The background of the slide is a dark grid of small, colorful dots in shades of purple, blue, and cyan. The dots are arranged in a regular pattern, with some dots appearing brighter or more saturated than others, creating a subtle, shimmering effect.

J-PARC E61 experiment

Feb. 21, 2018

Tomoyo Yoshida
(Tokyo Institute of Technology)

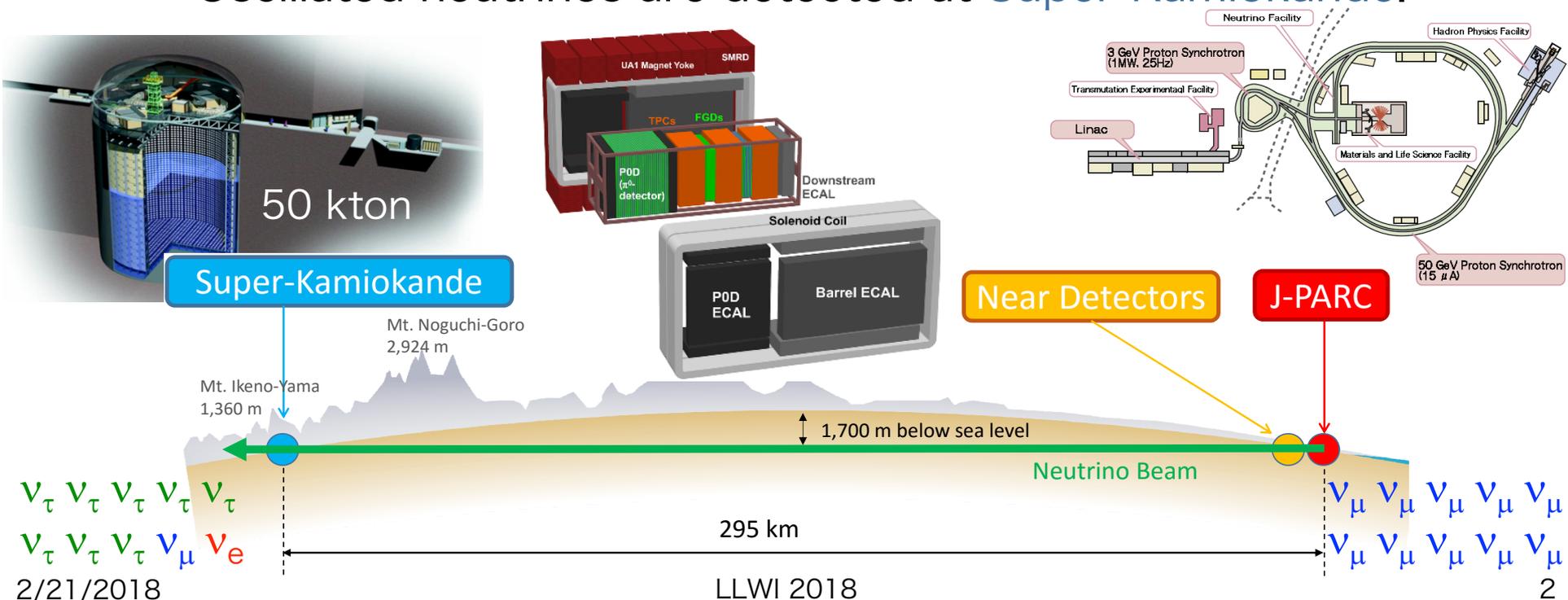
For J-PARC E61 collaboration

Lake Louise Winter Institute

T2K experiment

One of the world-leading neutrino oscillation experiments

- (Anti-) muon neutrino beam is generated at **J-PARC**.
- Beam property and neutrino interaction are measured at **near detectors**.
- Oscillated neutrinos are detected at **Super-Kamiokande**.

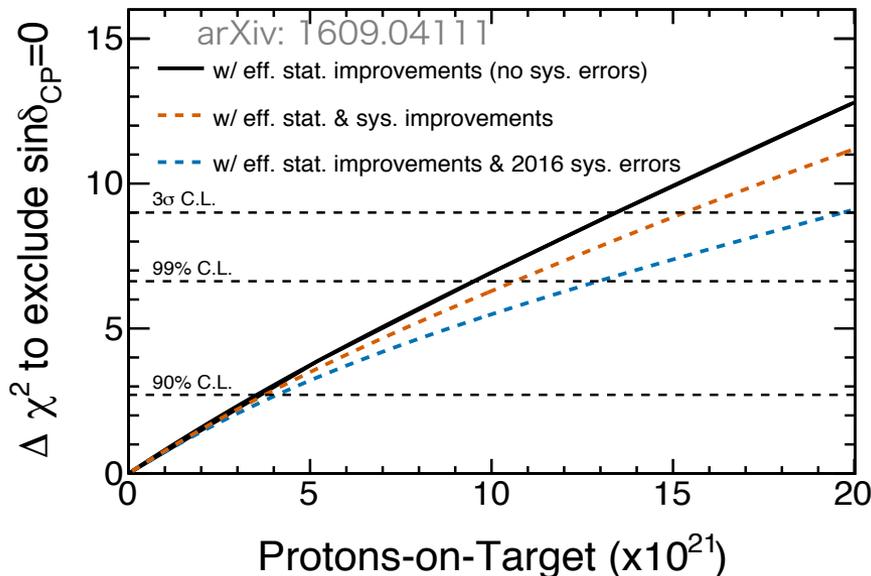


Future LBL experiments in Japan

T2K-II (~2025)

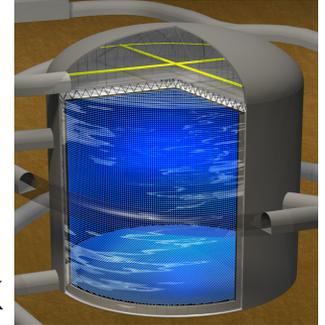
Proposed T2K extension
from original 7.8×10^{21} POT
to 20×10^{21} POT

Exclude CP conservation
at 3σ at maximal



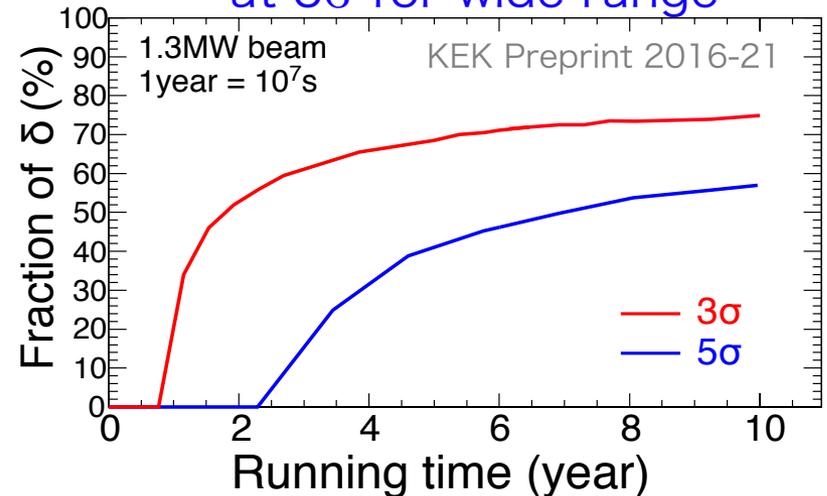
Hyper-Kamiokande (2026-)

Proposed
water Cherenkov
detector



260 kton / tank

Exclude CP conservation
at 5σ for wide range

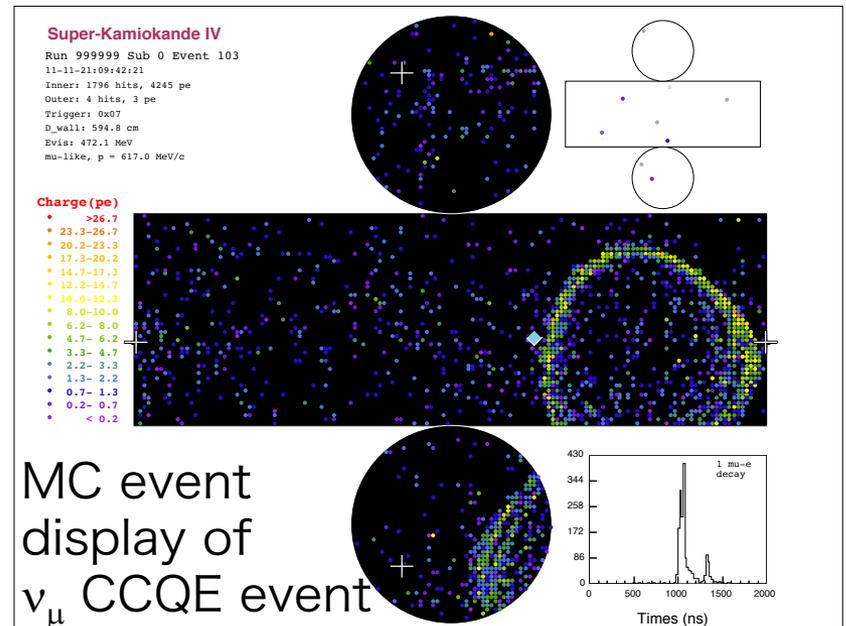
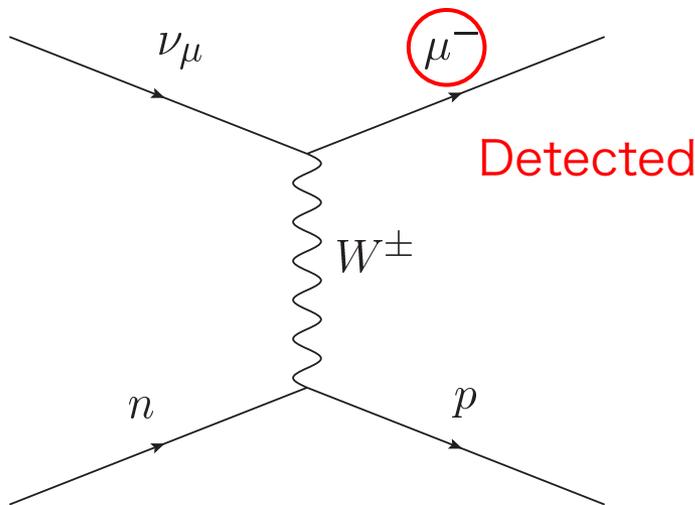


Long-baseline experiments will enter the precision era.

Neutrino energy reconstruction

Energy reconstruction is crucial since the oscillation probability depends on neutrino energy.

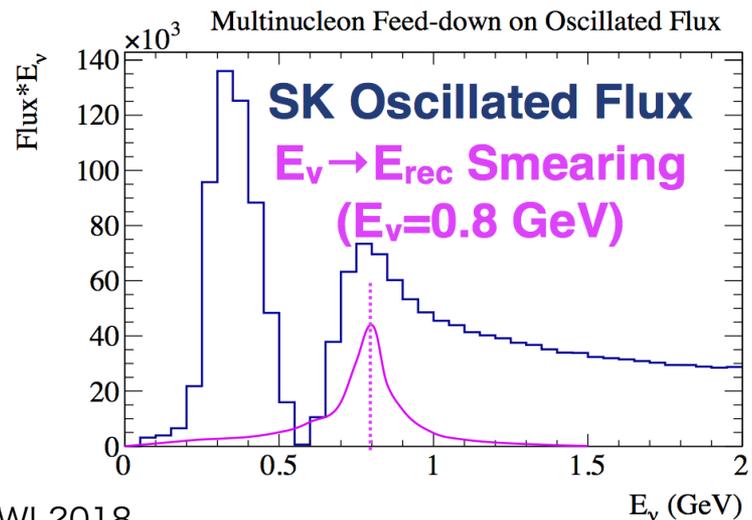
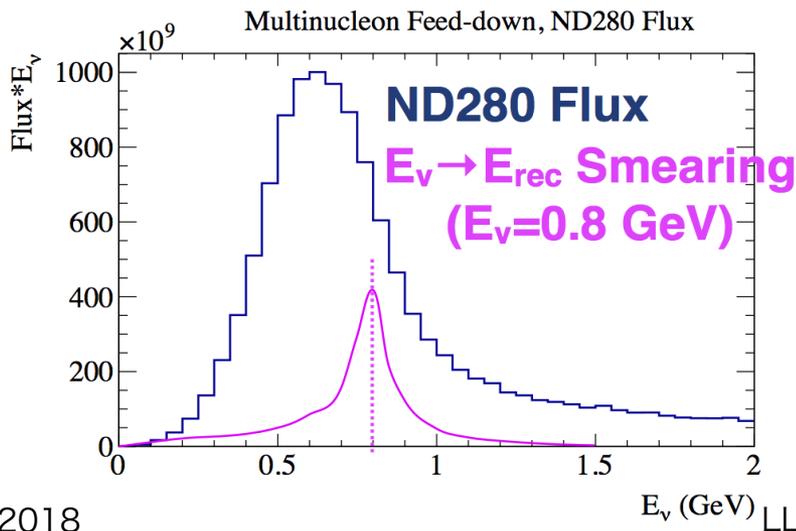
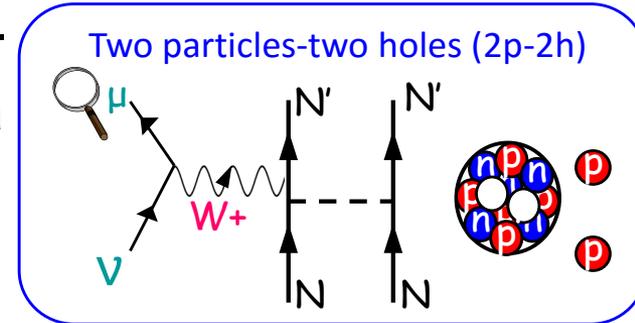
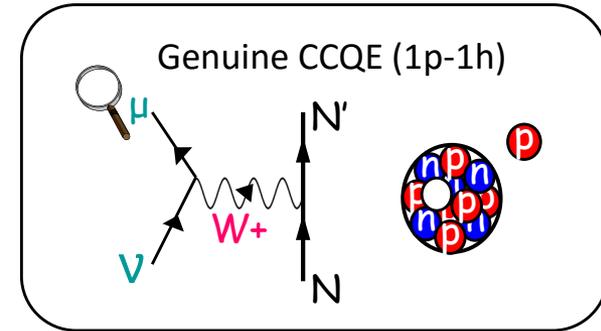
- **Single ring** events are selected at Super-K.
- Neutrino energy is reconstructed from momentum and direction of the **charged lepton**, assuming charged current quasi-elastic (**CCQE**) kinematics.



Interaction model uncertainties

Reducing model uncertainties will be more and more important in future.

- **Multi-nucleon interaction** components can **bias** neutrino energy reconstruction.
- Although this smearing may not be seen significantly at near detectors, the bias might be crucial at far detector, as neutrino oscillate.

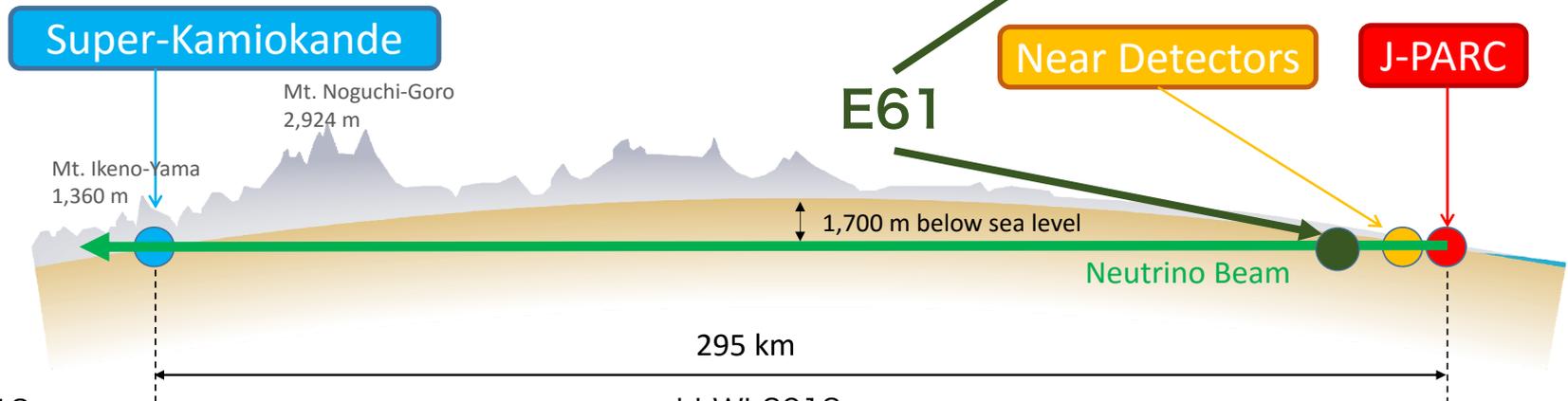
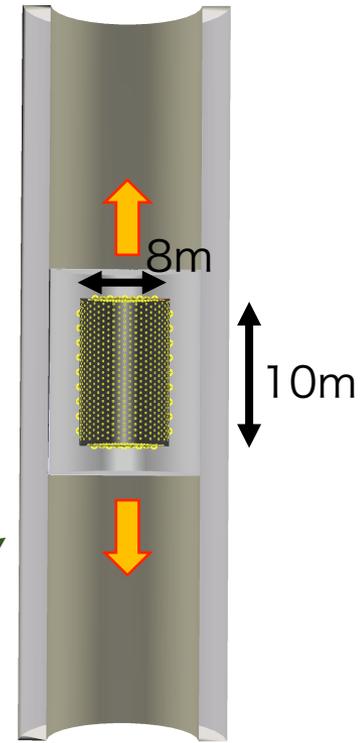


arXiv:
1611.07770

J-PARC E61 detector proposal

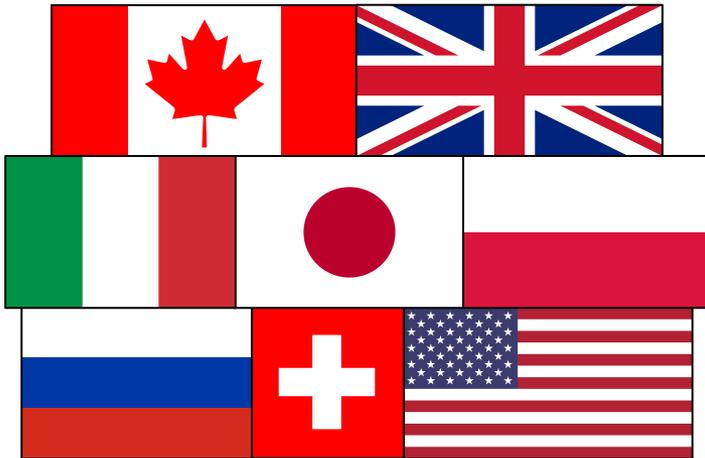
An **intermediate water Cherenkov** detector at 1 – 2 km downstream of J-PARC neutrino beam

- An instrumented volume **moves vertically** within a 50 m tall water pit
- Optically separated inner and outer detectors



J-PARC E61 collaboration

- 2 collaborations, NuPRISM and TITUS merged.
- 110 collaborators from 35 institutes in 8 countries as of Jan. 2018



A. Abramowski,³² A. Ajmi,¹⁰ M. Barbi,²² G. Barker,³³ L. Berns,²⁷ S. Bhadra,³⁵ J. Bian,³ A. Blondel,⁵ S. Boyd,³³ A. Bravar,⁵ C. Bronner,²⁵ A. Buchowicz,³² G. Catanesi,⁸ C. Checchia,¹⁰ J. Coleman,¹⁶ G. Collazuol,¹⁰ G. Cowan,⁴ G. De Rosa,⁹ T. Dealtry,¹⁵ F. Di Lodovico,²¹ E. Drakopoulou,⁴ T. Feusels,² A. Finch,¹⁵ G.A. Fiorentini,³⁵ G. Fiorillo,⁹ D. Grant,¹ D. Hadley,³³ L. Haegel,⁵ M. Hartz,^{13,31} N. Hastings,²² R. Helmer,³¹ R. Henderson,³¹ T. Ishida,^{6,*} M. Ishitsuka,²⁹ B. Jamieson,³⁴ H. Kakuno,²⁸ T. Katori,²¹ M. Khabibullin,¹² A. Khotjantsev,¹² S. King,²¹ A. Konaka,³¹ C. Kopfer,¹ L. Kormos,¹⁵ K. Kowalik,¹⁸ Y. Kudenko,^{12,†} R. Kurjata,³² M. Kuze,²⁷ M. La Posta,²² J. Lagoda,¹⁸ M. Laveder,¹⁰ T. Lindner,³¹ A. Longhin,¹⁰ P. Loverre,¹¹ X. Lu,²⁰ P. Lu,³¹ L. Ludovici,¹¹ K. Mahn,¹⁷ M. Malek,²³ L. Maret,⁵ J. Martin,³⁰ J. Marzec,³² N. McCauley,¹⁶ A. Mefodiev,¹² C. Metelko,¹⁶ M. Mezzetto,¹⁰ O. Mineev,¹² S. Nakayama,^{25,‡} M. Needham,⁴ Y. Nishimura,²⁶ F. Nova,²⁴ H. O’Keeffe,¹⁵ A. Owen,²¹ G. Pastuszek,³² C. Pidcott,²³ E. Pinzon Guerra,³⁵ S. Playfer,⁴ A. Pritchard,¹⁶ N. Prouse,²¹ B. Quilain,¹⁴ C. Riccio,⁹ B. Richards,²¹ E. Rondio,¹⁸ B. Rossi,⁹ A.C. Ruggeri,⁹ A. Rychter,³² M. Scott,³¹ T. Sekiguchi,^{6,*} M. Smy,³ H. Sobel,³ T. Sumiyoshi,²⁸ S. Suvorov,¹² M. Taani,⁴ R. Tacik,²² H.K. Tanaka,^{25,‡} H.A. Tanaka,^{30,§} L.F. Thompson,²³ T. Towstego,³⁰ C. Vilela,¹⁹ V. Volkov,¹² J. Walker,³⁴ M. Wascko,⁷ M. Wilking,¹⁹ J. Wilson,²¹ N. Yershov,¹² T. Yoshida,²⁷ J. Zalipska,¹⁸ K. Zaremba,³² G. Zarnecki,¹⁸ M. Ziembicki,³² and S. Zsoldos²¹

(The E61 Collaboration)

¹ University of Alberta, Centre for Particle Physics, Department of Physics, Edmonton, Alberta, Canada

² University of British Columbia, Department of Physics and Astronomy, Vancouver, British Columbia, Canada

³ University of California, Irvine, Department of Physics and Astronomy, Irvine, California, U.S.A.

⁴ University of Edinburgh, School of Physics and Astronomy, Edinburgh, United Kingdom

⁵ University of Geneva, Section de Physique, DPNC, Geneva, Switzerland

⁶ High Energy Accelerator Research Organization (KEK), Tsukuba, Ibaraki, Japan

⁷ Imperial College London, Department of Physics, London, United Kingdom

⁸ INFN Sezione di Bari and Università e Politecnico di Bari, Dipartimento Interuniversitario di Fisica, Bari, Italy

⁹ INFN Sezione di Napoli and Università di Napoli, Dipartimento di Fisica, Napoli, Italy

¹⁰ INFN Sezione di Padova and Università di Padova, Dipartimento di Fisica, Padova, Italy

¹¹ INFN Sezione di Roma and Università di Roma “La Sapienza”, Roma, Italy

¹² Institute for Nuclear Research of the Russian Academy of Sciences, Moscow, Russia

¹³ Kavli Institute for the Physics and Mathematics of the Universe (WPI),

Today Institutes for Advanced Study, University of Tokyo, Kashiwa, Chiba, Japan

¹⁴ Kyoto University, Department of Physics, Kyoto, Japan

¹⁵ Lancaster University, Physics Department, Lancaster, United Kingdom

¹⁶ University of Liverpool, Department of Physics, Liverpool, United Kingdom

¹⁷ Michigan State University, Department of Physics and Astronomy, East Lansing, Michigan, U.S.A.

¹⁸ National Centre for Nuclear Research, Warsaw, Poland

¹⁹ State University of New York at Stony Brook, Department of Physics and Astronomy, Stony Brook, New York, U.S.A.

²⁰ Oxford University, Department of Physics, Oxford, United Kingdom

²¹ Queen Mary University of London, School of Physics and Astronomy, London, United Kingdom

²² University of Regina, Department of Physics, Regina, Saskatchewan, Canada

²³ University of Sheffield, Department of Physics and Astronomy, Sheffield, United Kingdom

²⁴ STFC, Rutherford Appleton Laboratory, Harwell Oxford, and Daresbury Laboratory, Warrington, United Kingdom

²⁵ University of Tokyo, Institute for Cosmic Ray Research, Kamioka Observatory, Kamioka, Japan

²⁶ University of Tokyo, Institute for Cosmic Ray Research, Research Center for Cosmic Neutrinos, Kashiwa, Japan

²⁷ Tokyo Institute of Technology, Department of Physics, Tokyo, Japan

²⁸ Tokyo Metropolitan University, Department of Physics, Tokyo, Japan

²⁹ Tokyo University of Science, Department of Physics, Noda, Japan

³⁰ University of Toronto, Department of Physics, Toronto, Ontario, Canada

³¹ TRIUMF, Vancouver, British Columbia, Canada

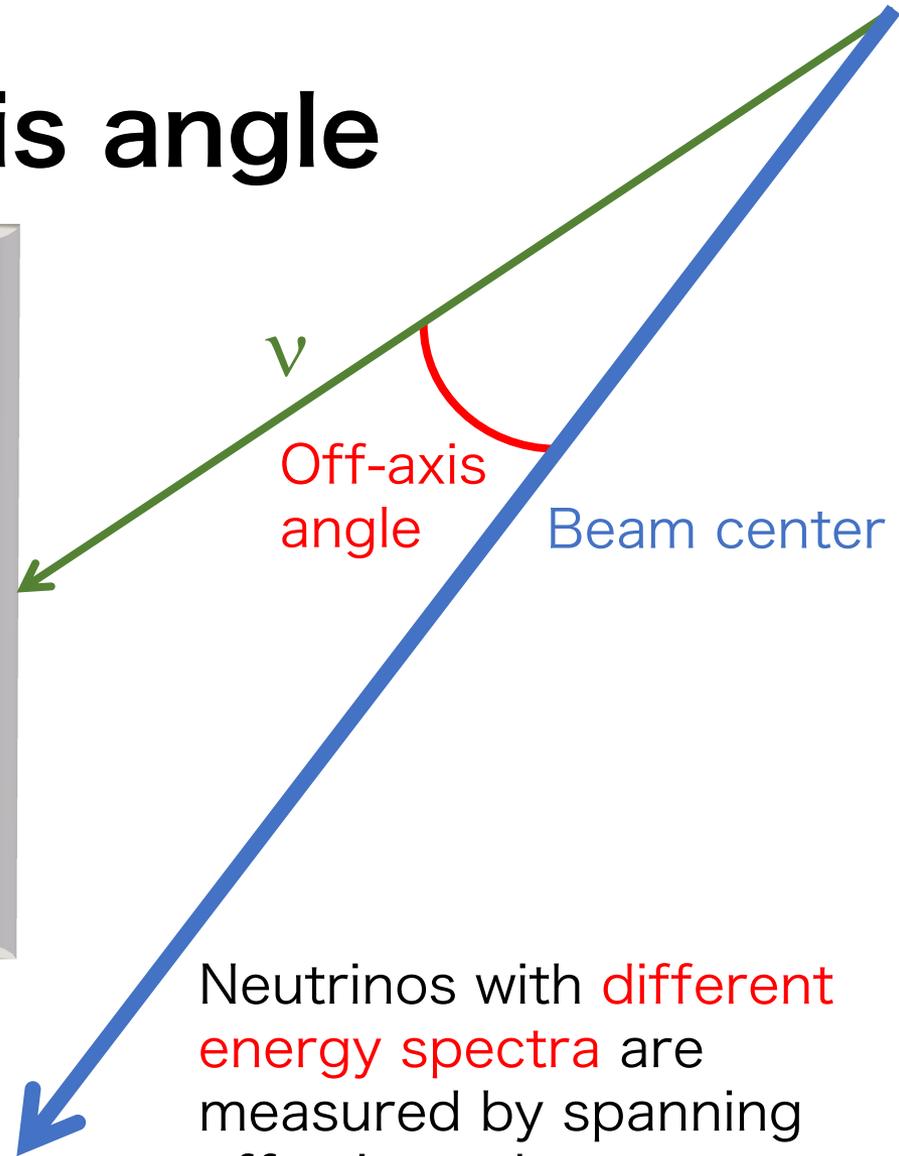
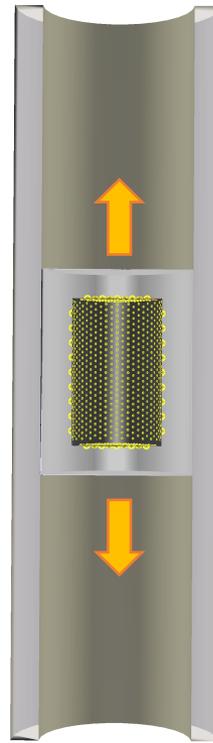
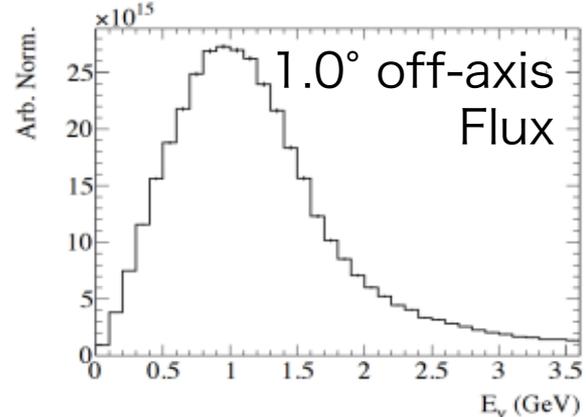
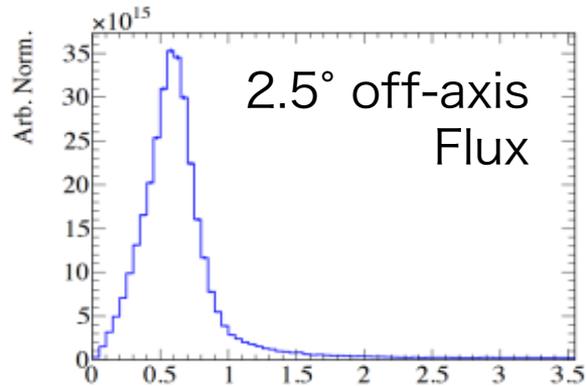
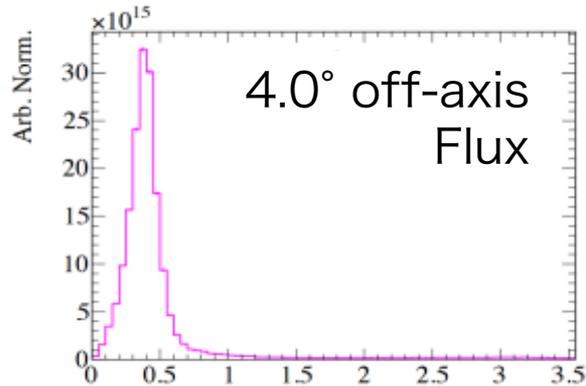
³² Warsaw University of Technology, Institute of Radioelectronics and Multimedia Technology, Warsaw, Poland

³³ University of Warwick, Department of Physics, Coventry, United Kingdom

³⁴ University of Winnipeg, Department of Physics, Winnipeg, Manitoba, Canada

³⁵ York University, Department of Physics and Astronomy, Toronto, Ontario, Canada

Spanning off-axis angle

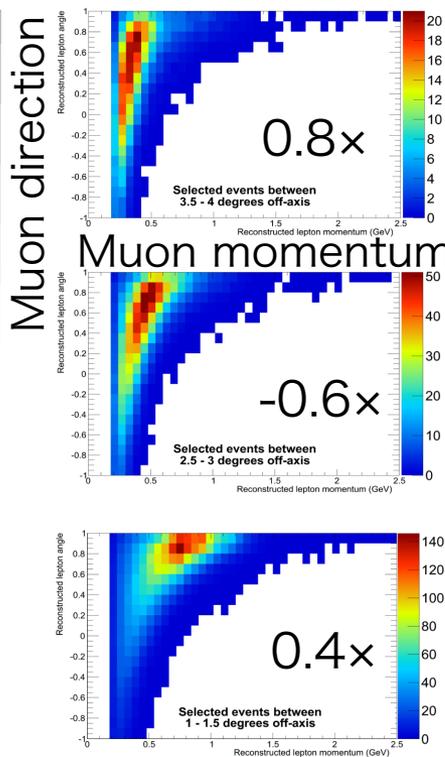
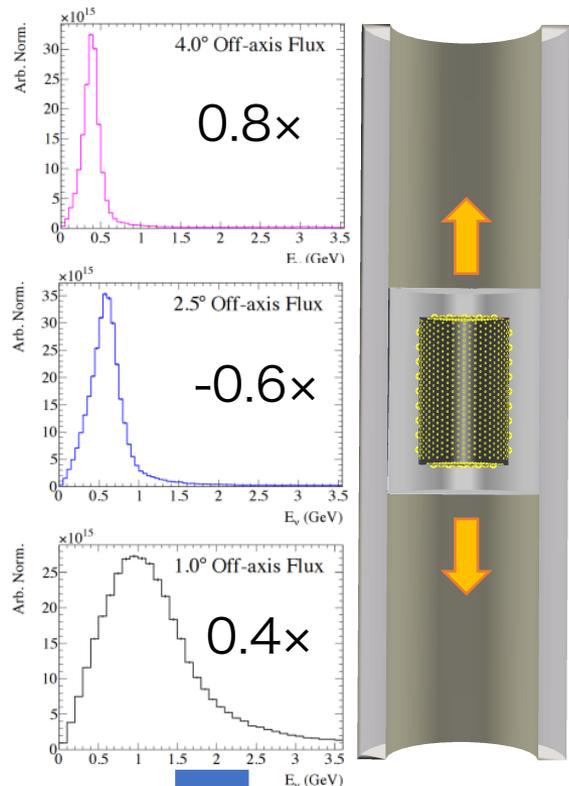


Neutrinos with **different energy spectra** are measured by spanning off-axis angle

$$E_\nu = \frac{m_\pi^2 - m_\mu^2}{2E_\pi(1 - \beta_\pi \cos \theta)}$$

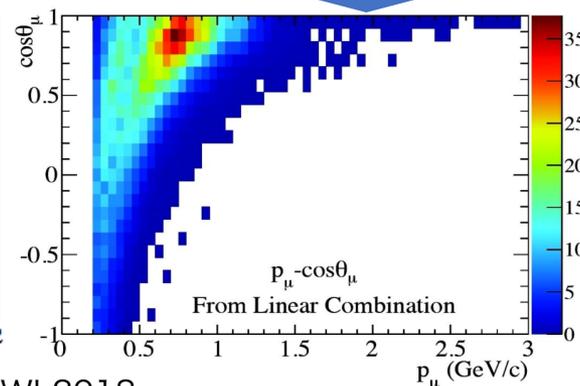
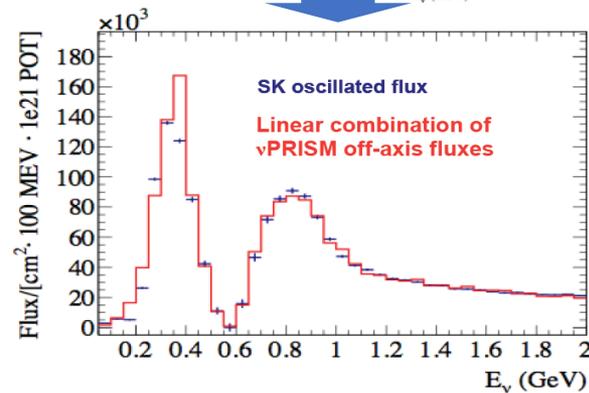
Linear combination of off-axis bins

1. Separate detector volume to 30 off-axis slices.
2. Take a linear combination of off-axis slices to reproduce desired spectrum.



3. Take the same linear combination of observed variables to expect their distributions corresponding to that spectrum.

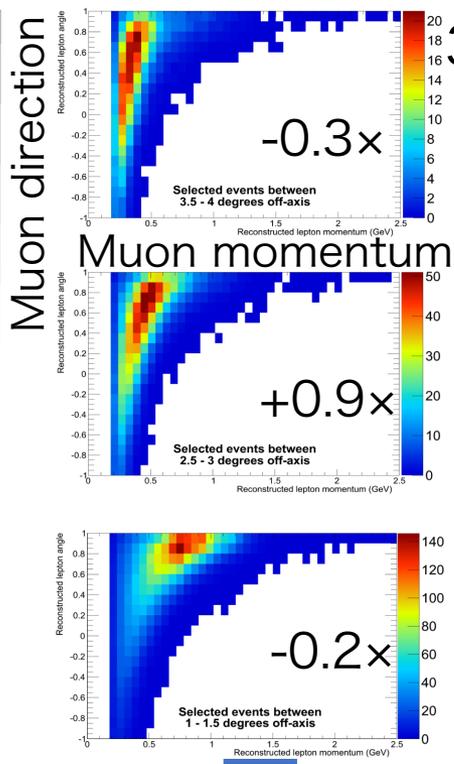
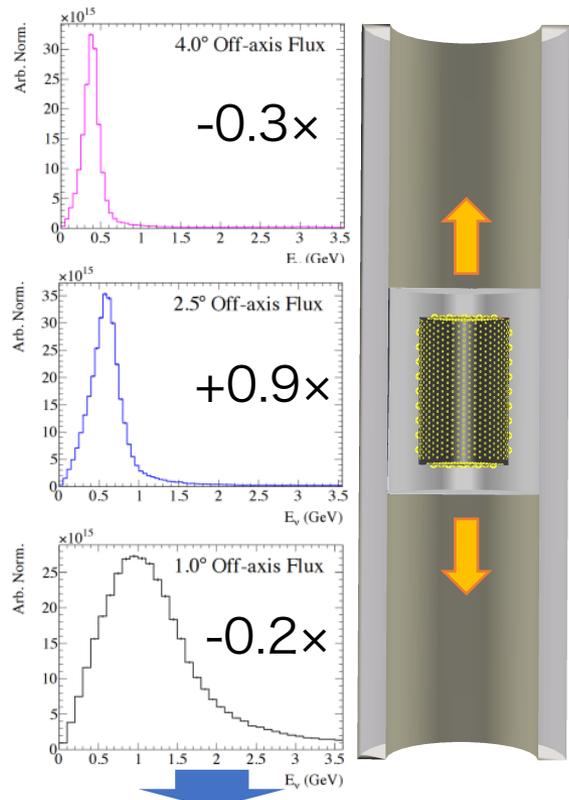
Example 1:
Oscillated far detector spectrum



Enables data-to-data fit to extract oscillation parameters, independent of interaction models

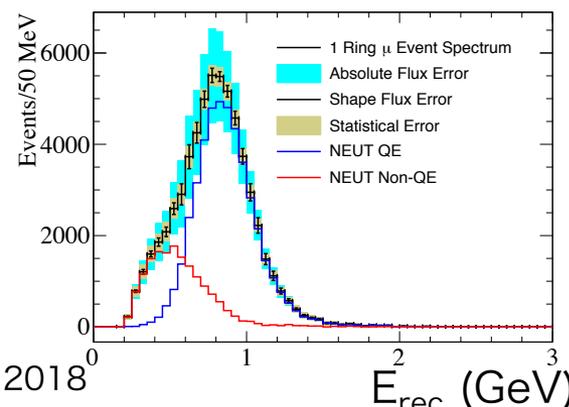
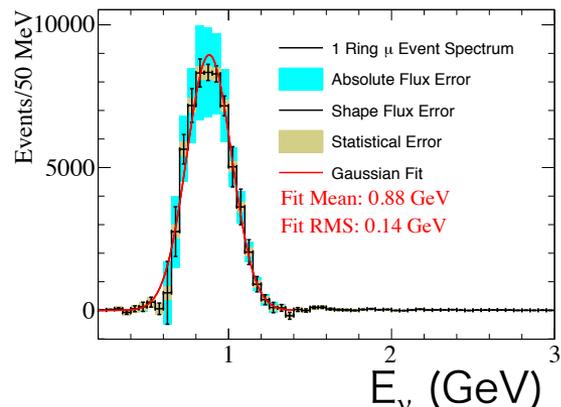
Linear combination of off-axis bins

1. Separate detector volume to 30 off-axis slices.
2. Take a linear combination of off-axis slices to reproduce desired spectrum.



3. Take the same linear combination of observed variables to expect their distributions corresponding to that spectrum.

Example 2: Gaussian

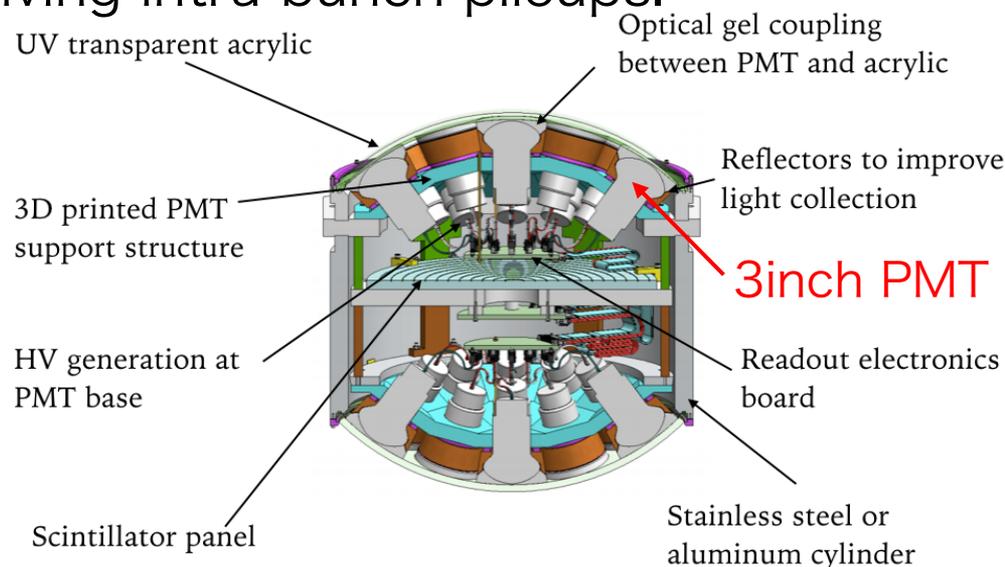


Can study neutrino interaction using pseudo-monochromatic beam

Multi-PMT developments

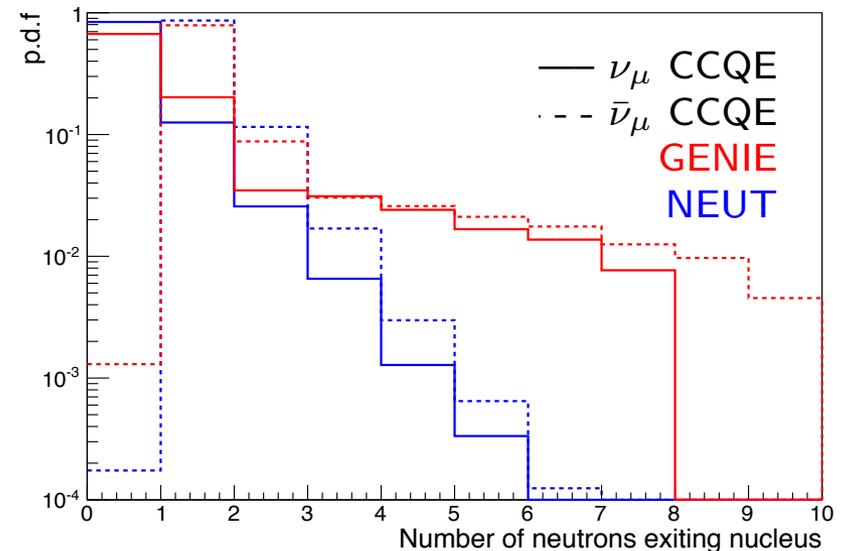
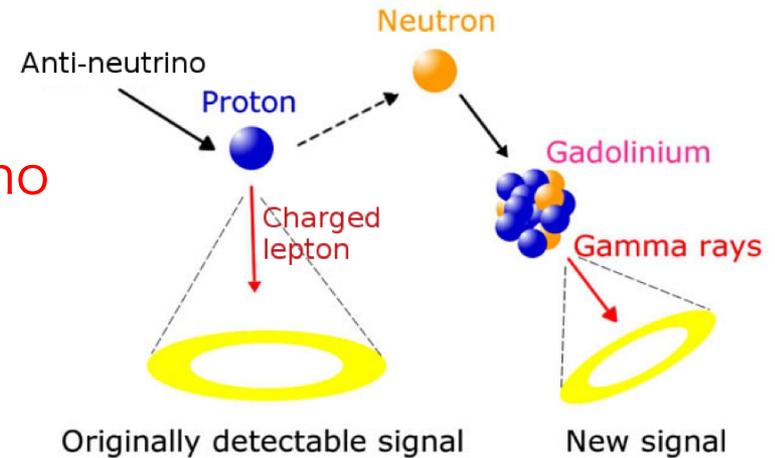
The E61 baseline design is equipped with multi-PMT modules with 27% of photo-cathode coverage.

- **Fine granularity** of Cherenkov ring images improves event reconstruction.
- **Timing resolution** is improved by using smaller PMTs, which is important for resolving intra-bunch pileups.
- R&D is ongoing.
 - PMT, acrylic, gel, ...
 - Electronics
 - Prototype production
- Geant4-based full detector simulation is developed.
- Event reconstruction development is underway.



Gd loading option

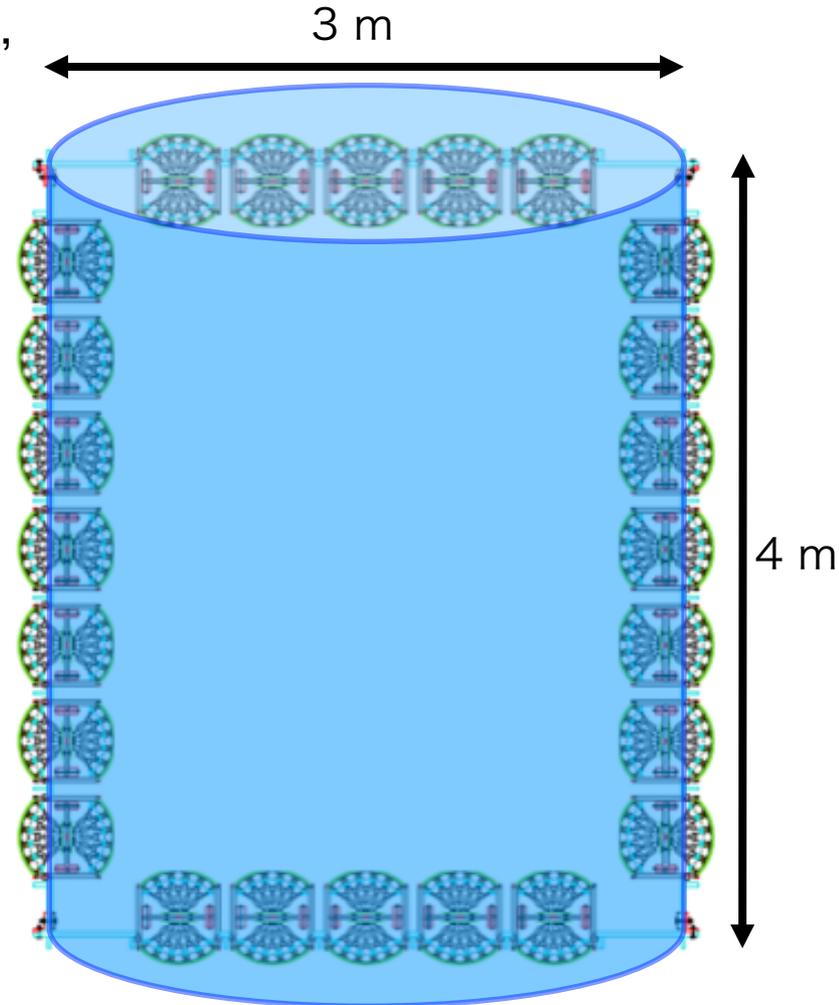
- Super-Kamiokande plans to load Gd in the water.
 - Statistical **neutrino/anti-neutrino separation** would be possible.
- However, there is no measurement of neutrino-induced neutron capture by Gd in water Cherenkov detector.
- Gd loading to E61 detector would help **constraining** neutron emission rates and tagging efficiency.



Staged approach — Beam test

- Before constructing full detector, **small-scale initial-phase detector** is considered to confirm the **detector performance** and to **calibrate** the detector precisely by using charged particle beam.
- We are investigating test beam facility at Fermilab.


 e, μ, π, p



Prospects

- We aim to take neutrino beam data before Hyper-Kamiokande operation starts.
- E61 received J-PARC Stage-1 status in 2016.
 - Scientific merit is recognized.
 - Developing concrete design is required for construction.



Year	2017				2018				2019				2020				2021				2022				2023				2024				2025				2026			
Quarter	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4	1	2	3	4
Test Experiment Design																																								
Test Experiment Construction																																								
Test Experiment Operation																																								
Full Detector Design																																								
Full Detector Construction																																								
Full Detector Operation																																								

Summary

- Reducing **neutrino interaction model uncertainties** will be more and more important in future oscillation experiments.
- J-PARC E61 experiment aims to give a **data-driven method** to convert observable final state kinematics to neutrino energy.
- R&D of **multi-PMT** module is ongoing, with initial prototypes in production.
- **Analysis scheme** based on multi-PMT is being developed using full detector simulation and event reconstruction algorithm.
- The construction of an **initial phase of the detector** has been proposed to ensure detector performance.

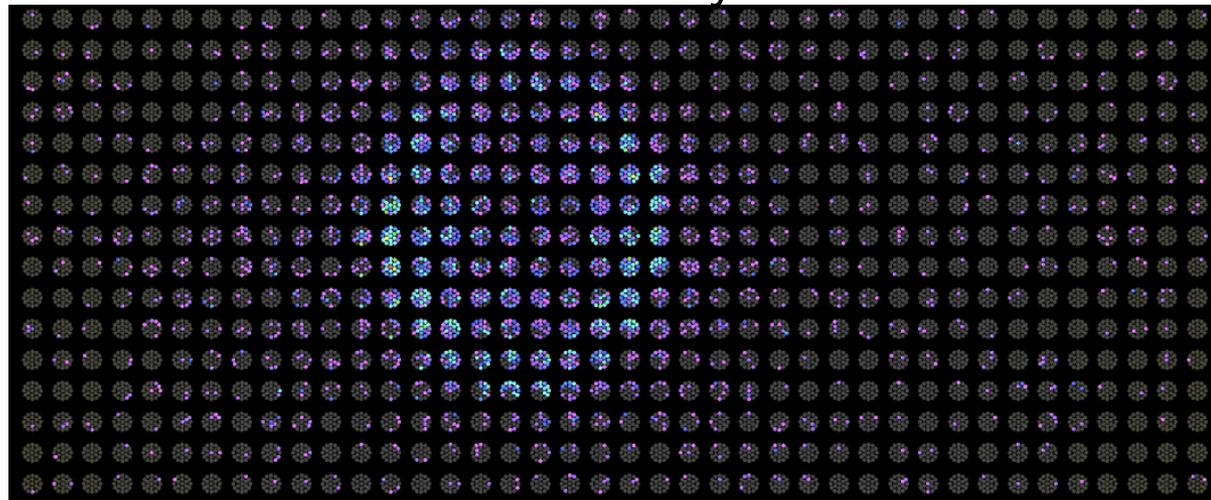
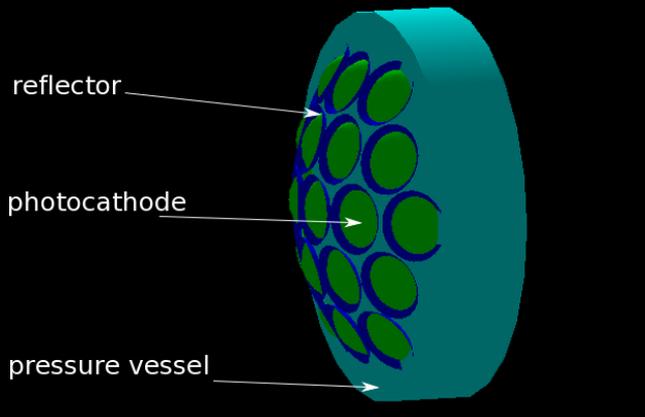
Back up

Software development

Detector simulation and event reconstruction algorithm are developed to study detector optimization and physics sensitivities.

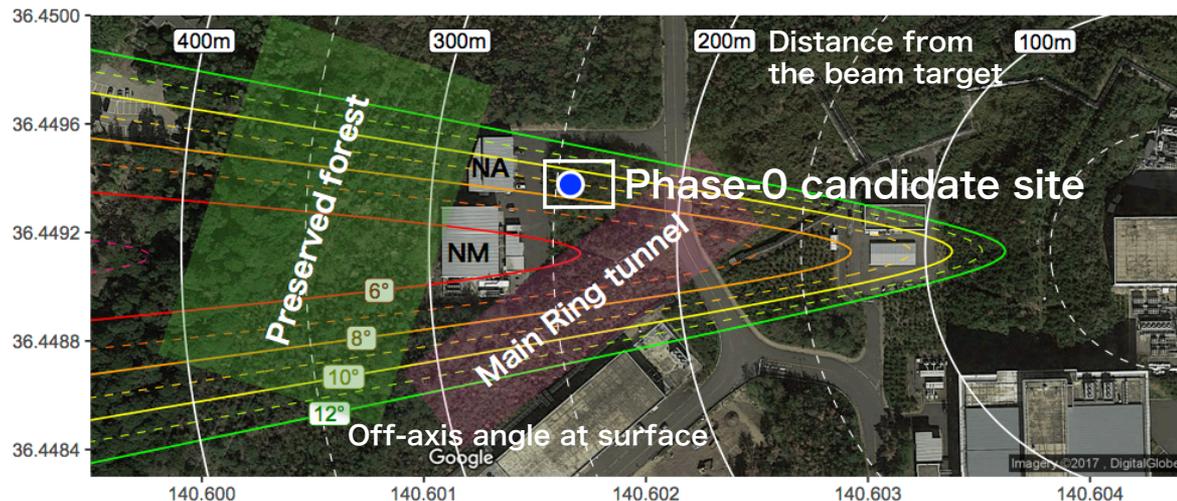
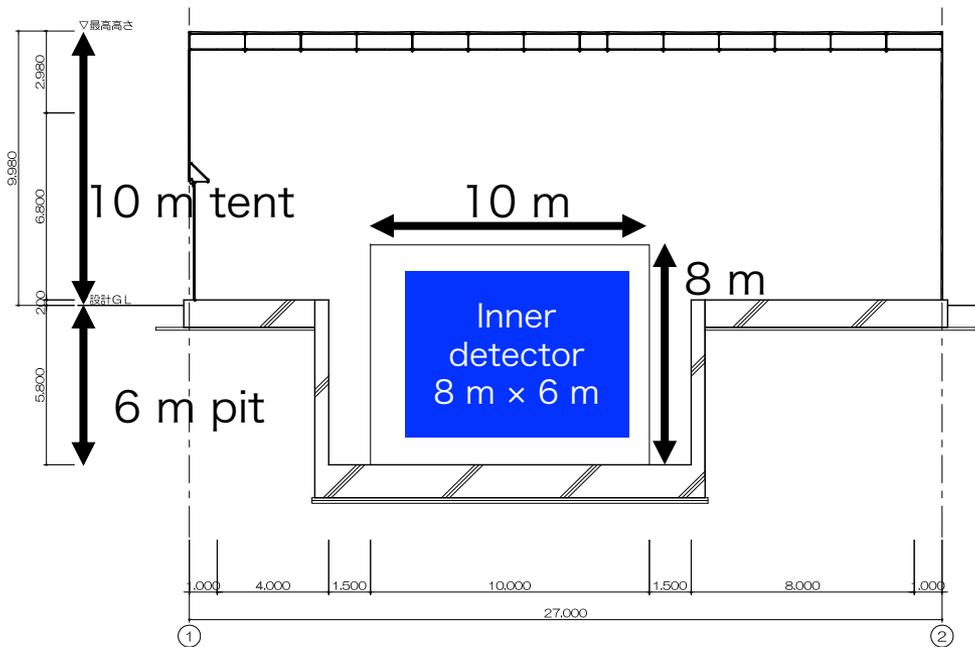
- Development of Geant4-based detector simulation is in cooperation with Hyper-Kamiokande collaboration. Multi-PMT is implemented.
- Event reconstruction is based on the algorithm used in T2K. Implementation of multi-PMT is underway.

inner detector half of nuPRISM mPMT



Staged approaches — Phase-0

- To demonstrate detector performance and calibration methods by a surface detector
- ν_e cross section and neutron multiplicity measurements are expected.
- Most detector components will be reused in Phase-1 operation inside the 50m-deep pit.



T2K systematics

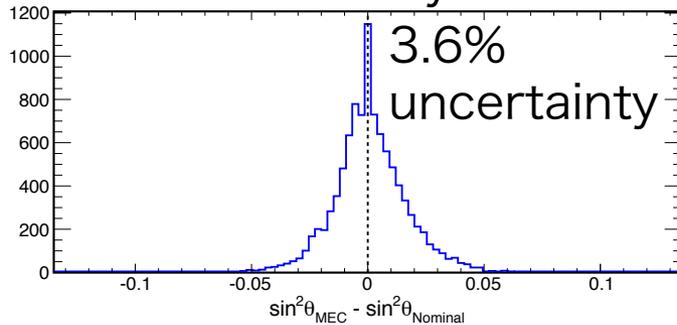
Current T2K systematic errors taken from PRD 96, 092006 (2017)

	$\delta N_{\text{SK}}/N_{\text{SK}}$ [%]			
	1-ring μ		1-ring e	
	ν mode	$\bar{\nu}$ mode	ν mode	$\bar{\nu}$ mode
Flux + Cross section	2.9	3.5	4.2	4.7
Final state interaction + secondary interaction + photonuclear effect at SK	1.5	2.1	2.5	3.0
SK detector	3.9	3.4	2.4	2.5
Total	5.1	5.3	5.5	6.5

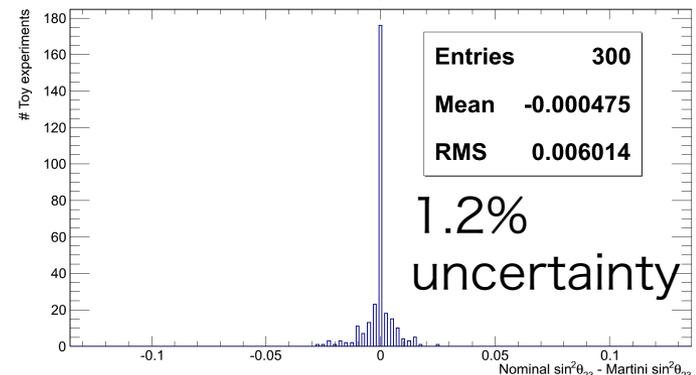
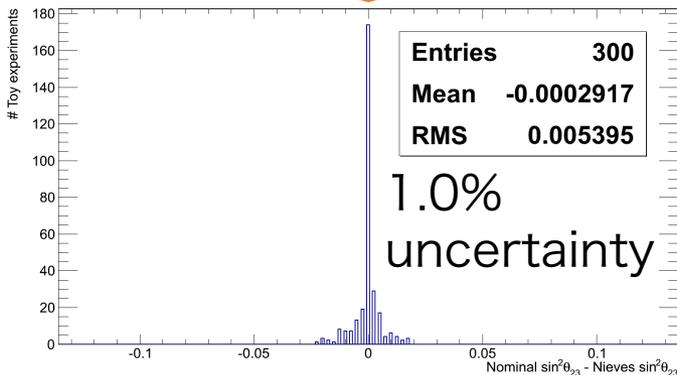
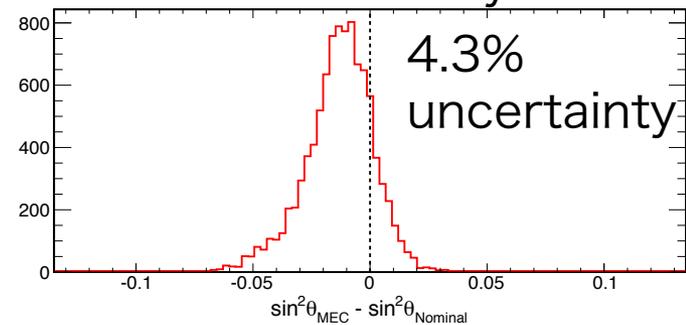
E61 contributions to ν_μ disappearance analysis

- E61 analysis is less sensitive to neutrino interaction models

Nieves model fitted by nominal model



Martini model fitted by nominal model



- Most of flux uncertainties cancel between E61 and SK

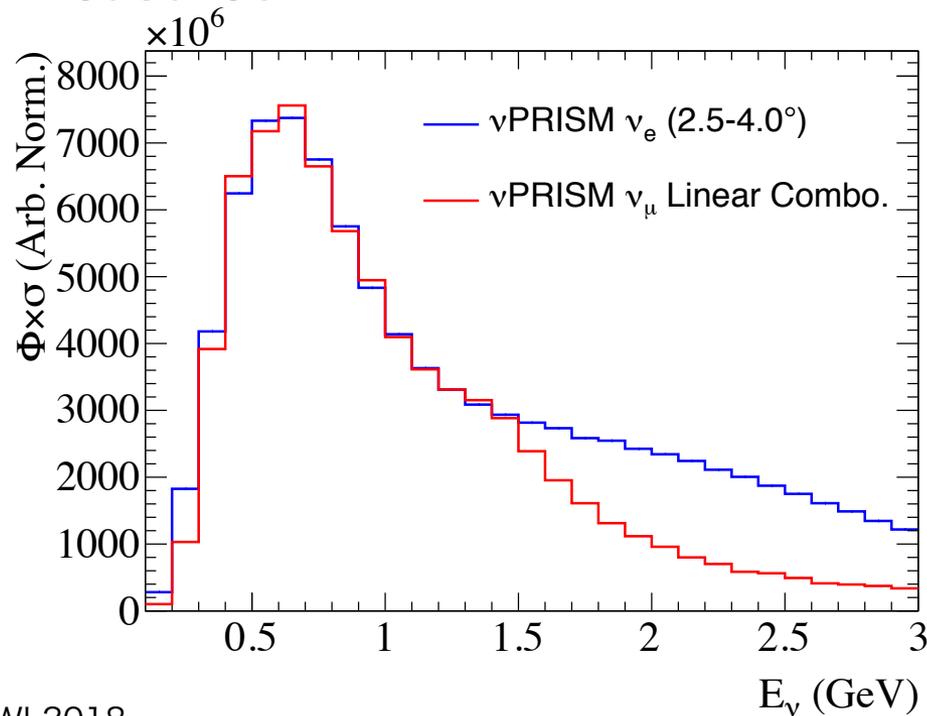
E61 contributions to ν_e appearance analysis

E61 provides a data-driven constraint on $\sigma(\nu_e)/\sigma(\nu_\mu)$ to 2–3%

- 1-ring ν_e candidates are selected with the purity of ~60% out of 1% beam ν_e contamination.
- By taking linear combination of ν_μ spectra to match ν_e spectrum, $\sigma(\nu_e)/\sigma(\nu_\mu)$ will be measured.

$$\frac{N_{\nu_e}}{N_{\nu_\mu}} = \frac{\Phi_{\nu_e} \sigma_{\nu_e}}{\Phi_{\nu_\mu} \sigma_{\nu_\mu}}$$

Measured Common flux uncertainties are canceled



E61 contributions to ν_e appearance analysis

Intrinsic ν_e and neutral current background at the far detector are constrained by measurement at 2.5° off axis

