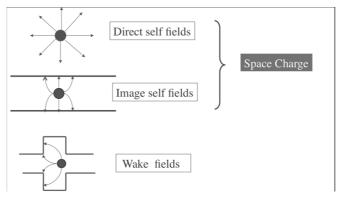
# Space Charge in Linear Machines Massimo.Ferrario@LNF.INFN.IT



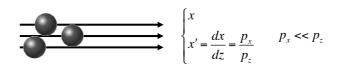


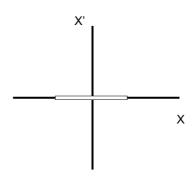
Egham – September 6<sup>th</sup> 2017



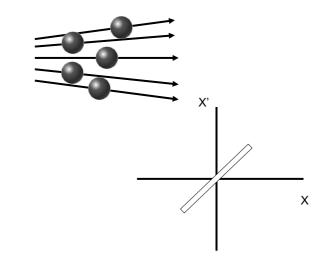
## **OUTLINE**

- The rms emittance concept
- rms envelope equation
- Space charge forces
- Space charge induced emittance oscillations
- Matching conditions and emittance compensation

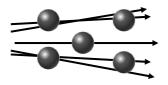


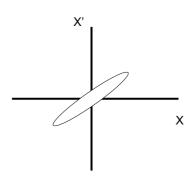


# Trace space of a laminar beam



Trace space of non laminar beam



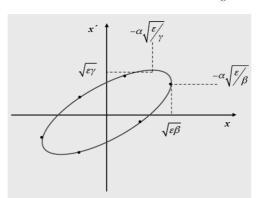


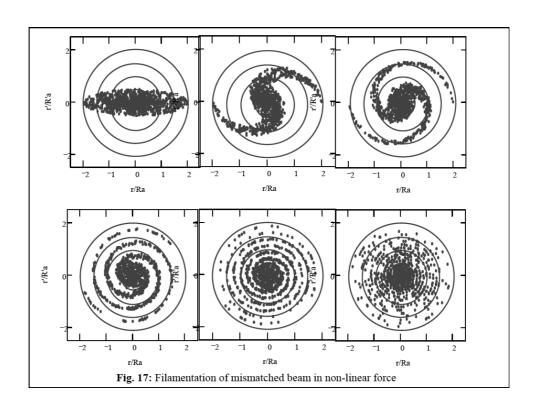
Geometric emittance:

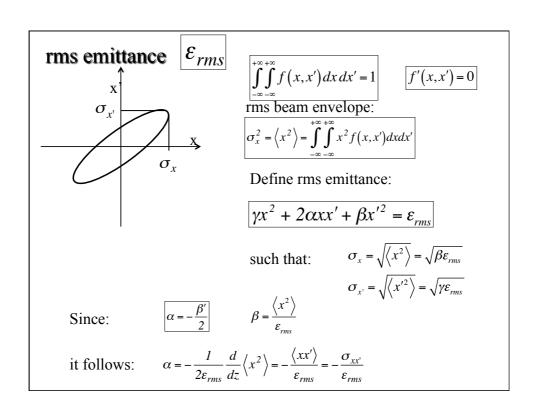
Ellipse equation:  $\gamma x^2 + 2\alpha x x' + \beta x'^2 = \varepsilon_g$ 

Twiss parameters:  $\beta \gamma - \alpha^2 = 1$   $\beta' = -2\alpha$ 

Ellipse area:  $A = \pi \varepsilon_g$ 







$$\sigma_{x} = \sqrt{\langle x^{2} \rangle} = \sqrt{\beta \varepsilon_{rms}}$$

$$\sigma_{x'} = \sqrt{\langle x^{'2} \rangle} = \sqrt{\gamma \varepsilon_{rms}}$$

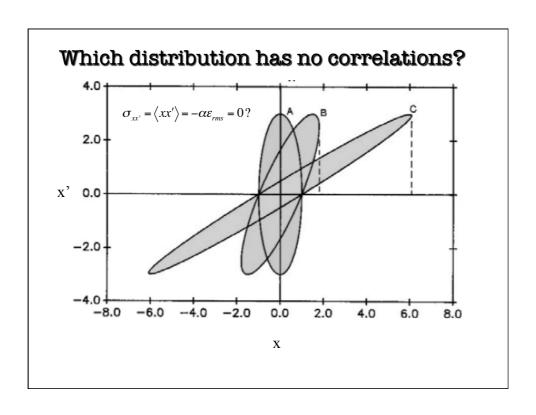
$$\sigma_{xx'} = \langle xx' \rangle = -\alpha \varepsilon_{rms}$$

It holds also the relation:  $\gamma \beta - \alpha^2 = I$ 

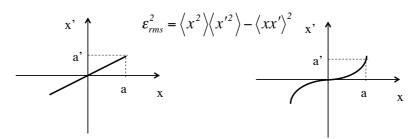
Substituting 
$$\alpha, \beta, \gamma$$
 we get  $\frac{\sigma_{x'}^2}{\varepsilon_{rms}} \frac{\sigma_x^2}{\varepsilon_{rms}} - \left(\frac{\sigma_{xx'}}{\varepsilon_{rms}}\right)^2 = 1$ 

We end up with the definition of rms emittance in terms of the second moments of the distribution:

$$\varepsilon_{rms} = \sqrt{\sigma_{x}^{2}\sigma_{x'}^{2} - \sigma_{xx'}^{2}} = \sqrt{\left(\left\langle x^{2}\right\rangle \left\langle x'^{2}\right\rangle - \left\langle xx'\right\rangle^{2}\right)}$$



What does rms emittance tell us about phase space distributions under linear or non-linear forces acting on the beam?



Assuming a generic x, x' correlation of the type:  $x' = Cx^n$ 

When 
$$n = 1 = \infty$$
  $\epsilon_{rms} = 0$ 

$$\epsilon_{rms}^2 = C^2 \left( \left\langle x^2 \right\rangle \left\langle x^{2n} \right\rangle - \left\langle x^{n+1} \right\rangle^2 \right)$$
When  $n \neq 1 = \infty$   $\epsilon_{rms} \neq 0$ 

## Constant under linear transformation only

$$\frac{\mathrm{d}}{\mathrm{d}z}\langle x^2\rangle\langle x'^2\rangle - \langle xx'\rangle^2 = 2\langle xx'\rangle\langle x'^2\rangle + 2\langle x^2\rangle\langle x'\rangle\langle x''\rangle - 2\langle xx''\rangle\langle xx'\rangle = 0$$

For linear transformations,  $x'' = -k_x^2 x$ , and the right-hand side of the equation is

$$2k_x^2\langle x^2\rangle\langle xx'\rangle - 2\langle x^2\rangle\langle xx'\rangle k_x^2 = 0,$$

SO

$$\frac{\mathrm{d}}{\mathrm{d}z}\langle x^2\rangle\langle x'^2\rangle - \langle xx'\rangle^2 = 0$$

And without acceleration:

$$x' = \frac{p_x}{p_z}$$

## Normalized rms emittance: $\varepsilon_{n,rms}$

Canonical transverse momentum:  $p_x = p_z x' = m_o c \beta \gamma x'$   $p_z \approx p$ 

$$\varepsilon_{n,rms} = \frac{1}{m_o c} \sqrt{\sigma_x^2 \sigma_{p_x}^2 - \sigma_{xp_x}^2} = \frac{1}{m_o c} \sqrt{\left(\left\langle x^2 \right\rangle \left\langle p_x^2 \right\rangle - \left\langle x p_x \right\rangle^2\right)} = \left(\left\langle \beta \gamma \right\rangle \varepsilon_{rms}\right)$$

Liouville theorem: the density of particles n, or the volume V occupied by a given number of particles in phase space  $(x,p_x,y,p_v,z,p_z)$  remains invariant under conservative forces.

$$\frac{dn}{dt} = 0$$

It hold also in the projected phase spaces  $(x,p_x),(y,p_y)(,z,p_z)$  provided that there are no couplings

## Limit of single particle emittance

Limits are set by Quantum Mechanics on the knowledge of the two conjugate variables  $(x,p_x)$ . According to Heisenberg:

$$\sigma_{x}\sigma_{p_{x}} \geq \frac{\hbar}{2}$$

This limitation can be expressed by saying that the state of a particle is not exactly represented by a point, but by a small uncertainty volume of the order of  $\hbar^3$  in the 6D phase space.

In particular for a single electron in 2D phase space it holds:

$$\varepsilon_{n,rms} = \frac{1}{m_o c} \sqrt{\sigma_x^2 \sigma_{p_x}^2 - \sigma_{xp_x}^2} \implies \begin{cases} = 0 & \text{classical limit} \\ \ge \frac{1}{2} \frac{\hbar}{m_o c} = \frac{\lambda_c}{2} = 1.9 \times 10^{-13} m & \text{quantum limit} \end{cases}$$

Where  $\lambda_c$  is the reduced Compton wavelength.

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## **Envelope Equation without Acceleration**

Now take the derivatives:

$$\frac{d\sigma_{x}}{dz} = \frac{d}{dz}\sqrt{\langle x^{2}\rangle} = \frac{1}{2\sigma_{x}}\frac{d}{dz}\langle x^{2}\rangle = \frac{1}{2\sigma_{x}}2\langle xx'\rangle = \frac{\sigma_{xx'}}{\sigma_{x}}$$

$$\frac{d^{2}\sigma_{x}}{dz^{2}} = \frac{d}{dz}\frac{\sigma_{xx'}}{\sigma_{x}} = \frac{1}{\sigma_{x}}\frac{d\sigma_{xx'}}{dz} - \frac{\sigma_{xx'}^{2}}{\sigma_{x}^{3}} = \frac{1}{\sigma_{x}}(\langle x'^{2}\rangle + \langle xx'\rangle) - \frac{\sigma_{xx'}^{2}}{\sigma_{x}^{3}} = \frac{\sigma_{x'}^{2} + \langle xx''\rangle}{\sigma_{x}} - \frac{\sigma_{xx'}^{2}}{\sigma_{x}^{3}}$$

And simplify: 
$$\sigma_x'' = \frac{\sigma_x^2 \sigma_{x'}^2 - \sigma_{xx'}^2}{\sigma_x^3} + \frac{\langle xx'' \rangle}{\sigma_x} = \frac{\varepsilon_{rms}^2}{\sigma_x^3} + \frac{\langle xx'' \rangle}{\sigma_x}$$

We obtain the rms envelope equation in which the rms emittance enters as defocusing pressure like term.

$$\sigma_x'' - \frac{\langle xx'' \rangle}{\sigma_x} = \frac{\varepsilon_{rms}^2}{\sigma_x^3}$$

$$\frac{\varepsilon_{rms}^2}{\sigma_x^3} \approx \frac{T}{V} \approx P$$

## Beam Thermodynamics

Kinetic theory of gases defines temperatures in each directions and global as:

$$k_B T_x = m \langle v_x^2 \rangle$$
  $T = \frac{1}{3} (T_x + T_y + T_z)$   $E_k = \frac{1}{2} m \langle v^2 \rangle = \frac{3}{2} k_B T$ 

Definition of beam temperature in analogy:

$$k_B T_{beam,x} = \gamma m_o \langle v_x^2 \rangle \qquad \qquad \langle v_x^2 \rangle = \beta^2 c^2 \langle x'^2 \rangle = \beta^2 c^2 \sigma_{x'}^2 = \beta^2 c^2 \frac{\varepsilon_{rms}^2}{\sigma_x^2} = \beta^2 c^2 \frac{\varepsilon_{rms}}{\beta_x}$$

$$\begin{aligned} k_B T_{beam,x} &= \gamma m_o \left\langle v_x^2 \right\rangle = \gamma m_o \beta^2 c^2 \frac{\varepsilon_{rms}^2}{\sigma_x^2} = \gamma m_o \beta^2 c^2 \frac{\varepsilon_{rms}}{\beta_x} \\ P_{beam,x} &= n k_B T_{beam,x} = n \gamma m_o \beta^2 c^2 \frac{\varepsilon_{rms}^2}{\sigma_x^2} = N_T \gamma m_o \beta^2 c^2 \frac{\varepsilon_{rms}^2}{\sigma_L \sigma_x^2} \end{aligned}$$

$$k_B T_{beam,x} = \gamma m_o \beta^2 c^2 \frac{\varepsilon_{rms}}{\beta_x}$$

Property	Hot beam	Cold beam
ion mass (m <sub>0</sub> )	heavy ion	light ion
ion energy (βγ)	high energy	low energy
beam emittance (ε)	large emittance	small emittance
lattice properties $(\gamma_{x,y} \approx 1/\beta_{x,y})$	strong focus (low β)	high β
phase space portrait	hot beamx	cold beam x,

Electron Cooling: Temperature relaxation by mixing a hot ion beam with co-moving cold (light) electron beam.

Particle Accelerators 1973, Vol. 5, pp. 61-65

#### EMITTANCE, ENTROPY AND INFORMATION

P. M. LAPOSTOLLE al d'Études des Télécommunications, Issy-les-Moulineaux, France

$$S = kN\log(\pi\varepsilon)$$

By means of the beam temperature concept one can also define the beam emittance at the source called thermal emittance. Assuming that electrons are in equilibrium with the cathode temperature  $T_c=T_{beam}$  and  $\gamma=1$  the thermal emittance is given by:  $\varepsilon_{ih,rmi}^{cot}=\sigma_x\sqrt{\frac{k_BT_c}{m_ec^2}}$  which, per unit rms spot size at the cathode, is  $\varepsilon_{ih,rmi}=0.3$  µm/mm at  $T_c=2500$  K. For comparison in a photocatode illuminated by a laser pulse with photon energy  $\hbar\omega$  the expression for the variance of the transverse momentum of the emitted electrons is given by  $\sigma_{p_c}=\sqrt{\frac{m_c}{3}(\hbar\omega-\phi_{eff})}$  where  $\phi_{eff}=\phi_w-\phi_{Schootly}$ ,  $\phi_w$  being the material work function and  $\phi_{Schootly}$  the Schottky work function [19]. The corresponding thermal emittance is  $\varepsilon_{ih,rms}^{ab}=\sigma_x\sqrt{\frac{\hbar\omega-\phi_{eff}}{3m_cc^2}}$  that, with the typical parameters of a Copper photocathode illuminated by a UV laser, gives a thermal emittance per unit spot size of about 0.5 µm/mm.

# Envelope Equation with Linear Focusing

$$\sigma_{x}'' - \frac{\left\langle xx''\right\rangle}{\sigma_{x}} = \frac{\varepsilon_{rms}^{2}}{\sigma_{x}^{3}}$$

Assuming that each particle is subject only to a linear focusing force, without acceleration:  $x'' + k_x^2 x = 0$ 

take the average over the entire particle ensemble  $\langle xx'' \rangle = -k_x^2 \langle x^2 \rangle$ 

$$\sigma_x'' + k_x^2 \sigma_x = \frac{\varepsilon_{rms}^2}{\sigma_x^3}$$

We obtain the rms envelope equation with a linear focusing force in which, unlike in the single particle equation of motion, the rms emittance enters as defocusing pressure like term.

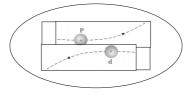
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# Space Charge: what does it mean?

The net effect of the **Coulomb** interactions in a multi-particle system can be classified into two regimes:

1) Collisional Regime ==> dominated by binary collisions caused by close particle encounters ==> Single Particle Effects



2) Space Charge Regime ==> dominated by the self field produced by the particle distribution, which varies appreciably only over large distances compare to the average separation of the particles ==> Collective Effects

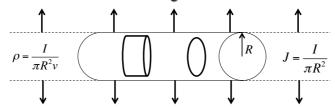






31/08/17

# Continuous Uniform Cylindrical Beam Model



$$\int \varepsilon_o E \cdot dS = \int \rho dV$$

Gauss's law 
$$\int \varepsilon_o E \cdot dS = \int \rho dV$$

$$E_r = \frac{I}{2\pi\varepsilon_o R^2 v} r \quad \text{for } r \leq R$$

$$E_r = \frac{I}{2\pi\varepsilon_o v} \frac{1}{r} \quad \text{for } r > R$$

$$B_\theta = \frac{\beta}{c} E_r$$
Ampere's law 
$$\int B \cdot dl = \mu_o \int J \cdot dS$$

$$B_\theta = \mu_o \frac{Ir}{2\pi R^2} \quad \text{for } r \leq R$$

$$B_\theta = \mu_o \frac{I}{2\pi R^2} \quad \text{for } r > R$$

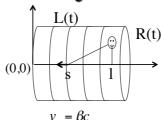
$$B_{\vartheta} = \frac{\beta}{c} E_{\vartheta}$$

$$\int B \cdot dl = \mu_o \int J \cdot dS$$

$$B_{\theta} = \mu_o \frac{Ir}{2\pi R^2} \quad \text{for} \quad r \le R$$

$$B_{\theta} = \mu_o \frac{I}{2\pi r} \quad \text{for} \quad r > R$$

# Bunched\_Uniform Cylindrical Beam Model



Longitudinal Space Charge field in the bunch moving frame:

$$\tilde{\rho} = \frac{Q}{\pi R^2 \tilde{L}} \qquad \qquad \tilde{E}_z(\tilde{s}, r = 0) = \frac{\tilde{\rho}}{4\pi\varepsilon_o} \int\limits_0^R \int\limits_0^{2\pi} \int\limits_0^{\tilde{L}} \frac{\left(\tilde{l} - \tilde{s}\right)}{\left[\left(\tilde{l} - \tilde{s}\right)^2 + r^2\right]^{3/2}} \ r dr d\varphi d\tilde{l}$$

$$\tilde{E}_z(\tilde{s},r=0) = \frac{\tilde{\rho}}{2\varepsilon_0} \left[ \sqrt{R^2 + (\tilde{L} - \tilde{s})^2} - \sqrt{R^2 + \tilde{s}^2} + \left(2\tilde{s} - \tilde{L}\right) \right]$$

Radial Space Charge field in the bunch moving frame by series representation of axisymmetric field:

$$\tilde{E}_r(r,\tilde{s}) \cong \left[\frac{\tilde{\rho}}{\varepsilon_0} - \frac{\partial}{\partial \tilde{s}}\tilde{E}_z(0,\tilde{s})\right] \frac{r}{2} + \left[\cdots\right] \frac{r^3}{16} +$$

$$\tilde{E}_r(r,\tilde{s}) = \frac{\tilde{\rho}}{2\varepsilon_0} \left[ \frac{(\tilde{L} - \tilde{s})}{\sqrt{R^2 + (\tilde{L} - \tilde{s})^2}} + \frac{\tilde{s}}{\sqrt{R^2 + \tilde{s}^2}} \right] \frac{r}{2}$$

#### **Lorentz Transformation to the Lab frame**

$$\begin{split} E_z &= \tilde{E}_z \\ E_r &= \gamma \tilde{E}_r \end{split} \qquad \qquad \tilde{L} = \gamma L \implies \tilde{\rho} = \frac{\rho}{\gamma} \end{split}$$

$$E_z(\theta,s) = \frac{\rho}{\gamma 2\varepsilon_\theta} \left[ \sqrt{R^2 + \gamma^2 (L-s)^2} - \sqrt{R^2 + \gamma^2 s^2} + \gamma \left(2s - L\right) \right]$$

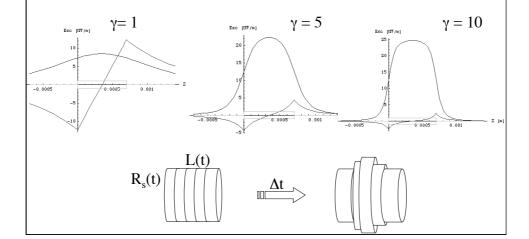
$$E_r(r,s) = \frac{\gamma \rho}{2\varepsilon_0} \left[ \frac{(L-s)}{\sqrt{R^2 + \gamma^2 (L-s)^2}} + \frac{s}{\sqrt{R^2 + \gamma^2 s^2}} \right] \frac{r}{2}$$

It is still a linear field with r but with a longitudinal correlation s

# Bunched Uniform Cylindrical Beam Model

$$E_z(0, s, \gamma) = \frac{I}{2\pi\gamma\varepsilon_0 R^2 \beta c} h(s, \gamma)$$

$$E_r(r, s, \gamma) = \frac{Ir}{2\pi\varepsilon_0 R^2 \beta c} g(s, \gamma)$$



## Lorentz Force

$$F_r = e(E_r - \beta cB_{\vartheta}) = e(1 - \beta^2)E_r = \frac{eE_r}{\gamma^2}$$

is a linear function of the transverse coordinate

$$\frac{dp_r}{dt} = F_r = \frac{eE_r}{\gamma^2} = \frac{eIr}{2\pi\gamma^2 \varepsilon_0 R^2 \beta c} g(s, \gamma)$$

The attractive magnetic force , which becomes significant at high velocities, tends to compensate for the repulsive electric force. Therefore space charge defocusing is primarily a non-relativistic effect. Using  $R=2\sigma_x$  for a uniform distribution:

$$F_{x} = \frac{eIx}{8\pi\gamma^{2}\varepsilon_{0}\sigma_{x}^{2}\beta c}g\left(s,\gamma\right)$$

## Envelope Equation with Space Charge

Single particle transverse motion:

$$\frac{dp_{x}}{dt} = F_{x} \qquad p_{x} = p \ x' = \beta \gamma m_{o} c x'$$

$$\frac{d}{dt}(px') = \beta c \frac{d}{dz}(p \ x') = F_{x}$$

$$x'' = \frac{F_{x}}{\beta c p}$$

$$F_{x} = \frac{eIx}{2\pi \gamma^{2} \varepsilon_{0} \sigma_{x}^{2} \beta c} g(s, \gamma)$$

$$k_{sc} = \frac{2I}{I_{A}} g(s, \gamma)$$

$$I_{A} = \frac{4\pi \varepsilon_{o} m_{o} c^{3}}{e}$$

Now we can calculate the term  $\langle xx'' \rangle$  that enters in the envelope equation

$$\sigma_{x}'' = \frac{\varepsilon_{rms}^{2}}{\sigma_{x}^{3}} + \frac{\langle xx'' \rangle}{\sigma_{x}}$$

$$x'' = \frac{k_{sc}}{\sigma_{x}^{2}} x$$

$$\langle xx'' \rangle = \frac{k_{sc}}{\sigma_{x}^{2}} \langle x^{2} \rangle = k_{sc}$$

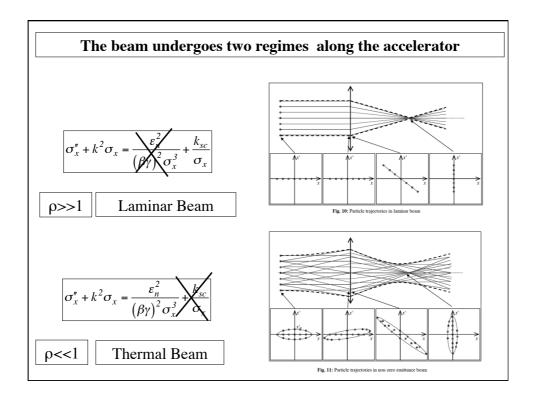
Including all the other terms the envelope equation reads:

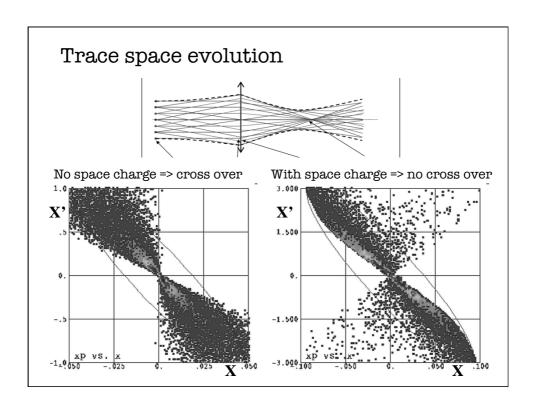
Space Charge De-focusing Force

$$\sigma_x'' + k^2 \sigma_x = \frac{\varepsilon_n^2}{(\beta \gamma)^2 \sigma_x^3} + \frac{k_{sc}}{\sigma_x}$$
Emittance Pressure

**External Focusing Forces** 

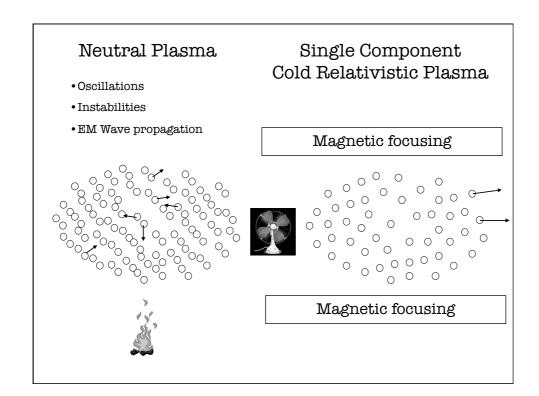
Laminarity Parameter: 
$$\rho = \frac{(\beta \gamma)^2 k_{sc} \sigma_x^2}{\varepsilon_n^2}$$





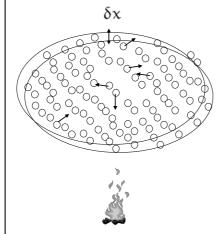
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Surface charge density

$$\sigma = e n \delta x$$



Surface electric field

$$E_x = -\sigma/\epsilon_0 = -e \, n \, \delta x/\epsilon_0$$

Restoring force

$$m\frac{d^2\delta x}{dt^2} = e E_x = -m \omega_p^2 \delta x$$

Plasma frequency

$$\omega_p^{\ 2} = \frac{n \ e^2}{\varepsilon_0 \ m}$$

Plasma oscillations

$$\delta x = (\delta x)_0 \, \cos{(\omega_p \, t)}$$

$$\sigma'' + k_s^2 \sigma = \frac{k_{sc}(s, \gamma)}{\sigma}$$

Equilibrium solution:

$$\sigma_{eq}(s,\gamma) = \frac{\sqrt{k_{sc}(s,\gamma)}}{k_{s}}$$

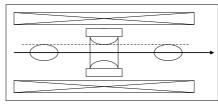
Small perturbation:

$$\sigma(\zeta) = \sigma_{eq}(s) + \delta\sigma(s)$$

$$\delta\sigma''(s) + 2k_s^2\delta\sigma(s) = 0$$

Single Component Relativistic Plasma

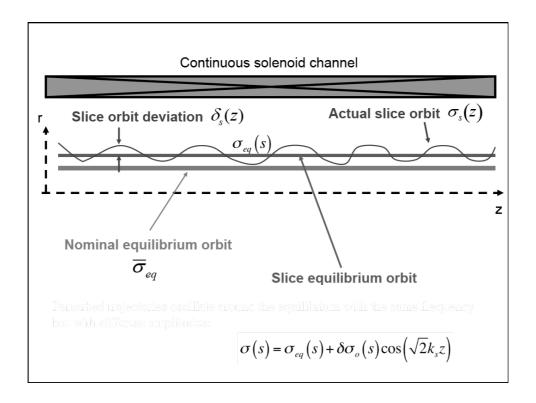
$$k_s = \frac{qB}{2mc\beta\gamma}$$

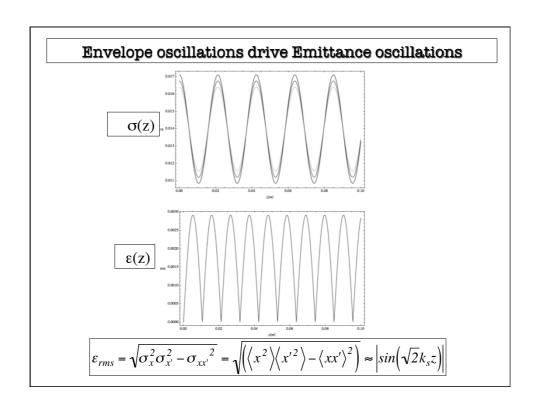


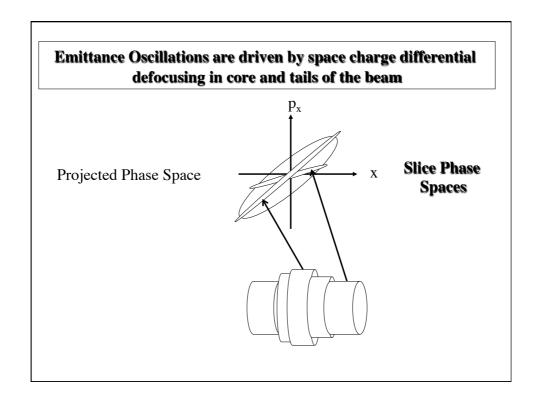
$$\delta\sigma(s) = \delta\sigma_o(s)\cos(\sqrt{2}k_s z)$$

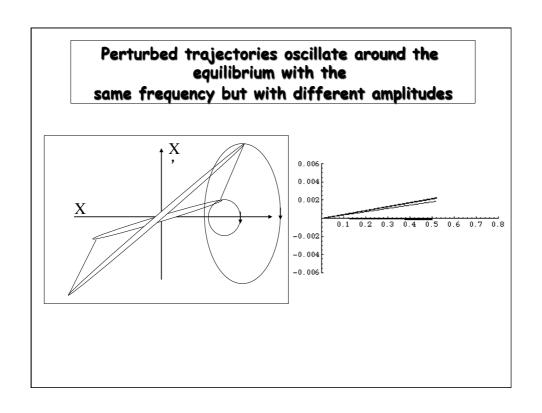
Perturbed trajectories oscillate around the equilibrium with the same frequency but with different amplitudes:

$$\sigma(s) = \sigma_{eq}(s) + \delta\sigma_{o}(s)\cos(\sqrt{2}k_{s}z)$$





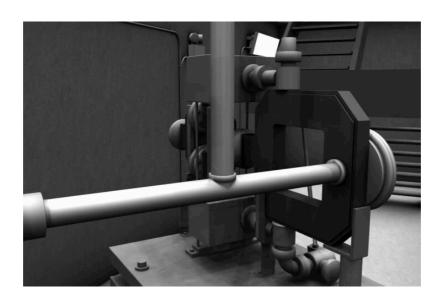


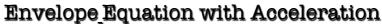


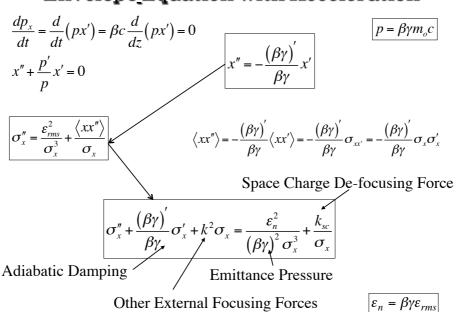
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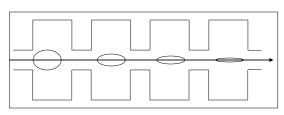
# High Brightness Photo-Injector







## Beam subject to strong acceleration



$$\sigma_x'' + \frac{\gamma'}{\gamma}\sigma_x' + \frac{k_{RF}^2}{\gamma^2}\sigma_x = \frac{\varepsilon_n^2}{\gamma^2\sigma_x^3} + \frac{k_{sc}^o}{\gamma^3\sigma_x}$$

We must include also the RF focusing force:

$$k_{RF}^2 = \frac{{\gamma'}^2}{2}$$

$$k_{sc}^{o} = \frac{2I}{I_A} g(s, \gamma)$$

$$\sigma_x'' + \frac{\gamma'}{\gamma}\sigma_x' + \frac{k_{RF}^2}{\gamma^2}\sigma_x = \frac{\varepsilon_n^2}{\gamma^2\sigma_x^3} + \frac{k_{sc}^o}{\gamma^3\sigma_x}$$

$$\gamma = 1 + \alpha z$$
 ==>  $\gamma'' = 0$ 

Looking for an "equilibrium" solution  $\sigma_{inv} = \sigma_o \gamma^n$  ==> all terms must have the same dependence on  $\gamma$ 

$$\sigma'_{inv} = n\sigma_o \gamma^{n-1} \gamma'$$

$$\sigma_{inv}'' = n(n-1)\sigma_o \gamma^{n-2} \gamma'^2$$

$$\sigma'_{inv} = n\sigma_o \gamma^{n-1} \gamma'$$

$$\sigma''_{inv} = n(n-1)\sigma_o \gamma^{n-2} \gamma'^2$$

$$n(n-1)\sigma_o \gamma^{n-2} \gamma'^2 + n\sigma_o \gamma^{n-2} \gamma'^2 + k_{RF}^2 \sigma_o \gamma^{n-2} = \frac{k_{sc}^o}{\sigma_x} \gamma^{-3-n}$$

$$n - 2 = -3 - n \Rightarrow n = -\frac{1}{2}$$

$$\sigma_x'' + \frac{\gamma'}{\gamma}\sigma_x' + \frac{k_{RF}^2}{\gamma^2}\sigma_x = \frac{\varepsilon_n^2}{\gamma^2\sigma_x^3} + \frac{k_{sc}^o}{\gamma^3\sigma_x}$$

$$\gamma = 1 + \alpha z$$
 ==>  $\gamma'' = 0$ 

Looking for an "equilibrium" solution  $\sigma_{inv} = \sigma_o \gamma^n$  ==> all terms must have the same dependence on  $\gamma$ 

$$\rho >> 1 \Rightarrow n = -\frac{1}{2}$$

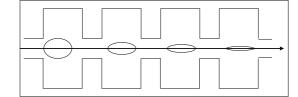
$$\sigma_q = \frac{\sigma_o}{\sqrt{\gamma}}$$

$$\rho << 1 \Rightarrow n = 0$$

$$\sigma_{\varepsilon} = \sigma_{\sigma}$$

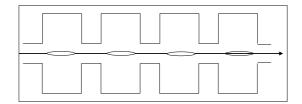
## Space charge dominated beam (Laminar)



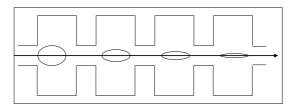


## Emittance dominated beam (Thermal)

$$\sigma_{\varepsilon} = \sqrt{\frac{2\varepsilon_n}{\gamma'}}$$



$$\sigma_q = \frac{1}{\gamma'} \sqrt{\frac{2I}{I_A \gamma}}$$



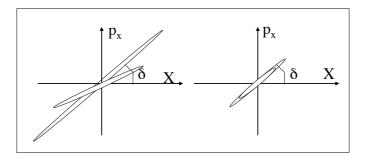
This solution represents a beam equilibrium mode that turns out to be the transport mode for achieving minimum emittance at the end of the emittance correction process

# An important property of the laminar beam

$$\sigma_q = \frac{1}{\gamma'} \sqrt{\frac{2I}{I_A \gamma}}$$

$$\sigma_q' = -\sqrt{\frac{2I}{I_A \gamma^3}}$$

Constant phase space angle:  $\delta = \frac{\gamma \sigma_q}{\sigma_q} = -\frac{\gamma}{2}$ 

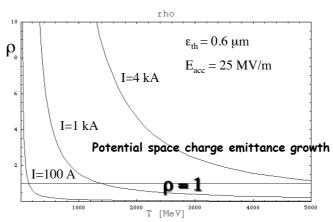


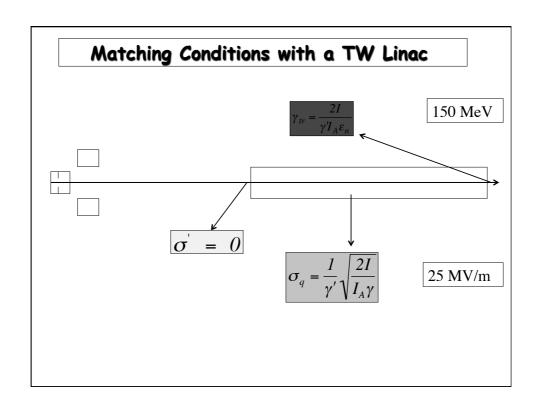


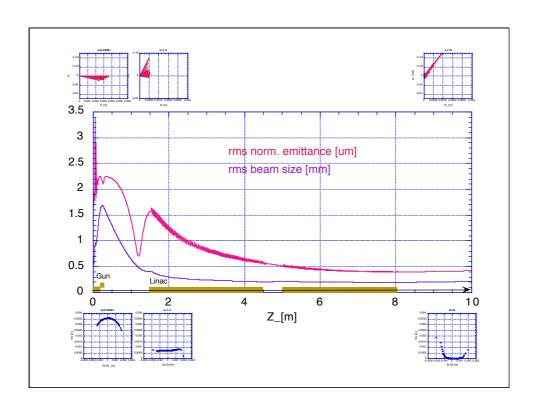
$$\rho = \frac{2I\sigma^2}{\gamma I_A \varepsilon_n^2} = \frac{2I\sigma_q^2}{\gamma I_A \varepsilon_n^2} = \frac{4I^2}{\gamma'^2 I_A^2 \varepsilon_n^2 \gamma^2}$$

# Transition Energy ( $\rho$ =1)









### Emittance Compensation for a SC dominated beam: Controlled Damping of Plasma Oscillations

- $\bullet$   $\epsilon_n$  oscillations are driven by Space Charge
- \* propagation close to the laminar solution allows control of  $\boldsymbol{\epsilon}_n$  oscillation "phase"
- $\epsilon_n$  sensitive to SC up to the transition energy

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