

Hopes and dreams for $B \rightarrow D^{(*)} \tau \nu$ angular analysis

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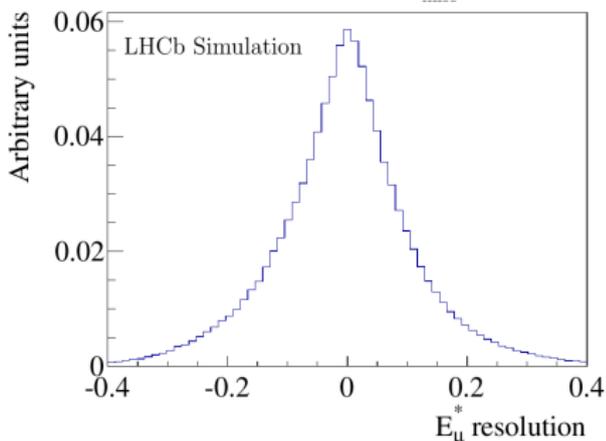
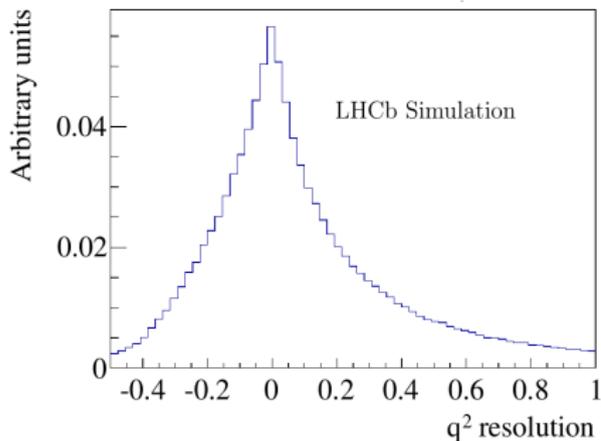
Nikhef

April 28, 2016

Introduction

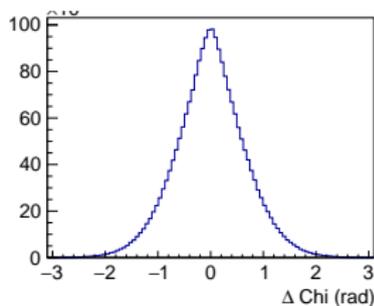
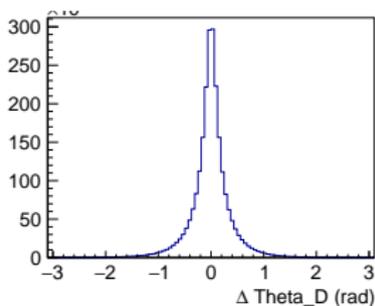
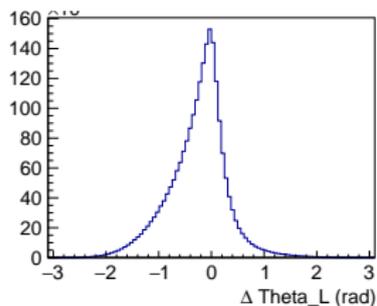
- $B \rightarrow D^{(*)} \ell \nu$ matrix element fully described by 2 (4) kinematic variables
- Most common choice: $q^2 + \text{angles}$
- For muonic $\mathcal{R}(D^*)$ analysis, we fit q^2 , muon energy, missing mass squared
 - These partially describe matrix element \rightarrow we already have some information
- At present we assume SM kinematic distributions for $\mathcal{R}(D^*)$
- What additional information should we try to fit?
- What physics should we try to measure?
- Disclaimer: talk is almost entirely opinions, hopes and speculation

Reconstruction ($\tau \rightarrow \mu\nu\nu$)



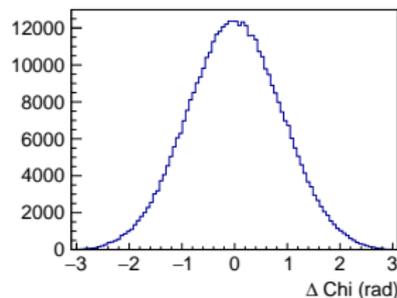
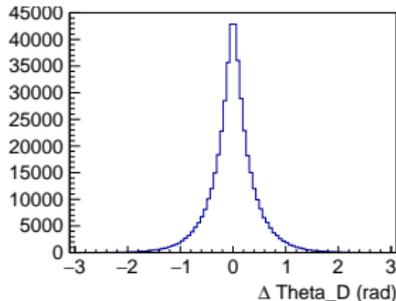
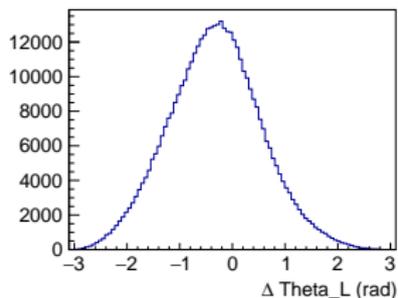
- Take $(\gamma\beta_z)_B = (\gamma\beta_z)_{D^*\mu}$
- Have approximation for rest frame with $\sim 15 - 20\%$ precision
- Can use this to calculate angles

Angular resolutions for $B \rightarrow D^* \mu \nu$



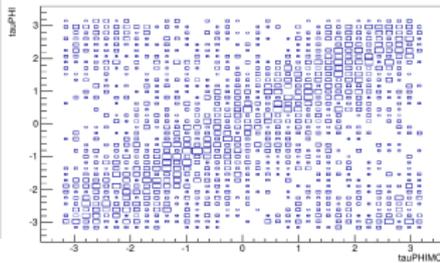
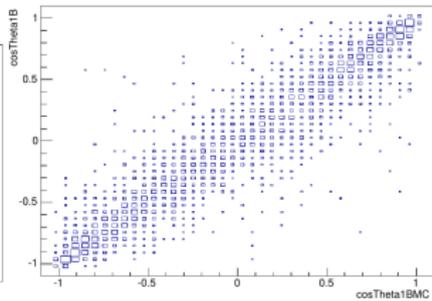
- Before taus, first look at angular resolution for $B \rightarrow D^* \mu \nu$
- Pretty wide, but have something to work with

Angular resolutions for $B \rightarrow D^* \tau \nu$ ($\tau \rightarrow \mu \nu \nu$)



- Angular resolution for $B \rightarrow D^* \tau \nu$
- Tau decay results in loss of information
 - Theta L gets very messy
 - Theta D still “good”
 - Chi worsens

Angular resolutions for $B \rightarrow D^* \tau \nu$ ($\tau \rightarrow \pi \pi \pi \nu$)



- Situation similar for $\tau \rightarrow \pi \pi \pi \nu$ mode
- Different reconstruction method:
 - Can reconstruct kinematics up to quadratic ambiguities using B and τ mass constraints + both vertex positions
 - Average over ambiguities
- Less information lost in tau decay, so Theta L a bit better

What can we do?

- Unfolding this seems a nightmare (as does background subtraction) → we are unlikely to publish corrected q^2 / angular distributions for signal
- But we can fit the data
 - Templates we fit already include effects of resolution, acceptance, tau decay, ...
 - We choose what physics model goes into templates, and can vary this in the fit
 - Fit parameters defined in terms of underlying physics
 - We already do this for systematic uncertainties

What to measure

- First need to see if the excess holds up!
- Afterwards, in order of ambition:
 - Does measured value change allowing NP operators?
 - Can enhancement be accommodated by theory uncertainty?
 - Is enhancement pure (pseudo)scalar?
 - If not, is it pure vector/axial/tensor/...?
 - Or a combination of operators?
 - Can we fit the full matrix element?

Can enhancement be accommodated by theory uncertainty?

- SM expectation for $\mathcal{R}(D^{(*)})$ relies on calculation for scalar form factor
 - Lepton mass suppressed \rightarrow can't measure in $B \rightarrow D^{(*)} \mu \nu$
- In flavour community, calculation seems to be trusted relative to current excess
 - But does the wider community believe it if we start making a serious new physics claim?

Scalar form factor

- Trying to measure scalar form factor doesn't seem so implausible
 - If no new (pseudo)scalar physics, and form factor agrees with prediction
→ model independent SM exclusion
 - Uncertainty from QED corrections?
- Testing SM only hypothesis → constrain other form factors from $B \rightarrow D^{(*)} \mu \nu$
- CLN vs general z expansion / BGL?

More complex fits

- We don't really know how far we can push this
- Where we actually have statistical power, how many parameters we can vary simultaneously
- Will slowly try progressively more complex fits, more dimensions
- The dream: 5D? fit to data, floating all operators + form factors in Z expansion
 - Sounds difficult, though

Conclusion

- We should explore what we can measure from the $B \rightarrow D^{(*)} \tau \nu$ kinematic distributions
- Our fit already incorporates some of this information
- We have some machinery to perform the fit
- In principle, we have access to the full distribution, with some resolution
- How far can we push this?