## Particle Motion in EM Fields Lecture 1



Image: Andrew Khosravani, 2016
http://richannel.org/collections/2016/particle-accelerators-for-humanity
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## Where are we now?

- So far, you should already have learned about:
- The basic types of accelerators
- Some useful hints at types of magnets and basic beam parameters
- Electromagnetic fields
- Maxwell's equations
- The Lorentz force
- We are getting a general idea that particles are bent and focused by magnets. Now we will explore the fundamental physics of how particles behave when they experience electromagnetic fields.


## Aims for this lecture

From the CAS Syllabus:

1. Arrive at a general description of particle motion in EM fields
2. Understand what "maps" are, and how they relate to particle motion and simulation
3. Derive some basic maps from the equations of motion

Particle in an EM field
How can we describe the motion?



## Motion of Charged Particles

- Let's start by looking at the motion of single particles.

- "...in principle, there are only two steps in the analysis of any dynamical system. The first step is to write down the equations of motion; and the second step is to solve them" [A. Wolski, pp. 59 ]
- Why not just use Newton's laws \& Lorentz force?

$$
\begin{equation*}
\frac{d \mathbf{p}}{d t}=\mathbf{F} \quad(1.1) \quad \mathbf{F}=q(\mathbf{E}+\mathbf{v} \times \mathbf{B}) \tag{1.2}
\end{equation*}
$$

- In an accelerator, magnets and rf cavities are generally defined along a trajectory (i.e. we know where they are in distance, not time). The Hamiltonian approach is lets us use this fact.
- The motion of particles in electromagnetic fields is conservative, and similar to a harmonic oscillator with perturbations. We have lots of mathematical tools to treat this in the Hamiltonian formalism.
- Ultimately, it makes our lives easier.



## Hamiltonian (straight beam line)

- The Hamiltonian represents the total energy of the particle
- We need a Hamiltonian that gives (1.1) and (1.2) when substituted into (1.3) and (1.4)
$H=H\left(x_{i}, p_{i} ; t\right)$
We propose the following Hamiltonian for a relativistic charged particle moving in an electromagnetic field:
$H=c \sqrt{(\mathbf{p}-q \mathbf{A})^{2}+m^{2} c^{2}}+q \phi$
But this is still defined in terms of time... so we want to change it.


## Hamiltonian (straight beam line)

$$
\begin{equation*}
H=c \sqrt{(\mathbf{p}-q \mathbf{A})^{2}+m^{2} c^{2}}+q \phi \tag{1.5}
\end{equation*}
$$

We don't need to go through every step here, but here's what we do.
STEP 1: We take (1.5) and change the independent variable to $z$, the distance along the beam line

$$
\left(x, p_{x}\right) \quad\left(y, p_{y}\right) \quad(t,-E)
$$

We get (try this at home...), use $\mathrm{H}=$ total energy E

$$
\begin{equation*}
H_{\text {new }}=-p_{z}=-\sqrt{\frac{(E-q \phi)^{2}}{c^{2}}-\left(p_{x}-q A_{x}\right)^{2}-\left(p_{y}-q A_{y}\right)^{2}-m^{2} c^{2}}-q A_{z} \tag{1.6}
\end{equation*}
$$

Note that if the $E$ and $B$ fields are static, the Hamiltonian is independent of time and the total energy of the particle is constant. Happy days.

## Straight Beamline Hamiltonian

## We're not going to go through all the steps...



Blue: "reference" trajectory
Red: actual particle trajectory
STEP 2: We choose new (canonical) variables for the position \& momentum that stay small as the particle moves along the beam line and scale by reference momentum Po (subscript '0' denotes reference)

STEP 3: We define new (canonical) longitudinal variables.

## Variables for Beam Dynamics

Eventually, our Hamiltonian with independent variable 's' along the beamline, becomes:

$$
\begin{equation*}
H=\frac{\delta}{\beta_{0}}-\sqrt{\left(\delta+\frac{1}{\beta_{0}}-\frac{q \phi}{c P_{0}}\right)^{2}-\left(p_{x}-a_{x}\right)^{2}-\left(p_{y}-a_{y}\right)^{2}-\frac{1}{\beta_{0}^{2} \gamma_{0}^{2}}}-a_{z} \tag{1.7}
\end{equation*}
$$

$$
\mathbf{a}=\frac{q}{P_{0}} \mathbf{A}
$$

$$
\delta=\frac{E}{c P_{0}}-\frac{1}{\beta_{0}}
$$

$$
\left(x, p_{x}, y, p_{y}, z, \delta\right)
$$

Scaled vector potential
Energy deviation
Co-ordinates \& momenta

- Using the Hamiltonian, we can get the equations of motion for a particle in a (straight) beam line.
- Usually these equations are too complex to solve exactly, so we have to make some approximations.


## Hamiltonian (curved beam line)



- Now we complicate things slightly with curved reference frame
- We want to measure $s$ along a path which curves with the trajectory of a particle on a curved orbit, which we call the reference trajectory
- Spoiler alert: simply turns out as a factor in front of the straight line Hamiltonian

In accelerator physics we ask: "What are the particles' generalized coordinates when they reach a certain point in space?"

First, we convert to a non-inertial reference frame. We use the 'Frenet-Serret' co-ordinate system


Particle motion is described with respect to a reference orbit in the noninertial frame (x, y, s). This co-ordinate system is known as Frenet-Serret

## - First, we convert to 'Frenet-Serret' co-ordinate system

$\hat{s}(s)=\frac{d \vec{r}_{0}(s)}{d s} \quad$ Tangent unit vector to closed orbit
$\hat{x}(s)=-\rho(s) \frac{d \hat{s}(s)}{d s} \quad$ Unit vector perpendicular to tangent vector
$\hat{y}(s)=\hat{x}(s) \times \hat{y}(s) \quad$ Third unit vector...
Particle trajectory: $\vec{r}(s)=\vec{r}_{0}(s)+x \hat{x}(s)+y \hat{y}(s)$
nb. the reference frame moves WITH the particle

$$
H=e \phi+c \sqrt{m^{2} c^{2}+\frac{\left(p_{s}-e A_{s}\right)^{2}}{(1+x / \rho)^{2}}+\left(p_{x}-e A_{x}\right)^{2}+\left(p_{y}-e A_{y}\right)^{2}}
$$

- As before, we change the independent variable from to s

$$
\text { The new conjugate phase space variables are } \quad\left(x, p_{x}, y, p_{y}, t,-H\right)
$$

And the new Hamiltonian (s-dependent) is $\tilde{H}=-p_{s}$

$$
\tilde{H}=-(1+x / \rho)\left[\frac{(H-e \phi)^{2}}{c^{2}}-m^{2} c^{2}-\left(p_{x}-e A_{x}\right)^{2}-\left(p_{y}-e A_{y}\right)^{2}\right]^{1 / 2}-e A_{s}
$$

Which is time-independent (if also $\phi, A$ are time-independent)

Expanding the Hamiltonian to second order in $\mathrm{p}_{\mathrm{x}}, \mathrm{p}_{\mathrm{y}}$

$$
\tilde{H} \approx-p(1+x / \rho)+\frac{1+x / \rho}{2 p}\left[\left(p_{x}-e A_{x}\right)^{2}+\left(p_{y}-e A_{y}\right)^{2}\right]^{1 / 2}-e A_{s}
$$

$H-e \phi=E \quad$ is the total particle energy
$p=\sqrt{E^{2} / c^{2}-m^{2} c^{2}}$ is the total particle momentum

## Where are we now?

So far, we have been looking in general for any Vector potential A Scalar potential $\phi$

But in reality, we (usually) use electric fields to accelerate particles and magnetic fields to bend, focus and manipulate the beams.

So we need to know
"which magnetic fields can we really create?"
(We'll come back to the Hamiltonian later...)

## Magnetic Fields

- Maxwell's equations, time independent, no sources, so: $\vec{J}=0$

$$
\begin{array}{ll}
\nabla \times \vec{B}=0 & \vec{B}=\mu_{0} \vec{H} \\
\nabla \cdot \vec{B}=0 &
\end{array}
$$

- We'll "guess" that the following obeys these equations:
- A constant vertical field $B_{z}$, and

$$
B_{y}+i B_{x}=C_{n}(x+i y)^{n-1}
$$

- n is an integer $>0, \mathrm{C}$ is a complex number
- (real part understood)


## Does this obey Maxwell in free space?

Now apply $\quad \frac{\partial}{\partial x}+i \frac{\partial}{\partial y} \quad$ to each side of $\quad B_{y}+i B_{x}=C_{n}(x+i y)^{n-1}$
LHS: $=\frac{\partial B_{y}}{\partial x}-\frac{\partial B_{x}}{\partial y}+i\left(\frac{\partial B_{x}}{\partial x}+\frac{\partial B_{y}}{\partial y}\right)$
$=[\nabla \times \vec{B}]_{z}+i \nabla \cdot \vec{B} \quad$ Where we know $\mathrm{B}_{z}$ is constant.

RHS: $\quad=(n-1)(x+i y)^{n-2}+i^{2}(n-1)(x+i y)^{n-2}=0$

$$
\therefore \quad \nabla \times \vec{B}=0 \text { and } \nabla \cdot \vec{B}=0
$$

So we find that as expected, the field $\quad B_{y}+i B_{x}=C_{n}(x+i y)^{n-1}$ satisfies Maxwell's equations in free space

## Multipole fields

In the usual notation:
$B_{y}+i B_{x}=B_{r e f} \sum_{n=1}^{\infty}\left(b_{n}+i a_{n}\right)\left(\frac{x+i y}{R_{r e f}}\right)^{n-1}$
$\mathrm{b}_{n}$ are "normal multipole coefficients" (LEFT) and $a_{n}$ are "skew multipole coefficients" (RIGHT) 'ref' means some reference value
$\mathrm{n}=1$, dipole field
$n=2$, quadrupole field
$n=3$, sextupole field

Images: A. Wolski, https://cds.cern.ch/record/1333874


## Multipole Magnets




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Image: Wikimedia commons


Image: STFC


Image: Danfysik

So we now have:
An idea of how to get the Hamiltonian and

An idea of which types of magnetic fields we might encounter

## Lecture 2

FYI... in the next lecture we will:
4. Understand the approach to compute linear and non-linear maps
5. Derive and look at transfer matrices for main types of magnets used in accelerators
5. Get a glimpse (only) of non-linear dynamics (covered further later \& at the Advanced CAS school)

## References

Beam Dynamics:

- A. Wolski, "Beam Dynamics in High Energy Particle Accelerators", Imperial College Press, 2014.
- S. Y. Lee, "Accelerator Physics", 3rd Edition, World Scientific, 2011.
- K. Brown, SLAC-75-rev-4 (1982); SLAC-91-rev-2 (1977)

Electromagnetism:

- J. D. Jackson, Classical Electrodynamics, 3rd Ed, Wiley \& sons (1999).

Hamiltonian Mechanics:

- H. Goldstein et al., Classical Mechanics (3rd ed.). Addison-wesley (2001).

Most of the images/animations l've used here:
http://richannel.org/collections/2016/particle-accelerators-for-humanity

## Particle Motion in EM Fields Lecture 2

## From Lecture 1: How can we describe the motion?



## What are maps?

- Maps are the basic mathematical instruments to transport particles through (arbitrary) EM fields.
- They require some input and give an output.
- A transfer map is a statement of how dynamical variables change at different points. They are abstract objects, and can be matrices, Taylor series, Lie Transforms or other objects.
- In this case we talk about particle co-ordinates but we could propagate any mathematical object (beam sizes, spin, etc...) in similar way

$$
\begin{aligned}
& \vec{x}=\left(x, p_{x}, y, p_{y}, z, \delta\right) \\
& \vec{x}\left(s_{1}\right)=\vec{M}\left(\vec{x}\left(s_{0}\right)\right)
\end{aligned}
$$

[^0] be aware that the concept is more general. Maps can be based on symplectic integrators (usually done in modern tracking programs), and other objects.


## Maps for Circular Machines

Accelerators are made up of beamline elements, each with their own linear and nonlinear fields, they might be mis-alignmed, mis-powered etc...

If one tried to write down the entire Hamiltonian for that system, it would be pretty darn complicated!

So instead, we take a piecewise approach:

1. First compute the maps for individual beamline elements using a local coordinate system that is appropriate to the element.
2. Then the maps are combined to produce a one-turn map, and then we can do our analysis on that map.

## Map Definition and Conventions

- If the equations of motion are linear, we can write the map as a matrix.
- We can understand many off the effects of dynamics of particles by using linear equations.
- (Even though, in fact, the equations of motion are non-linear in reality, even in a drift space!)


## Transfer matrices <br> $$
\vec{x}(s)=M\left(s \mid s_{0}\right) \vec{x}\left(s_{0}\right) \quad \vec{x}(s)=\binom{x(s)}{p_{x}(s)}
$$

Where $M$ is the 'transfer matrix'.

The effect of a succession of drifts \& lenses can be found by multiplying their transfer matrices...

$$
\vec{x}\left(s_{n}\right)=M_{n}\left(s_{n} \mid s_{n-1}\right) \ldots M_{3}\left(s_{3} \mid s_{2}\right) M_{2}\left(s_{2} \mid s_{1}\right) M_{1}\left(s_{1} \mid s_{0}\right) \vec{x}\left(s_{0}\right)
$$

We could do this for a whole ring, but usually can exploit some symmetry (superperiod or cell)

## Drift space

- Let's consider the simplest case, a drift. There are no electric or magnetic fields and we can set potentials to zero. $\quad \phi=0, \mathbf{A}=0$
- The Hamiltonian is then: $H=\frac{p_{x}^{2}}{2}+\frac{p_{y}^{2}}{2}+\frac{\delta^{2}}{2 \beta_{0}^{2} \gamma_{0}^{2}}$
- Where we've expanded to second order, and dropped terms of 3rd and higher order.
- WHY? Because we want to construct LINEAR maps.
- It is also possible to have NON-LINEAR maps (see later)


## Wait, let's go back a step...

The actual Hamiltonian in a drift (no potentials) is:

$$
H=\frac{\delta}{\beta_{0}}-\sqrt{\left(\delta+\frac{1}{\beta_{0}}\right)^{2}-p_{x}^{2}-p_{y}^{2}-\frac{1}{\beta_{0}^{2} \gamma_{0}^{2}}}
$$

## Expanding to second order gives:

$$
H=\neq 1+\frac{p_{x}^{2}}{2}+\frac{p_{y}^{2}}{2}+\frac{\delta^{2}}{2 \beta_{0}^{2} \gamma_{0}^{2}}+O(3)
$$

- And we can drop the 3rd and higher order terms (and zeroth order as this doesn't contribute to the dynamics)
- This is called the paraxial approximation and you will hear people talking about it. It's worth knowing that we do this, and (for instance) whether the simulation code you use takes this approximation or not.
- We COULD solve the equations of motion and then approximate, but it is useful to start with a hamiltonian (even an approximate one) as it has conserved quantities


## Back to our drift space...

$$
H=\frac{p_{x}^{2}}{2}+\frac{p_{y}^{2}}{2}+\frac{\delta^{2}}{2 \beta_{0}^{2} \gamma_{0}^{2}}
$$

- Using Hamilton's equations (1.3) and (1.4) to find equations of motion:

$$
\begin{aligned}
& \frac{d x}{d s}=\frac{\delta H}{\delta p_{x}}=p_{x} \\
& \frac{d p_{x}}{d s}=-\frac{\delta H}{\delta x}=0
\end{aligned}
$$

(for y are the same format...)

$$
\begin{aligned}
& \frac{d z}{d s}=\frac{\partial H}{\partial \delta}=\frac{\delta}{\beta_{0}^{2} \gamma_{0}^{2}} \\
& \frac{d \delta}{d s}=-\frac{\delta H}{\delta z}=0
\end{aligned}
$$

Solve exactly to get

$$
\begin{aligned}
& x_{1}=x_{0}+L p_{x 0} \\
& p_{x 1}=p_{x 0} \\
& y_{1}=y_{0}+L p_{y 0} \\
& p_{y 1}=p_{y 0}
\end{aligned}
$$

$$
z_{1}=z_{0}+\frac{L}{\beta_{0}^{2} \gamma_{0}^{2}} \delta_{0}
$$

$$
\delta_{1}=\delta_{0}
$$

## In matrix form

- We can express this linear map as a matrix

$$
\begin{gathered}
\vec{x}_{1}=R_{d r i j t} \vec{x}_{0} \\
\vec{x}=\left(\begin{array}{c}
x \\
p_{x} \\
y \\
p_{y} \\
z \\
\delta
\end{array}\right) \quad R_{\text {drift }}=\left(\begin{array}{cccccc}
1 & L & 0 & 0 & 0 & 0 \\
0 & 1 & 0 & 0 & 0 & 0 \\
0 & 0 & 1 & L & 0 & 0 \\
0 & 0 & 0 & 1 & 0 & 0 \\
0 & 0 & 0 & 0 & 1 & \frac{L}{\beta_{0}^{2} \gamma_{0}^{2}} \\
0 & 0 & 0 & 0 & 0 & 1
\end{array}\right)
\end{gathered}
$$

## Quadrupole magnet

- Let's try a non-trivial example, and one that is commonly used in accelerators: the quadrupole
- Quadrupoles provide transverse focusing


Image: STFC

## Quadrupole magnet

- Remember our equation for magnetic multipoles?

$$
B_{y}+i B_{x}=B_{r e f} \sum_{n=1}^{\infty}\left(b_{n}+i a_{n}\right)\left(\frac{x+i y}{R_{r e f}}\right)^{n-1}
$$

NB. When working with the Hamiltonian, we need the vector potential, which can be written:

$$
\mathbf{A}=\left(0,0, A_{z}\right)
$$

$$
A_{z}=-B_{r e f} \operatorname{Re} \sum_{n=1}^{\infty}\left(b_{n}+i a_{n}\right) \frac{(x+i y)^{n}}{n R_{r e f}^{n-1}}
$$

This becomes... for $\mathrm{n}=2$ (quadrupole):

$$
A_{z}=-\frac{B_{r e f} b_{2}}{2 R_{r e f}}\left(x^{2}-y^{2}\right)
$$

NB. For a pure multipole field, we normally use the multipole fields normalised by $q / P_{0}$

So, we can define the normalised multipole strength, $k$ :

$$
k_{n-1}=\frac{q}{P_{0}} \frac{\partial^{n-1} B_{y}}{\partial x^{n-1}}=(n-1)!\frac{B_{r e f}}{R_{r e f}^{n}} b_{n}
$$

So for our quadrupole, $\mathrm{n}=2$, we have: $\quad k_{1}=\frac{q}{P_{0}} \frac{\partial B_{y}}{\partial x}=\frac{B_{r e f}}{R_{\text {ref }}} b_{2}$
And we get the magnetic field and vector potentials:

$$
\mathbf{b}=\left(k_{1} y, k_{1} x, 0\right) \quad \mathbf{a}=\left(0,0,-\frac{k_{1}}{2}\left(x^{2}-y^{2}\right)\right)
$$

37 this is the bit we need next...

- Now we go back down the Hamiltonian rabbit hole...

$$
\mathbf{a}=\left(0,0,-\frac{k_{1}}{2}\left(x^{2}-y^{2}\right)\right)
$$

Oh wait, it's not so bad... it's just like the drift, but with the new potential!

$$
H_{\text {quad }}=\frac{p_{x}^{2}}{2}+\frac{p_{y}^{2}}{2}+\frac{\delta^{2}}{2 \beta_{0}^{2} \gamma_{0}^{2}}+\frac{k_{1}}{2}\left(x^{2}-y^{2}\right)
$$

- Now we can solve the equations of motion again...

$$
\begin{gathered}
H_{\text {quad }}=\frac{p_{x}^{2}}{2}+\frac{p_{y}^{2}}{2}+\frac{\delta^{2}}{2 \beta_{0}^{2} \gamma_{0}^{2}}+\frac{k_{1}}{2}\left(x^{2}-y^{2}\right) \\
\frac{d x}{d s}=\frac{\delta H}{\delta p_{x}}=p_{x}
\end{gathered} \frac{\frac{d y}{d s}=\frac{\delta H}{\delta p_{y}}=p_{y} \quad \frac{d z}{d s}=\frac{\partial H}{\partial \delta}=\frac{\delta}{\beta_{0}^{2} \gamma_{0}^{2}}}{\frac{d p_{x}}{d s}=-\frac{\delta H}{\delta x}=k_{1} x} \begin{array}{ll}
\frac{d p_{y}}{d s}=-\frac{\delta H}{\delta y}=-k_{1} y & \frac{d \delta}{d s}=-\frac{\delta H}{\delta z}=0 \\
\text { etc } \ldots
\end{array}
$$

## Quadrupole transfer matrix

Through a quad of length $L$, strength $\mathrm{k}_{1}$

$$
\vec{x}_{1}=R_{\text {quad }} \vec{x}_{0} \quad \omega=\sqrt{k_{1}}
$$

$$
\vec{x}=\left(\begin{array}{c}
x \\
p_{x} \\
y \\
p_{y} \\
z \\
\delta
\end{array}\right) \quad R_{\text {quad }}=\left(\begin{array}{cccccc}
\cos (\omega L) & \frac{\sin (\omega L)}{\omega} & 0 & 0 & 0 & 0 \\
-\omega \sin (\omega L) & \cos (\omega L) & 0 & 0 & 0 & 0 \\
0 & 0 & \cosh (\omega L) & \frac{\sinh (\omega L)}{\omega} & 0 & 0 \\
0 & 0 & \omega \sinh (\omega L) & \cosh (\omega L) & 0 & 0 \\
0 & 0 & 0 & 0 & 1 & \frac{L}{\beta_{0}^{2} \gamma_{0}^{2}} \\
0 & 0 & 0 & 0 & 0 & 1
\end{array}\right)
$$

## Let's look at the $\mathrm{x}, \mathrm{px}$ part

$$
R_{\text {quad }}=\left(\begin{array}{cc}
\cos (\omega L) & \frac{\sin \omega L}{\omega} \\
-\omega \sin (\omega L) & \cos (\omega L)
\end{array}\right)
$$

If $\omega=\sqrt{k_{1}}$ is real so $\mathrm{k}_{1}>0$, the motion is a harmonic oscillator
(but in accelerators, the particle won't undergo a whole oscillation just in a single quadrupole)

But note that when $\mathrm{k}_{1}<0$ the motion is defocusing (while in the vertical it becomes focusing)
i.e. we can't focus in $x$ and $y$ at the same time with a single quadrupole, the other plane is always defocusing.

## But we can get around that (later)...



## Thin lens approximation

- If we view a quadrupole as having length $L_{q} \rightarrow 0$
- $f$ is a constant.

$$
\begin{array}{ll}
k_{1} L_{q} \rightarrow 1 / f & k_{1}=\frac{q}{P_{0}} \frac{\partial B_{y}}{\partial x} \\
R_{F}^{x}=\left(\begin{array}{cc}
1 & 0 \\
-1 / f & 1
\end{array}\right) & R_{F}^{y}=\left(\begin{array}{cc}
1 & 0 \\
1 / f & 1
\end{array}\right)
\end{array}
$$

focusing in x-plane reversed (defocusing) in y-plane

thin lens focusing "F" quadrupole

## Sneak peak: AG focusing - thin lens

For infinitesimally short lenses, we can recover most of the physics

$$
K(s)= \pm \delta(s) / f \quad \text { where } f \text { is the focal length. }
$$

In the 'thin lens' approximation:

$$
\begin{aligned}
M=\left(\begin{array}{ll}
1 & d \\
0 & 1
\end{array}\right) & \left(\begin{array}{cc}
1 & 0 \\
1 / f & 1
\end{array}\right)\left(\begin{array}{ll}
1 & d \\
0 & 1
\end{array}\right)\left(\begin{array}{cc}
1 & 0 \\
-1 / f & 1
\end{array}\right) \\
& =\left(\begin{array}{cc}
1-\frac{d}{f}-\frac{d^{2}}{f^{2}} & 2 d+\frac{d^{2}}{f} \\
-\frac{d}{f^{2}} & 1+d / f
\end{array}\right)
\end{aligned}
$$

Focusing \& defocusing with a drift between doesn't cancel out. This is what gives us 'alternating gradient' focusing

## Particle in AG focusing



## Dipole magnet

- In a dipole magnet, the field is (ideally) uniform
- But it's not so simple as the reference trajectory is curved - so that the dynamical variables stay small

The vector potential in curvilinear co-ordinates is

$$
A_{x}=0 \quad A_{y}=0 \quad A_{s}=-B_{0}\left(x-\frac{x^{2}}{(x+\rho)}\right)
$$

let's look at the resulting linear transfer matrix (after using the 2nd order Hamiltonian as usual)

## Transfer Matrix for a Dipole <br> Pure sector dipole

$$
R_{\text {dipole }}=\left(\begin{array}{cc}
\cos \theta & \rho \sin \theta \\
-(1 / \rho) \sin \theta & \cos \theta
\end{array}\right) \quad \begin{gathered}
k_{x}=1 / \rho^{2} \\
\theta=L / \rho
\end{gathered}
$$

in the deflecting plane...


A few things to note:
Because of curvilinear co-ordinate system it should look like a drift... but it looks like there is focusing (weak focusing)

## Solenoid



- Produces a uniform field parallel to beam direction
- Can provide transverse focusing, particle capture etc and
- Also used in detectors for PP experiments

$$
\begin{aligned}
& \mathbf{B}=\left(0,0, B_{0}\right) \\
& \mathbf{A}=\left(\frac{-B_{0} y}{2}, \frac{B_{0} x}{2}, 0\right)
\end{aligned}
$$



## Transfer Matrix for a Solenoid

$R_{s o l}=\left(\begin{array}{cc|cc|cc}\cos ^{2}(\omega L) & \frac{1}{2 \omega} \sin (2 \omega L) & \frac{1}{2} \sin (2 \omega L) & \frac{1}{\omega^{2}} \sin (2 \omega L) & 0 & 0 \\ -\frac{\omega}{2} \sin (2 \omega L) & \cos ^{2}(\omega L) & -\omega \sin ^{2}(\omega L) & \frac{1}{2} \sin (2 \omega L) & 0 & 0 \\ -\frac{1}{2} \sin (2 \omega L) & -\frac{1}{\omega} \sin ^{2}(2 \omega L) & \cos ^{2}(\omega L) & \frac{1}{2 \omega} \sin (2 \omega L) & 0 & 0 \\ \omega \sin ^{2}(\omega L) & -\frac{1}{2} \sin (2 \omega L) & -\frac{\omega}{2} \sin (2 \omega L) & \cos ^{2}(\omega L) & 0 & 0\end{array}\right)$
There is coupling between horizontal \& vertical motion
(And we have lost some higher order effects (chromaticity) by using 2nd order Hamiltonian)
$\omega_{s}=k_{s}=\frac{q}{P_{0}} \frac{B_{0}}{2}$

## Now let's look at the equation of motion

- In a circular accelerator... with linear magnetic field components. Hamilton's equations of motion* are:

$$
\frac{d x}{d s}=\frac{\partial H}{\partial p_{x}}, \quad \frac{d p_{x}}{d s}=-\frac{\partial H}{\partial x}
$$

$$
\frac{d y}{d s}=\frac{\partial H}{\partial p_{y}}, \frac{d p_{y}}{d s}=-\frac{\partial H}{\partial y}
$$

With transverse magnetic fields we can show (scaled \& in (x,s,y)) :

$$
\begin{aligned}
& \vec{B}=B_{x}(x, y) \hat{x}+B_{y}(x, y) \hat{y} \\
& B_{x}=-\frac{1}{(1+x / \rho)} \frac{\partial A_{s}}{\partial y} \quad B_{y}=-\frac{1}{(1+x / \rho)} \frac{\partial A_{s}}{\partial x}
\end{aligned}
$$

Betatron equations of motion become: (neglect higher order terms)

$$
x^{\prime \prime}-\frac{\rho+x}{\rho^{2}}=\frac{B_{y}}{B \rho} \frac{p_{0}}{p}\left(1+\frac{x}{\rho}\right)^{2} \quad y^{\prime \prime}=-\frac{B_{x}}{B \rho} \frac{p_{0}}{p}\left(1+\frac{x}{\rho}\right)^{2}
$$

*neglecting synchrotron motion

## Getting to Hill's equation (2)

So we have these equations:

$$
x^{\prime \prime}-\frac{\rho+x}{\rho^{2}}=\frac{B_{y}}{B \rho} \frac{p_{0}}{p}\left(1+\frac{x}{\rho}\right)^{2} \quad y^{\prime \prime}=-\frac{B_{x}}{B \rho} \frac{p_{0}}{p}\left(1+\frac{x}{\rho}\right)^{2}
$$

Expand the B field to first order in $\mathrm{x}, \mathrm{y}$ :

$$
\begin{gathered}
B_{y}=-B_{0}+\frac{\partial B_{y}}{\partial x} x \quad B_{x}=\frac{\partial B_{y}}{\partial x} y \\
\frac{B_{0}}{B \rho}=\frac{1}{\rho} \quad \text { ie. dipole field defines the closed orbit } \\
x^{\prime \prime}+K_{x}(s) x=0 \\
y^{\prime \prime}+K_{y}(s) y=0
\end{gathered} \quad \begin{aligned}
& K_{x}=1 / \rho^{2}-K_{1}(s) \\
& K_{y}=K_{1}(s)
\end{aligned} \quad K_{1}(s)=\frac{1}{B \rho} \frac{\partial B_{1}}{\partial x}
$$

nb. in a quadrupole $K_{x}=-K_{y}$

## Hill's Equation

Hill's equation is a linearised equation of motion describing particle oscillations:

$$
\frac{d^{2} x}{d s^{2}}+k_{x}(s) x=0 \quad \frac{d^{2} y}{d s^{2}}+k_{y}(s) y=0
$$

Where k changes along the path, and $B_{1}(s)=\partial B_{y} / \partial x$

$$
\begin{aligned}
& k_{x}(s)=\frac{1}{\rho^{2}}-\frac{B_{1}(s)}{B \rho} \quad k_{y}(s)=\frac{B_{1}(s)}{B \rho} \\
& \text { evaluated at the closed orbit }
\end{aligned}
$$

Focusing functions are periodic over length L , ie. $\quad K_{x, y}(s+L)=K_{x, y}(s)$ nb. In a quadrupole: $\quad k_{x}(s)=-\frac{B_{1}(s)}{B \rho}$
Following similar notation to S. Y. Lee, Accelerator Physics, pp. 41

## Non-Linear Maps

- We will talk about this qualitatively only...


## Particle Motion in EM Fields

Hopefully, we have done the following:

1. Arrived at a general description of particle motion in EM fields
2. Understood what "maps" are, and how they relate to particle motion and simulation
3. Derived some basic maps from the equations of motion
4. Understood the approach to compute linear and non-linear maps
5. Derived and looked at transfer matrices for main types of magnets used in accelerators
6. Got a glimpse of non-linear dynamics

## References

Beam Dynamics:

- A. Wolski, "Beam Dynamics in High Energy Particle Accelerators", Imperial College Press, 2014.
- S. Y. Lee, "Accelerator Physics", 3rd Edition, World Scientific, 2011.
- K. Brown, SLAC-75-rev-4 (1982); SLAC-91-rev-2 (1977)

Electromagnetism:

- J. D. Jackson, Classical Electrodynamics, 3rd Ed, Wiley \& sons (1999).

Hamiltonian Mechanics:

- H. Goldstein et al., Classical Mechanics (3rd ed.). Addison-wesley (2001).

Most of the images/animations l've used here:
http://richannel.org/collections/2016/particle-accelerators-for-humanity

## Additional Material

Let's check if the following solves Hill's equation...

$$
x^{\prime \prime}+k x=0
$$

$x=\sqrt{\beta(s) \varepsilon} \cos \left(\phi(s)+\phi_{0}\right)$
Substitute $\quad w=\sqrt{\beta} \quad \phi=\phi(s)+\phi_{0}$
\& differentiate...
$x^{\prime}=\sqrt{\varepsilon}\left\{w^{\prime}(s) \cos \phi-\frac{d \phi}{d s} w(s) \sin \phi\right\} \quad$ nb. we need: $\frac{d \phi}{d s}=\frac{1}{\beta(s)}=\frac{1}{w^{2}(s)}$

Differentiate again...
$x^{\prime \prime}=\sqrt{\varepsilon}\{w^{\prime \prime}(s) \cos \phi \underbrace{\frac{w^{\prime}(s)}{w^{2}(s)} \sin \phi+\frac{w^{\prime}(s)}{w^{2}(s)} \sin \phi}_{=0}-\frac{1}{w^{3}} \cos \phi\}$

$$
\begin{array}{ll}
\text { Sub into Hill's... } \\
\sqrt{\sqrt{\varepsilon}\left\{w^{\prime \prime}(s) \cos \phi-\frac{1}{w^{3}} \cos \phi\right\}+k w \sqrt{\varepsilon} \cos \phi=0} \quad \text { gives... } & w^{\prime \prime}(s)-\frac{1}{w^{3}}+k w=0 \\
& \text { 'envelope equation' } \\
& \frac{1}{2} \beta 6 \beta^{\prime \prime}-\frac{1}{4} \beta^{\prime 2}+k \beta^{2}=1
\end{array}
$$

## The usual approach - transfer matrices

Express solution in matrix form...

$$
\vec{x}(s)=M\left(s \mid s_{0}\right) \vec{x}\left(s_{0}\right) \quad \vec{x}(s)=\binom{x(s)}{x^{\prime}(s)}
$$

Where $M$ is the 'transfer matrix'.

We already know (because we showed)

$$
\frac{d \phi(s)}{d s}=\frac{1}{w^{2}}
$$

$x=w \sqrt{\varepsilon} \cos \left(\phi(s)+\phi_{0}\right)$
Take derivative for $\mathrm{x}^{\prime}$...

$$
x^{\prime}=w^{\prime} \sqrt{\varepsilon} \cos \left(\phi(s)+\phi_{0}\right)-\frac{\sqrt{\varepsilon}}{w} \sin \left(\phi(s)+\phi_{0}\right)
$$

reminder...

$$
\frac{d(\cos (f(x))}{d x}=-\sin (f(x)) \frac{d f(x)}{d x}
$$

$$
\binom{x\left(s_{2}\right)}{x^{\prime}\left(s_{2}\right)}=\left(\begin{array}{ll}
a & b \\
c & d
\end{array}\right)\binom{x\left(s_{1}\right)}{x^{\prime}\left(s_{1}\right)}
$$

Trace two rays... 'cosine like' $\quad \phi=0 \quad$ 'sine like' $\quad \phi=\pi / 2$

$$
\begin{aligned}
x & =w \sqrt{\varepsilon} \cos \left(\phi(s)+\phi_{0}\right) \\
x^{\prime} & =w^{\prime} \sqrt{\varepsilon} \cos \left(\phi(s)+\phi_{0}\right)-\frac{\sqrt{\varepsilon}}{w} \sin \left(\phi(s)+\phi_{0}\right)
\end{aligned}
$$

Yields 4 simultaneous equations so we can solve for $a, b, c, d \ldots$
$M_{12}=\left(\begin{array}{cc}\mu=\phi_{2}-\phi_{1} \\ \frac{w_{2}}{w_{1}} \cos \mu-w_{2} w_{1}{ }^{\prime} \sin \mu & w_{1} w_{2} \sin \mu \\ -\frac{1+w_{1} w_{1}{ }^{\prime} w_{2} w_{2}{ }^{\prime}}{w_{1} w_{2}} \sin \mu-\left(\frac{w_{1}{ }^{\prime}}{w_{2}}-\frac{w_{2}{ }^{\prime}}{w_{1}}\right) \cos \mu & \frac{w_{1}}{w_{2}} \cos \mu+w_{1} w_{2}{ }^{\prime} \sin \mu\end{array}\right)$

## You will see this later...

Simplify by considering a period or 'turn', and w's are equal.

$$
M_{\text {period }}=\left(\begin{array}{cc}
\cos \mu-w w^{\prime} \sin \mu & w^{2} \sin \mu \\
-\frac{1+w^{2} w^{\prime 2}}{w^{2}} \sin \mu & \cos \mu+w w^{\prime} \sin \mu
\end{array}\right)
$$

If we define the so-called 'Twiss' or 'Courant-Snyder' parameters:

$$
\begin{gathered}
\beta=w^{2} \quad \alpha=-\frac{1}{2} \beta^{\prime} \quad \gamma=\frac{1+\alpha}{\beta} \\
M_{\text {period }}=\left(\begin{array}{cc}
\cos \mu+\alpha \sin \mu & \beta \sin \mu \\
-\gamma \sin \mu & \cos \mu-\alpha \sin \mu
\end{array}\right)
\end{gathered}
$$

(sorry that we are reusing symbols again... these are NOT the relativistic parameters)


[^0]:    In the Introductory school, we will focus on matrices (1st order maps), but you should

