

CORE+ Telescope

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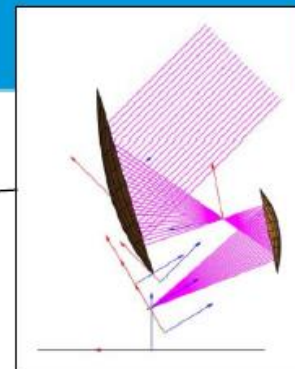
Comments on

- Where we are
- Baffles + Stop
- Focal Surface Shape
 - ARC

Telescope configuration trade-off

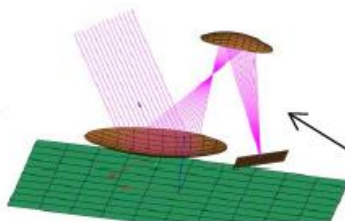
Four configurations considered:

- Gregorian option 1



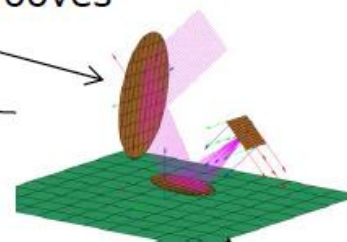
Selected option:
Fits in V-Grooves
and can easily be
mounted

- Gregorian option 2



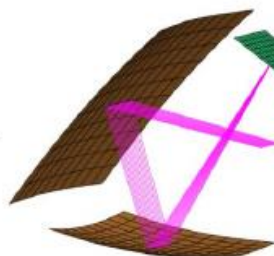
Does not fit
in V-Grooves

- Open Dragone



Focal plane too high.
Complex Thermo-Mechanical
accommodation

- Cross-Dragone
with $F \sim 2$

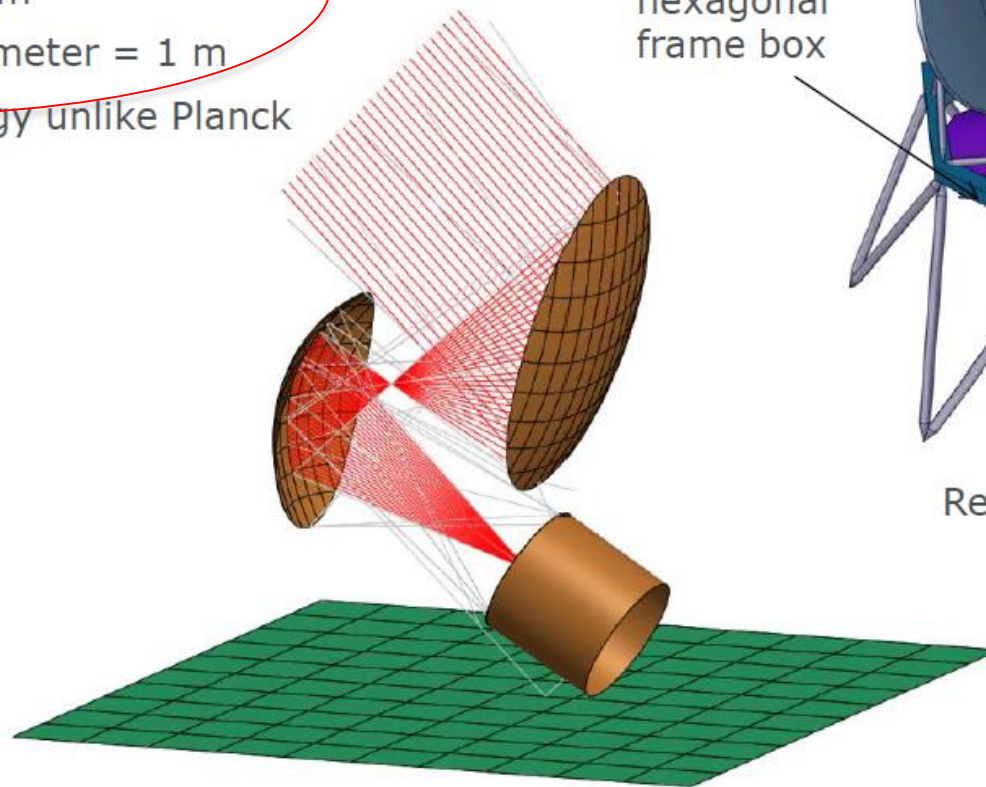
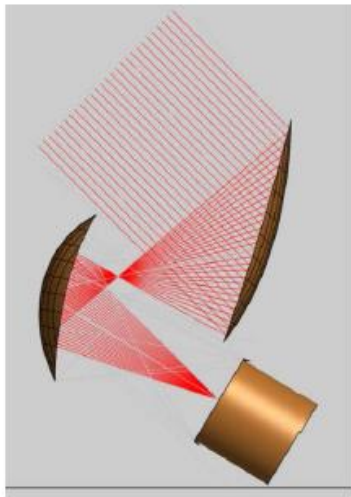


Large secondary mirror

Telescope baseline

Gregorian configuration:

- Aperture = 1.2 m - $F/D \sim 2$.
- Primary Mirror 1.5X1.2 m
- Secondary reflector diameter = 1 m
- Monolithic SiC technology unlike Planck

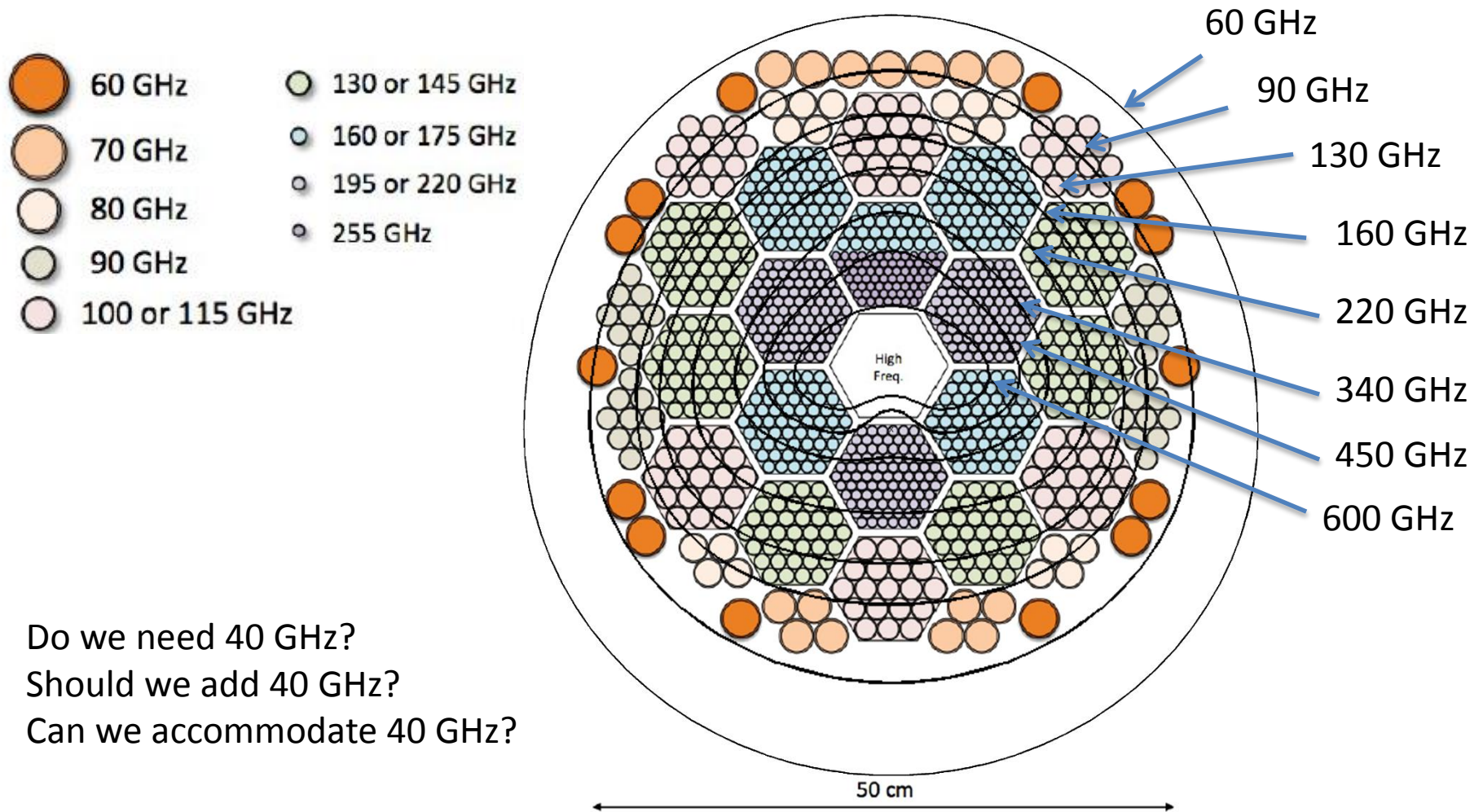


hexagonal
frame box

FPU baffle

Reflectors frame box

Focal plane layout

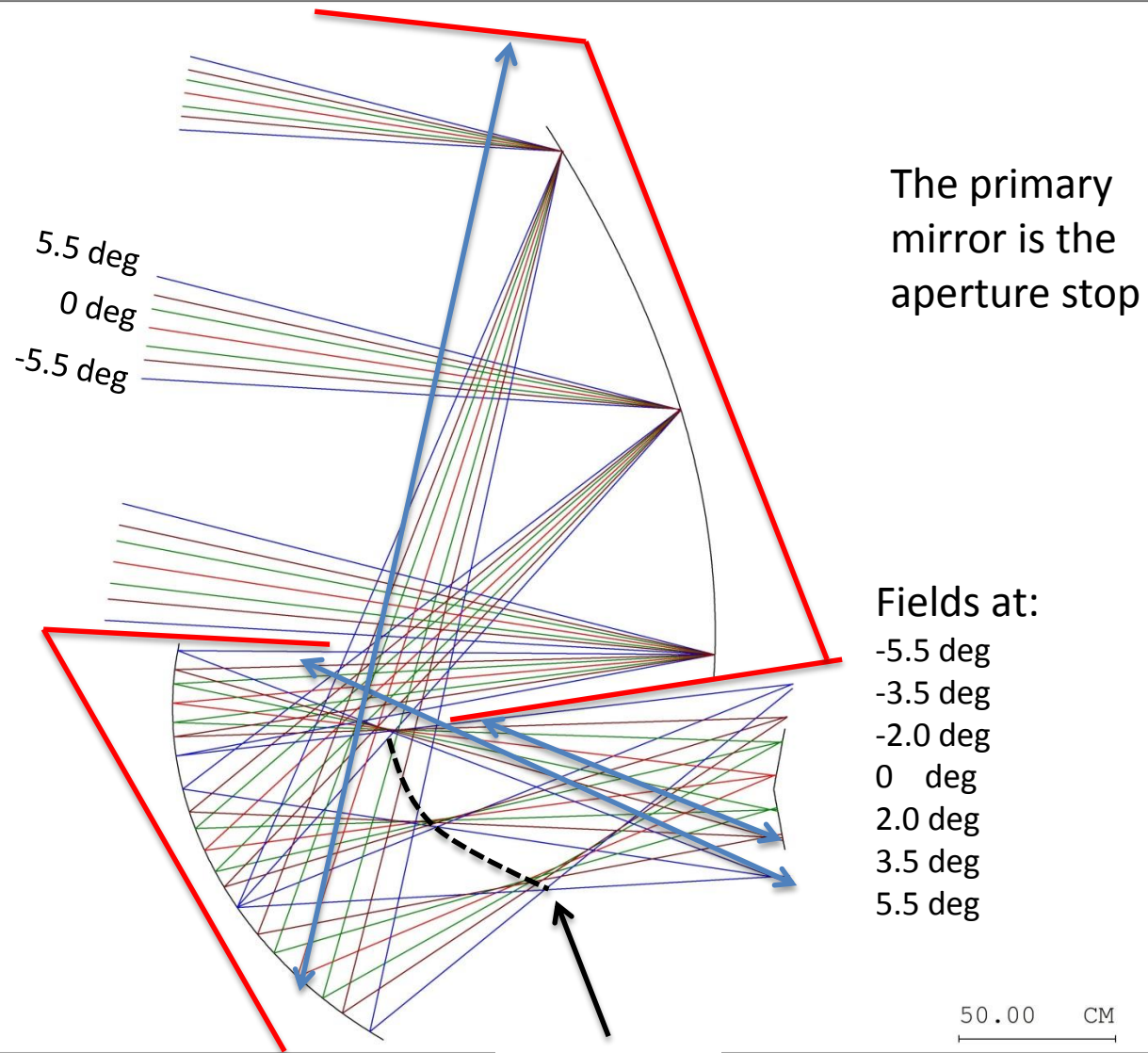


Do we need 40 GHz?
Should we add 40 GHz?
Can we accommodate 40 GHz?

Figure 3: Sketch of the focal plane of CORÉ+ Light. Contours are Strehl= 0.8 for 60, 90, 130, 160, 220, 340, 450, and 600 GHz.

1.5 m Aperture Gregorian, Optimized Dragone, with Aspherics

- Baffles are necessary
- Note compactness of system
- Need GRASP including baffles



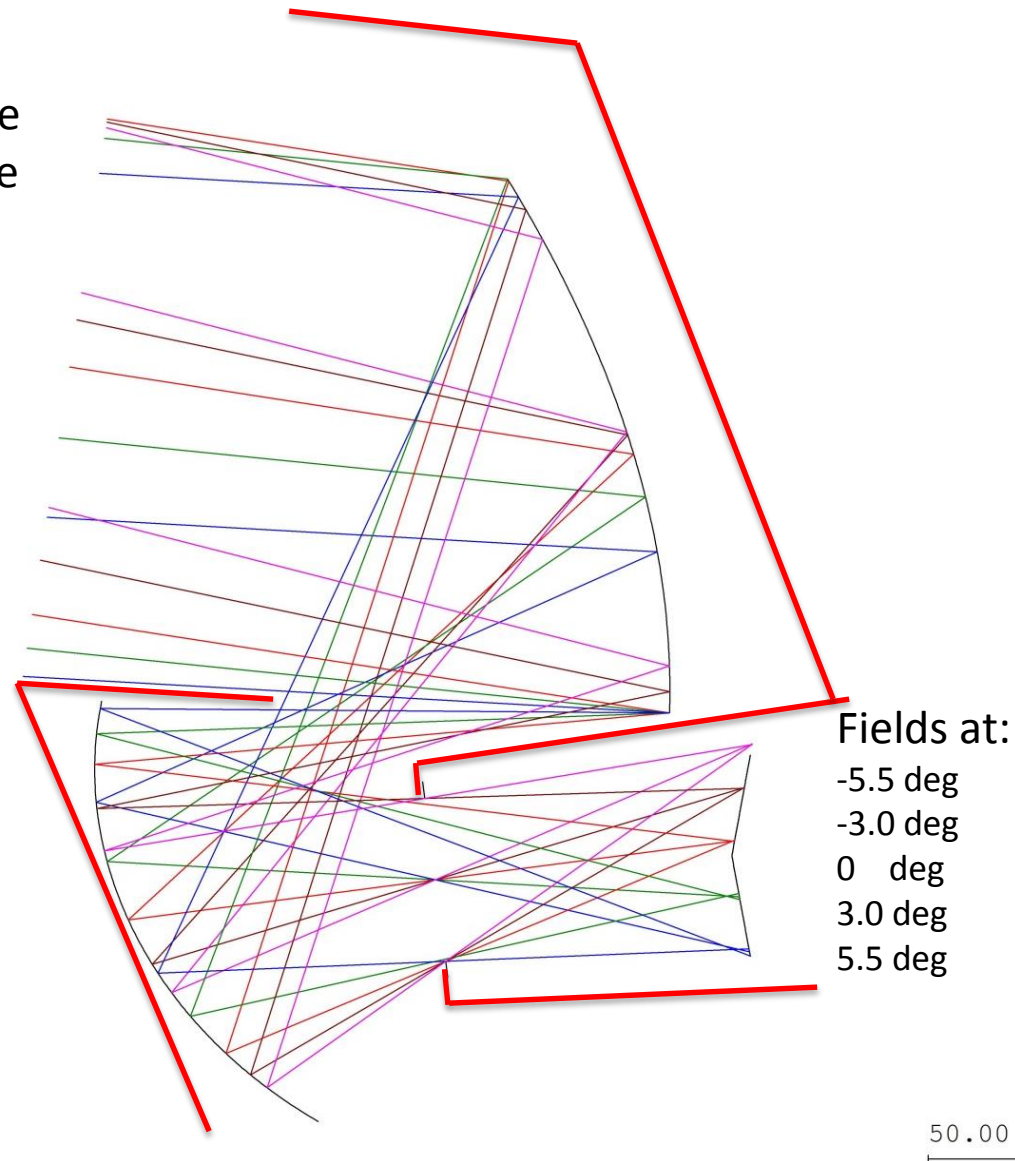
COrE gregorian 1.5meter f/2

Image of
Primary

03-May-16

Cold Stop Enforced (for most of the rays)

Vignetting: Some rays are limited by the stop, some by the primary.



Change in pupil footprint area due to vignetting

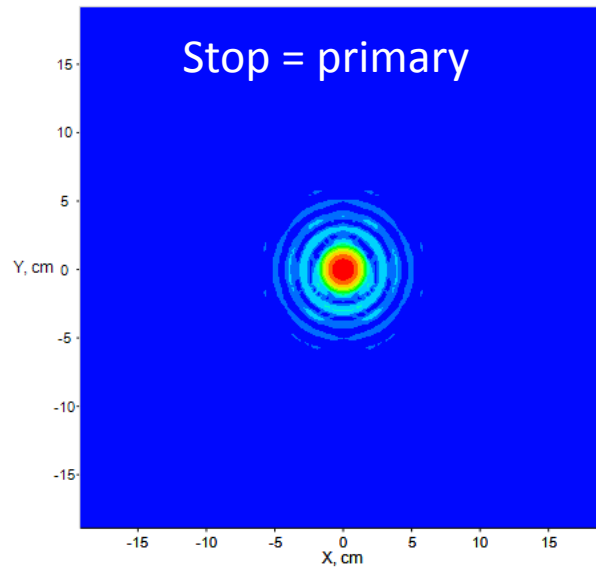
Field Angle (degrees)	Area relative to unvignetted 1.5m pupil
0	1.00
+3 (-3) elevation	1.03; (0.89)
5.5 (-5.5) elevation	1.01; (0.79)
+3.5 in azimuth	0.98
+5.5 in azimuth	0.88

Change in Strehl with Cold Stop; 60 GHz

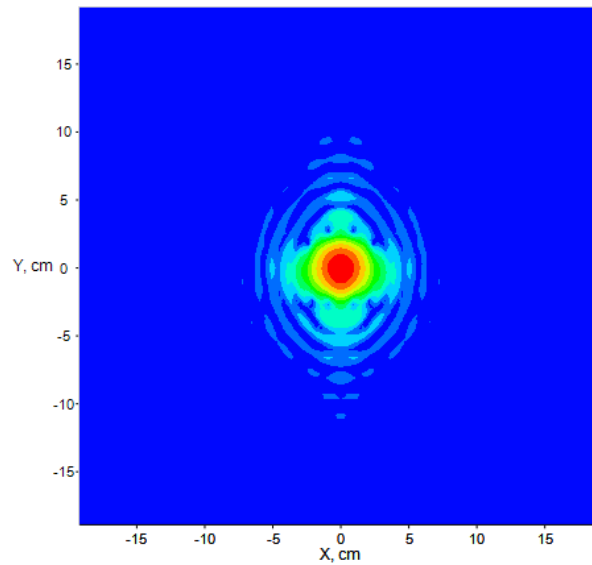
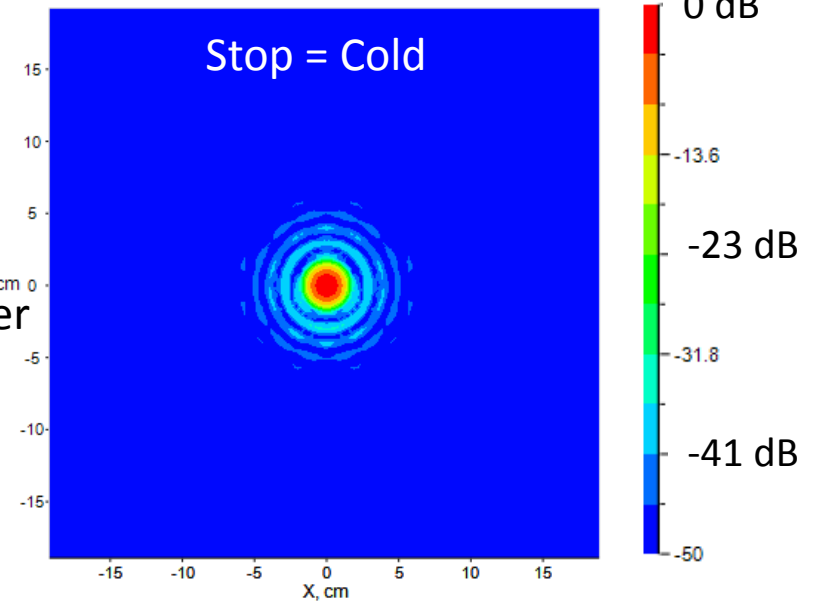
El. Field angle (degrees)	No stop	Stop 90 cm from focal plane
0	0.99	0.99
5.5	0.81	0.84
-5.5	0.84	0.85
Az. Field angle (degrees)		
5.5	0.79	0.825

Strehl is the ratio of peak energy to peak energy of diffraction limited system system. It is a single point measure. It gives no information about beam shape.

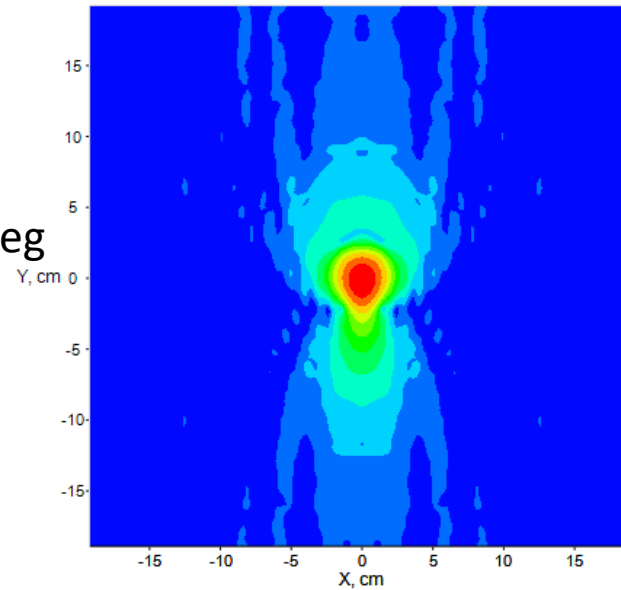
PSFs



60 GHz, Center
Strehl = .99

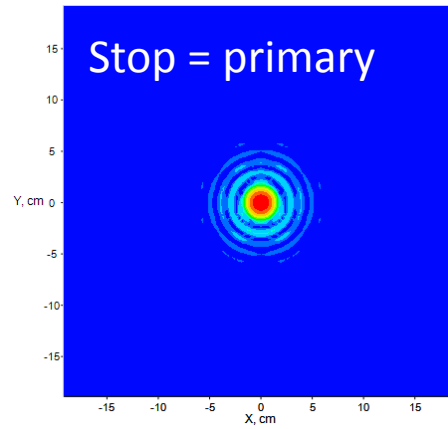


60 GHz, 5.5 deg
elevation

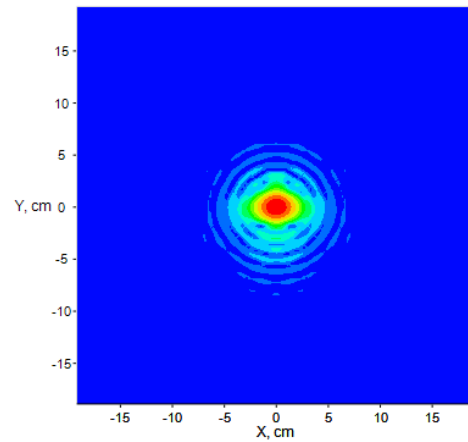
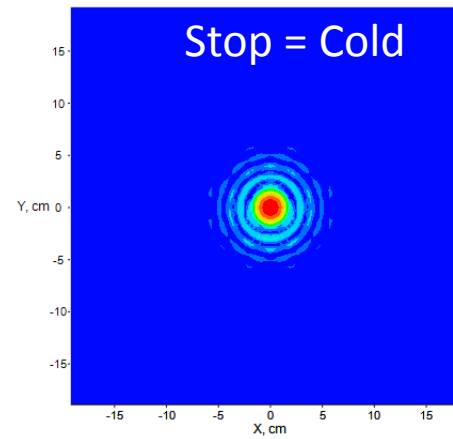


Color scale dB
at focal plane,
normalized to 0.

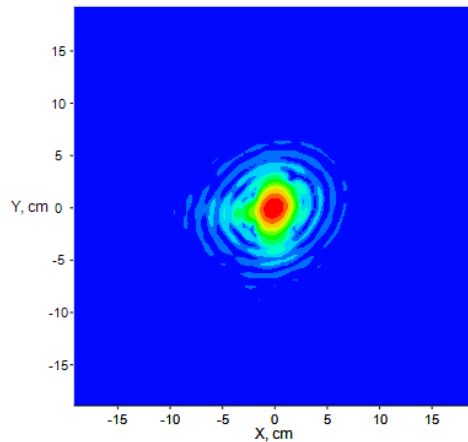
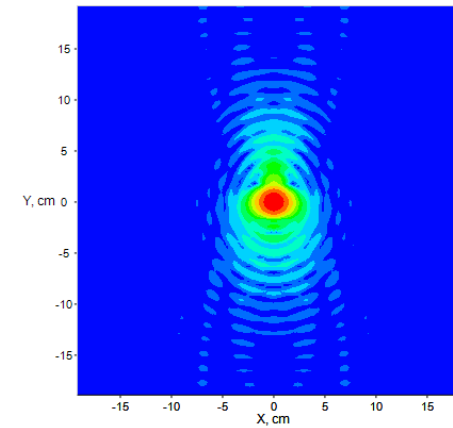
PSFs



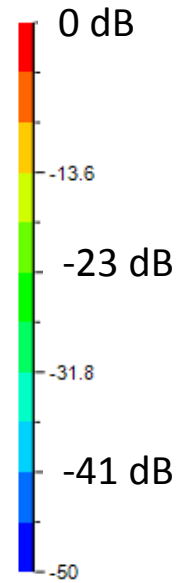
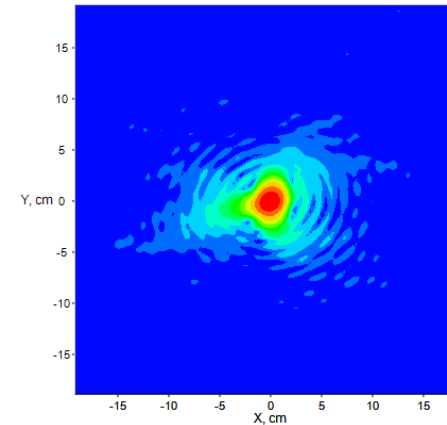
60 GHz, Center



60 GHz, -5.5 deg
elevation

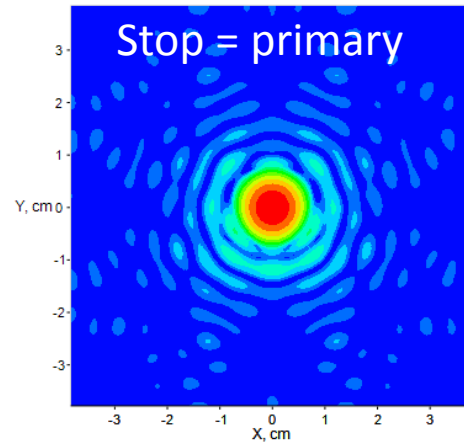


60 GHz, 5.5 deg
azimuth

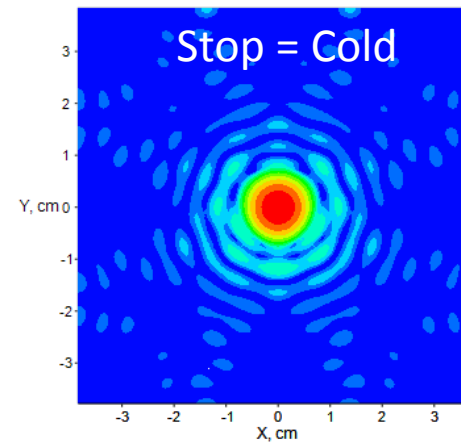


Color scale dB
at focal plane,
normalized to 0.

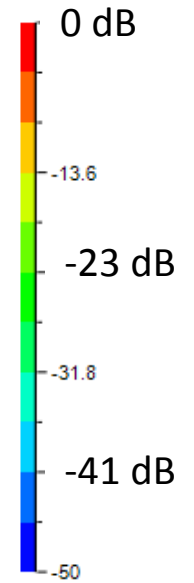
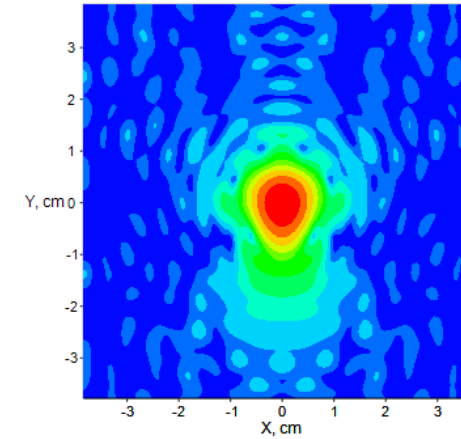
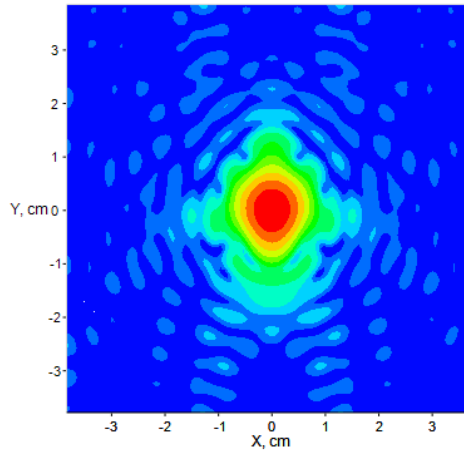
PSFs



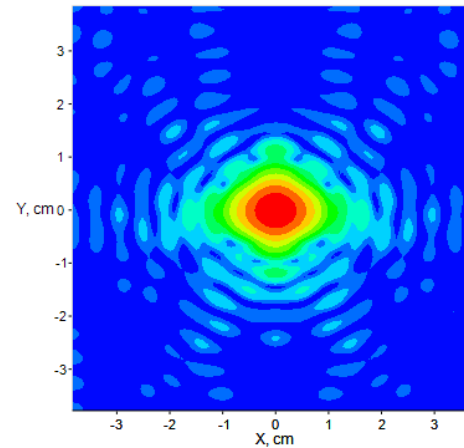
150 GHz, Center



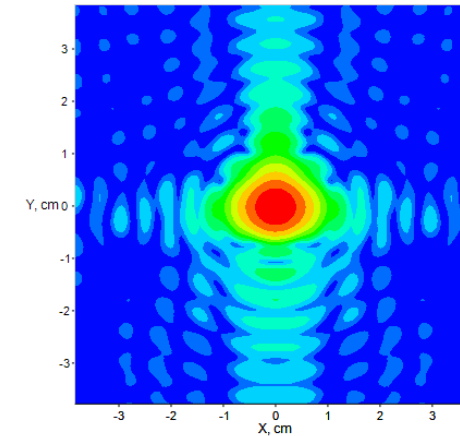
150 GHz, 3 deg
elevation



Color scale dB
at focal plane,
normalized to 0.



150 GHz, -4 deg
elevation

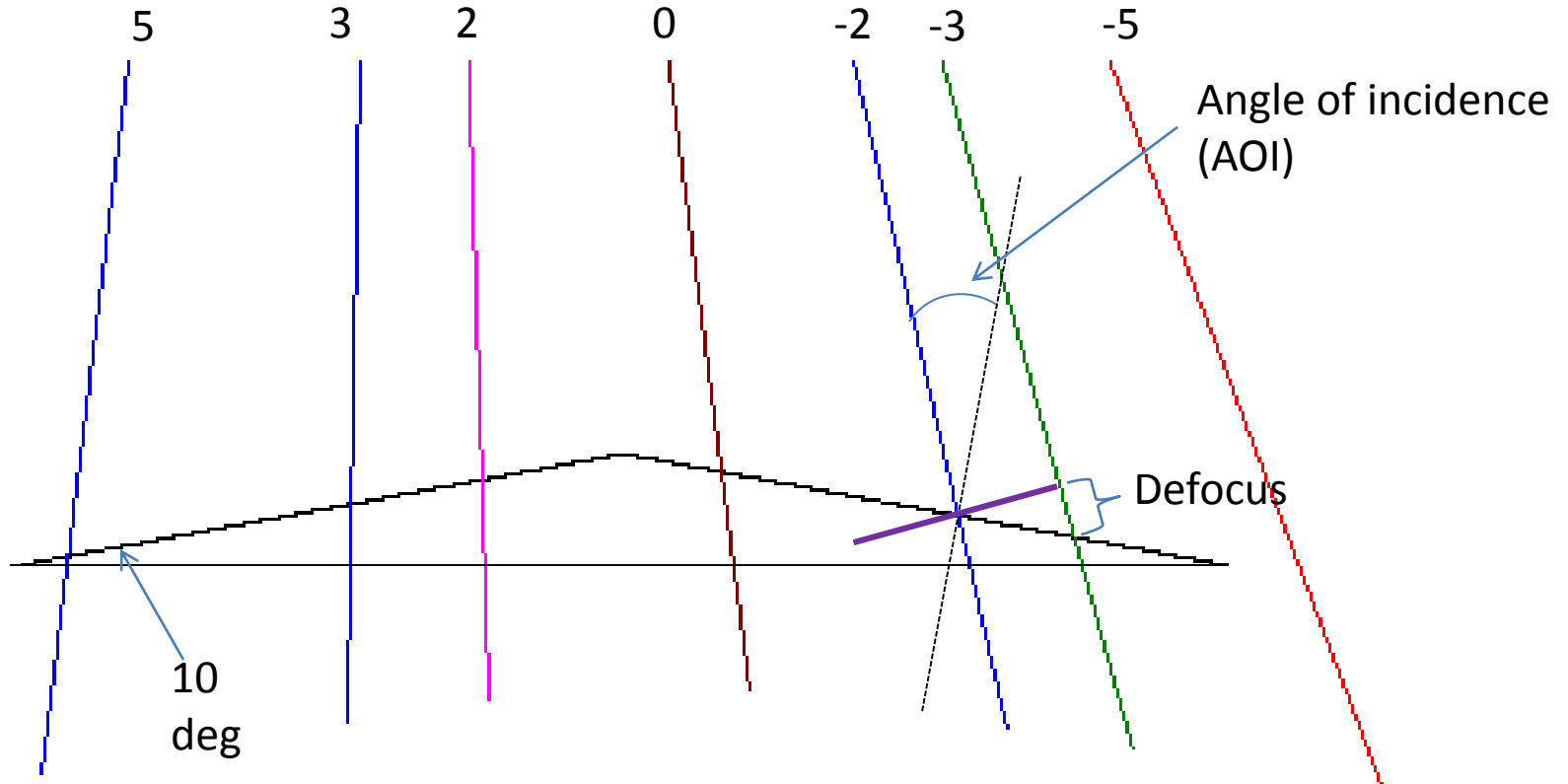


Didn't yet check

- Polarization – should use GRASP
- Higher Frequencies

Conic, Non-telecentric, Focal Surface: What if we tilt the arrays?

Chief rays for 7
field angles.

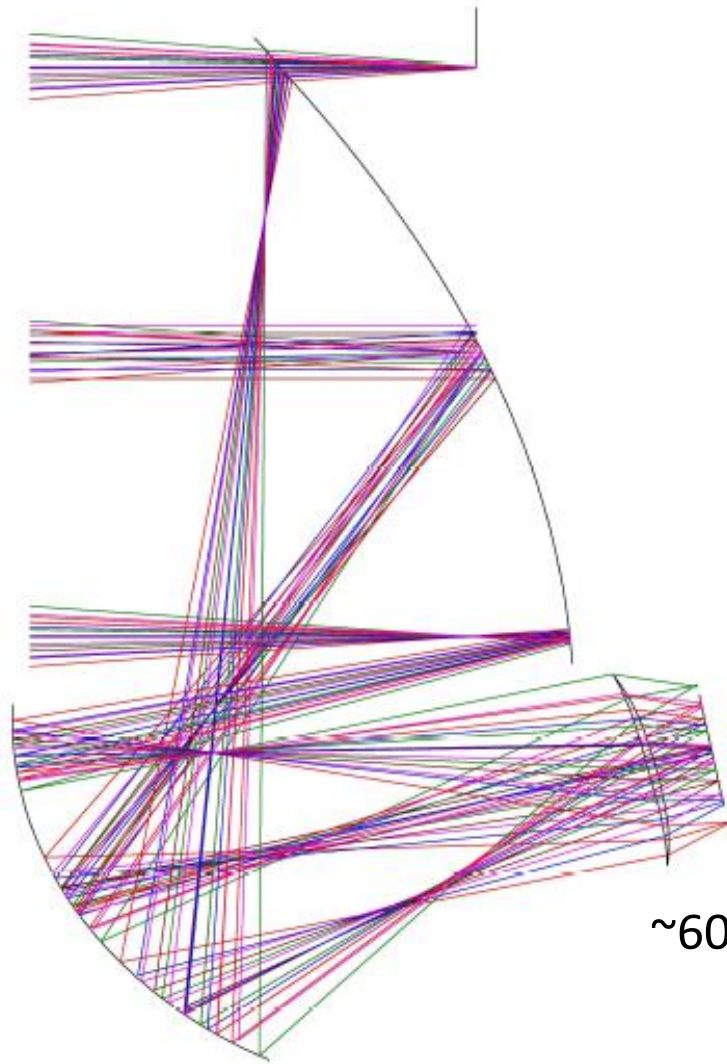


Results

Field (degrees)	Freq. (GHz)	AOI (deg)	Defocus (cm)		Strehl at focus	Strehl				average delta strehl/cm
			4"	3"		4" outer	4" inner	3" outer	3" inner	
-5	60	30.9	3.150	2.347	0.89	0.81	0.72	0.85	0.78	0.03
-3	150	25.9	2.419	1.803	0.93	0.57	0.60	0.70	0.740	0.13
-2	300	23.3	2.031	1.514	0.90	0.22	0.40	0.35	0.62	0.28
0	600	17.9	1.214	0.904	0.90	0.28	0.23	0.45	0.38	0.53
2	300	8.6	0.216	0.161	0.71	0.70	0.71	0.70	0.71	0.025
3	150	11.5	0.231	0.172	0.81	0.80	0.81	0.81	0.81	0.01
5	60	17.39	1.136	0.846	0.86	0.88	0.83	0.87	0.84	0.02

Need to steer the beams

Or – Use a Lens?



Alumina lens ~60 cm diameter

$n=3.1$

Flat focal plane

Fields shown = ± 5 deg

Telecentric within 10 deg

$F\# = \sim 1.8$

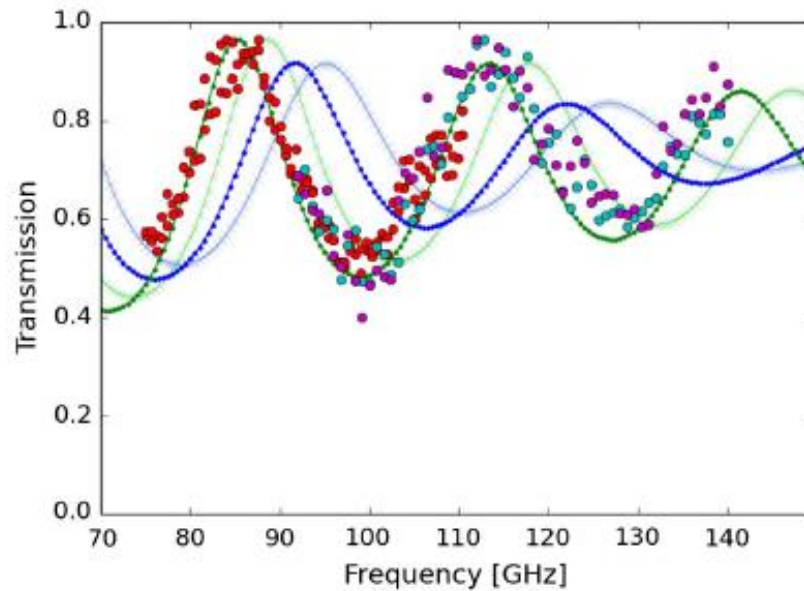
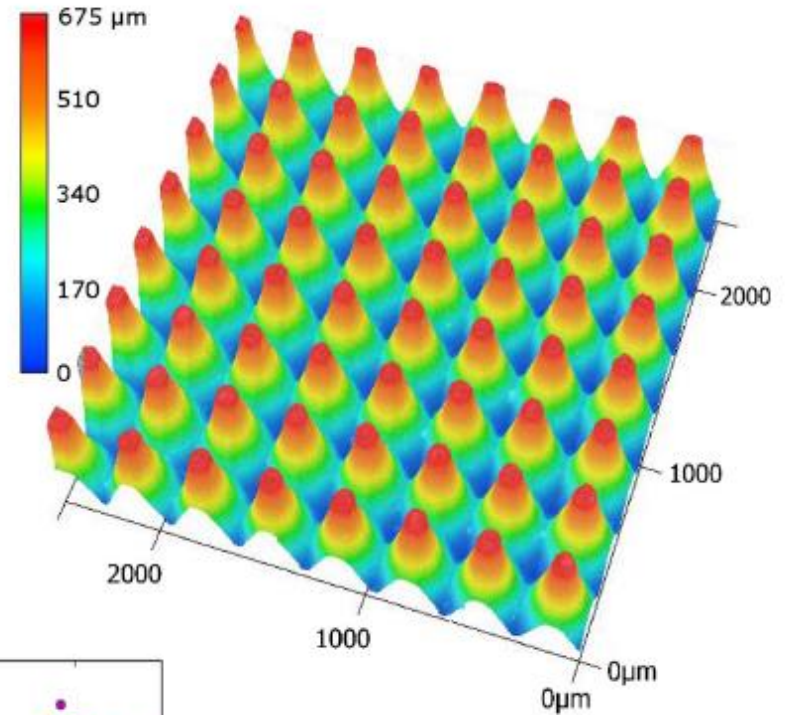
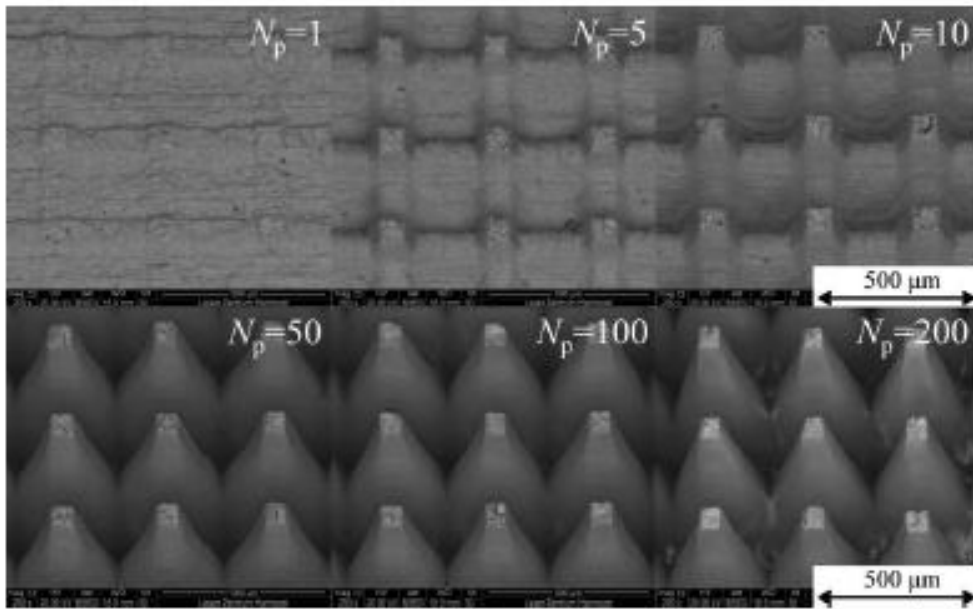
Strehl ratios similar to $F\#=2$

(requires more detailed study)

~60 cm diameter

46.30 CM

Broadband ARC – Laser Ablation



Matsumura et al. 2016

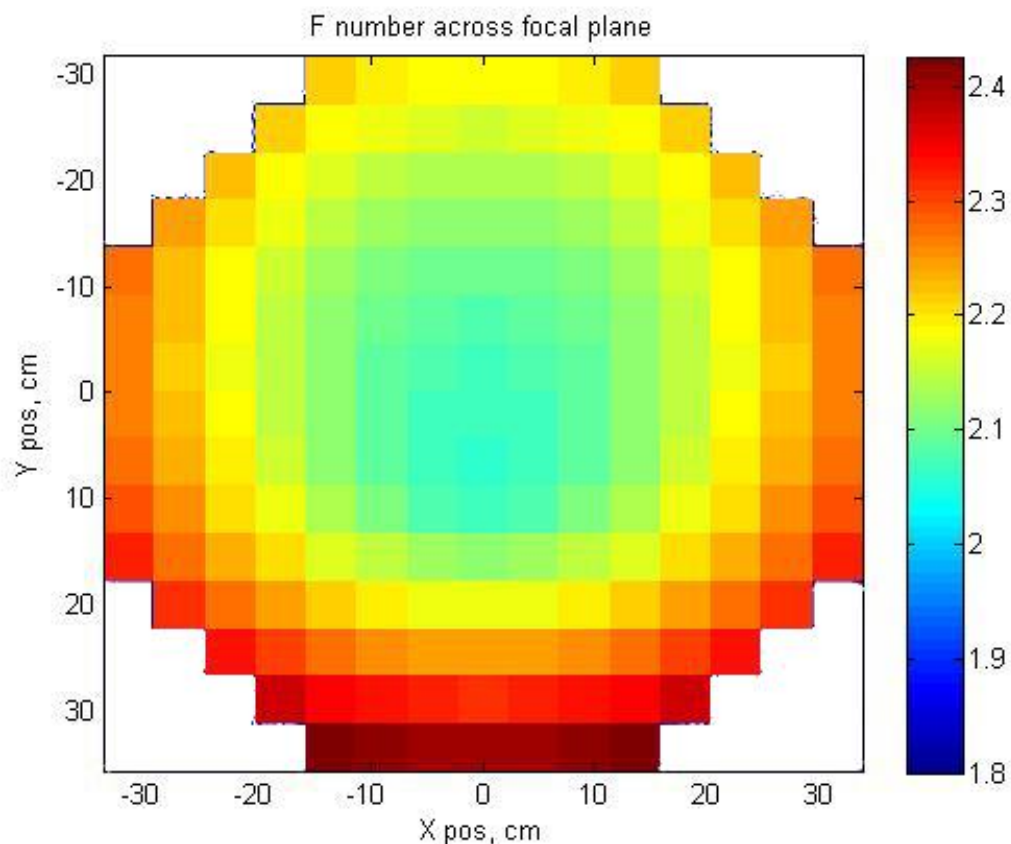
Summary

- So far: 1.5 m; Should we look at 1.2 m?
- Low T baffles/stop
 - Baffles: OK
 - Stop: questionable
 - Do more detailed GRASP for polarization and far sidelobes
- Focal Plane:
 - Steer the beam
 - Use lens?

Backup Slides

To Do

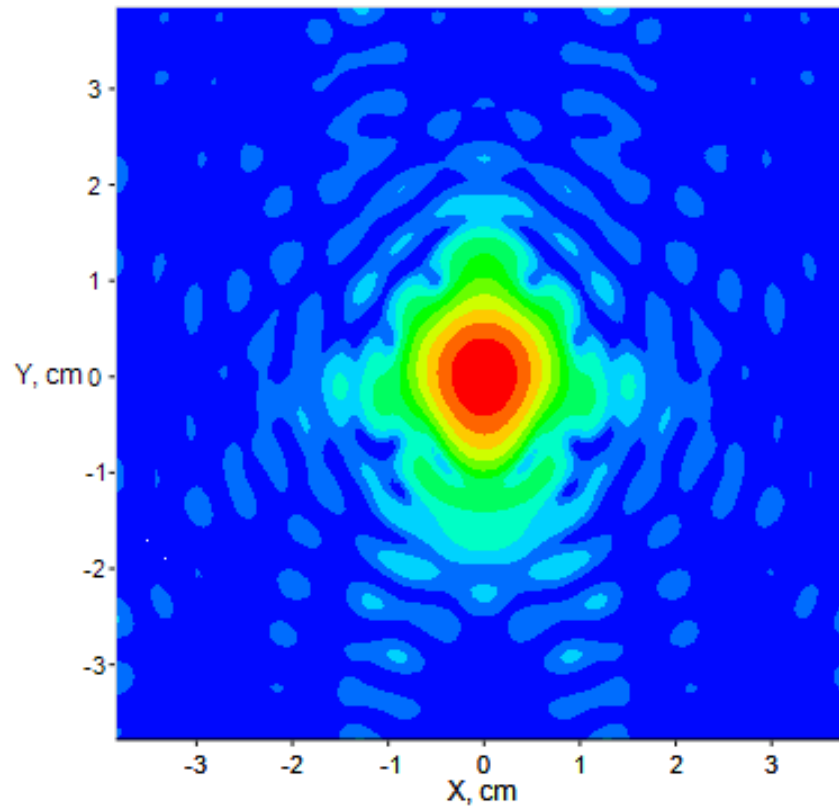
F-number across focal plane



For comparison
EBEX2013 had $f/\#$
from 1.86 – 2.04

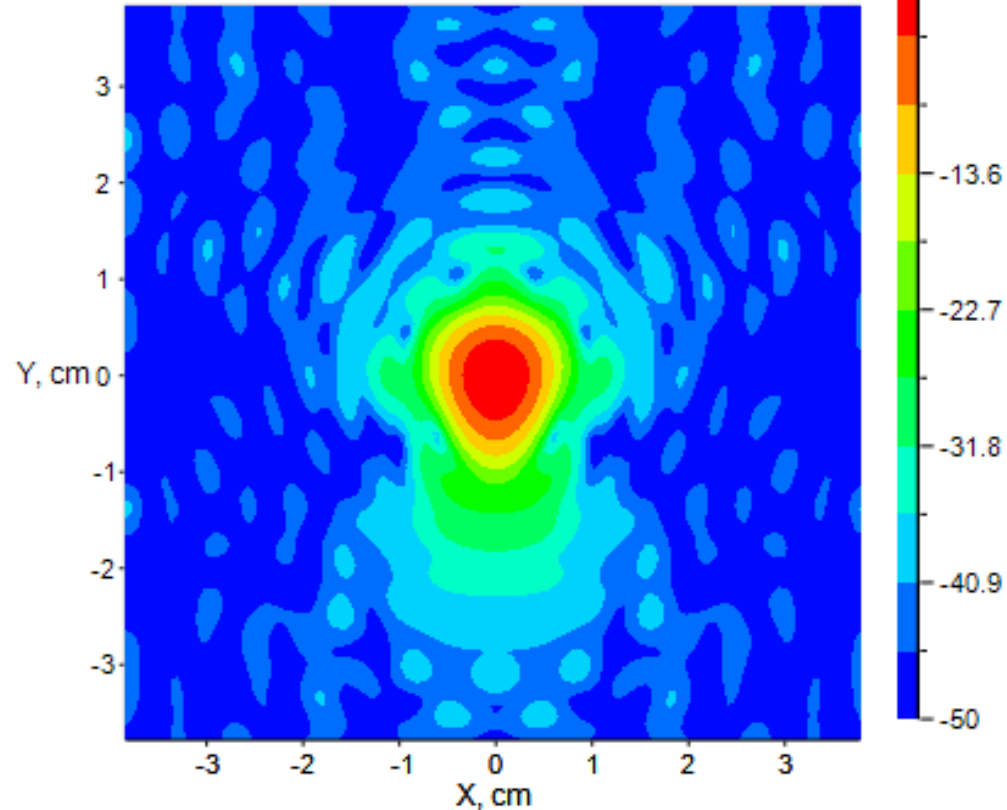
PSFs, at 150 GHz, 3 degrees elevation

Stop = primary



Strehl = 0.81

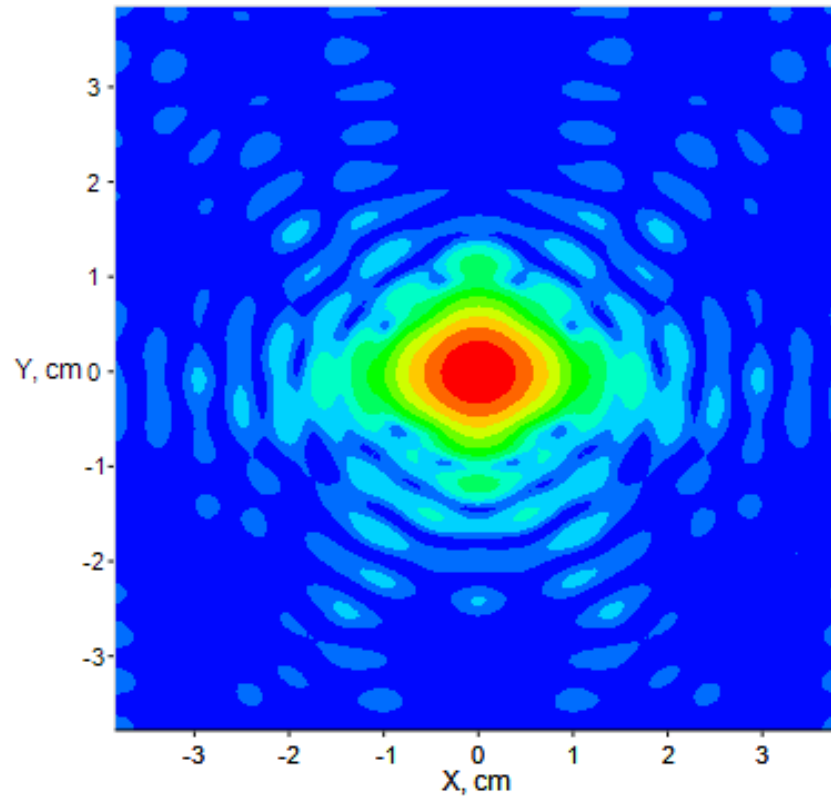
Stop = Cold



Strehl = 0.83

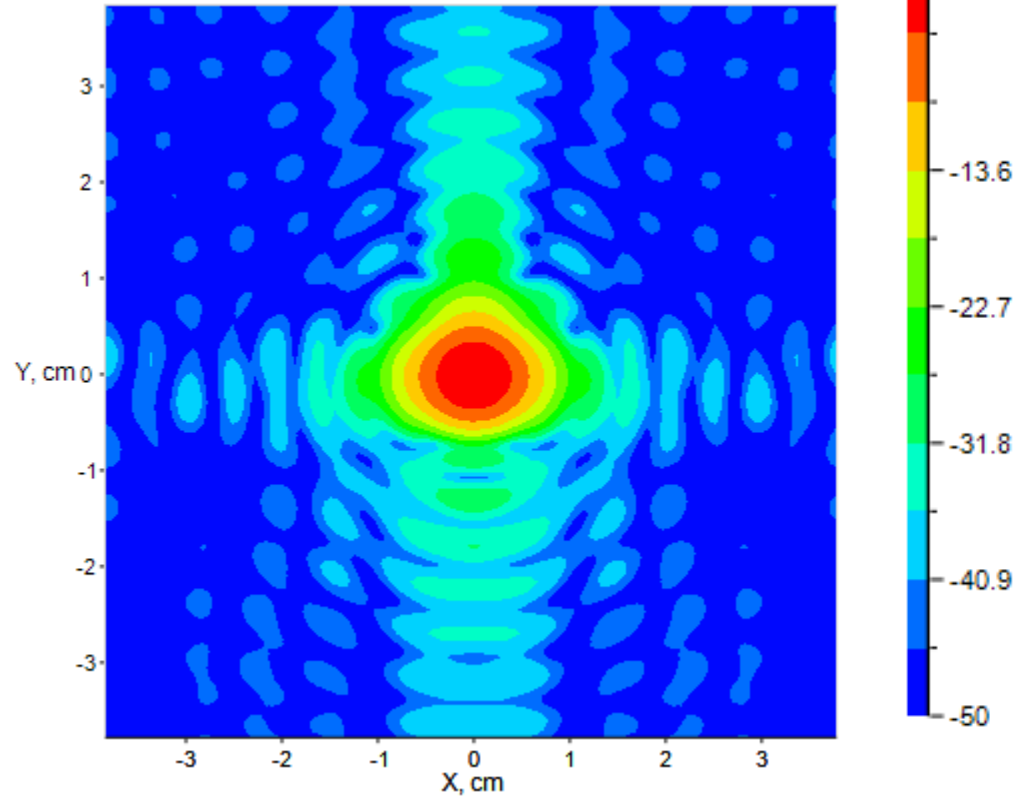
PSFs, at 150 GHz, -4 degrees elevation

Stop = primary



Strehl = 0.80

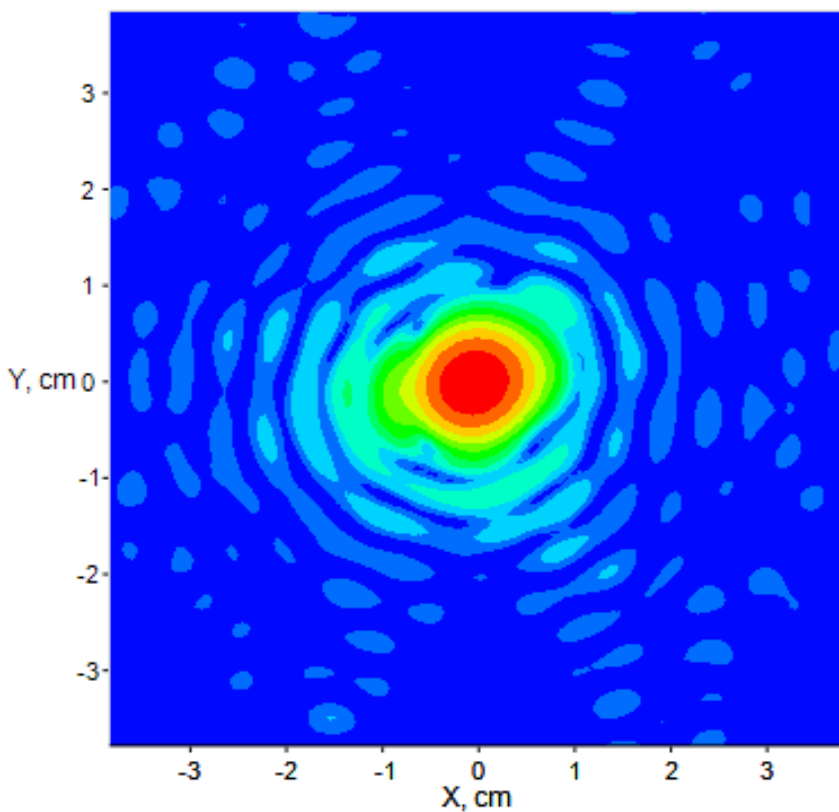
Stop = Cold



Strehl = 0.81

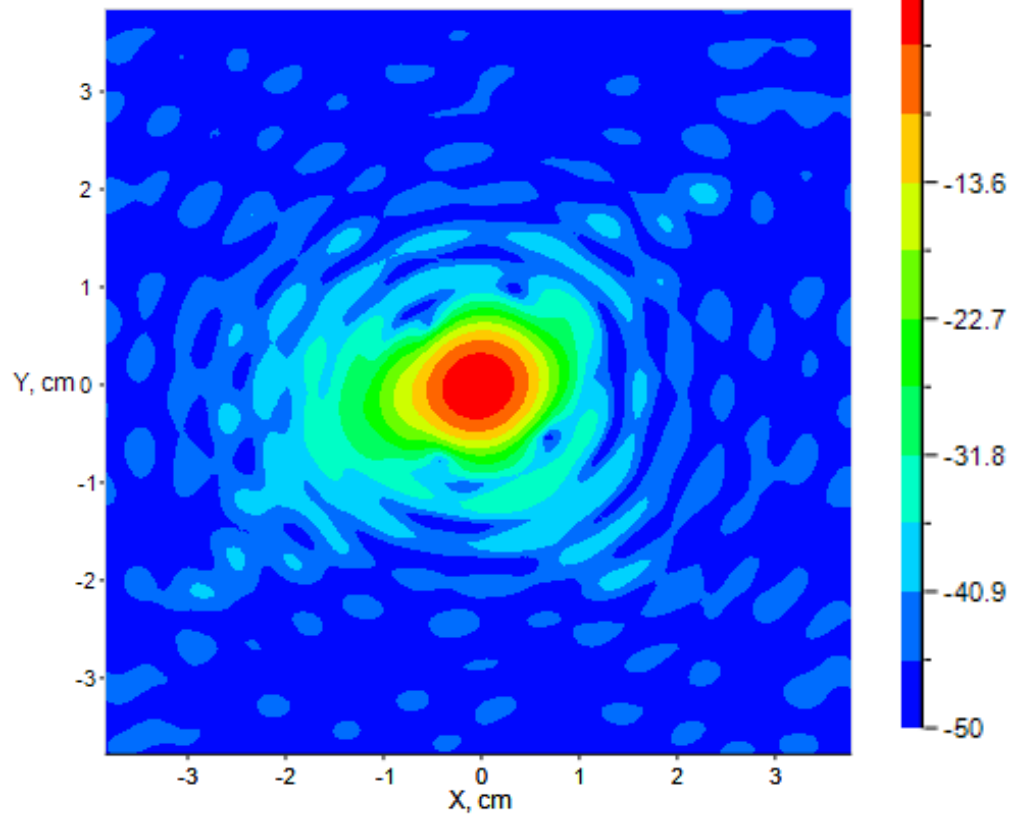
PSFs, at 150 GHz, 3 degrees azimuth

Stop = primary



Strehl = 0.92

Stop = Cold

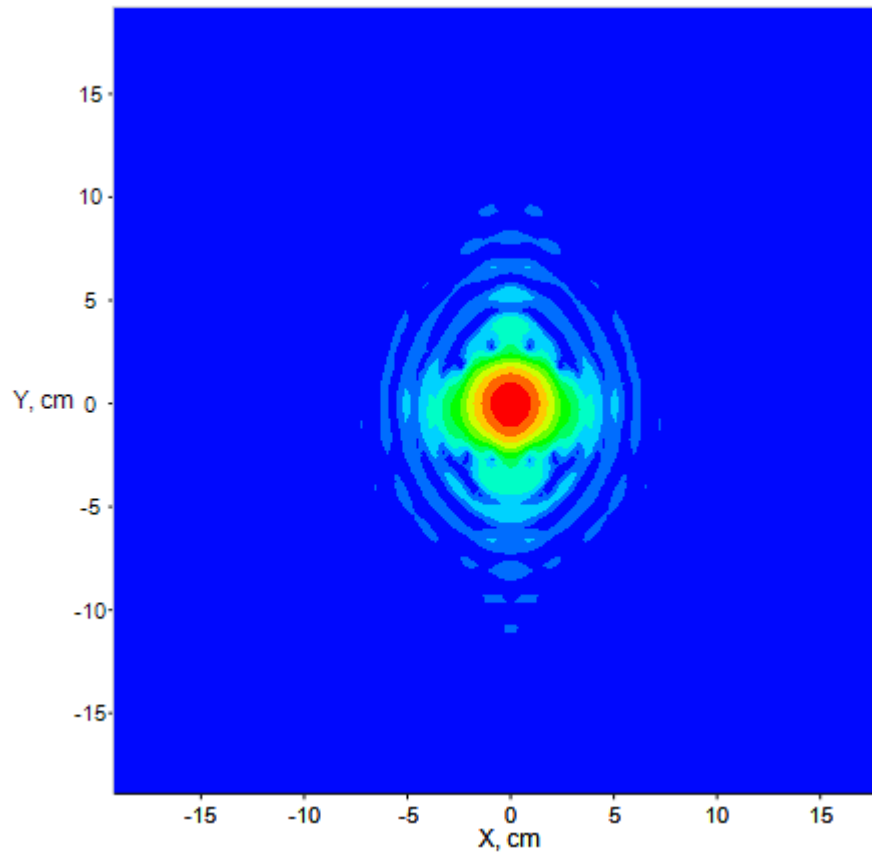


Strehl = 0.91

PSFs, at 60 GHz, 5.5 degrees elevation

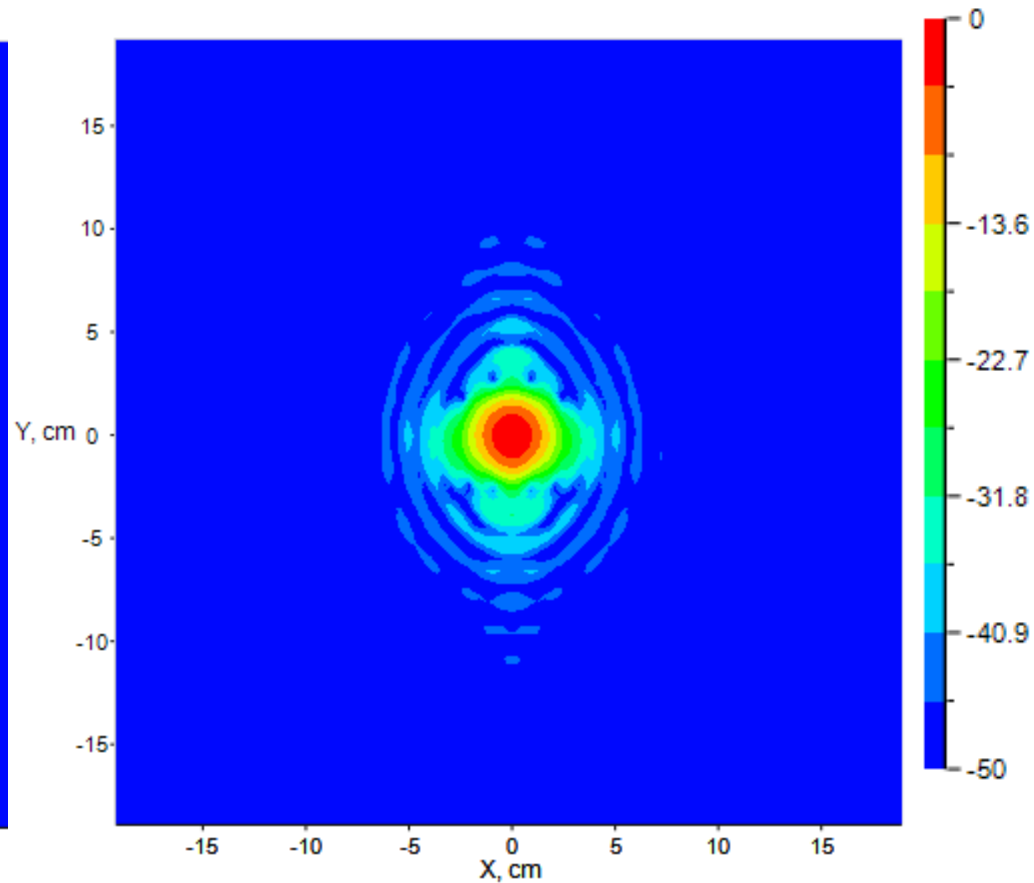
Polarized input, primary is stop.

Vertical polarization,
 $E \parallel$ to Y axis



Strehl = 0.813

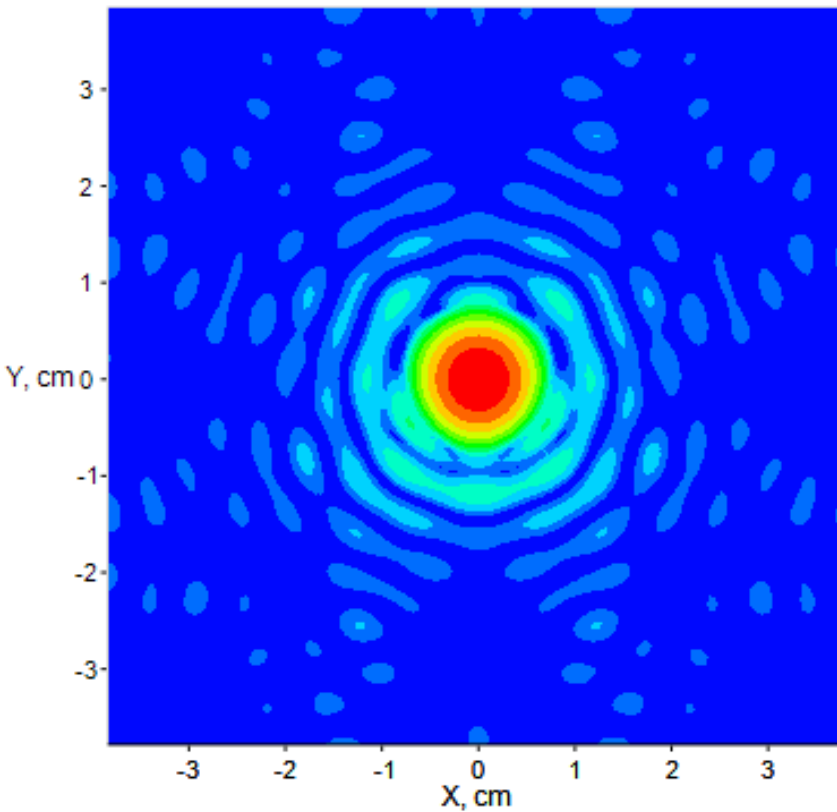
Horizontal polarization



Strehl = 0.812

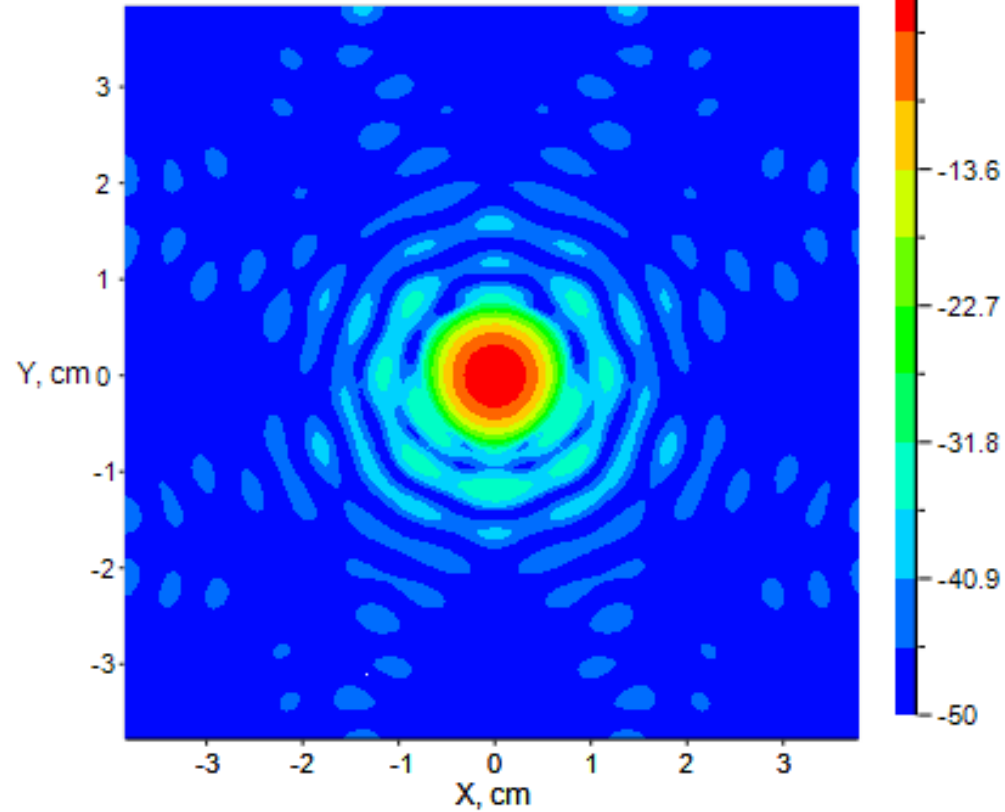
PSFs at 150 GHz, 0 degrees

Stop = primary



Strehl = .99 for both

Stop = Cold



Color scale dB at focal plane,
normalized to 0.