

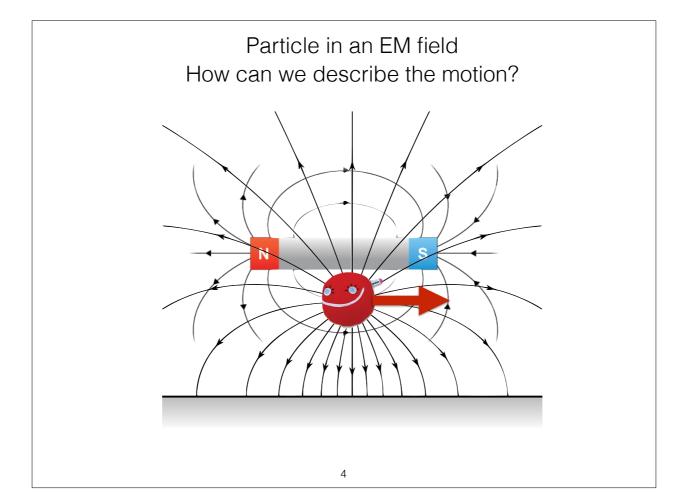
## Aims for this lecture

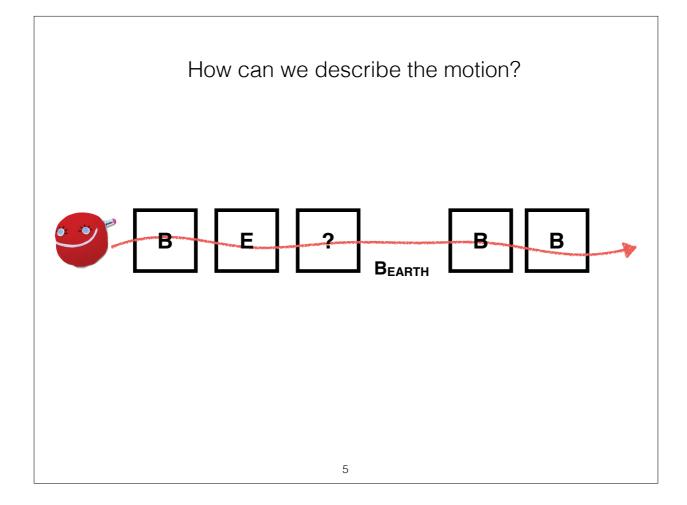
From the CAS Syllabus:

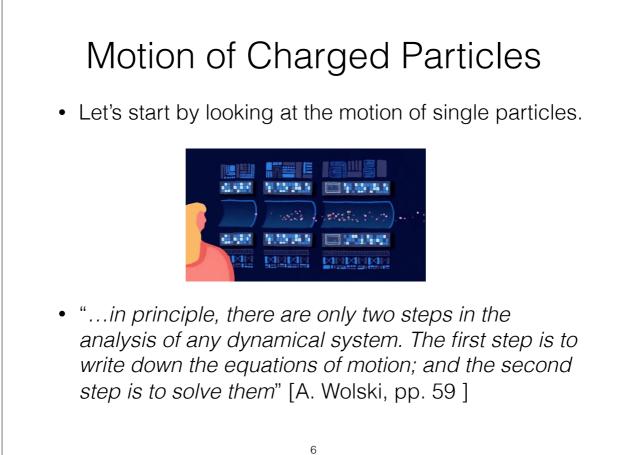
1. Arrive at a general description of particle motion in EM fields

2. Understand what "maps" are, and how they relate to particle motion and simulation

3. Derive some basic maps from the equations of motion



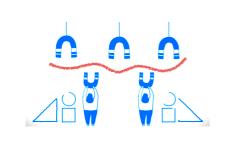




• Why not just use Newton's laws & Lorentz force?

$$\frac{d\mathbf{p}}{dt} = \mathbf{F} \qquad (1.1) \qquad \mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B}) \qquad (1.2)$$

- In an accelerator, magnets and rf cavities are generally defined along a trajectory (i.e. we know where they are in distance, not time). The Hamiltonian approach is lets us use this fact.
- The motion of particles in electromagnetic fields is conservative, and similar to a harmonic oscillator with perturbations. We have lots of mathematical tools to treat this in the *Hamiltonian formalism*.
- Ultimately, it makes our lives easier.



If we know the Hamiltonian and Hamilton's equations, we can find the equations of motion for a dynamical system.

Our first goal: find out the Hamiltonian!

# Hamiltonian (straight beam line)

- The Hamiltonian represents the total energy of the particle
- We need a Hamiltonian that gives (1.1) and (1.2) when substituted into (1.3) and (1.4)

 $H = H(x_i, p_i; t)$ 

Hamilton's equations  

$$\frac{dx_i}{dt} = \frac{\delta H}{\delta p_i} \qquad (1.3)$$

$$\frac{dp_i}{dt} = -\frac{\delta H}{\delta x_i} \qquad (1.4)$$

We propose the following Hamiltonian for a relativistic charged particle moving in an electromagnetic field:

$$H = c\sqrt{\left(\mathbf{p} - q\mathbf{A}\right)^2 + m^2 c^2} + q\phi \quad (1.5)$$

But this is still defined in terms of time... so we want to change it.

Hamiltonian (straight beam line)

(1.5) 
$$H = c\sqrt{(\mathbf{p} - q\mathbf{A})^2 + m^2c^2} + q\phi$$

We don't need to go through every step here, but here's what we do.

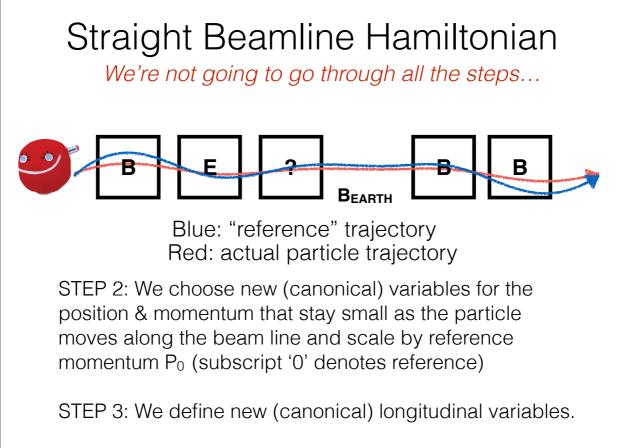
STEP 1: We take (1.5) and change the independent variable to z, the distance along the beam line

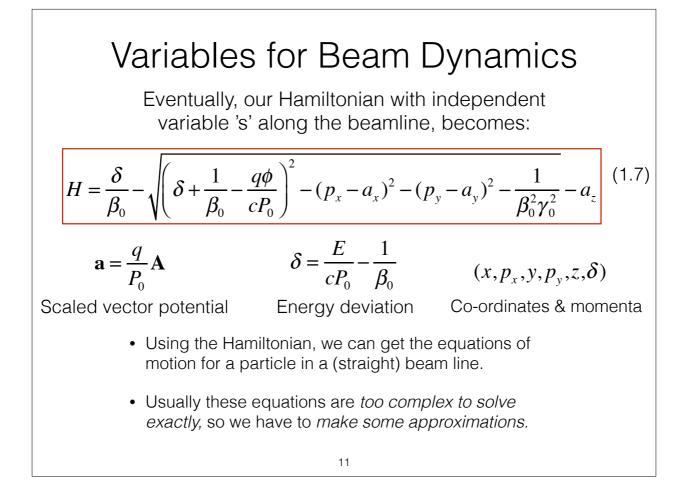
 $(x, p_x)$   $(y, p_y)$  (t, -E)

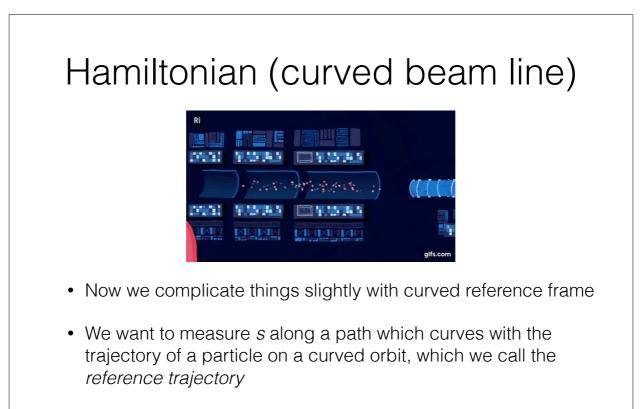
We get (try this at home...), use H= total energy E

$$H_{new} = -p_z = -\sqrt{\frac{(E - q\phi)^2}{c^2} - (p_x - qA_x)^2 - (p_y - qA_y)^2 - m^2c^2} - qA_z$$
(1.6)

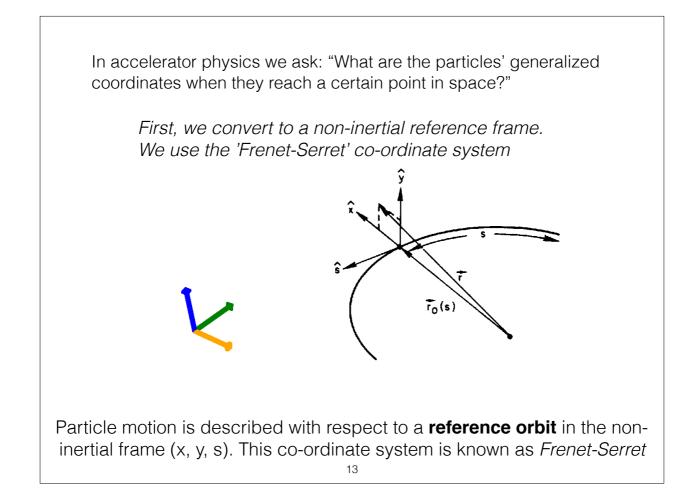
Note that if the E and B fields are static, the Hamiltonian is independent of time and the total energy of the particle is constant. Happy days.







• Spoiler alert: simply turns out as a factor in front of the straight line Hamiltonian



• First, we convert to 'Frenet-Serret' co-ordinate system  

$$\hat{s}(s) = \frac{d\tilde{r}_0(s)}{ds}$$
 Tangent unit vector to closed orbit  
 $\hat{x}(s) = -\rho(s)\frac{d\hat{s}(s)}{ds}$  Unit vector perpendicular to tangent vector  
 $\hat{y}(s) = \hat{x}(s) \times \hat{y}(s)$  Third unit vector...  
Particle trajectory:  $\vec{r}(s) = \vec{r}_0(s) + x\hat{x}(s) + y\hat{y}(s)$   
nb. the reference frame moves WITH the particle  
 $H = e\phi + c\sqrt{m^2c^2 + \frac{(p_s - eA_s)^2}{(1 + x/\rho)^2} + (p_s - eA_s)^2 + (p_y - eA_y)^2}$ 

• As before, we change the independent variable from t to s

The new conjugate phase space variables are  $(x, p_x, y, p_y, t, -H)$ And the new Hamiltonian (s-dependent) is  $\tilde{H} = -p_s$   $\tilde{H} = -(1 + x/\rho) \left[ \frac{(H - e\phi)^2}{c^2} - m^2 c^2 - (p_x - eA_x)^2 - (p_y - eA_y)^2 \right]^{1/2} - eA_s$ Which is time-independent (if also  $\phi, A$  are time-independent) Expanding the Hamiltonian to second order in  $p_x, p_y$   $\tilde{H} \approx -p(1 + x/\rho) + \frac{1 + x/\rho}{2p} \left[ (p_x - eA_x)^2 + (p_y - eA_y)^2 \right]^{1/2} - eA_s$   $H - e\phi = E$  is the total particle energy  $p = \sqrt{E^2/c^2 - m^2c^2}$  is the total particle momentum

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### Magnetic Fields

• Maxwell's equations, time independent, no sources, so:  $\vec{J} = 0$ 

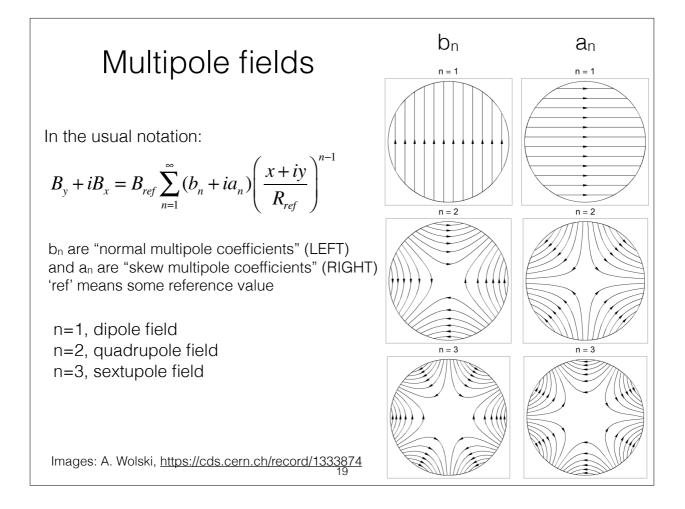
 $\nabla \times \vec{B} = 0 \qquad \qquad \vec{B} = \mu_0 \vec{H}$  $\nabla \cdot \vec{B} = 0$ 

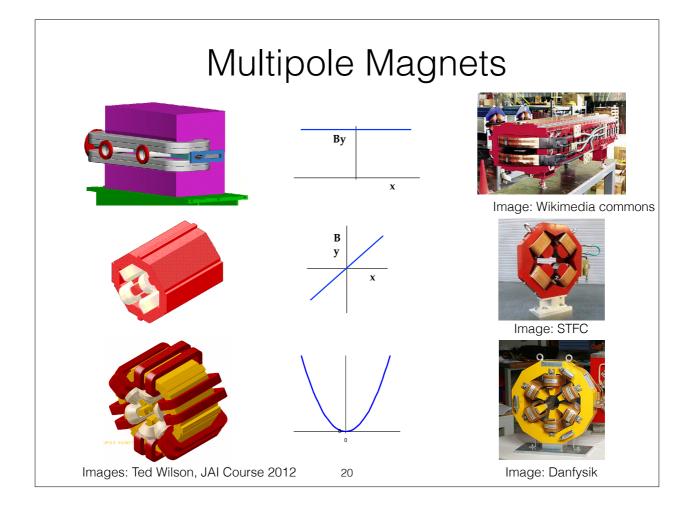
- We'll "guess" that the following obeys these equations:
- A constant vertical field B<sub>z</sub>, and

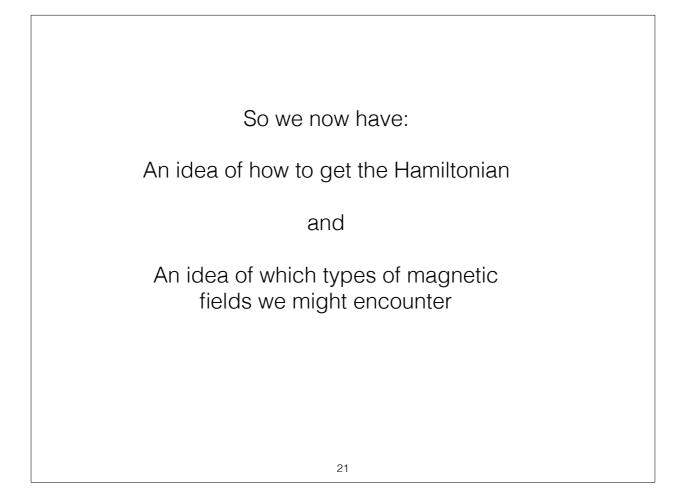
$$B_{v} + iB_{x} = C_{n}(x + iy)^{n-1}$$

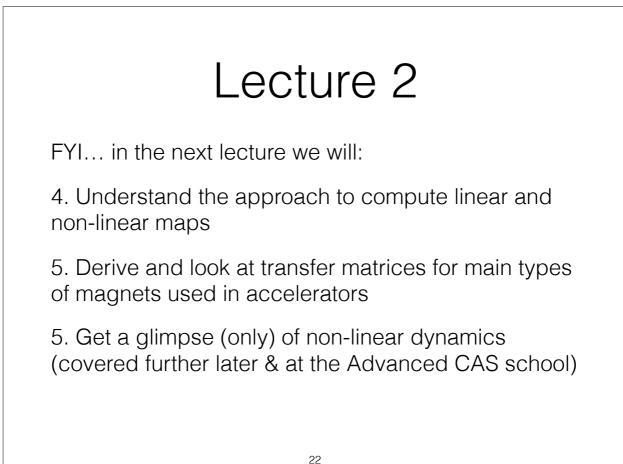
- n is an integer > 0, C is a complex number
- (real part understood)

Does this obey Maxwell in free space?  
Now apply 
$$\frac{\partial}{\partial x} + i \frac{\partial}{\partial y}$$
 to each side of  $B_y + iB_x = C_n (x + iy)^{n-1}$   
LHS:  $= \frac{\partial B_y}{\partial x} - \frac{\partial B_x}{\partial y} + i \left( \frac{\partial B_x}{\partial x} + \frac{\partial B_y}{\partial y} \right)$   
 $= \left[ \nabla \times \vec{B} \right]_z + i \nabla \cdot \vec{B}$  Where we know  $B_z$  is constant.  
RHS:  $= (n-1)(x+iy)^{n-2} + i^2(n-1)(x+iy)^{n-2} = 0$   
 $\therefore \quad \nabla \times \vec{B} = 0$  and  $\nabla \cdot \vec{B} = 0$   
So we find that as expected, the field  $B_y + iB_x = C_n (x+iy)^{n-1}$   
satisfies Maxwell's equations in free space









### References

Beam Dynamics:

- A. Wolski, "Beam Dynamics in High Energy Particle Accelerators", Imperial College Press, 2014.
- S. Y. Lee, "Accelerator Physics", 3rd Edition, World Scientific, 2011.
- K. Brown, SLAC-75-rev-4 (1982); SLAC-91-rev-2 (1977)

Electromagnetism:

• J. D. Jackson, Classical Electrodynamics, 3rd Ed, Wiley & sons (1999).

Hamiltonian Mechanics:

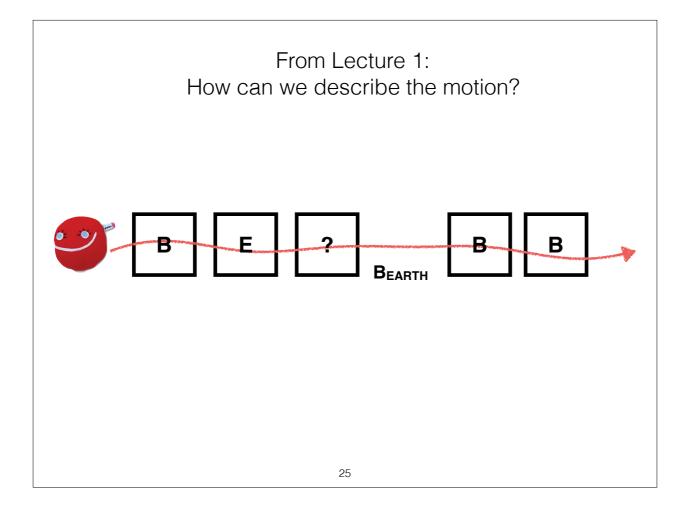
• H. Goldstein et al., Classical Mechanics (3rd ed.). Addison-wesley (2001).

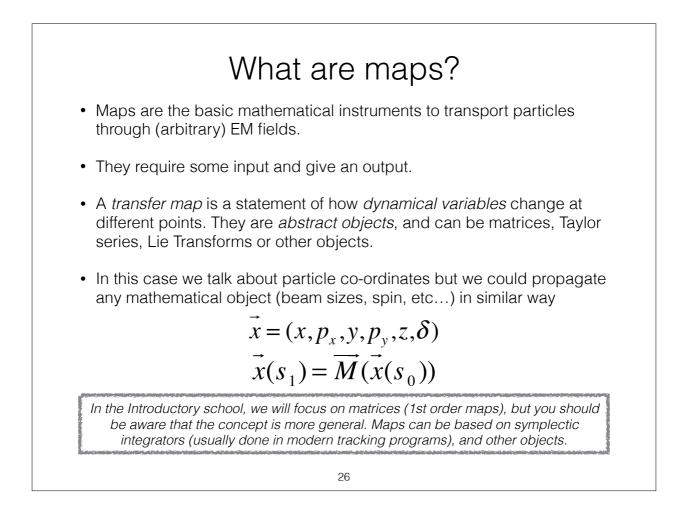
Most of the images/animations I've used here: <u>http://richannel.org/collections/2016/particle-accelerators-for-humanity</u>

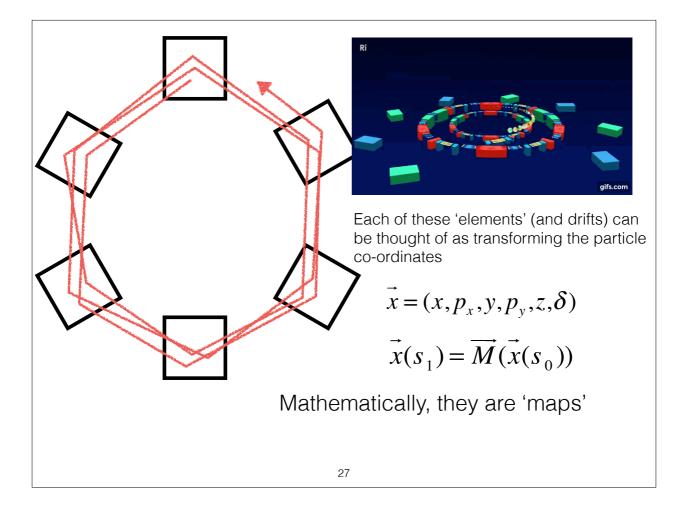
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## Particle Motion in EM Fields Lecture 2

Dr. Suzie Sheehy John Adams Institute for Accelerator Science University of Oxford







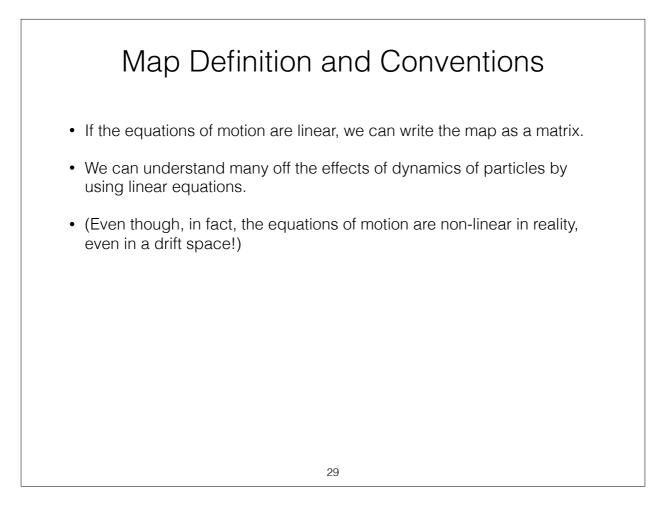
# Maps for Circular Machines

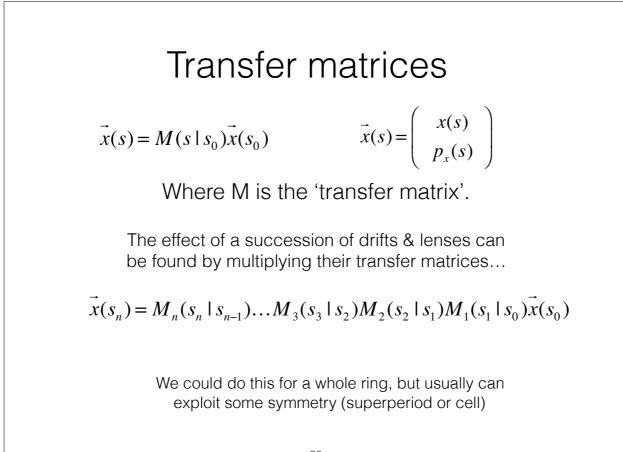
Accelerators are made up of beamline elements, each with their own linear and nonlinear fields, they might be mis-alignmed, mis-powered etc...

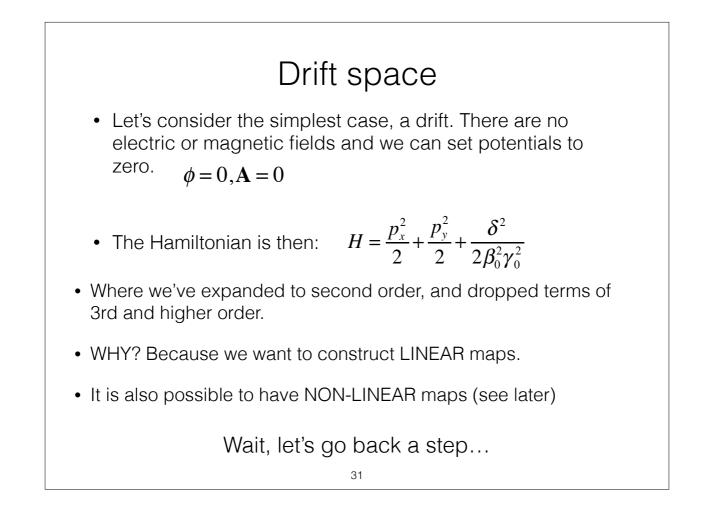
If one tried to write down the entire Hamiltonian for that system, it would be pretty darn complicated!

So instead, we take a piecewise approach:

- 1. First compute the maps for individual beamline elements using a local coordinate system that is appropriate to the element.
- 2. Then the maps are combined to produce a one-turn map, and then we can do our analysis on that map.







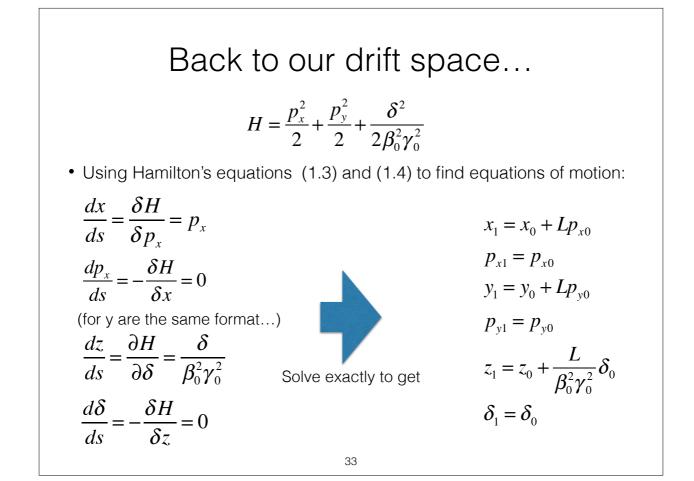
The actual Hamiltonian in a drift (no potentials) is:

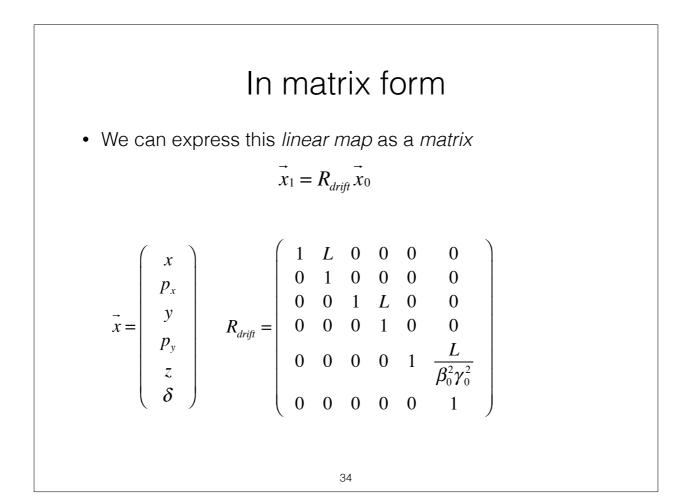
$$H = \frac{\delta}{\beta_0} - \sqrt{\left(\delta + \frac{1}{\beta_0}\right)^2 - p_x^2 - p_y^2 - \frac{1}{\beta_0^2 \gamma_0^2}}$$

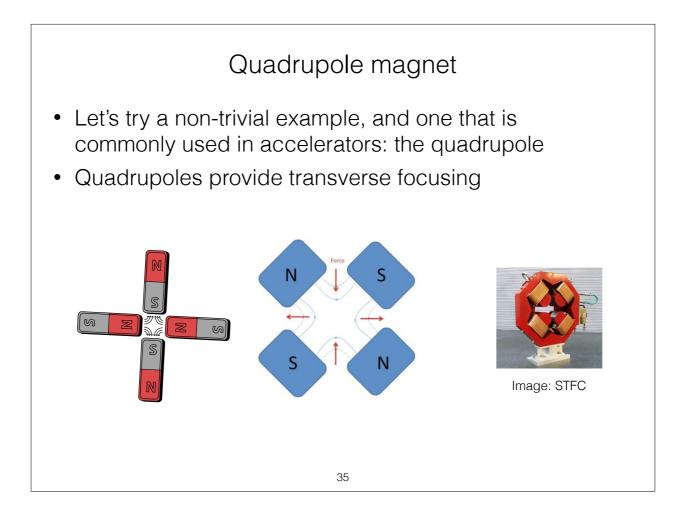
Expanding to second order gives:

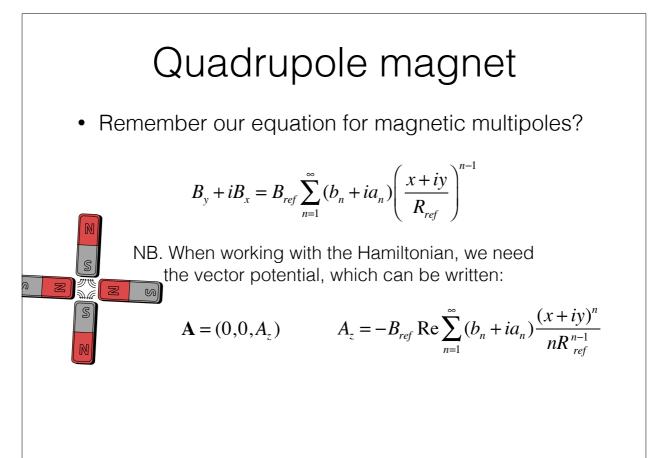
$$H = 1 + \frac{p_x^2}{2} + \frac{p_y^2}{2} + \frac{\delta^2}{2\beta_0^2\gamma_0^2} + Q(3)$$

- And we can drop the 3rd and higher order terms (and zeroth order as this doesn't contribute to the dynamics)
- This is called the *paraxial approximation* and you will hear people talking about it. It's worth knowing that we do this, and (for instance) whether the simulation code you use takes this approximation or not.
- We COULD solve the equations of motion and then approximate, but it is useful to start with a hamiltonian (even an approximate one) as it has conserved quantities









This becomes... for n=2 (quadrupole):

$$A_{z} = -\frac{B_{ref}b_{2}}{2R_{ref}}(x^{2} - y^{2})$$

NB. For a pure multipole field, we normally use the multipole fields normalised by  $q/P_0$ 

So, we can define the normalised multipole strength, k:

$$k_{n-1} = \frac{q}{P_0} \frac{\partial^{n-1} B_y}{\partial x^{n-1}} = (n-1)! \frac{B_{ref}}{R_{ref}^n} b_n$$

So for our quadrupole, n=2, we have:  $k_1 = \frac{q}{P_0} \frac{\partial B_y}{\partial x} = \frac{B_{ref}}{R_{ref}} b_2$ And we get the magnetic field and vector potentials:

**b** = 
$$(k_1 y, k_1 x, 0)$$
 **a** =  $(0, 0, -\frac{k_1}{2}(x^2 - y^2))$ 

<sup>37</sup> this is the bit we need next...

• Now we go back down the Hamiltonian rabbit hole...

$$\mathbf{a} = (0, 0, -\frac{k_1}{2}(x^2 - y^2))$$

Oh wait, it's not so bad... it's just like the drift, but with the new potential!

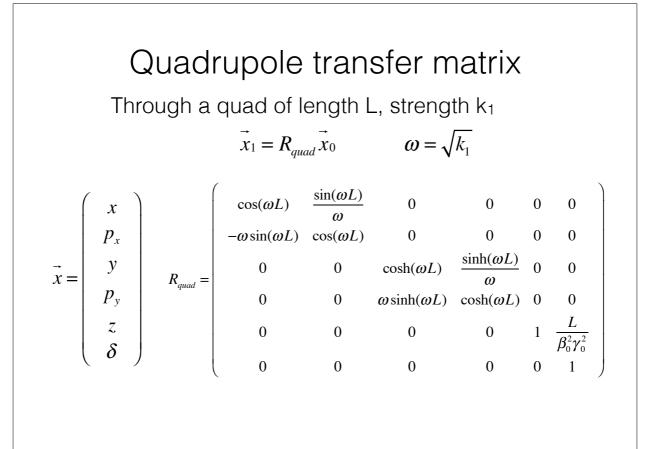
$$H_{quad} = \frac{p_x^2}{2} + \frac{p_y^2}{2} + \frac{\delta^2}{2\beta_0^2\gamma_0^2} + \frac{k_1}{2}(x^2 - y^2)$$

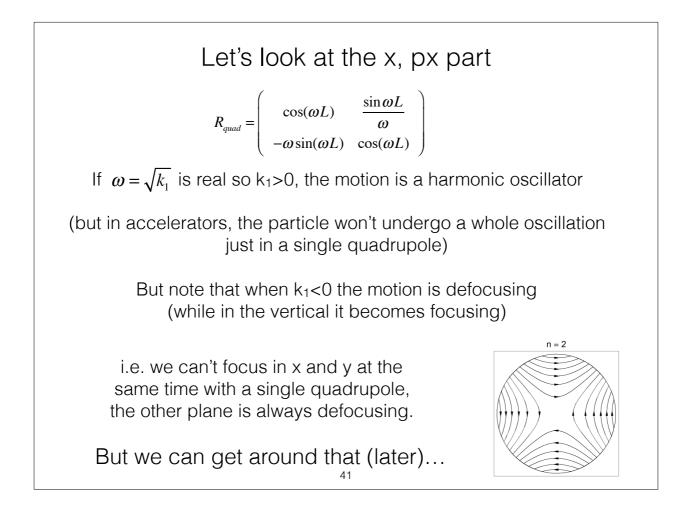
nb. we've made the paraxial approximation again

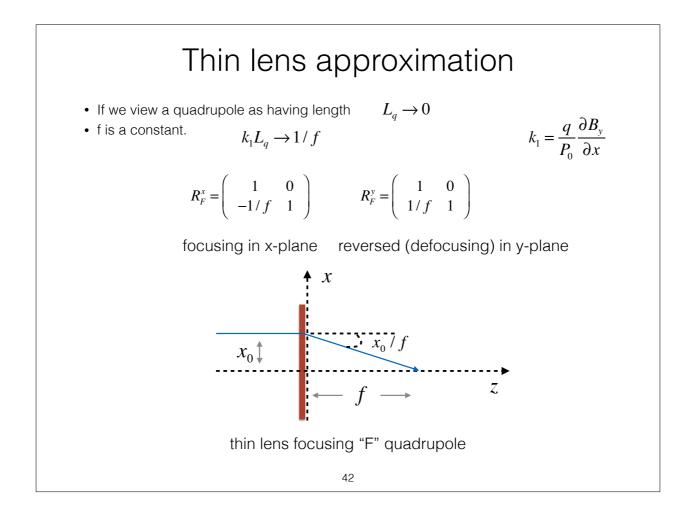
• Now we can solve the equations of motion again...

$$H_{quad} = \frac{p_x^2}{2} + \frac{p_y^2}{2} + \frac{\delta^2}{2\beta_0^2\gamma_0^2} + \frac{k_1}{2}(x^2 - y^2)$$

$$\frac{dx}{ds} = \frac{\delta H}{\delta p_x} = p_x \qquad \frac{dy}{ds} = \frac{\delta H}{\delta p_y} = p_y \qquad \frac{dz}{ds} = \frac{\partial H}{\partial \delta} = \frac{\delta H}{\beta_0^2 \gamma_0^2}$$
$$\frac{dp_x}{ds} = -\frac{\delta H}{\delta x} = k_1 x \qquad \frac{dp_y}{ds} = -\frac{\delta H}{\delta y} = -k_1 y \qquad \frac{d\delta}{ds} = -\frac{\delta H}{\delta z} = 0$$
etc...

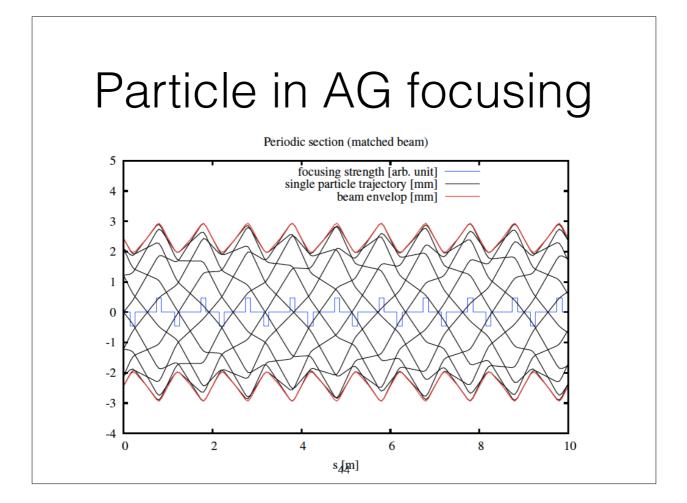


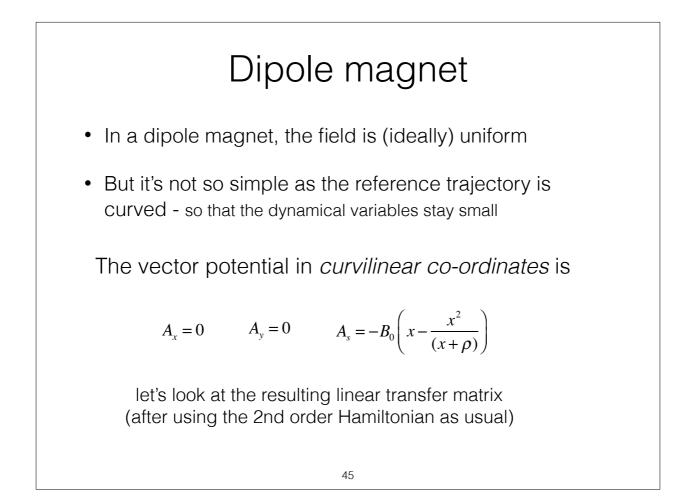


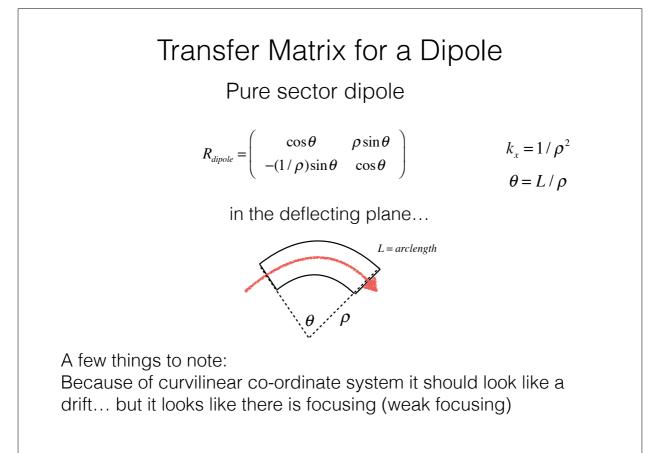


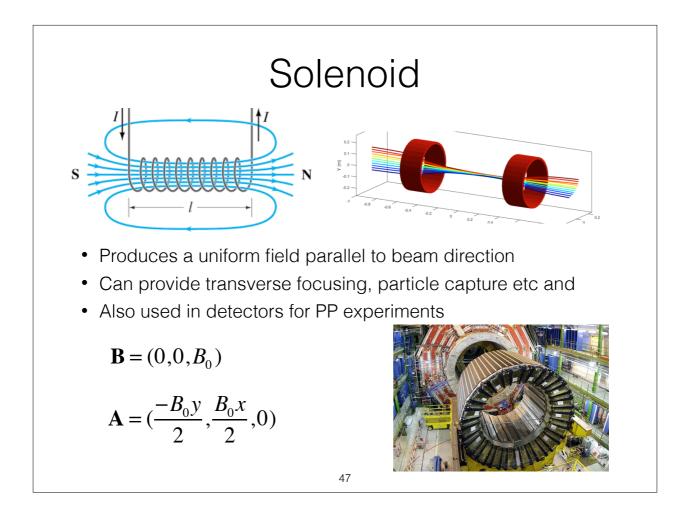
### Sneak peak: AG focusing - thin lens

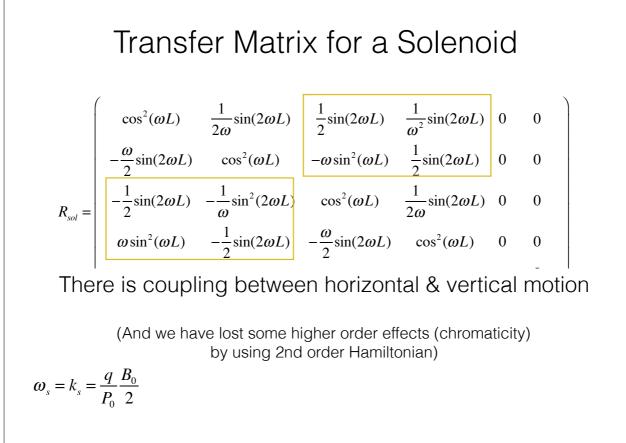
For infinitesimally short lenses, we can recover most of the physics  $K(s) = \pm \delta(s)/f \quad \text{where f is the focal length.}$ In the 'thin lens' approximation:  $M = \begin{pmatrix} 1 & d \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 \\ 1/f & 1 \end{pmatrix} \begin{pmatrix} 1 & d \\ 0 & 1 \end{pmatrix} \begin{pmatrix} 1 & 0 \\ -1/f & 1 \end{pmatrix}$   $= \begin{pmatrix} 1 - \frac{d}{f} - \frac{d^2}{f^2} & 2d + \frac{d^2}{f} \\ -\frac{d}{f^2} & 1 + d/f \end{pmatrix}$ Focusing & defocusing with a drift between doesn't cancel out. This is what gives us 'alternating gradient' focusing

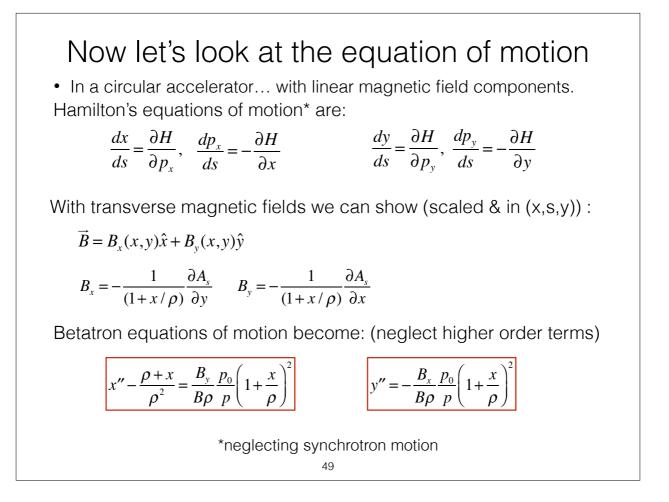


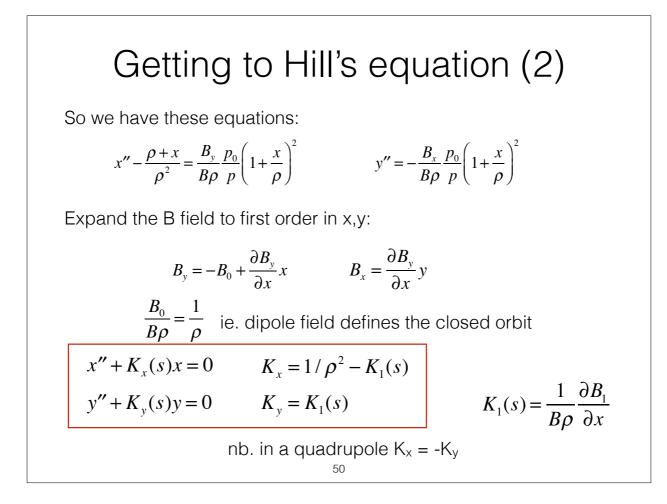


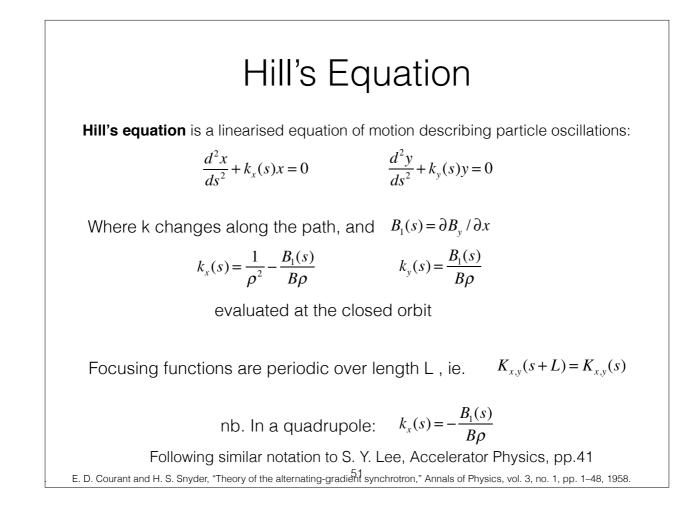












### Non-Linear Maps

• We will talk about this qualitatively only...

## Particle Motion in EM Fields

Hopefully, we have done the following:

1. Arrived at a general description of particle motion in EM fields

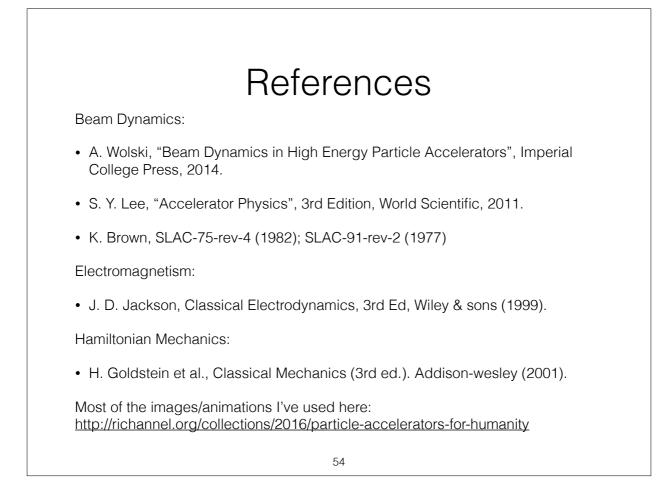
2. Understood what "maps" are, and how they relate to particle motion and simulation

3. Derived some basic maps from the equations of motion

4. Understood the approach to compute linear and non-linear maps

5. Derived and looked at transfer matrices for main types of magnets used in accelerators

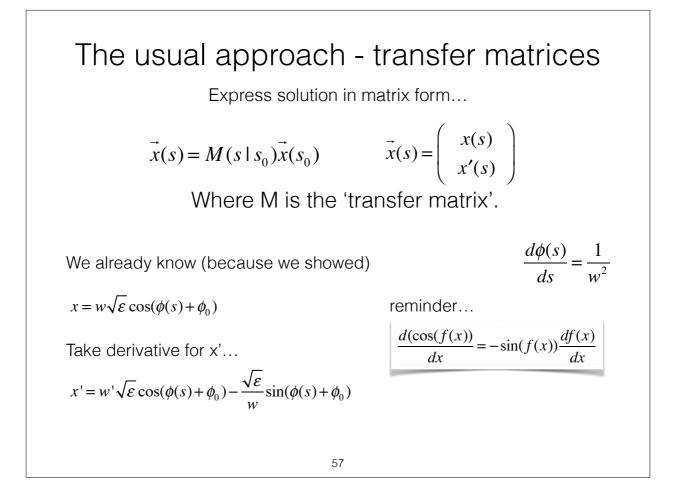
5. Got a glimpse of non-linear dynamics



# Additional Material

Let's check if the following solves Hill's equation... 
$$x'' + kx = 0$$
  

$$\boxed{x = \sqrt{\beta(s)\varepsilon} \cos(\phi(s) + \phi_0)}$$
Substitute  $w = \sqrt{\beta}$   $\phi = \phi(s) + \phi_0$   
& differentiate...  
 $x' = \sqrt{\varepsilon} \left\{ w'(s) \cos\phi - \frac{d\phi}{ds} w(s) \sin\phi \right\}$  nb. we need:  $\frac{d\phi}{ds} = \frac{1}{\beta(s)} = \frac{1}{w^2(s)}$   
Differentiate again...  
 $x'' = \sqrt{\varepsilon} \left\{ w''(s) \cos\phi - \frac{w'(s)}{w^2(s)} \sin\phi + \frac{w'(s)}{w^2(s)} \sin\phi} - \frac{1}{w^3} \cos\phi \right\}$   
Sub into Hill's...  
 $\sqrt{\varepsilon} \left\{ w''(s) \cos\phi - \frac{1}{w^3} \cos\phi \right\} + kw\sqrt{\varepsilon} \cos\phi = 0$  gives...  
 $\frac{w''(s) - \frac{1}{w^3} + kw = 0}{\frac{1}{2}\beta\beta'' - \frac{1}{4}\beta'^2 + k\beta^2 = 1}$ 



$$\begin{pmatrix} x(s_2) \\ x'(s_2) \end{pmatrix} = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \begin{pmatrix} x(s_1) \\ x'(s_1) \end{pmatrix}$$
  
Trace two rays...  
'cosine like'  $\phi = 0$  'sine like'  $\phi = \pi/2$   
 $x = w\sqrt{\varepsilon}\cos(\phi(s) + \phi_0)$   
 $x' = w'\sqrt{\varepsilon}\cos(\phi(s) + \phi_0) - \frac{\sqrt{\varepsilon}}{w}\sin(\phi(s) + \phi_0)$   
Yields 4 simultaneous equations so we can solve for a,b,c,d...  
 $\mu = \phi_2 - \phi_1$   
 $M_{12} = \begin{pmatrix} \frac{w_2}{w_1}\cos\mu - w_2w_1'\sin\mu & w_1w_2\sin\mu \\ -\frac{1+w_1w_1'w_2w_2'}{w_1w_2}\sin\mu - \left(\frac{w_1'}{w_2} - \frac{w_2'}{w_1}\right)\cos\mu & \frac{w_1}{w_2}\cos\mu + w_1w_2'\sin\mu \end{pmatrix}$   
<sub>58</sub>

### You will see this later...

Simplify by considering a period or 'turn', and w's are equal.

 $M_{period} = \begin{pmatrix} \cos \mu - ww' \sin \mu & w^2 \sin \mu \\ -\frac{1 + w^2 w'^2}{w^2} \sin \mu & \cos \mu + ww' \sin \mu \end{pmatrix}$ 

If we define the so-called 'Twiss' or 'Courant-Snyder' parameters:

$$\beta = w^{2} \qquad \alpha = -\frac{1}{2}\beta' \qquad \gamma = \frac{1+\alpha}{\beta}$$
$$M_{period} = \begin{pmatrix} \cos\mu + \alpha\sin\mu & \beta\sin\mu \\ -\gamma\sin\mu & \cos\mu - \alpha\sin\mu \end{pmatrix}$$

(sorry that we are reusing symbols again... these are NOT the relativistic parameters) 59