Kinetic freeze out from a local anisotropic fluid

Hot Quarks 2016

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Content

- Reminder on Cooper-Frye freeze out
- Motivation for local anisotropy
- Effects of the local anisotropy on:
 - 1. transverse particle spectra
 - 2. flow coefficients
 - 3. HBT radii (time permitting)
- Conclusion

Need for kinetic freeze out

- Inside QGP fluid-like behavior with small $l_{mfp} \rightarrow$ hydrodynamics works
- Expanding fireball of conserved global quantities
- Particles reaching the detector do not interact with each other, large/infinite $l_{mfp} \rightarrow$ kinetic theory works

 $\leftarrow \text{ kinetic freeze out here} \rightarrow \text{ kinetic theory} \\ \rightarrow \text{ kinetic theory} \\ \rightarrow \text{ Hadrons} \\ \rightarrow \text{ Hydro} \\ \hline QGP \\ \rightarrow \text{ ideal Hydro} \\ \hline QGP \\ \rightarrow \text{ aniso. / dissipative Hydro} \\ \hline r \\ \hline early \text{ stages} \\ \rightarrow \text{ pre-equilibrium} \\ \hline \end{array}$

Need to glue two very different types of theoretical models

Cooper-Frye freeze out

- Description of the medium turns suddenly from fluid to a kinetic one, while passing through a hypersurface Σ
- On-shell phase space distribution $f(x^{\mu}, \vec{p})$
- Particles inside a fluid element: $dN = dV \int_{-\infty}^{\infty} d^3 \vec{p} f(E_p)$
- Lorentz invariant particle spectrum:

$$E_p \frac{d^3 N}{d^3 \vec{p}} = \frac{1}{(2\pi)^3} \int_{\Sigma} f(x^{\mu}, \vec{p}) p^{\mu} d^3 \sigma_{\mu}(x^{\mu})$$

Problems of (naive) Cooper-Frye freeze out

- Negative contributions to the particle spectrum due to Σ regions where $d\sigma_{\mu}d\sigma^{\mu} < 0$ (see talk by Oliinychenko)
- Computed observables depend on choice of freeze out parameter

Not physical, Nature performs a "smooth" transition between the two asymptotic models

But Cooper-Frye recipe is much too attractive to be discarded

$$E_p \frac{d^3 N}{d^3 \vec{p}} = \frac{1}{(2\pi)^3} \int_{\Sigma} f(x^{\mu}, \vec{p}) p^{\mu} d^3 \sigma_{\mu}(x^{\mu})$$

Motivation for an anisotropic distribution function

- Nonrelativistic studies show:
 - Hypersonic flow through a nozzle into vacuum will end in an anisotropic distribution function
 - Characterized by two effective temperatures T_{\parallel} and T_{\perp}
- Anisotropic hydrodynamics can improve transition from pre-equilibrium to hydrodynamics (see talk by Bazow)
- Generalization to relativistic HIC case ...

Anisotropic free $Z(\mathbf{B}, \mathbf{p}) = \int_{\mathbf{Q}} \int_{\mathbf{T}} f(\tau, \mathbf{x}) + \delta f(\mathbf{x}, \mathbf{y}) d\mathbf{x}$

- <u>Idea</u>: implement an anisotropy along the radial direction in the phase space distribution $f(\tau, \mathbf{x}, \mathbf{p}) = f_{aniso}(\mathbf{p}, \Lambda(\tau, \mathbf{x}), \boldsymbol{\xi}(\tau, \mathbf{x}))$
- Tool: Romatschke-Strickland distribution

$$f_{aniso}^{LRF}(x^{\mu}, \vec{p}, \Lambda, \xi) \propto exp\left(-\frac{\sqrt{(p \cdot u)^2 f_{\text{miss}}^{LRF}(\vec{x}^{\mu}) \left(\frac{\sqrt{p^2 + \xi(\mathbf{x}, \tau)p_z^2}}{\Lambda(x^{\mu})}\right)}{\Lambda(x^{\mu})}\right)^{\xi = \frac{\langle p_T^2 \rangle}{2\langle p_\tau^2 \rangle} - 1}$$

- Boost into lab-frame
- Consequence: additional parameter ξ
- In order to generate higher effective temperature/pressure in radial direction, need to choose $\xi < 0$
- <u>Next step</u>: insert into Cooper-Frye Integral

Taken from: Michael Strickland's Quark Matter 2015 Talk

ε

 $-1 < \xi < 0$



Technical details

- Cylindrical lab-frame coordinates
- Blast-wave like fluid velocity profile:

$$u^{\phi} = u^{\eta} = 0 \quad ; u^{r} = \bar{u}_{max} \frac{r}{R} \left(1 + 2\sum_{n} V_{n} \cos(n\phi) \right)$$

•
$$R = 10 fm, \ \tau_{fo} = 7.5 \frac{fm}{c}, \ V_2 = V_3 = 0.05$$

- Particle mass: 140 MeV
- Anisotropic temperature $\Lambda(x^{\mu})$, anisotropy parameter $\xi(x^{\mu})$

Transverse momentum spectrum for different:

 ξ values

 (Λ,ξ) -pairs



- ξ decreasing more fast particles, due to higher effective pressure
- Varying "temperatures"
 Λ at freeze out, but
 nearly same spectrum

Anisotropic flow coefficients

Fourier series in azimuthal angle ϕ_p $d^2N = 1 \ dN \left[1 \ \sum_{n=1}^{\infty} 2^n d^n \right]$

$$\frac{d^2 N}{d^2 \vec{p_t}} = \frac{1}{2\pi} \frac{dN}{p_t dp_t} \left[1 + \sum_{n=1}^{\infty} 2v_n \cos(n(\phi_p - \psi_n)) \right]$$

Fourier coefficients:

$$v_n = \frac{\int_0^{2\pi} d\phi_p \cos(n\phi_p) \frac{d^2 N^{aniso}}{d\vec{p_t}}}{\int_0^{2\pi} d\phi_p \frac{d^2 N^{aniso}}{d\vec{p_t}}}$$

Elliptic flow v_2 for different

 ξ values

 (Λ,ξ) -pairs



- Higher anisotropy leads to smaller anisotropic flow
- v_3 and v_4 follow same trend

 Variation of parameters gives nearly the same observables

Hanbury-Brown Twiss radii

• Computation of HBT radii: $R_{out}^2 = \langle (\tilde{x} - v_{out} \ \tilde{t})^2 \rangle$

$$R_{side}^2 = \langle (\tilde{y})^2 \rangle$$
$$R_{long}^2 = \langle (\tilde{z})^2 \rangle$$

• Averaging:
$$\langle X \rangle = \frac{\int d^4x \ X \ S(x^{\mu}, \vec{k})}{\int d^4x \ S(x^{\mu}, \vec{k})}$$

• Source-function: $S(x^{\mu}, \vec{p}) = \int p^{\mu} d\Sigma_{\mu} f(\vec{p}, x'^{\mu}) \delta^{(4)}(x'^{\mu} - x^{\mu})$

HBT radii for different anisotropy parameter values



- R_{out} varies due to higher radial thermal motion
- R_{long} nearly constant

HBT radii for different (Λ,ξ)pairs



- Small variation of Rout
- R_{side} very sensitive to parameter pair

Conclusion

• Introduce an anisotropic phase space distribution

 \rightarrow Get additional parameter ξ

- Lower sensitivity of observables on parameters at freeze out
- Option for a smoother transition from hydro to kinetic theory
- Further reading:
 - 1. N. Borghini, SF, C. Lang; Eur. Phys. J. C 75 (2015) 275
 - 2. Work in preparation

Thank you for your attention