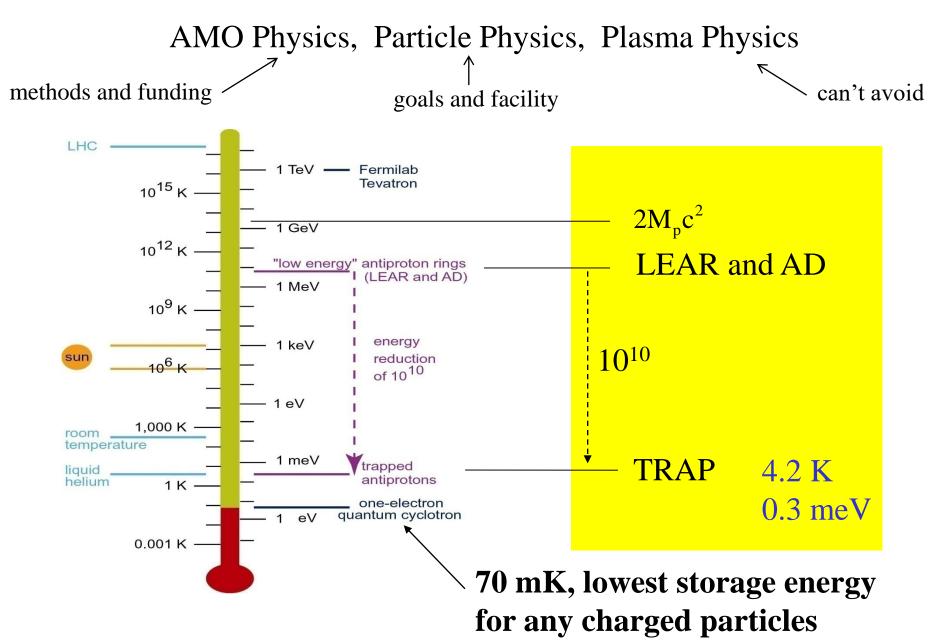
# Low Energy Tests of the Standard Model and its Fundamental Symmetries

### Gerald Gabrielse Leverett Professor of Physics, Harvard University

60 years of since parity violation

Inspired by the experiment of Wu, and the proposal of Lee and Yang,
 → small-scale, low-energy experiments
 to investigate the particles, interactions, and symmetries of the universe,
 to test and help develop our most fundamental theoretical descriptions.

### **Low Energy Particle Physics**



# New in 2017 Center for Fundamental Physics at Low Energy

(cfp.physics.northwestern.edu)

Specializing in small-scale, low-energy experiments

- to investigate the particles, interactions, and symmetries of the universe
- to test and help develop our most fundamental theoretical descriptions.

Exciting opportunities available for

- New faculty members -- need promise to do Wu-like experiments
- CFP postdocs need aspire to do Wu-like experiments
- Graduate students need to desire a Wu-like adventure

Founding director: G.G. First faculty search starts: Fall 2016 (contact G.G.)

# **Stringent Low Energy Tests of the Standard Model and Its Symmetries**

Illustrate with experiments that have 3 kinds of objectives

1. Testing the Standard Model's most precise prediction by making the most precise measurement of a property of an elementary particle

electron magnetic dipole moment

2. Testing very different predictions of the Standard Model and Supersymmetry (and other) models

electron electric dipole moment

also neutron, proton, Hg

Testing the most fundamental symmetry of the Standard Model
 q/m for antiproton and proton mag. moments of e<sup>+</sup> and e<sup>-</sup>

# → The Great Triumph and the Great Frustration of Modern Physics

Start with frustration

Embarrassing, Unsolved Mystery: How did our Matter Universe Survive Cooling After the Big Bang?



Gabrielse

Big bang → equal amounts of matter and antimatter created during hot time

As universe cools  $\rightarrow$  antimatter and matter annihilate

**Big Questions:** 

- How did any matter survive?
- How is it that we exist?

Our experiments are looking for evidence of any way that antiparticles and particles may differ



# **Our "Explanations" are Not so Satisfactory**



### **Baryon-Antibaryon Asymmetry in Universe is Not Understood**

### **Standard "Explanation"**

- CP violation
- Violation of baryon number
- Thermodynamic non-equilibrium Sakarov

### Alternate

- CPT violation
- Violation of baryon number
- Thermo. equilib. Bertolami, Colladay, Kostelecky, Potting Phys. Lett. B 395, 178 (1997)

Why did a universe made of matter survive the big bang? Makes sense look for answers to such fundamental questions in the few places that we can hope to do so very precisely.



Bigger problem: don't understand dark energy within 120 orders of magnitude



# Why Compare H and H (or P and P)?

### **Reality is Invariant – symmetry transformations**

- P parity
- **CP** charge conjugation, parity
- **CPT** charge conjugation, parity, and time reversal

### **CPT Symmetry**

### → Particles and antiparticles have

- same mass
- opposite charge
- → Atom and anti-atom have

 $\rightarrow$  same structure

**Looking for Surprises** 

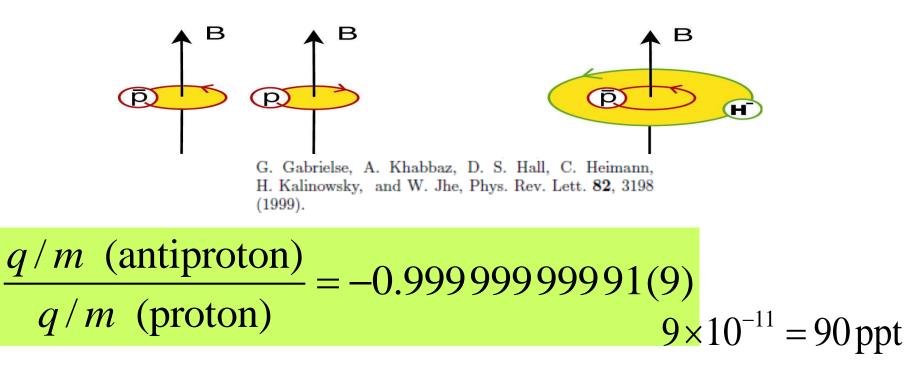
- simple systems
- extremely high accuracy
- comparisons will be convincing

- same magnetic moment
- same mean life

- reasonable effort
- FUN

# **High Precision Tests of CPT Invariance**

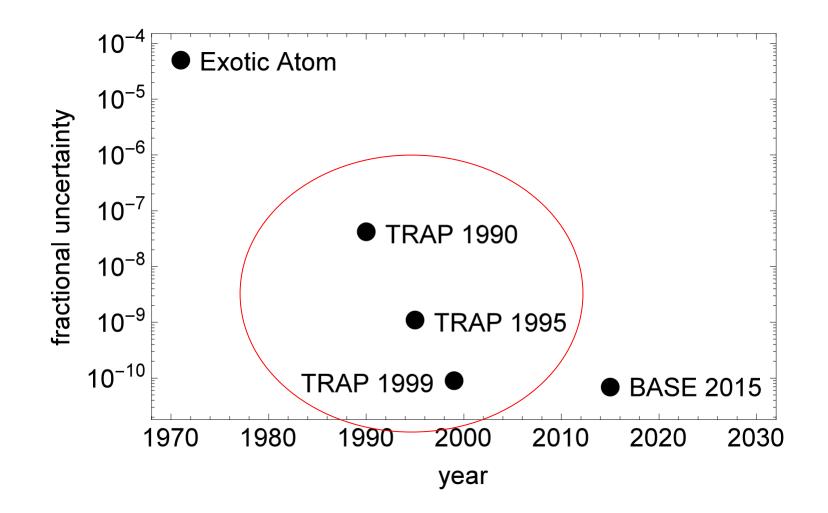
The Most Precise CPT Test with Baryons  $\rightarrow$  by TRAP at CERN



(most precise result of CERN's antiproton program before the AD)

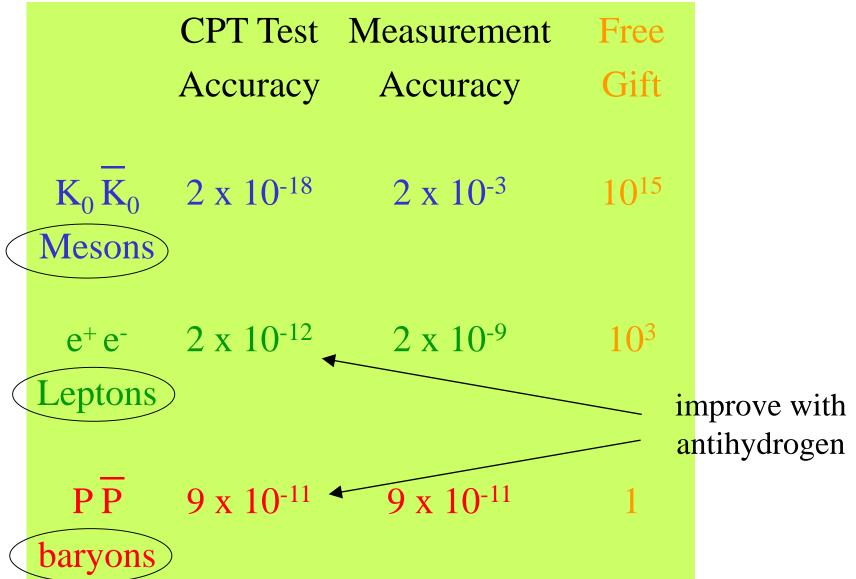
Goal at the AD: Make CPT tests that approach or exceed this precision

# **Uncertainty in Comparison of Q/M for the Antiproton and Proton**



# **Comparing the CPT Tests**

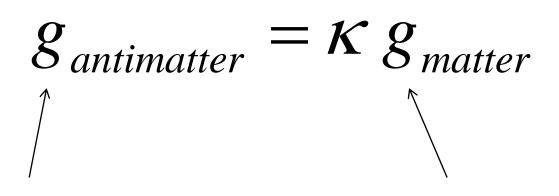
### Warning – without CPT violation models it is hard to compare



3 fundamentally different types of particles

# **Direct Comparison of Antimatter and Matter Gravity**

Does antimatter and matter accelerate at the same rate in a gravitational field?



acceleration due to gravity for antimatter acceleration due to gravity for matter

# The Most Precise Experimental Answer is "Yes" → to at lease a precision of 1 part per million

Gravitational red shift for a clock:  $\Delta \omega / \omega = g h / c^2$ 

→ Antimatter and matter clocks run at different rates if g is different for antimatter and matter

for tensor gravity (would be 1 for scalar gravity) Hughes and Holzscheiter, Phys. Rev. Lett. 66, 854 (1991).

grav. pot. rnergy difference between empty flat space time and inside of hypercluster of galaxies

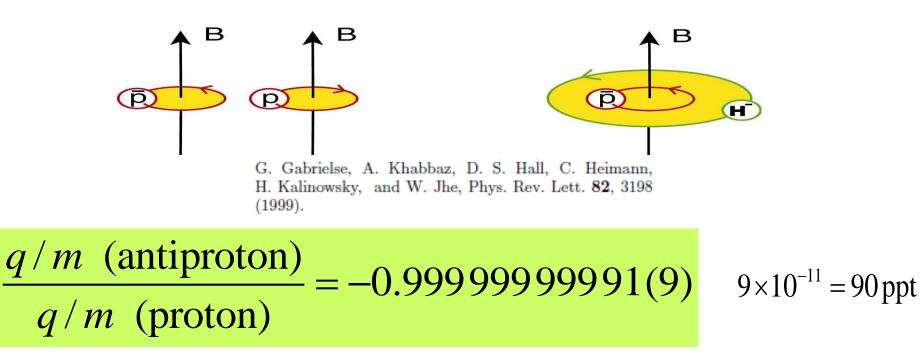
Experiment: TRAP Collaboration, Phys. Rev. Lett. 82, 3198 (1999).

$$\frac{\Delta \omega_c}{\omega_{cc}} < 10^{-10} \qquad --> \qquad \kappa = 1 \pm (< 10^{-6})$$

Comparable limit to that on neutrinos and antineutrinos 1987A

# **Comparison of an Antimatter and Matter Clock**

The Most Precise CPT Test with Baryons  $\rightarrow$  by TRAP at CERN



(most precise result of CERN's antiproton program)

Goal at the AD: Make CPT test that approaches and exceed this precision

# Hard to Get the Part per Million Precision of the Redshift Limit with Antihydrogen and Hydrogen

 $g_{antimatter} = \kappa g_{matter}$ 

Our TRAP gravitational redshift:

$$\frac{\Delta \omega_c}{\omega_{\partial c}} < 10^{-10}$$

 $--> 0.999999 < \kappa < 1.000001$ 

10<sup>8</sup>

ALPHA trapped antihydrogen released (2013):  $-110 < \kappa < 110$ 

(no mention direct redshift comparison)

# Gravitational Redshift Comparison is Ignored citing an unpublished rational for a Fermilab gravity measurement proposal (not approved)

Direct Observation Limits on Antimatter Gravitation

arXiv 0808.3929

Mark Fischler, Joe Lykken, and Tom Roberts<sup>†</sup>

May 20, 2008

- Perhaps CPT violations in the electromagnetic clocks cancel the CPT violation for gravity 
   — not likely
- If gravity would have a finite range then using the local supercluster of galaxies would not be appropriate <u>adds violations</u>
- Use of gravitational potential energy isn't sound **—** not needed
  - → can use metric perturbation to flat space that must vanish at infinity to ensure that matter and antimatter look the same away from gravitational sources

# How Much Better Could the Gravitational Comparison Be?

If we improve the charge-to-mass ratio measurement by a factor of 100 → gravitation comparison will be 100 times more stringent

BASE is working on this.

We still hope to contribute but do not have enough time and people yet.

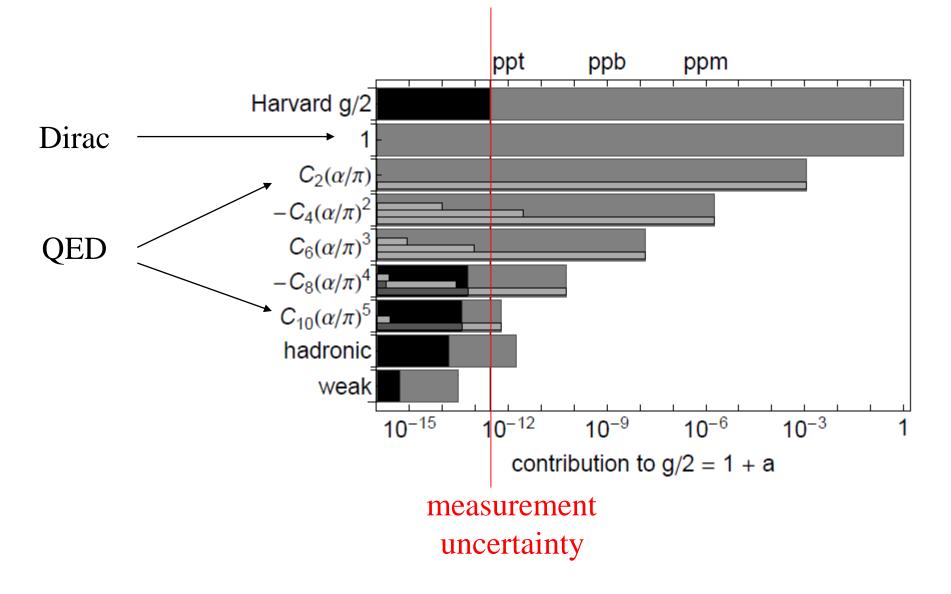
# **Electron Magnetic Dipole Moment**

$$\vec{\mu} = \mu \; \frac{\vec{S}}{\hbar/2}$$

### **Standard Model's Most Precise Prediction**

$$\underbrace{\stackrel{e\hbar}{2m}}_{=} \underbrace{-\frac{\mu}{\mu_{B}} = 1 + C_{2}\left(\frac{\alpha}{\pi}\right) + C_{4}\left(\frac{\alpha}{\pi}\right)^{2} + C_{6}\left(\frac{\alpha}{\pi}\right)^{3} + C_{8}\left(\frac{\alpha}{\pi}\right)^{4} + C_{10}\left(\frac{\alpha}{\pi}\right)^{5} + \dots \\ + a_{hadronic} + a_{weak} + a_{new physics} \\ \hline \alpha = \frac{1}{4\pi\varepsilon_{0}} \frac{e^{2}}{\hbar c} \approx \frac{1}{137} \\ \hline \text{Dirac} \qquad 1 \\ \hline QED \qquad C_{2} = 0.500\ 000\ 000\ 000\ 000\ (exact) \\ C_{4} = -0.328\ 478\ 444\ 002\ 55\ (33) \longleftarrow \text{essentially} \\ C_{6} = 1.181\ 234\ 016\ 815\ (11) \longleftarrow \text{exact} \\ C_{8} = -1.909\ 7\ (20) \\ C_{10} = 9.16\ (0.57). \\ \hline \text{Hadronic} \qquad a_{e}^{\text{hadronic}} = 1.677(16) \times 10^{-12} \\ \hline \text{Weak} \qquad a_{weak} \qquad \text{smaller} \\ \hline \end{array}$$

# **Probing 10th Order and Hadronic Terms**



David Hanneke G.G.

a.S.

and the second s

Shannon Fogwell

......

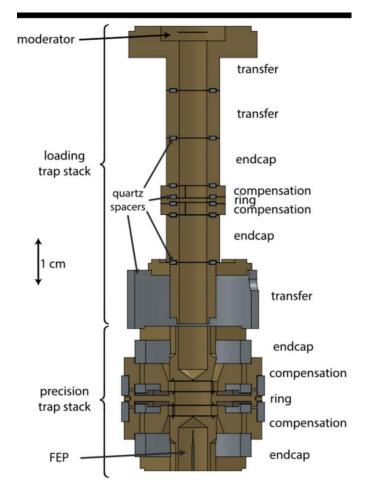
# **Need Good Students and Stable Funding**

20 years8 theses

Elise Novitski Joshua Dorr Shannon Fogwell Hogerheide David Hanneke Brian Odom, Brian D'Urso, Steve Peil, Dafna Enzer, Kamal Abdullah Ching-hua Tseng Joseph Tan

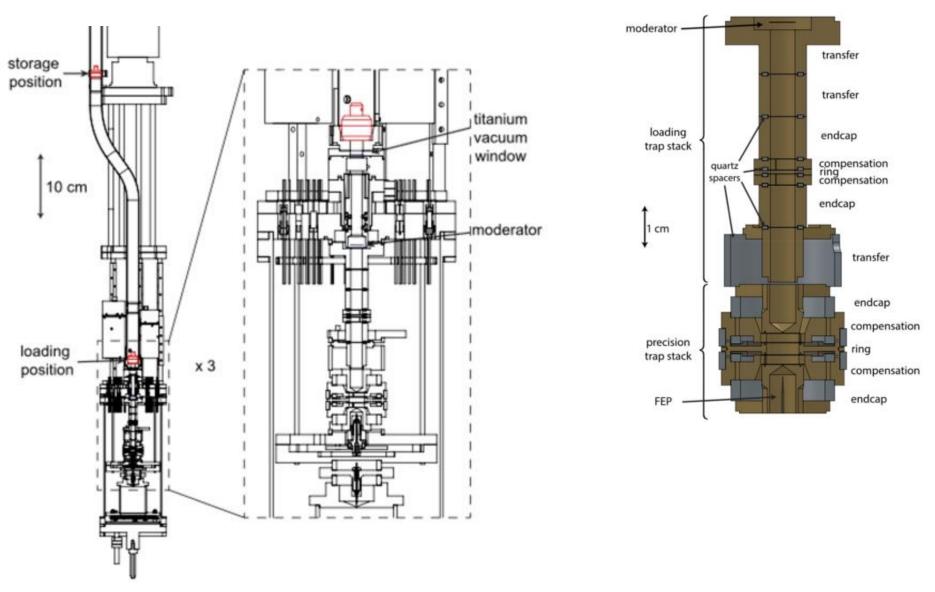
# **Current Team and Trap** $\rightarrow$ e<sup>-</sup> and e<sup>+</sup>



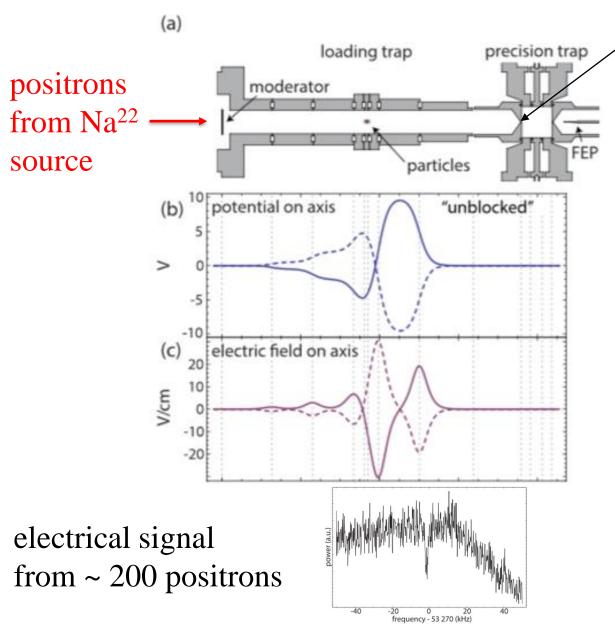


Positron – electron trap
→ to compare magnetic moments of the positron and electron

# **Capturing Positrons from a "Student Source"**



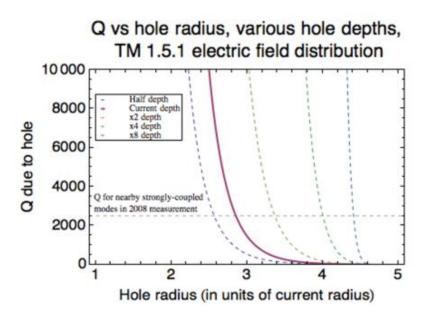
# **Efficient Trapping of Positrons**

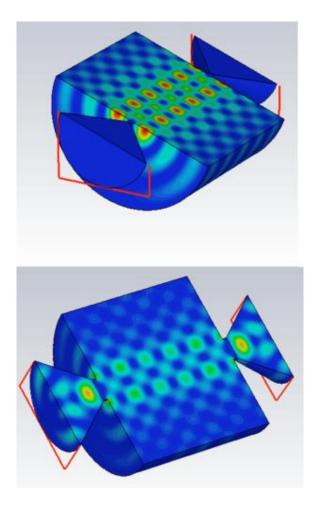


Small hole is giving us trouble

- Need small to keep low cavity loss (high Q)
- Need large enough to let positrons through

### **More on the Hole**



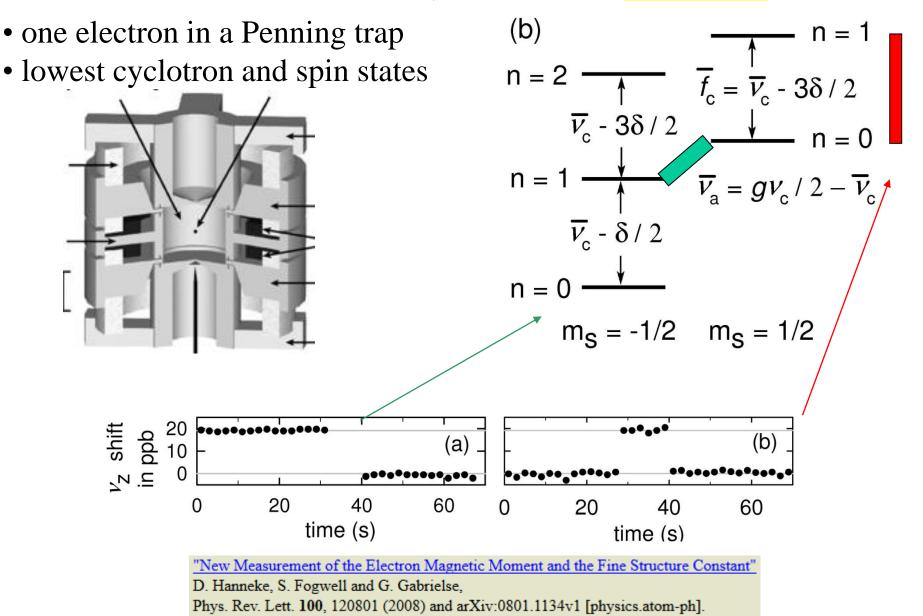


### **Quantum Measurement of the Electron Magnetic Moment**

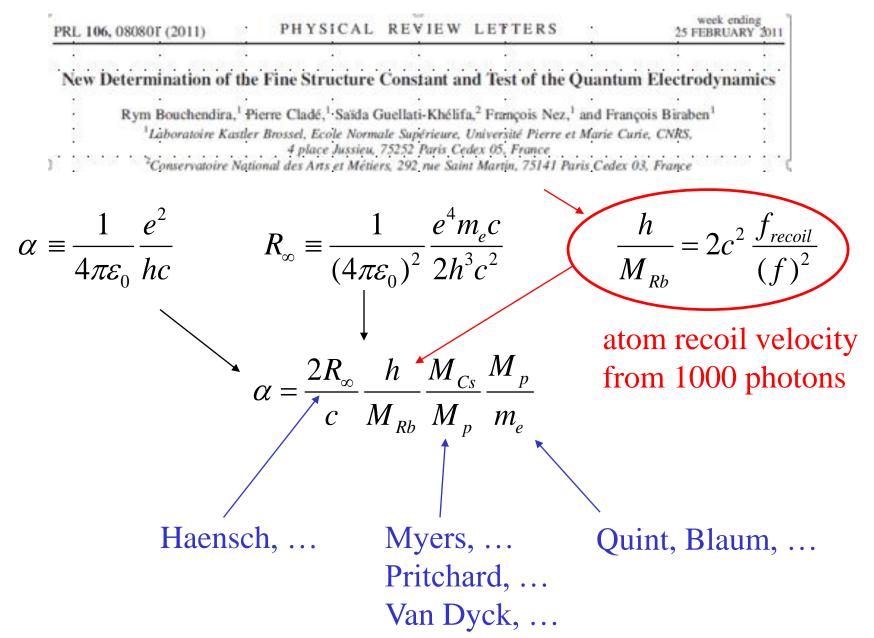
$$E = m_{s}\hbar\omega_{s} + (n+1/2)\hbar\omega_{c} \qquad \vec{\mu} = \mu \frac{S}{\hbar/2}$$
Spin flip energy:  $\hbar\omega_{s} = -\vec{\mu} \cdot \vec{B} = -2\mu B \qquad \qquad \vec{\omega}_{s} = -\frac{\mu}{\mu_{B}}$ 
Cyclotron energy:  $\hbar\omega_{c} = \hbar \frac{eB}{m} = 2\mu_{B}B \qquad \qquad \vec{\omega}_{c} = -\frac{\mu}{\mu_{B}}$ 
(the magnetometer)
Bohr magneton  $\frac{e\hbar}{2m}$ 

Need to resolve the quantum states of the cyclotron motion  $\rightarrow$  Relativistic shift is 1 part in 10<sup>9</sup> per quantum level

# **Most Precisely Measured Property of an Elementary Particle (**2.8×10<sup>-13</sup>**)**



### **SM Prediction Needs an Independent** α

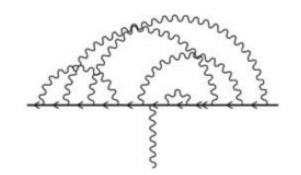


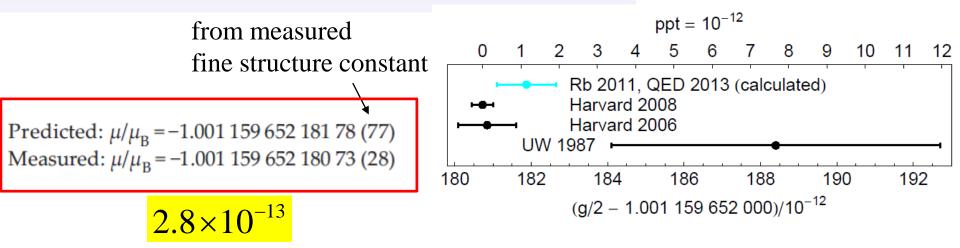
# The standard model's greatest triumph

Gerald Gabrielse

The standard model predicts the electron magnetic moment to an astonishing accuracy of one part in a trillion.

**Gerald Gabrielse** is the George Vasmer Leverett Professor of Physics at Harvard University in Cambridge, Massachusetts.





www.physicstoday.org

### **Test for Physics Beyond the Standard Model**

$$-\frac{\mu}{\mu_B} = \frac{g}{2} = 1 + a_{QED}(\alpha) + \delta a_{SM:Hadronic+Weak} + \delta a_{New Physics}$$

### **Does the electron have internal structure?**

S. J. Brodsky and S. D. Drell. Anomalous Magnetic Moment and Limits on Fermion Substructure. *Phys. Rev. D*, 22:2236 – 2243, 1980.

 $m^*$  = total mass of particles bound together to form electron

$$R < 5 \times 10^{-19} m \qquad m^* > \frac{m}{\sqrt{\delta a}} = 360 \ GeV/c^2 \qquad \begin{array}{limited by the uncertainty in independent $\alpha$ value} \\ R < 2 \times 10^{-19} m \qquad m^* > \frac{m}{\sqrt{\delta a}} = 1 \ TeV/c^2 \qquad \text{if our uncertainty} \\ \text{was the only limit} \end{array}$$

Not bad for an experiment done at 100 mK, but LEP does better $R < 2 \times 10^{-20} m$  $m^* > 10.3 TeV/c^2$ LEP contact interaction limit> 20,000,000 electron masses of binding energy

# **Electron-Positron Summary**

Already the most precise test of the standard model

Soon should be the most precise test of the standard model's most fundamental CPT symmetry (compare electron and positron)

Not so easy to improve on a magnetic moment already determined to 3 parts in 10<sup>13</sup>, but progress continues toward a big improvement

One-electron Q-bit work has just restarted

# **Despite the Great Success of the Standard Model The Standard Model Cannot be the Whole Story**

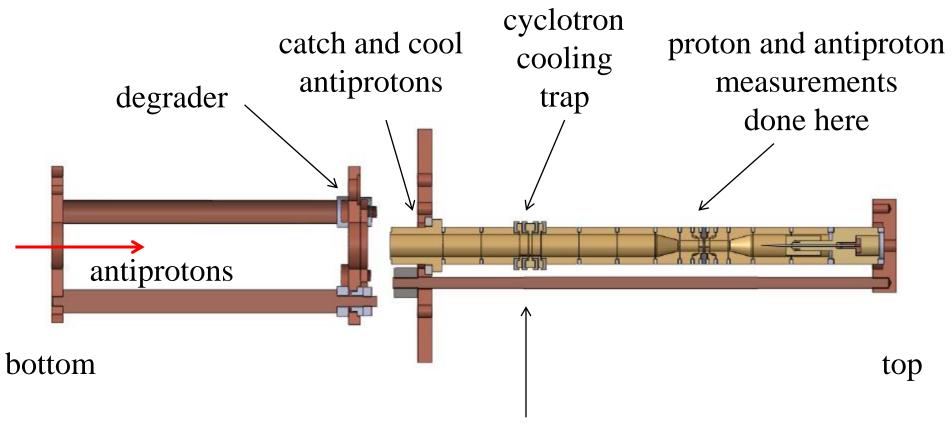
- Cannot explain how a matter universe exists (baryon imbalance is an unsolved mystery)
- Gravity does not fit well (can't be renormalized)
- Cannot explain inflation
- Cannot explain dark energy

The standard model is the great success and great frustration of fundamental particle physics

# Proton and Antiproton Magnetic Moments are Much Smaller

Harder: nuclear magneton rather than Bohr magneton  $\mu_N/\mu_B = m_e/m_p \sim 1/2000$ 

### **For Magnetic Moments: Three Antiproton Traps**



#### more precise measurements will take place here

Located within a self-shielding superconducting solenoid → we invented in part to deal with magnetic noise at CERN

#### **680 Times Improved Pbar to P Comparison**

Selected for a Viewpoint in Physics PHYSICAL REVIEW LETTERS

week ending 29 MARCH 2013

#### S

#### **One-Particle Measurement of the Antiproton Magnetic Moment**

J. DiSciacca,<sup>1</sup> M. Marshall,<sup>1</sup> K. Marable,<sup>1</sup> G. Gabrielse,<sup>1,\*</sup> S. Ettenauer,<sup>1</sup> E. Tardiff,<sup>1</sup> R. Kalra,<sup>1</sup> D. W. Fitzakerley,<sup>2</sup> M. C. George,<sup>2</sup> E. A. Hessels,<sup>2</sup> C. H. Storry,<sup>2</sup> M. Weel,<sup>2</sup> D. Grzonka,<sup>3</sup> W. Oelert,<sup>3,4</sup> and T. Sefzick<sup>3</sup>

(ATRAP Collaboration)

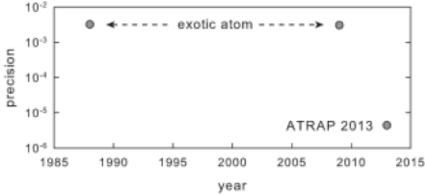
<sup>1</sup>Department of Physics, Harvard University, Cambridge, Massachusetts 02138, USA
<sup>2</sup>Department of Physics and Astronomy, York University, Toronto, Ontario M3J 1P3, Canada
<sup>3</sup>IKP, Forschungszentrum Jülich GmbH, 52425 Jülich, Germany
<sup>4</sup>Institut für Physik, Johannes Gutenberg Universität Mainz, D-5509 Mainz, Germany
(Received 21 January 2013; published 25 March 2013)

For the first time a single trapped antiproton  $(\bar{p})$  is used to measure the  $\bar{p}$  magnetic moment  $\mu_{\bar{p}}$ . The moment  $\mu_{\bar{p}} = \mu_{\bar{p}}S/(\hbar/2)$  is given in terms of its spin S and the nuclear magneton  $(\mu_N)$  by  $\mu_{\bar{p}}/\mu_N = -2.792\,845 \pm 0.000\,012$ . The 4.4 parts per million (ppm) uncertainty is 680 times smaller than previously realized. Comparing to the proton moment measured using the same method and trap electrodes gives  $\mu_{\bar{p}}/\mu_p = -1.000\,000 \pm 0.000\,005$  to 5 ppm, for a proton moment  $\mu_p = \mu_p S/(\hbar/2)$ , consistent with the prediction of the *CPT* theorem.

$$\mu_p/\mu_p = -1.000\,000 \pm 0.000\,005$$
 [5.1 ppm]

PRL 110, 130801 (2013)

 $\mu_{\bar{p}}/\mu_p = -0.9999992 \pm 0.0000044$  [4.4 ppm],



### **Comparing to Other CPT Tests**

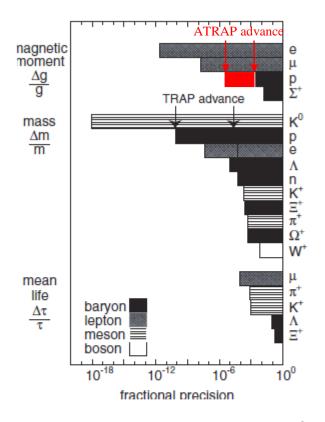


Figure 1: CPT Tests (primarily from the Particle Data Group compilation). Charge-to-mass ratio comparisons are included in "mass" measurements.

- Already one of the most precise antimatter-matter comparisons
- Will be one of the most precise tests if we improve by an additional 1000 to 10,000

### Stringent Low Energy Tests of the Standard Model and Its Symmetries

Illustrate with experiments that have 3 kinds of objectives

1. Testing the Standard Model's most precise prediction by making the most precise measurement of a property of an elementary particle electron dipole

electron magnetic dipole moment

2. Testing very different predictions of the Standard Model and Supersymmetry (and other) models

electron electric dipole moment

also neutron, proton, Hg

Testing the most fundamental symmetry of the Standard Model
 q/m for antiproton and proton mag. moments of e<sup>+</sup> and e<sup>-</sup>

### **Electron Electric Dipole Moment**

- A most precise test of extensions to the standard model
- 12 times more precise than previous measurements

Magnetic moment: 
$$\vec{\mu} = \mu \frac{\vec{S}}{\hbar/2}$$
  
Well measured  
(just reviewed)

Electric dipole moment:  $\vec{d} = d \frac{\vec{S}}{\hbar/2}$ Does this also exist? Why is it interesting?

### **12-Fold More Sensitive Measurement of the Electron Electric Dipole Moment**

#### Gerald Gabrielse Leverett Professor Physics, Harvard University



#### Advanced Cold Molecule EDM

HOW ROUND IS THE ELECTRON

Science

ACME Collaboration: Jacob Baron, Wesley C. Campbell, David DeMille, John M. Doyle, Gerald Gabrielse, Yulia V. Gurevich, Paul W. Hess, Nicholas R. Hutzler, Emil Kirilov, Ivan Kozyryev, Brendon R. O'Leary, Cristian D. Panda, Maxwell F. Parsons, Elizabeth S. Petrik, Ben Spaun, Amar C. Vutha, Adam D. West

#### Science 343, 269 (2014)



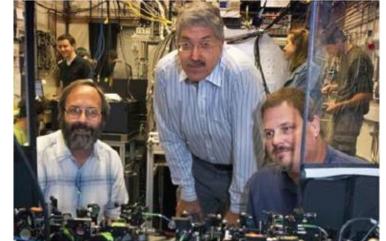
### **ACME Collaboration**

Gerald

Gabrielse

(Harvard)

# Joint effort of 3 research groups



David DeMille (Yale)

ACME PhD



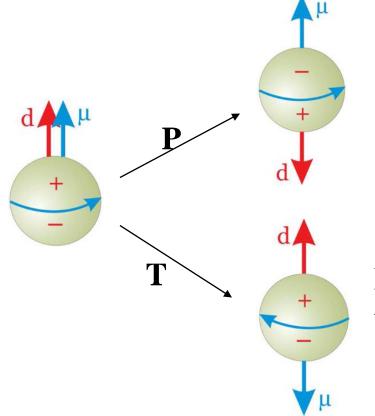
Brendon O'Leary (Paul Hess) Jacob Baron Elizabeth Petrik

Earlier: Amar Vutha, Yulia Gurevich, Emil Kirilov, Ivan Kozyreyv, Wes Campbell

### **Particle EDM Requires Both P and T Violation**

## Magnetic moment: $\vec{\mu} = \mu \frac{\vec{S}}{\hbar/2}$

Electric dipole Moment:  $\vec{d} = d \frac{\vec{S}}{\hbar/2}$ 



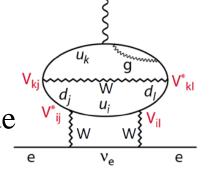
If reality is invariant under parity transformations  $P \rightarrow d = 0$ 

If reality is invariant under time reversal transformations T  $\rightarrow d = 0$ 

#### **Standard Model of Particle Physics Predicts a Non-zero Electron EDM**

Standard model:  $d \sim 10^{-38}$  e-cm

Too small to measure by orders of magnitude best measurement:  $d \sim 2 \times 10^{-27} \text{ e-cm}^{-1}$ 



four-loop level in perturbation theory

M. Pospelov and I. B. Khriplovich, "Electric dipole moment of the W boson and the electron in the Kobayashi-Maskawa model," Sov. J. Nucl. Phys. **53**, 638–640 (1991).

Weak interaction couples quark pairs (generations)

CKM matrix relates to d, s, b quarks (Cabibbo-Kabayashi-Maskawa matrix)

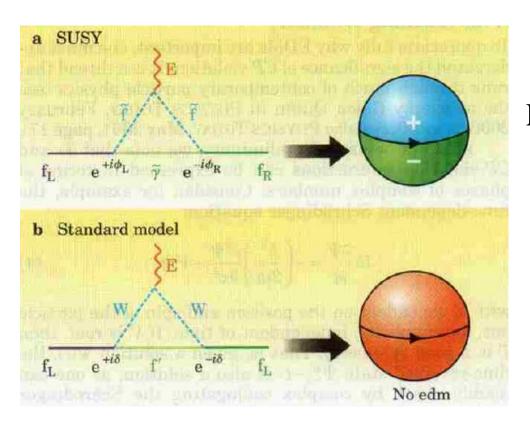
 $\begin{pmatrix} u \\ d' \end{pmatrix}, \begin{pmatrix} c \\ s' \end{pmatrix}, \begin{pmatrix} t \\ b' \end{pmatrix}$ 

almost the unit matrix

0.227 0.973 0.042 0.999

# → Much Bigger, Measureable Electron EDM

#### An example



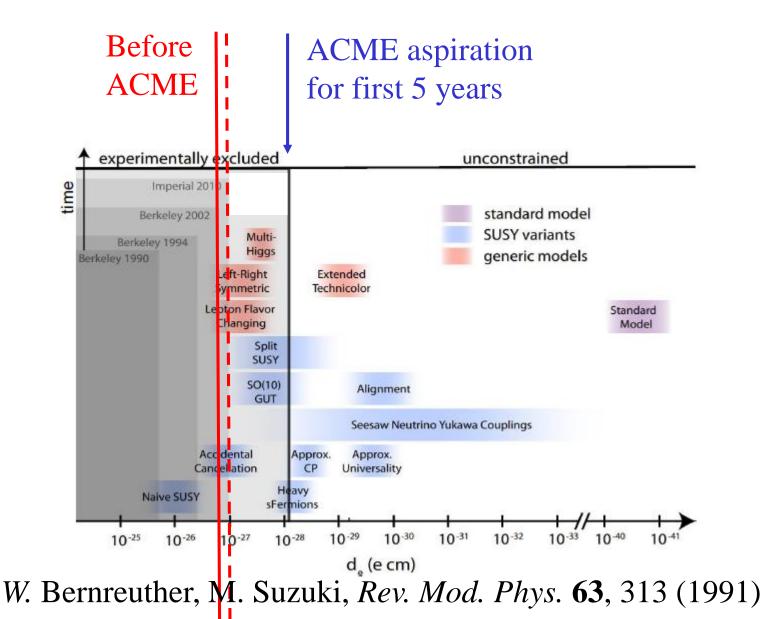
#### Low order contribution → larger moment

Gabrielse

#### Low order contribution → vanishes

From Fortson, Sandars and Barr, Physics Today, 33 (June 2003)

### **Before Our ACME Measurement of Electron EDM**



#### **Before ACME: No Particle EDM Had Yet Been Detected**

#### Electron EDM limit

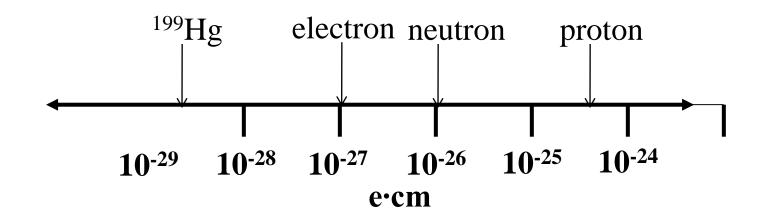
Commins, ... PRL **88**, 071805 (2002)

#### Neutron EDM limit

IIL Grenoble, PRL **97**, 131801 (2006)

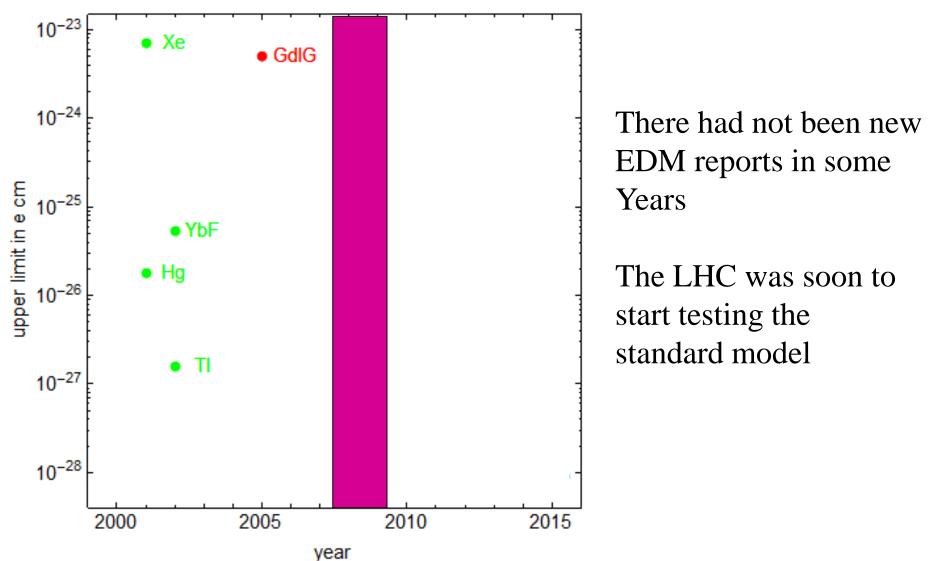
#### Proton EDM limit

Heckel, Fortson, ... PRL **102**, 101601 (2009)  $|d_e| \le 1.6 \times 10^{-27} e \text{ cm}$ Hinds, 2011 1.0  $|d_n| < 2.9 \times 10^{-26} e$  cm  $|d_p| < 7.9 \times 10^{-25} \ e \,\mathrm{cm}$ from  $|d(^{199}\text{Hg})| < 3.1 \times 10^{-29} \ e \text{ cm}$  $|d_n| < 5.8 \times 10^{-26} \ e \,\mathrm{cm}$ also sets

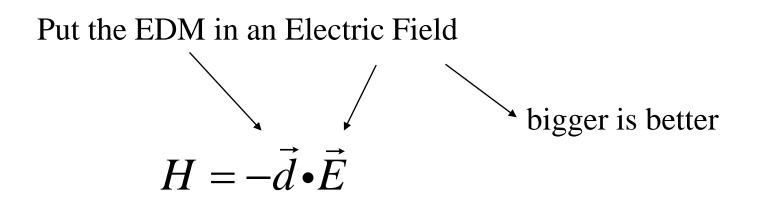


### **Electron EDM Measurements Before ACME**

ACME started 2007 - 2009



#### **How to Measure an Electron EDM**



Measure the energy shift for the system

### Cannot Use Electric Field Directly on an Electron or Proton

Electric field would accelerate an electron out of the apparatus

Simple E and B can be used for neutron EDM measurement (neutron has magnetic moment but no net charge)

Electron EDM are done within atoms and molecules (first molecular ion measurement is now being attempted)

#### **Schiff Theorem – for Electron in an Atom or Molecule**

Schiff (1963) – no atomic or molecular EDM (i.e. linear Stark effect)

- from electron edm
- nonrelativistic quantum mechanics limit

Sandars (1965) – can get atomic or molecular EDM (i.e. linear Stark effect)

- from electron edm
- relativistic quantum mechanics
- get significant enhancement for large Z

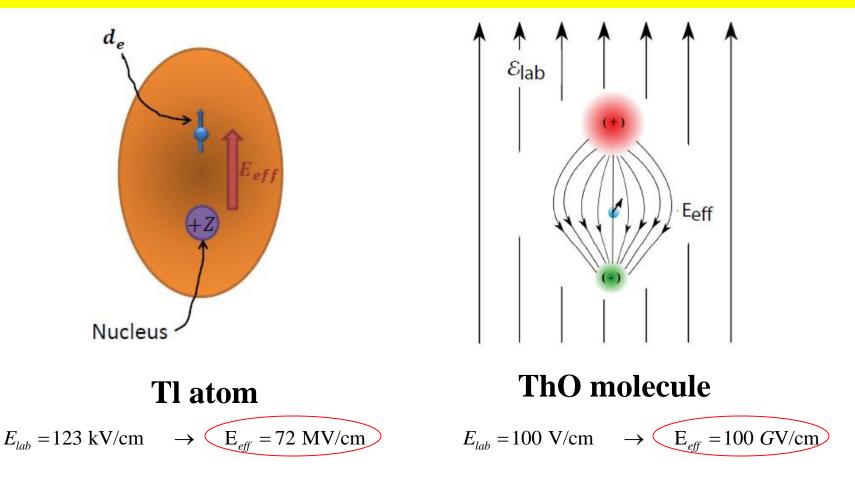
#### Commins, Jackson, DeMille (2007) – intuitive explanation Schiff → Lorentz contraction of the electron EDM in lab frame

Schiff, Phys. Rev. Lett. 132, 2194 (1963);

Sandars, Phys. Rev. Lett. 14, 194 (1965); *ibid* 22, 290 (1966).

Commins, Jackson, DeMille, Am. J. Phys. 75, 532 (2007).

### Why Use a Molecule? → To Make Largest Possible Electric Field on Electron



Molecule can be more easily polarized using nearby energy levels with opposite parity (not generally available in atoms)

### **Promising Molecules**

PHYSICAL REVIEW A 78, 010502(R) (2008)

#### Prospects for an electron electric-dipole moment search in metastable ThO and ThF+

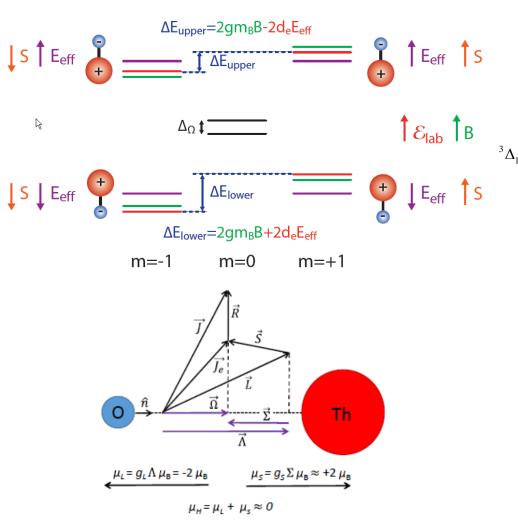
Edmund R. Meyer\* and John L. Bohn

JILA, NIST, and University of Colorado, Department of Physics, Boulder, Colorado 80309-0440, USA (Received 1 May 2008; published 11 July 2008) Molecular calc. project on atomic basis

Molecule	Published	Old [14]	New		
BaF	7.4 [9]	5.1	6.1	_	
YbF Imperia	al 26 [10]	43	32	The 11's and	-
HgF	99 [8]	68	95	Thallium Experime	
PbF Oklahor	ma –29 [8]	-36.6	-31	used 70 N	
a(1) PbO Yale	26.2 [11]	3.2 [22]	23		
HI <sup>+</sup>	0.34 [13]	0.57	0.34		
HfF <sup>+</sup> JILA	24 [15]	18	30		
ThO Harvard - Yale N/A		N/A	104	GV/cm	89
ThF <sup>+</sup>	N/A	N/A	90	_	

high Z

### **ThO H Metastable State**



#### **Omega Doublet**

- Nearly degenerate (300 kHz) opposite party)
- Change internal field direction with no lab field change
- V/cm electric field saturates

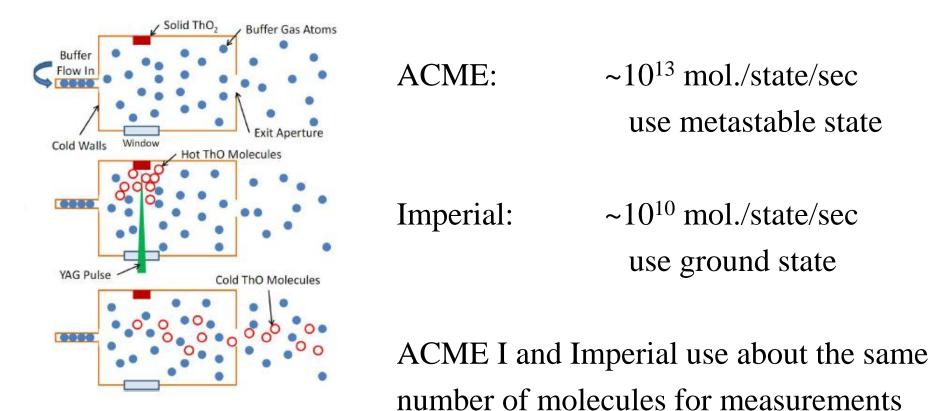
**Tiny magnetic moment** 0.01 Bohr magnetons

long lived (> 1.8 ms)

#### diode lasers, TDA, fiber amplifiers

### **Disadvantage of ThO: Use an excited state**

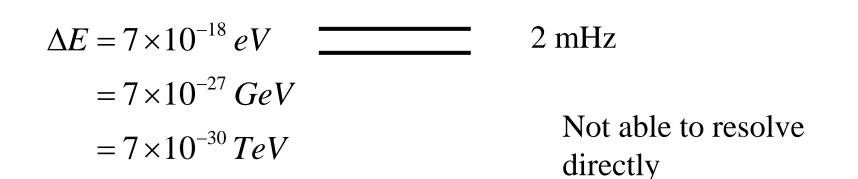
Solution: intense ablation source with Ne buffer gas cooling



Specialty of the Doyle group:

\*Hutzler et al. A cryogenic beam of refractory, chemically reactive molecules with expansion cooling. PCCP (2011).

### **Despite Huge Electric Field the EDM Gives Tiny Shift of Energy Levels**



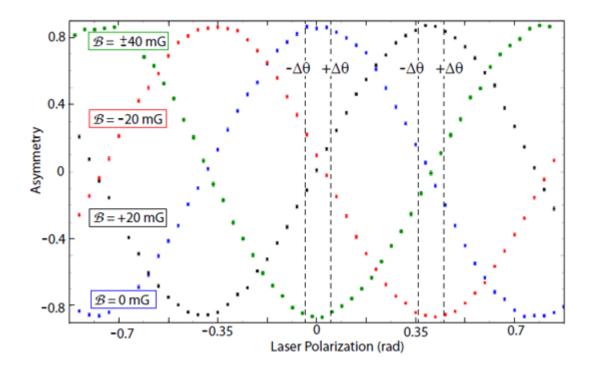
Example is for an electron edm equal the ACME upper limit.

#### **Detect the Small Phase Shift**

set by choice of direction of set by choice of dark state the first of the two orthogonal detection laser polarizations  $|m=1\rangle+e^{i(\phi_{O}+\phi_{I}+\phi)}$  $|m=1\rangle+e^{\iota\phi_O}|m=-1\rangle$  $m=-1\rangle$ time in E, B  $\phi \propto (\mu \mathbf{B} + \mathbf{d} \mathbf{E}) \tau$ X +

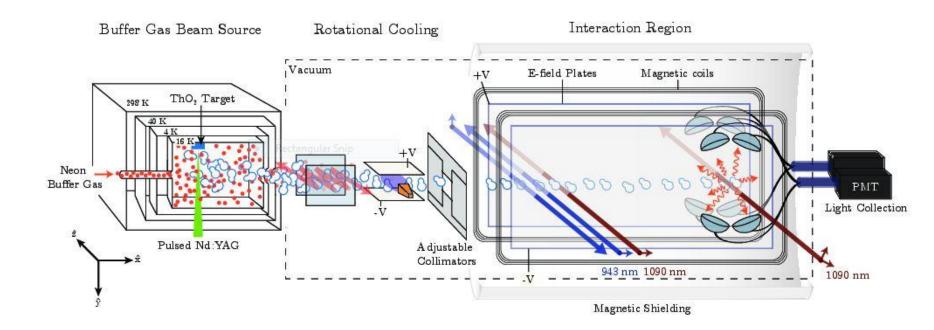
> $T = 1.1 ms \rightarrow \phi = 11 \times 10^{-6} = 0.6 \times 10^{-3}$  degrees Example is for an electron edm equal the ACME upper limit.

#### **Observed Fringes**

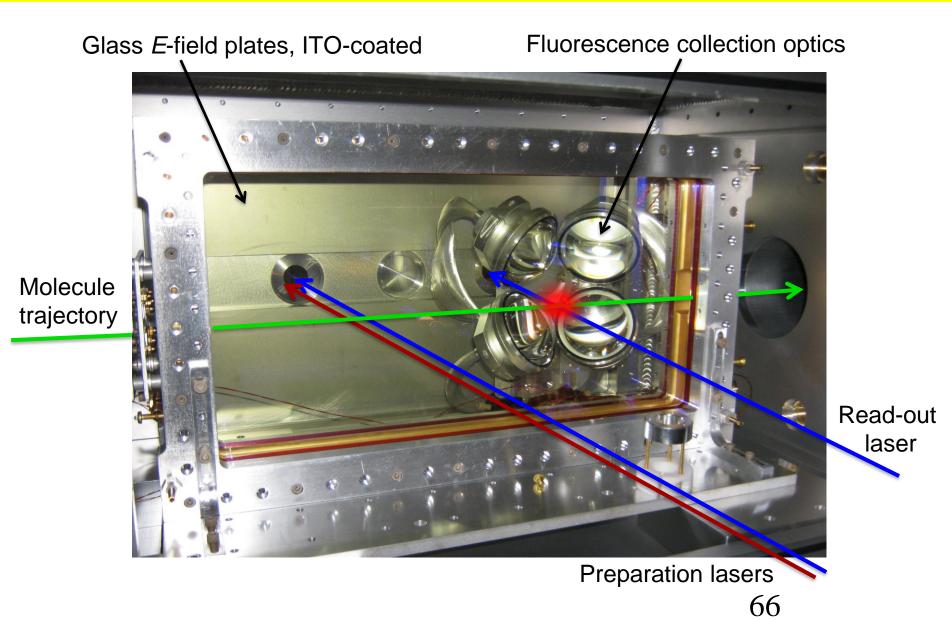


Sit on zero crossing to maximize phase sensitivity

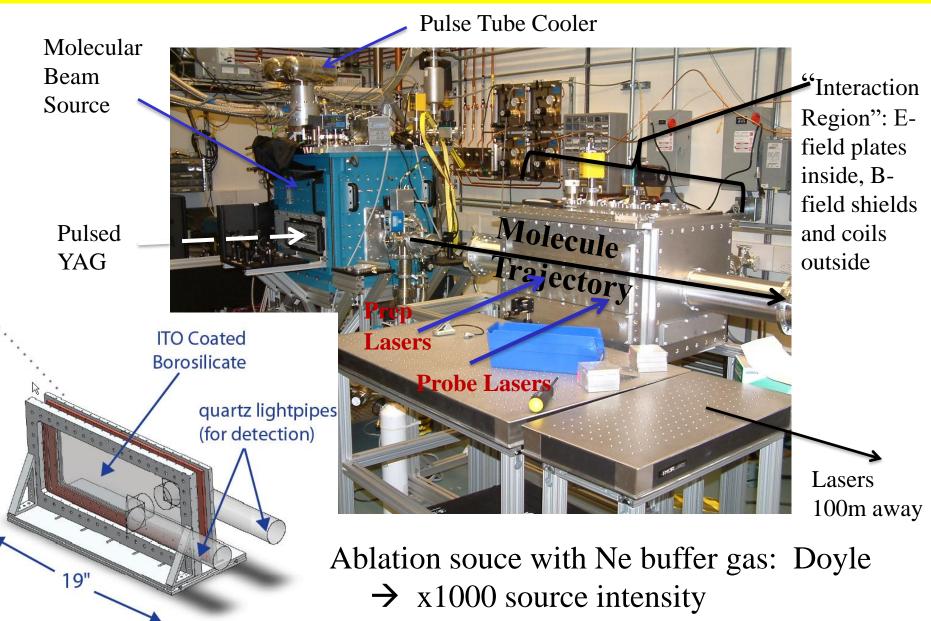
### **Schematic of Experiment**



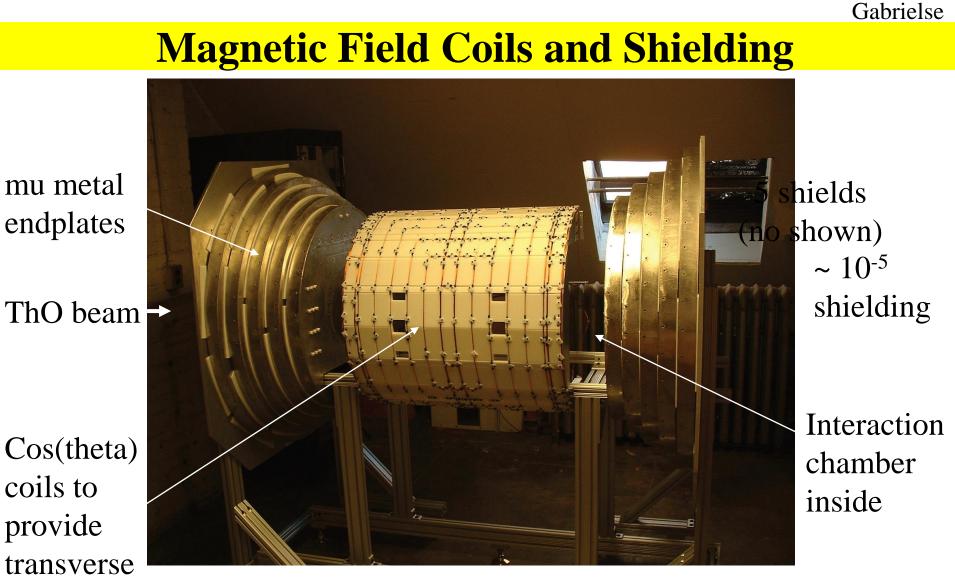
### **Pump -- Evolve in E, B -- Probe**



### **ThO Molecular Beam**



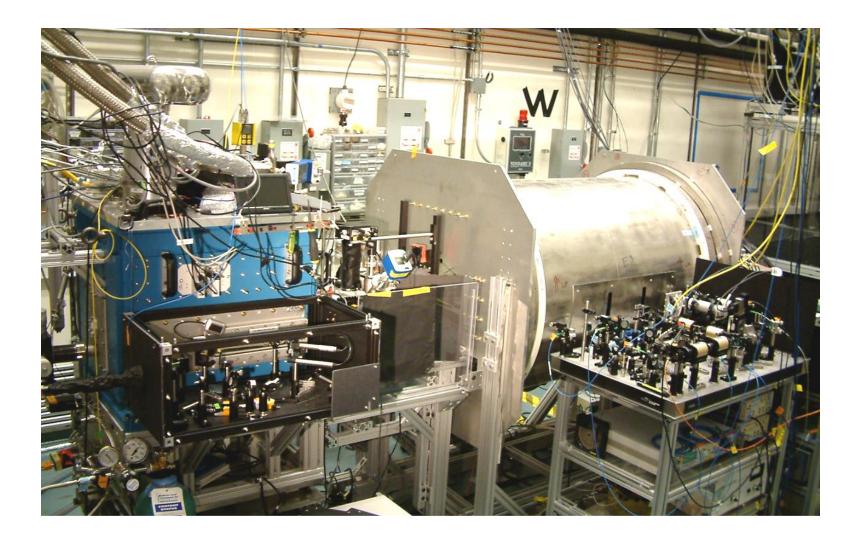
allows use of an excited state

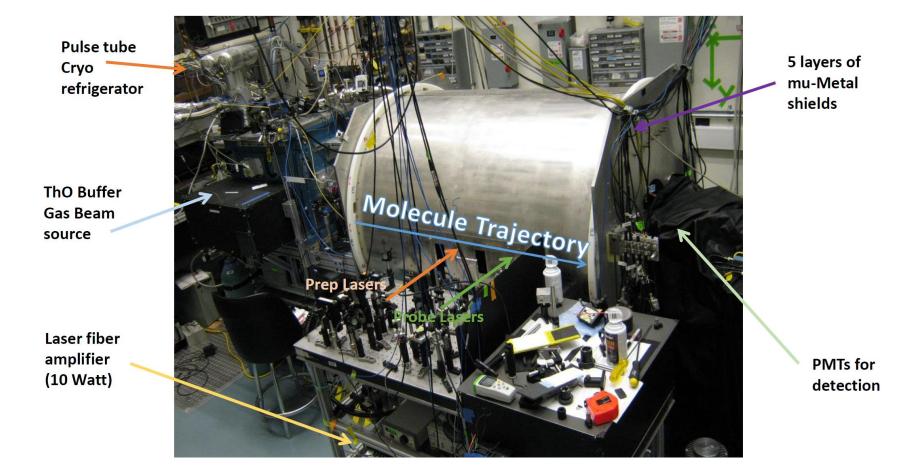


200 mG with uniformity of 10<sup>-3</sup> over 26 cm

B field

### **Molecular Beam Apparatus**





#### **Lasers are 100 Meters Away**

#### Harvard Jefferson Building Harvard LISE Building

100 m optical fibers

ThO Beam (2 floors down)

Lasers, Iodine Clock, Comb

### **Many Lasers**

One of several optical tables w/ ~15 lasers total, modulators, locking electronics, fibers spanning across two buildings



"control room"



# Fast polarization chopping

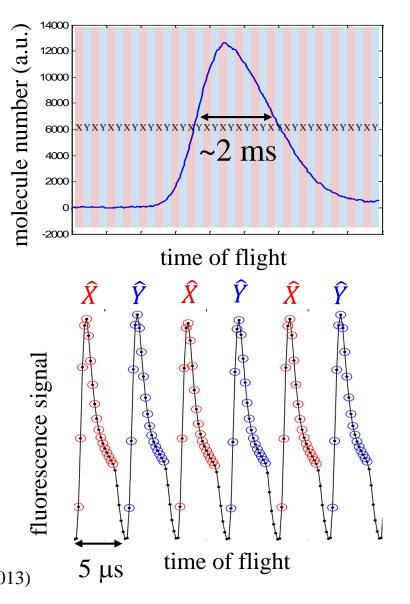
• Fluorescence signal is proportional to phase *and* molecule number:

 $S_X = N_0 \sin^2(\phi)$  $S_Y = N_0 \cos^2(\phi)$ 

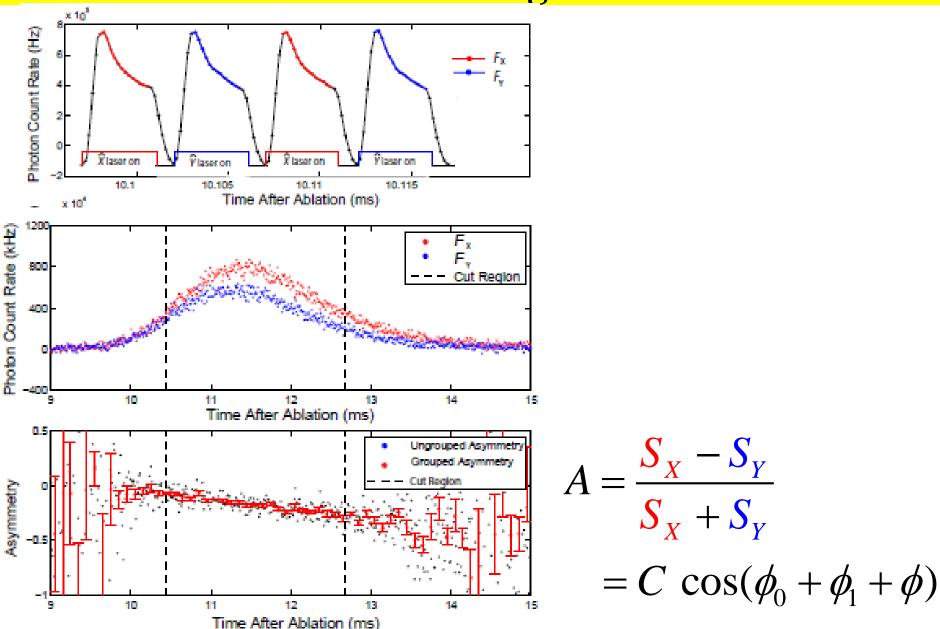
- Rapidly switch probe laser polarization
- Form asymmetry, which is immune to molecule number fluctuations:

$$A = \frac{S_Y - S_x}{S_Y + S_x} = \mathcal{C} \cos(2\phi)$$

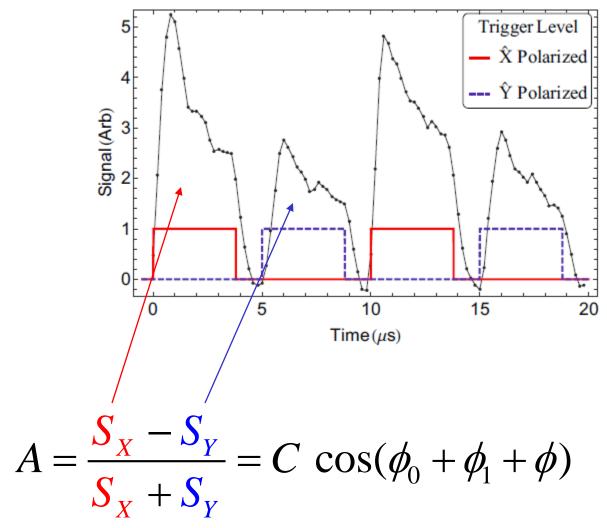
• Achieve shot-noise limited sensitivity.\* \*E. Kirilov *et al.*, PRA **88**, 013844 (2013)



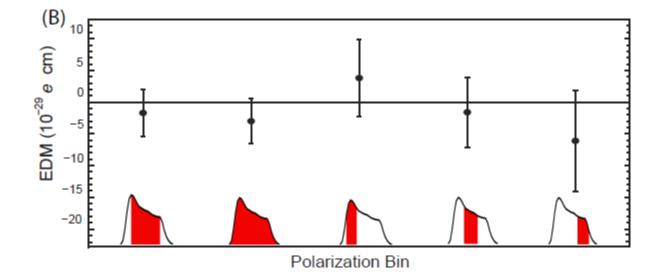
#### **Fast Polarization Switching to Detect the Phase**



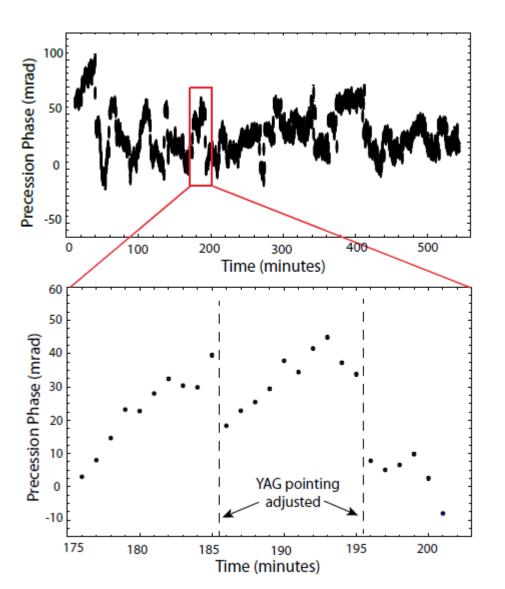
#### **Detect Final Phase Using Two Linear Polarizations**



## **Integrate Over Various Times Intervals**



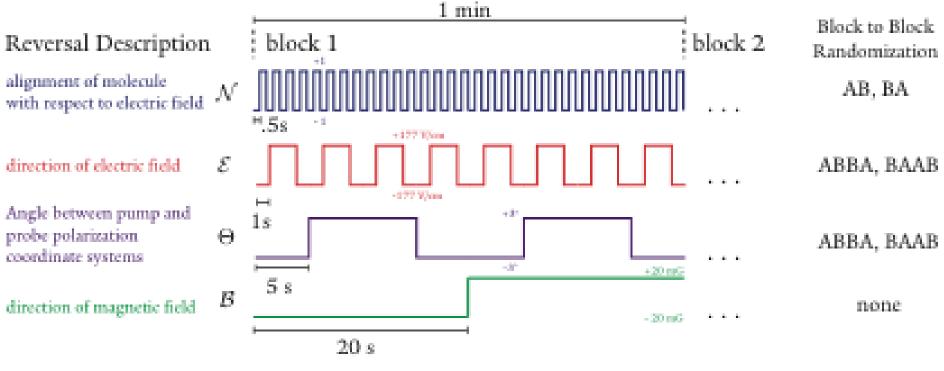
### **Phase Slowly Drifts** as molecular velocity distribution changes



compare EDM limit : 11 µrad

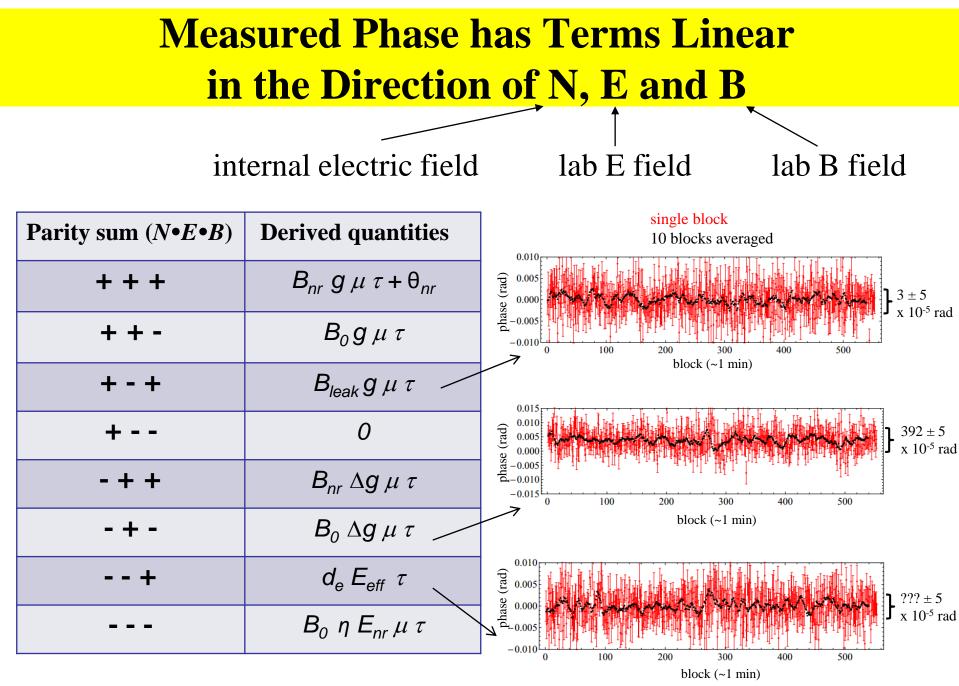
→ Look for much more rapid In phase

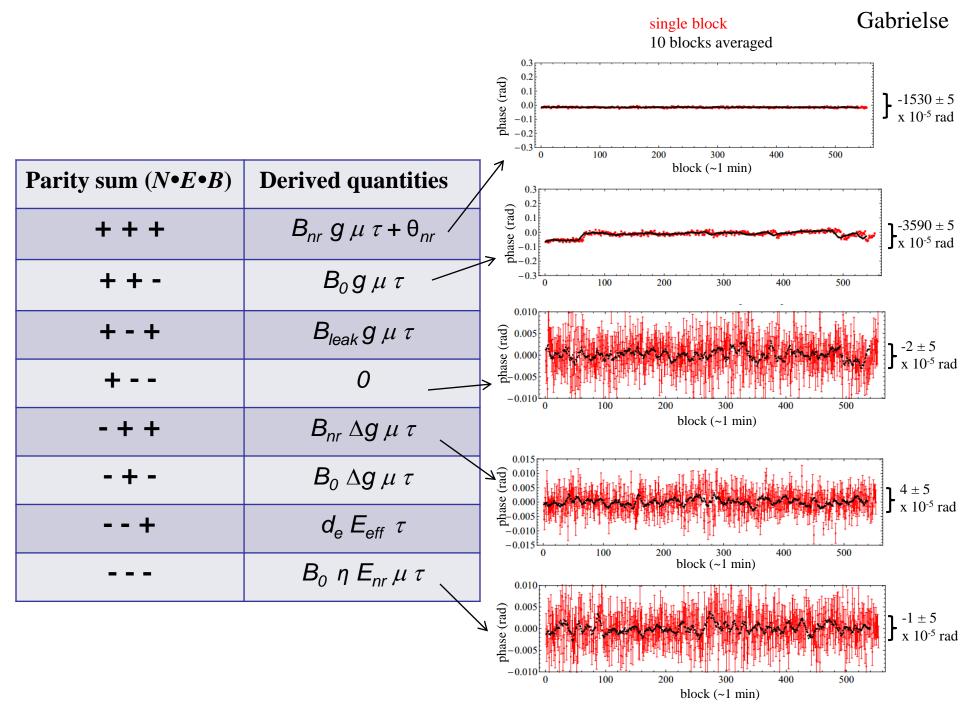
## **Fast Switches**



Degeneracy of States: 4

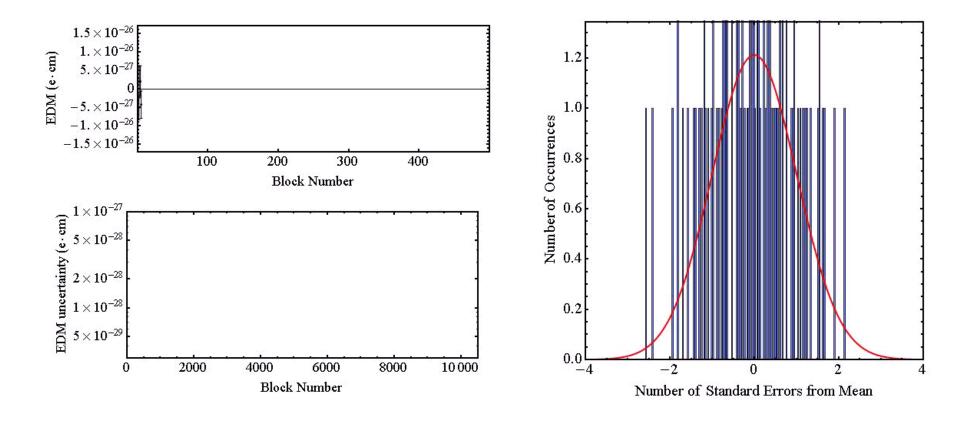
Minimize the time over which the beam, etc. could change





### EDM Measurements – 2013 data

- 10,000 blocks of data  $\rightarrow$  200,000 independent EDM measurements
- ~2 weeks of integration time



### **Uncertainties**

- ~ 40 systematics checks
- Where possible we exaggerated the effect (e.g B gradients)

Parameter	Shift	Uncertainty
$\mathcal{E}^{\mathrm{nr}}$ correction	-0.81	0.66
$\begin{array}{l} \Omega_{\rm r}^{\mathcal{NE}} \mbox{ correction} \\ \phi^{\mathcal{E}} \mbox{ correlated effects} \end{array}$	-0.03	1.58
$\phi^{\mathcal{E}}$ correlated effects	-0.01	0.01
$\phi^{\mathcal{N}}$ correlation		1.25
Non-Reversing $\mathcal{B}$ -field $(\mathcal{B}_z^{\mathrm{nr}})$		0.86
Transverse $\mathcal{B}$ -fields $(\mathcal{B}_x^{\mathrm{nr}}, \mathcal{B}_y^{\mathrm{nr}})$		0.85
$\mathcal{B}$ -Field Gradients		1.24
Prep./Read Laser Detunings		1.31
$\tilde{\mathcal{N}}$ Correlated Detuning		0.90
$\mathcal{E}$ -field Ground Offset		0.16
Total Systematic	-0.85	3.24
Statistical		4.80
Total Uncertainty		5.79

TABLE I. Systematic and statistical errors for  $\omega^{\mathcal{NE}}$ , in units of mrad/s. All errors are added in quadrature. In EDM units, 1.3 mrad/s  $\approx 10^{-29}$  e cm.

## Need Effective Electric Field (from Theory) to Extract EDM

# 104 GV/cm E. R. Meyer, J. L. Bohn, Prospects for an electron electric-dipole moment search in metastable ThO and ThF<sup>+</sup>. *Phys. Rev. A* 78, 010502 (2008).

- 84 GV/cm L. V. Skripnikov, A. N. Petrov, A. V. Titov, Theoretical study of ThO for the electron electric dipole moment search. J. Chem. Phys. 139, 221103 (2013).
- 75.6 GV/cm (3%) ← preprint arrived from India this morning
   T. Flieg and M.K. Nayak
   relativistic, configuration interaction

#### Order of Magnitude Smaller Limit on the Electric Dipole Moment of the Electron

ACME Collaboration: Jacob Baron, Wesley C. Campbell, David DeMille, John M. Doyle, Gerald Gabrielse, Yulia V. Gurevich, Paul W. Hess, Nicholas R. Hutzler, Emil Kirilov, Ivan Kozyryev, Brendon R. O'Leary, Cristian D. Panda, Maxwell F. Parsons, Elizabeth S. Petrik, Ben Spaun, Amar C. Vutha, Adam D. West

(Submitted on 28 Oct 2013 (v1), last revised 7 Nov 2013 (this version, v2))

The Standard Model (SM) of particle physics fails to explain dark matter and why matter survived annihilation with antimatter following the Big Bang. Extensions to the SM, such as weak-scale Supersymmetry, may explain one or both of these phenomena by positing the existence of new particles and interactions that are asymmetric under time-reversal (T). These theories nearly always predict a small, yet potentially measurable  $(10^{-27} \cdot 10^{-30} e \text{ cm})$  electron electric dipole moment (EDM,  $d_e$ ), which is an asymmetric charge distribution along the spin ( $\vec{S}$ ). The EDM is also asymmetric under T. Using the polar molecule thorium monoxide (ThO), we measure  $d_e = (-2.1 \pm 3.7_{stat} \pm 2.5_{syst}) \times 10^{-29} e \text{ cm}$ . This corresponds to an upper limit of  $|d_e| < 8.7 \times 10^{-29} e \text{ cm}$  with 90 percent confidence, an order of magnitude improvement in sensitivity compared to the previous best limits. Our result constrains T-violating physics at the TeV energy scale.



```
d_e = (-2.1 \pm 3.7_{
m stat} \pm 2.5_{
m syst}) 	imes 10^{-29} \; e cm.
```

 $|d_{e}| < 8.7 imes 10^{-29} \ e$  cm

### We actually constrain the EDM and C<sub>S</sub>

$$-d_{e}\mathcal{E}_{eff} - W_{S}C_{S}$$

$$C_{S} = (-1.3 \pm 3.0) \times 10^{-9}$$

$$C_{S} = (-1.3 \pm 3.0) \times 10^{-9}$$
Assuming d=0

From molecular calculation

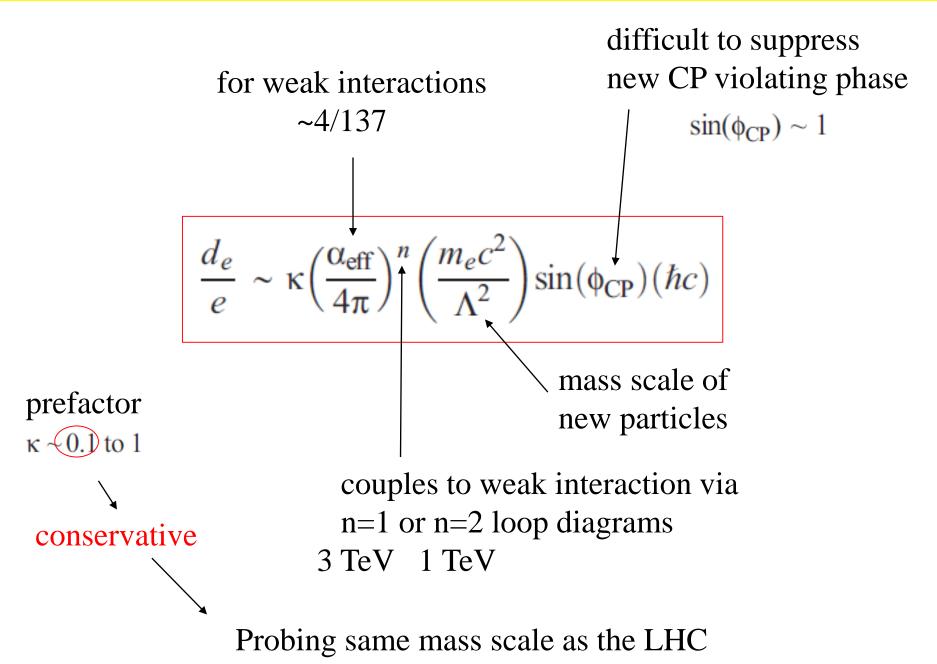
## Sensitivity to Other CP Violating Observables illustrated in Recent Hg EDM Measurement

arXiv.org > physics > arXiv:1601.04339					Search of An
	i i i i i i i i i i i i i i i i i i i				
Physics > Atomic Physics					
Reduced Limit on the Permane	nt Electri	c Dipole	Momen	t of $^{199}$ Hg	
B. Graner, Y. Chen, E. G. Lindahl, B. R. Heckel					

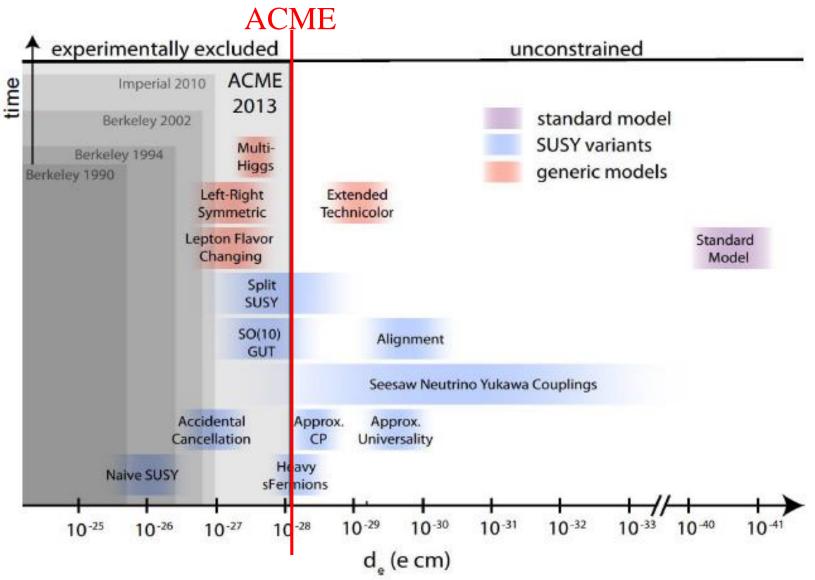
TABLE III. Limits on CP-violating observables from the <sup>199</sup>Hg EDM limit. Each limit is based on the assumption that it is the sole contribution to the atomic EDM. In principle, the result for  $\mathbf{d}_n$  supercedes [11] as the best neutron EDM limit.

Quantity	Expression	Limit	Ref.
$\mathbf{d}_n$	$\mathbf{S}_{Hg}/(1.9~\mathrm{fm}^2)$	$1.6 \times 10^{-26} \ e \cdot \mathrm{cm}$	[21]
$\mathbf{d}_p$	$1.3 \times \mathbf{S}_{Hg} / (0.2 \text{ fm}^2)$	$2.0 \times 10^{-25} e \cdot \mathrm{cm}$	[21]
$ar{g}_0$	${f S}_{Hg}/(0.135 \ e \cdot { m fm}^3)$	$2.3 \times 10^{-12}$	[5]
$ar{g}_1$	$\mathbf{S}_{Hg}/(0.27 \ e \cdot \mathrm{fm}^3)$	$1.1 \times 10^{-12}$	[5]
$ar{g}_2$	$\mathbf{S}_{Hg}/(0.27 \ e \cdot \mathrm{fm}^3)$	$1.1 \times 10^{-12}$	[5]
$ar{ heta}_{QCD}$	$ar{g}_0/0.0155$	$1.5 \times 10^{-10}$	[22, 23]
$(\widetilde{d}_u - \widetilde{d}_d)$	$\bar{g}_1/(2 \times 10^{14} \mathrm{cm}^{-1})$	$5.7 \times 10^{-27} \text{ cm}$	[25]
$C_S$	$\mathbf{d}_{Hg}/(5.9 \times 10^{-22} \ e \cdot \mathrm{cm})$	$1.3 \times 10^{-8}$	[15]
$C_P$	$\mathbf{d}_{Hg}/(6.0 \times 10^{-23} \ e \cdot \mathrm{cm})$	$1.2 \times 10^{-7}$	[15]
$C_T$	$\mathbf{d}_{Hg}/(4.89 \times 10^{-20} \ e \cdot \mathrm{cm})$	$1.5 \times 10^{-10}$	see text

### **Constraining New Physics on the 1 to 3 TeV Scale**



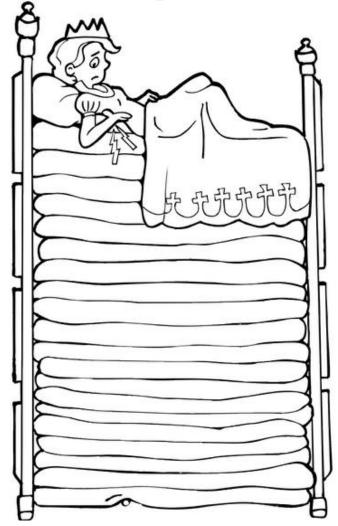
### **2014 ACME Electron EDM Measurement**



W. Bernreuther, M. Suzuki, Rev. Mod. Phys. 63, 313 (1991)

## How Big is 8 x 10<sup>-29</sup> e cm?

How sensitive was our princess to the hidden pea?



Scale size of the polarization cloud around the electron  $\rightarrow$  earth

Shift in earth center by 2 nm

earth-sized polarization cloud around electron (scale classical electron radius)

### **Relationship to LHC Physics**

The LHC is exciting and important but EDMs also play a role

- should get an improved electron EDM on the LHC time scale
- If the LHC sees new particles, is CP violation involved?
- If the LHC sees nothing, EDM game is the only one in town
- Would be great to use LHC results and ours together to see what we have learned together about Standard Model extensions

#### ATLAS SUSY Searches\* - 95% CL Lower Limits

Status: SUSY 2013

	Model	e, μ, τ, γ	Jets	E <sup>miss</sup> T	∫£dt[ft	<sup>-1</sup> ] Mass limit	Reference
Inclusive Searches	$ \begin{array}{l} \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \text{MSUGRA/CMSSM} \\ \overline{q}q, \overline{q} \rightarrow q \overline{x}_{1}^{0} \\ \overline{g}\mathcal{B}, \mathcal{B} \rightarrow q q \overline{x}_{1}^{0} \\ \overline{g}\mathcal{B}, \mathcal{B} \rightarrow q q \overline{x}_{1}^{1} \\ \overline{g}\mathcal{B}, \mathcal{B} \rightarrow q q (\mathcal{U}/\mathcal{U}/\mathcal{W}) \overline{\chi}_{1}^{0} \\ \overline{g}\mathcal{B}, \mathcal{B} \rightarrow q q (\mathcal{U}/\mathcal{U}/\mathcal{W}) \overline{\chi}_{1}^{0} \\ \overline{g}\mathcal{MSB} (\overline{z} \text{ NLSP}) \\ \overline{G}\mathcal{M} (\text{bino NLSP}) \\ \overline{G}\mathcal{GM} (\text{wino NLSP}) \\ \overline{G}\mathcal{GM} (\text{higgsino-hino NLSP}) \\ \overline{G}\mathcal{GM} (\text{higgsino-hino NLSP}) \\ \overline{G}\mathcal{GM} (\text{higgsino NLSP}) \\ \overline{G}\mathcal{GM}$	$\begin{array}{c} 0 \\ 1  e, \mu \\ 0 \\ 0 \\ 1  e, \mu \\ 2  e, \mu \\ 2  e, \mu \\ 1 - 2  \tau \\ 2  \gamma \\ 1  e, \mu + \gamma \\ \gamma \\ 2  e, \mu (Z) \\ 0 \end{array}$	2-6 jets 3-6 jets 7-10 jets 2-6 jets 2-6 jets 3-6 jets 0-3 jets 0-2 jets - 1 b 0-3 jets mono-jet	Yes Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.3 20.3 20.3 20.3 20.3 20.3 4.7 20.7 4.8 4.8 4.8 4.8 5.8 10.5	4,8     1.7 TeV     m(q)=m(g)       8     1.2 TeV     ary m(q)       9     1.1 TeV     ary m(q)       9     740 GeV     m(t)=0 GeV       9     740 GeV     m(t)=0 GeV       8     1.1 TeV     ary m(q)       8     1.13 TeV     m(t)=0 GeV       8     1.12 TeV     m(t)=0 GeV       8     1.12 TeV     m(t)=0 GeV       8     1.12 TeV     m(t)=0 GeV       8     1.24 TeV     tape<15	ATLAS-CONF-2013-047 ATLAS-CONF-2013-052 1308.1841 ATLAS-CONF-2013-047 ATLAS-CONF-2013-047 ATLAS-CONF-2013-047 ATLAS-CONF-2013-052 ATLAS-CONF-2013-025 1209.0753 ATLAS-CONF-2012-144 1211.1167 ATLAS-CONF-2012-152 ATLAS-CONF-2012-152 ATLAS-CONF-2012-147
3 <sup>rd</sup> gen. <u>8</u> med.	$B \rightarrow b \bar{b} \bar{v}_{d}^{0}$ $B \rightarrow t \bar{t} \bar{v}_{1}^{0}$ $B \rightarrow t \bar{t} \bar{v}_{1}^{1}$ $B \rightarrow b \bar{t} \bar{v}_{1}^{1}$	0 0 0-1 e,μ 0-1 e,μ	3 b 7-10 jets 3 b 3 b	Yes Yes Yes Yes	20.1 20.3 20.1 20.1	8         1.2 TeV         m(花):5600 GeV           8         1.1 TeV         m(花):350 GeV           8         1.34 TeV         m(花):350 GeV           8         1.3 TeV         m(花):300 GeV	ATLAS-CONF-2013-061 1308.1841 ATLAS-CONF-2013-061 ATLAS-CONF-2013-061
3 <sup>24</sup> gen. squarks direct production	$\begin{array}{l} b_1b_1, \ b_1 \rightarrow b\overline{t}_1^n \\ b_1b_1, \ b_1 \rightarrow t\overline{t}_1^n \\ \overline{t}_1\overline{t}_1(\text{light}), \ \overline{t}_1 \rightarrow b\overline{t}_1^n \\ \overline{t}_1\overline{t}_1(\text{light}), \ \overline{t}_1 \rightarrow b\overline{t}_1^n \\ \overline{t}_1\overline{t}_1(\text{light}), \ \overline{t}_1 \rightarrow b\overline{t}_1^n \\ \overline{t}_1\overline{t}_1(\text{medum}), \ \overline{t}_1 \rightarrow t\overline{t}_1^n \\ \overline{t}_1\overline{t}_1(\text{medum}), \ \overline{t}_1 \rightarrow t\overline{t}_1^n \\ \overline{t}_1\overline{t}_1(\text{medum}), \ \overline{t}_1 \rightarrow t\overline{t}_1^n \\ \overline{t}_1\overline{t}_1, \ \overline{t}_1 \rightarrow t\overline{t}_1^n \\ \overline{t}_1\overline{t}_1, \ \overline{t}_1 \rightarrow c\overline{t}_1^n \\ \overline{t}_1\overline{t}_1, \ \overline{t}_1 \rightarrow c\overline{t}_1^n \\ \overline{t}_1\overline{t}_1, \ \overline{t}_1 \rightarrow c\overline{t}_1^n \\ \overline{t}_2\overline{t}_1(\text{neturn}) \text{IGMSB}) \\ \overline{t}_2\overline{t}_2\overline{t}_2, \ \overline{t}_2 \rightarrow \overline{t}_1 + Z \end{array}$	0 2 e, µ (SS) 1-2 e, µ 2 e, µ 2 e, µ 0 1 e, µ 0 1 e, µ 3 e, µ (Z)	2 b 0-3 b 1-2 b 0-2 jets 2 jets 2 b 1 b 2 b 1 ono-jet/c-t 1 b 1 b	Yes Yes Yes Yes Yes Yes Yes Yes Yes	20.1 20.7 4.7 20.3 20.3 20.1 20.7 20.5 20.3 20.7 20.7	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	1308.2631 ATLAS-CONF-2013-007 1208.4305, 1209.2102 ATLAS-CONF-2013-048 ATLAS-CONF-2013-048 ATLAS-CONF-2013-057 ATLAS-CONF-2013-024 ATLAS-CONF-2013-025 ATLAS-CONF-2013-025
EW direct	$ \begin{array}{c} \tilde{t}_{\perp,R}\tilde{t}_{\perp,R},\tilde{t} \rightarrow \tilde{\alpha}_{1}^{0} \\ \tilde{x}_{1}^{-1}\tilde{x}_{1}^{-1},\tilde{x}_{1}^{-1} \rightarrow \tilde{\nu}\nu(\ell r) \\ \tilde{x}_{1}^{-1}\tilde{x}_{1}^{-1},\tilde{x}_{1}^{-1} \rightarrow r\nu(r r) \\ \tilde{x}_{1}^{-1}\tilde{x}_{0}^{0} \rightarrow \tilde{t}_{1}\nu\tilde{a}_{1}^{0}\ell(r r), \ell r\tilde{r}_{L}\ell(r r) \\ \tilde{x}_{1}^{-1}\tilde{x}_{0}^{0} \rightarrow W\tilde{x}_{1}^{0}\ell(r r) \\ \tilde{x}_{1}^{-1}\tilde{x}_{0}^{0} \rightarrow W\tilde{x}_{1}^{0}h\tilde{x}_{1}^{0} \end{array} $	2 e, µ 2 e, µ 2 τ 3 e, µ 3 e, µ 1 e, µ	0 - 0 2 b	Yes Yes Yes Yes Yes Yes	20.3 20.3 20.7 20.7 20.7 20.7 20.3	ž         85-315 GeV         m(t <sup>2</sup> <sub>1</sub> )=0 GeV $\chi_1^+$ 125-450 GeV         m(t <sup>2</sup> <sub>1</sub> )=0 GeV, m(t <sup>2</sup> , $\bar{\tau}$ )=0.5(m(t <sup>2</sup> <sub>1</sub> )+m(t <sup>2</sup> <sub>1</sub> )) $\chi_1^+$ 180-330 GeV         m(t <sup>2</sup> <sub>1</sub> )=0 GeV, m(t <sup>2</sup> , $\bar{\tau}$ )=0.5(m(t <sup>2</sup> <sub>1</sub> )+m(t <sup>2</sup> <sub>1</sub> )) $\chi_1^+, \chi_2^0$ 600 GeV         m(t <sup>2</sup> <sub>1</sub> )=0 GeV, m(t <sup>2</sup> , $\bar{\tau}$ )=0.5(m(t <sup>2</sup> <sub>1</sub> )+m(t <sup>2</sup> <sub>1</sub> )) $\chi_1^+, \chi_2^0$ 600 GeV         m(t <sup>2</sup> <sub>1</sub> )=m(t <sup>2</sup> <sub>2</sub> ), m(t <sup>2</sup> <sub>1</sub> )=0.5(m(t <sup>2</sup> <sub>1</sub> )+m(t <sup>2</sup> <sub>1</sub> )) $\chi_1^+, \chi_2^0$ 315 GeV         m(t <sup>2</sup> <sub>1</sub> )=m(t <sup>2</sup> <sub>2</sub> ), m(t <sup>2</sup> <sub>1</sub> )=0.5(m(t <sup>2</sup> <sub>1</sub> )+m(t <sup>2</sup> <sub>1</sub> )) $\chi_1^+, \chi_2^0$ 285 GeV         m(t <sup>2</sup> <sub>1</sub> )=m(t <sup>2</sup> <sub>2</sub> ), m(t <sup>2</sup> <sub>1</sub> )=0, skeptons decoupled	ATLAS-CONF-2013-049 ATLAS-CONF-2013-049 ATLAS-CONF-2013-028 ATLAS-CONF-2013-035 ATLAS-CONF-2013-035 ATLAS-CONF-2013-035
Long-lived particles	Direct $\tilde{x}_1^+ \tilde{x}_1^-$ prod., long-lived $\tilde{x}_1^+$ Stable, stopped g R-hadron GMSB, stable $\tau, \tilde{x}_1^0 \rightarrow \tau(a, \bar{\mu}) + \tau(c)$ GMSB, $\tilde{x}_1^0 \rightarrow \gamma \tilde{G}$ , long-lived $\tilde{x}_1^0$ $\bar{q}q, \tilde{x}_1^0 \rightarrow qq\mu$ (RPV)	0	1 jet 1-5 jets - -	Yes Yes Yes -	20.3 22.9 15.9 4.7 20.3	X <sup>+</sup> 270 GeV         m(t <sup>2</sup> <sub>1</sub> )-m(t <sup>2</sup> <sub>2</sub> )=160 MeV, $\tau(t21)=0.2 \text{ ns}$ 8         832 GeV         m(t <sup>2</sup> <sub>1</sub> )-m(t <sup>2</sup> <sub>2</sub> )=160 MeV, $\tau(t21)=0.2 \text{ ns}$ 8         832 GeV         m(t <sup>2</sup> <sub>1</sub> )=10 GeV, 10 µs< $\tau(t2)<0.0 \text{ s}$ 10         475 GeV         10-tanµ<50	ATLAS-CONF-2013-069 ATLAS-CONF-2013-057 ATLAS-CONF-2013-058 1304.6310 ATLAS-CONF-2013-092
RPV	$ \begin{array}{l} \label{eq:LFV} LFV \; pp \!$	$\begin{array}{c} 2 \ e, \mu \\ 1 \ e, \mu + \tau \\ 1 \ e, \mu \\ \phi_e \\ \gamma \\ 0 \\ 2 \ e, \mu \\ \gamma \end{array} \\ \begin{array}{c} 0 \\ 0 \\ 2 \ e, \mu \\ (SS) \end{array}$	7 jets 7 jets 6-7 jets 0-3 b	- Yes Yes Yes - Yes	4.6 4.7 20.7 20.7 20.3 20.7	Pr         1.61 TeV         Z <sub>211</sub> =0.10, Z <sub>122</sub> =0.05           Pr         1.1 TeV         Z <sub>211</sub> =0.10, Z <sub>122</sub> =0.05           Q. g         1.2 TeV         m(Q)=m(g), Cr <sub>1,27</sub> Q. g         760 GeV         m(Q)=m(g), Cr <sub>1,27</sub> X <sub>1</sub> 760 GeV         m(t <sup>2</sup> <sub>1</sub> )>300 GeV, Z <sub>122</sub> >0           X <sub>1</sub> 350 GeV         m(t <sup>2</sup> <sub>1</sub> )>B0(GeV, Z <sub>122</sub> >0           8         880 GeV         BR(t)=BR(b)=BR(c)=0%	1212.1272 1212.1272 ATLAS-CONF-2012.140 ATLAS-CONF-2013-096 ATLAS-CONF-2013-096 ATLAS-CONF-2013-091 ATLAS-CONF-2013-007
Other		$2 e, \mu$ (SS) 0 $\sqrt{s} = 8 \text{ TeV}$ partial data	4 jets 1 b mono-jet √s = full o	Yes Yes 8 TeV data	4.6 14.3 10.5	sgluon         100-287 GeV         incl. limit from 1110.2693           sgluon         800 GeV         m(x)<80 GeV, limit ok:687 GeV for D8	1210.4826 ATLAS-CONF-2013-051 ATLAS-CONF-2012-147

\*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 a theoretical signal cross section uncertainty.

### $\int \mathcal{L} dt = (4.6 - 22.9) \text{ fb}^{-1} \quad \sqrt{s} = 7, 8 \text{ TeV}$

https://twiki.cern.ch/twiki/pub/AtlasPublic/CombinedSummaryPlots/AtlasSearchesSUSY\_SUSY2013.pdf

## Lots of Theory Papers in Reaction to the ACME Limit

~ 40 papers in a couple months

#### Theoretical Prediction and Impact of Fundamental Electric Dipole Moments

#### Sebastian A.R. Ellis, Gordon L. Kane

(Submitted on 29 May 2014)

The predicted Standard Model (SM) electric dipole moments (EDMs) of electrons and guarks are tiny, providing an important window to observe new physics. Theories beyond the SM typically allow relatively large EDMs. The EDMs depend on the relative phases of terms in the effective Lagrangian of the extended theory, which are generally unknown. Underlying theories, such as string/M-theories compactified to four dimensions, could predict the phases and thus EDMs in the resulting supersymmetric (SUSY) theory. Earlier one of us, with collaborators, made such a prediction and found, unexpectedly, that the phases were predicted to be zero at tree level in the theory at the unification or string scale  $\sim {\cal O}(10^{16}~{
m GeV})$ . Electroweak (EW) scale EDMs still arise via running from the high scale, and depend only on the SM Yukawa couplings that also give the CKM phase. Here we extend the earlier work by studying the dependence of the low scale EDMs on the constrained but not fully known fundamental Yukawa couplings. The dominant contribution is from two loop diagrams and is not sensitive to the choice of Yukawa texture. The electron EDM should not be found to be larger than about  $5 imes 10^{-30}e$  cm, and the neutron EDM should not be larger than about  $5 imes 10^{-29}e$  cm. These values are quite a bit smaller than the reported predictions from Split SUSY and typical effective theories, but much larger than the Standard Model prediction. Also, since models with random phases typically give much larger EDMs, it is a significant testable prediction of compactified M-theory that the EDMs should not be above these upper limits. The actual EDMs can be below the limits, so once they are measured they could provide new insight into the fundamental Yukawa couplings of leptons and quarks. We comment also on the role of strong CP violation. EDMs probe fundamental physics near the Planck scale.

## EDM should be just smaller than our limit

### "Testable prediction of compatified M-theory"

The dominant contribution is from two loop diagrams and is not sensitive to the choice of Yukawa texture. The electron EDM should not be found to be larger than about  $5\times 10^{-30}e$  cm, and the

#### should not be ... larger than

Gabrielse

### **One Baryogenesis Model**

Search or

arXiv.org > hep-ph > arXiv:1406.0517

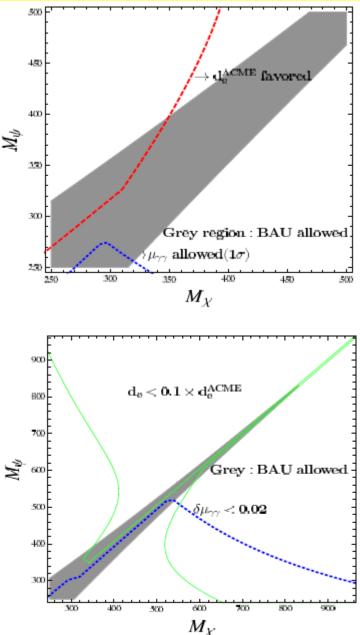
High Energy Physics - Phenomenology

#### Electroweak Baryogenesis, Electric Dipole Moments, and Higgs Diphoton Decays

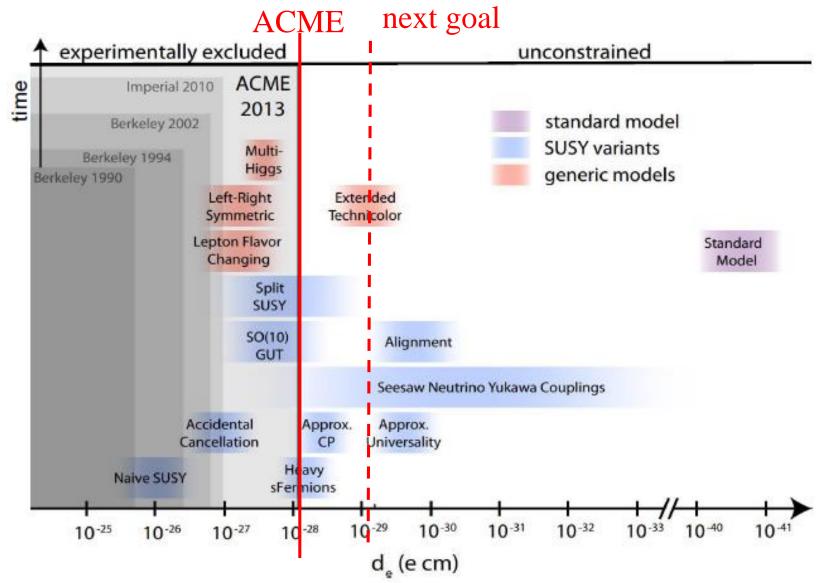
Wei Chao, Michael J. Ramsey-Musolf

(Submitted on 2 Jun 2014)

We study the viability of electroweak baryogenesis in a two Higgs doublet model scenario augmented by vector-like, electroweakly interacting fermions. Considering a limited, but illustrative region of the model parameter space, we obtain the observed cosmic baryon asymmetry while satisfying present constraints from the non-observation of the permanent electric dipole moment (EDM) of the electron and the combined ATLAS and CMS result for the Higgs boson diphoton decay rate. The observation of a non-zero electron EDM in a next generation experiment and/or the observation of an excess (over the Standard Model) of Higgs to diphoton events with the 14 TeV LHC run or a future  $e^+e^-$  collider would be consistent with generation of the observed baryon asymmetry in this scenario.



### **ACME** > Nearing Data Taking for Generation II



W. Bernreuther, M. Suzuki, Rev. Mod. Phys. 63, 313 (1991)

## **STIRAP, etc. Improvements**

PHYSICAL REVIEW A 93, 052110 (2016)

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#### Stimulated Raman adiabatic passage preparation of a coherent superposition of ThO $H^3\Delta_1$ states for an improved electron electric-dipole-moment measurement

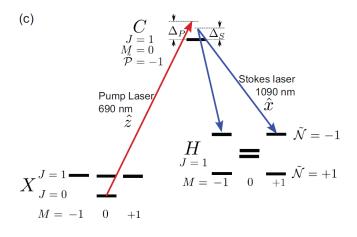
C. D. Panda,<sup>1,\*</sup> B. R. O'Leary,<sup>2</sup> A. D. West,<sup>2</sup> J. Baron,<sup>1</sup> P. W. Hess,<sup>1,†</sup> C. Hoffman,<sup>1,‡</sup> E. Kirilov,<sup>2,§</sup> C. B. Overstreet,<sup>1,¶</sup> E. P. West,<sup>1</sup> D. DeMille,<sup>2</sup> J. M. Doyle,<sup>1</sup> and G. Gabrielse<sup>1</sup>

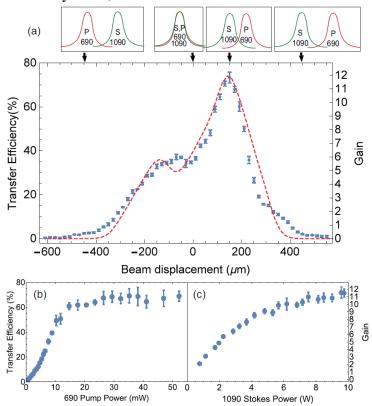
<sup>1</sup>Department of Physics, Harvard University, Cambridge, Massachusetts 02138, USA

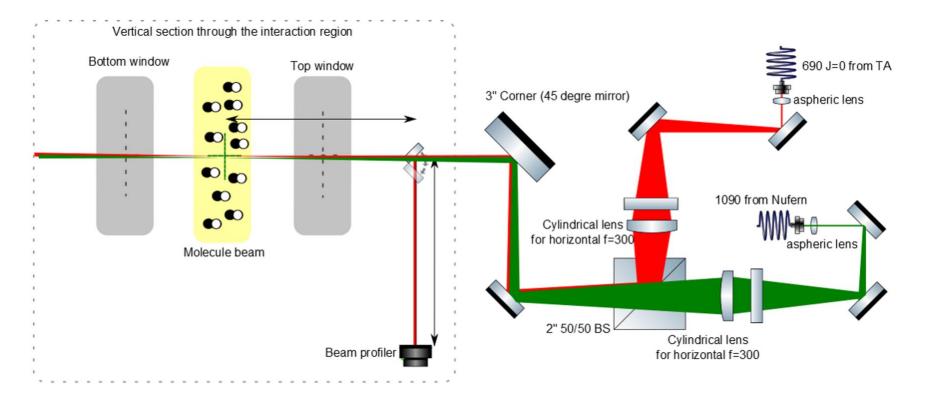
Department of Physics, Harvara University, Cambriage, Massachusetts 02138, USA

<sup>2</sup>Department of Physics, Yale University, New Haven, Connecticut 06511, USA

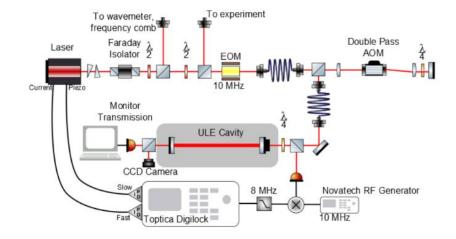
(Received 28 March 2016; published 16 May 2016)

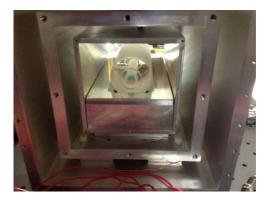


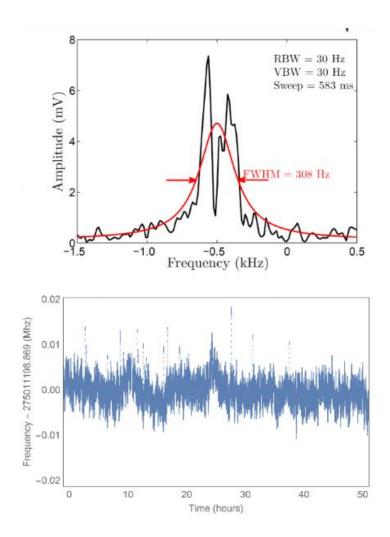




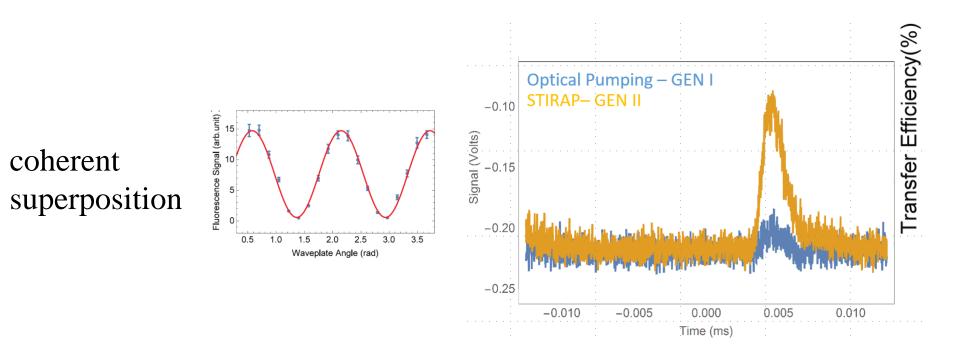
### **Stable Narrow Lasers for STIRAP**







### **STIRAP Excitation – 12 Times Increase in Signal**



## **Usable Molecular Flux Improvement Factors**

What did not work so far: electrostatic focusing  $\rightarrow$  made x-rays

#### What is working so far:

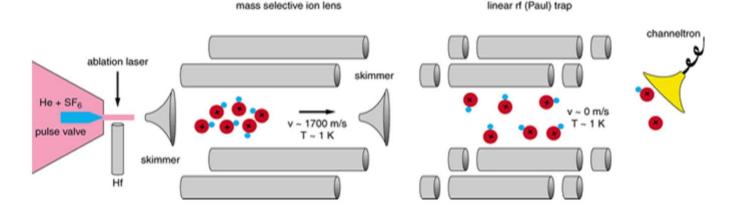
STIRAP	12
Light pipes rather than optical fibers	2.5
Higher quantum efficiency for 512 nm rather than 690 nm:	2
Improved solid angle	8
Total	~ 500
Statistical precision improvement:	~ 20

What may be close: thermochemical source rather than ablation

## Other Electron EDM Measurement Aspirations to Probe Below Our 10<sup>-28</sup> e cm

• Imperial College: YbF molecules (Hinds) 1 x 10<sup>-27</sup> e cm

• JILA: Trapped HfH<sup>+</sup>, HfH<sup>+</sup>, PtH<sup>+</sup> molecular ions (Cornell, Ye)



- Penn. State: extremely cold Cs and Rb atoms (Weiss)
- U. Texas Austin: cold trapped Cs (Heinzen)

Seal(child)

## **Other Important EDM Measurements**

#### Neutron EDM – earlier talk this session

#### Mercury EDM -- recent progress

#### arXiv.org > physics > arXiv:1601.04339

Physics > Atomic Physics

#### Reduced Limit on the Permanent Electric Dipole Moment of <sup>199</sup> Hg

B. Graner, Y. Chen, E. G. Lindahl, B. R. Heckel

(Submitted on 17 Jan 2016 (v1), last revised 13 Apr 2016 (this version, v3))

This paper describes the results of the most recent measurement of the permanent electric dipole moment (EDM) of neutral <sup>199</sup> Hg atoms. Fused silica vapor cells containing enriched <sup>199</sup> Hg are arranged in a stack in a common magnetic field. Optical pumping is used to spin-polarize the atoms orthogonal to the applied magnetic field, and the Faraday rotation of near-resonant light is observed to determine an electric-field-induced perturbation to the Larmor precession frequency. Our results for this frequency shift are consistent with zero; we find the corresponding <sup>199</sup> Hg EDM

 $d_{Hg} = (-2.20 \pm 2.75_{stat} \pm 1.48_{syst}) \times 10^{-30} e \cdot \text{cm}$ . We use this result to place a new upper limit on the <sup>199</sup> Hg EDM  $|d_{Hg}| < 7.4 \times 10^{-30} e \cdot \text{cm}$  (95\% C.L.), improving our previous limit by a factor of 4. We also discuss the implications of this result for various CP-violating observables as they relate to theories of physics beyond the standard model.

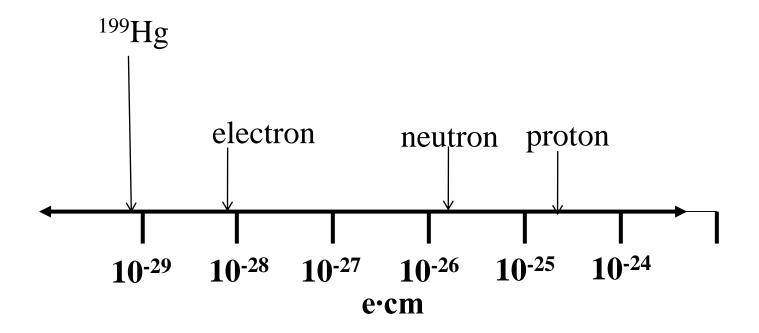
## **Sensitivity to Many CP Violating Observables illustrated in Recent Hg EDM Measurement**

arXiv.org > physics > arXiv:1601.04339					Search of An
	i i i i i i i i i i i i i i i i i i i				
Physics > Atomic Physics					
Reduced Limit on the Permane	nt Electri	c Dipole	Momen	t of $^{199}$ Hg	
B. Graner, Y. Chen, E. G. Lindahl, B. R. Heckel					

TABLE III. Limits on CP-violating observables from the <sup>199</sup>Hg EDM limit. Each limit is based on the assumption that it is the sole contribution to the atomic EDM. In principle, the result for  $\mathbf{d}_n$  supercedes [11] as the best neutron EDM limit.

Quantity	Expression	Limit	Ref.
$\mathbf{d}_n$	$\mathbf{S}_{Hg}/(1.9~\mathrm{fm}^2)$	$1.6 \times 10^{-26} \ e \cdot \mathrm{cm}$	[21]
$\mathbf{d}_p$	$1.3 \times \mathbf{S}_{Hg} / (0.2 \text{ fm}^2)$	$2.0 \times 10^{-25} e \cdot \mathrm{cm}$	[21]
$ar{g}_0$	${f S}_{Hg}/(0.135~e\cdot{ m fm}^3)$	$2.3 \times 10^{-12}$	[5]
$ar{g}_1$	$\mathbf{S}_{Hg}/(0.27 \ e \cdot \mathrm{fm}^3)$	$1.1 \times 10^{-12}$	[5]
$ar{g}_2$	$\mathbf{S}_{Hg}/(0.27 \ e \cdot \mathrm{fm}^3)$	$1.1 \times 10^{-12}$	[5]
$ar{ heta}_{QCD}$	$ar{g}_0/0.0155$	$1.5 \times 10^{-10}$	[22, 23]
$(\widetilde{d}_u - \widetilde{d}_d)$	$\bar{g}_1/(2 \times 10^{14} \mathrm{cm}^{-1})$	$5.7 \times 10^{-27} \text{ cm}$	[25]
$C_S$	$\mathbf{d}_{Hg}/(5.9 \times 10^{-22} \ e \cdot \mathrm{cm})$	$1.3 \times 10^{-8}$	[15]
$C_P$	$\mathbf{d}_{Hg}/(6.0 \times 10^{-23} \ e \cdot \mathrm{cm})$	$1.2 \times 10^{-7}$	[15]
$C_T$	$\mathbf{d}_{Hg}/(4.89 \times 10^{-20} \ e \cdot \mathrm{cm})$	$1.5 \times 10^{-10}$	see text

### **Still No Particle EDM Has Yet Been Detected**



also there are limits on other parameter

## **Other EDM Experiments are Also Important**

Other electron EDM measurements

- Check ACME result
- Different systematics
- If nonzero, atoms are more calculable
- Isotopes offer the chance to check and perhaps cancel systematics and and structure dependence

Neutron and Nuclear EDM

• Sensitive to other sources of T violation

Proton proposed (method more like ion trap method, next talk)

• Will it be possible to get needed sensitivity?

### **Summary**

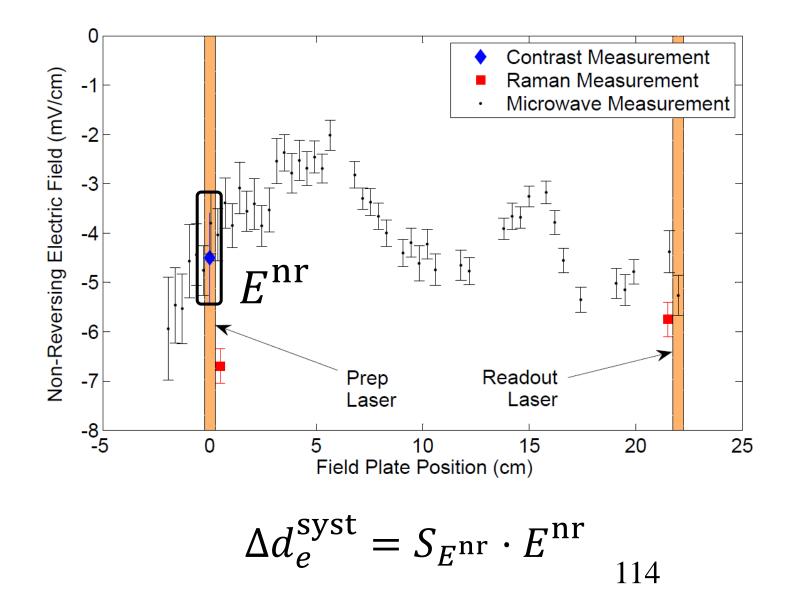
#### **Electron Electric Dipole Moment**

Despite a 12-Fold Improved Measurement
→ No electron edm yet

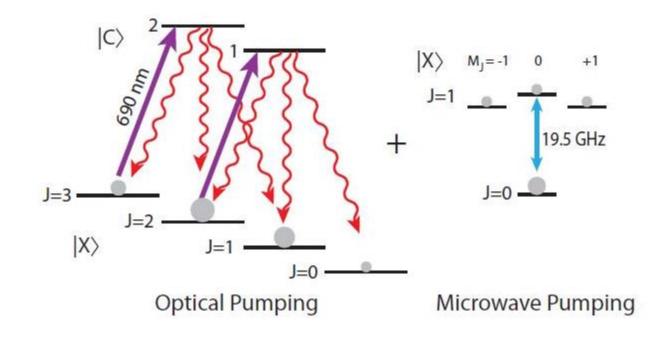
Probing for New Physics at TeV scales and higher
→ comparable or higher than the LHC

Substantial improvement in EDM precision seems possible
 → We are plunging on. > 10x improvement seems very likely

## **Patch Potential** *E*<sup>nr</sup>



## Prepare 17% of Molecules in the J=1 Ground State





# Molecule and photon losses



Parameter	Symbol	Estimate
Beam yield: molecules/(2 quantum states)/pulse Ground state enhancement	N <sub>beam</sub> g	~2 x 10 <sup>11</sup> 1.6
Beam forward velocity	V_f	200 m/s
Beam divergence	$\Omega_b$	0.36 sr
Solid angle of beam detected	Ω <sub>d</sub>	4 x 10 <sup>-5</sup> sr
Beam length before interaction region	$L_0$	130 cm
Beam length in interaction region	L	22 cm
Coherence time = $L/y_f$	t <sub>e</sub>	1.1 ms
H state lifetime	t <sub>H</sub>	$\geq$ 1.8 ms
Surviving <i>H</i> state fraction = $Exp[-t_c/t_H]$ Beam collisional losses (100% = no loss)	f c	0.55 70%
State preparation efficiency	e <sub>p</sub>	5%
Geometric collection efficiency	ee	13%
Quantum efficiency of detector: PMT	eg	10%
Expected photon counts/pulse = $\frac{N_{beam}(\Omega_d/\Omega_b) f c g e_p e_g e_q}{N_{beam}(\Omega_d/\Omega_b) f c g e_p e_g e_q}$	$S_0$	~8 x 10 <sup>3</sup>
D.1		50 1 (

# **Statistical Comparison of ACME and Imperial**

Statistical sensitivity:

$\delta d_e = \frac{1}{25}$	$\frac{\hbar}{\sqrt{1+2}}$	$7 \ge 1.7 \ge 24$		
internal electric field time	time nce counting			
	ACME ThO	<b>Imperial YbF</b>	$\checkmark$	
Effective E field	100 GV/cm	14 GV/cm	7	
Coherence time	1.1 ms	0.65 ms	1.7	
Photons/second*	1000 x 50 =50,000	500 x 25 =12,500	41/2	
Precision in same tin Time for same preci		24 $(24)^2 \sim 600$		

\*Our molecule source is more intense, allowing us to use a metastable state rather than the ground state (as needed in ThO)

### **Berry's Phase (Geometrical Phases)**

Spatial inhomogeneities in the applied electric and magnetic fields, which appear as time-varying fields in the rest frame of molecules in the beam, can give rise to geometric phase-induced systematic effects [32, 33]. We have used the

 $\delta d_{\rm e}({
m sys}) \ll 10^{-32} e \,{
m cm}$ 

<u>"Search for the Electric Dipole Moment of the Electron with Thorium Monoxide"</u>
 A.C. Vutha, W.C. Campbell, Y.V. Gurevich, N.R. Hutzler, M. Parsons,
 D. Patterson, E. Petrik, B. Spaun, J.M. Doyle, G. Gabrielse, and D. DeMille,
 J. Phys. B. At. Mol. Opt. Phys. 43 074007 (2010).

[34] Vutha A and DeMille D 2009 arXiv:0907.5116