

Higgs Drawbacks

- ❖ So with the addition of a Higgs boson around 125 GeV particle physics could be “complete”
 - Like Mendeleev’s table for chemistry, but **not understood**. By itself, the Higgs is very unsatisfactory:
 - Why are the couplings to the fermions what they are?
 - ▶ Dumb luck (aka landscape)?
 - What is the link to gravity?
 - What about Dark Matter?
 - Why does the Higgs break the symmetry?
 - Why are there 3....?

Hunting for Answers

❖ Get more information

- Measure particles and their interactions in detail
 - Precision measurements (e.g. LHCb)
- Observe new particles or interactions
 - Search in new areas in “phase space”

❖ Find the underlying pattern(s)

- Hypothesize, build models
 - Internally consistent? Consistent with data?
 - Suggestions on where to look

Experiment

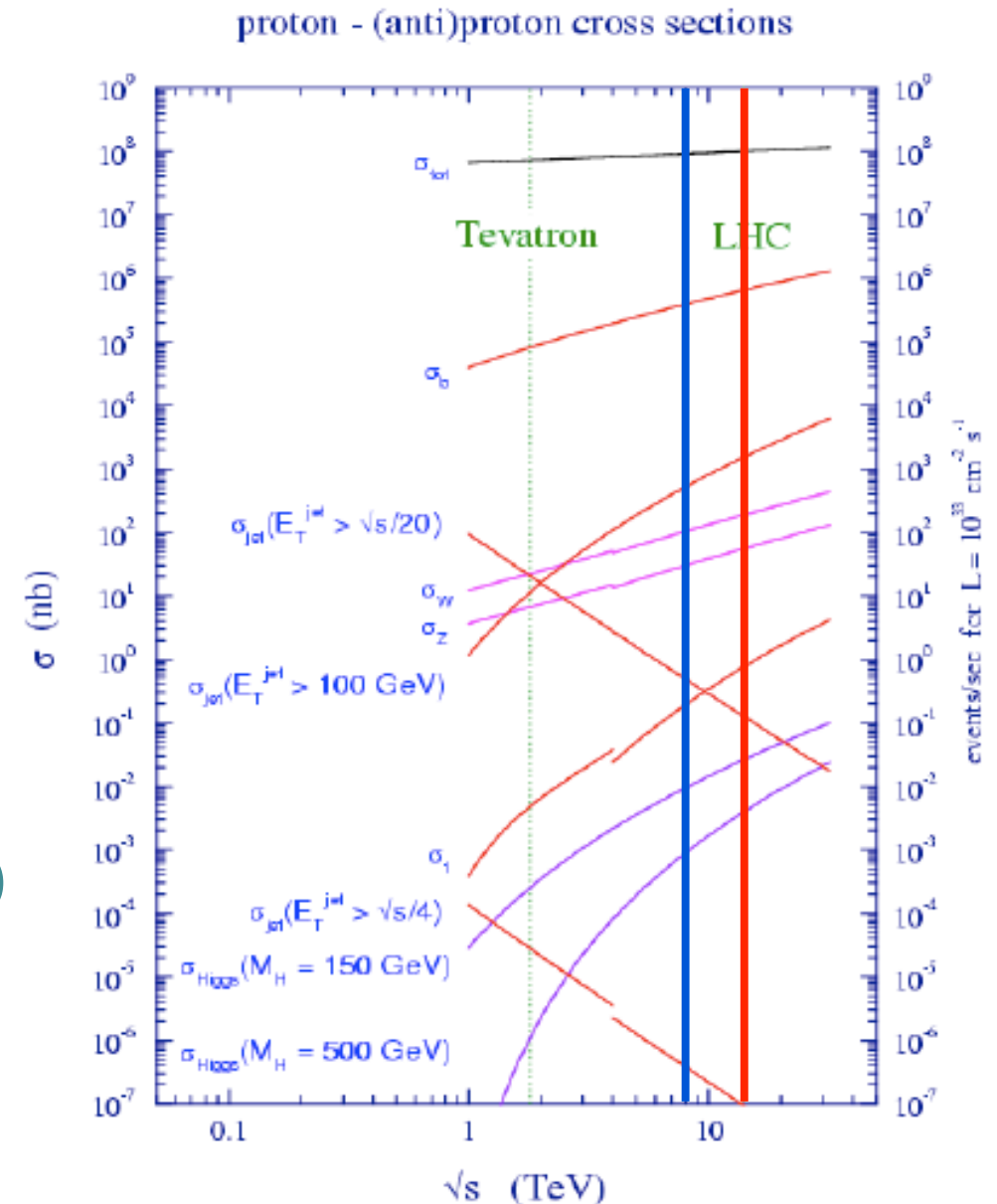
Theory

Where to Start?

- ❖ BSM physics **must** couple to SM (if it helps with the hierarchy problem), but is it
 - Resonant?
 - Does it have new massive particles decaying to electrons, muons, quarks, W, Z,...?
 - “SM-like”?
 - Same but includes some new long-lived particles in the decay chain... (e.g. dark matter candidate)
 - No new “particles” in reach
 - Hidden or too heavy or.... don't exist
 - Are there new interactions?

Physics @ LHC

- ❖ LHC opened a new era:
 - Tevatron was mega-W
 - LHC is
 - Giga-W
 - Giga-Z
 - Top factory (~giga-top)
 - Higgs factory (mega-Higgs)
 - New physics factory?

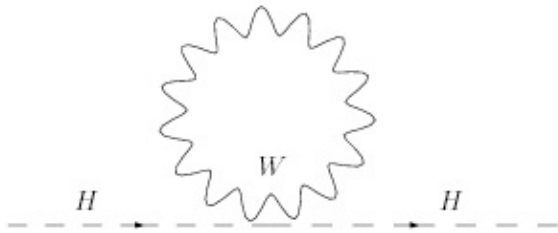


Experimental Searches

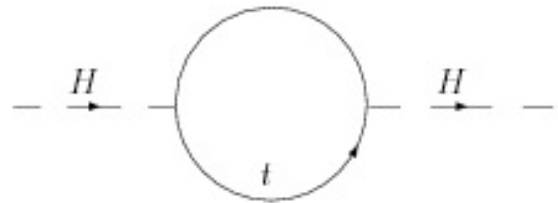
- ❖ By final state, so main questions are
 - Does the new physics produce dark matter?
 - Something we basically know exists and interacts weakly at best with SM
 - ➔ Yes: signatures contain missing transverse energy
 - ➔ No: MET not generic signature
 - Are there new interactions?
 - ➔ No: we know how to calculate everything
 - ➔ Yes: strong (resonances) or very weak (long-lived particles) or...?
- ❖ e.g. SUSY is (Yes,No) if R-parity, technicolor (No,Yes)....

With Dark Matter

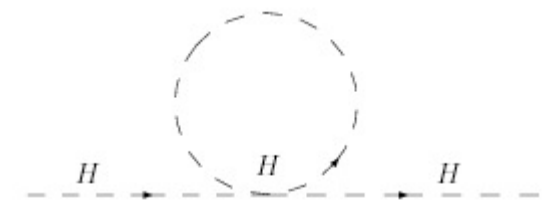
(Super)Symmetry Solution



$$\longrightarrow \frac{1}{16\pi^2} g^2 E^2$$



$$\longrightarrow \frac{3}{16\pi^2} y_t^2 E^2$$



$$\longrightarrow \frac{1}{16\pi^2} \lambda E^2$$

❖ If for every fermion there is a partner boson and vice-versa

- Loops cancel each other

❖ Symmetry cannot be exact (no bosonic electron observed)

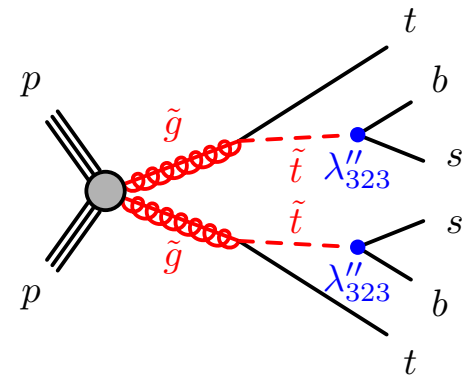
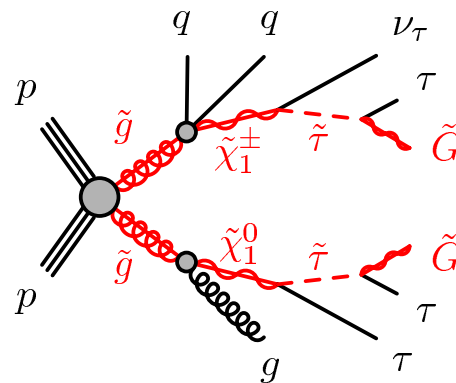
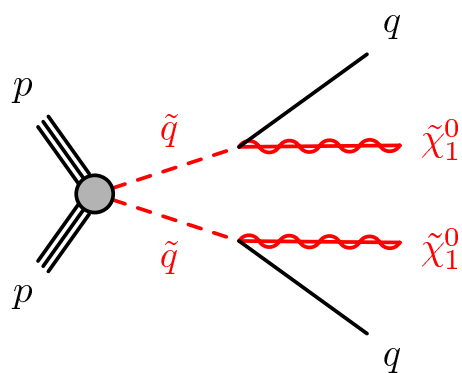
- Symmetry breaking leads to “residual” Higgs mass

❖ This is supersymmetry

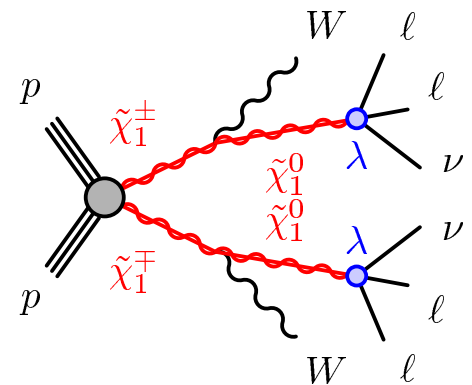
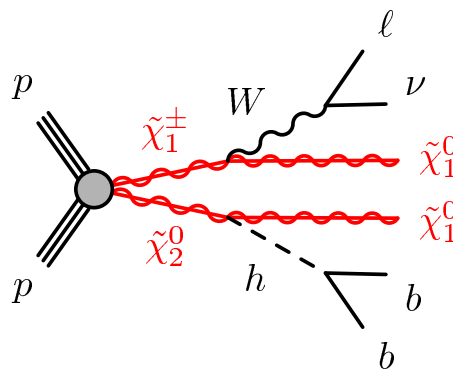
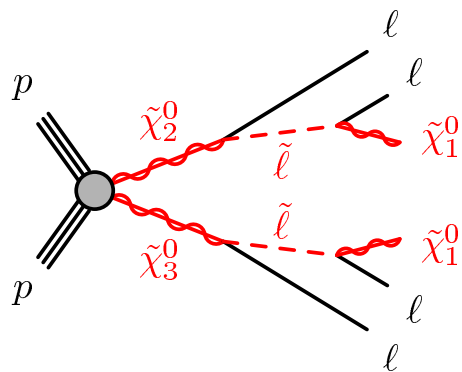
Canonical SUSY

❖ Wide range of signatures

- Strong production... (large cross-section)



- ... or weak production

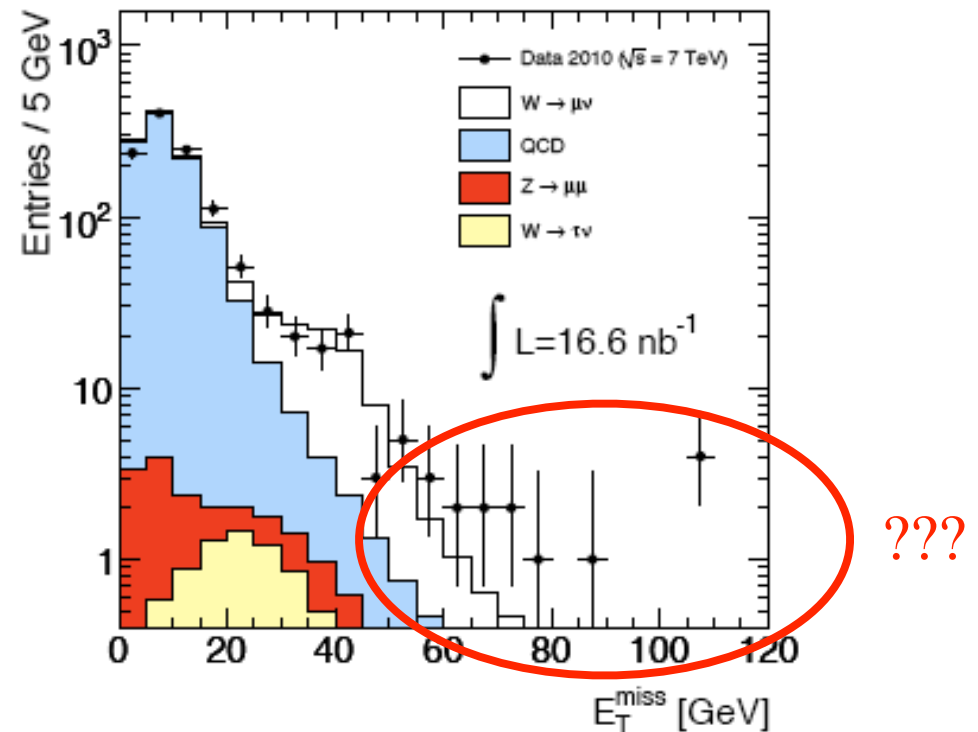
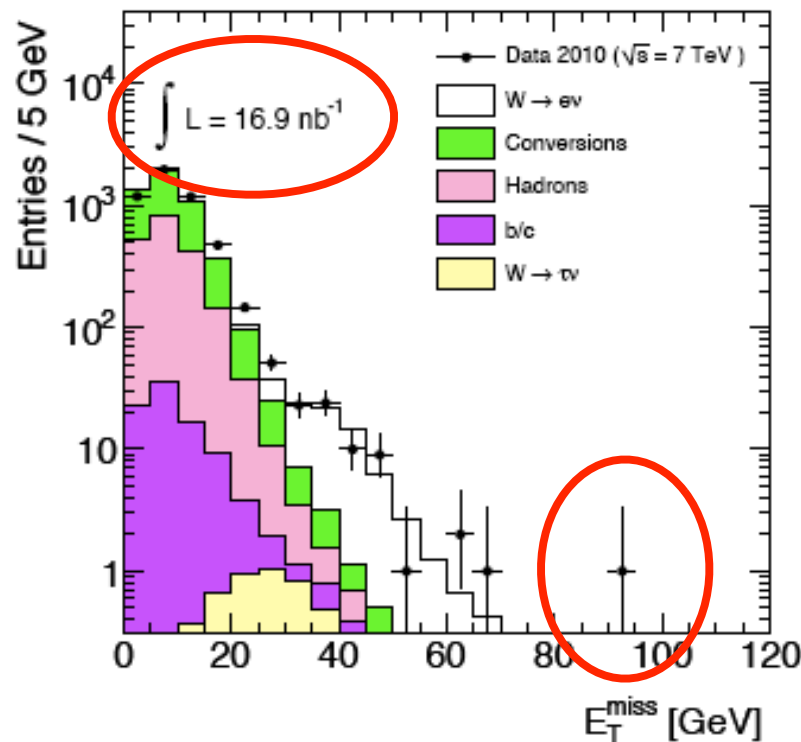


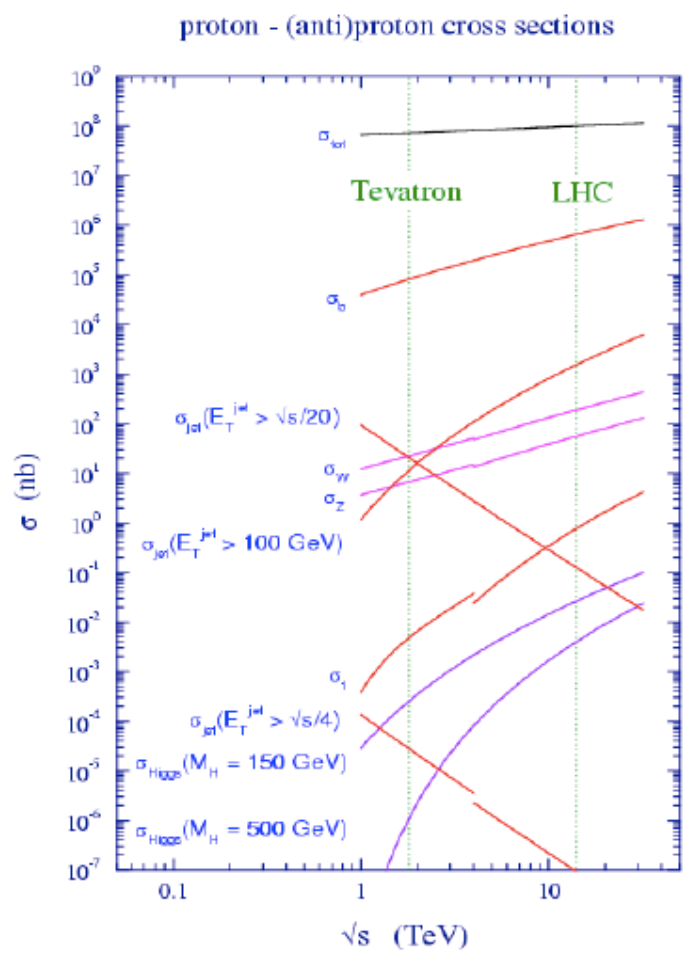
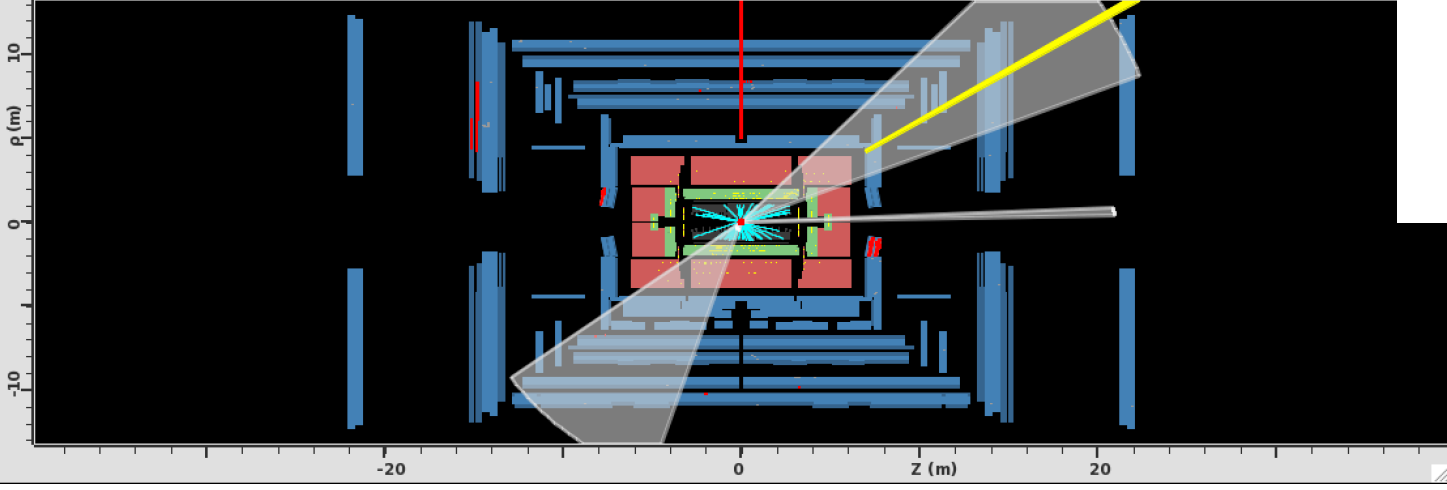
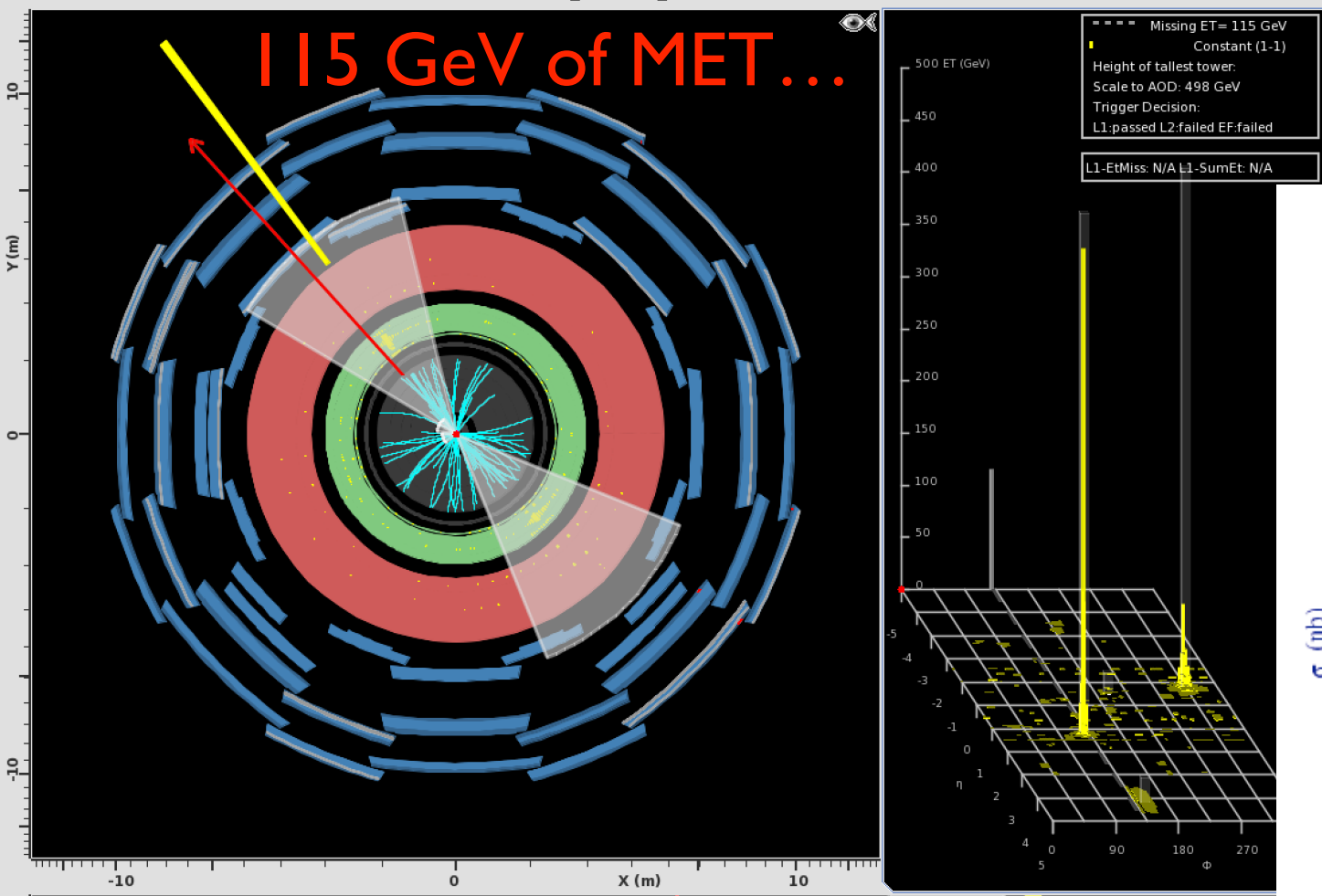
RPV

Missing ET

- ❖ “Evil” variable: - Σ (everything else)
- Need to understand “everything else”
- Good benchmark: leptonic W boson decays

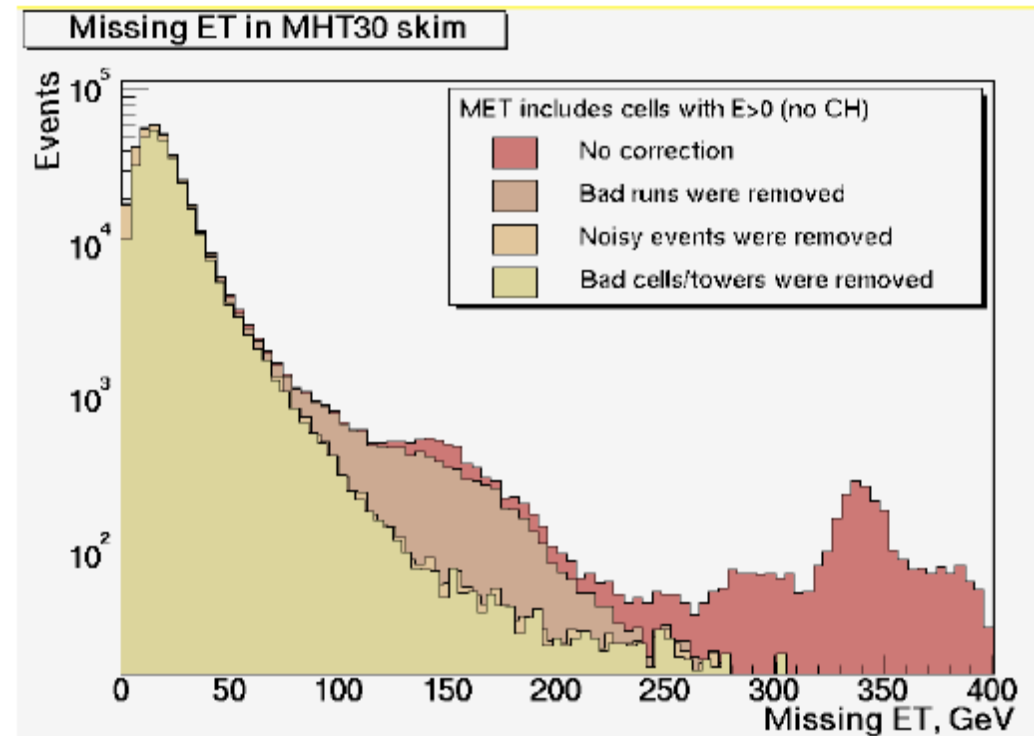
Early 2010



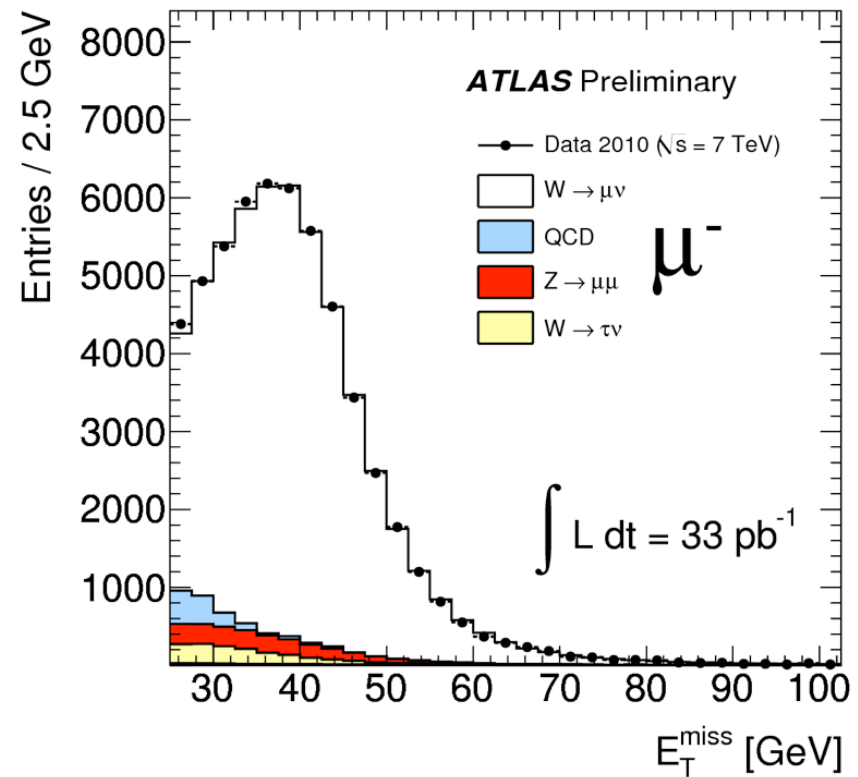
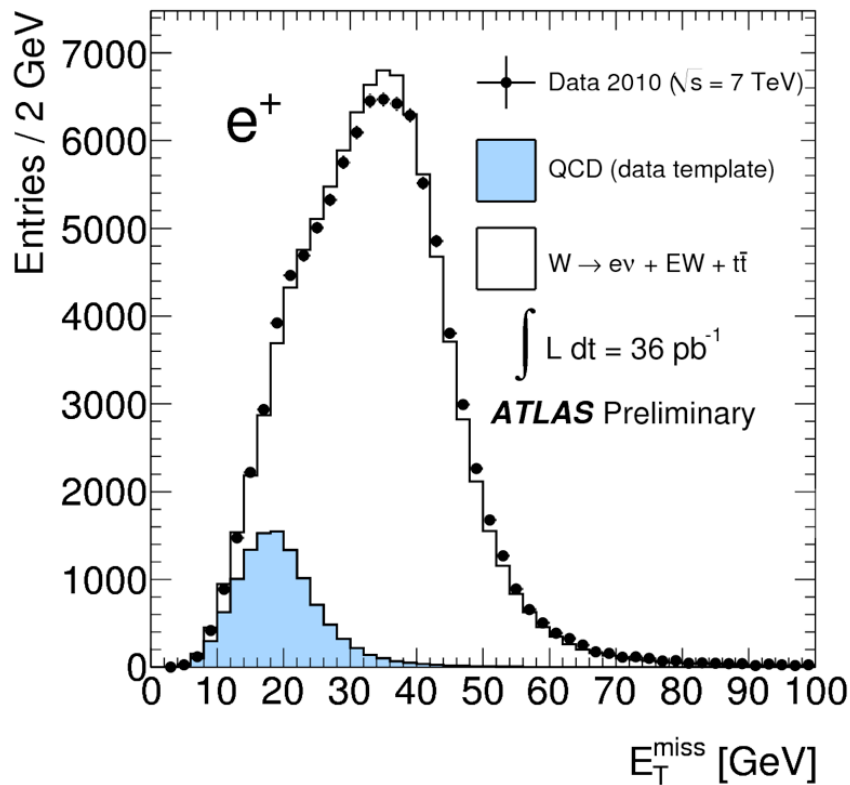


❖ Analyses using MET are particularly sensitive

- Requires the full calorimeter to behave, and calorimeter is generally the most sensitive subdetector (analog, ~ 16 bits)
- Easy: basic DQ (high voltage trip, etc.)
- Hard: low frequency
- Can't spot a 10^{-5} Hz (once a day) effect online or in first pass DQ
- But can be biggest part of dataset after cuts!



❖ With “cleaning”, QCD evaluated from data,...

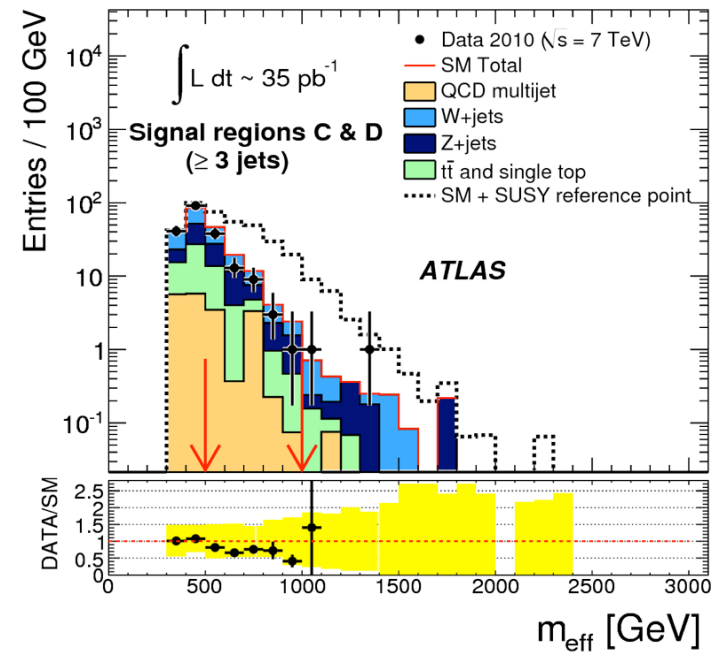
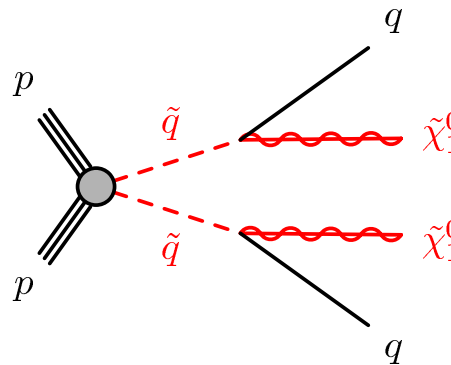
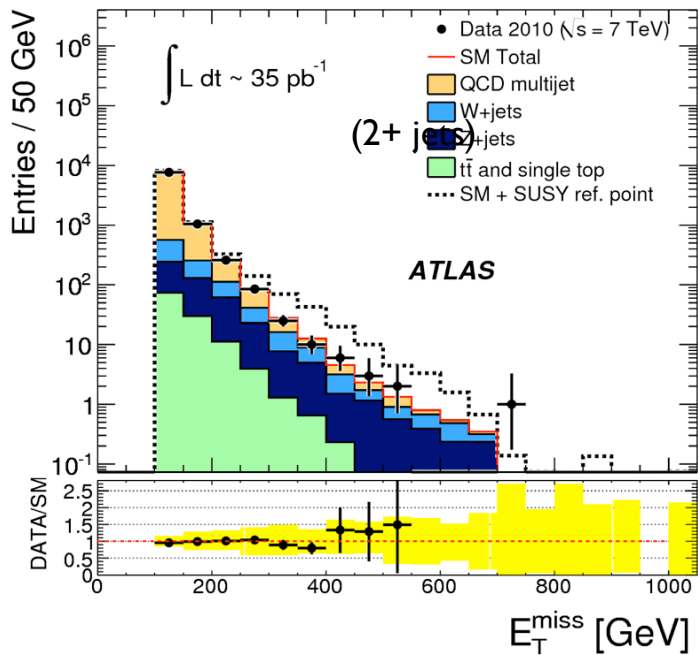


❖ Already ~200k clean $W \rightarrow \ell\nu$ events in 2010

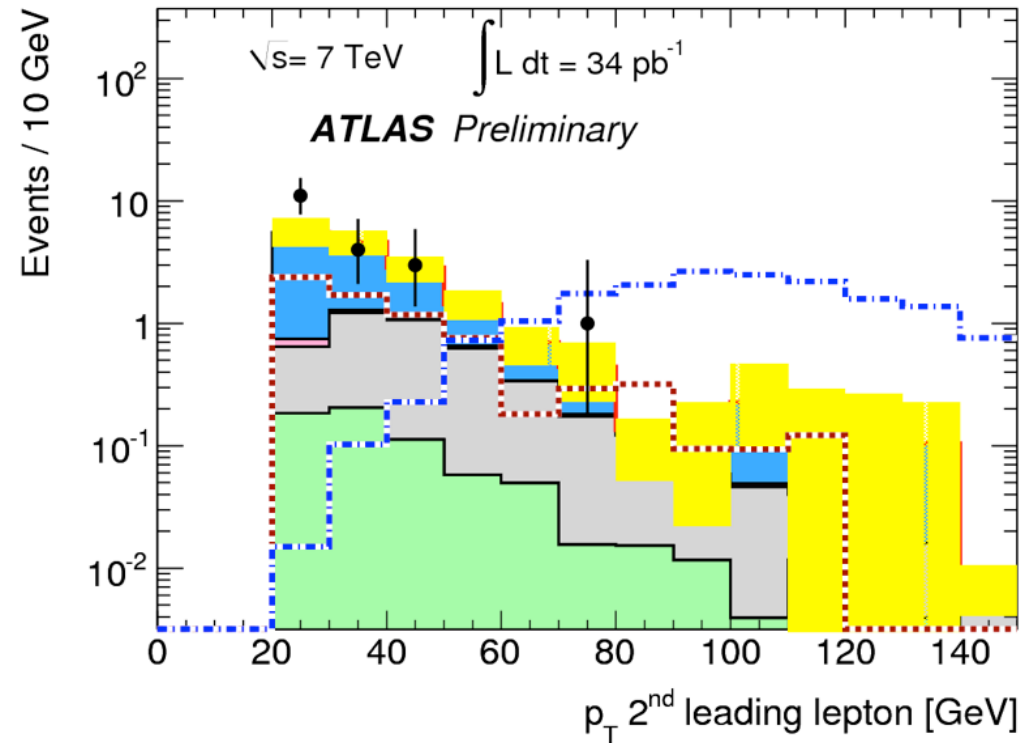
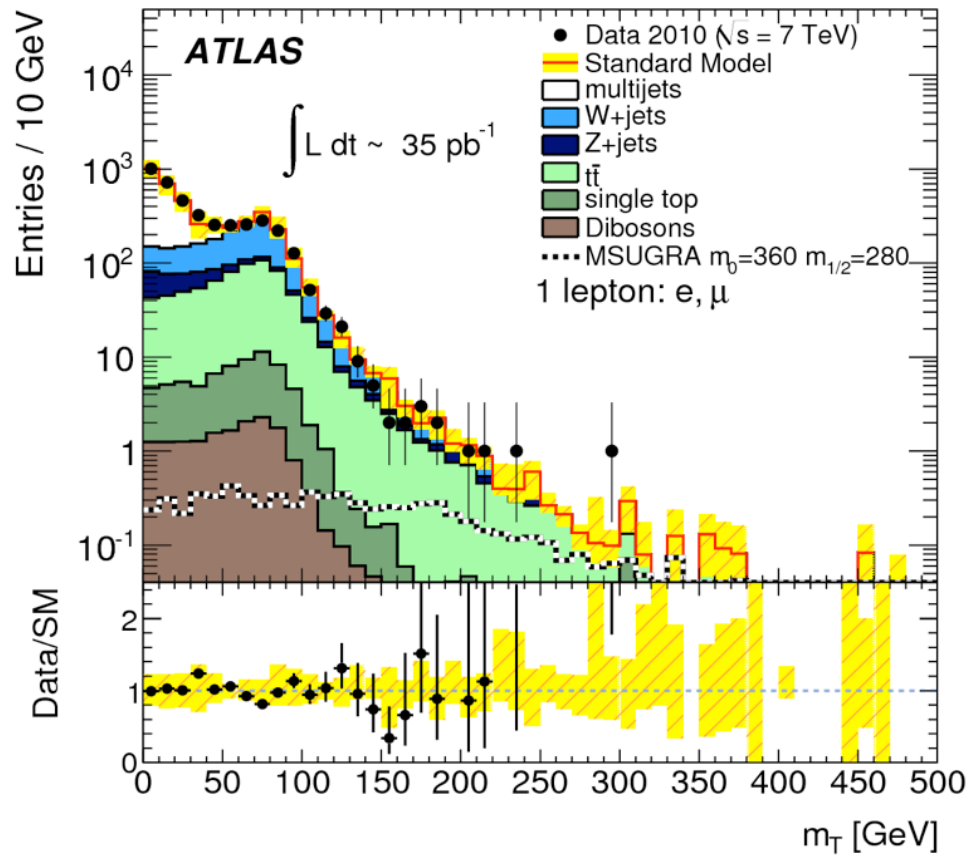
- Almost a billion now

SUSY as a Benchmark

- ❖ Hadron collider \Rightarrow produce squarks and gluinos decaying to jets + MET
 - Optimize jet p_T & MET cuts for different scenarios, since gluinos produce more jets than squarks
 - Use M_{eff} to discriminate, measure of event Q^2



❖ Leptons in decay chains....



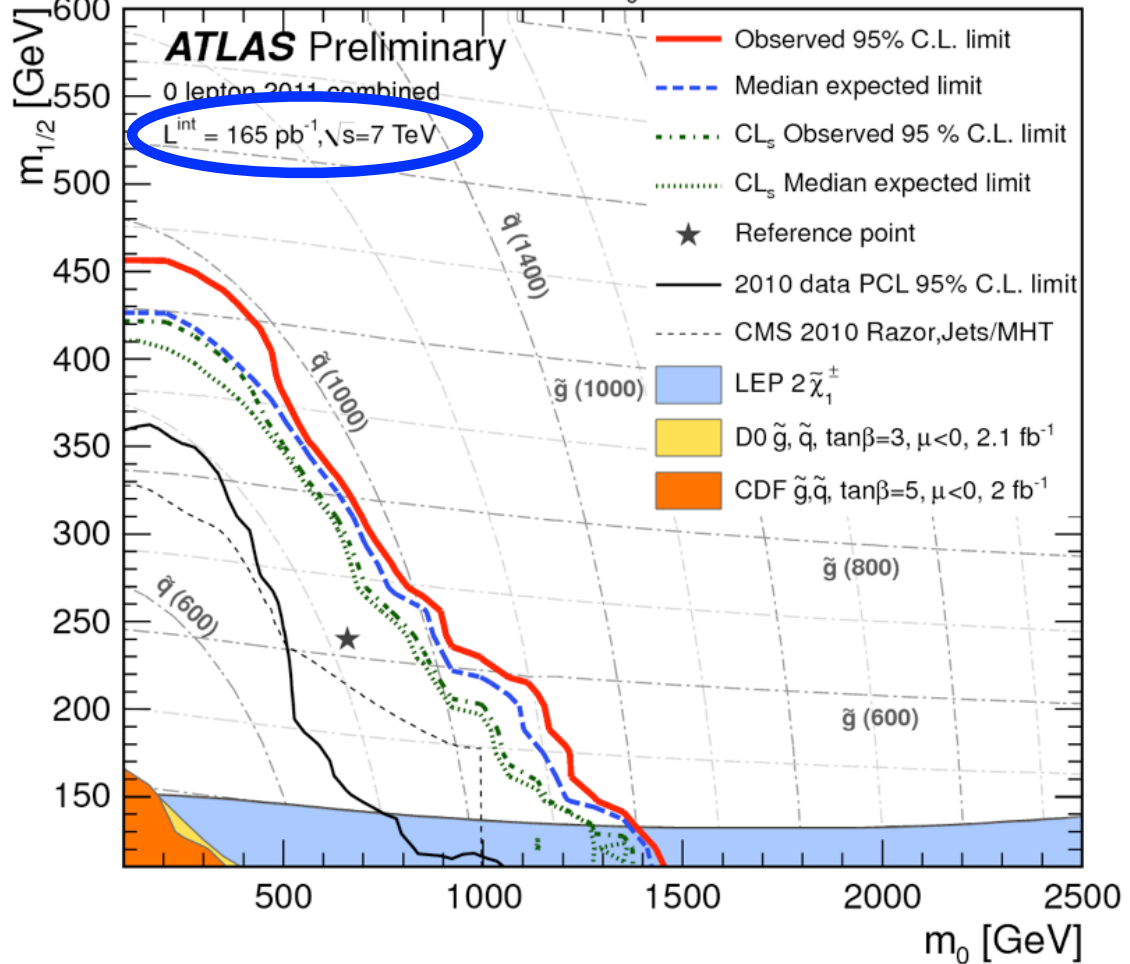
All Praise COM Energy!

ATLAS SUSY Searches* - 95% CL Lower Limits
Status: Feb 2015

Model	$\epsilon, \mu, \tau, \gamma$	Jets	E_{T}^{miss}	$[\mathcal{L} d\Omega(\text{fb}^{-1})]$	Mass limit	Reference	
Inclusive Searches	MSUGRA/CMSSM	0-2 jets	Yes	20.3	\tilde{g}, \tilde{q}	1.7 TeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-2 jets	Yes	20.3	\tilde{g}, \tilde{q}	250 GeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$ (compressed)	1 γ	0-1 jet	Yes	20.3	\tilde{g}, \tilde{q}	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-2 jets	Yes	20.3	\tilde{g}, \tilde{q}	1.33 TeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q} + \gamma$	0-2 jets	Yes	20	\tilde{g}, \tilde{q}	1.2 TeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q} + \gamma + \gamma$	0-3 jets	Yes	20	\tilde{g}, \tilde{q}	1.32 TeV	
	GMSB (NLSIP)	$1.2 + 0.1 \ell$	0-2 jets	Yes	20.3	\tilde{g}, \tilde{q}	1.6 TeV
	GGM (bino NLSIP)	2γ	-	Yes	20.3	\tilde{g}, \tilde{q}	1.28 TeV
	GGM (stau NLSIP)	$1 e, \mu + \gamma$	-	Yes	4.8	\tilde{g}, \tilde{q}	619 GeV
	GGM (higgsino-bino NLSIP)	$1 b$	Yes	4.8	\tilde{g}, \tilde{q}	900 GeV	
3 ν gen. squarks direct production	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-3 jets	Yes	20.1	\tilde{g}, \tilde{q}	1.25 TeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}, \tilde{q}	1.1 TeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}, \tilde{q}	1.34 TeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-1 e, μ	3 b	Yes	20.1	\tilde{g}, \tilde{q}	1.3 TeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-2 jets	Yes	20.1	\tilde{g}, \tilde{q}	100-620 GeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{g}, \tilde{q}	275-440 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	1 e, μ	1-2 b	Yes	4.7	\tilde{g}, \tilde{q}	290-560 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-1 e, μ	1-2 b	Yes	20.3	\tilde{g}, \tilde{q}	90-191 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-1 e, μ	1-2 b	Yes	20	\tilde{g}, \tilde{q}	215-530 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0 mono-jet+tag	20.3	\tilde{g}, \tilde{q}	90-240 GeV		
EW direct	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	3 e, μ (Z)	1 b	Yes	20.3	\tilde{g}, \tilde{q}	150-580 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	3 e, μ (Z)	1 b	Yes	20.3	\tilde{g}, \tilde{q}	290-800 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2 e, μ	0	Yes	20.3	\tilde{g}, \tilde{q}	90-325 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2 e, μ	0	Yes	20.3	\tilde{g}, \tilde{q}	140-465 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2 e, μ	0	Yes	20.3	\tilde{g}, \tilde{q}	100-350 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	3 e, μ	0	Yes	20.3	\tilde{g}, \tilde{q}	700 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2-3 e, μ	0-2 jets	Yes	20.3	\tilde{g}, \tilde{q}	420 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2 e, μ	0-2 b	Yes	20.3	\tilde{g}, \tilde{q}	250 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	4 e, μ	0	Yes	20.3	\tilde{g}, \tilde{q}	620 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	1 $\mu, \text{dipl. vtx.}$	-	Yes	20.3	\tilde{g}, \tilde{q}	435 GeV
Long-lived particles	Direct $\tilde{t} \rightarrow t\tilde{q}$ prod. long-lived \tilde{t}	Disapp. trk	1 jet	Yes	20.3	\tilde{t}	278 GeV
	Stable, stopped \tilde{t} R-hadron	0-1 jets	Yes	27.9	\tilde{t}	832 GeV	
	Stable \tilde{t} R-hadron	trk	-	19.1	\tilde{t}	1.27 TeV	
	GMSB, stable \tilde{t} , $\tilde{t} \rightarrow t\tilde{q}, \mu \rightarrow e\tilde{q}$	1 μ	-	18.1	\tilde{t}	537 GeV	
	GMSB, $\tilde{t} \rightarrow t\tilde{q}$, long-lived \tilde{t}	2 γ	-	Yes	20.3	\tilde{t}	435 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	1 $\mu, \text{dipl. vtx.}$	-	Yes	20.3	\tilde{g}, \tilde{q}	1.0 TeV
	LFV $pp \rightarrow \tilde{q} + X, \tilde{q} \rightarrow q + \mu$	2 e, μ	-	4.6	\tilde{q}	1.81 TeV	
	LFV $pp \rightarrow \tilde{q} + X, \tilde{q} \rightarrow q + \tau$	1 $e, \mu + \tau$	-	4.6	\tilde{q}	1.1 TeV	
	Binneer RPV CMSSM	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{g}, \tilde{q}	1.35 TeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	3 $e, \mu + \tau$	-	Yes	20.3	\tilde{g}, \tilde{q}	750 GeV
RPV	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	3 $e, \mu + \tau$	-	Yes	20.3	\tilde{g}, \tilde{q}	450 GeV
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	0-3 jets	Yes	20.3	\tilde{g}, \tilde{q}	916 GeV	
	$\tilde{g}, \tilde{q} \rightarrow q\tilde{q}$	2 e, μ (SS)	0-3 b	Yes	20.3	\tilde{g}, \tilde{q}	850 GeV
	Scalar charm, $\tilde{t} \rightarrow c\tilde{q}$	0	2 e	Yes	20.3	\tilde{t}	200 GeV

*Only a selection of the available mass limits on new states or phenomena is shown. All limits quoted are observed minus 1 σ theoretical signal cross section uncertainty.

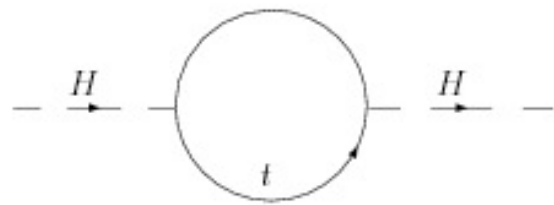
MSUGRA/CMSSM: $\tan\beta = 10, A_0 = 0, \mu > 0$



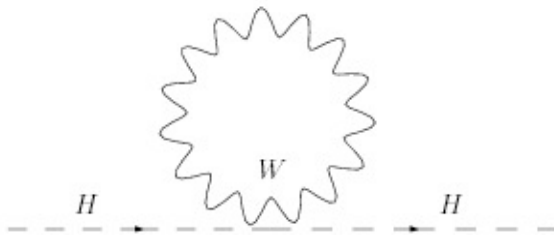
Tevatron blown away... 8 (2016) hours of LHC data

But...

We've Found a Higgs!



$$\longrightarrow \frac{3}{16\pi^2} y_t^2 E^2$$



$$\longrightarrow \frac{1}{16\pi^2} g^2 E^2$$



$$\longrightarrow \frac{1}{16\pi^2} \lambda E^2$$

❖ If new scale, these go to the new scale...

❖ To ~cancel these, need to primarily compensate for

- Top

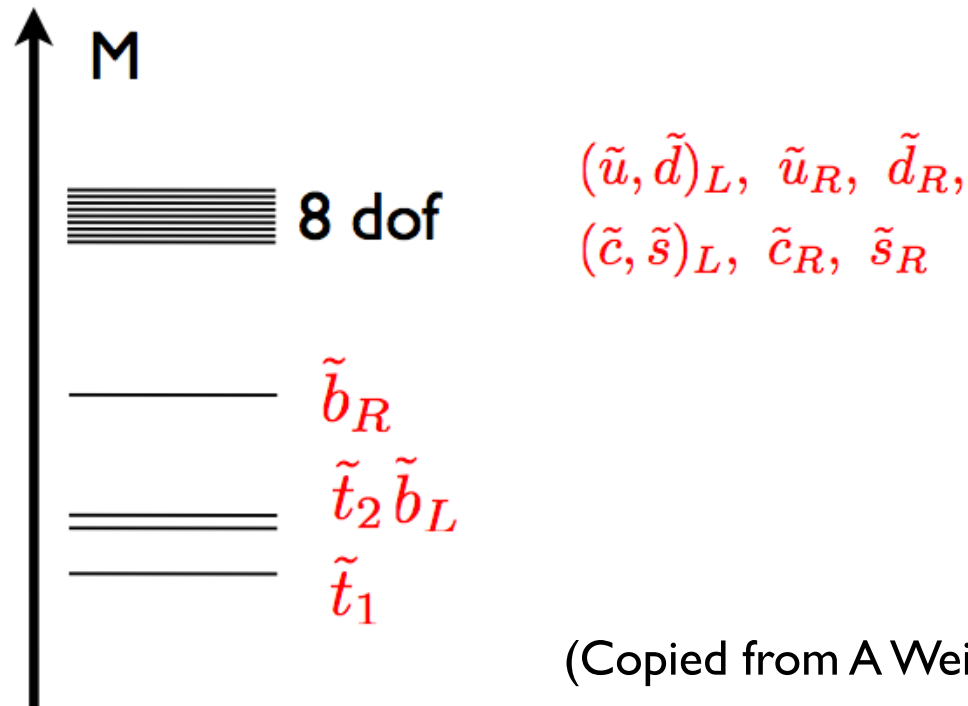
- W/Z

- H

➔ **Discovery of the light Higgs refocuses new physics search**

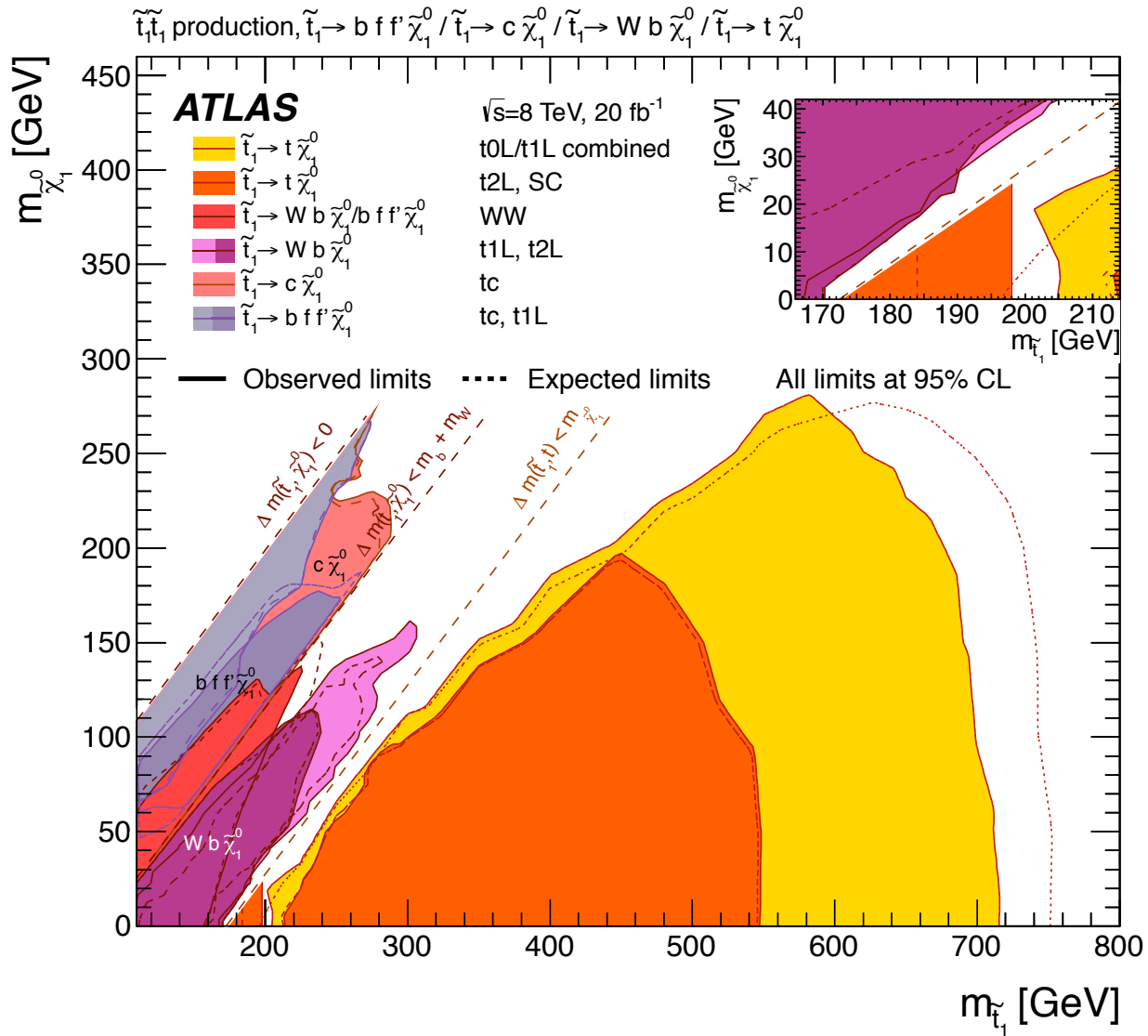
SUSY and the Higgs

- ❖ For SUSY, 125 GeV is rather heavy!
 - Need light higgsinos, stops, sbottoms... but heavy “light” squarks \Rightarrow “natural SUSY”
 - Stop at the forefront!

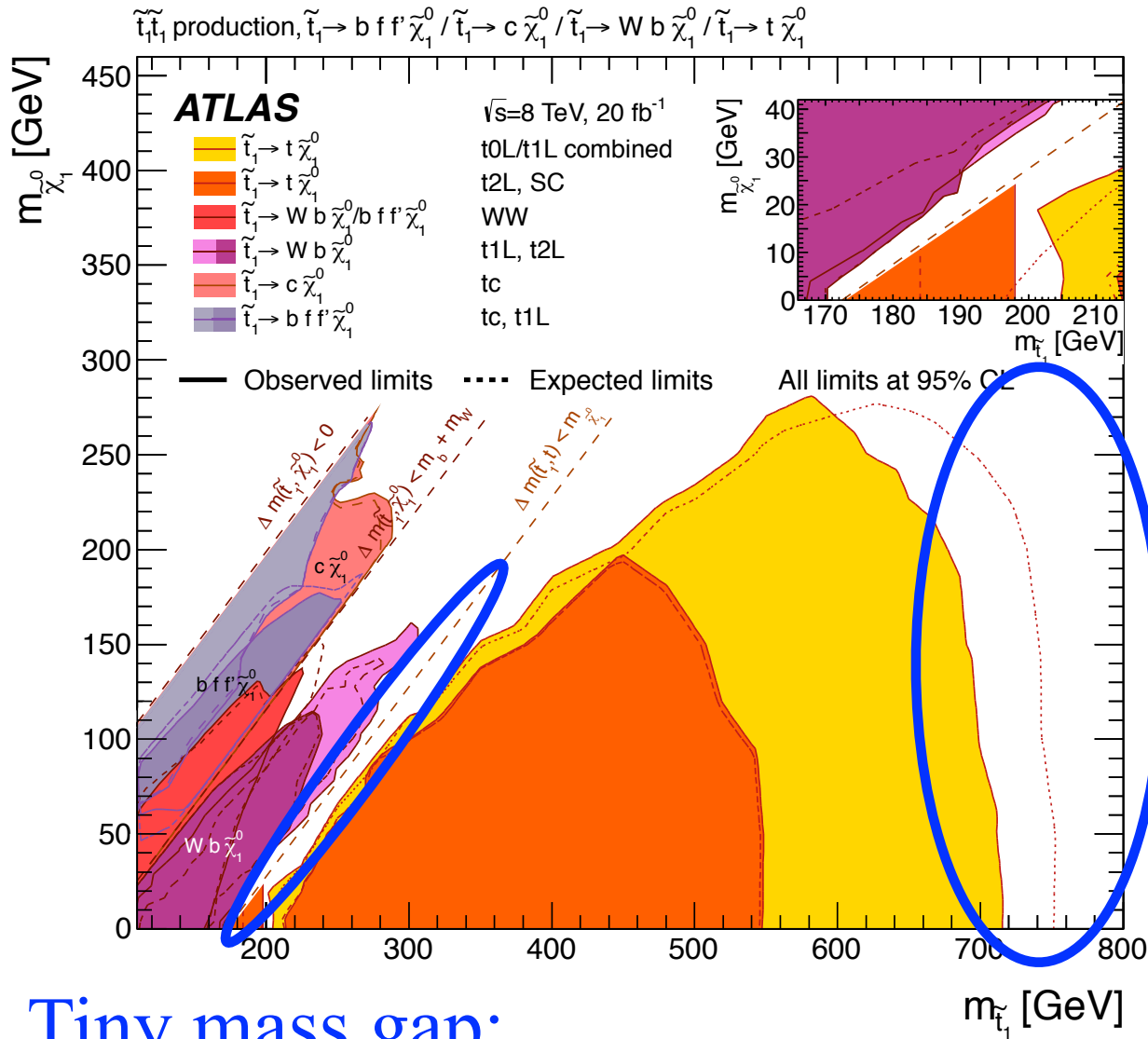


(Copied from A Weiler)

Stop Searching Anatomy



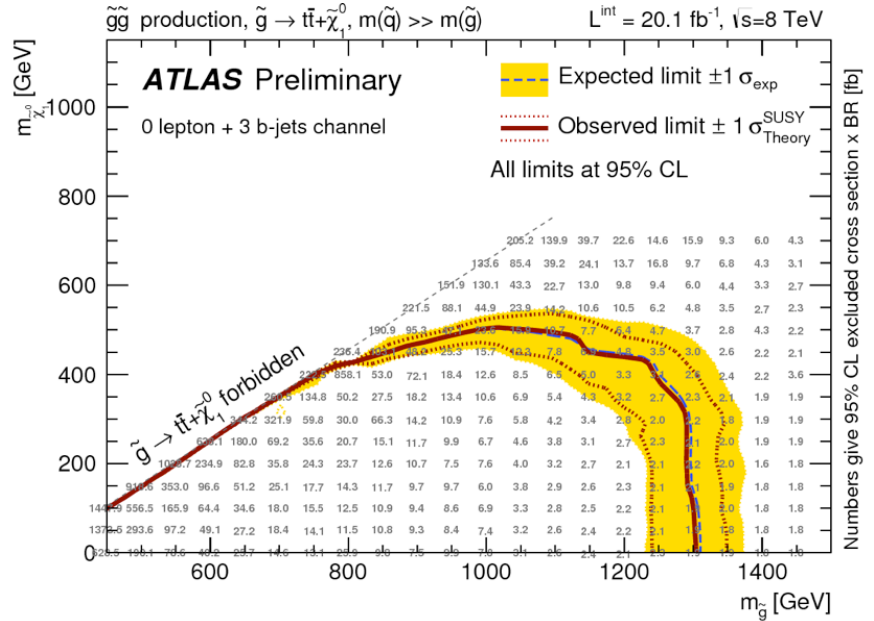
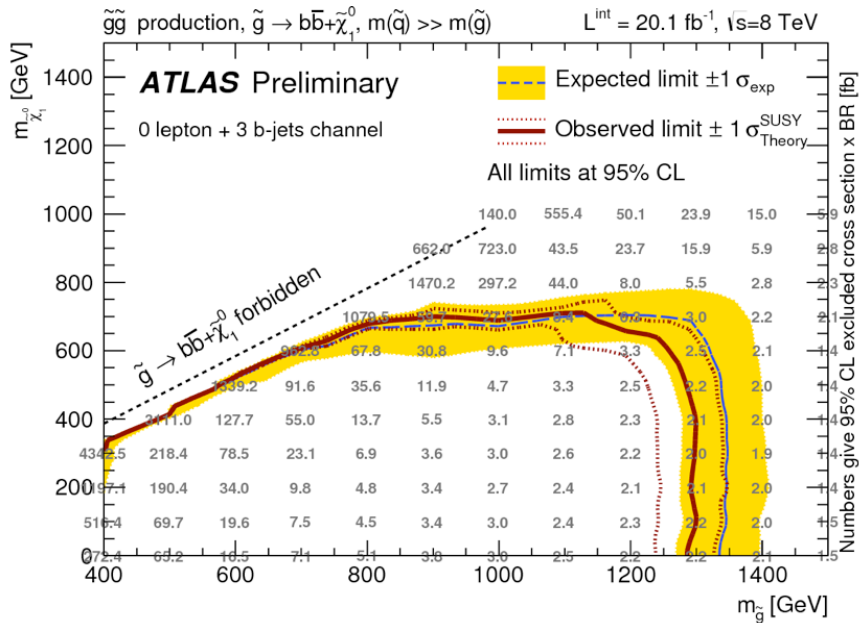
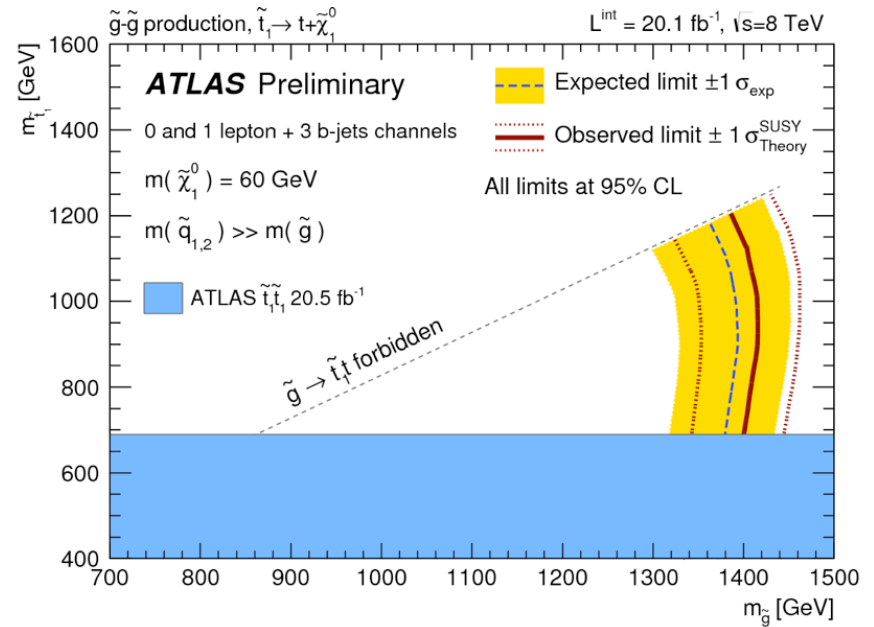
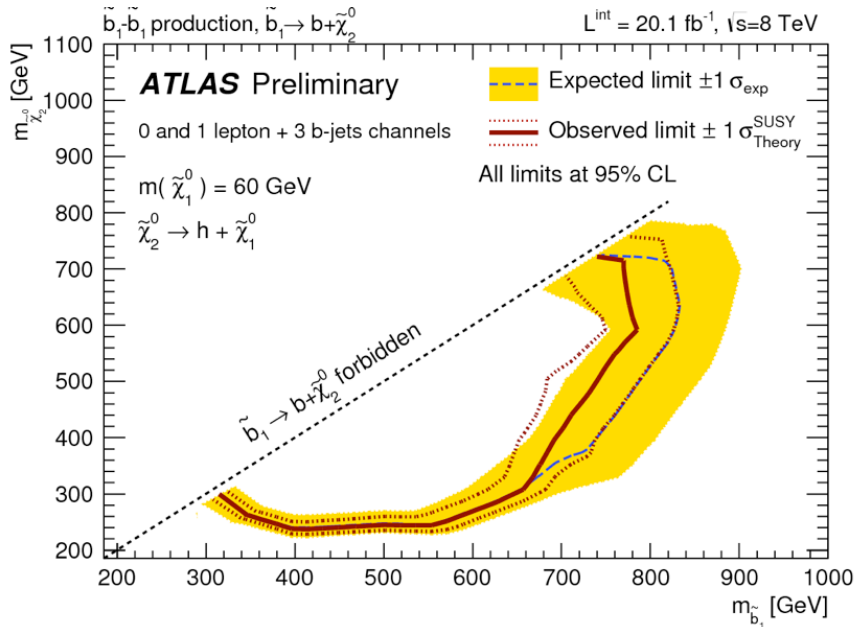
Stop Searching Anatomy



High mass:
run out of
cross-section

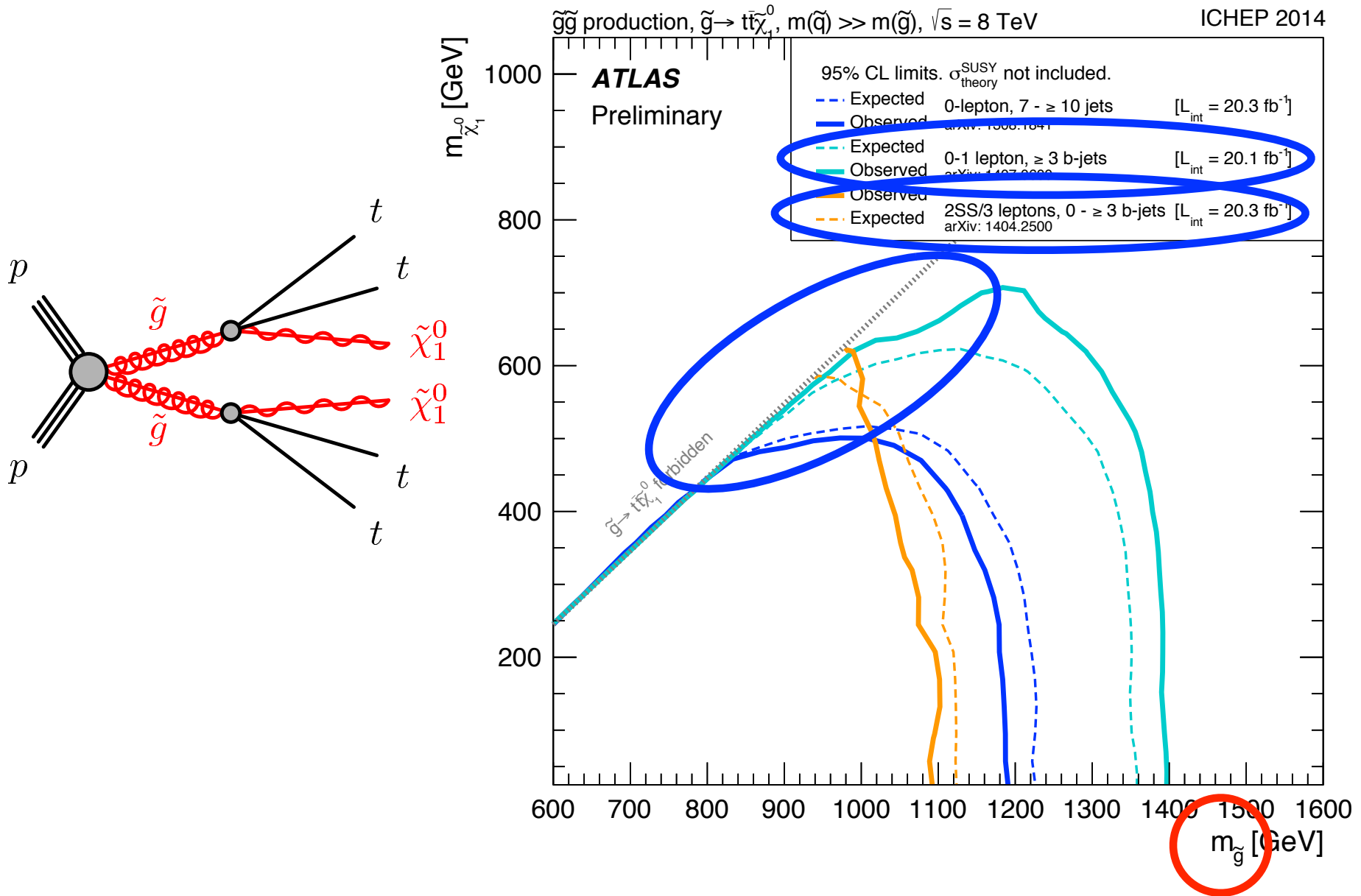
Tiny mass gap:
soft decay products

Many Many Limits... Sigh



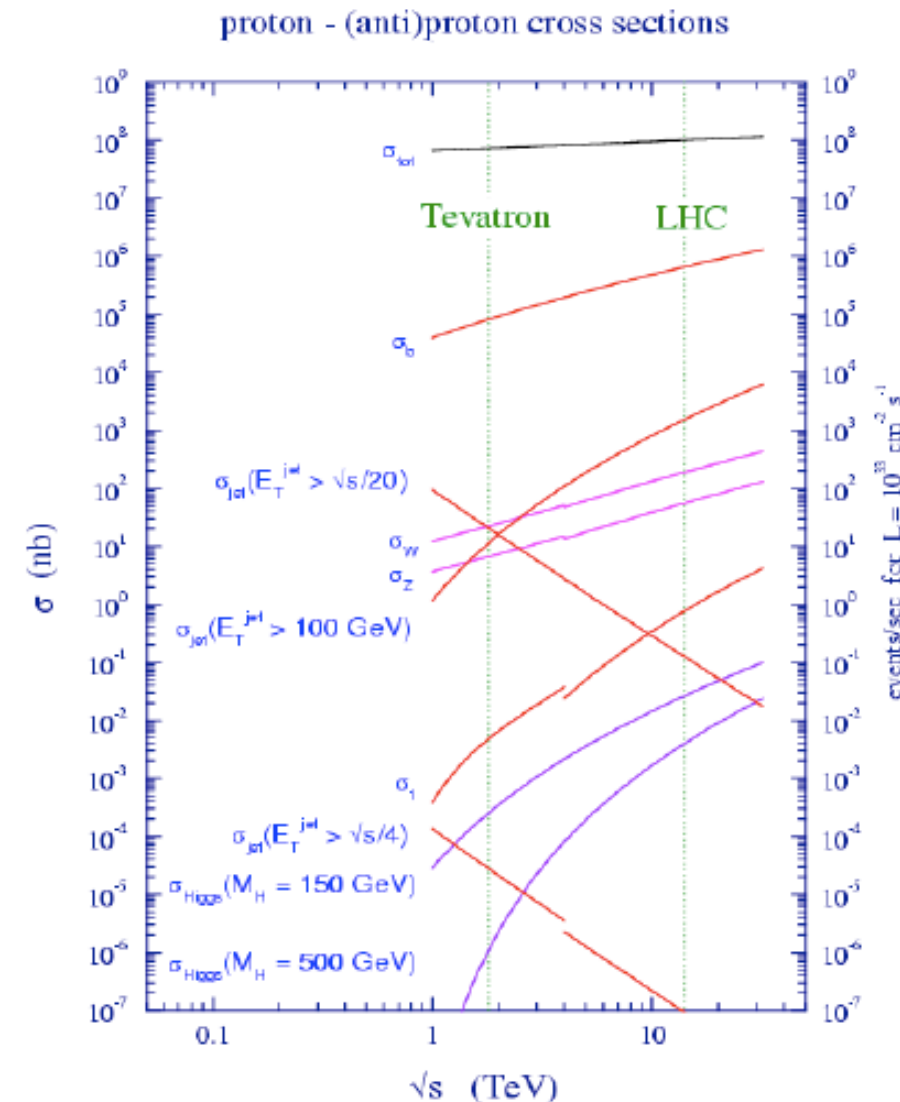
Stop Searching Anatomy

(Off-shell intermediate stop)

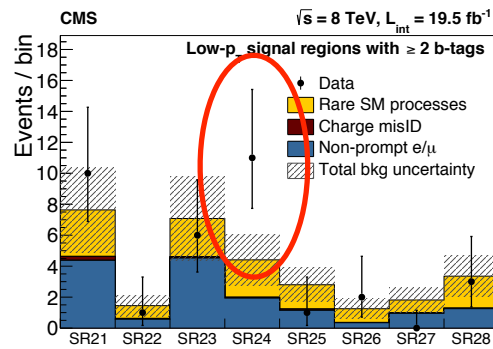
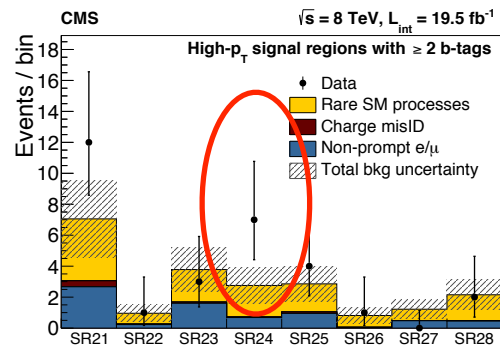


Same-Sign Leptons

- ❖ At hadron colliders, leptons signify something interesting happened
 - E.g. Z production
- ❖ Same-sign leptons even more interesting? Lower background?
 - $W^\pm W^\pm$
 - but also B/D meson oscillations
 - mostly low p_T
 - and wrong charge measurement
- ❖ With lower background, access to smaller cross-sections, smaller mass gaps
 - At the cost of small branching ratio



Same Sign Lepton Excesses

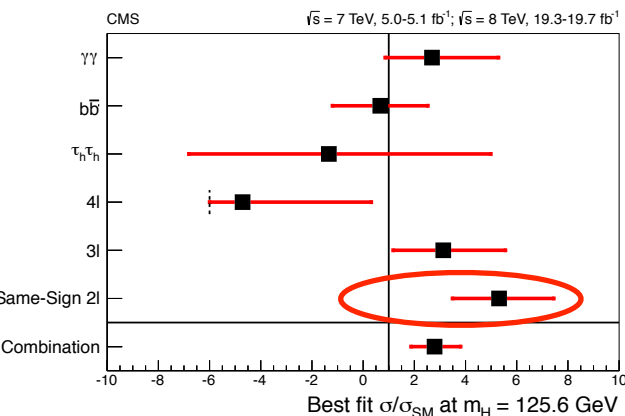


SR1b _{1bin}	Total	ee	eμ	μμ
Observed events	10	6	4	0
Total expected background events	4.7 ± 2.1	1.4 ± 0.8	2.1 ± 1.1	1.2 ± 0.4
Components of the background				
<i>t</i> tV, <i>t</i> tH, <i>t</i> Z and <i>t</i> t <i>t</i>	2.5 ± 1.7	0.6 ± 0.3	1.2 ± 1.0	0.7 ± 0.3
Dibosons and tribosons	0.9 ± 0.4	0.10 ± 0.04	0.3 ± 0.1	0.5 ± 0.3
Fake leptons	$0.8^{+1.2}_{-0.8}$	$0.4^{+0.7}_{-0.4}$	$0.4^{+0.5}_{-0.4}$	< 0.1
Charge-flip electrons	0.5 ± 0.1	0.3 ± 0.1	0.3 ± 0.1	–
$p(s=0)$	0.07	0.01	0.18	0.50

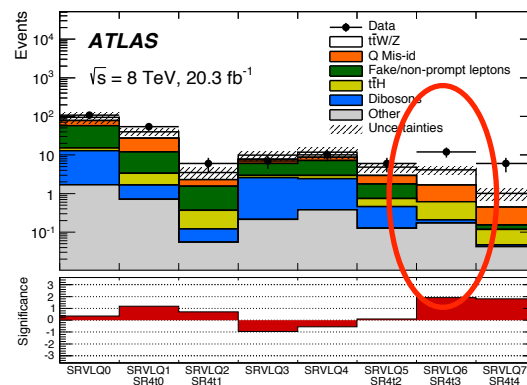
CMS (SUSY), <http://arxiv.org/abs/1311.6736>

(24 signal regions in paper)

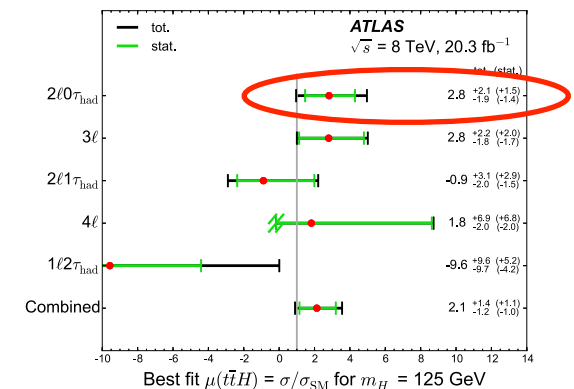
ATLAS (SUSY), <http://arxiv.org/abs/1404.2500> (5 signal regions in paper)



CMS (ttH), <http://arxiv.org/abs/1408.1682>



ATLAS (TT), <http://arxiv.org/abs/1504.04605>



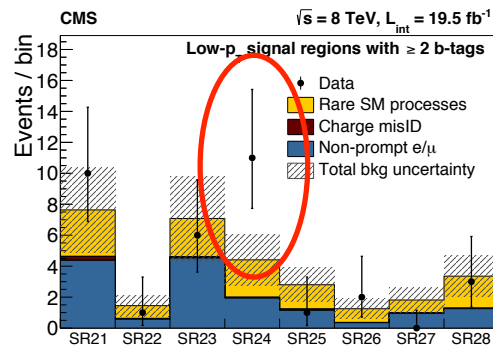
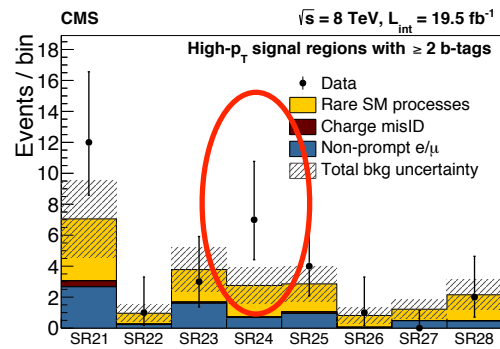
ATLAS (ttH), <http://arxiv.org/abs/1506.05988>

It certainly looks like multiple analyses looking at same sign leptons and b-jets see excesses!

Could it be SUSY? E.g. $\tilde{t}_R \rightarrow t + \tilde{B} \rightarrow t + (\tilde{W}^\pm + W^\mp)$

Huang et al, <http://arxiv.org/abs/1507.01601>

Same Sign Lepton Excesses

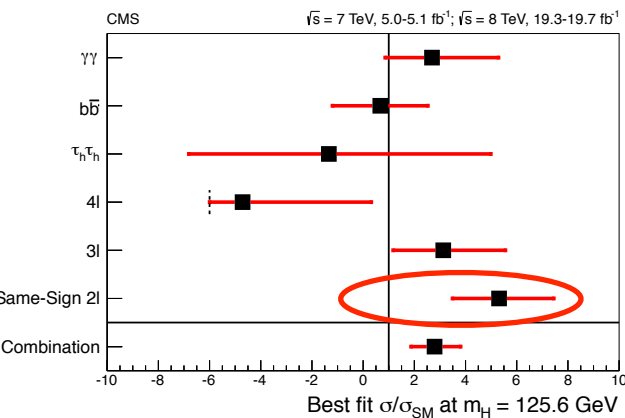


SR1b 1_{bin}	Total	ee	$e\mu$	$\mu\mu$
Observed events	10	6	4	0
Total expected background events	4.7 ± 2.1	1.4 ± 0.8	2.1 ± 1.1	1.2 ± 0.4
Components of the background				
$t\bar{t}V$, $t\bar{t}H$, tZ and $t\bar{t}t$	2.5 ± 1.7	0.6 ± 0.3	1.2 ± 1.0	0.7 ± 0.3
Dibosons and tribosons	0.9 ± 0.4	0.10 ± 0.04	0.3 ± 0.1	0.5 ± 0.3
Fake leptons	$0.8^{+1.2}_{-0.8}$	$0.4^{+0.7}_{-0.4}$	$0.4^{+0.5}_{-0.4}$	< 0.1
Charge-flip electrons	0.5 ± 0.1	0.3 ± 0.1	0.3 ± 0.1	–
$p(s=0)$	0.07	0.01	0.18	0.50

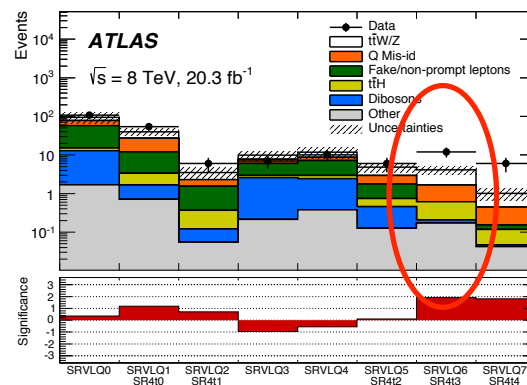
CMS (SUSY), <http://arxiv.org/abs/1311.6736>

(24 signal regions in paper)

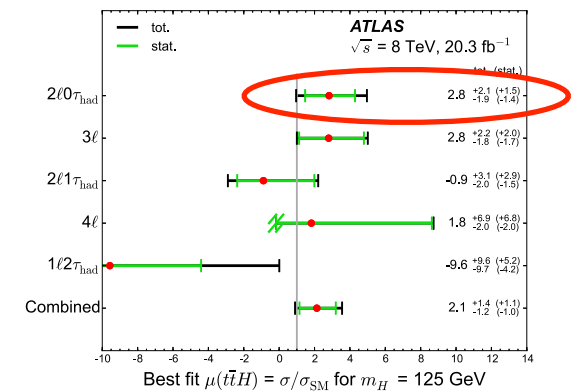
ATLAS (SUSY), <http://arxiv.org/abs/1404.2500> (5 signal regions in paper)



CMS (ttH), <http://arxiv.org/abs/1408.1682>



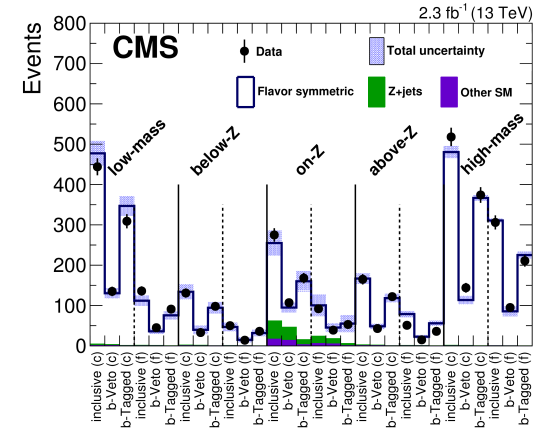
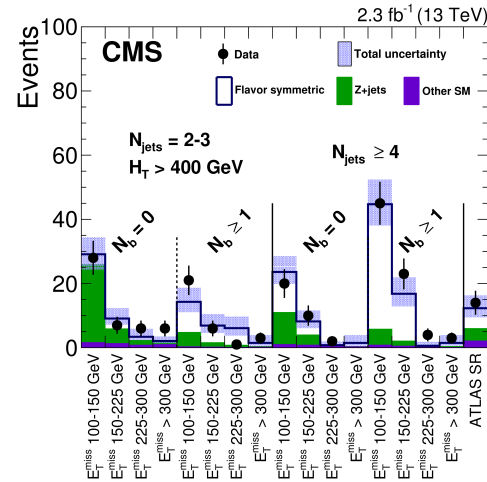
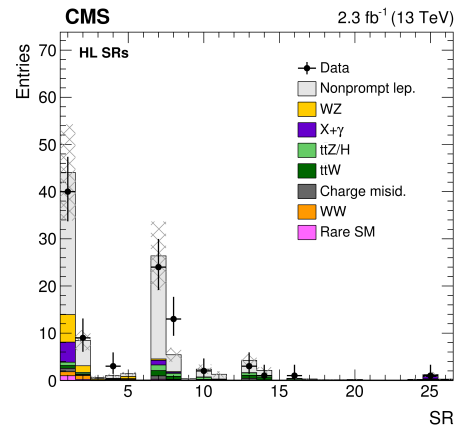
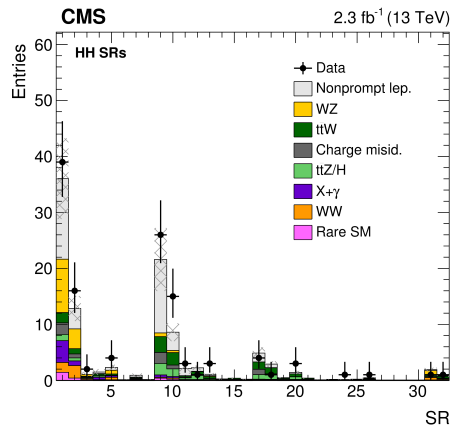
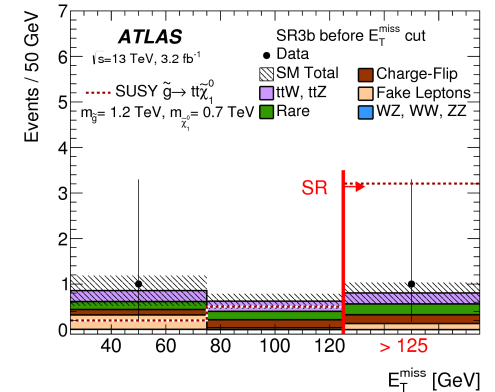
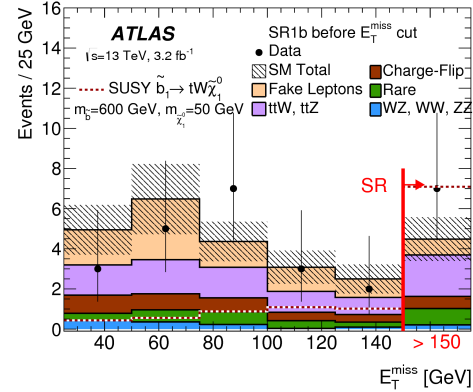
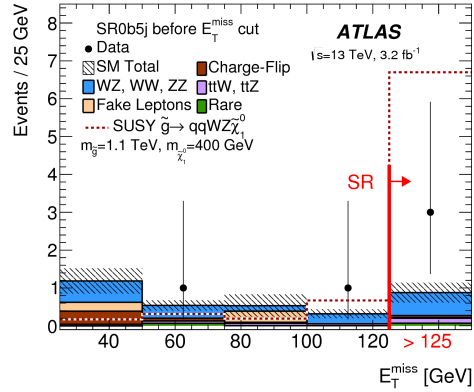
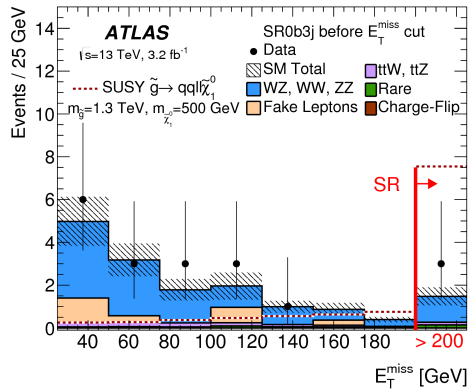
ATLAS (TT), <http://arxiv.org/abs/1504.04605>



ATLAS (ttH), <http://arxiv.org/abs/1506.05988>

The ATLAS analyses are correlated, and same for CMS
 So, ~2 analyses and excesses are $< 3 \sigma$
 Worth keeping an eye on? Sure.

13 TeV

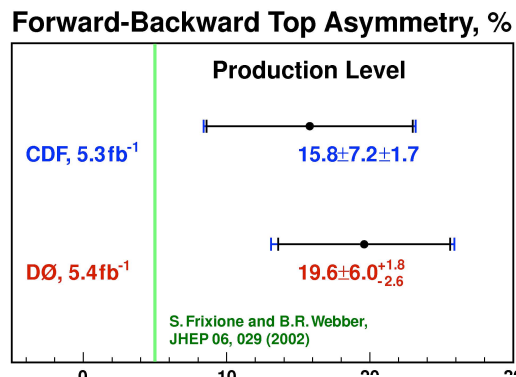


Not much there.... so far

Anecdotes From the Field (II)

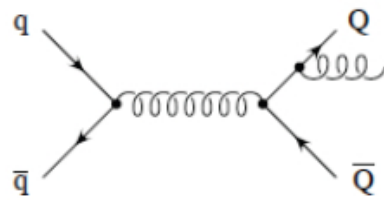
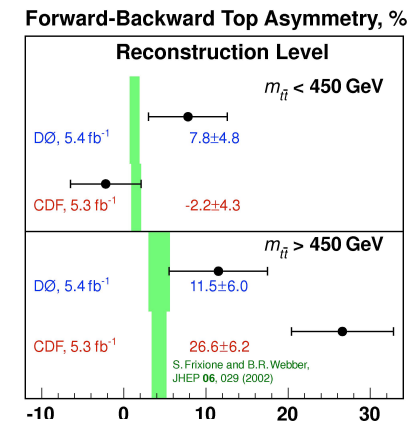
❖ $t\bar{t}$ charge asymmetry at the Tevatron

- At Feynman diagram level, NLO effect (Tevatron is proton-anti-proton collider)

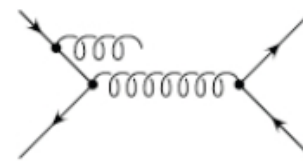


<http://arxiv.org/abs/1107.4995>

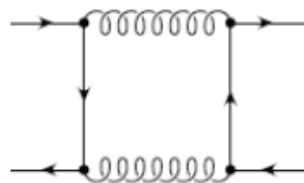
Ca. 2010, big fuss:
much larger than SM!



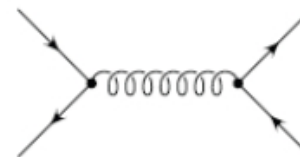
(a) Final state gluon bremsstrahlung FSR



(b) Initial state gluon bremsstrahlung ISR



(c) double virtual gluon exchange

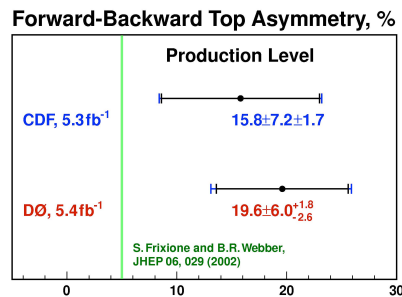


(d) Born diagram

Anecdotes From the Field (II)

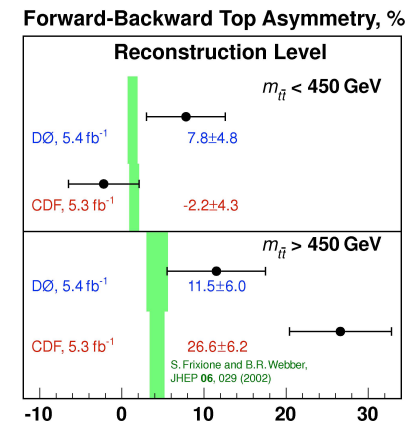
❖ ttbar charge asymmetry at the Tevatron

- At Feynman diagram level, NLO effect (Tevatron is proton-anti-proton collider)



<http://arxiv.org/abs/1107.4995>

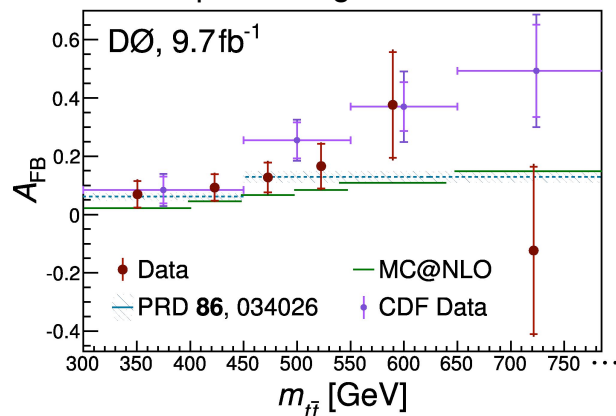
Ca. 2010, big fuss:
much larger than SM!



- In real life, already exists at \sim LO!

- Shown it is there in Pythia: parton shower, recoils! <http://arxiv.org/abs/1205.1466>

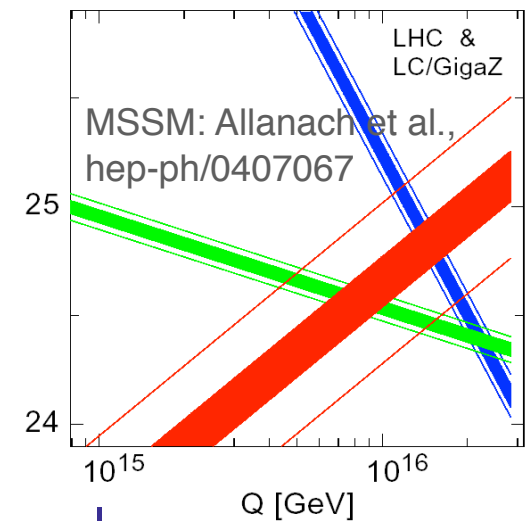
<http://arxiv.org/abs/1405.0421>



(My) current conclusion:
no BSM physics here: just (N)NLO
+ non-perturbative effects at work

Not SUSY?

- ❖ SUSY theories (and others with full or partial set of SM-partners) have a number of attractive features
 - “Explanation” for low Higgs mass (and sometimes EWSB)
 - Gauge coupling unification (often)
 - Dark matter candidate (if introduce a new parity, natural in UED, ~ad-hoc in SUSY)
 - No new interactions (often)
- ❖ But answering those questions comes at a large cost
 - Many new particles, with masses and mixing angles
 - Need to explain why mass scale is so low (or high), spin?



Not SUSY?

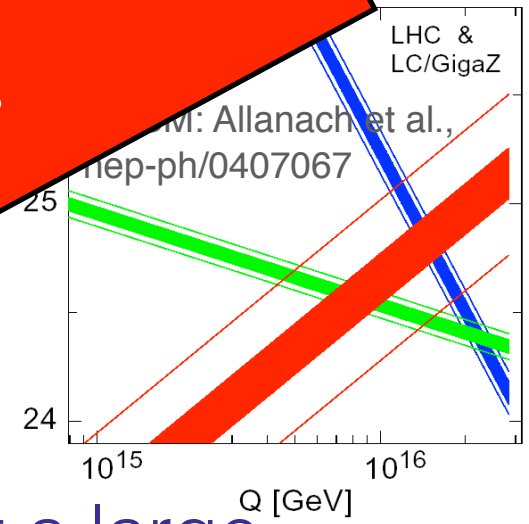
❖ SUSY theories (and others with full or partial set of SM-partners) have a number of attractive features

- “Explanation” for low Higgs mass (and so for the hierarchy problem)
- Gauge coupling unification (often)
- Dark matter candidate (if R -parity, natural in UED)
- No new interactions

❖ But an important question comes at a large scale

- Many new particles, with masses and mixing angles
- Need to explain why mass scale is so low (or high), spin?

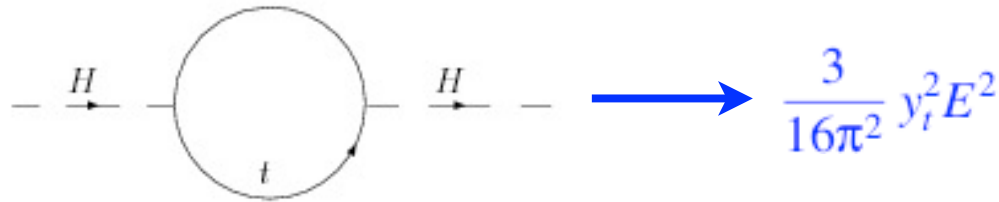
Dinosaurs on Venus?



Less Ambitious

Giving up on Dark Matter

- ❖ Electroweak-scale WIMPs fit the data well
 - But maybe hard/impossible to produce at colliders
- ❖ Or dark matter not WIMPs at all
- ❖ Back to problem #1:



The diagram shows a circular loop of top quarks (labeled 't') with two external Higgs boson lines (labeled 'H'). A blue arrow points from the diagram to the mathematical expression $\frac{3}{16\pi^2} y_t^2 E^2$.

➡ Top partner!

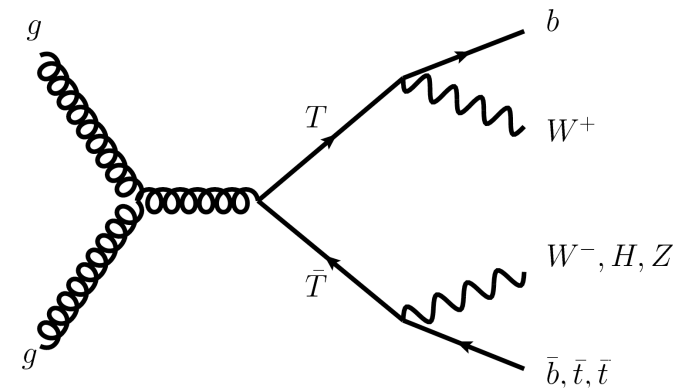
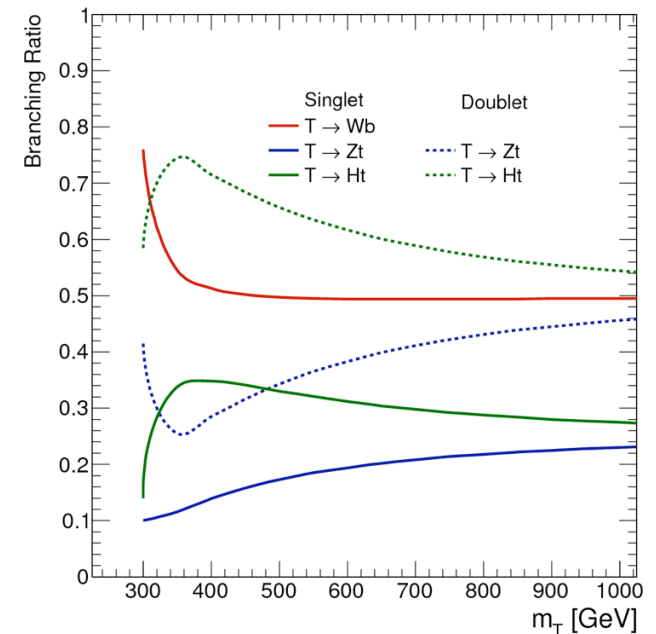
Singlets, Doublets, ...

❖ Vector-like top partners (still fermions) less constrained by flavor....

- Opens up decay modes
- Top partner partners:
 - $T^{5/3}$
 - ...

❖ Rich set of signatures

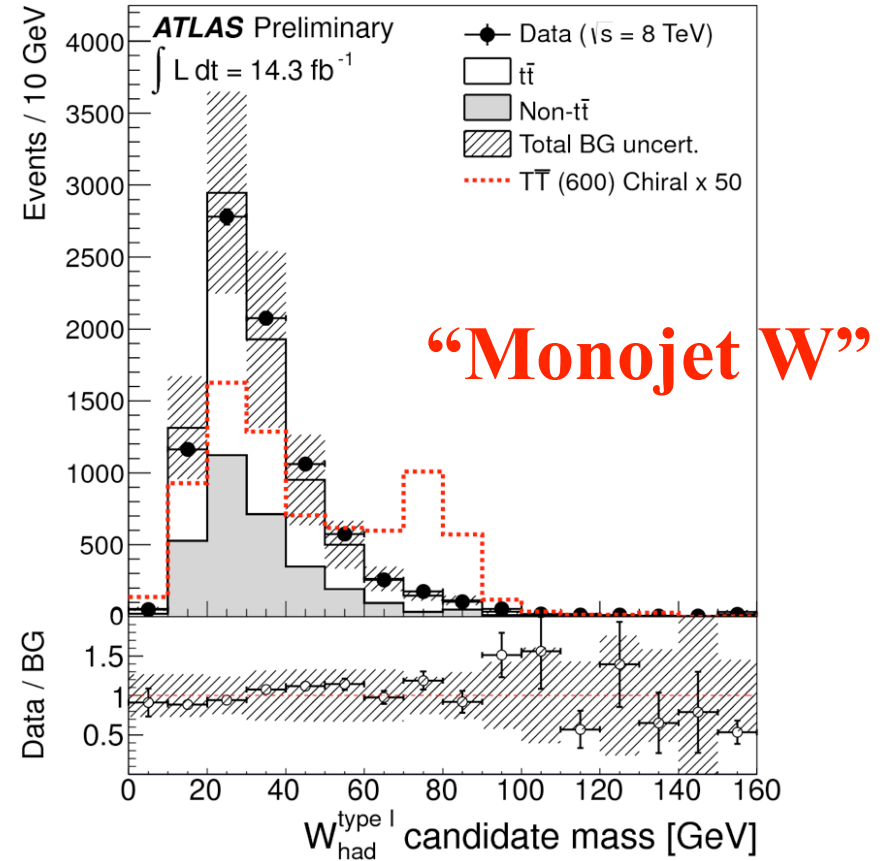
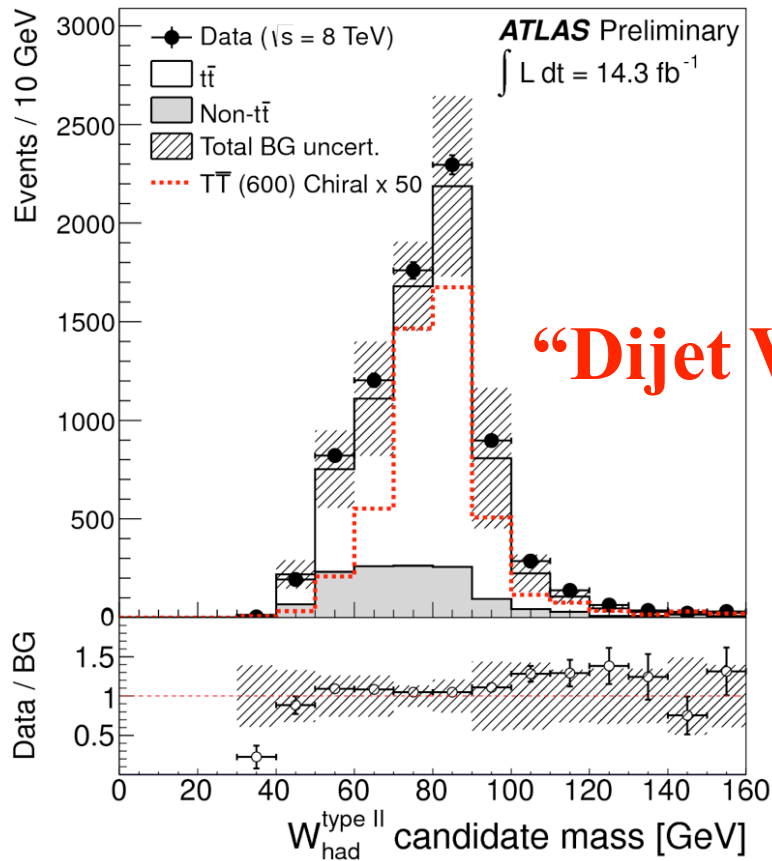
- Just no huge MET
- At least not systematically



W's Can Be Light

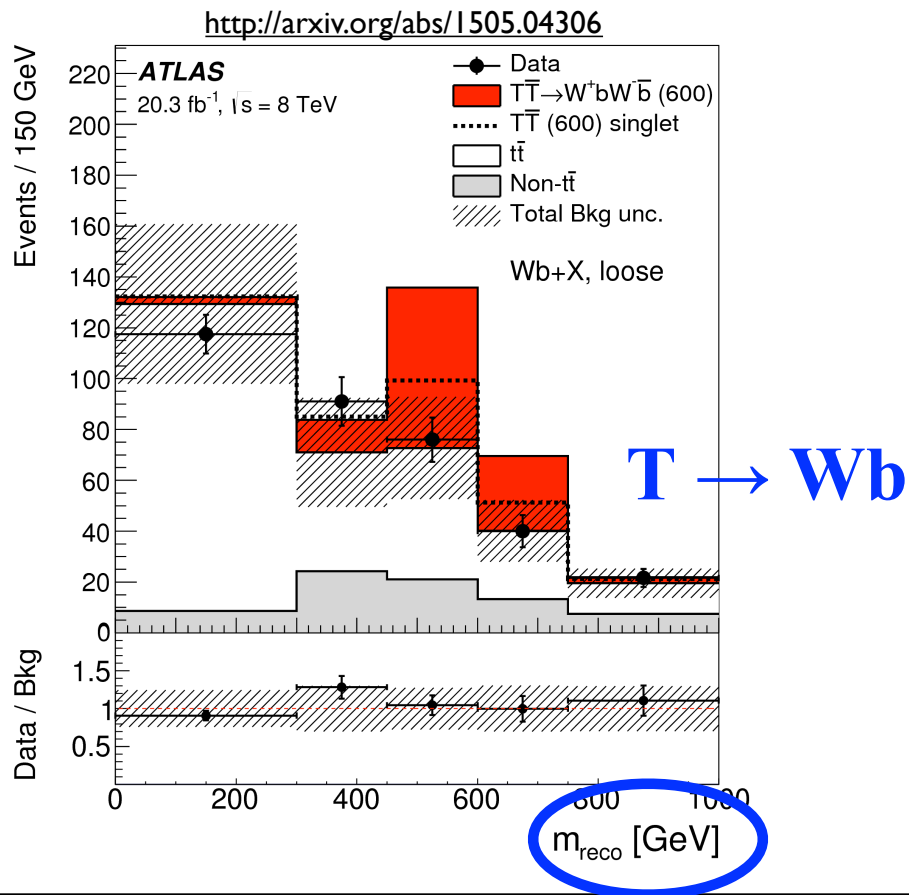
❖ $T \rightarrow Wb$ with $m_T \sim 600$ GeV

➔ W will be boosted, and if decays hadronically \rightarrow single jet



Wb versus Ht

- ❖ $T \rightarrow Wb$ yields the same final state as $t \rightarrow Wb$
 - Need to discriminate, e.g. reconstruct m_T

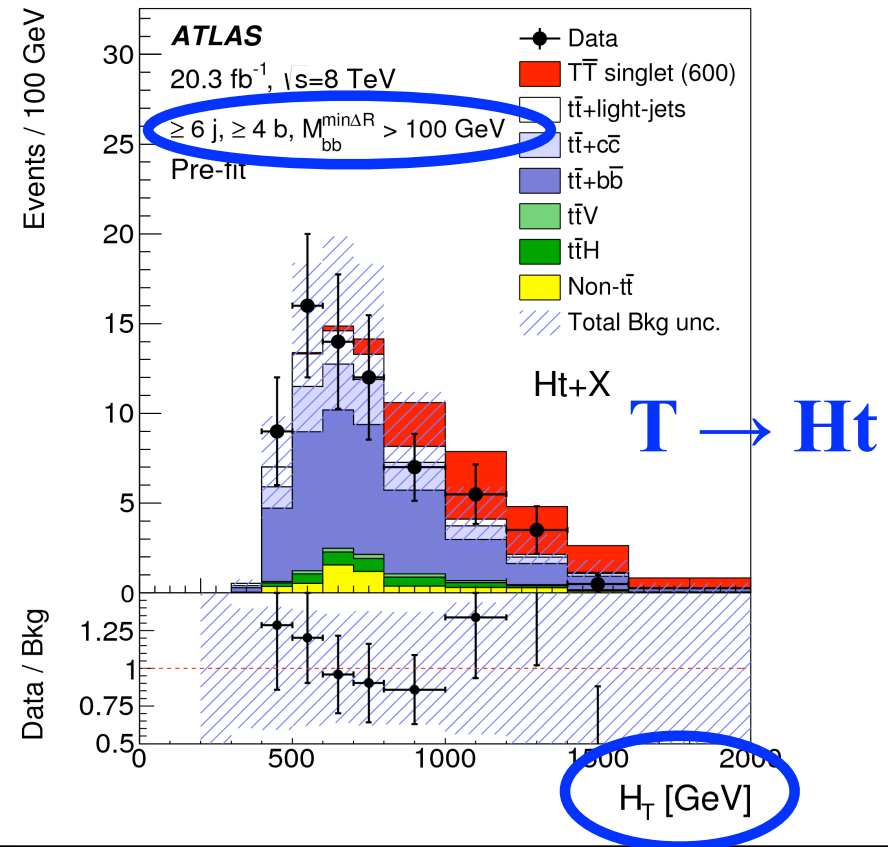
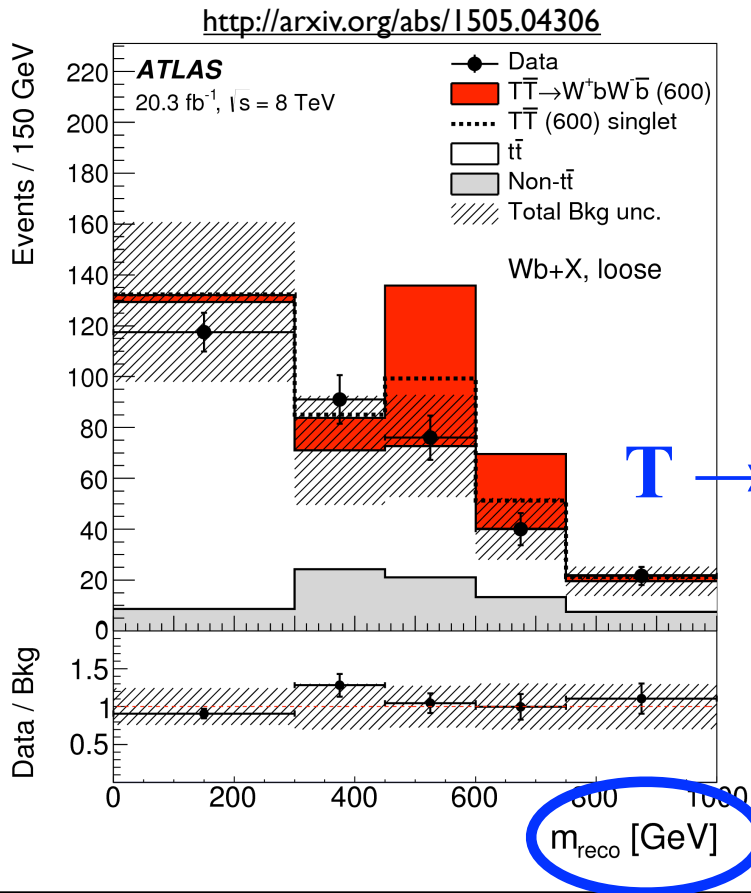


Wb versus Ht

❖ $T \rightarrow Wb$ yields the same final state as $t \rightarrow Wb$

- Need to discriminate, e.g. reconstruct m_T

❖ $T \rightarrow Ht$: $t\bar{t}HH$, so $WWbbbb$

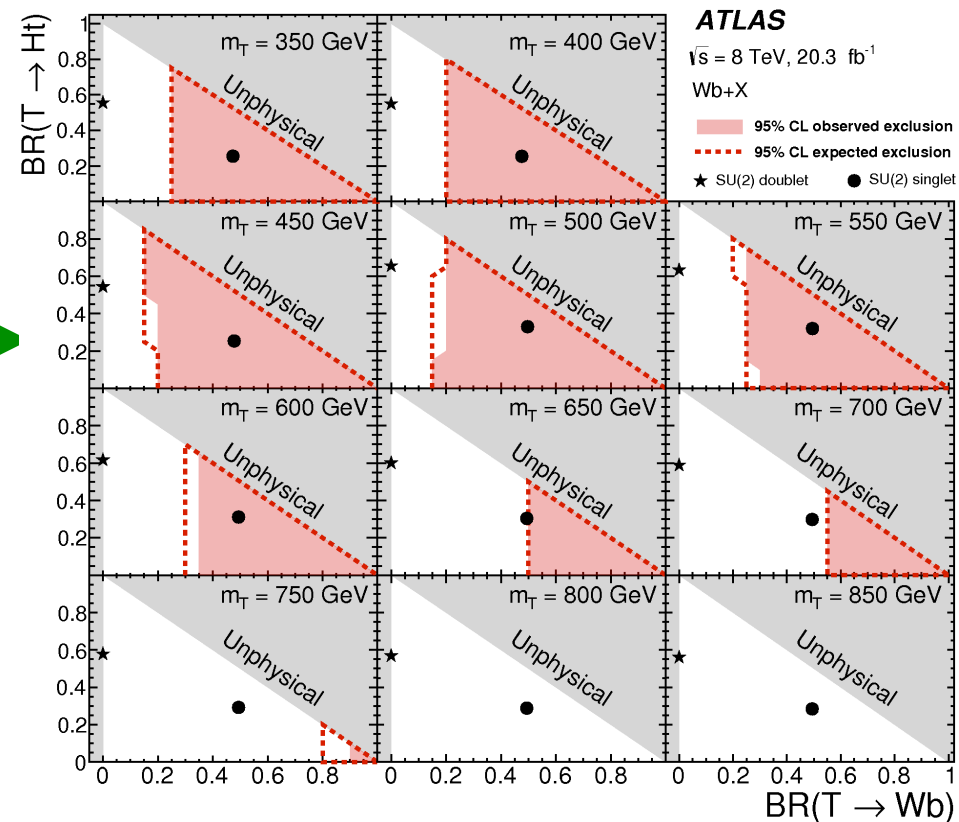
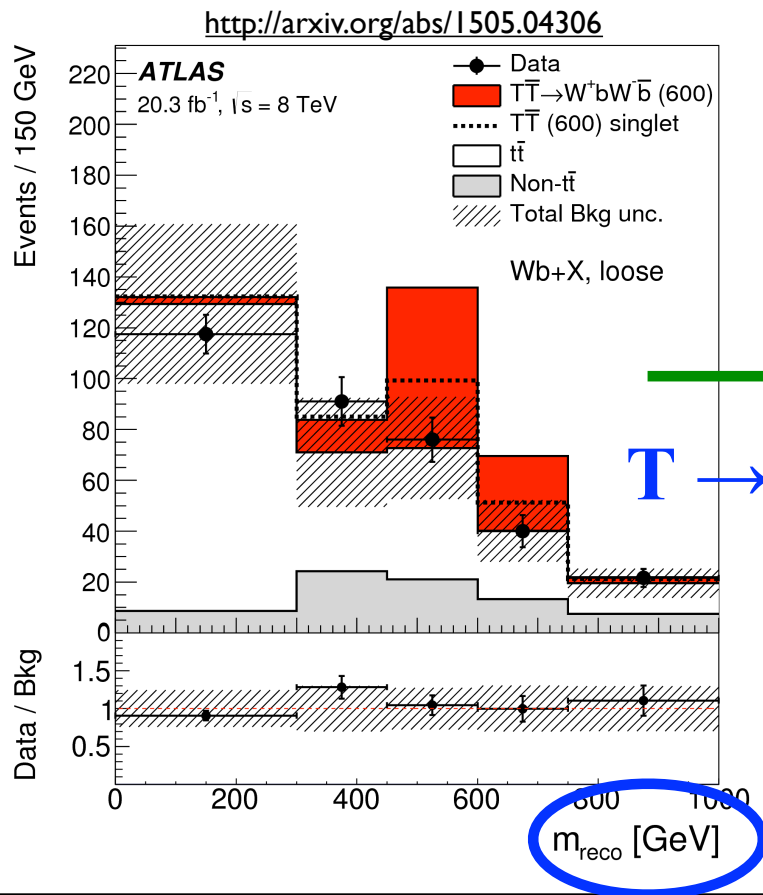


Wb versus Ht

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❖ $T \rightarrow Ht$: $t\bar{t}HH$, so $WWb\bar{b}\bar{b}\bar{b}$

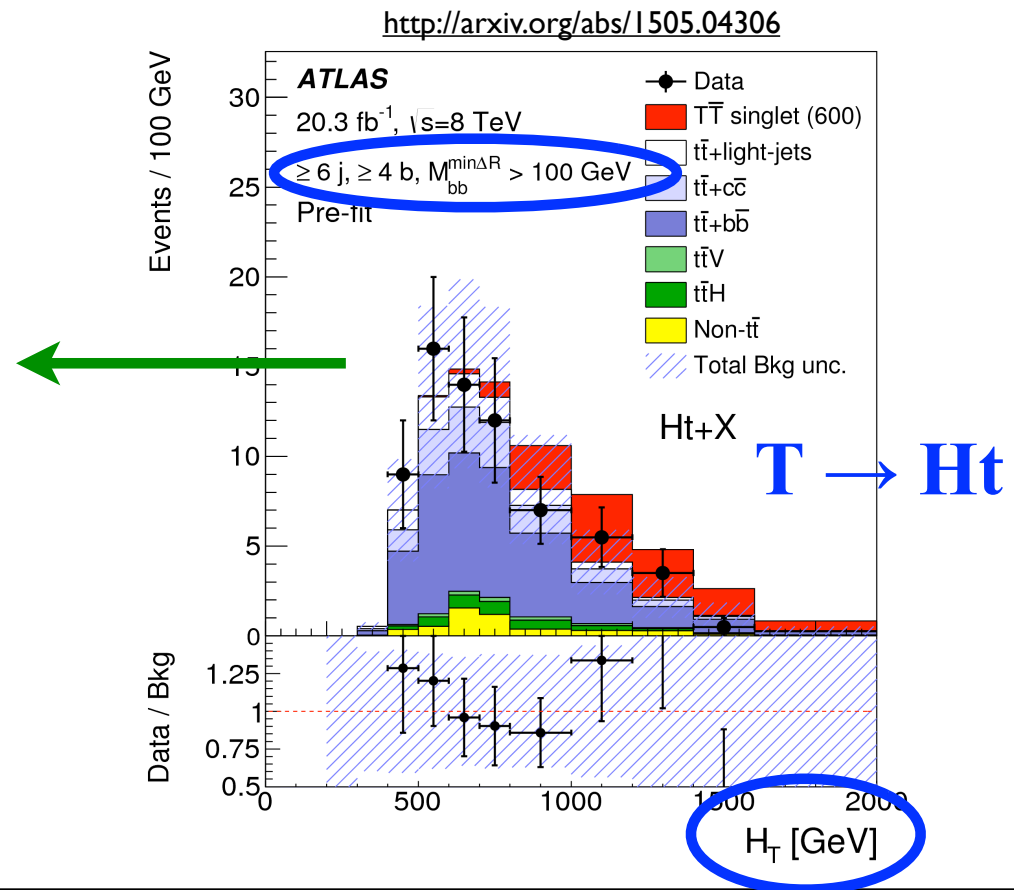
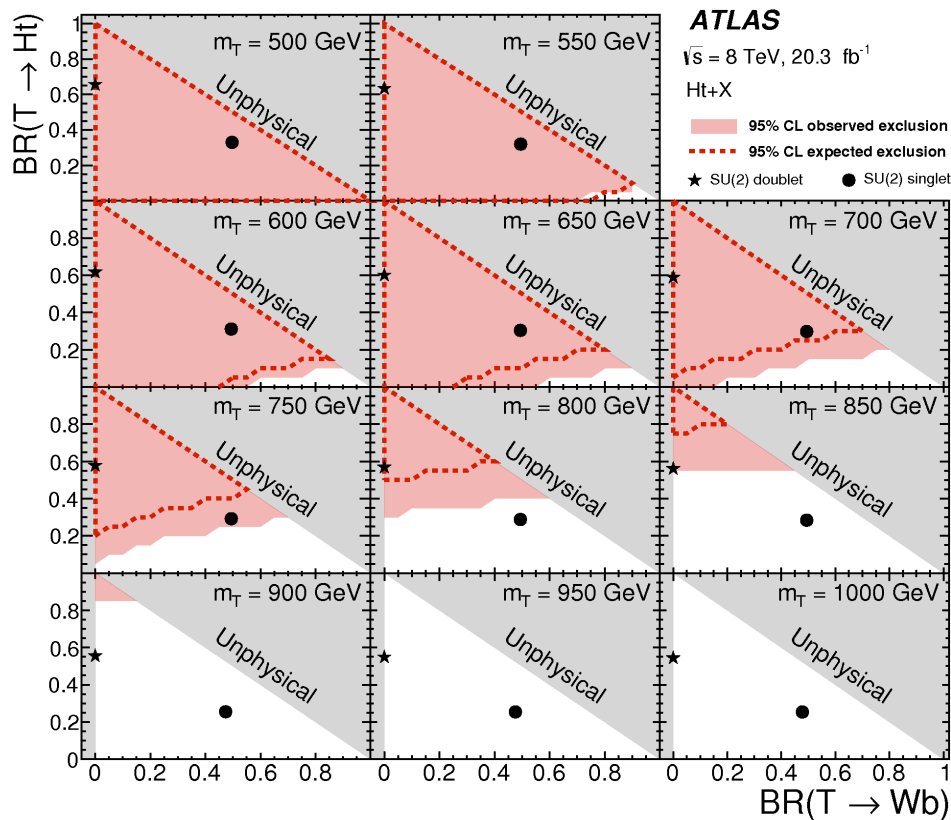


Wb versus Ht

❖ $T \rightarrow Wb$ yields the same final state as $t \rightarrow Wb$

- Need to discriminate, e.g. reconstruct m_T

❖ $T \rightarrow Ht$: $ttHH$, so $WWbbbb$

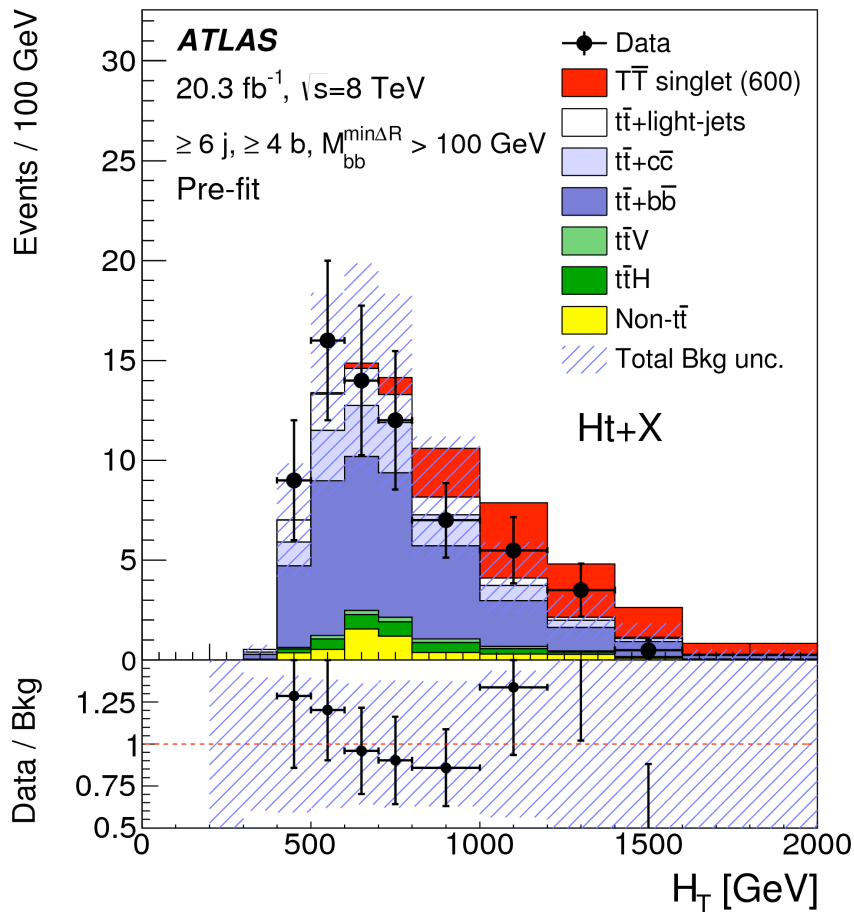


Systematic Uncertainties

- ❖ Statistical uncertainties are easy: with limited number of events (and experiments), precision on a measurement is limited
- ❖ Systematic uncertainties vastly more complex
 - Example: measure a cross-section: $\sigma = \frac{N_{\text{events}}}{LA\epsilon}$
 - L is the integrated luminosity, A the acceptance, ϵ the efficiency
 - Statistical uncertainty comes from N_{events}
 - Systematic uncertainties arise from limited knowledge of L, A and ϵ
 - ▶ L is estimated from Van der Meer scans
 - ▶ A typically depends on parton distribution functions
 - ▶ efficiency is a convolution of many experimental uncertainties

Example

❖ H_T is the sum of scalar energies of jets, leptons,...



- If the jet energy scale is different between data and MC, comparison is wrong
- If the jet energy scale dependence on jet energy is wrong, distort shape
- etc.

❖ But how do I determine the jet energy scale uncertainty?

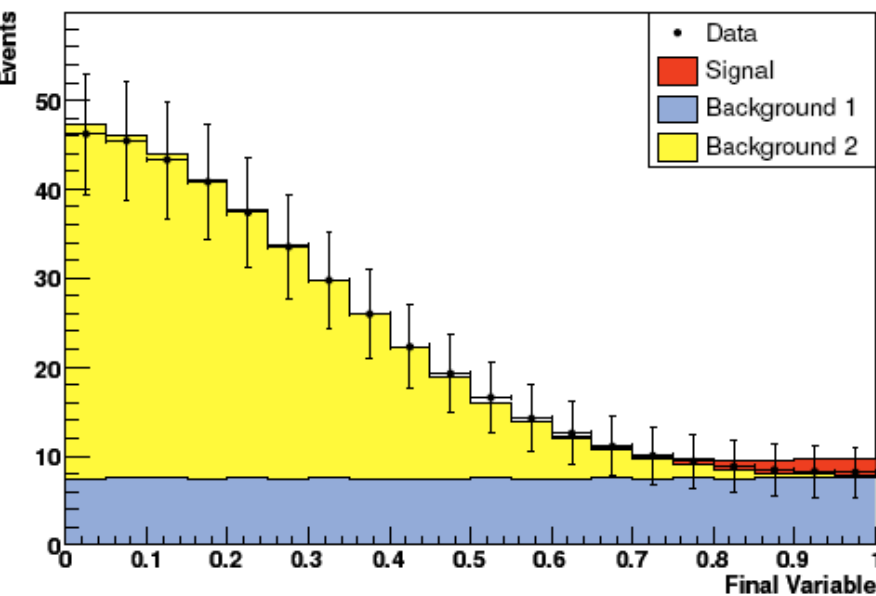
- testbeams (single pions)
- dijet balance
- γ/Z +jet balance
- ...

Systematics Profiling

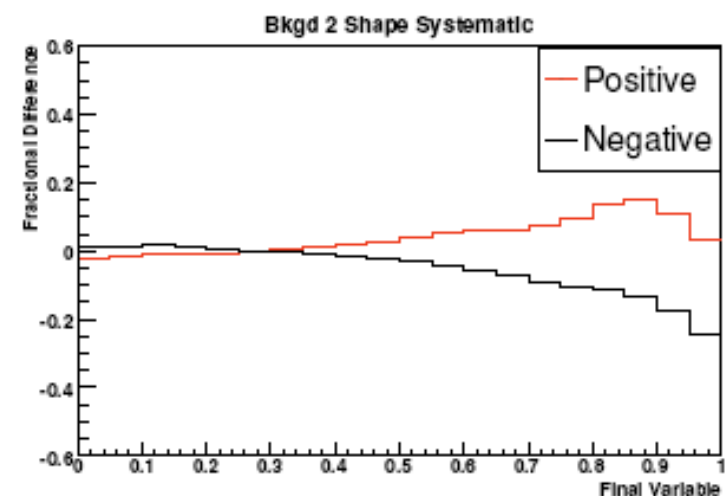
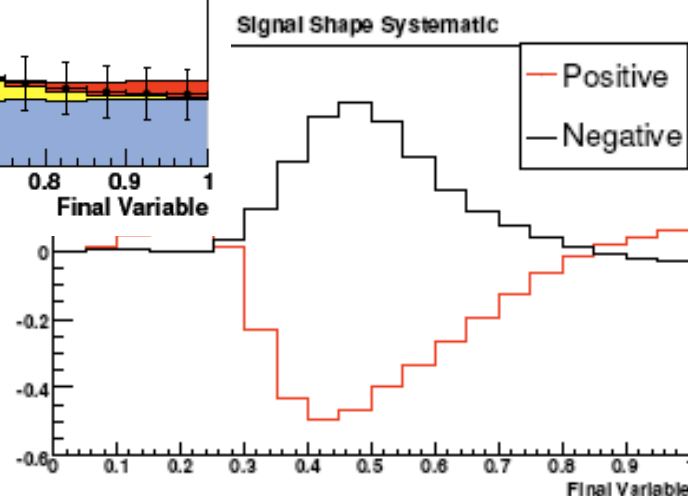
- ❖ Systematic uncertainties are propagated through the full analysis chain to the discriminating distribution
 - E.g. we repeat the analysis with jet energy scale shifted up & down by 1σ
 - Some systematic uncertainties affect shape (jet/lepton/photon reconstruction efficiency, energy scale and resolution, p_T distributions, background models), others only normalization (lepton reconstruction efficiencies and momentum calibration, background normalizations, theoretical cross-sections and luminosity)
 - Systematic uncertainties are treated as nuisance parameters when fitting signal+background to the data
 - I.e. modify signal and background shape
 - Can be fixed, or allowed to change

Systematics Profiling

- ❖ Nuisance parameters tend to be correlated, but not 100%, among backgrounds
 - Can affect rates, shapes, or both (in any distribution), and often asymmetric and non-gaussian



Toy Example (W. Fisher)

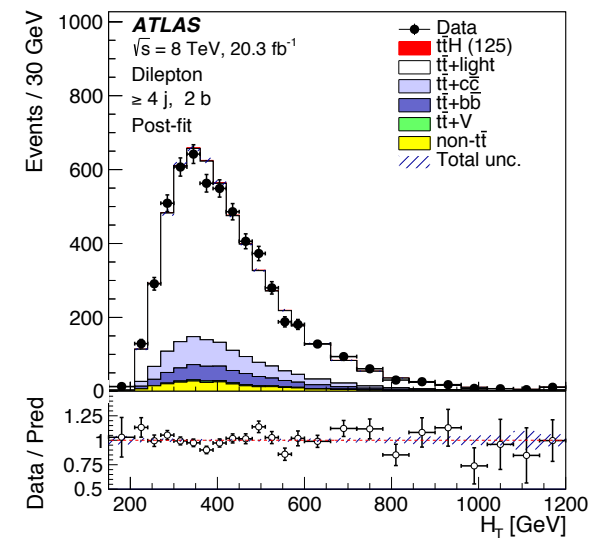
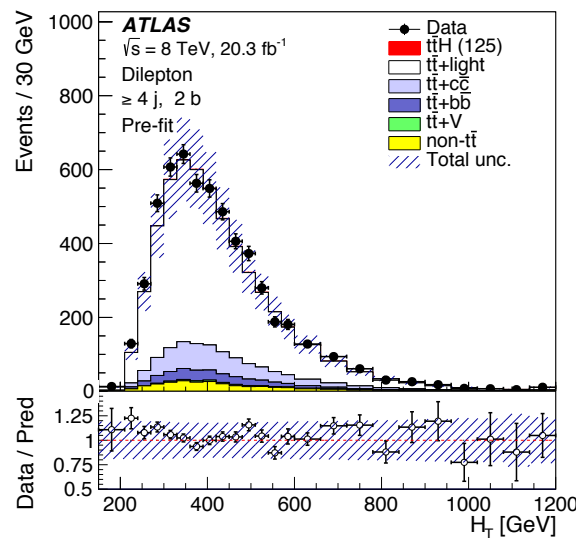
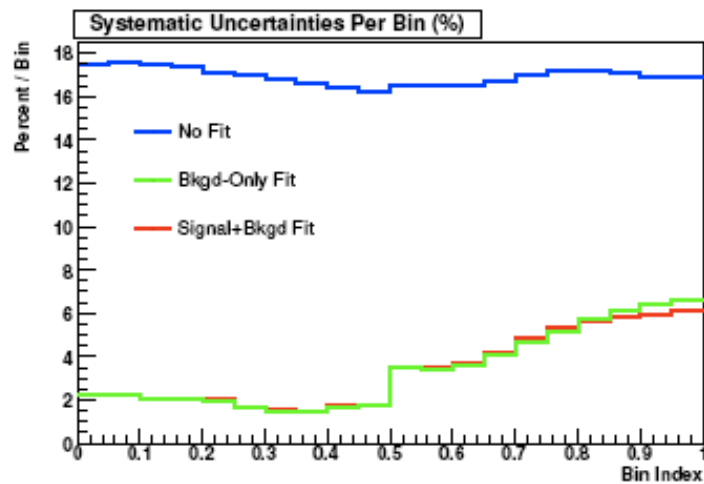


- ❖ Generate pseudo-experiments (events in bins according to poisson), then for each experiment vary nuisance parameters
 - Variations in background (& S+B) prediction
 - Compare results to data using log-likelihood ratio
- ❖ We can maximize likelihood ratio as a function of nuisance parameters → constrain them
 - I.e. use full shape of distribution(s) to see which background uncertainties are over/underestimated
 - Of course limited to size of statistical fluctuations
 - Can remove bins with large S/B if needed
 - Mostly important if uncertainties lead to similar shape distortions
 - Want enough background-rich phase space in fit!
 - Even include control regions

❖ Test example:

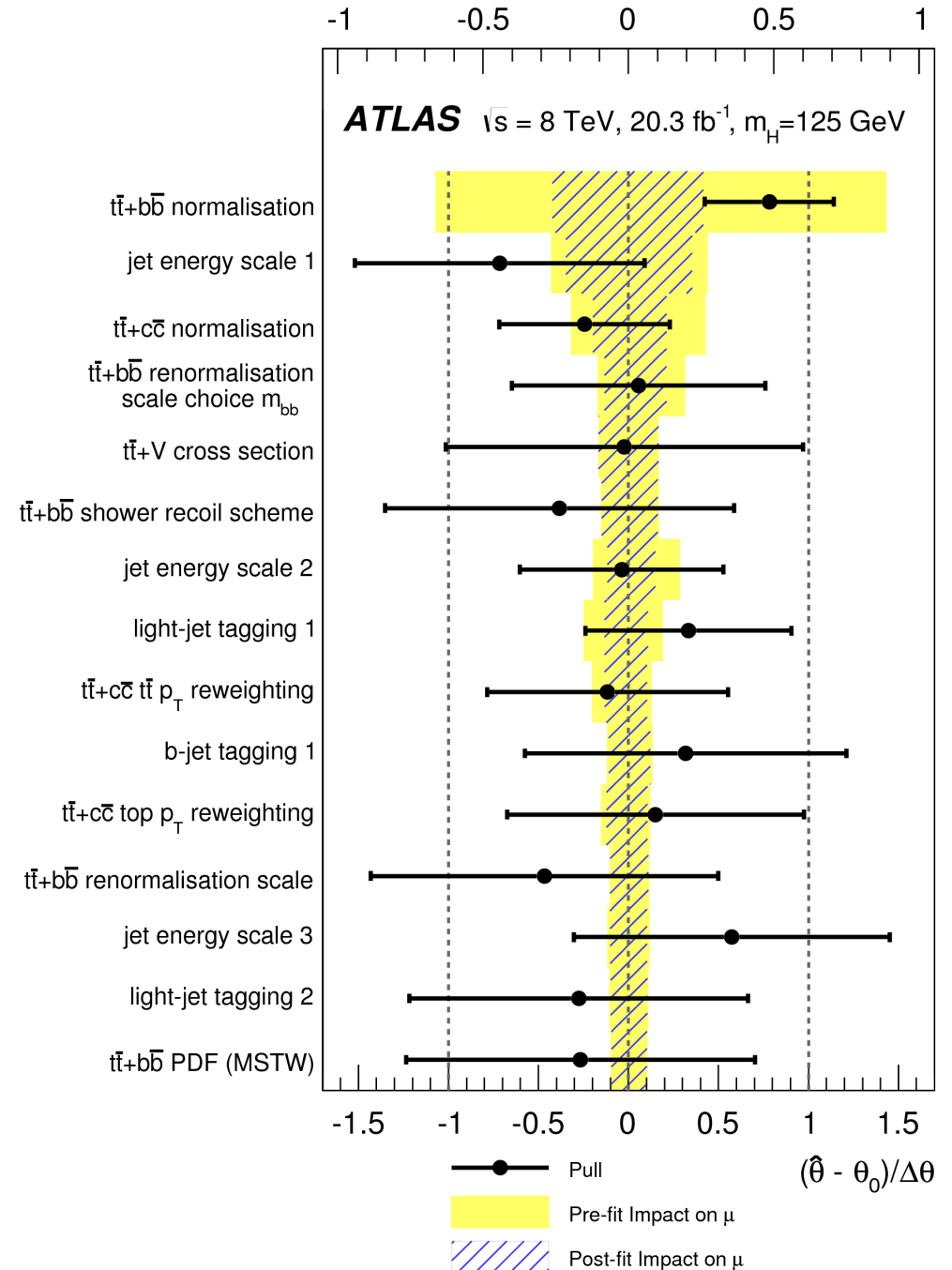
- Data constructed to disagree with background-only hypothesis (wrong estimates for background uncertainties)
- But to agree with background-only better than signal+background
 - Improvement quite spectacular (by construction in example)

ATLAS ttH search: [arXiv:1503.05066](https://arxiv.org/abs/1503.05066)



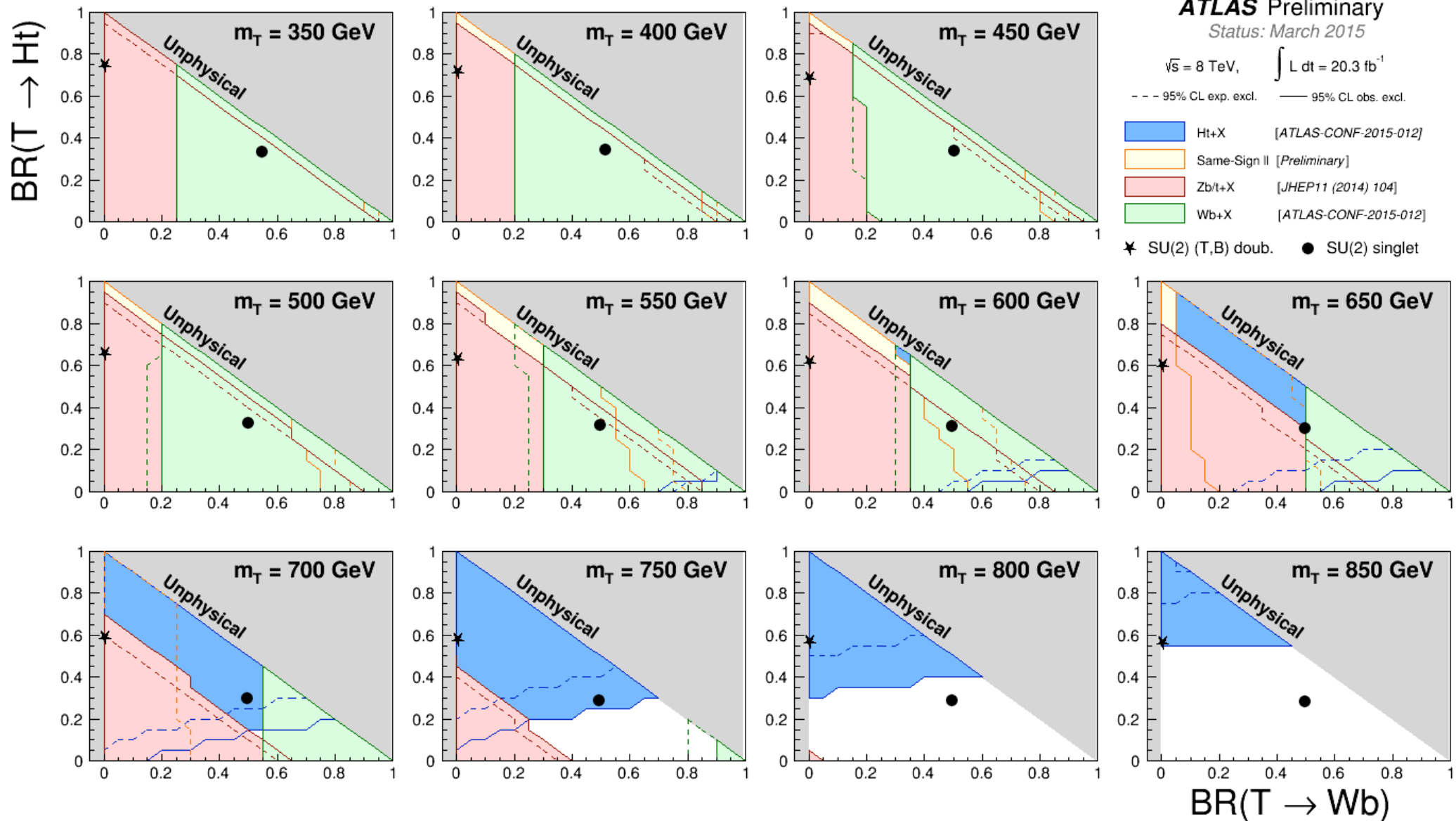
Fit Results

- ❖ Need to compare starting point and results
 - Pathologies due to lack of MC stats in some areas, strong correlations, ...
- ❖ Crucial to design analysis with good control regions the fit can use to address least understood systematics



ATLAS ttH search: [arXiv:1503.05066](https://arxiv.org/abs/1503.05066)

All Together Now



Presented Differently

