

SUSY/BSM II



Mario Martínez



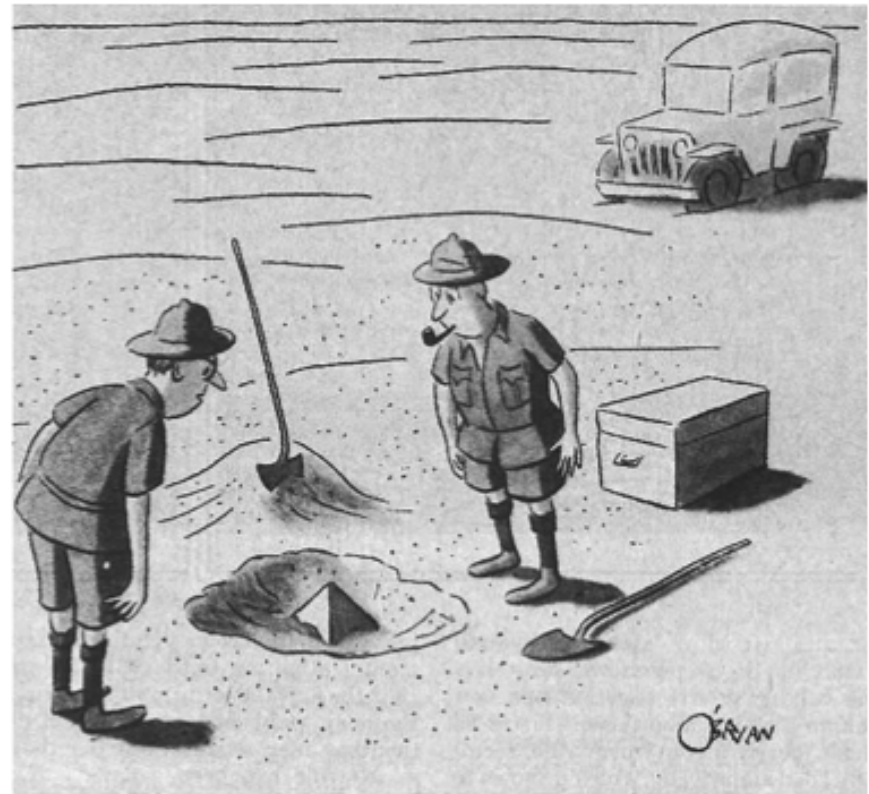
HASCO SUMMER SCHOOL 2016

Outline for Part II

- How to do/read a search
- Statistical facts

Overview of SUSY searches

- Inclusive Searches
- 3rd Generation Squarks
- EWK Searches
- SUSY Higgs
- $B_s \rightarrow \mu\mu$
- GMSB SUSY
- RVP Searches



"This could be the discovery of the century. Depending, of course, on how far down it goes."

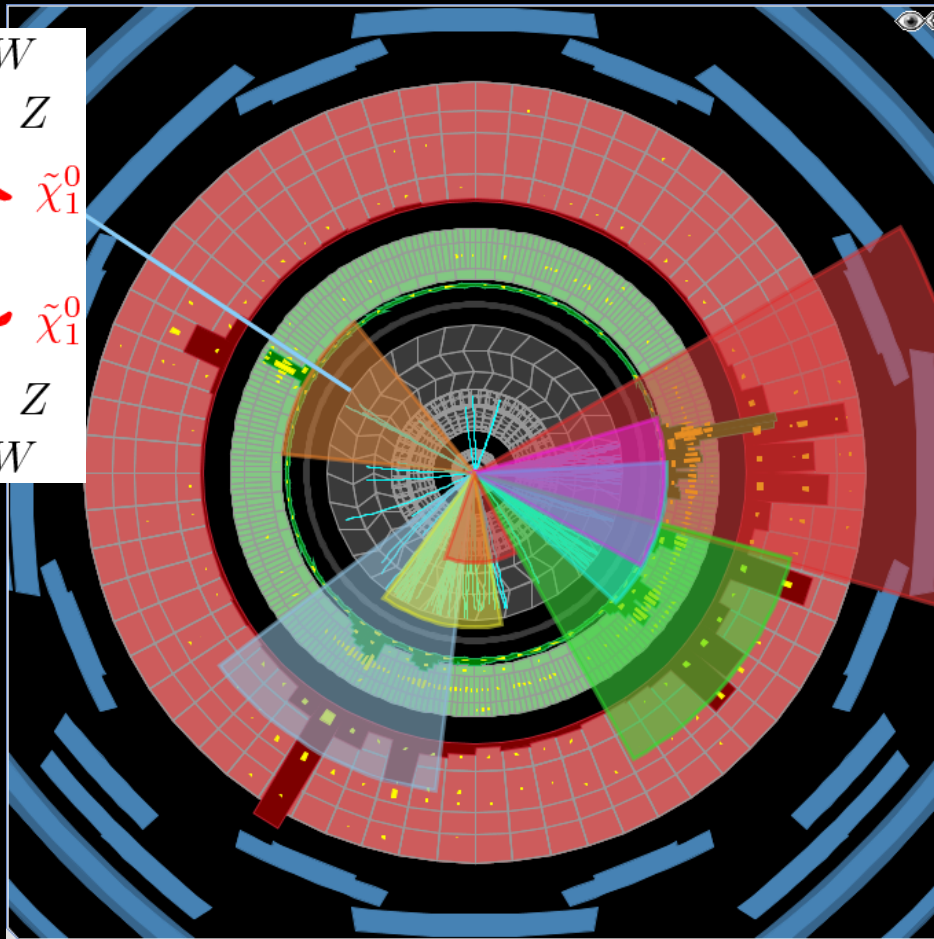
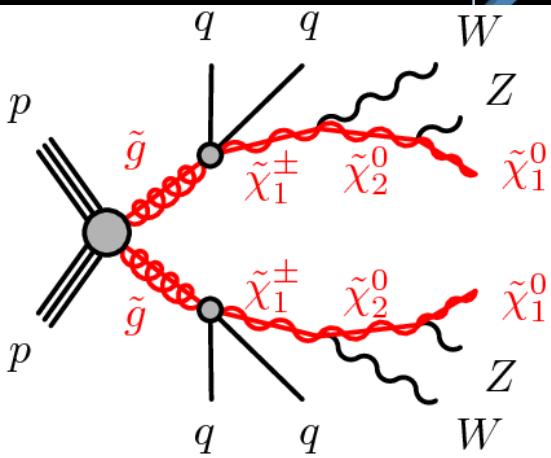
*Disclaimer: completely unbalanced set of results from CMS and ATLAS
No attempt to have latest results everywhere*

Basic Guidelines on how to perform a Search

Object reconstruction

Background Estimation

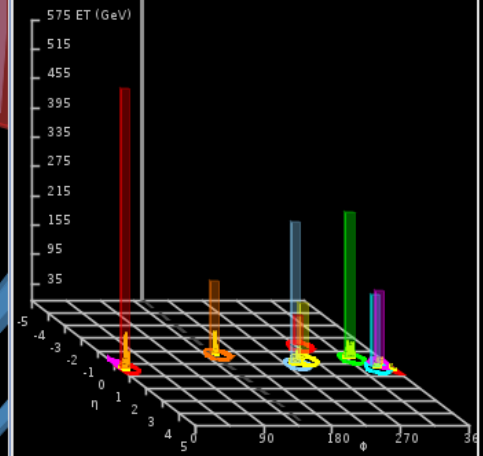
Likelihood Fits



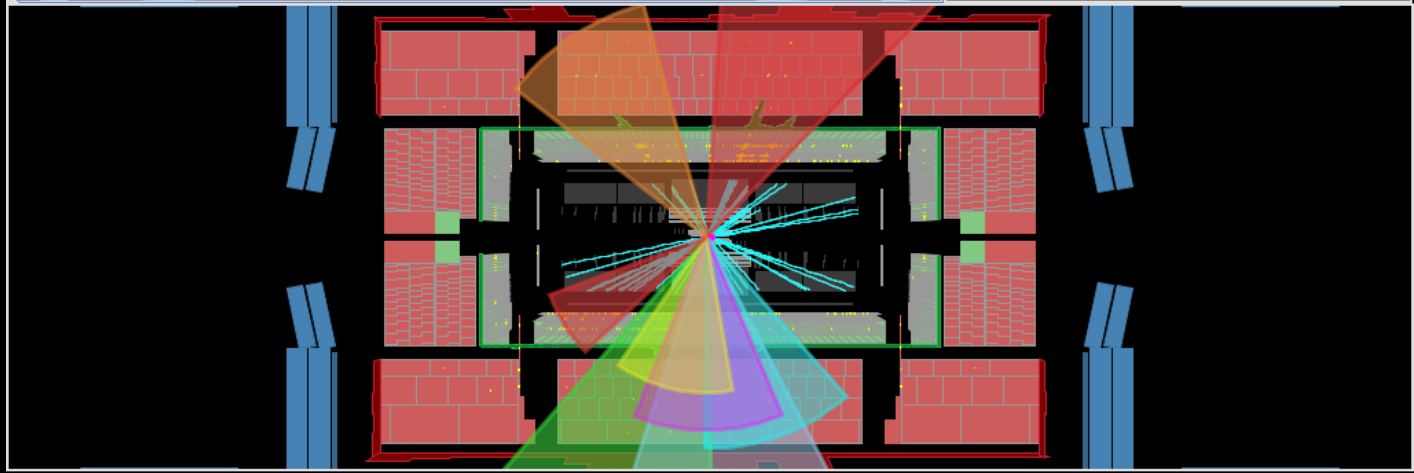

ATLAS
 EXPERIMENT

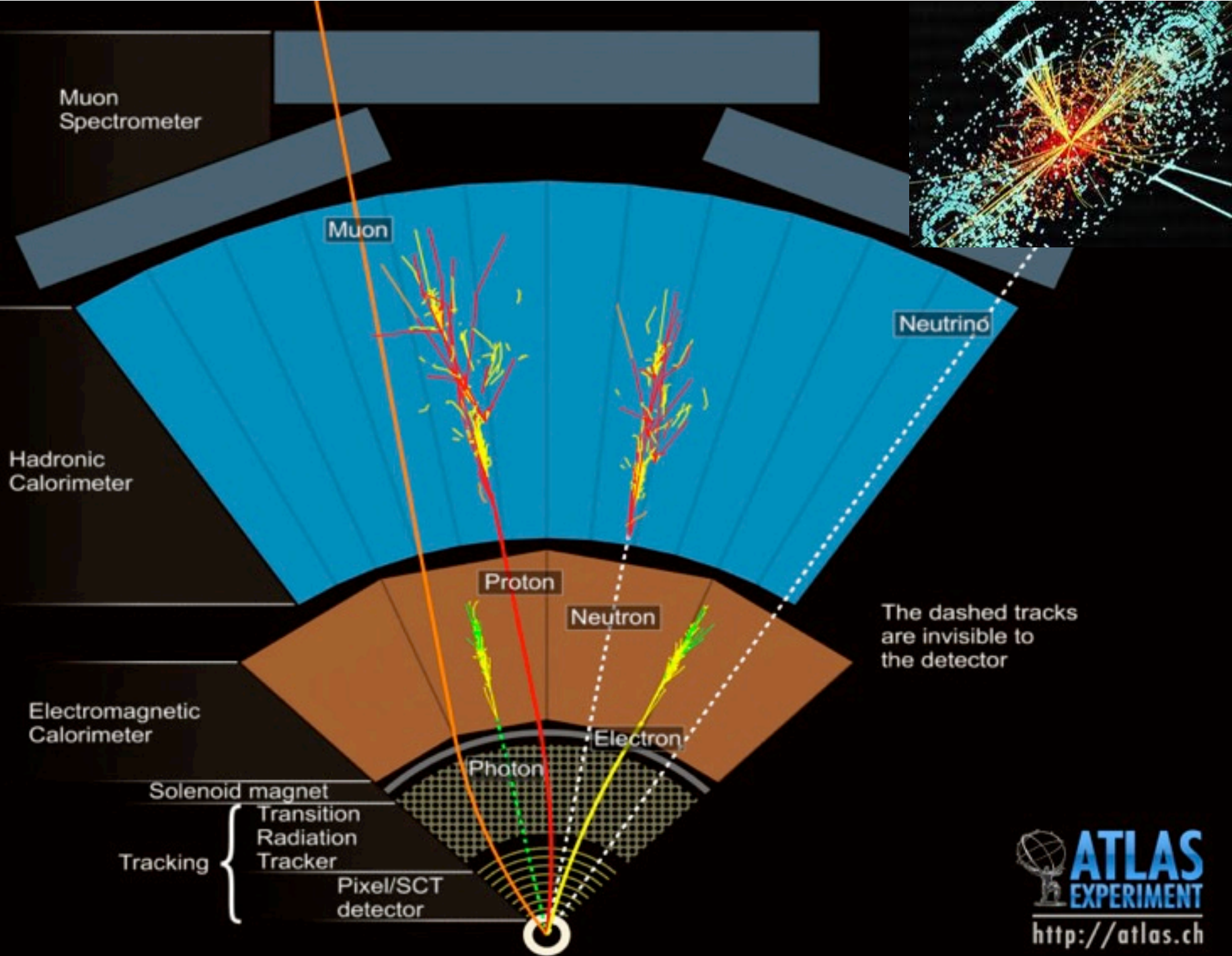
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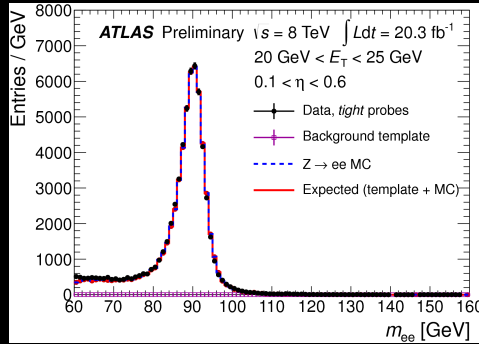
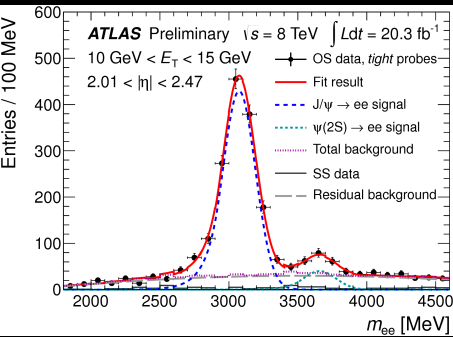
?



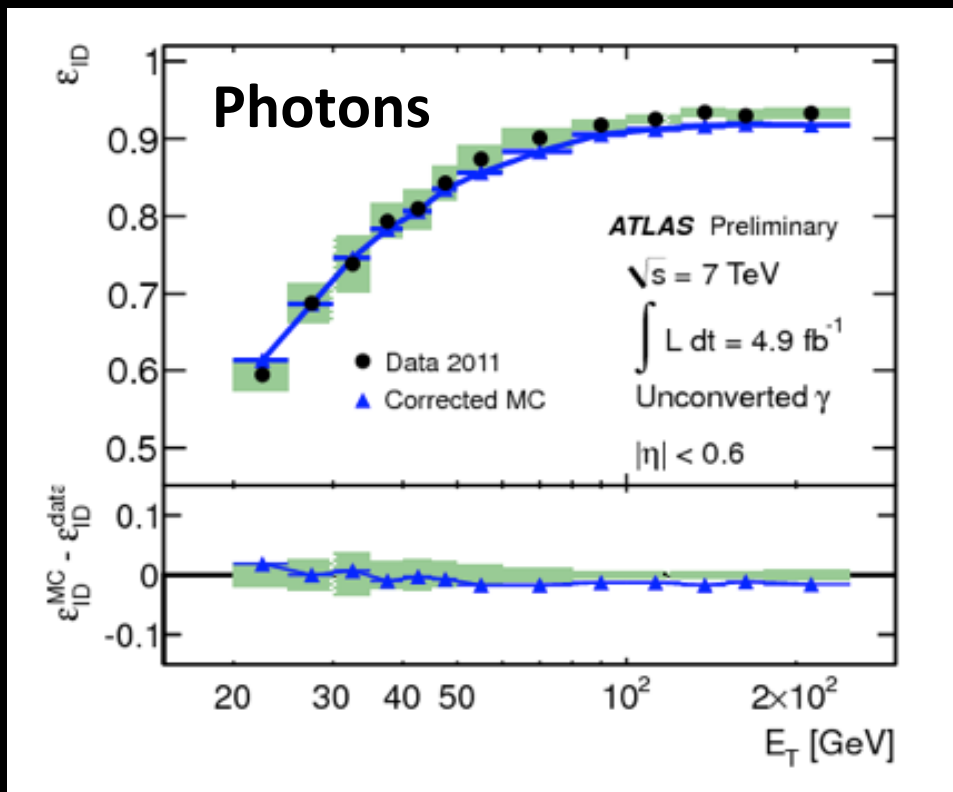
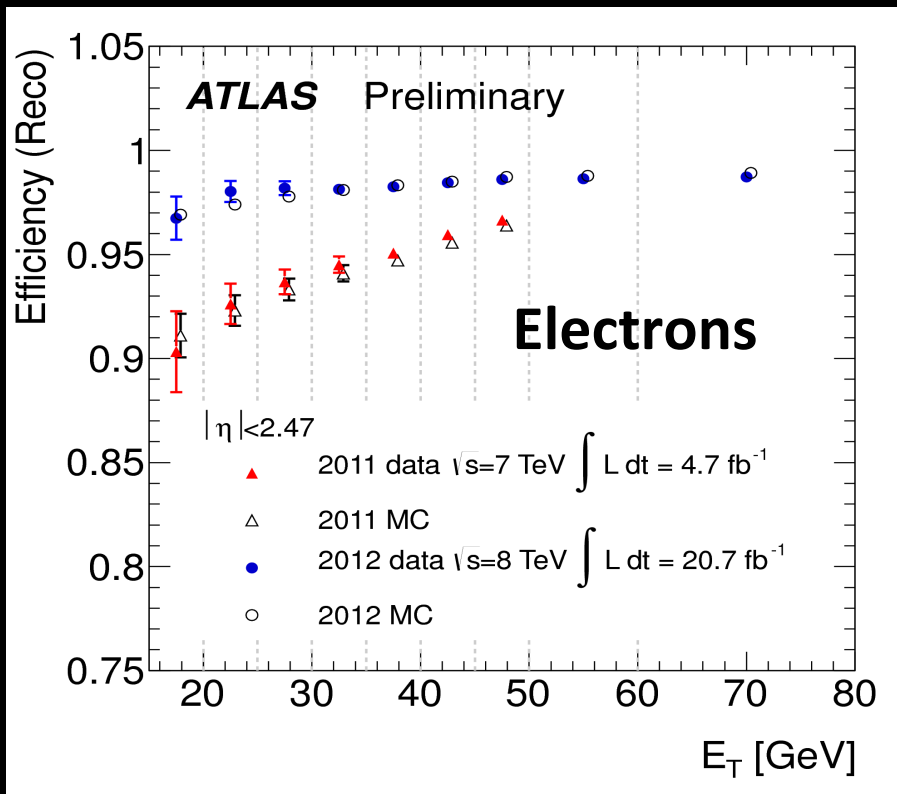
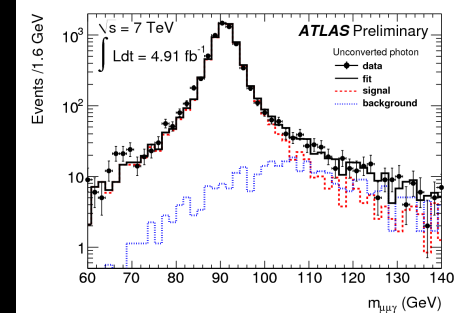
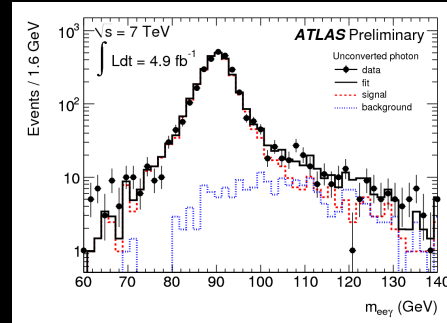


Building Blocks

$Z \rightarrow ee, J/\Psi \rightarrow ee \dots$

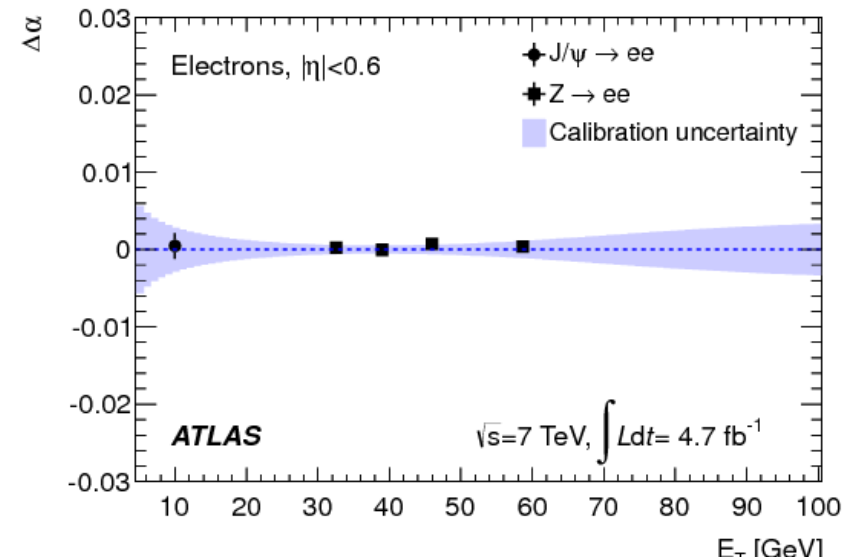
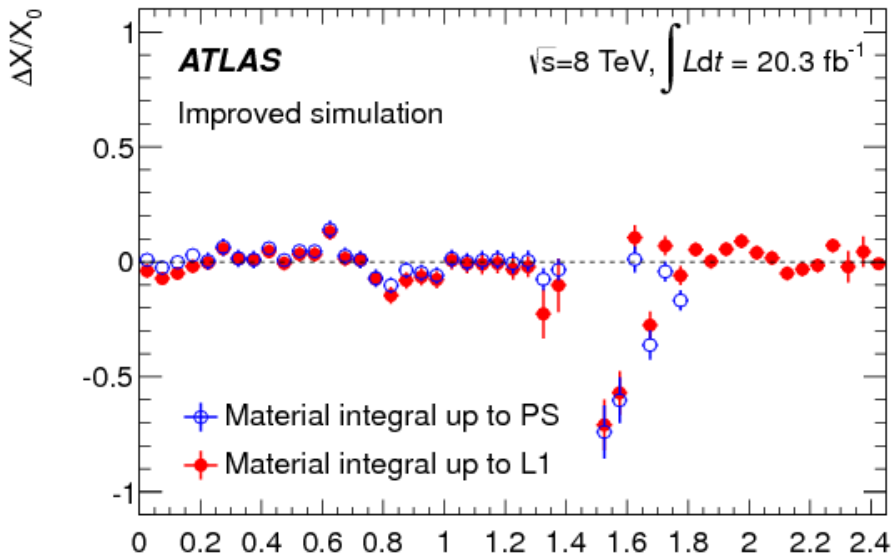
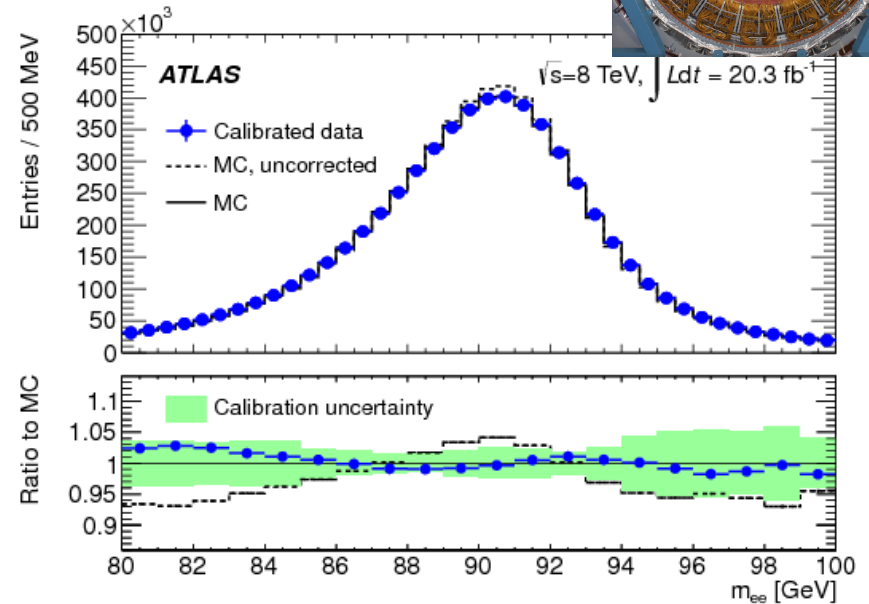
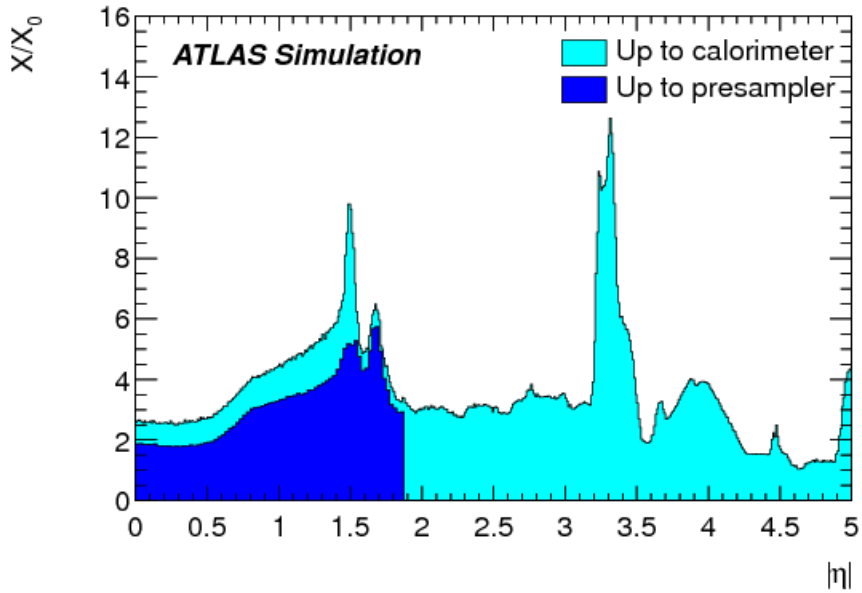
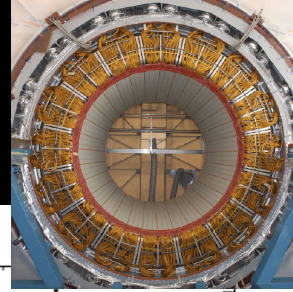


eey and $\mu\mu\gamma$

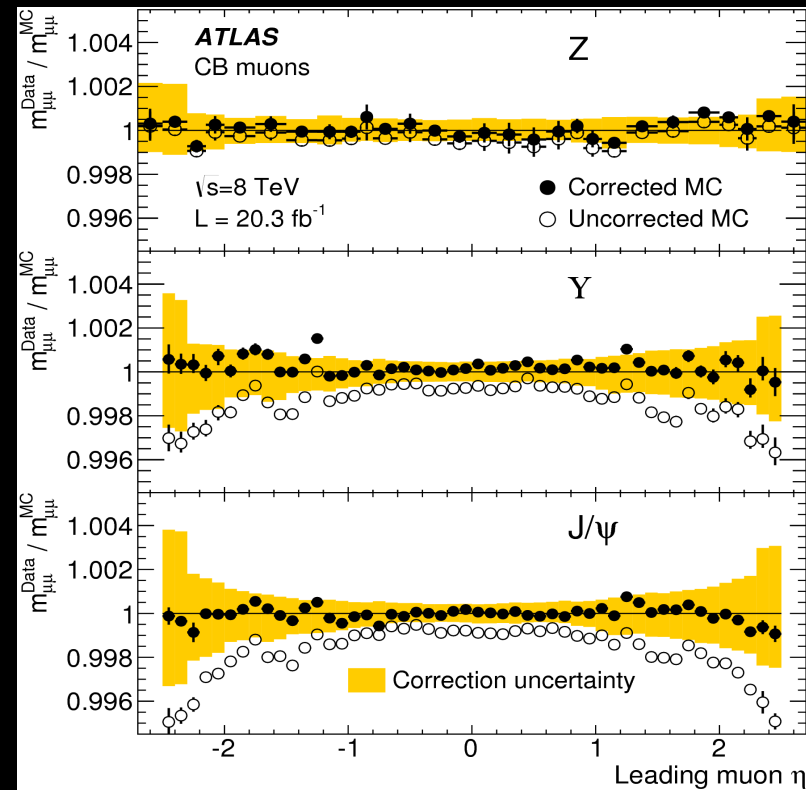


Detector Material Building Blocks

EM absolute scale

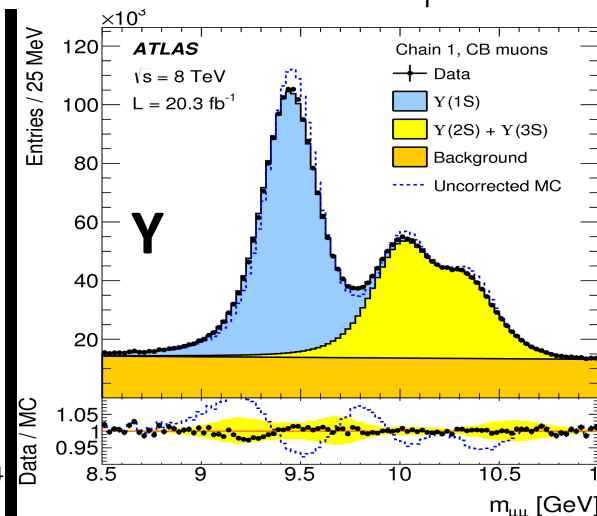
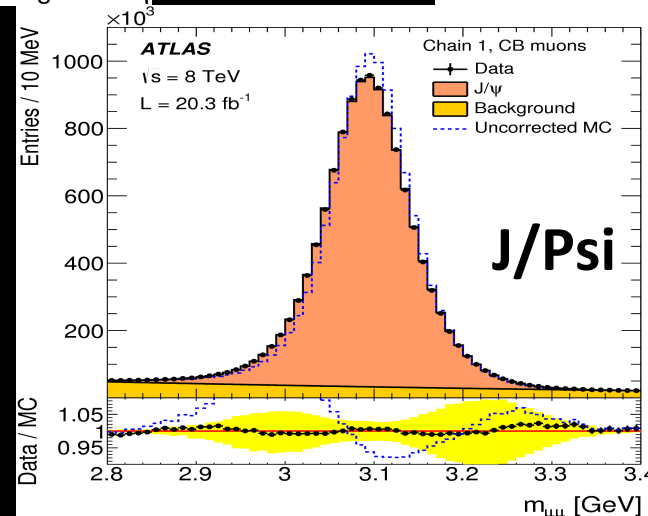
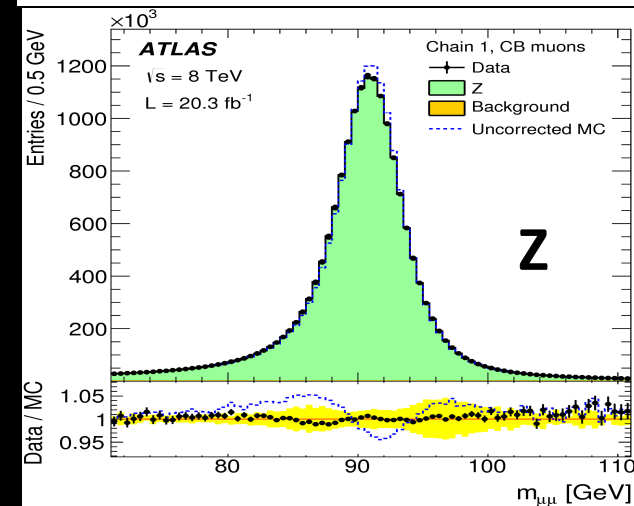
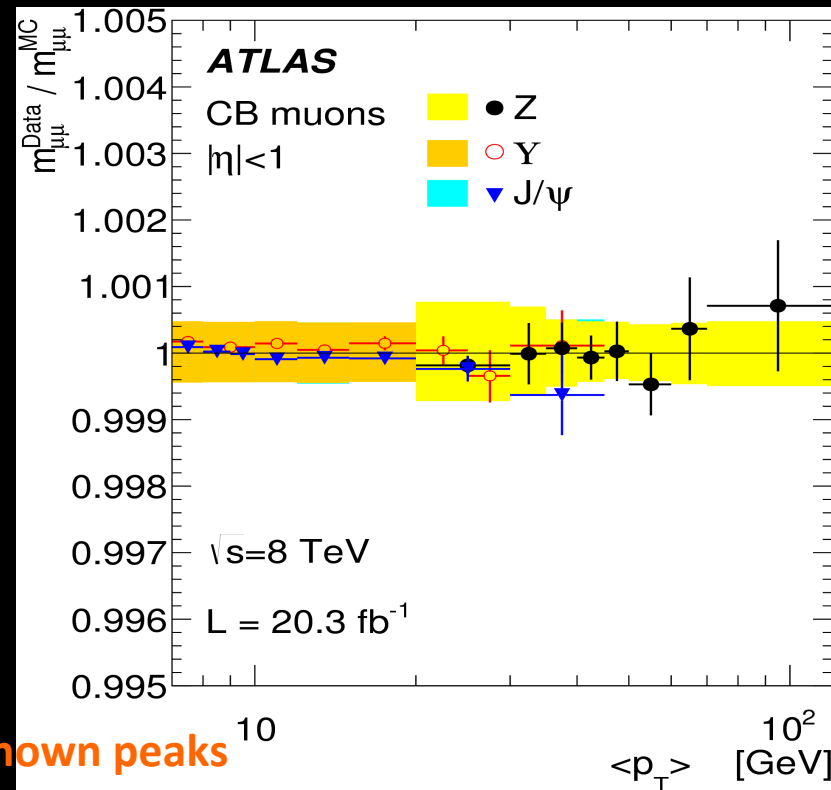


Building Blocks



Alignment of trackers and muon chambers

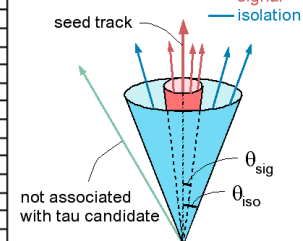
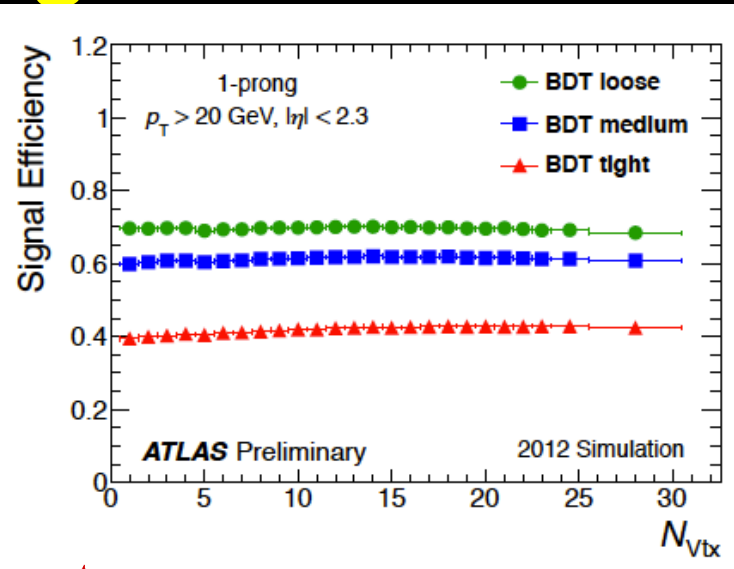
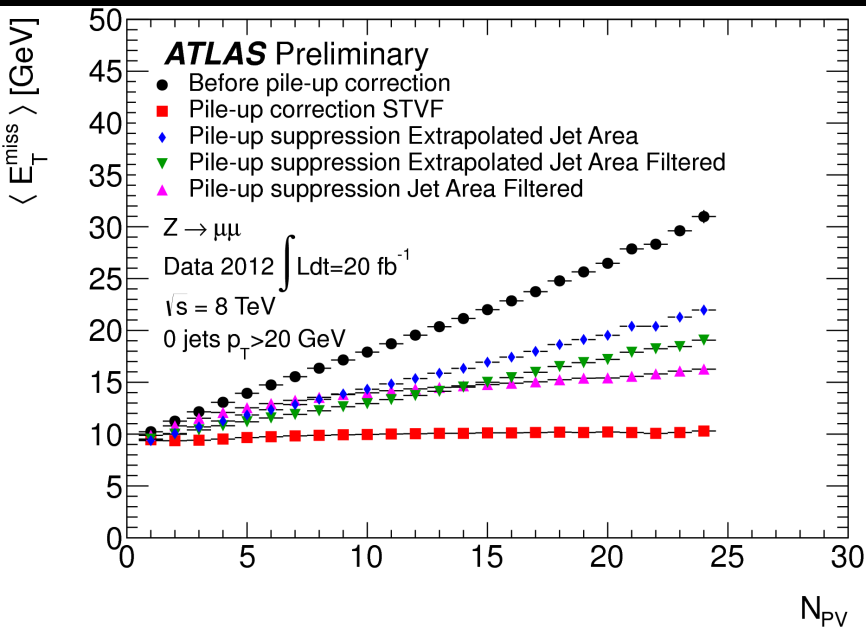
Using well-known peaks



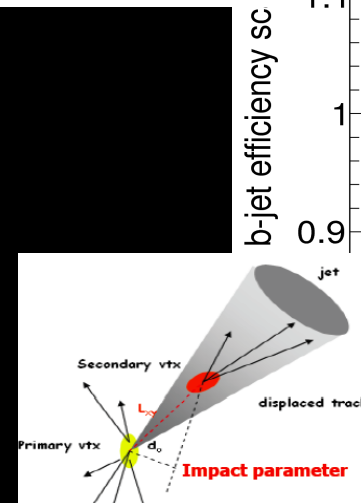
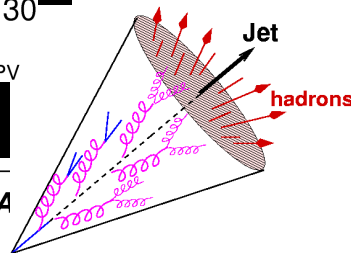
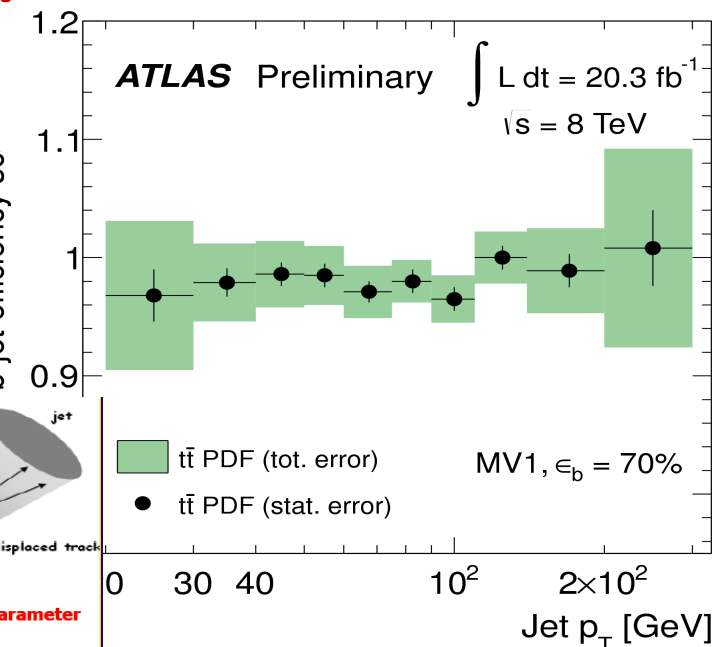
Missing E_T vs PILEUP

Building Blocks

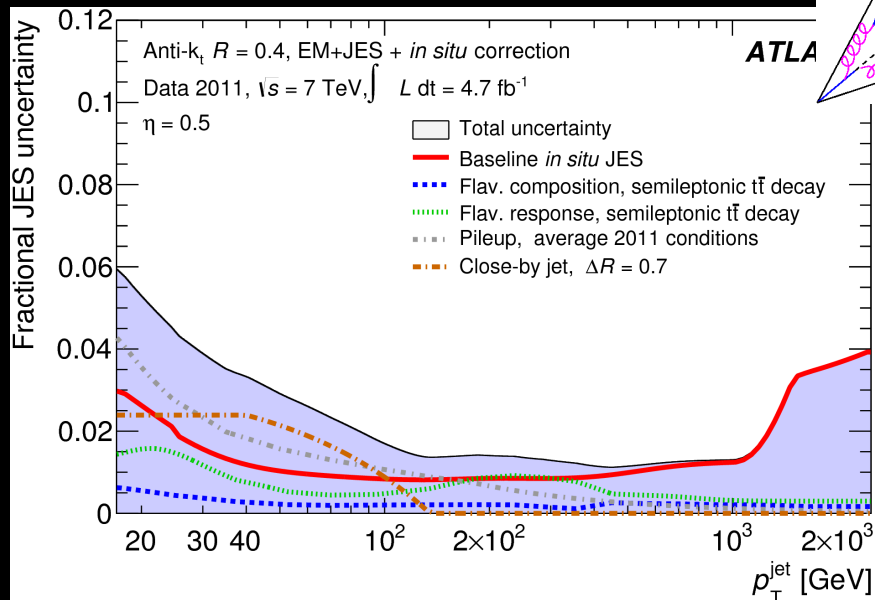
TAU ID vs PILEUP



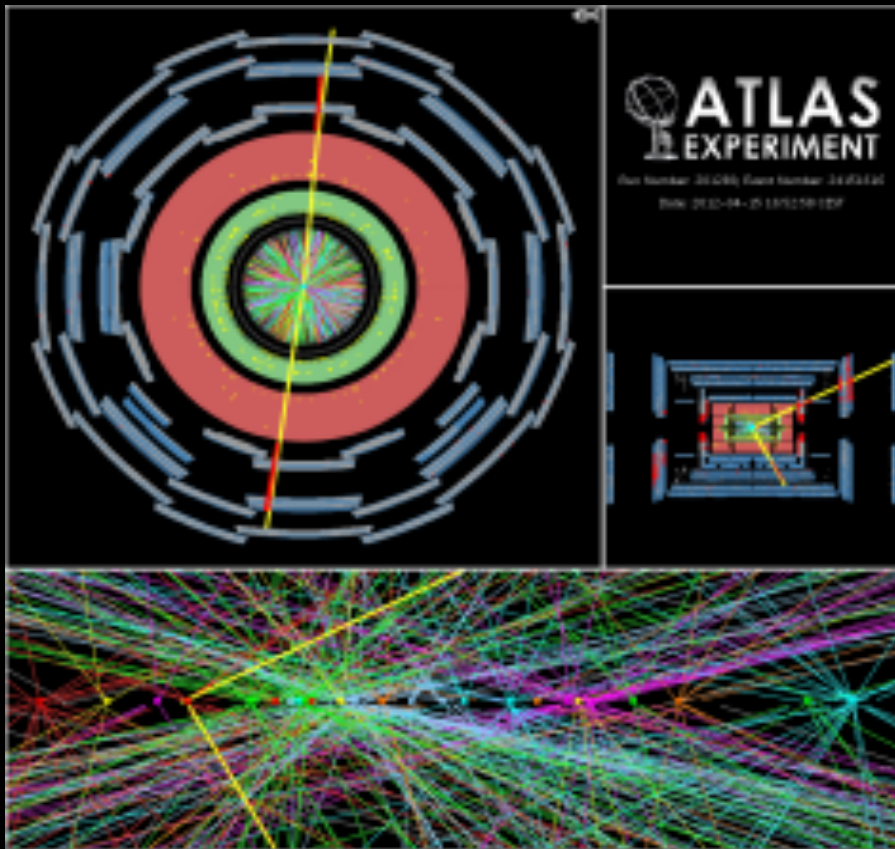
B-JET TAGGING EFFICIENCY



JET ENERGY SCALE UNCERTAINTY

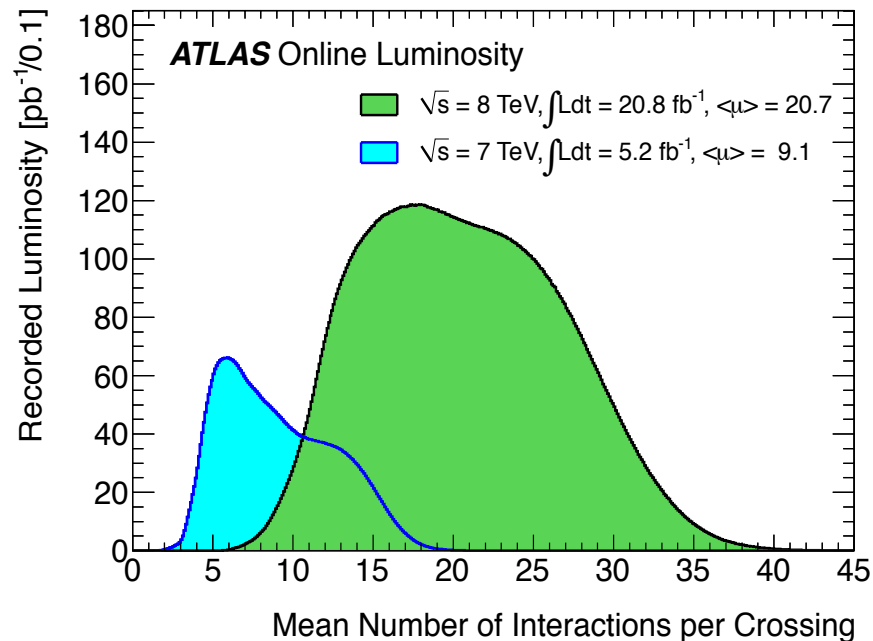


Multiple Interactions

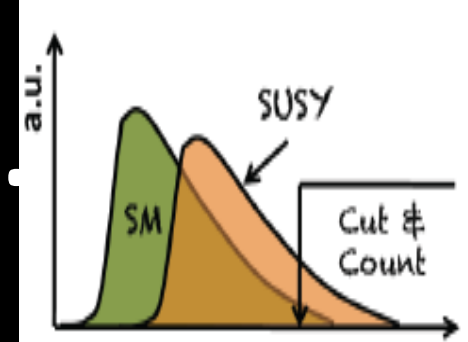


Up to 40 interactions / crossing
(requires enormous efforts to understand the reconstruction of the physics objects...)

$Z \rightarrow \mu\mu$ events with
20 interactions on top

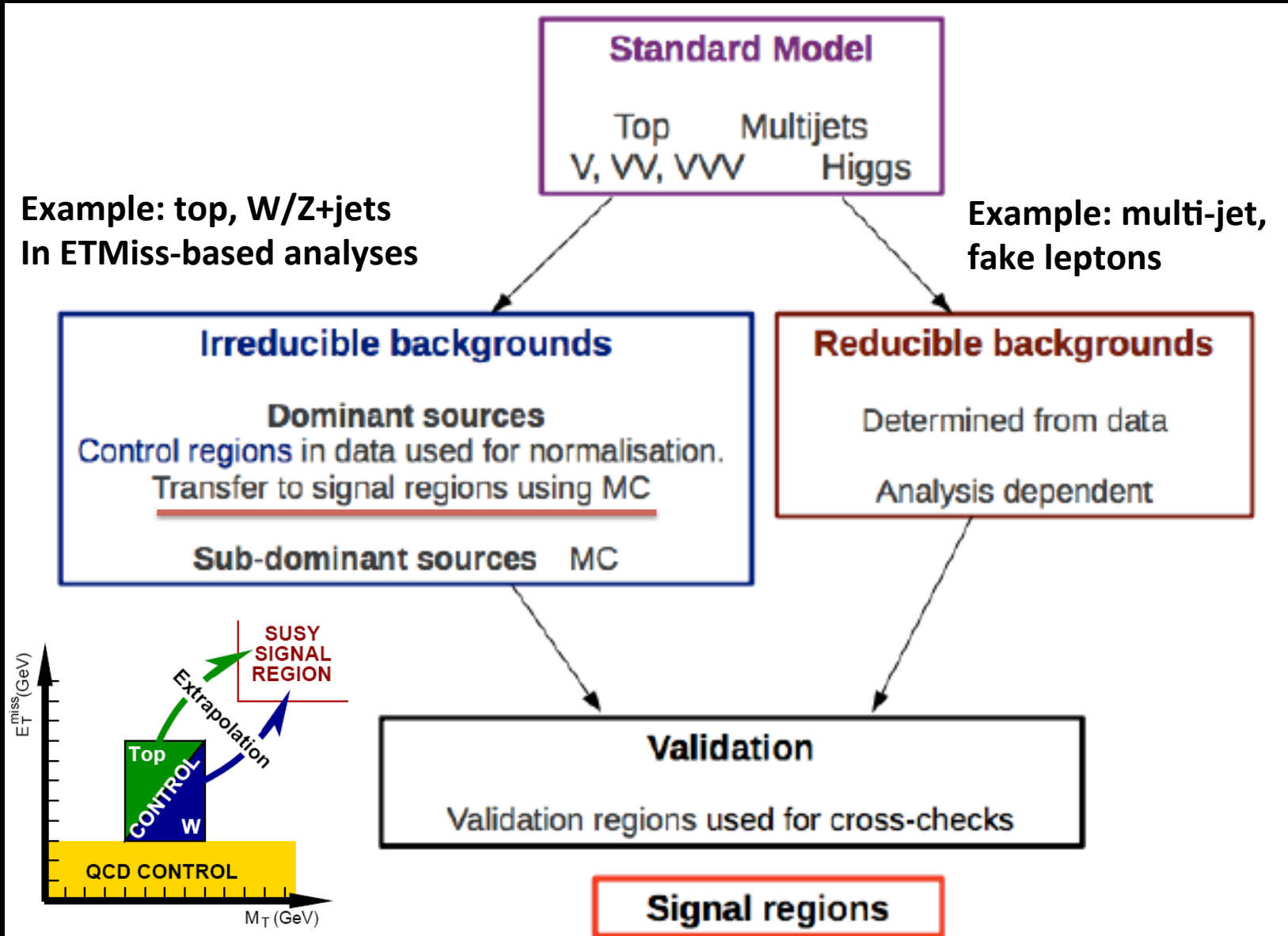


How to make your search...



1. Find good discriminant(s) \rightarrow signal region (blind it!)
2. Determine your SM background + related systematics
 1. As much as possible from data
 2. If taken from theory/simulations \rightarrow define control regions in data (orthogonal to your signal regions) to constrain the predictions with data (and to reduce systematics from models)
 3. Validate your predictions in regions close to the signal region (similar kinematics) where you do not expect new physics
 4. Convince your collaborators all is under control and open box
3. Use a sophisticated Likelihood fit to determine whether your data is statistically consistent with a background only hypothesis in the signal region taking into account correlations of systematic uncertainties, etc..
 1. Buy a ticket to Stockholm or compute exclusion limits @ 95% CL

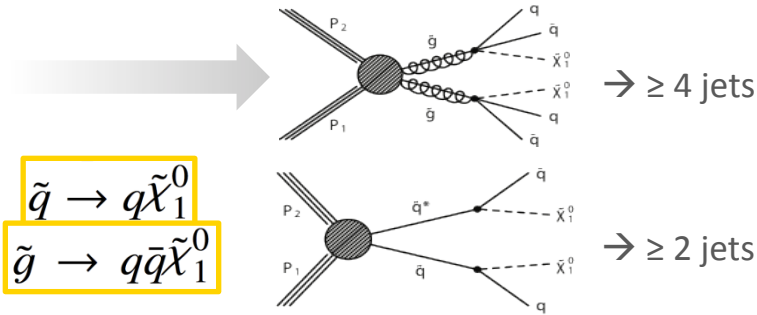
Background strategy



An example: 0-lepton signature

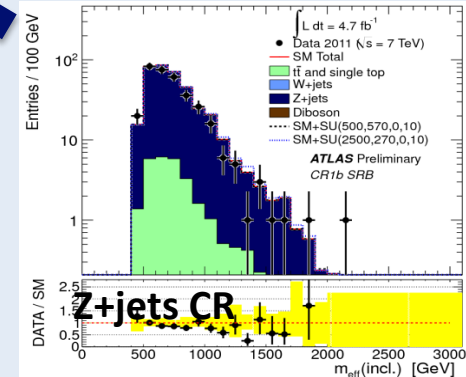
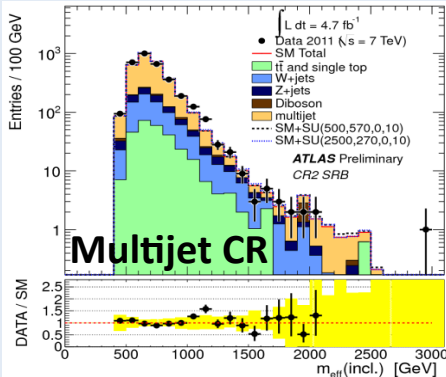
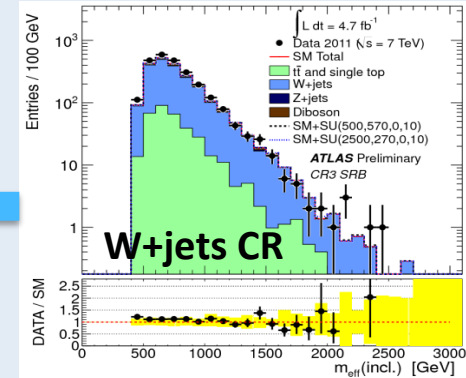
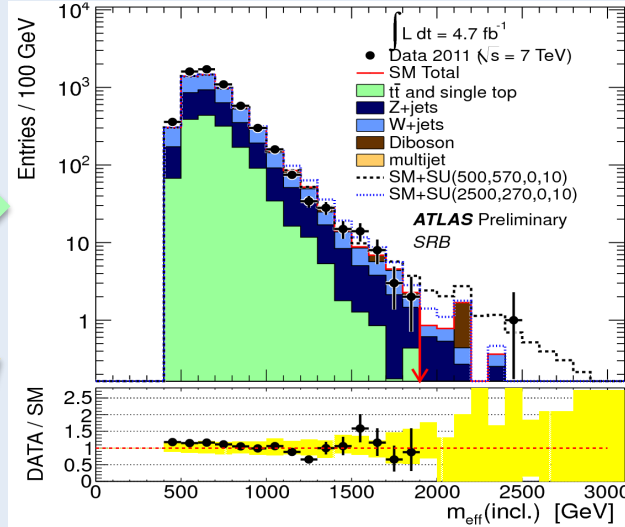
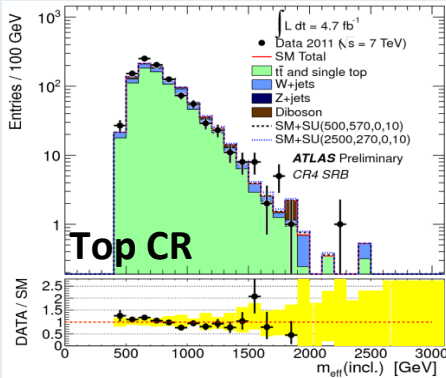
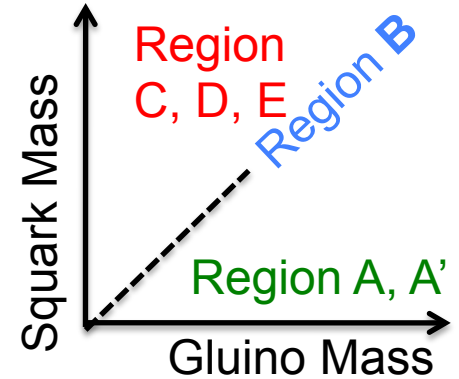
- Searches in inclusive jets + E_t^{miss} events

– from 2 (A) to 6 (E) jets



Expect significant
"effective mass"

$$\sum_{\text{jet}} E_{T,\text{jet}} + E_T^{\text{miss}}$$



Normalizations obtained in all CR and extrapolated to signal regions simultaneously by combined maximum likelihood fit

The mother of all fits...

- Combine all info in a global fit. Likelihood based on CR and SR (mutually exclusive)

$$L(n | \mu, b, \theta) = P_{SR} \times P_Z \times P_W \times P_{Top} \times P_{QCD} \times C_{syst}$$

b =background

θ = systematics treated as nuisance parameters with Gaussian

n =Number of observed events in data

μ = SUSY signal strength to be tested

- Each region described with a **Poisson p.d.f**
- Statistical and systematic uncertainties on the expected values included in the fit as nuisance parameters → typically constrained by a **Gaussian function** with width corresponding to the size of the uncertainty considered
 - correlations between these parameters are taken into account
- Inputs:** Transfer factors (c), N events for data in SR(s) and CR(b_j)

$$P_{SR} = P(n | \lambda_S(\mu, b, \theta)) = \mu \cdot c_{sR \rightarrow SR}(\theta) \cdot s + \sum_j c_{jR \rightarrow SR}(\theta) \cdot b_j$$

$$P_i = P(n | \lambda_i(\mu, b, \theta)) = \mu \cdot c_{sR \rightarrow iR}(\theta) \cdot s + \sum_j c_{jR \rightarrow iR}(\theta) \cdot b_j$$

$\lambda(\mu, b, \theta)$ = expected number of events

Statistical facts

Notes on statistics

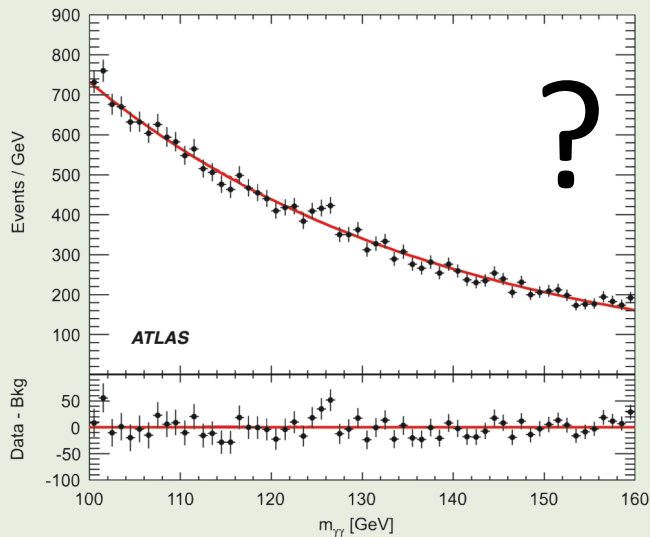
Using Higgs discovery for illustration

Notes on Statistical Significance

PHYSICAL REVIEW LETTERS™

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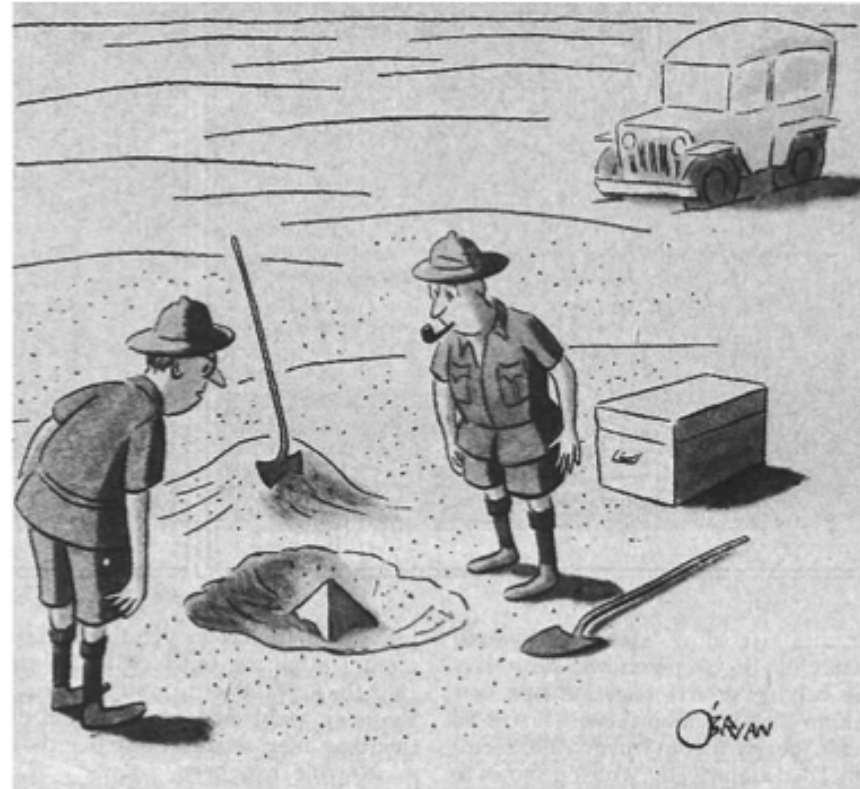
Articles published week ending 16 MARCH 2012



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APS
physics

Volume 108, Number 11



"This could be the discovery of the century. Depending, of course, on how far down it goes."

Likelihood ratio

$$L(\mu, \theta) = f_b \phi_b(m_{\gamma\gamma}) + f_s \phi_s(m_{\gamma\gamma})$$

$$f_s \propto \mu$$

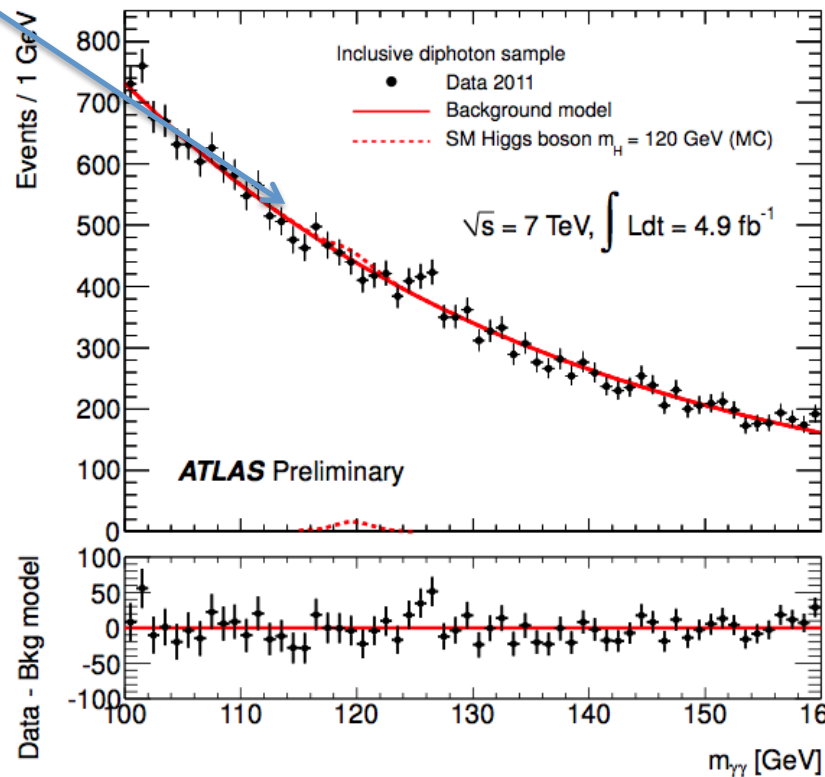
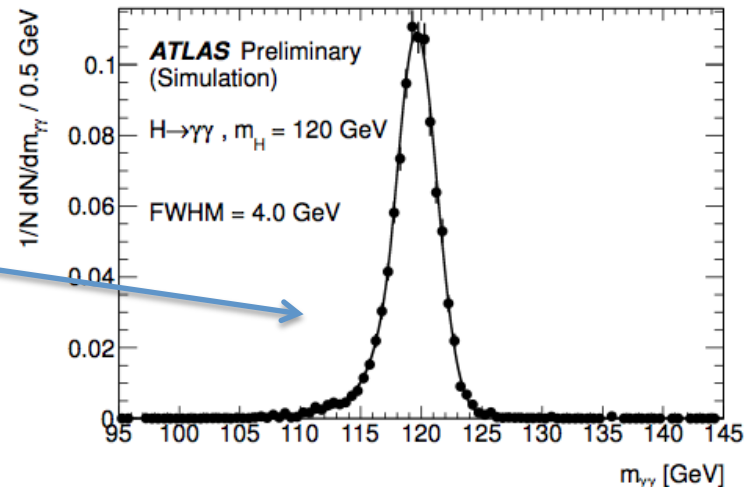
$$n_s = \mu \sigma_s^{\text{visible}}$$

Nuisance parameters

$$\lambda_\mu = \lambda(\mu, \theta) = \frac{L(\mu, \hat{\theta}(\mu))}{L(\hat{\mu}, \hat{\theta})}$$

$$q_\mu = -2 \ln \lambda_\mu$$

$$p_\mu = \int_{q_{\text{obs}}}^{\infty} f(q_\mu | \mu) dq_\mu$$



Only background ?

Test of “null” hypothesis of no signal

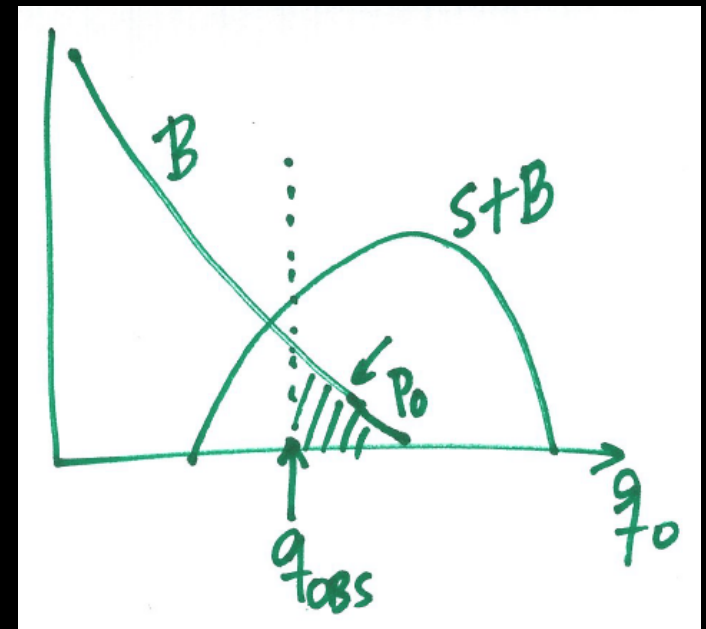


If a real signal appears ... $p_0 \rightarrow 0$
(once $p_0 < 2.87 \times 10^{-7} \rightarrow$ Discovery)

$$\lambda_0 = \lambda(0, \theta) = \frac{L(0, \hat{\theta}(0))}{L(\hat{\mu}, \hat{\theta})}$$

$$q_0 = -2 \ln \lambda_0$$

$$p_0 = \int_{q_{obs}}^{\infty} f(q_0 | 0) dq_0$$



Only background ?

Test of “null” hypothesis of no signal

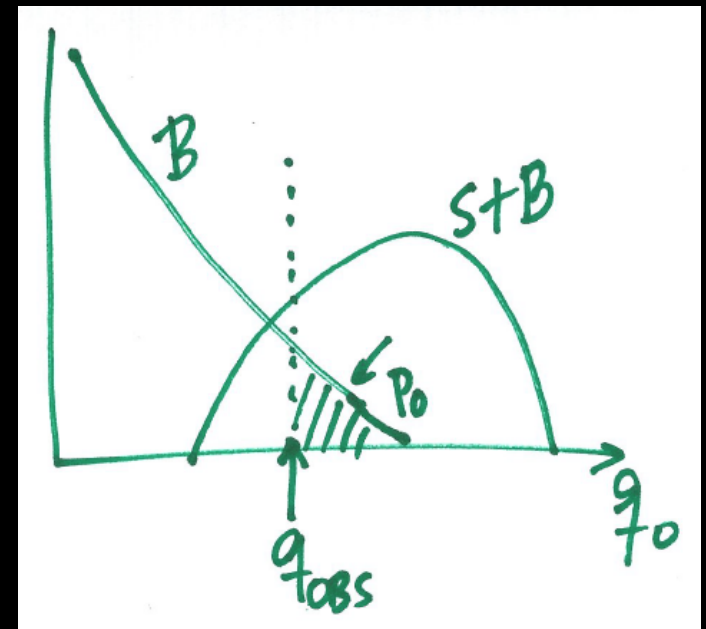


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$$q_0 = -2 \ln \lambda_0$$

$$p_0 = \int_{q_{obs}}^{\infty} f(q_0 | 0) dq_0$$



CL_s

(do not exclude your signal...)

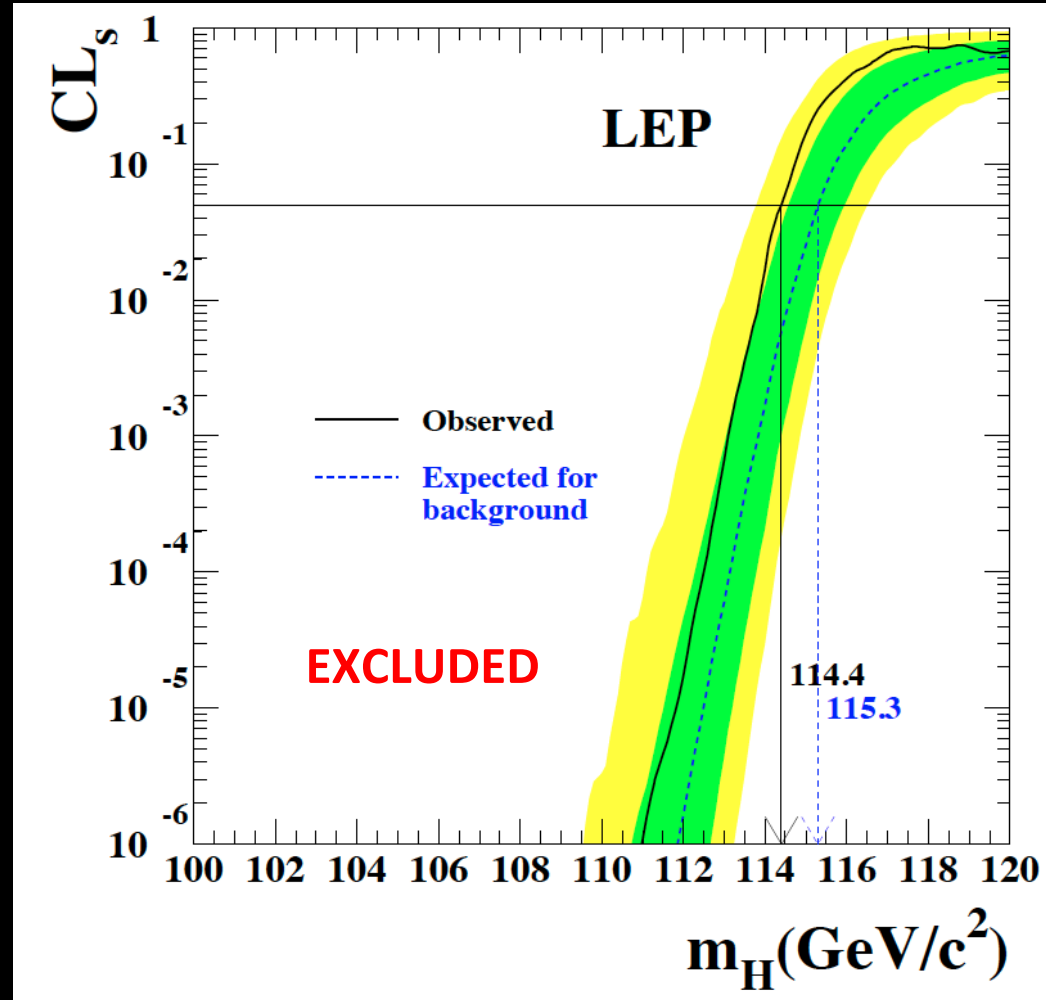


$$CL_s = \frac{p_s}{1 - p_b}$$

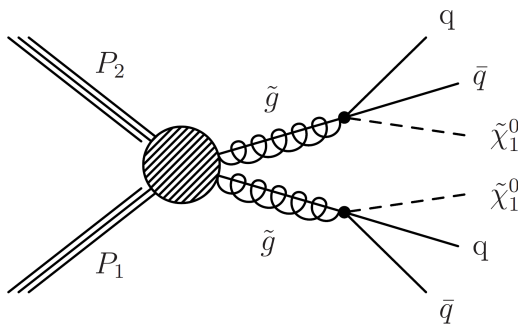
In the case of very small signals (limited sensitivity) the use of p_s to exclude signals can lead to false exclusions if the data fluctuates down....

In these cases it is better to use CLs ... which is conservative in the exclusion

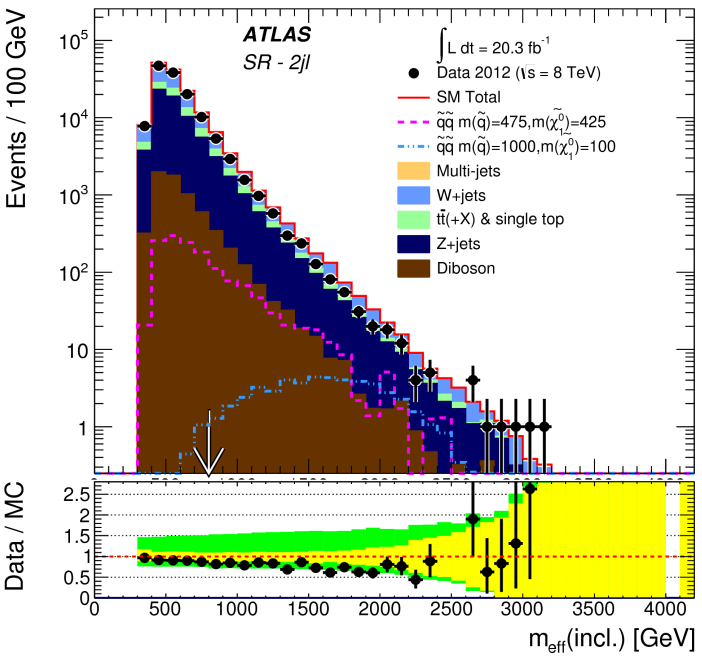
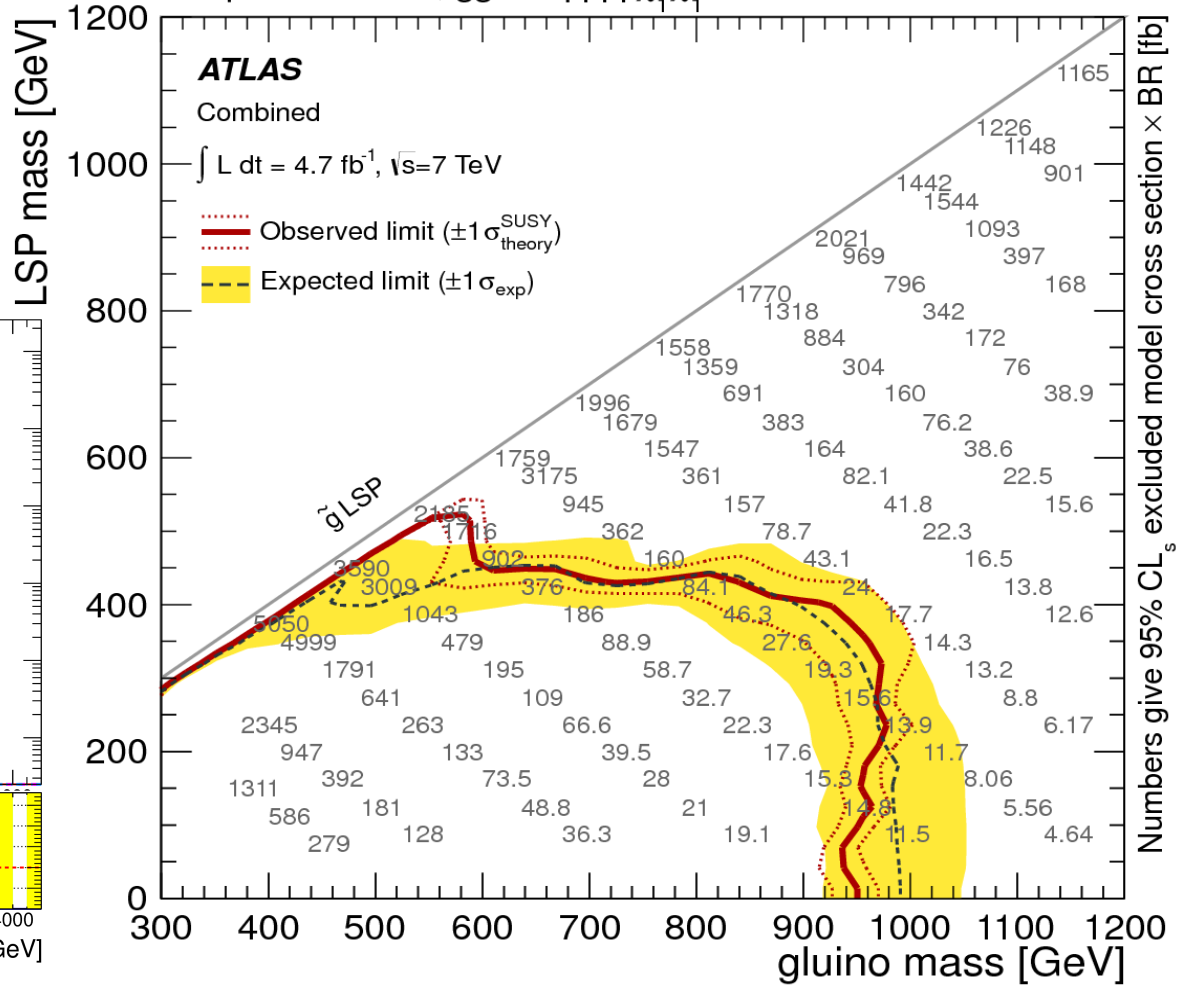
If $CL_s < 0.05 \rightarrow$ excluded at 95% CL



Typical SUSY exclusion plot..



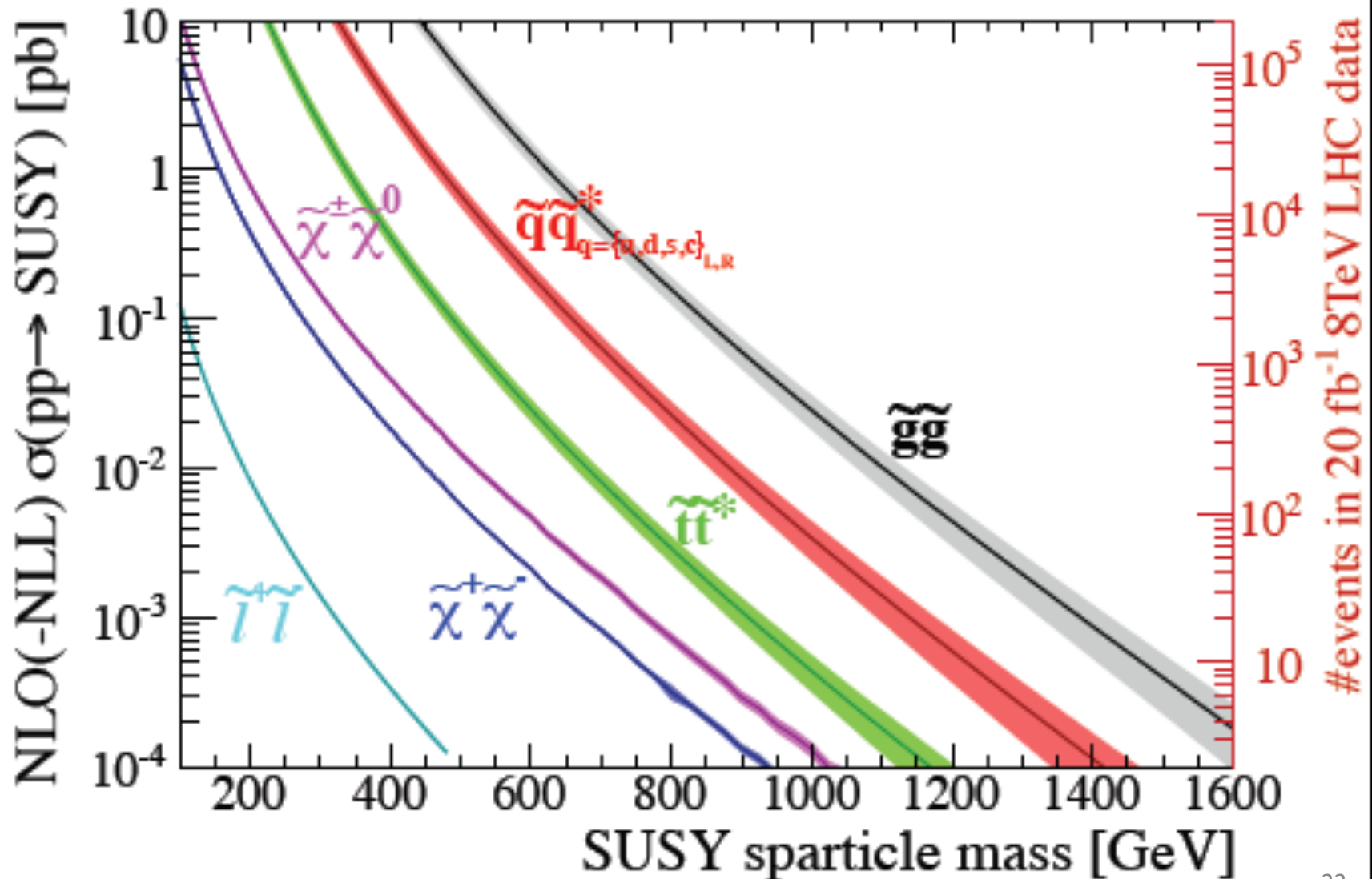
Simplified model, $\tilde{g}\tilde{g} \rightarrow q\bar{q}q\bar{q} \tilde{\chi}_1^0 \tilde{\chi}_1^0$



Expected: use SM prediction as data (yellow band reflects uncertainties in SM prediction)
Observed: what the data tells you (dashed band depends on the model uncertainties)
Numbers inside: 95% CL signal exclusion (if your signal is larger than this.. You excluded it)

SUSY Cross Sections @ LHC (8 TeV)

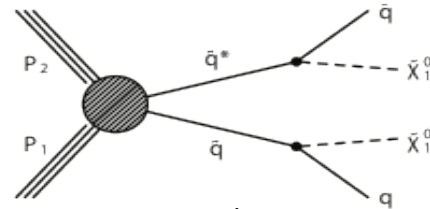
LPCC SUSY σ WG



Inclusive Searches

Multiple-jets and large E_T^{miss}

[JHEP09\(2014\)176](#)

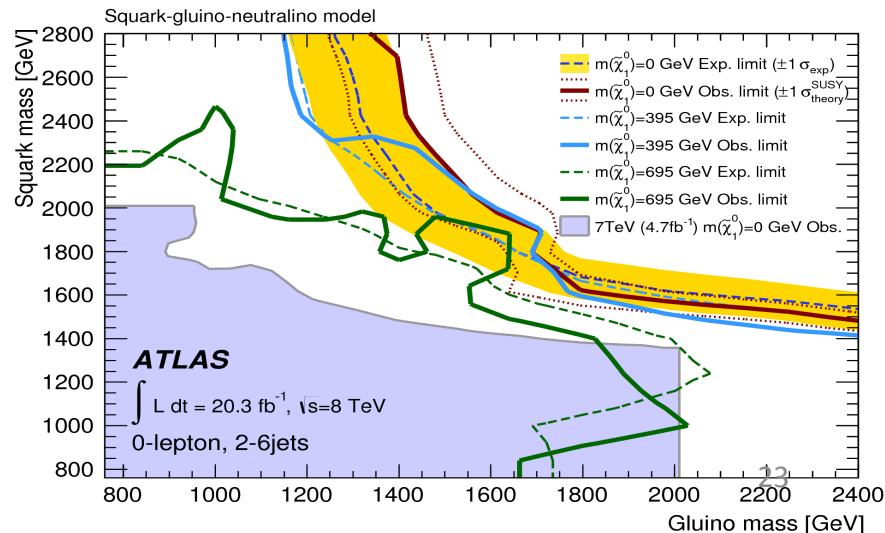
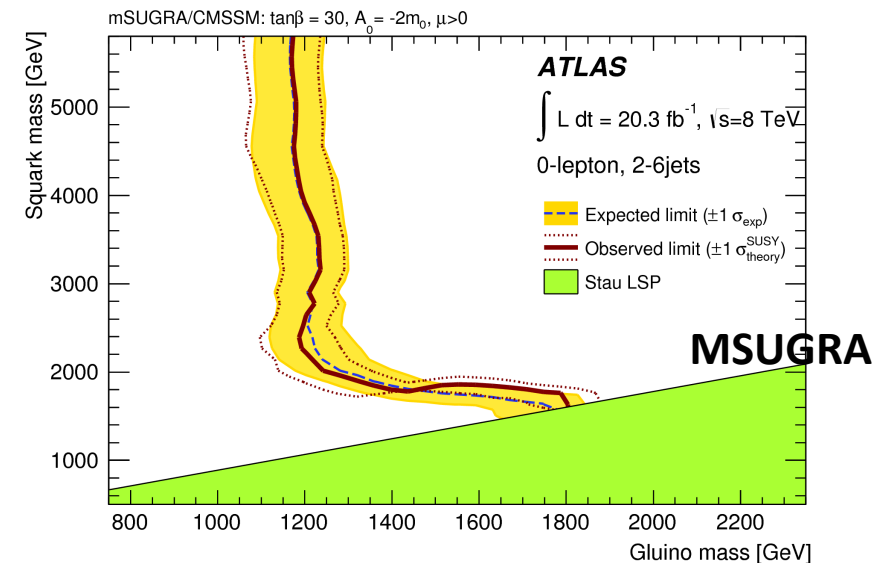
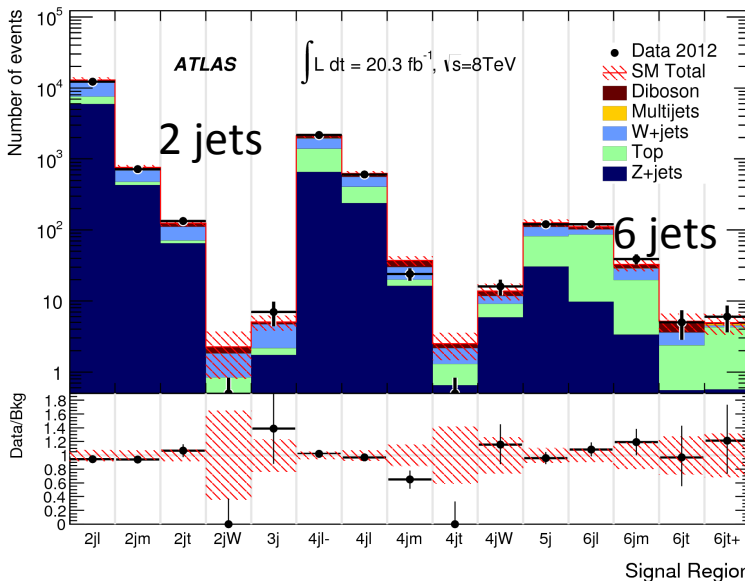
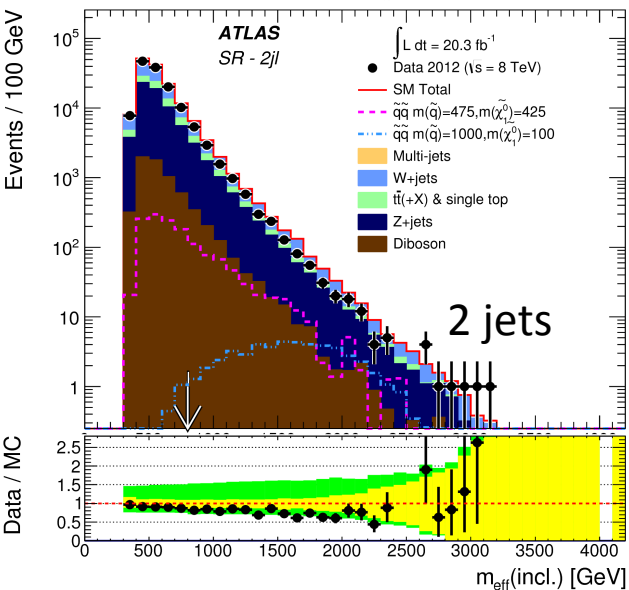
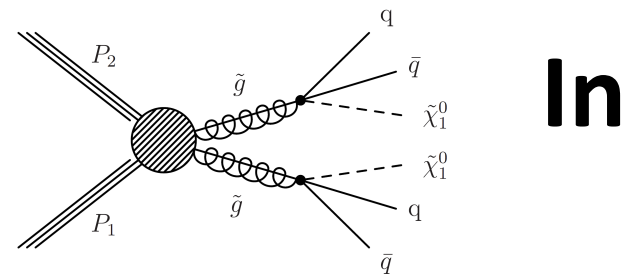


$$m_{\text{eff}} = E_T^{\text{miss}} + H_T$$

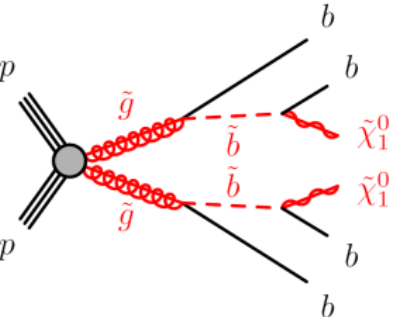
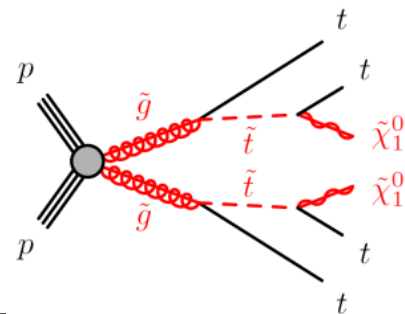
$$H_T = \sum_{\text{jets}} p_T$$

Background from
1 lepton b-tag/veto
& photon CRs with
similar jet and
kinematic selections
as in SRs

**SIMPLIFIED
SCENARIO**



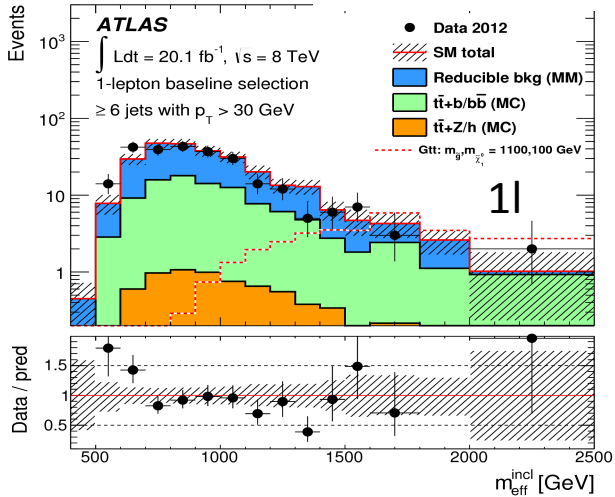
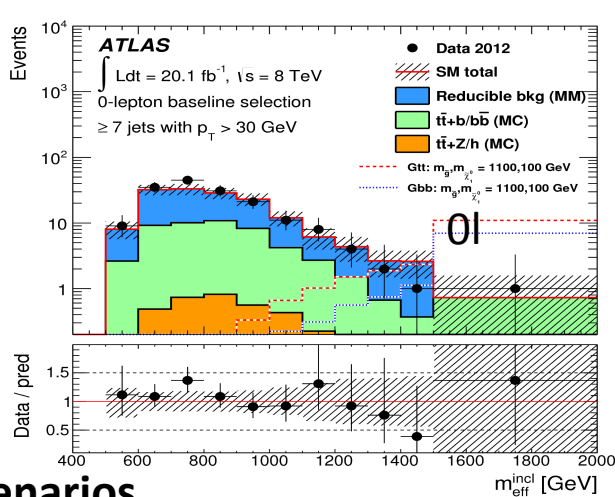
Glauino-mediated Stop/Sbottom Production



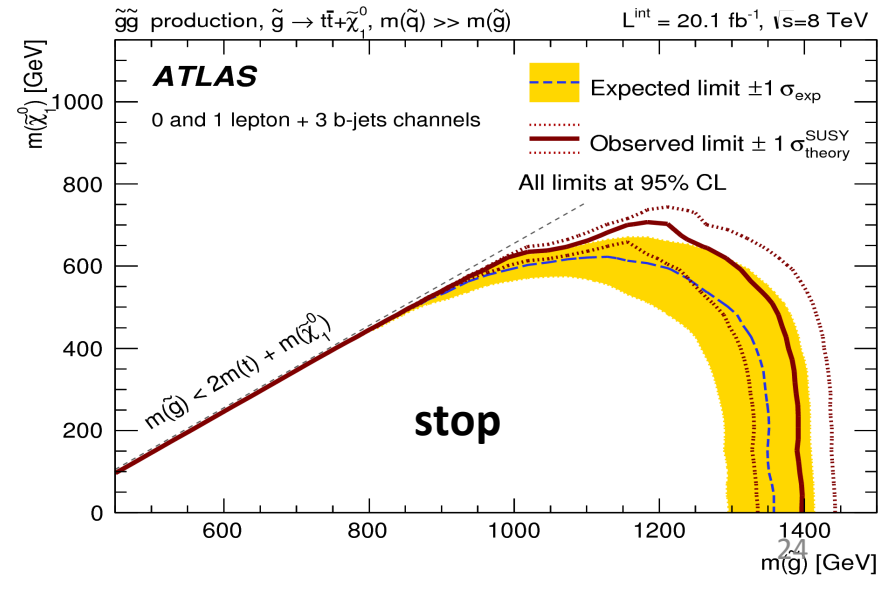
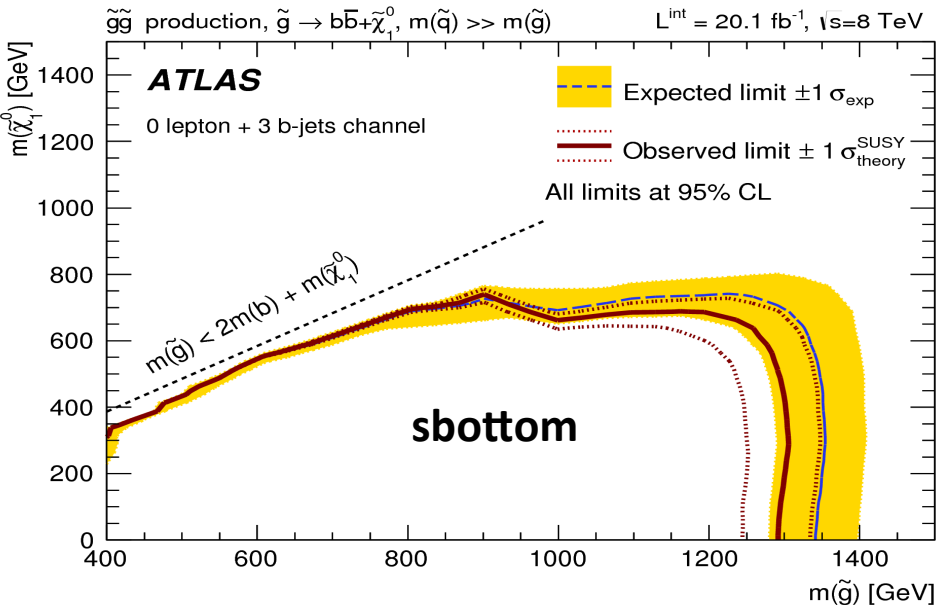
0/1 lepton + 3b-jets + E_t^{miss}

Background dominated by $tt+bb$, $tt+V$ and $tt+fakes$

most powerful for very heavy gluinos



Interpreted in simplified scenarios

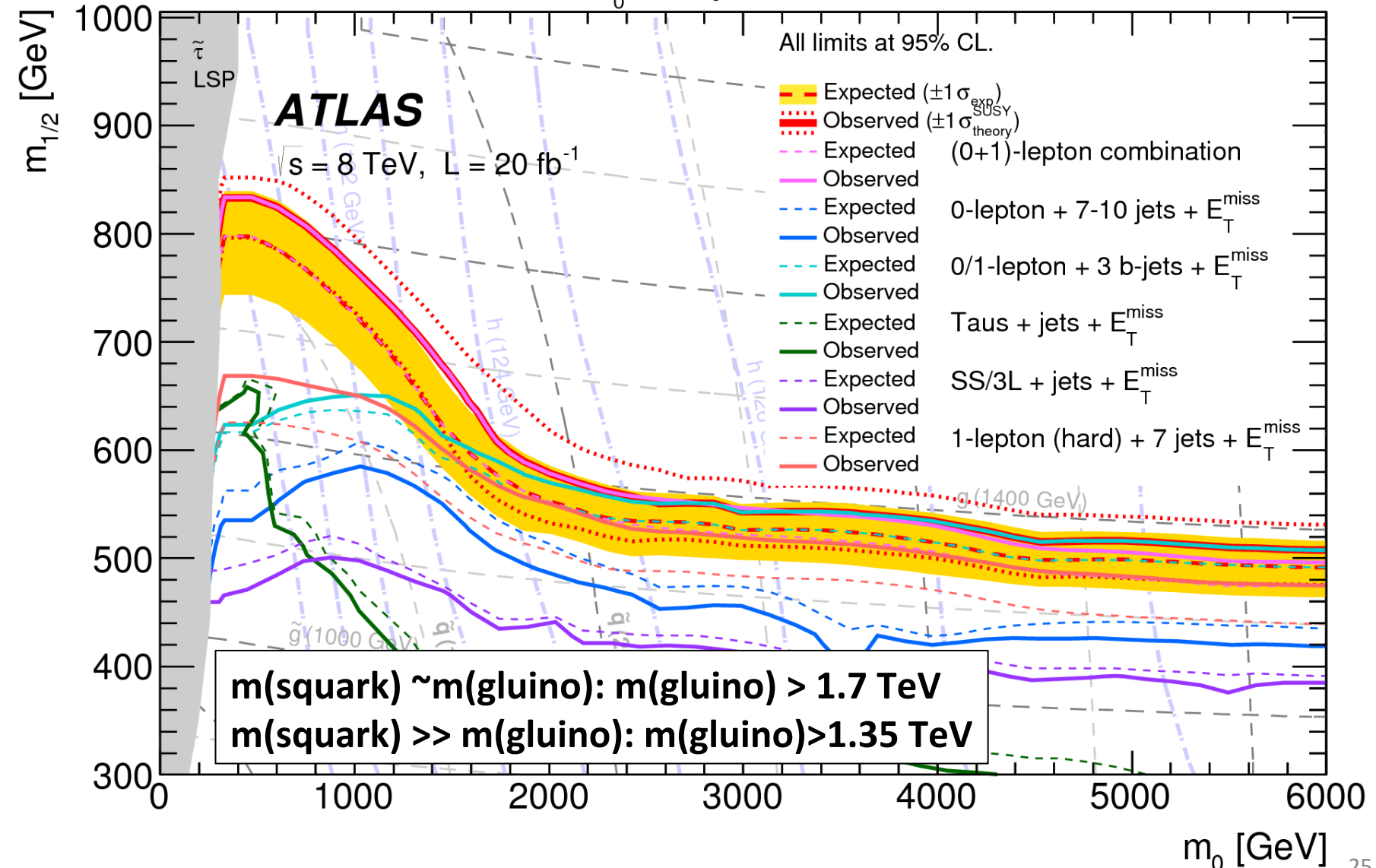


MSUGRA Scenario

[arXiv:1507.05525](https://arxiv.org/abs/1507.05525)

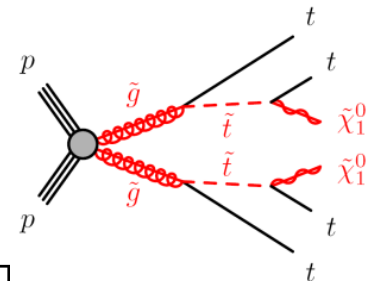
Run-1 summary on inclusive squark and gluino searches

MSUGRA/CMSSM: $\tan(\beta) = 30$, $A_0 = -2m_0$, $\mu > 0$

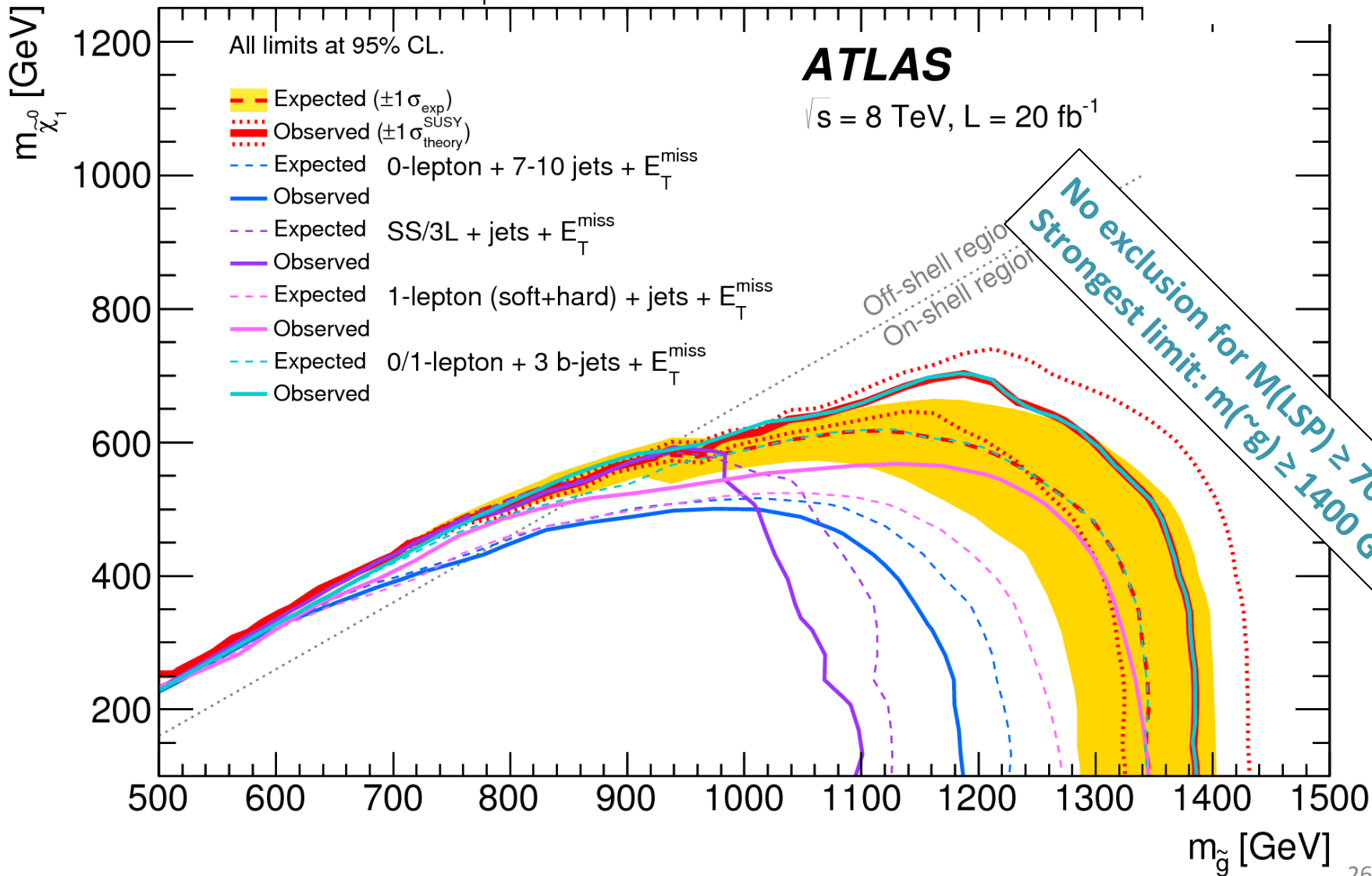


Part of the model plane accommodates a lightest neutral scalar Higgs boson mass of 125 GeV

Simplified model

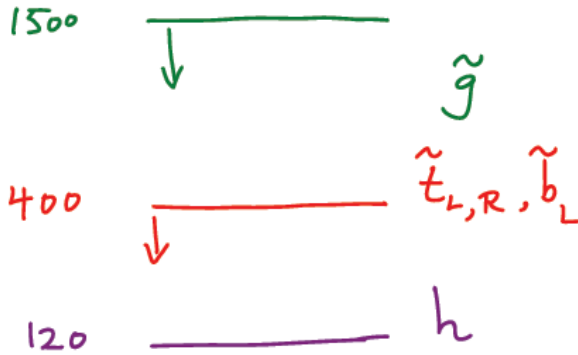


$\tilde{g}\tilde{g}$ production, $\tilde{g} \rightarrow tt^{(*)}\tilde{\chi}_1^0$; $m(\tilde{t}) \gg m(\tilde{g})$, including up to five-body decays



“Natural SUSY 2012”

Compulsory Natural SUSY



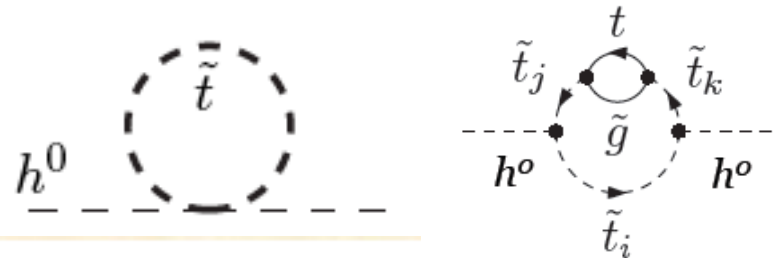
Unavoidable tunings: $\left(\frac{400}{m_{\tilde{t}}}\right)^2, \left(\frac{4m_{\tilde{t}}}{M_{\tilde{g}}}\right)^2$

N. Arkani-Hamed talk at CERN Oct. 2012

→ Light higgsinos

→ Light stop ($t_1 < 1 \text{ TeV}$)

→ Light gluinos ($< 1-2 \text{ TeV}$)



$$\frac{m_H^2}{2} = -|\mu|^2 + \dots + \delta m_H^2$$

$$\delta m_H^2|_{\text{stop}} \cong -\frac{3y_t^2}{8\pi^2} (m_{Q_3}^2 + m_{U_3}^2 + |A_t|^2) \ln\left(\frac{\Lambda}{\text{TeV}}\right)$$

$$\delta m_H^2|_{\text{gluino}} \cong -\frac{2y_t^2}{\pi^2} \left(\frac{\alpha_s}{\pi}\right) |M_3|^2 \ln^2\left(\frac{\Lambda}{\text{TeV}}\right)$$

$$\begin{pmatrix} \tilde{t}_1 \\ \tilde{t}_2 \end{pmatrix} = \begin{pmatrix} \cos \theta_t & \sin \theta_t \\ -\sin \theta_t & \cos \theta_t \end{pmatrix} \begin{pmatrix} \tilde{t}_L \\ \tilde{t}_R \end{pmatrix}$$

**One light stop and sbottom
....rest of sparticles can be
decoupled....**

$$\begin{pmatrix} \tilde{t}_L \\ \tilde{b}_L \end{pmatrix} \quad \tilde{t}_R \quad \tilde{b}_R$$

(same weak isospin multiplet)

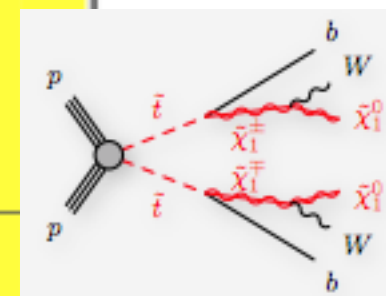
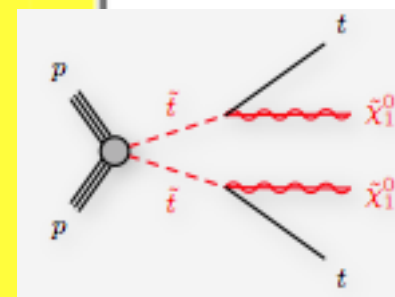
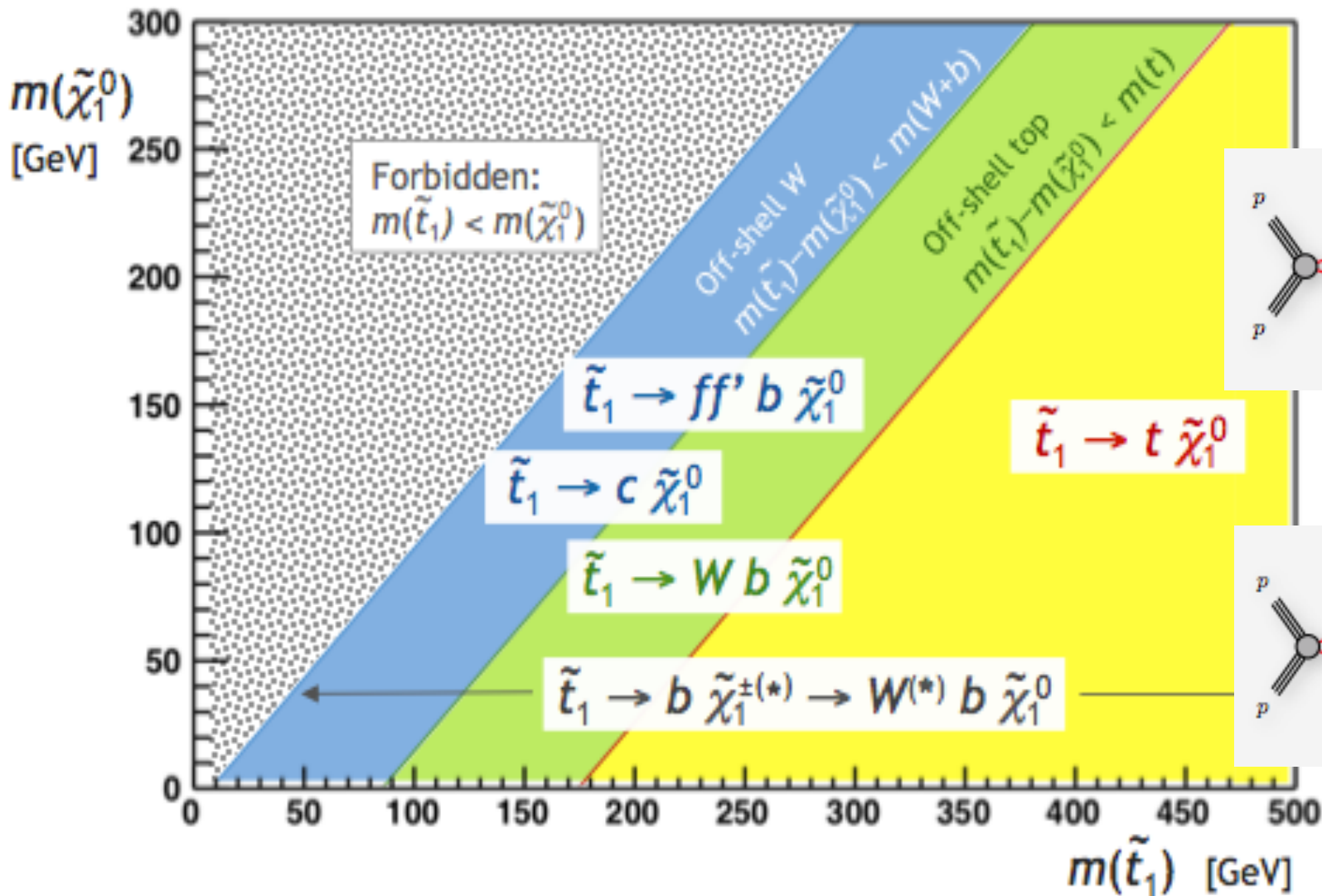
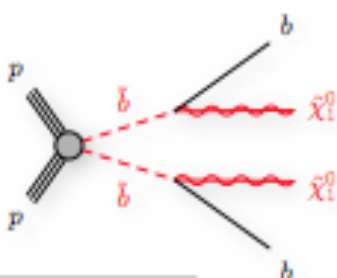
Direct Stop/Sbottom

In the scenario with TeV gluinos / squarks (1st/2nd generations)
 All the attention is put now in searches for stop/sbottom
 Multiple channels according to the decays

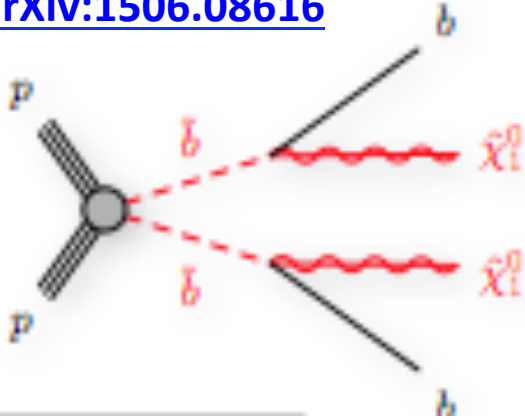
$$\tilde{b}_1 \rightarrow b \tilde{\chi}_1^0$$

$$\tilde{b}_1 \rightarrow t \tilde{\chi}_1^\pm$$

$$\tilde{b}_1 \rightarrow b \tilde{\chi}_2^0 \rightarrow b h(Z) \tilde{\chi}_1^0$$



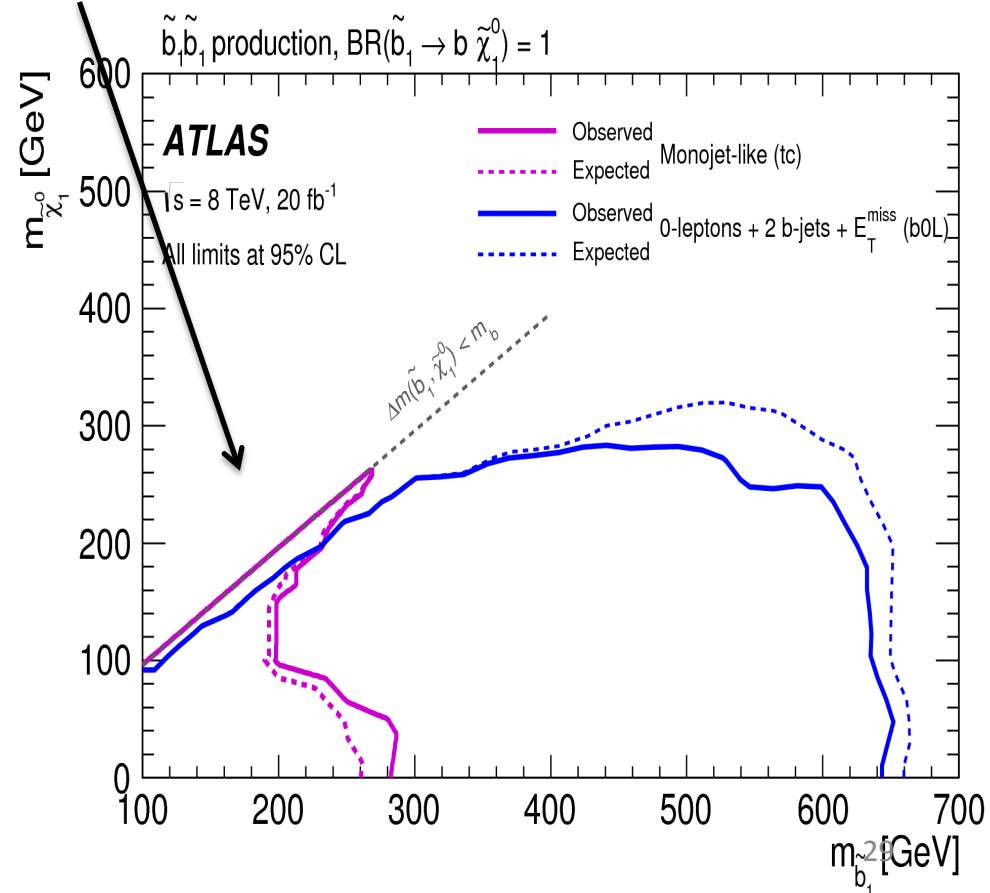
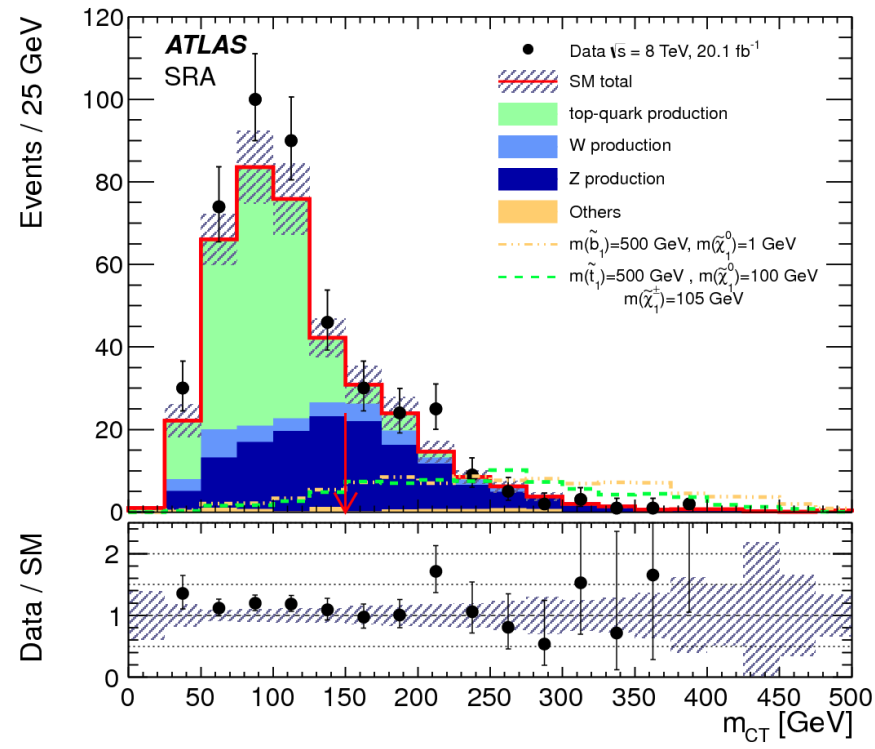
Sbottom Direct Production



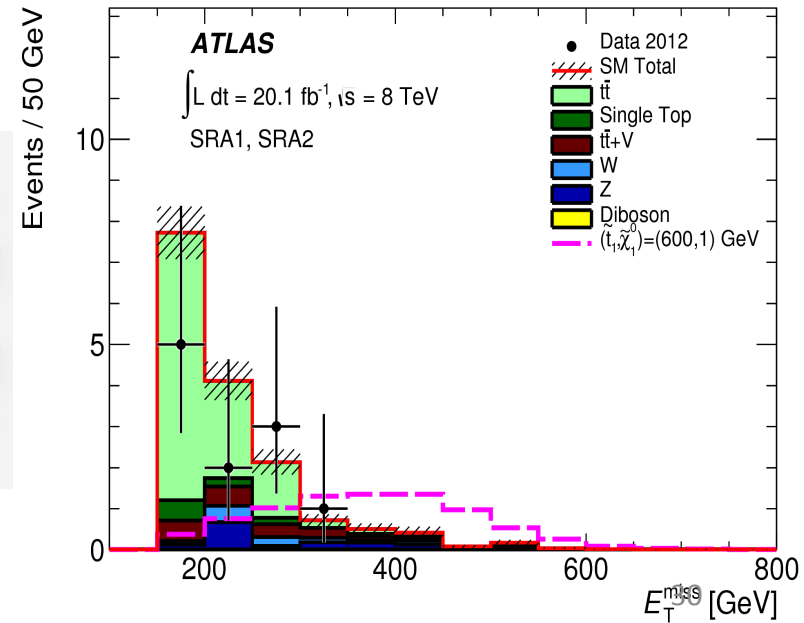
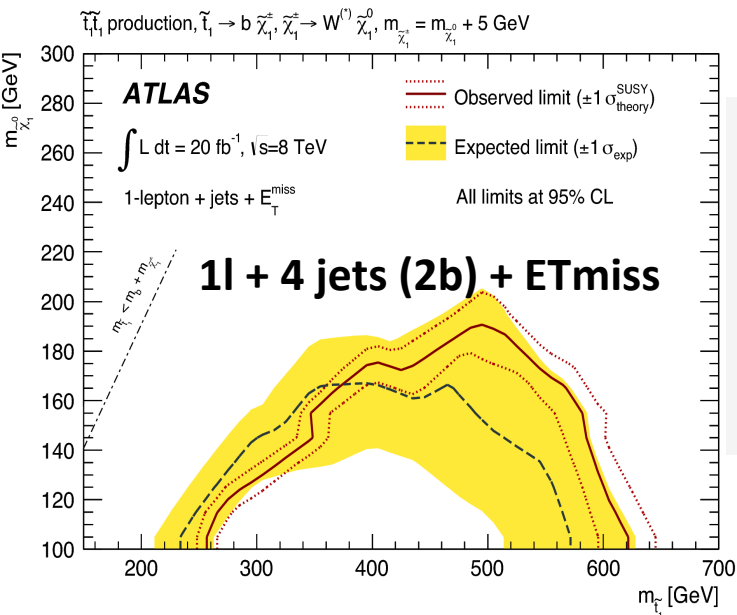
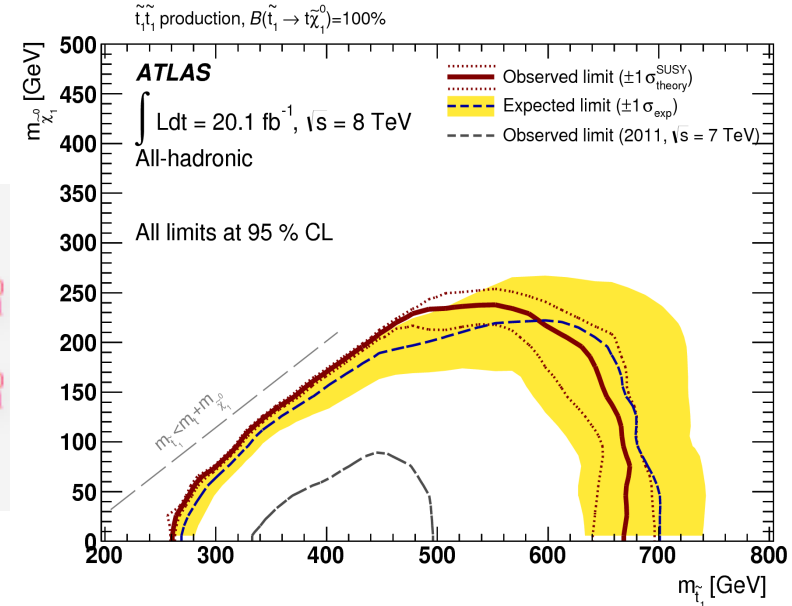
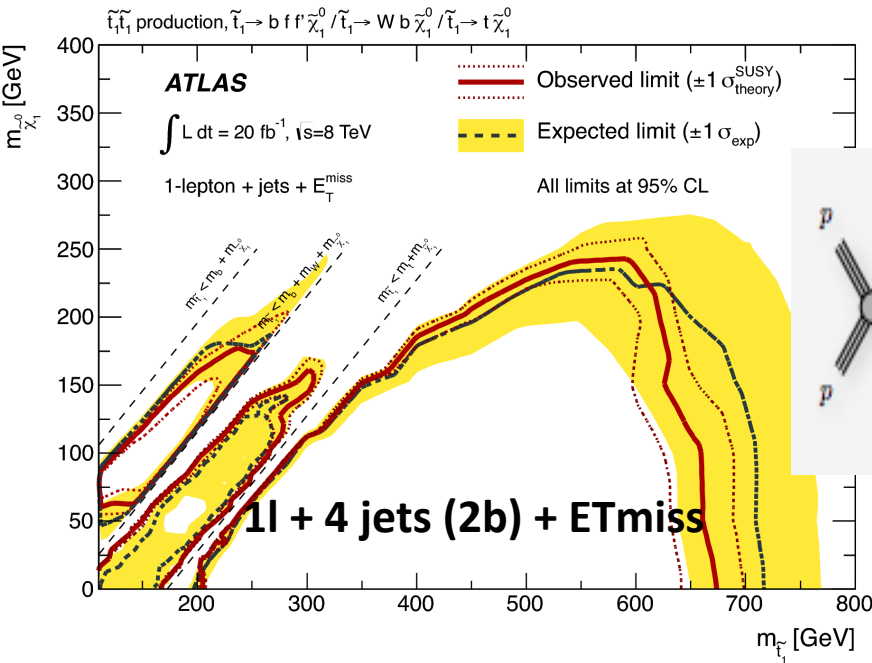
Large E_T^{miss} and 2 b-jets
Discriminating variable MCT

Additional selections to target also
compressed scenarios (assisted with ISR jets)

$$m_{\text{CT}}^2(v_1, v_2) = [E_T(v_1) + E_T(v_2)]^2 - [\mathbf{p}_T(v_1) - \mathbf{p}_T(v_2)]^2$$



Direct Stop



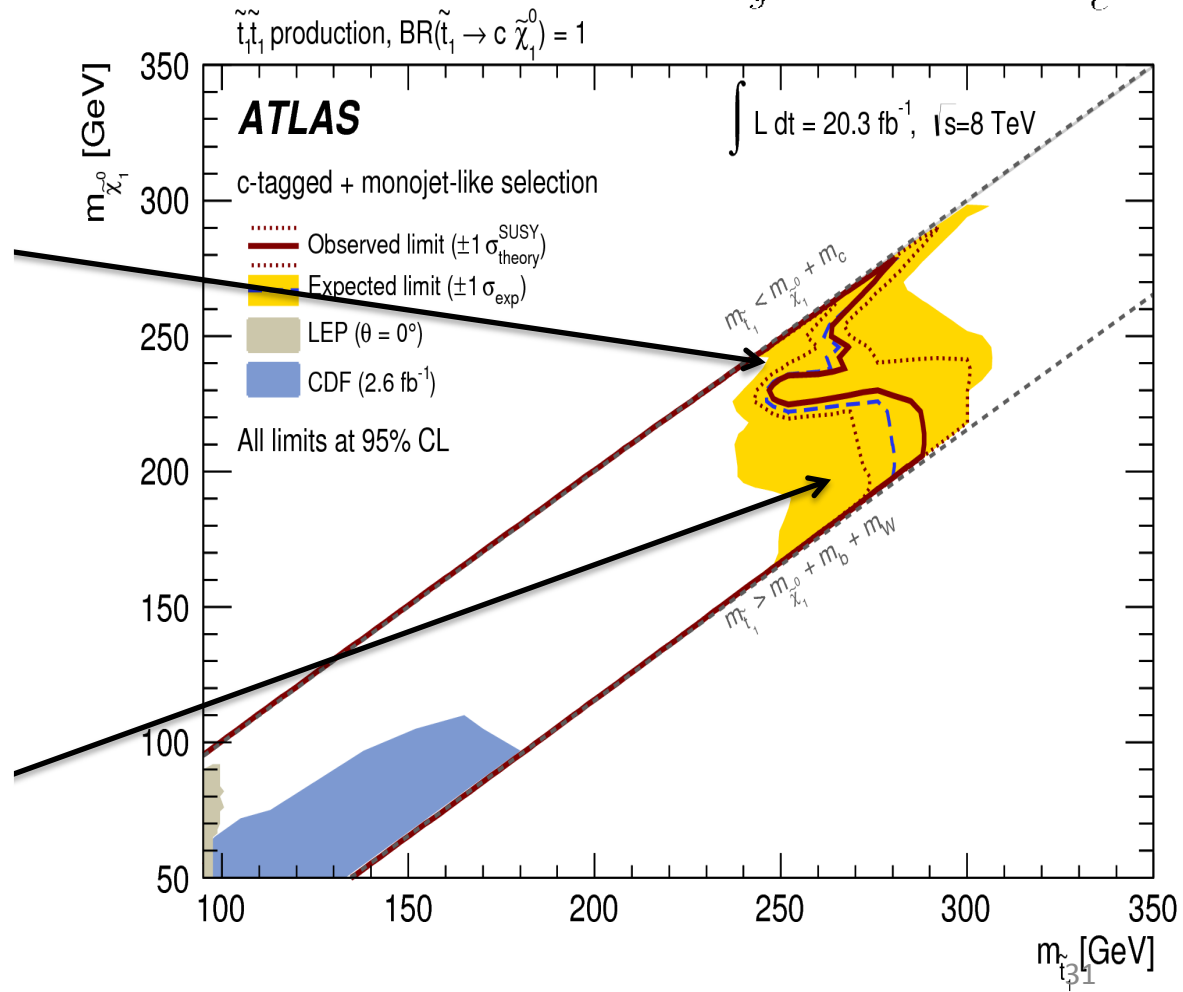
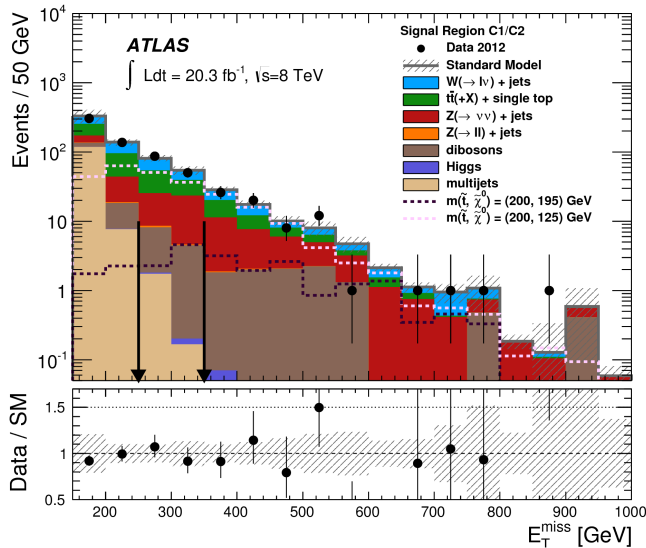
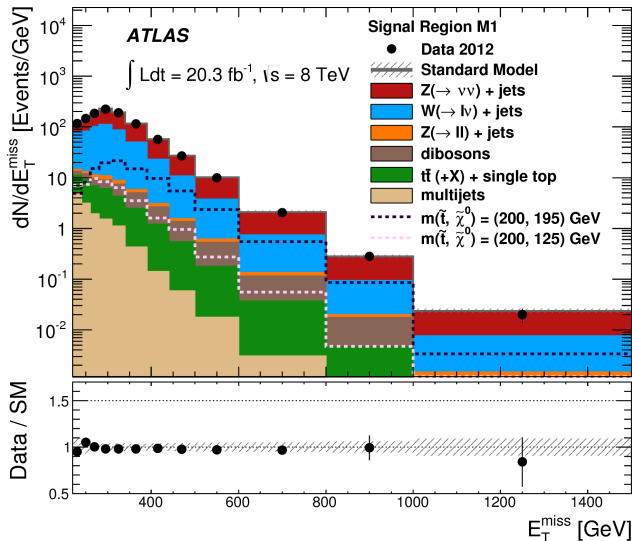
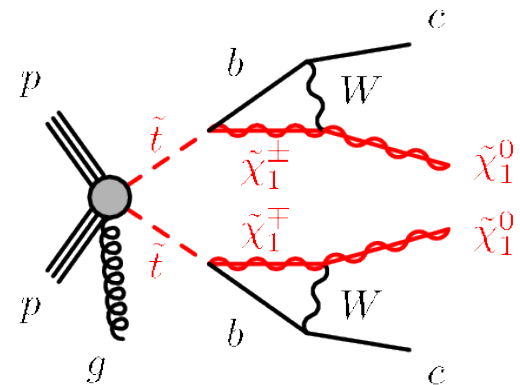
Direct Stop

Two selections:

Monojet for compressed scenario

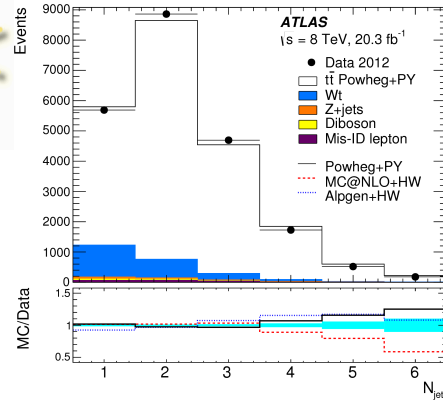
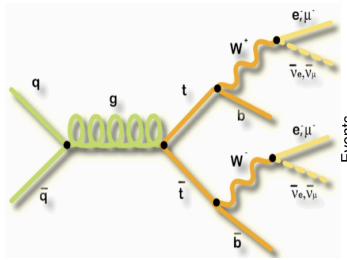
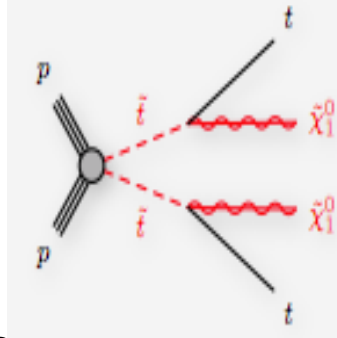
Large E_T^{miss} and c-jets

Very light stop

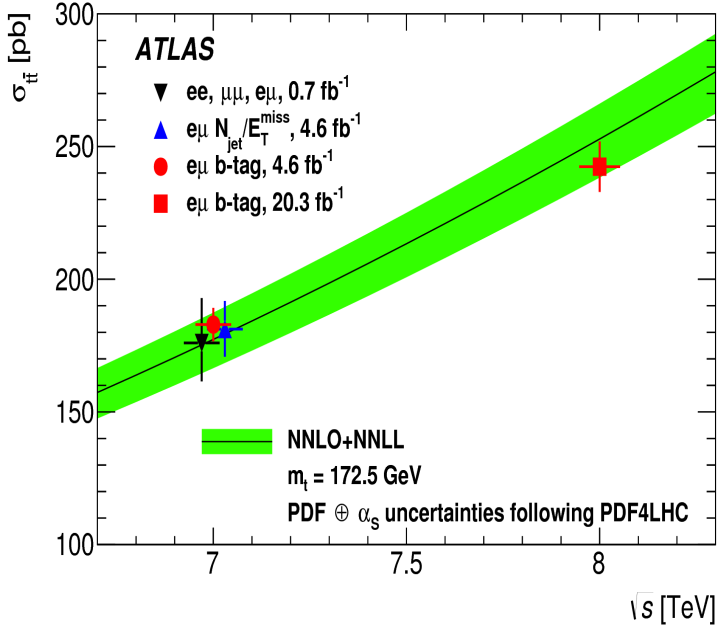


Further stop constrains

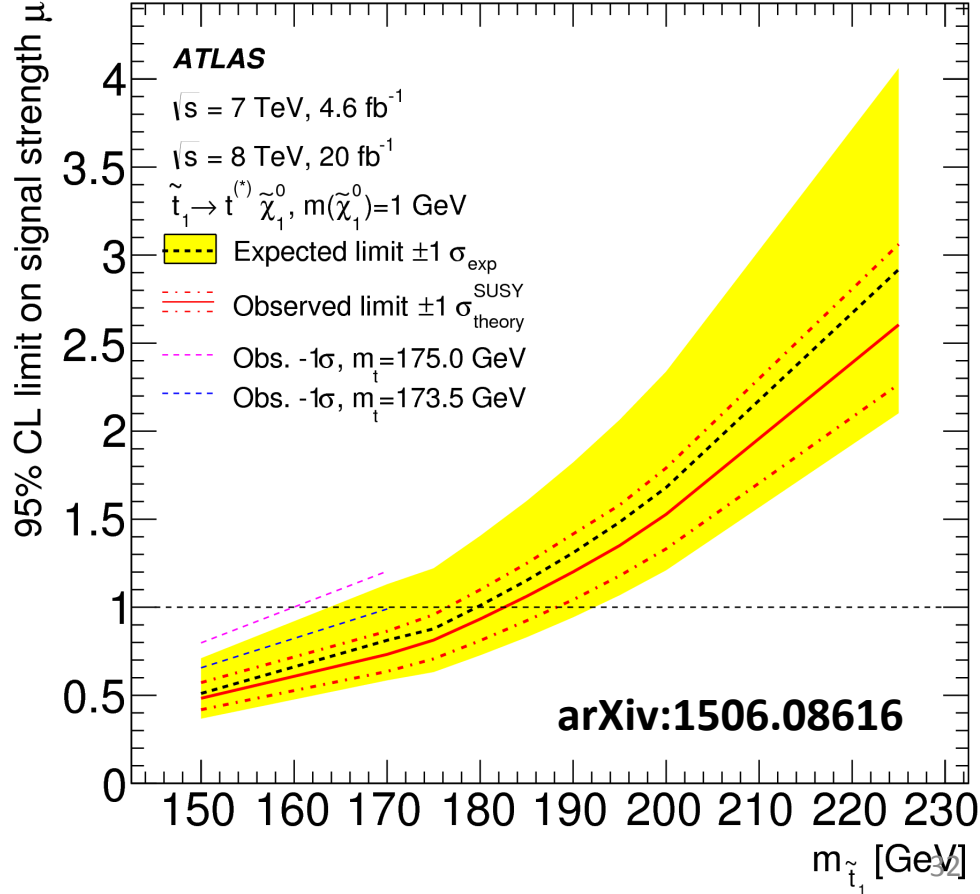
(using the top cross section measurements)



Dilepton (e-μ) channel with exactly one and two b-jets



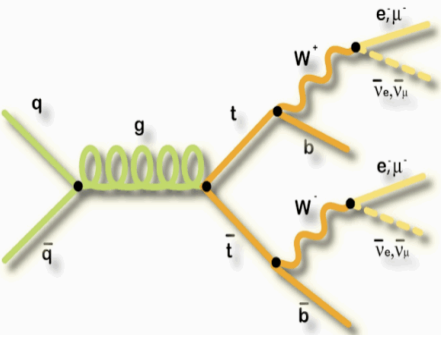
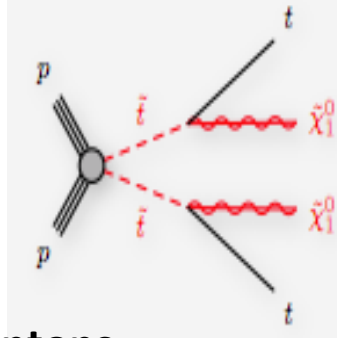
For stop masses close to the top mass the separation of SUSY signal and background is very challenging → use top cross section



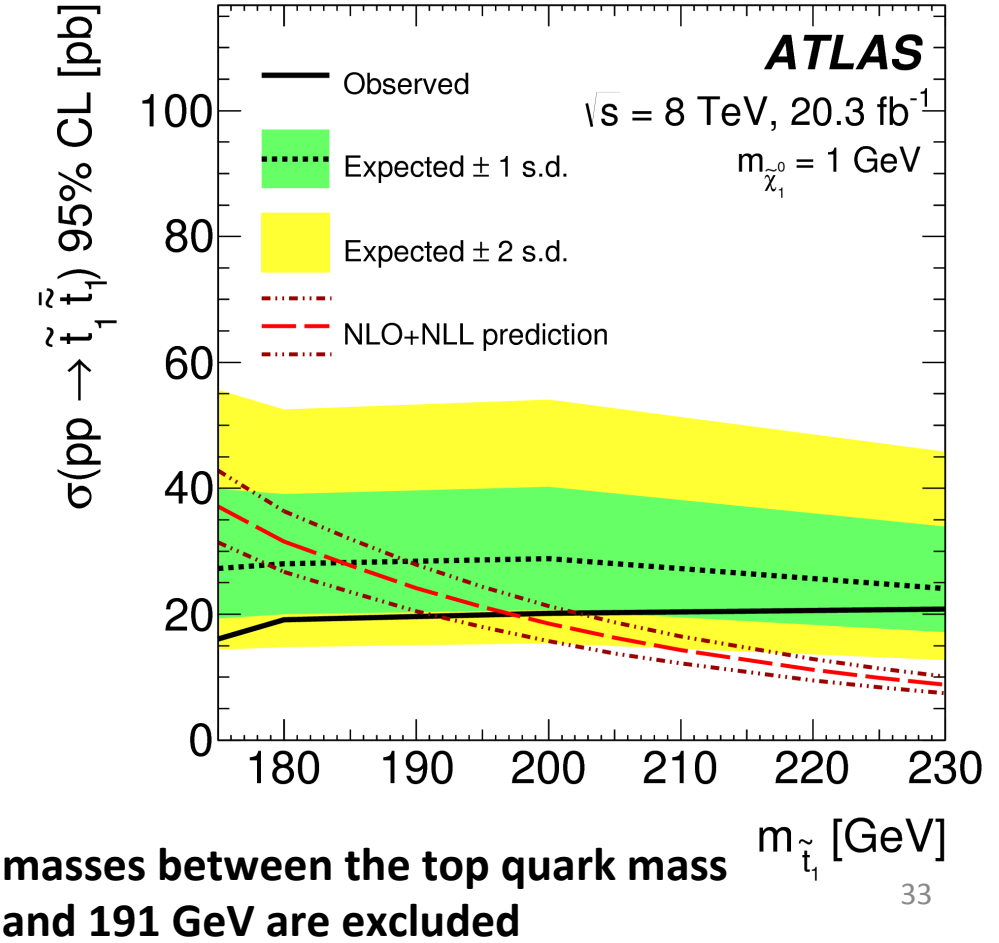
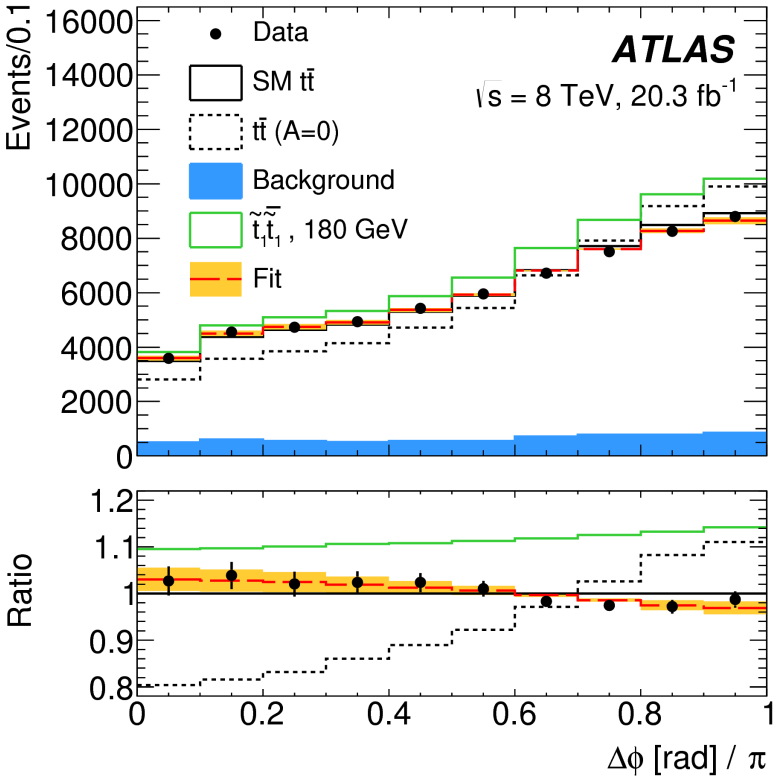
arXiv:1506.08616

Further stop constrains

(using the top-(anti)top spin correlations)

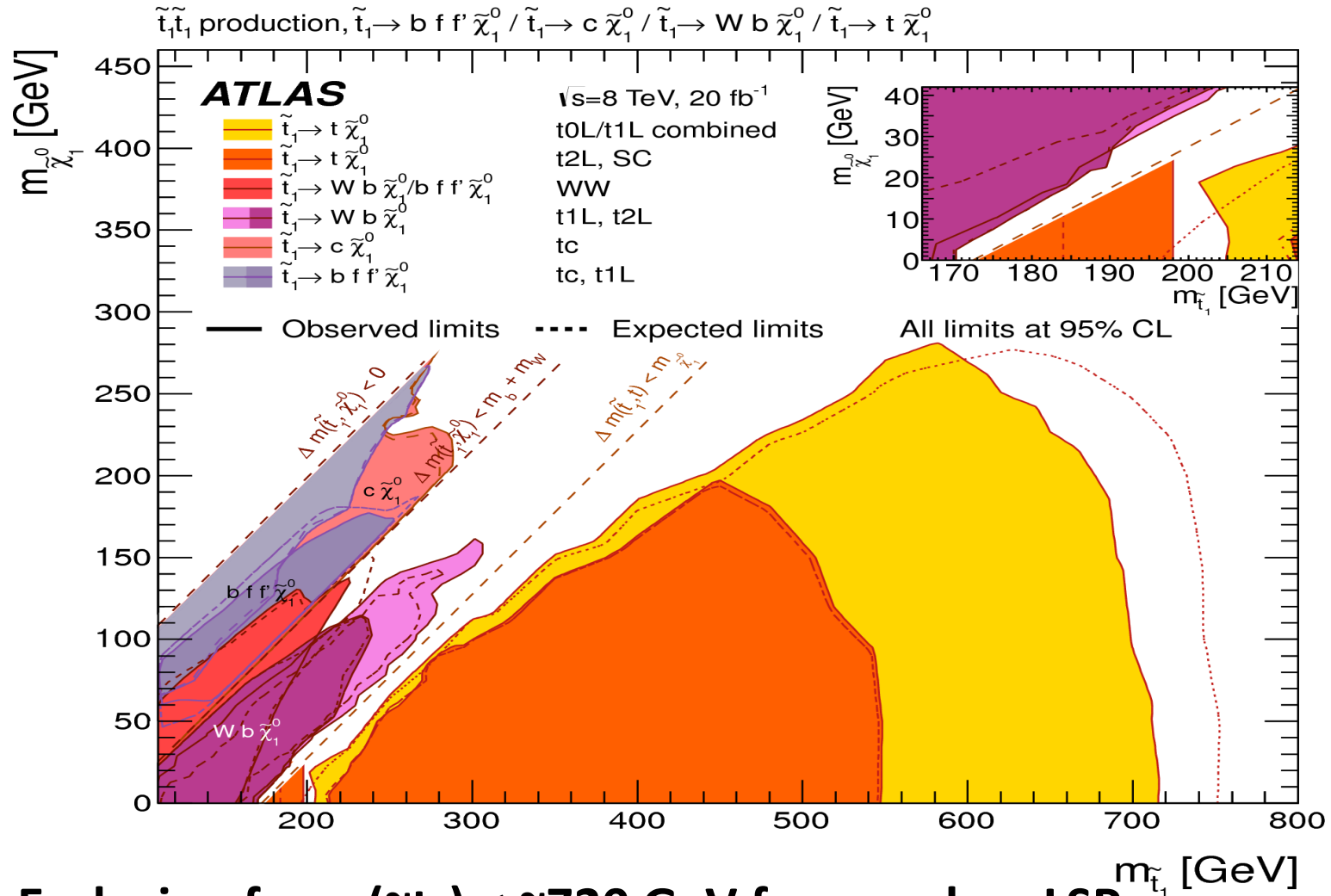


using azimuthal separation of leptons from top-(anti)top decays in the final state



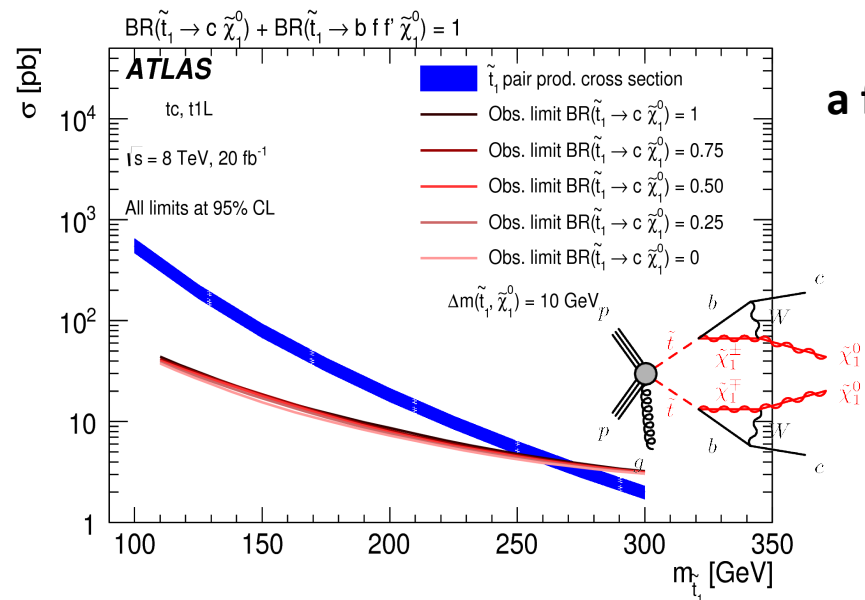
Summary Searches for Stop

(different mass hierarchies, simplified models)

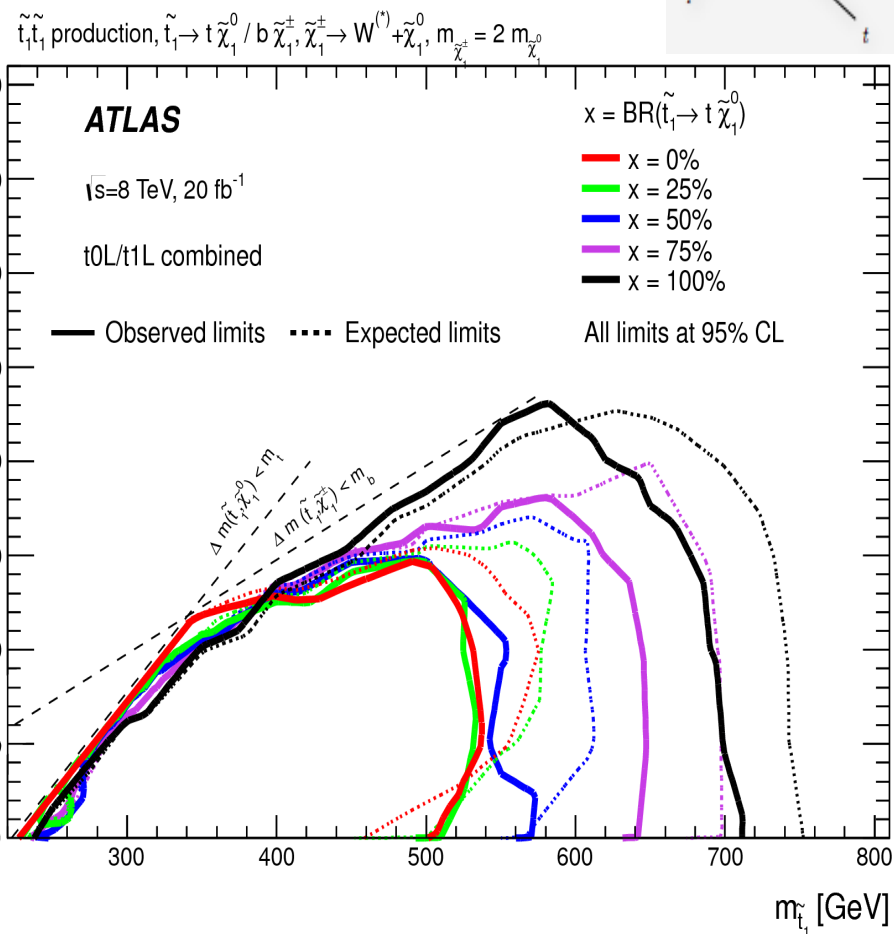
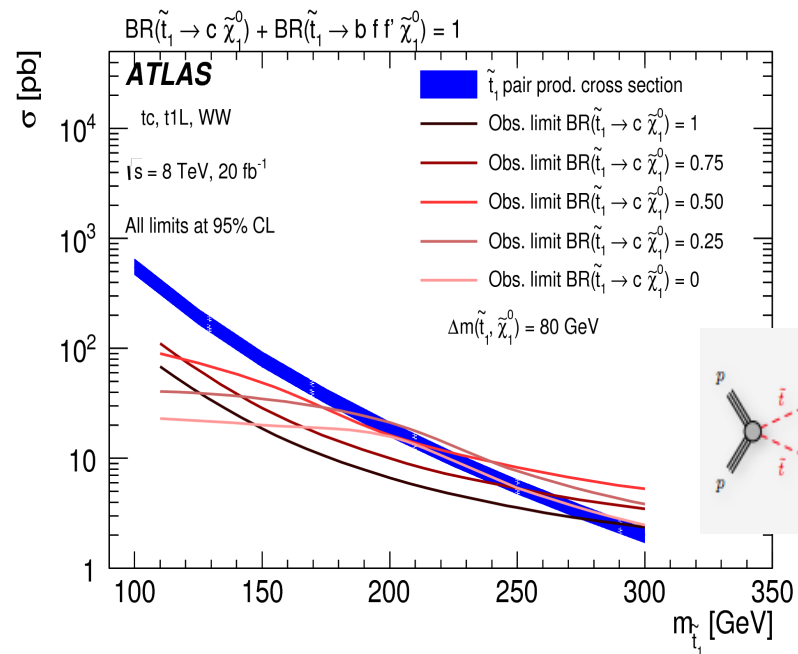
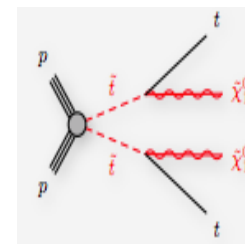


Exclusion for $m(\tilde{t}_1) < \sim 720 \text{ GeV}$ for massless LSP
 Exclusion up to $m(\text{LSP}) \sim 280 \text{ GeV}$

Branching Ratio Dependence



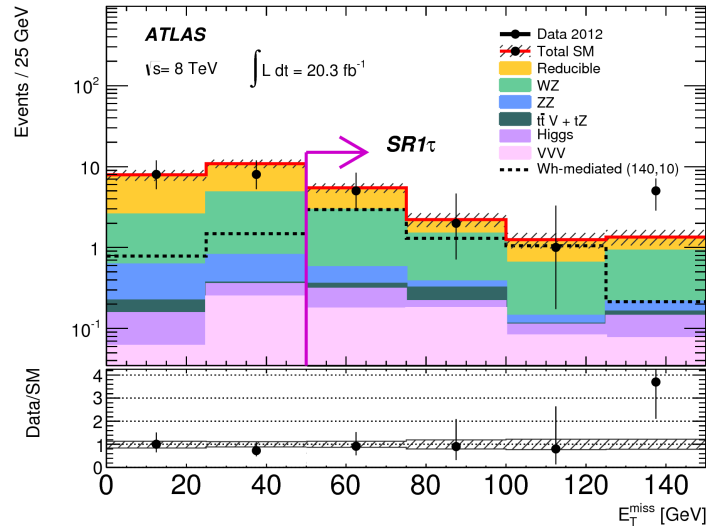
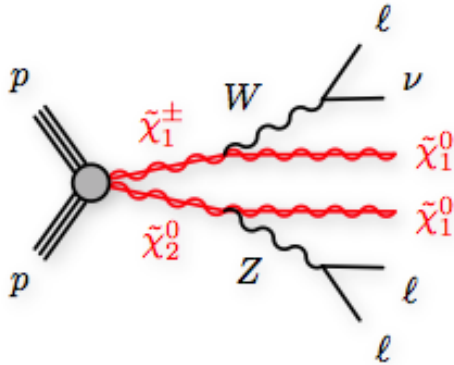
Exclusion limits are computed as a function of the decay Br for the stop



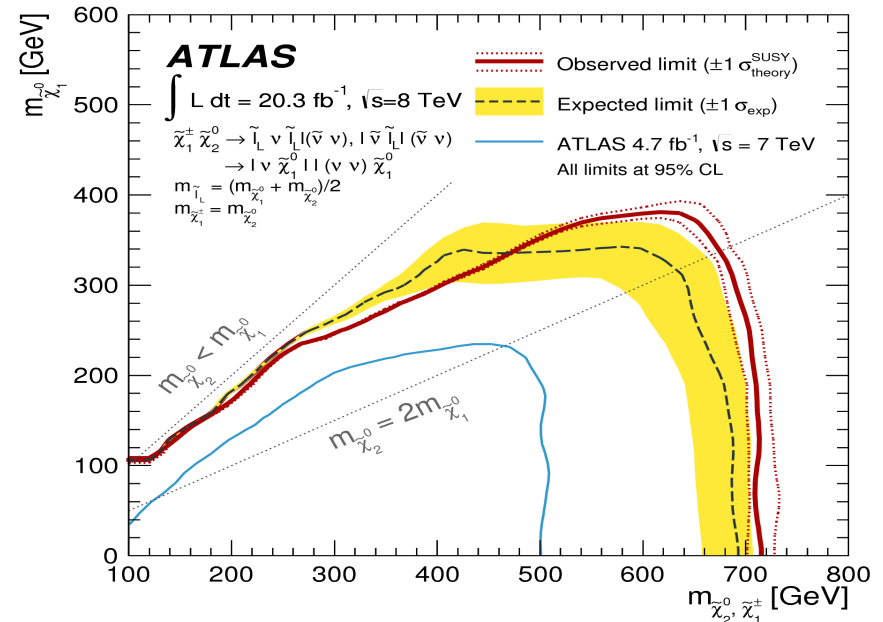
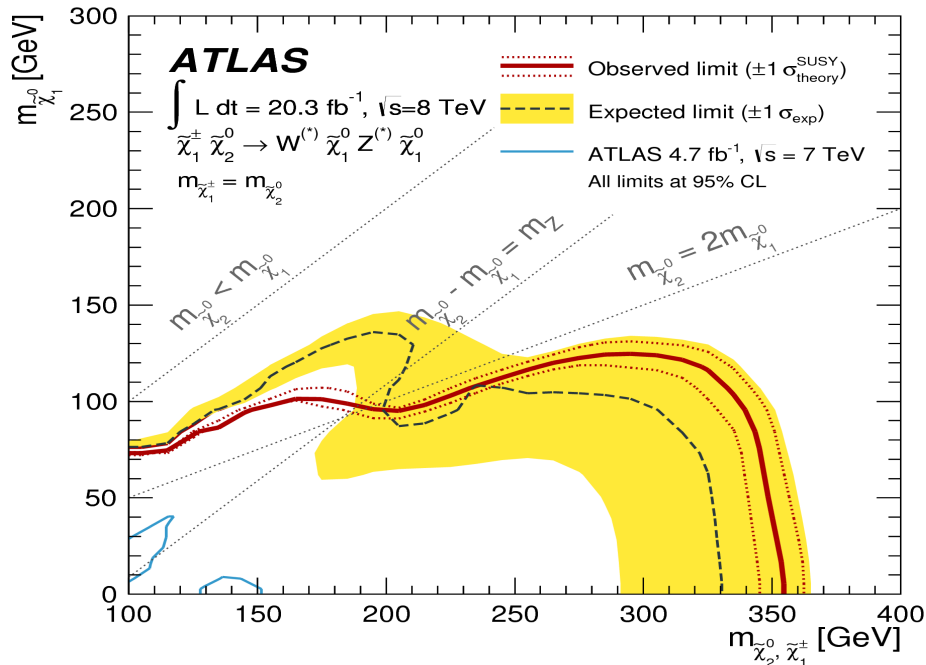
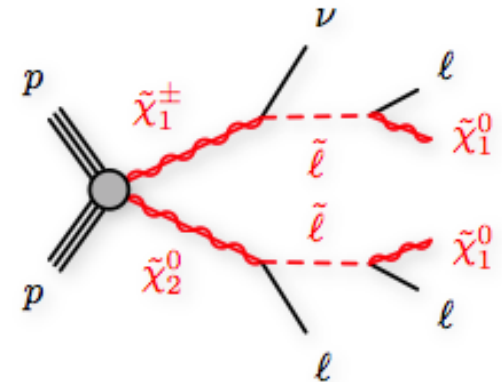
As expected, large variations are observed (not in the case of compressed scenario)

Chargino/Neutralino

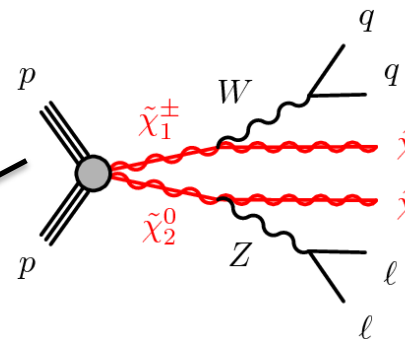
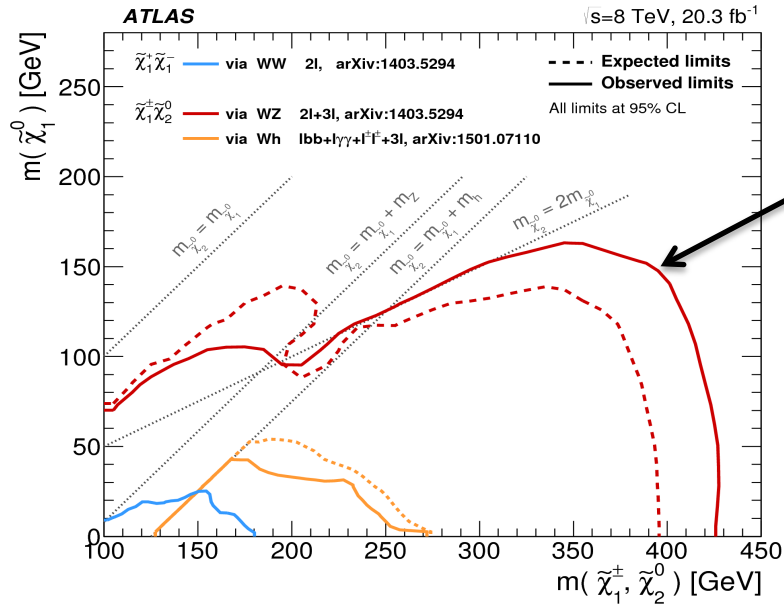
3 leptons + E_T^{miss}



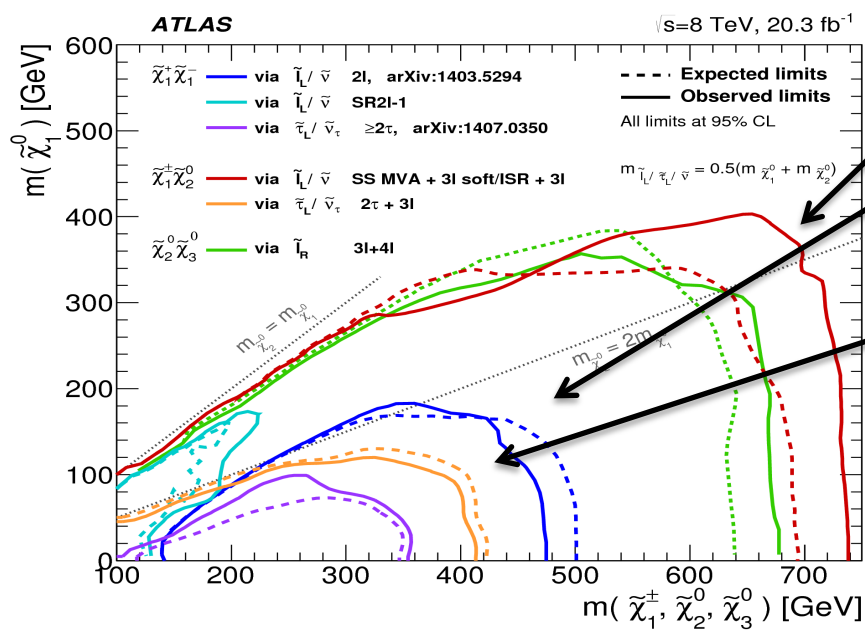
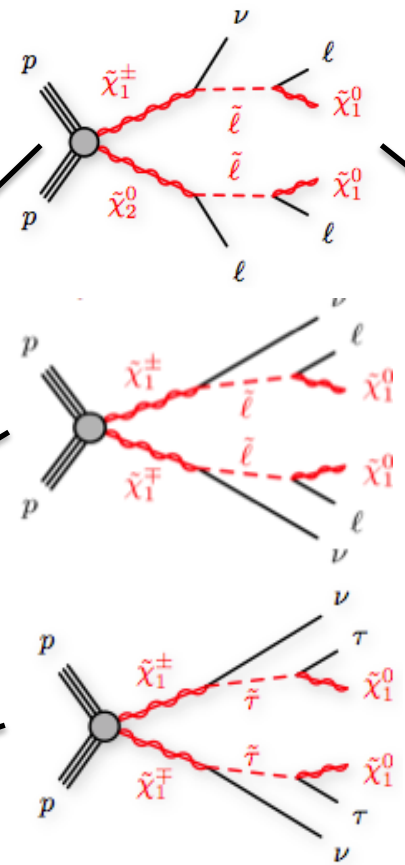
Background mostly WZ, ZZ followed by ttV and tribosons



EWK production



2 or 3 leptons + E_T^{miss}



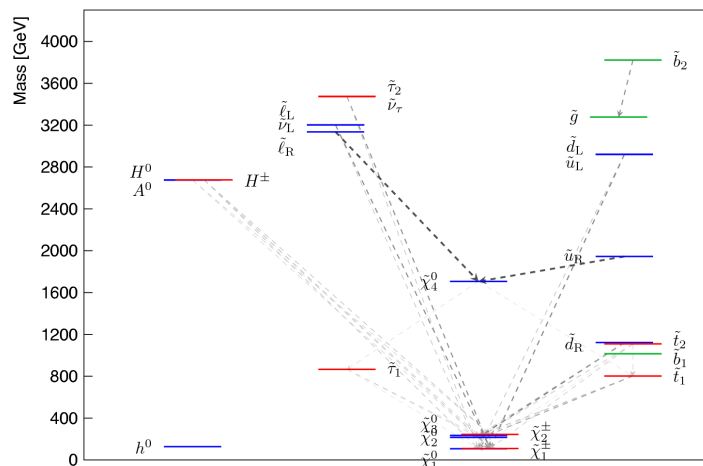
Chargino mass > ~700 GeV
Neutralino mass > ~380 GeV

Very sensitive to the details of the scenario considered

Up to 22 different ATLAS searches considered

Analysis	Ref.	Category
0-lepton + 2-6 jets + E_T^{miss}	[57]	Inclusive
0-lepton + 7-10 jets + E_T^{miss}	[58]	
1-lepton + jets + E_T^{miss}	[59]	
$\tau(\tau/\ell)$ + jets + E_T^{miss}	[60]	
SS/3-leptons + jets + E_T^{miss}	[61]	
0/1-lepton + 3b-jets + E_T^{miss}	[62]	Third generation
Monojet	[63]	
0-lepton stop	[64]	
1-lepton stop	[55]	
2-leptons stop	[65]	
Monojet stop	[66]	
Stop with Z boson	[67]	Electroweak
2b-jets + E_T^{miss}	[68]	
$tb + E_T^{\text{miss}}$, stop	[56]	
th	[69]	
2-leptons	[53]	
2- τ	[54]	Other
3-leptons	[52]	
4-leptons	[70]	
Disappearing Track	[71]	
Long-lived particle	[72,73]	
$H/A \rightarrow \tau^+\tau^-$	[74]	

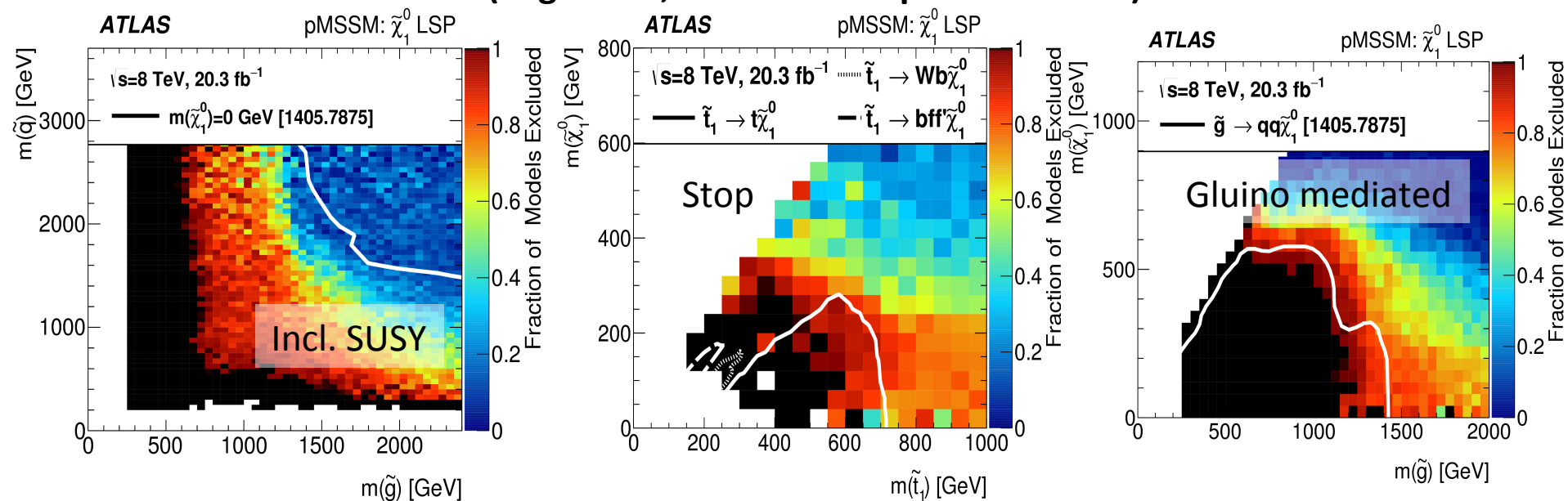
One example of particle spectra



The results are interpreted in the context of 19-parameter

Parameter	Min value	Max value	Note
$m_{\tilde{t}_1} (= m_{\tilde{t}_2})$	90 GeV	4TeV	Left-handed slepton (first two gens.) mass
$m_{\tilde{e}_1} (= m_{\tilde{e}_2})$	90 GeV	4TeV	Right-handed slepton (first two gens.) mass
$m_{\tilde{t}_3}$	90 GeV	4TeV	Left-handed stau doublet mass
$m_{\tilde{e}_3}$	90 GeV	4TeV	Right-handed stau mass
$m_{\tilde{Q}_1} (= m_{\tilde{Q}_2})$	200 GeV	4TeV	Left-handed squark (first two gens.) mass
$m_{\tilde{u}_1} (= m_{\tilde{u}_2})$	200 GeV	4TeV	Right-handed up-type squark (first two gens.) mass
$m_{\tilde{d}_1} (= m_{\tilde{d}_2})$	200 GeV	4TeV	Right-handed down-type squark (first two gens.) mass
$m_{\tilde{Q}_3}$	100 GeV	4TeV	Left-handed squark (third gen.) mass
$m_{\tilde{u}_3}$	100 GeV	4TeV	Right-handed top squark mass
$m_{\tilde{d}_3}$	100 GeV	4TeV	Right-handed bottom squark mass
$ M_1 $	0 GeV	4TeV	Bino mass parameter
$ M_2 $	70 GeV	4TeV	Wino mass parameter
$ \mu $	80 GeV	4TeV	Bilinear Higgs mass parameter
M_3	200 GeV	4TeV	Gluino mass parameter
$ A_t $	0 GeV	8TeV	Trilinear top coupling
$ A_b $	0 GeV	4TeV	Trilinear bottom coupling
$ A_\tau $	0 GeV	4TeV	Trilinear τ lepton coupling
M_A	100 GeV	4TeV	Pseudoscalar Higgs boson mass
$\tan\beta$	1	60	Ratio of the Higgs vacuum expectation values

Fraction of pMSSM models excluded (in general, similar to simplified models)



**Hard to believe in SUSY if the
Higgs sector stays with just h_0**

...Looking for extra Higgs particles ...

MSSM Higgs Sector

Two Higgs doublets are required in supersymmetric theories to generate mass for both "up" type and "down" type quarks and charged leptons.

$$H_d = \begin{pmatrix} (v_d + \phi_d^0 + i\chi_d^0)/\sqrt{2} \\ \phi_d^- \end{pmatrix} \quad H_u = \begin{pmatrix} \phi_u^+ \\ (v_u + \phi_u^0 + i\chi_u^0)/\sqrt{2} \end{pmatrix}$$

normalization: $v^2 \equiv v_d^2 + v_u^2 = 4m_W^2/g^2 = (246\text{GeV})^2$

$$\frac{v_u}{v_d} = \tan \beta, \quad \beta \in (0, \frac{\pi}{2})$$

Two parameter : M_A and $\tan\beta$

$$m_{H^\pm}^2 = m_A^2 + m_W^2$$

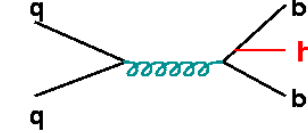
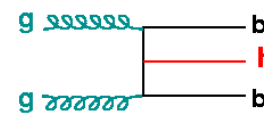
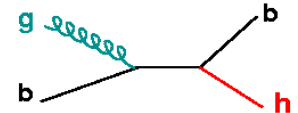
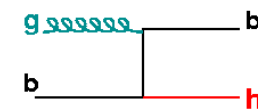
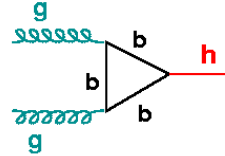
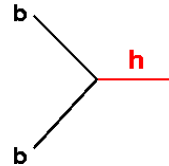
$$m_{h^0, H^0}^2 = \frac{1}{2} \left(m_{A^0}^2 + m_Z^2 \mp \sqrt{(m_{A^0}^2 + m_Z^2)^2 - 4m_Z^2 m_{A^0}^2 \cos^2 2\beta} \right)$$

MSSM Higgs

$H_U, H_D \longrightarrow h, H, A$ and H^\pm

$$\tan \beta = \langle H_U \rangle / \langle H_D \rangle$$

$$M(h) < 135 \text{ GeV}$$



Tree level : M_A and $\tan \beta$ as parameters

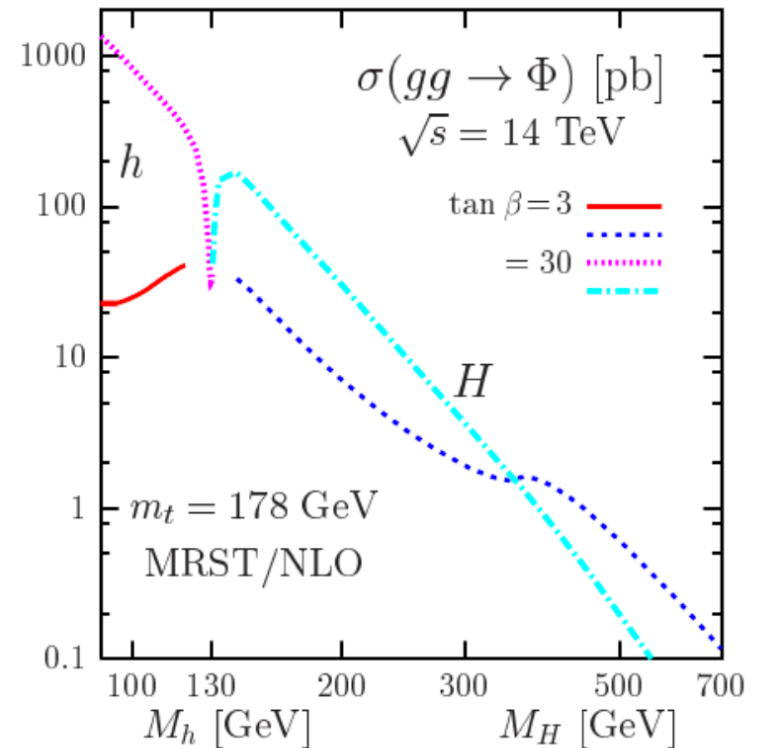
MSSM Higgs production cross section boosted compared to SM at large $\tan \beta$

$$\sigma(b\bar{b}A) \times \text{BR}(A \rightarrow b\bar{b}) \simeq \sigma(b\bar{b}A)_{\text{SM}} \frac{\tan^2 \beta}{(1 + \Delta_b)^2} \times \frac{9}{(1 + \Delta_b)^2 + 9}$$

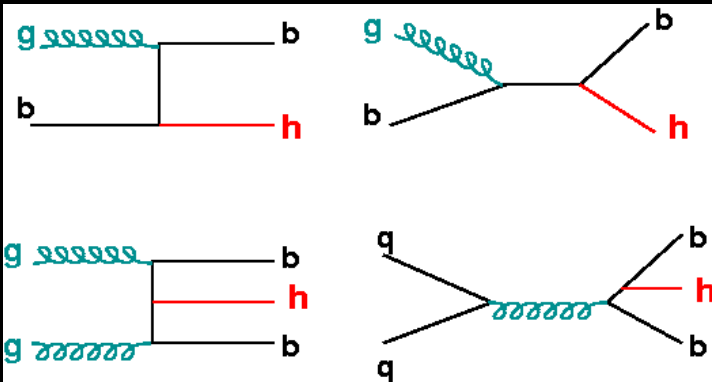
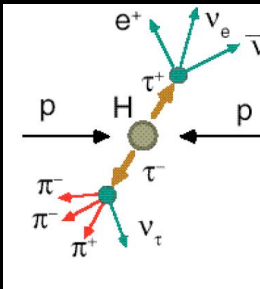
$$\sigma(gg, b\bar{b} \rightarrow A) \times \text{BR}(A \rightarrow \tau^+\tau^-) \simeq \sigma(gg, b\bar{b} \rightarrow A)_{\text{SM}} \frac{\tan^2 \beta}{(1 + \Delta_b)^2 + 9}$$

At low masses

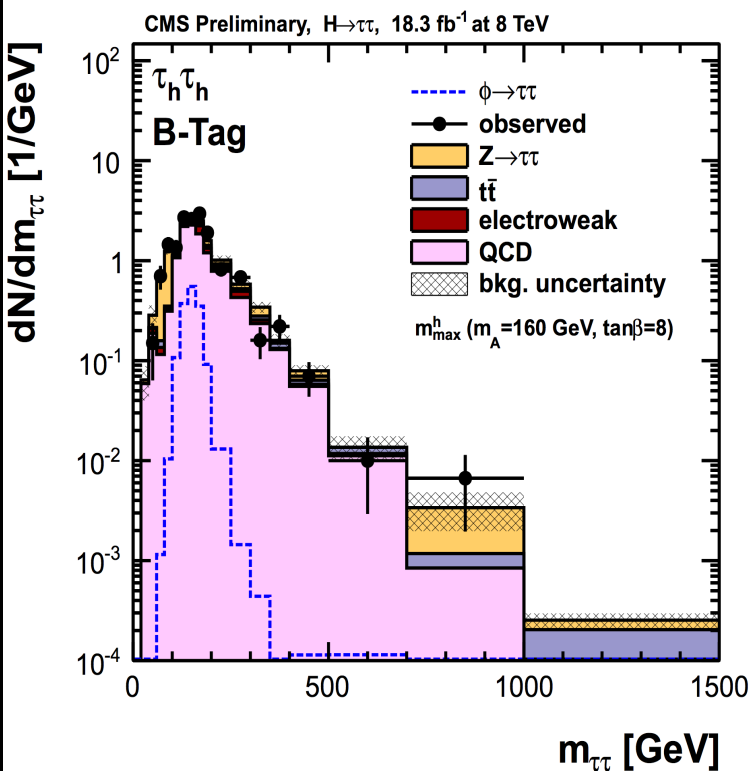
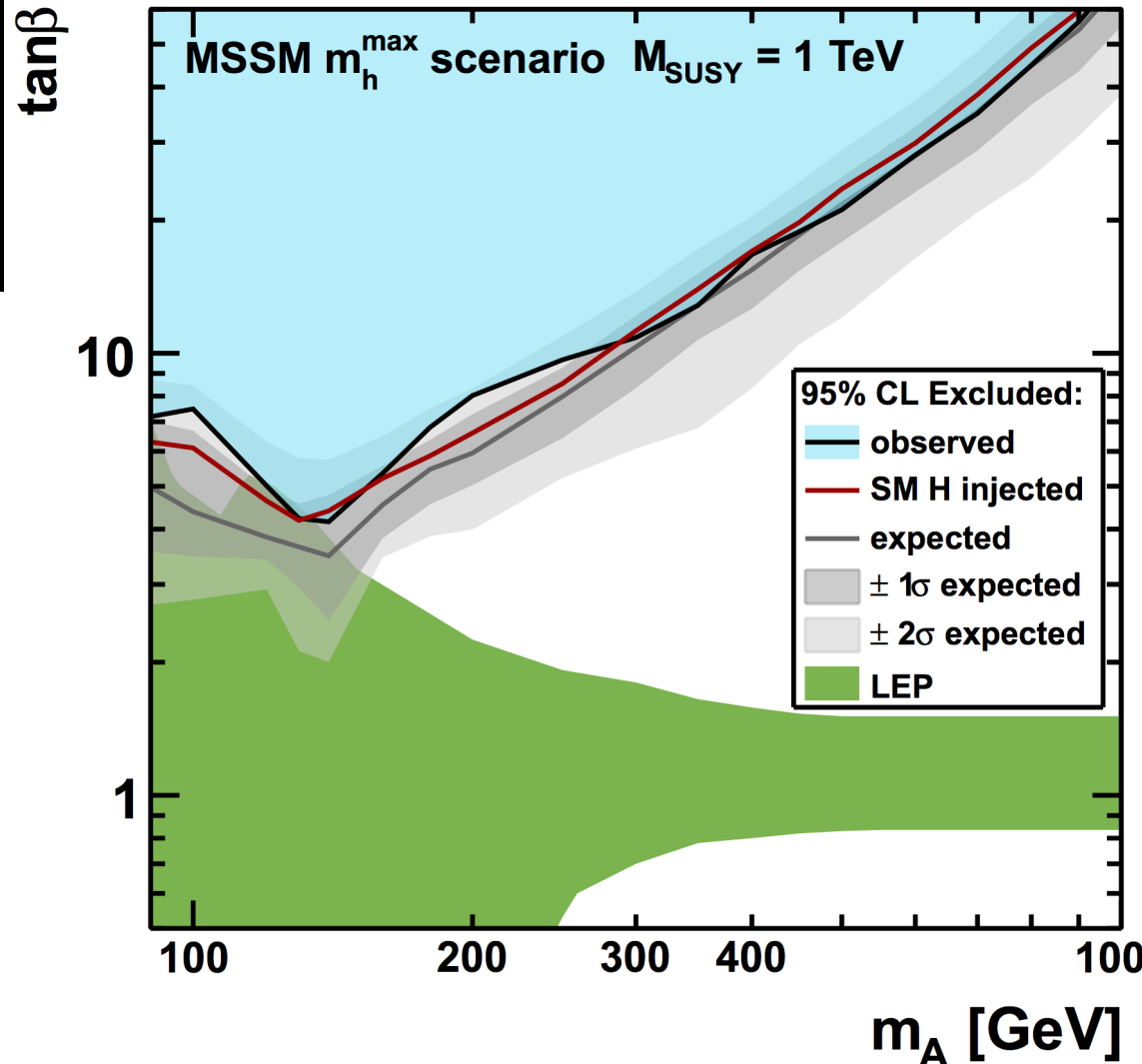
$\text{Br}(h \rightarrow b\bar{b}) \sim 90\%$, $\text{Br}(h \rightarrow \tau\tau) \sim 10\%$



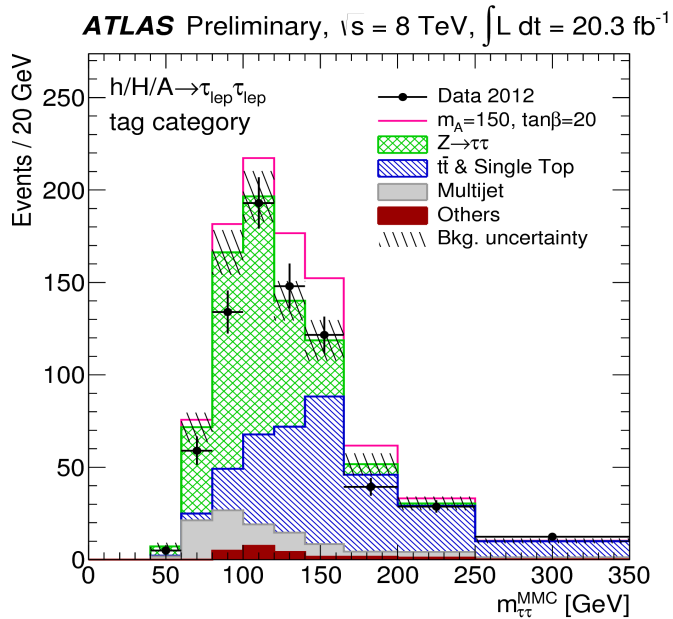
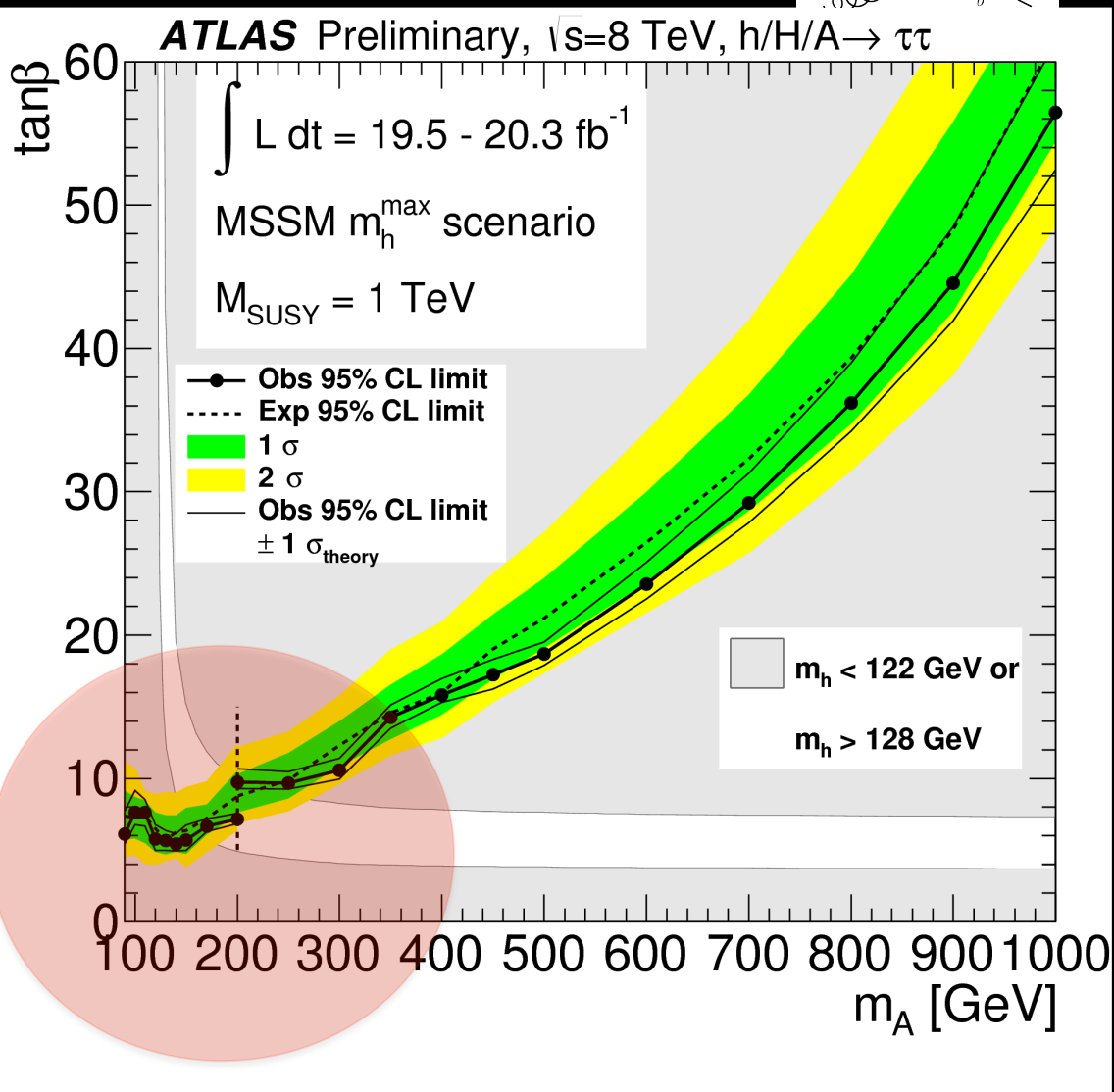
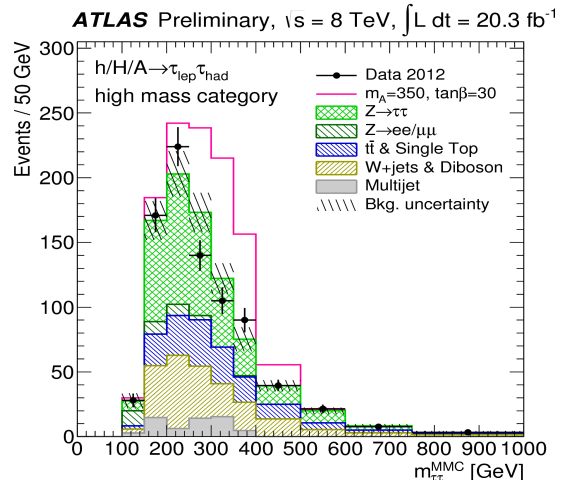
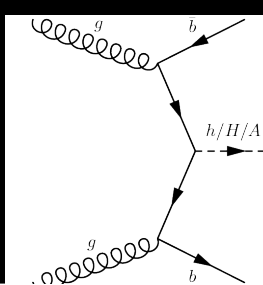
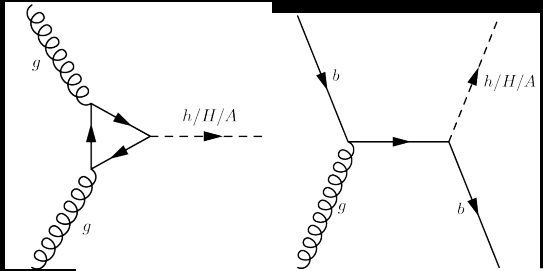
MSSM Neutral Higgs



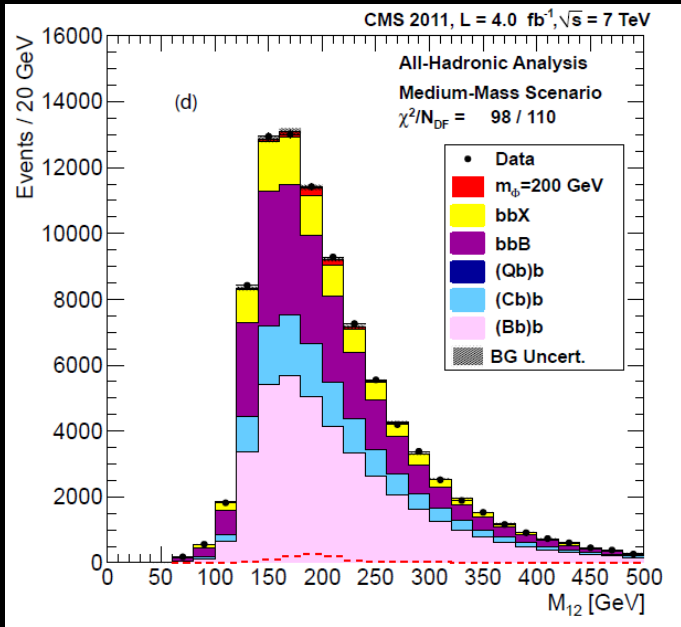
CMS Preliminary, $H \rightarrow \tau\tau$, 4.9 fb^{-1} at 7 TeV, 19.7 fb^{-1} at 8 TeV



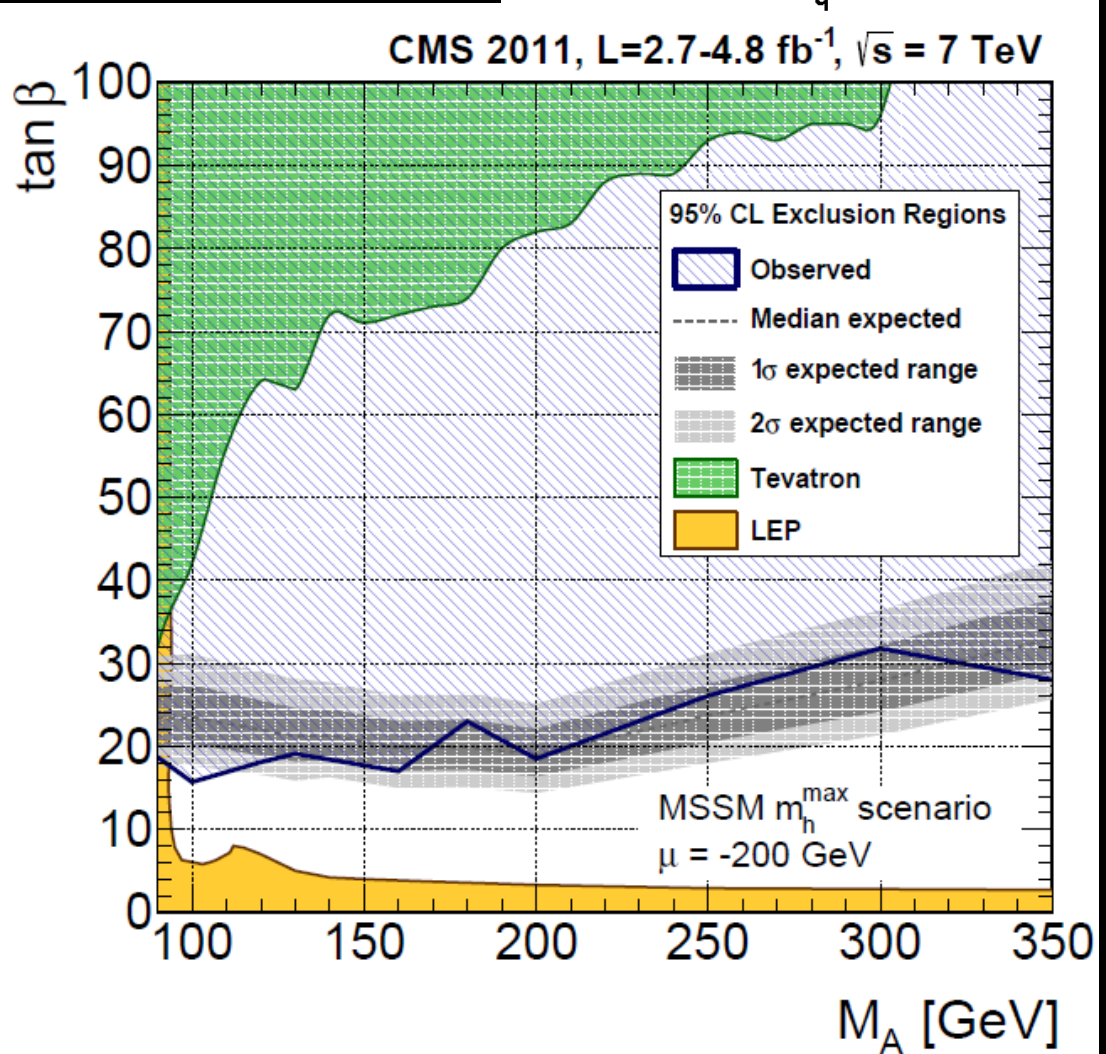
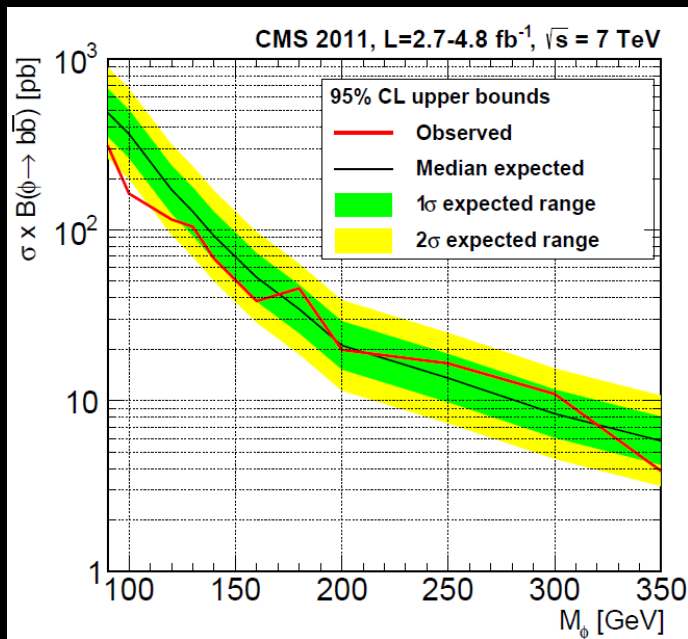
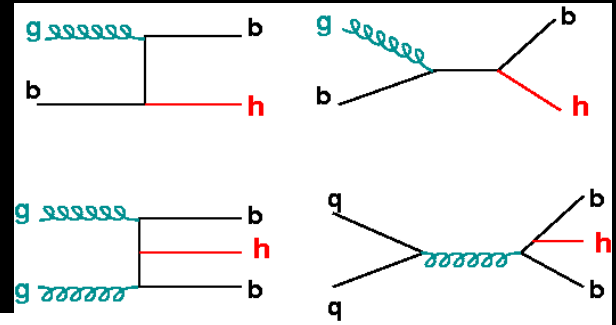
MSSM Neutral Higgs



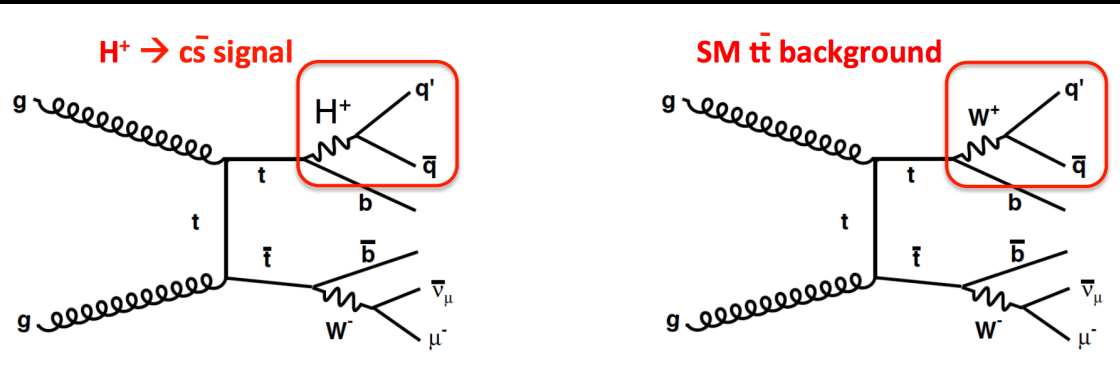
MSSM Higgs



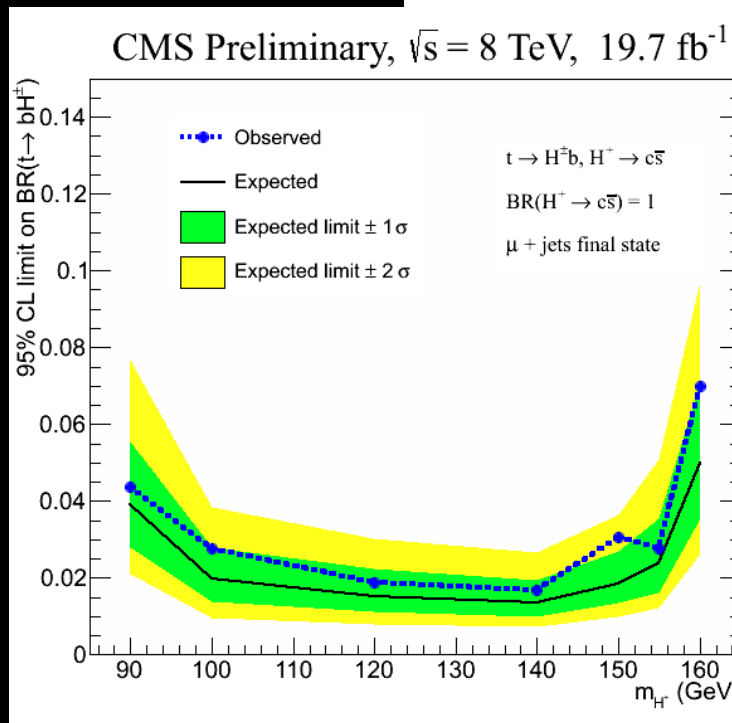
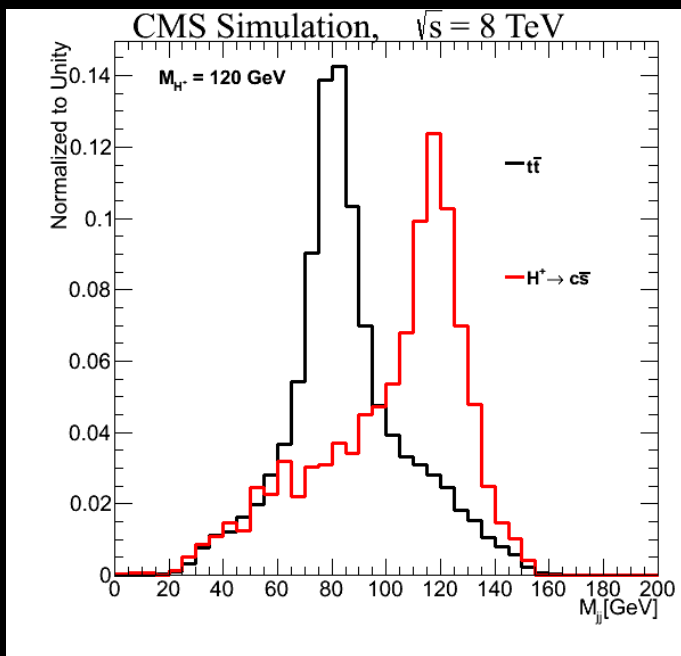
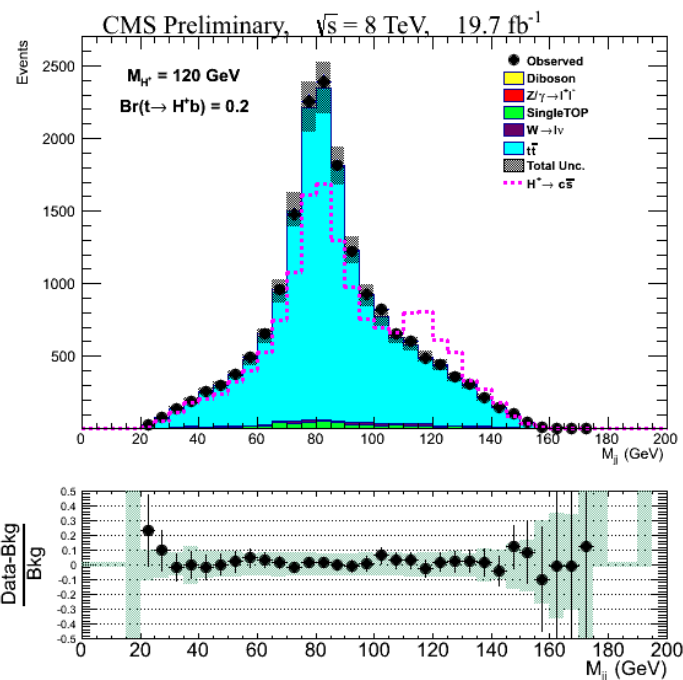
$h \rightarrow bb$

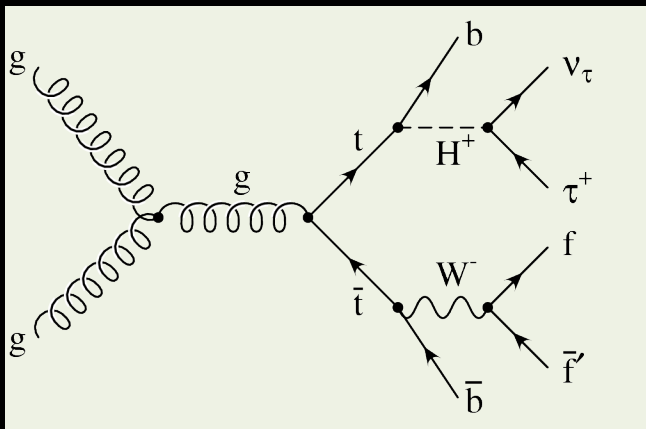


$$H^+ \rightarrow c\bar{s}$$



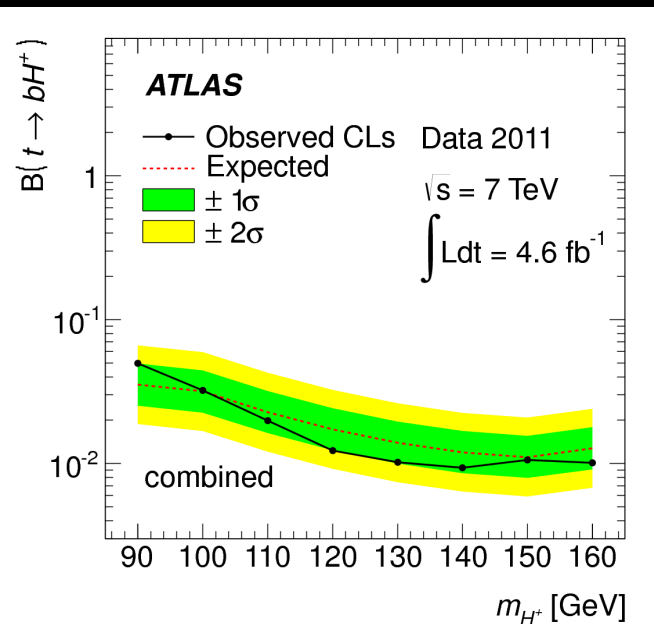
Expressed in terms of limits on BR in top decays



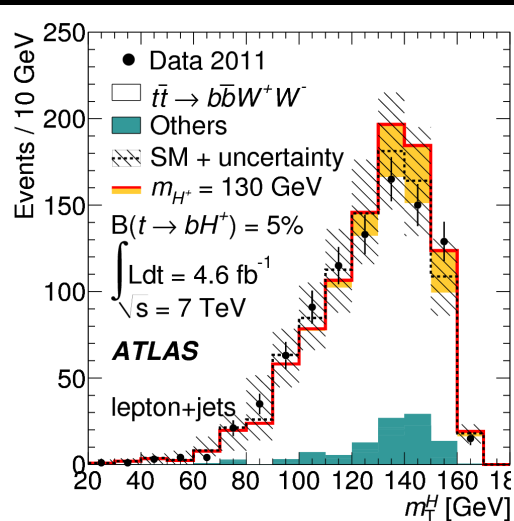


$$H^+ \rightarrow \tau \nu$$

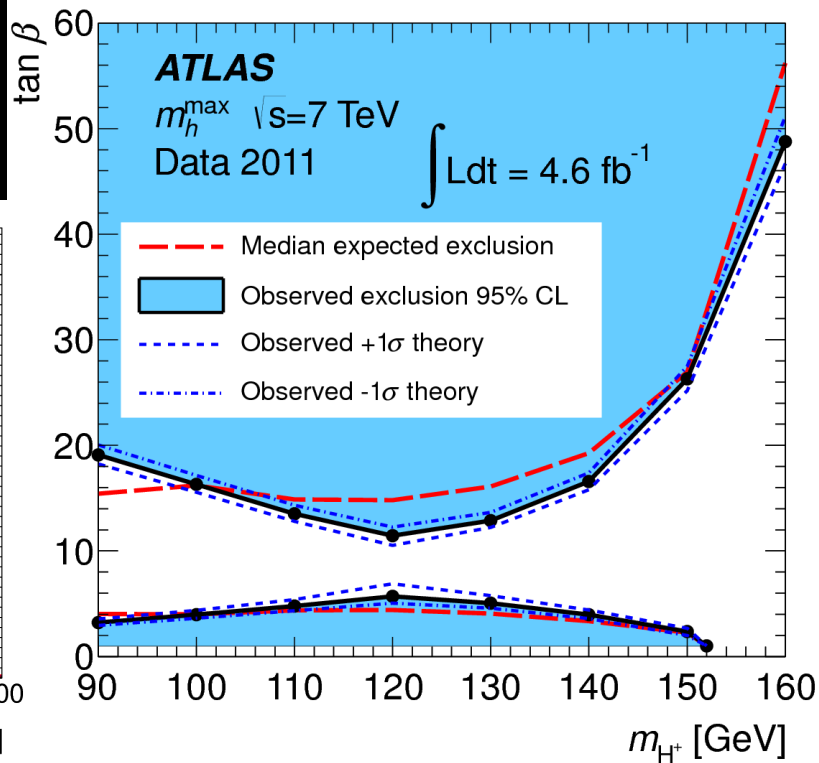
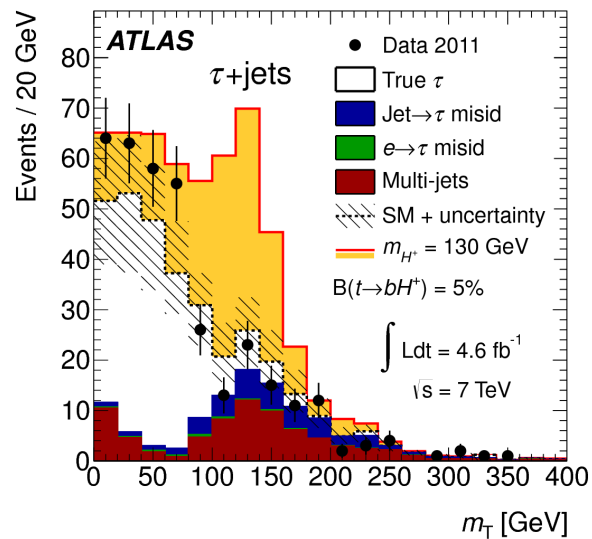
Limits on top BR and SUSY MSSM parameters



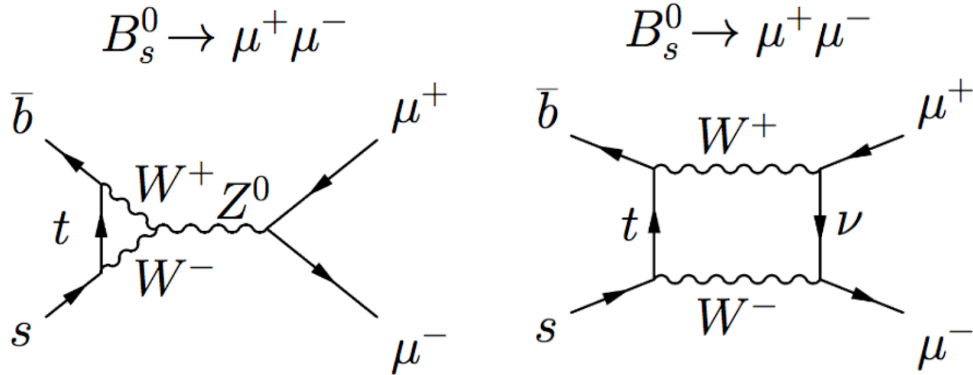
lepton + jets



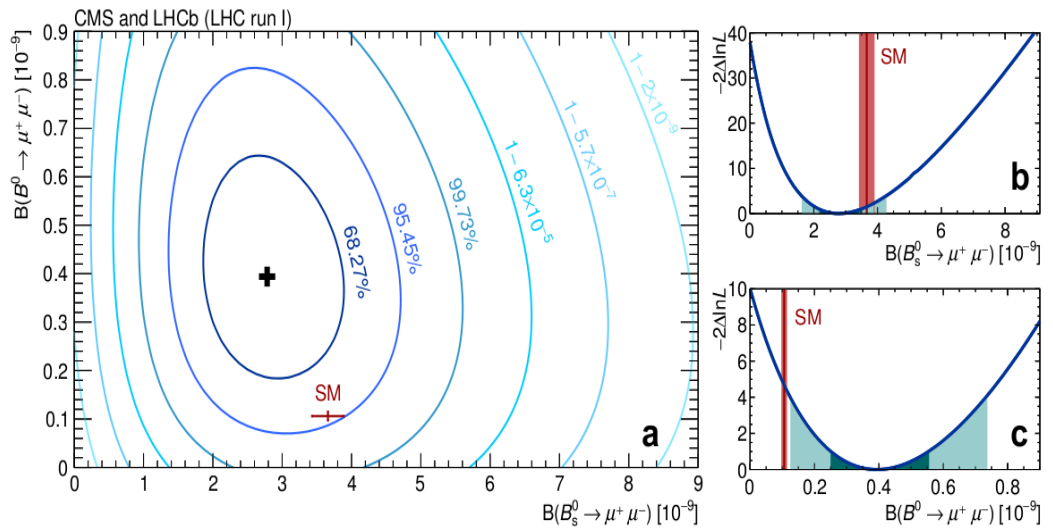
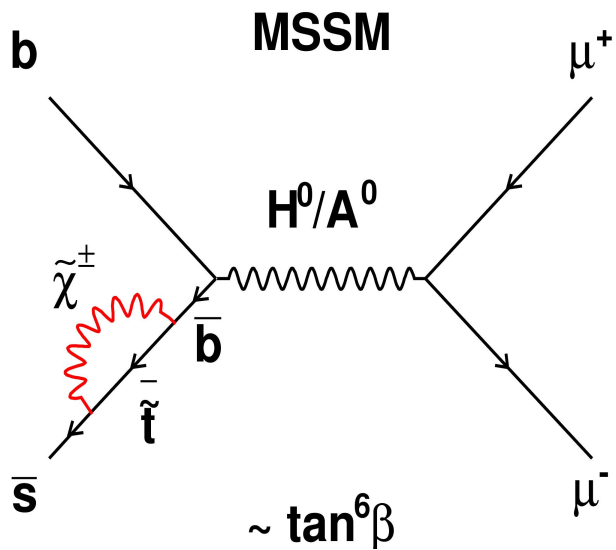
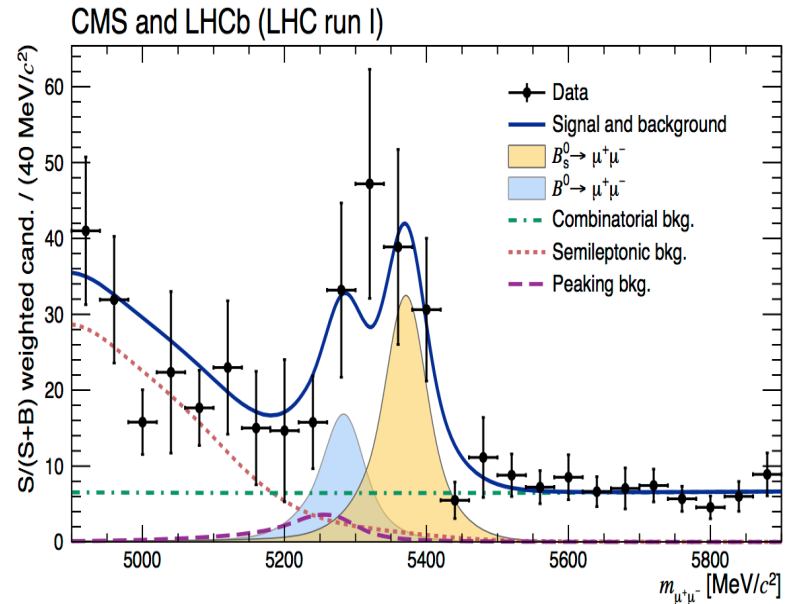
Tau-ID + jets



$B_s \rightarrow \mu^+ \mu^-$



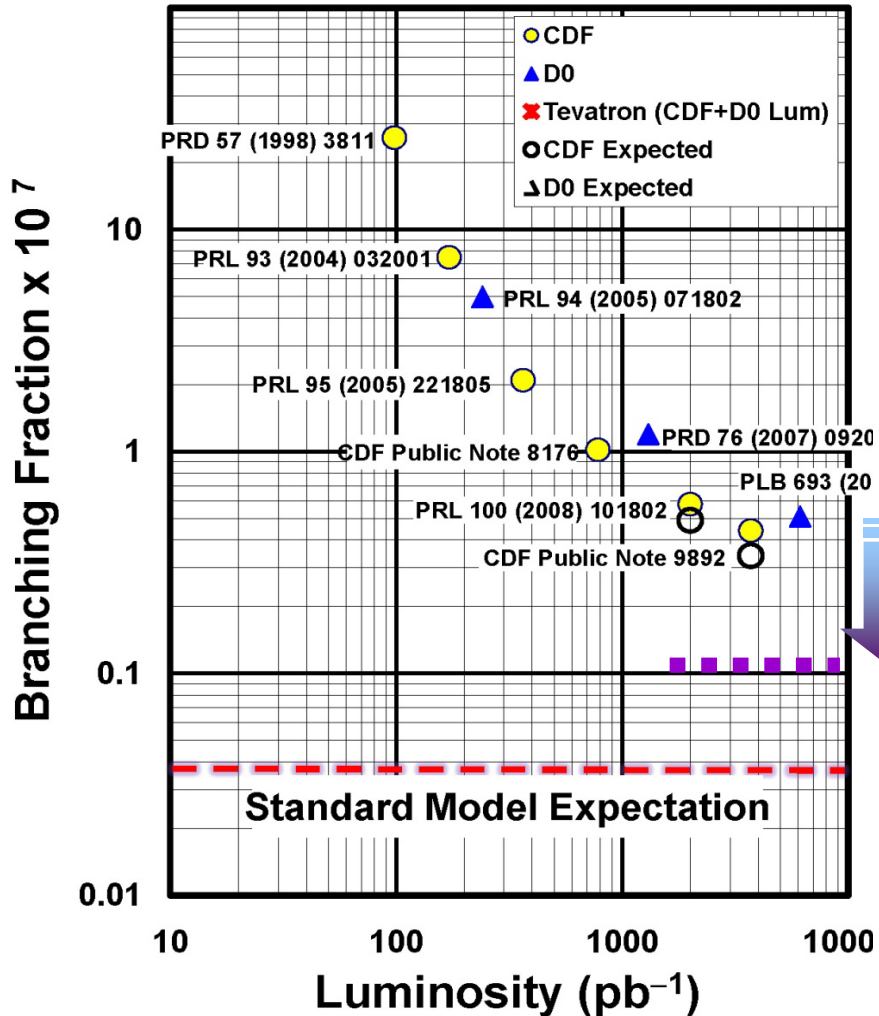
SM prediction $\Rightarrow \text{Br}(B_s \rightarrow \mu^+ \mu^-) \sim 3.42 \times 10^{-9}$
 SUSY \rightarrow big enhancement $\rightarrow (\sim \tan^6 \beta)$



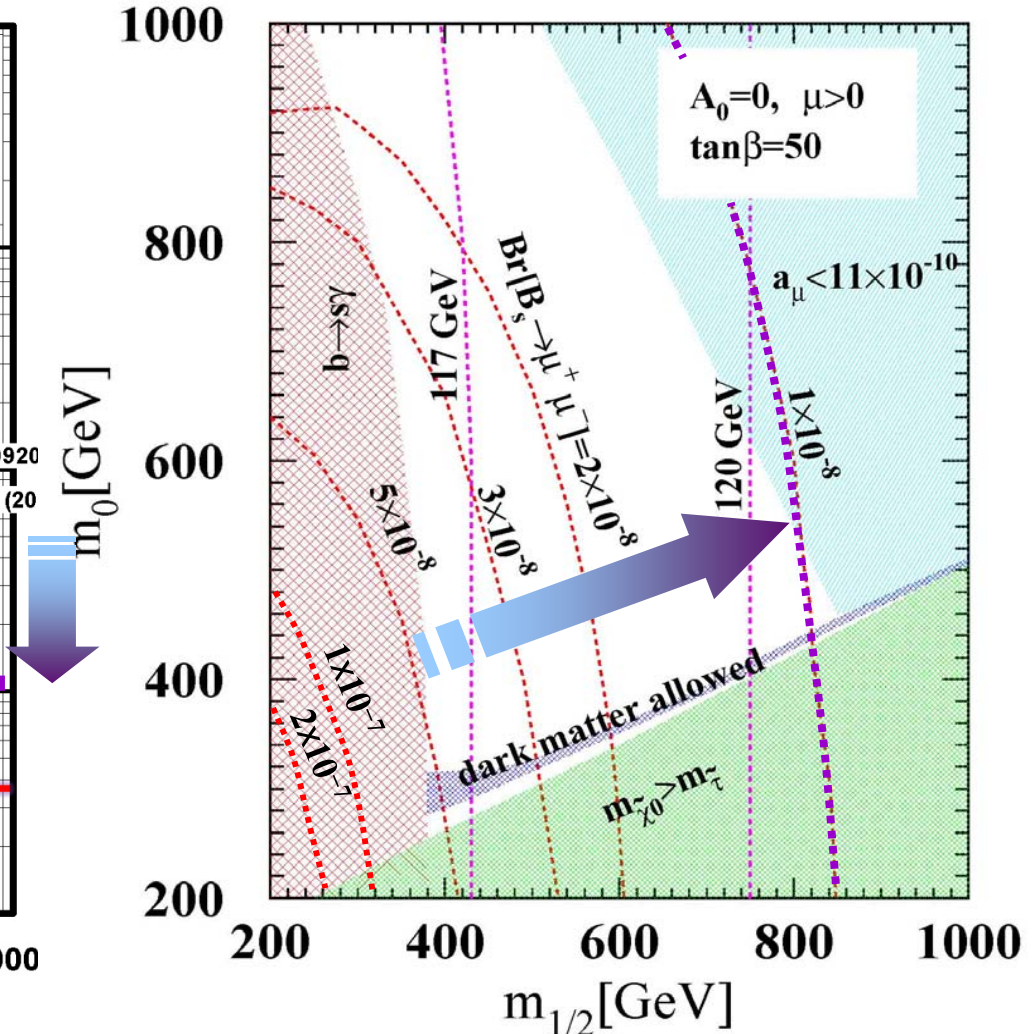
Consistent with SM predictions

$\mathcal{B}(B_s \rightarrow \mu\mu)$ and Cosmological Connection

95% CL Limits on $\mathcal{B}(B_s \rightarrow \mu\mu)$



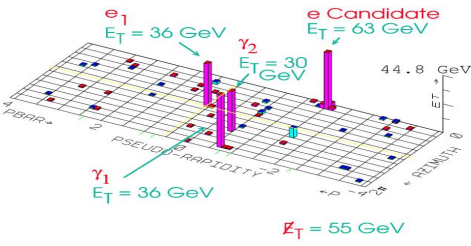
mSUGRA at $\tan\beta = 50$
 Arnowitt, Dutta, et al., PLB 538 (2002) 121



GMSB

(gauge mediated SUSY breaking)

Snowmass p8 spectra

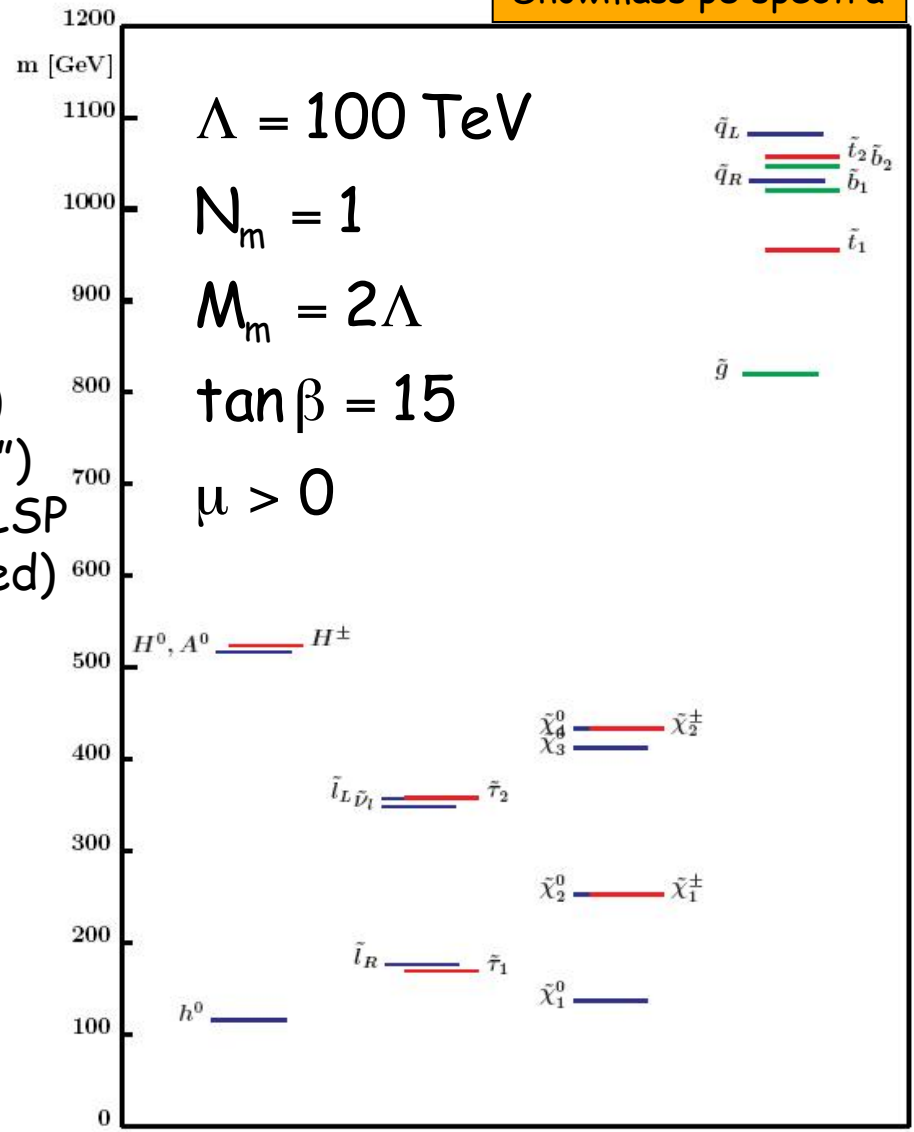


- SUSY breaking at scale Λ (10 -100 TeV)
- Mediated by Gauge Fields ("messengers")
- Gravitino is very light (\sim KeV) and the LSP
- Neutralino is NLSP (its lifetime not fixed)

$$\chi_1^0 \rightarrow \tilde{G} \gamma$$

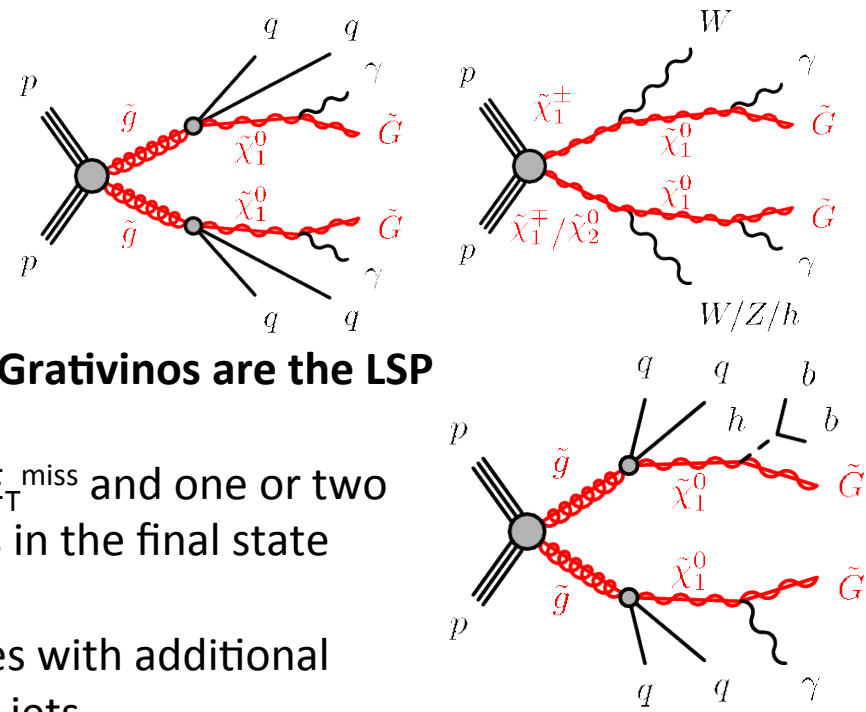
- Rp conservation (2 NLSP in final state)

$$\gamma\gamma + \cancel{E}_T + X$$



(taken from N. Ghodbane et al., hep-ph/0201233)

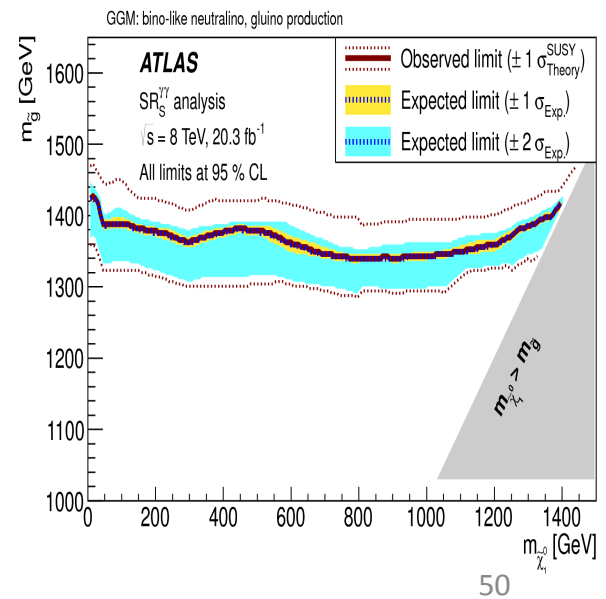
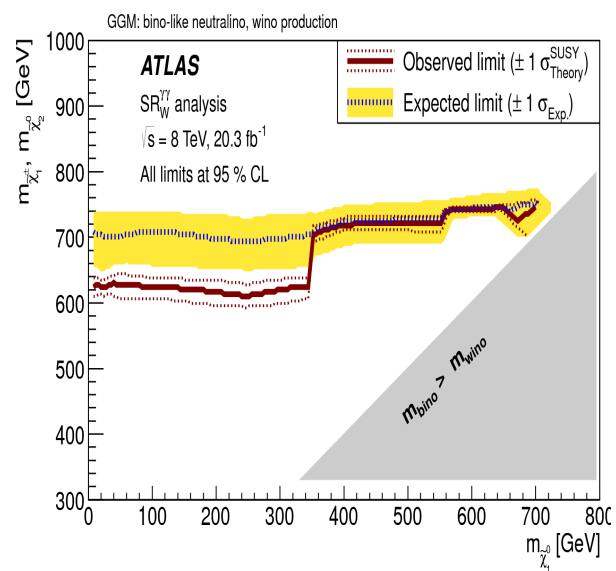
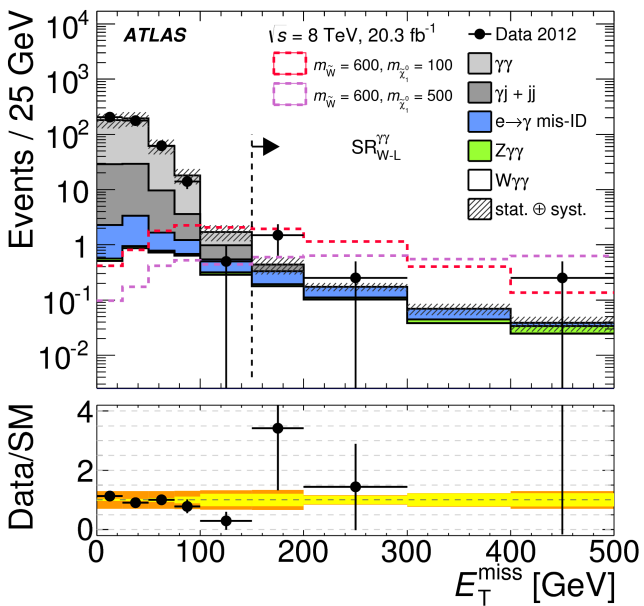
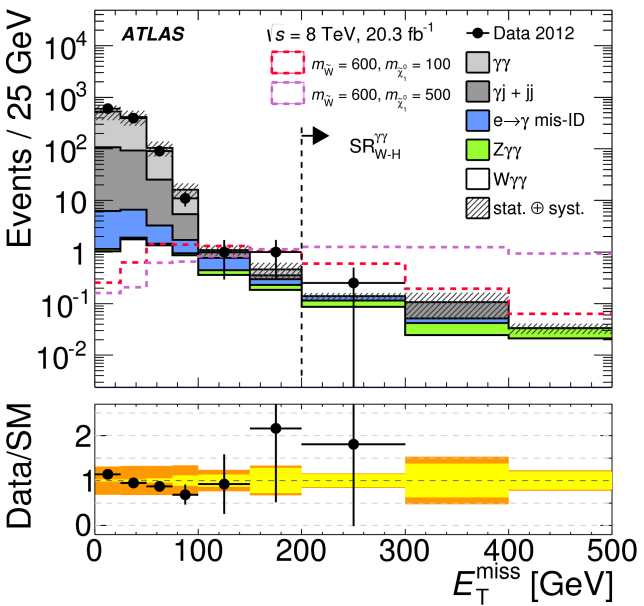
GMSB



Scenario in which Gravitinos are the LSP

Events with large E_T^{miss} and one or two energetic photons in the final state

Different topologies with additional leptons, jets and b-jets.





GMSB Gravitino

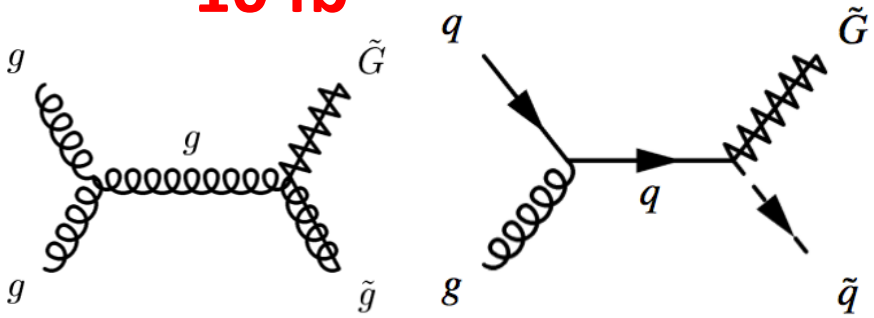
8 TeV
10 fb⁻¹

Monojet analysis
ATLAS-CONF-2012-147

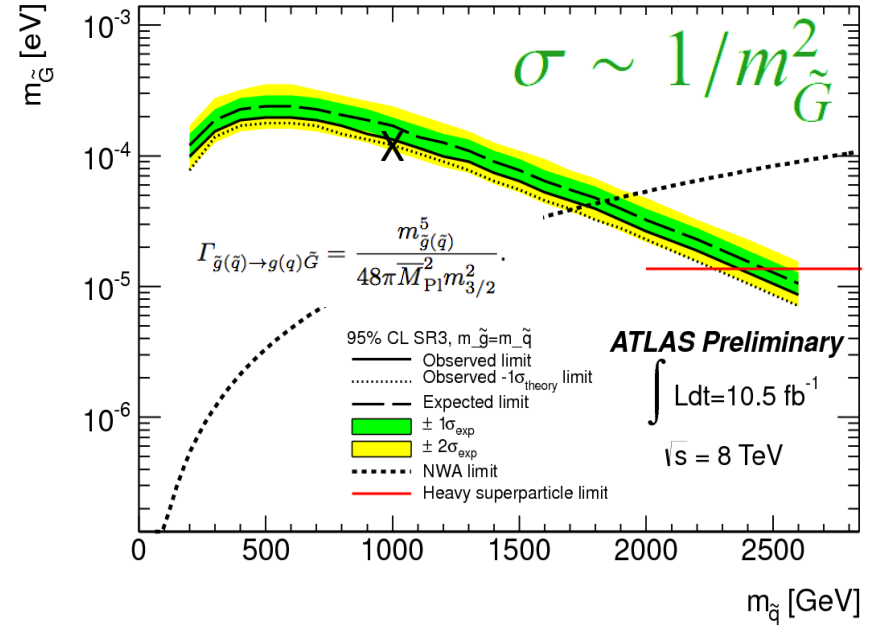
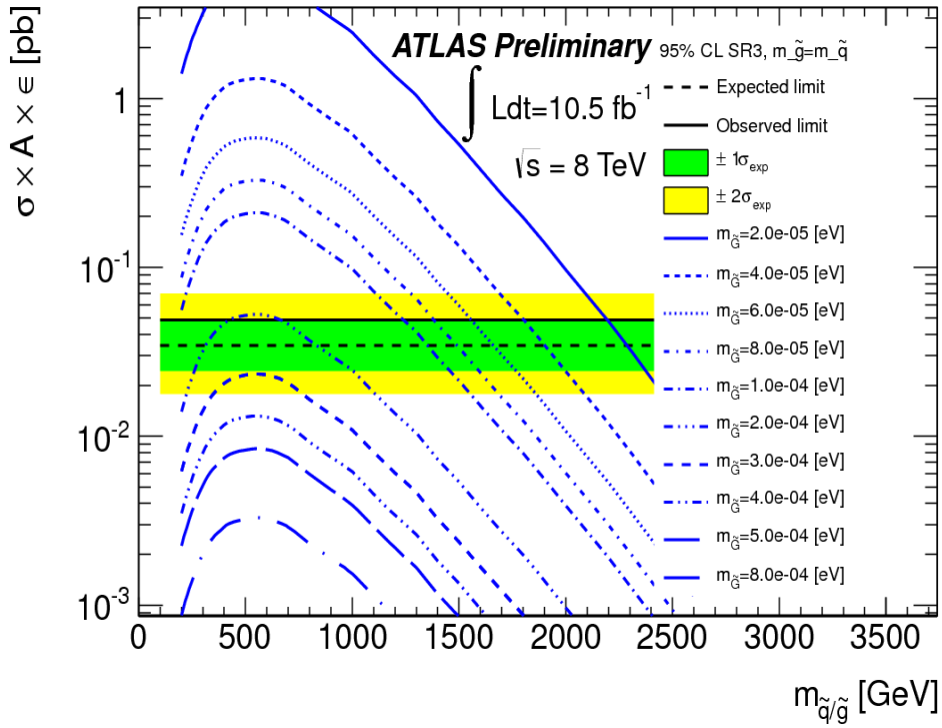
$$m_{3/2} = \langle F \rangle / \sqrt{3} \overline{M}_{\text{Pl}}$$

Interpreted in terms of GMSB
gravitino+squark/gluino production

gluinos (squarks) decay to
gluon (quark) plus Gravitino (100%)

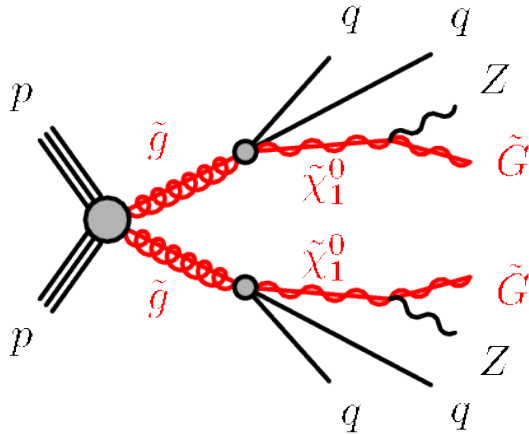


Best limits to date on the gravitino mass

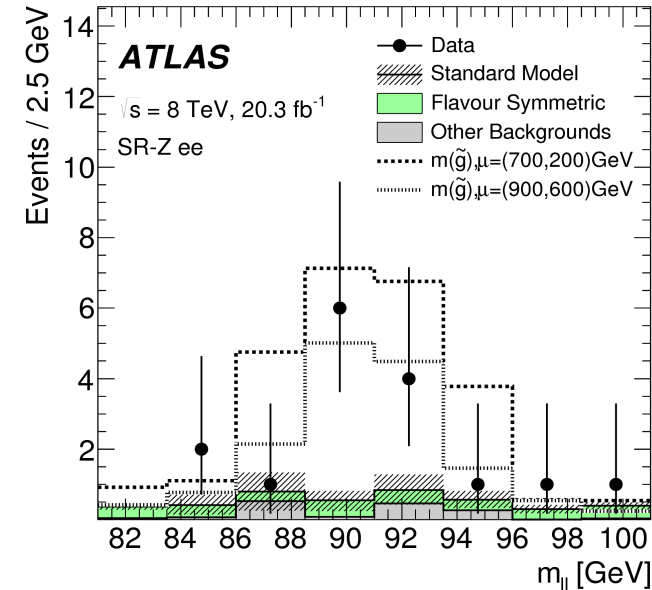
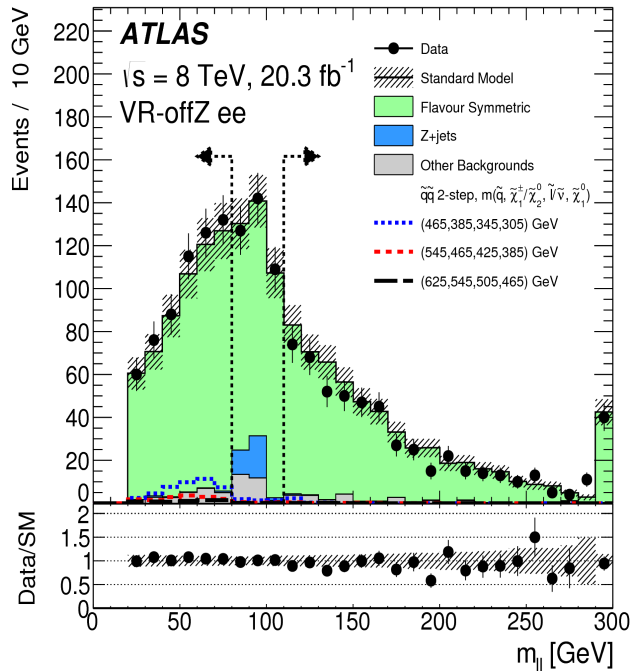
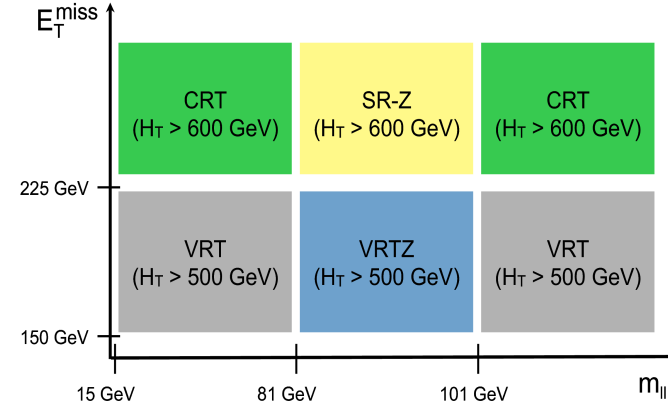
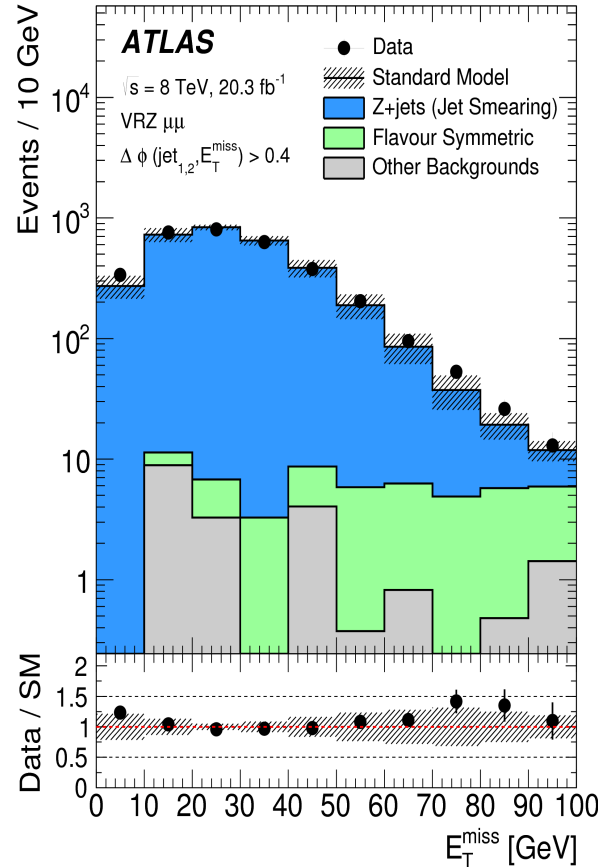


$m_{3/2} > 10^{-4} \text{ eV} \rightarrow \sqrt{F} > 640 \text{ GeV}$
(LEP limit 240 GeV)

SUSY searches with on-shell Zs



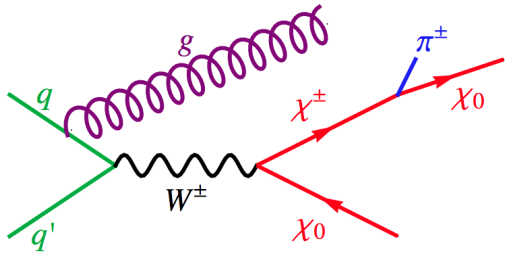
On-shell Z Large H_T and E_T^{miss}



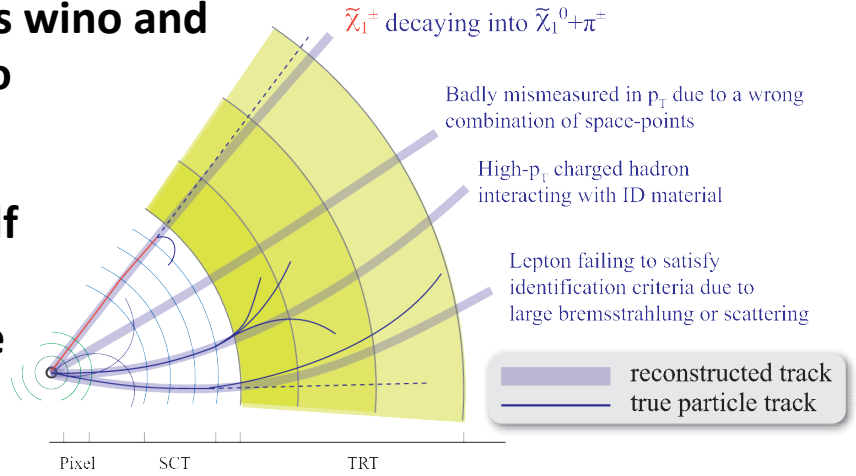
3 sigma effect

Disappearing track

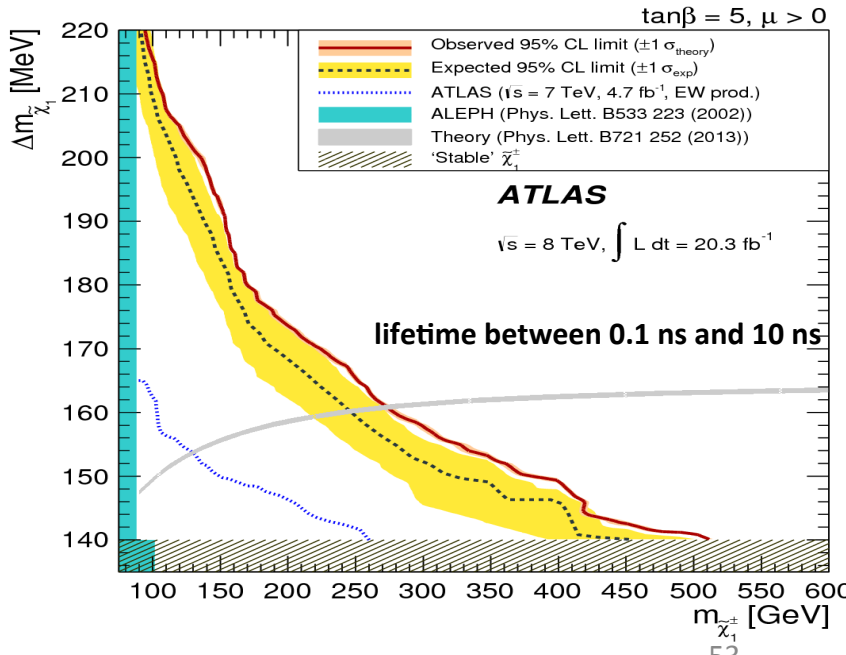
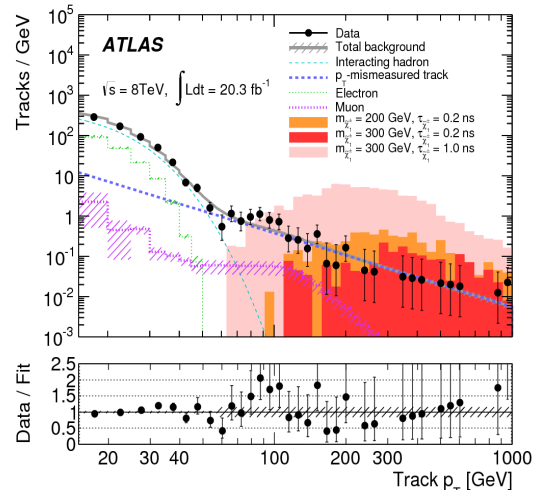
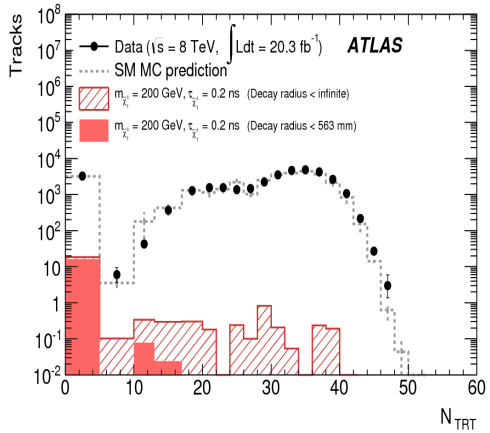
In AMSB model where the lightest chargino (N)LSP is wino and nearly mass-degenerate with the lightest neutralino



The signal reveals itself as disappearing tracks in the outer ID volume



At least one jet with $PT > 90$ GeV, $E_{miss} > 90$ GeV
 Lepton (electron and muon) vetoes
 One good isolated track with $pT > 15$ GeV
 and less than 5 hits in the TRT



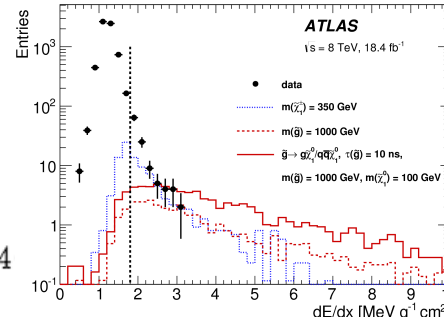
A chargino mass below 270 GeV is excluded at 95% CL

Heavy meta-stable Particles

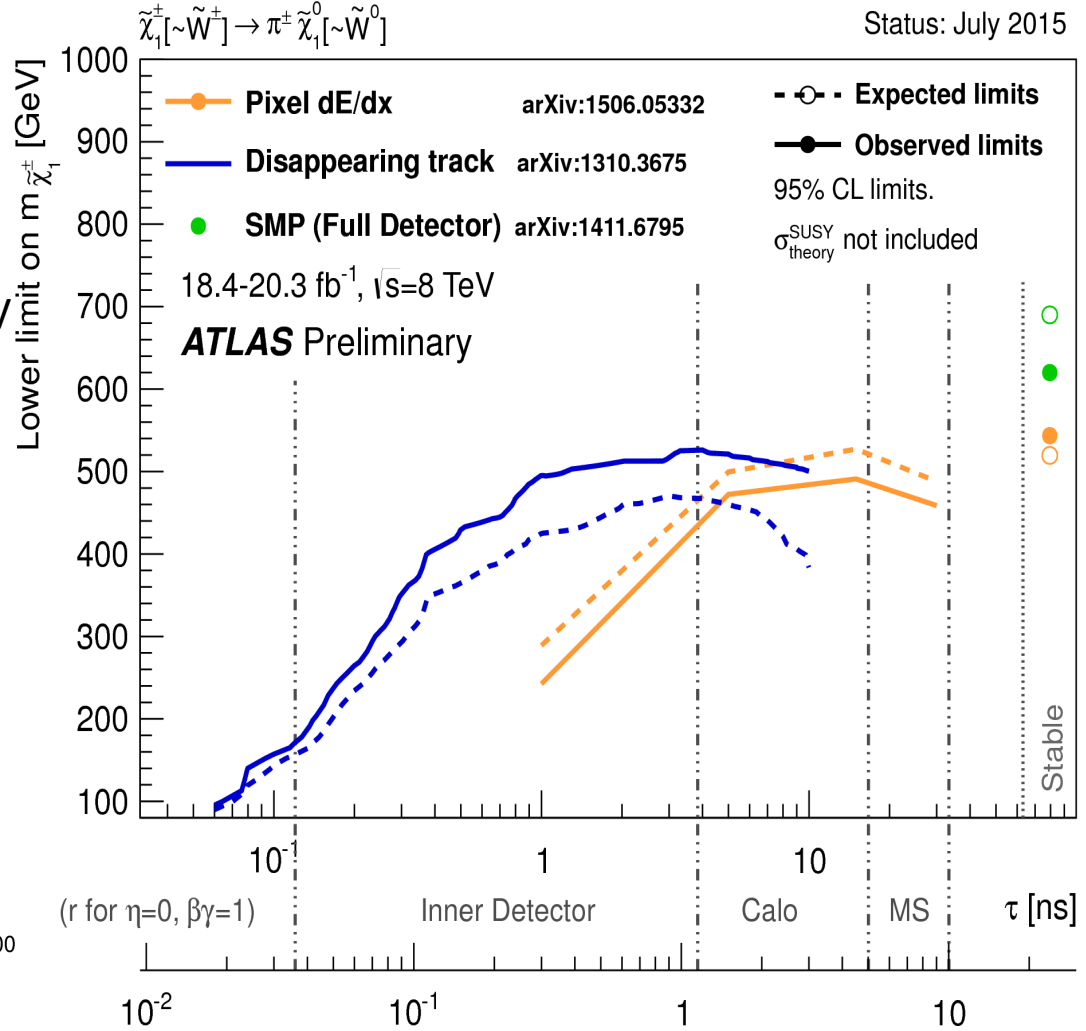
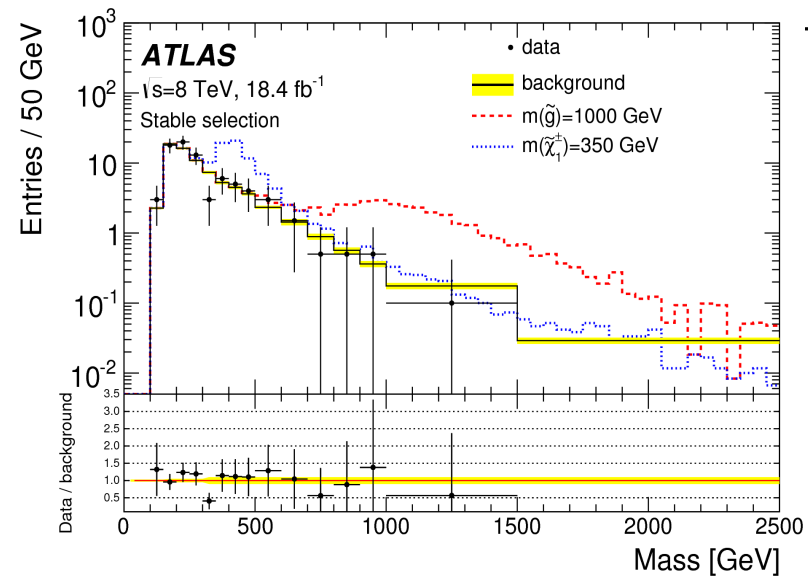
Using dE/dx in pixel detector
to extract the mass of the particle

$$dE/dx_{MPV}(\beta\gamma) = \frac{p_1}{\beta p_3} \ln[1 + (|p_2|\beta\gamma)^{p_5}] - p_4$$

Searching for highly ionizing particles
(R-hadrons, charginos in AMSB)



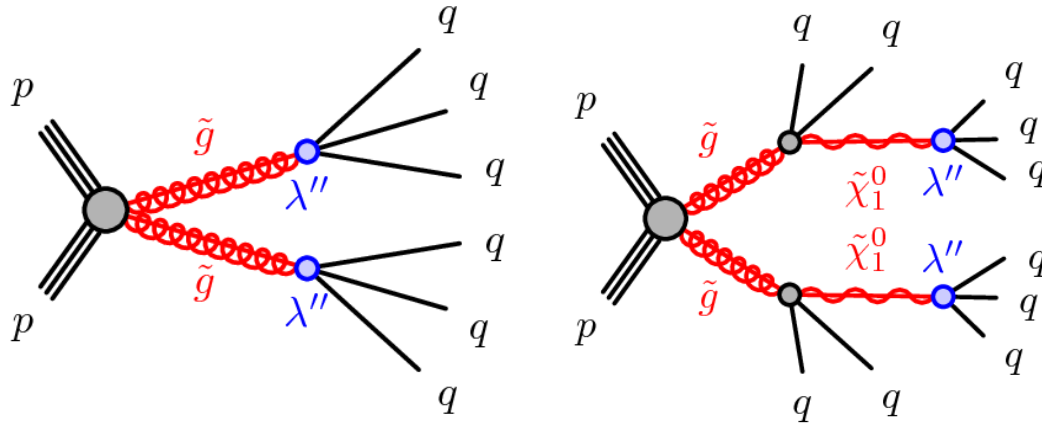
- E_{miss} > 100 GeV
- Lepton (electron and W-muon) vetoes
- One good isolated track with p_T > 80 GeV
- Highly ionizing track



Stable (metastable) charginos with masses smaller than 534 GeV (483 GeV @ 15 ns τ) are excluded⁵⁴

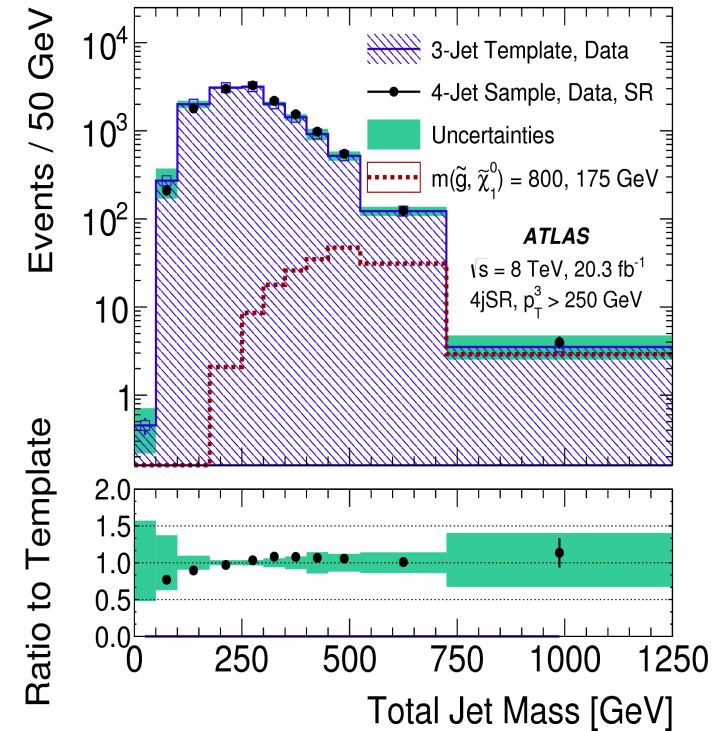
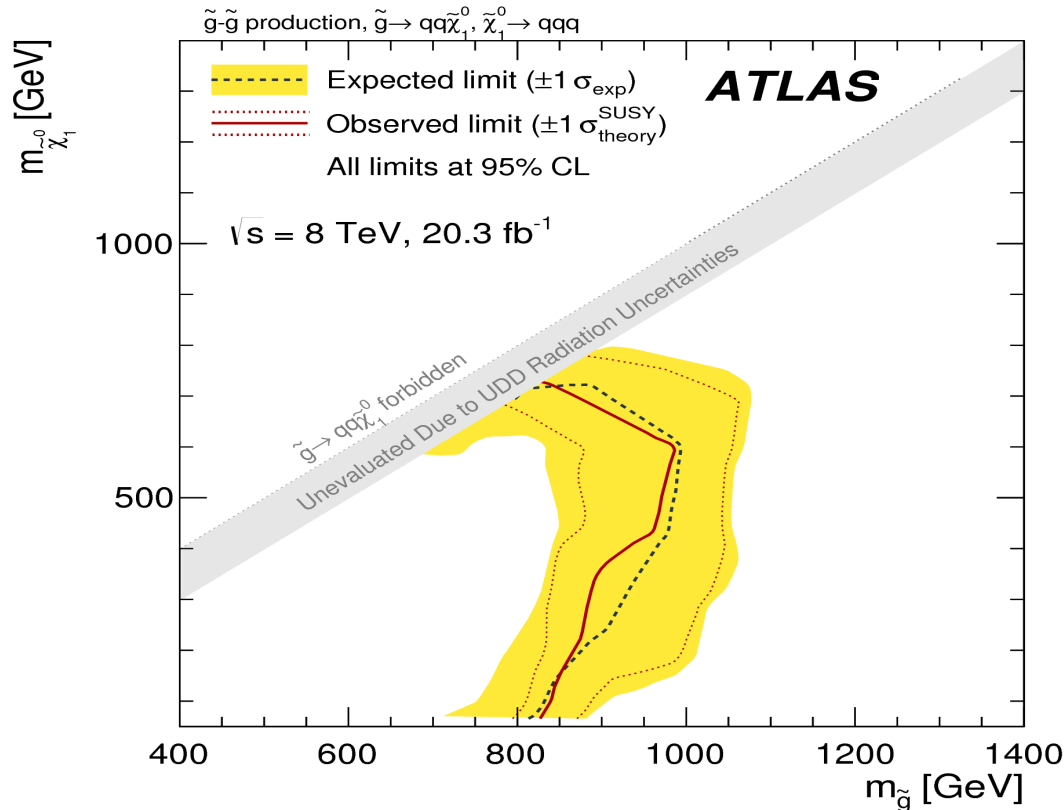
One example of RVP search

Phys. Rev. D 91, 112016 (2015)

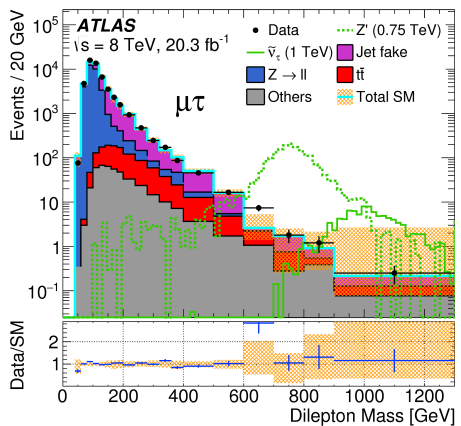
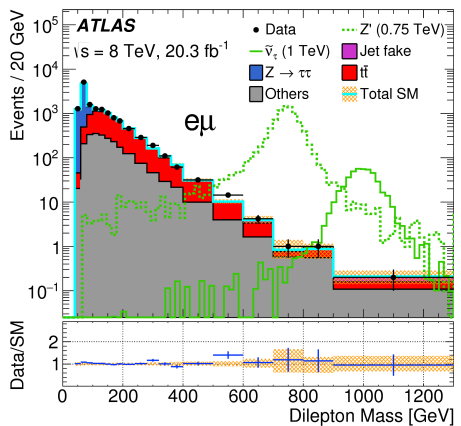
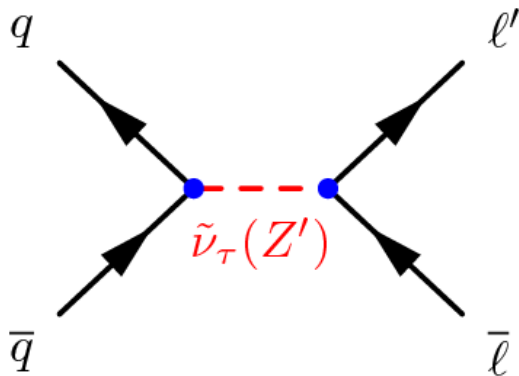


Searching for an excess of events in multijet final states

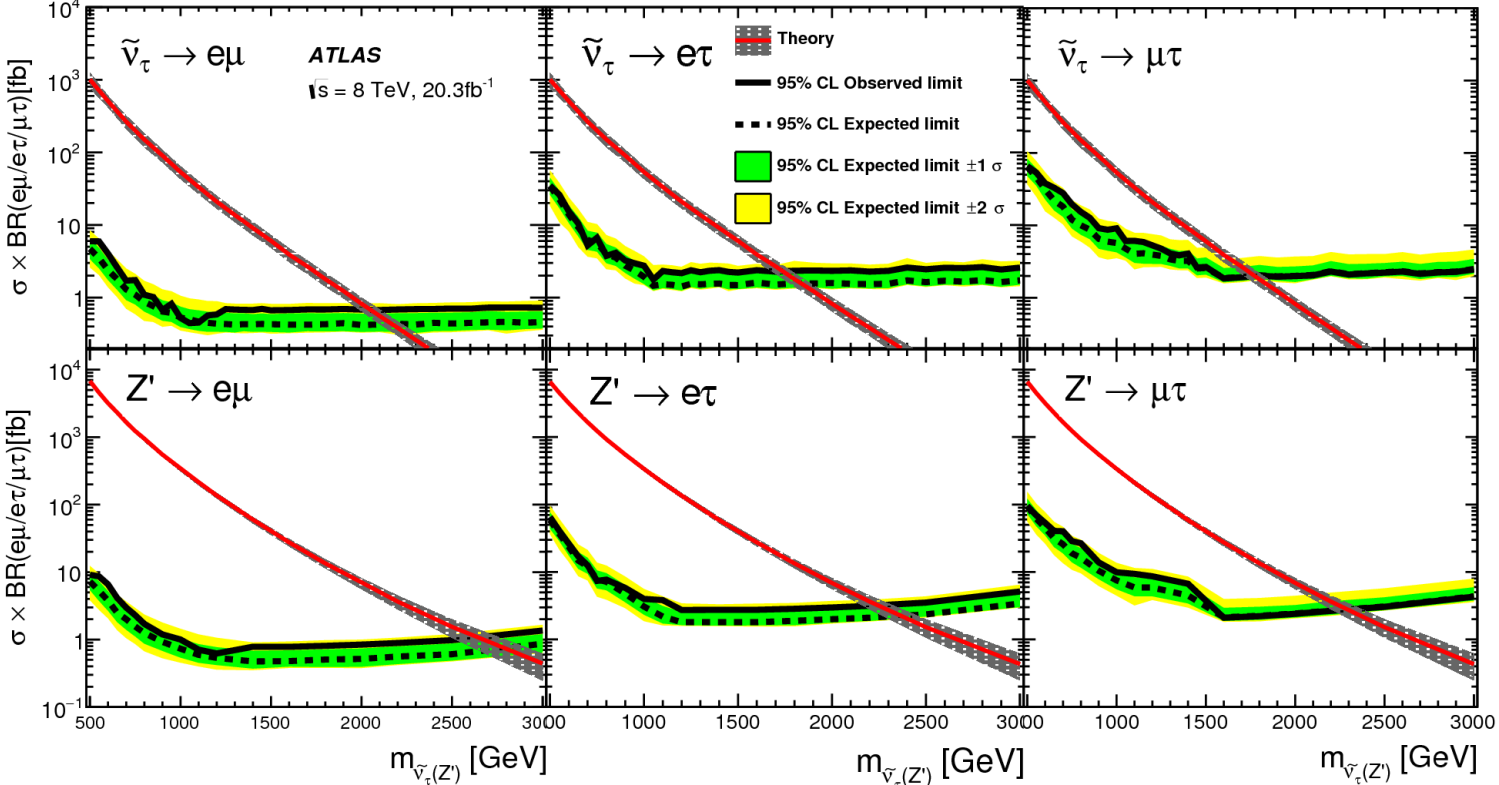
→ Computing the inv. mass of mega jets reflecting the presence of a very massive particle



RPV ($e\mu$, $e\tau$, $\mu\tau$)



Nature does not like this...



ATLAS SUSY Searches* - 95% CL Lower Limits

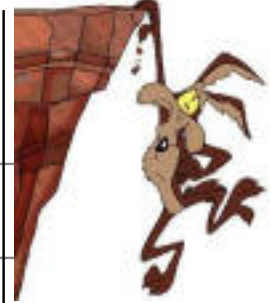
Status: July 2015

ATLAS Preliminary

$\sqrt{s} = 7, 8 \text{ TeV}$

Model	e, μ, τ, γ	Jets	E_T^{miss}	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	$\sqrt{s} = 7 \text{ TeV}$	$\sqrt{s} = 8 \text{ TeV}$	Reference
Inclusive Searches	MSUGRA/CMSSM	0-3 $e, \mu/1-2 \tau$	2-10 jets/3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.8 TeV	
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{q}	850 GeV	
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}\tilde{\chi}_1^0$ (compressed)	mono-jet	1-3 jets	Yes	20.3	\tilde{q}	100-440 GeV	
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow \tilde{q}(L\ell/L\nu/\nu\nu)\tilde{\chi}_1^0$	2 e, μ (off-Z)	2 jets	Yes	20.3	\tilde{q}	780 GeV	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{q}\tilde{q}\tilde{\chi}_1^0$	0	2-6 jets	Yes	20.3	\tilde{g}	1.33 TeV	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{q}\tilde{q}\tilde{\chi}_1^{\pm} \rightarrow \tilde{q}\tilde{q}W^{\pm}\tilde{\chi}_1^0$	0-1 e, μ	2-6 jets	Yes	20	\tilde{g}	1.26 TeV	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{q}\tilde{q}(L\ell/L\nu/\nu\nu)\tilde{\chi}_1^0$	2 e, μ	0-3 jets	-	20	\tilde{g}	1.32 TeV	
	GMSB ($\tilde{\ell}$ NLSP)	1-2 $\tau + 0-1 \ell$	0-2 jets	Yes	20.3	\tilde{g}	1.6 TeV	$m(\tilde{\chi}_1^0)=0 \text{ GeV}$ $\tan\beta > 20$ $c\tau(\text{NLSP}) < 0.1 \text{ mm}$
	GGM (bino NLSP)	2 γ	-	Yes	20.3	\tilde{g}	1.29 TeV	$m(\tilde{\chi}_1^0) < 900 \text{ GeV}, c\tau(\text{NLSP}) < 0.1 \text{ mm}, \mu < 0$
	GGM (higgsino-bino NLSP)	γ	1 b	Yes	20.3	\tilde{g}	1.3 TeV	$m(\tilde{\chi}_1^0) < 850 \text{ GeV}, c\tau(\text{NLSP}) < 0.1 \text{ mm}, \mu > 0$ $m(\text{NLSP}) > 430 \text{ GeV}$
GGM (higgsino NLSP)	2 e, μ (Z)	2 jets	Yes	20.3	\tilde{g}	850 GeV	$m(\tilde{G}) > 1.8 \times 10^{-1} \text{ eV}, m(\tilde{g})=m(\tilde{q})=1.5 \text{ TeV}$	
Gravitino LSP	0	mono-jet	Yes	20.3	$F^{1/2}$ scale	865 GeV		
3rd gen. $\tilde{g}, \text{ med.}$	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{b}\tilde{\chi}_1^0$	0	3 b	Yes	20.1	\tilde{g}	1.25 TeV	$m(\tilde{\chi}_1^0) < 400 \text{ GeV}$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^0$	0	7-10 jets	Yes	20.3	\tilde{g}	1.1 TeV	$m(\tilde{\chi}_1^0) < 350 \text{ GeV}$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow t\tilde{t}\tilde{\chi}_1^{\pm}$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.34 TeV	$m(\tilde{\chi}_1^0) < 400 \text{ GeV}$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow b\tilde{t}\tilde{\chi}_1^0$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}	1.3 TeV	$m(\tilde{\chi}_1^0) < 300 \text{ GeV}$
3rd gen. squarks direct production	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow b\tilde{\chi}_1^0$	0	2 b	Yes	20.1	\tilde{b}_1	100-620 GeV	$m(\tilde{\chi}_1^0) < 90 \text{ GeV}$
	$\tilde{b}_1\tilde{b}_1, \tilde{b}_1 \rightarrow t\tilde{\chi}_1^{\pm}$	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{b}_1	275-440 GeV	$m(\tilde{\chi}_1^{\pm}) = 2 m(\tilde{\chi}_1^0)$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\chi}_1^{\pm}$	1-2 e, μ	1-2 b	Yes	4.7/20.3	\tilde{t}_1	110-167 GeV	$m(\tilde{\chi}_1^{\pm}) = 2m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0) = 55 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow Wb\tilde{\chi}_1^0$ or $t\tilde{\chi}_1^0$	0-2 e, μ	0-2 jets/1-2 b	Yes	20.3	\tilde{t}_1	90-191 GeV	$m(\tilde{\chi}_1^0) = 1 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow c\tilde{\chi}_1^0$	0	mono-jet/c-tag	Yes	20.3	\tilde{t}_1	90-240 GeV	$m(\tilde{t}_1) - m(\tilde{\chi}_1^0) < 85 \text{ GeV}$
	$\tilde{t}_1\tilde{t}_1$ (natural GMSB)	2 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_1	150-580 GeV	$m(\tilde{\chi}_1^0) > 150 \text{ GeV}$
	$\tilde{t}_2\tilde{t}_2, \tilde{t}_2 \rightarrow \tilde{t}_1 + Z$	3 e, μ (Z)	1 b	Yes	20.3	\tilde{t}_2	290-600 GeV	$m(\tilde{\chi}_1^0) > 200 \text{ GeV}$
EW direct	$\tilde{\ell}_{L,R}\tilde{\ell}_{L,R}, \tilde{\ell} \rightarrow \tilde{\chi}_1^0$	2 e, μ	0	Yes	20.3	$\tilde{\ell}$	90-325 GeV	$m(\tilde{\chi}_1^0) = 0 \text{ GeV}$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow \tilde{\ell}\nu(\tilde{\nu})$	2 e, μ	0	Yes	20.3	$\tilde{\chi}_1^{\pm}$	140-465 GeV	$m(\tilde{\chi}_1^0) = 0 \text{ GeV}, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^{\pm} \rightarrow \tilde{\nu}\nu(\tilde{\nu}\tilde{\nu})$	2 τ	-	Yes	20.3	$\tilde{\chi}_1^{\pm}$	100-350 GeV	$m(\tilde{\chi}_1^0) = 0 \text{ GeV}, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^0 \rightarrow \tilde{\ell}_L\tilde{\nu}_L^{\pm}(\tilde{\nu}\nu), \tilde{\ell}\tilde{\nu}_L^{\pm}(\tilde{\nu}\nu)$	3 e, μ	0	Yes	20.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^0$	700 GeV	$m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0) = 0, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^{\pm}) + m(\tilde{\chi}_1^0))$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^0 \rightarrow W\tilde{\chi}_1^0 Z\tilde{\chi}_1^0$	2-3 e, μ	0-2 jets	Yes	20.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^0$	420 GeV	$m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0) = 0, \text{ sleptons decoupled}$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^0 \rightarrow W\tilde{\chi}_1^0 h\tilde{\chi}_1^0, h \rightarrow b\tilde{b}/WW/\tau\tau/\gamma\gamma$	e, μ, γ	0-2 b	Yes	20.3	$\tilde{\chi}_1^{\pm}, \tilde{\chi}_1^0$	250 GeV	$m(\tilde{\chi}_1^{\pm}) = m(\tilde{\chi}_1^0), m(\tilde{\chi}_1^0) = 0, \text{ sleptons decoupled}$
	$\tilde{\chi}_2^0\tilde{\chi}_3^0, \tilde{\chi}_2^0 \rightarrow \tilde{\ell}_R\tilde{\ell}$	4 e, μ	0	Yes	20.3	$\tilde{\chi}_2^0, \tilde{\chi}_3^0$	620 GeV	$m(\tilde{\chi}_2^0) = m(\tilde{\chi}_3^0), m(\tilde{\chi}_1^0) = 0, m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_2^0) + m(\tilde{\chi}_1^0))$
	GGM (wino NLSP) weak prod.	1 $e, \mu + \gamma$	-	Yes	20.3	\tilde{W}	124-361 GeV	$c\tau < 1 \text{ mm}$
	Long-lived particles	Direct $\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ prod., long-lived $\tilde{\chi}_1^{\pm}$	Disapp. trk	1 jet	Yes	20.3	$\tilde{\chi}_1^{\pm}$	270 GeV
Direct $\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}$ prod., long-lived $\tilde{\chi}_1^{\pm}$		dE/dx trk	-	Yes	18.4	$\tilde{\chi}_1^{\pm}$	482 GeV	$m(\tilde{\chi}_1^{\pm}) - m(\tilde{\chi}_1^0) \sim 160 \text{ MeV}, \tau(\tilde{\chi}_1^{\pm}) < 15 \text{ ns}$
Stable, stopped \tilde{g} R-hadron		0	1-5 jets	Yes	27.9	\tilde{g}	832 GeV	$m(\tilde{\chi}_1^0) = 100 \text{ GeV}, 10 \mu\text{s} < \tau(\tilde{g}) < 1000 \text{ s}$
Stable \tilde{g} R-hadron		trk	-	-	19.1	\tilde{g}	1.27 TeV	1411.6795
GMSB, stable $\tilde{\tau}, \tilde{\chi}_1^0 \rightarrow \tilde{\tau}(\tilde{e}, \tilde{\mu}) + \tau(e, \mu)$		1-2 μ	-	-	19.1	$\tilde{\chi}_1^0$	537 GeV	$10 < \tan\beta < 50$
GMSB, $\tilde{\nu}_1^0 \rightarrow \gamma\tilde{G}$, long-lived $\tilde{\chi}_1^0$		2 γ	-	Yes	20.3	$\tilde{\chi}_1^0$	435 GeV	$2 < \tau(\tilde{\chi}_1^0) < 3 \text{ ns}, \text{ SPS8 model}$
$\tilde{g}\tilde{g}, \tilde{G} \rightarrow ee/\mu\nu/\mu\mu\nu$		displ. $ee/\mu\nu/\mu\mu\nu$	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$7 < c\tau(\tilde{\chi}_1^0) < 740 \text{ mm}, m(\tilde{g}) = 1.3 \text{ TeV}$
GGM $\tilde{g}\tilde{g}, \tilde{\chi}_1^0 \rightarrow Z\tilde{G}$	displ. vtx + jets	-	-	20.3	$\tilde{\chi}_1^0$	1.0 TeV	$6 < c\tau(\tilde{\chi}_1^0) < 480 \text{ mm}, m(\tilde{g}) = 1.1 \text{ TeV}$	
RPV	LFV $pp \rightarrow \tilde{\nu}_\tau + X, \tilde{\nu}_\tau \rightarrow e\mu/\tau\mu/\mu\tau$	$e\mu, e\tau, \mu\tau$	-	-	20.3	$\tilde{\nu}_\tau$	1.7 TeV	$A'_{311} = 0.11, \lambda_{132/133/233} = 0.07$
	Bilinear RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{q}, \tilde{g}	1.35 TeV	$m(\tilde{g}) = m(\tilde{q}), c\tau_{LSP} < 1 \text{ mm}$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}, \tilde{\chi}_1^{\pm} \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow ee\tilde{\nu}_\mu, e\mu\tilde{\nu}_e$	4 e, μ	-	Yes	20.3	$\tilde{\chi}_1^{\pm}$	750 GeV	$m(\tilde{\chi}_1^0) > 0.2 \times m(\tilde{\chi}_1^{\pm}), \lambda_{121} = 0$
	$\tilde{\chi}_1^{\pm}\tilde{\chi}_1^{\mp}, \tilde{\chi}_1^{\pm} \rightarrow W\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow \tau\tau\tilde{\nu}_e, e\tau\tilde{\nu}_\tau$	3 $e, \mu + \tau$	-	Yes	20.3	$\tilde{\chi}_1^{\pm}$	450 GeV	$m(\tilde{\chi}_1^0) > 0.2 \times m(\tilde{\chi}_1^{\pm}), \lambda_{133} = 0$ $\text{BR}(j) = \text{BR}(b) = \text{BR}(c) = 0\%$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{q}$	0	6-7 jets	-	20.3	\tilde{g}	917 GeV	$m(\tilde{\chi}_1^0) = 600 \text{ GeV}$
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{q}$	0	6-7 jets	-	20.3	\tilde{g}	870 GeV	
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow \tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{g}	850 GeV	
	$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow bs$	0	2 jets + 2 b	-	20.3	\tilde{t}_1	100-308 GeV	
$\tilde{t}_1\tilde{t}_1, \tilde{t}_1 \rightarrow b\tilde{\ell}$	2 e, μ	2 b	-	20.3	\tilde{t}_1	0.4-1.0 TeV	$\text{BR}(\tilde{t}_1 \rightarrow b\ell/\mu) > 20\%$	
Other	Scalar charm, $\tilde{c} \rightarrow c\tilde{\chi}_1^0$	0	2 c	Yes	20.3	\tilde{c}	490 GeV	$m(\tilde{\chi}_1^0) < 200 \text{ GeV}$

No hint for SUSY ?



1209.2102, 1407.0583
1506.08616
1407.0608
1403.5222
1403.5222

1403.5294
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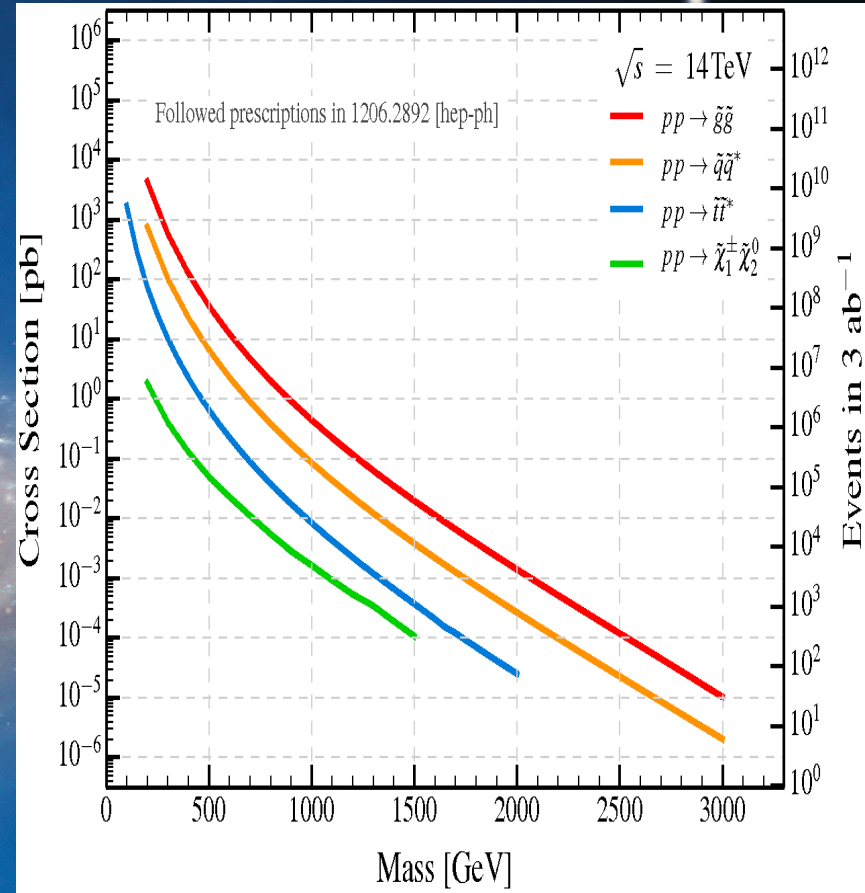
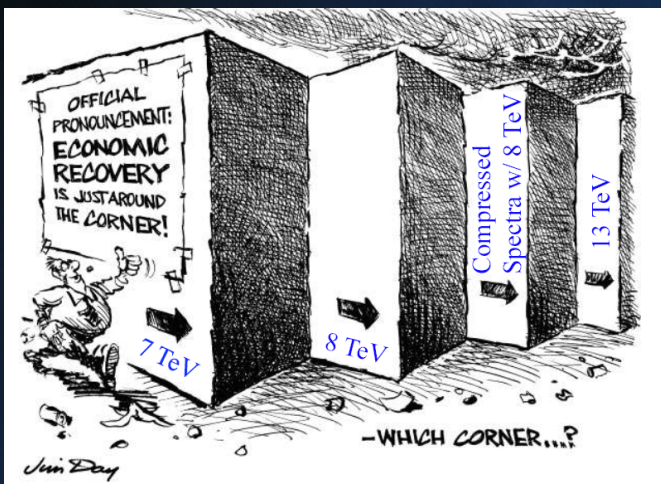
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1310.6584
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1411.6795
1409.5542
1504.05162
1504.05162

1503.04430
1404.2500
1405.5086
1405.5086
1502.05686
1502.05686
1404.250

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10^{-1} 1 Mass scale [TeV]

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1σ theoretical signal cross section uncertainty.



The discovery of New Physics requires more energy and more data

End Part II