

***Development of Particle Detectors at TIFR  
for  
Underground Experiments-  
Then and Now***

***Naba K Mondal, INO Cell, TIFR***

# **Prelude**

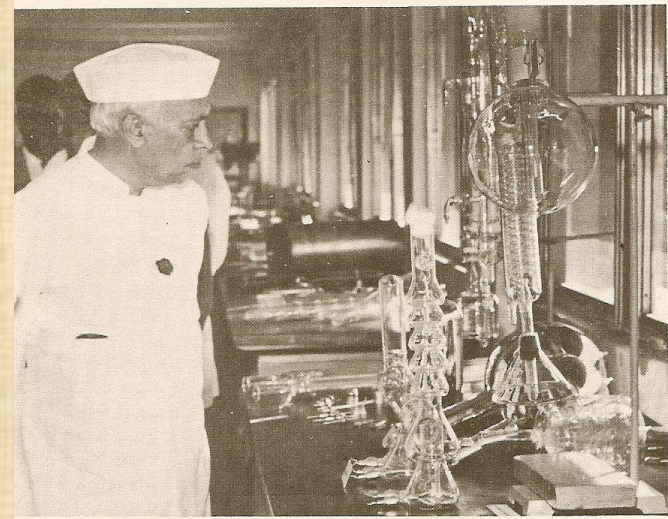
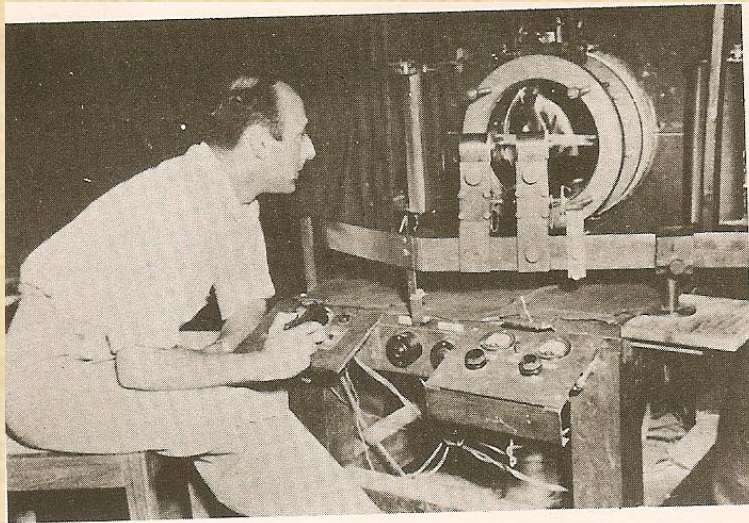
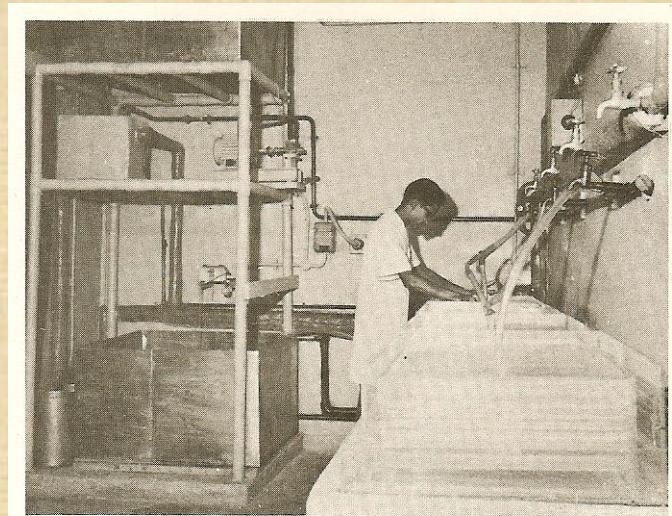
## **Cosmic Ray Research Unit at Bangalore**

- ***Seeds for Experimental High Energy Physics research at TIFR were really sown even before TIFR came into existence on 1<sup>st</sup> June, 1945.***
- ***In December, 1944 Homi Bhabha flew two telesopes with GM counters sandwiched with lead plates using US airplanes stationed at Bangalore during the war.***
  - ***First measurement of high altitude meson intensities at equatorial latitude.***
- ***Built a 12" circular cloud chamber similar to one available at Blackett's Lab in UK.***
  - ***Used it to study the scattering characteristics of mesons***



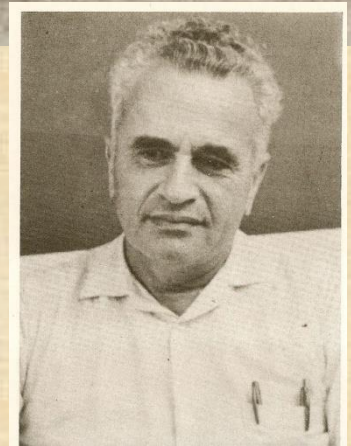
## *Early Particle physics experiments at TIFR*

- *Cloud Chambers.*
- *G.M. Counters.*
- *Emulsions.*





# *Conference on Elementary Particles at TIFR in 1950*







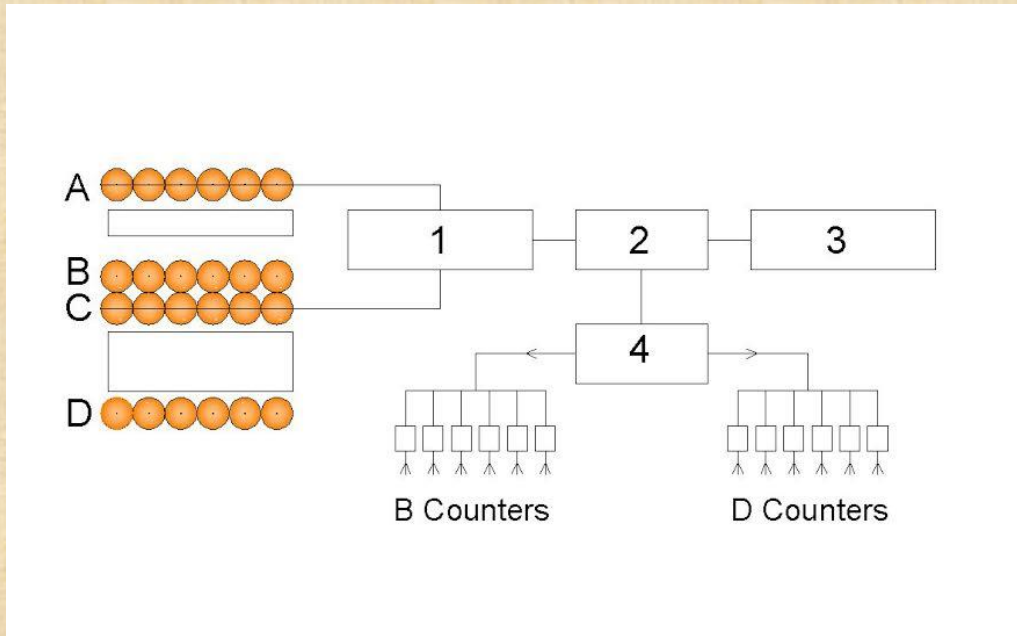
## *1951 Start of KGF activities*



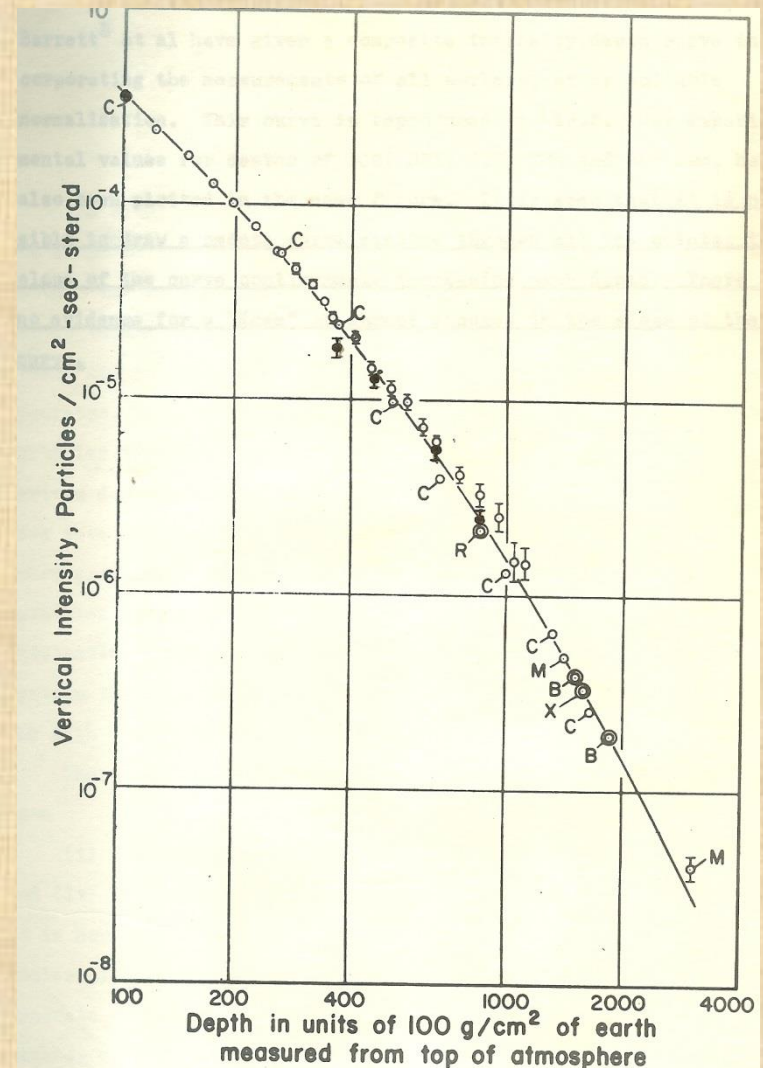
**First set of measurements were carried out at Nundydurg Mines**



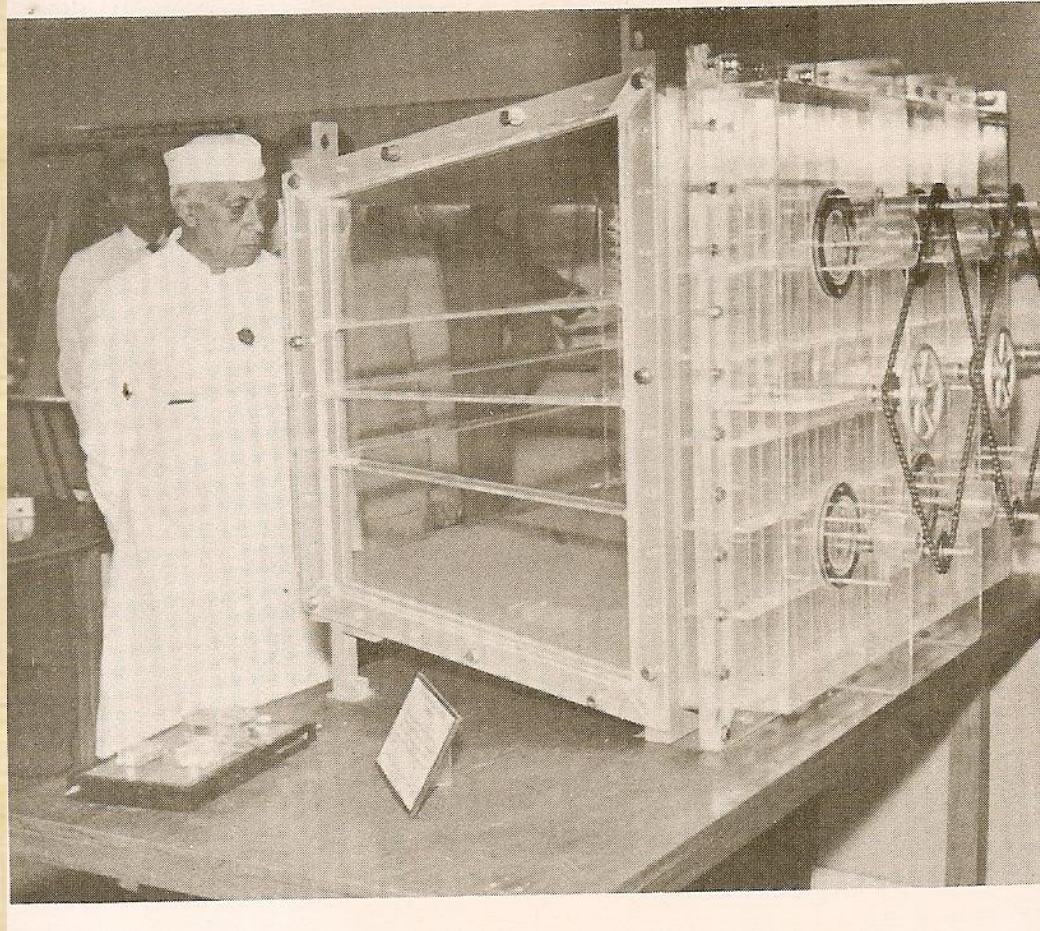
# KGF 1951-1954 - GM Counter array upto a depth of 900 ft ( $E_{\mu} \sim 200$ GeV)



1. Double coincidence circuits.
2. Cathode follower.
3. Film advancing circuits.
4. Inverter.



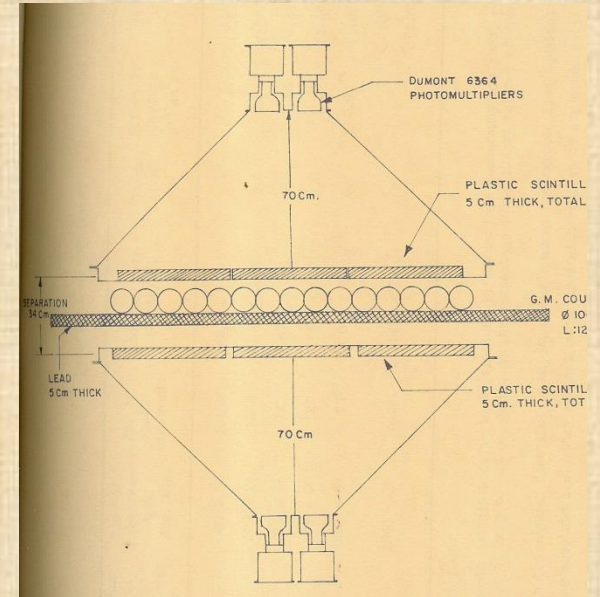
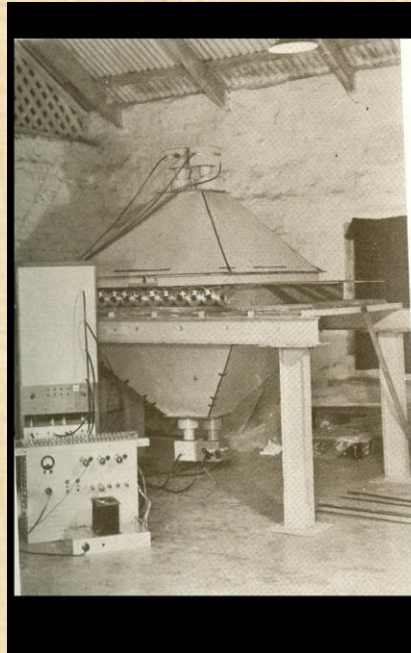
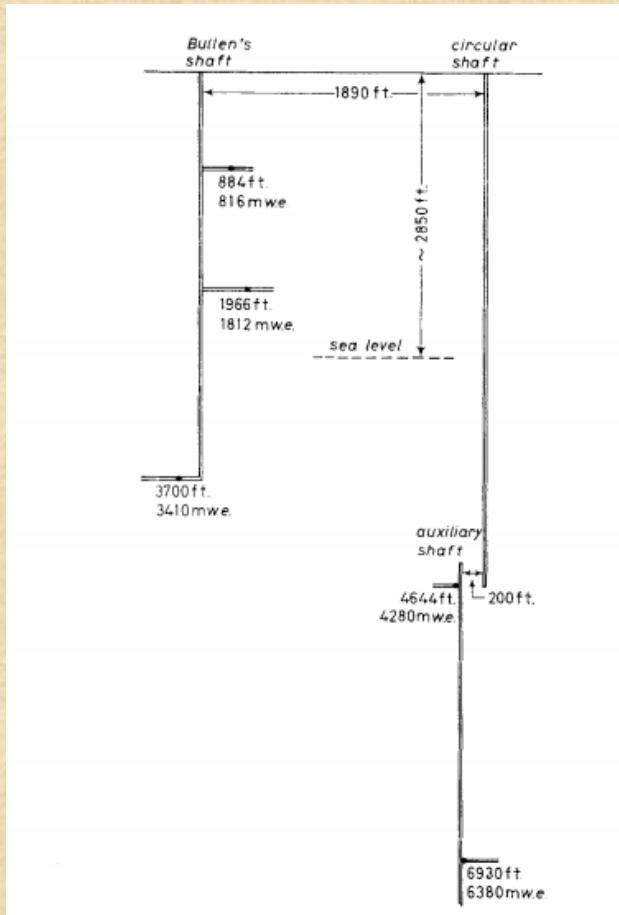
# *Large Cloud Chamber*





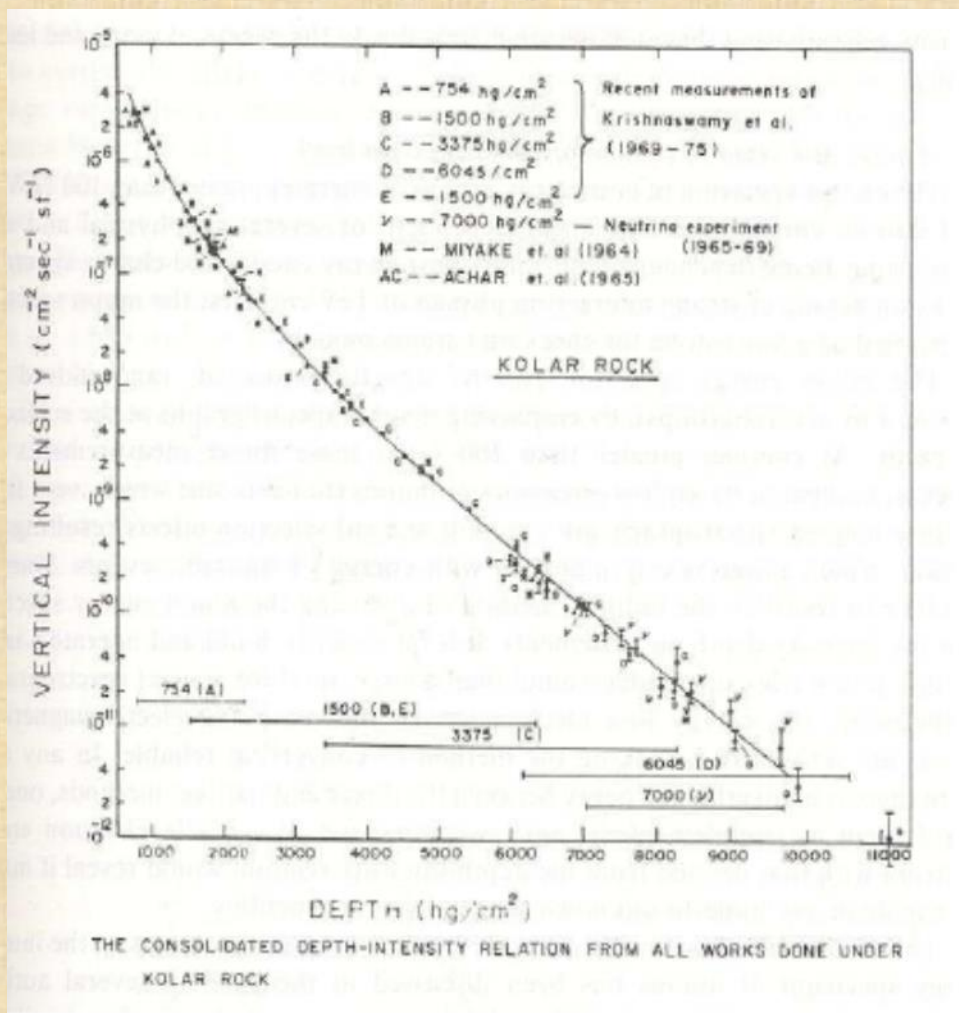
# KGF 1960-1963

## Up to a depth of 2.7 km. ( 800 – 8400 mwe)



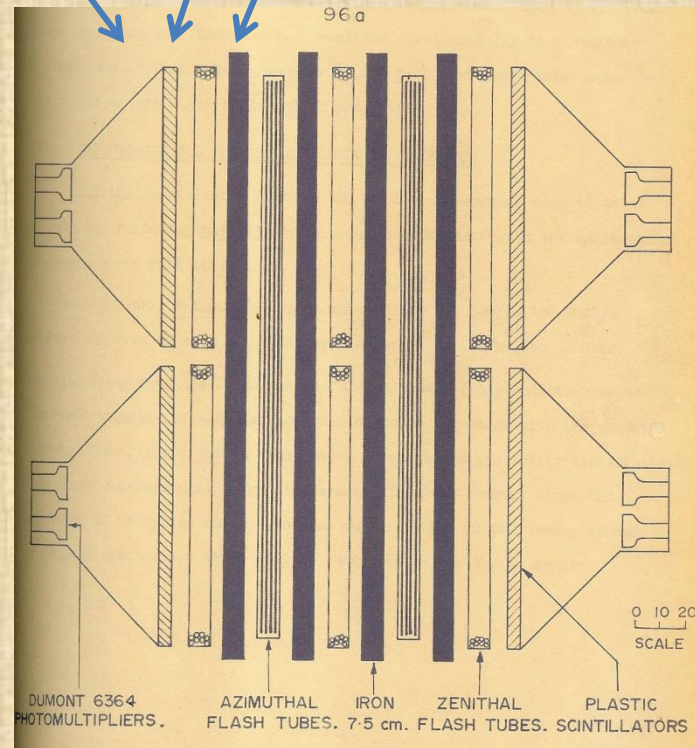
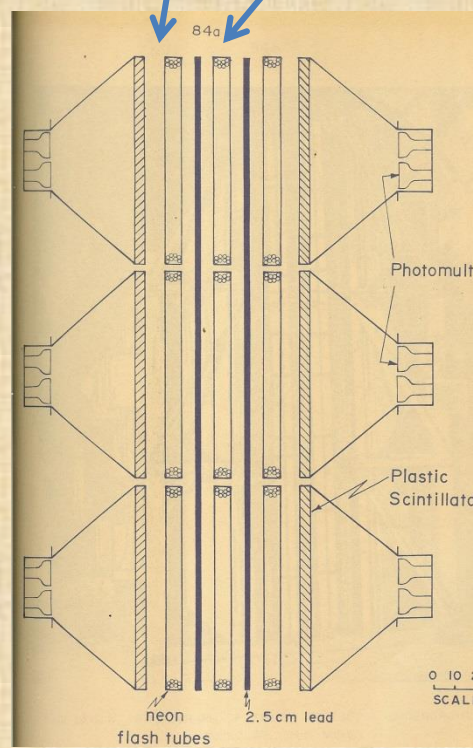
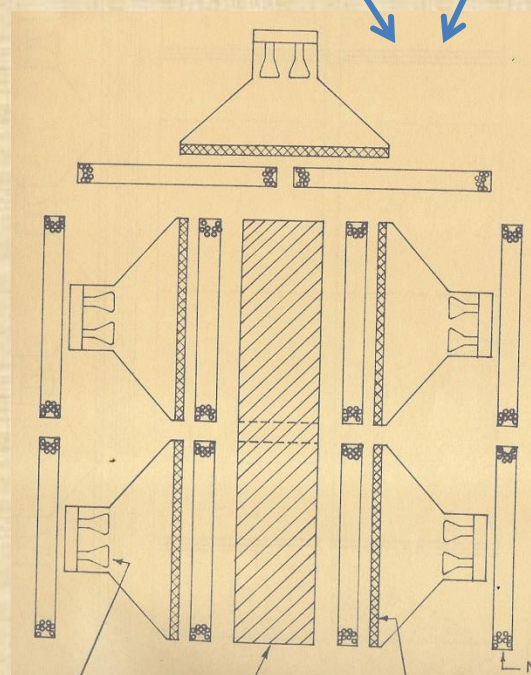
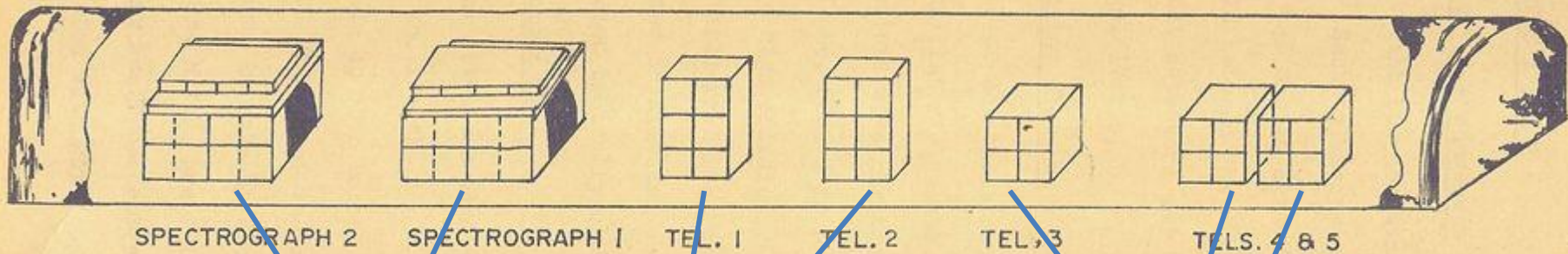
Miyake, Narasimham, Ramanamurthy

# Most comprehensive Depth-intensity curve



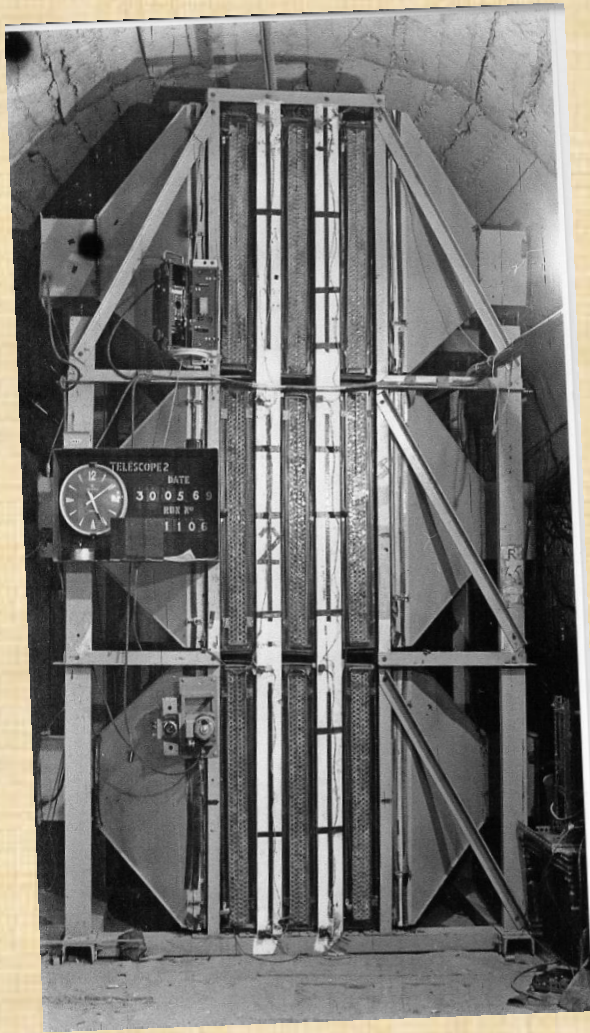


# Neutrino Experiment at KGF 1965 onwards





# Atmospheric neutrino detection in 1965



Atmospheric neutrino detector  
at Kolar Gold Field –1965

## DETECTION OF MUONS PRODUCED BY COSMIC RAY NEUTRINO DEEP UNDERGROUND

C. V. ACHAR, M. G. K. MENON, V. S. NARASIMHAM, P. V. RAMANA MURTHY  
and B. V. SREEKANTAN,

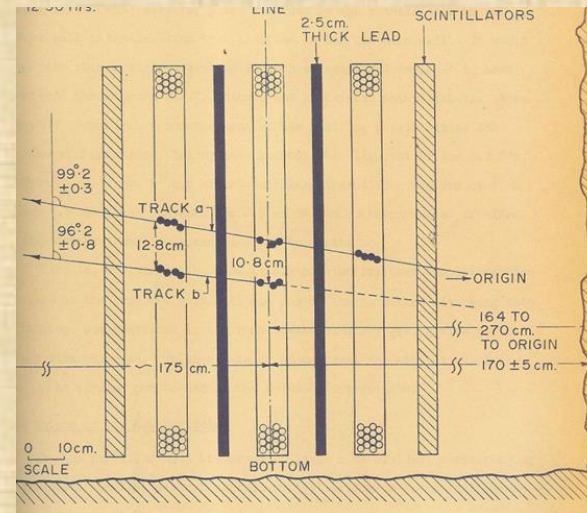
*Tata Institute of Fundamental Research, Colaba, Bombay*

K. HINOTANI and S. MIYAKE,  
*Osaka City University, Osaka, Japan*

D. R. CREED, J. L. OSBORNE, J. B. M. PATTISON and A. W. WOLFENDALE  
*University of Durham, Durham, U.K.*

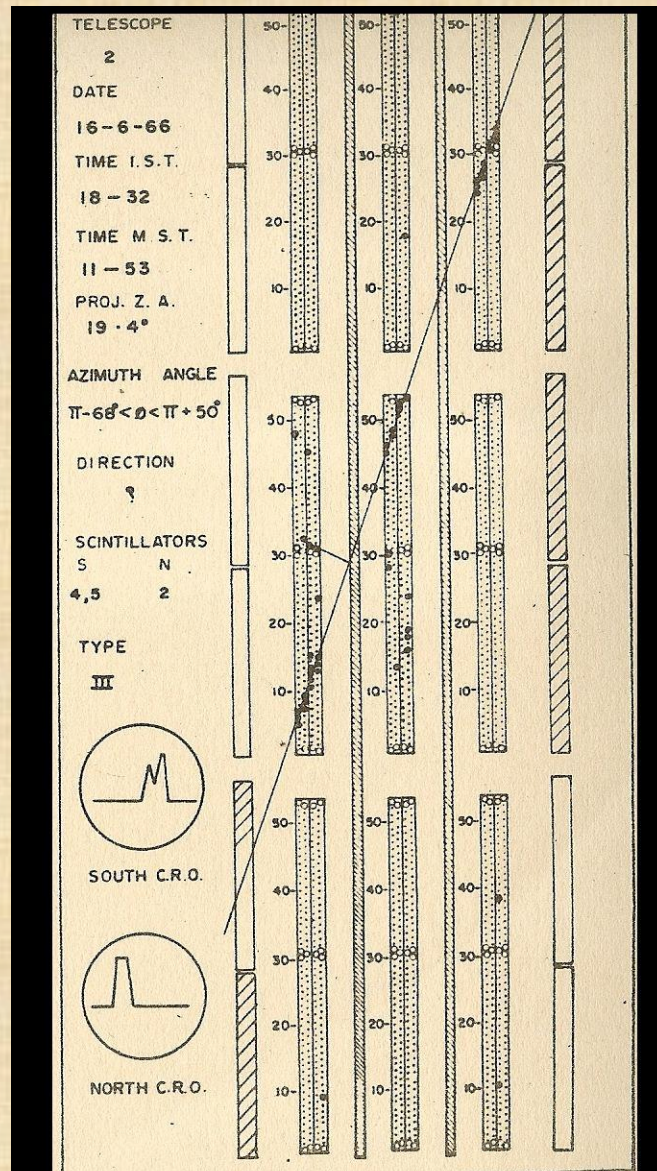
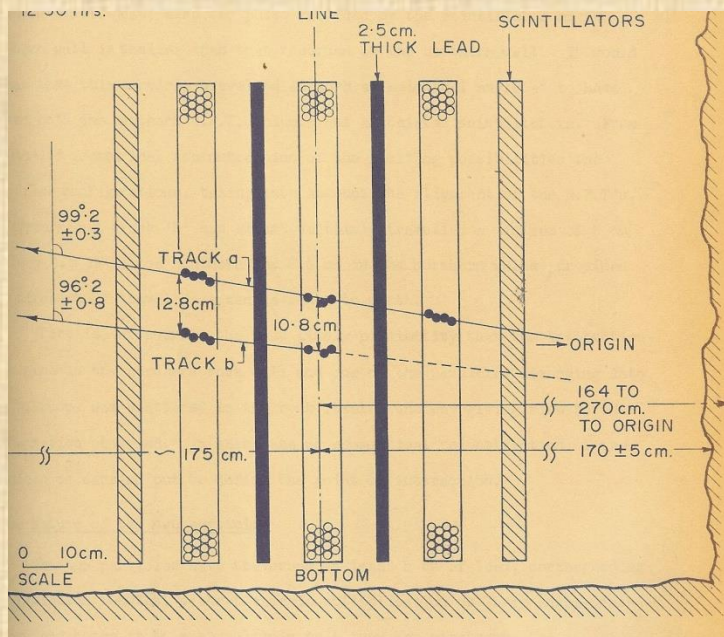
Received 12 July 1965

Physics Letters 18, (1965) 196, dated 15th Aug 1965

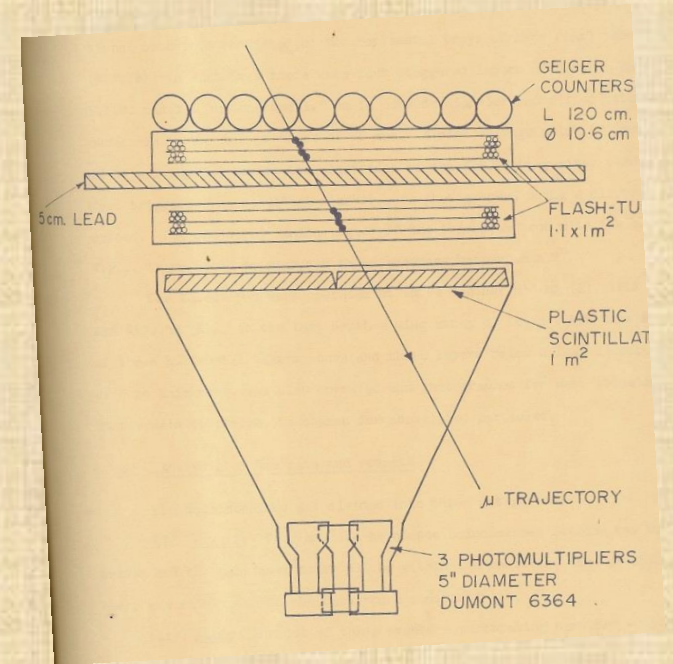
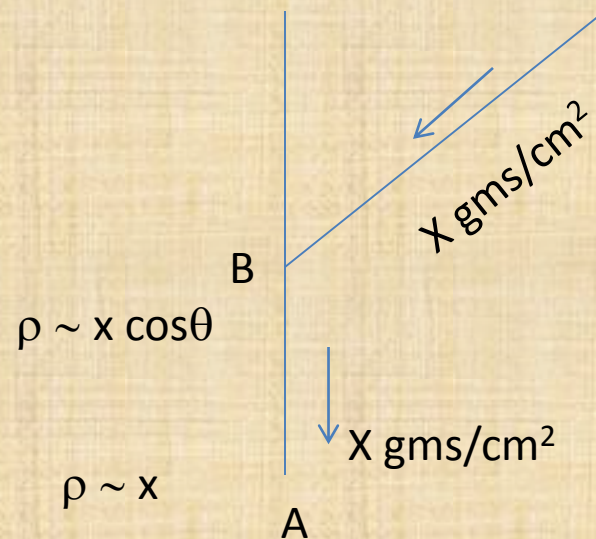




# Neutrino Events at KGF



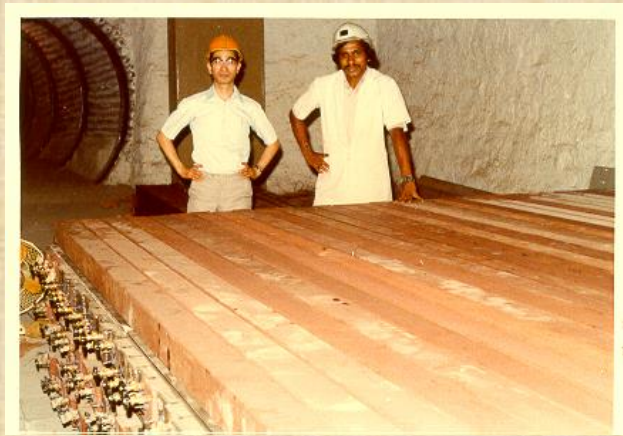
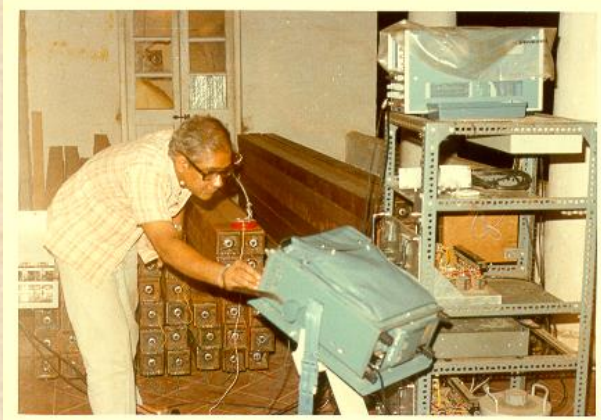
# Muon Angular Distribution Deep Underground



- At very small residual thickness of the atmosphere, ( $< 200 \text{ gms/cm}^2$ ), the density of air is proportional to the thickness of the atmosphere itself.
- Decay probability is inversely proportional to density and thus increases as  $\sec\theta$
- Any deviation from this  $\sec\theta$  will suggest new source of atmospheric muons other than pion/kaon decay



# ***KGF Phase I Nucleon Decay Detector (1979-1992)***





# *KGF Phase-I Nucleon Decay Detector*





# ***ICOBAN 1982***





# *Inauguration of KGF Phase-II Nucleon Decay Laboratory (1983)*





# ***KGF Phase-II Nucleon decay Experiment 1984-1992***



# Neutrino Events

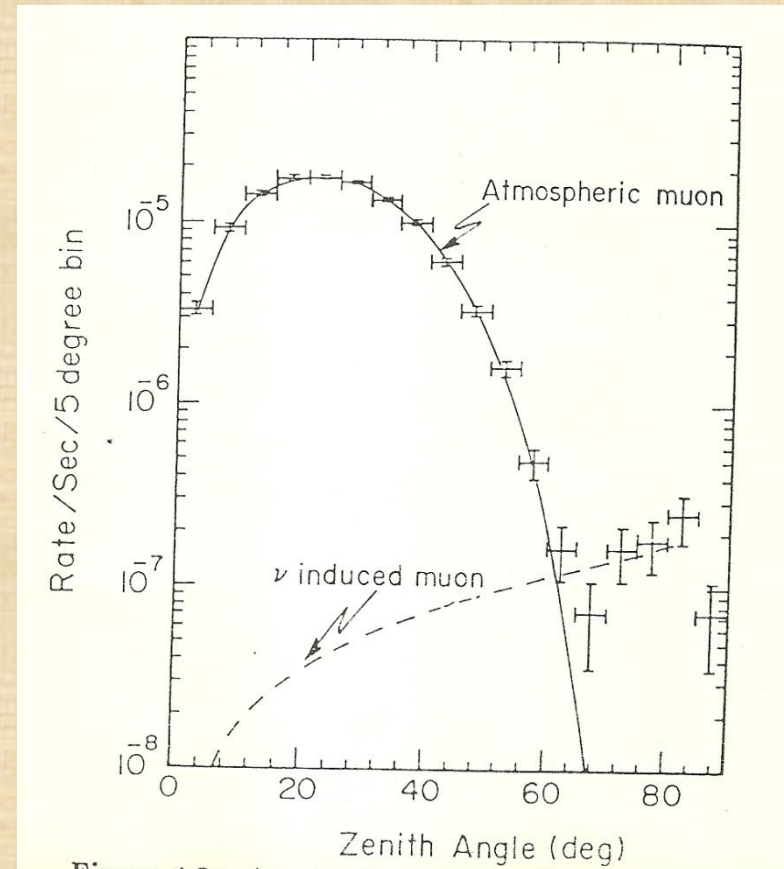
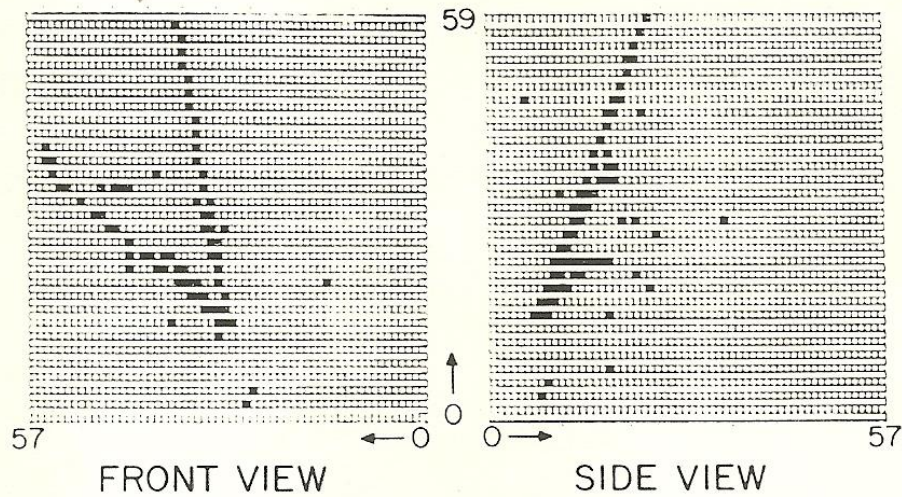


Figure 1.0

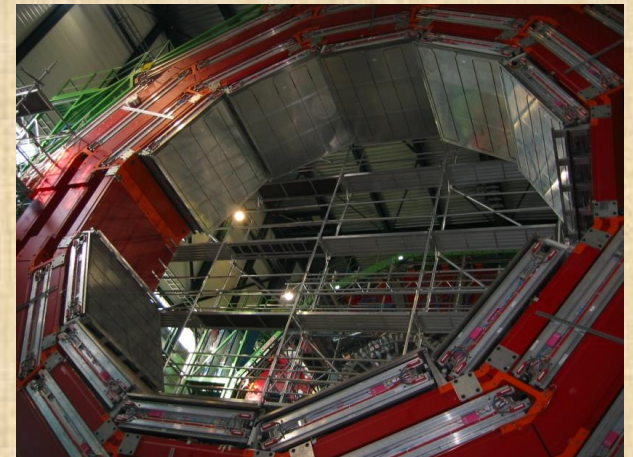
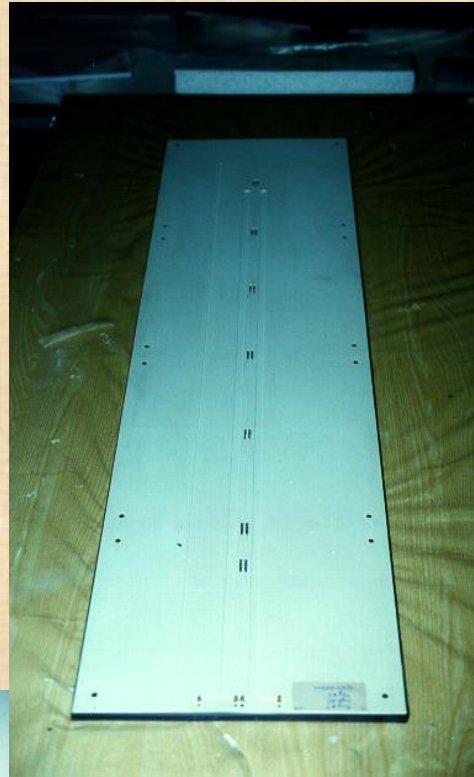
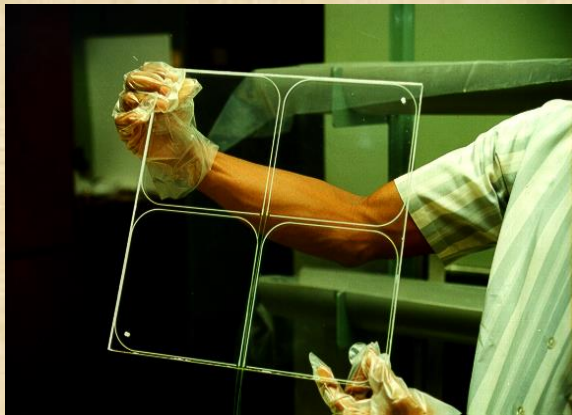


# ***DØ Detector fabrication***

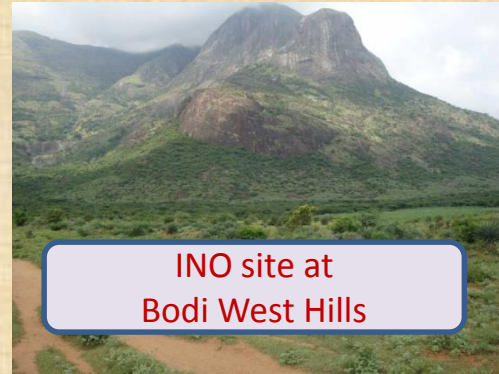
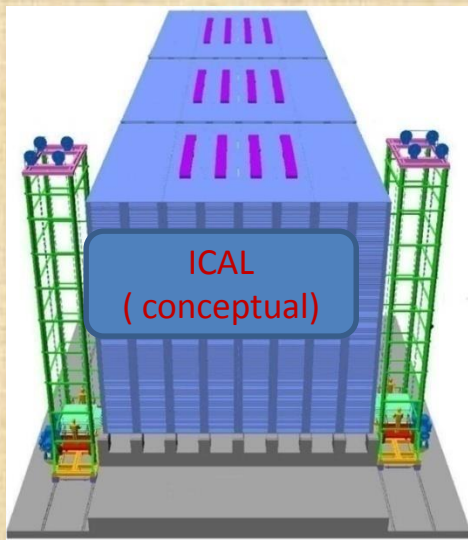




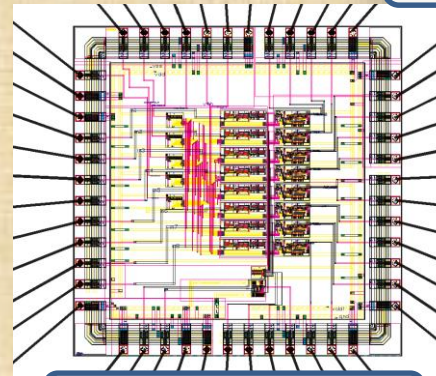
# ***CMS Detector Fabrication work***





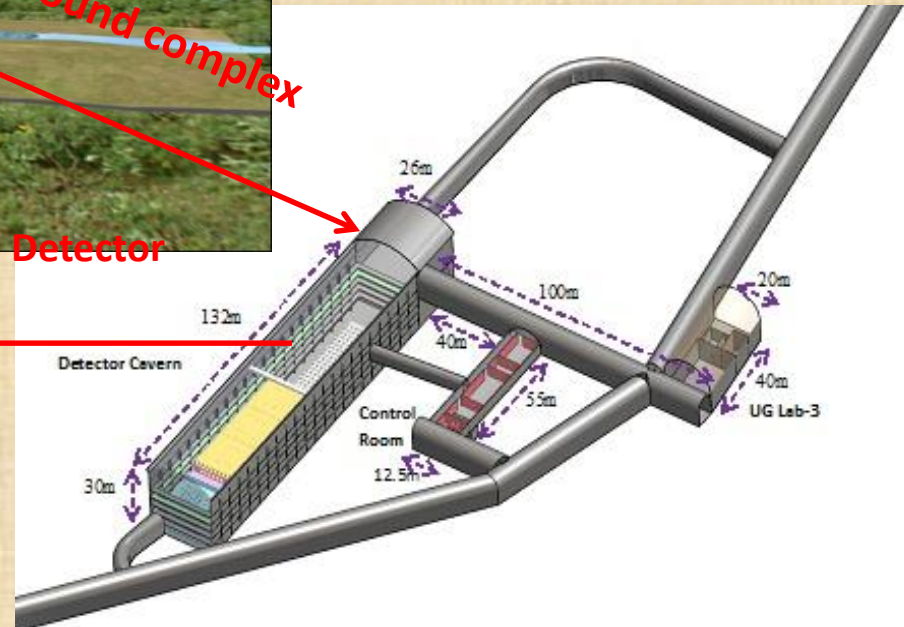
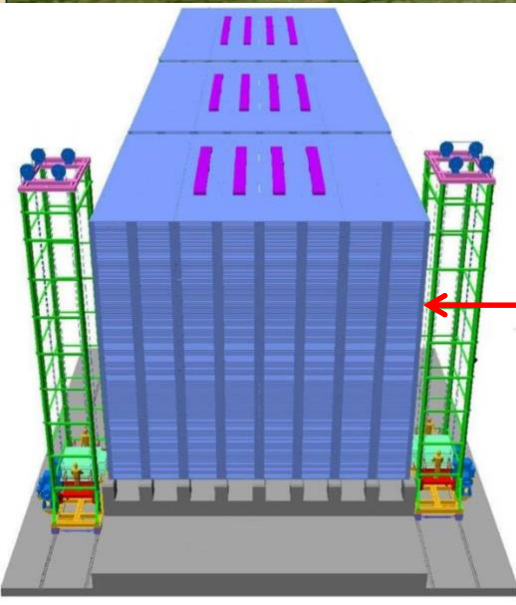
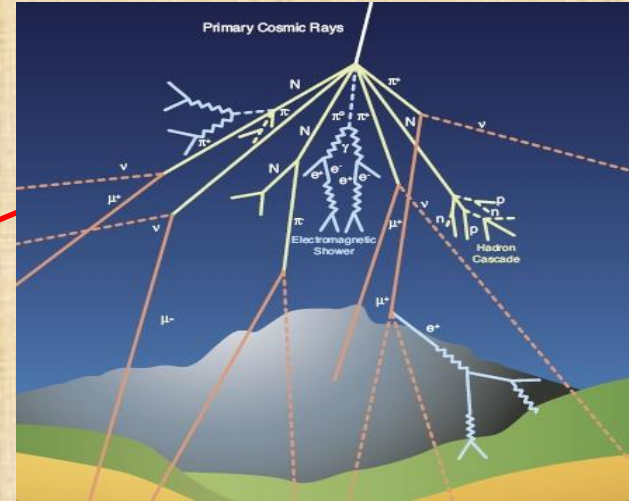
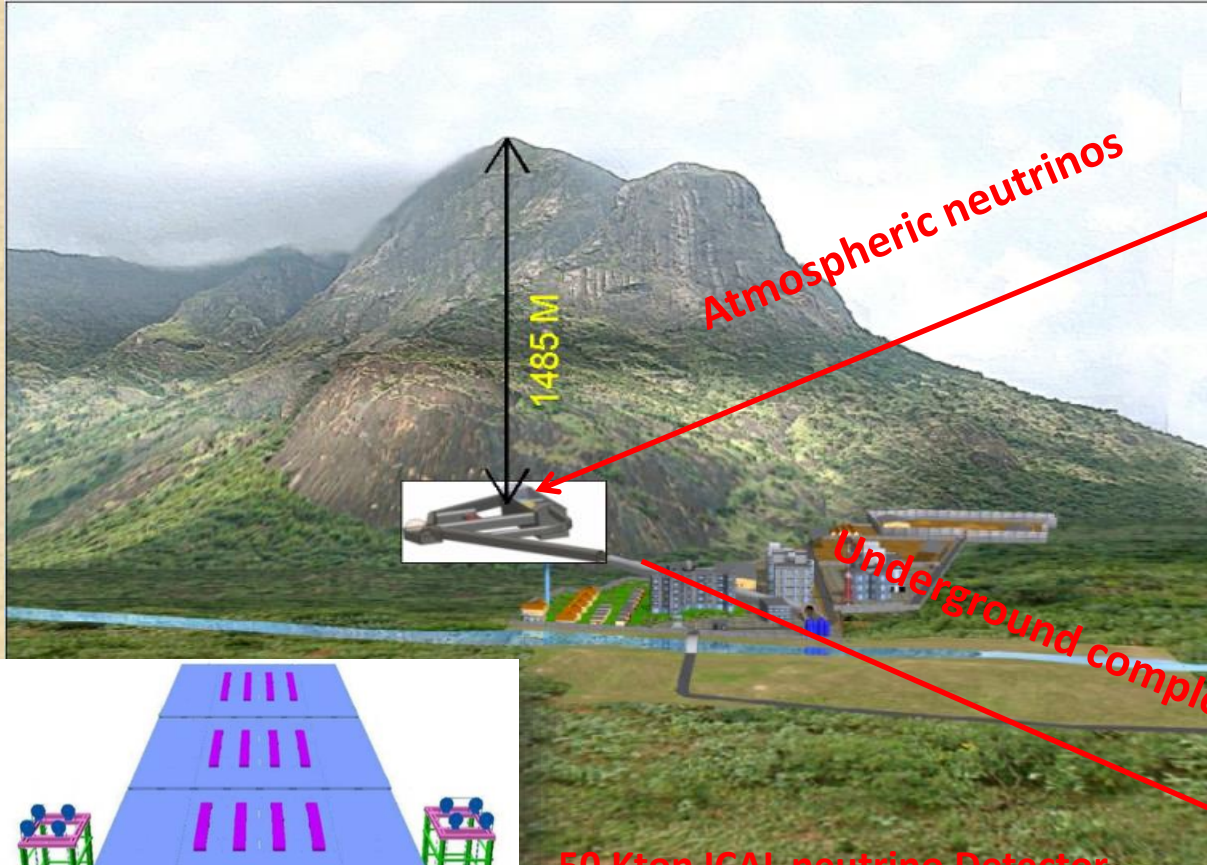


# India-Based Neutrino Observatory (INO)



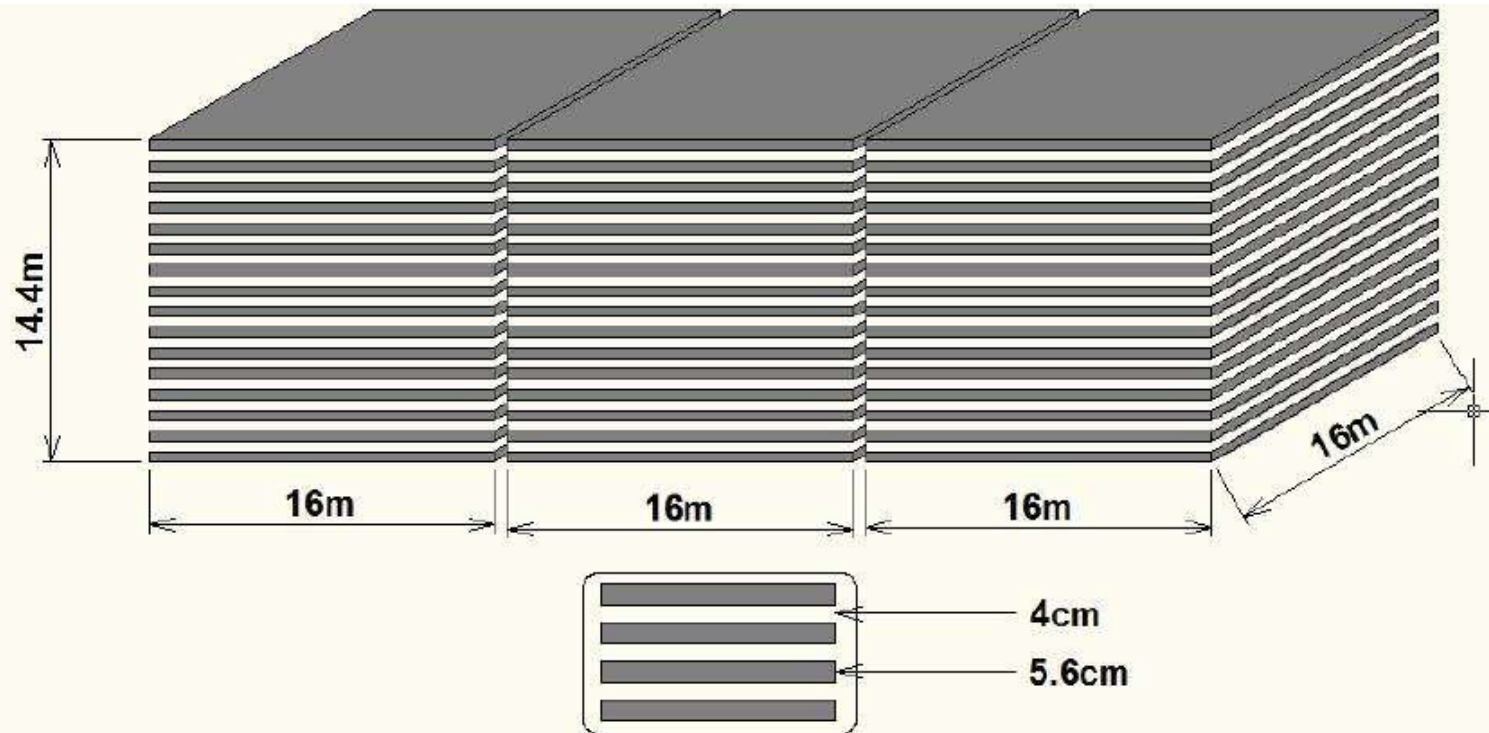


# INO site at BodiHills

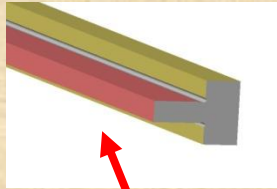




# ***INO-ICAL Detector***



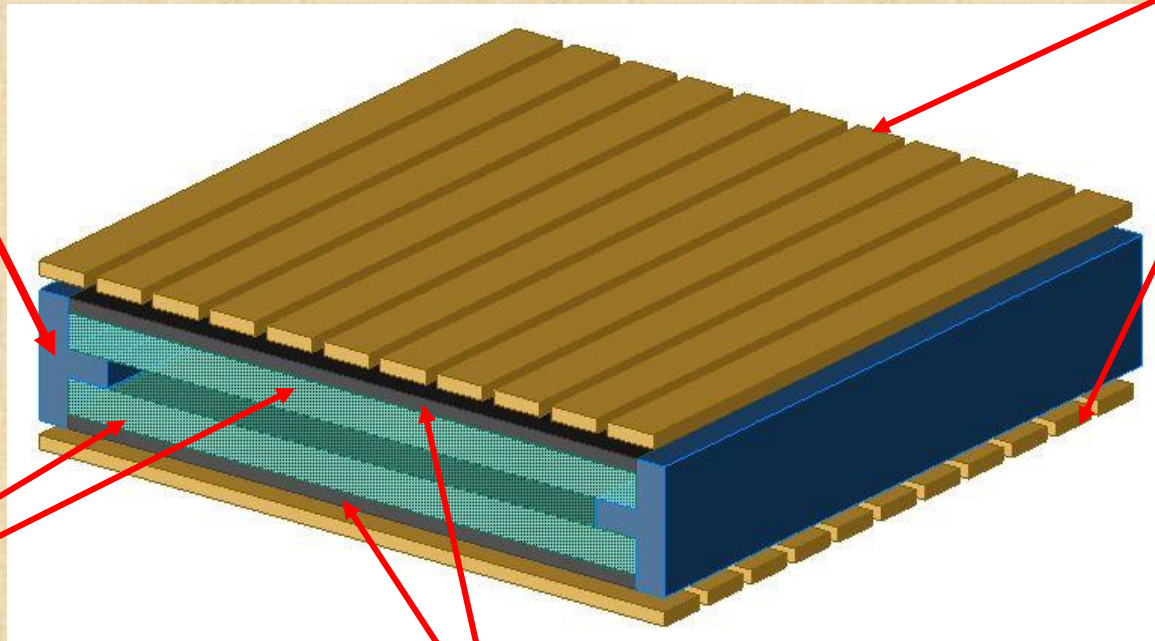
# Construction of RPC



*Two 2 mm thick float Glass  
Separated by 2 mm spacer*

*2 mm thick spacer*

*Pickup strips*

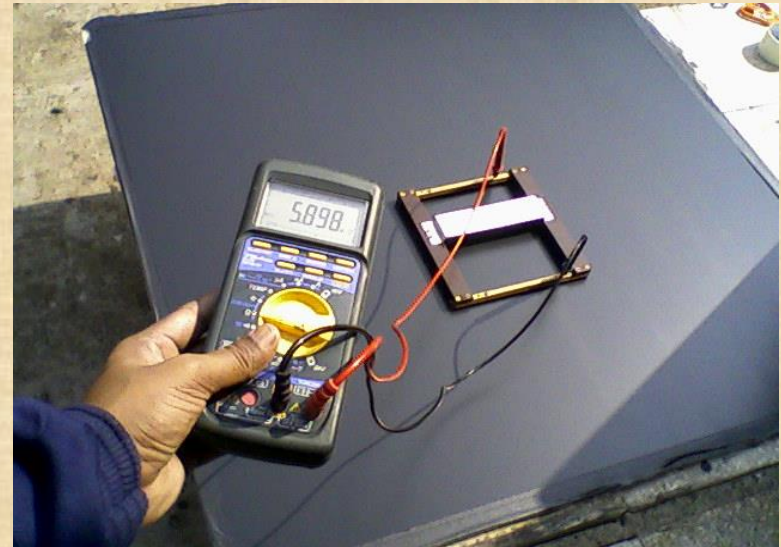
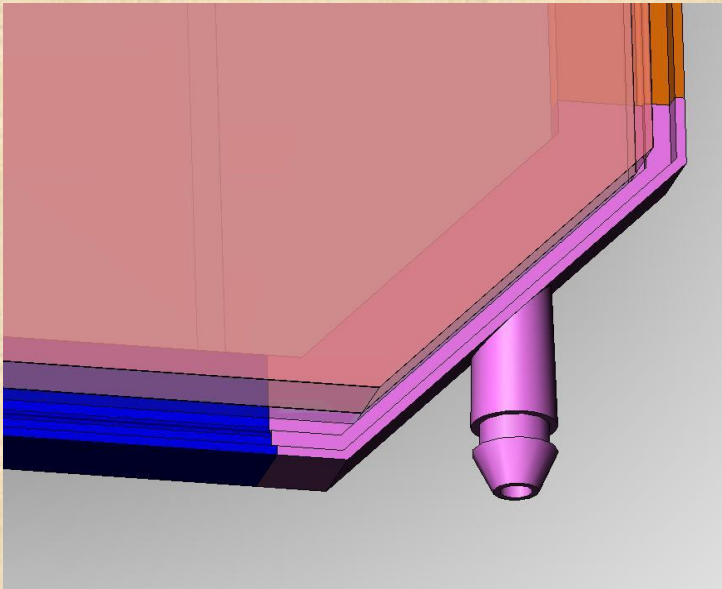
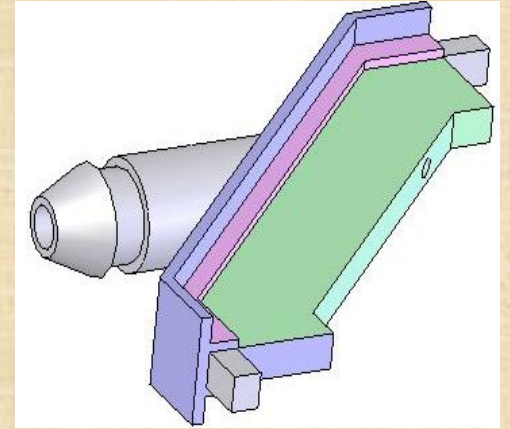
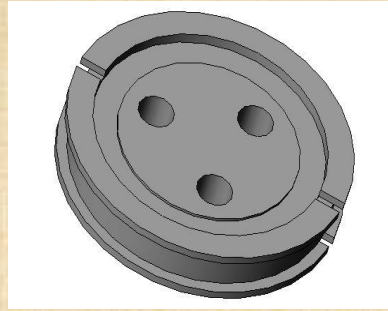
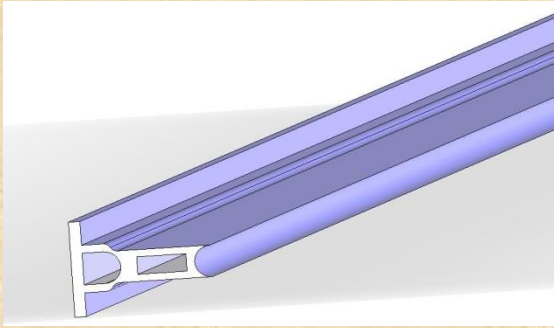


*Glass plates*

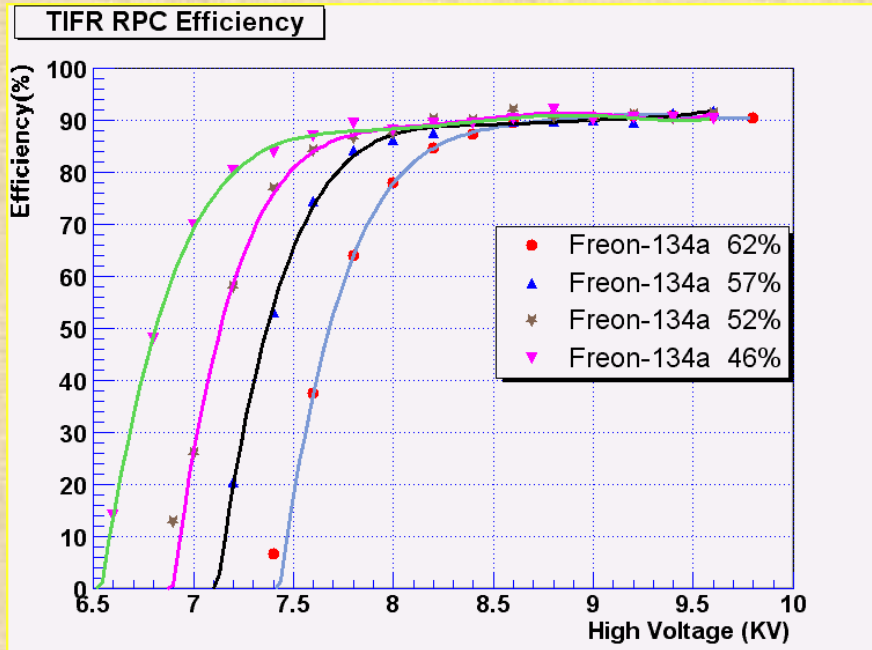
*Resistive coating on the outer surfaces of glass*



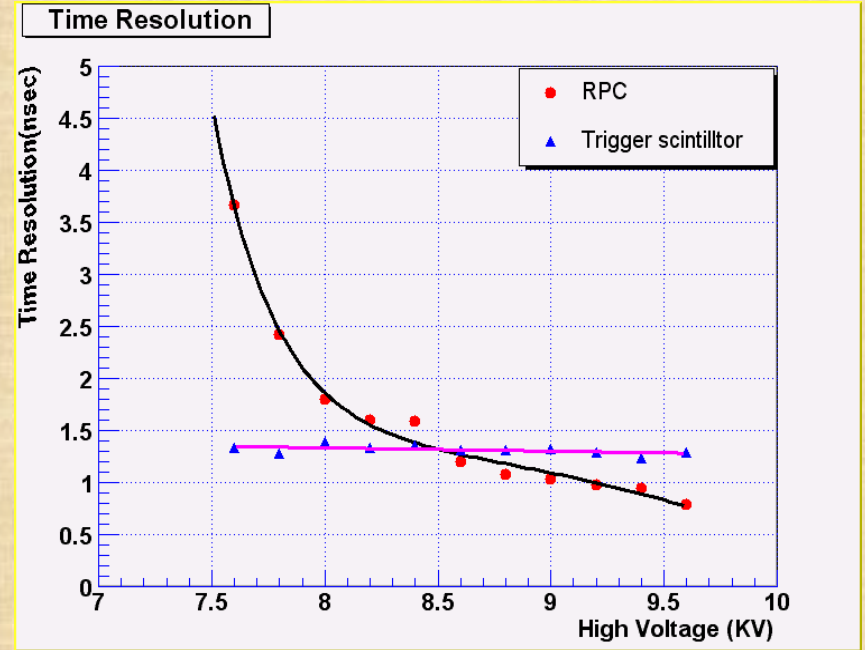
# RPC building blocks



## Early results on RPC efficiencies and time resolution



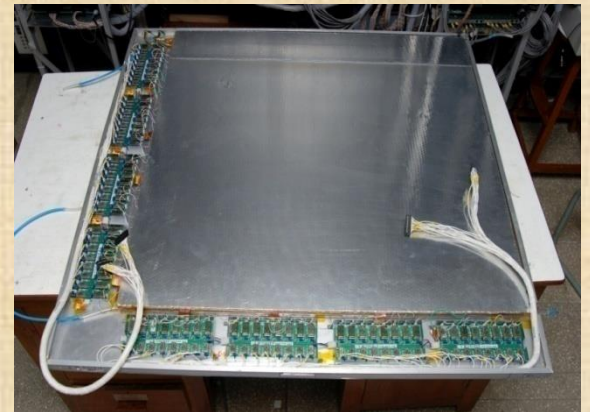
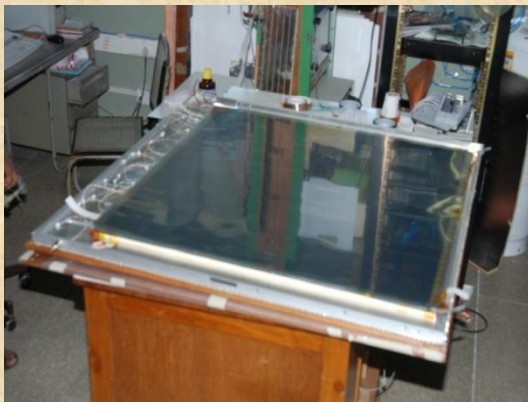
**Efficiency**



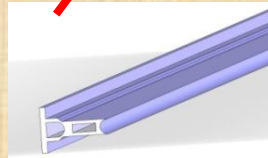
**Time resolution**



# ***Fabrication of 1m x 1m RPCs***



# Final RPC Frontier - Making of 2m x 2m RPCs





# ***RPC fabrication at Asahi Float Glass Co.***



## *A journey through RPC road*



**200 cm x 200 cm**

**30 cm x 30 cm**

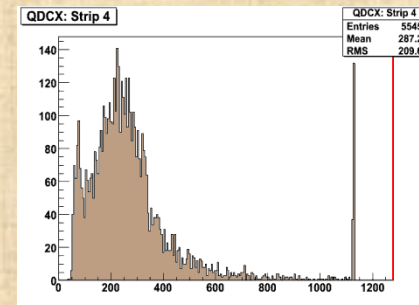
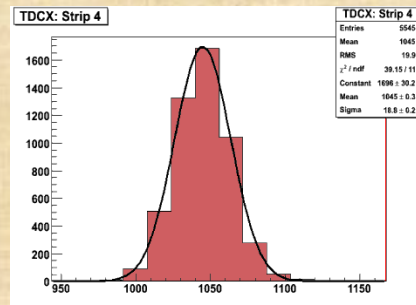
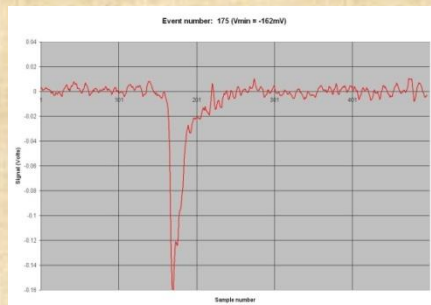
**10 cm x 30 cm**



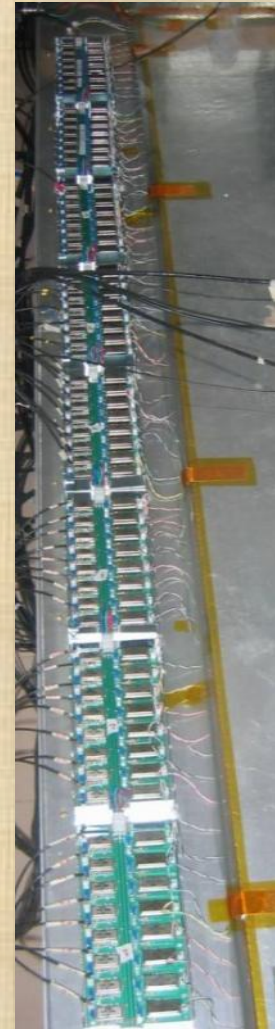
**100 cm x 100 cm**



# Prototype RPC Stack at TIFR tracking Muons

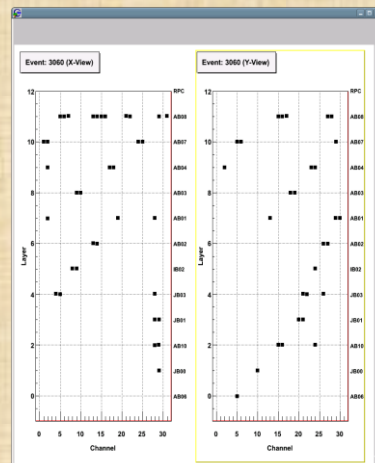
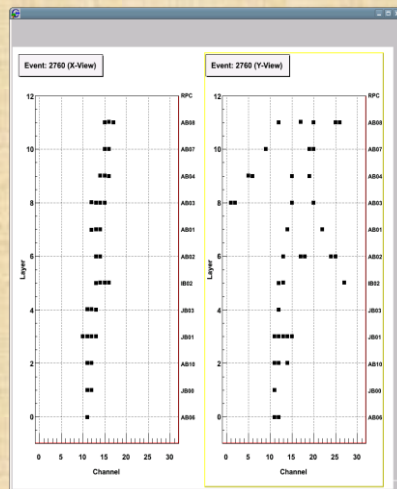
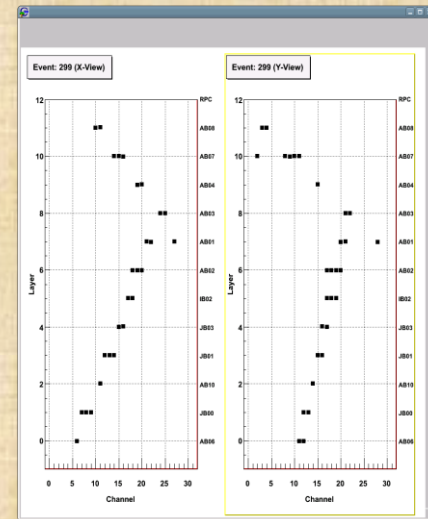
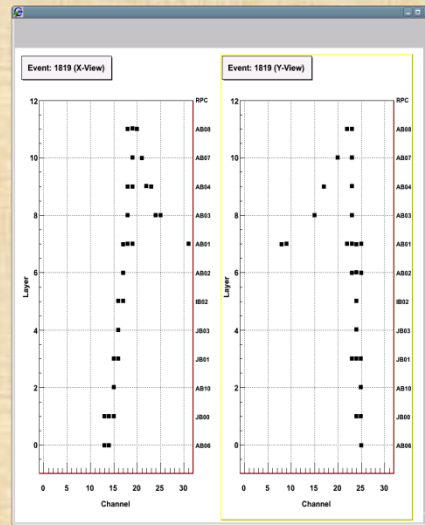
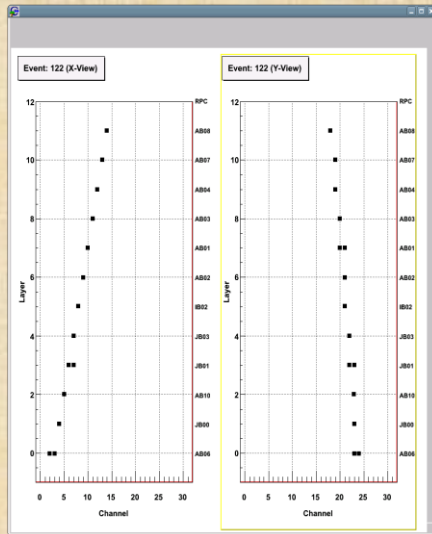


## ***2m x 2m glass RPC test stand***





# cosmic ray tracks in the RPC stand

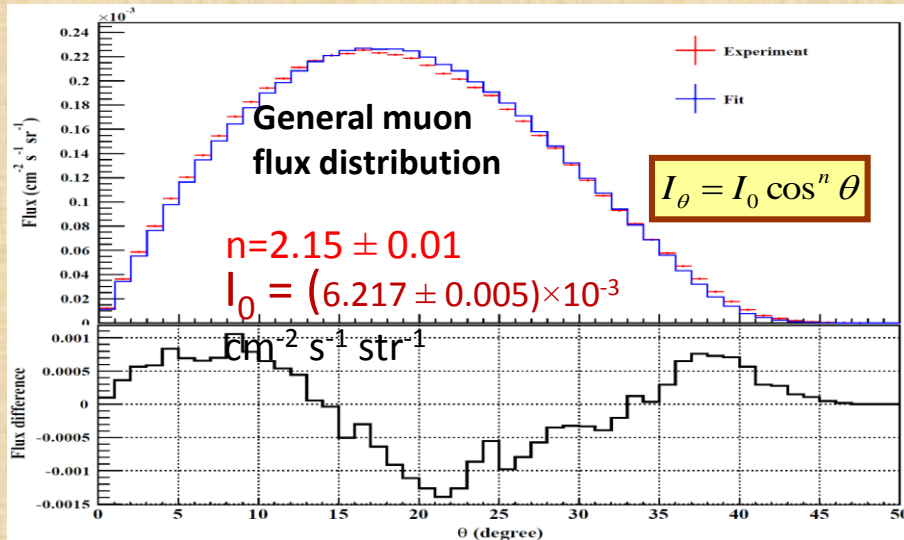
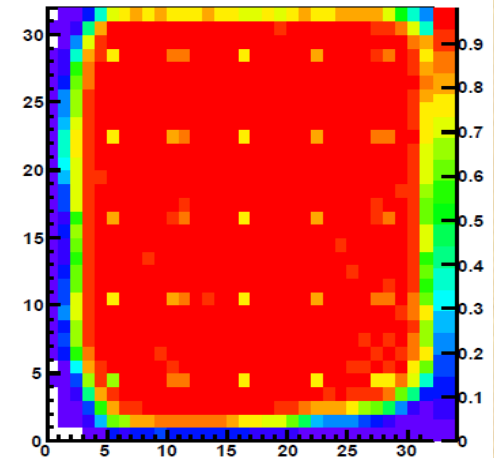
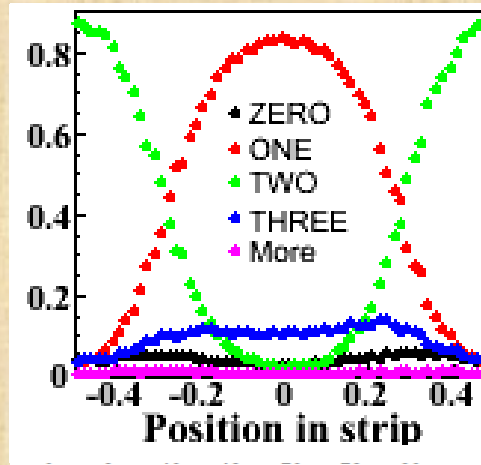


**Demonstrate the Tracking Capability of the RPC system**

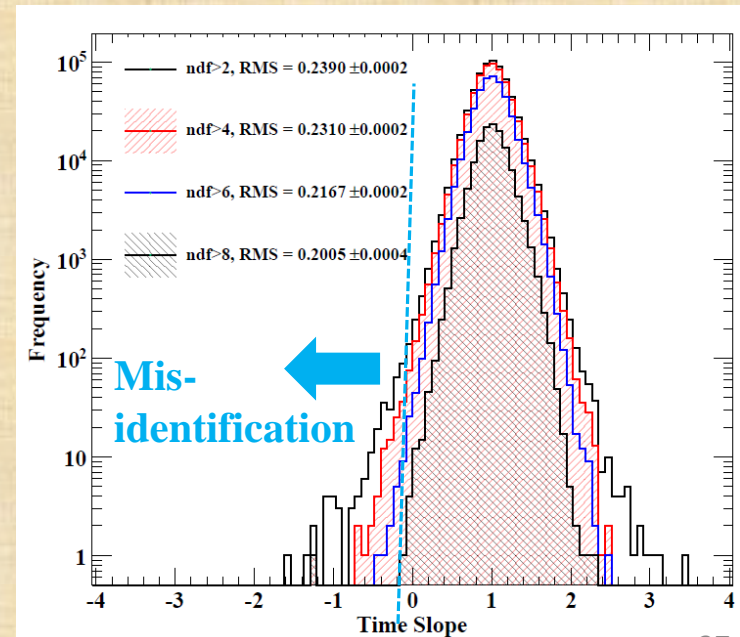
# Running Prototype RPC Stack at TIFR



Zenith angle of muon, measurement of cosmic muon flux as well as its angular dependency



Input to detector simulation and digitisation



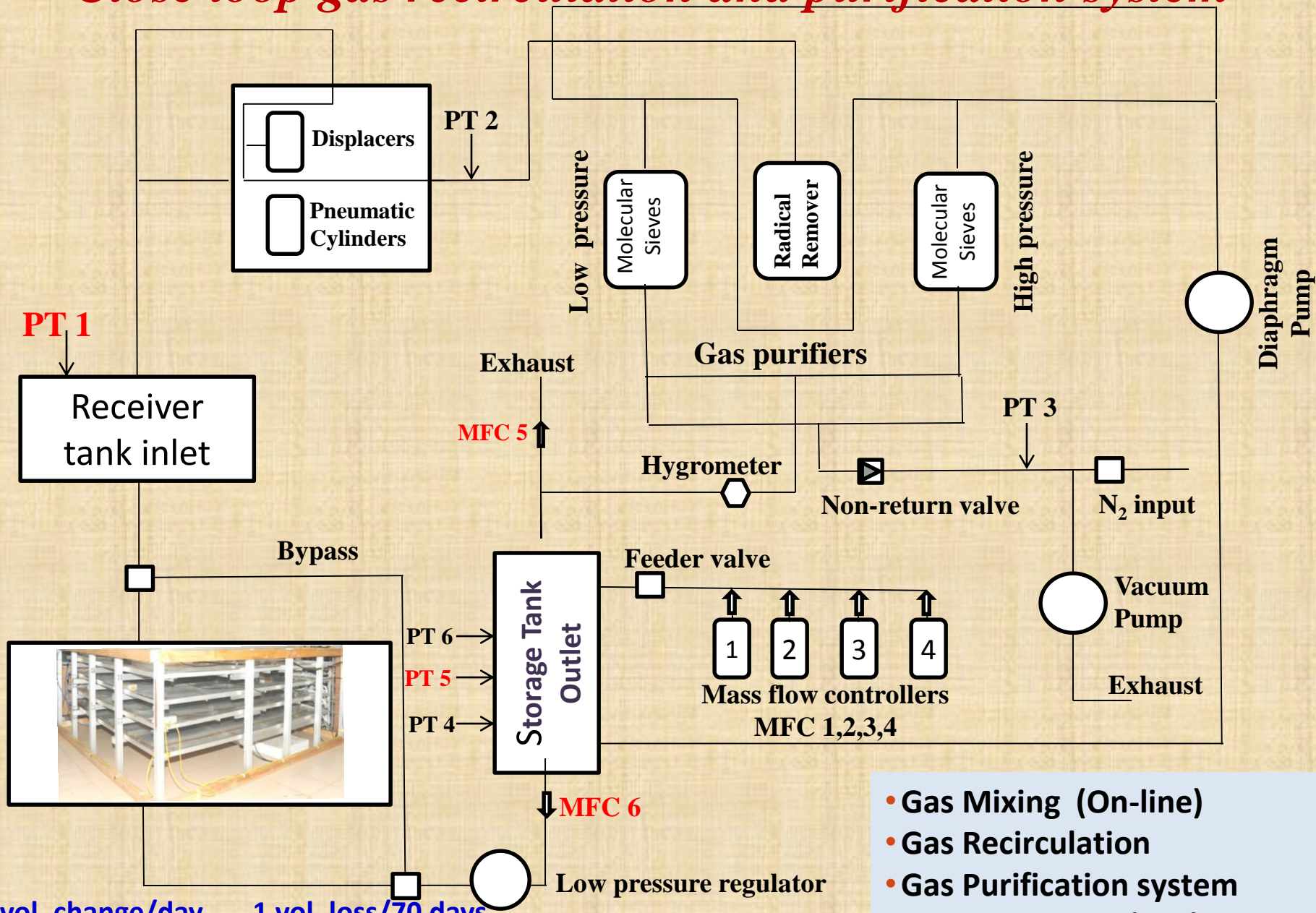
Distinction of up/down Muon



# *Newly developed gas recirculation system*



# Close loop gas recirculation and purification system

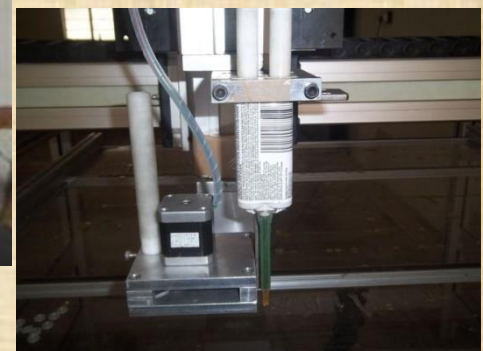


1 vol. change/day    1 vol. loss/70 days  
Automated pressure control, 1-3mbar above atm

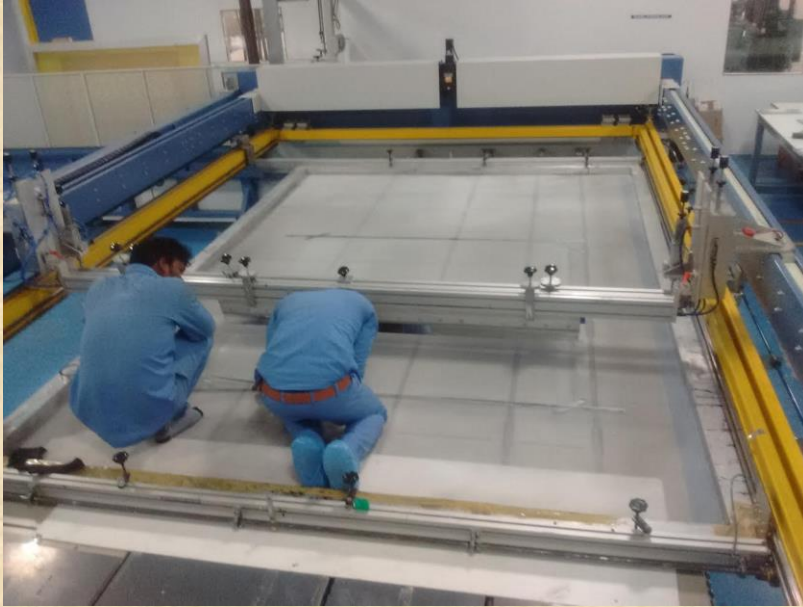
- Gas Mixing (On-line)
- Gas Recirculation
- Gas Purification system
- Control System (PLC)



# Automatic RPC gap making



# *Industrial production of RPC*





# *Running of Prototype RPC Stack at Madurai*

Operational since last one year



# ***Publications RPC R & D***

## ***List of publications:***

- ***JCAP, 07, 2012, 033***
- ***NIM A678, 105, 2012***
- ***NIM, A694, 126, 2012***
- ***NIM A661,64, 2012***
- ***NIM A661,68, 2012***
- ***NIM A661,73, 2012***
- ***NIM A661,77, 2012***
- ***NIM A661, 234, 2012***
- ***NIM A602, 784, 2009***
- ***NIM A602, 744, 2009***
- ***NIM A 602, 845, 2009***
- ***NIM A 602, 835, 2009***
- ***NIM A 701, 153, 2013***
- ***NIM A 736, 13, 2014***
- ***JINST 11, 2016***