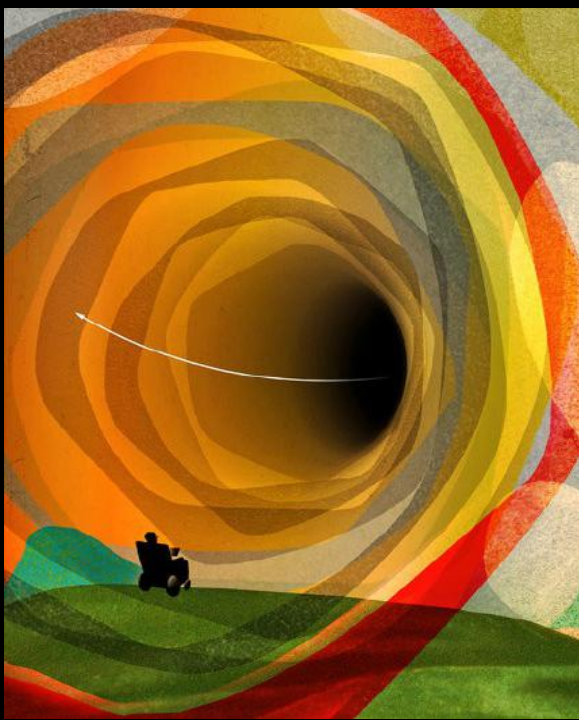


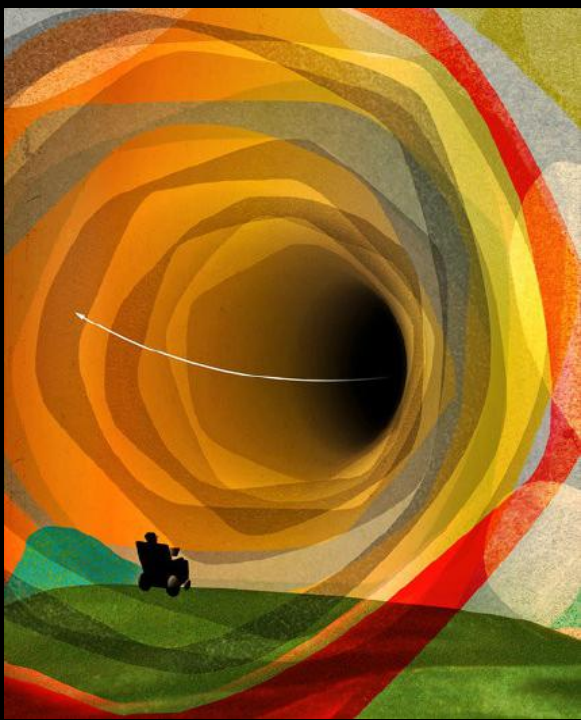
NOVA, DUNE



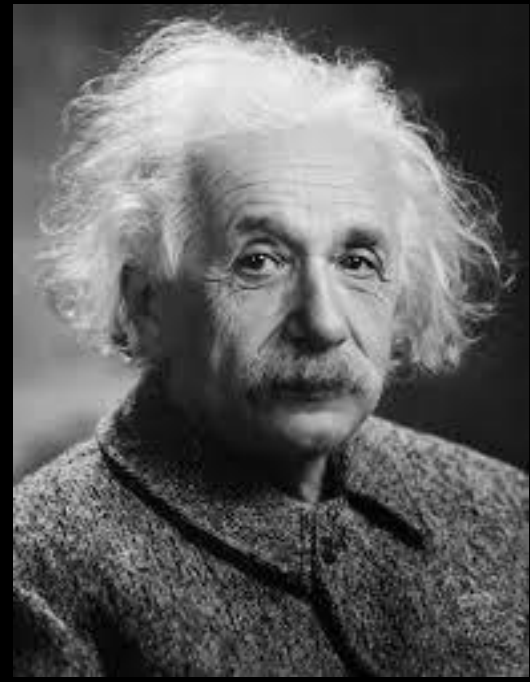


How about gravity?

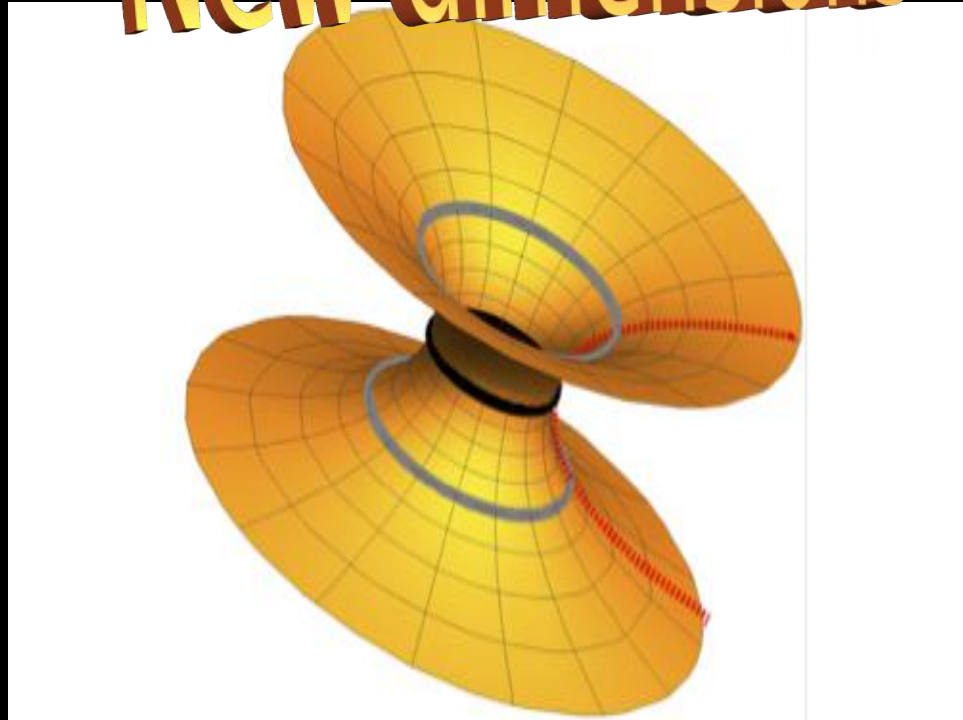




Curved space Gravity



New dimensions



Warped flavor

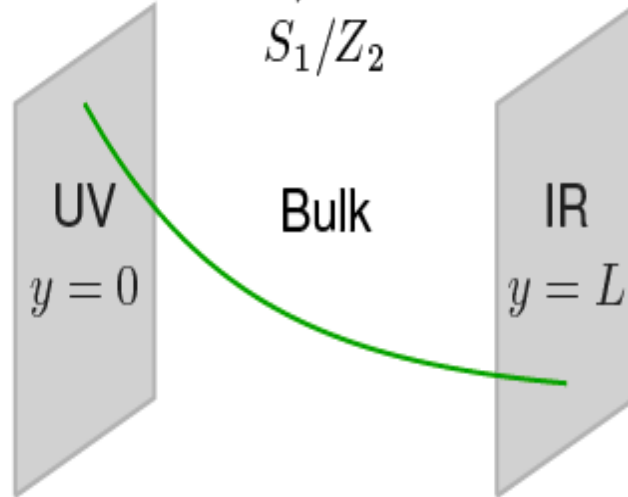
Chen et al arXiv:1509.06683

JHEP01(2016)007

Randall-Sundrum brane-world

$$ds^2 = e^{-2ky} \eta_{\mu\nu} dx^\mu dx^\nu - dy^2$$

S_1/Z_2



5D field theory on a slice of AdS_5

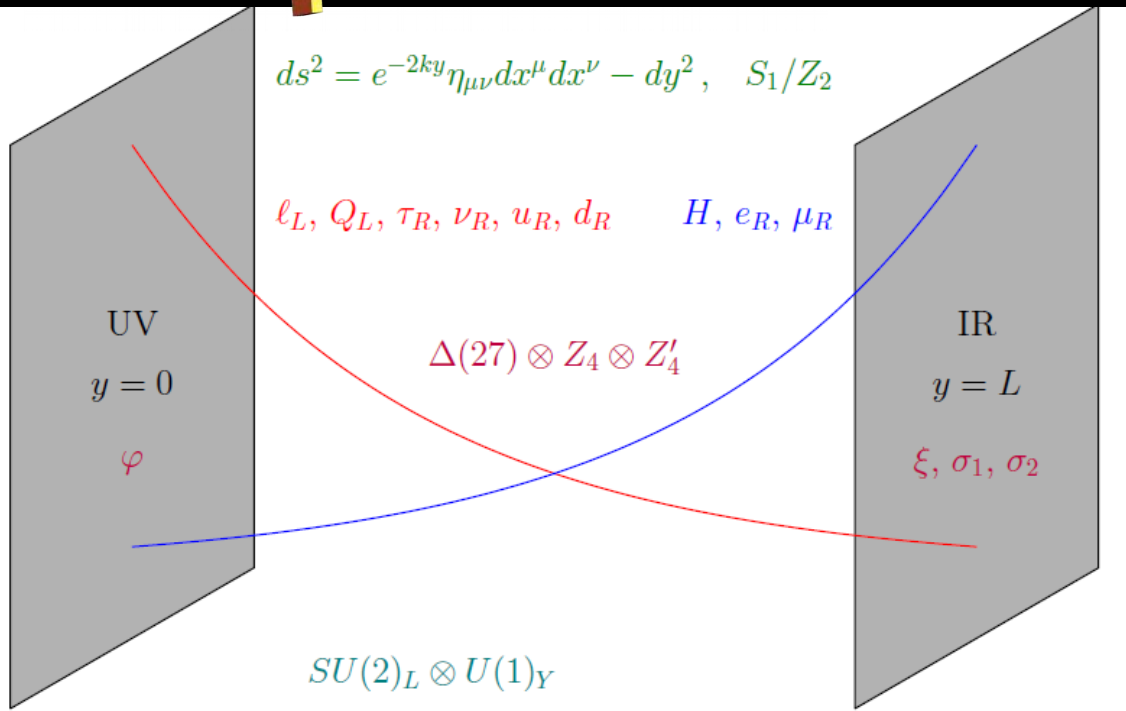
Energy scales warped down as one approaches IR

Localizing the Higgs at the IR brane and setting $kL \simeq 35$ explains the smallness of the EW scale!

Warped flavor

Chen et al arXiv:1509.06683

JHEP01(2016)007

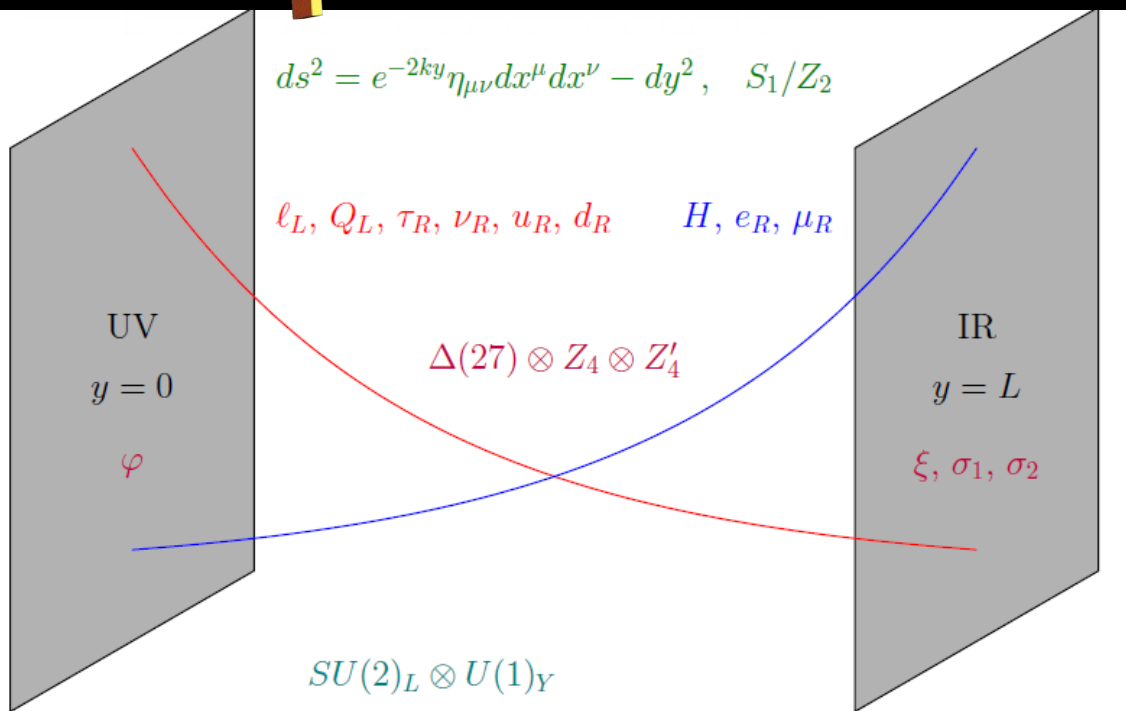


Mass hierarchies in principle accounted for by judicious bulk parameter choices

Warped flavor

Chen et al arXiv:1509.06683

JHEP01(2016)007



Mass hierarchies in principle accounted for by judicious bulk parameter choices

$$\sin^2 \theta_{12} \cos^2 \theta_{13} = 1/3$$

4 neutrino mixing angles & CP phase in terms of 2

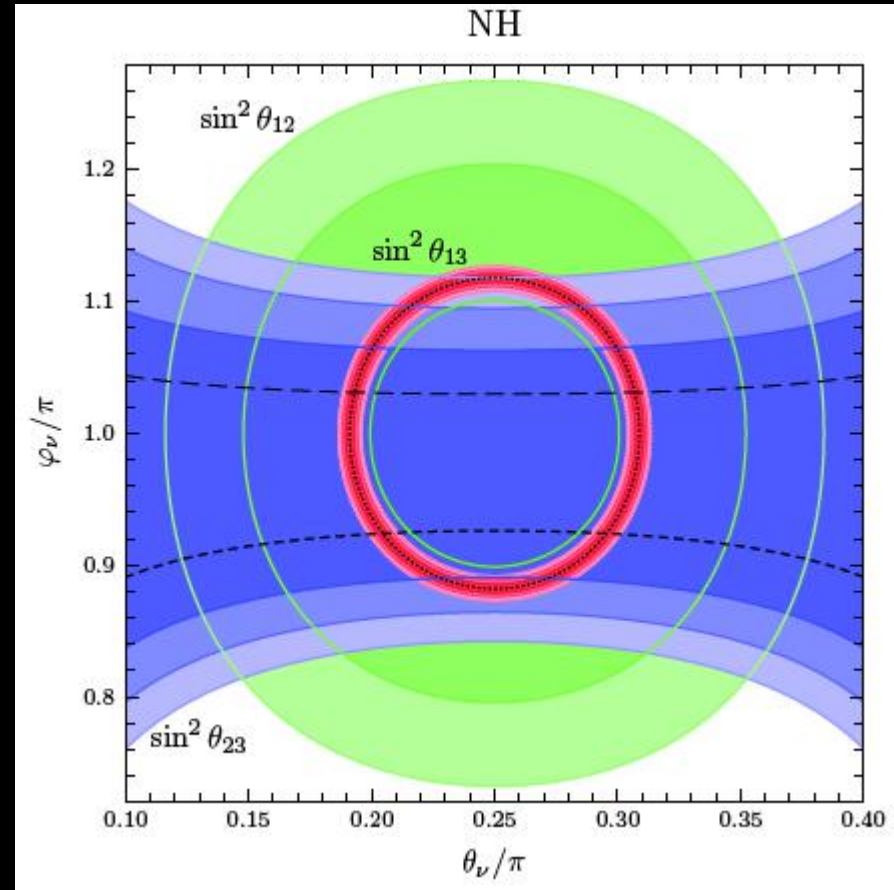
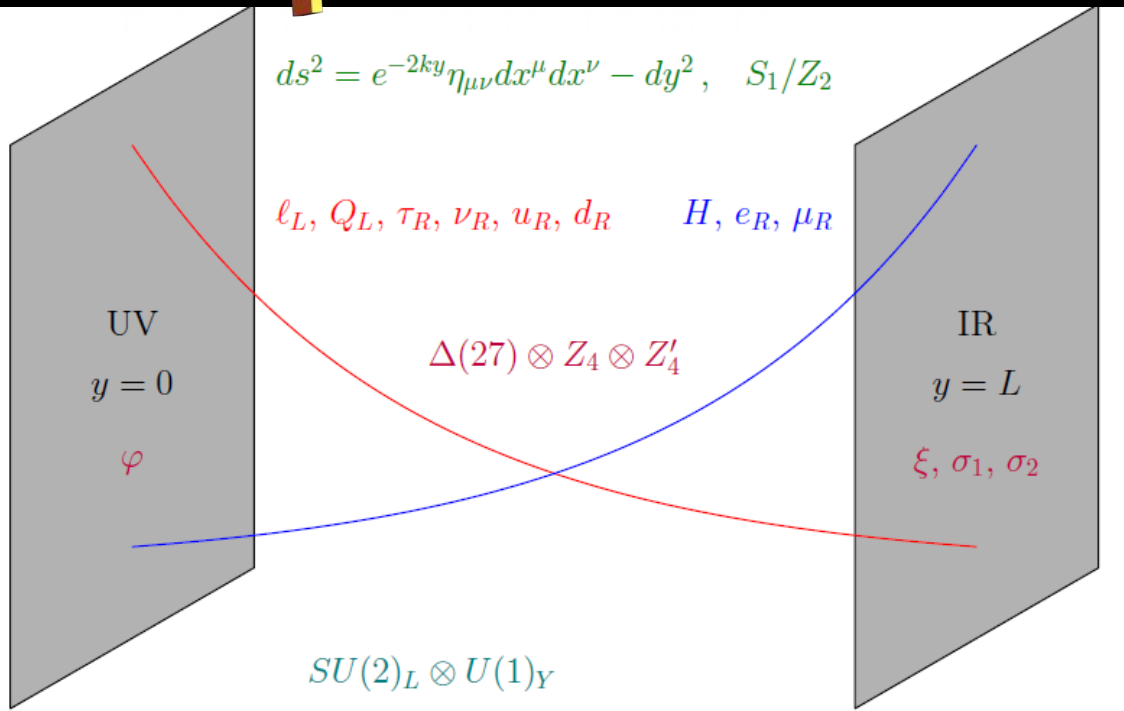
$$\sin^2 \theta_{23} = \frac{1 - \sin 2\theta_\nu \sin(\pi/6 - \varphi_\nu)}{2 - \sin 2\theta_\nu \cos \varphi_\nu}$$

$$J_{CP} = -\frac{1}{6\sqrt{3}} \cos 2\theta_\nu .$$

Warped flavor

Chen et al arXiv:1509.06683

JHEP01(2016)007



Mass hierarchies in principle accounted for by judicious bulk parameter choices

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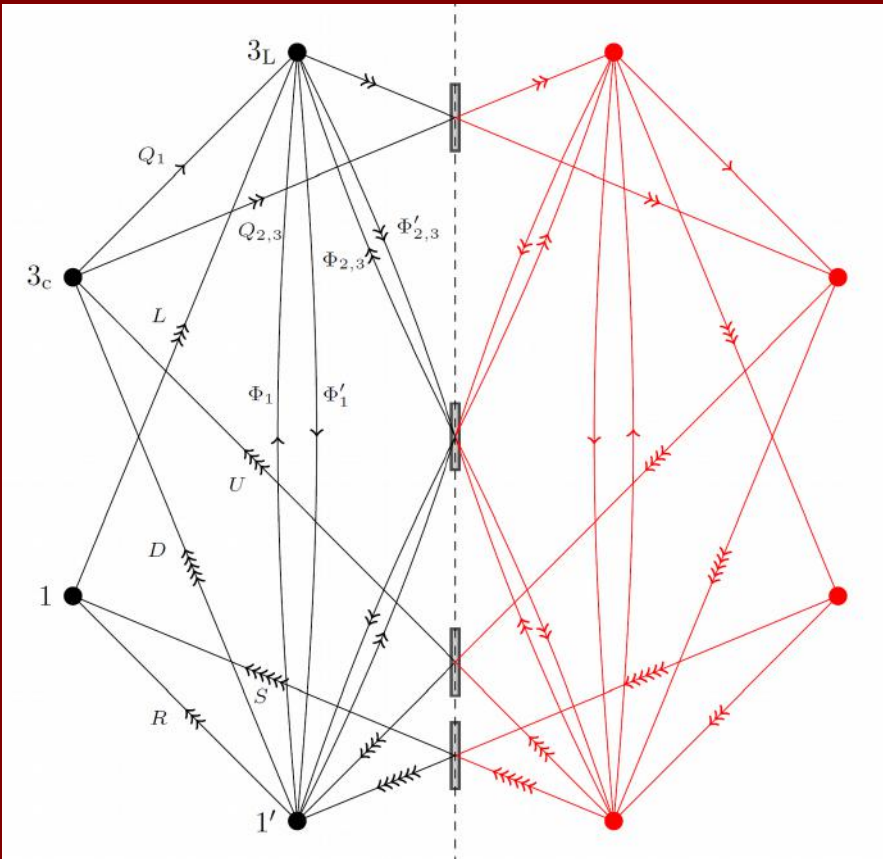
$$\sin^2 \theta_{12} \cos^2 \theta_{13} = 1/3$$

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$$J_{CP} = -\frac{1}{6\sqrt{3}} \cos 2\theta_\nu .$$

String completion of 331

Addazi et al Phys.Lett. B759 (2016) 471-478



string completion of EW model with **No** conventional GUT embedding

Quiver set-up free of gauge and string anomalies **iff** extra RH states

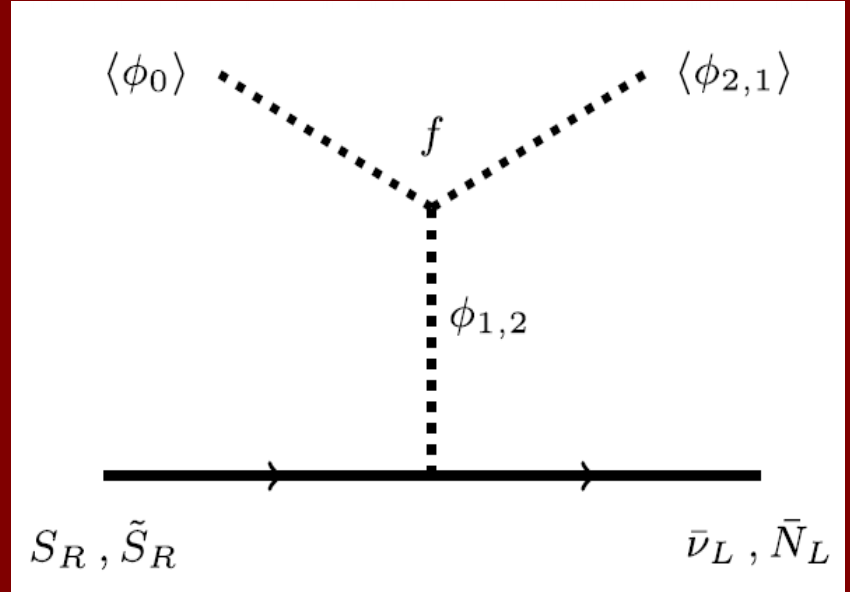
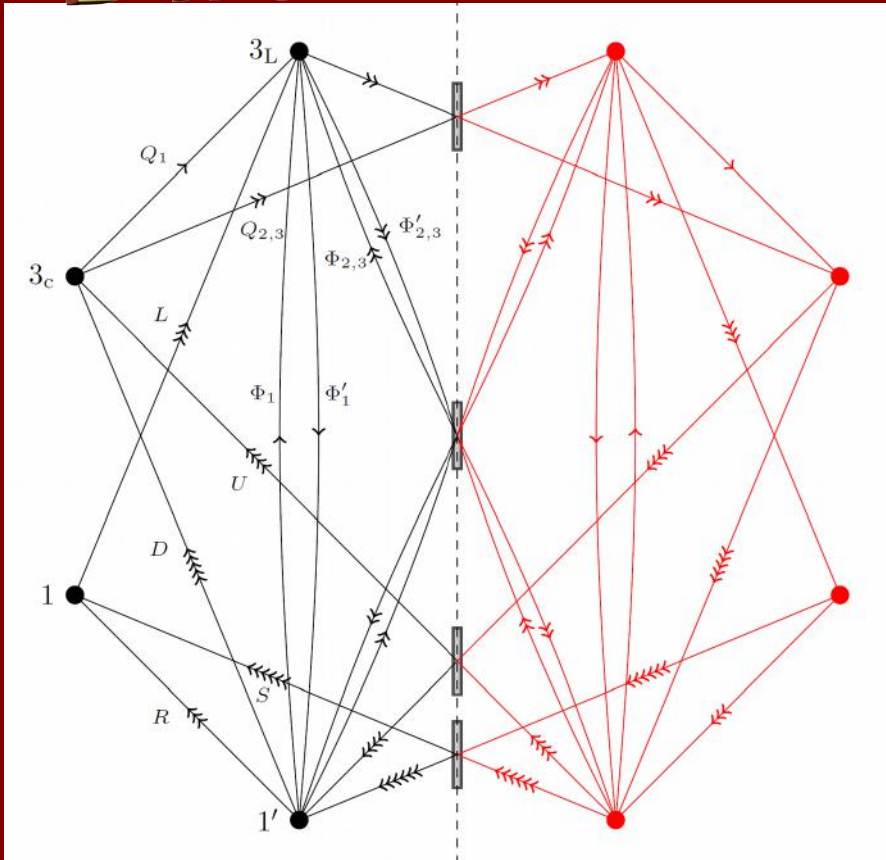
no p decay, **no** RPV ...

neutron-antineutron oscillations added from exotic instantons

Dirac seesaw

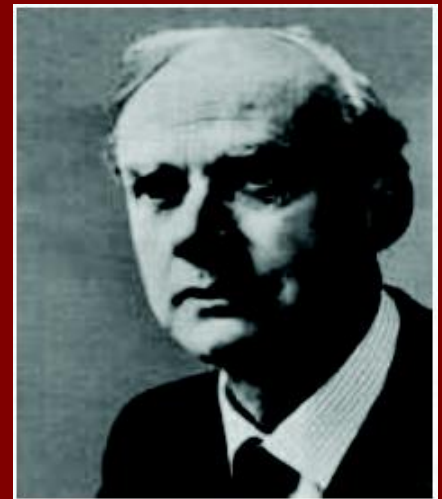
Addazi et al arXiv:1604.02117

From strings

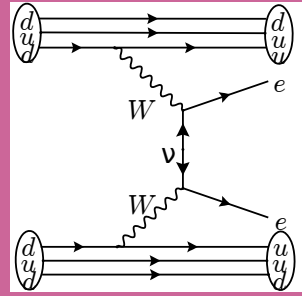
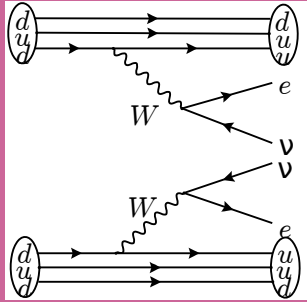


Physics Letters B 755 (2016) 363–366

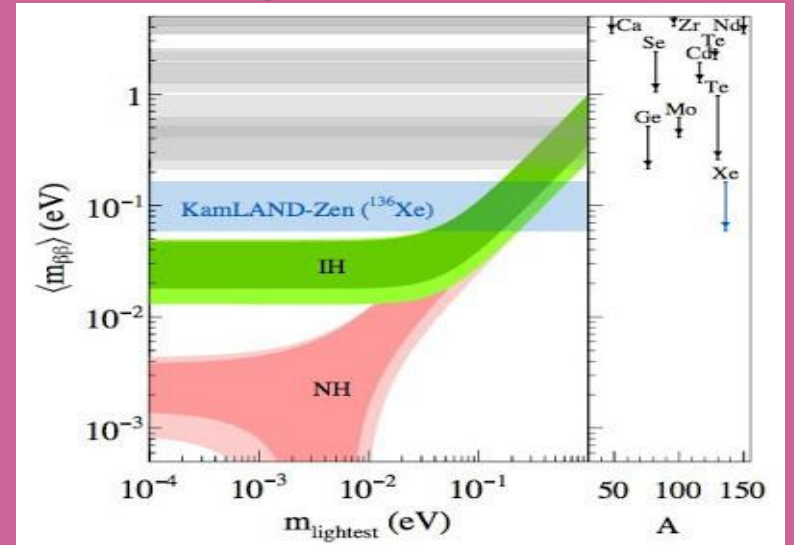
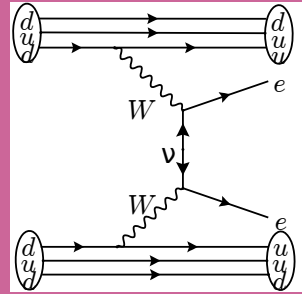
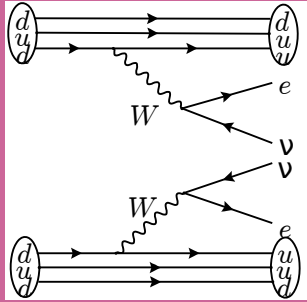
arXiv:1605.08362



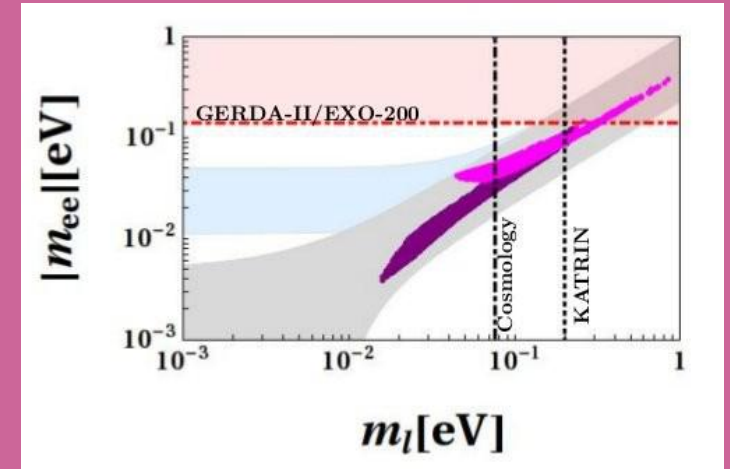
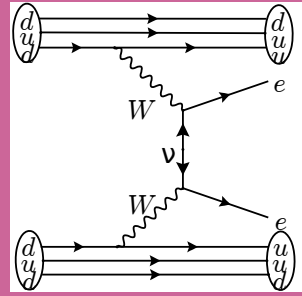
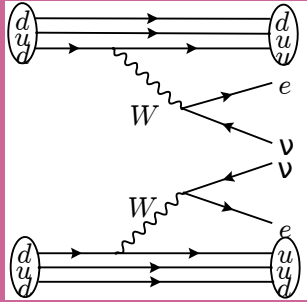
Neutrinoless double beta decay



Neutrinoless double beta decay



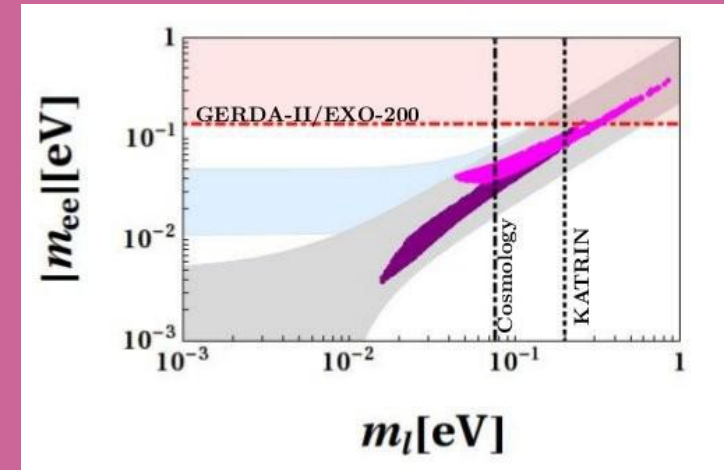
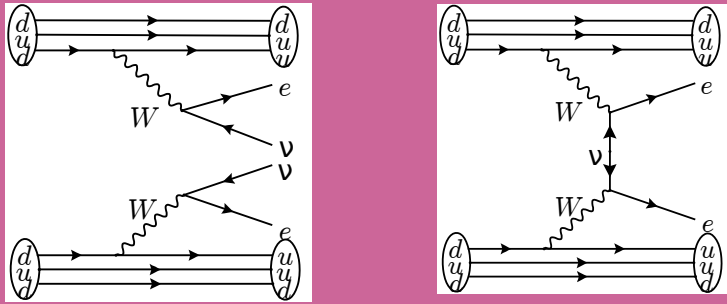
Neutrinoless double beta decay



Flavor Sensitivity

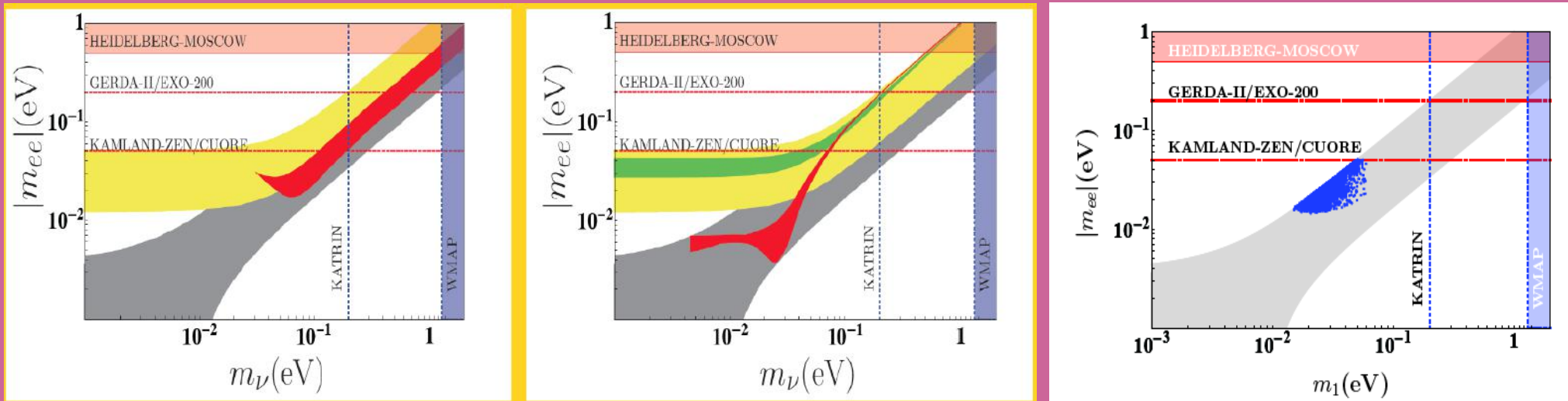
Bonilla et al arXiv:1411.4883

Neutrinoless double beta decay



Bonilla et al arXiv:1411.4883

Flavor Sensitivity



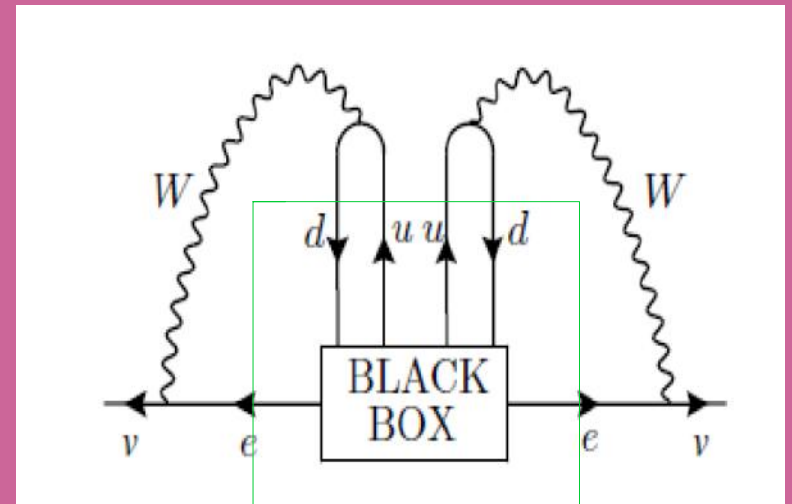
Dorame et al
NPB861 (2012) 259-270

Valle
PhysRevD.86.056001

King et al Phys. Lett. B 724 (2013) 68



The Black Box



Schechter, JWFV 82

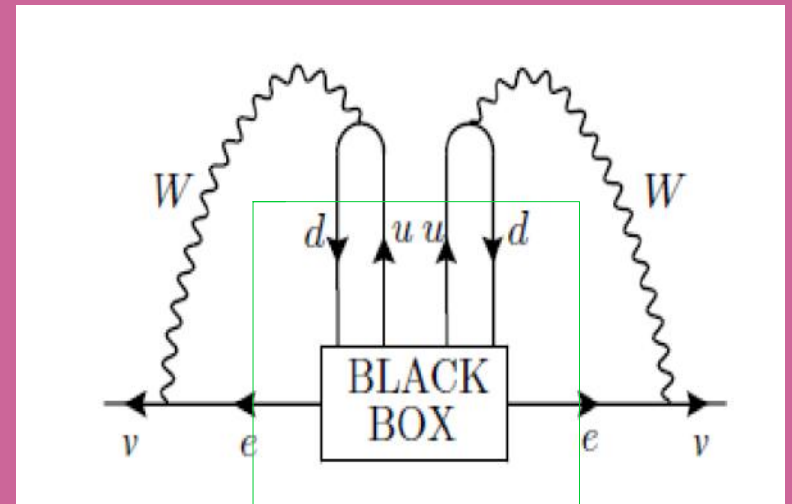
Lindner et al JHEP 1106 (2011) 091



The Black Box

*Even if mediated by
short-range mechanism ...*

Heavy mediators at the LHC



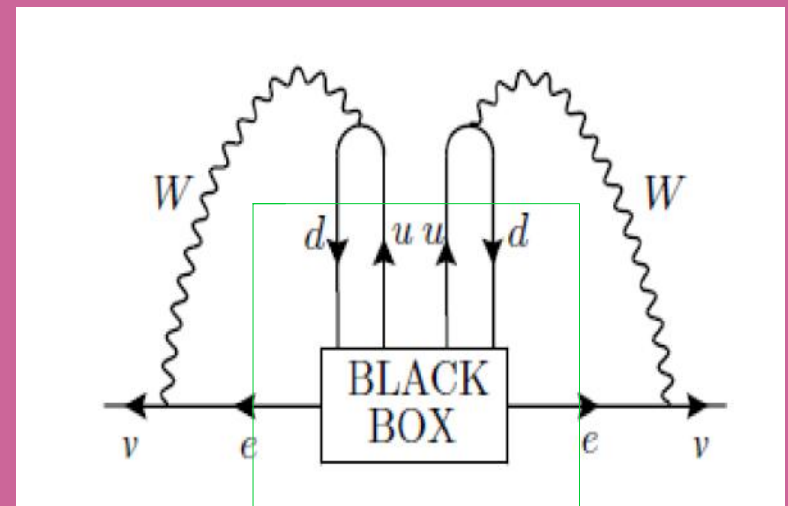
Schechter, JWFV 82

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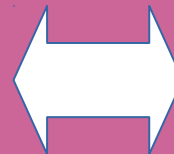
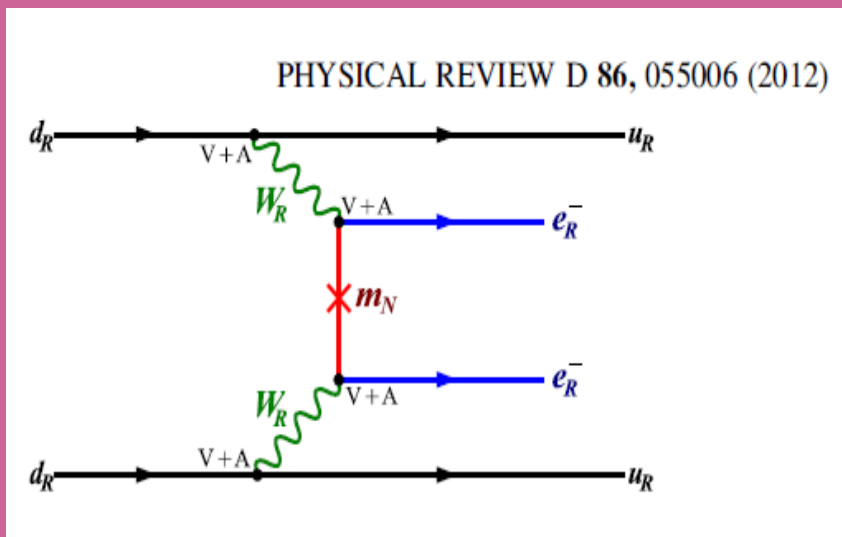
The Black Box

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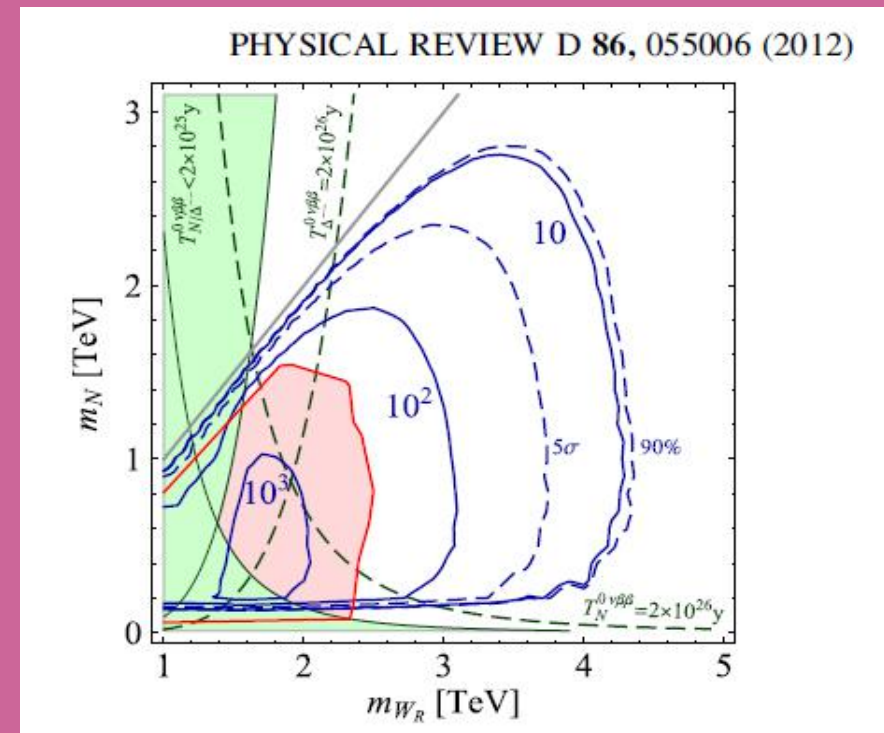


Schechter, JWFV 82
Lindner et al JHEP 1106 (2011) 091

Heavy mediators at the LHC



Valle

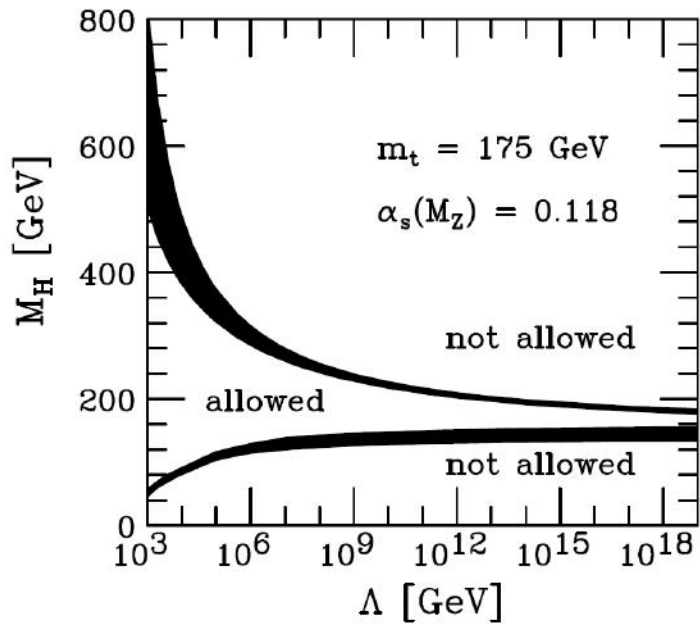




SM vacuum

Valle

SM vacuum

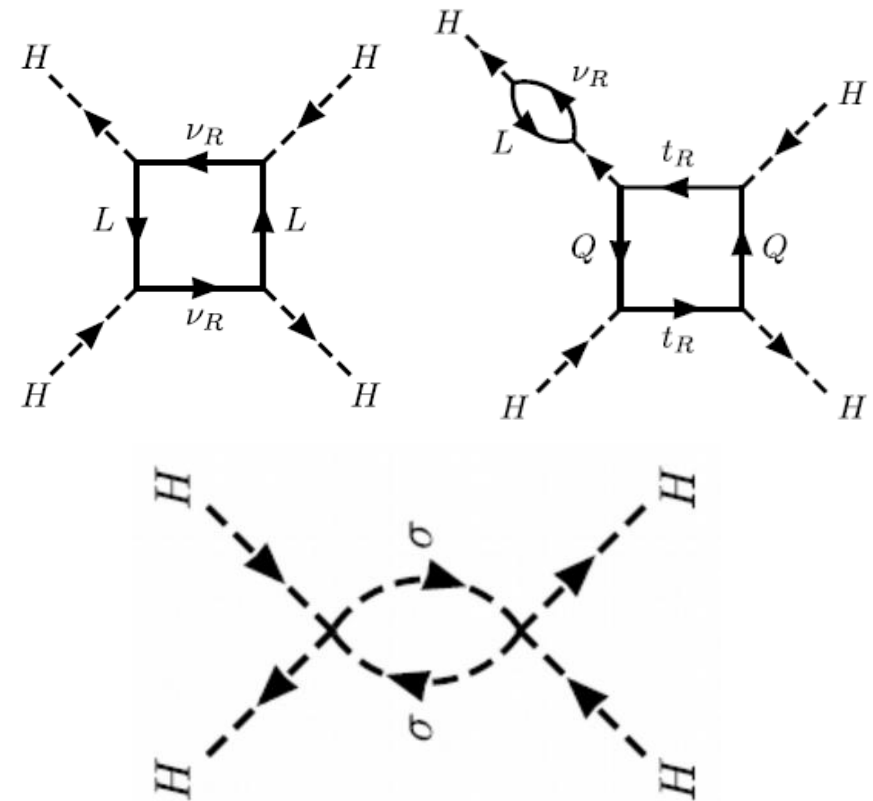
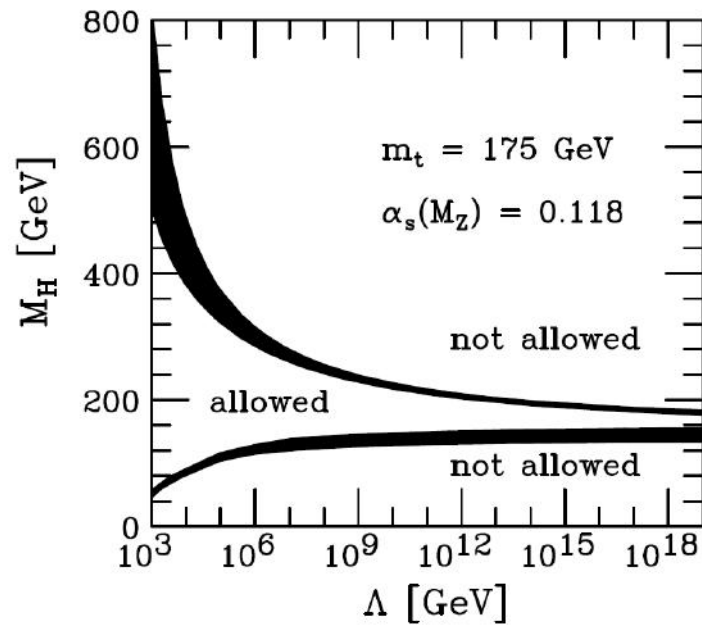


Valle



SM vacuum and neutrinos

Physics Letters B 756 (2016) 345–349



Valle

Neutrino mass and invisible Higgs decays at the LHCCesar Bonilla,^{1,*} Jorge C. Romão,^{2,†} and José W. F. Valle^{1,‡} $H_i \rightarrow JJ$ and $H_2 \rightarrow 2H_1 \rightarrow 4J$ $\left(\text{when } m_{H_1} < \frac{m_{H_2}}{2}\right).$

arXiv:1502.01649

channel	ATLAS	CMS
$\mu_{\gamma\gamma}$	1.17 ± 0.27	$1.14^{+0.26}_{-0.23}$
μ_{WW}	$1.00^{+0.32}_{-0.29}$	0.83 ± 0.21
μ_{ZZ}	$1.44^{+0.40}_{-0.35}$	1.00 ± 0.29
$\mu_{\tau^+\tau^-}$	$1.4^{+0.5}_{-0.4}$	0.91 ± 0.27
$\mu_{b\bar{b}}$	$0.2^{+0.7}_{-0.6}$	0.93 ± 0.49

Neutrino mass and invisible Higgs decays at the LHC

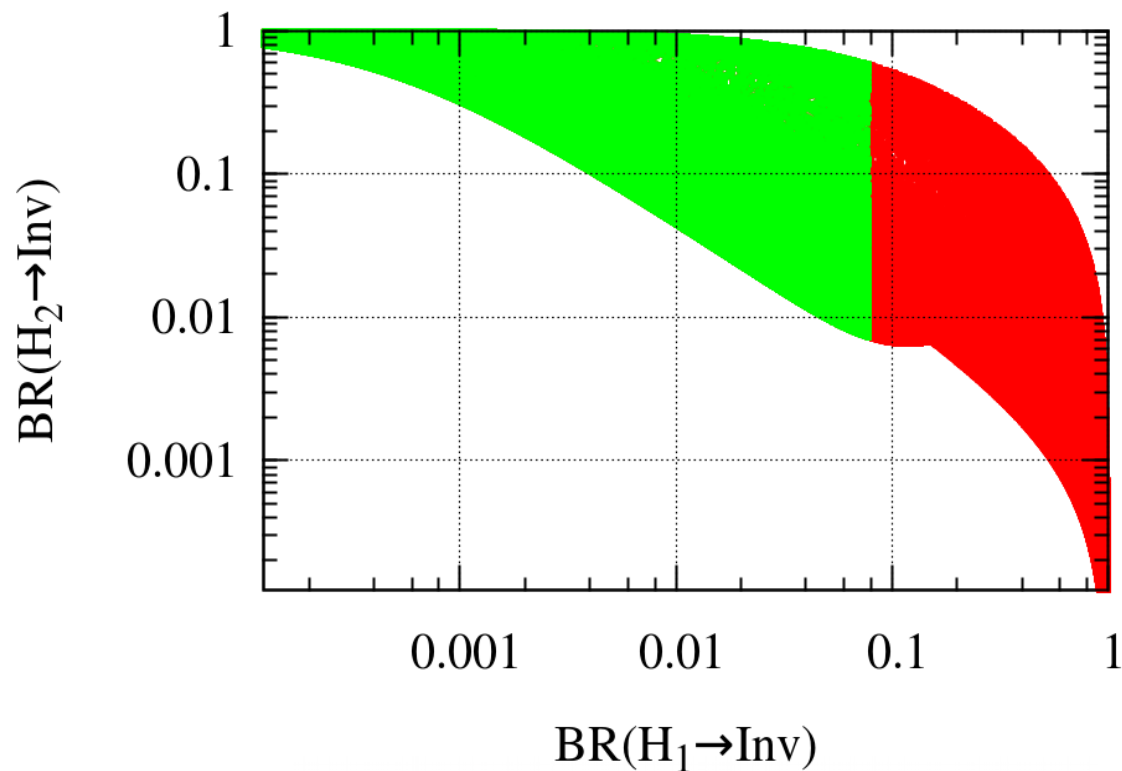
Cesar Bonilla,^{1,*} Jorge C. Romão,^{2,†} and José W.F. Valle^{1,‡}

$v_\sigma=3$ TeV

$$H_i \rightarrow JJ \quad \text{and} \quad H_2 \rightarrow 2H_1 \rightarrow 4J$$

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Valle

**Neutrino as
higgs benchmark**

Why neutrinos



The road to new physics



Why neutrinos



The road to new physics



diphoton



Why neutrinos



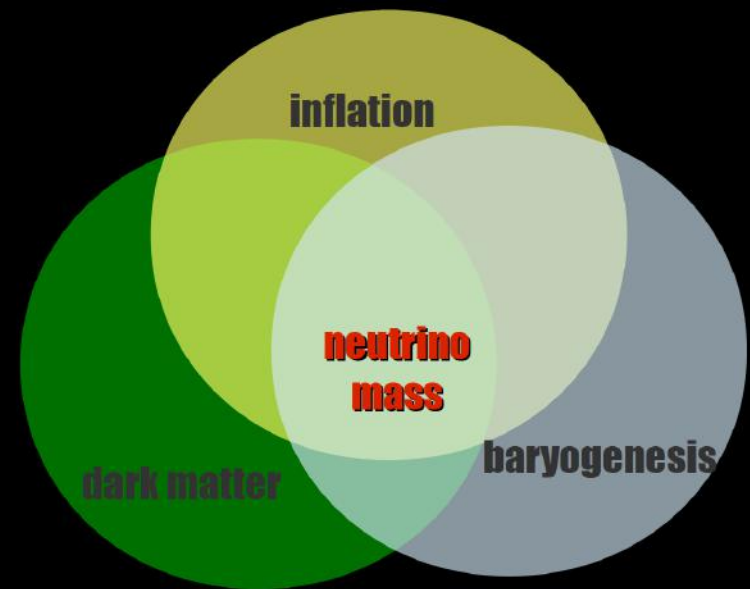
Basic cosmo & astro probe

May explain cosmo puzzles

The road to new physics



diphoton



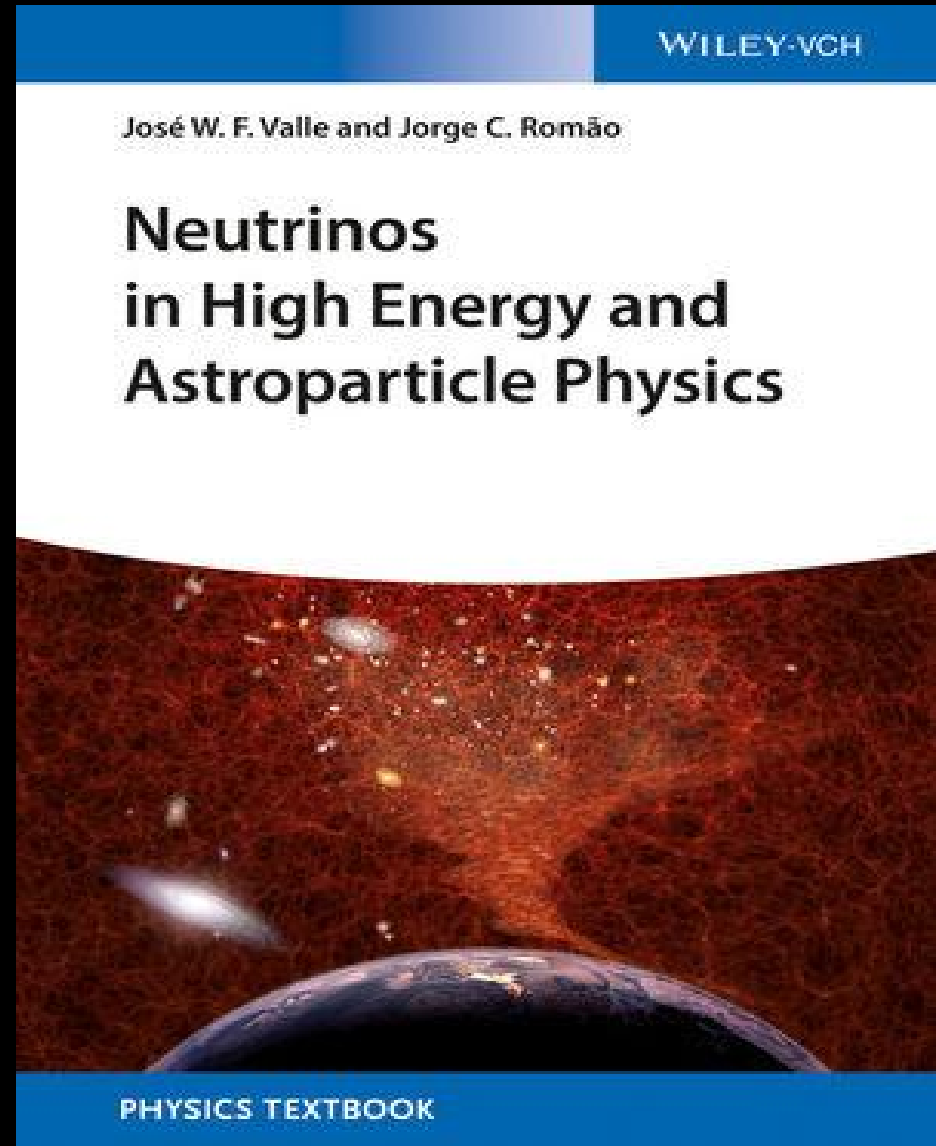
Thank you

<http://t.co/yUSLiF1cGX>

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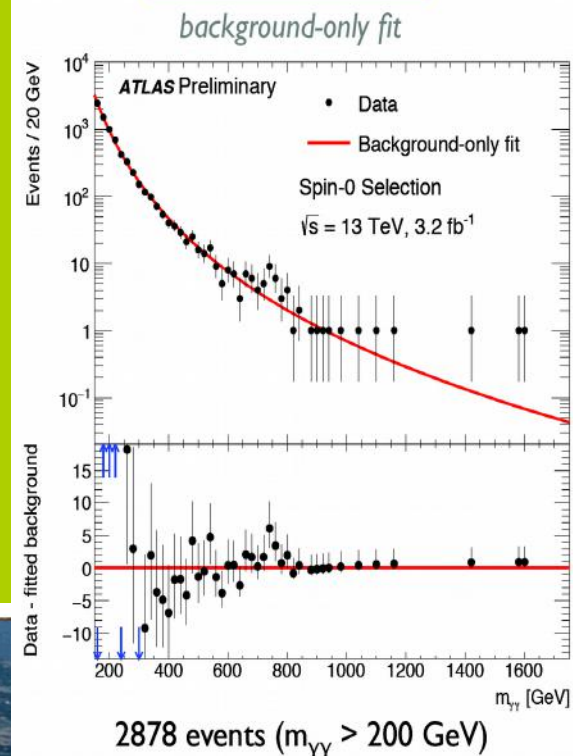
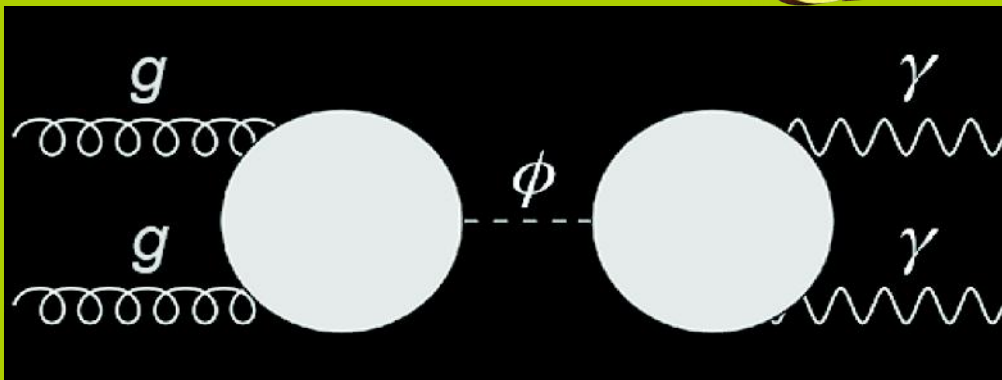
*All you ever
Wanted to know
About neutrinos
As guide to
New physics*

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458 pages
2015



Backup slides

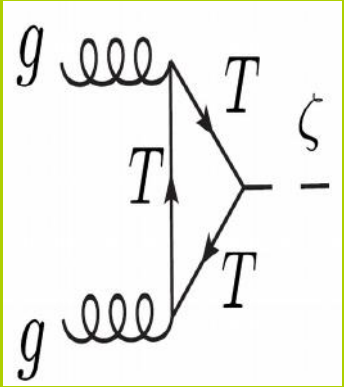
Diphoton anomaly



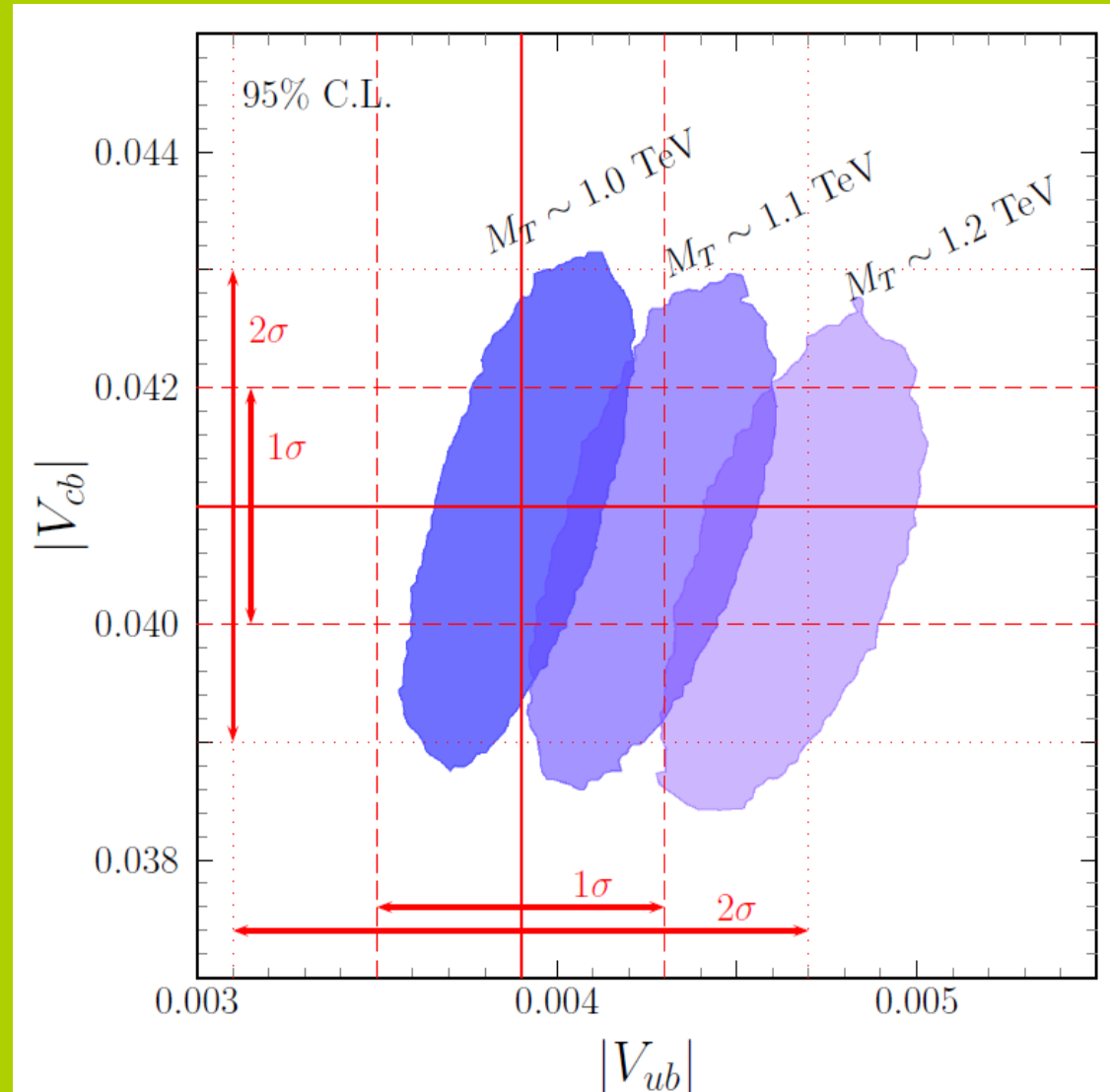
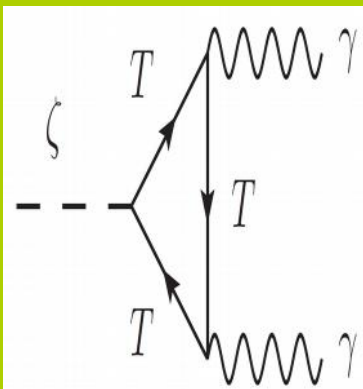
Large Hadron Collider



Diphoton anomaly



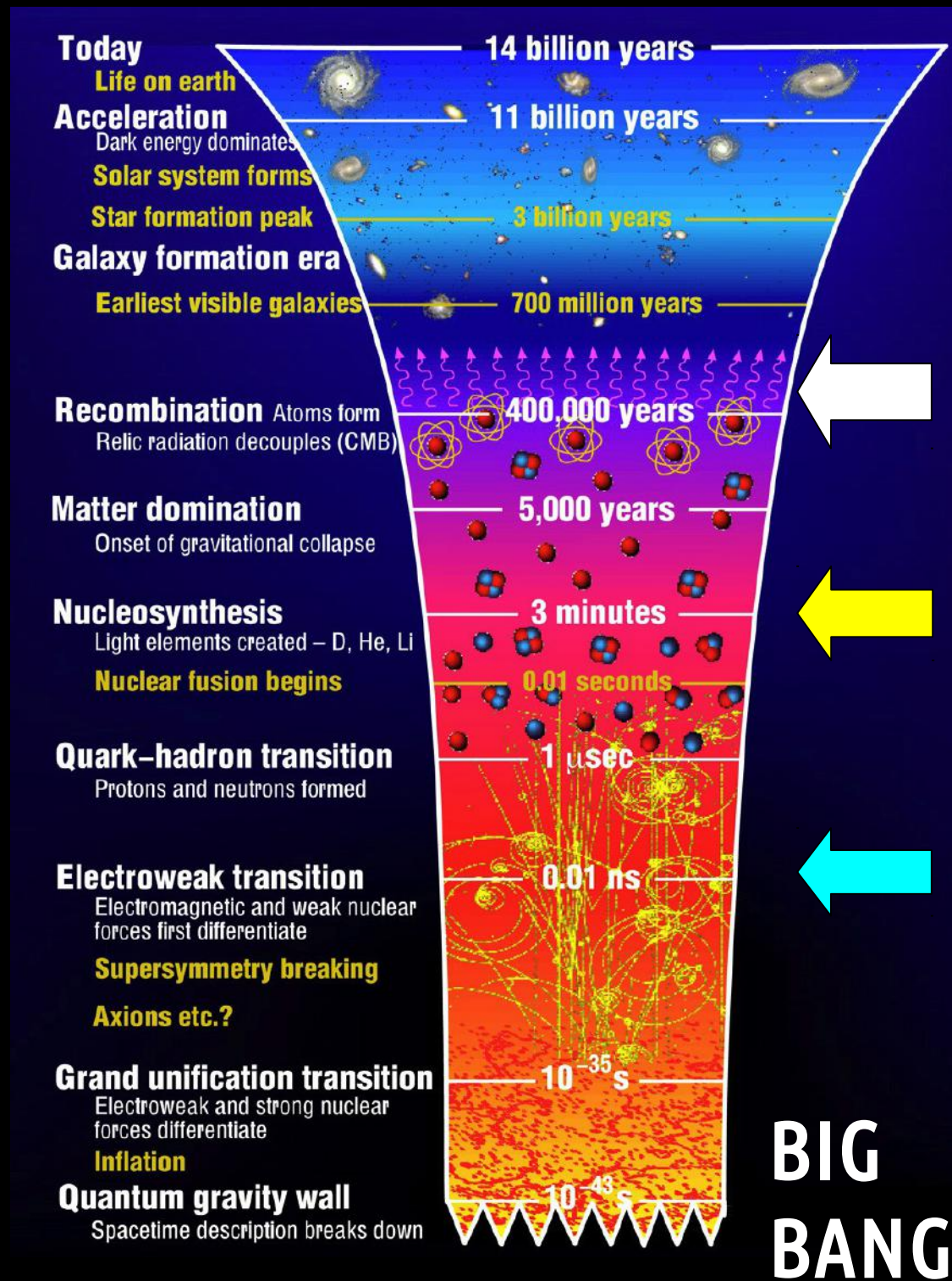
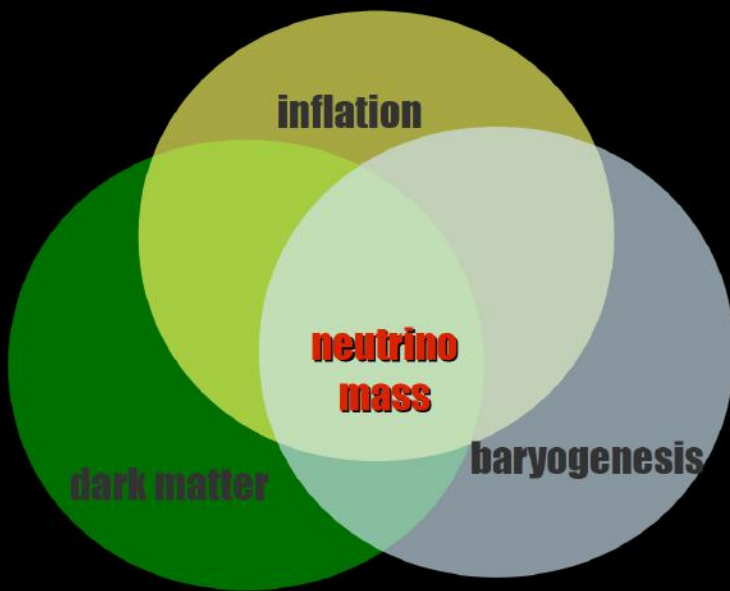
As a flavon



Neutrinos affect the CMB and large scale structure in the Universe ...

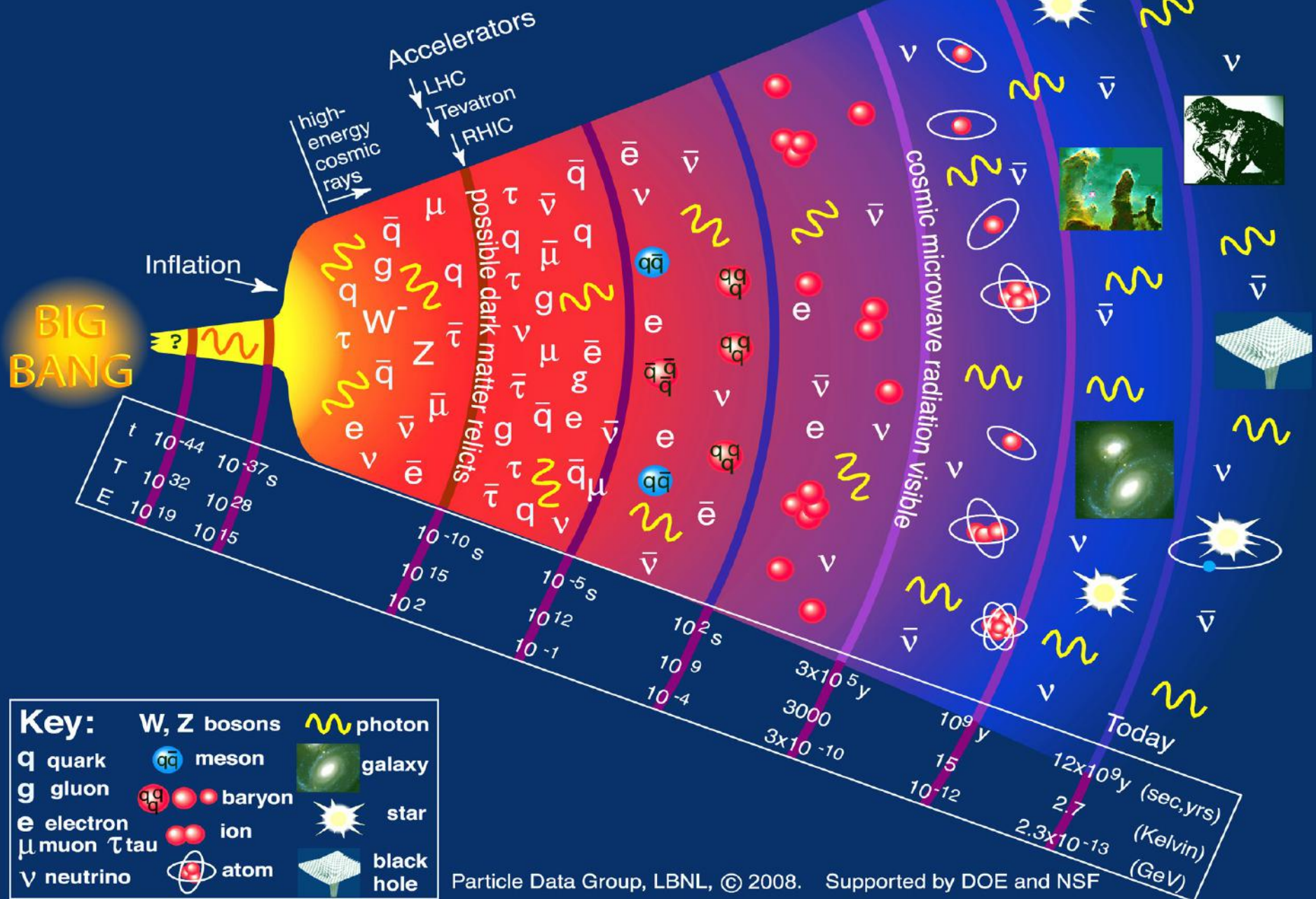
are key in the synthesis of light elements

can “probe” the Universe earlier than photons ...



understanding the largest from the SM q, l ...

History of the Universe



Seesaw inflation & majoron dark matter

$$\sigma = \frac{1}{\sqrt{2}} (\langle \sigma \rangle + \rho + iJ)$$

NEUTRINO MASSES

DARK MATTER

INFLATON

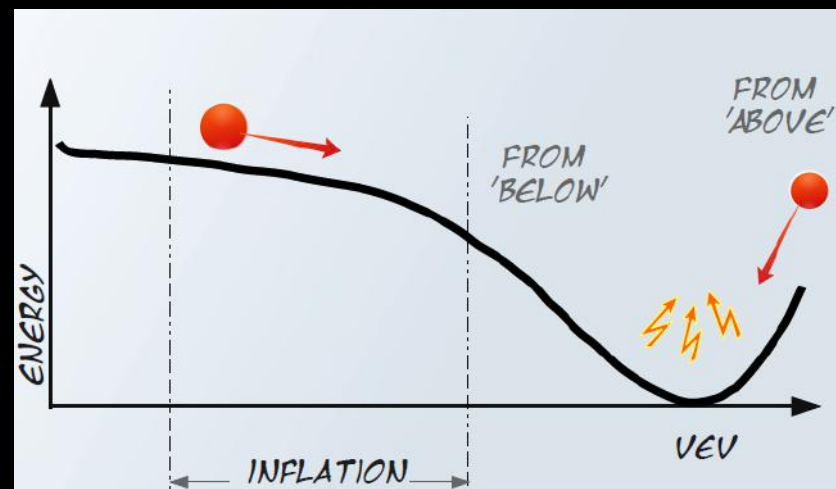
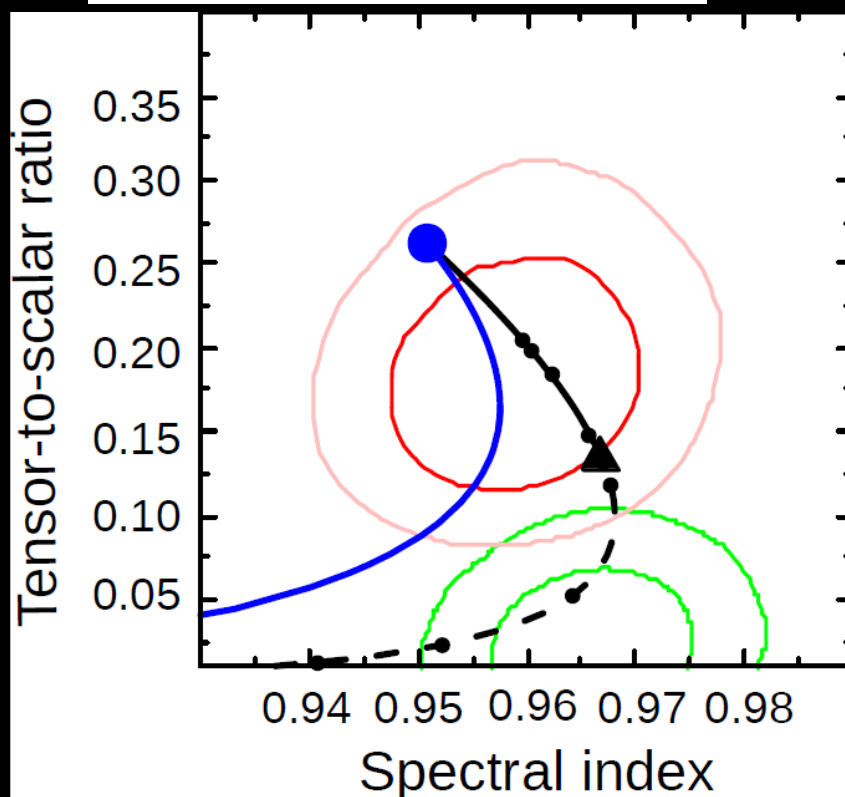
Boucenna et al arXiv:1405.2332

PRD90 (2014) 05502

type-I seesaw **Leptogenesis**

Aristizabal et al arXiv:1405.4706

Quartic versus Higgs Inflation

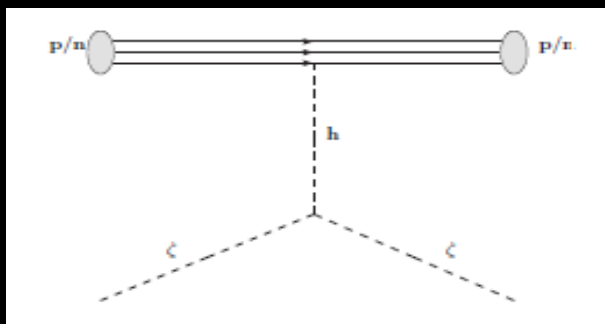
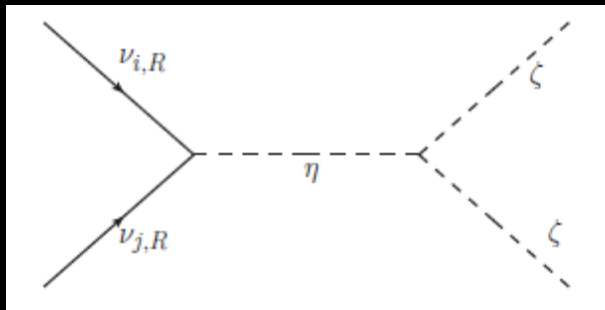
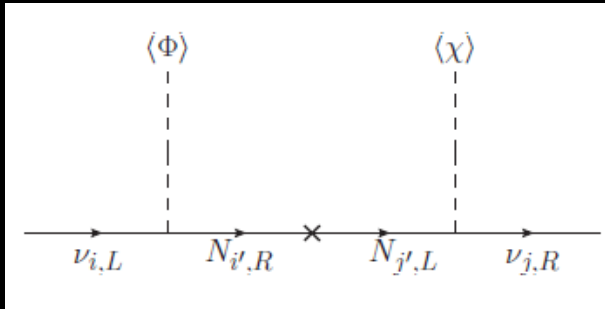


<http://arxiv.org/pdf/1502.00612v1>

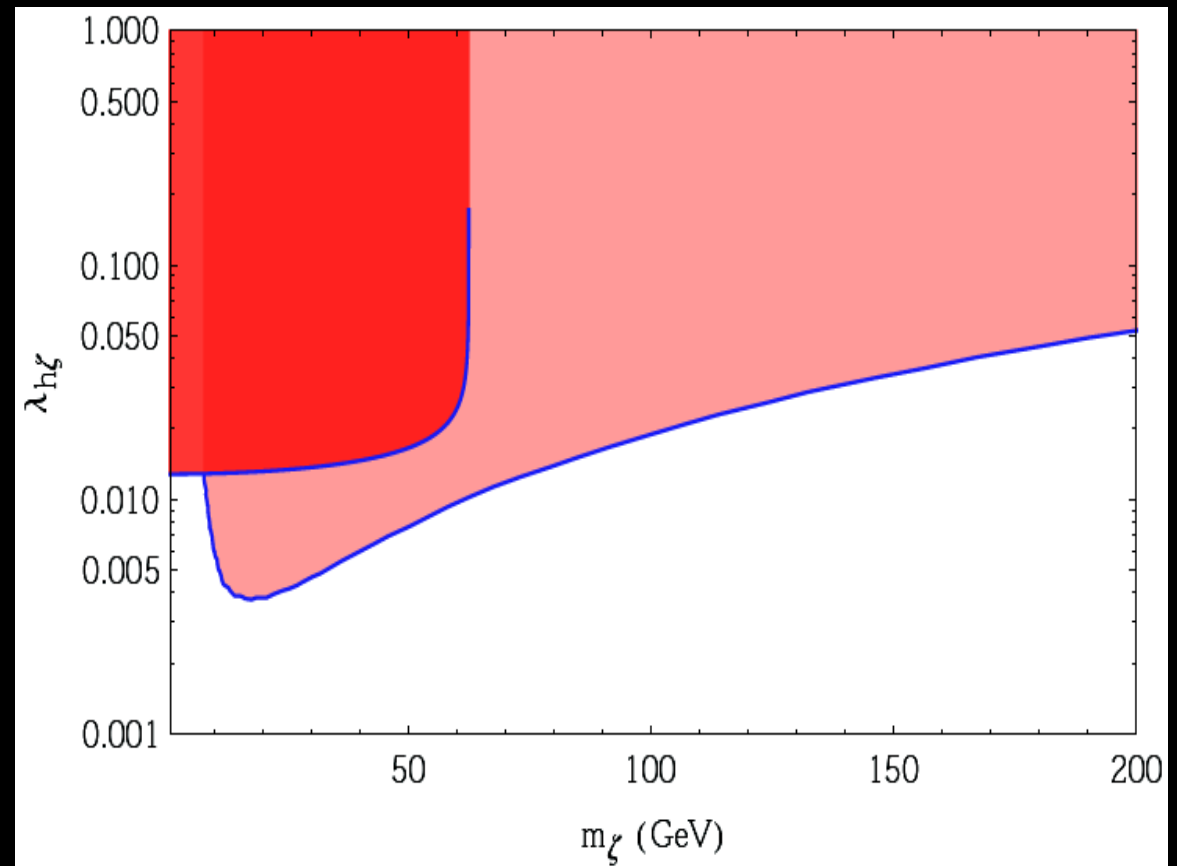
Dark Matter Stability from Dirac nature of neutrinos

arXiv:1606.04543

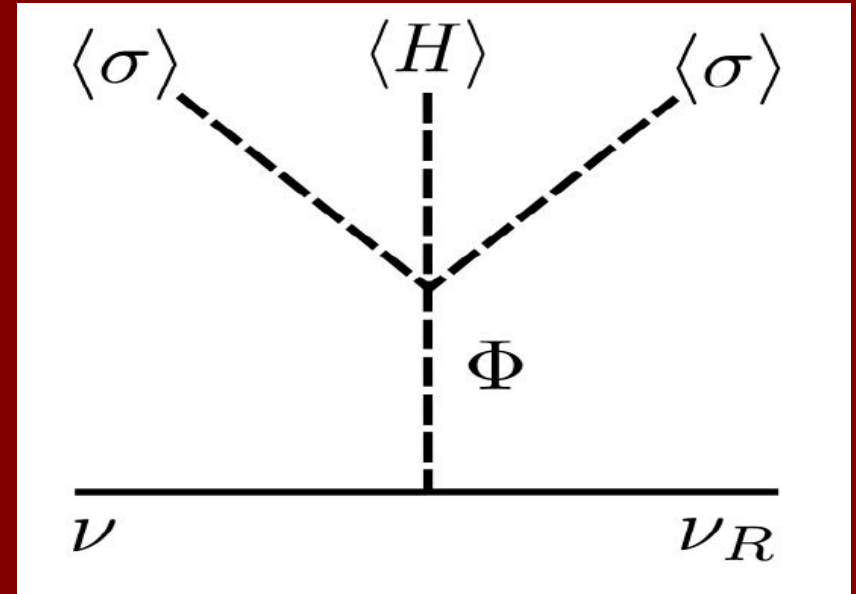
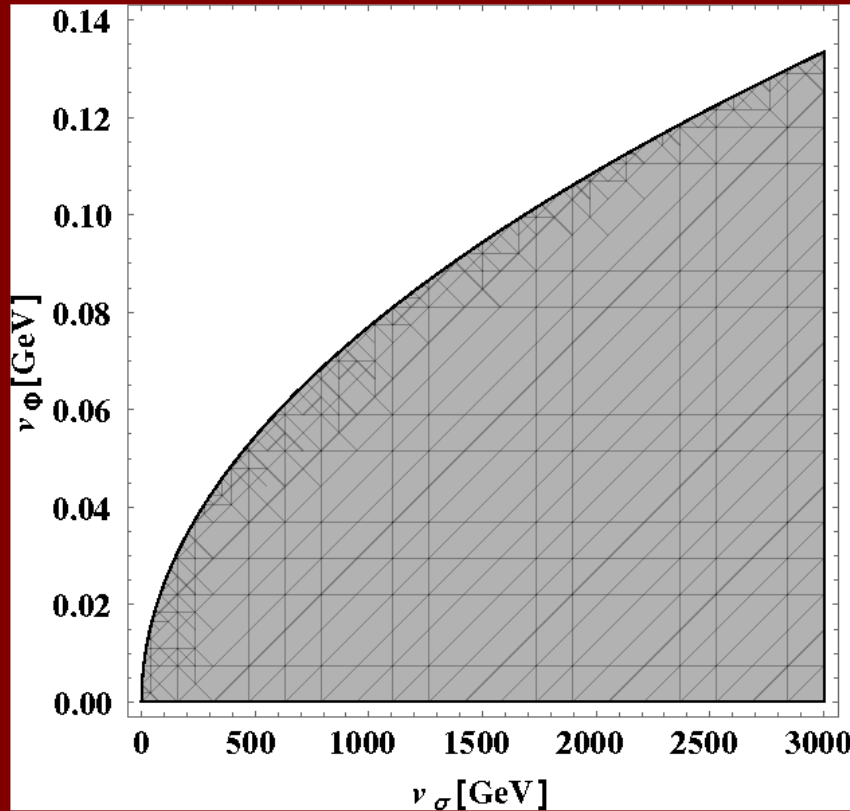
arXiv:1606.06904



Lepton Quarticity versus Lepton number

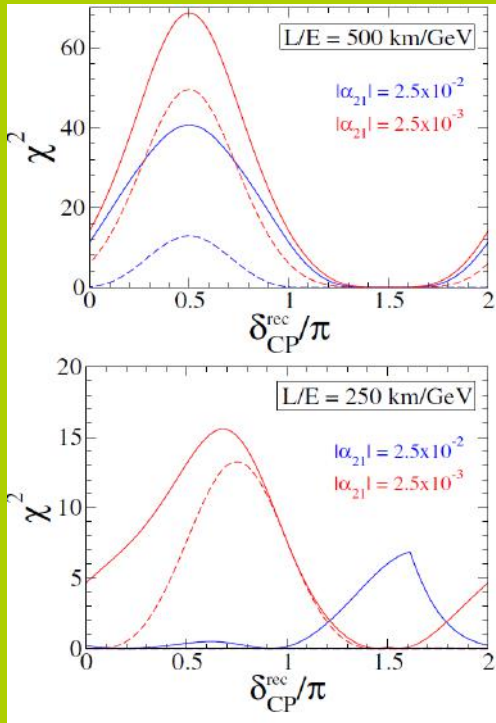


Diracon type-II seesaw



arXiv:1605.08362

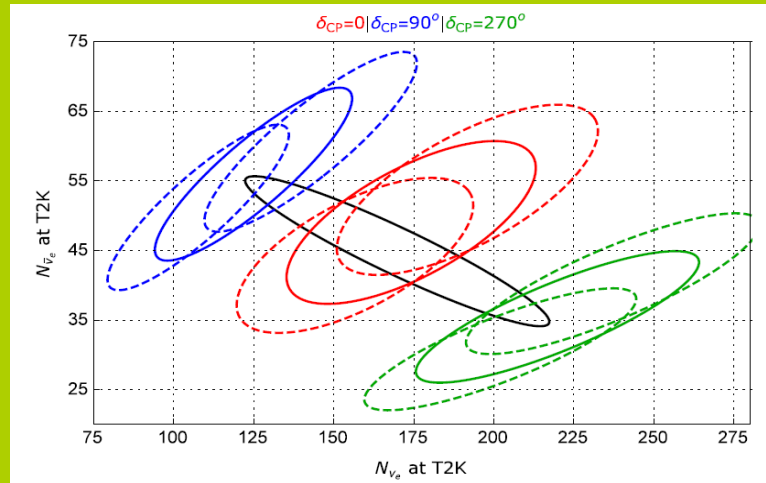
Non unitary lepton mixing & CP



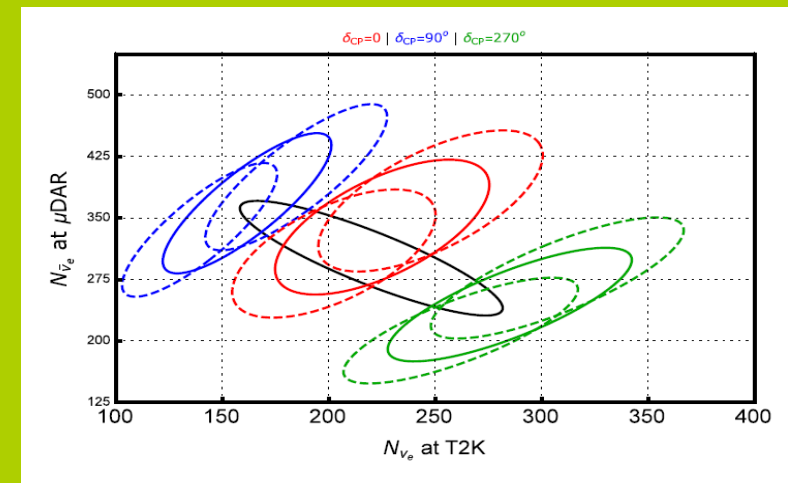
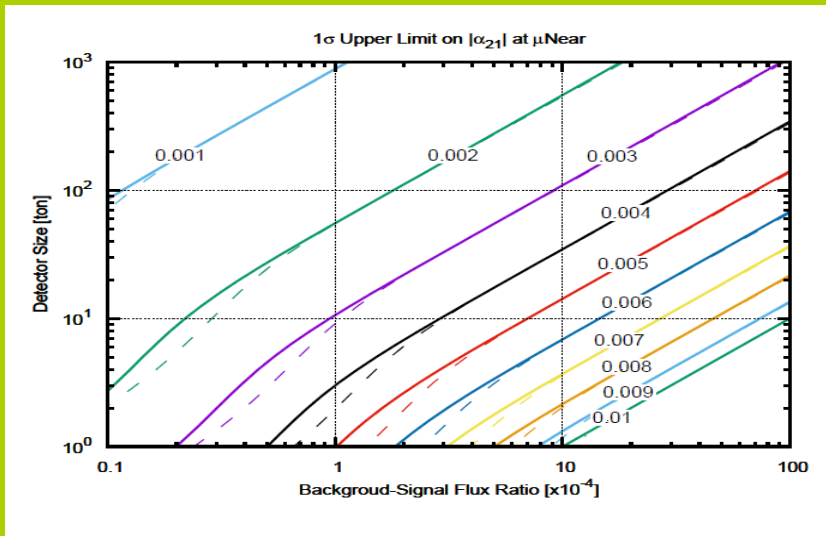
Schechter & JV PRD22 (1980) 2227 & PDG
 Rodejohann, JV Phys.Rev. D84 (2011) 073011
<http://dx.doi.org/10.1103/PhysRevD.92.053009>

arXiv:1604.05690

ArXiv:1605.01670

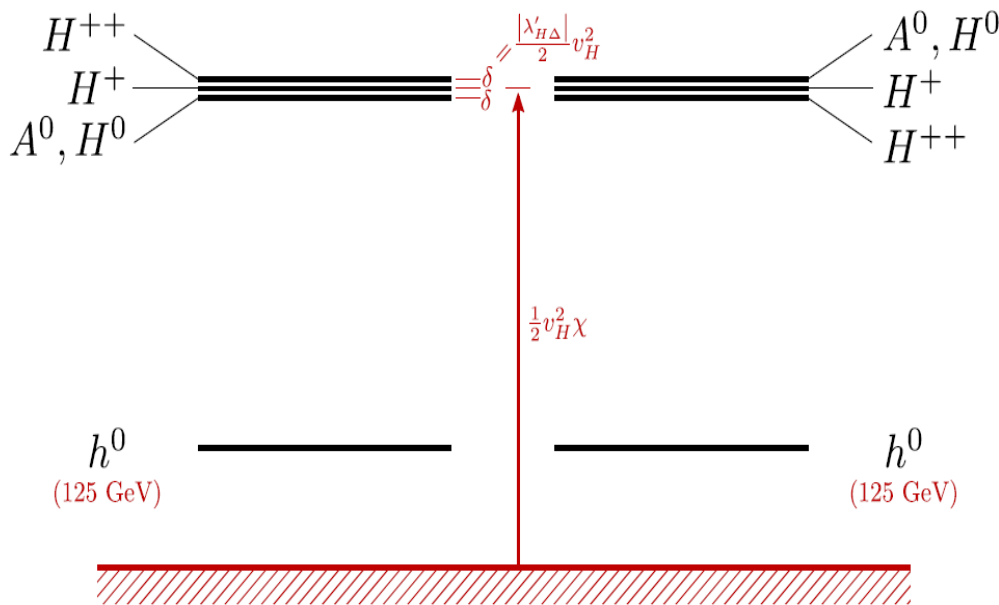
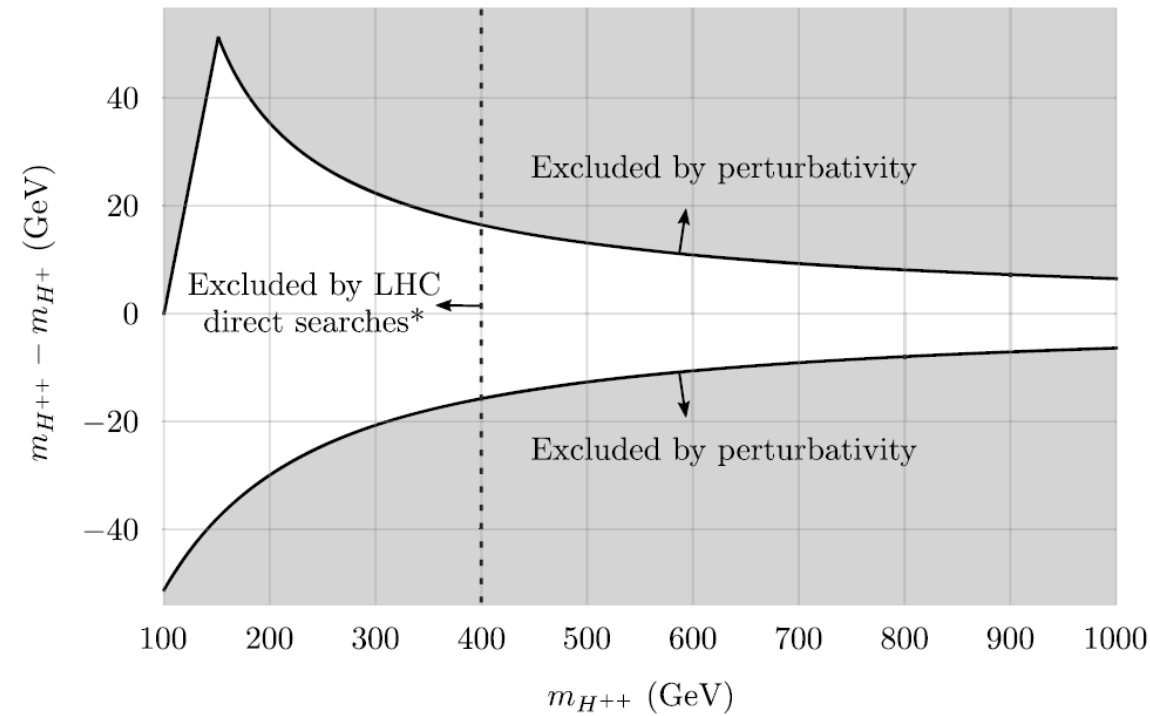


atm



Valle

Consistency of the triplet seesaw model revisited

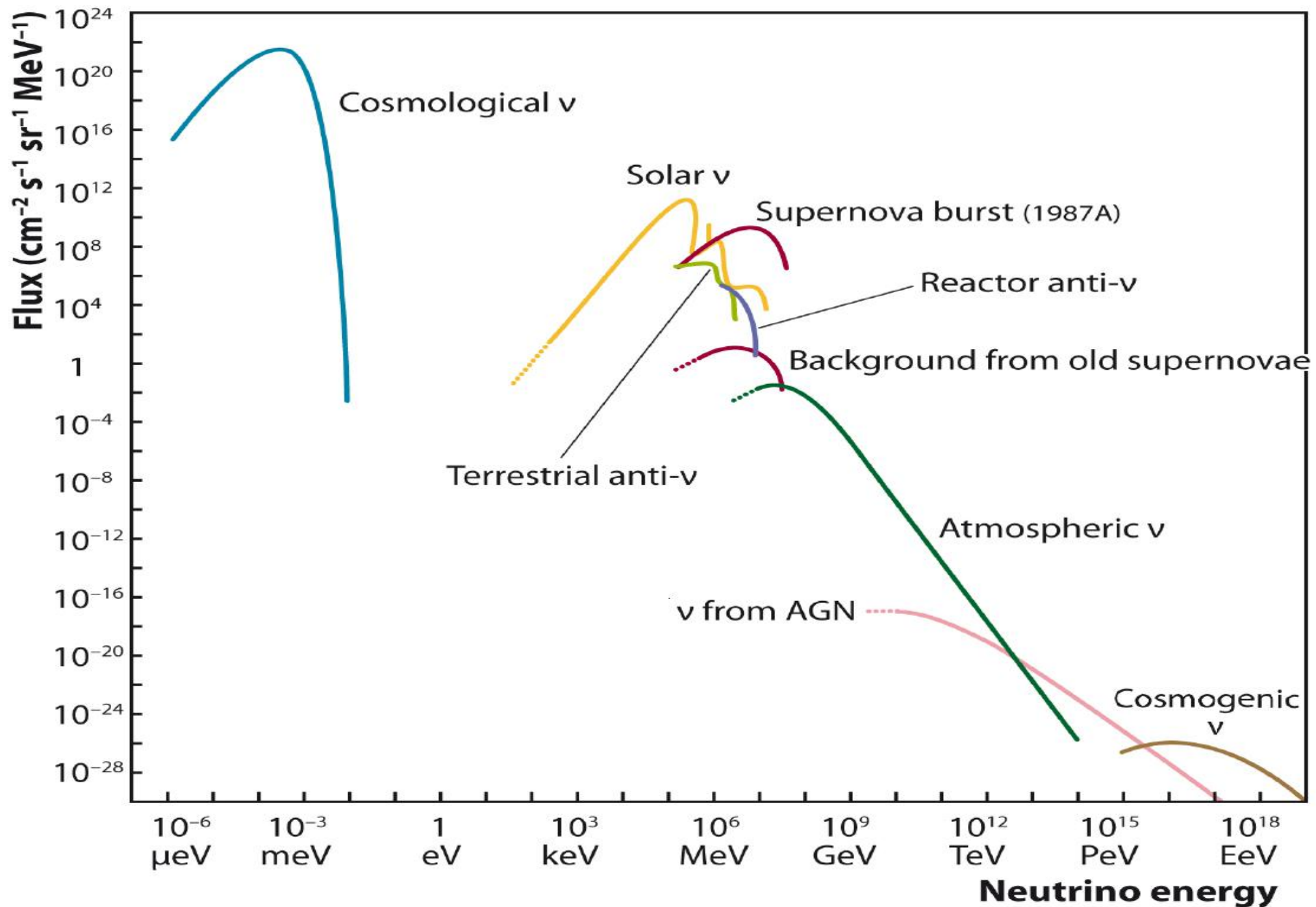


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**Neutrino as
higgs benchmark**

neutrino sources



neutrino cross sections

