

Simplified readout for CYGNUS-TPC:

a massive directional dark matter detector

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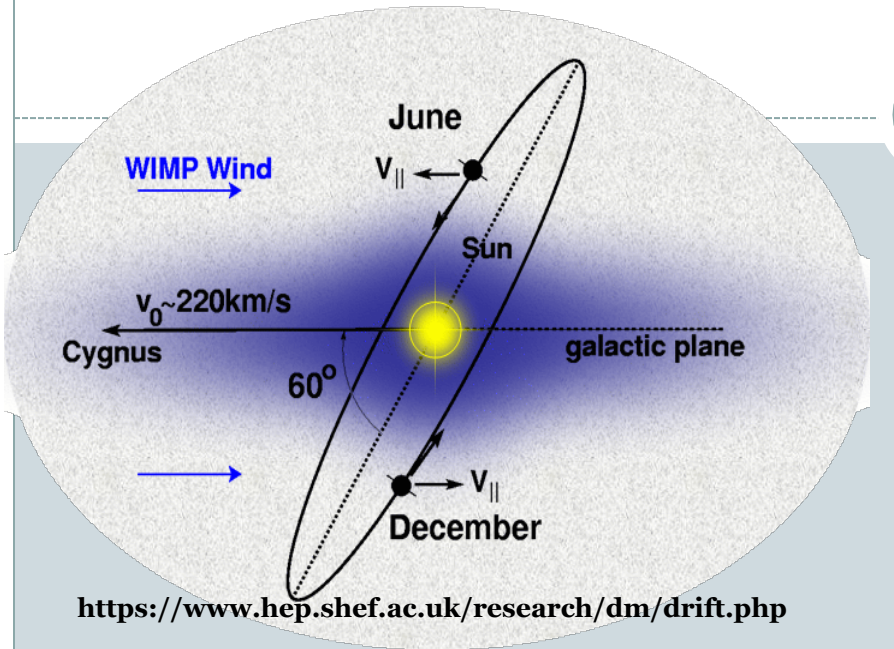
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The University of Sheffield
(DRIFT collaboration)



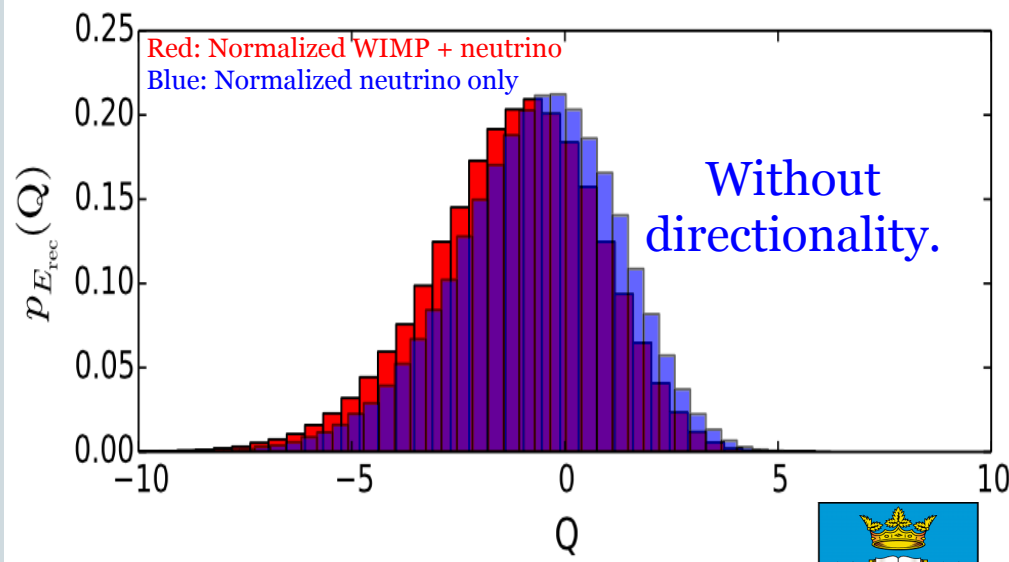
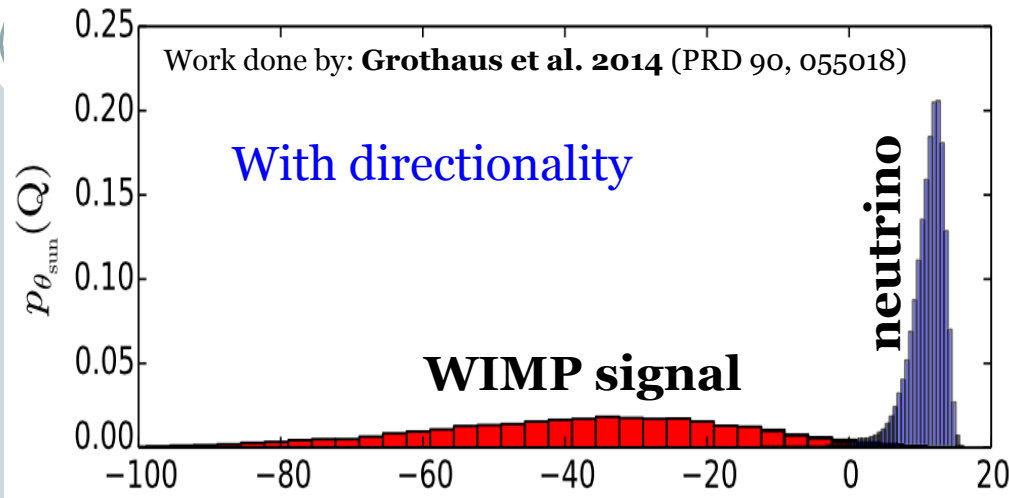
TUESDAY, 22ND MARCH 2016

Our Motivation



<https://www.hep.shef.ac.uk/research/dm/drift.php>

- WIMP wind coming from CYGNUS
- WIMP induced recoil tracks should point back to CYGNUS
- Will need massive (in tons) directional detector to be able to go beyond the neutrino floor.
- Plus enough physics run time.
See: O'Hare et al. (2015) PRD 92, 063518.
- No "neutrino floor" for directional DM detectors Grothaus et al. 2014.

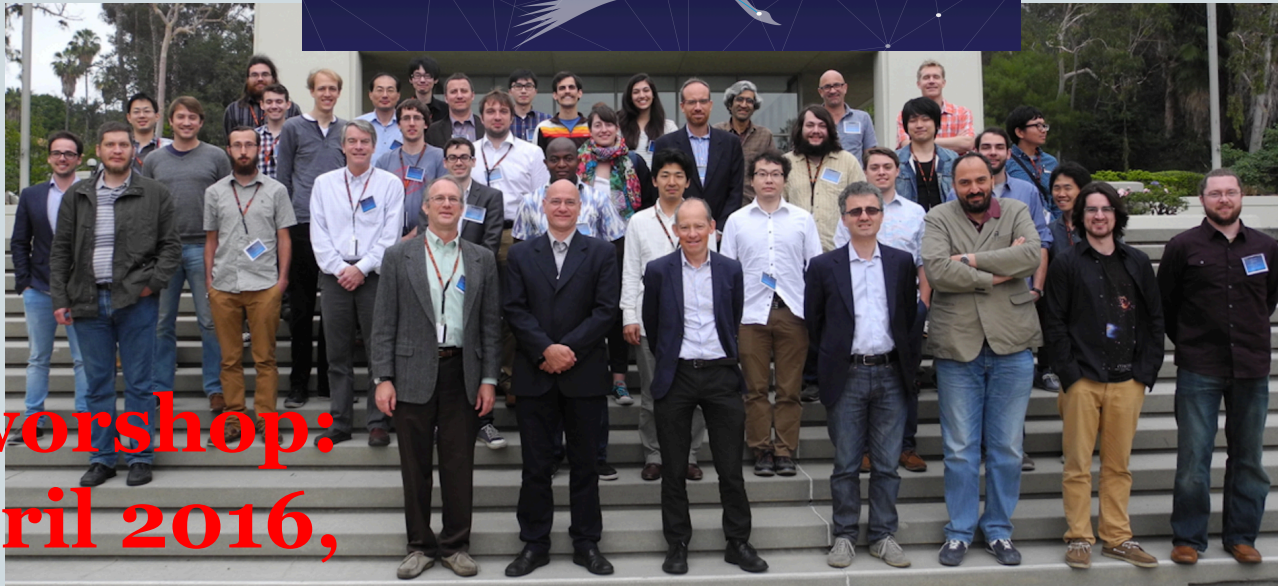


- Study done for 6 GeV WIMP in CF₄ TPC



Plans for CYGNUS-TPC

Los Angeles, USA
June 2015.



**Next workshop:
7-8 April 2016,
Italy.**

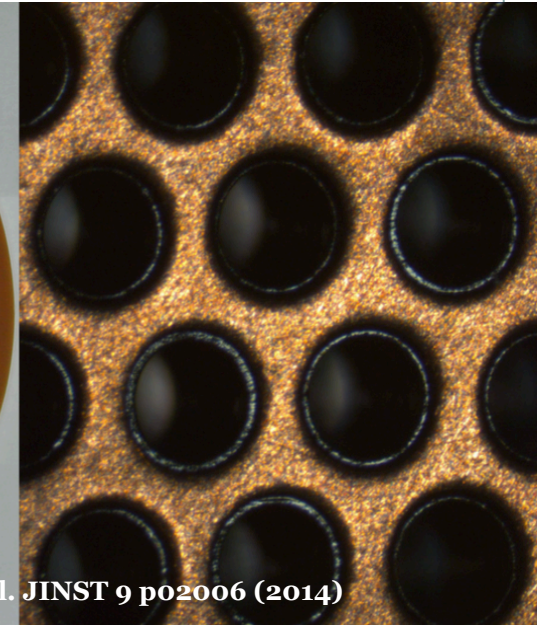
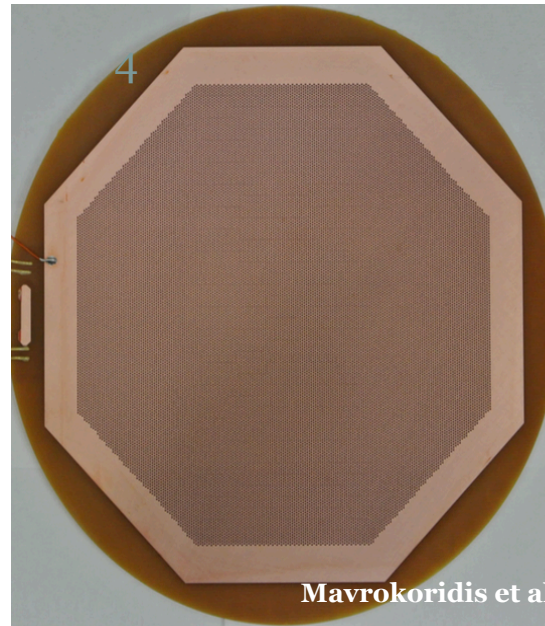
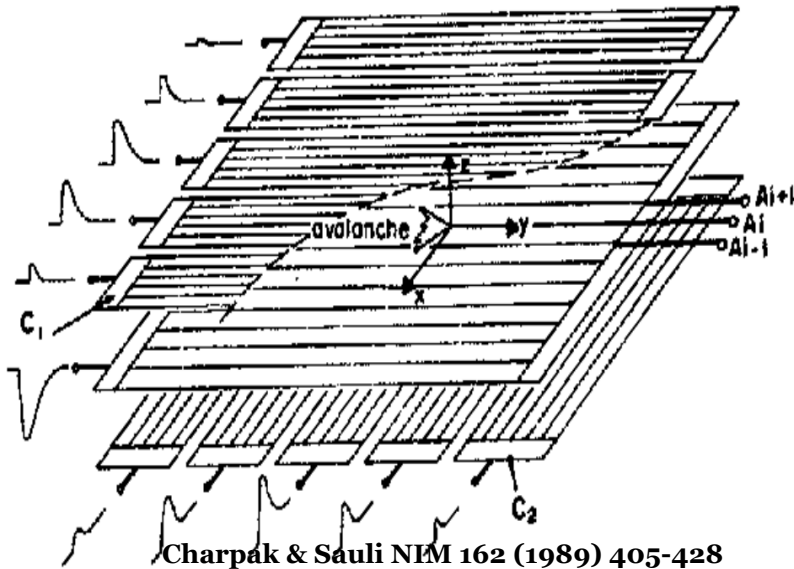
**Lots of phone
meetings.**

Boulby, UK
Jan. 2016

**CYGNUS-TPC WORKSHOP
2016.**



Possible readout for CYGNUS



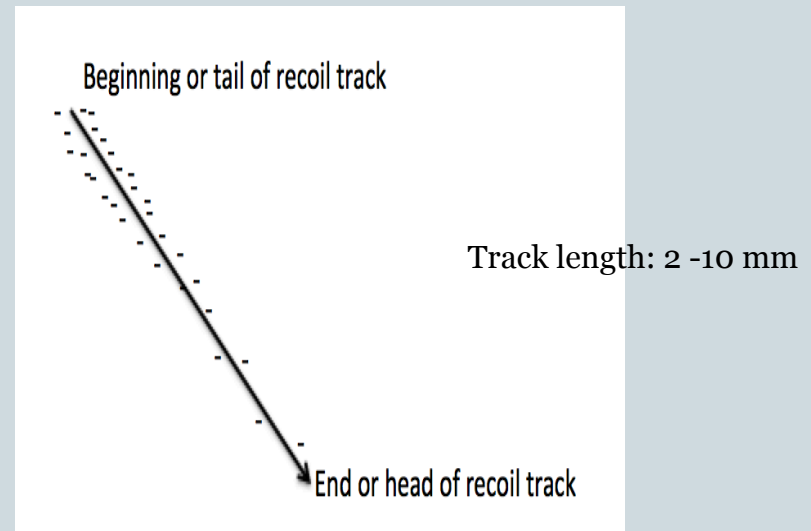
- ❑ Conventional MWPC: limited to ~ 1 mm resolution in x-direction.
- ❑ More wires: means more money for readout electronics.
- ❑ Any favored readout for CYGNUS should be sensitive to signal head-tail effect.
- ❑ Robust, good gas gain: $10^4 - 10^6$ in pure CF_4 .
- ❑ Configuration can be double ThGEM or tripple ThGEM.
- ❑ Can be mesh based: Micromegas etc.

DRIFT-II_d DETECTOR AND HEAD-TAIL

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- ❑ Fiducial volume: $0.8 \times 0.8 \times 0.8 \text{ m}^3$.
- ❑ Fiducial mass: 137 g.
- ❑ Gas: 30:10:1 Torr of $\text{CS}_2:\text{CF}_4:\text{O}_2$ gas mixture.
- ❑ Drift field: $\sim 600 \text{ Vcm}^{-1}$.



- ❑ Consists of two back-to-back gas TPC
- ❑ One 1 m^2 central cathode, two 1 m^2 MWPC readouts.

For details on HT measurement in DRIFT-II_d: *AstroPart. Phys.* **31 (2009) 261-266**

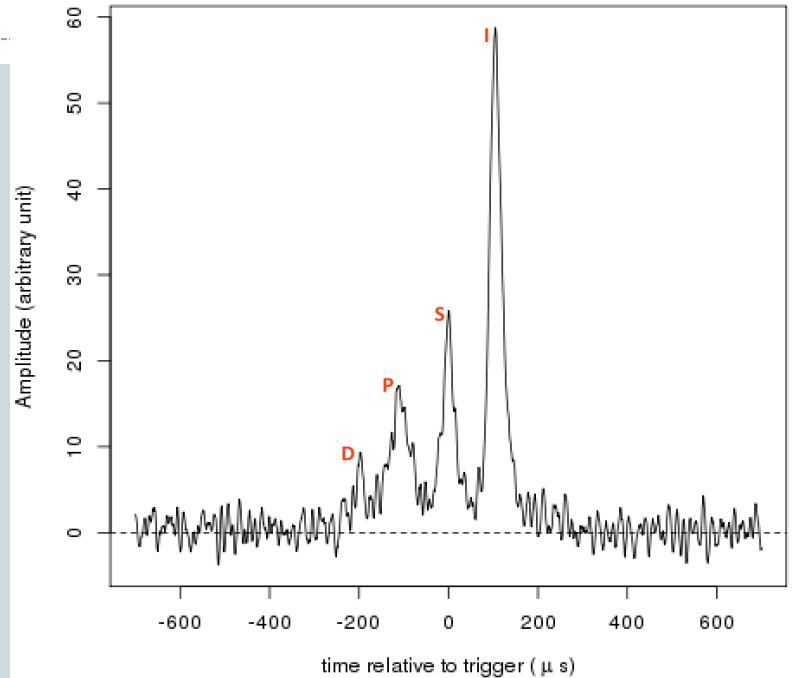
Fiducialisation in DRIFT

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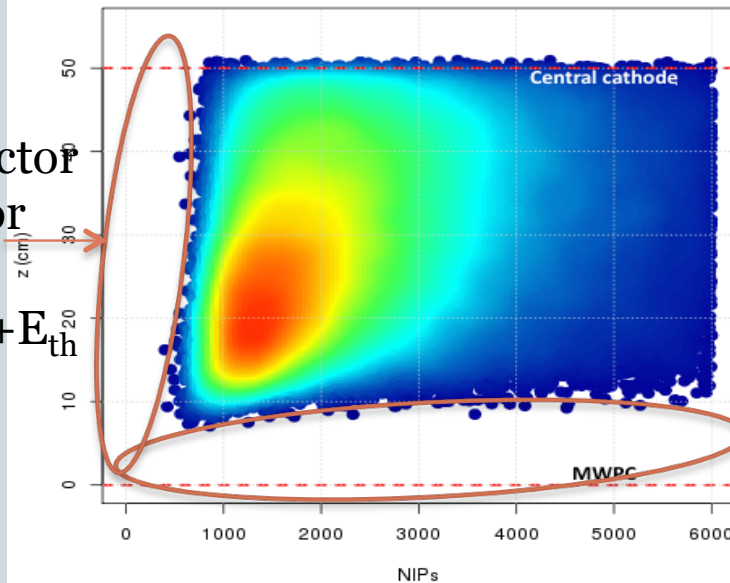
- Added 1 Torr of O₂ in 30+10 Torr of CS₂+CF₄ gas mixture.
- This results to three minority carriers and one main charge cloud.

$$z = (t_I - t_P) \left(\frac{v_I v_P}{v_P - v_I} \right)$$

See: Battat et al. 2015 PDU 9-10, 1-7.



Due to detector efficiency for low energy events in $z + E_{th}$

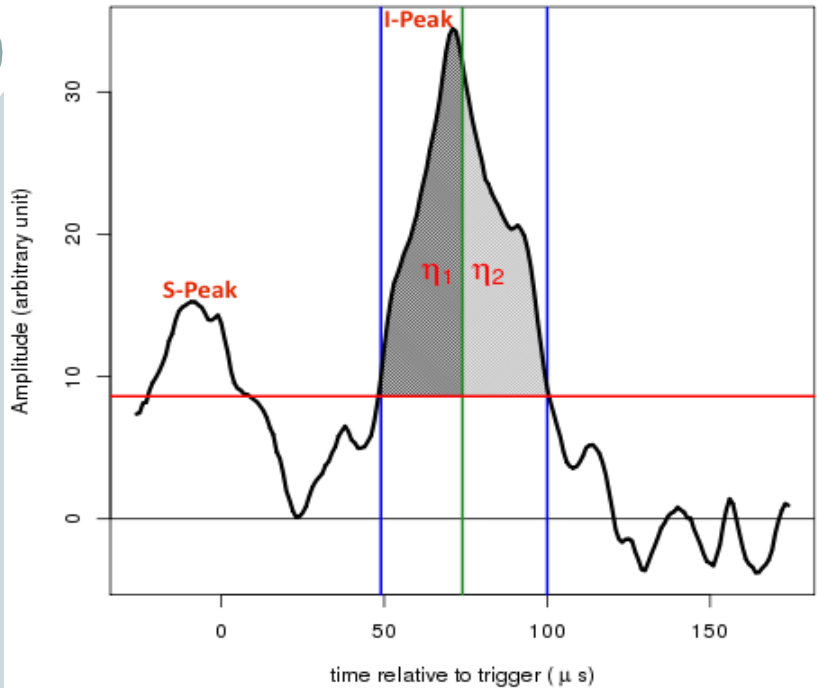
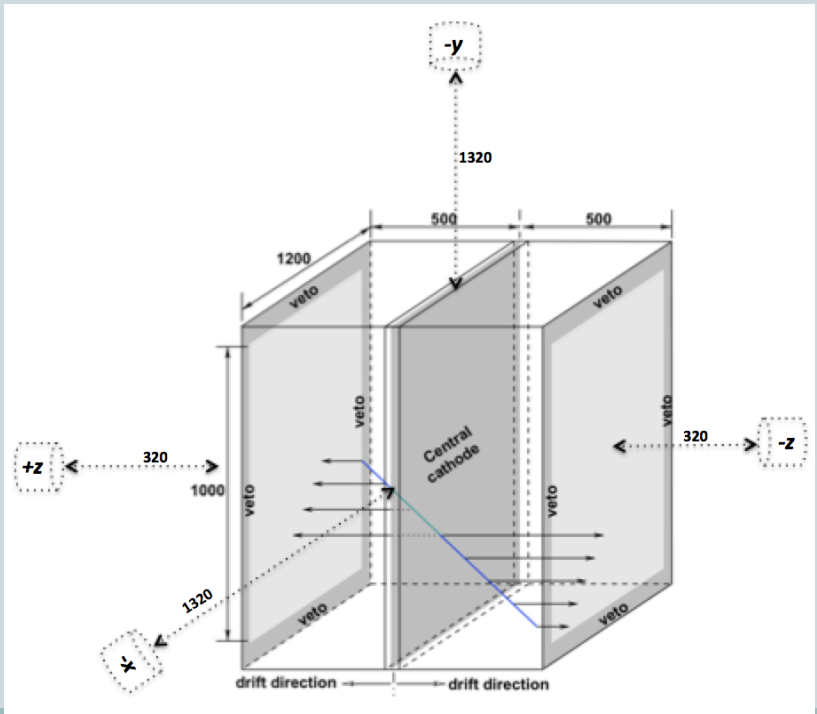


Low-z unfiducialised events were rejected in analysis. Low-z events are unfiducialised due to lack of peak separation.

Directionality after fiducialisation

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- DRIFT-IId was exposed to fast neutrons from Cf-252 source from different directions.
- Measured asymmetry in charge distribution along recoil tracks known as HT effect.



- More charge in the beginning (TAIL) than the end (HEAD) of nuclear recoil tracks.

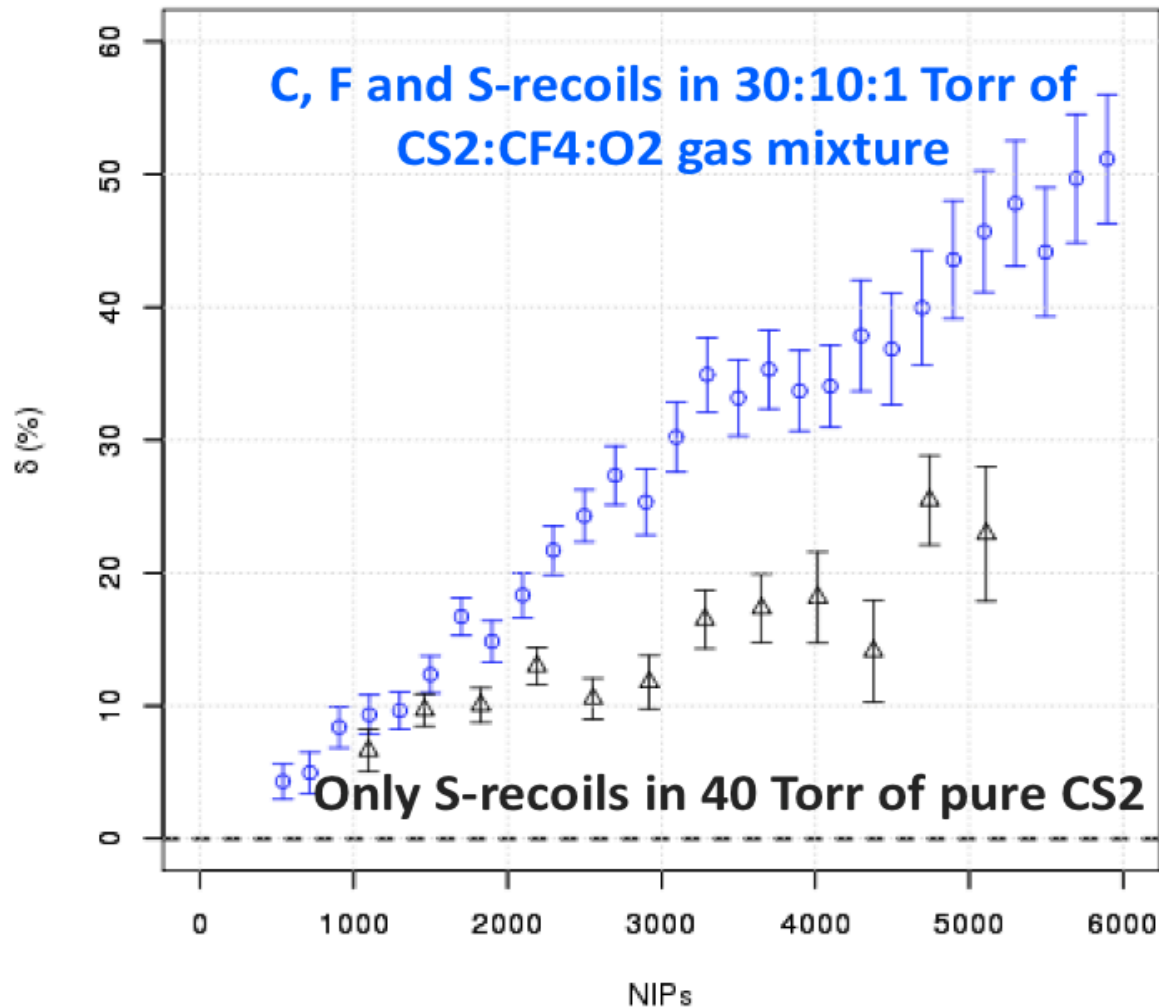
See: *AstroPart. Phys.* 31 (2009) 261-266.

- When η_1 is the TAIL and η_2 is the HEAD of a recoil track, HT asymmetry parameter α is:

$$\alpha = \frac{\eta_1}{\eta_2}$$



Head-tail results



$$\delta = \left[\frac{200 |\langle \alpha_L \rangle - \langle \alpha_R \rangle|}{\langle \alpha_L \rangle + \langle \alpha_R \rangle} \right]_{op}$$

- ❑ Sensitive to HT at lower energy compared old result from only S-recoils is due to lower analysis threshold.
- ❑ 50 keV fluorine equivalent recoil energy is 1055 NIPs.

See: Hitachi RPC 77 (2008) 1311-1317

For the pure CS₂ results, see: *AstroPart. Phys.* 31 (2009) 261-266



Neutron events for 3-sigma H-T

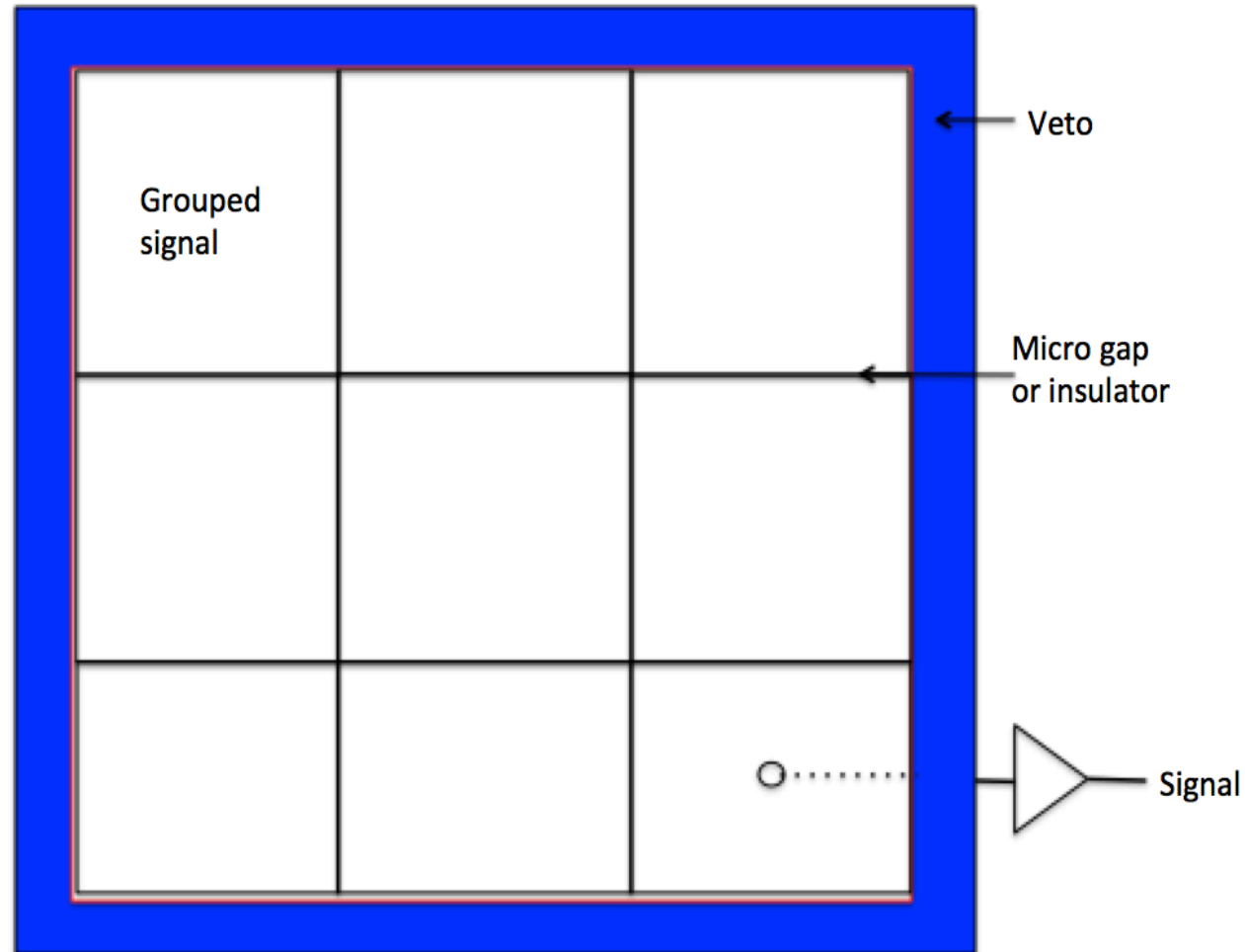
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- ❑ HT significance was computed for different numbers of events and was used to put a constraint on the number of neutron events required for HT detection.
- ❑ Considering only events from the optimal directions (+z and -z runs).
- ❑ It was found that ~320 neutron events will be enough for 3-sigma head-tail detection.
- ❑ MC will be required to estimate the number of events for a given mass of WIMP that will produce this result.

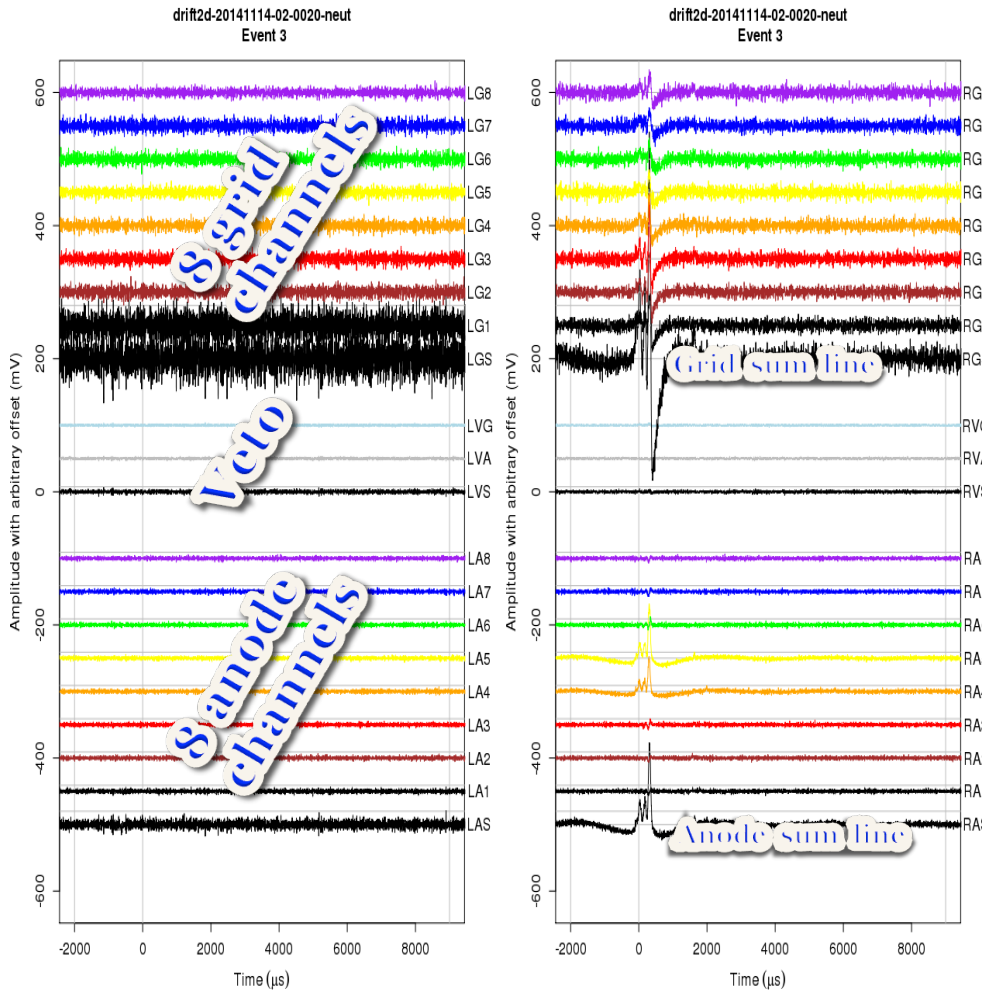


Can reduce cost of readout

- ❑ Can group signal even with ThGEMs.
- ❑ Though, smaller gaps will increase the capacitance.
- ❑ Micro gaps needed for alpha discrimination.
- ❑ Can include fiducialisation and retain head-tail.



Test analysis for a simplified readout

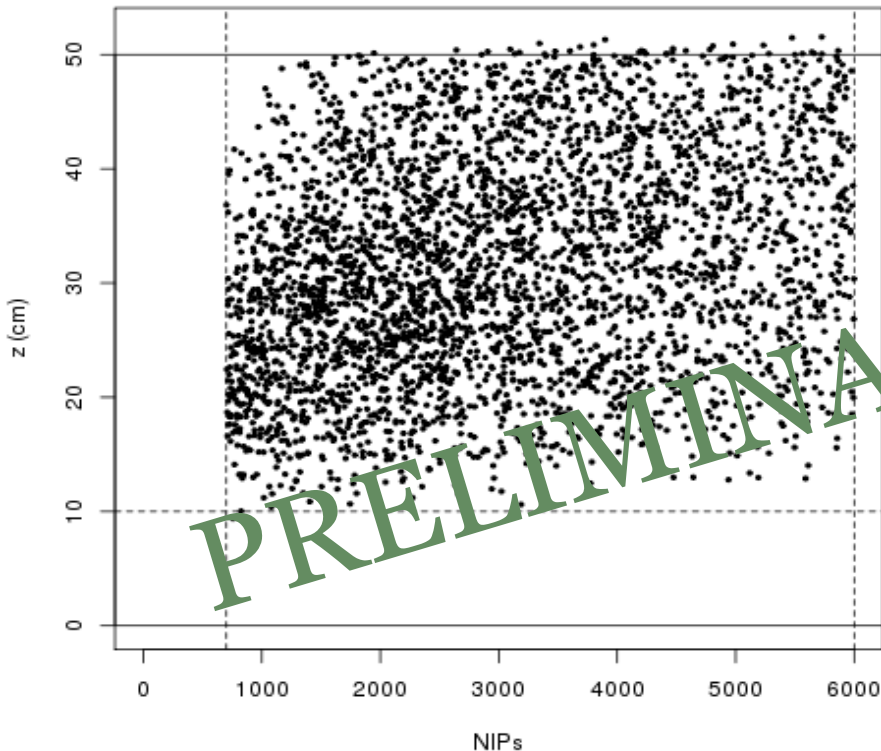


- ❑ MWPC in the simplified mode:
- ❑ Grouped **all** 448 anode wires and
- ❑ Grouped **all** 448 grid wires in DRIFT.
- ❑ Reduced the number of readout channels to 4 and the vetos.
- ❑ One anode, one grid channel from each detector.
- ❑ Analyzed neutron events and
- ❑ Background events in this new grouped data mode.
- ❑ Turned off number of hits cut (can be active in analysis of Cygnus data).

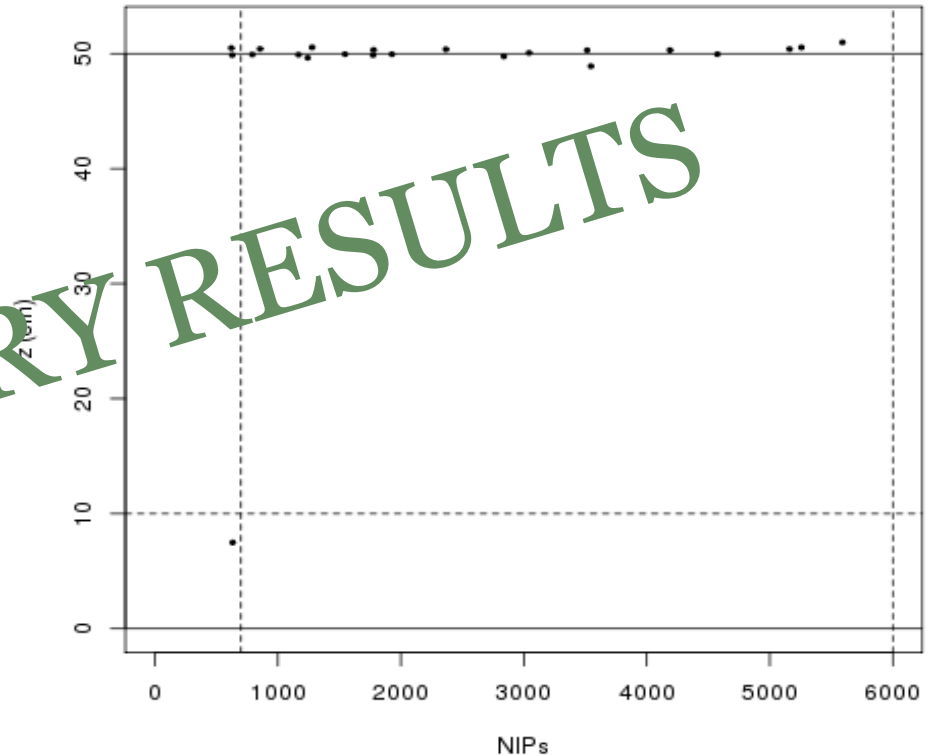
Normal DRIFT DAQ configuration.

Results: simplified readout analysis

z vs NIPs, neutron events, analysed only sum-line
0.8 live-time days, Rate: 4486 \pm 30 events/day(3589 events).



z vs NIPs, background run (shielded), analysed only sum-line
31.6 live-time days, Rate: 0.7 \pm 0.1 events/day.



PRELIMINARY RESULTS

- Still sensitive to neutron induced recoils.
- No background event seen over 31.6 live-time days, gammas and electron recoils may show up for longer runs.
- Optimization of this method requires more work.

CONCLUSION

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- ❑ Preliminary results suggest that the simplified and low cost readout could work for massive directional dark matter detectors for instance CYGNUS-TPC.
- ❑ This can be achieved without compromising the detector's head-tail sensitivity.
- ❑ It can also be fiducialised using existing technologies.
- ❑ Though gamma events and electron recoils are potential problems of this new set-up.

