



Accelerator operation and luminous region scenarios

G. Arduini, R. Tomas for the EDQ Working Group

3rd ECFA HL-LHC Meeting Experiment Workshop – Aix les Bains - 03/10/2016

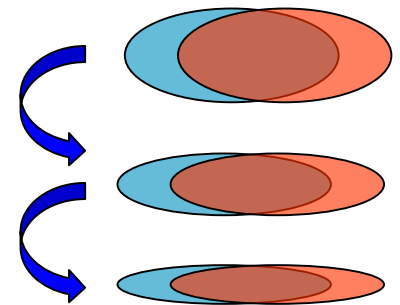
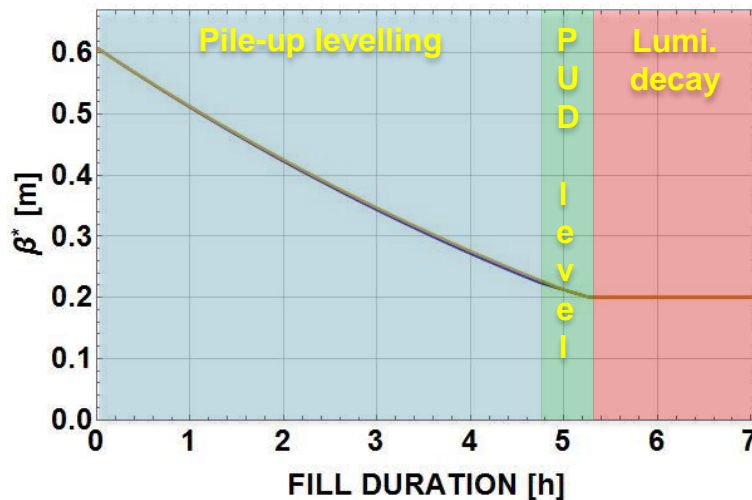
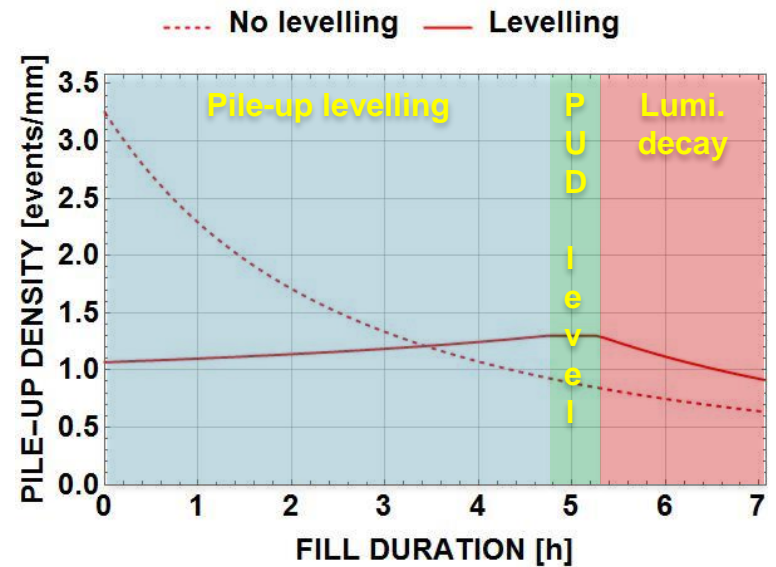
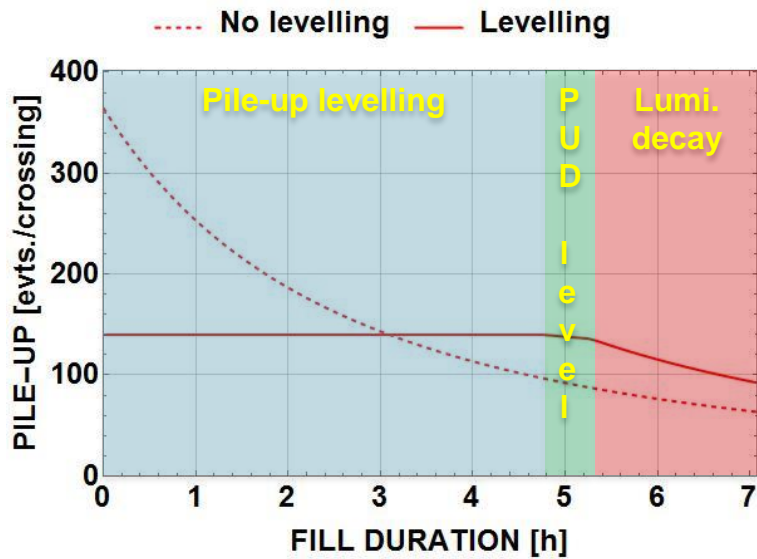
Mode of operation

$$L = \frac{kN_b^2 f \gamma}{4\pi \beta^* \varepsilon^*} F$$

- Operation at **pile-up/pile-up density limit** (set by the experiments) by choosing parameters that allow **higher than design pile-up (140 events) / pile-up density (<1.3 events/mm)**:
 - Beam brightness and in particular bunch population to sustain burn-off over long periods → **LHC Injector Upgrade**
 - Maximize number of bunches to minimize pile-up → **25 ns**
 - **Low β^* optics**
 - **Large crossing angle** to minimize the beam-beam effects
 - Fight the **reduction factor F** by **crab crossing**
- Improve **'Machine Efficiency'** → minimize the number of unscheduled beam aborts

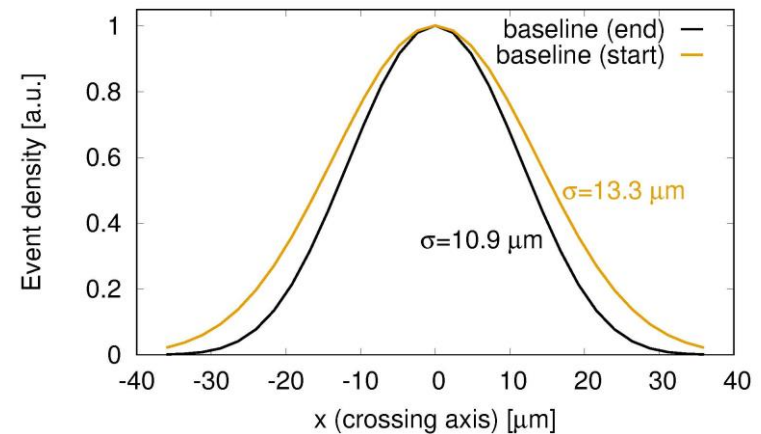
Mode of operation

$$L = \frac{kN_b^2 f \gamma}{4\pi \beta^* \epsilon^*} F$$



Luminous region

- 3D distribution of the collision event vertices
- Transverse sizes of the luminous region depends on:
 - β -function at the interaction point (varying during levelling $\rightarrow \beta^*$ is our nominal levelling mechanism for IP1 and IP5)
 - Transverse emittance
 - Transverse separation (levelling by separation is our nominal levelling mechanism in IP2 and IP8)
 - Crossing angle and crab cavity voltage



Luminous region

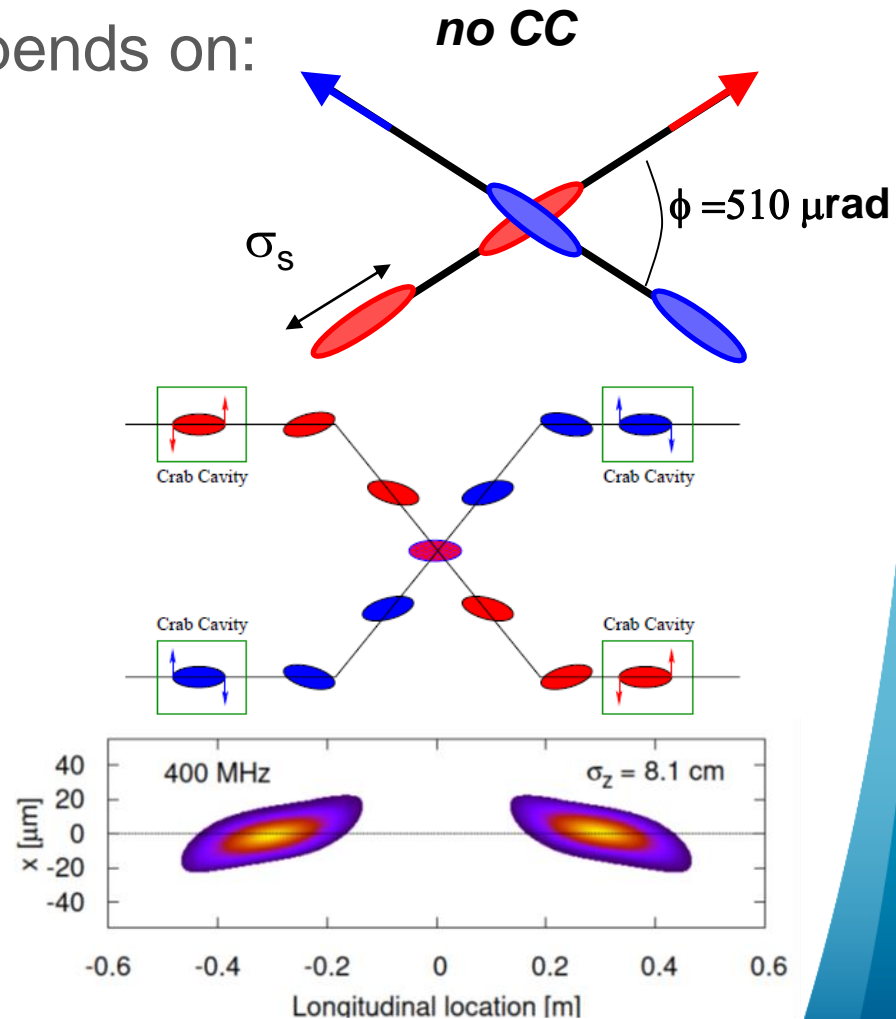
- 3D distribution of the collision event vertices

- Longitudinal size σ_{LRZ} depends on:

- Crossing angle

- Crab cavity voltage

- Crab cavity RF curvature



Luminous region

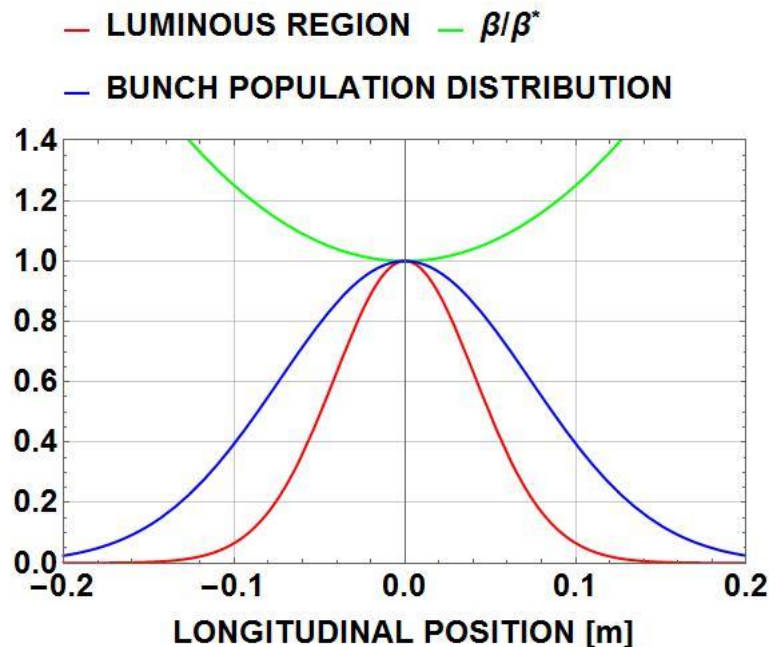
- 3D distribution of the collision event vertices

- Longitudinal size σ_{LRZ} depends on:

- Bunch length σ_z (for head on collisions $\beta^* \gg \sigma_z \rightarrow \sigma_{LRZ} = \frac{\sigma_z}{\sqrt{2}}$)

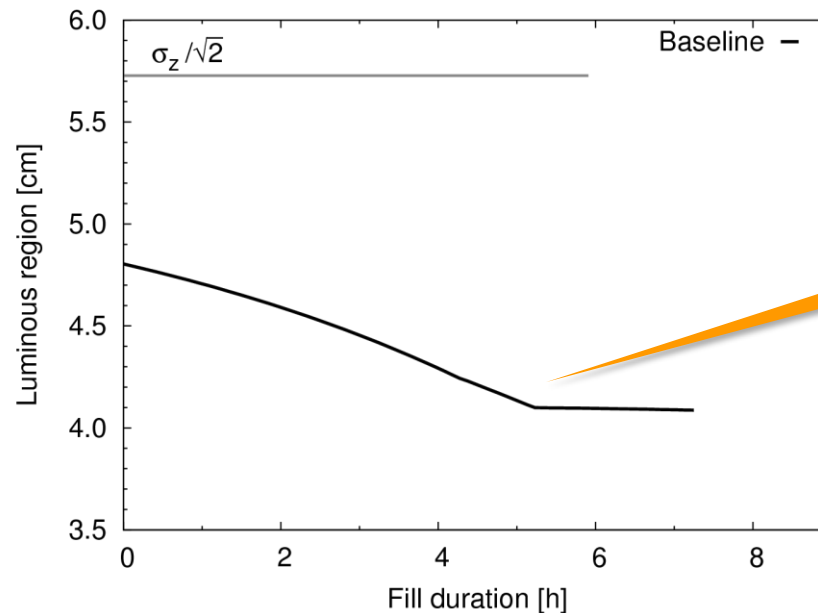
- Hourglass effect: varies during the fill because of the β^* levelling

Baseline scenario
Gaussian bunch longitudinal distribution (slightly pessimistic)
End of levelling



Luminous region

- r.m.s. length of the longitudinal distribution evolves during the levelling process. Because of the evolution of β^*



Minimum β^*

Baseline scenario

Pile-up

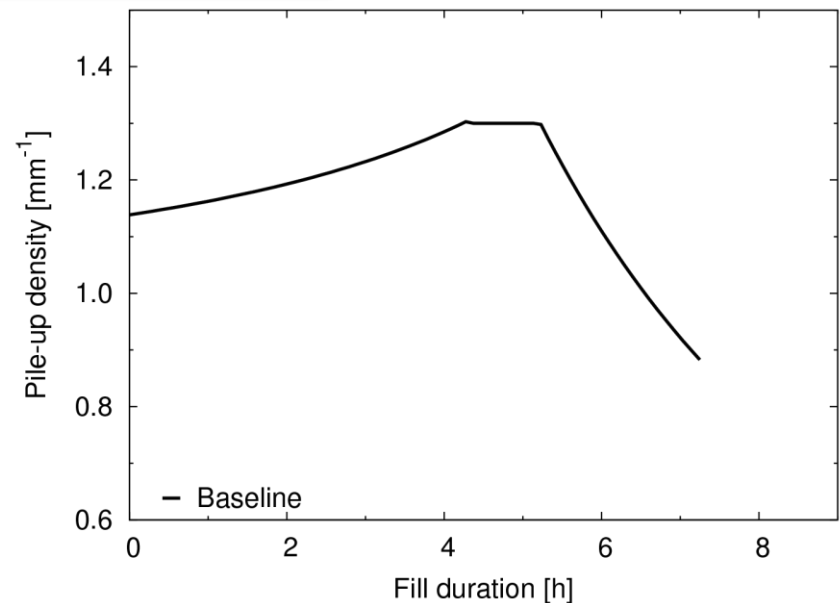
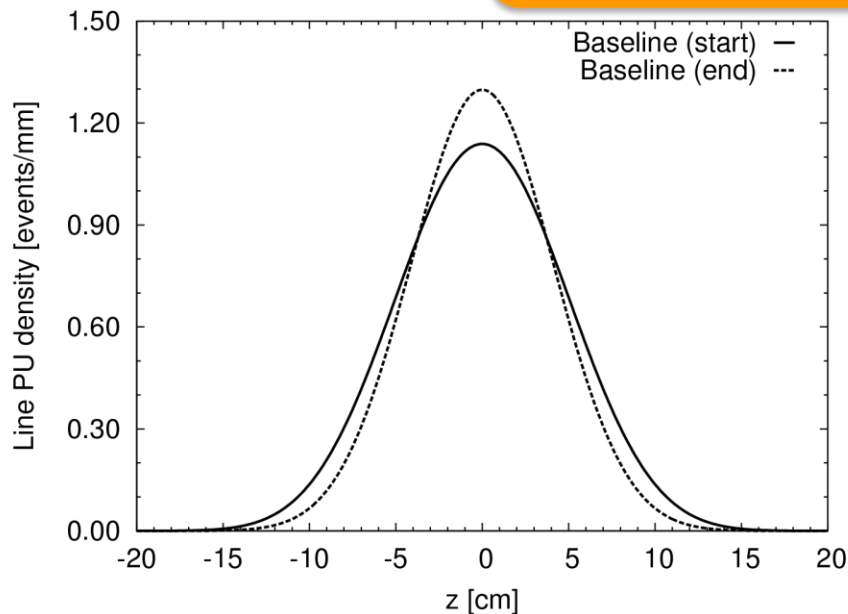
- Pile-up [events/crossing]:
 - A certain number of events can occur in a single bunch crossing
 - The number of events per crossing μ is distributed according to Poisson distribution with average $\langle\mu\rangle$. We call **pile-up** the value

$$\langle\mu\rangle(t) = \sigma_{inel} L_{inst}(t)$$

Line Pile-up Density

- Line Pile-up density [events/crossing/mm]:
 - Longitudinal density of the event distribution, evolves during the levelling process

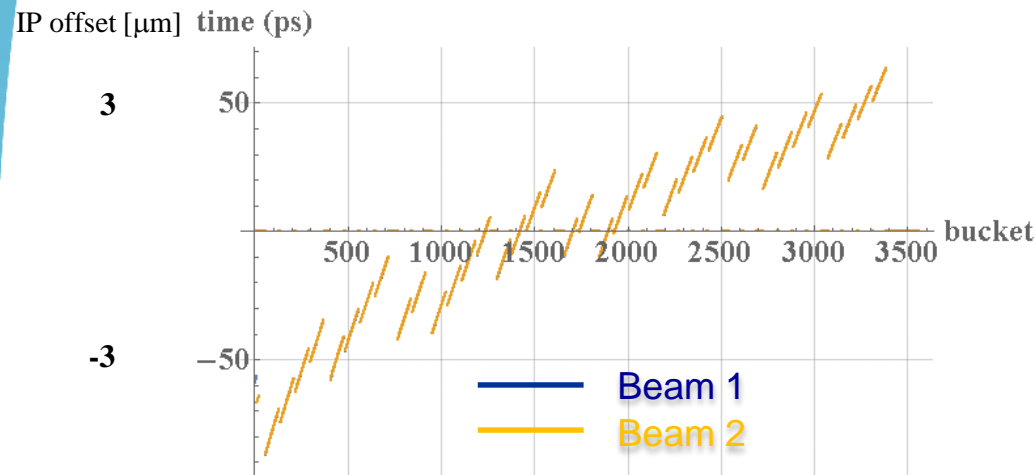
$$\frac{d\langle\mu\rangle}{ds}(s,t) = \sigma_{inel} \frac{dL_{inst}(s,t)}{ds}$$



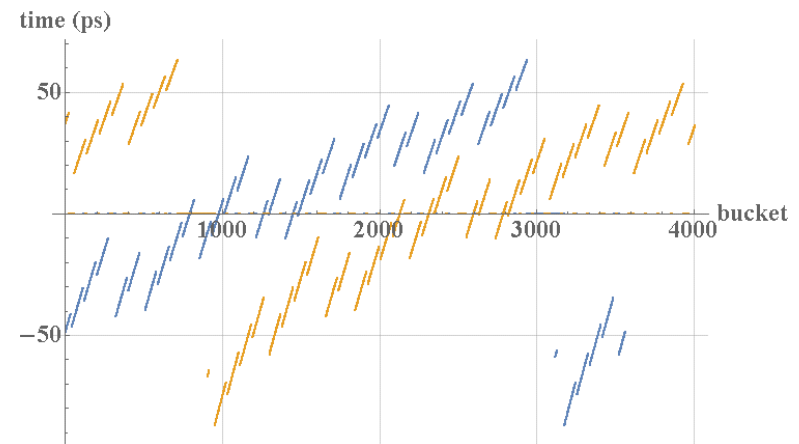
Baseline scenario

Other features

- In order to operate at double current as compared to LHC with the available klystron power we will have to operate in the so-called “full detuning mode” the main 400 MHz RF system → **Modulation of the bunch arrival time over the machine turn $\sim \pm 70$ ps**



IP1/5 – symmetric for symmetric filling patterns



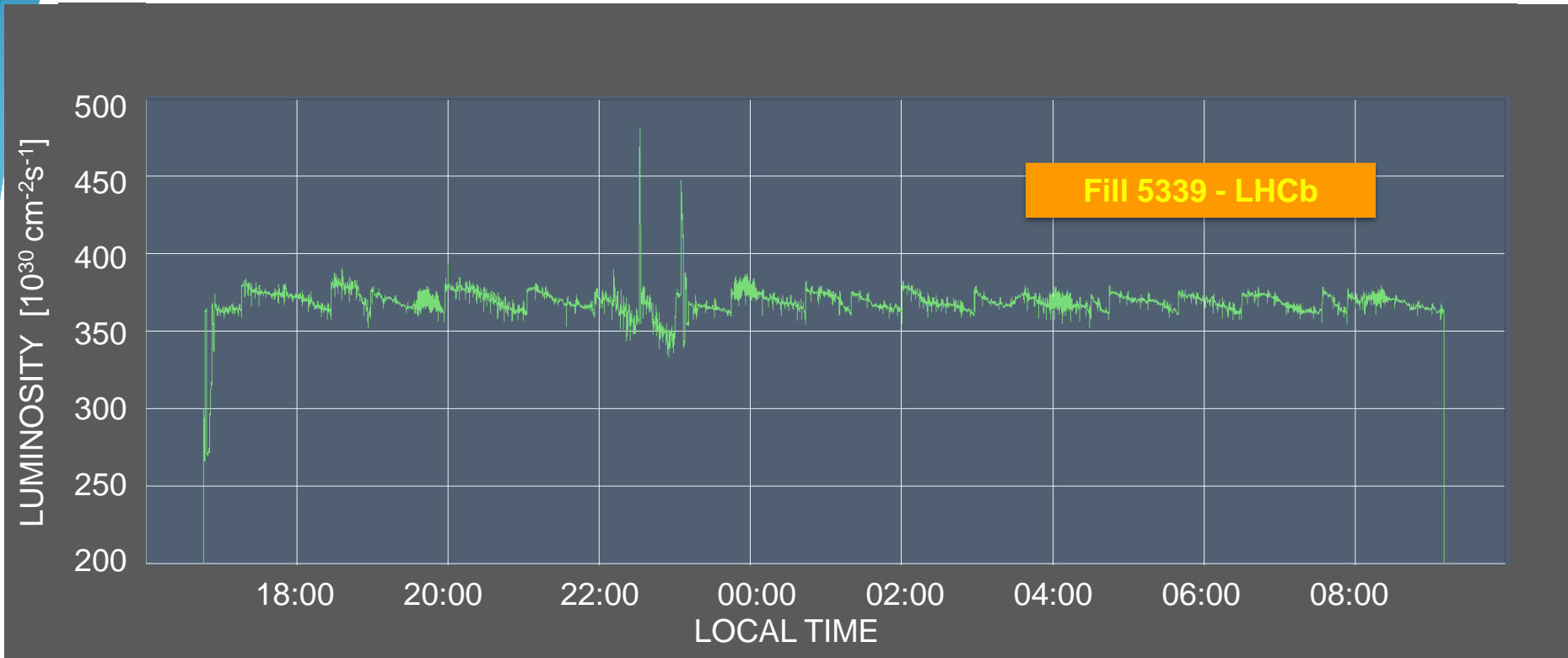
IP2/8 – asymmetric even for symmetric filling patterns

- modulation of the transverse position** of the collision point at **IP1-5** in the crossing plane and **modulation of the longitudinal position** of the collision point in **IP2-8** ($\sim \pm 1$ cm) → $\sim 1\%$ lumi. modulation

Levelling

- Any form of levelling implies:
 - Getting the observable(s) on which we need to level
 - Having a variable to act on the observable
- At present we are counting on:
 - Luminosity
- In the future we will need information on
 - Luminosity \sim proportional to pileup for a given number of colliding pairs
 - Pile-up density (or luminosity density: i.e. r.m.s. luminous region length and luminosity)

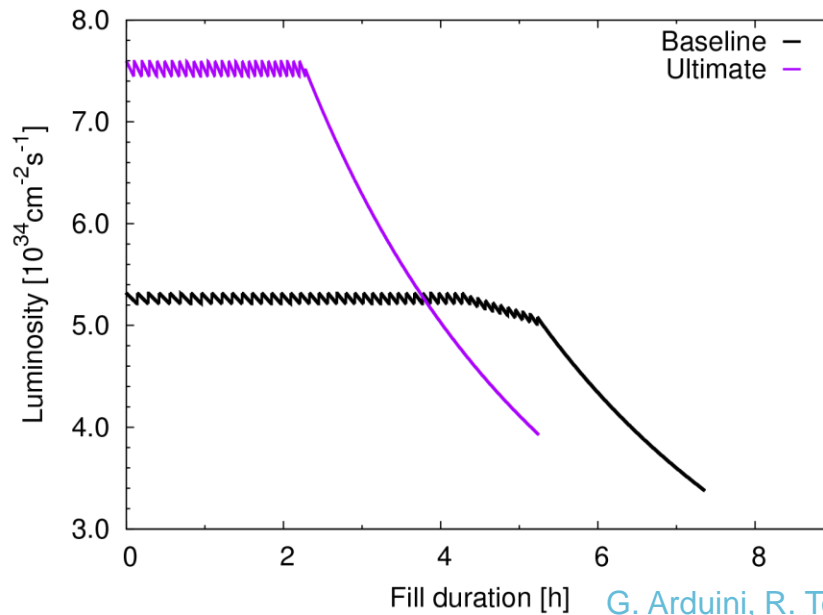
How does it look like today?



- Levelling by separation within $\pm 2\%$
- β^* levelling more complex but first MD results positive
→ **Need to gain operational experience**

Levelling

- Granularity of 2% assumed for HL-LHC fill simulations:
 - Possible hybrid schemes with feedforward based on expected evolution and feedback based on levelling by separation could be considered → we need to work out the details

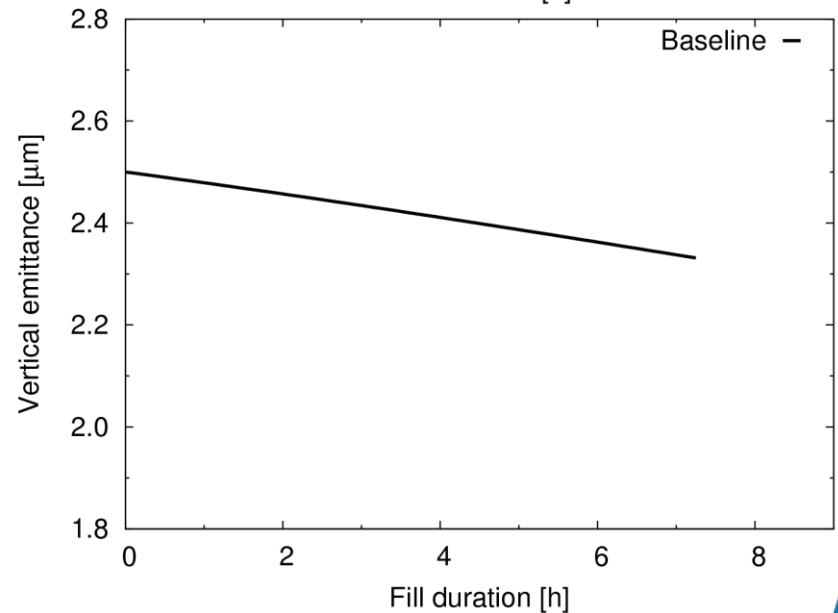
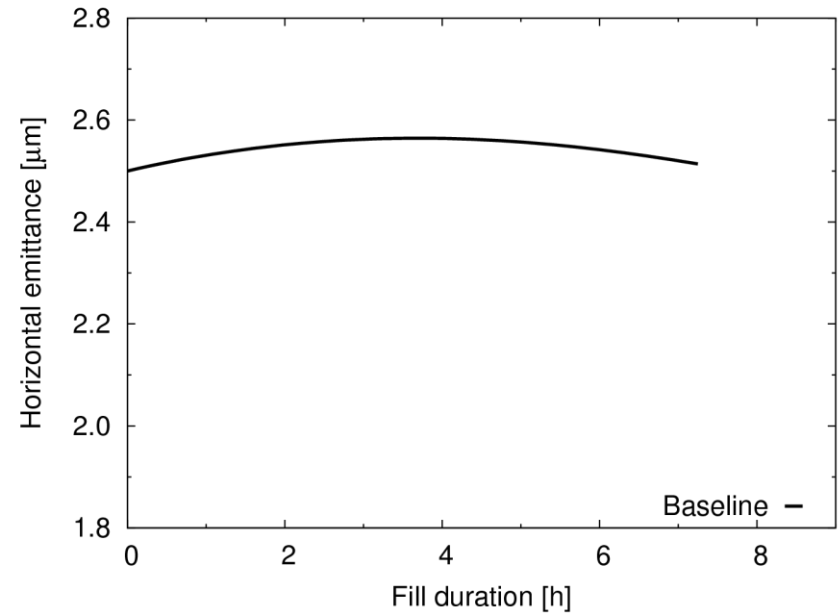
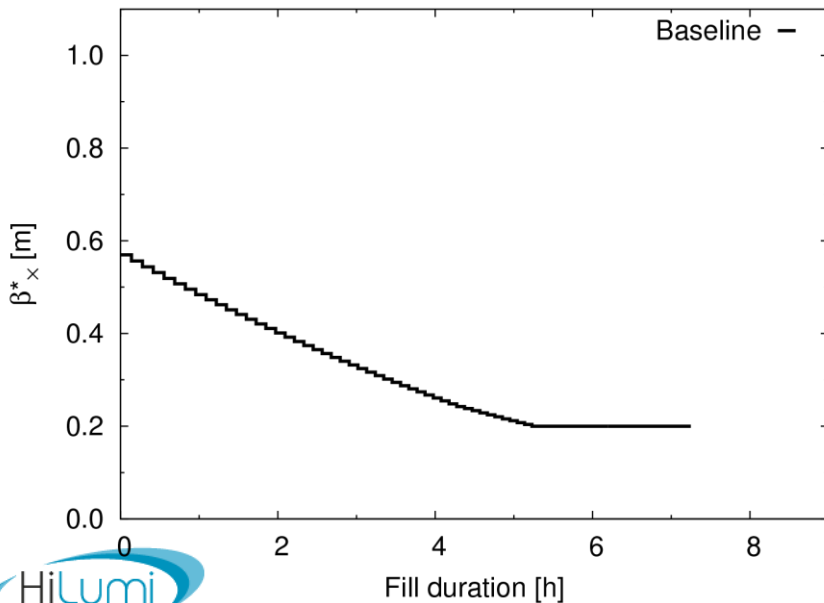


HL-LHC Parameters

Parameter	Nominal	HL-LHC updated
Bunch population N_b [10^{11}]	1.15	2.2
Number of bunches	2808	2748
Beam current [A]	0.58	1.12
Stored Beam Energy [MJ]	362	677
Full crossing angle [μrad]	285	510
Beam separation [σ]	9.4	12.5
Min β^* [m]	0.55	0.2
Normalized emittance ε_n [μm]	3.75	2.5
r.m.s. bunch length [m]	0.075	0.081
Peak Luminosity (w/o CC) [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	1.2 (1.2)	12.6 (6.5)
Max. Luminosity [$10^{34} \text{ cm}^{-2}\text{s}^{-1}$]	1	5.3
Levelled Pile-up Pile-up density [evt. evt./mm]	26/0.2	140/1.3

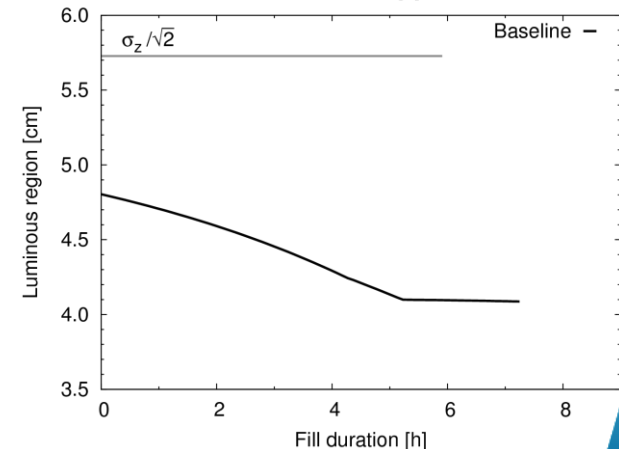
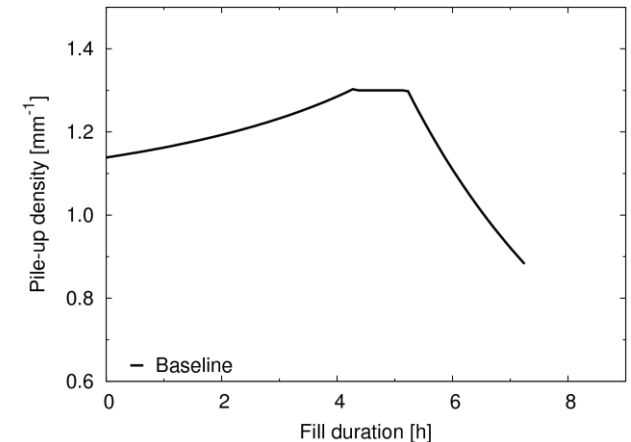
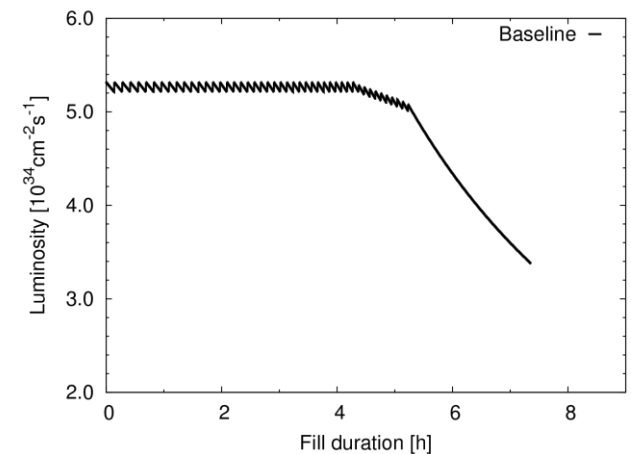
Baseline

- Round optics:
 - Constant crossing angle $510 \mu\text{rad}$
 - Min $\beta^* = 20 \text{ cm}$
 - Partial ($\sim 75\%$) compensation of crossing angle with two CCs per IP side per beam ($2 \times 3.4 \text{ MV}$)



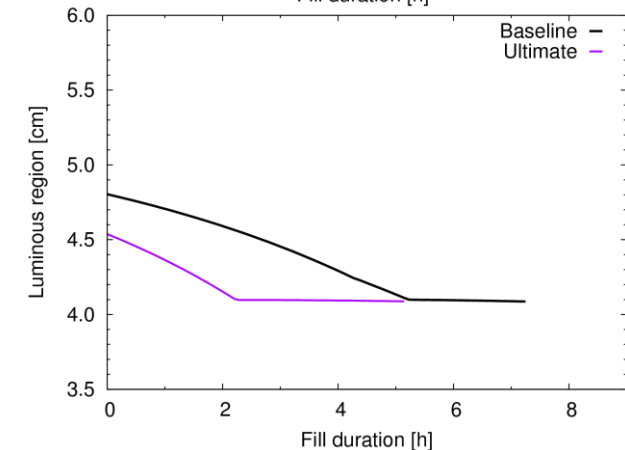
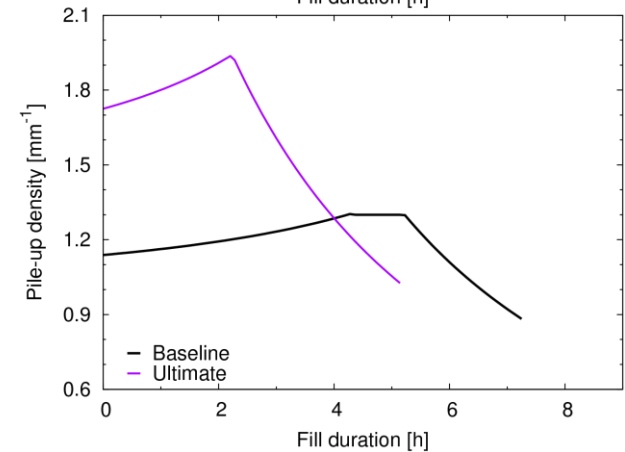
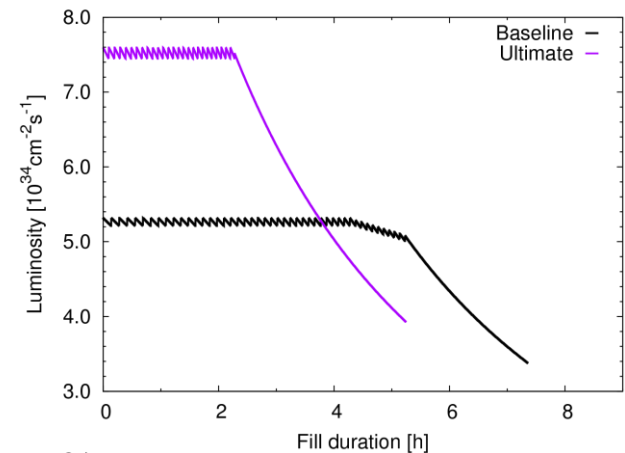
Baseline

- Levelled luminosity at $5.3 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$ for **pile-up = 140 events/xing**
- **Peak line pile-up density limited to 1.3 events/mm/xing.**
- **3000 fb^{-1} during HL-LHC period**
- Assumptions:
 - Min. turn-around time: 3 h
 - Performance Efficiency: 50% (60.1 % in 2016 – so far)
 - 160 days of proton physics (this number increases in Run 5 (200 days) and Run 6 (220 days))



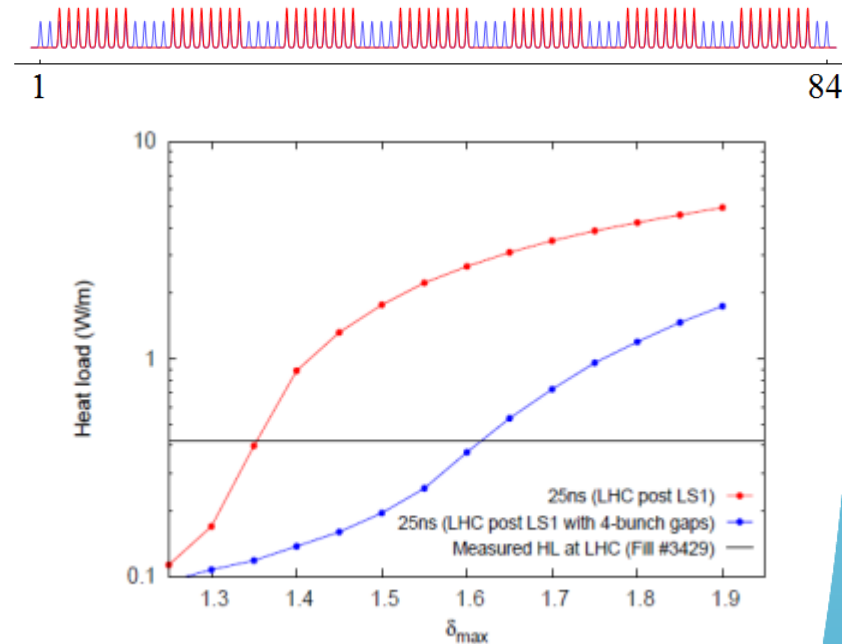
Ultimate

- Differences w.r.t. baseline
 - levelling at **pile-up = 200 events/xing**
→ Luminosity levelling at $7.6 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$
 - Peak **line pile-up density up to 1.9 events/mm/xing**
 - Beam parameters are the same
- **Up to 4000 fb^{-1} during HL-LHC period** compatibly with engineering margins assuming **58% performance efficiency** (2016 so far ~60%)

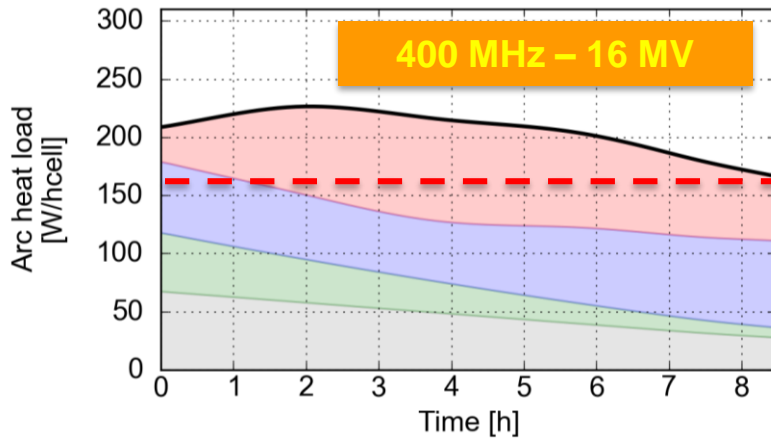


Back-up scenario (e-cloud mitigation): 8b+4e

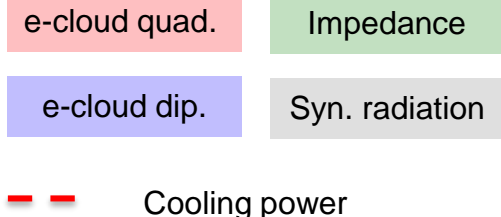
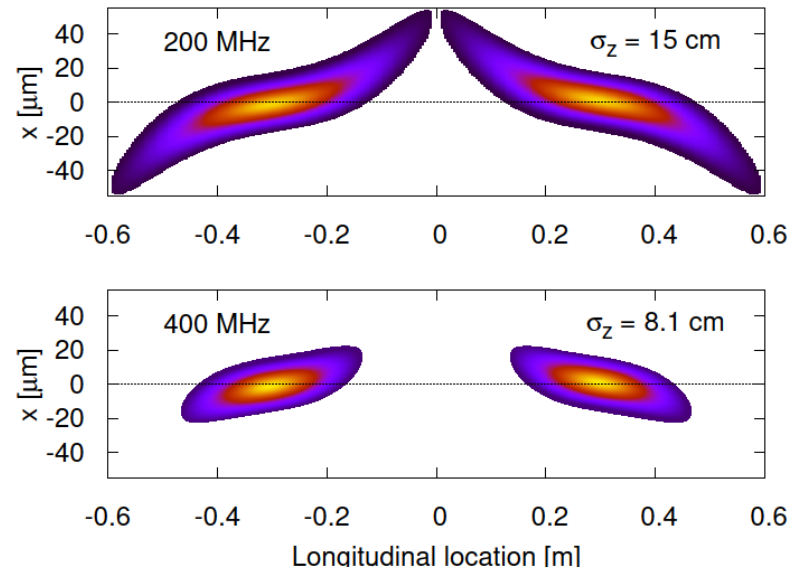
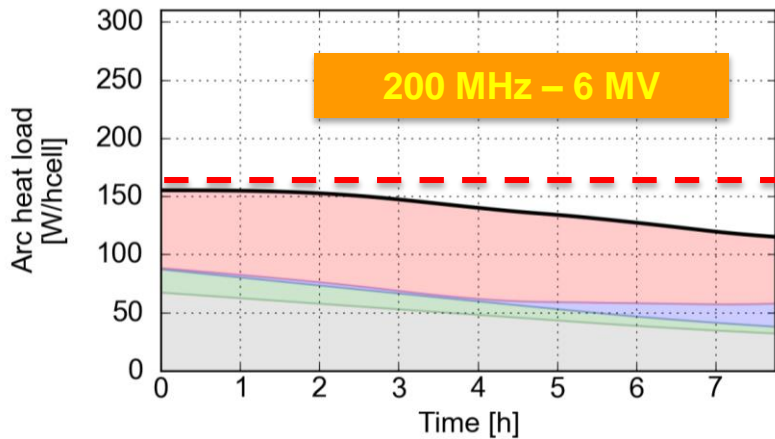
- Relies on scrubbing to suppress electron cloud in the dipoles (heat load and beam stability)
- Alternatives to mitigate electron cloud:
 - 'ad-hoc' 25 ns filling schemes to minimize electron cloud build-up (e.g. 8b+4e scheme)
 - No additional HW but reduced number of bunches (1968 as compared to 2748)



Alternative scenario (e-cloud mitigation): 200 MHz

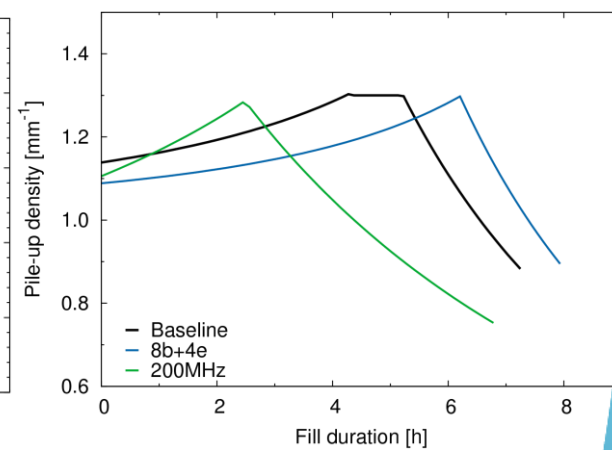
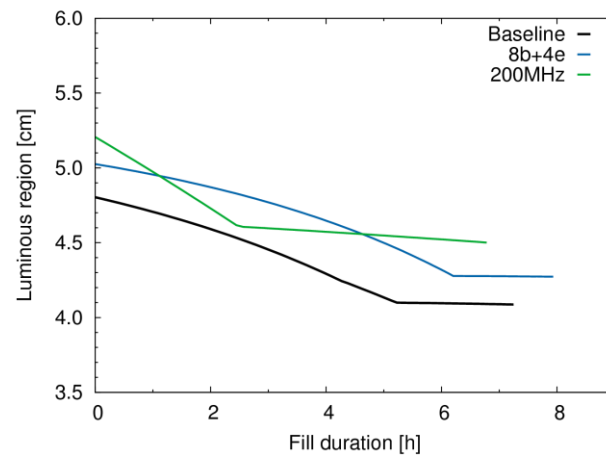
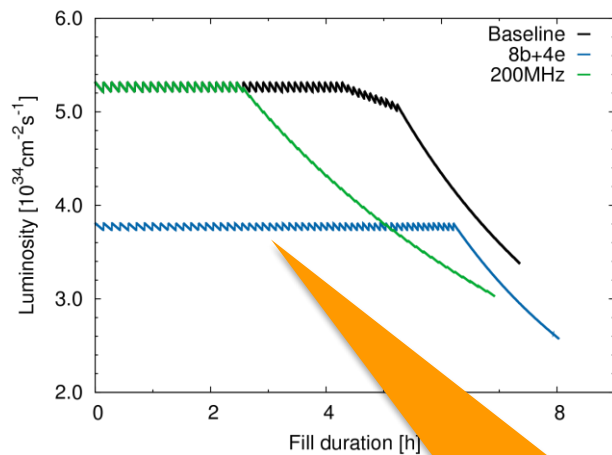


- In case scrubbing cannot reduce $SEY_{dip} = SEY_{quad}$ below 1.40 \rightarrow Heat load exceeds the available cooling capacity for 400 MHz system
- Long bunches provided by 200 MHz system would allow operating within the cryo limits
- Crab cavities less efficient \rightarrow stronger effect of the RF curvature
- **Requires new HW – not in baseline**



Luminous region/performance 8b+4e and 200MHz

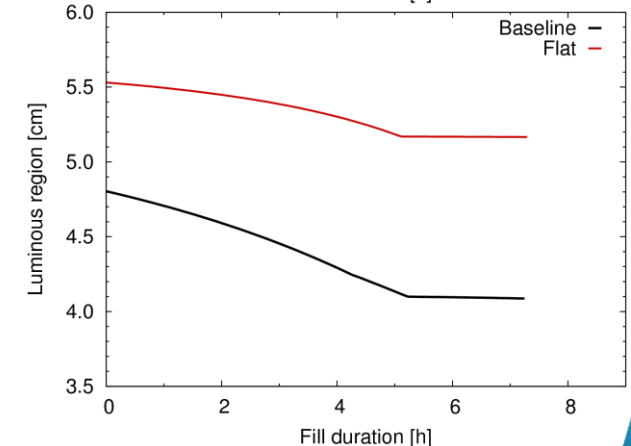
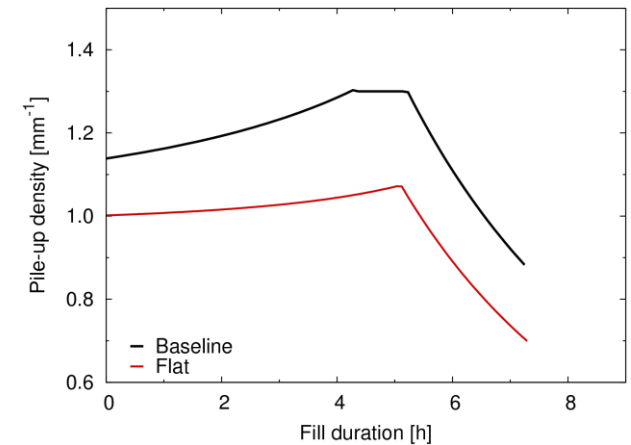
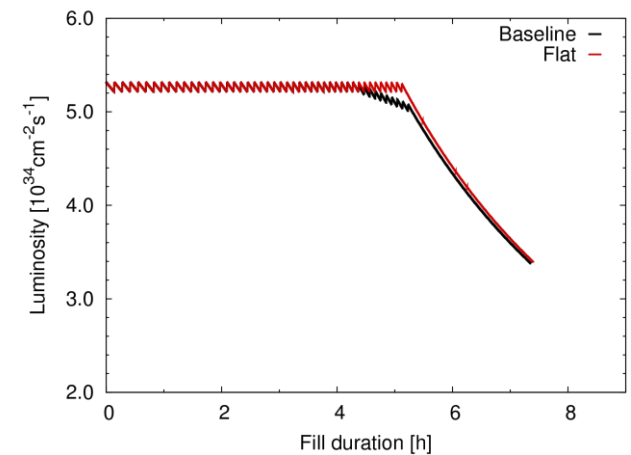
- Scenarios for mitigation of electron cloud effects:
 - 8b+4e → -25% integrated luminosity wrt baseline
 - 200 MHz → -14% integrated luminosity wrt baseline



8b+4e: Lower number of bunches – 5% larger bunch population

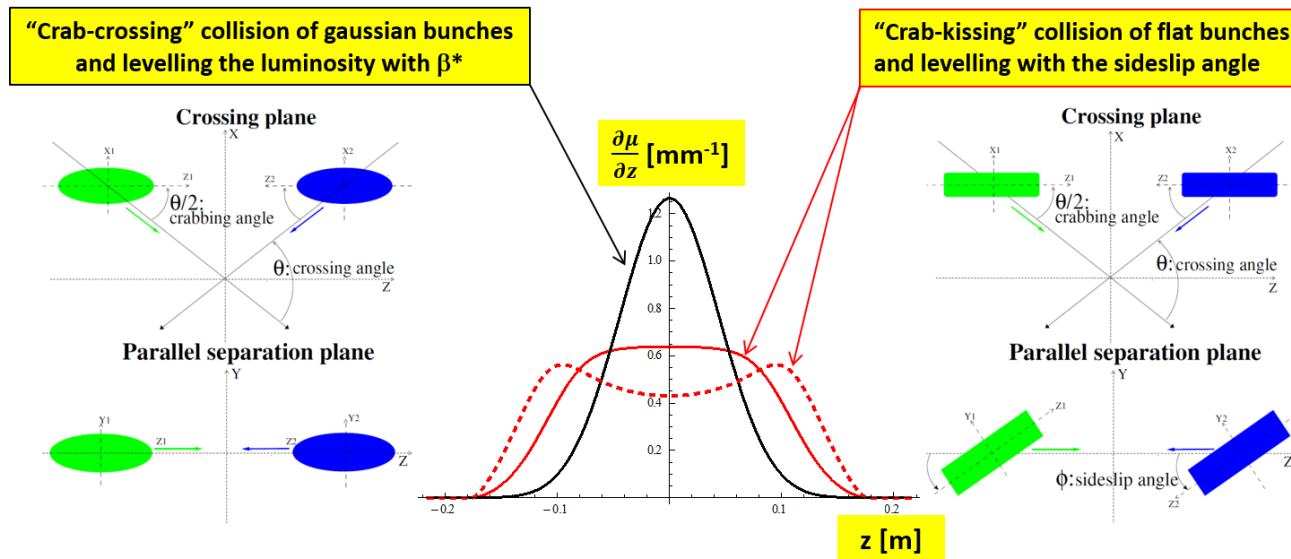
Flat Optics

- Flat optics with $\beta^* = 40 \text{ cm} / 15 \text{ cm}$ and 12σ beam separation → pushed scenario to be proven by experience with flat optics (not baseline). Might require beam-beam long range compensation.
- Lower peak pile-up density w.r.t. baseline
- Integrated luminosity comparable to nominal but reduction of the peak pile-up density



Crab Kissing

- Scenario studied (S. Fartoukh) to reduce pile-up density. Requires:
 - Crab-cavities in both $||$ and X-sing planes (levelling with both sets of CCs)
 - Flat optics
 - Bunch charge distribution as flat as possible for more PU density flatness (800 MHz or longitudinal profile shaping by RF noise)
 - Requires additional HW (not in baseline)**



- Provides the lowest peak pile-up density of ~ 0.6 - 0.65 events/xing/mm with little reduction on integrated luminosity ($\sim 5\%$)

Summary Table

Scenario	Max. PU [events/xing]	Max. PU density [events /xing/mm]	Luminous Region long. r.m.s. size [cm]	$\Delta L_{\text{int}}/L_{\text{nom}}$ [%]	Additional HW [Y/N]
Nominal	140	1.3	<4.8	0	N
Ultimate	200	1.9	<4.5	+33	N
8b+4e	140	1.3	<5	-25	N
200 MHz	140	1.3	<4.8	-14	Y
Flat	140	1.1	<5.5	0	May be
Crab kissing	140	0.6-0.65	<7	-5	Y

Summary

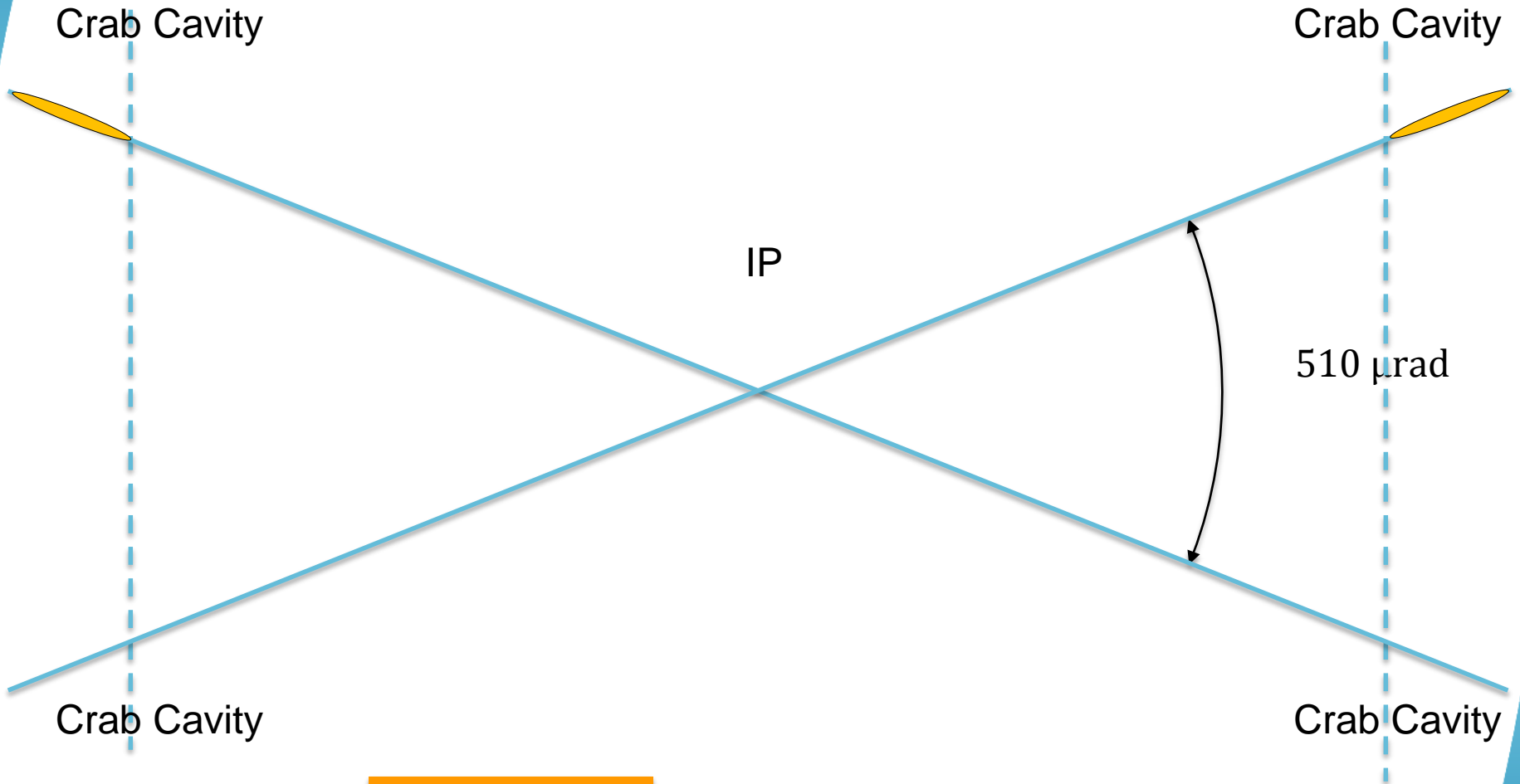
- An overview of the parameters affecting luminous region has been sketched
- The range of parameters for the nominal and ultimate parameters has been given
- A few **non baseline** scenarios have been sketched with the aim of providing the luminous region parameter space for the estimate of the performance of the detectors
- **Basis for understanding detector performance and devise improved scenarios**



Thank you for your attention!

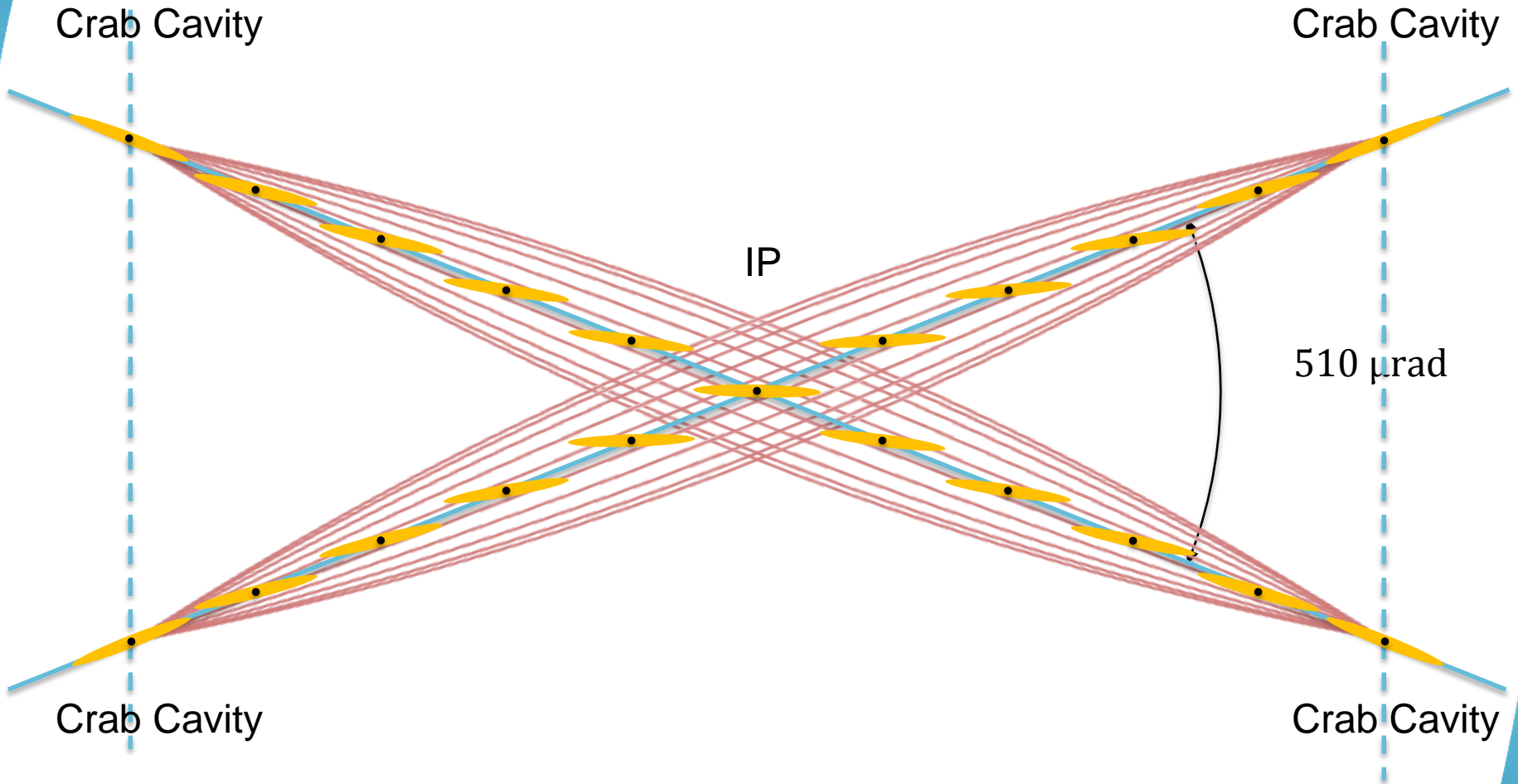
Acknowledgements: EDQ Working Group Members

Other features

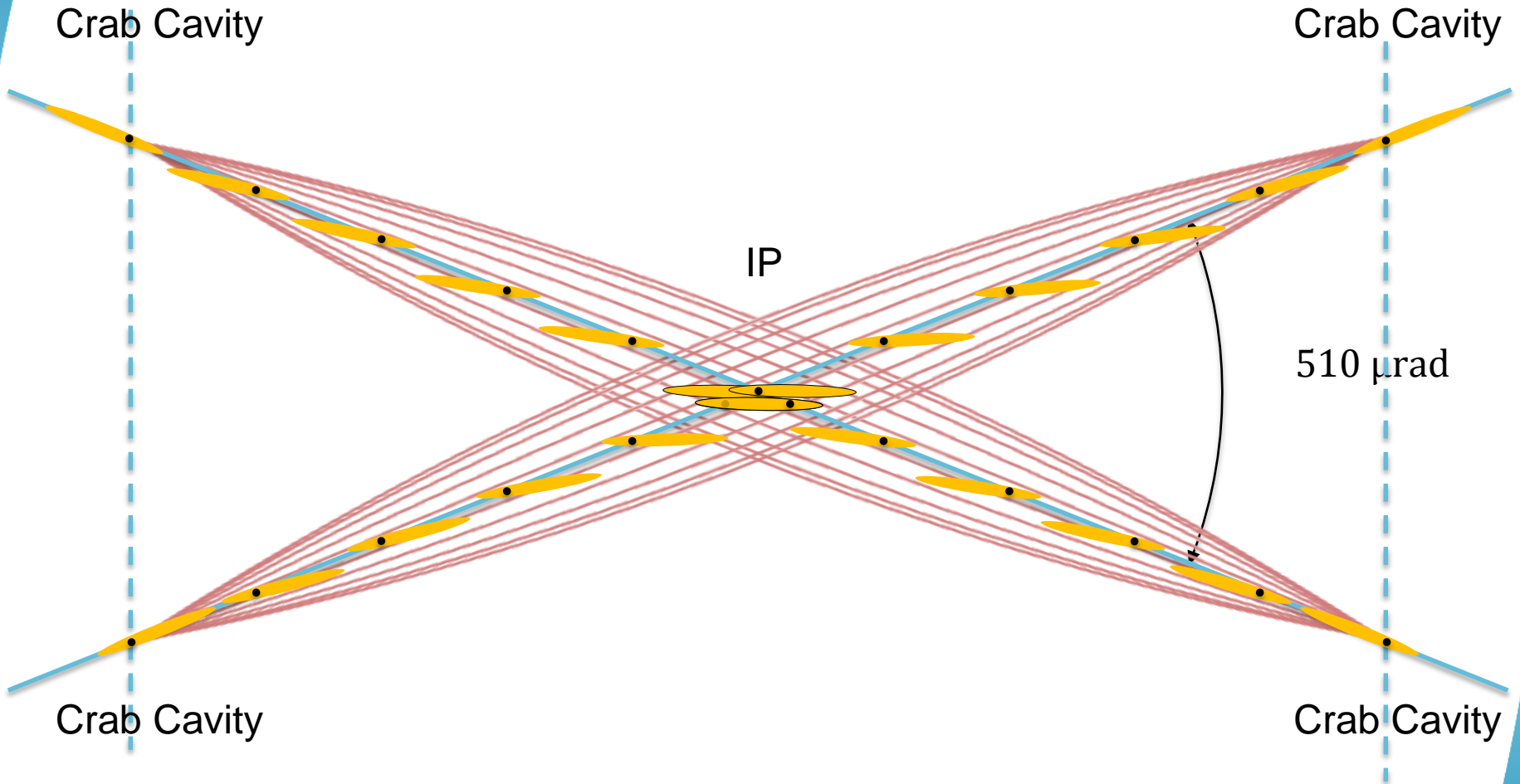


E. Jensen

Other features



Other features



Baseline

