



Development and  
performance of  
**Silicon-drift detectors**  
and their potentials for  
**soft x-ray**  
advanced light sources  
applications

Andrea Vacchi

INFN Trieste

XVI International Conference  
on Science, Arts and Culture  
International Conference  
ON  
**SESAME**  
In Honour of Paolo Budinich  
29 August - 2 September 2016  
Veli Lošinj, Croatia

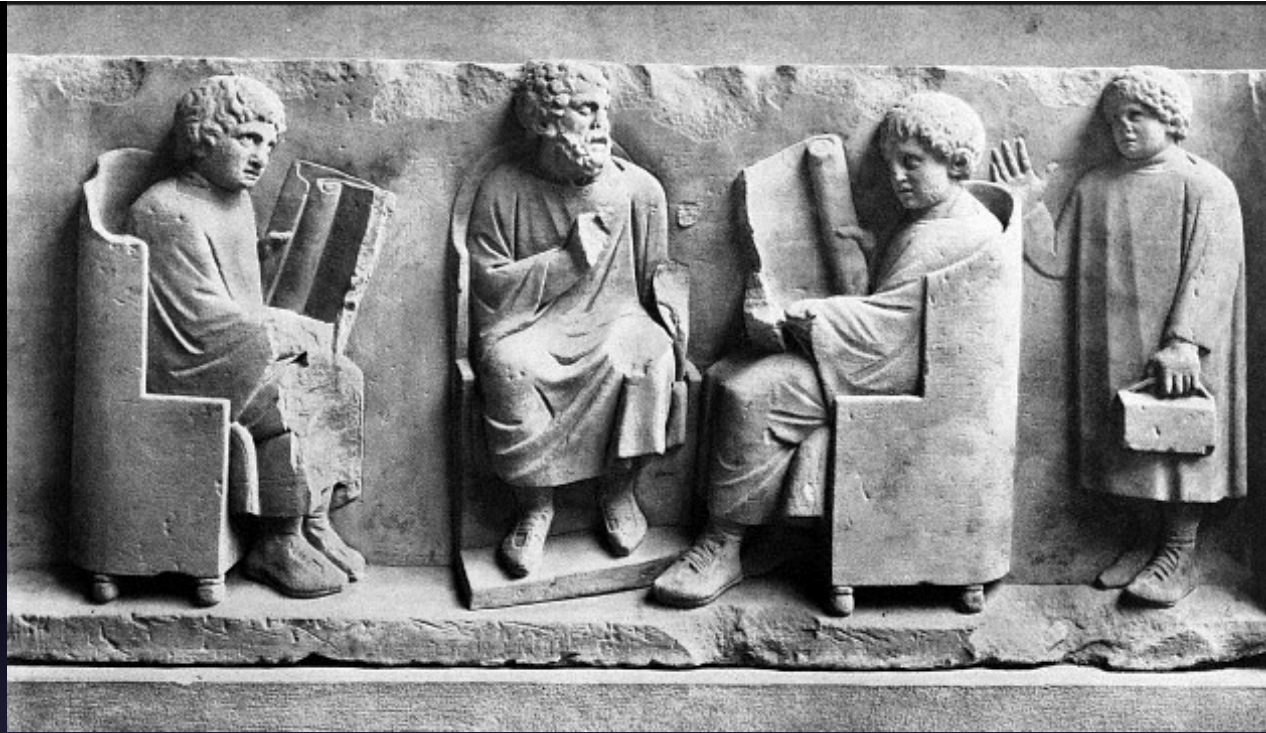
8/30/16



# outline

- Few introductory remarks
  - The due historical framing
  - Our daily SDD; working principle
  - tracking & ALICE detector more than  $1 \text{ m}^2$  since 2007
  - Large Area & soft X-rays:
  - Towards X-ray astrophysics
  - X-ray astronomy LOFT et.al.
  - Some recent results
- 
- G. Zampa presentation at this conference





# what's new?

Know-how diffusion;  
parallel developments;  
efficiency & optimization;  
large impact



# REDSOX collaboration state of the art SDD

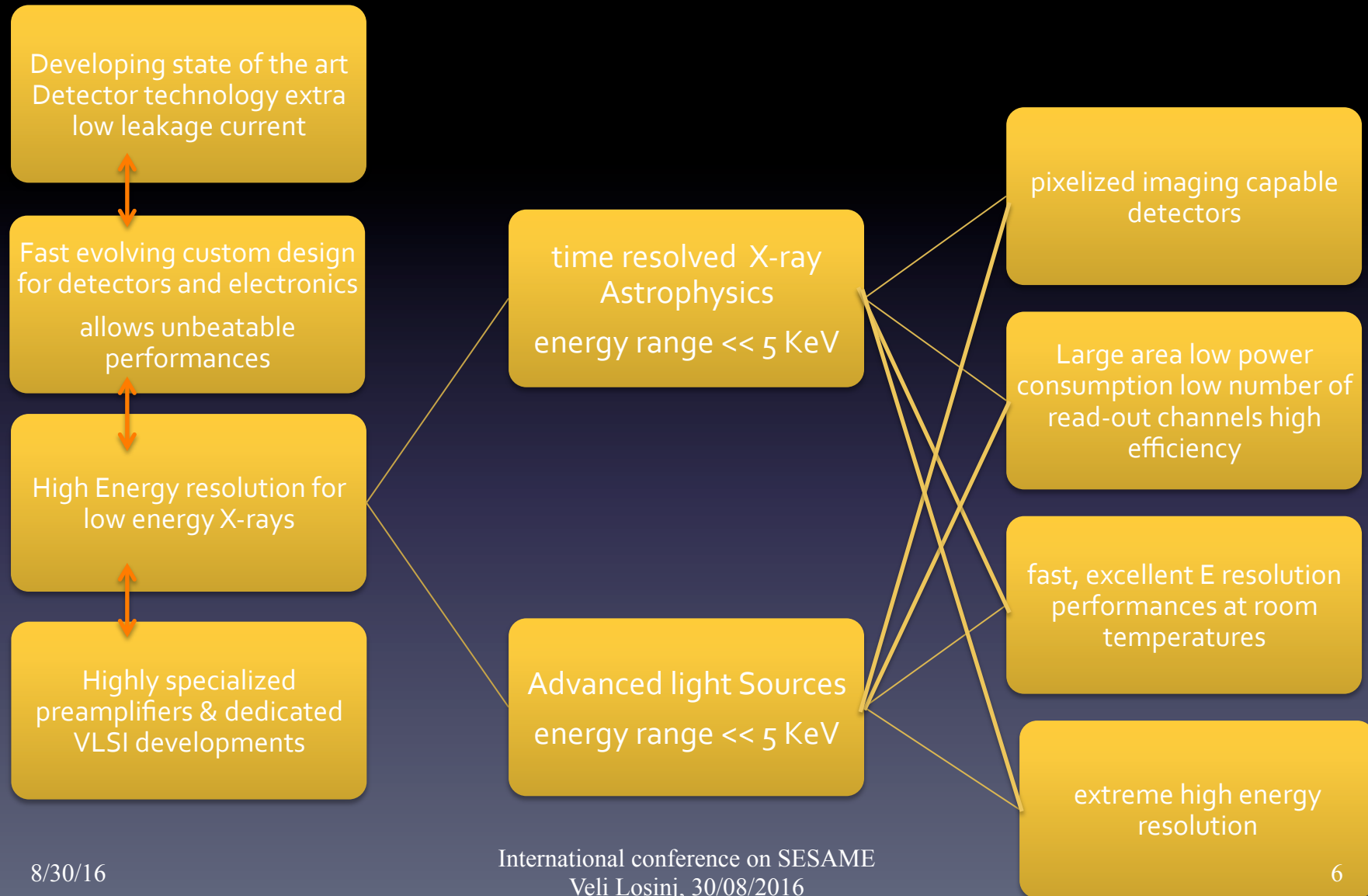


2nd Meeting  
on Silicon Drift  
Detectors for Low  
Energy X-Ray  
Applications  
9-11 May 2016,  
Palazzo Natta,  
COMO





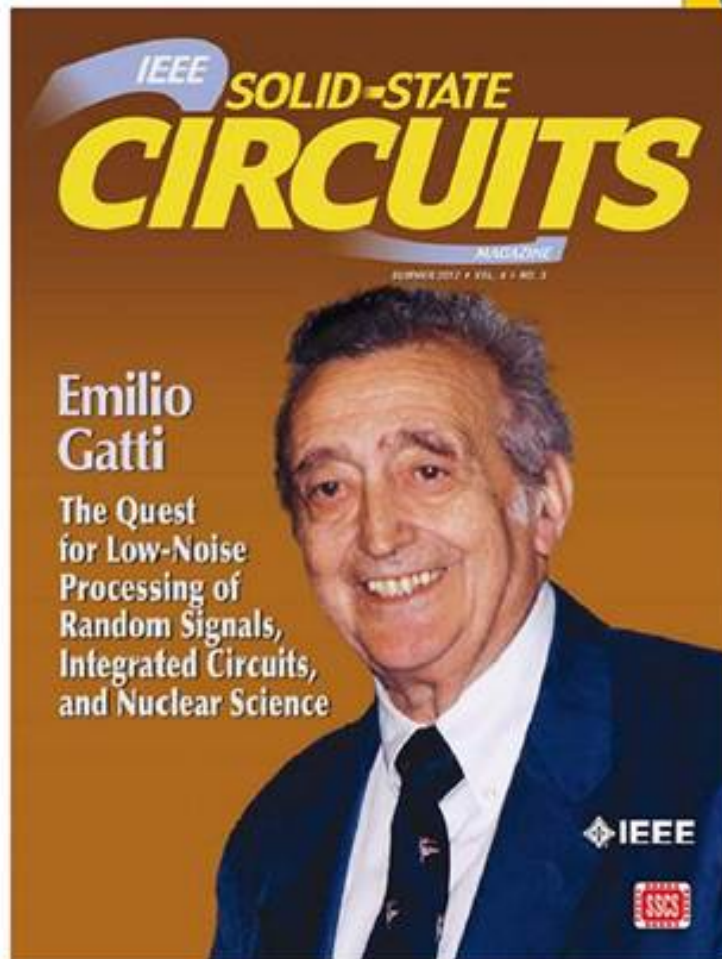
# continuous flow of front edge developments in two technologically hungry scientific fields



# Foreword:

- **E. Gatti and P. Rehak** have proposed Silicon Drift Detectors (SDD) in 1983. It is now 33 years that this very versatile and elegant device evolves in different directions. The SDD operation principle has given rise to the development of a variety of solutions adapted to very challenging specific needs.
- All this was made possible by the planar technology introduced by **J. Kemmer** dedicated to the field of high performance detectors.
- *The first large area SDD, for particles tracking in a high multiplicity environment, was published in 1990, while in 2007 more than 1 m<sup>2</sup> of SDD was made operative in the internal tracking system of the LHC-ALICE experiment, demonstrating the possibility of industrial mass production with high production yield.*
- **Far from being exhausted, large area SDD's continue to evolve.**





**4 ottobre 2012**

In occasione dell'uscita del numero speciale a lui dedicato,

alla presenza e in onore di **Emilio Gatti**

**Erik H.M. Heijne** terrà il seminario:

**Imaging and imagination:  
the use of nanoelectronics for particle  
physics experiments**

Maggiori dettagli sono disponibili al sito: [www.del.polimi.it](http://www.del.polimi.it)

**Ore 14.15  
Sala Conferenze**

Dipartimento di Elettronica  
e Informazione  
Politecnico di Milano

POLITECNICO DI MILANO



DIPARTIMENTO DI  
ELETRONICA E  
INFORMAZIONE

PAVEL REHAK



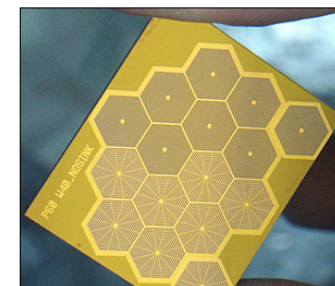
**BROOKHAVEN**  
NATIONAL LABORATORY *Features Media & Communications Office*

[News Releases](#) [Features](#) [Photos](#) [Videos](#) [Fact Sheets](#) [Lab History](#) [Contacts](#)

By Satya Shanmugham | December 19, 2008

## 25th Anniversary of Rehak's and Gatti's Innovative Detector

Brookhaven's Instrumentation Division develops state-of-the-art instrumentation required for experimental research programs at BNL and throughout the world. Much of their work is focused on detectors that track particles after a collision or detectors that identify the energy and location of absorbed x-rays. Some of the current silicon detector developments are for RHIC and ATLAS upgrades, NSLS and NSLS-II, and NASA missions, to name a few.



1983

8

# Dr. Josef Kemmer



**KETEK**  
Creative Detector Solutions

[HOME](#) [PRODUCTS](#) [COMPANY](#) [DOWNLOADS](#) [NEWS](#) [CONTACT](#)



1980 Publishes planar technology with ion implantation in Nuclear Instruments and Methods (NIM)

## HISTORY

Development of the planar technology for the fabrication of silicon radiation detectors by Dr. Josef Kemmer at the Technical University of Munich (TUM).

Fabrication of the first SDDs by Dr. Josef Kemmer at the TUM utilizing the planar process. The detector principle was proposed by Emilio Gatti from Politecnico di Milano and Pavel Rehak from Brookhaven National Laboratory.

Formation of Ketek GmbH by Dr. Josef Kemmer.

Main business of KETEK is technology transfer to industry and to research institutes. Contribution to some of the world's most advanced detector projects in high energy physics and space research like XMM Newton.

Development and launch of the first generation of commercial SDD modules. Introduction of KETEK SDD to the XRF and EDX market within a strong co-operation with Röntec.

Development of monolithic and discrete multi-channel SDD systems.

Manufacturing and delivery of the 1000th commercial SDD Presentation of the 3rd generation of SDD with an energy resolution  $< 128\text{eV}$  at Mn K $\alpha$ .  
Initiation of the electronic product line with the KETEK Analytical X-ray Acquisition System (AXAS) with own development and production.

The MER-Rover in NASA's Mars Lander carries six KETEK alpha particle detectors and a customized SDD.





# PATENT Josef KEMMER 1980

United States Patent [19]

[11] 4,442,592

Kemmer

[45] Apr. 17, 1984

[54] PASSIVATED SEMICONDUCTOR PN JUNCTION OF HIGH ELECTRIC STRENGTH AND PROCESS FOR THE PRODUCTION THEREOF

[76] Inventor: Josef Kemmer, No. 41 D, Hauptstrasse, 8041 Haimhausen, Fed. Rep. of Germany

[21] Appl. No.: 225,069

[22] Filed: Jan. 14, 1981

[30] Foreign Application Priority Data

Jan. 31, 1980 [DE] Fed. Rep. of Germany ..... 3003391

[51] Int. Cl.<sup>3</sup> ..... H01L 31/18

[52] U.S. Cl. .... 29/572; 29/576 B; 29/578; 148/1.5; 148/187

[58] Field of Search ..... 148/1.5, 187; 29/571, 29/572, 569 R, 576 B, 578; 156/643, 662

[56] References Cited

U.S. PATENT DOCUMENTS

3,769,109 10/1973 MacRae et al. .... 148/1.5 X

3,891,468 6/1975 Ito et al. .... 148/1.5

Primary Examiner—Brian E. Hearn

Assistant Examiner—David A. Hey

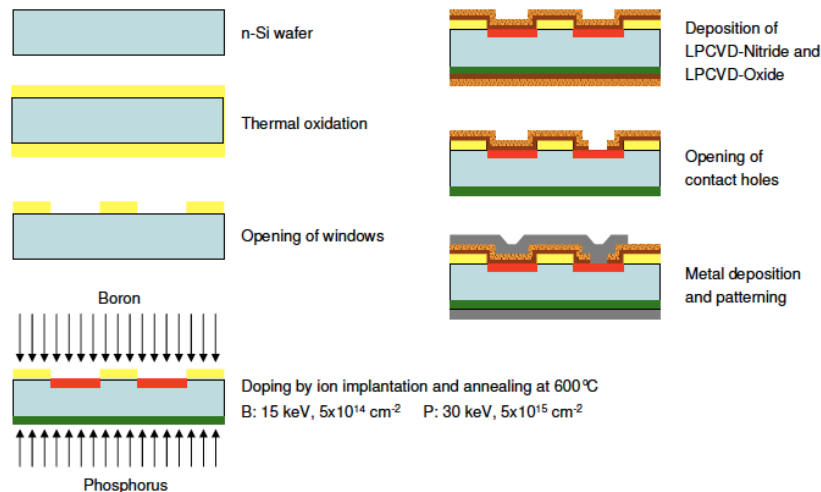
Attorney, Agent, or Firm—John C. Smith, Jr.

[57] ABSTRACT

A passivated semiconductor pn junction is provided which has a high electric strength, one area being heavily doped and being very thin, in particular for radiation detectors. The pn junction has an edge zone at which a depletion zone (surface channel) is provided underneath the passivation layer.

12 Claims, 2 Drawing Figures

## Planar Technology



## PATENT Josef KEMMER 1980

4,442,592

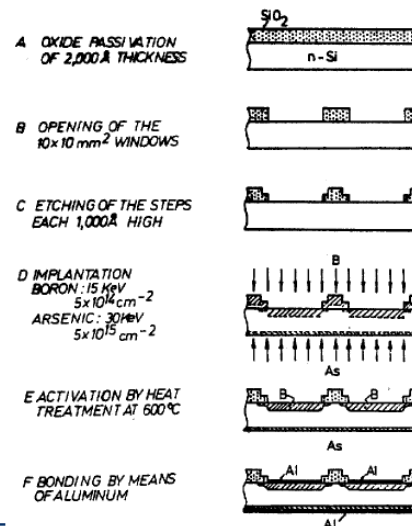
CRITICAL STEPS :

OXIDE GROWTH USING HCl (not mentioned)

OXIDE STEPS

REAR IMPLANT

PATTERNING of Al in CONTACTS (not mentioned)



ence on SESAME

Veli Losinj, 30/08/2016

# The First Large-area Silicon Drift Detector (1991 )

Nuclear Instruments and Methods in Physics Research A306 (1991) 187–193  
North-Holland

187

1991

## Performance of the UA6 large-area silicon drift chamber prototype

A. Vacchi

*The Rockefeller University, New York, NY, USA*

A. Castoldi, S. Chinnici, E. Gatti, A. Longoni, F. Palma and I

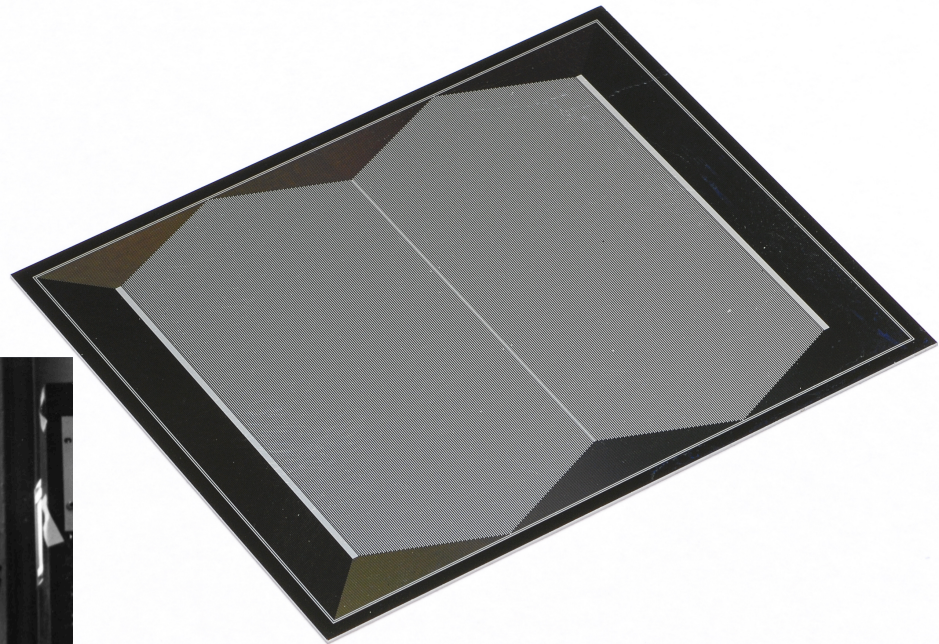
*Politecnico di Milano, Dipartimento di Elettronica and Centro di Elettronica Quantistica*

P. Rehak

*Brookhaven National Laboratory, Upton, NY, USA*

J. Kemmer

*Facultät für Physik der Technischen Universität, Munich, Germany*



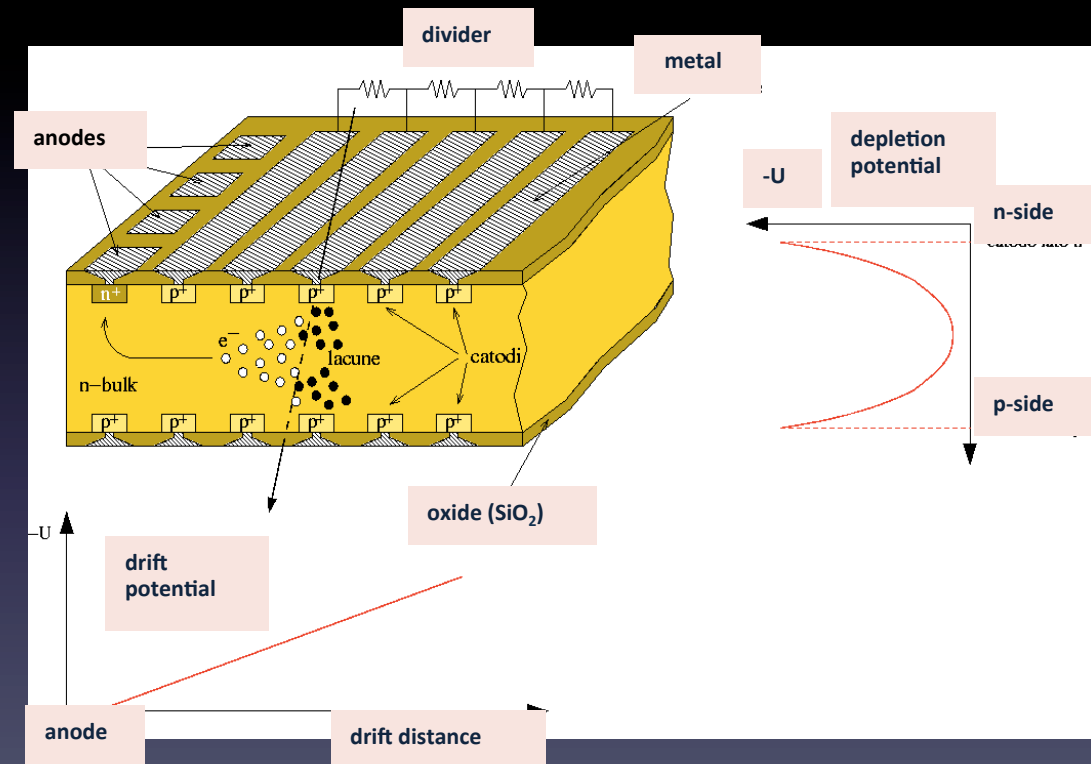
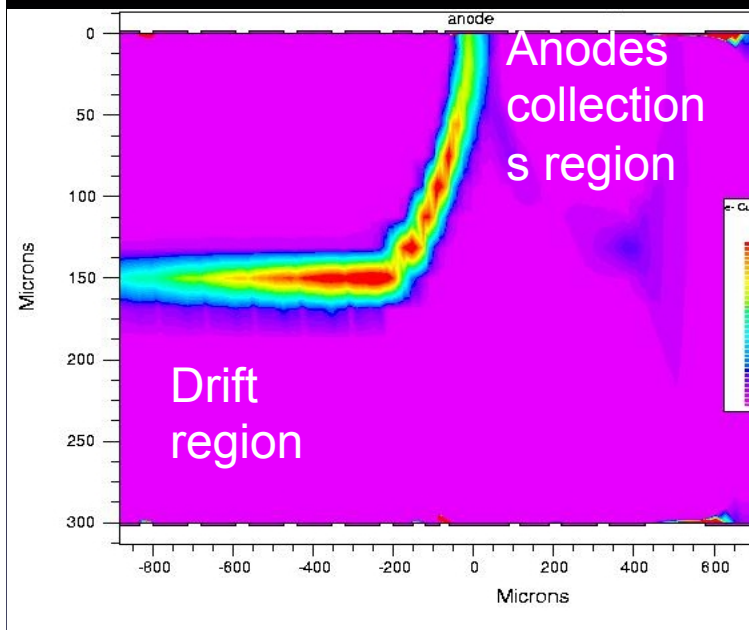
8/30/16

International conference on SESAME  
Zeljko Losinji, 20/08/2016

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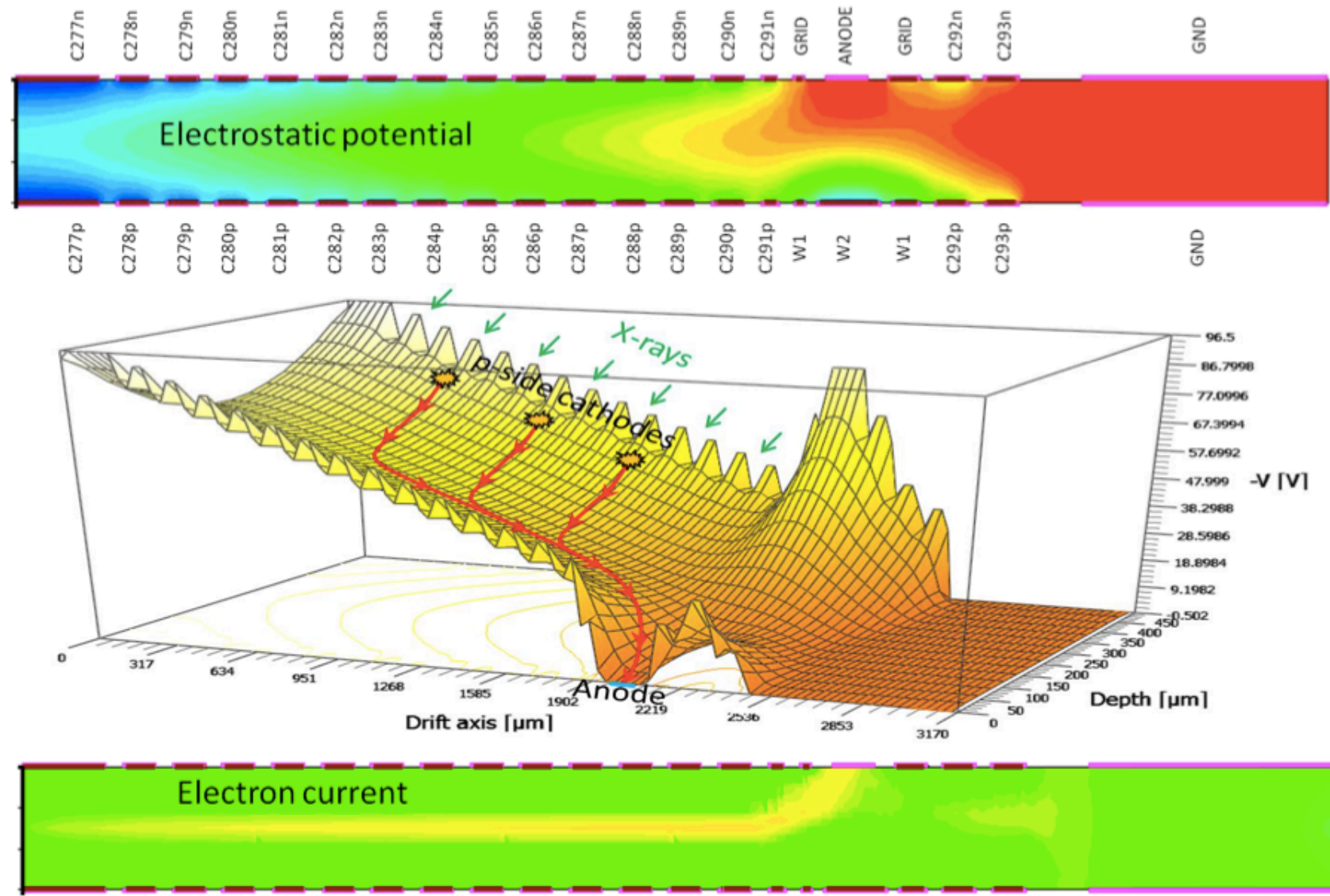


# Silicon Drift Detectors (SDD) Basic Principle

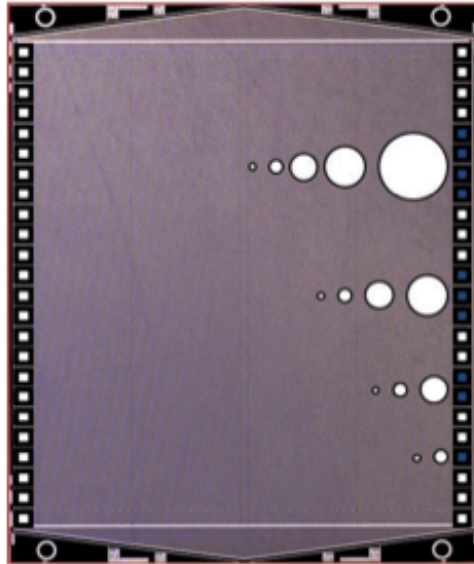


- due to their collection electrode geometry, anodes, SDD have excellent noise performance and are well suited for low energy X-ray spectroscopy applications

# The linear Silicon Drift Detector



# The linear SDD signal and charge diffusion



The charge cloud has a Gaussian shape. Its width depends on the drift time which, for a constant electric field, is a linear function of the drift distance:

$$D = \frac{k_B T}{q} \mu \quad \text{diffusion coefficient}$$

$$v = \mu E \quad \text{drift velocity}$$

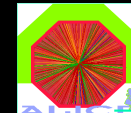
$$t = \frac{x}{v} \quad \text{drift time}$$

$$\sigma = \sqrt{2Dt + \sigma_0^2} = \sqrt{2 \cdot \frac{k_B T}{q} \mu \cdot \frac{x}{\mu E} + \sigma_0^2} = \sqrt{2 \frac{k_B T}{qE} x + \sigma_0^2}$$

For low energy X-ray photons it is possible to take  $\sigma_0 \approx 0$

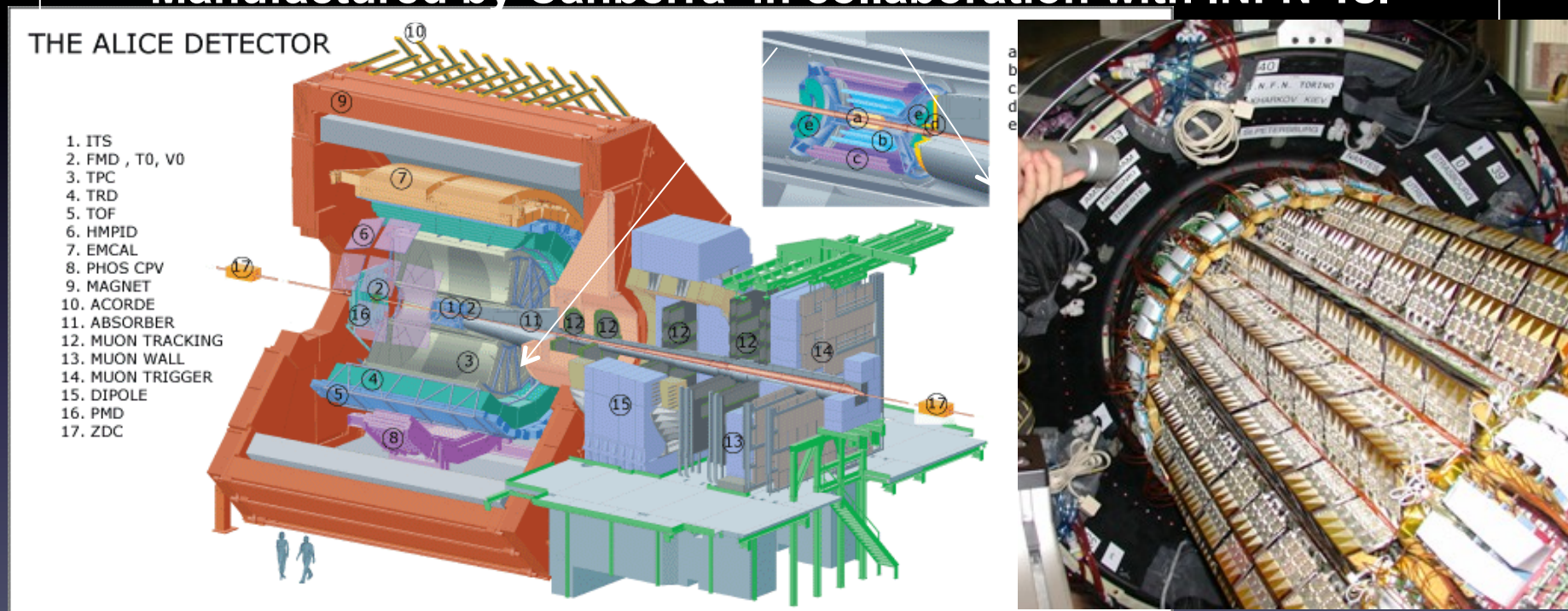
- Measuring the charge cloud lateral dimension it is possible to derive the drift length with a coarse resolution
- Several anodes (N) are involved in the signal acquisition: the electronic noise contribution to the energy resolution is multiplied by  $\sqrt{N}$
- For the best energy resolution the anode pitch should be large in order to have at most 2 anodes integrating the charge
- For imaging purposes the anode pitch should be small to better estimate the drift distance from the signal shape at the anodes (no prompt trigger with photons!!!)





- production yield better than 60% (>90% for last batches)

260 detectors ( $\sim 1.37 \text{ m}^2$ ) operational in LHC since 2007 .  
Manufactured by Canberra in collaboration with INFN-Ts.



*Nuclear collisions at the LHC have allowed experimenters to study strongly interacting matter in unprecedented conditions of temperature and density and with a much enhance range of probes unveiling new features of the quark-gluon plasma in these extreme conditions.*





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A. Rashevsky et al. / Nuclear Instruments and Methods in Physics Research A 461 (2001) 133–138

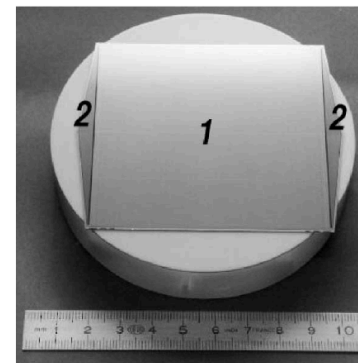


Fig. 1. Picture showing a SDD of the design ALICE-D1B. The numbers correspond to (1) sensitive area; (2) guard area.

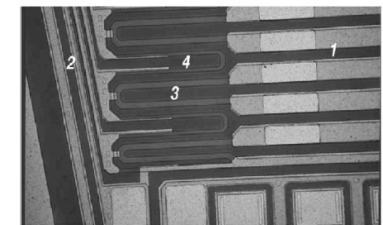


Fig. 2. Details of the H.V. integrated divider region: (1) drift cathode; (2) guard cathode; (3) divider resistor for the drift cathodes; (4) divider resistor for the guard cathodes.

mental conditions, if one does not foresee special solutions in the detector design to attenuate this influence [4,5]. From this point of view, in case of the SDD, the punch-through phenomenon is the most critical. When the voltage difference between adjacent  $p^+$  cathodes reaches a critical value  $U_{pt}$ , a

[www.elsevier.nl/locate/nima](http://www.elsevier.nl/locate/nima)

## Characteristics of the ALICE

A. Rashevsky<sup>a,\*</sup>, V. Bonvicini<sup>a</sup>, P. Burger<sup>b</sup>, P.  
R. Hernández-Montoya<sup>c,1</sup>, A. Kolojvari<sup>c</sup>  
C. Piemonte<sup>a</sup>, F. Tosello<sup>c</sup>, A.

<sup>a</sup>Università di Trieste, c/o Area di Ricerca, Palazzo

<sup>b</sup>Canberra Semiconductor, NV,

<sup>c</sup>INFN Sezione di To

For the ALICE Coll.

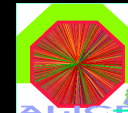


(SDD) with an active area of  $7.0 \times 7.0$  cm<sup>2</sup>. The development of the SDD





# SDD for LHC–ALICE experiment



## Wafer type:

- 5" Neutron Transmutation Doped  
<111> 3 kΩ·cm, thickness 300 μm

## Area:

- Sensitive:  $7.02 \times 7.53 \approx 53 \text{ cm}^2$ ,  
divided in 2 drift regions
- total:  $7.25 \times 8.76 \text{ cm}^2$ , (ratio = 0.83)

## Each drift region:

- has a length of 35 mm
- has 291 cathodes biased by an integrated voltage divider
- has 256 anodes – pitch of 294 μm
- has 3 lines of 33 MOS charge injectors for the drift velocity calibration

## Guard region:

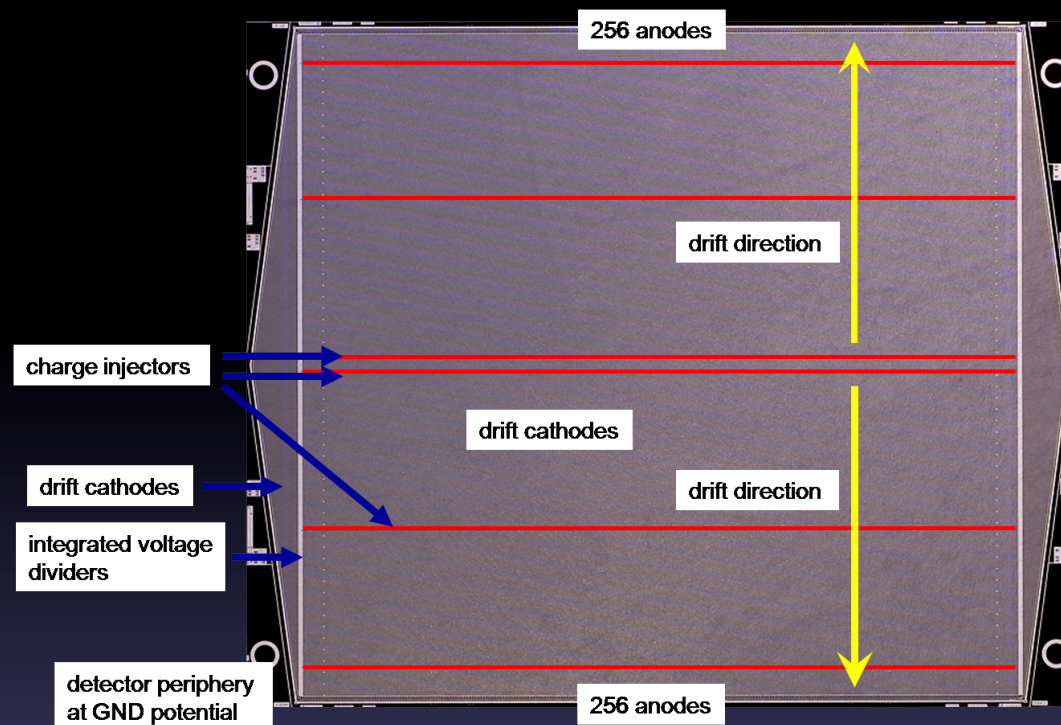
- independent voltage dividers

## Integrated dividers:

- Equivalent resistance of all voltage dividers  $R_{\text{tot}} = 4781 \text{ k}\Omega$

## Each anode:

- has a very small capacitance of  
 $\sim 100 \text{ fF}$
- reads an area of 10 mm<sup>2</sup>

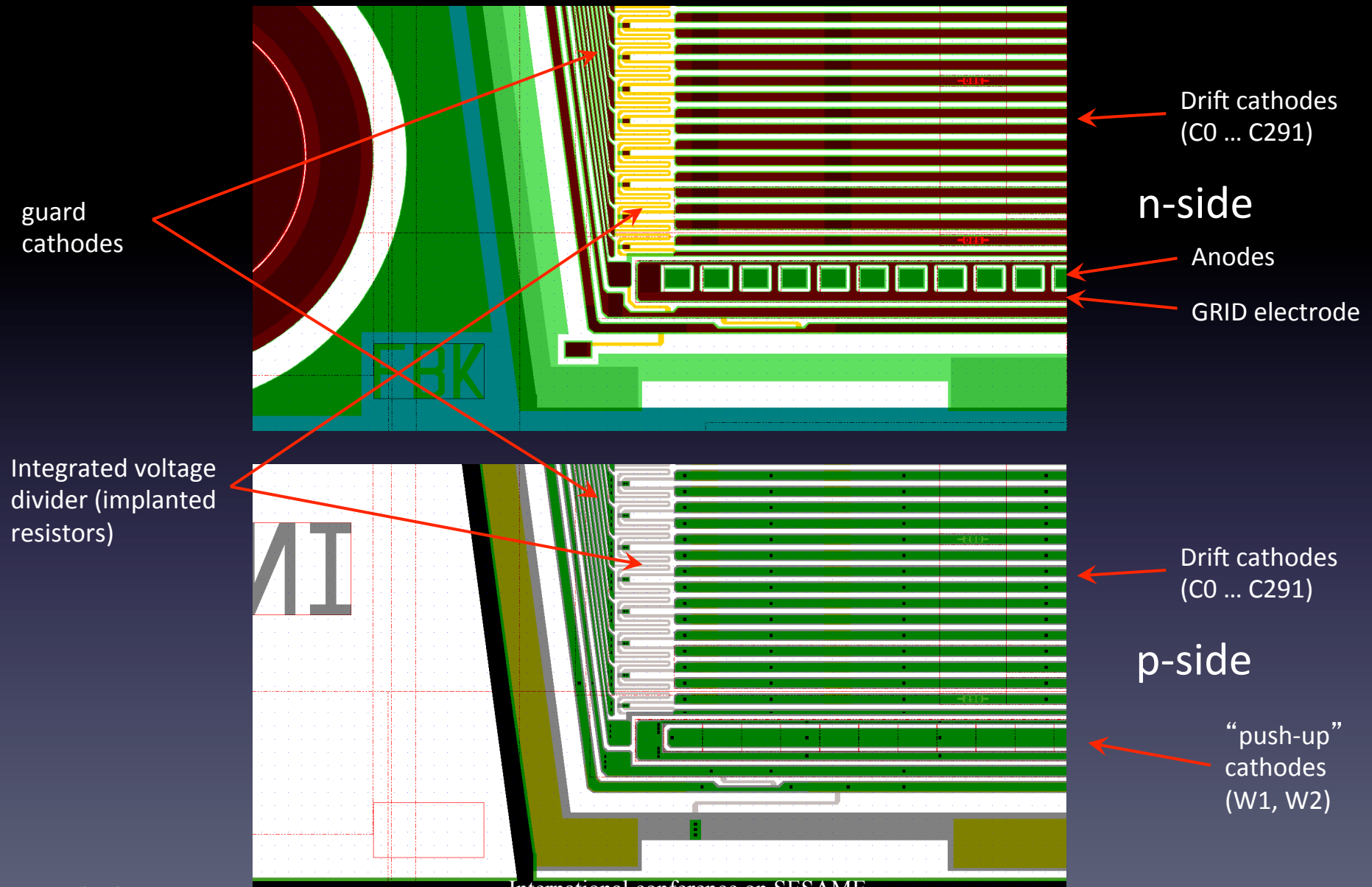


## Specs for the detector :

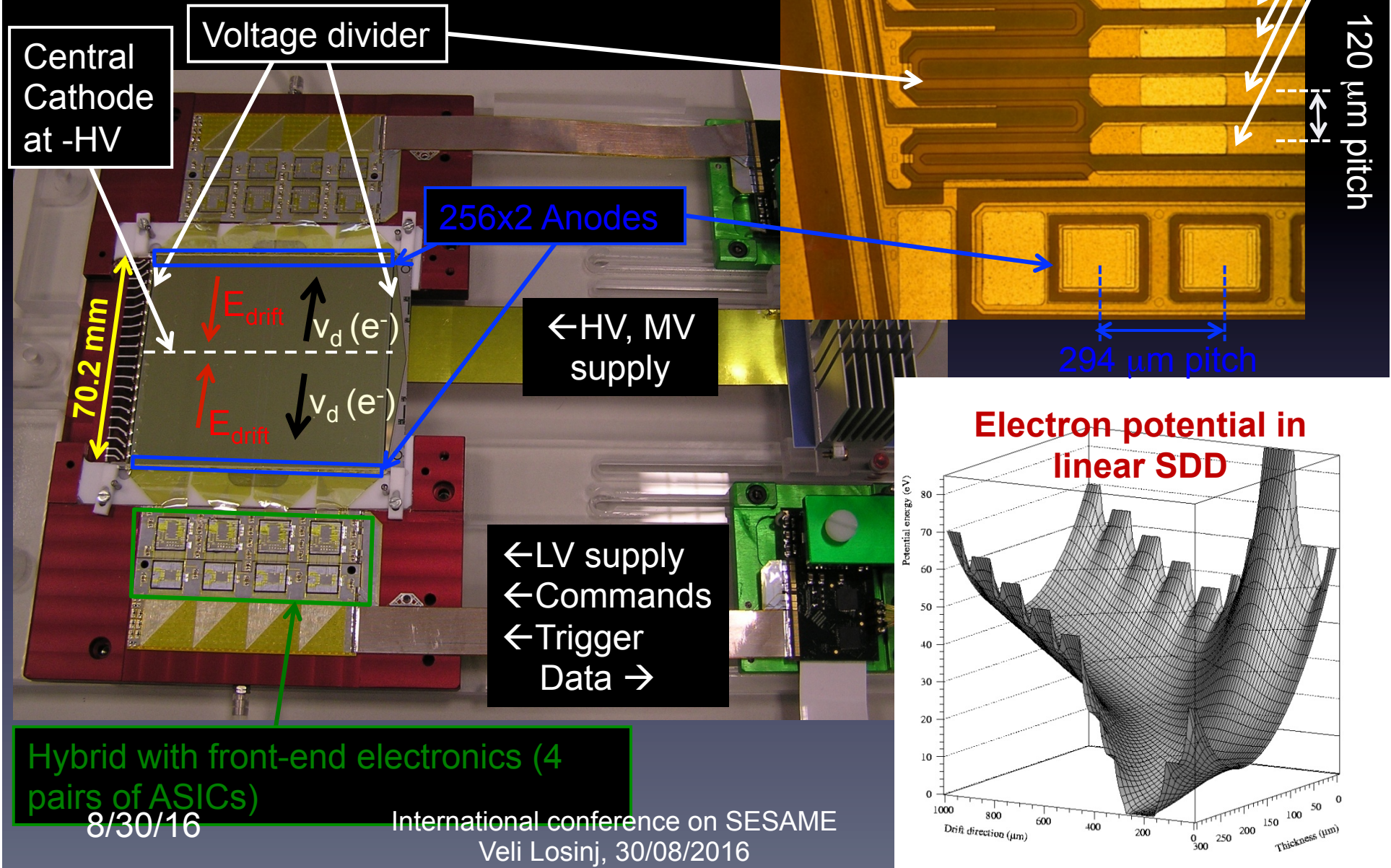
- HV bias: -2.4 kV, 8V/cathode  $E = 670 \text{ V/cm}$
- 35mm in a drift time of 4.3 μs,  $v_d = 8 \mu\text{m/ns}$
- total current on the voltage dividers  $\sim 0.48 \text{ mA}$
- on board power consumption: 1.15 W



# SDD layout details and terminology



# SDD sensor

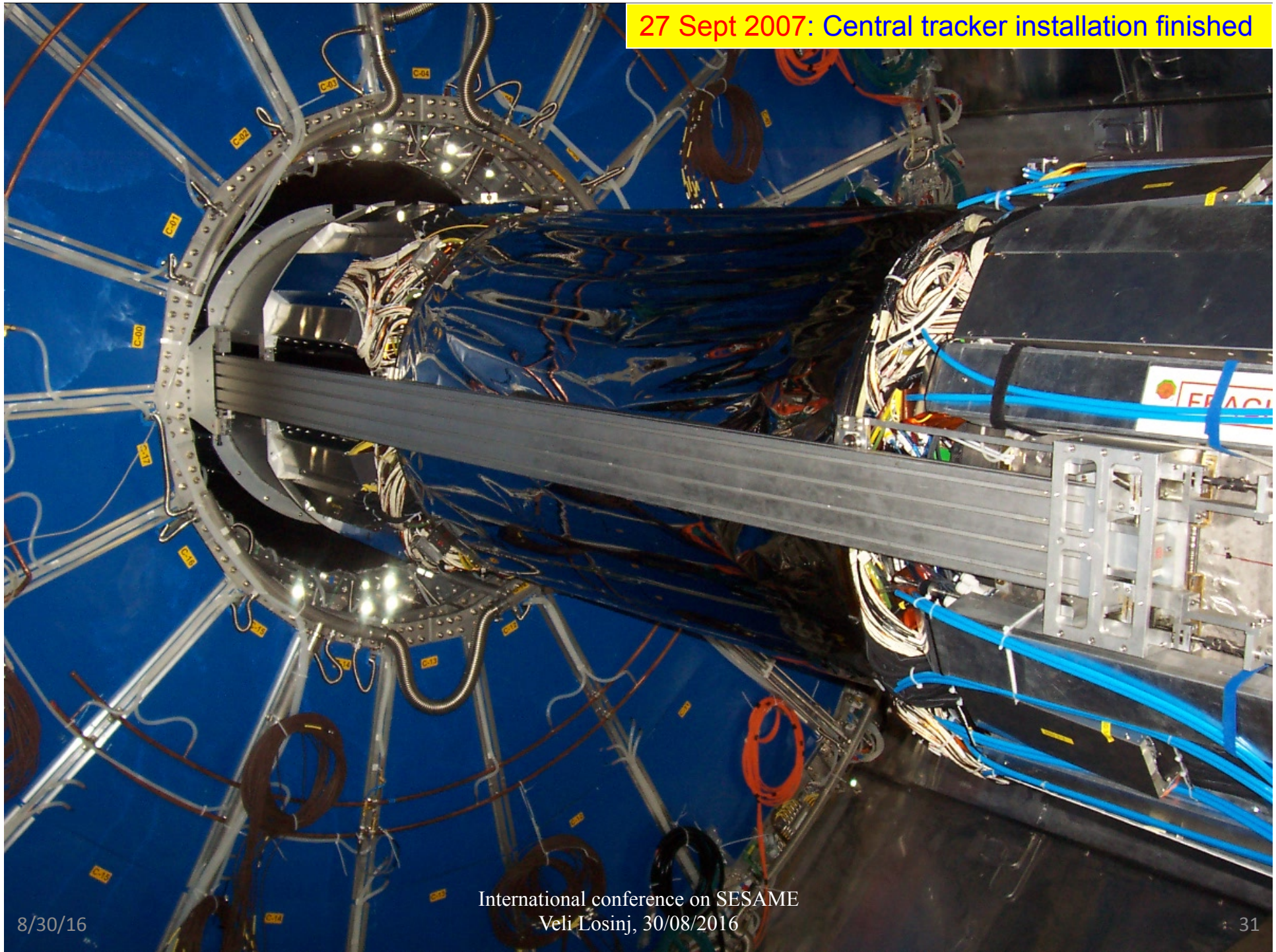


8/30/16

International conference on SESAME  
Veli Losinj, 30/08/2016



27 Sept 2007: Central tracker installation finished



8/30/16

International conference on SESAME  
Veli Losinj, 30/08/2016

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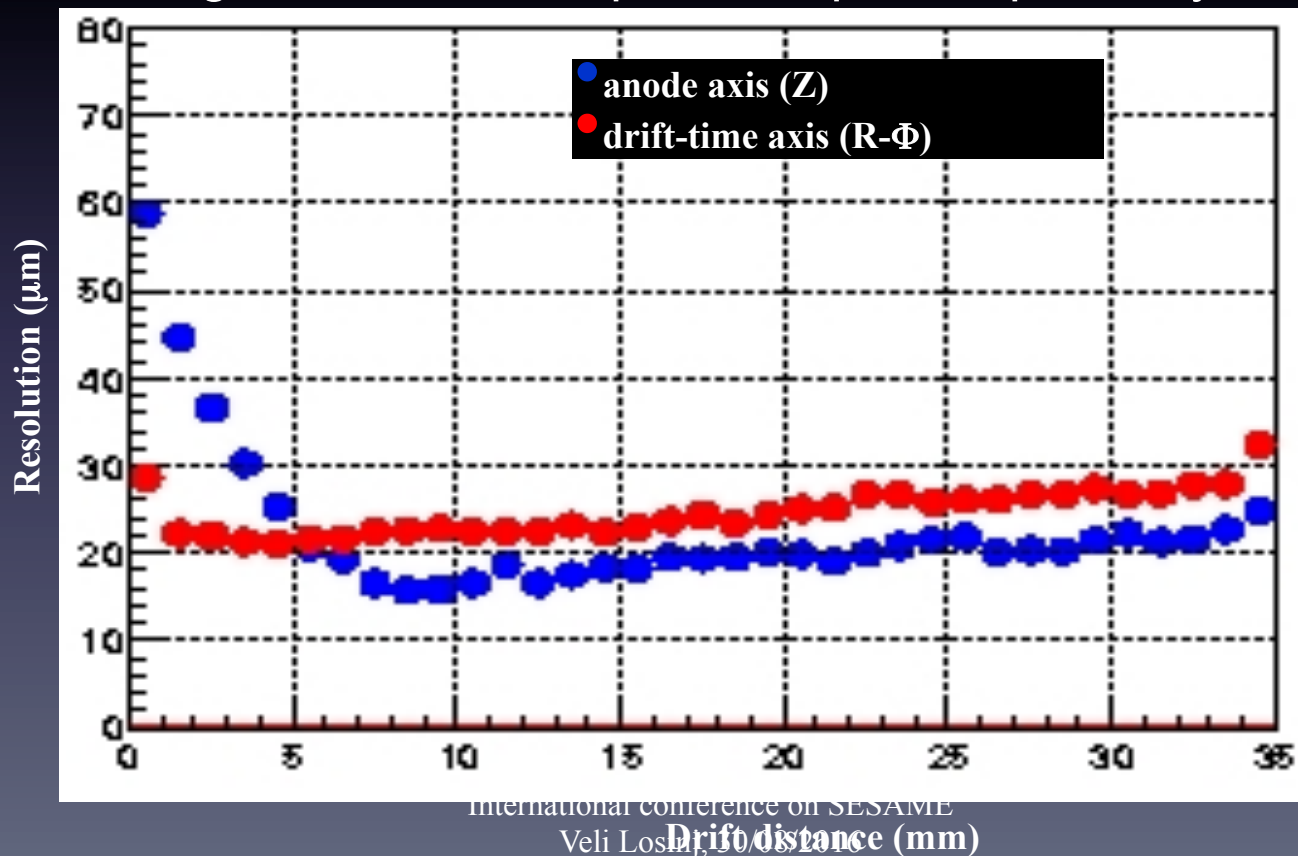


## The spatial precision

along the drift direction, is better than  $30\text{ }\mu\text{m}$  over the whole detector surface.

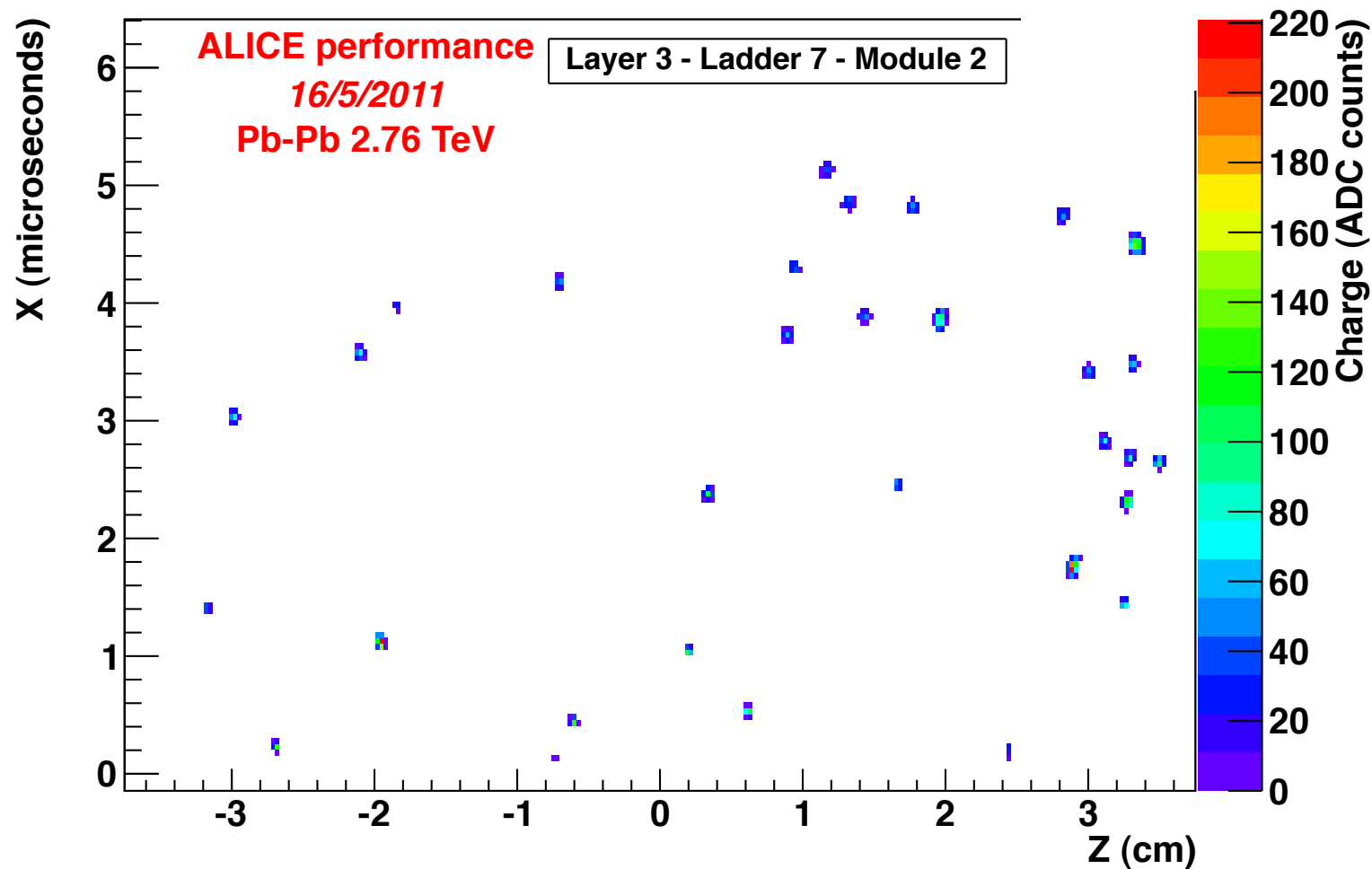
The precision along the anode, is better than  $30\text{ }\mu\text{m}$  over 94% of the detector surface and reaches  $60\text{ }\mu\text{m}$  close to the anodes, where a fraction of clusters affects only one anode.

The average values are  $35\text{ }\mu\text{m}$  x  $25\text{ }\mu\text{m}$  respectively.

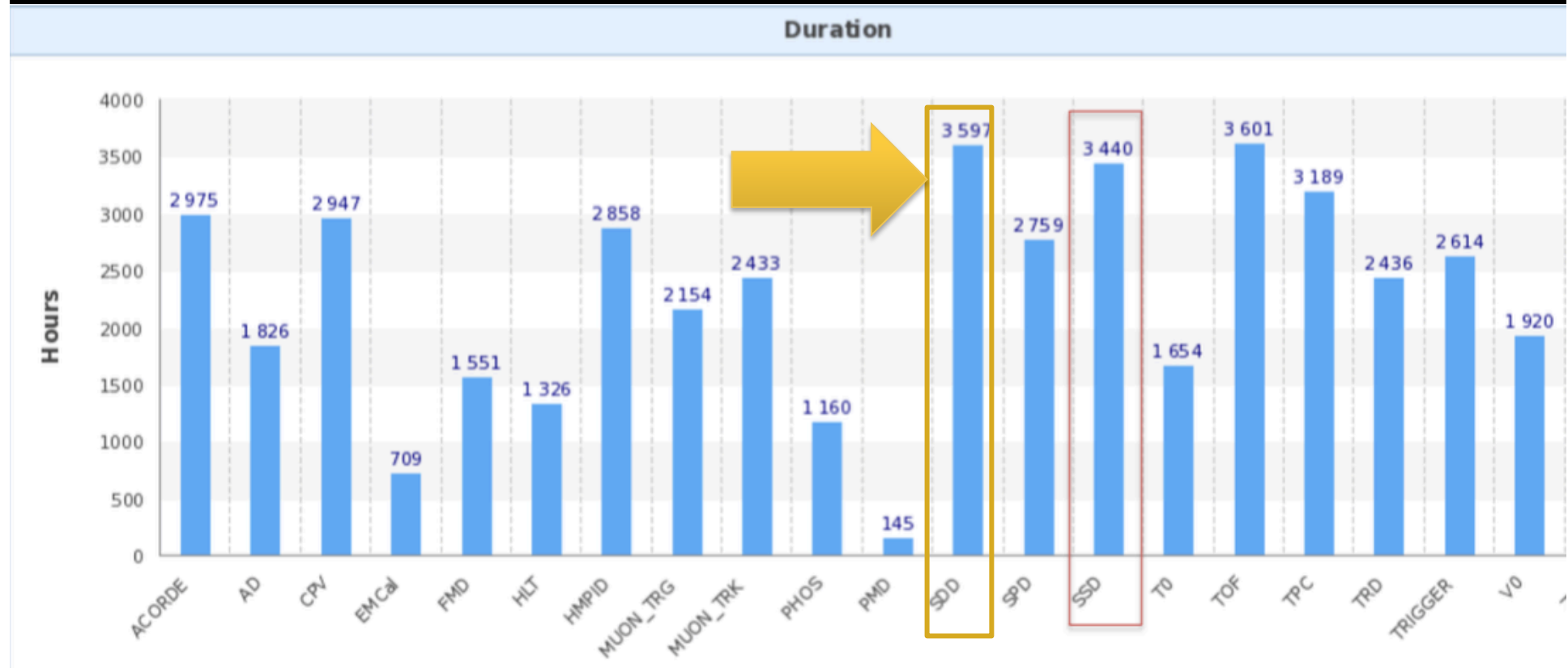


*test beam data*

# High multiplicity two coordinates delivered Clean!



# Alice SDD detection system run active time since 2016 year beginning





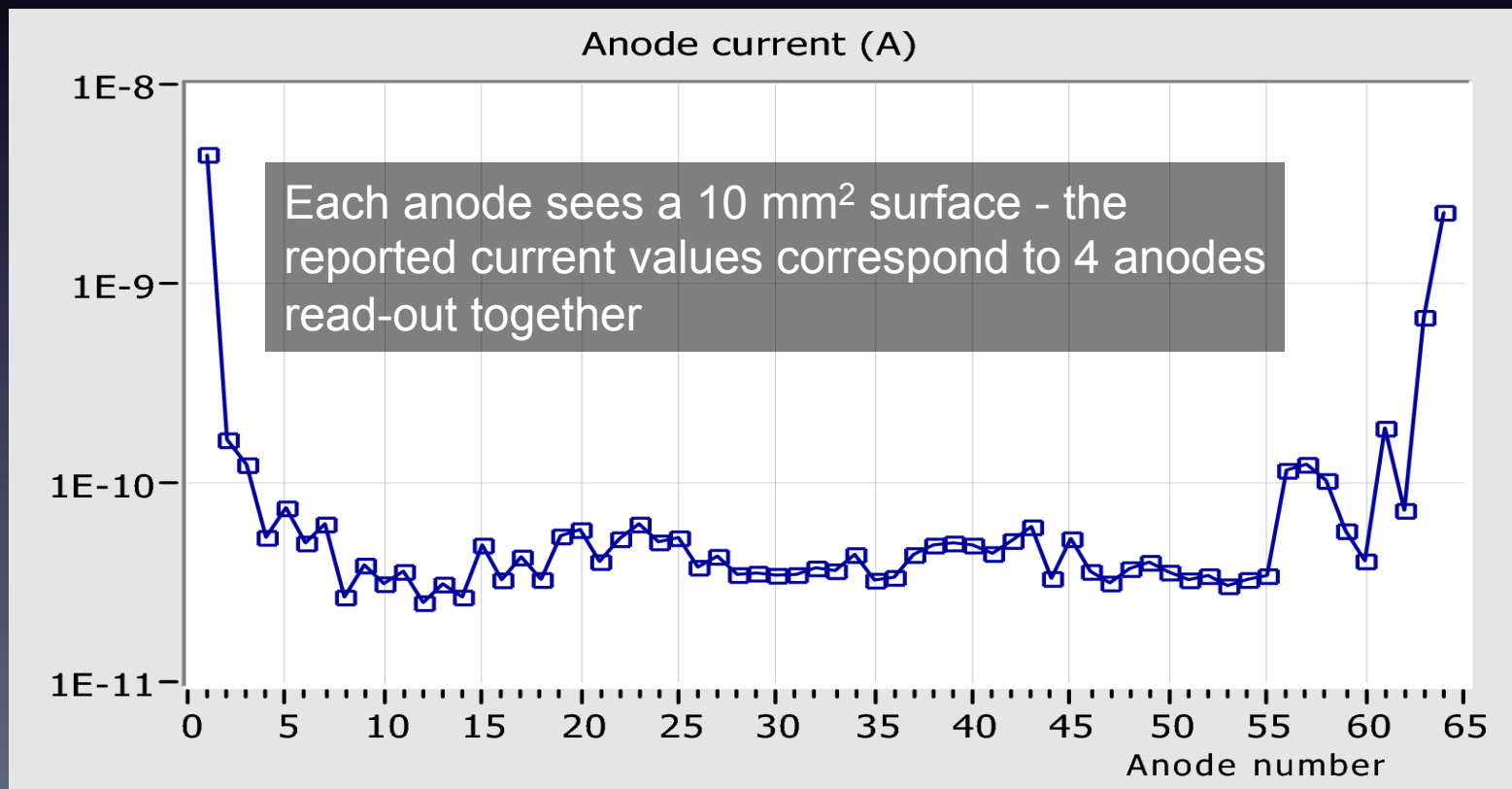
**suitable for low-energy X-ray spectroscopy applications:**

**+ anode capacitance is about 50 fF,**

**+ the leakage current at the anode measured at room temperature is very low.**

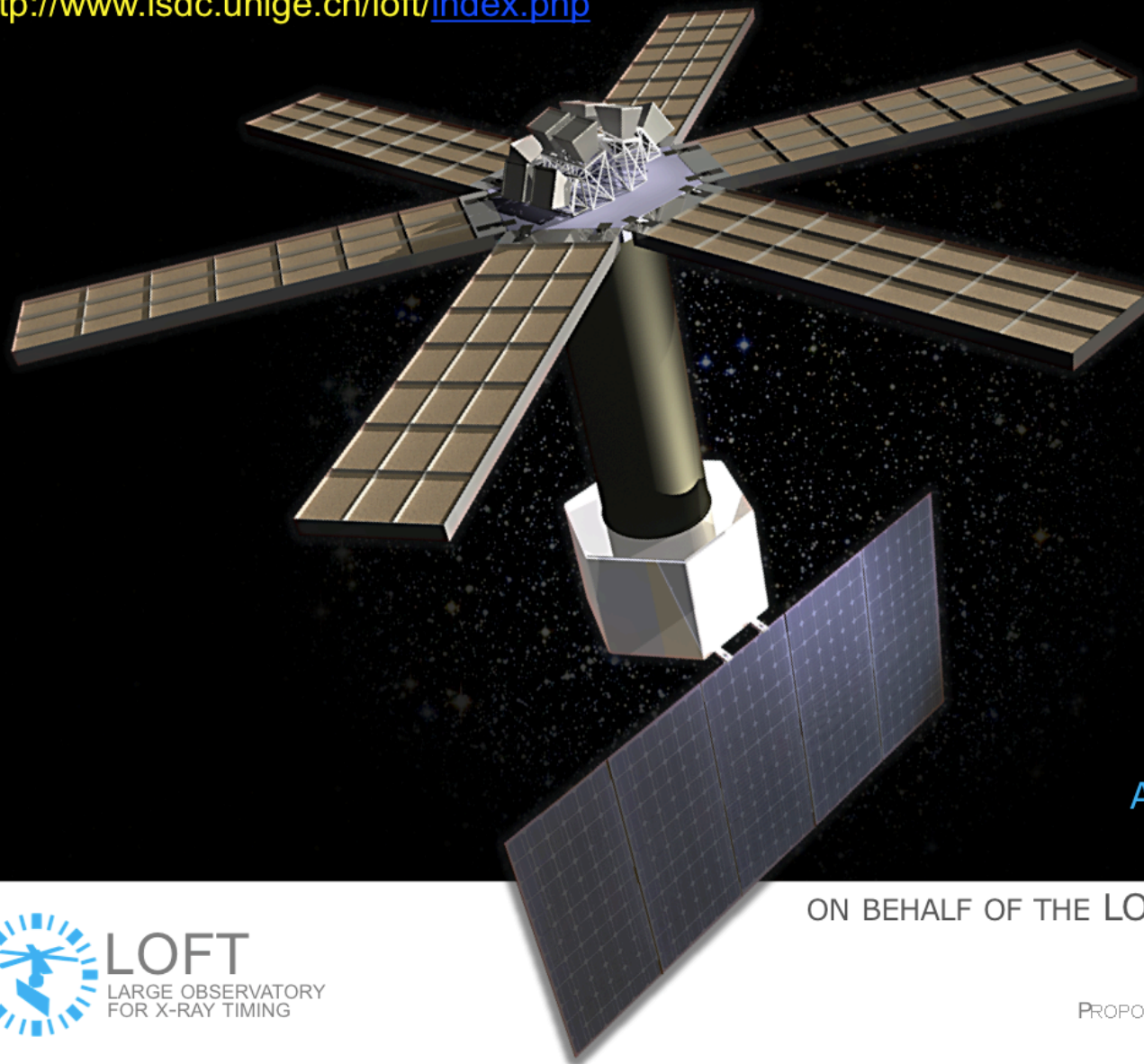
**This allows a very-low noise contribution from the front-end electronics.**

**More than two order of magnitudes larger sensitive area than standard spectroscopic SDDs, this detector can open the way to application areas of the X-ray spectroscopy that require wide surface coverage.**



# LOFT concept

<http://www.isdc.unige.ch/loft/index.php>



MARCO FEROCI  
INAF, ROME

LUIGI STELLA  
INAF, ROME

ANDREA VACCHI  
INFN, TRIESTE

ON BEHALF OF THE LOFT CONSORTIUM

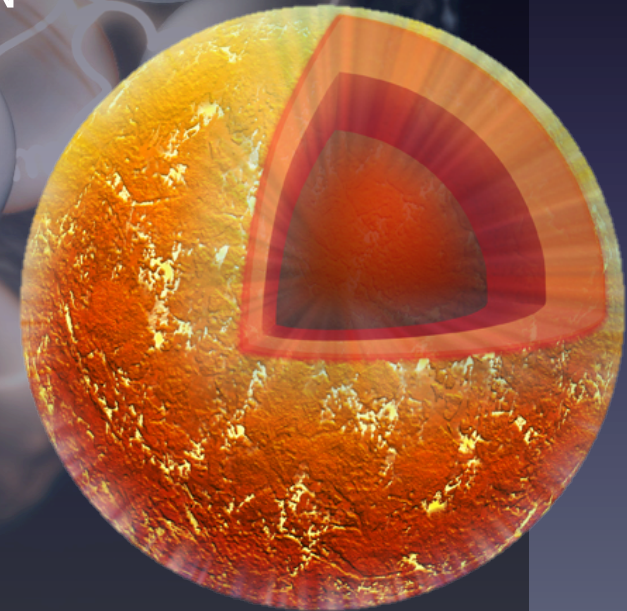
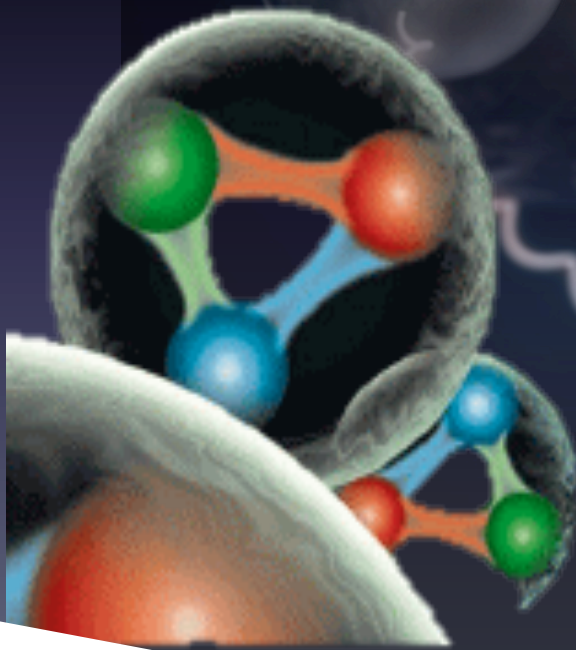


ASI COSMIC VISION M4  
PROPOSALS PRESENTATION EVENT

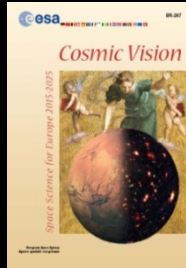
The strong force determines the state of nuclear matter - from atomic nuclei to neutron stars.

It is a major problem within modern physics.

PROGRESS IS DRIVEN  
BY **LABORATORY  
EXPERIMENT** AND  
**ASTROPHYSICAL  
OBSERVATION.**

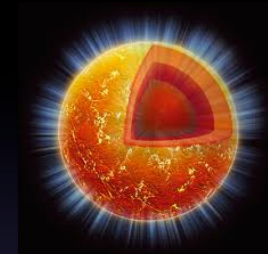






## LOFT ADDRESSES THE COSMIC VISION THEME *"Matter Under Extreme Conditions"*

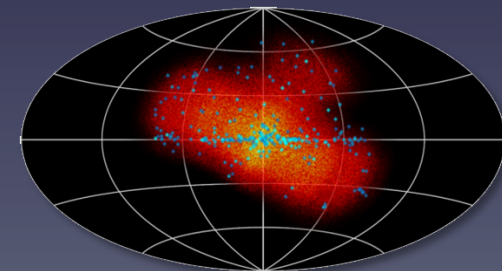
Probe the state of matter at supra  
nuclear densities in Neutron Stars  
(**"Dense Matter"**)



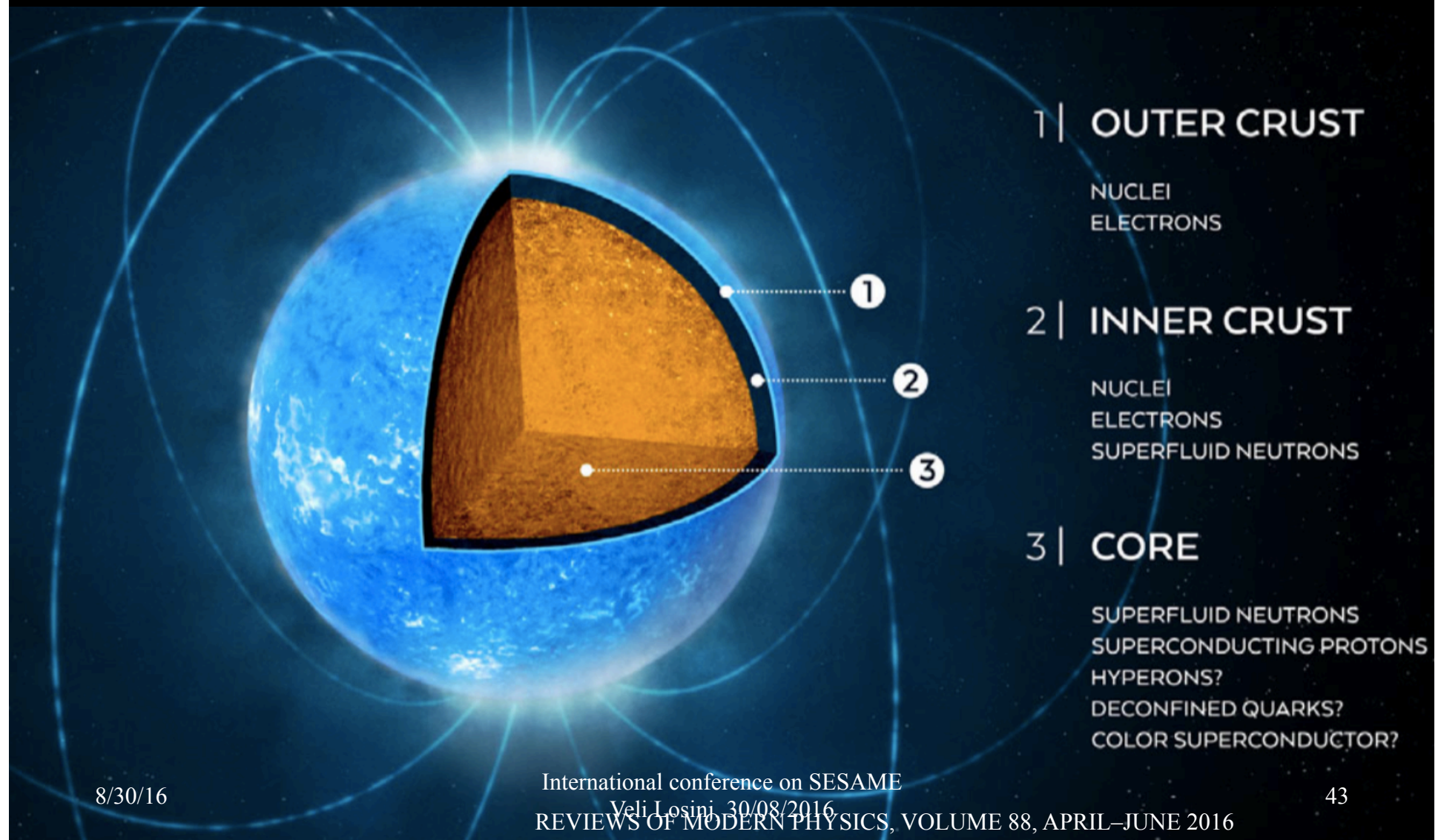
Probe gravity theory in the very strong  
field environment of Black Holes  
(**"Strong Gravity"**)

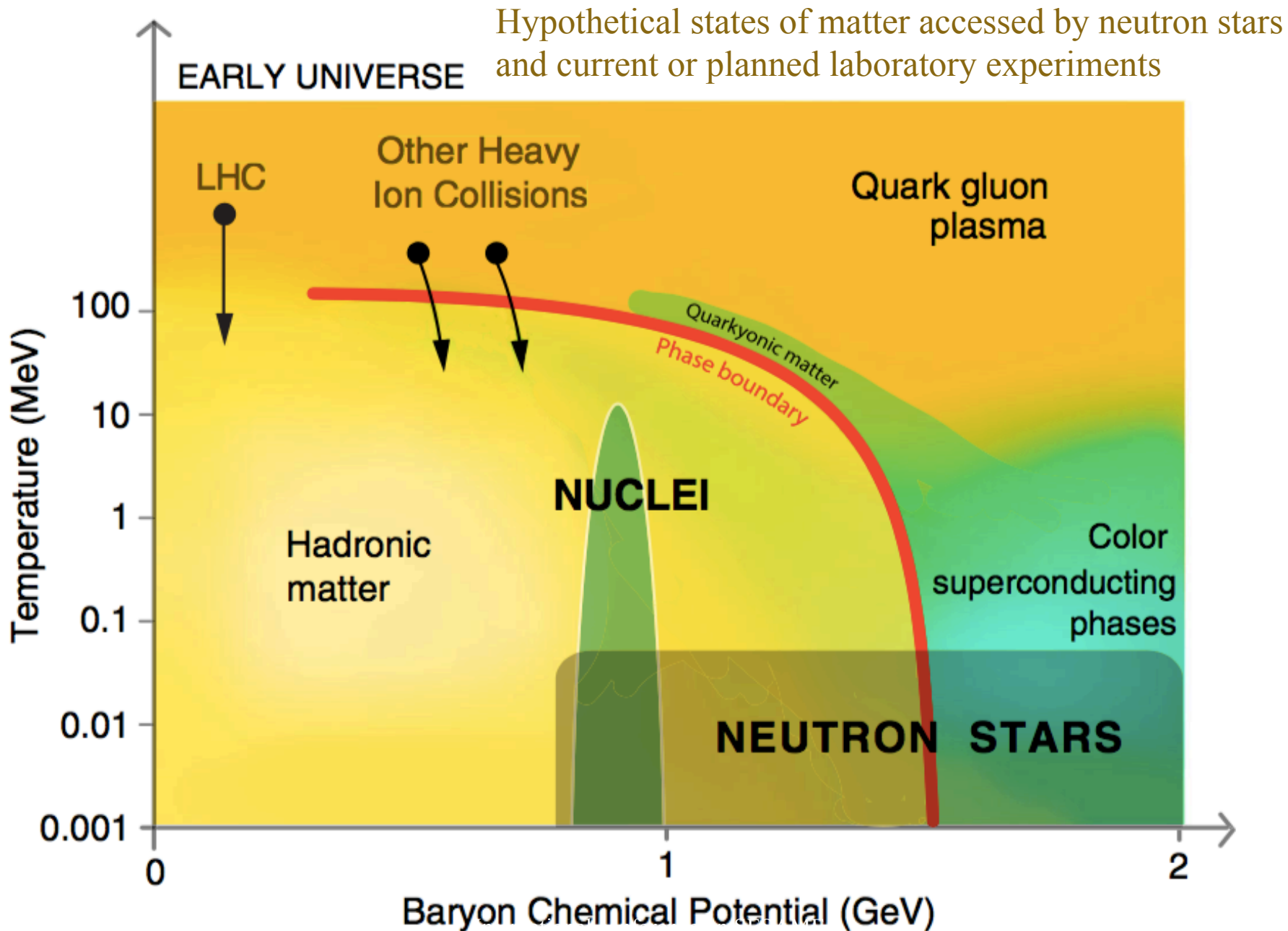


Probe physics of hundreds of galactic  
and bright extragalactic cosmic sources  
(**"Observatory Science"**)



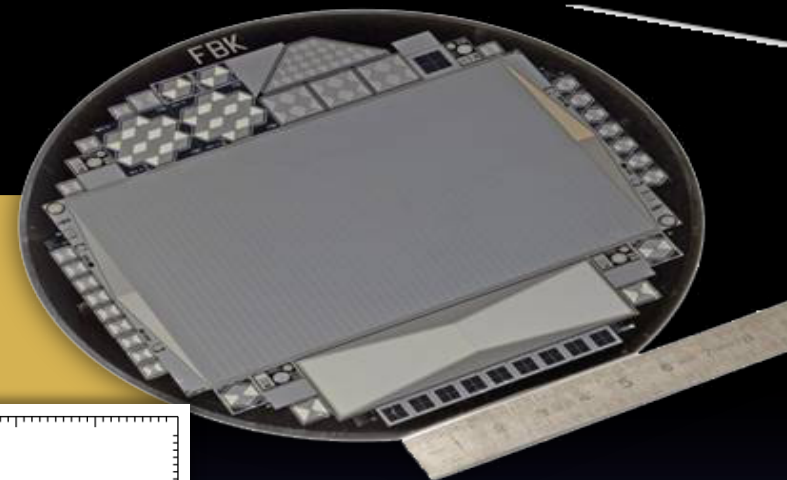
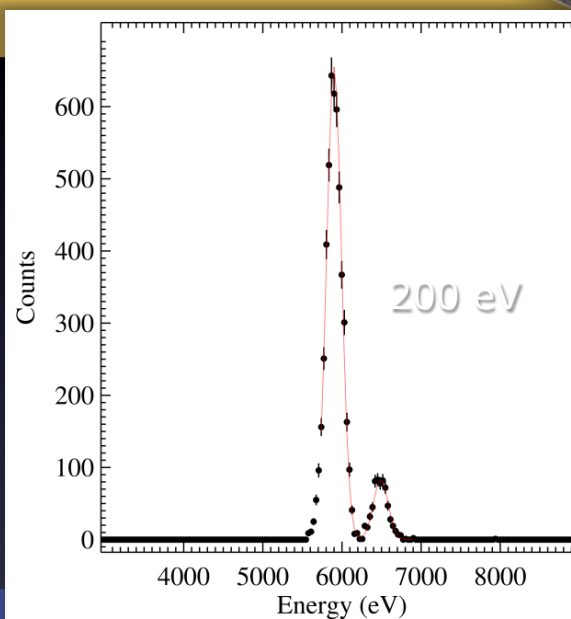
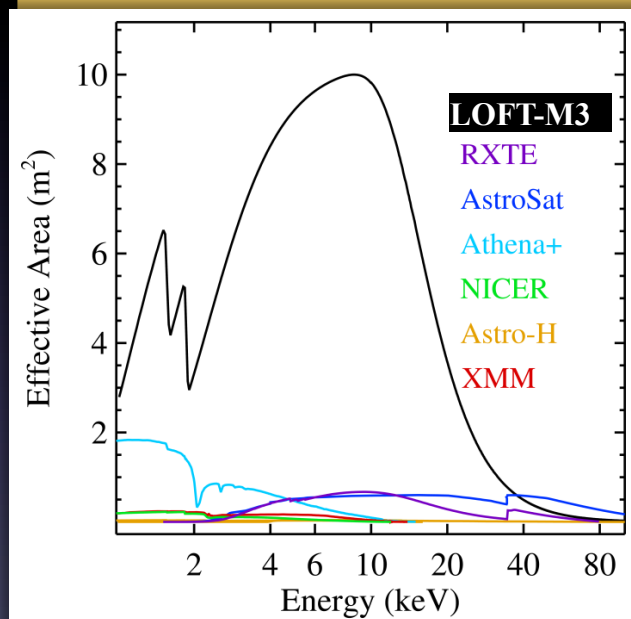
# Measuring the neutron star equation of state using x-ray timing



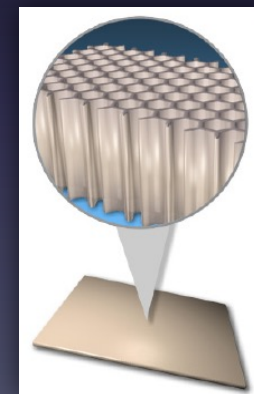




## Large Collecting Area



SDD Detectors Heritage



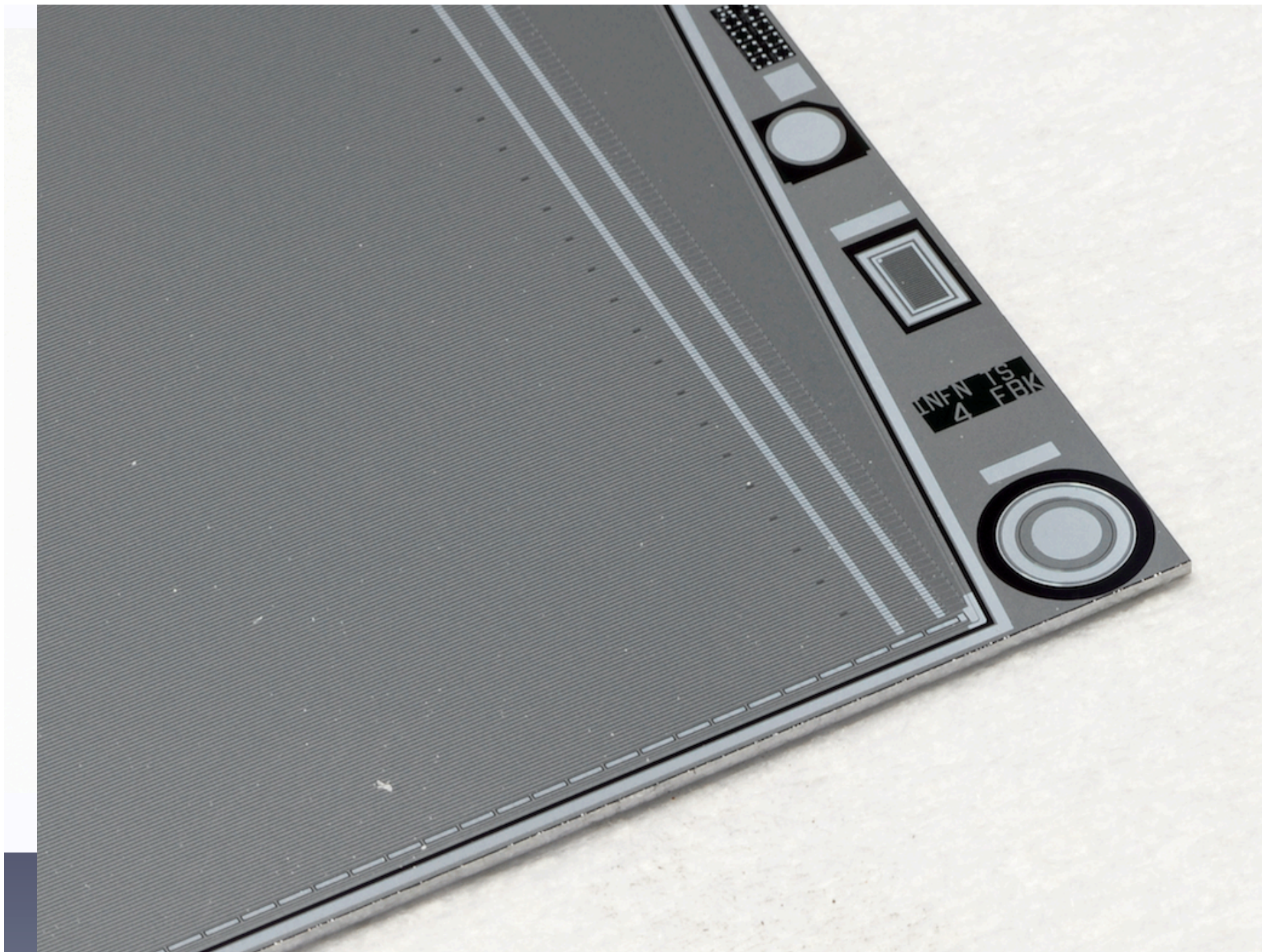
Microchannel Plate Collimators

## Good Energy Resolution (XMM-class)







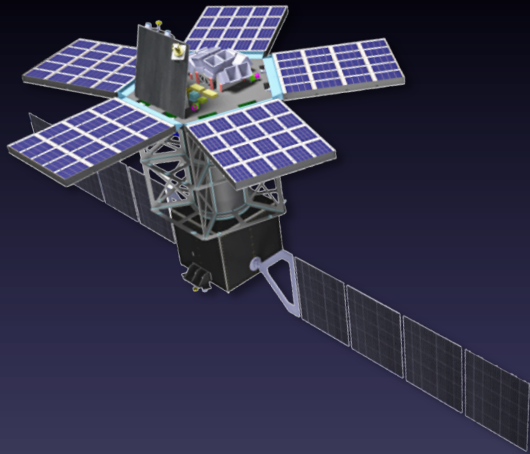




# THREE POSSIBLE MISSION APPROACHES

## LOFT

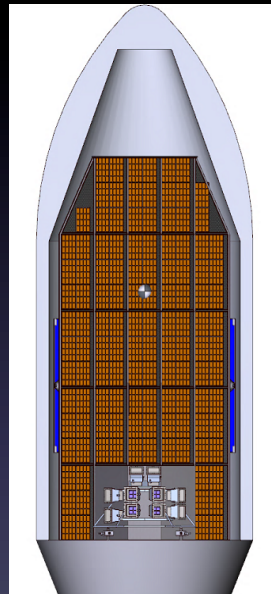
Large Observatory For x-ray Timing (ESA)



*Bright sources: Large Collimated Area*

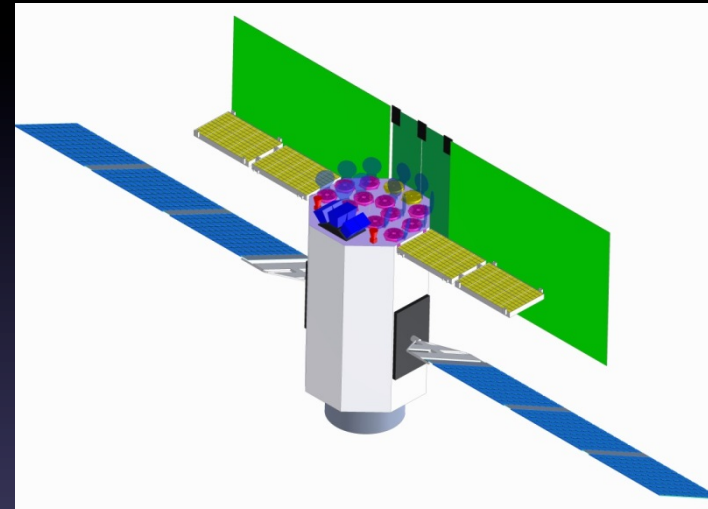
## LOFT-P

LOFT-Probe  
(NASA)



## eXTP

enhanced X-ray Timing and  
Polarization mission (CAS)



*Weak/soft sources: Collimated  
Area + Telescopes.  
And Polarimeter*



### LAD

#### 40 arrays SDD

Energy band: 2-50 keV

Collimated FOV 1deg FWHM

**Time resolution: 1us**

**Energy res.: 200eV@ 6keV**

Sensitivity: 0.01 uCrab ( $10^4$ s)

**Effective area: 3.4m<sup>2</sup>@6keV**

## Payload

### Soft Focusing Array

**11 arrays:** focal length 4.5m,

Diameter 450 mm, SDD

Energy band: 0.5-20 keV

**FOV: 12 arcmin**

Time resolution: 10 us

Energy res.: 180eV@6keV

**Angular resolution: 1 arcmin**

Sensitivity: 0.16uCrab ( $10^4$ s)

**Effective area: 6500cm<sup>2</sup>@6keV**

**0.9 m<sup>2</sup> between 1-2 keV**

### Polarimetry Focusing Array

**2 arrays** focal length 4.5 m,  
diameter 450 mm

**Energy band: 2-10 keV**

FOV: 12 arcmin

Time resolution: 500us

Energy res: 1.8keV@6keV

**Angular resolution: 15 arcsec**

Sensitivity: 5 uCrab ( $10^4$ s)

**Effective area: 250cm<sup>2</sup>@2keV**

### WFM

**3 arrays, SDD**

Energy band: 2-50 keV

FOV: 1.33PI

Time resolution: 2 us

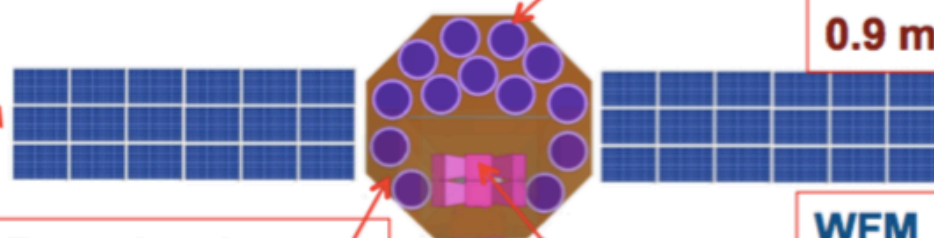
Energy resolution:

300eV@6keV

Angular resolution: 4.5 arcmin

Sensitivity: 3uCrab ( $2 \times 10^4$ s)

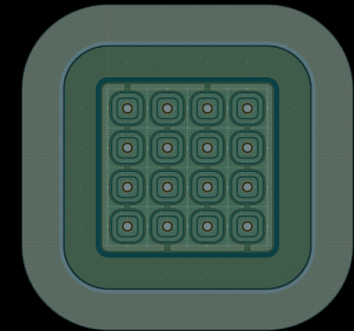
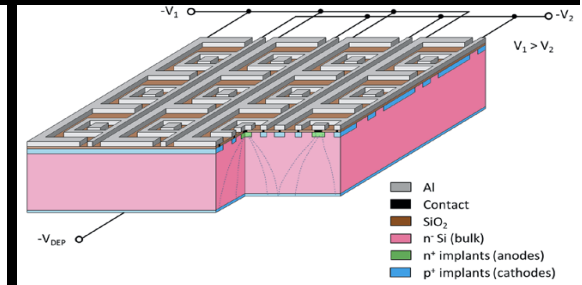
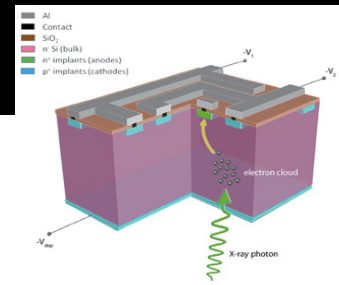
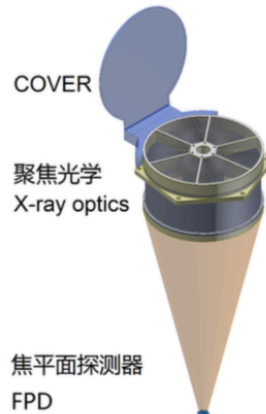
Effective area: 170cm<sup>2</sup>@6keV



# Pixel SDD

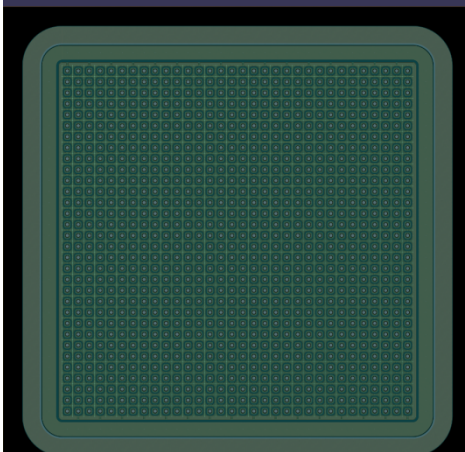
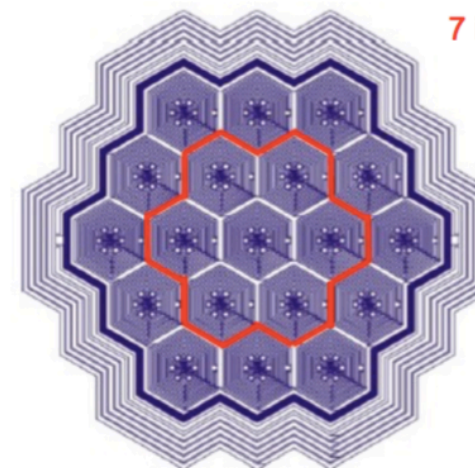
- SFA for spectroscopic and timing measurement with imaging capability.

Focal length	4.5 m
Area	>500 cm <sup>2</sup> @ 6 keV, 1 module >6000 cm <sup>2</sup> @ 2-6 keV, in total
Energy range	0.5~20 keV
Field of View	+/-6 arcmin
Angular res.	HEW≤1', W90≤3'
Energy res.	<180eV@6keV
Timing res.	10 us



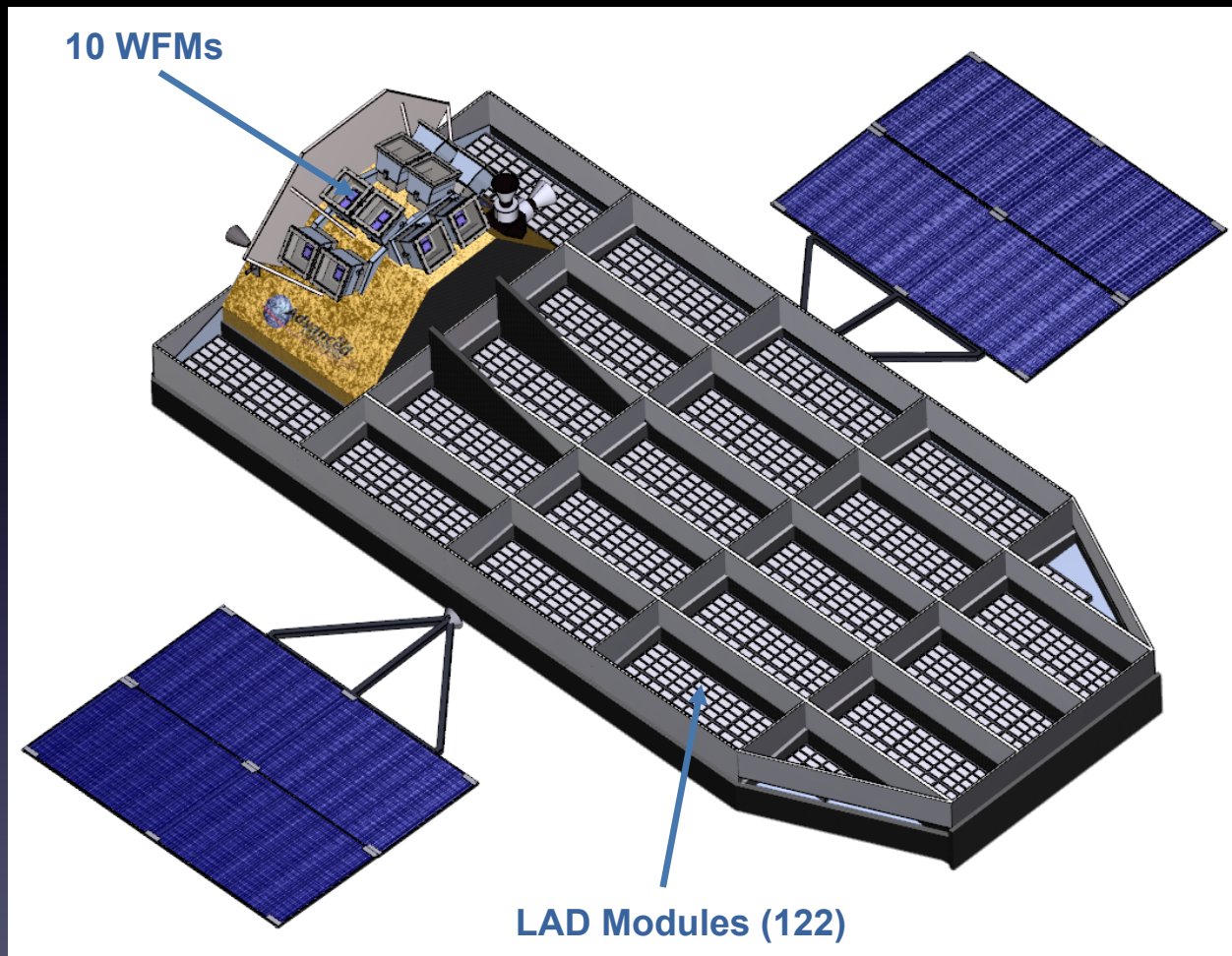
- The baseline for the SFA focal plane detector (FPD) is a 7 pixels SDD. The pixel size shall not exceed 3'.
- It is agreed that the baseline FEE is based on discrete components.

7 instead of 19 pixels





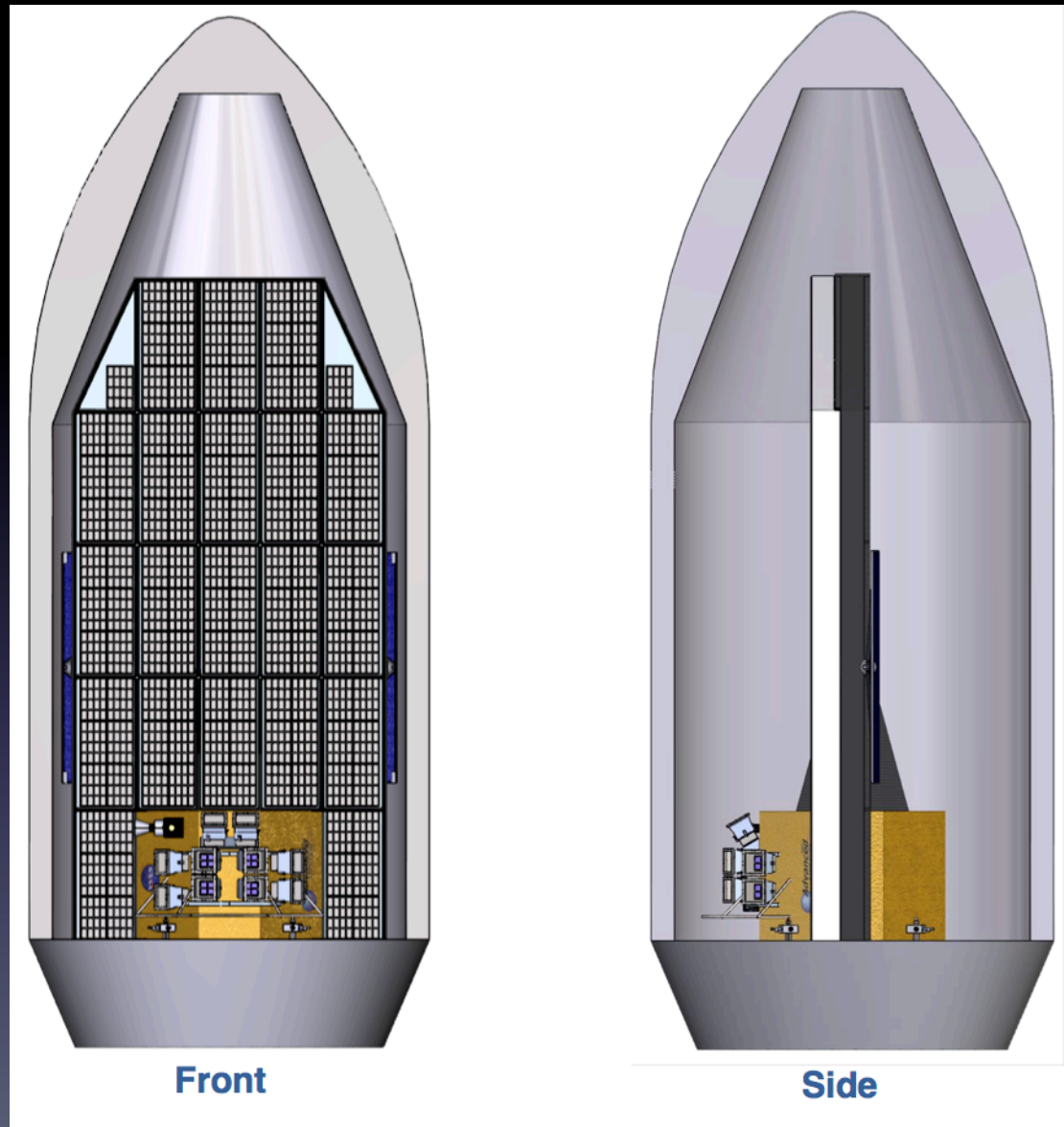
## LOFT-P



Credits: NASA/MSFC and ACO

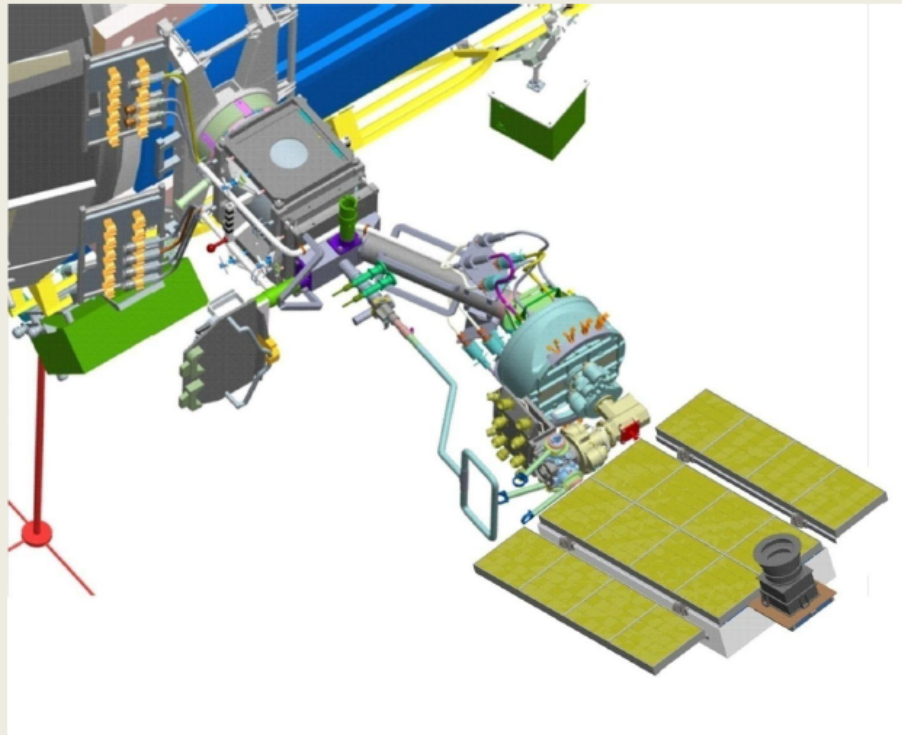


# LOFT-P



LOFT-pathfinder  
1-m<sup>2</sup> experiment on ISS (Russian segment) to be launched in 2020

## Rotary pointing system on the Russian segment of the ISS



At the moment the Canadian telescopes successively operates





### eXTP

- Pre-selected in China, currently in phase 0/A. Possible selection in 2016 or 2017, for a launch in 2024-2025. Participation of the whole LOFT consortium + INAF/OAB + MPE.

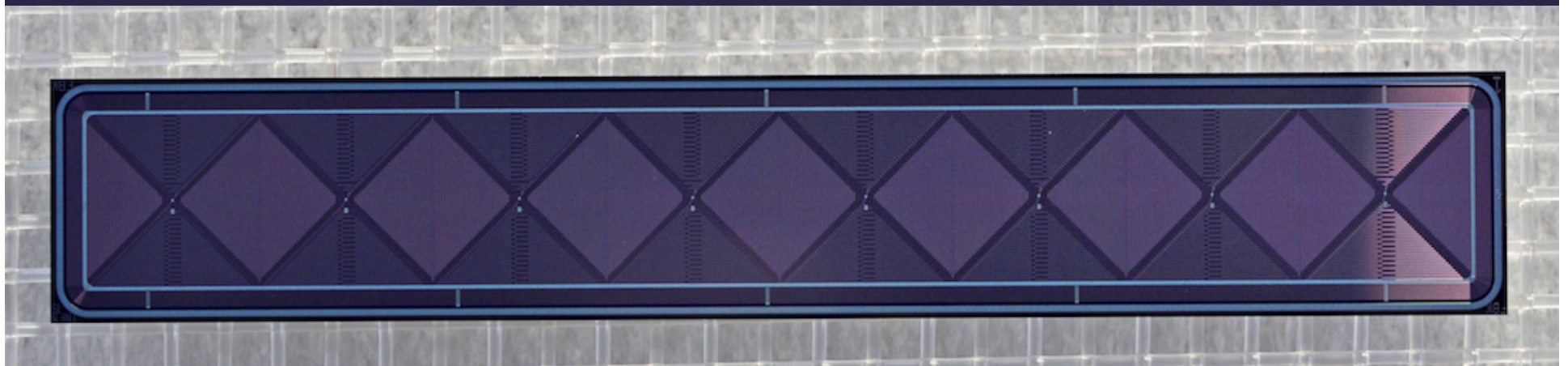
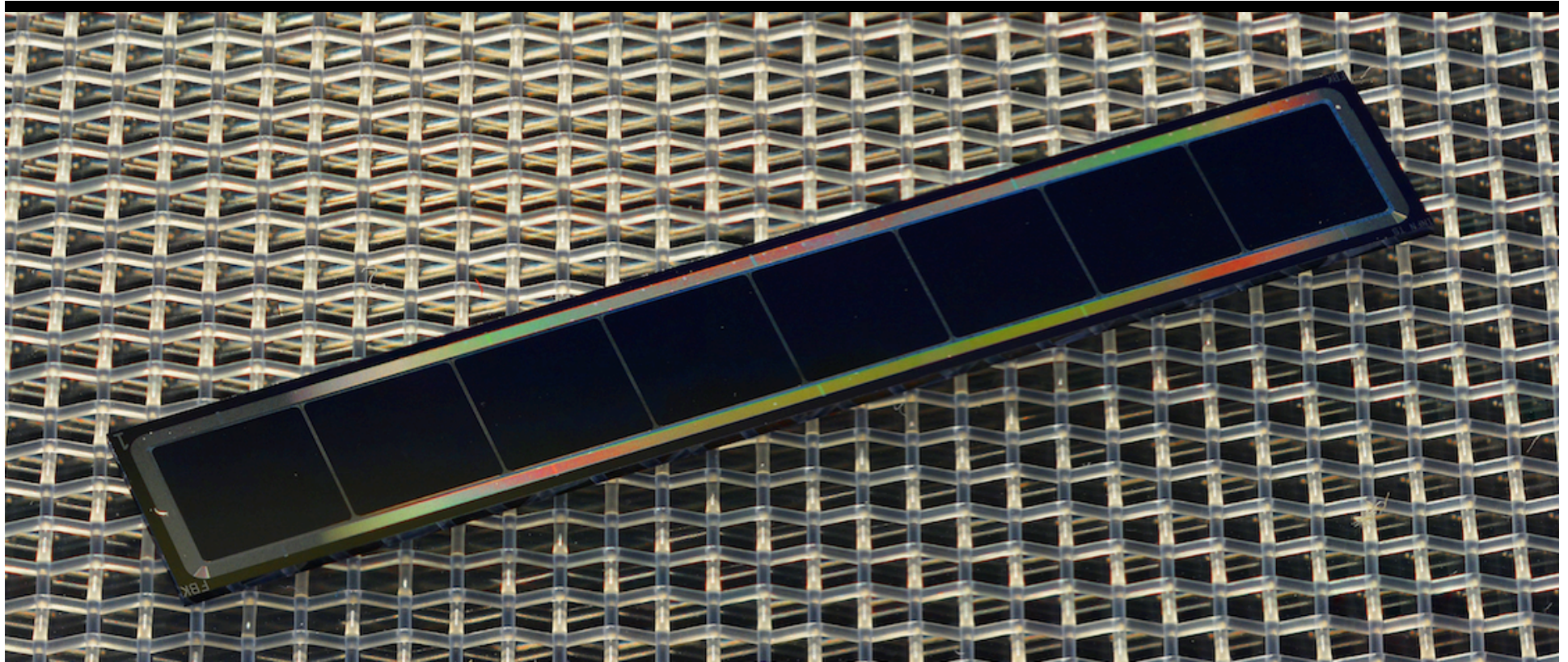
### LOFT-P

- Preparation activities for the Decadal Survey 2020. NASA/MSFC-funded mission study currently ongoing. Probe-class. Possible notional mission call end-2016/early-2017. Real mission call early-2020's. Launch in late-2020's.

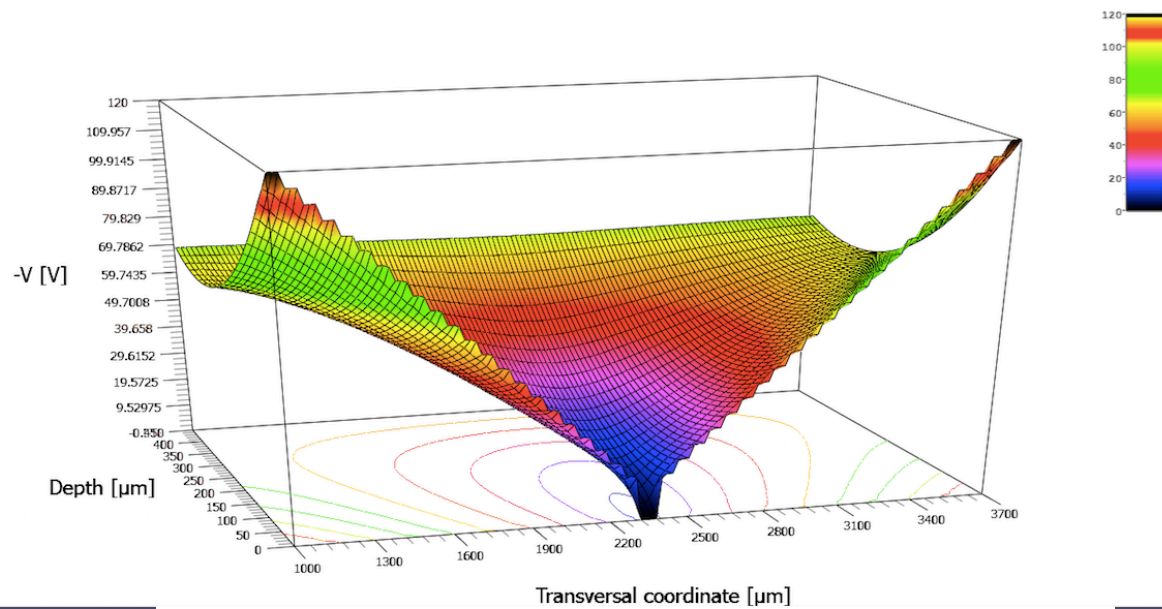
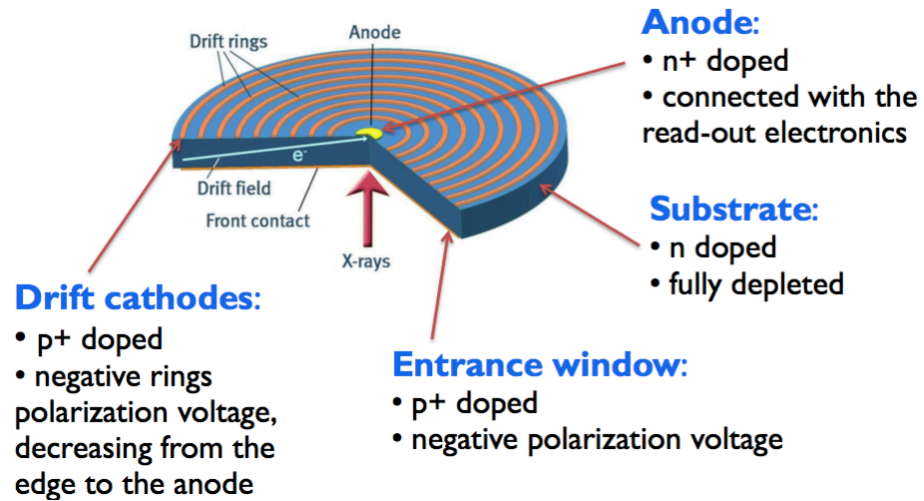
### MVN-M2 on ISS

- Experiment led by IKI (Academy of Sciences, Moscow). PI: M. Pavlinski. ROSCOMOS industrial partner: Energya. Approved and funded in Russia. Phase 0/A/B1: Jan 2016 – Nov 2017. Launch end-2020.



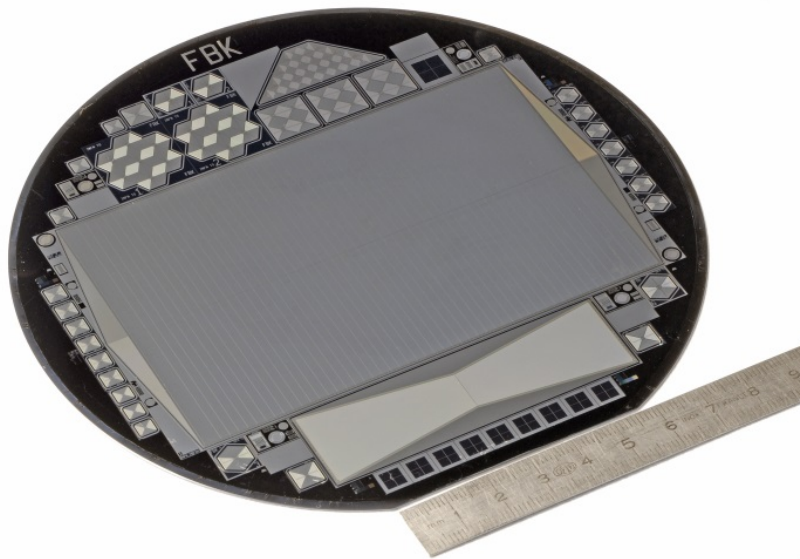
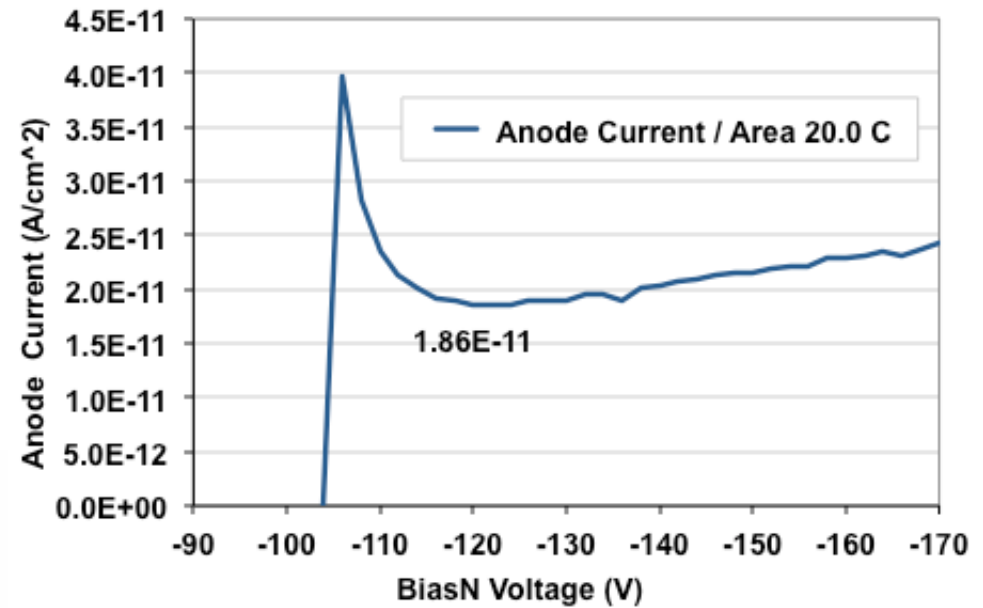








# FBK Trento Low leakage current



The SDD, with an active area of 13 mm, has been manufactured by optimizing the production processes in order to reduce the anode current, successfully reaching leakage current densities between 19 pA/cm and 25 pA/cm at 20 C.

TABLE I  
X-RAY SDD PERFORMANCE AT ROOM TEMPERATURE

Reference	year	Detector area (mm <sup>2</sup> )	Temperature (°C)	5.9 keV <sup>55</sup> Fe FWHM (eV)	ENC (e <sup>-</sup> r.m.s.)	Peaking time (μs)
This work	2015	13	+20	136	7.4	1.4
[8]	2014	13	+21	141	8.6	0.8
[9]	2012	25	+25	260	27	0.1
[10]	2001	10	+25	300	--	--
[11]	1996	3.5	+27	220	21	0.5
[12]	1994	2	+24	(374)	41	--
[13]	1994	1.5	+20	327	--	7
[13]	1994	0.5	+20	267	--	7
[14]	1992	78	R.T.	940	110	0.25

room temperature state of the art, our collaboration

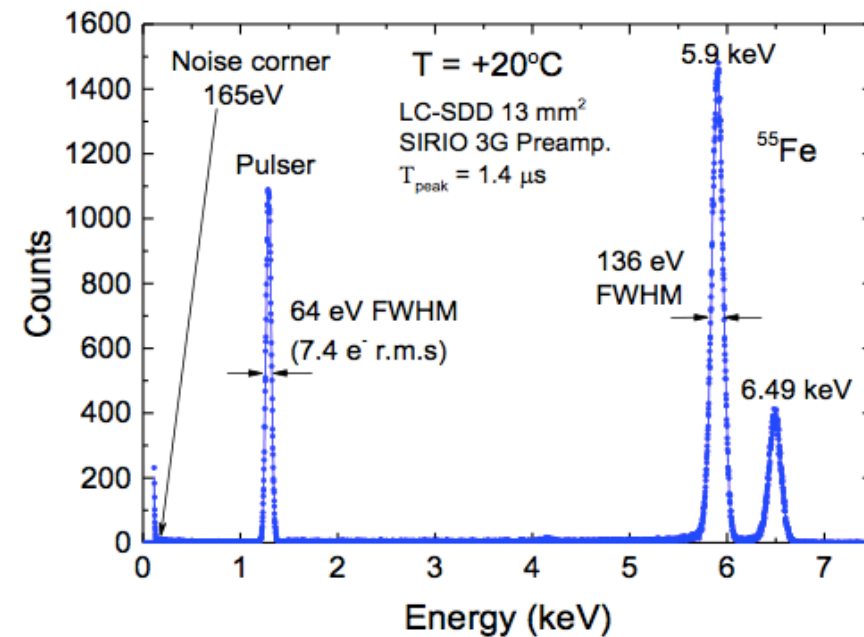


Figure 6 <sup>55</sup>Fe acquired at +20°C and optimum peaking time (1.4 μs). The pulser line width is 64 eV FWHM, corresponding to 7.4 electrons r.m.s..

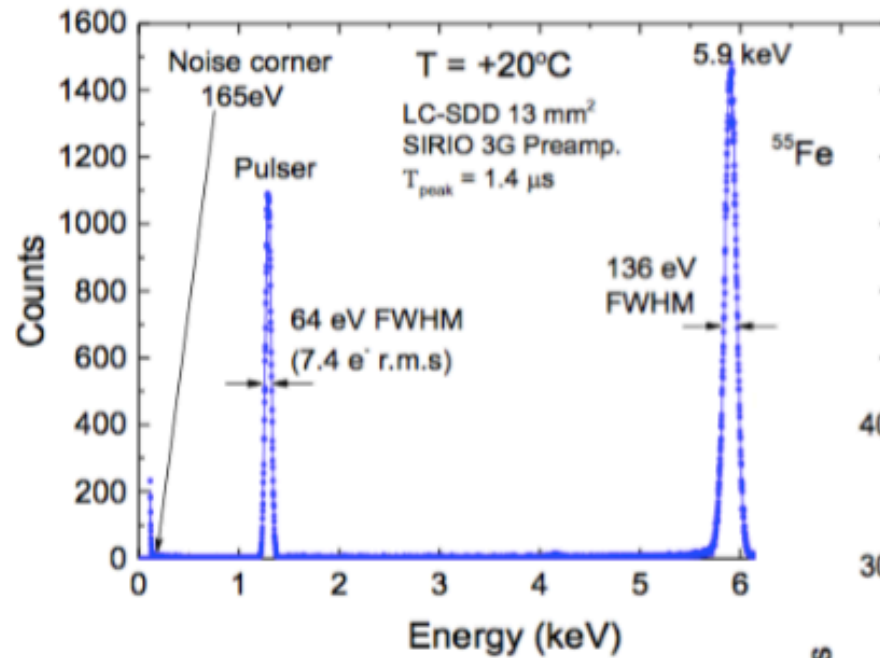


Figure 6  $^{55}\text{Fe}$  acquired at  $+20^\circ\text{C}$  and optimum peaking time. The pulser line width is 64 eV FWHM, corresponding to 7.4 e<sup>-</sup> r.m.s..

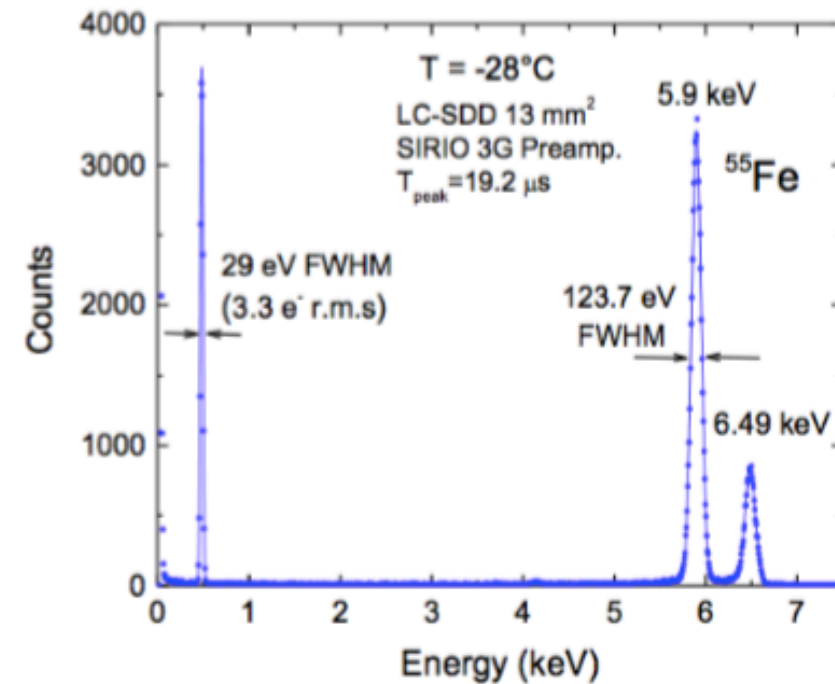
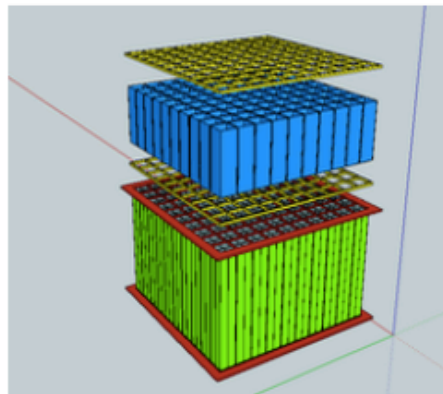
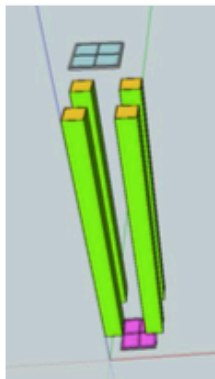
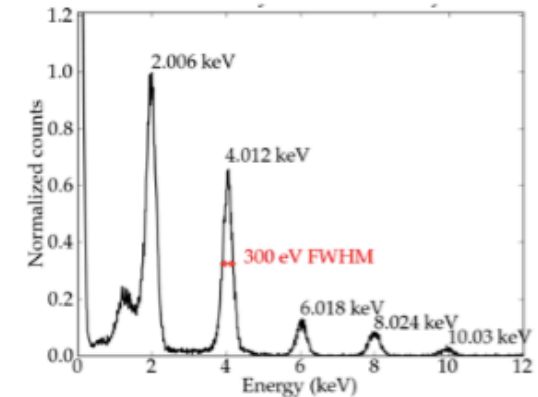
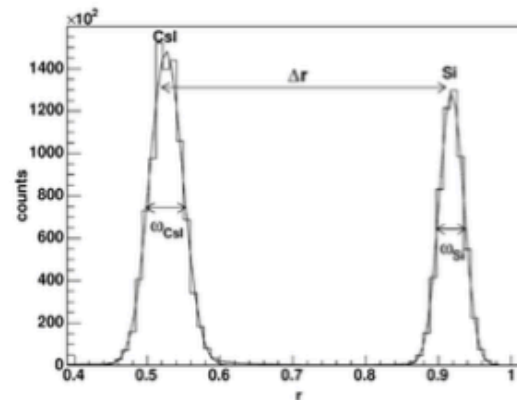
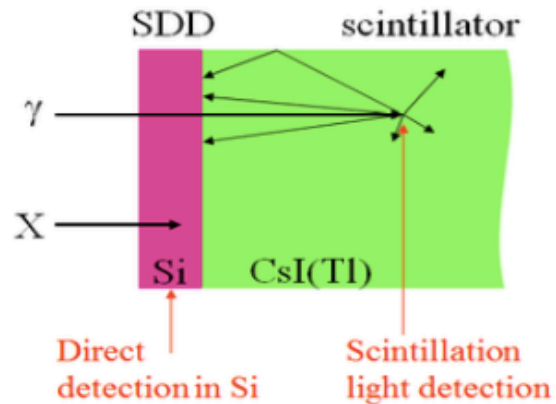


Figure 7.  $^{55}\text{Fe}$  acquired at  $-28^\circ\text{C}$  and optimum peaking time (19.2  $\mu\text{s}$ ).. The pulser line width is 29 eV FWHM, corresponding to 3.3 electrons r.m.s..

5/10/16

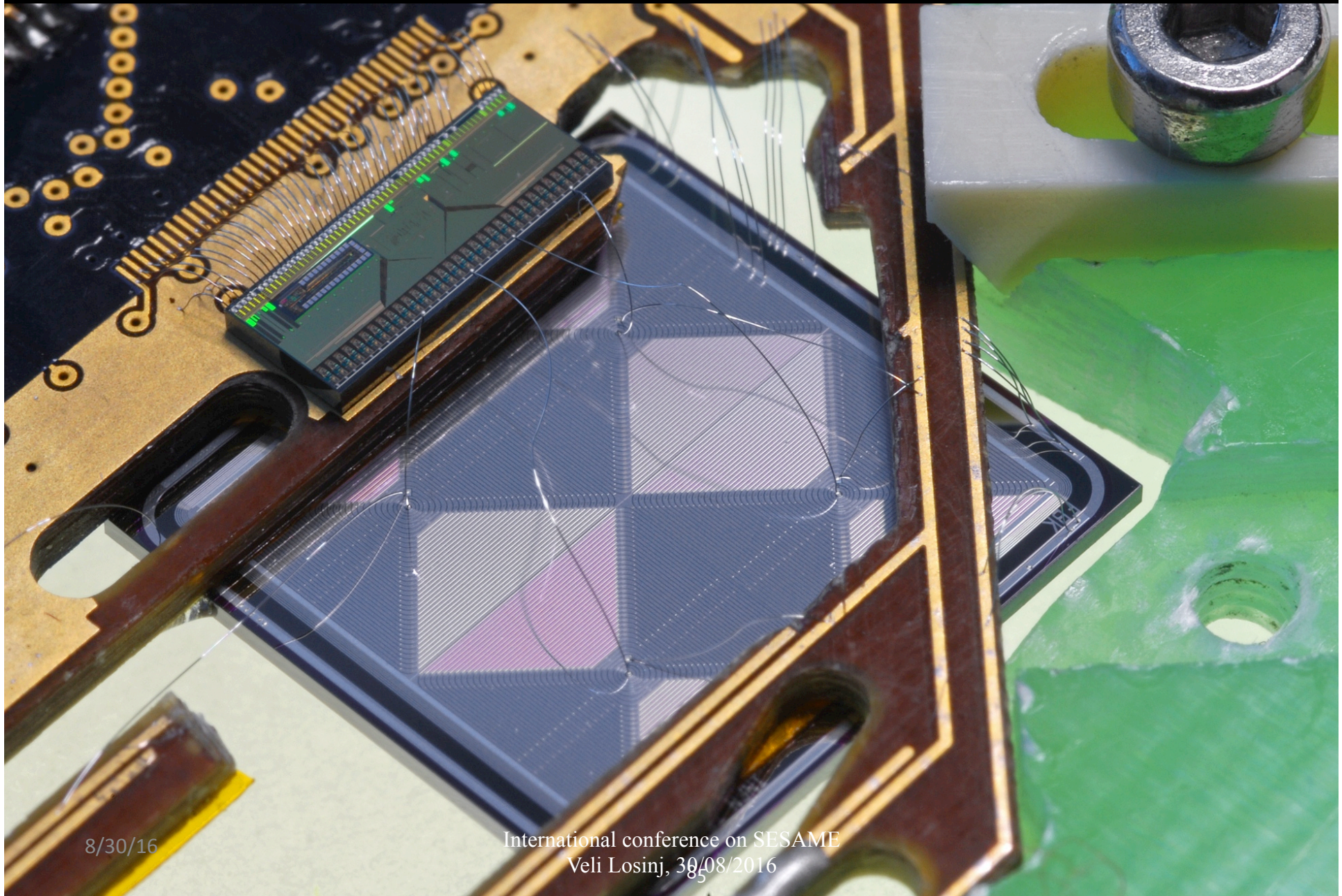


# The X-Gamma-rays spectrometer (XGS)



Energy band (keV)	2-20000
Detection principle	Siswitch
Low-energy detector	Silicon Drift Detector
High energy detector	CsI(Tl)
Separation of energy losses in Si and CsI(Tl)	Pulse shape analysis
Number of modules	25
Size of each module (mm <sup>3</sup> )	130x130x100
Lateral passive shielding/module	0.5 mm W
Slat collimator/module	0.5 mm W
Overall collimator FOV (FWHM)	
Average useful area in the SXI FOV (cm <sup>2</sup> )	~2000



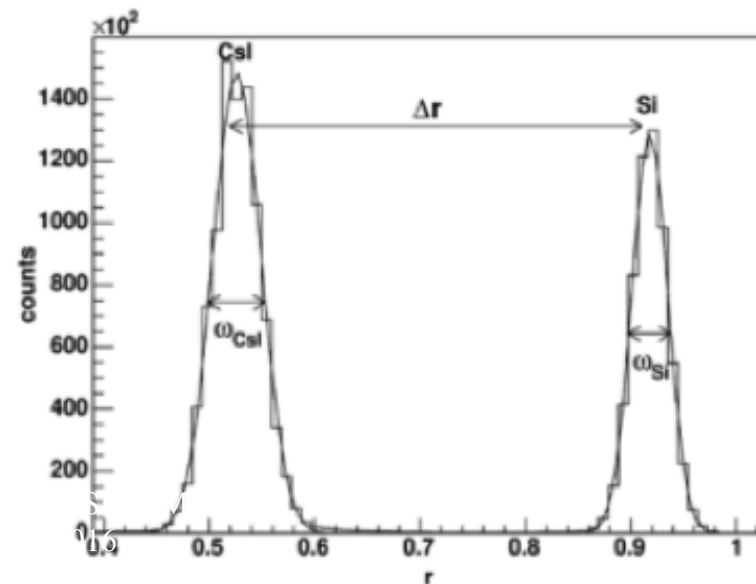
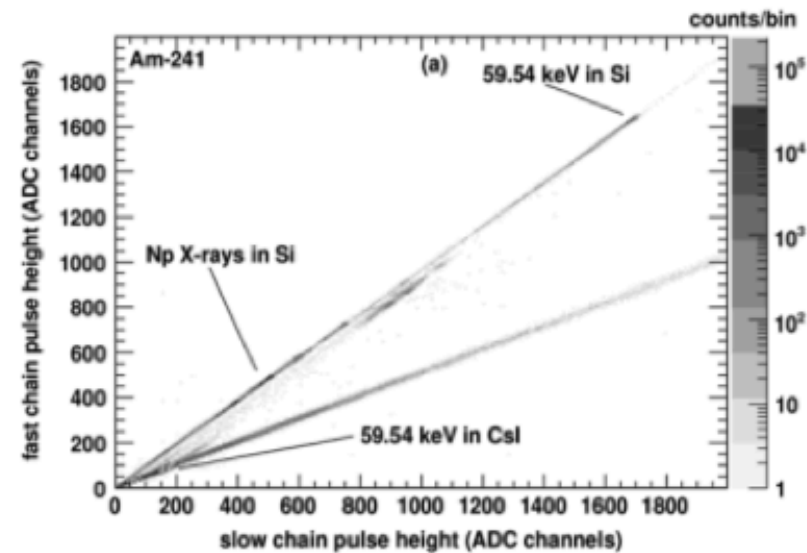
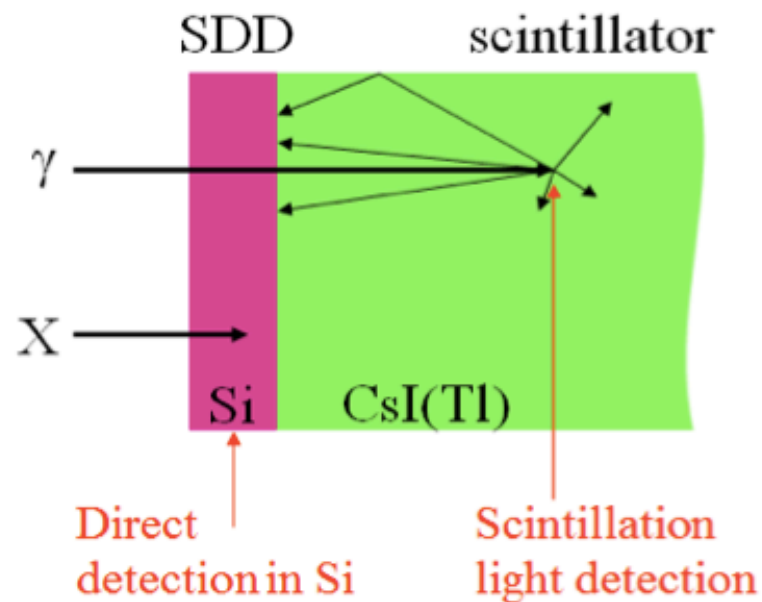


8/30/16

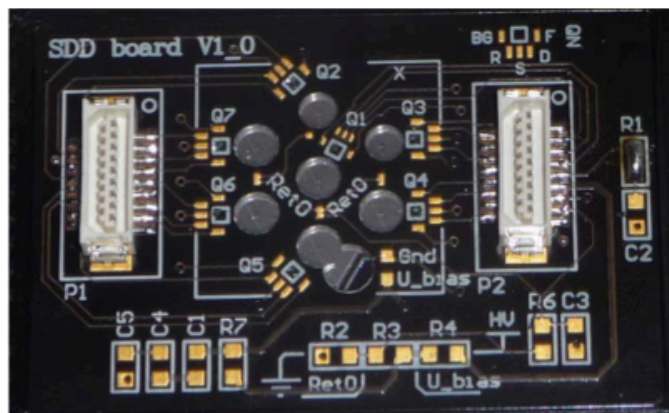
International conference on SESAME  
Veli Losinj, 30/08/2016



# XGS – X and Gamma-ray Spectrometer detection principle

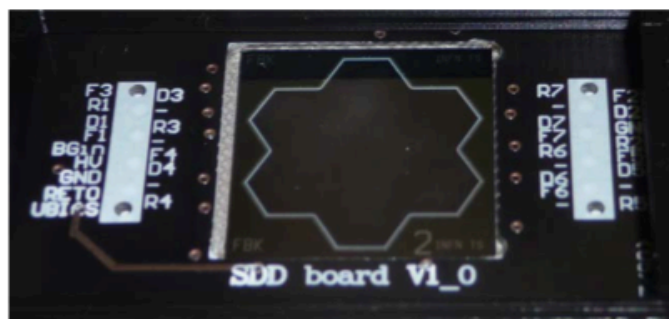
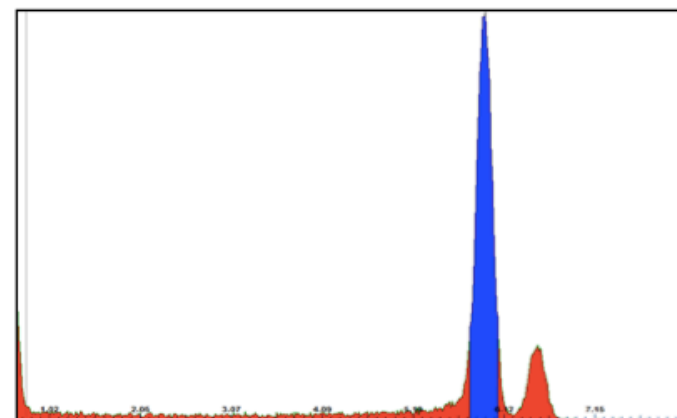






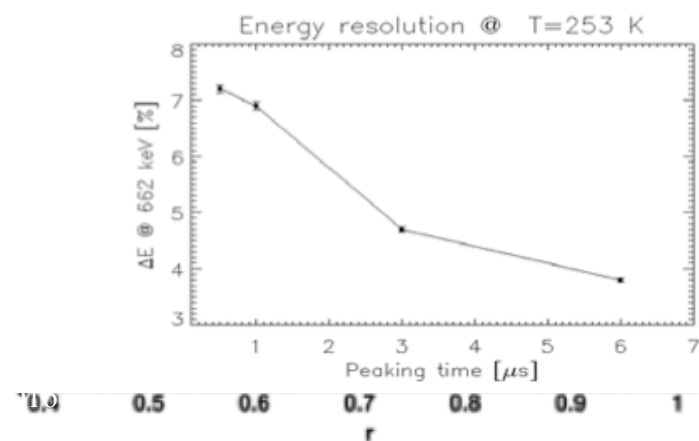
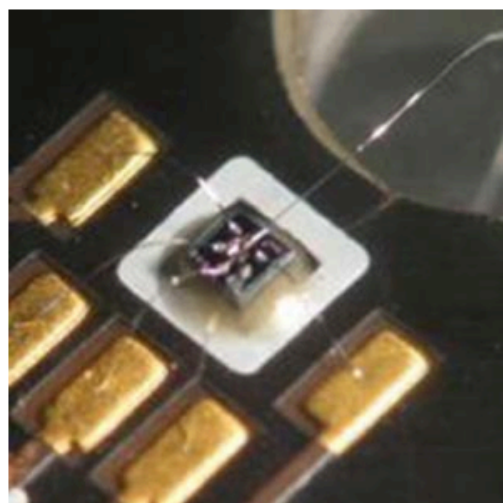
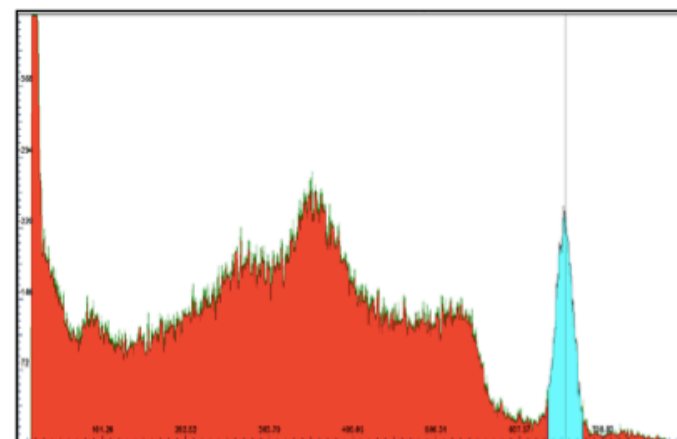
SDD only  
Fe55  
@253K

170 eV

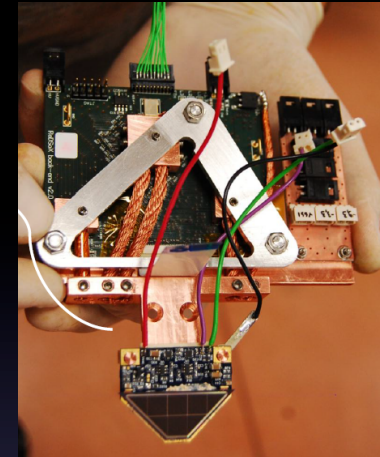
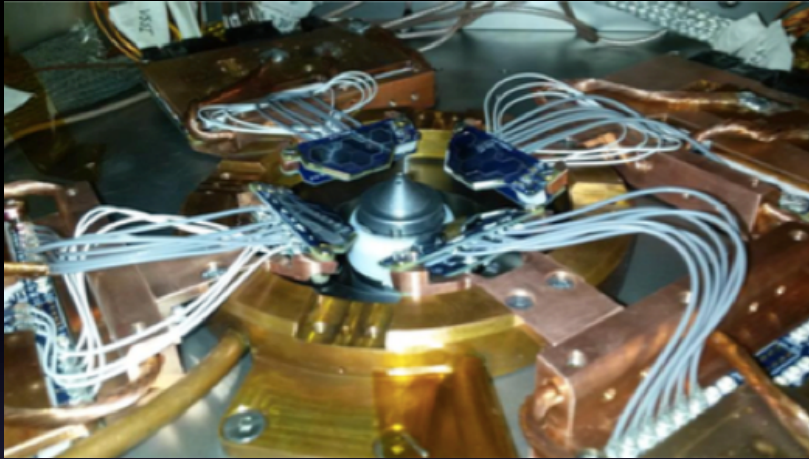


SDD+CsI  
Cs137  
@253K

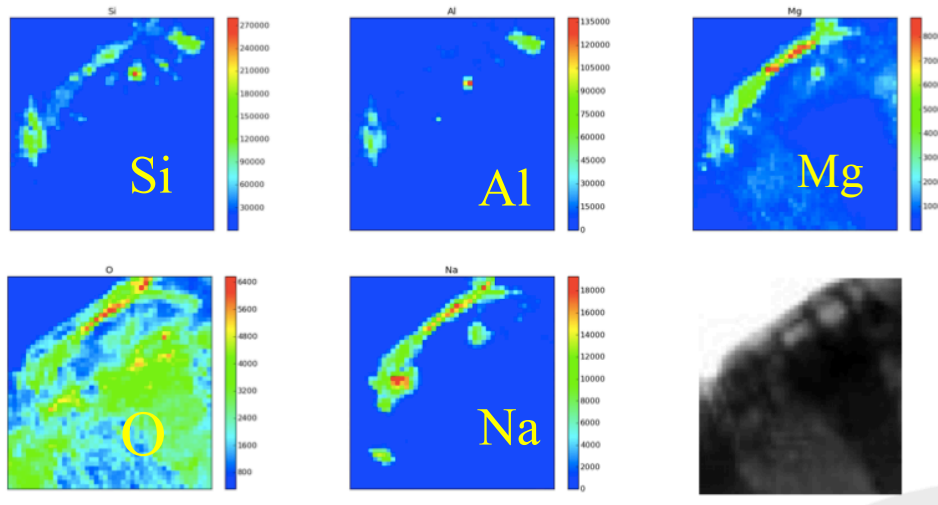
3.8%



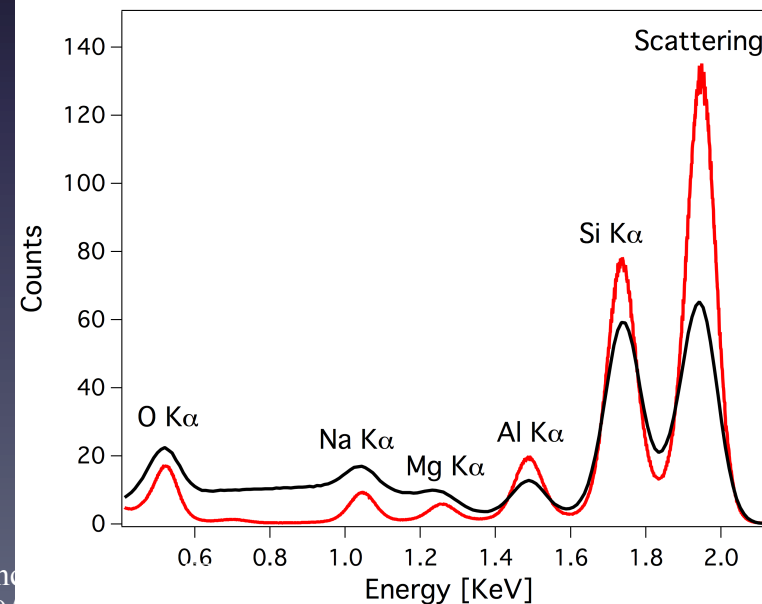
TwimMic (Elettra) beam line  
maps of a soybean root section  
acquired with the new detector system,  
realized within the REDSOX collaboratio



6 SDD cells



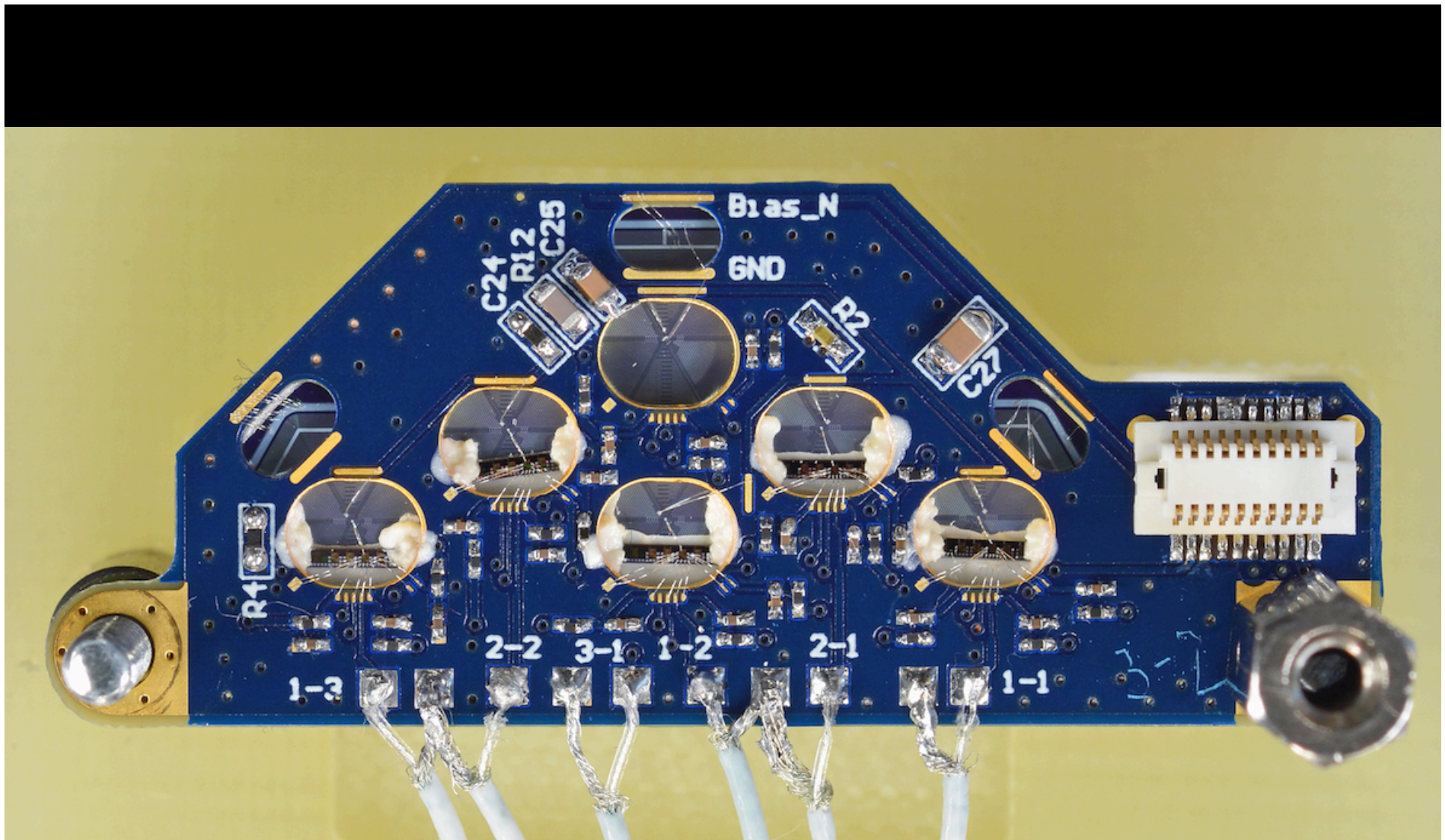
Al-Mg-Si-O-Na spectrum of a soybean root section  
acquired with the new TwinMic detector system.



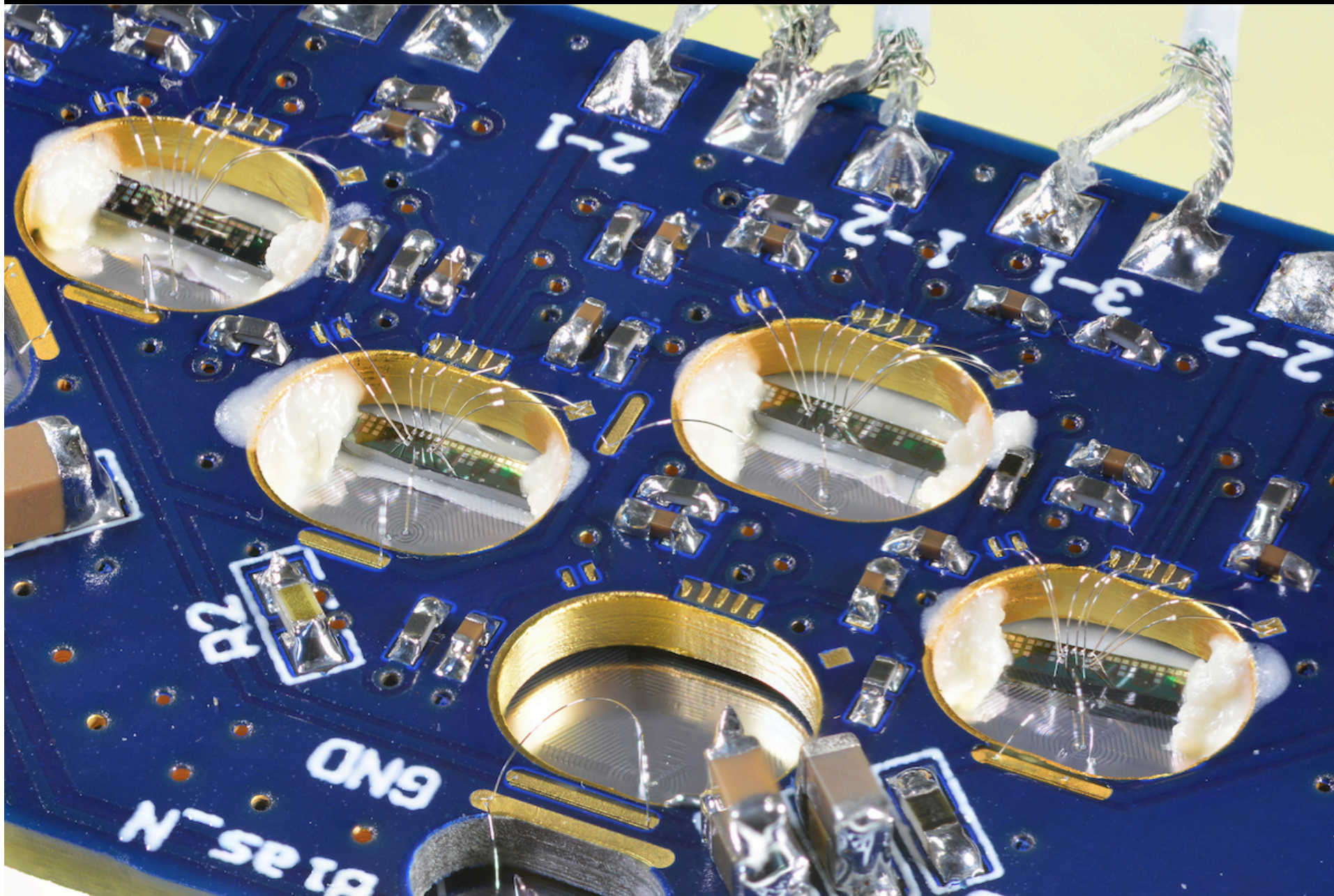














Bruxelles, 30<sup>th</sup> May, 2016

For the International Center for Synchrotron-light for Experimental Science and Applications  
in the Middle East

Khaled Toukan  
Prof. Khaled Toukan,  
Director

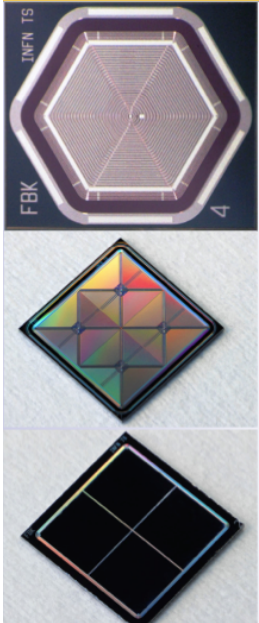
For Istituto Nazionale di Fisica Nucleare

Ferroni  
Prof. Fernando Ferroni  
President





# Our INFN collaboration specialized in the development of state of the art Silicon Drift Detectors will carry out the work for SESAME



## Outline

- 1) front edge unique results
- 2) working principle
- 3) SESAME detector

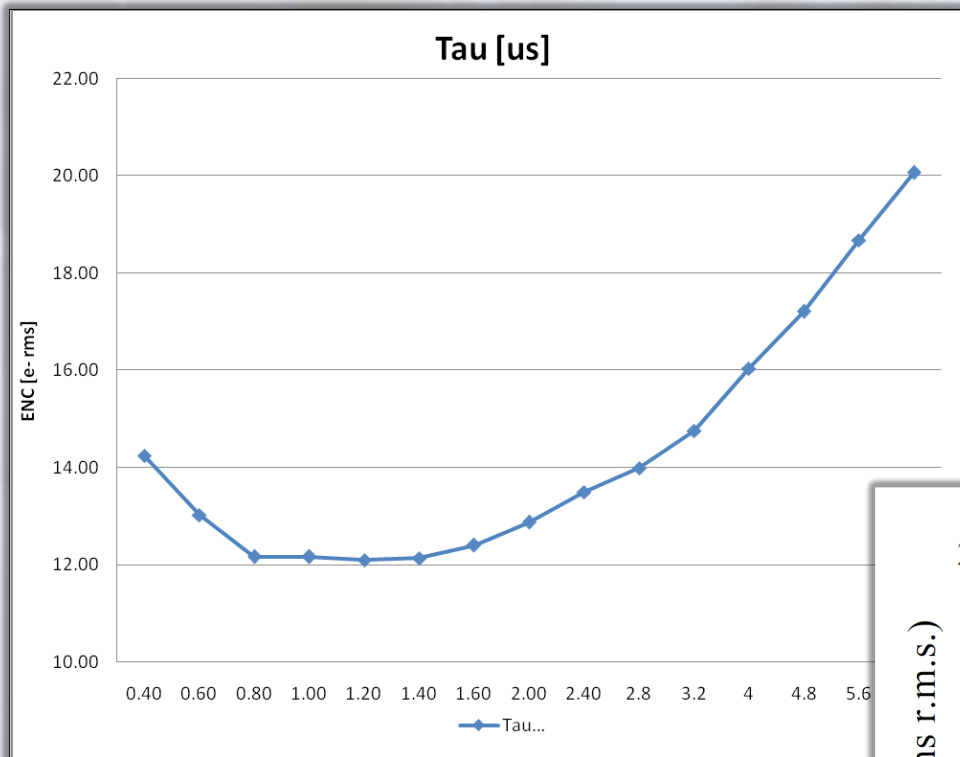


INFN Trieste  
FBK Trento  
Elettra Sincrotrone Trieste  
INAF IASF INFN Roma 2  
INFN INAF Bologna  
INFN Pavia  
INFN TIFPA Trento  
ICTP Mlab Trieste  
INFN Firenze Labec  
Poli Milano El.&Det.

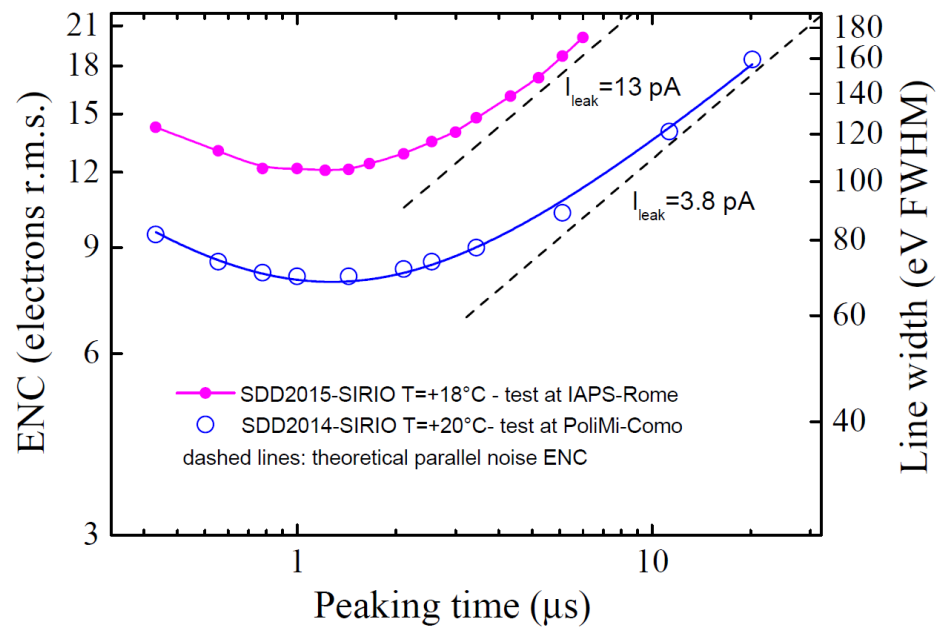


# The new detectors for XAFS beamlines



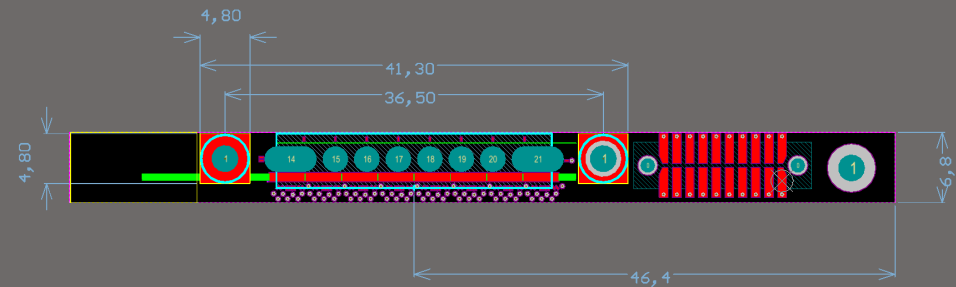
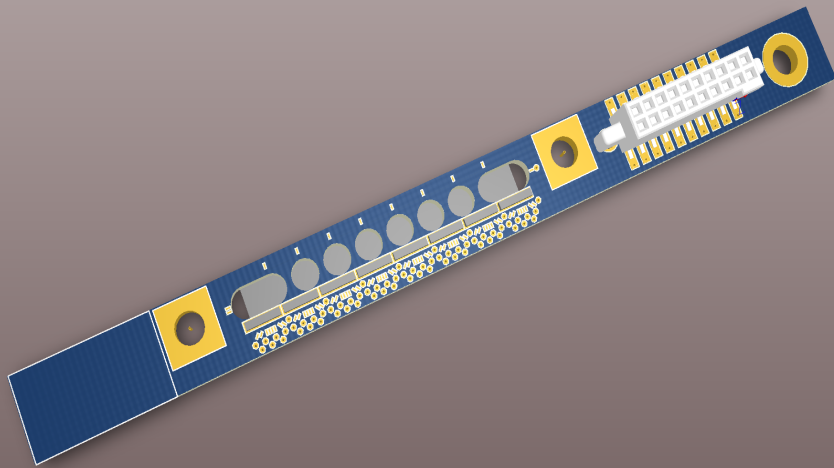


ENC vs  $\tau$  @ 18° C  
 Trapezoidal shaping (flat-top 200ns)  
 Pulse amplitude 4.98 keV (eq.).

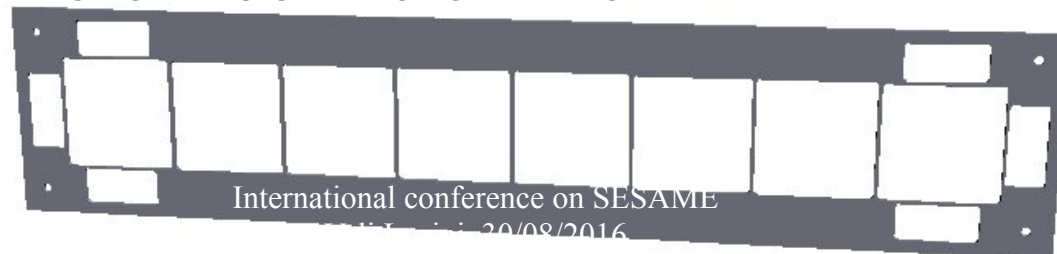




# Prototype – 2 : work in progress

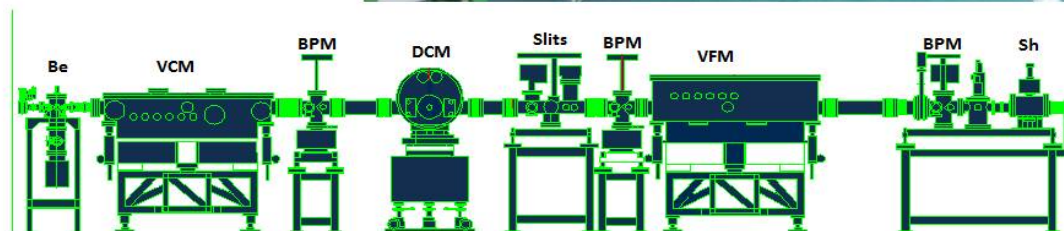


- Tungsten collimator 100 micron thick



# SESAME XAFS-Beamline (ROBL-ESRF)

Beamline Components set up in lab  
Under Test with ESRF experts



XAFS and  $\mu$ -XAFS  
(X-ray absorption fine structures),

XAFS beam-line combines X-ray absorption spectroscopy and X-ray diffraction to provide chemically specific structural information of materials.

Journal of Physics: Conference Series 689 (2016) 012017

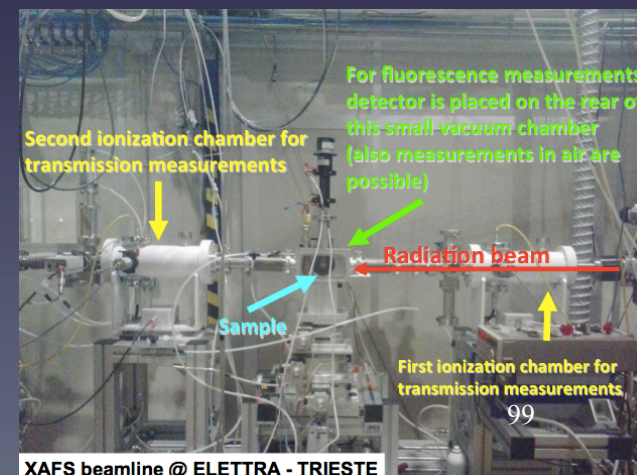
among the XAFS applications: cultural heritage,  
environmental analysis



8/30/16

Paleolithic paintings from caves in Spain

International conference on SESAME  
Veli Losinj, 30/08/2016



XAFS beamline @ ELETTRA - TRIESTE

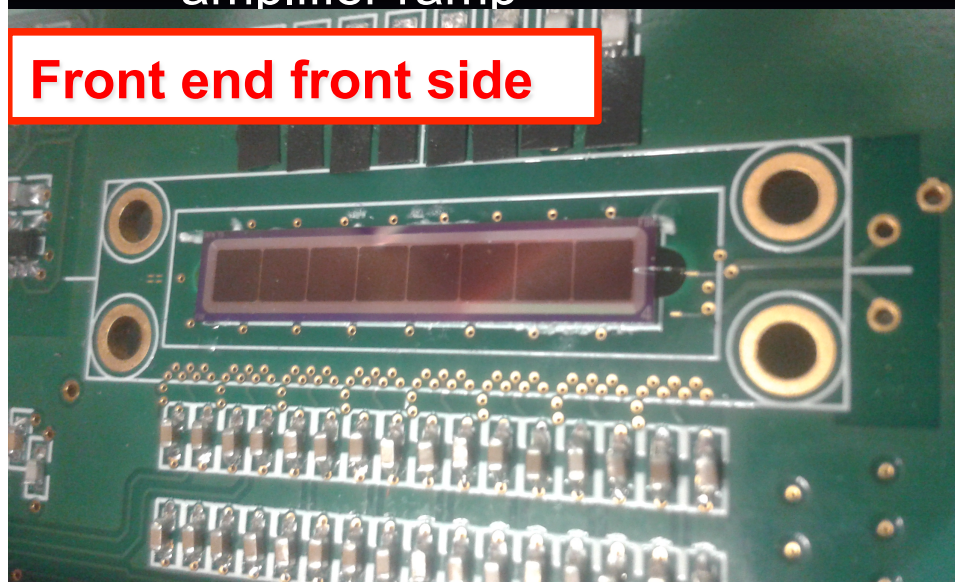


# SESAME Prototype - 0 : the first beam test at XAFS Elettra

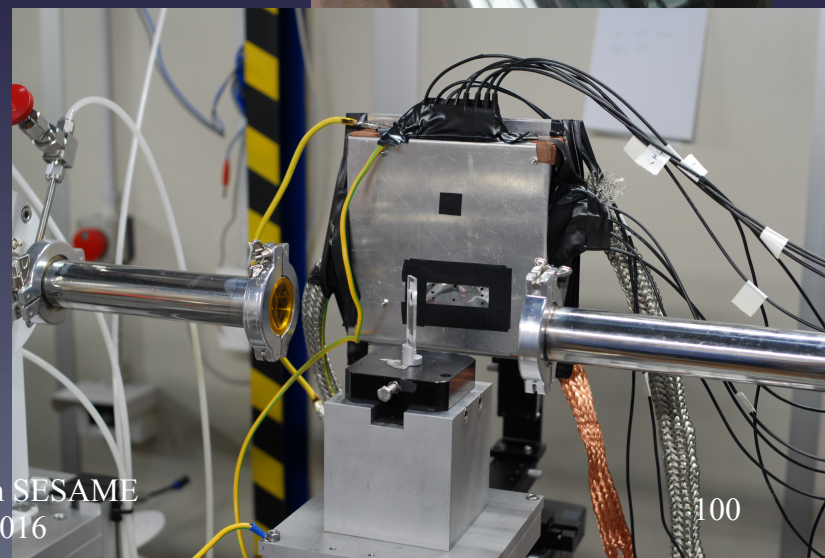
The prototype comprises:

1 SDD with 8 cells mounted on a  
font end electronics (by INFN-Ts)  
Signal output unplified pre-  
amplifier ramp

**Front end front side**



back end electronics equipped  
with 8 ADCs for each independent  
channel and an FPGA (by Elettra)





Thank you  
for your attention  
[vacchi@ts.infn.it](mailto:vacchi@ts.infn.it)

