

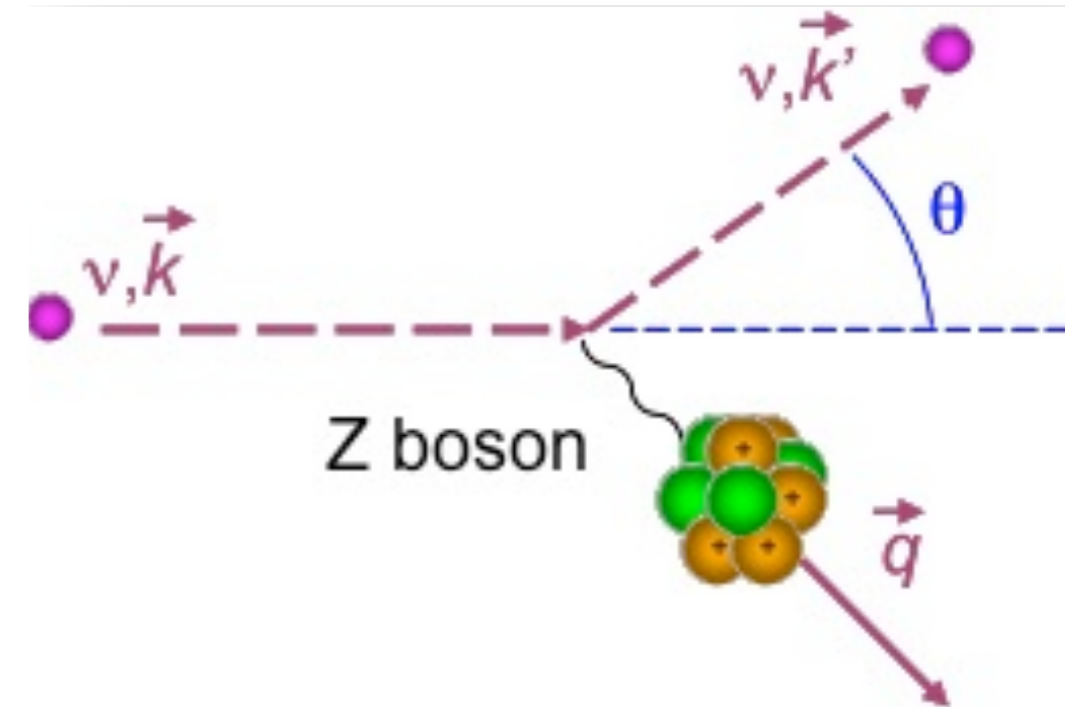


$\nu$  Scattering

Phil Barbeau, Duke University

# Coherent $\nu$ -Nucleus Scattering

- Predicted in 1974 with the realization of the weak neutral current: as yet unobserved
- Neutrino scatters coherently off all Nucleons  $\rightarrow$  cross section enhancement:  $\sigma \propto \mathbf{N}^2$
- Initial and final states must be identical: Neutral Current elastic scattering
- Nucleons must recoil in phase  $\rightarrow$  low momentum transfer  $qR < 1 \rightarrow$  very low energy nuclear recoil



D. Z. Freedman, PRD 9 (5) 1974

# Why Measure Coherent $\nu$ -Nucleus Scattering?

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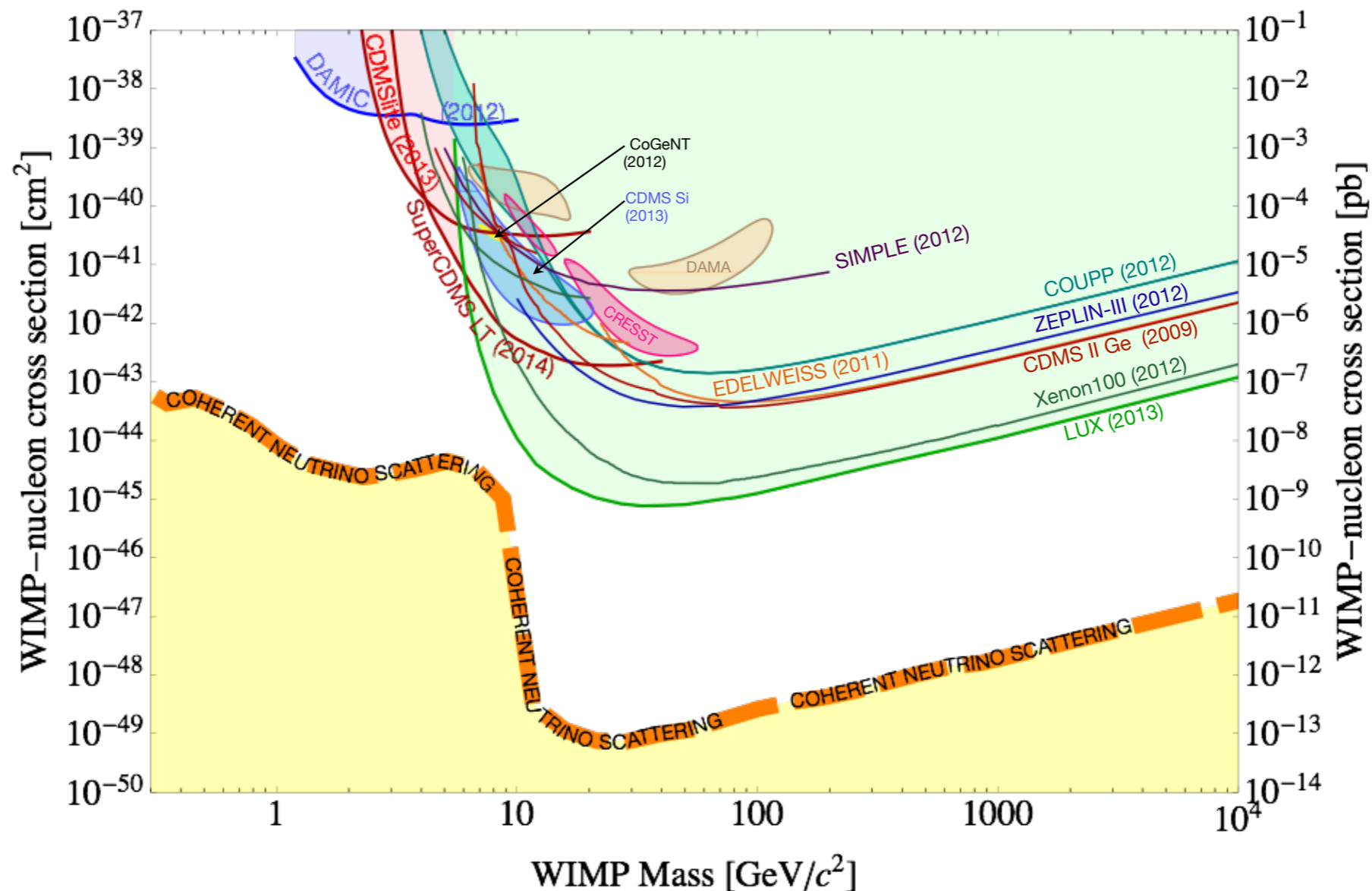
- Largest  $\sigma$  in Supernovae dynamics. We should measure it to validate the models

**J.R. Wilson, PRL 32 (74) 849**



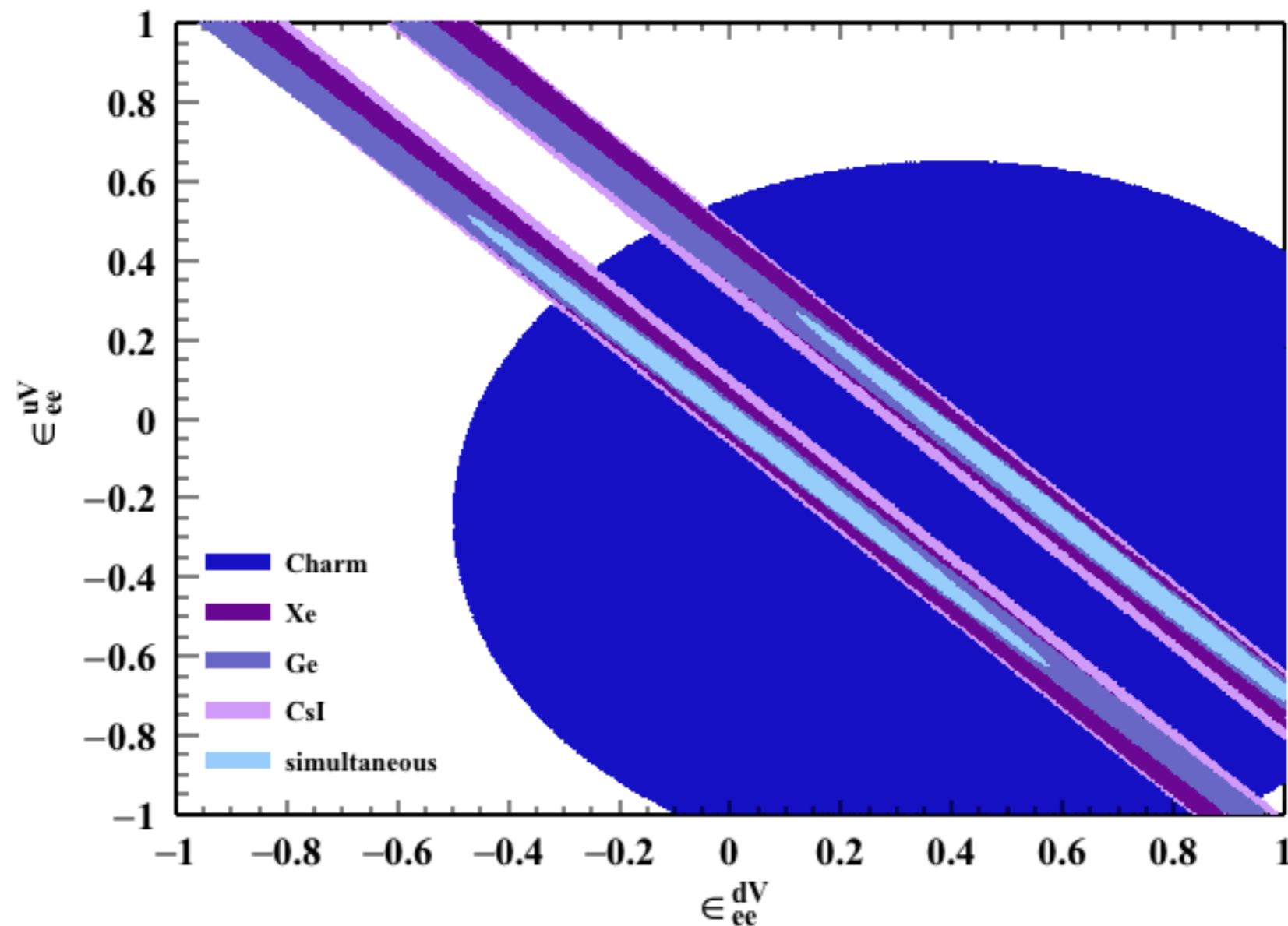
# Why Measure Coherent $\nu$ -Nucleus Scattering?

- CEvNS is an irreducible background from WIMP searches, and should be measured in order to validate background models and detector responses.



# Why Measure Coherent $\nu$ -Nucleus Scattering?

- By measuring the relative rates on several nuclear targets we dramatically extend the sensitivity of searches for Non-Standard  $\nu$  Interactions **K. Scholberg, Phys.Rev.D73:033005,2006**  
**J. Barranco et al., JHEP0512:021,2005**



# Why Measure Coherent $\nu$ -Nucleus Scattering?

- A high- $\sigma$ , neutral current detector would be a clean way to search for sterile  $\nu$ 's

**A. Drukier & L. Stodolsky, PRD 30 (84) 2295**

- The development of a coherent neutrino scattering detection capability provides perhaps the best way to explore any sterile neutrino sector that could be uncovered with ongoing experiments.

**A. J. Anderson et al., PRD 86 013004 (2012)**

- Coherent  $\sigma$  proportional to  $Q_w^2$ . A precision test of  $\sigma$  is a sensitive test of new physics above the weak scale.  $M_{\text{top}}$  and  $M_{\text{higgs}}$  are known  $\rightarrow$  Remaining theoretical uncertainties  $\sim 0.2\%$

**L. M. Krauss, PLB 269, 407**

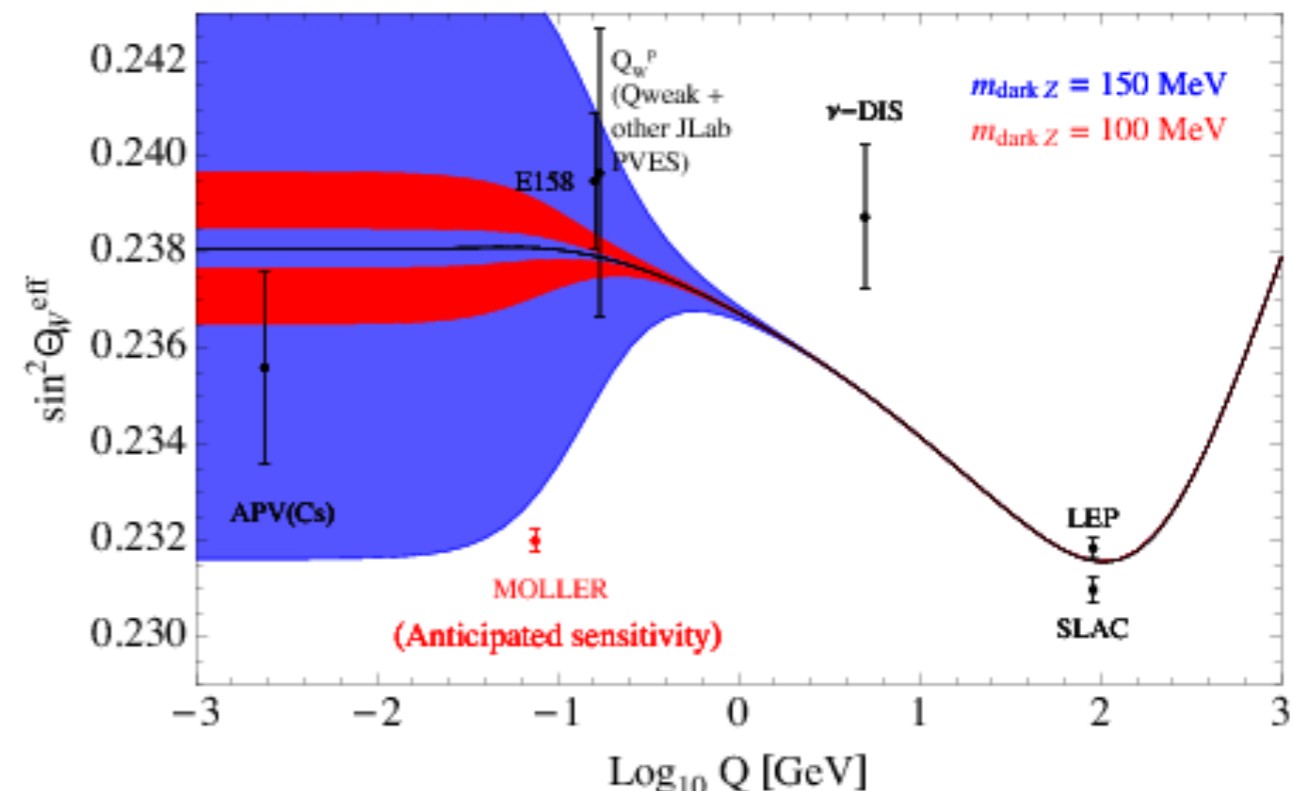
$$\sigma_{\text{coh}} \sim \frac{G_f^2 E^2}{4\pi} (Z(4 \sin^2 \theta_w - 1) + N)^2$$

- Neutrino Magnetic Moments

**A. C. Dodd, et al., PLB 266 (91), 434**

- Measuring the neutron distribution functions (Form Factors)

**K. Patton, et al., PRC 86, 024216**



MOLLER Collaboration, arXiv:1411.4088

# COHERENT SNS



Duke University  
Indiana University  
ITEP  
LANL  
LBNL  
MEPhI

NC Central University  
NC State University  
New Mexico State University  
ORNL  
SNL

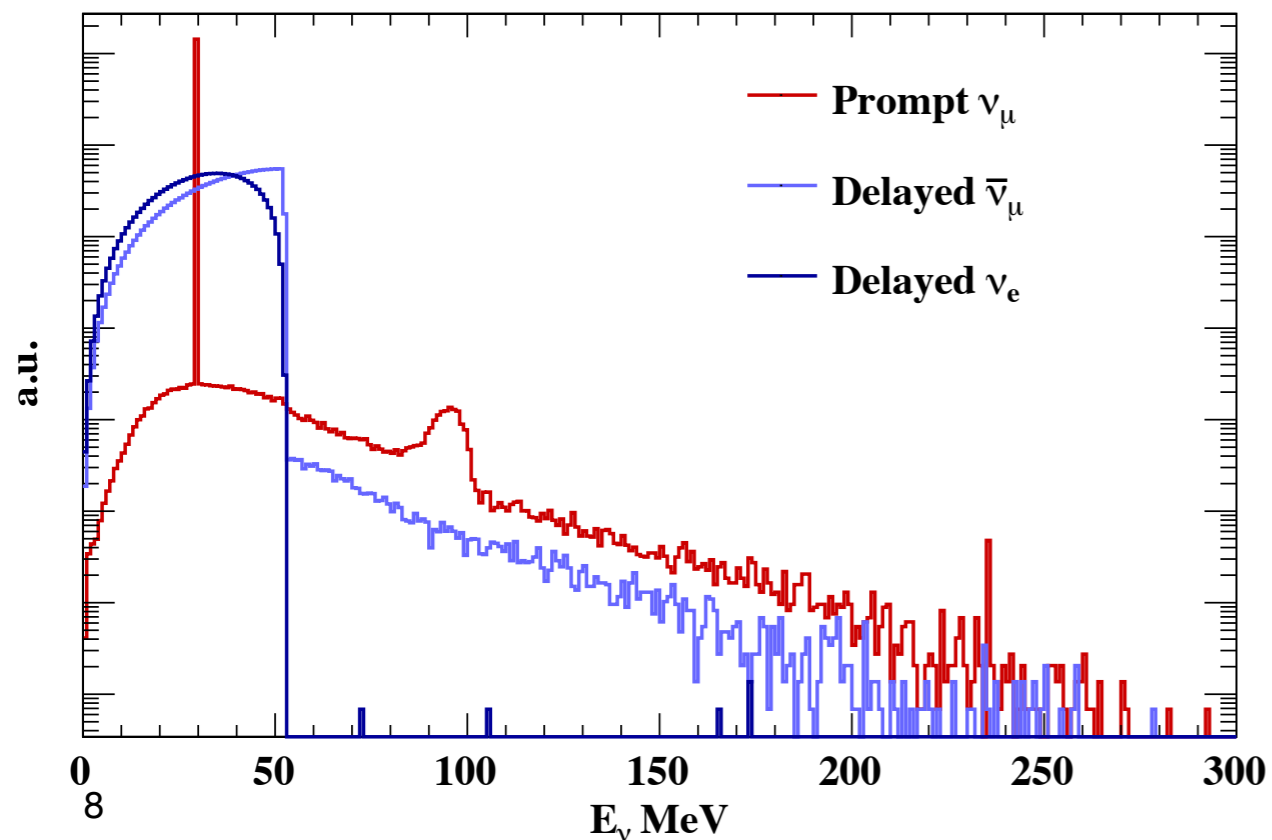
TUNL  
UC Berkeley  
University of Chicago  
University of Florida  
University of Tennessee  
University of Washington

# The Spallation Neutron Source

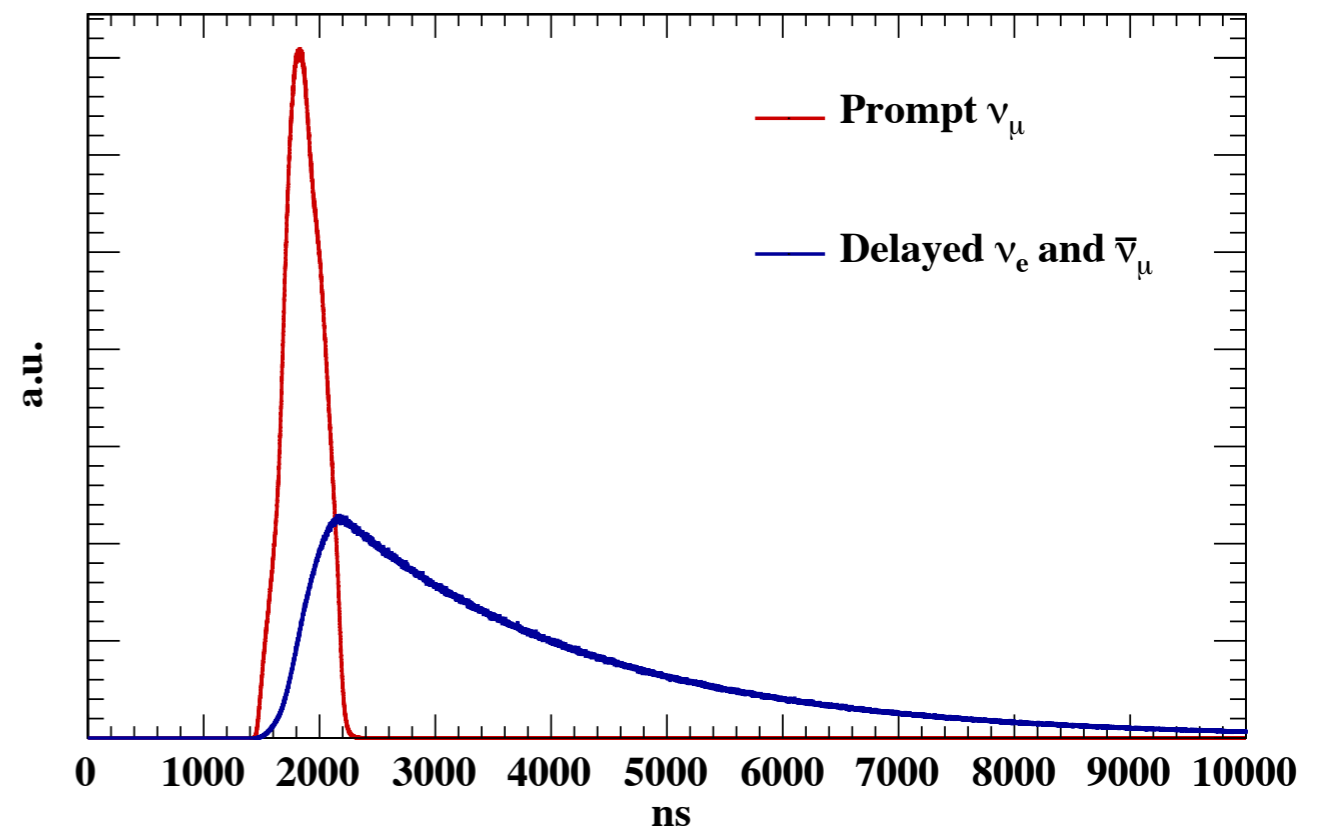
- Pion Decay-at-Rest Neutrino Source
- $\nu$  flux  $4.3 \times 10^7 \nu \text{ cm}^{-1} \text{ s}^{-1}$  at 20 m
- Pulsed: 800 ns full-width at 60 Hz



**<1% contamination from non-CEvNS scatters**



**$\sim 4 \times 10^{-5}$  background reduction**

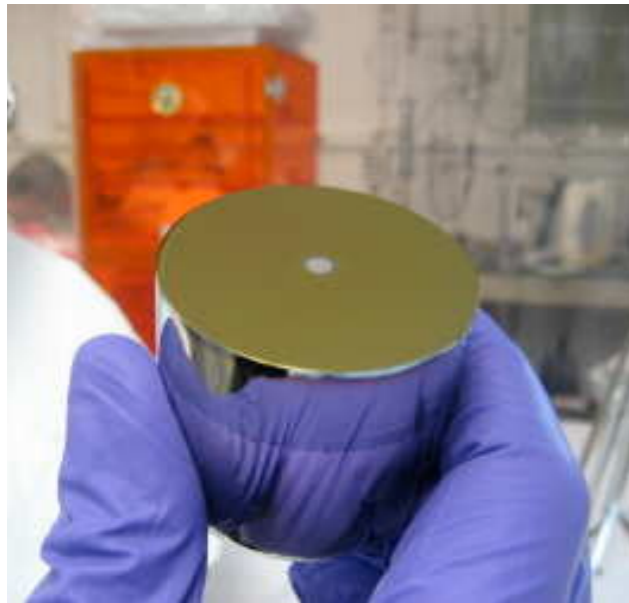




# How to Make an Unambiguous Measurement

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- Observe the pulsed  $\nu$  time-structure
- Observe the  $2.2 \mu\text{s}$  characteristic decay of muon decay  $\nu$ 's
- Observe the  $N^2$  cross section behavior between targets



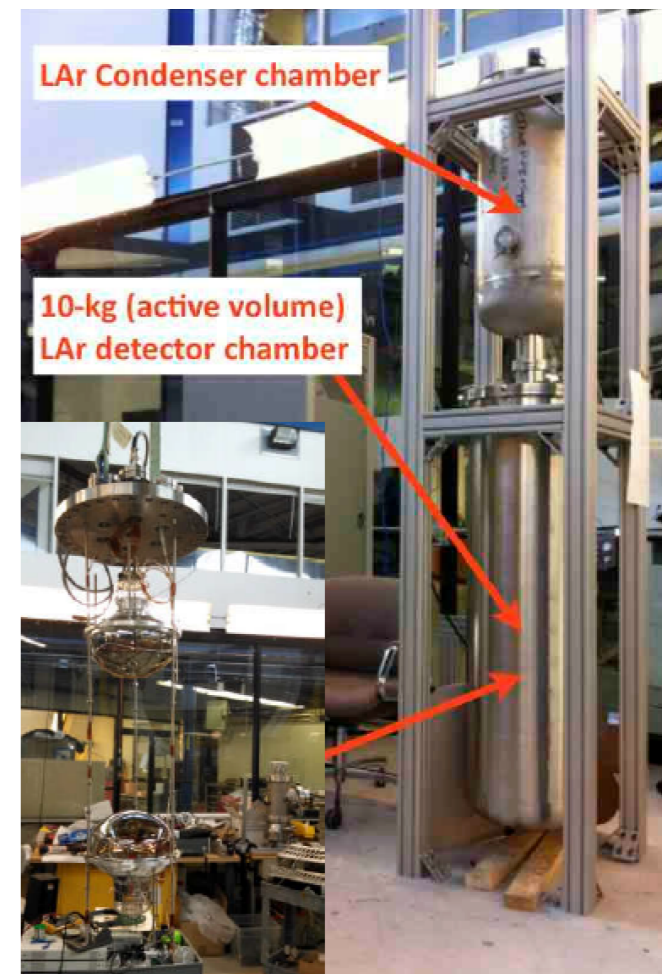
P-Type Point  
Contact HPGe



Low-Background  
CsI[Na]



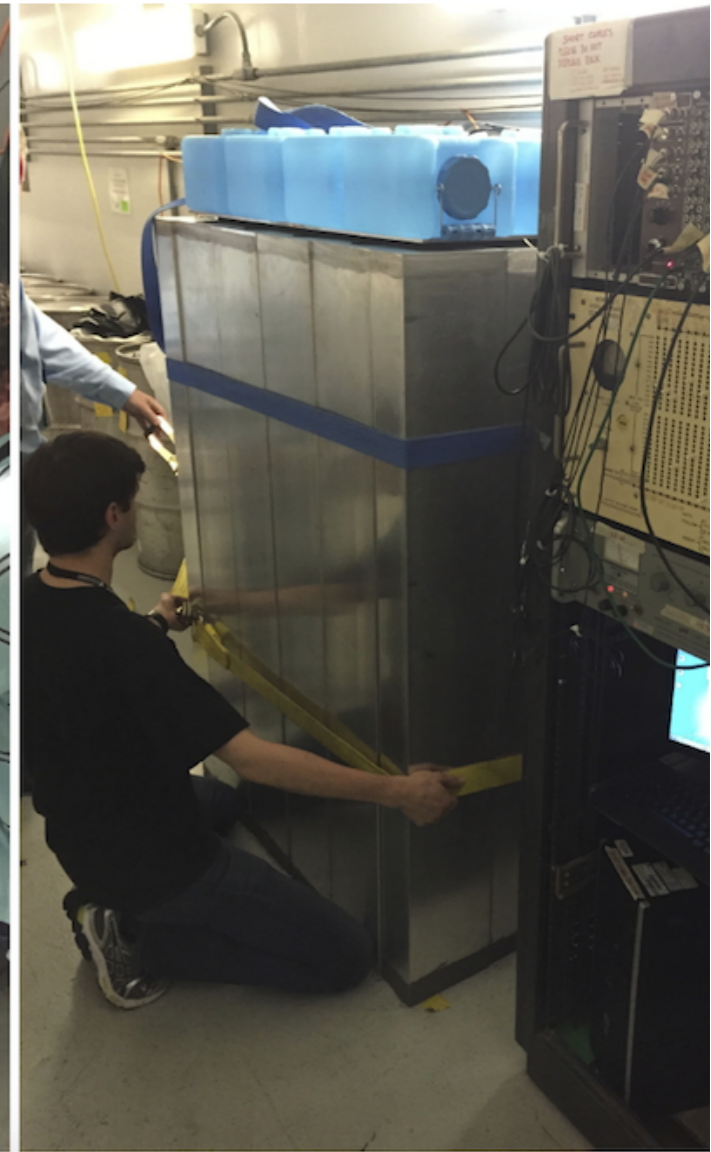
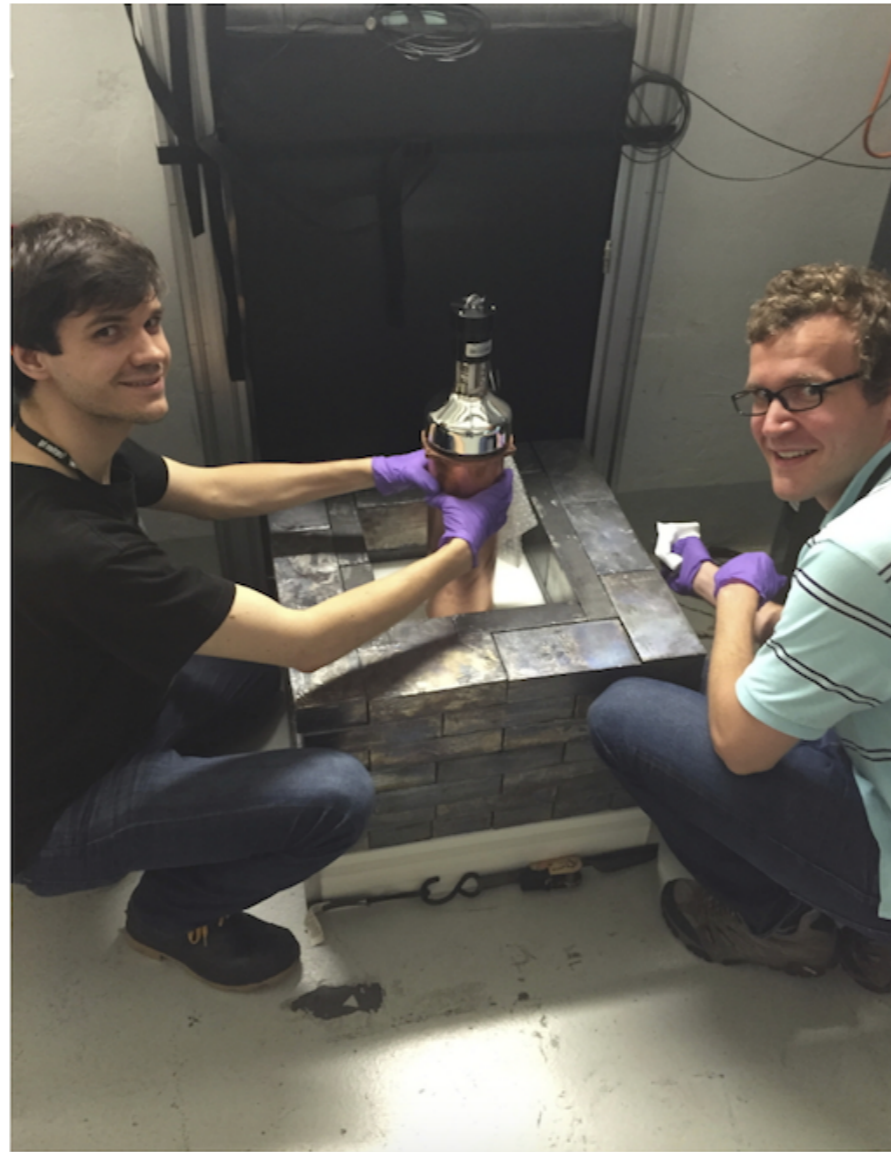
NaI[Tl]



Single  
Phase LAr

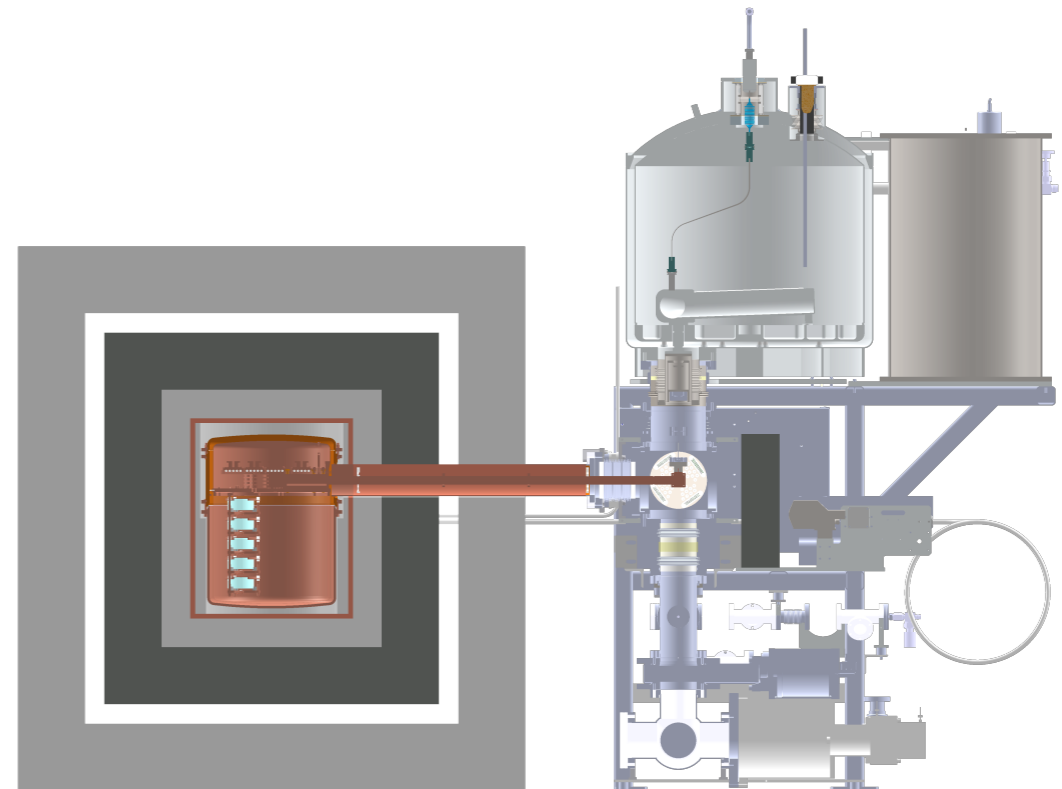
# Detector Subsystems: CsI[Na]

- 14 kg low-background CsI[Na] crystal
- Large N: 74, 78
- Already installed at SNS



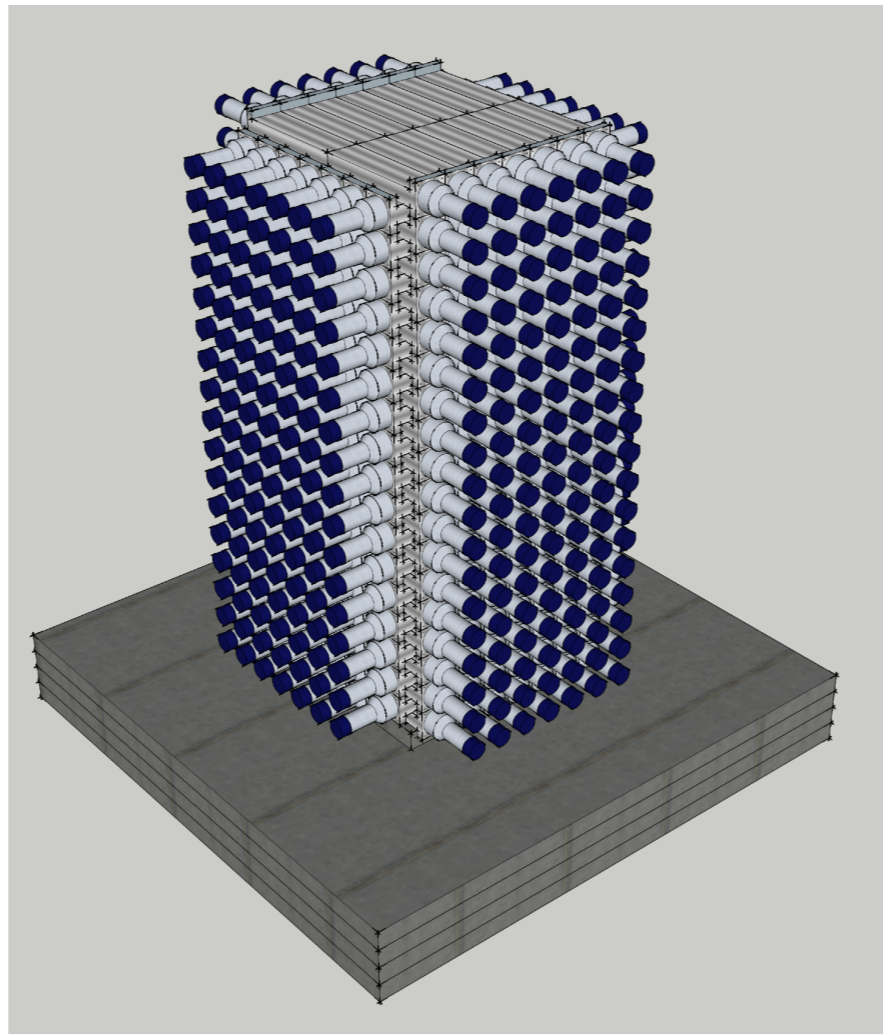
# Detector Subsystems: HPGe PPCs

- Repurposed MAJORANA DETECTORS
- 5-10kg PPC detector mass
- Smaller N: 38-44
- Excellent resolution at low energies
- Well-measured quenching factor



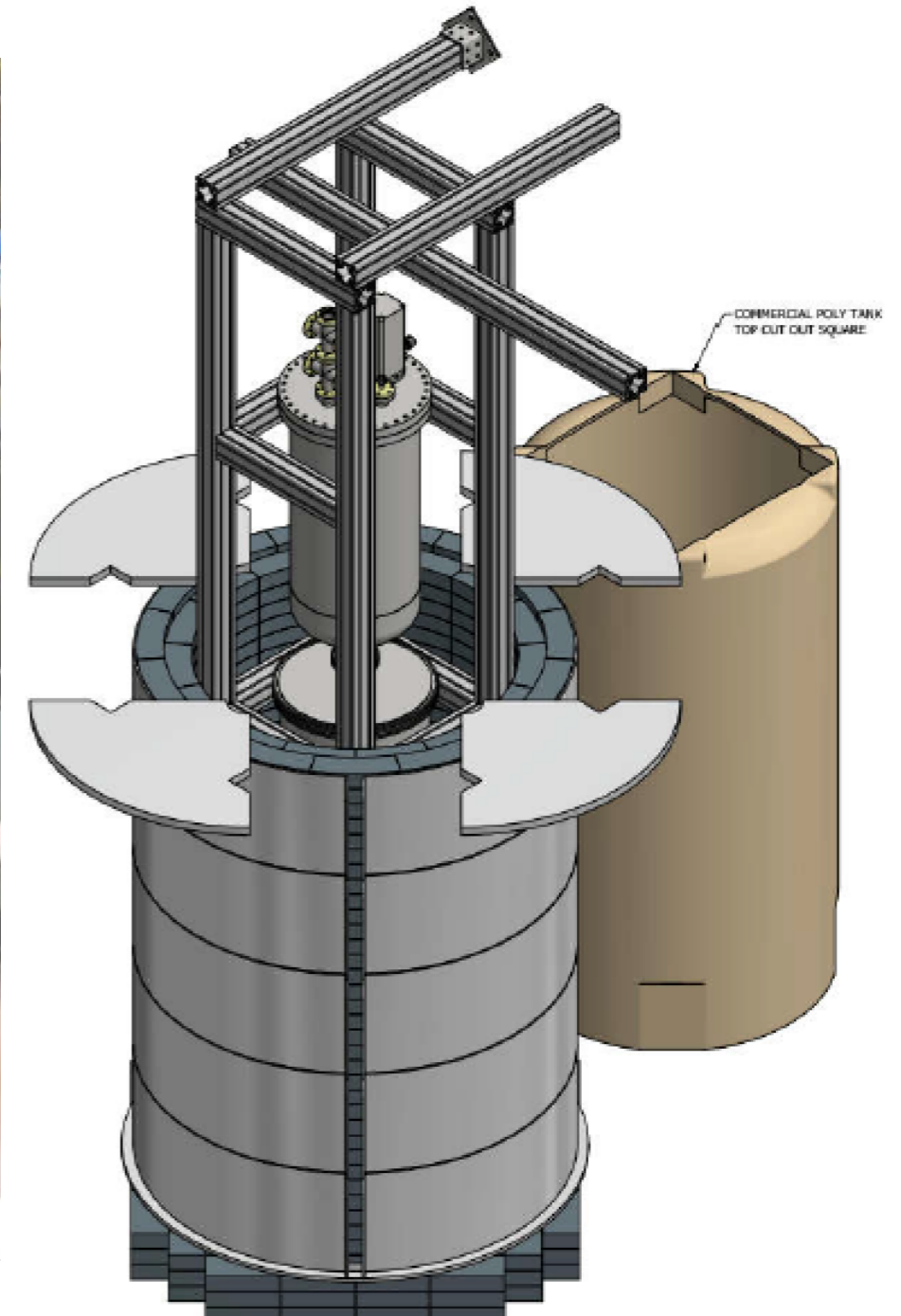
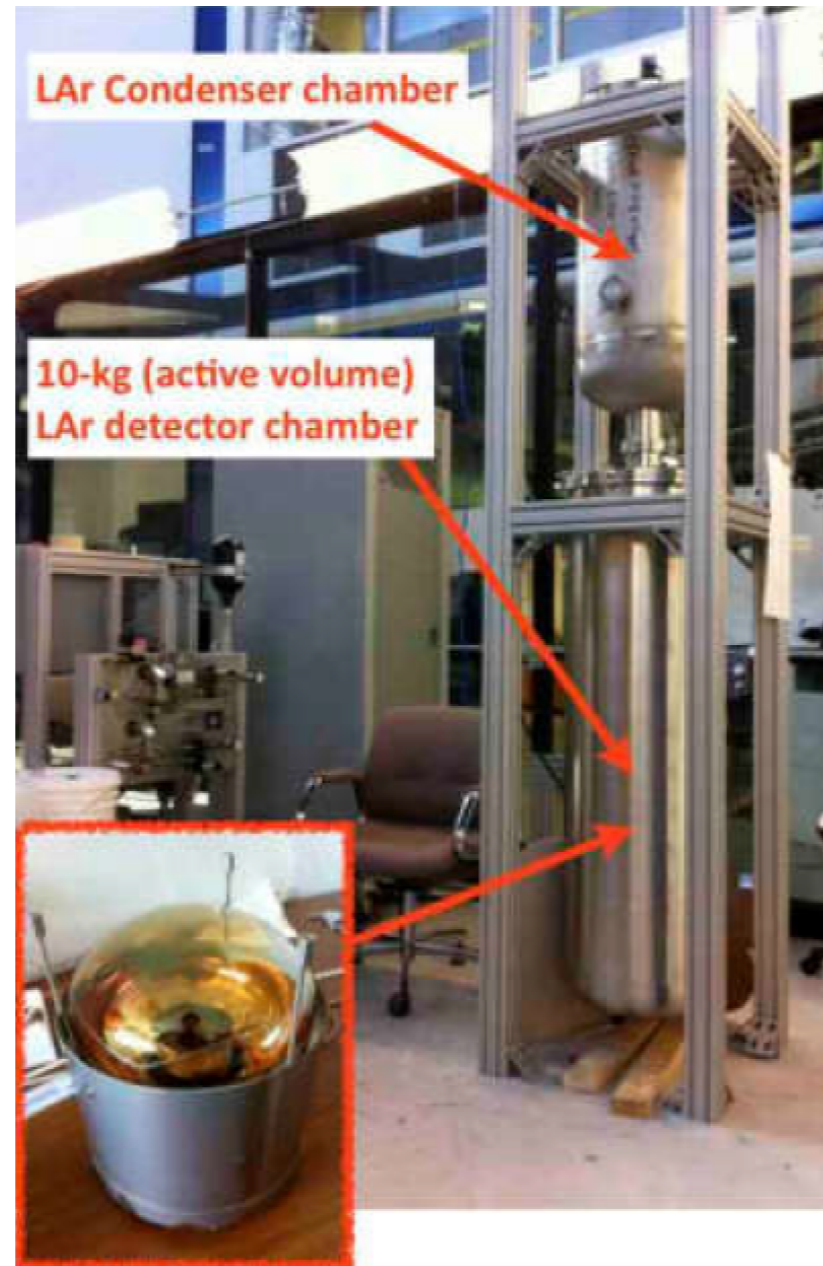
# Detector Subsystems: NaI[Tl]

- Initial deployment  
185 kgs
- Up to 9 tons in  
hand
- $N = 23$  for Na
- Instrumentation  
tests underway at  
Duke and UW
- QF understood



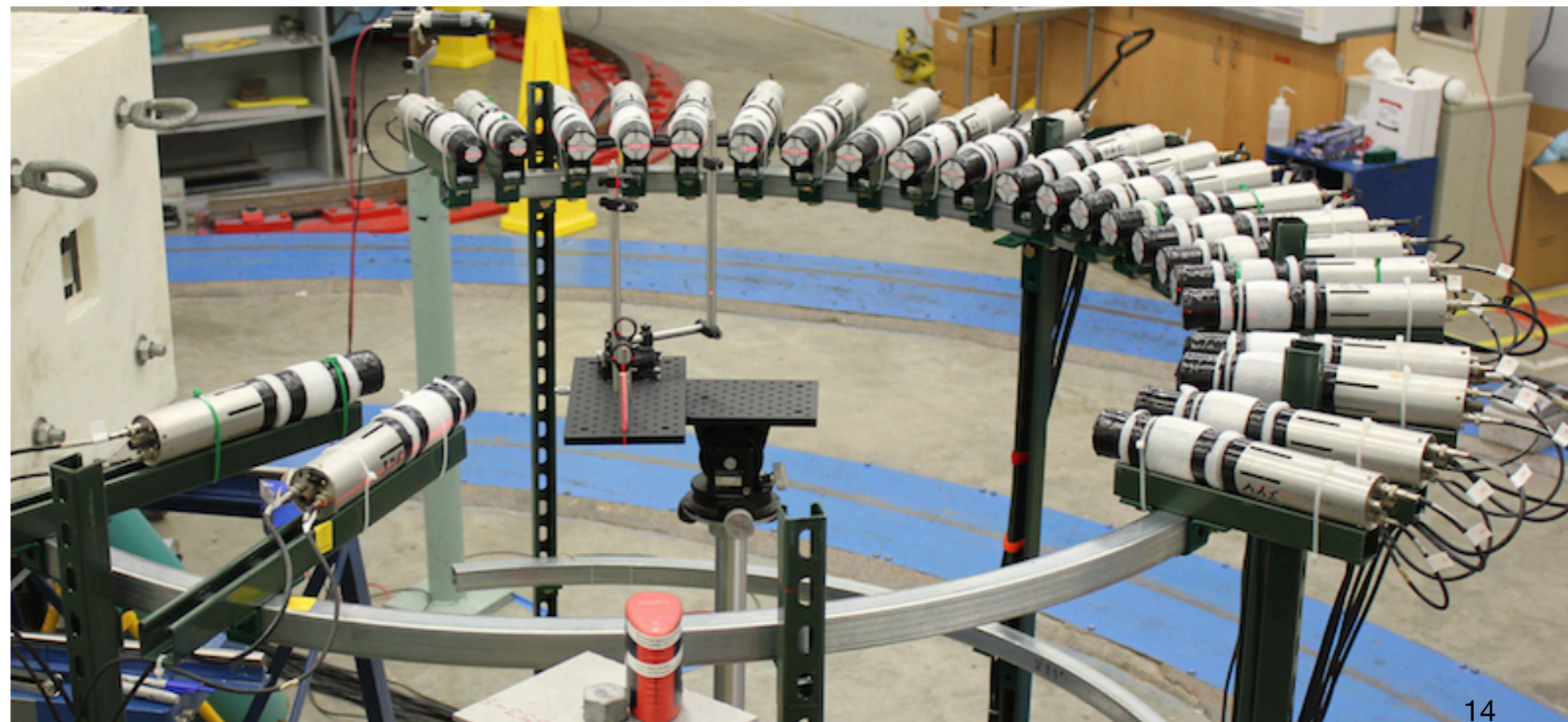
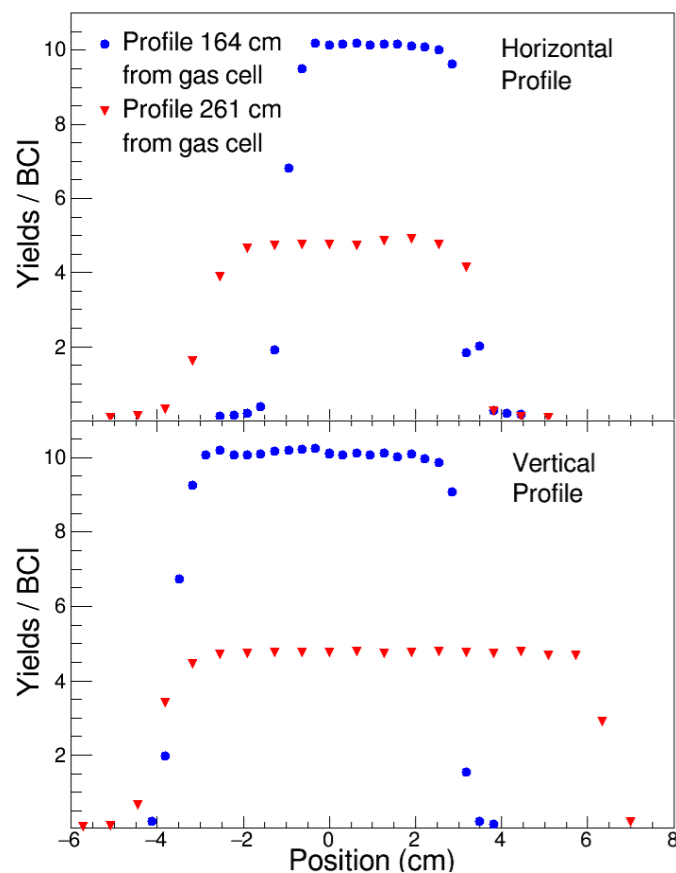
# Detector Subsystems: Single Phase LAr

- Medium N: 40
- CENNS-10  
Detector under  
consideration
- QF also known



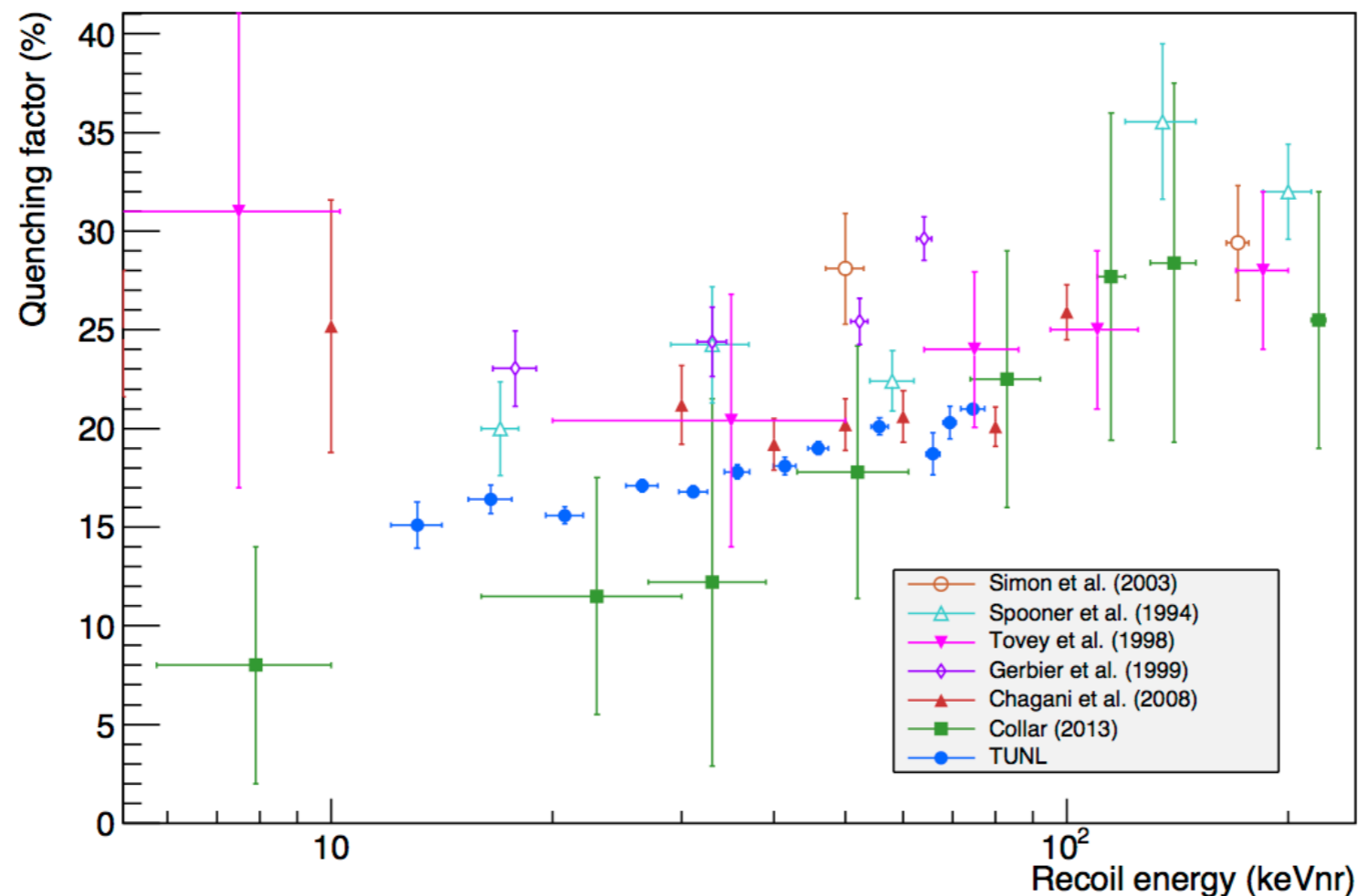
# Quenching Factor Measurements

- A facility has been developed at Duke/TUNL to enable the precision calibration of all of these detectors. *CsI(Na)* and *NaI(Tl)* data in the can. *Quenching factor uncertainties are the dominant uncertainty on the cross-sections, after the beam flux.*
- The neutron beam is tunable (20 keV - 3 MeV), Monochromatic (3 keV width), collimated (1.5 cm) and pulsed (2 ns)



# Quenching Factor Measurements

- The story of the quenching factors of Na recoils goes back a long way. High precision measurements recently performed by Duke and Princeton confirm ~ 15%.
- Recently remeasured CsI[Na] with encouraging results.



# NaI[Tl]: Two primary measurement goals

- CEvNS on Na
- The electron neutrino Charged-Current interaction on  $^{127}\text{I}$

Isotope	Reaction Channel	Source	Experiment	Measurement ( $10^{-42} \text{ cm}^2$ )	Theory ( $10^{-42} \text{ cm}^2$ )
$^2\text{H}$	$^2\text{H}(\nu_e, e^-)pp$	Stopped $\pi/\mu$	LAMPF	$52 \pm 18(\text{tot})$	54 (IA) (Tatara <i>et al.</i> , 1990)
$^{12}\text{C}$	$^{12}\text{C}(\nu_e, e^-)^{12}\text{N}_{\text{g.s.}}$	Stopped $\pi/\mu$	KARMEN	$9.1 \pm 0.5(\text{stat}) \pm 0.8(\text{sys})$	9.4 [Multipole](Donnelly and Peccei, 1979)
		Stopped $\pi/\mu$	E225	$10.5 \pm 1.0(\text{stat}) \pm 1.0(\text{sys})$	9.2 [EPT] (Fukugita <i>et al.</i> , 1988).
		Stopped $\pi/\mu$	LSND	$8.9 \pm 0.3(\text{stat}) \pm 0.9(\text{sys})$	8.9 [CRPA] (Kolbe <i>et al.</i> , 1999b)
	$^{12}\text{C}(\nu_e, e^-)^{12}\text{N}^*$	Stopped $\pi/\mu$	KARMEN	$5.1 \pm 0.6(\text{stat}) \pm 0.5(\text{sys})$	5.4-5.6 [CRPA] (Kolbe <i>et al.</i> , 1999b)
		Stopped $\pi/\mu$	E225	$3.6 \pm 2.0(\text{tot})$	4.1 [Shell] (Hayes and S, 2000)
		Stopped $\pi/\mu$	LSND	$4.3 \pm 0.4(\text{stat}) \pm 0.6(\text{sys})$	
	$^{12}\text{C}(\nu_\mu, \nu_\mu)^{12}\text{C}^*$	Stopped $\pi/\mu$	KARMEN	$3.2 \pm 0.5(\text{stat}) \pm 0.4(\text{sys})$	2.8 [CRPA] (Kolbe <i>et al.</i> , 1999b)
	$^{12}\text{C}(\nu, \nu)^{12}\text{C}^*$	Stopped $\pi/\mu$	KARMEN	$10.5 \pm 1.0(\text{stat}) \pm 0.9(\text{sys})$	10.5 [CRPA] (Kolbe <i>et al.</i> , 1999b)
$^{12}\text{C}(\nu_\mu, \mu^-)X$	Decay in Flight	LSND	$1060 \pm 30(\text{stat}) \pm 180(\text{sys})$	1750-1780 [CRPA] (Kolbe <i>et al.</i> , 1999b)	
				1380 [Shell] (Hayes and S, 2000)	
				1115 [Green's Function] (Meucci <i>et al.</i> , 2004)	
$^{12}\text{C}(\nu_\mu, \mu^-)^{12}\text{N}_{\text{g.s.}}$	Decay in Flight	LSND	$56 \pm 8(\text{stat}) \pm 10(\text{sys})$	68-73 [CRPA] (Kolbe <i>et al.</i> , 1999b)	
				56 [Shell] (Hayes and S, 2000)	
$^{56}\text{Fe}$	$^{56}\text{Fe}(\nu_e, e^-)^{56}\text{Co}$	Stopped $\pi/\mu$	KARMEN	$256 \pm 108(\text{stat}) \pm 43(\text{sys})$	264 [Shell] (Kolbe <i>et al.</i> , 1999a)
$^{71}\text{Ga}$	$^{71}\text{Ga}(\nu_e, e^-)^{71}\text{Ge}$	$^{51}\text{Cr}$ source	GALLEX, ave.	$0.0054 \pm 0.0009(\text{tot})$	0.0058 [Shell] (Haxton, 1998)
		$^{51}\text{Cr}$	SAGE	$0.0055 \pm 0.0007(\text{tot})$	
		$^{37}\text{Ar}$ source	SAGE	$0.0055 \pm 0.0006(\text{tot})$	0.0070 [Shell] (Bahcall, 1997)
$^{127}\text{I}$	$^{127}\text{I}(\nu_e, e^-)^{127}\text{Xe}$	Stopped $\pi/\mu$	LSND	$284 \pm 91(\text{stat}) \pm 25(\text{sys})$	210-310 [Quasi-particle] (Engel <i>et al.</i> , 1994)



# Backgrounds

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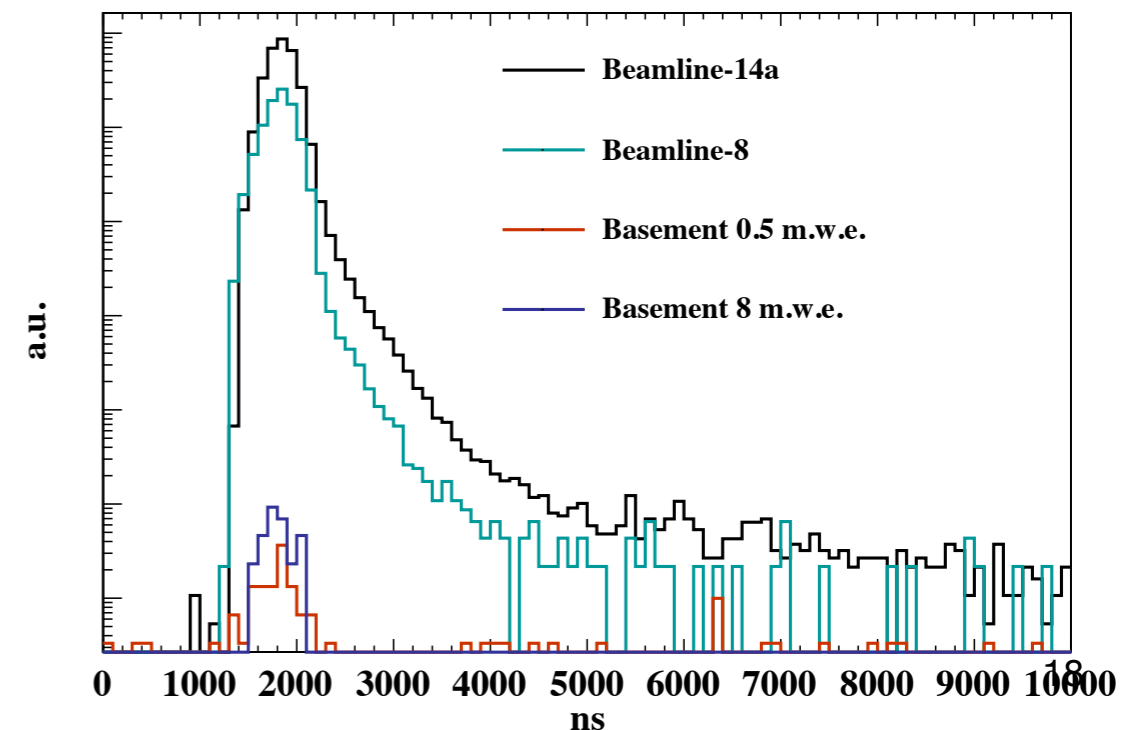
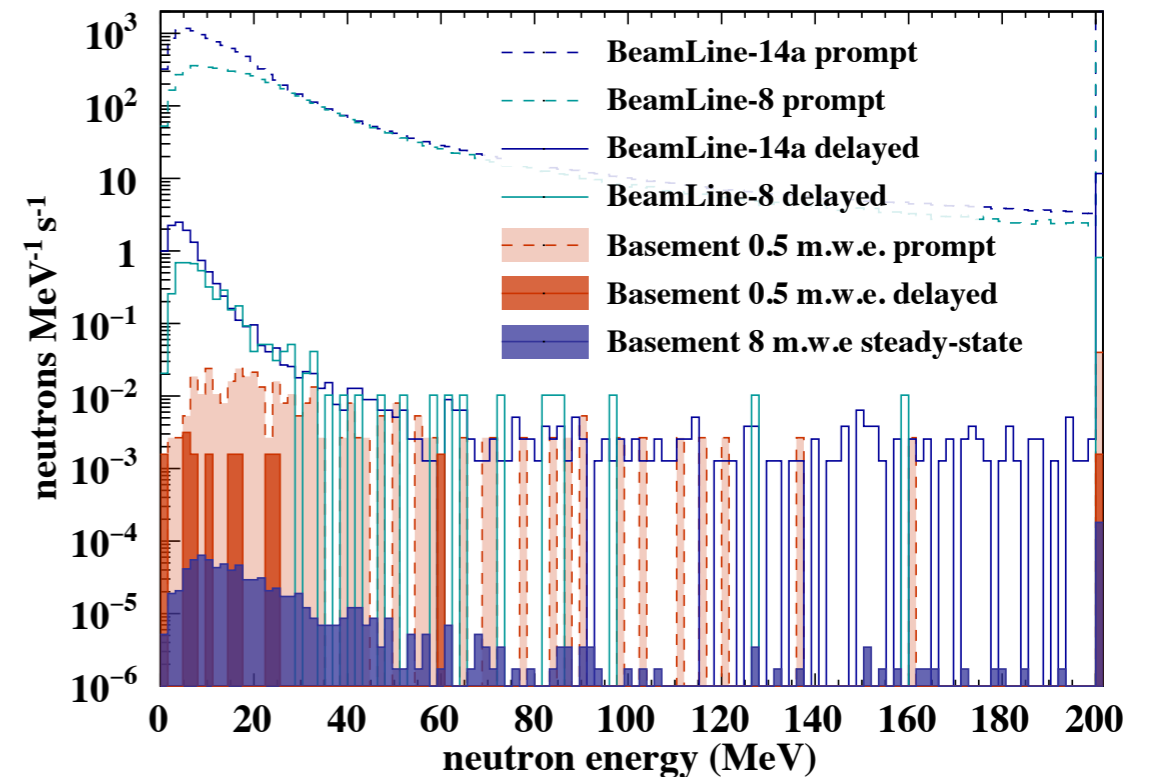
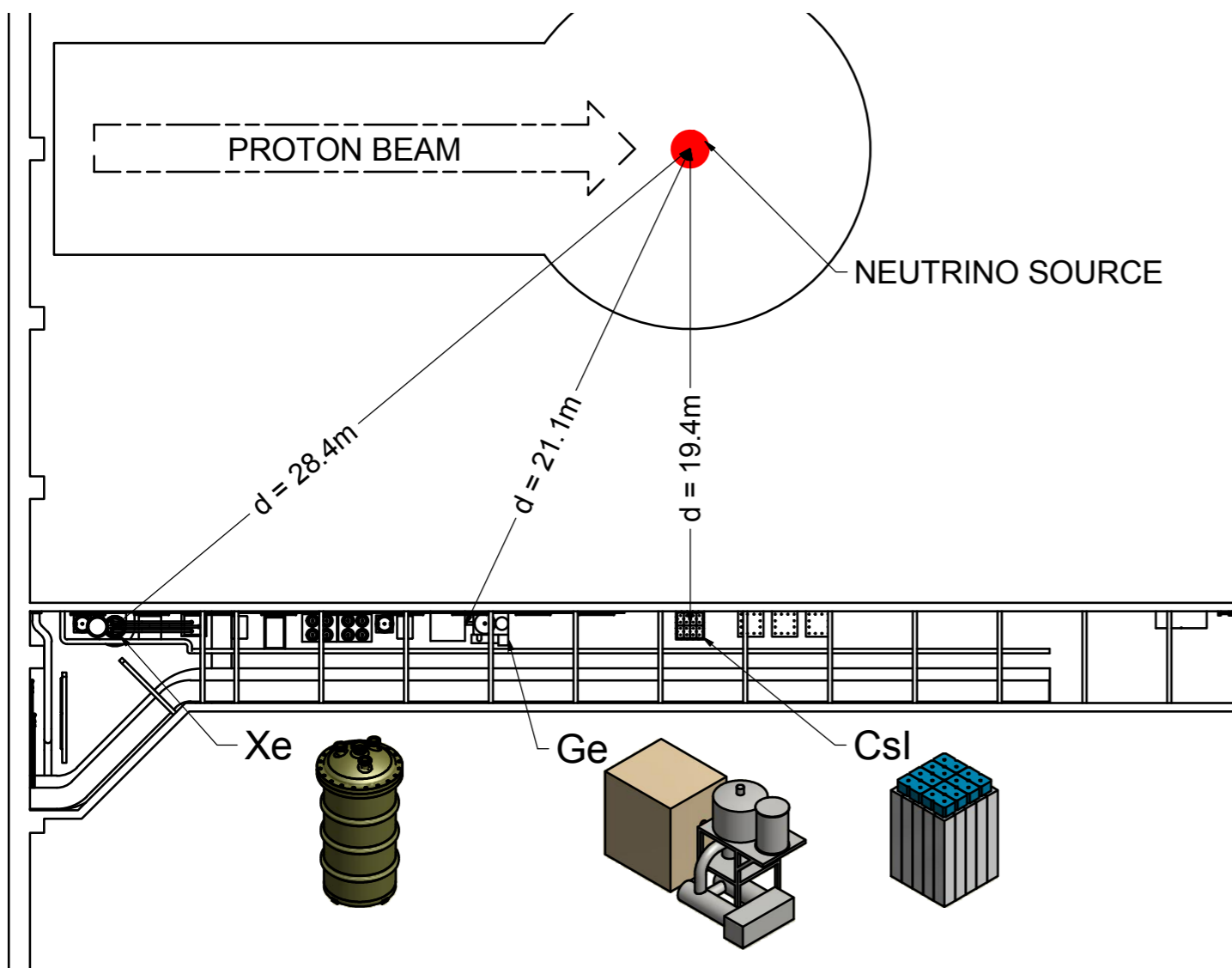
The SNS is a facility designed to produce neutrons ( $> 100$  MeV), that are pulsed with the same time structure of the neutrinos (**with the exception of the characteristic decay time of the muon**).



**Neutron image of the SNS target, through shielding**

# Hunting for a Background-Free Location: Neutrino Alley

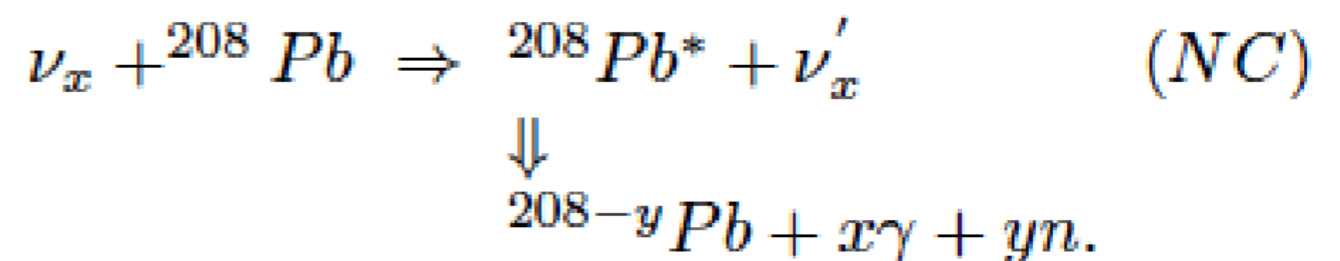
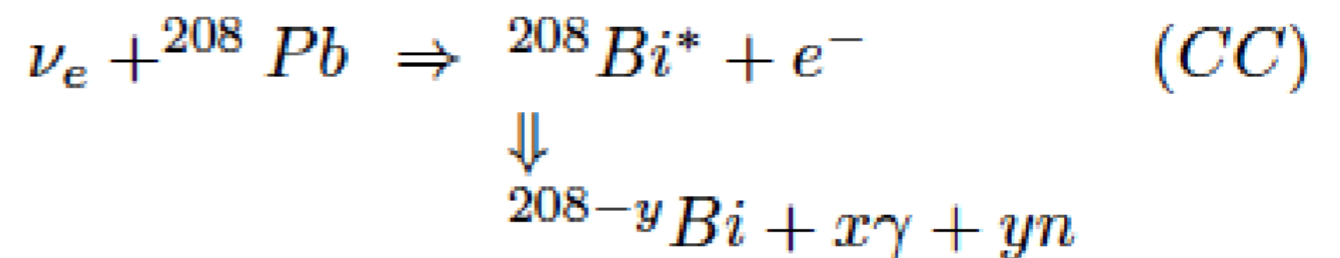
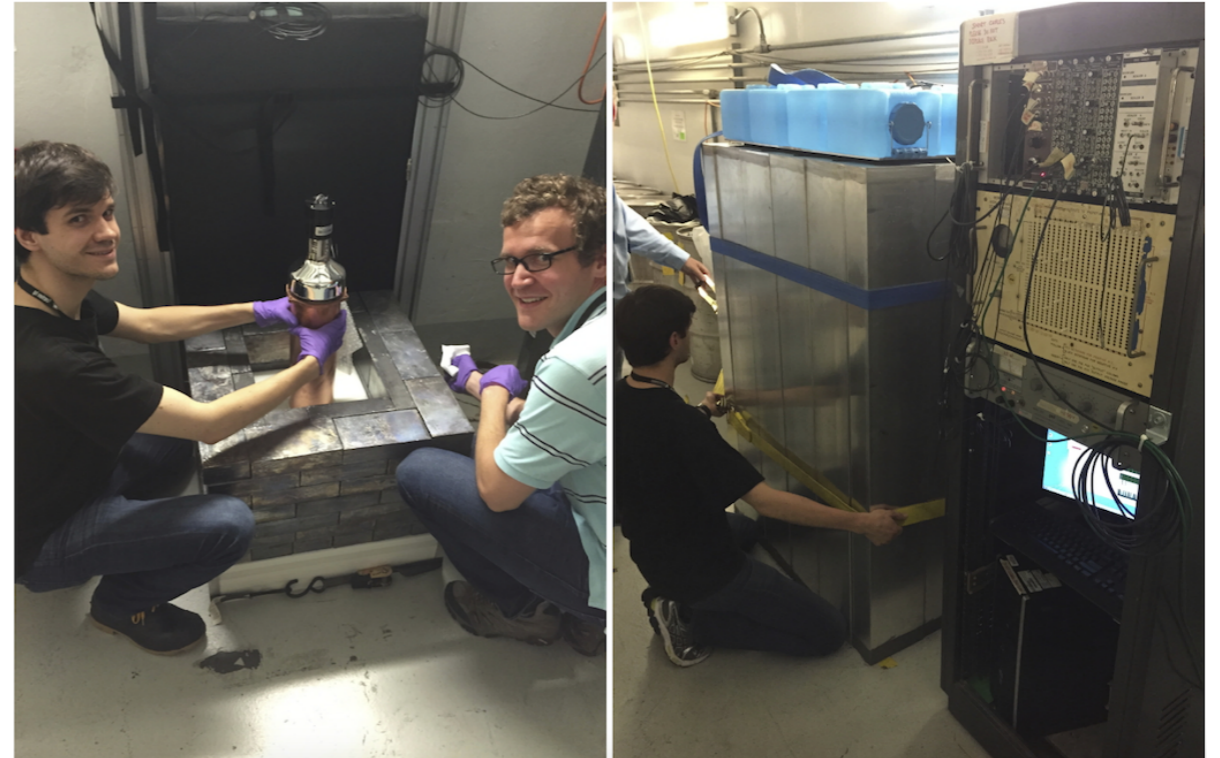
- Extensive background measurement campaign since 2013 points to the SNS basement as the optimal location ( $>10^4$  reduction)



# New Background: $\nu$ -induced neutrons (NINs)

- The detector shields use several tons of lead
- Neutrons can be produced near the detectors. They will be pulsed, and share the  $2.2 \mu\text{s}$  decay time of the  $\nu$ 's
- Need to measure this  $\sigma$  and optimize the shields

## CsI(Na) detector and shield



# NINs: Other uses

- NINs from Pb are fundamental mechanism for detection in HALO supernova neutrino detector [1]
- NIN interactions may influence nucleosynthesis in certain astrophysical environments [2]

[1] C.A. Duba *et al.* J.Phys.Conf.Series 136 (2008)

[2] Y-Z. Qian *et al.*, Phys. Rev. C 55 (1997)

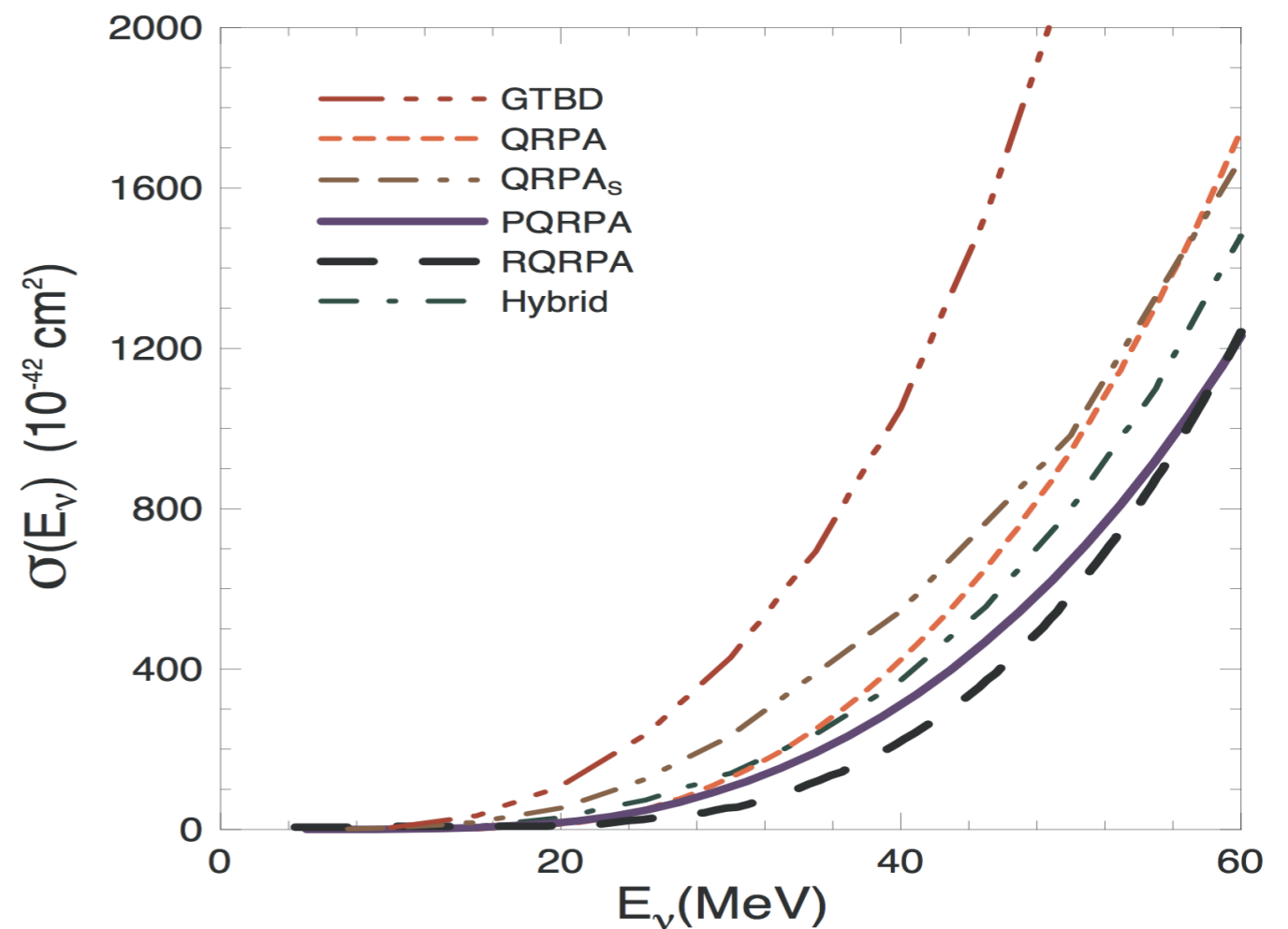
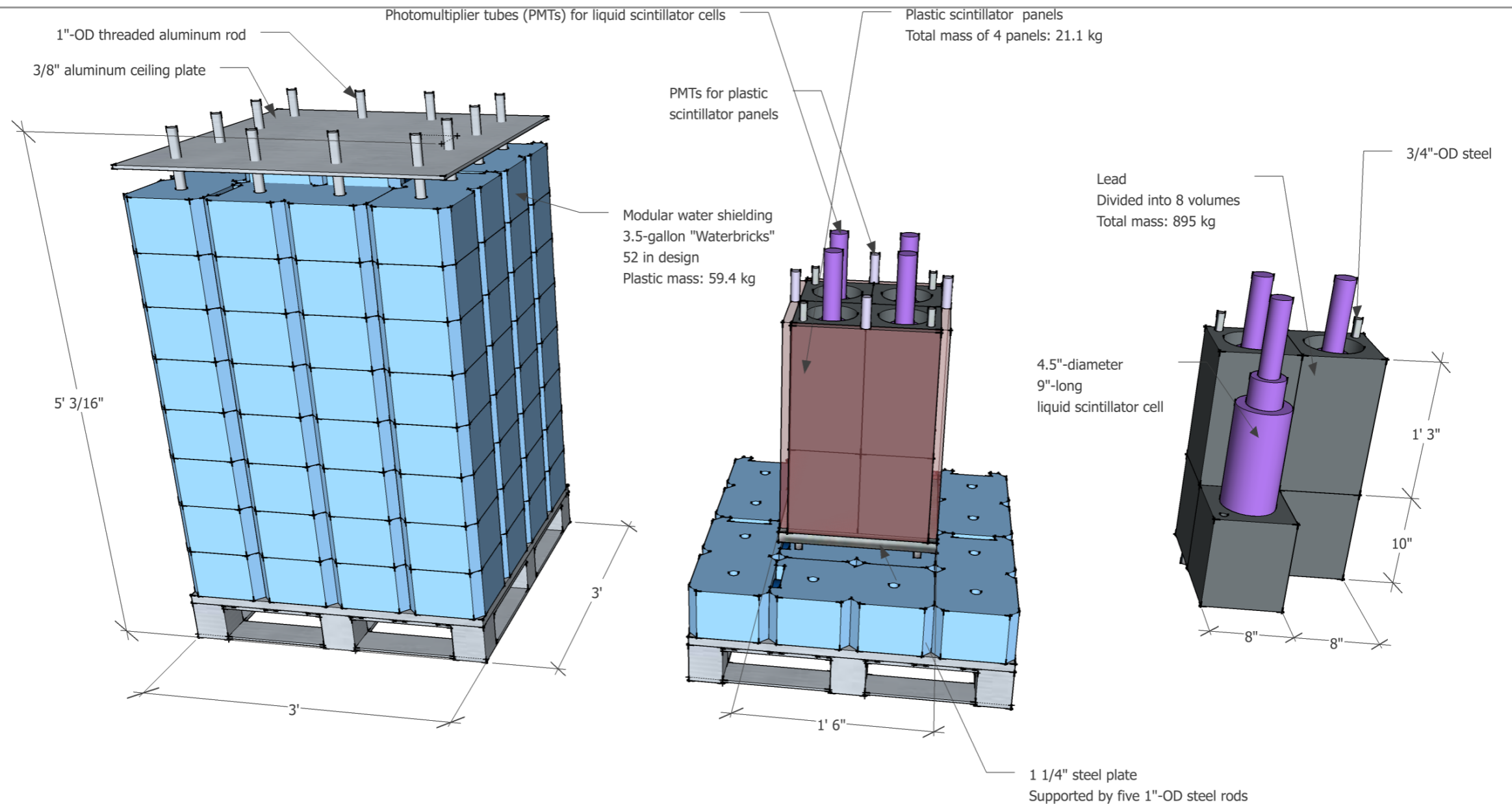


Figure from A.R. Samana and C.A. Bertulani, Phys. Rev. C (2008)

# Measuring the $\nu$ -induced Neutrons



- Several palletized (mobile) targets with LS detectors delivered to the SNS
- Will measure neutrino-induced-neutrons on Pb, Fe and Cu

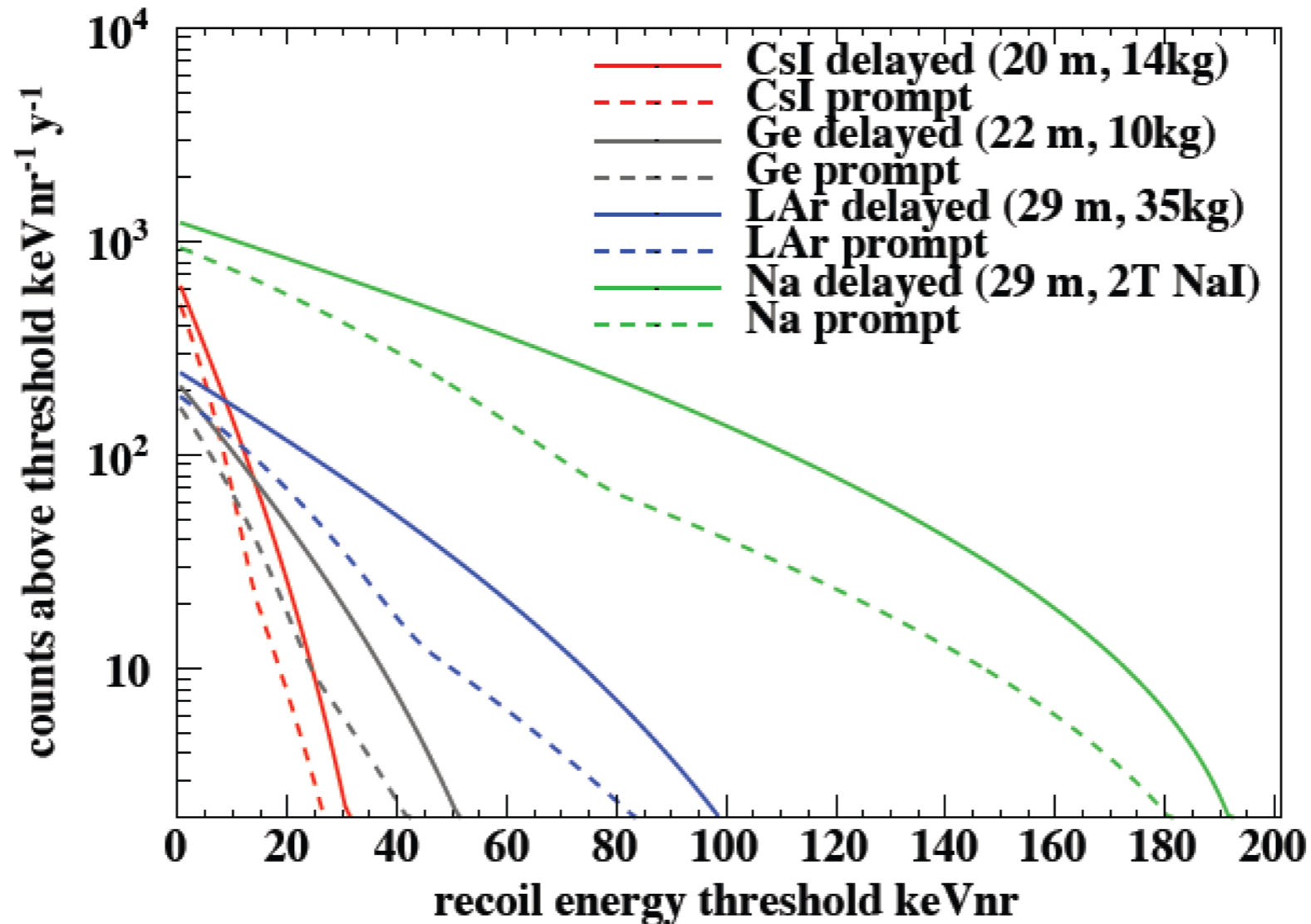
# Measuring the $\nu$ -induced Neutrons

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- The three on-site “neutrino-cubes” also provide nice, compact laboratories for other studies: NaI[TI] CEvNS and  $\nu_e$  CC on  $^{127}\text{I}$

# Expected Signals

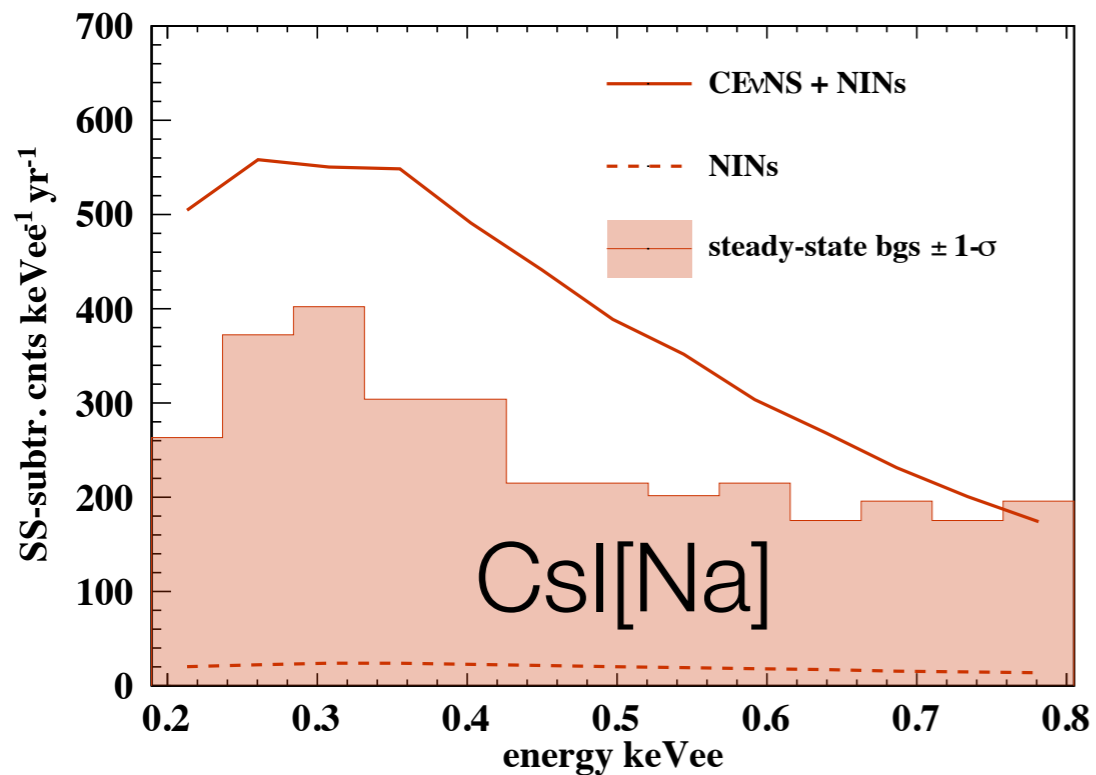


1.2  $\mu$ s cut used to differentiate prompt and delayed neutrinos

Rates depend on detector thresholds and quenching factors.

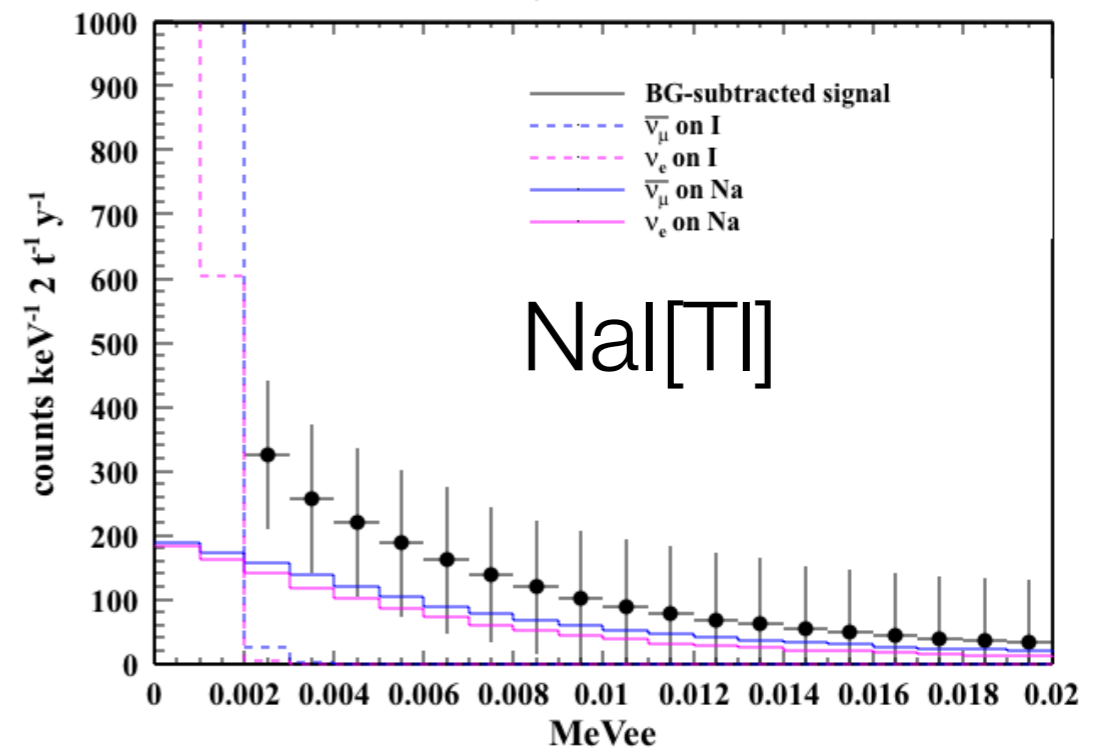
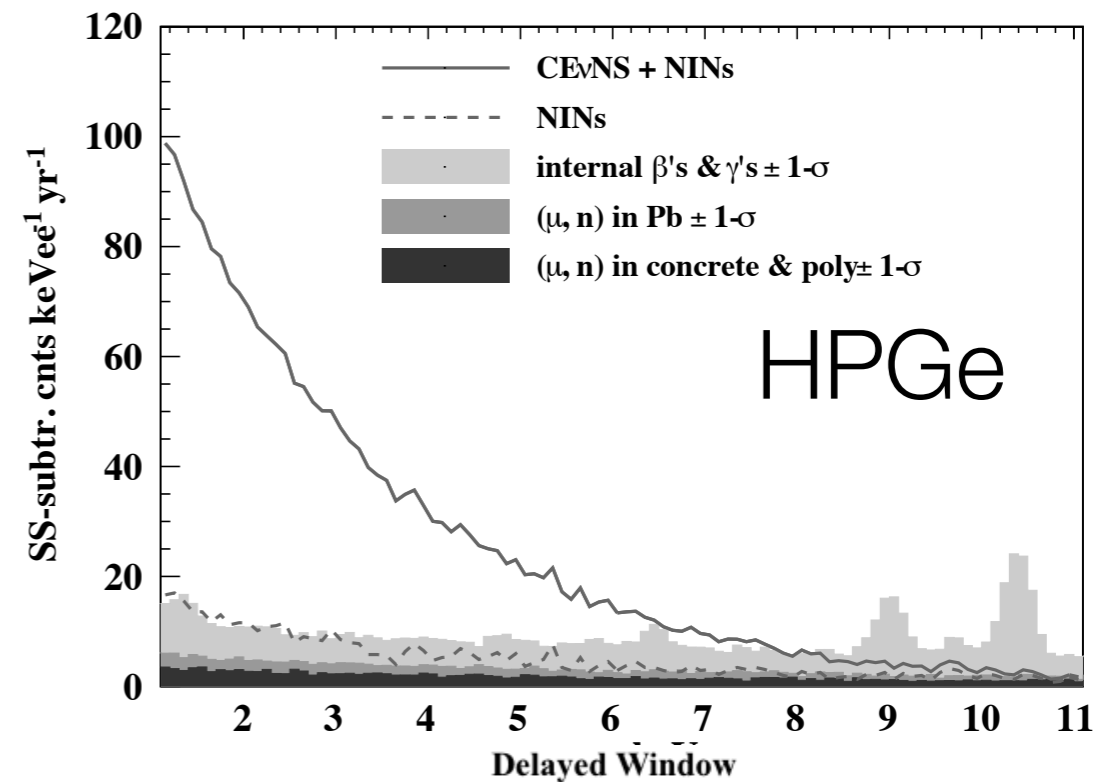
Thresholds and energy resolution effects not included.

# Expected Signals



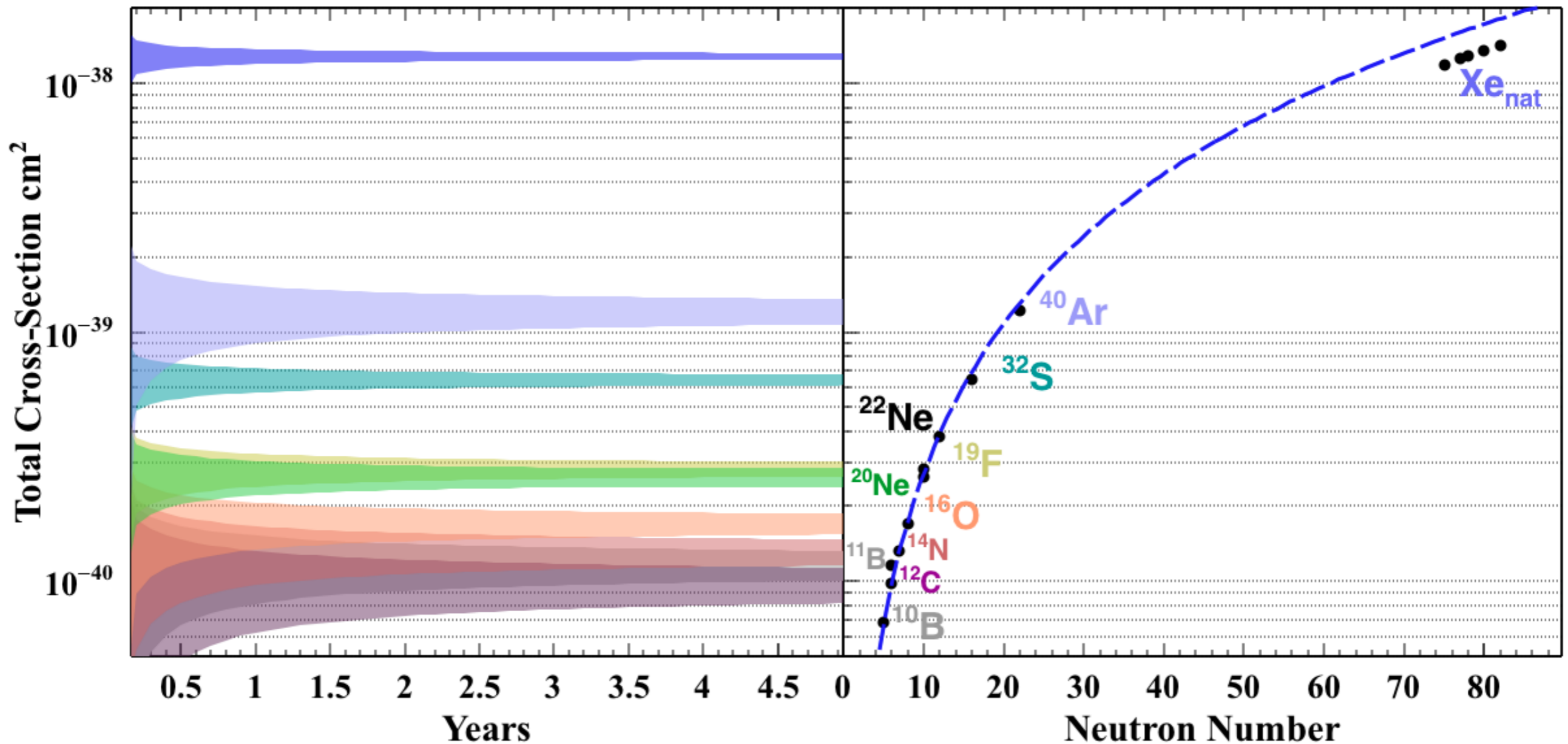
Steady-state background measured with anti-coincident triggers

NIN production rates inform the optimal shielding designs





# Many Detectors working in Concert



- Statistics and systematics limited.
- 10% beam flux uncertainty not included

# Summary



- A new collaboration has formed in 2013, combining the efforts of several groups that have been aiming towards a coherent neutrino-nucleus scattering measurement.
- Background studies indicate the basement as the optimal location
- CsI[Na] is in operation
- Several detectors to measure the  $\nu$ -induced induced neutron emission cross-sections on Pb, Fe and Cu in operation
- This will allow us to confirm that the signal is beam-related (**pulsed nature**), a result of  $\nu$ 's (**2.2  $\mu$ s decay**) and due to CEvNS ( **$\sigma \sim N^2$** )