



University of
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Summary and re-optimized Performance Document

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Introduction

Technical Proposal

Now



- SHiP is sailing very fast
- And changing a lot!
- The idea is to write a document for the re-optimisation phase

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New Simulation

First iteration (Intermediate document):

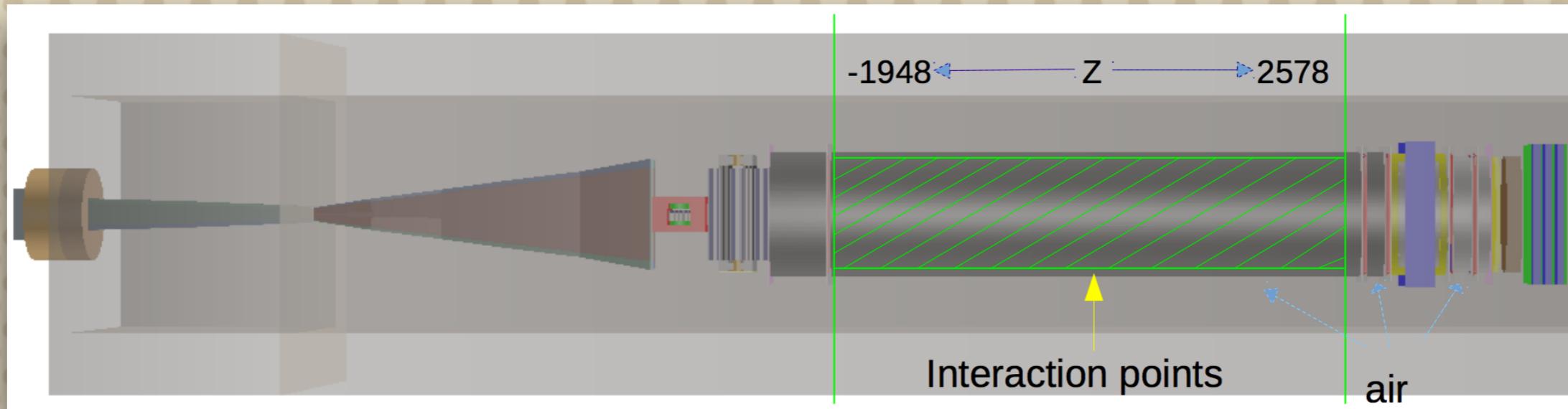
- Geometry: dimensions of the decay volume, starting and ending point, shape, ...
- Vacuum VS He: Looks like the driving parameter here is the cost
- PID performances: As Gaia pointed out we should have a strategy to decide on the requirements

We need to do performance studies with the re-optimized SHiP and updated cost -> "progress report" to the SPSC

These studies will be the input to a more advanced/longer document which will include all sub detector details

Current Status of Background Studies

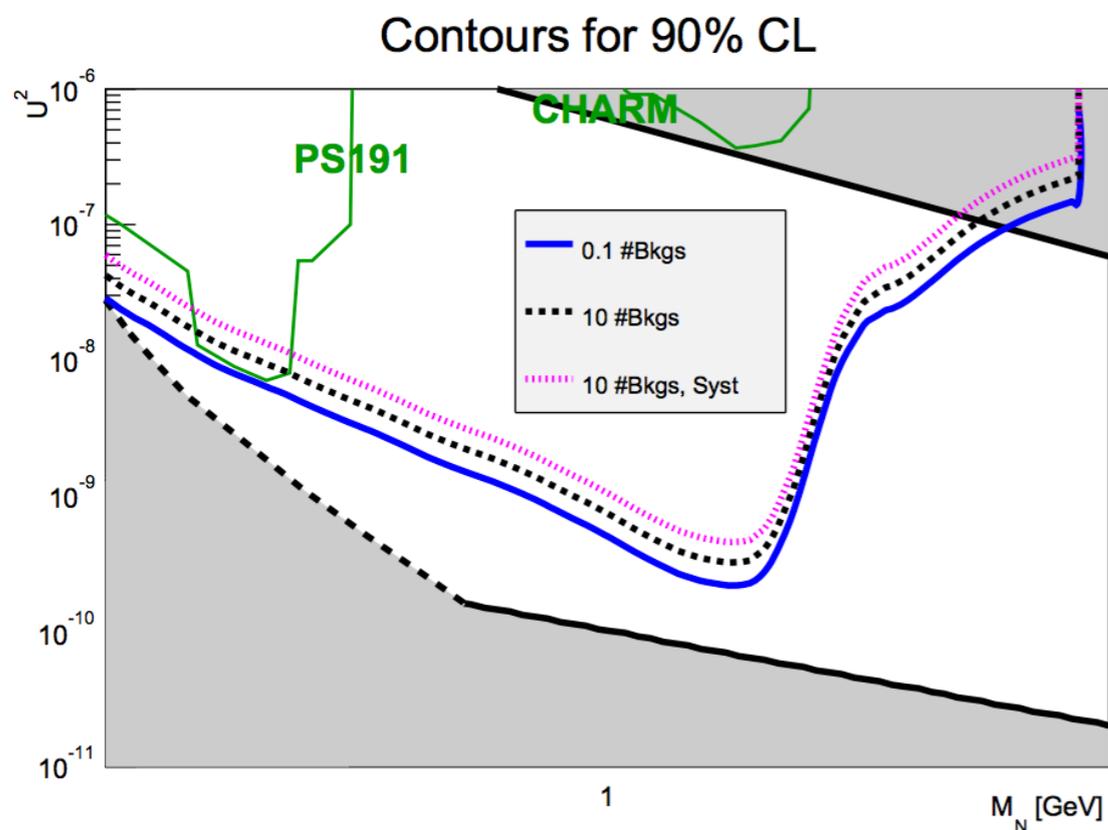
Iaroslava studied the impact of neutrino in the air



- Fully reconstructed signal $N \rightarrow \mu \pi$: 0(1) bkg events with or without veto
- Partially reconstructed signal: 40 events without veto, 4 with the veto, zero with the PID (see also Behzad)
- Take these numbers with care: we only studied in details HNLs, we cannot afford to narrow down our physics case

Status of Signals

- We are shifting from a paradigm where we had 0.1 expected total bkg, to a paradigm where we study exclusive final states
- This is good, but we need to study the implications for all signals, in collaboration with the people studying the various signals



- Old plot in the addendum should be updated with the re-optimized geometry and the new paradigm

Status of Signals

- R-parity neutralino (implemented in FairSHiP by Kostas)
- Dark Scalar (work in progress by Gaia)
- Axino (one presentation by Ki Young Choi, but very slow or no progress)
- ALPs (background studies from Walter), need implementation
- Dark Photon is one of our main channels, no FairSHiP implementation
- Dirac goldstino (Alexey/Oleg did studies with toys, we need FairSHiP implementation)

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“New” Signals

Photophobic Dark Photon :-)

Evidence for a Protophobic Fifth Force from ^8Be Nuclear Transitions

Jonathan L. Feng,¹ Bartosz Fornal,¹ Iftah Galon,¹ Susan Gardner,^{1,2}
Jordan Smolinsky,¹ Tim M. P. Tait,¹ and Philip Tanedo¹

[arXiv:1604.07411](https://arxiv.org/abs/1604.07411)

¹Department of Physics and Astronomy, University of California, Irvine, California 92697-4575 USA

²Department of Physics and Astronomy, University of Kentucky, Lexington, Kentucky 40506-0055 USA

We should understand what is our sensitivity since they cite SHiP in their paper

Strongly interacting Dark Matter

The SIMPlEst Miracle

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⁷Stanford Institute for Theoretical Physics, Stanford University, Stanford, CA 94305 USA

[arXiv:1411.3727](https://arxiv.org/abs/1411.3727)

Beautiful idea predicts 1GeV particles of the type we are looking AND Dark Matter copiously produced in SHiP

Dark Matter Searches

- Several models predicts large DM fluxes in SHiP-like experiments
- Andrey, Giovanni, Yandex are studying the idea of using the emulsion spectrometer for these searches
- This would be a major addition to our physics case
- Remember that if the mediator of the Dark Sector couples with DM and if DM particles are light enough the Dark Sector particles will not reach out detector

Re-optimization

- We need to update all background studies in the re-optimization configuration:
 - Muon inelastic (how about interaction in He?)
 - Cosmic muons
 - Neutrino scattering in the material
 - Combinatorial
- These backgrounds were not studied with the same details as the bkg neutrino in the air
- We should have a strategy to actually prove with data

Summary

- We are in the re-optimization phase and we need to have answer about Geometry, He/Vacuum, PID requirements
- We should make performance evaluation studies in the re-optimized scenario (also we should get on with these studies to make the decisions)
- Prepare a short 30ish page document spring next year with updated sensitivities and background

